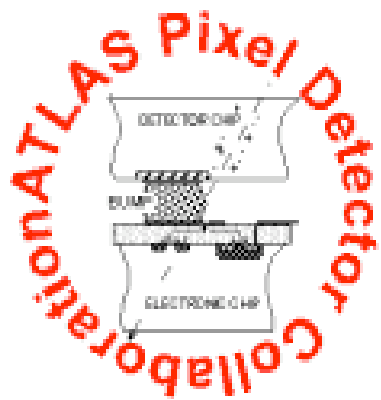


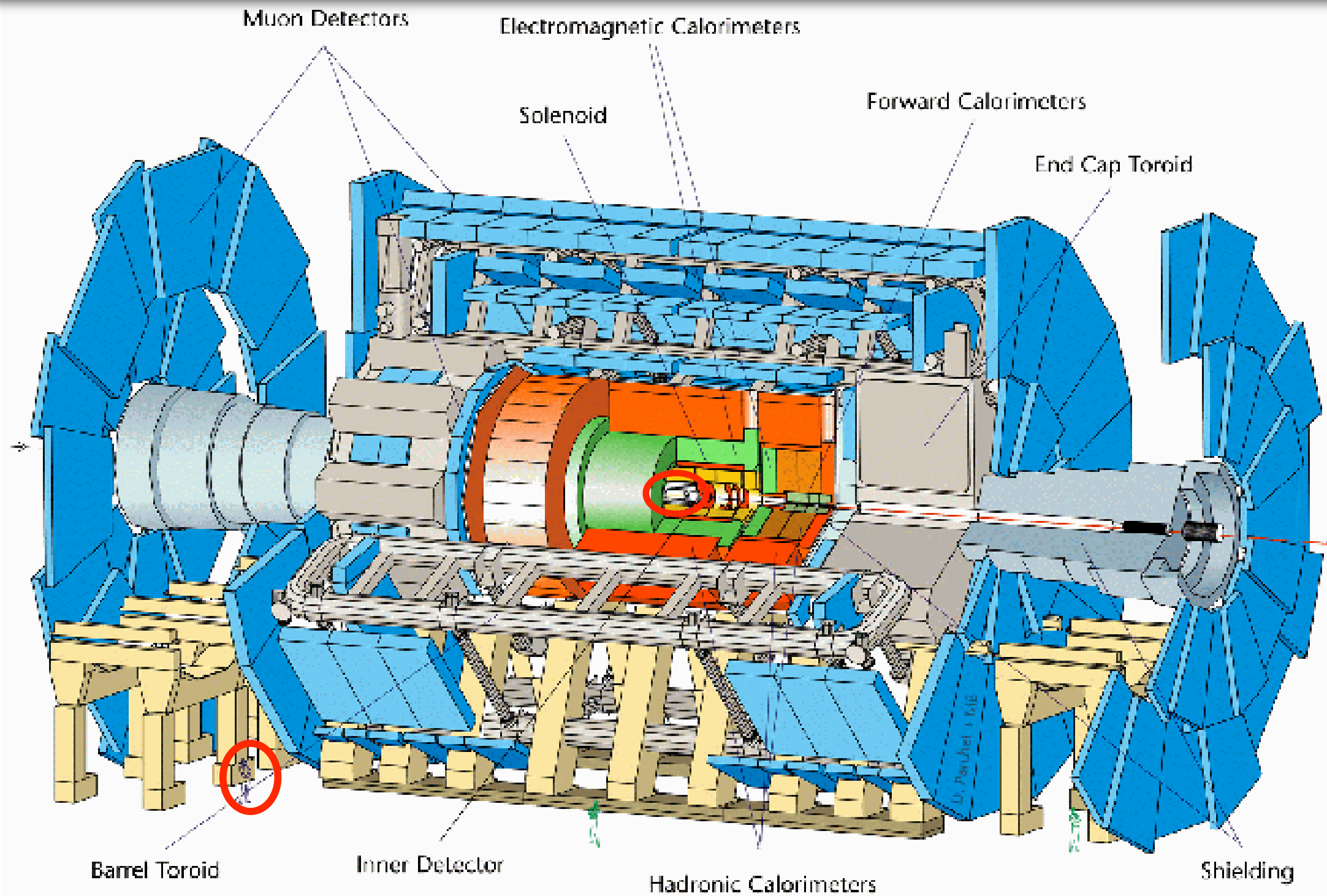
# Results from the commissioning of the ATLAS Pixel Detector

Lucia Masetti - Physikalisches Institut Uni Bonn  
for the ATLAS Pixel Collaboration

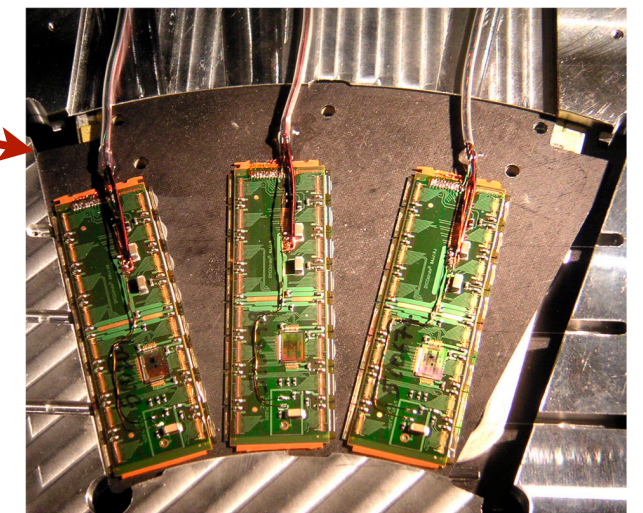
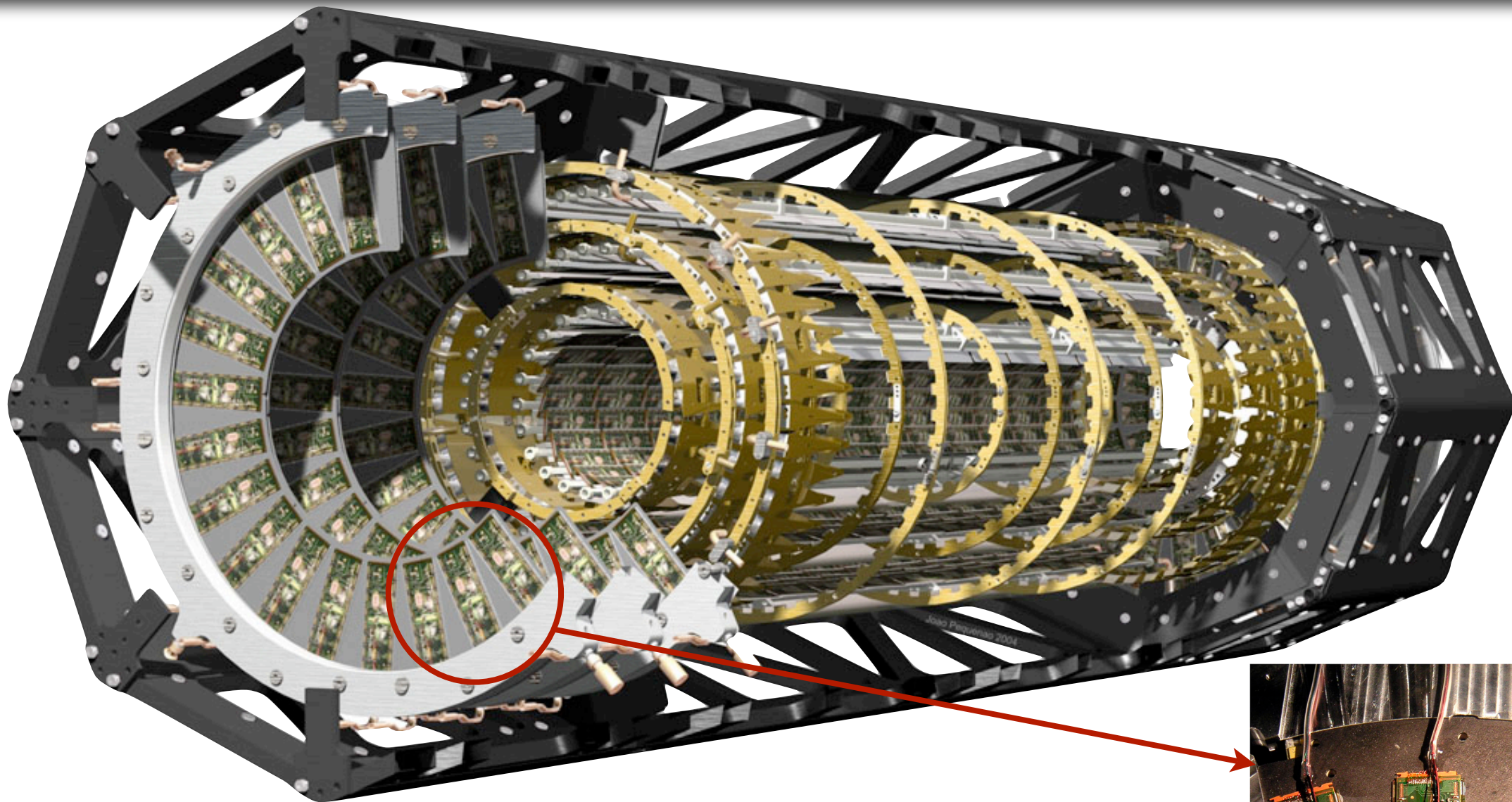
17 September 2008  
TWEPP 2008 - Naxos, Greece



