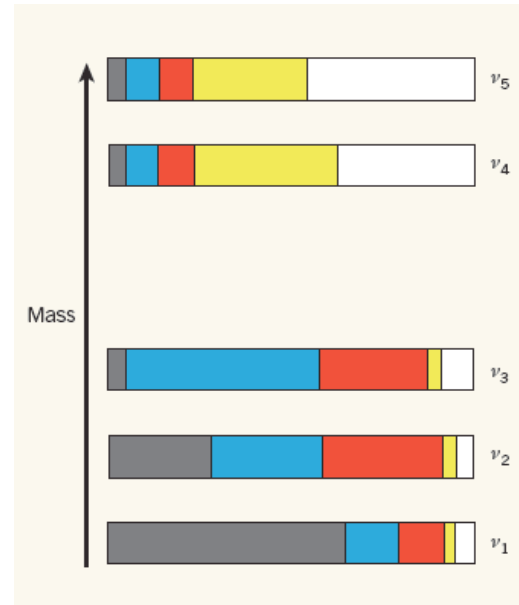
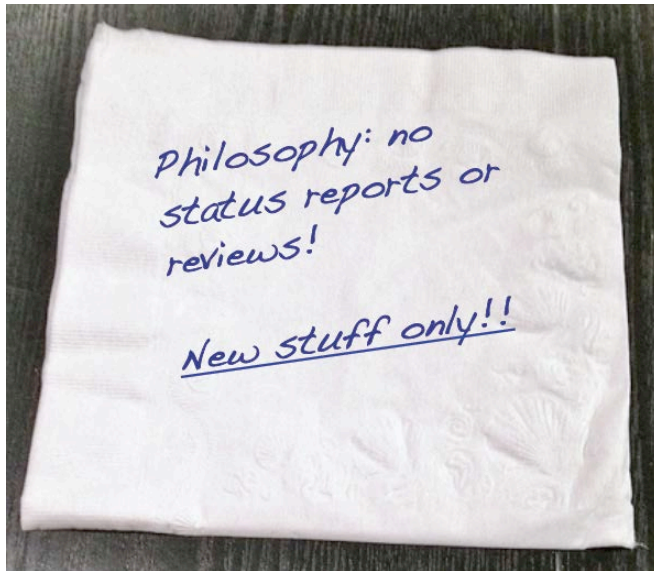


Probing Neutrino Oscillations At Very Short Baselines with Reactors and Radioactive Sources

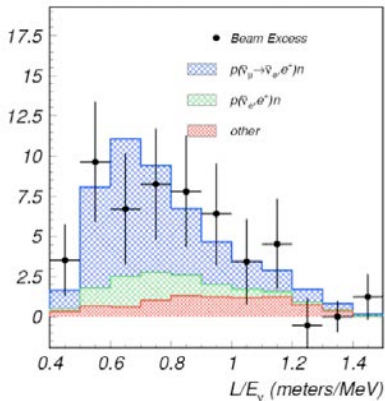


Karsten Heeger
University of Wisconsin

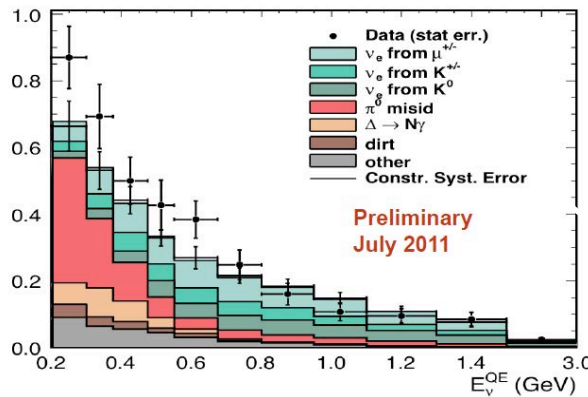
Aspen, February 8, 2013

Neutrino Anomalies and Sterile Neutrinos

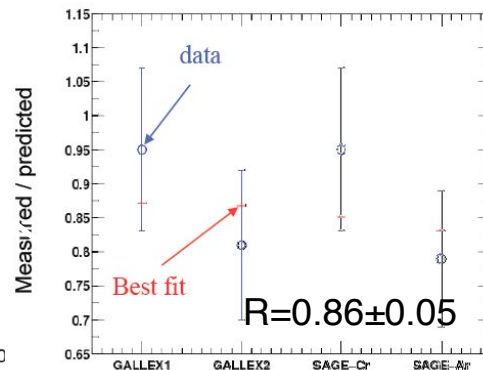
LSND



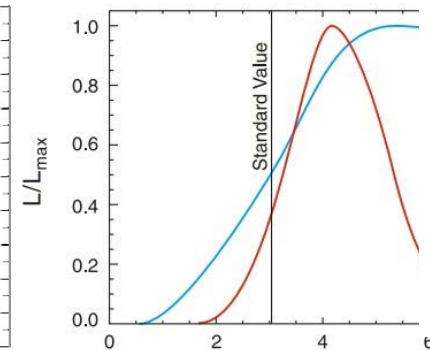
MiniBoone



Ga Source



Cosmology (WMAP)



Anomalies in 3-ν interpretation of global oscillation data

LSND ($\bar{\nu}_e$ appearance)

MiniBoone ($\bar{\nu}_e$ appearance)

Ga anomaly

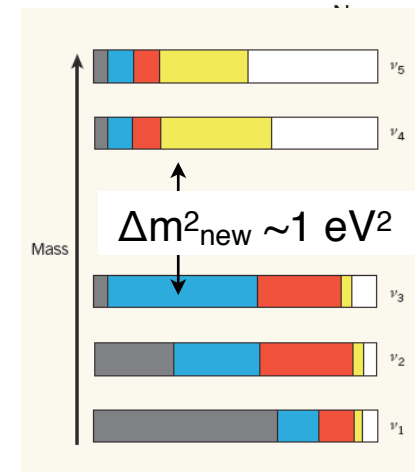
N_{eff} in cosmology

Short-baseline reactor anomaly ($\bar{\nu}_e$ disappearance)

new oscillation signal requires $\Delta m^2 \sim O(1\text{eV}^2)$ and $\sin^2 2\theta > 10^{-3}$
systematics or experimental effect? → need to test each effect

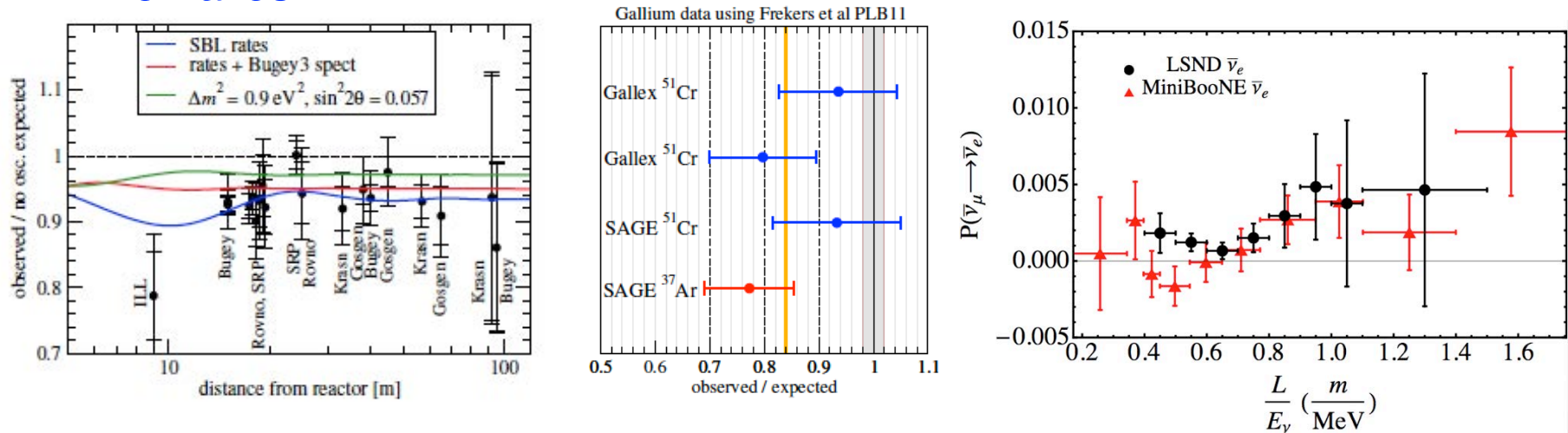
Reactor Anomaly

one of several anomalies, could go away with revised flux prediction based on “old” data, no modern reference experiment

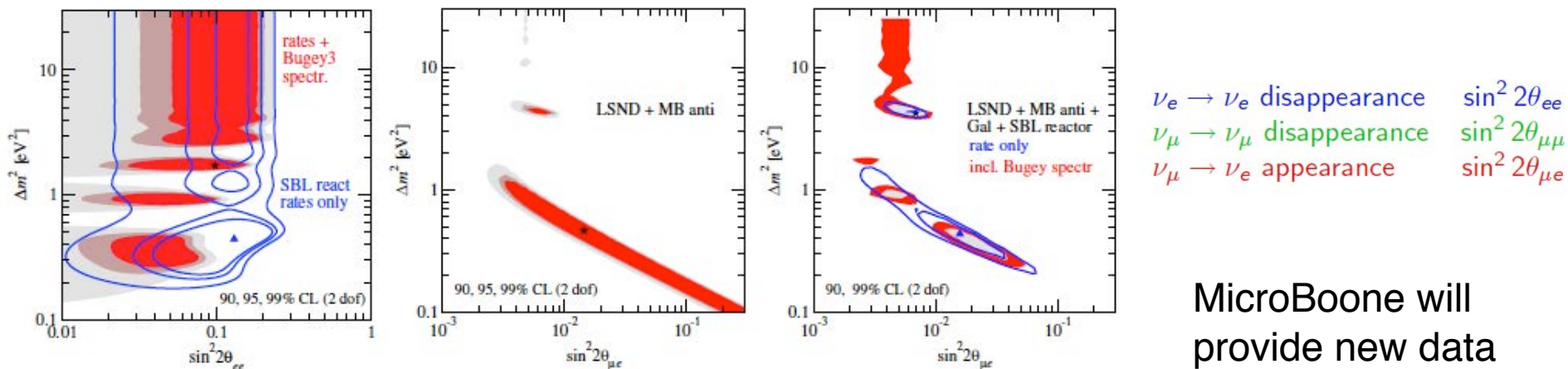


Beyond 3v - Sterile Neutrinos?

Anomalies



Are $\nu_e \rightarrow \nu_e$ and $\nu_\mu \rightarrow \nu_e$ consistent?



strong tension if all three are combined, tension also in 3+2 fit

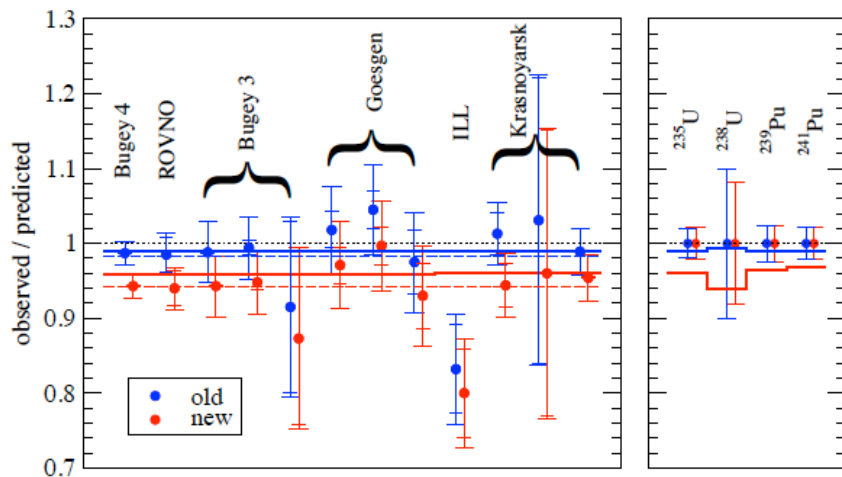
MicroBoone will provide new data starting ~2014

Reactor $\bar{\nu}$ Fluxes

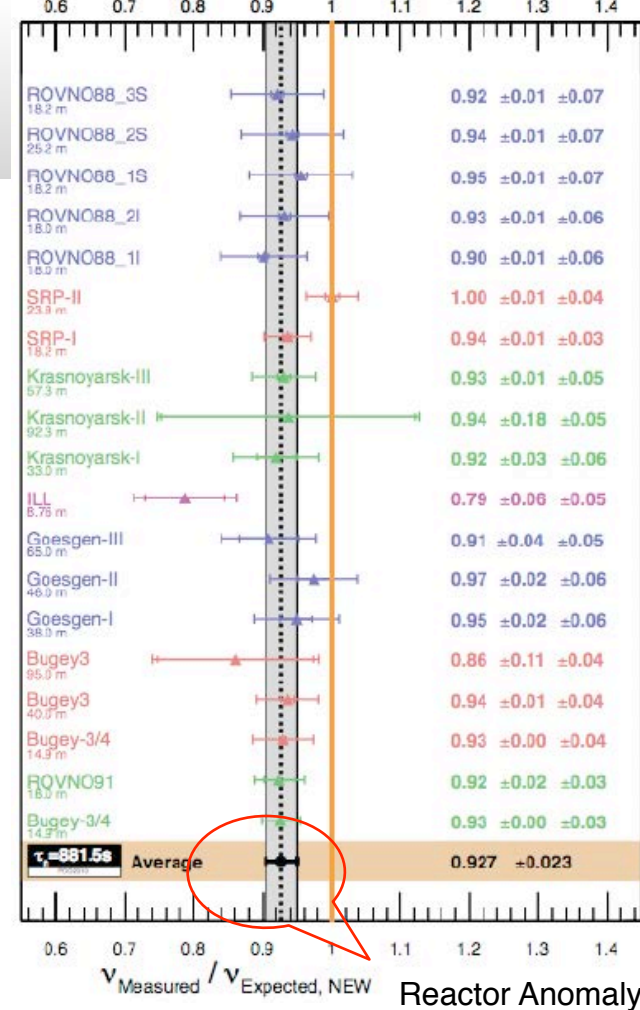
Theory Meets Experiment

Recently the reactor $\bar{\nu}_e$ fluxes have been recalculated

T.A. Mueller et al., [arXiv:1101.2663].; P. Huber, [arXiv:1106.0687].



re-evaluations find higher fluxes by about 3.5%



Two issues:

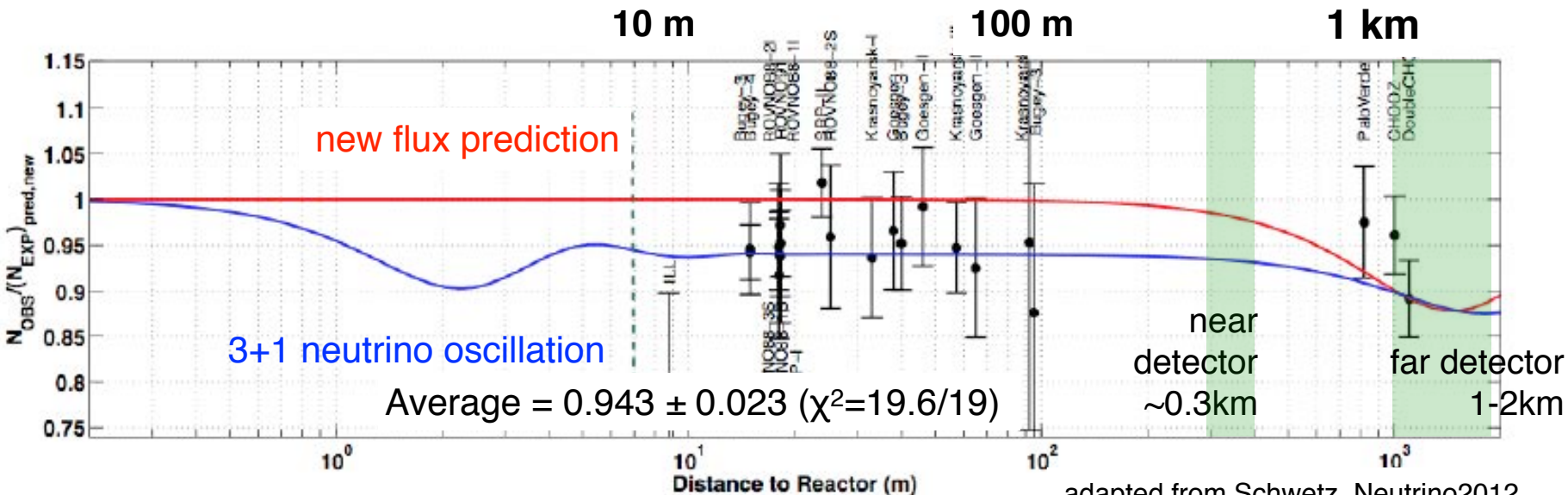
1. Model-dependence of physics determining the increase in the spectra?

