

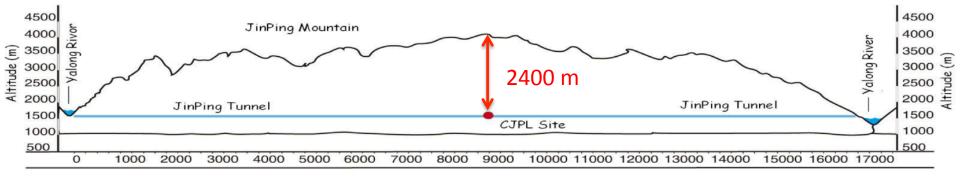
### Dark Matter Direct Detection Experiments in China

Kaixuan Ni Shanghai Jiao Tong University

### A brief history of DM direct detection in China...

- 1992-1996: Beijing (IHEP)-Roma-Saclay collaboration to search for dark matter using NaI detectors
- 1990's: IHEP joined the DAMA (NaI) experiment
- 2000's: Tsinghua University (THU) and IHEP joined KIMS (CsI) and TEXONO (Ge) experiments
- 2009: Shanghai Jiao Tong University (SJTU) joined XENON100 (LXe) experiment
- 2011: IHEP joined the DarkSide (LAr) experiment
- 2009-now: Development of the China Jinping underground laboratory (CJPL) by THU and Yalong River Company
- 2010-now: Two experiments, CDEX (Ge, led by THU) and PandaX (LXe, led by SJTU), being developed at CJPL
- 2010-now: R&D for new target detectors, including CINDMS (CsI) by IHEP.





Excavated through on Aug.8, 2008, Jinping tunnel is the world's second deepest transportation tunnel.