# The ATLAS Detector



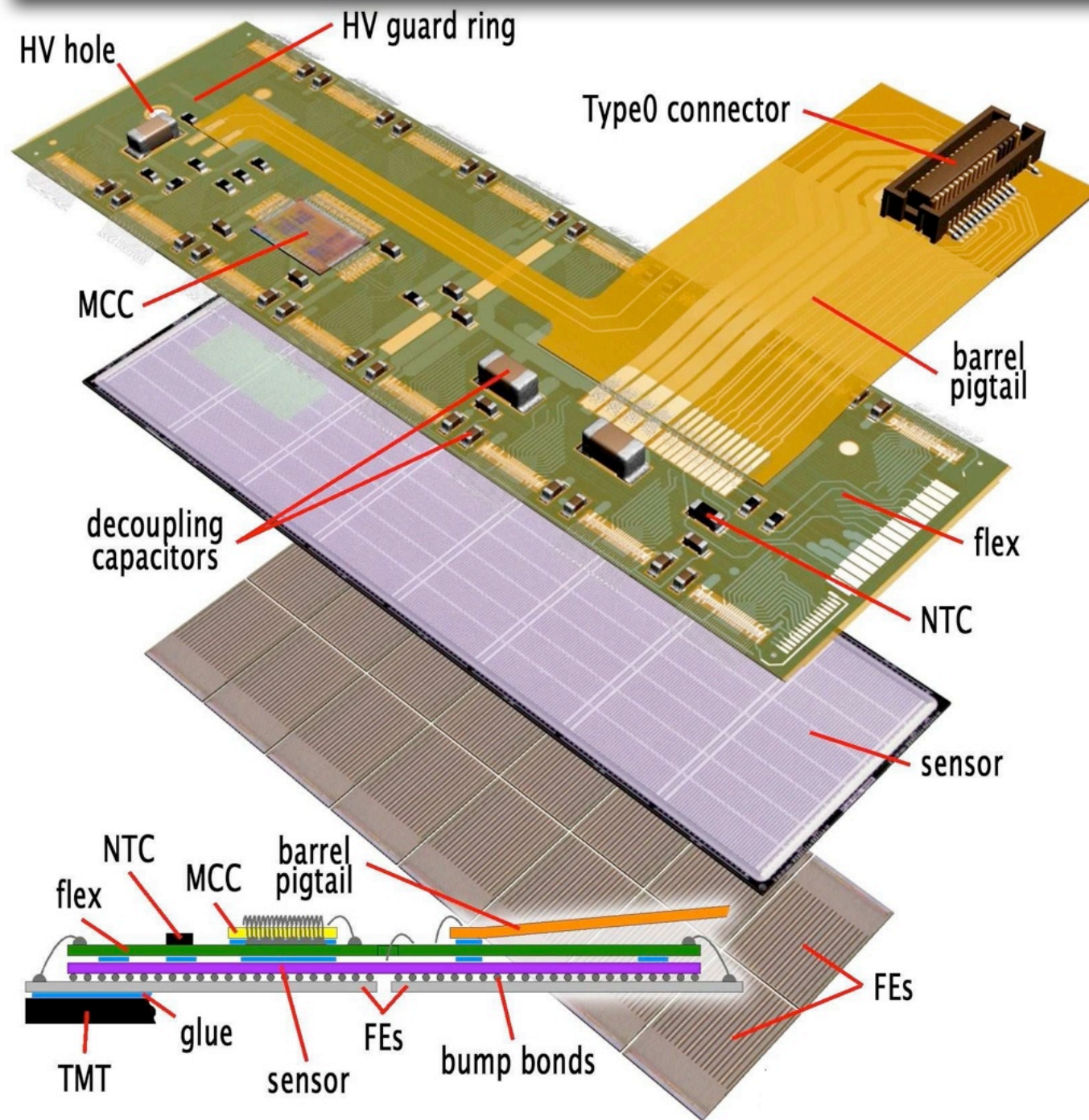
# The Pixel Detector



80M channels in 1744 modules, sensitive area 1.6 m<sup>2</sup>  
Coverage up to  $|\eta| < 2.5$  with 3 points  
Innermost layer at R=5cm



# The Pixel Module

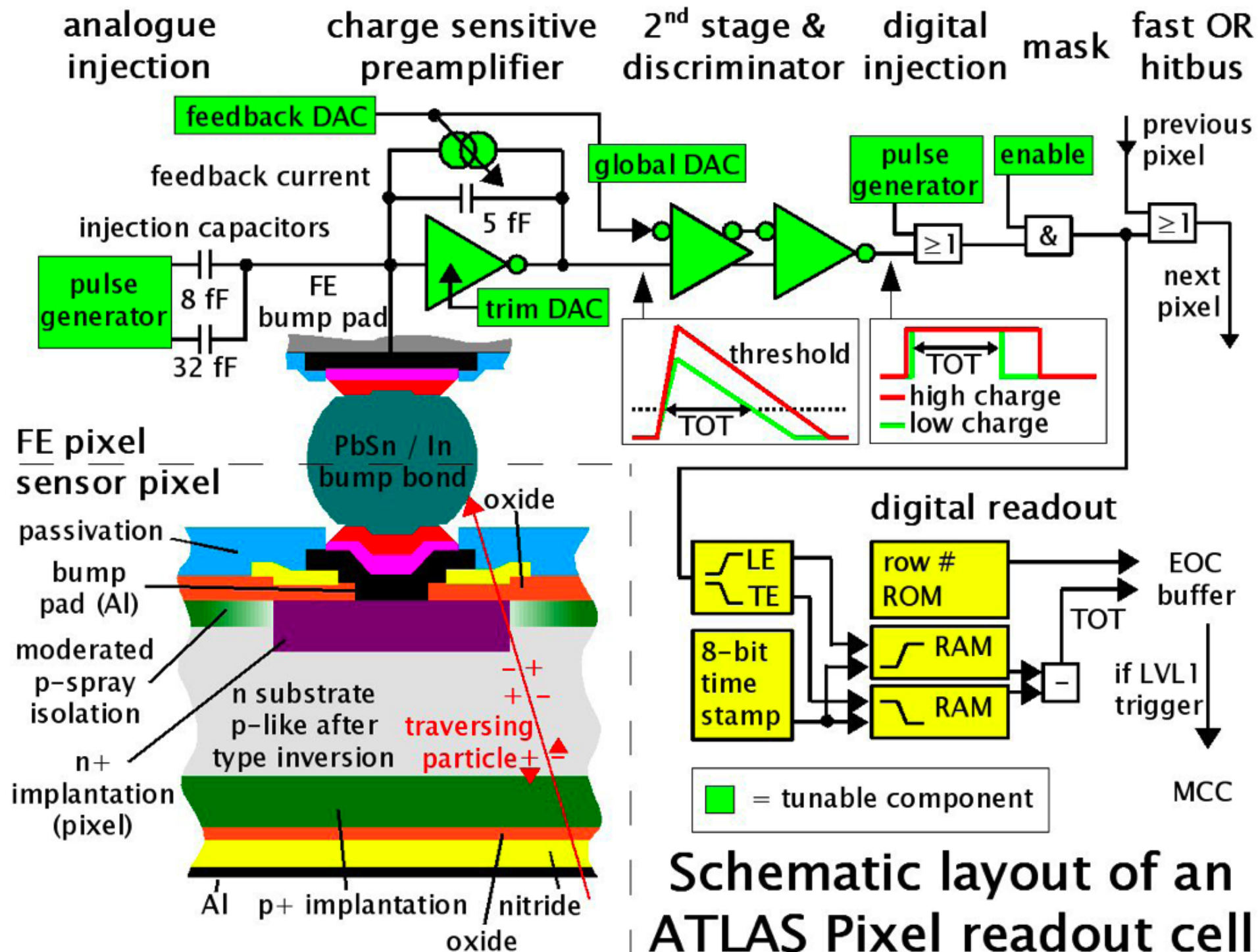


## Features

- $60.8 \times 16.4 \text{ mm}^2 \times 250 \text{ }\mu\text{m}$  active silicon volume
- $50 \text{ }\mu\text{m}$  ( $R\phi$ )  $\times$   $400 \text{ }\mu\text{m}$  ( $\eta$ ) pixels
- 16 Front-end chips, 2 bump bonding techniques
- Module Control Chip on flex hybrid: distribution of commands and event building
- Thickness:  $196 \text{ }\mu\text{m}$  FEs +  $100 \text{ }\mu\text{m}$  flex hybrid
- Tested to be radiation hard up to 500 kGy and  $10^{15} \text{ n}_{\text{eq}} / \text{cm}^2$