- SM physics for GT and Fermi Transitions
- some transitions are forbidden transitions, corrections unknown

2. Overall uncertainties in reactor antineutrino fluxes?

Ref: Mention et al, 1101.2755 (2012 upd)

θ_{13} Experiments and the Reactor Anomaly



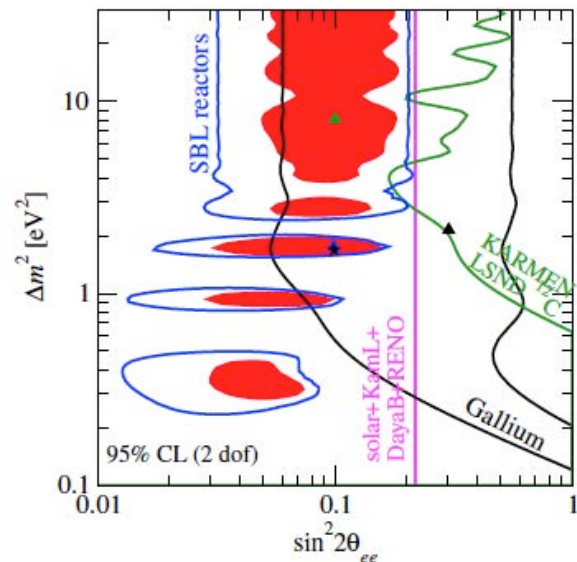
adapted from Schwetz, Neutrino2012

Absolute flux measurement

	Double Chooz	RENO	Daya Bay
detector syst	2.1%	1.5% (correlated)	1.9% (correlated)
reactor syst	1.8%	2.0% (correlated)	3.0% (correlated)

Precision spectral measurement

verify that we understand how to predict reactor \overline{v}_e spectrum



Source Experiments

Sterile ν Oscillation Searches with Radioactive Sources

Place source near or inside existing detector, search for $\bar{\nu}_e$ or ν_e disappearance.

Advantages

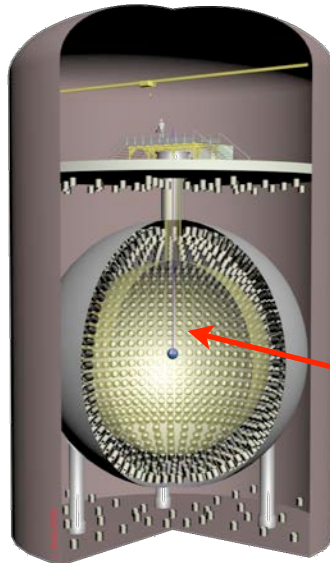
- baseline can be as short as needed
- detectors can be underground to minimize backgrounds
- demonstrate oscillation signature vs baseline and energy
- re-use existing, well-characterized detectors

Challenges

- construct suitable, intense radioactive source
- regulatory and licensing requirements for radioactive source

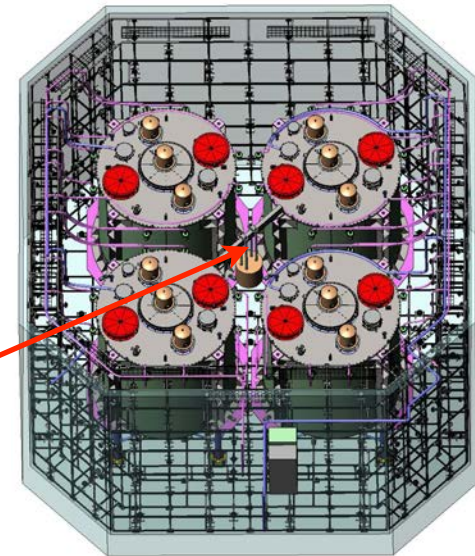
source inside
detector

e.g. CeLAND



source next to
detector

e.g. Daya Bay

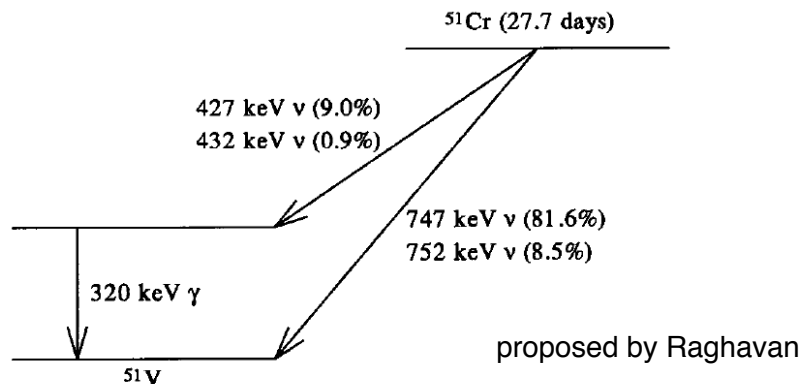


Sterile ν Searches with Very Short Baselines: Sources

A Variety of Sources and Detectors Are Feasible

Sources based on EC
(^{65}Zn , ^{51}Cr , ^{152}Eu , ^{37}Ar)

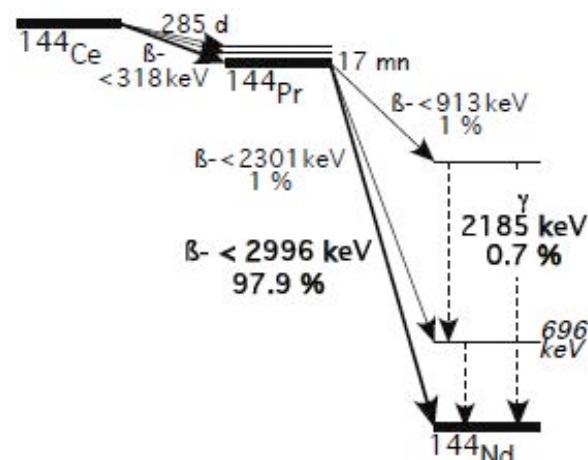
e.g. ^{51}Cr , mono-energetic, ν_e , 750 keV



Decay scheme of ^{51}Cr to ^{51}V through electron capture.

Sources based on beta-decays

e.g. ^{144}Ce - ^{144}Pr , $\bar{\nu}_e$, continuous spectrum



arxiv:1107.2335
Cribier et al

Detection Channels & Proposed Experiments

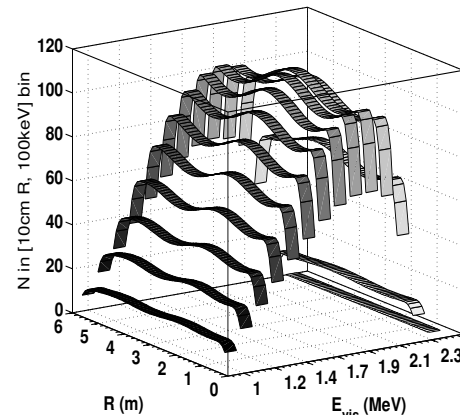
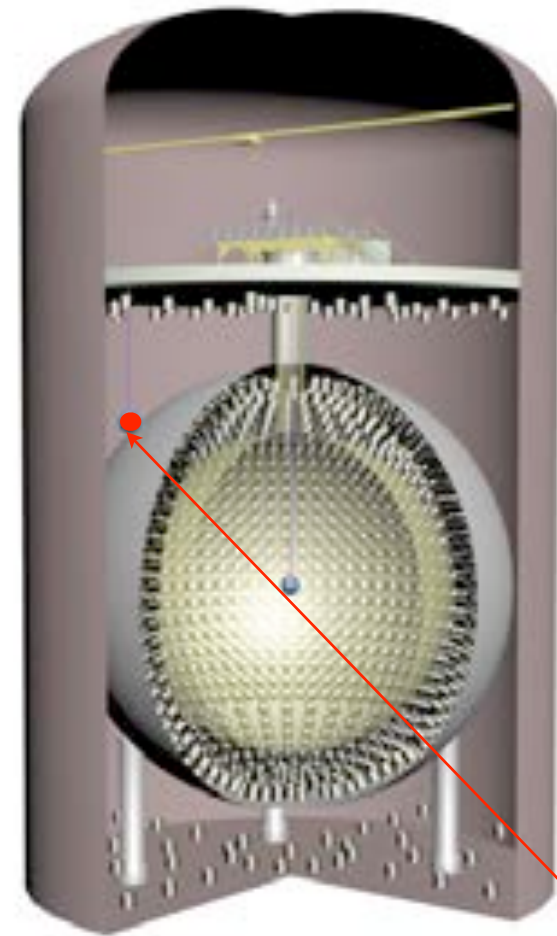
Elastic Scattering: Borexino, SNO+Cr

Charged Current: LENS-Sterile, Baksan, CeLAND, Borexino, Daya Bay

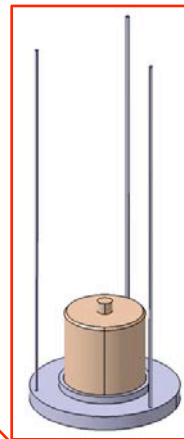
Neutral Current: RICOCHET

CeLAND

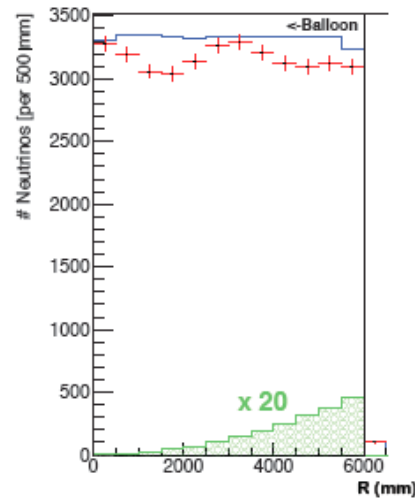
^{144}Ce source inside and outside KamLAND



