



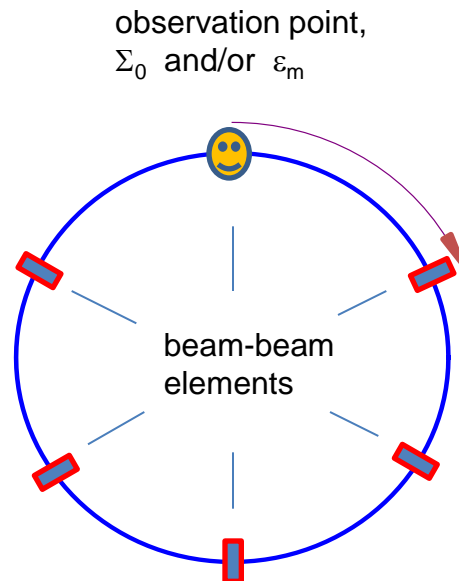
Fermilab Accelerator Physics Center

Computation of Eigen-Emittances (and Optics Functions!) from Tracking Data

Yuri Alexahin

Motivation

- "Express" space-charge modeling (e.g. with MADX beam-beam elements)



$$\Sigma^{(k)} = \mathbf{M}_{0 \rightarrow k} \cdot \Sigma^{(0)} \cdot \mathbf{M}_{0 \rightarrow k}^T$$

alternatively

$$\sigma_u^2 = \sum_m \beta_{um} \varepsilon_m, \quad u = x, y, z$$

does not require knowledge of Σ_k
in a simple kick-rotate scheme
(but does require Σ_i , $i=0, \dots, k-1$)

requires knowledge of β -functions,
with fixed β -s can be used as a
rough approximation

- Muon cooling channel design -

<http://map-docdb.fnal.gov/cgi-bin/ShowDocument?docid=4358>

Outline

- How to suppress halo contribution to covariance matrix in a self-consistent way to obtain right sizes for the space charge forces computation? - Multidimensional case please!
 - Iterative procedure for nonlinear fit of particle distribution in the phase space with a Gaussian or other smooth function.
- How to find the normal mode emittances (eigen-emittances) when optics functions are not known?
 - Eigen-emittances as well as optics functions can be determined from the covariance matrix.
- (Extremely fast & simple!) exponential fit of particle distribution when the optics functions are known – already implemented in MAD-X, but not described anywhere.

Definitions

Phase space vector:

$$\underline{z} = \{x, P_x, y, P_y, s - c\beta_0 t, \delta\}$$

Canonical momenta in units of the reference value $p_0 = mc\beta_0\gamma_0$:

$$P_x = (p_x + \frac{e}{c} A_x) / p_0$$

Energy deviation (disguised as momentum)

$$\delta = (\gamma - \gamma_0) / \beta_0^2 \gamma_0$$

Covariance matrix (Σ - matrix)

$$\Sigma_{i,j} = \frac{1}{N} \sum_{k=1}^N \zeta_i^{(k)} \zeta_j^{(k)}, \quad \zeta_i^{(k)} = z_i^{(k)} - \bar{z}_i, \quad \bar{z}_i = \frac{1}{N} \sum_{k=1}^N z_i^{(k)}, \quad i = 1, \dots, 6$$

Basic assumption: particle distribution is a function of quadratic form

$$\Phi(\underline{\zeta}) = (\underline{\zeta}, \Sigma^{-1} \underline{\zeta}) \equiv \sum_{i=1}^6 \zeta_i (\Sigma^{-1} \underline{\zeta})_i = \sum_{i,j=1}^6 \Sigma_{ij}^{-1} \zeta_i \zeta_j$$

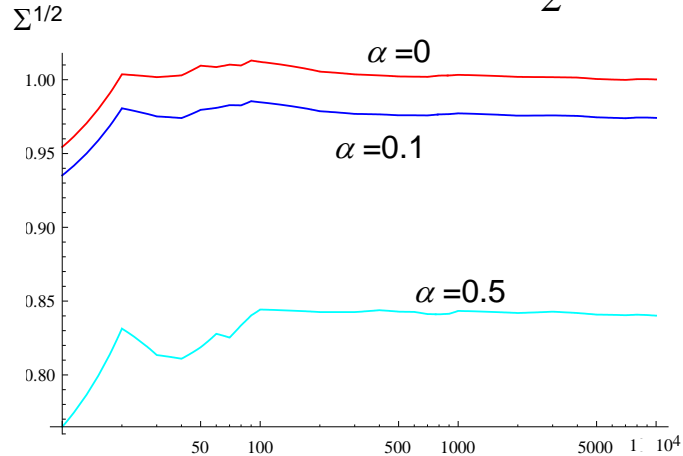
How to Suppress Halo Contribution?