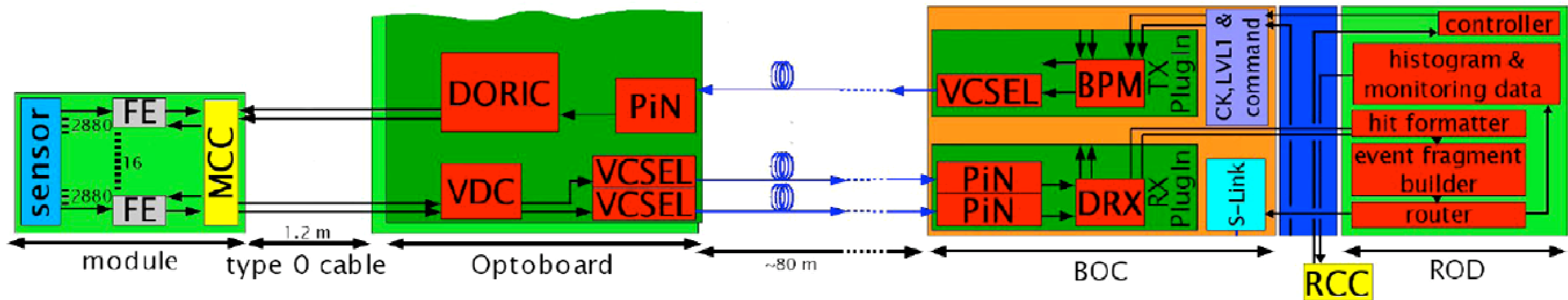


# Pixel readout cell



Schematic layout of an ATLAS Pixel readout cell

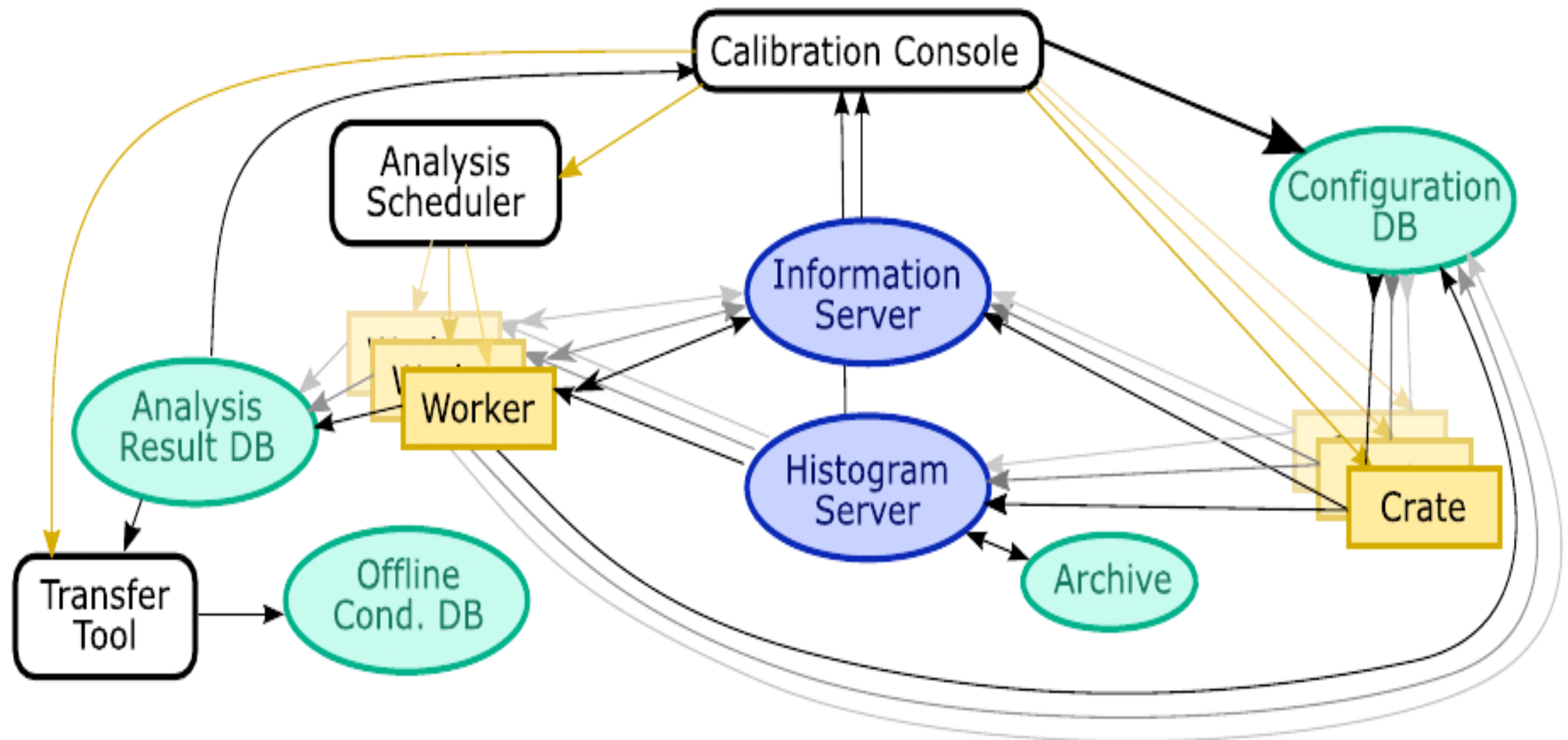
# The readout system



- Connection between on- and off-detector electronics via optical links
- **Tunable components:**
  - Voltage controlling the amplitude of the VCSEL current on the optoboard (1 per board = 6-7 modules)
  - Phase and threshold of each link (RX plugin)
- **Separate paths** for data and calibration hits (different occupancy)



# Calibration infrastructure



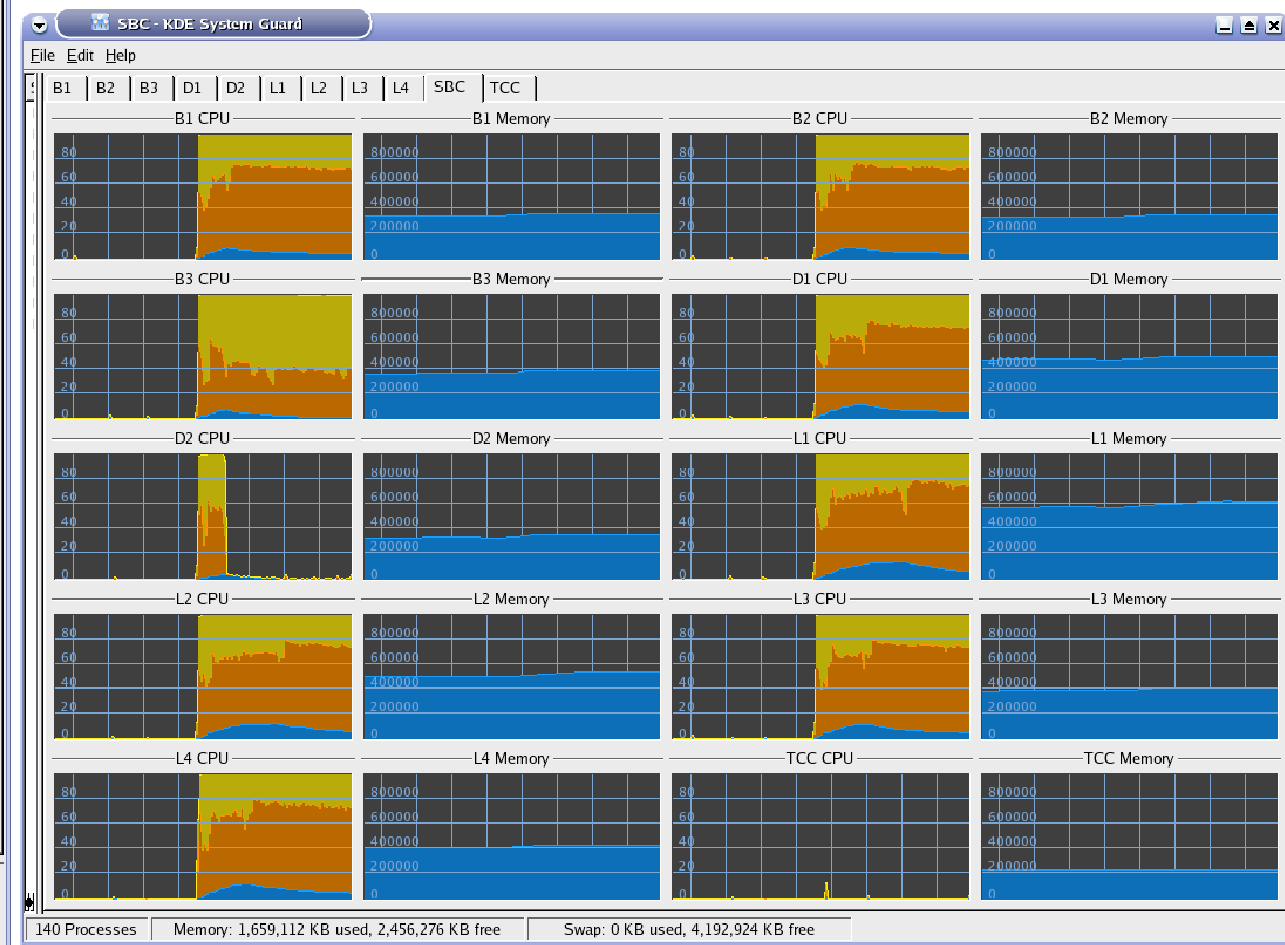
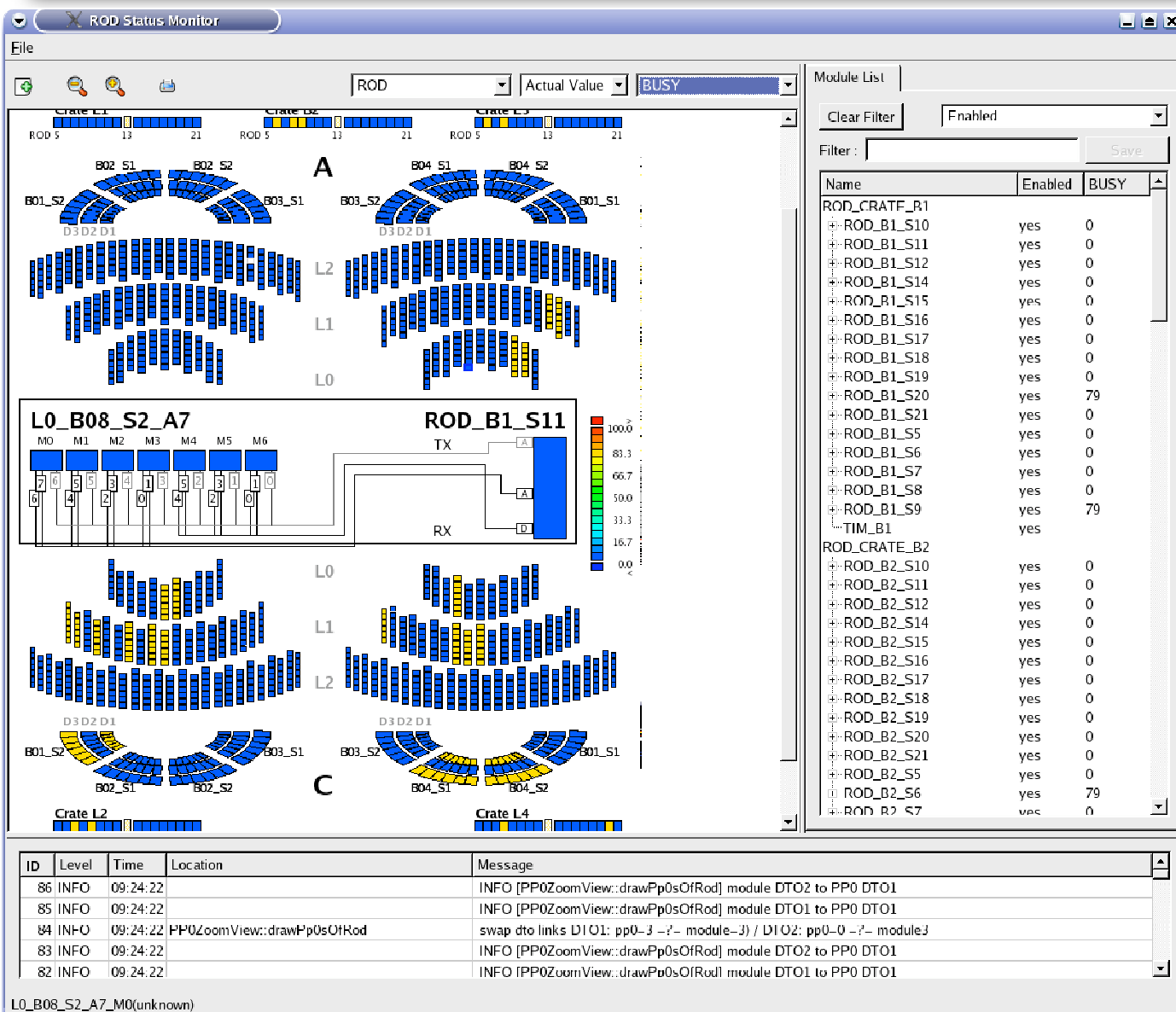
# Calibration procedure

- Distributed system steered by a single GUI
- **Scans** started through 9 single board computers (1 per VME crate) on the RODs
- ROD DSPs send commands to the modules, produce histograms from module data and perform fitting
- Final histograms retrieved via VME bus, stored on a histogram server and on disk and **analysed automatically** on a dedicated farm