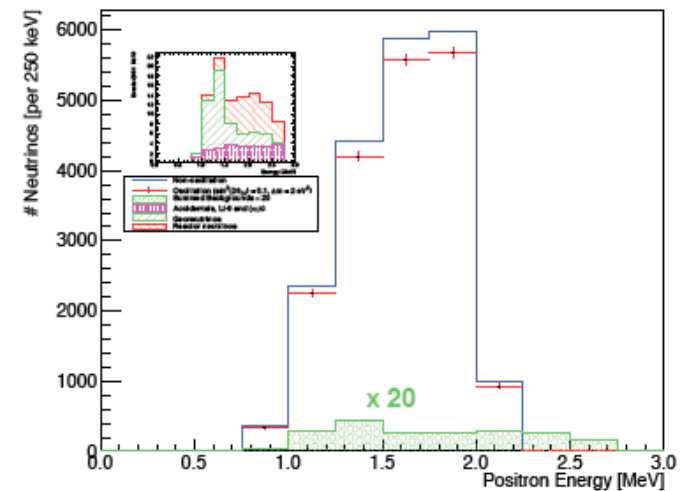
phase 1



Spatial event distribution



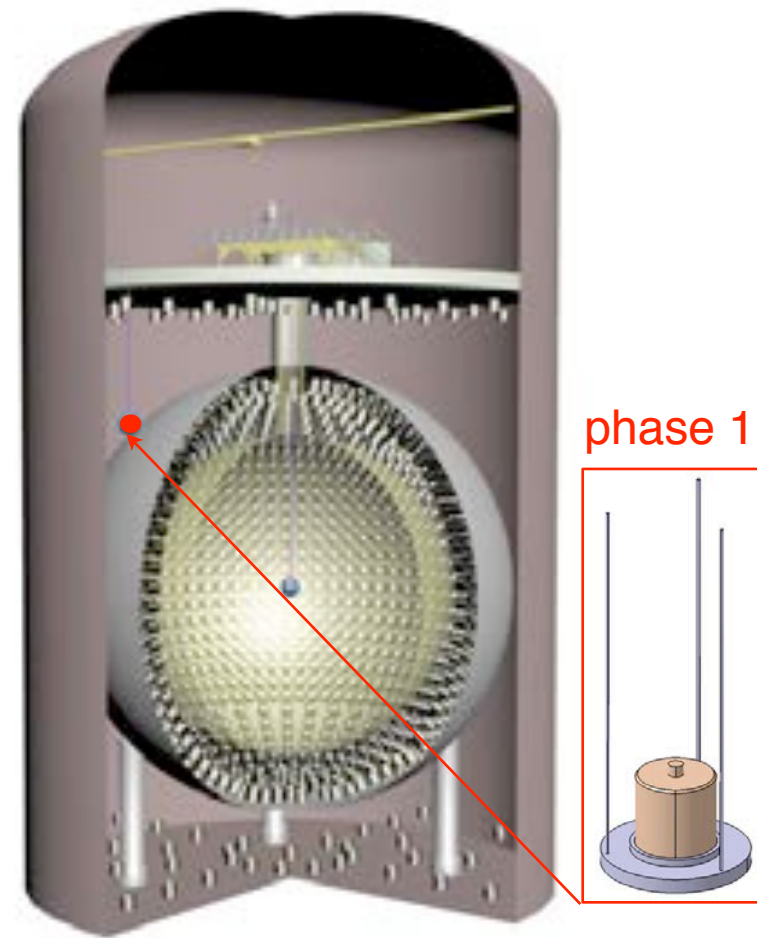
Energy Spectrum



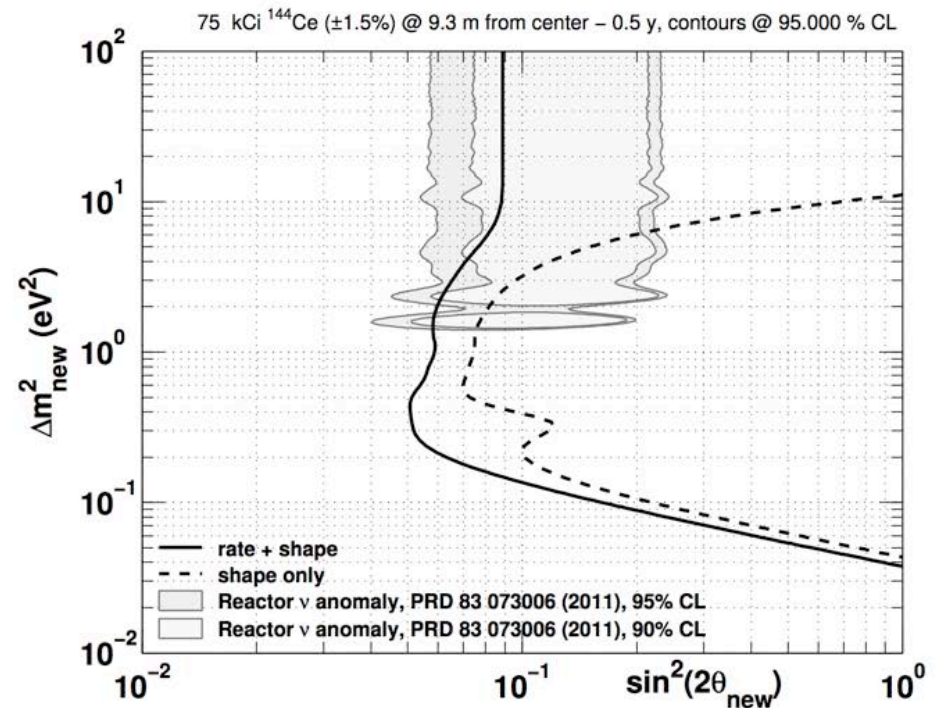
Ref: Lasserre

CeLAND

^{144}Ce source inside and outside KamLAND



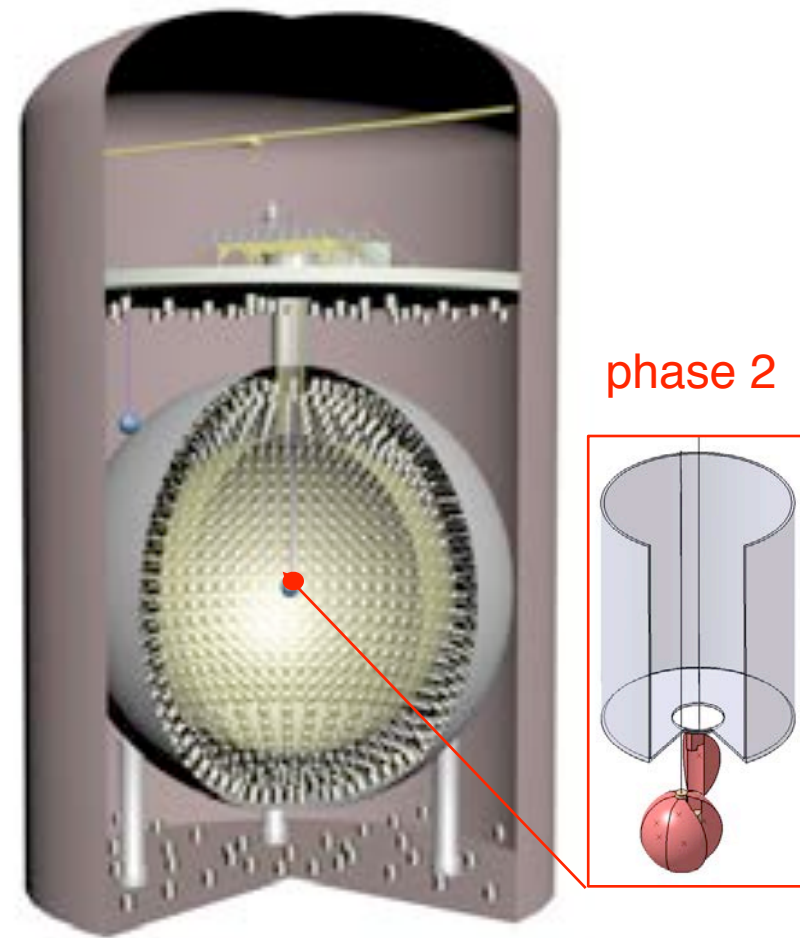
75kCi
0.5 yrs outside, 95%CL



Ref: Lasserre

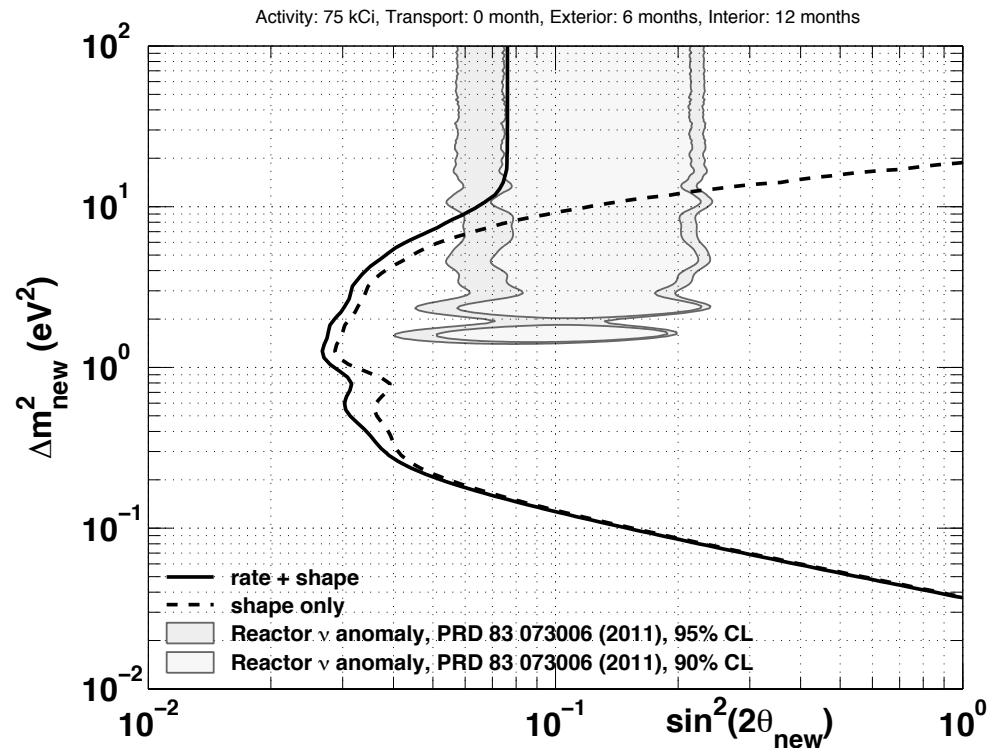
CeLAND

^{144}Ce source inside and outside KamLAND



75kCi

0.5 yr outside + 1 yr in center, 95%CL



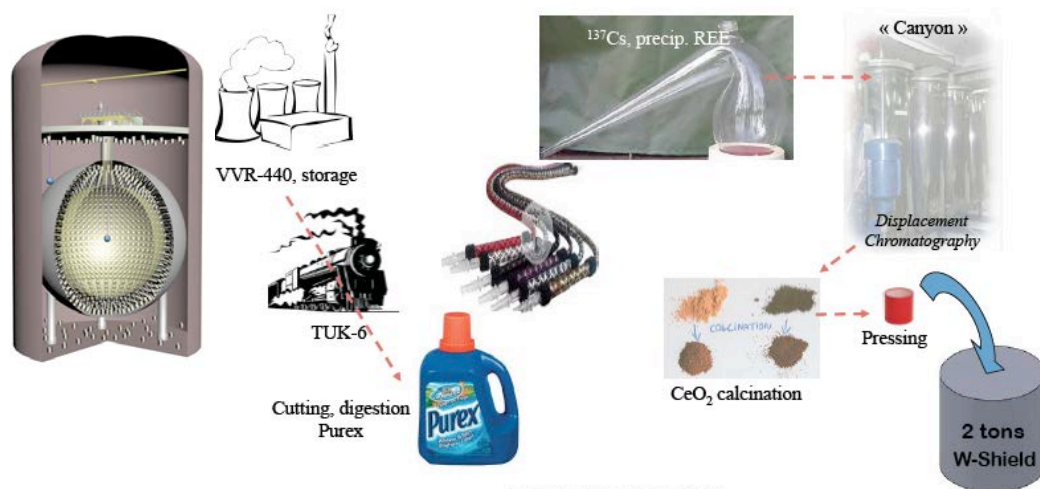
Ref: Lasserre

^{144}Ce - ^{144}Pr Source Production and CeLAND Status

^{144}Ce Production at PA Mayak: 2014

erc

75 kCi (2.77 PBq), 4 kg of CeO_2 ($\rho = 4 \text{ g/cm}^3$), 600 W



Th. Lasserre – vMass 2013

^{144}Ce Source Transportation: 2014/15

erc



- Ship radioprotection shield (Sh_{radio}) to PA Mayak (2014)
- Fit the ^{144}Ce source (Ce) inside first Sh_{radio} at Mayak (2014/15)
- Use certified container for further transportation (Conta)

Th. Lasserre – vMass 2013

44

Source Production

- agreements for production between CEA and Rosatom
- currently working on the source production (samples will be available soon), transportation (of course a real issue), activity measurement (<1% seems feasible)

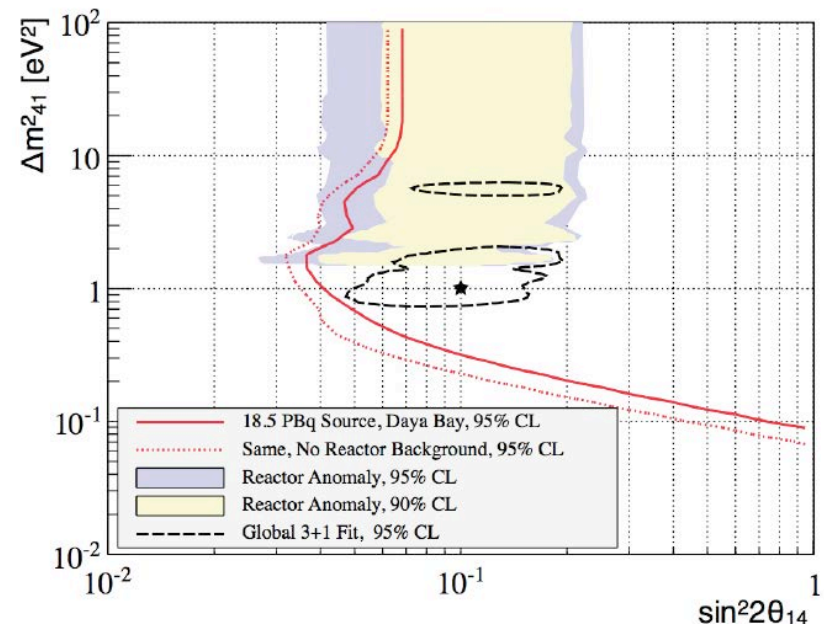
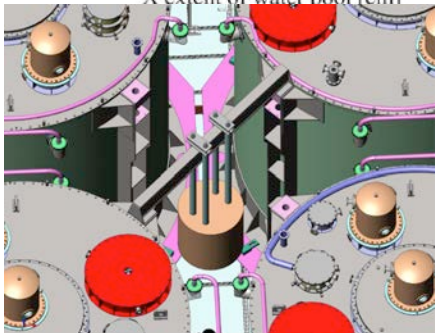
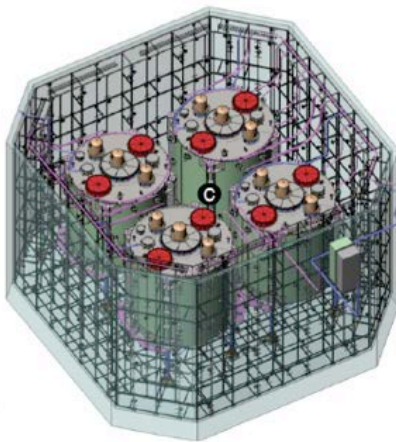
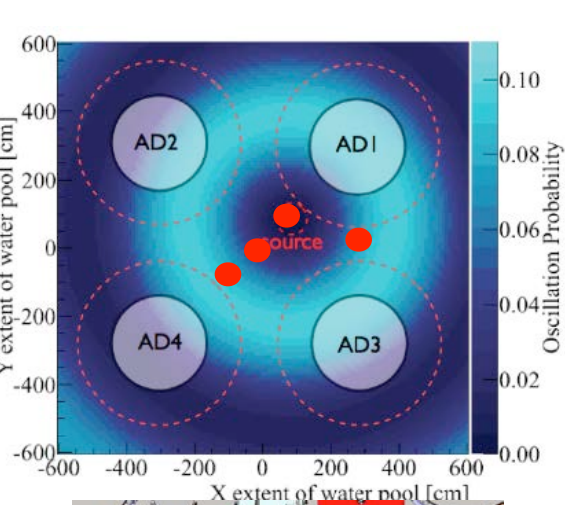
CeLAND Plans and Status

- deploy 75 kCi in KamLAND outer detector (phase 1) in 2015.
- aim for deployment in center with the same source if there is a hint (much more challenging)
- Japan working on authorizations, expect to know more in March at the collaboration meeting

Daya Bay Sterile Neutrino Search

^{144}Ce - ^{144}Pr source can Probe baselines from 1.5-8 m with an source in the water pool of the Daya Bay Far Hall

- Advantageous to place source outside detectors in water pool.
- Multiple detectors allow for control of systematics.

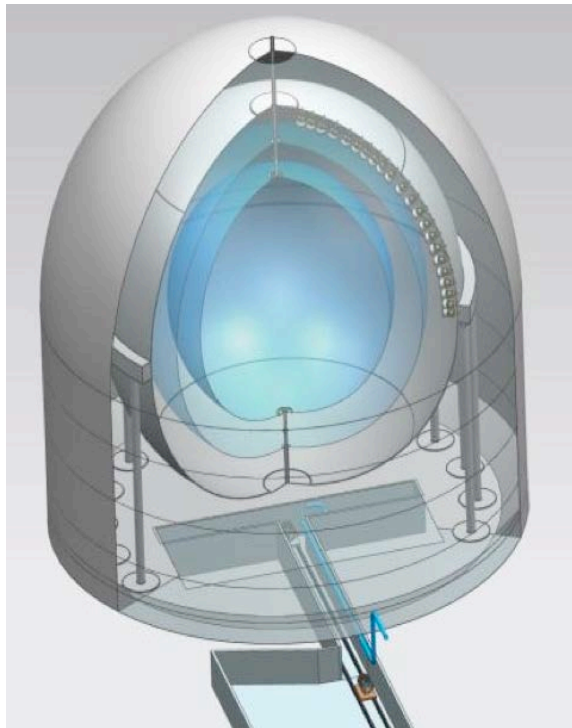


source experiment feasible in Far Hall of the Daya Bay experiment after θ_{13} measurement.