And to do this in a self-consistent way?

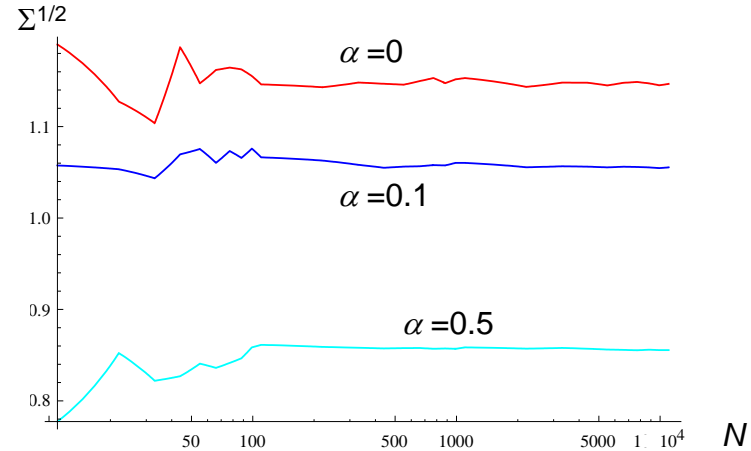
- a simple heuristic method is to introduce weights proportional to some degree of the distribution function. This leads to an iterative procedure

$$\bar{z}_i = \sum_{k=1}^N w_k z_i^{(k)} / \sum_{k=1}^N w_k, \quad \zeta_i^{(k)} = z_i^{(k)} - \bar{z}_i, \quad \Sigma_{i,j} = \sum_{k=1}^N w_k \zeta_i^{(k)} \zeta_j^{(k)} / \sum_{k=1}^N w_k, \quad (1)$$

For Gaussian $w_k = \exp[-\frac{\alpha}{2}(\zeta_i^{(k)}, \Sigma^{-1} \zeta_i^{(k)})]$, α being a fitting parameter ($0 < \alpha < 1$)



Square root of Σ from eq.(1) averaged over 25 realizations of 1D Gaussian distribution with $\sigma=1$ as function of the number of particles N .



Square root of Σ from eq.(1) averaged over 25 realizations of superposition of 1D Gaussian distributions with $\sigma=1$ (90%) and $\sigma=3$ (10%)

This method is imprecise and ambiguous \Rightarrow something based on a more solid foundation is needed.

Nonlinear Fit of the Klimontovich Distribution

$$G(\underline{z}) = \frac{1}{N} \sum_{k=1}^N \delta_{6D}(\underline{z} - \underline{z}^{(k)}) \equiv \frac{1}{N} \sum_{k=1}^N \prod_{i=1}^6 \delta(z_i - z_i^{(k)})$$

We want to approximate it with a smooth function, e.g. Gaussian

$$F(\underline{\zeta}) = \frac{\eta}{(2\pi)^{n/2} \sqrt{\det \Sigma}} \exp\left[-\frac{1}{2} (\underline{\zeta}, \Sigma^{-1} \underline{\zeta})\right]$$

where η is the fraction of particles in the beam core,
via the minimization problem

$$\int_{-\infty}^{\infty} \dots \int_{-\infty}^{\infty} |F - G|^2 dz_1 \dots dz_n = \int_{-\infty}^{\infty} \dots \int_{-\infty}^{\infty} (F^2 - 2FG) dz_1 \dots dz_n + \int_{-\infty}^{\infty} \dots \int_{-\infty}^{\infty} G^2 dz_1 \dots dz_n \rightarrow \min$$

or the maximization problem for the 1st term in the r.h.s. taken with the opposite sign

$$M(\bar{z}, \Sigma, \eta) = \int_{-\infty}^{\infty} \dots \int_{-\infty}^{\infty} (2FG - F^2) dz_1 \dots dz_n =$$

$$\frac{\eta}{(2\pi)^{n/2} \sqrt{\det \Sigma}} \left\{ \frac{2}{N} \sum_{k=1}^N \exp\left[-\frac{1}{2} (\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})\right] - \frac{\eta}{2^{n/2}} \right\} \rightarrow \max$$

For n=6 there is $n(n+3)/2+1=28$ fitting parameters – convergence too slow

Rigorous Iterative Procedure

By differentiating $M(\bar{z}, \Sigma, \eta)$ w.r.t. fitting parameters we recover equations which can be solved iteratively.