# Calibration monitoring



ROD Status Monitor

SBC Memory and CPU  
usage Monitor



L. Masetti

Commissioning of the ATLAS Pixel Detector

TWEPP 2008



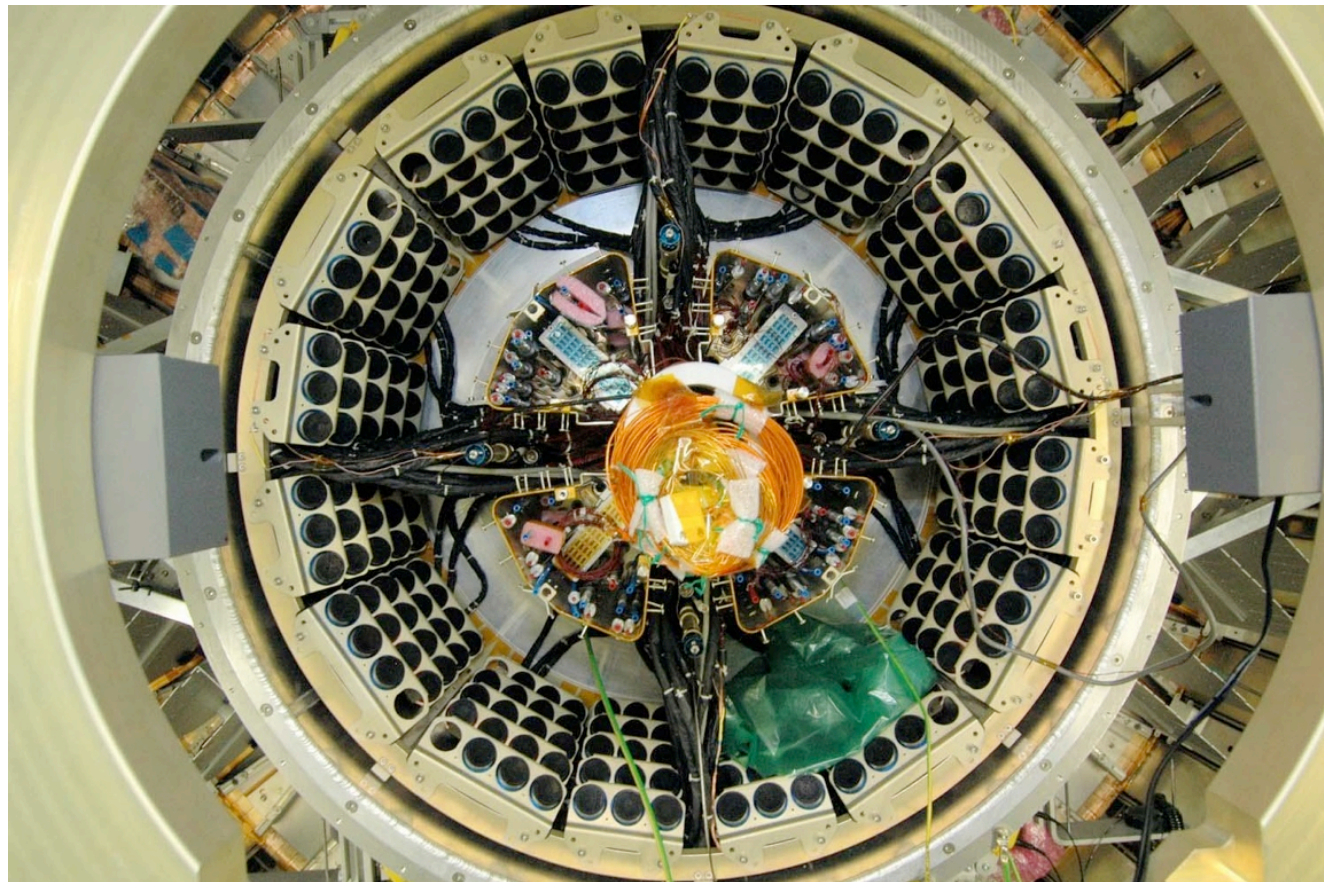
# Installation

- **March-June 2007:** Assembly and connectivity test: check mapping from modules up to PP0
- **July 2007:** Installation of pixel-package in the cavern, then wait for TRT and SCT to finish cabling
- **December 2007:** Finished services connection and test from PP2 outwards, replaced cooling exhaust pipes at PP1 damaged by corrosion
- **February 2008:** Connection at PP1 and connectivity test: low voltage, high voltage, fibres



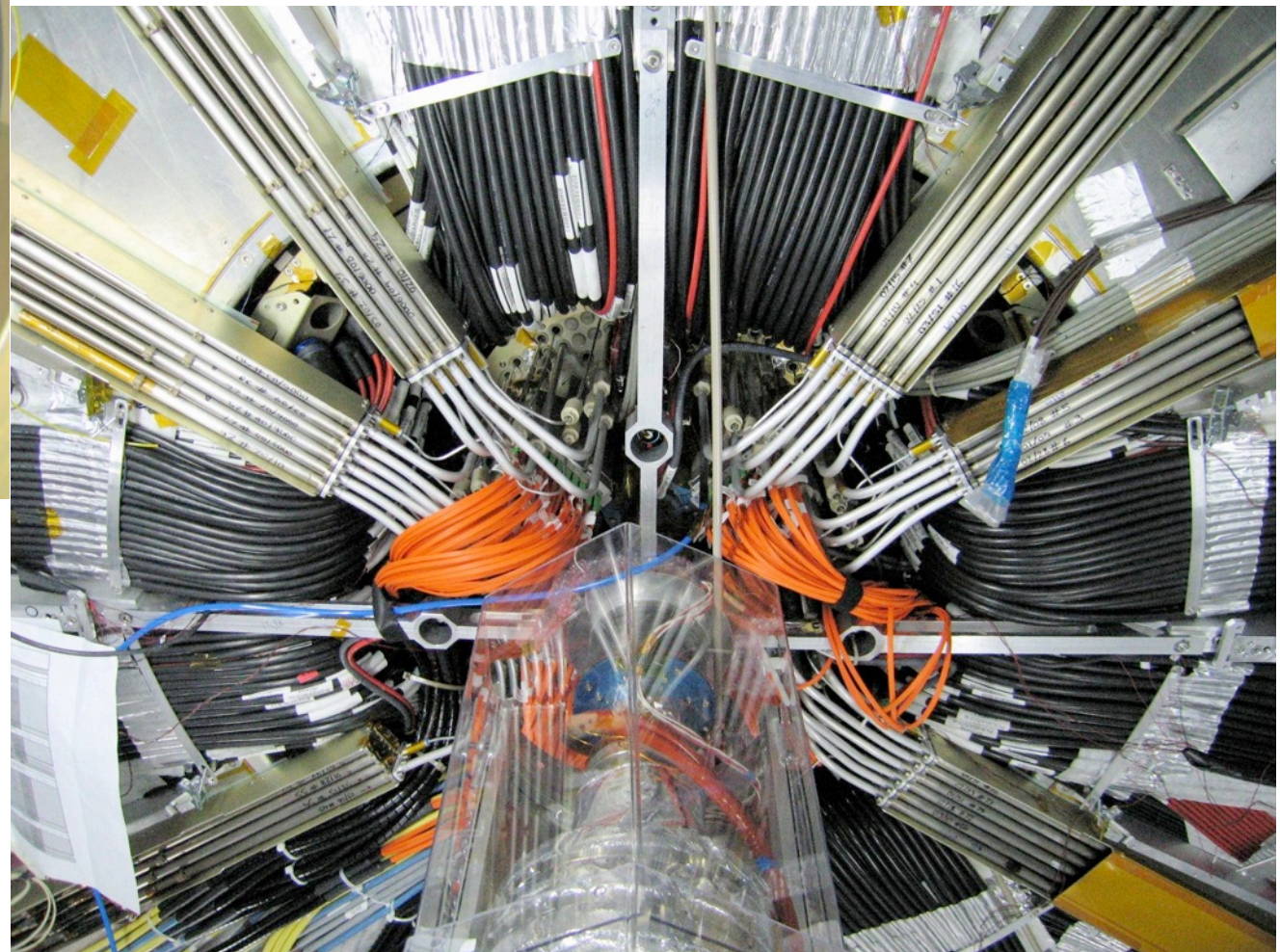


# PP1 connection



PP1 before...

... and after connection





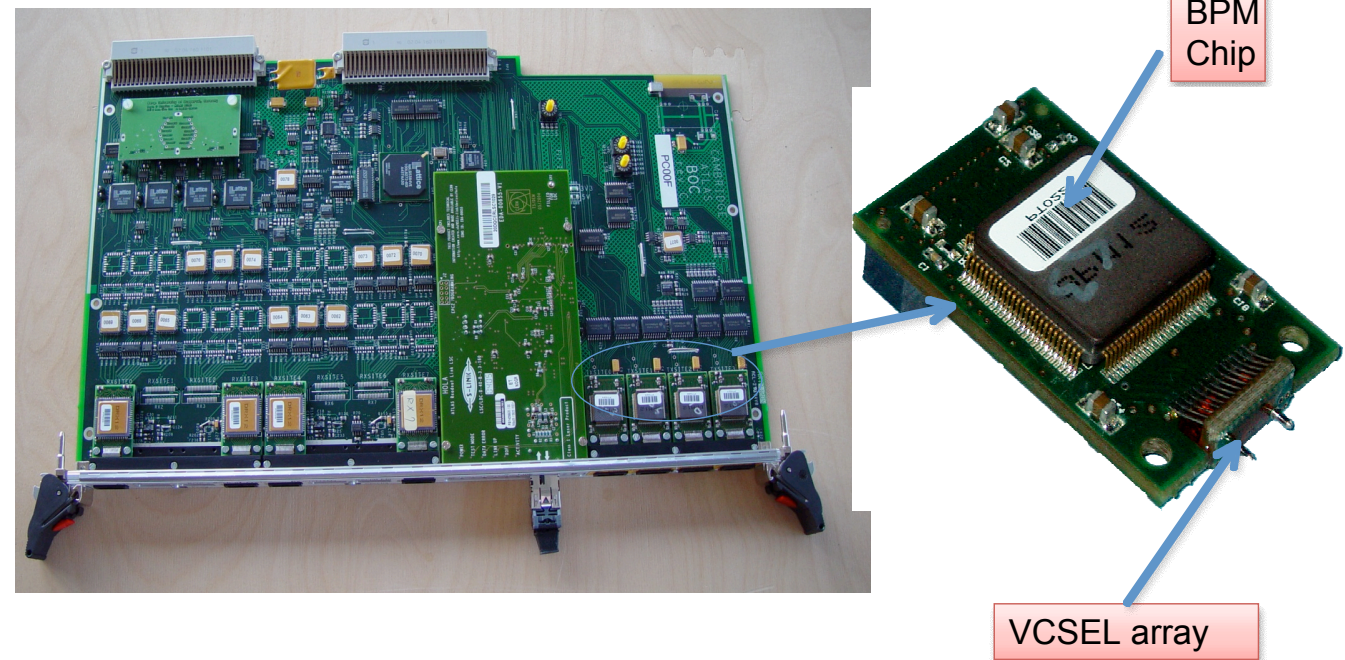
# Connectivity test in the pit

- **March 2008:** Found general fibre swap, reconnection at BOC needed (over Easter)
- **April 2008:** Checked light transmission to and from detector, that modules can be configured and that the sensor is biased
- No cooling available, could only configure one FE per module and scan 1/10 of its pixels