arXiv:1109.6036
Dwyer, Littlejohn, Vogel, KMH

SOX - Short Distance Oscillations with Borexino

^{51}Cr source next to Borexino

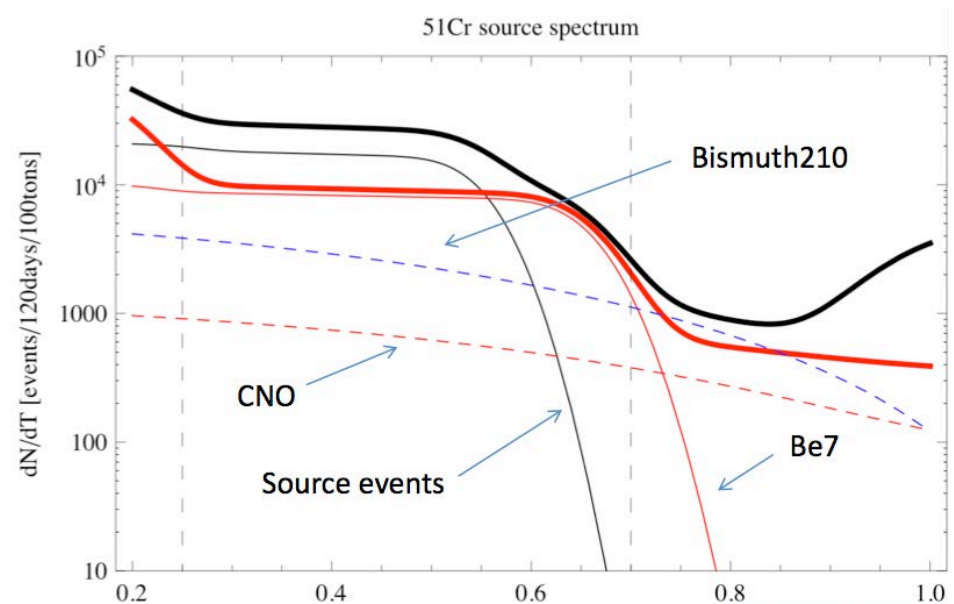


existing tunnel, source at 8.25 m from the LS target

10-15 MCi ^{51}Cr

Re-use Gallex 36 kg of enriched chromium (38%)
need add. enriched ^{50}Cr

ν_e elastic scattering signal versus distance



Detection as ^7Be solar ν

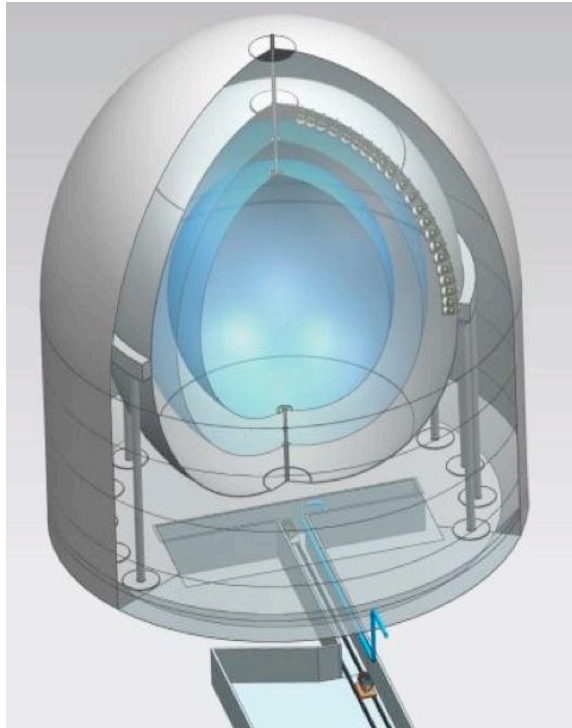
Well known background in 0.25-0.7 MeV:
solar ν 's & ^{210}Bi
1% fiducial volume error

Reactors (Petten, Ludmila, US)

$n_{th} \approx 10^{15} \text{ n/cm}^2/\text{sec}$
Space to accommodate ^{50}Cr

SOX - Short Distance Oscillations with Borexino

^{51}Cr source next to Borexino

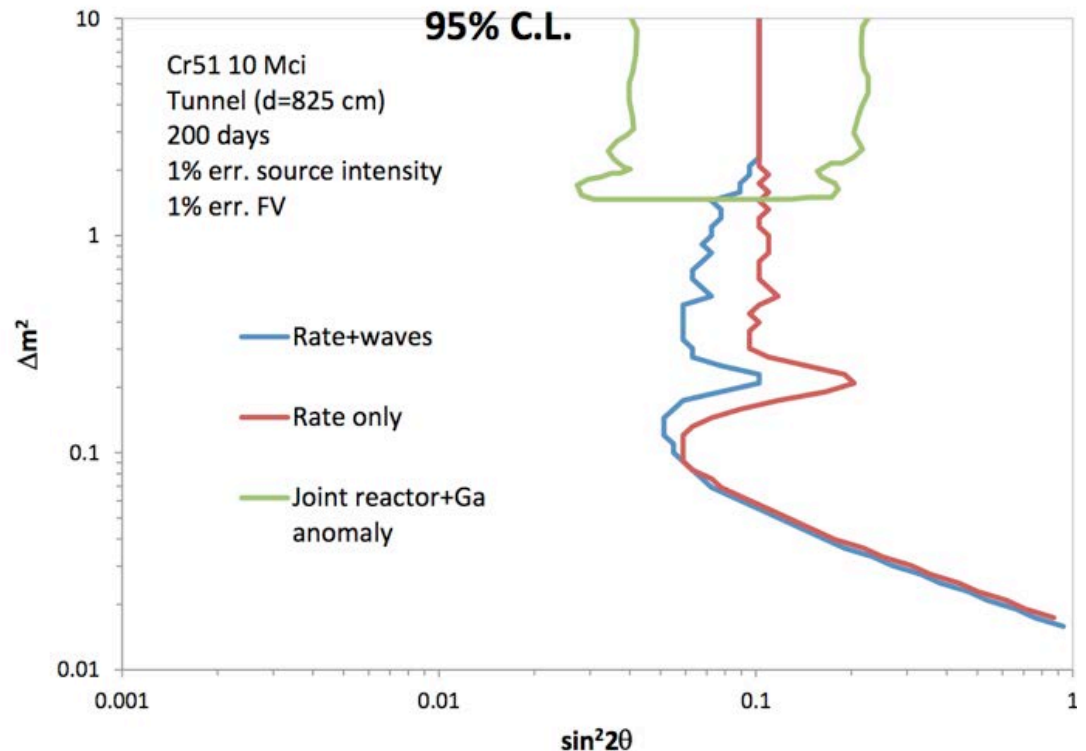


existing tunnel, source at 8.25 m from the LS target

10-15 MCi ^{51}Cr

Re-use Gallex 36 kg of enriched chromium (38%)
need add. enriched ^{50}Cr

^{51}Cr 10 MCi, 95% CL



Status

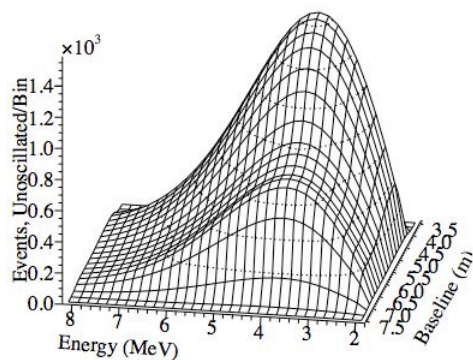
ERC funding obtained
source production by 2015

Reactor Experiments

Reactor Experiments at Very Short Baselines

Energy and Baseline Dependent Effect

Reactor Antineutrinos

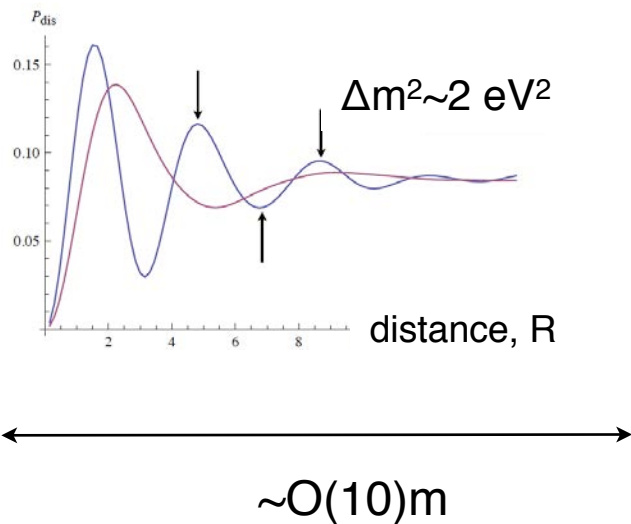
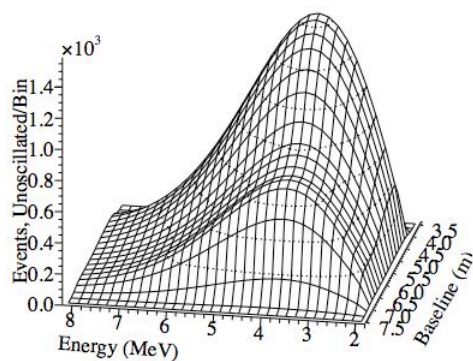


Reactor Experiments at Very Short Baselines

Energy and Baseline Dependent Effect

Reactor Antineutrinos

$\bar{\nu}_e$ Oscillation

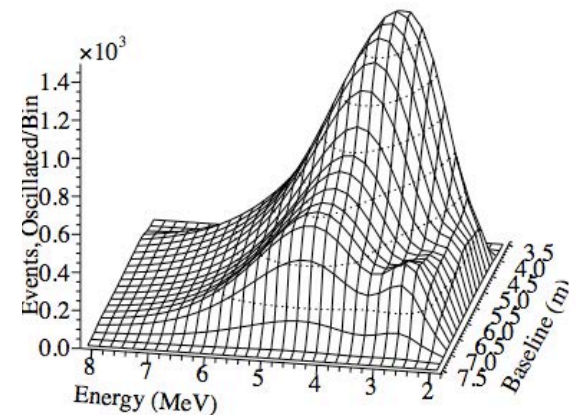
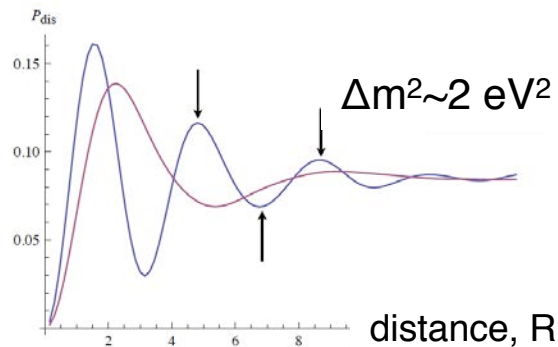
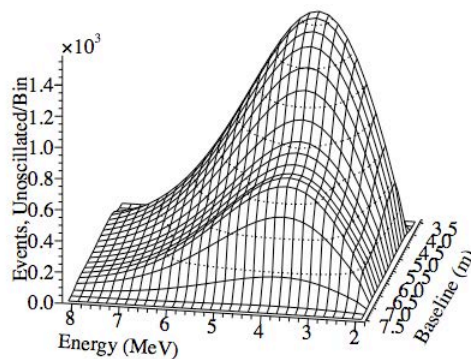


Reactor Experiments at Very Short Baselines

Energy and Baseline Dependent Effect

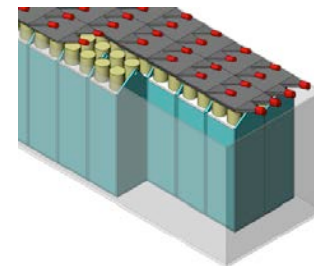
Reactor Antineutrinos

$\bar{\nu}_e$ Oscillation



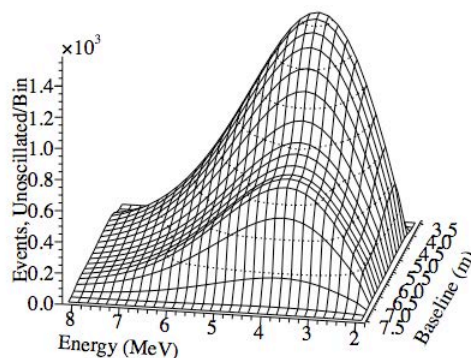
Reactor Experiments at Very Short Baselines

Energy and Baseline Dependent Effect

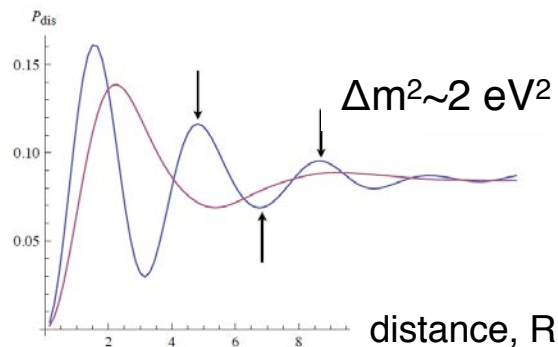


segmented,
extended
detector

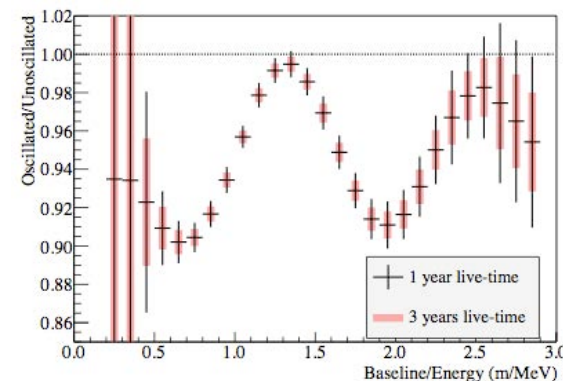
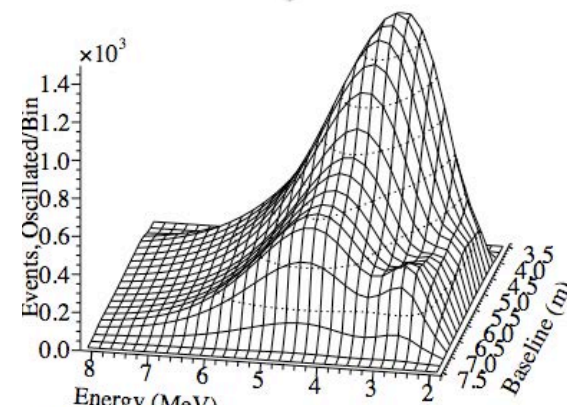
Reactor Antineutrinos



$\bar{\nu}_e$ Oscillation

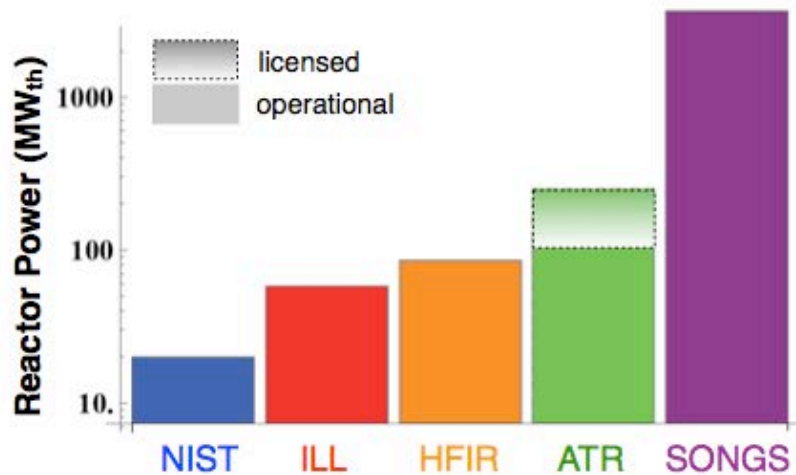


← ~O(10)m →



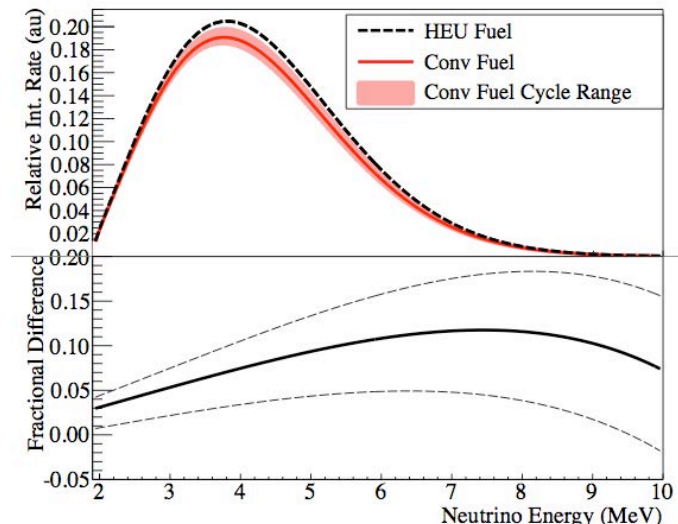
Reactor Experiments at Very Short Baselines