## Excavation and renovation of the CJPL underground lab (6x6x40 m<sup>3</sup>) for two experiments (CDEX/PandaX)





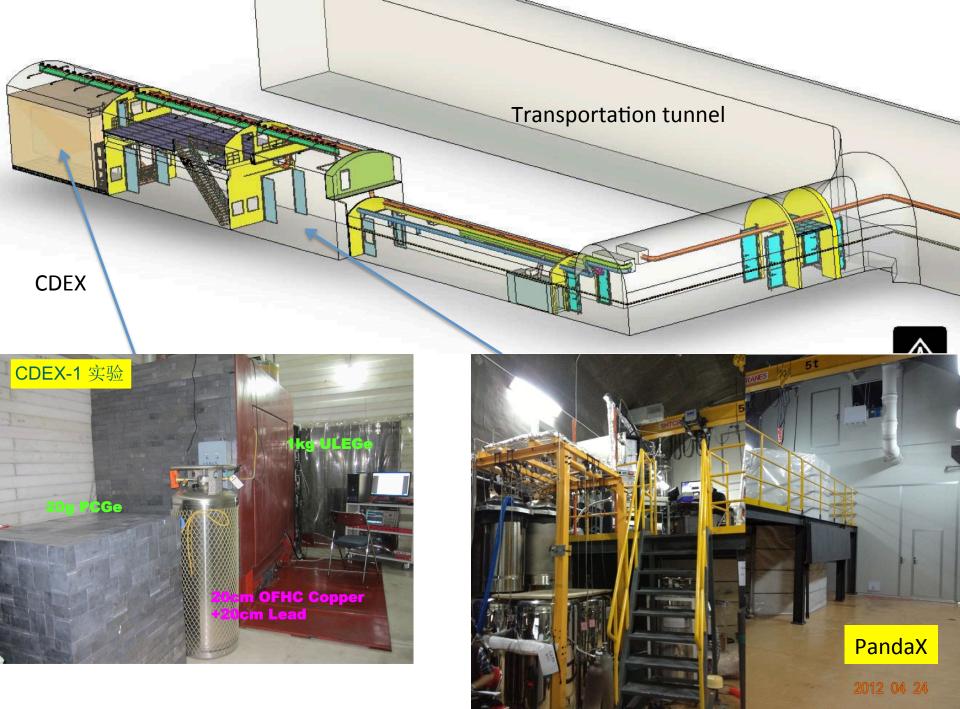


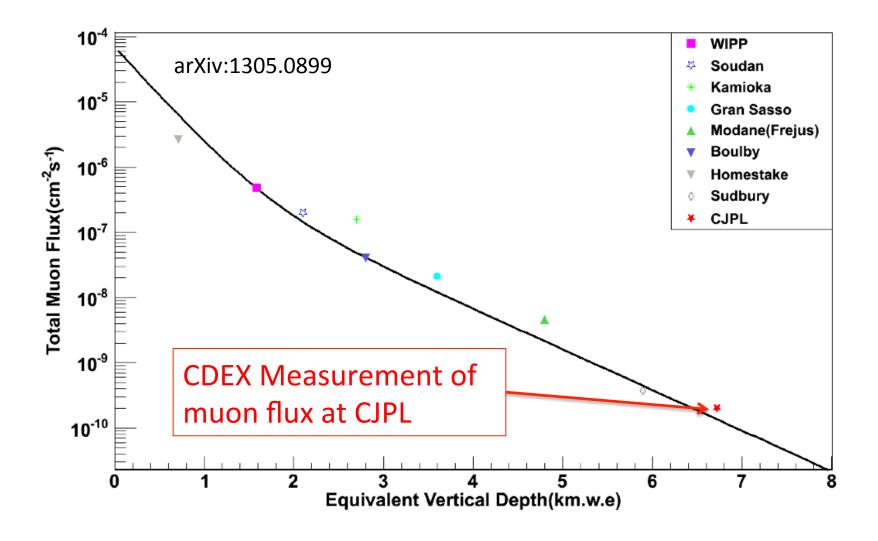






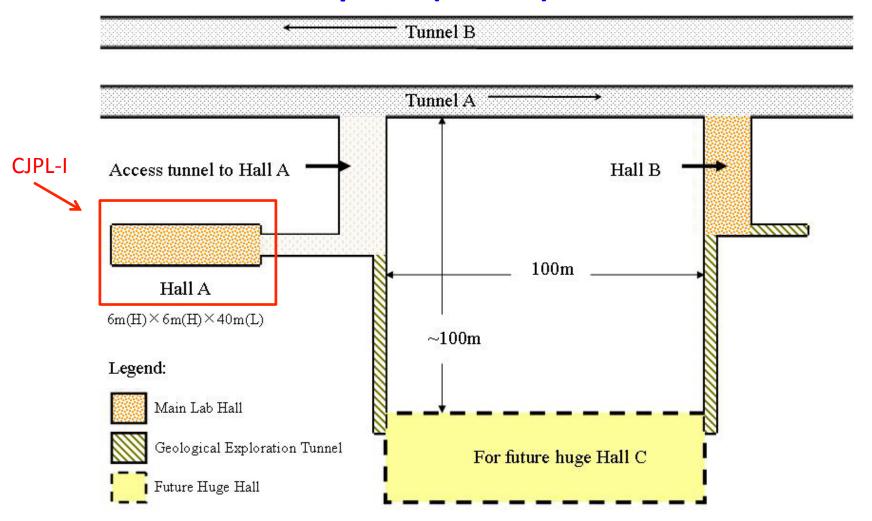






- CJPL muon flux: 2.0 x 10<sup>-10</sup> cm<sup>-2</sup> s<sup>-1</sup> (~60 m<sup>-2</sup> yr<sup>-1</sup>)
- A factor of 100 lower compared to the muon flux at Gran Sasso

### **Future plans (CJPL-II)**



CJPL-II: x20 more space than CJPL-1, 12mx12mx50mxN, under design and plan to be finished by 2015.



HOME

REGISTRATION

LODGING

BULLETINS

IMPORTANT DATES

MEETING ORGANIZATION/ TIMETABLES

SCIENTIFIC PROGRAM

PARALLEL SESSION TOPICS AND CONVENORS

CONTRIBUTED TALKS

### A Town Meeting for the 2nd-phase Development of the China Jinping Underground Laboratory

The town meeting for discussing the second phase development of the China Jinping Underground Laboratory has been moved to Asilomar to allow better coordination with TAUP2013. The meeting will begin at 9:00am on Sunday, September 8, the "arrival day" for TAUP participants, and continue until 5:00pm. Thus those attending TAUP who would also like to participate in the Town Meeting can do so by arranging one additional night in Asilomar; TAUP attendees who arrive in Asilomar on September 8 are also welcome to participate for a partial day. There is no registration fee. Basic information is below:

Date: Sept 8, 2013

Venue: Asilomar, CA, USA (Coordinated with TAUP2013)

Organization Committee: Jianmin Li (Chair, Tsinghua University, China)

Xiangdong Ji (U. Maryland/Shanghai Jiao Tong University, China)

Wick Haxton (LBNL & UC Berkeley, USA)

Joe Wang (LBNL, USA)

### **CDEX**: China Dark matter EXperiment

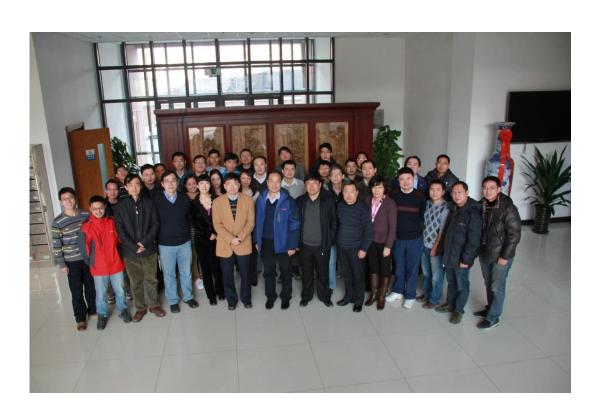
### CDEX target:

Direct detection of cold dark matter with Ton-scale Point-Contact Germanium (PCGe) array detectors with ultra-low energy threshold (<300 eV).

slides from Qian Yue (THU)

# China Darkmatter EXperiment (CDEX) Established in 2009

- Tsinghua University, THU
- Sichuan University, SCU
- Nankai Univeristy, NKU
- China Institute of Atomic Energy, CIAE
- Ertan Hydropower Company, EHDC
- Collaborate with TEXONO and KIMS group.











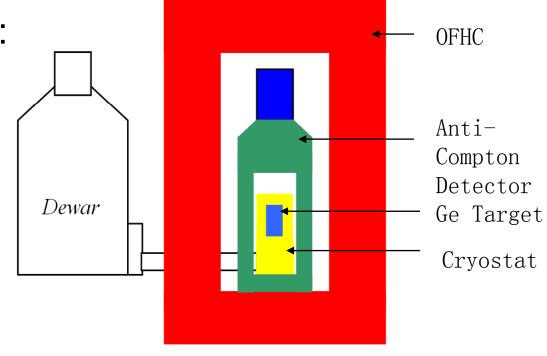


## CDEX-1kg @ CJPL

✓ Point-contact Ge array detector with ultra-low energy threshold (~300eV or less).

✓ Mass of Ge target: 20g, 1000g.

✓ Further ultra-pure crystal serve as active shielding and anti-compton detector.