# Optical sign-off results

- **Light power** on TX channels lower than measured during assembly: compatible with expected fibre attenuation
- **Dead TX channels:** 24 dead channels at the end of the sign-off. Channels dying at a very high rate (few per month)
- Behaviour compatible with ESD damage during production
- More channels dead in pixel than in SCT because lasers were kept on longer



# Cooling commissioning

- Cooling available **from April 25th**
- Tested behaviour of temperatures, back-pressure and heater power with modules off (0 W), normal configuration (4 W) and high power configuration (5.5 W)
- Better stabilisation with configured modules, some **instabilities with detector off**
- **3 leaky loops** found in the disks, but can be operated at least for commissioning
- 77/88 loops tested till **May 1st** when system had to be turned off due to **cooling plant accident**



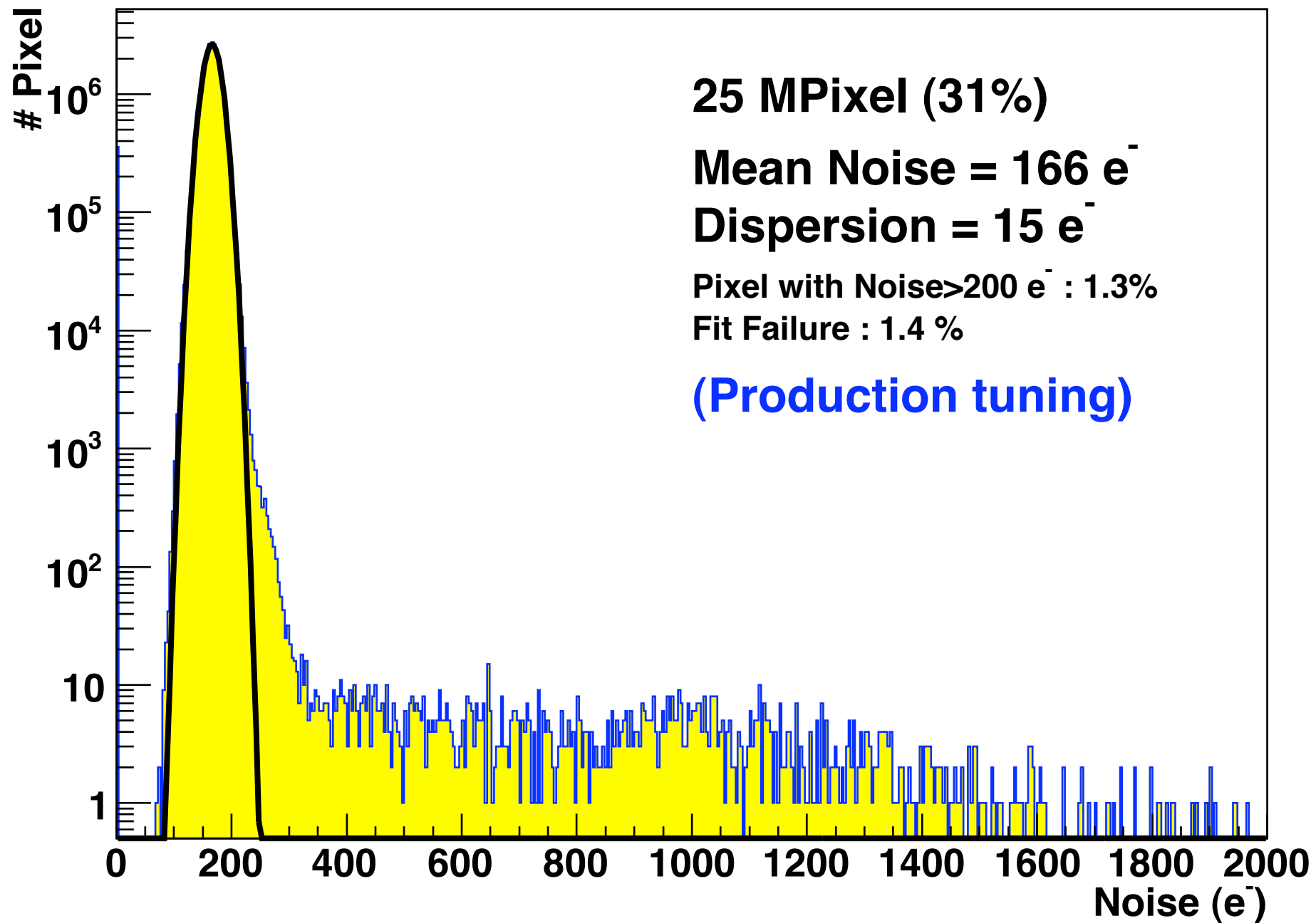
# Module status after sign-off

- 8 modules with HV open (mainly in the disks)
- 1 module without clock
- 2 swapped modules (corrected in DB)
- 4 badly tuned modules (to be retuned)
- 3 modules with suspect behaviour (like HV open)
- Modules with (suspect or proven) HV open are kept on for calibration and standalone data taking
- Threshold scan on about 900 modules (all pixels) showed **no significant increase of dead pixels wrt production**





# Pixel noise



# Commissioning restart

- **3 compressors were damaged** due to prolonged slippage in the magnetic coupling between the motor and the compressor shaft
- Compressors were repaired and 800 kg contaminated coolant was replaced
- Cooling working again in time for **beam-pipe bakeout at the end of July**
- Bakeout successful, but with coolant loss of 1 kg/hour
- **Cooling commissioning till August 11th**

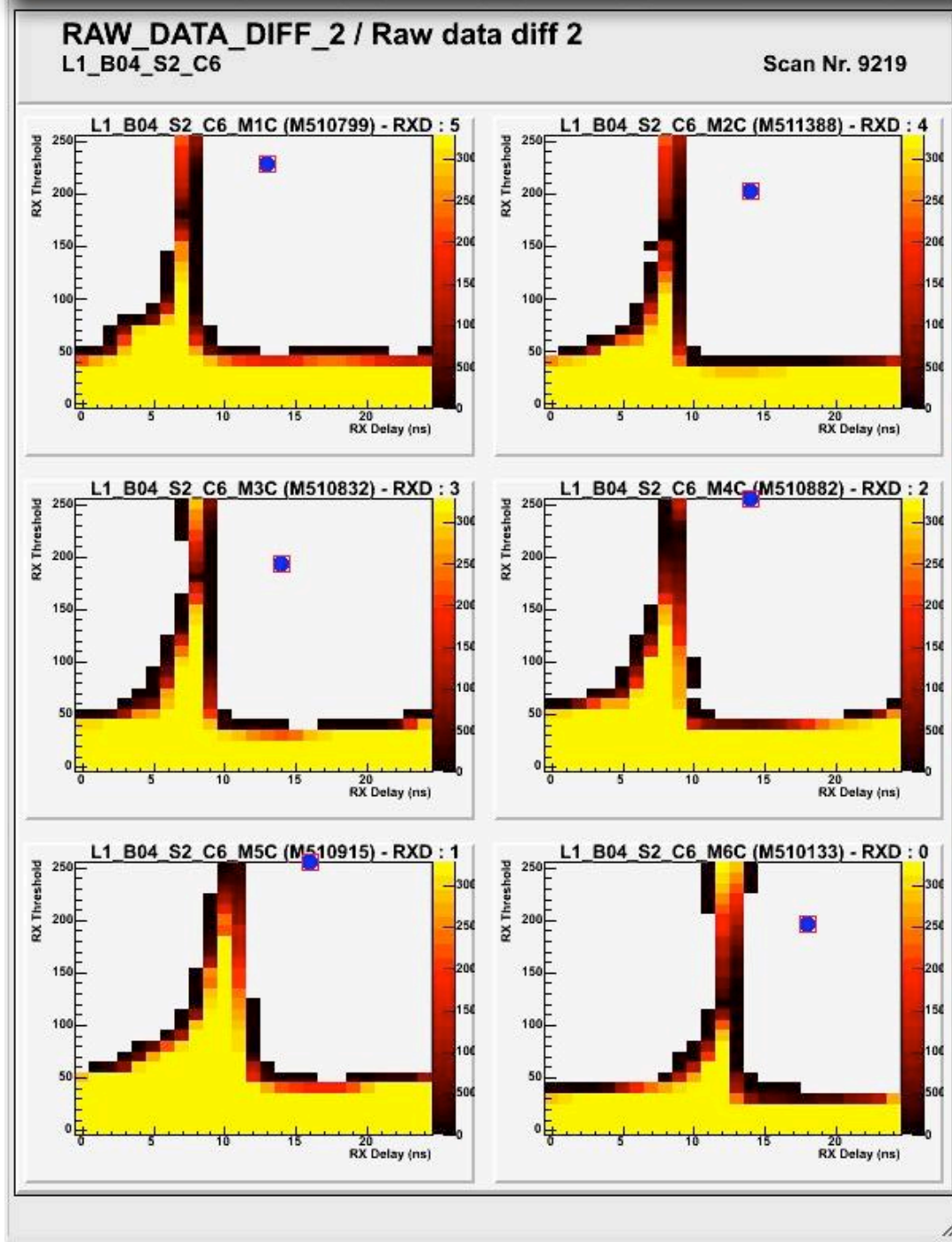


# Optolink tuning

- Modules operate at  $-2$  to  $4^{\circ}\text{C}$ , too cold for stable operation of optoboards
- **Heaters** can keep optoboards at room temperature
- Tried to run at  $10^{\circ}\text{C}$  to preserve laser channels from dying
- **Optolinks could only be tuned at  $20^{\circ}\text{C}$** , that was finally chosen as the operating temperature
- On **August 28th** the whole detector could be turned on and **1662/1744 modules** could be further tested (communication ok)