Reactor Power and Duty Cycle

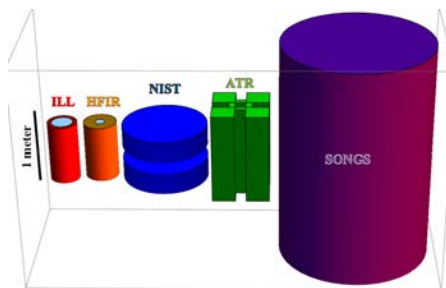


Reactor	Power (MW _{th})	Baselines (m)	Reactor On (Days)	Reactor Off (Days)	Down-Time
NIST	20	4-13	42	10	~32%
HFIR	85	6-8	24	18	~50%
ATR	250 (licensed) 110 (operational)	7-8 (restricted) 12-20 (full access)	48-56	14-21	~27%
ILL	58	7-9	50	41	~45%
SONGS	3438	24	639	60	8.6%

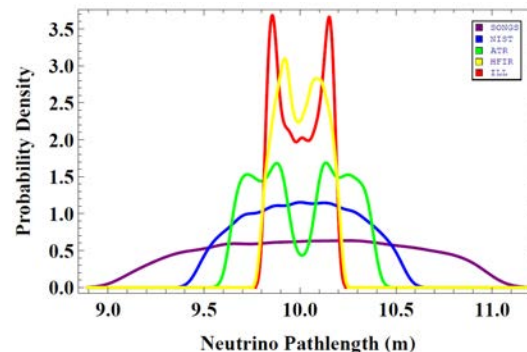
Reactor Fuel



Reactor Core Size



Pathlength spread at detector from core

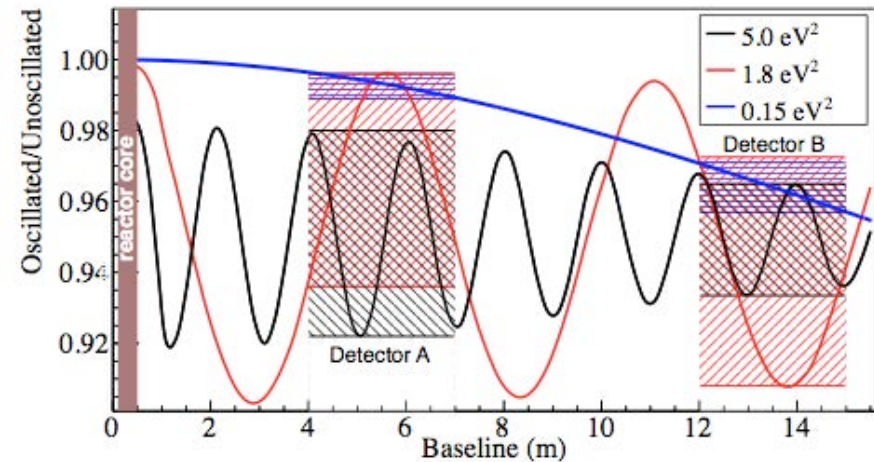
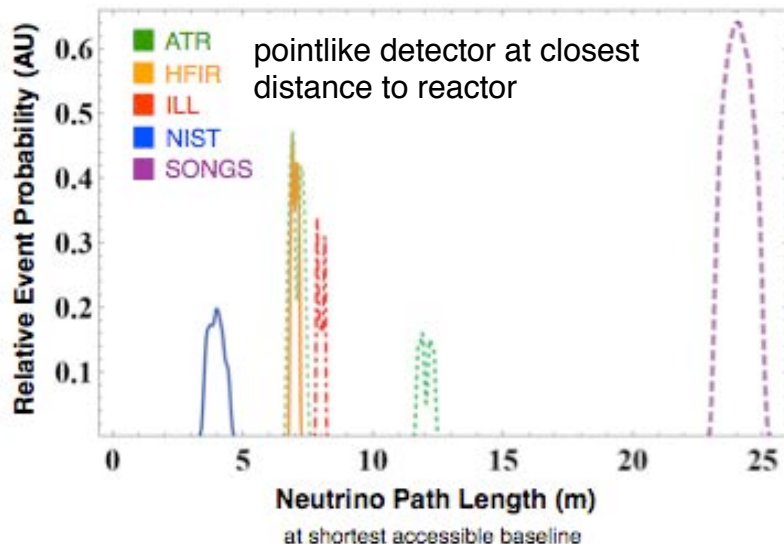


arXiv:1212.2182

Littlejohn, Mumm, Tobin, KMH

Reactor Experiments at Very Short Baselines

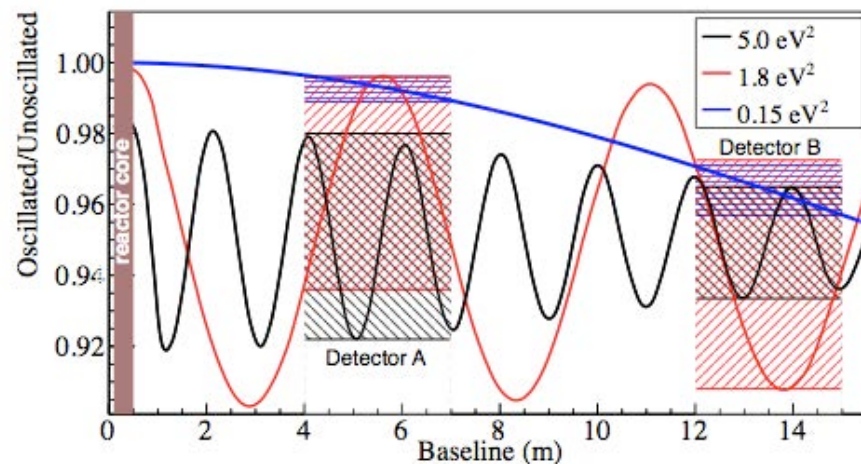
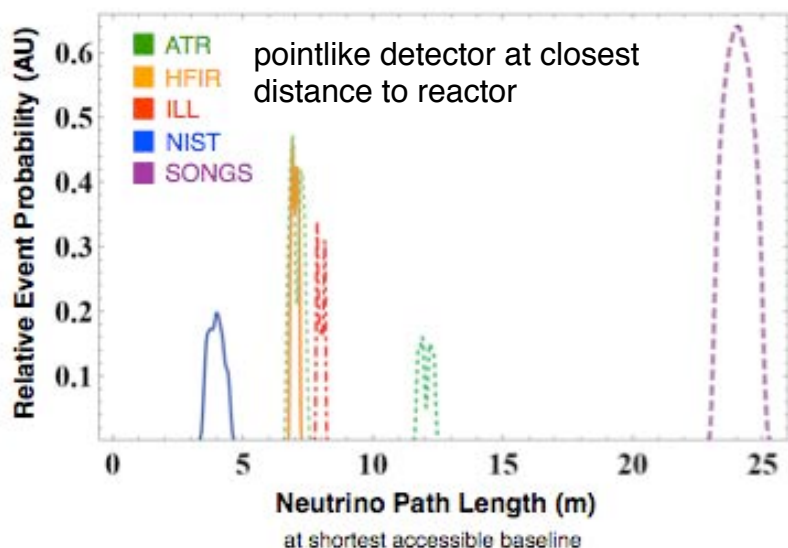
Reactor-Detector Distance: How close do we need to be?



detector samples different oscillations for different Δm^2 , \rightarrow multiple detectors useful

Reactor Experiments at Very Short Baselines

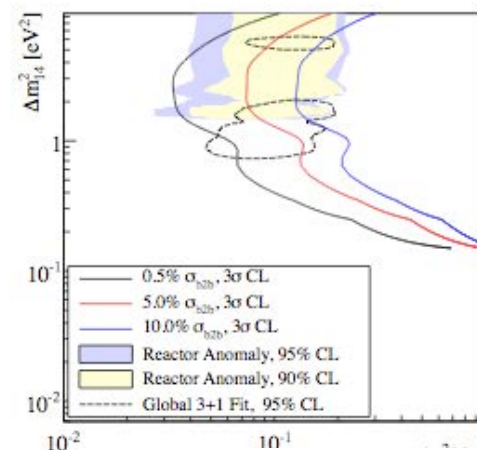
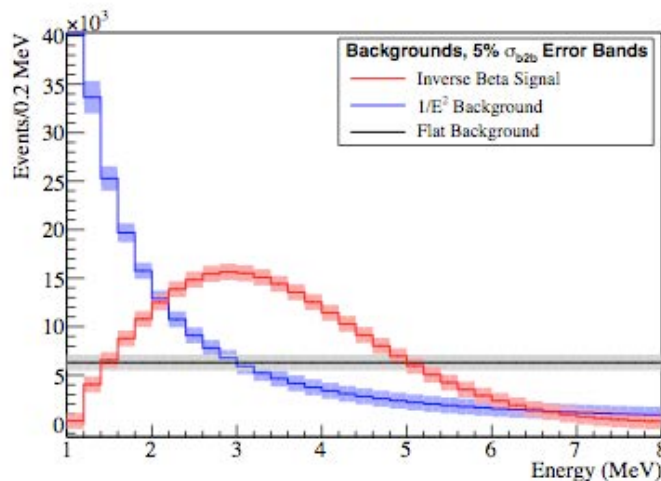
Reactor-Detector Distance: How close do we need to be?



detector samples different oscillations for different Δm^2 , \rightarrow multiple detectors useful

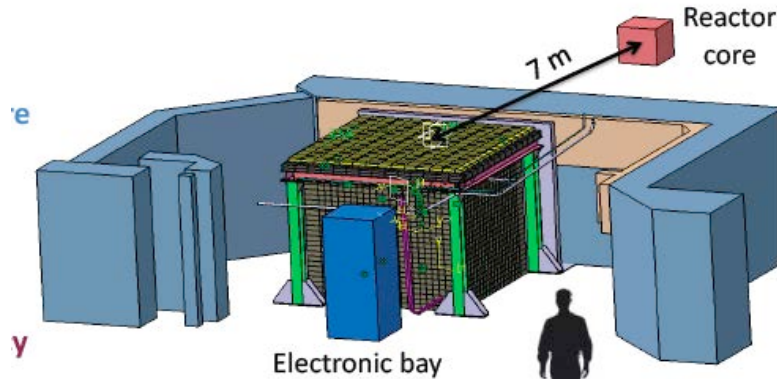
Backgrounds

significant challenge,
critical to know/measure
background distributions



Reactor Monitoring Experiments

NUCIFER at Osiris



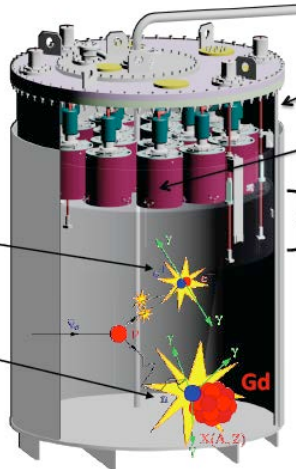
core: $\sigma \sim 0.3\text{m}$
baseline: 7m

"inverse β -decay" process

$$\bar{\nu}_e + p \rightarrow e^+ + n$$

Prompt e^+ signal

+
Delayed neutron signal ($\Delta t \sim 30 \mu\text{s}$)

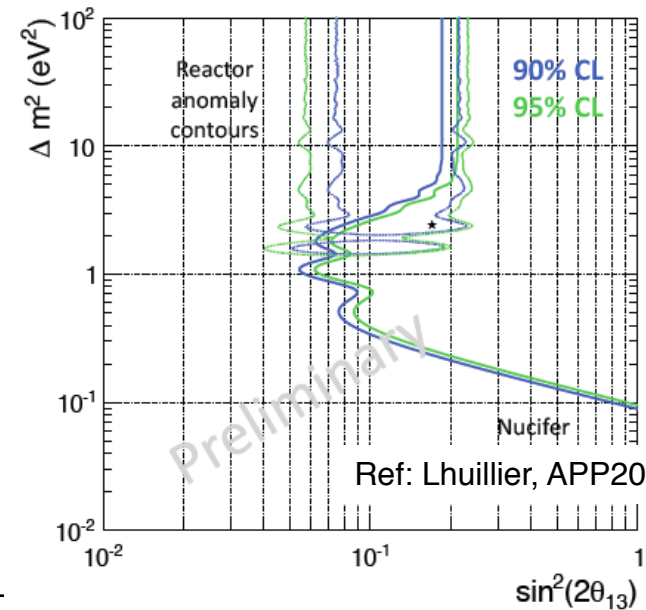
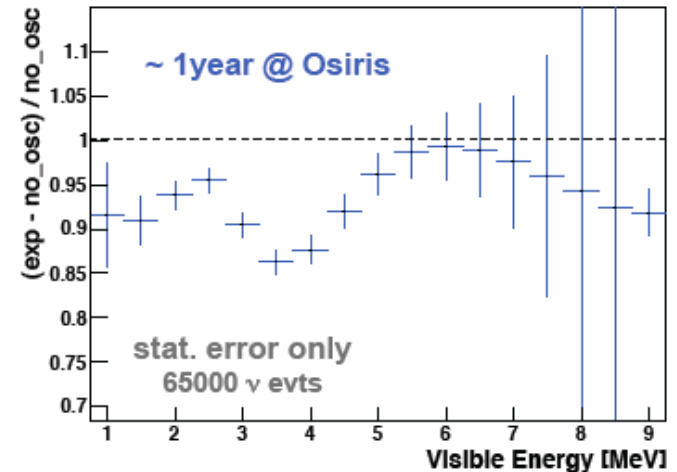


- Norm error = 4%
- 100 days full power @ Osiris
- S/B = 1 (?), assuming same shapes (worst case).
- E resol = $0.15 \cdot E$

Pre-industrial, unattended reactor neutrino monitor

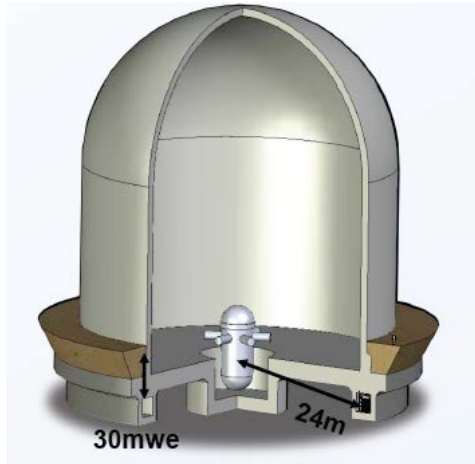
May be used to test reactor anomaly with compact core.
PSD R&D for background rejection.