For average values of coordinates the equations coincide with heuristic ones with $\alpha=1$

$$\bar{z}_i = \frac{\sum_{k=1}^N z_i^{(k)} \exp[-\frac{1}{2}(\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})]}{\sum_{k=1}^N \exp[-\frac{1}{2}(\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})]}, \quad \zeta_i^{(k)} = z_i^{(k)} - \bar{z}_i$$

$$\left(\frac{1}{N} \sum_{k=1}^N \dots \rightarrow \sum_{k=1}^N w_k \dots / \sum_{k=1}^N w_k \right)$$

for weighted particles)

We can keep η fixed (i.e. set the fraction of particles taken into account)

Then for Σ - matrix we get

$$\Sigma_{ij} = \frac{1}{N} \sum_{k=1}^N \zeta_i^{(k)} \zeta_j^{(k)} \exp[-\frac{1}{2}(\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})] / \left(\frac{1}{N} \sum_{k=1}^N \exp[-\frac{1}{2}(\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})] - \frac{\eta}{2^{n/2+1}} \right)$$

For $\eta \rightarrow 1$ some damping is necessary in $n=6$ case to avoid oscillations:

$$\Sigma^{(i)} = (1-d)\Sigma^{(i-1)} + d\Sigma^{(formula)}, \quad d \approx 0.8$$

(Mathematics is presented in the cited MAP note)

Rigorous Iterative Procedure (cont'd)

We can try to find the optimal fraction of particles η for the fit.

From equation $\frac{d}{d\eta} M(\bar{z}, \Sigma, \eta) = 0$ we get

$$\eta = \frac{2^{n/2}}{N} \sum_{k=1}^N \exp\left[-\frac{1}{2} (\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})\right], \quad \zeta_i^{(k)} = z_i^{(k)} - \bar{z}_i$$

Equations for average values of coordinates remain the same,

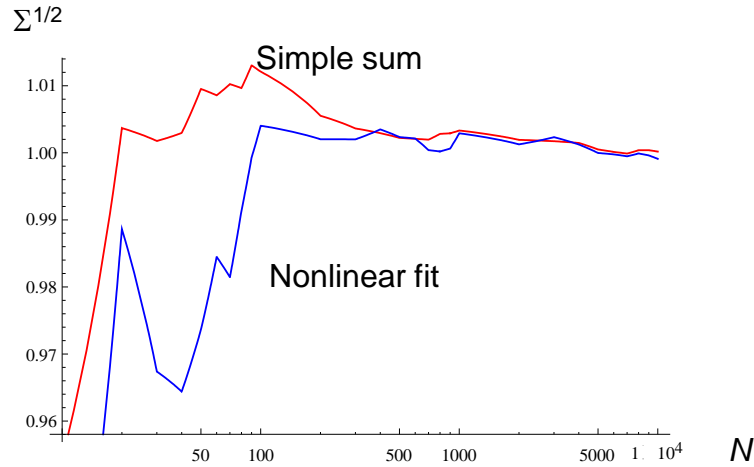
whereas for Σ - matrix we obtain expression with an extra factor of 2 (!) compared to the heuristic one

$$\Sigma_{ij} = 2 \sum_{k=1}^N \zeta_i^{(k)} \zeta_j^{(k)} \exp\left[-\frac{1}{2} (\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})\right] / \sum_{k=1}^N \exp\left[-\frac{1}{2} (\underline{\zeta}^{(k)}, \Sigma^{-1} \underline{\zeta}^{(k)})\right]$$

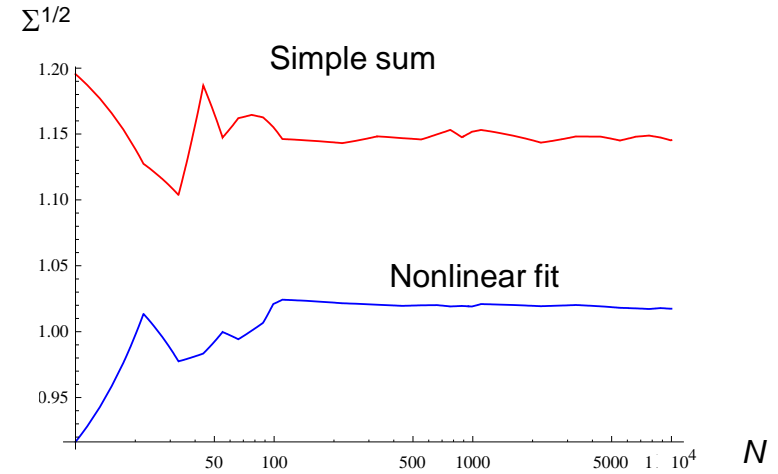
Damping is not necessary in this case.