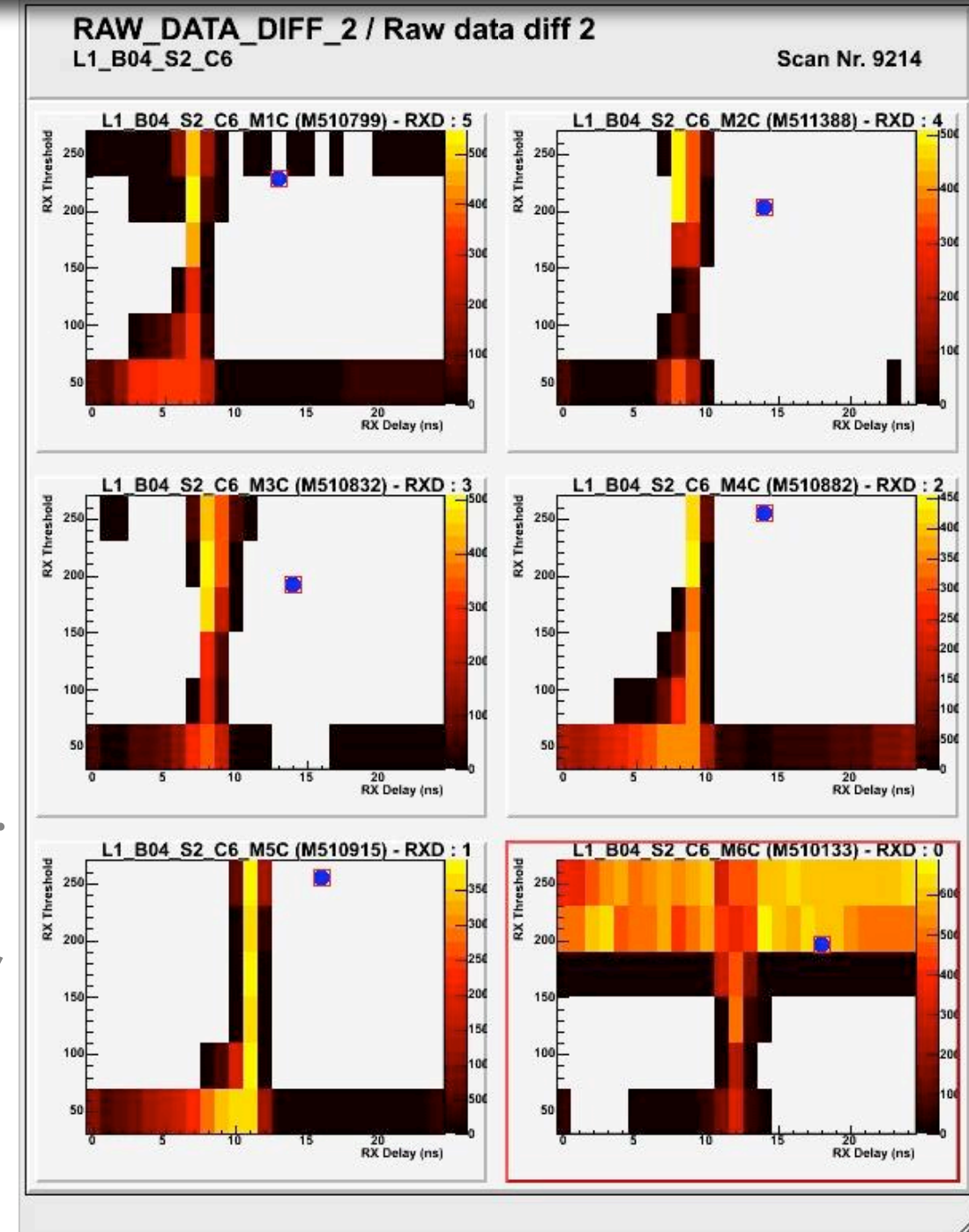


# Optolink tuning problems



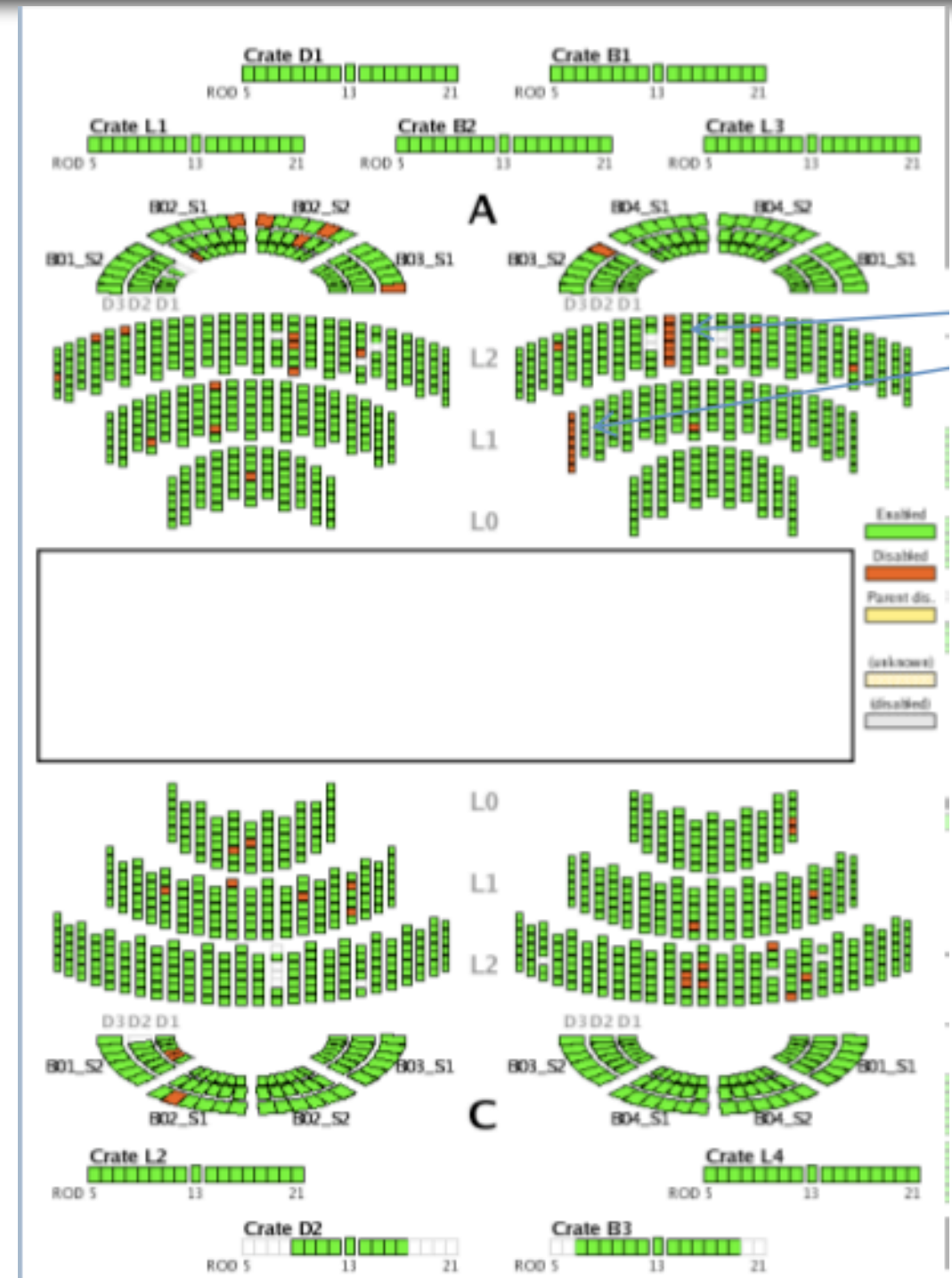
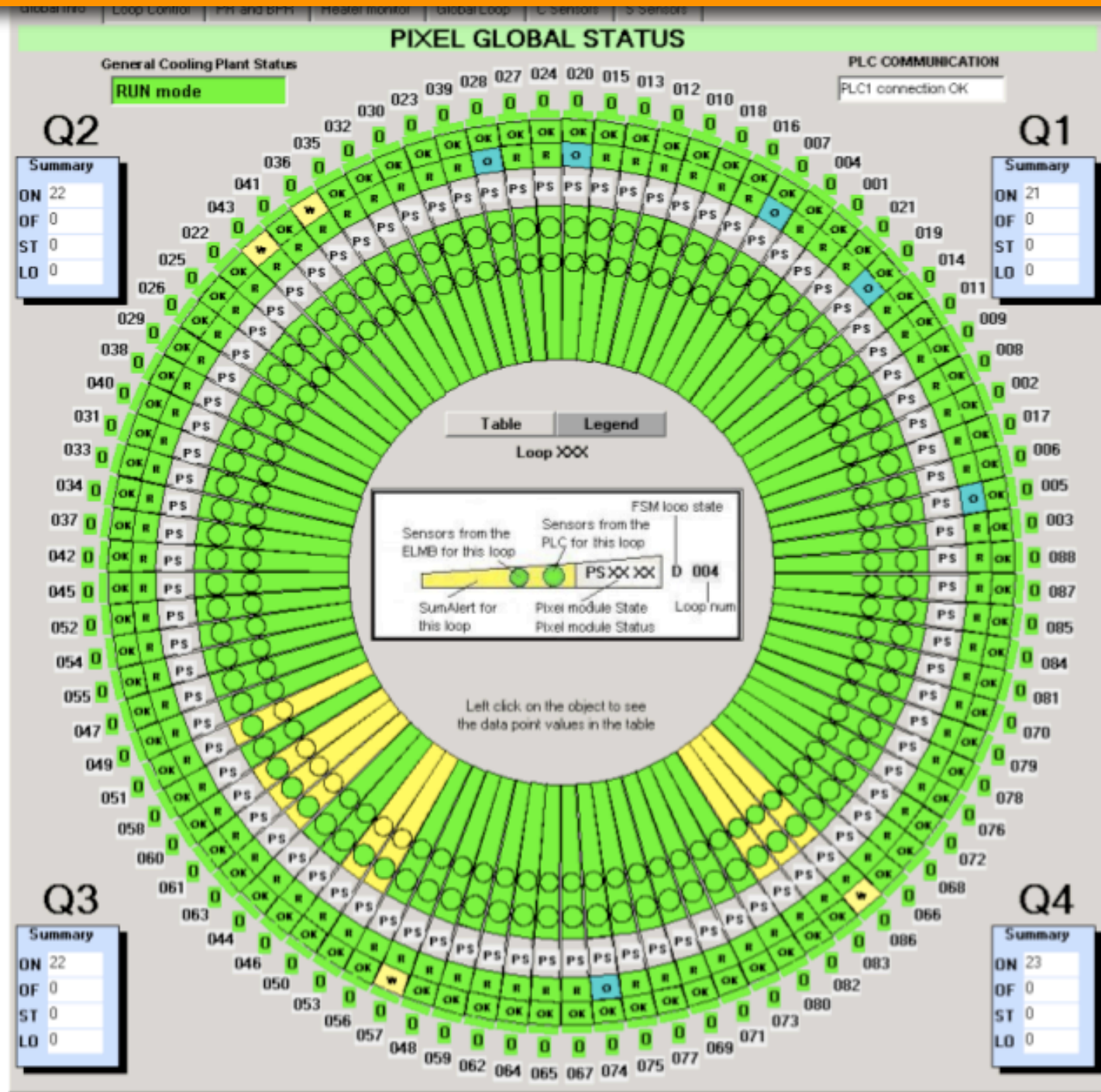
Optoboards  
at 10°C