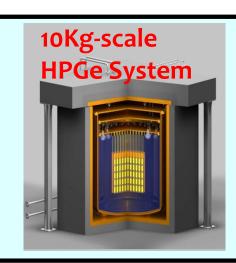


## CDEX-1 Shielding System





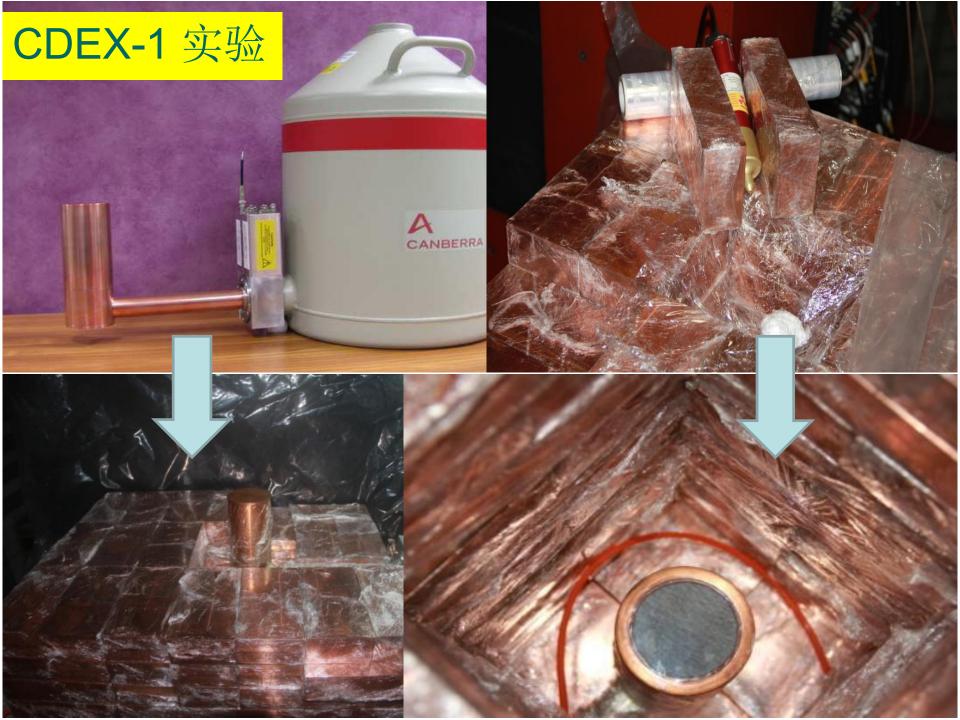
PE shielding room



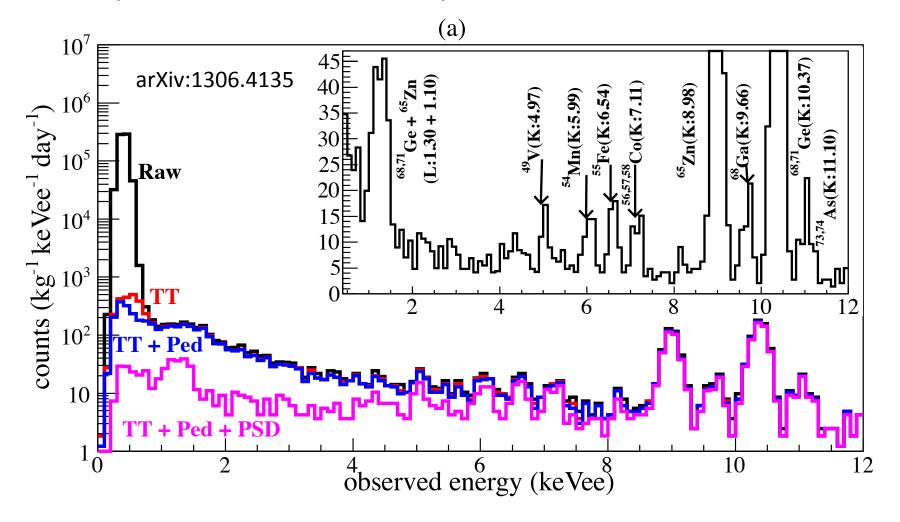
CDEX-1
Shielding
system





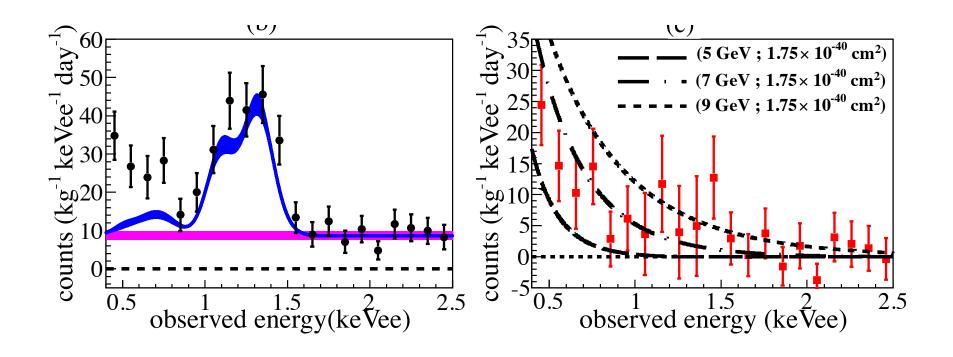


First dark matter search data from CDEX-1: background dominated by cosmogenic radioisotopes, such as <sup>68,71</sup>Ge and are clearly visible.

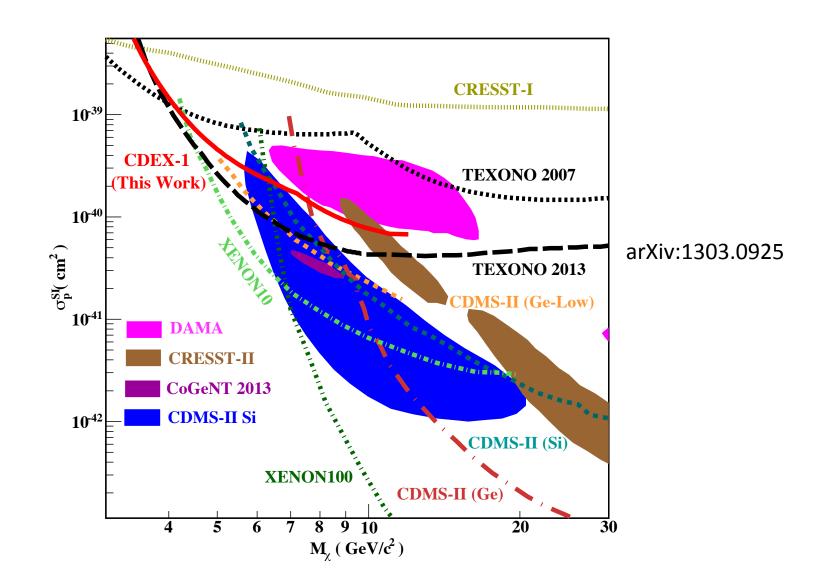


14.6 kg-day exposure400 eVee thresholdNo anti-compton vetoNo bulk/surface background discrimination

At the low energy (0.4-2.4 keVee), the L-shell contribution from the cosmogenic radio-isotopes are subtracted for WIMP analysis.