For $n=6$ in all cases just 20-30 iterations are required to achieve precision $\leq 10^{-6}$, it takes *Mathematica* ~13 seconds with $N=10^4$ on my home PC. For a Fortran or C code it will be a fraction of a second.

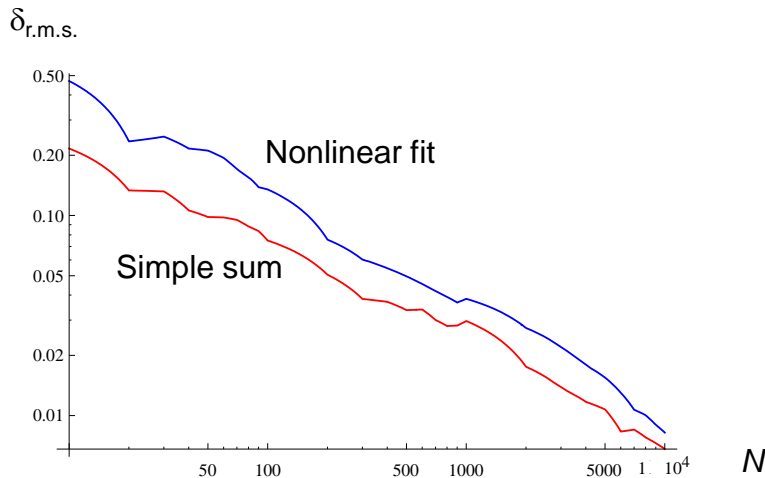
1D Precision Test



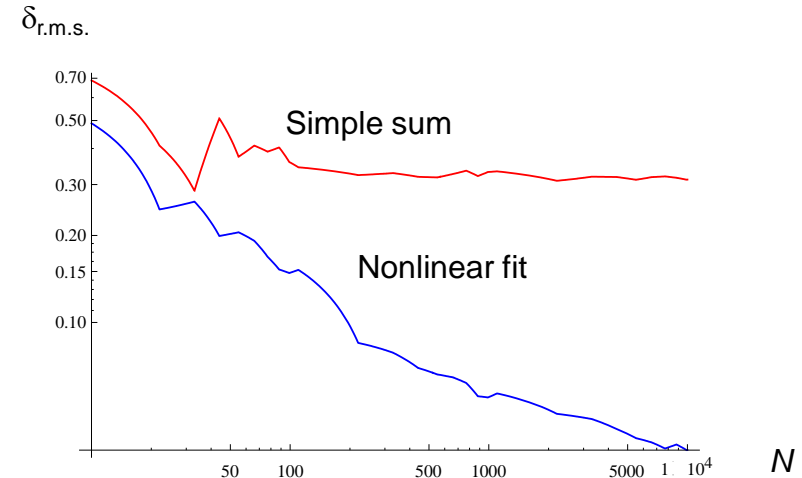
Square root of Σ averaged over 25 realizations of 1D Gaussian distribution with $\sigma=1$ as function of the number of particles N .



Square root of Σ averaged over 25 realizations of superposition of 1D Gaussian distributions with $\sigma=1$ (90%) and $\sigma=3$ (10%)