Different data  
patterns  
give different  
optimal  
threshold/  
phase settings.  
Can only tune  
with the faster  
one (left)



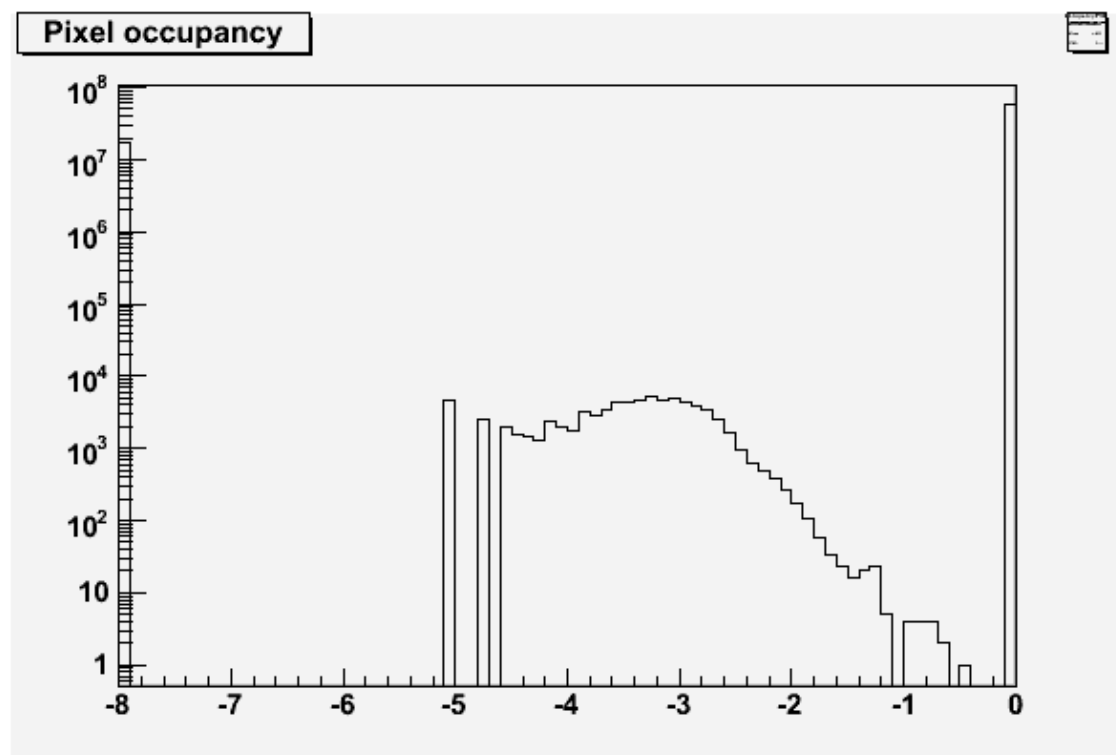


# 95% of detector working



# Noise runs

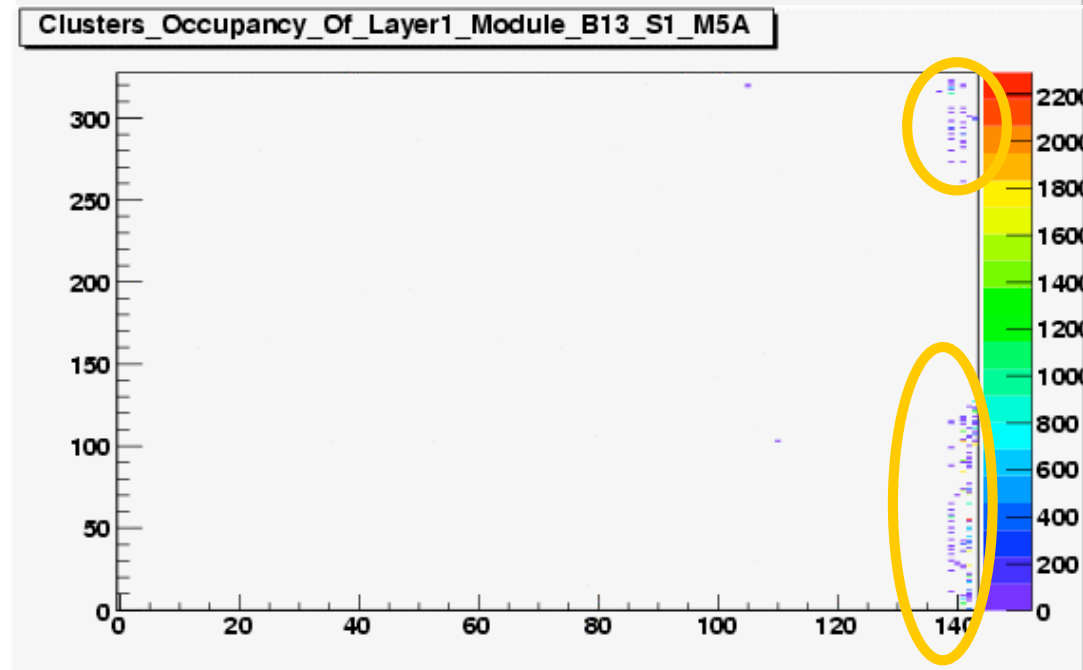
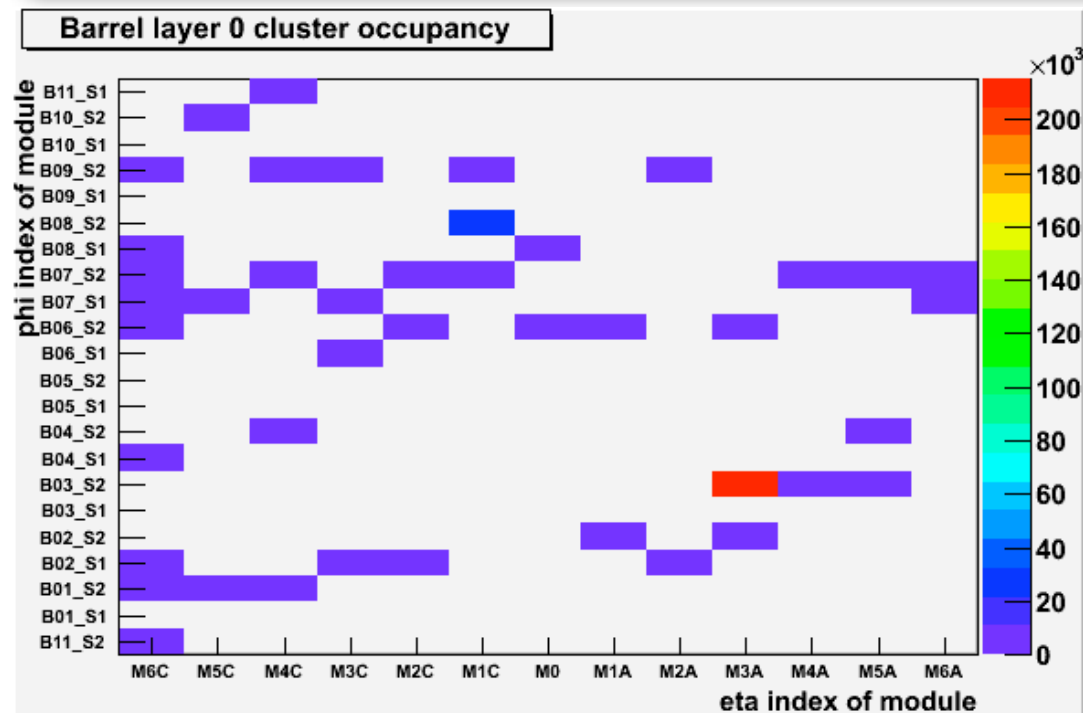
- First noise runs on whole detector taken end of August (1-2 million events)
- Disabled problematic channels for data taking, almost complete overlap with module selection for calibration



- Modules with HV open included to check their behaviour
- Less than 400 modules with noisy pixels (occupancy  $> 10^{-4}$ ), about **130000 pixels masked**
- 100 clusters/BC without masking

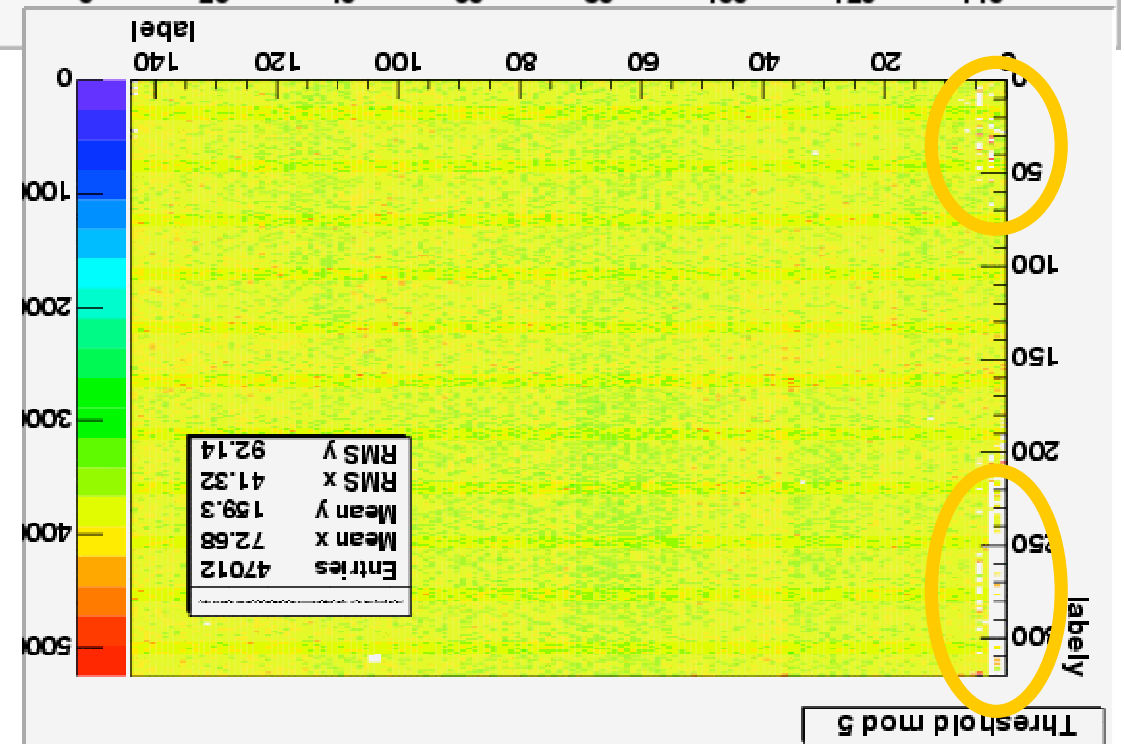
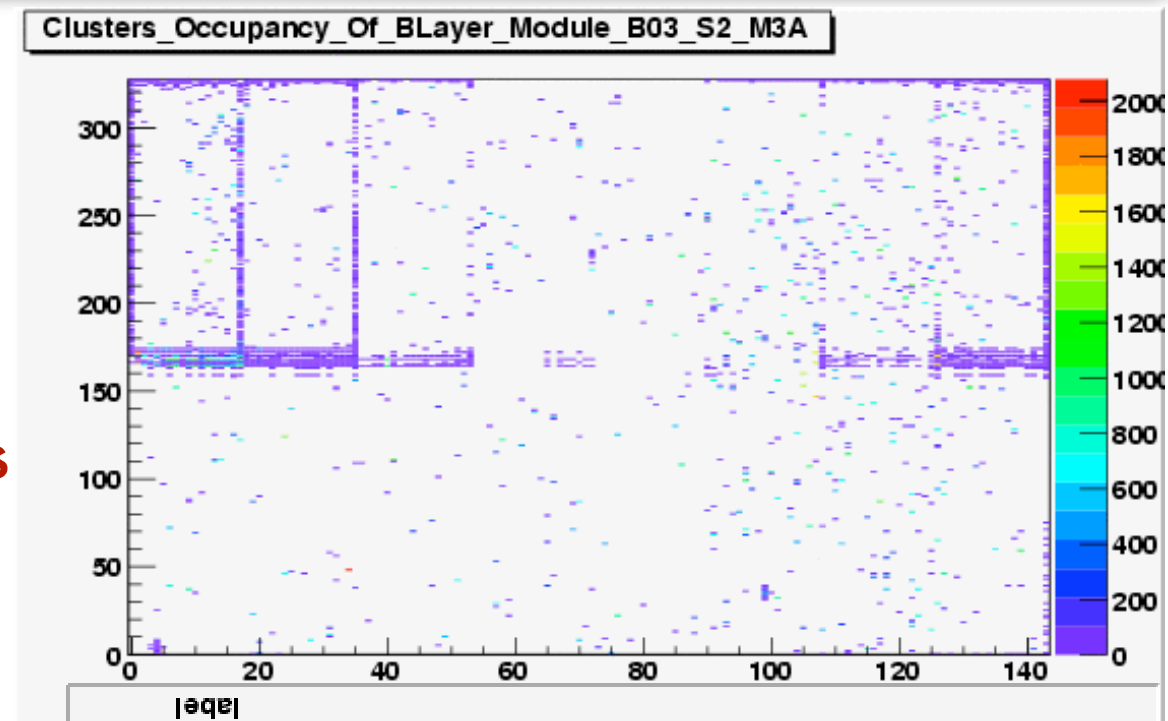