#### First CDEX limit reported based on the above data (arXiv:1306.4135)

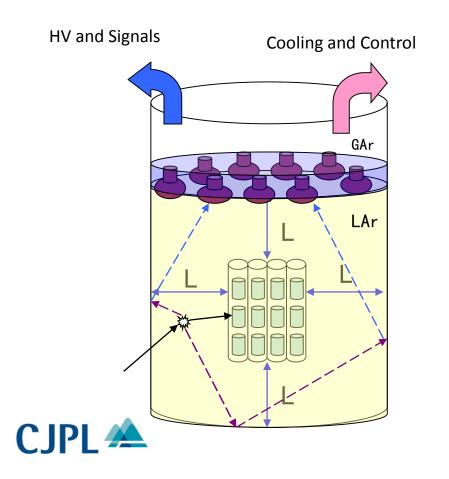


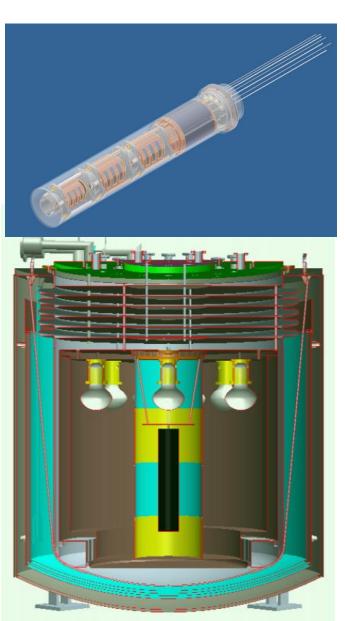
## CDEX-10kg Experiment

LAr: Passive shielding +Active shielding.

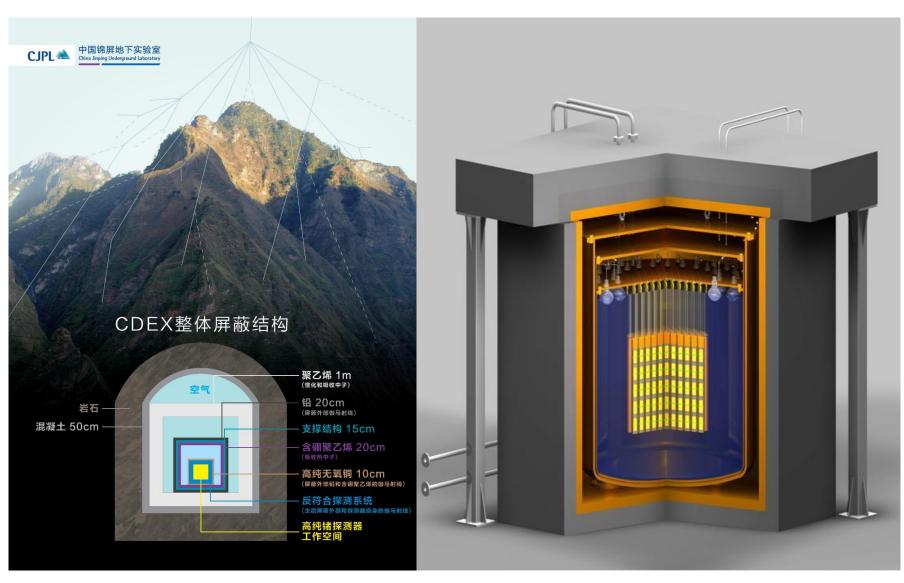
Ge: Encapsuled into copper vacuum tube.

WLS: Transfering 128nm light to~420nm light.





## CDEX-1T plan



### PandaX: Particle AND Astrophysical Xenon experiment

The goal is to build a large-mass two-phase xenon detector with ultra-low background for dark matter and neutrino-less double beta decay searches.

The initial experiment is optimized for scintillation light detection to enhance the sensitivity to low mass WIMPs, while has the capacity to upgrade to a ton-scale experiment.



## PANDAX at CJPL



Shanghai Jiao Tong University
Shanghai Institute of Applied Physics
Shandong University
Peking University

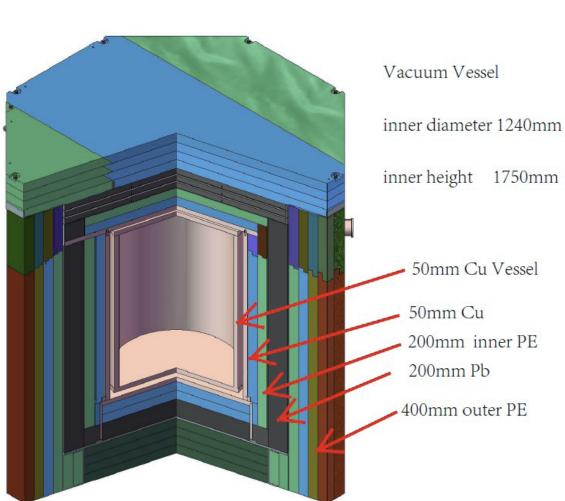


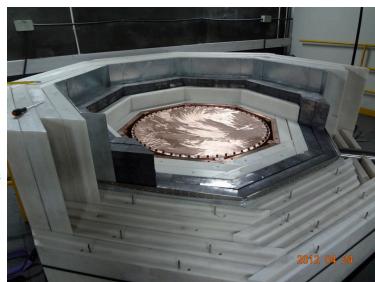
University of Michigan University of Maryland

Yalong River Hydropower Development Company

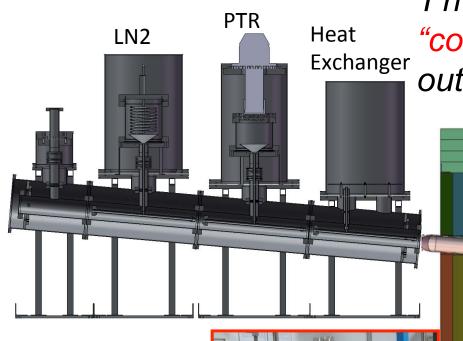
http://pandax.org

Passive shield, which can accommodate a ton-scale detector, is built for PandaX.