Expected E spectrum deformation
with anomaly best fit: $\Delta m^2 = 2.4 \text{ eV}^2$ & $\sin^2(2\theta) = 0.15$

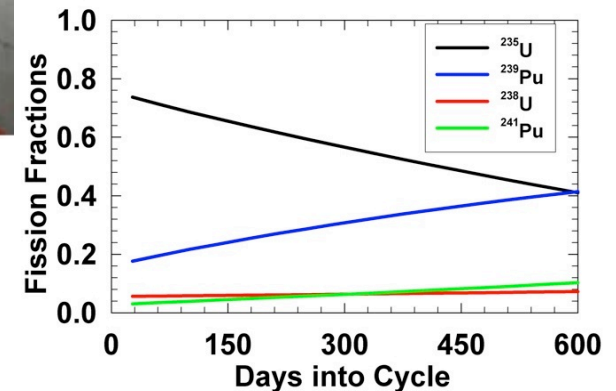
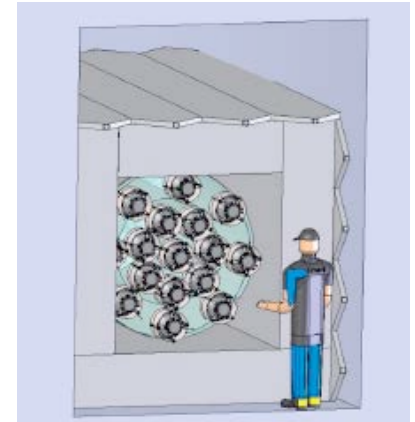


Reactor Monitoring Experiments

SONGS - San Onofre Nuclear Power Generating Station



core \varnothing : ~3m,
fixed baseline: 24m



For many years, applied antineutrino physics studies have used SONGS tendon galleries

Ideal for detector R&D and characterization, coherent scattering studies

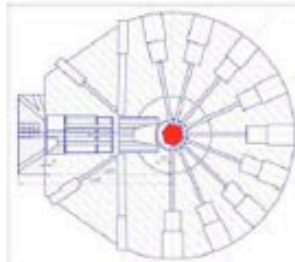
SONGS currently offline since early 2012. Unit 2 restart anticipated this summer. More frequent outages (background measurement opportunities) likely for several years

Ref: Bowden

New Short Baseline Reactor Experiments

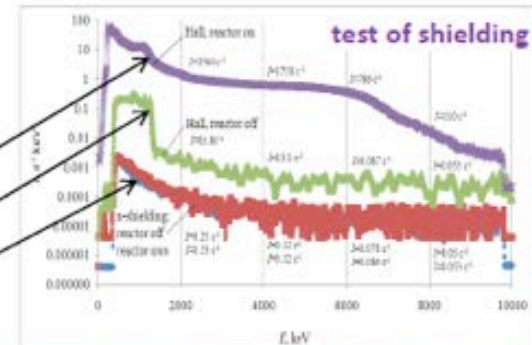
NEUTRINO-4 experiment

Preparation at WWR-M reactor (18 MW) in PNPI (Gatchina)



Reactor power - 18 MW
Size of active core – 0.6 m

reactor on without shielding
reactor off without shielding
reactor on/off with shielding



Installation of 2 sections test antineutrino detector with liquid scintillator (total volume 0.4 m³)

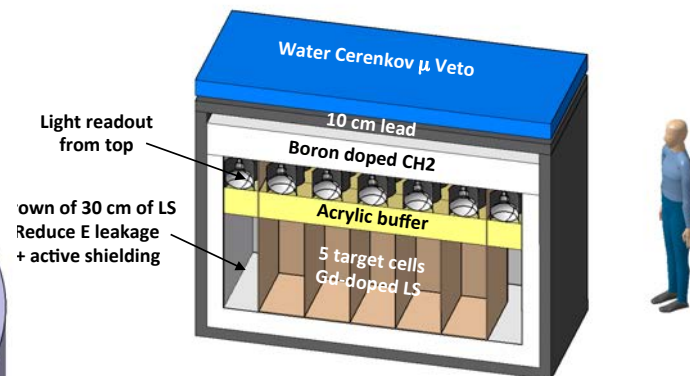
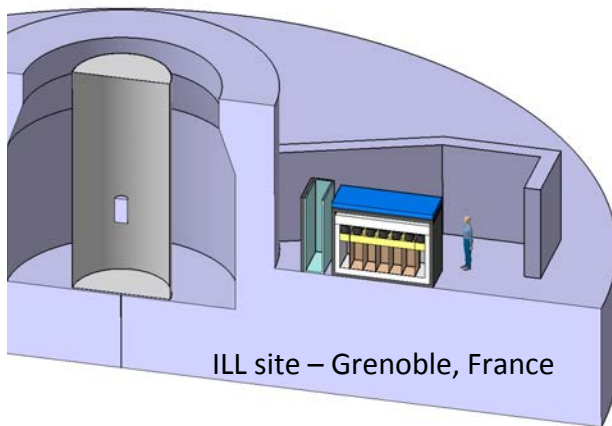
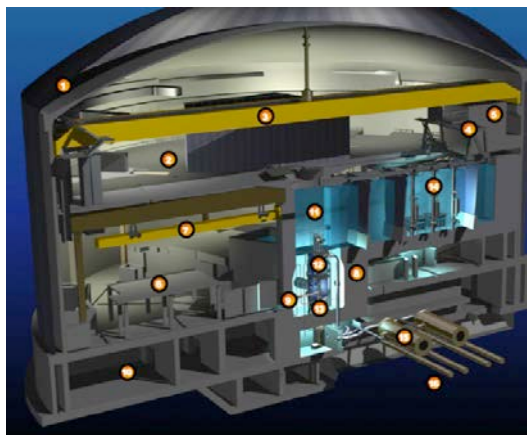


Installation of anticoincidence shielding from plastic scintillator 0.5x0.5x0.125 m³ with PMT (32 pieces)

A.Serebrov, PNPI

New Short Baseline Reactor Experiments

STEREO at ILL



Shape analysis +
3.5 % uncertainty on normalization

Reactor Site

50 MW compact core
($\phi=40\text{cm}$, $h=80\text{ cm}$)

Short baseline
[7-9] m

Pure ^{235}U spectrum

Background Rejection

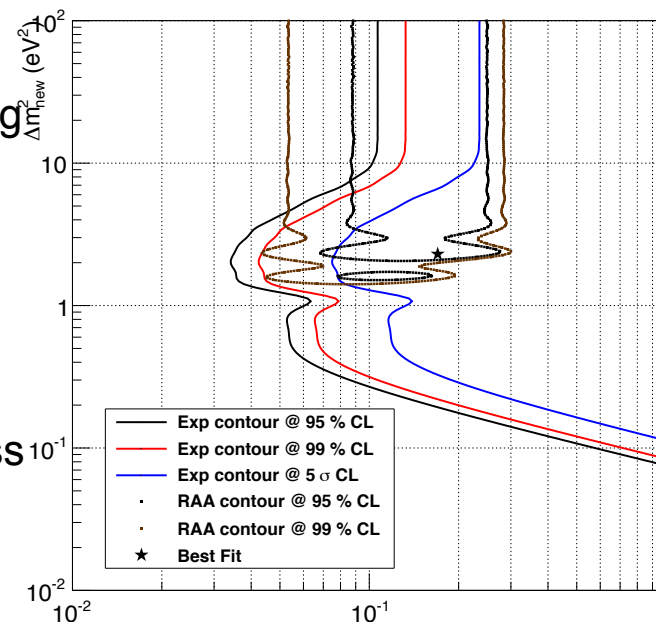
Large passive and active shielding
15 m.w.e. overburden

Pulse Shape Discrimination

Segmented detector

On-site measurements in progress

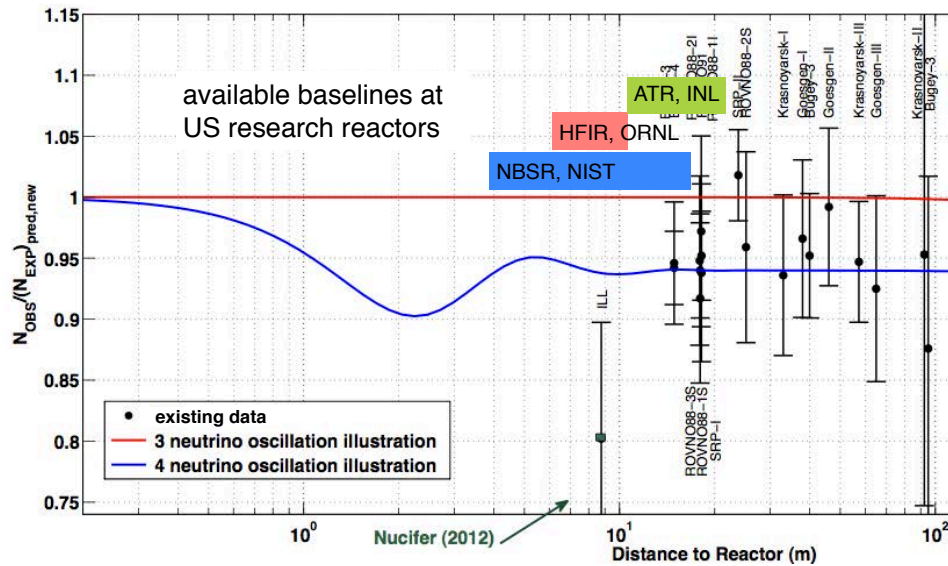
Aim for first data in 2015
Funding decision in 2013



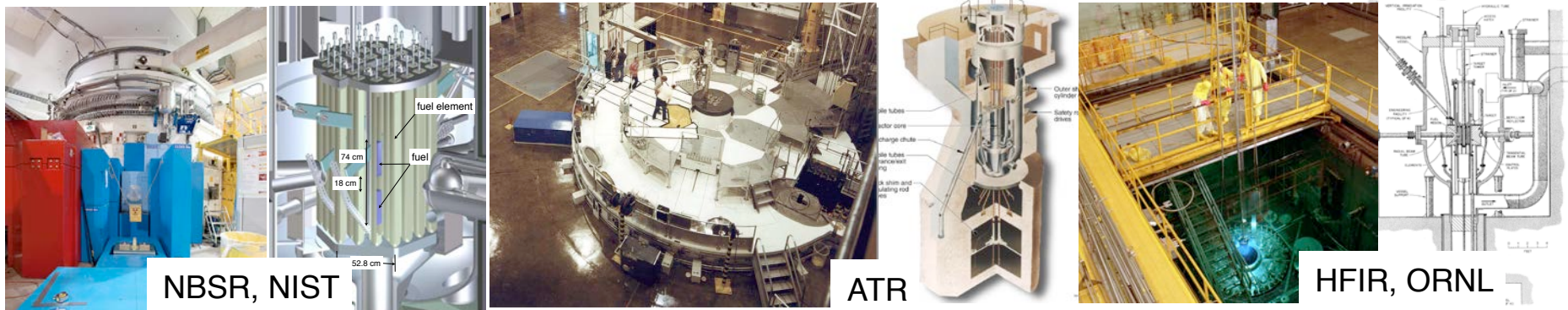
Ref: Lhuillier

Opportunities for Experiment at US Reactor?

Shortest Accessible Baselines

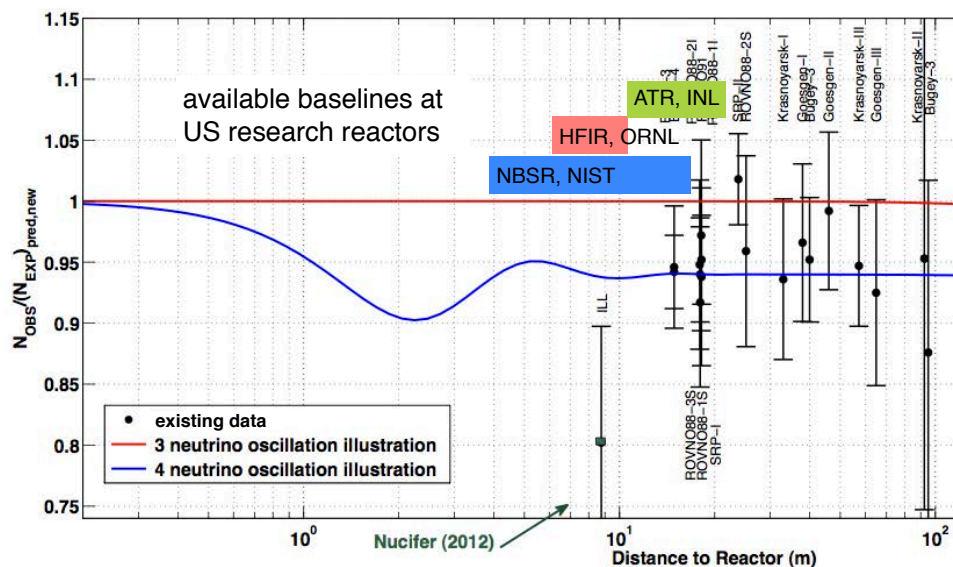


High-Power US Research Reactors

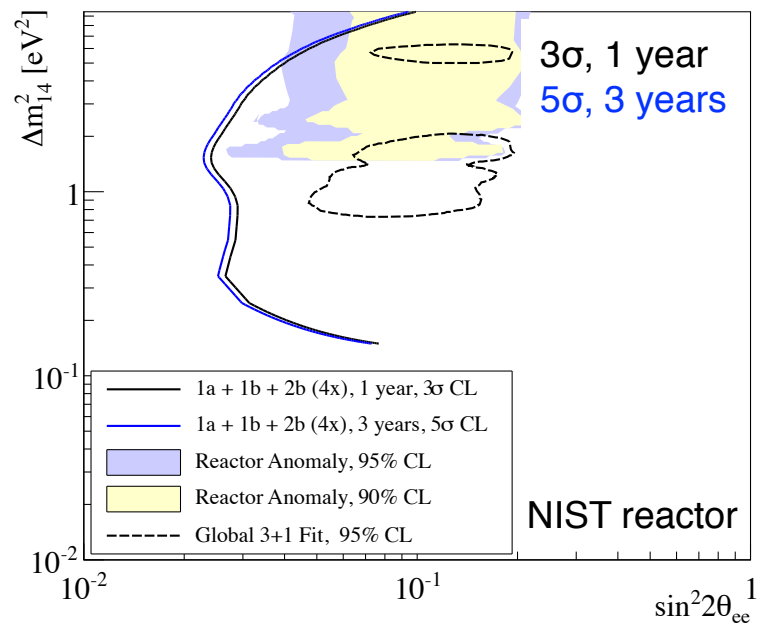


Opportunities for Experiment at US Reactor?