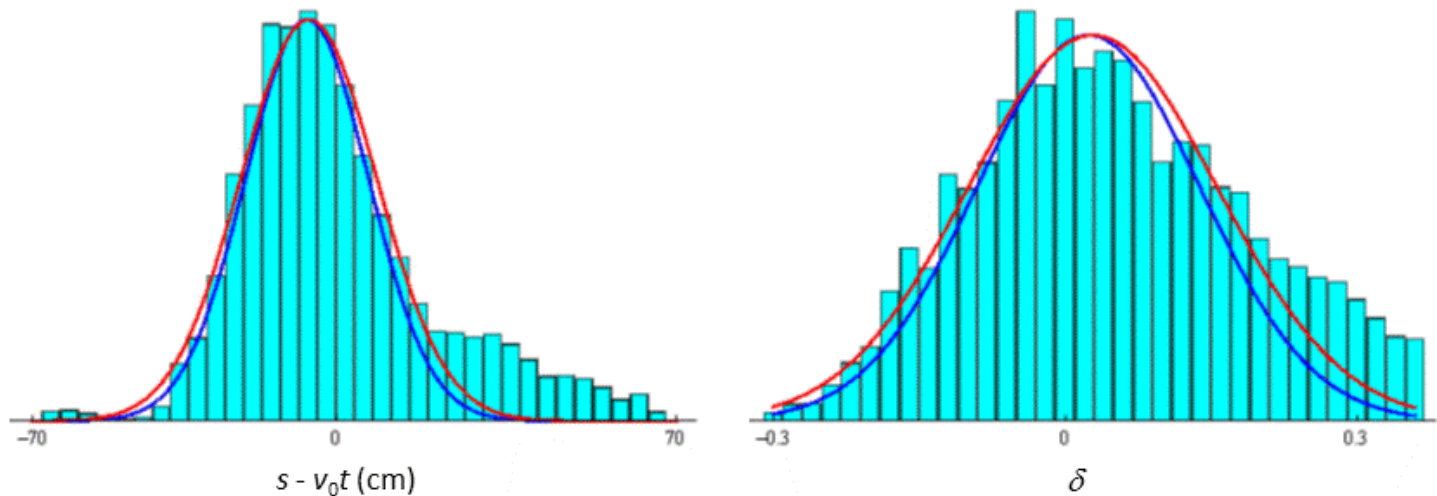


R.m.s. error in $\Sigma^{1/2}$ from above



R.m.s. error in $\Sigma^{1/2}$ from above

Realistic Example



Projections onto the longitudinal coordinate (left) and δ (right) of the original particle distribution (cyan bars) and of its Gaussian fit with $\eta = 1$ and $\eta = \eta_{\text{fit}}$ (red and blue solid lines respectively).

N.B. Projection of the Gaussian distribution onto the m^{th} axis in a multidimensional case is proportional to

$$\exp\left\{-\frac{1}{2}(\Sigma^{-1})_{mm}\zeta_m^2[2 - (\Sigma^{-1})_{mm}\Sigma_{mm}]\right\}$$

Eigen-Emittances from Σ - matrix

With Σ - matrix known, how to find the normal mode emittances?

- Σ - matrix has positive eigenvalues but they are useless unless the matrix of transformation to diagonal form is symplectic (generally not the case)
- solution suggested by theory developed by V.Lebedev & A.Bogacz :

Consider a product $\Omega = S\Sigma^{-1}$ of inverse Σ - matrix and symplectic unity matrix

$$S = \begin{pmatrix} 0 & 1 & 0 & 0 & 0 & 0 \\ -1 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 1 \\ 0 & 0 & 0 & 0 & -1 & 0 \end{pmatrix},$$

Matrix Ω has purely imaginary eigenvalues which are inverse eigen-emittances :

$$\lambda_{2m-1} = -\frac{i}{\mathcal{E}_m}, \quad \lambda_{2m} = \frac{i}{\mathcal{E}_m}, \quad m = 1, 2, 3$$

(Again, mathematics is presented in the cited MAP note)

Eigen-Vectors of Matrix Ω

Using real and imaginary parts of eigen-vectors $\underline{v}'_i \equiv \text{Re } \underline{v}_i$, $\underline{v}''_i \equiv \text{Im } \underline{v}_i$ as columns we can build a matrix:

$$\mathbf{V} = \{\underline{v}'_1, -\underline{v}''_1, \underline{v}'_3, -\underline{v}''_3, \underline{v}'_5, -\underline{v}''_5\}$$

which is symplectic, $\mathbf{V}^t \mathbf{S} \mathbf{V} = \mathbf{S}$, and brings Ω to diagonal form:

$$\mathbf{V}^{-1} \Omega \mathbf{V} = \mathbf{S} \Xi, \quad \Xi = \text{diag}\left(\frac{1}{\varepsilon_1}, \frac{1}{\varepsilon_1}, \frac{1}{\varepsilon_2}, \frac{1}{\varepsilon_2}, \frac{1}{\varepsilon_3}, \frac{1}{\varepsilon_3}\right).$$