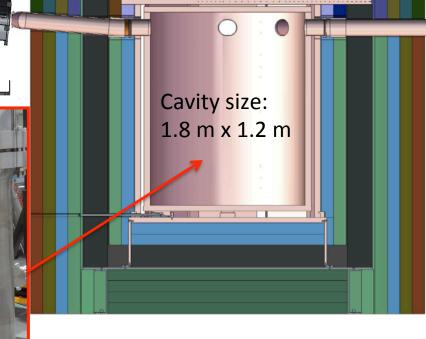






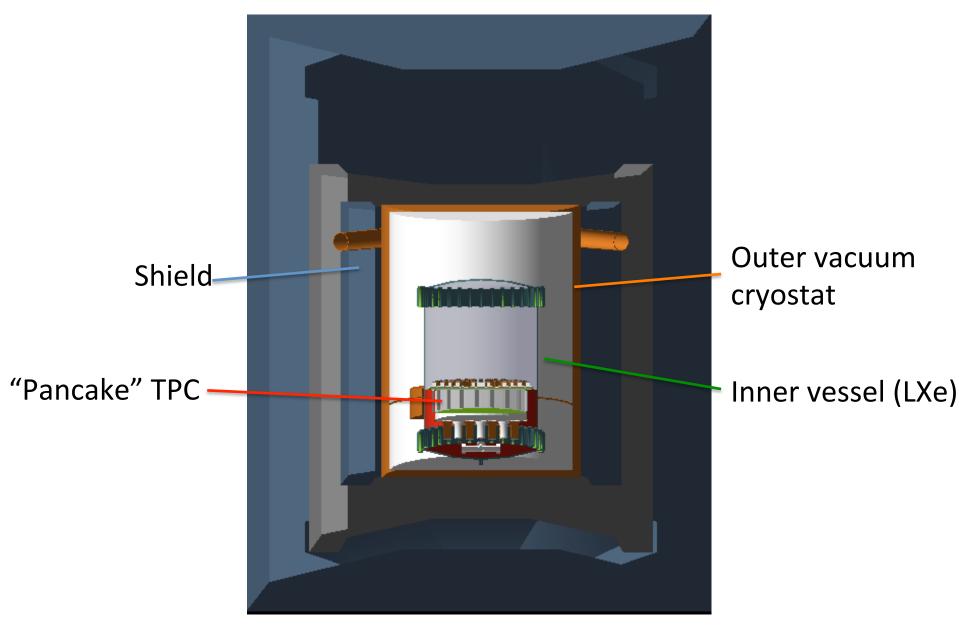
The cryogenic system "cooling bus" is located at outside of the shield.

An inner liquid xenon vessel with a "weir" structure to control the liquid level.