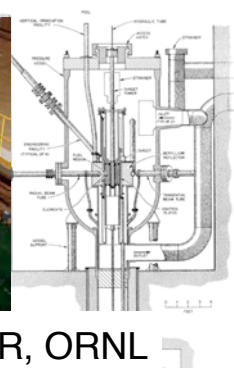
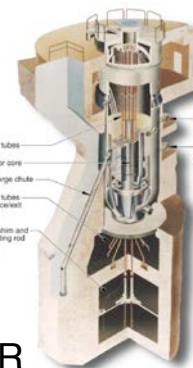
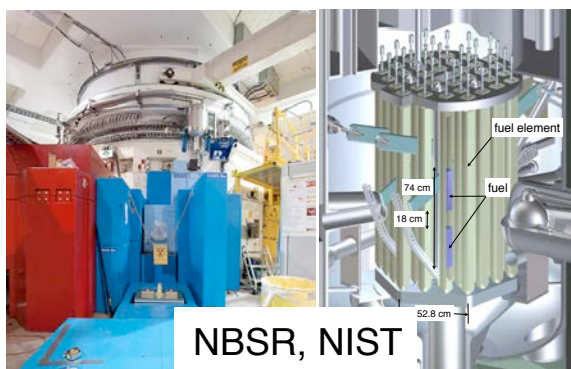
Shortest Accessible Baselines



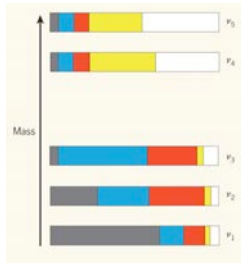
Discovery Potential



High-Power US Research Reactors

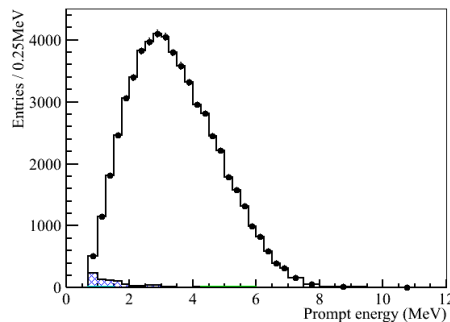


Scientific Opportunities for a US Reactor Experiment



Searches for new physics

- test **sterile ν hypothesis** and **short-baseline oscillations**
- probe and understand **reactor anomaly**
- neutrino **coherent scattering** and **magnetic moment** searches



Reactor cores, fuel, and antineutrino spectra

- precision studies of **reactor antineutrino spectra with HEU core**
- studying **HEU to LEU core conversion** at US research reactors



Background studies and detector development

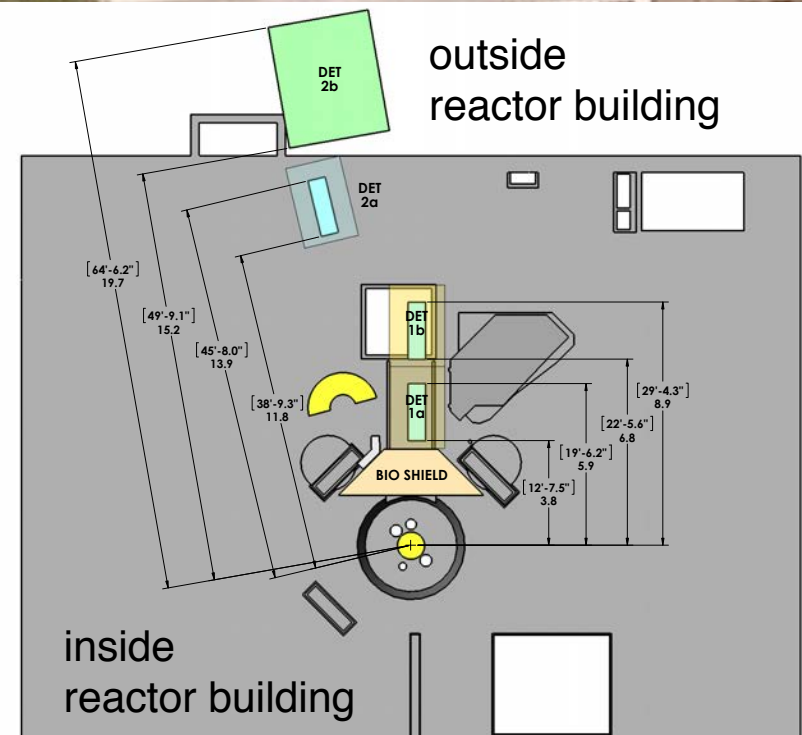
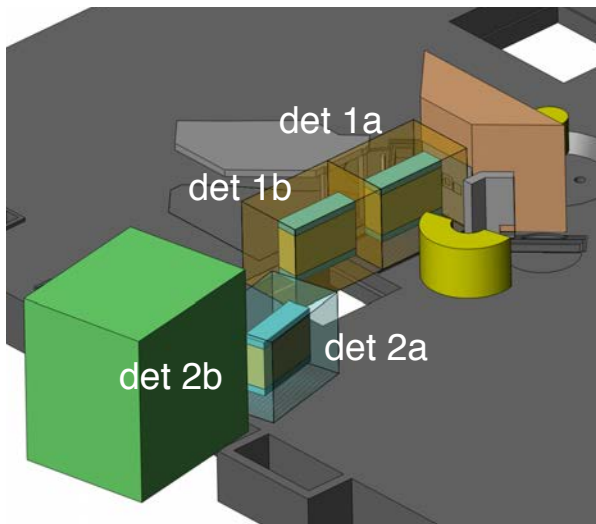
- demonstrate operation of **on-surface antineutrino detectors**
- synergies with **safeguard** and reactor monitoring
- develop **scintillators for neutron detection** with PSD (Gd and Li-doped, LAB vs water)

US community interested in pure and applied reactor antineutrino studies

NIST: An Example in the US

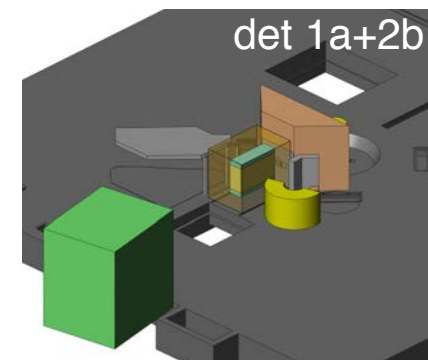
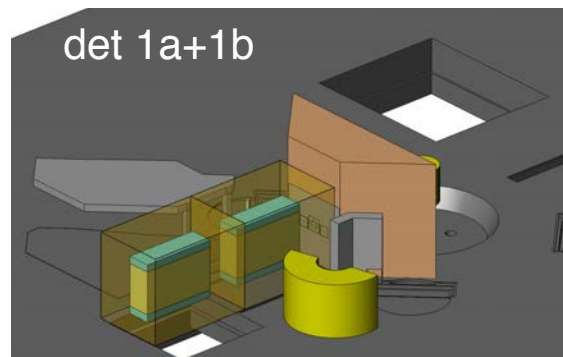
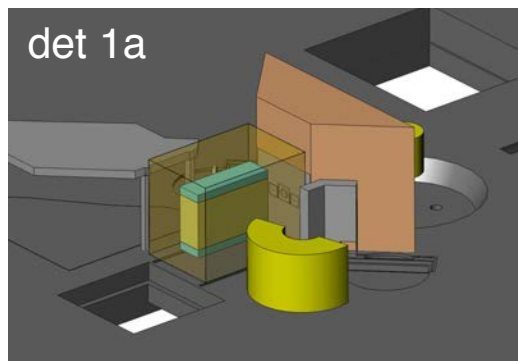
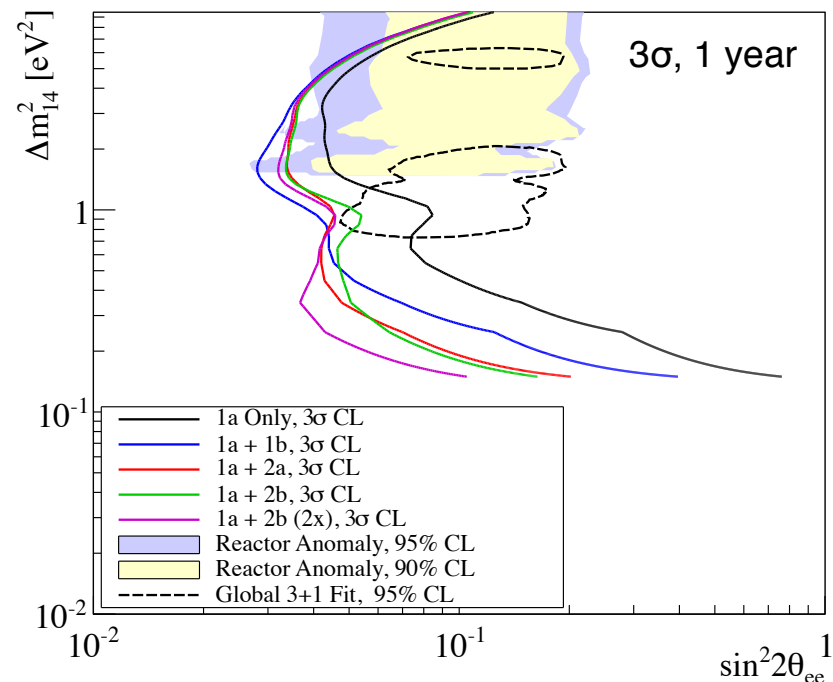
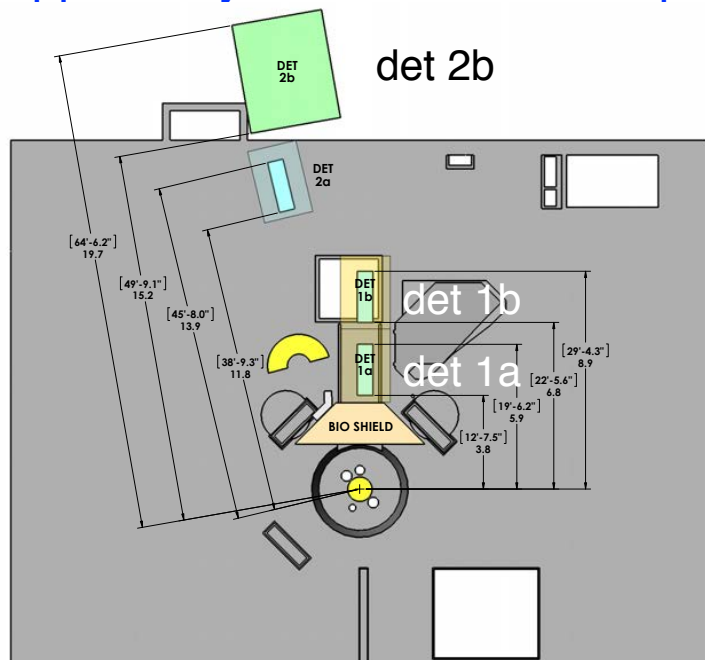


Possible detector baselines at NIST
3.8m, 6.8m, 11.8, 15.2m (up to 20m)



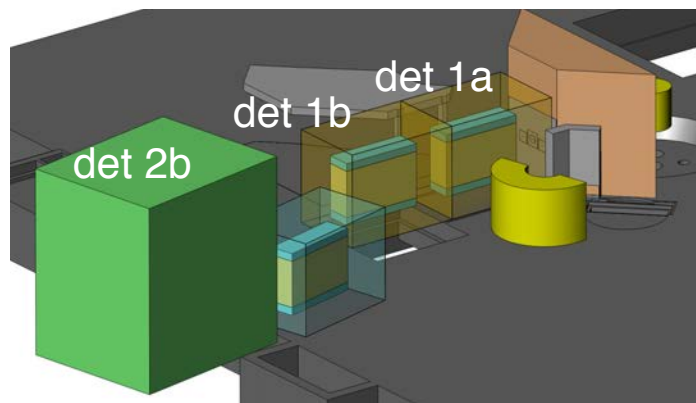
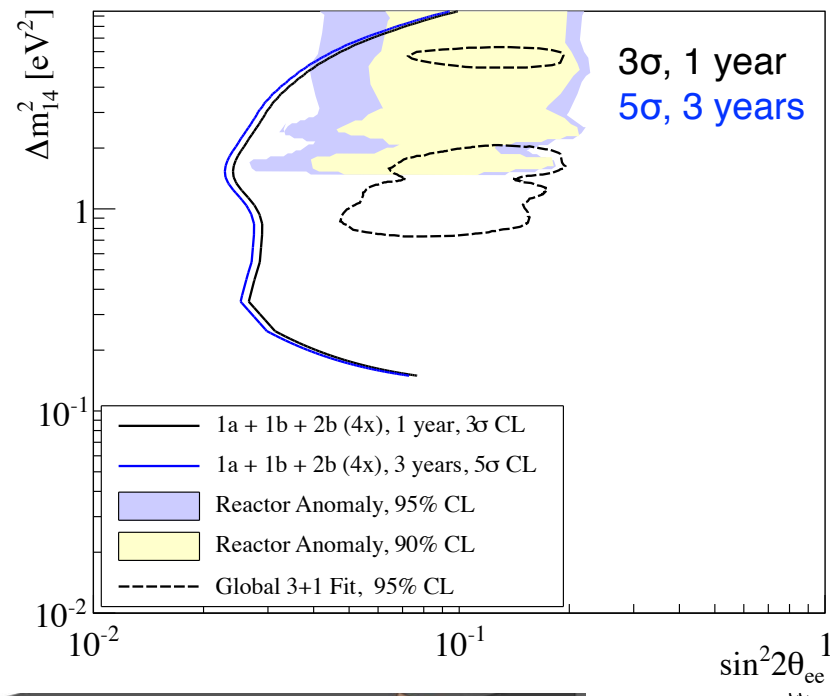
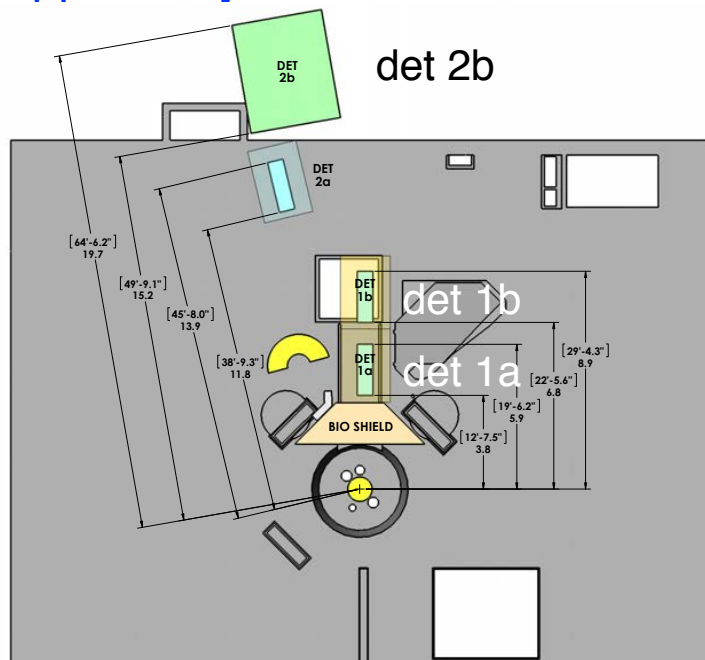
NIST: An Example in the US

Opportunity for a 2-Detector Experiment



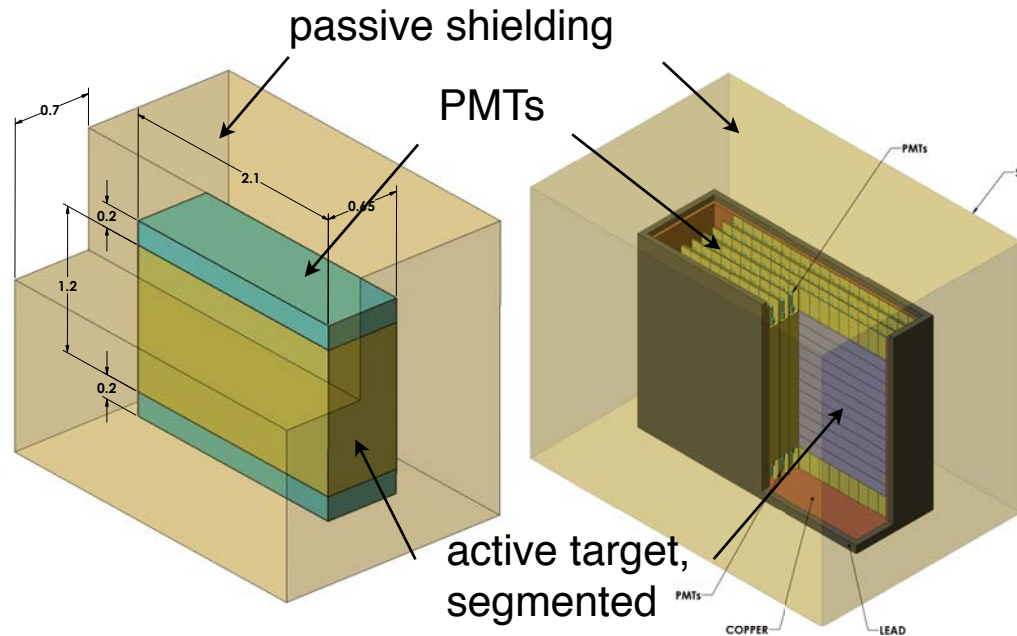
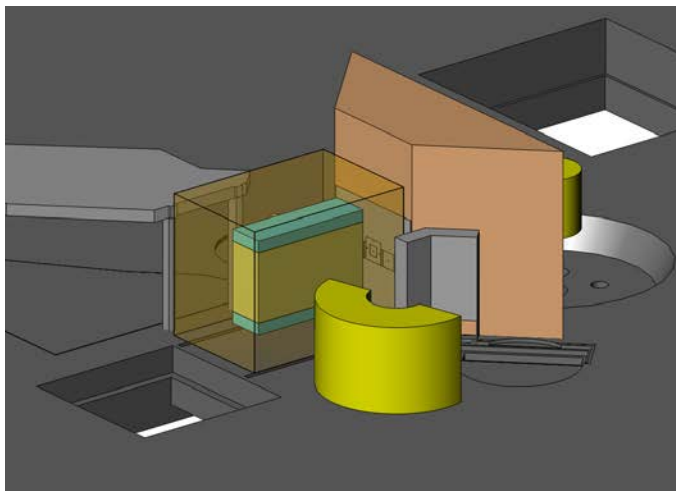
NIST: An Example in the US