# Modules without HV



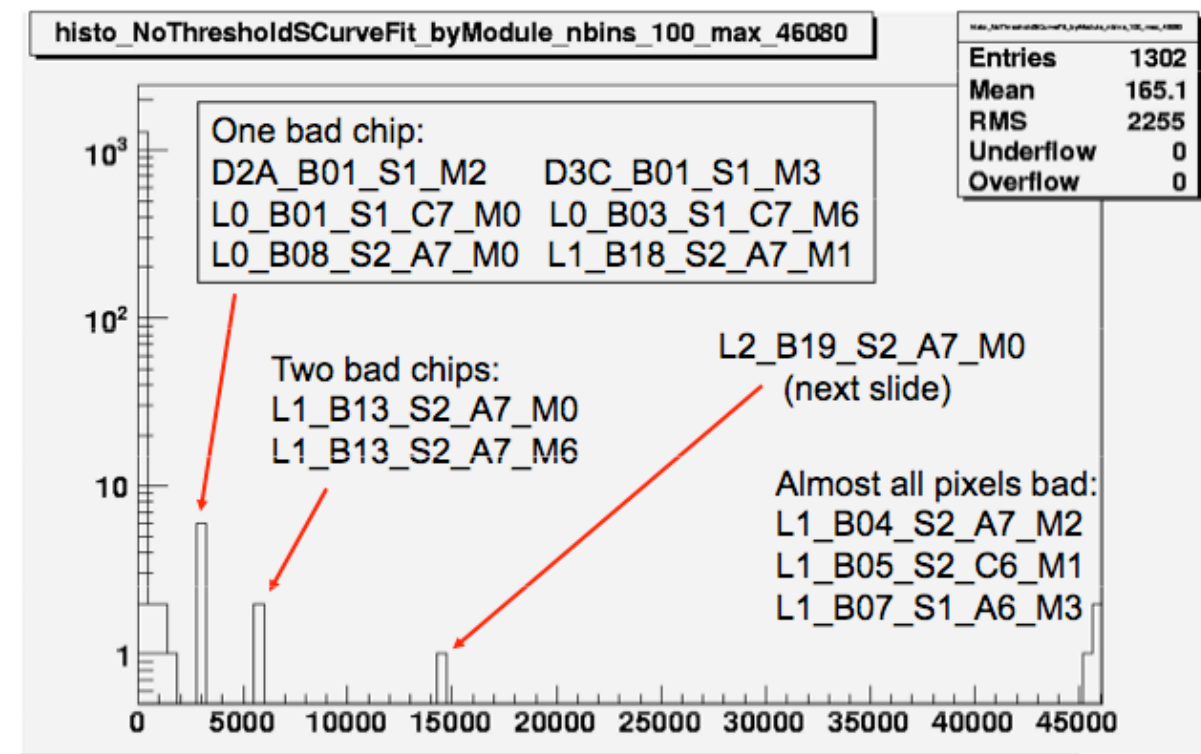
Most clusters  
from single  
noisy modules

Good  
agreement  
between noise  
data and  
threshold  
scan



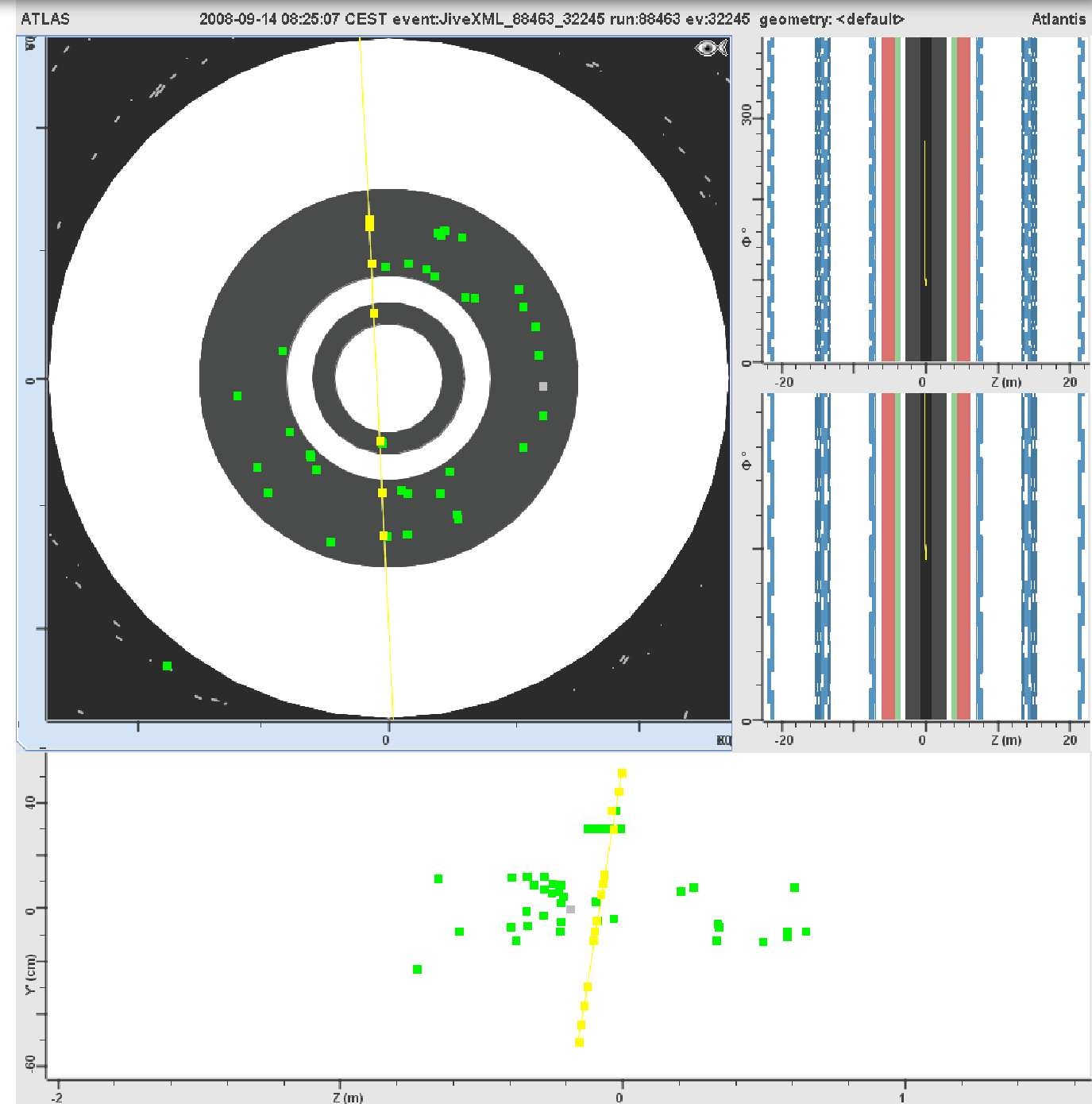
# Calibration with prod tuning

- **Plan for first September week:** complete collection of calibration data with production tuning for modules without communication problems
- **Threshold scans:** completed
- **In-time threshold scans:** completed
- **ToT scans:** partially done
- Number of **pixels failing S-curve fit** per module (see figure):
  - 3 modules almost completely bad
  - 8 modules with bad FEs + 1 with strange behaviour
  - About 30 other modules with few hundred bad pixels



# Combined cosmics runs

- **September 4th:** first combined cosmics runs with TRT
- Bad timing, no tracks with pixel hits
- **September 14th:** combined cosmics runs with SCT
- Changed time window for readout, reading 8 BCs
- **First track:** 7 pixel hits and 16 SCT hits
- Alignment studies ongoing, just few tracks available





# Conclusions

- Have to run **optoboards at 20°C** to get stable optolink tuning
- **95% of the Pixel detector** was calibrated and is running fine using production tuning
- 3 leaky cooling loops (36 modules) were also commissioned, but will not run this year
- New threshold and ToT tuning foreseen but not strictly necessary
- Debugging of excluded modules has lower priority for the moment
- **Identification of bad pixels** agrees well enough between calibration and noise runs
- **Cosmics tracks being collected**, but beam commissioning doesn't allow to have HV on all the time