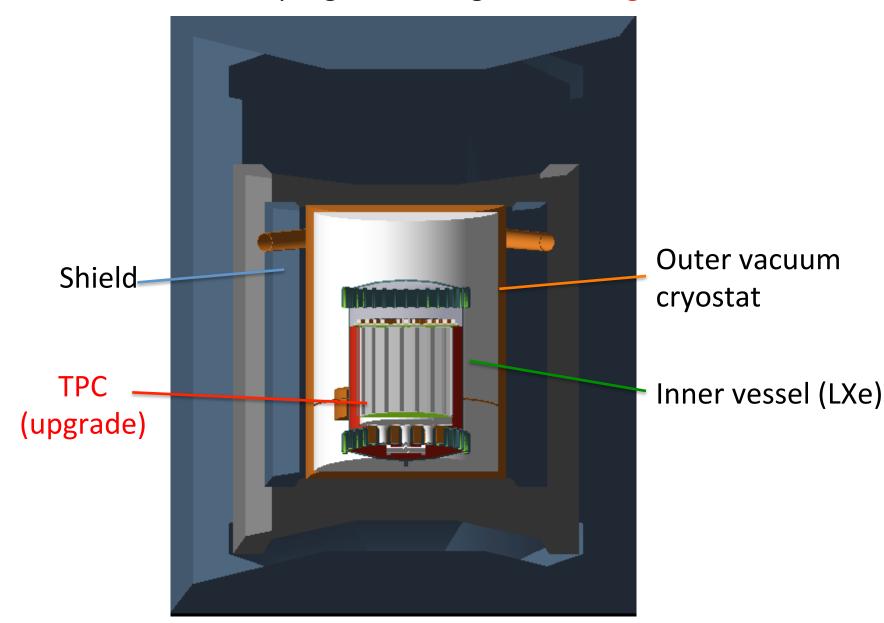


### PandaX will progress through three stages.



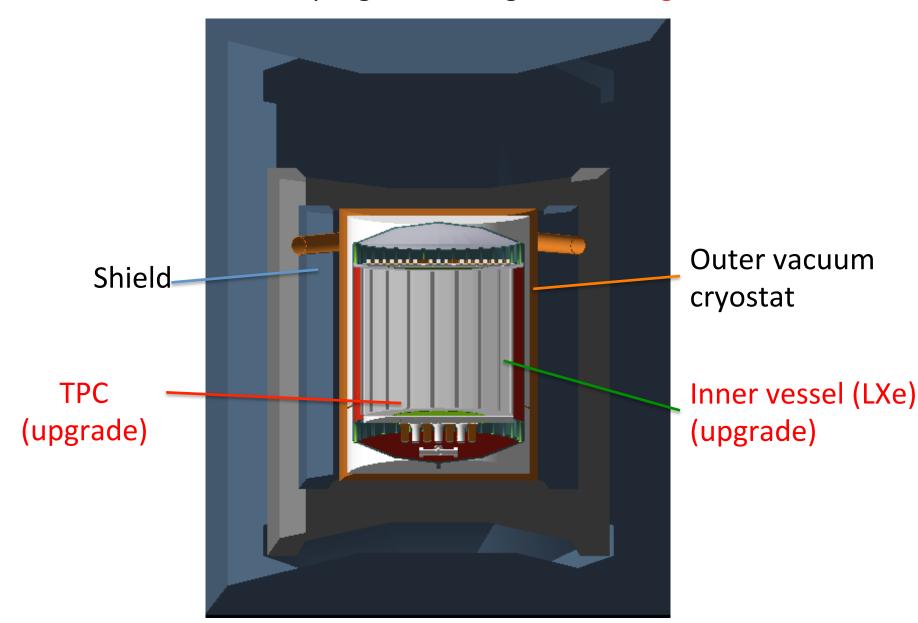
Stage-1a (125 kg target)

### PandaX will progress through three stages.



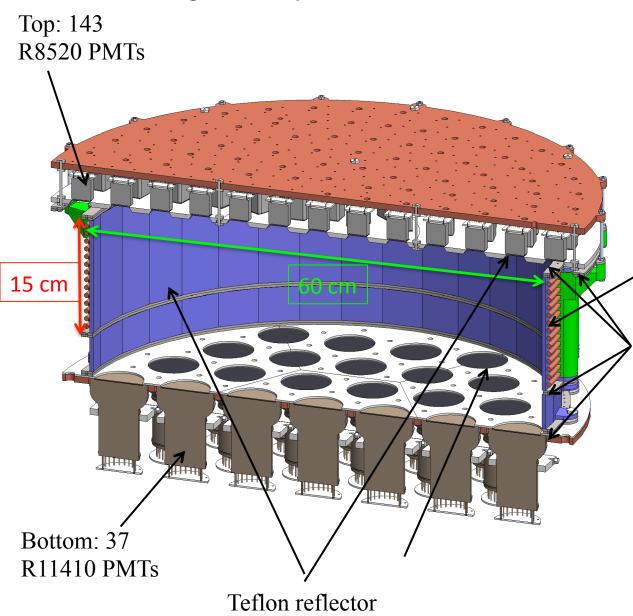
**Stage-1b** (500 kg target)

### PandaX will progress through three stages.



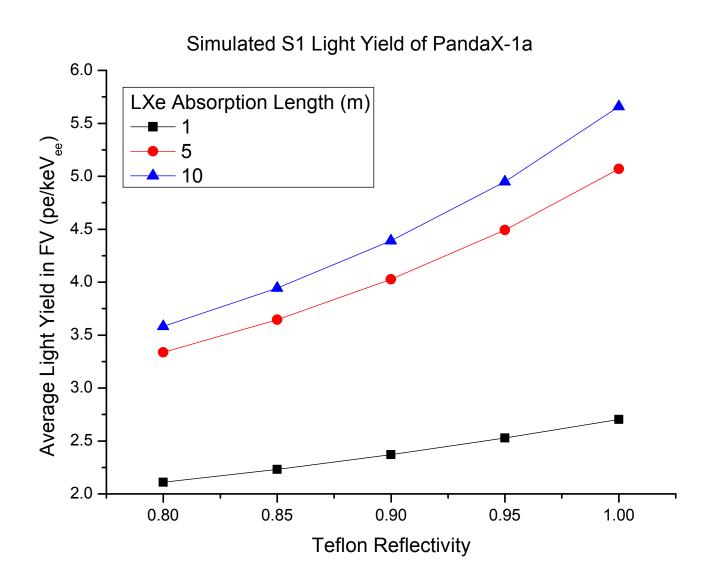
Stage-2 (2.5 ton target)

### Stage-1a: "pancake" TPC

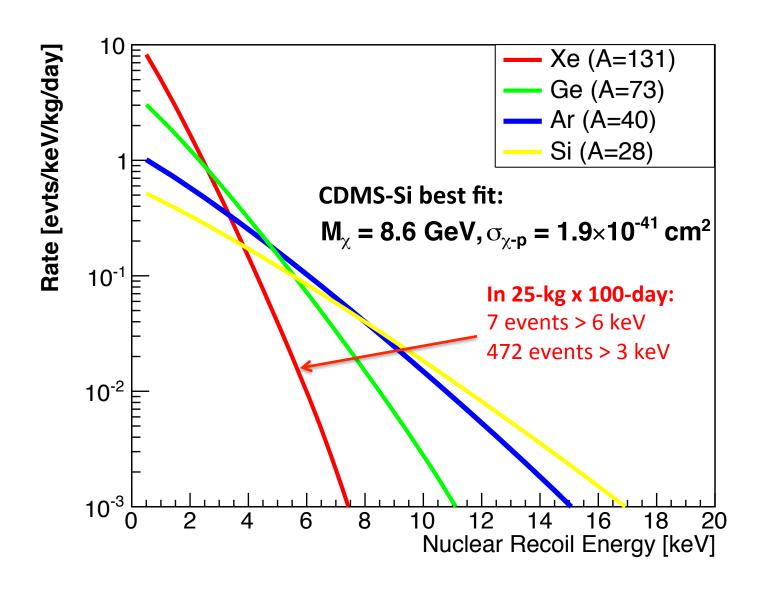




We expect to get 3.3~5.7 pe/keVee for 122 keV gammas at 1 kV/cm with at least 5 m of absorption length.