The quadratic form Φ takes the form:

$$\Phi = (\underline{\zeta}, \Sigma^{-1} \underline{\zeta}) \rightarrow (\underline{\xi}, \Xi \underline{\xi}) = \sum_{m=1}^3 \frac{\xi_{2m-1}^2 + \xi_{2m}^2}{\varepsilon_m} = 2 \sum_{m=1}^3 \frac{J_m}{\varepsilon_m}, \quad \underline{\xi} = \mathbf{V}^{-1} \underline{\zeta}$$

Eigen-vectors provide information on β - and dispersion functions :

$$\beta_{xm} = |(\underline{v}_{2m})_1|^2, \quad \beta_{ym} = |(\underline{v}_{2m})_3|^2, \quad \beta_{sm} = |(\underline{v}_{2m})_5|^2, \quad m = 1, 2, 3$$

$$D_x = \frac{x}{\delta} = \frac{V_{16}V_{55} - V_{15}V_{56}}{V_{66}V_{55} - V_{65}V_{56}}, \quad D_y = \frac{y}{\delta} = \frac{V_{36}V_{55} - V_{35}V_{56}}{V_{66}V_{55} - V_{65}V_{56}}.$$

Exponential Fit

If the optics is known – and therefore matrix V of eigen-vectors – we can find action variables of particles:

$$J_m = \frac{1}{2} (\xi_{2m-1}^2 + \xi_{2m}^2), \quad \underline{\xi} = \mathbf{V}^{-1} \underline{\zeta}$$

Ignoring the actual distribution in canonical angles we may look for distribution in the form

$$F = \frac{1}{(2\pi)^3 \varepsilon_1 \varepsilon_2 \varepsilon_3} \exp\left[-\sum_{m=1}^3 J_m / \varepsilon_m\right] = F_1 F_2 F_3$$

The fitting is reduced to 3 one-dimensional exponential fits!

Actually it is better to fit the integrated distribution function:

$$f_m(J_m) = \int_0^{2\pi} d\varphi \int_0^{J_m} F_m(x) dx = 1 - \exp[-J_m / \varepsilon_m]$$

Exponential Fit (cont'd)

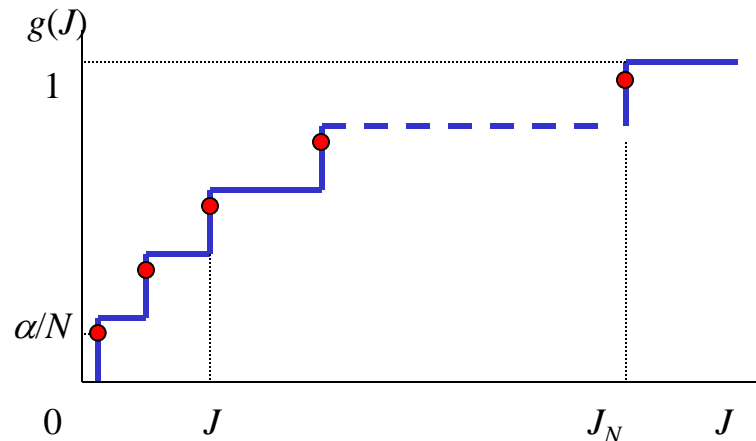
The corresponding part of the Klimontovich distribution (integrated over all other variables)

$$G(J) = \frac{1}{N} \sum_{k=1}^N \delta(J - J_k) \rightarrow g(J) = \int_0^J G(x) dx = \frac{1}{N} \sum_{k=1}^N \theta(J - J_k)$$

where $\theta(x)$ is an asymmetric Heaviside step-function

$$\theta(x) = \begin{cases} 0, & x < 0 \\ 0 < \alpha < 1, & x = 0 \\ 1, & x > 0 \end{cases}$$

Parameter α is empirically adjusted
($\alpha=0.1$ is slightly better than 1 or 0)



Exponential Fit (cont'd)

Now let us numerate particles in the order of increasing J_k

take $\log[1 - g(J)]$ at all $J=J_k$ and equate it with $\log[1 - f(J)] = -J / \varepsilon$

Taking simple average over all particles we get

$$\frac{1}{\varepsilon} = -\frac{1}{N} \sum_{k=1}^N \frac{1}{J_k} \log\left[1 - \frac{k-1+\alpha}{N}\right]$$

The only (complicated) thing to do is to re-order the particles!

This formula gives a precise result in absence of halo, but provides only moderate ($\sim 1/J$) suppression of tails.