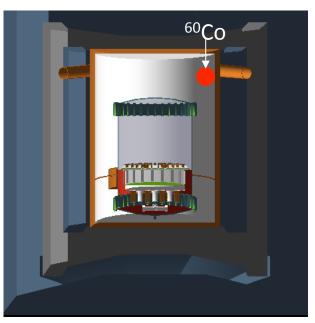


### Expected event rate for "light WIMPs".

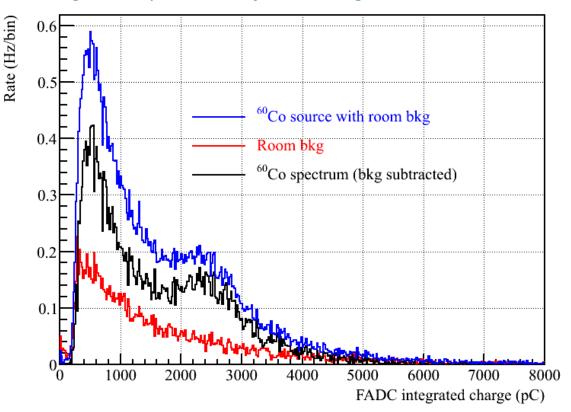


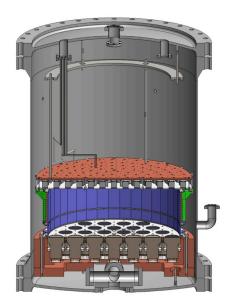
### First liquid xenon test run at CJPL (Apr-May, 2013)



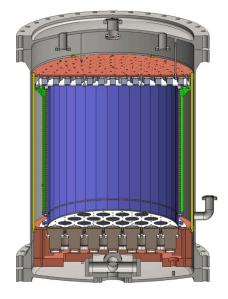


- >400 kg liquid xenon was filled into the detector and maintained stably during the test run.
- Achieved 35 SLPM circulation speed through the purification getter.
- Bottom PMT array operational to observe the first light in liquid xenon from background and source.



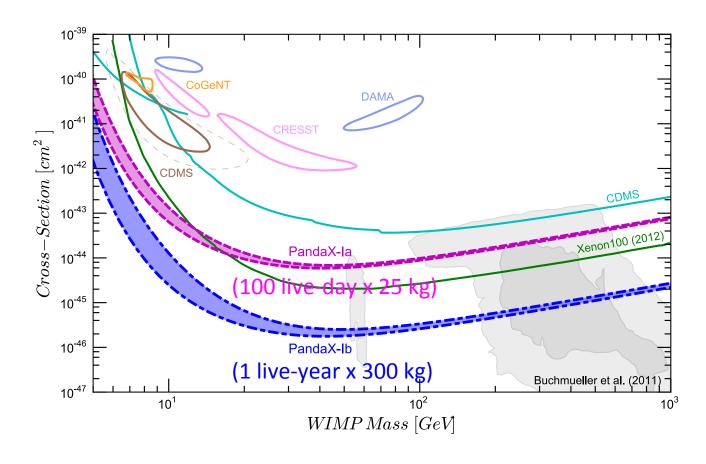


PandaX-1a



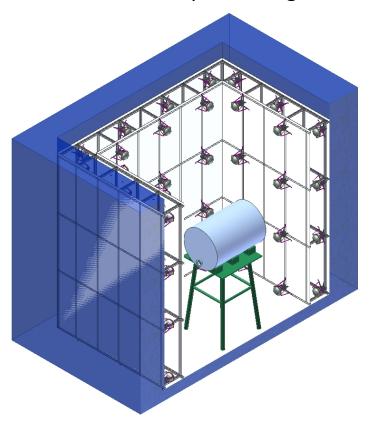
PandaX-1b

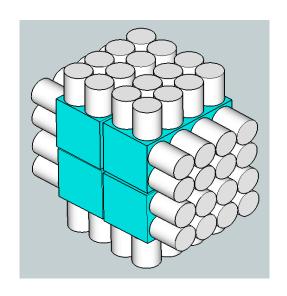
## Projected sensitivity for PandaX-1a and PandaX-1b detectors.



# CsI(Na) Dark Matter Search Experiment CINDMS

Detector conceptual design





Target: CsI(Na) dual light emitting crystal

Gamma background rejection based on pulse-shape discrimination

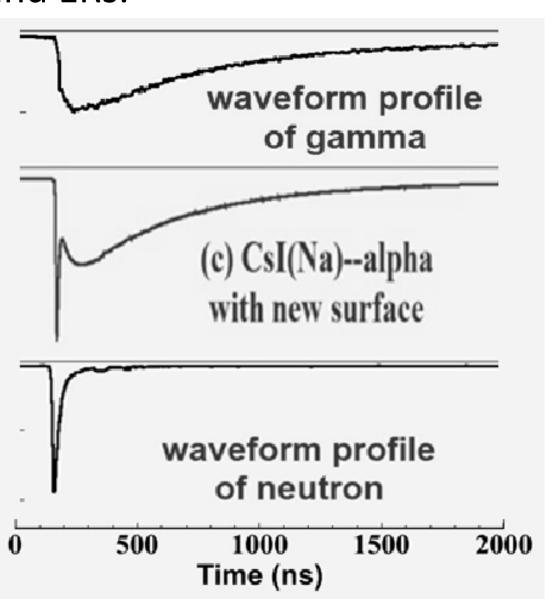
# The Ratio of light emitting from Na<sup>+</sup> and pure CsI is different for NR and ERs.

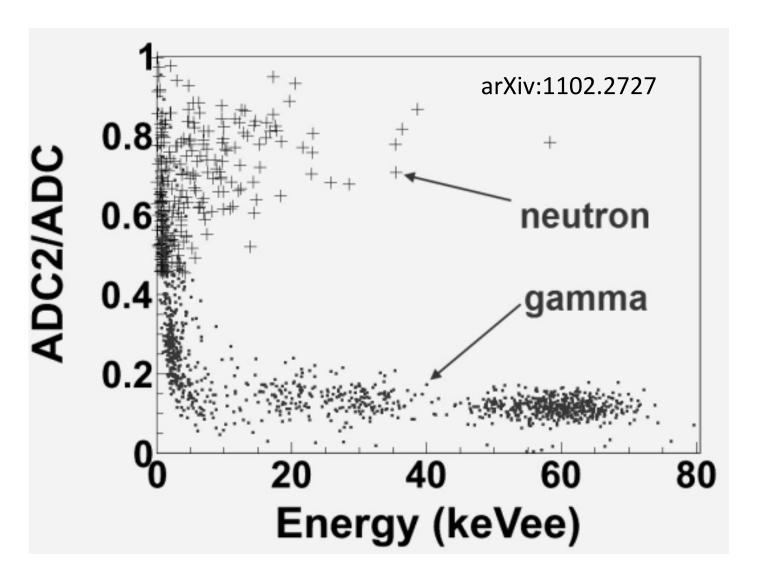
Na<sup>+</sup> light: slow and high yield CsI light: fast and low yield

Average spacing of doped Na ions is 77Å

electron recoils cover more Na due to large range ~ n\*1000Å

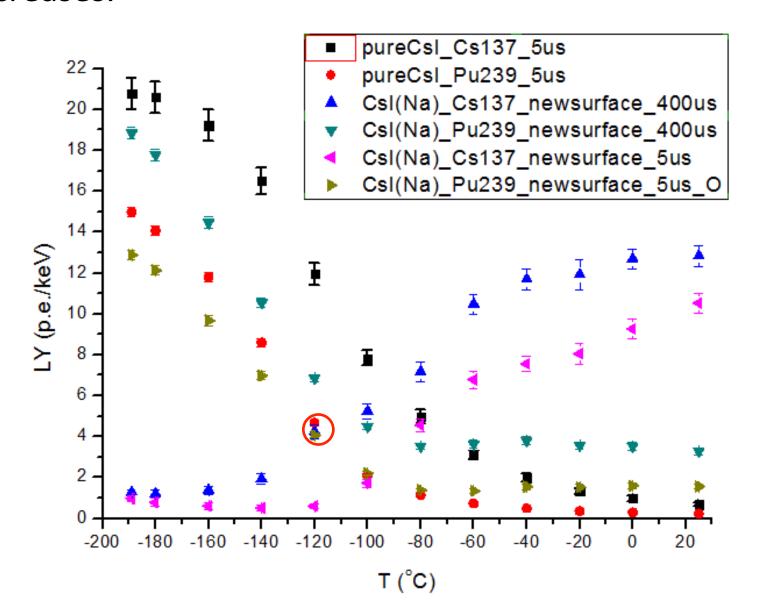
Nuclear recoils cover less Na due to small range ~ n\*100Å



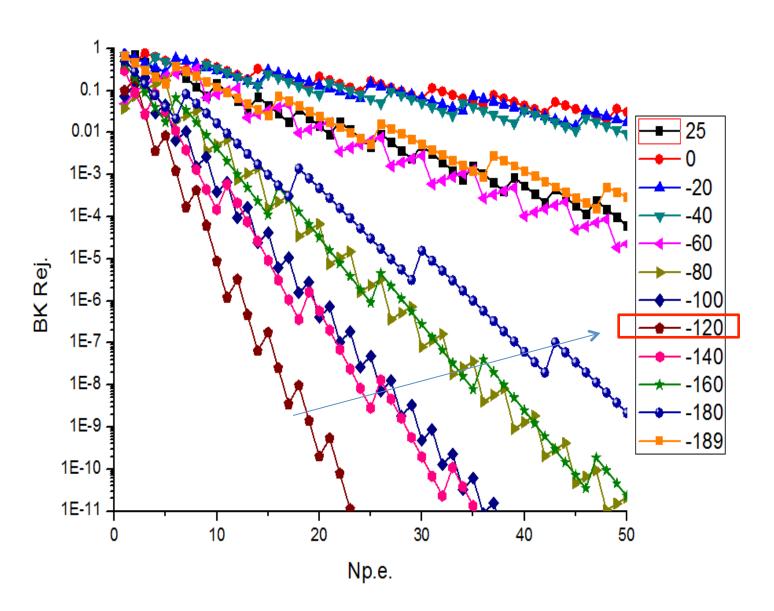


Problem: scintillation efficiency of pure CsI is very low at room temperature

The light yield pure CsI increases as temperature decreases.



### Optimized ER background rejection power at around -120 °C

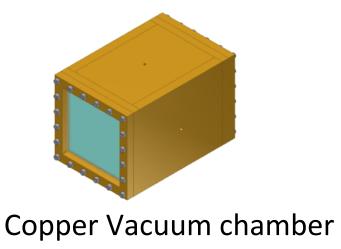


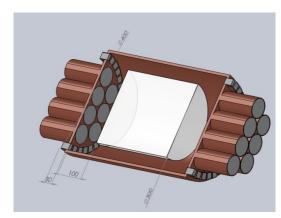
### Plans of CINDMS

2013-2014: verification of the technology

- funded by IHEP for the first 40-kg module to be installed at Daya Bay underground laboratory
- using ultra-pure water for extraction and recrystallization to produce ultralow background Csl(Na) crystals

2015 : ton-scale experiment





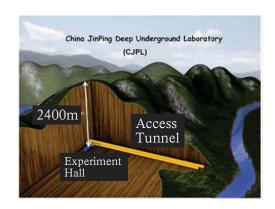
Active shield with LXe

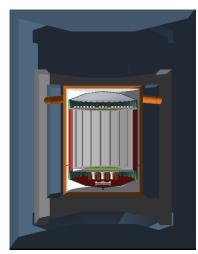
### Summary

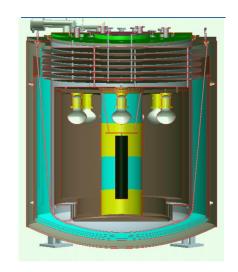
CDEX and PandaX are being constructed towards ton-scale dark matter experiments at CJPL.

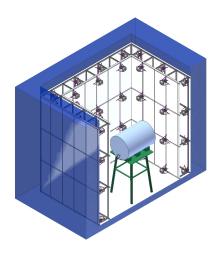
New techniques, such as CINDMS, are being developed to join the dark matter search.

CJPL provides the deepest underground environment for dark matter search. CJPL-II will host multi-target multi-ton scale experiments.









Thanks the above collaborations for their latest updates presented in this talk!