Opportunity for a Multi-Detector Experiment



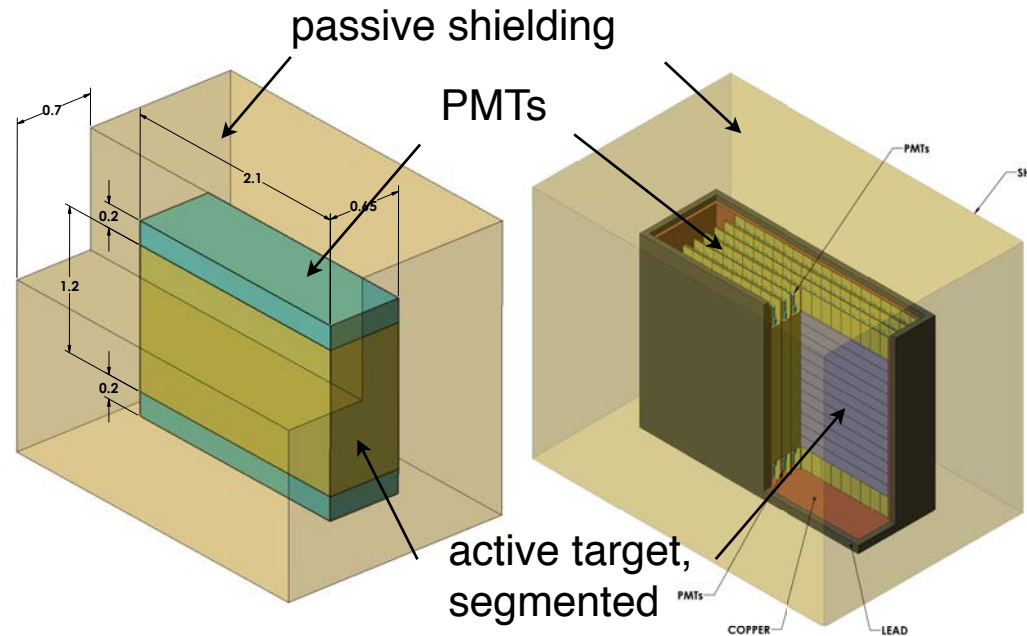
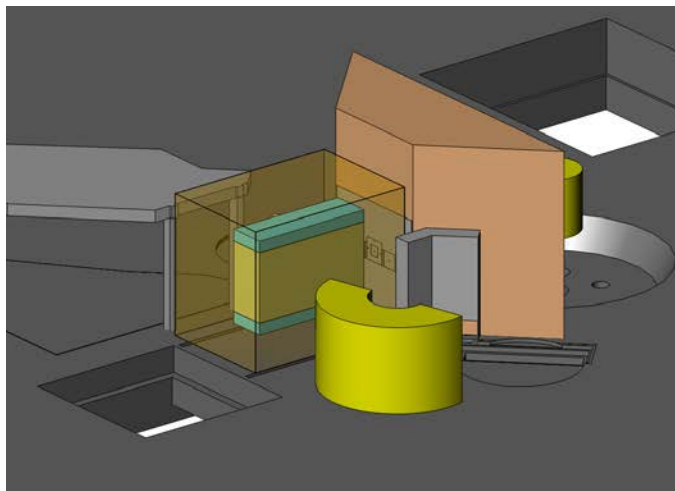
NIST: An Example in the US

Detector Module

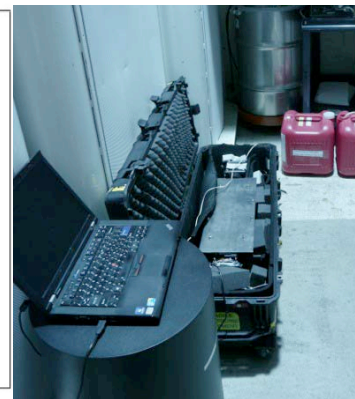
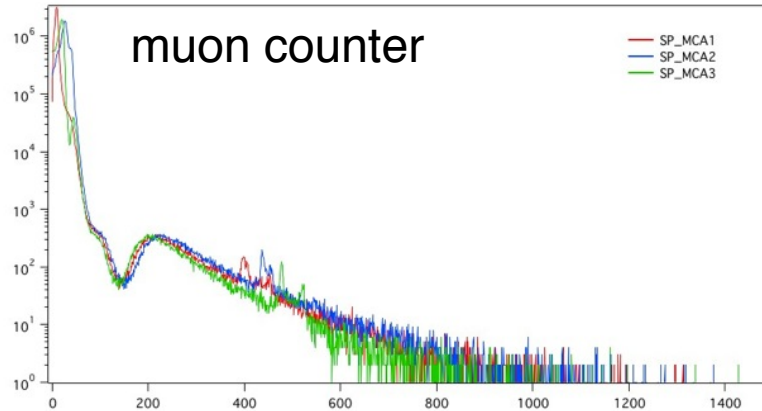
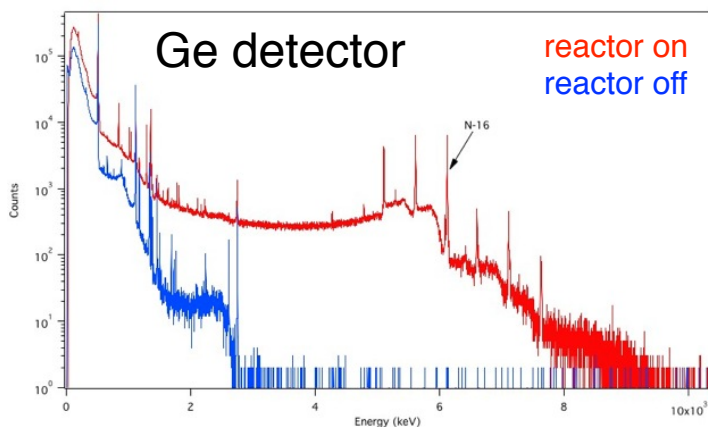


NIST: An Example in the US

Detector Module



Background studies near reactor in progress



Bowden, Cherwinka Mumm, Littlejohn, Pettus, KMH

Summary

Science Objectives with Discovery Potential

- For > 50 years reactor experiments have played an important role in neutrino physics, in both discoveries and precision measurements.
- Very short baseline (**L~10m**) measurements offer opportunities for precision studies of the **reactor spectra** and searches for **new physics**.

Reactors as a Unique Experimental Tool

- Reactor and source experiments provide a complimentary way to probe short baseline oscillations.
- **Reactors are only way to definitely understand reactor anomaly**
- Reactors allow multiple cross-checks (no special source production needed)
- **US has some of the most powerful research reactors.**

Challenges and Synergies

- **Background mitigation** and **systematic controls** will be key.
- Several R&D efforts. **US site can accommodate multiple detectors.**
- On-surface **reactor neutrino monitors** useful for **safeguards** applications.

Thanks to many colleagues for slides and material: N. Bowden, A. Derbin, Y. Kim, T. Lasserre, D. Lhuillier, M. Pallavicini, A. Serebrov, A. Starostin, M. Yeh, Y. Wang, H. Wong, et al..

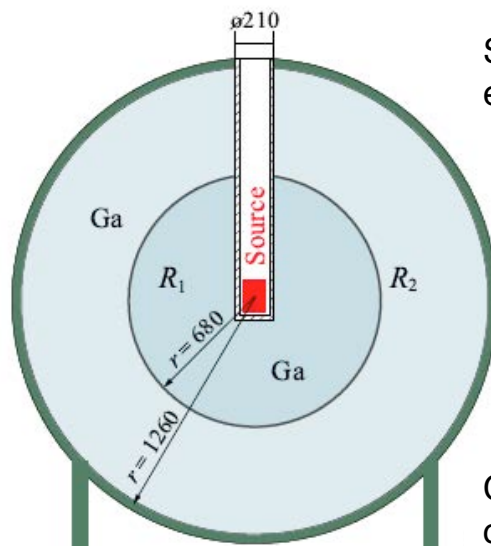
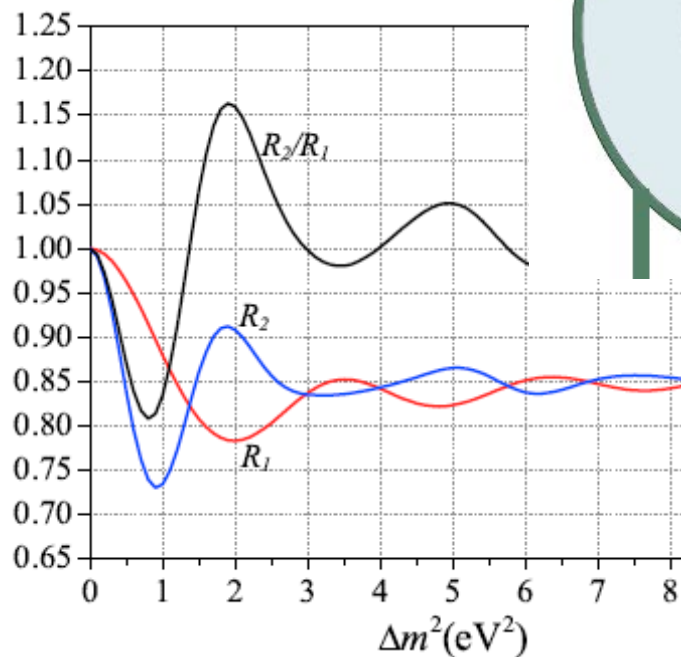
Disclaimer: Several R&D efforts worldwide. Could not present all in this talk. My apologies to those I missed.

End & Backup

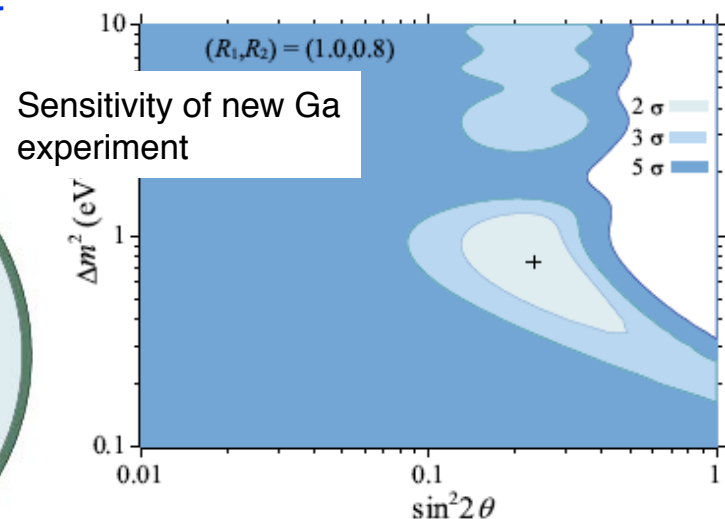
Short Baseline Oscillation with Ga Target

^{51}Cr Source inside Dual Metallic Ga Target

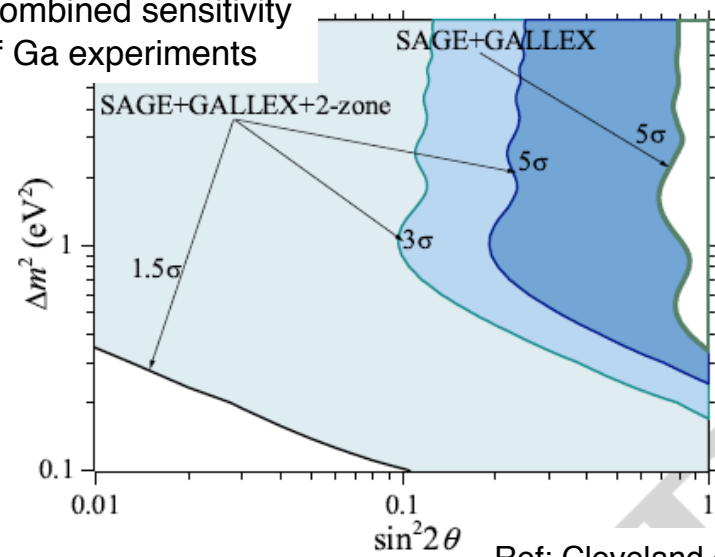
measure ratio of capture rates in R_1 and R_2



ratio of measured capture rates to predicted rate in inner and outer zones and their ratio R_2/R_1



Combined sensitivity of Ga experiments

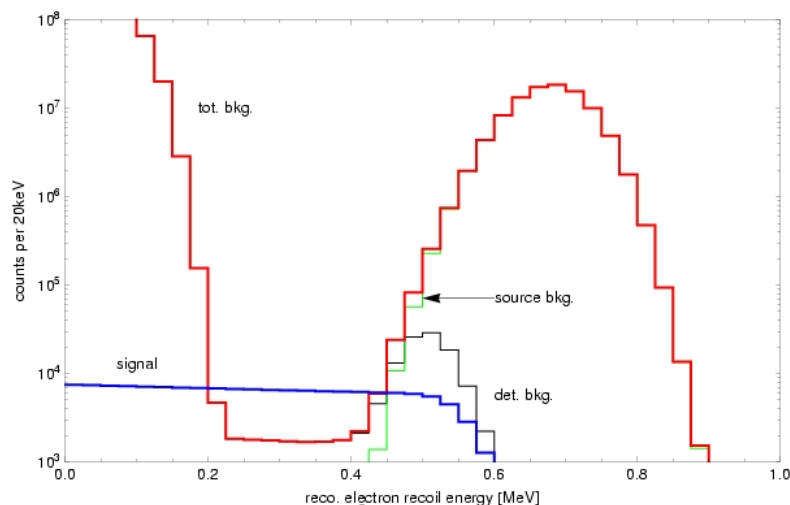


Ref: Cleveland et al.

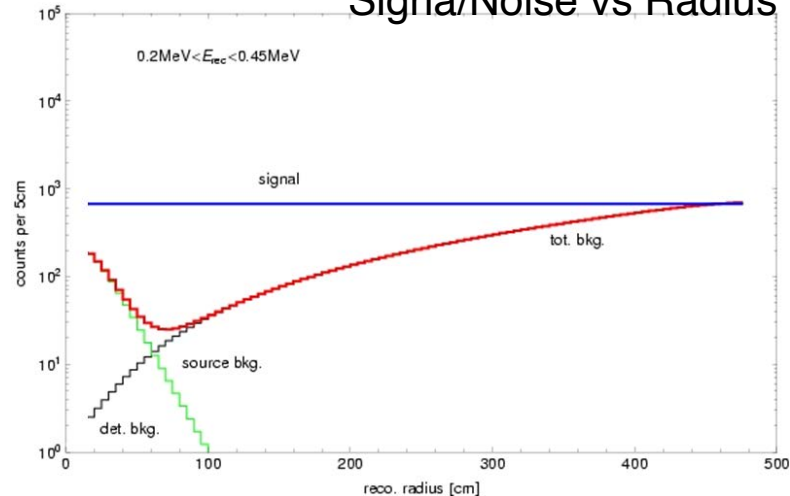
SNO-Cr Source Experiment

^{51}Cr source inside SNO+

Central Source Location in SNO+

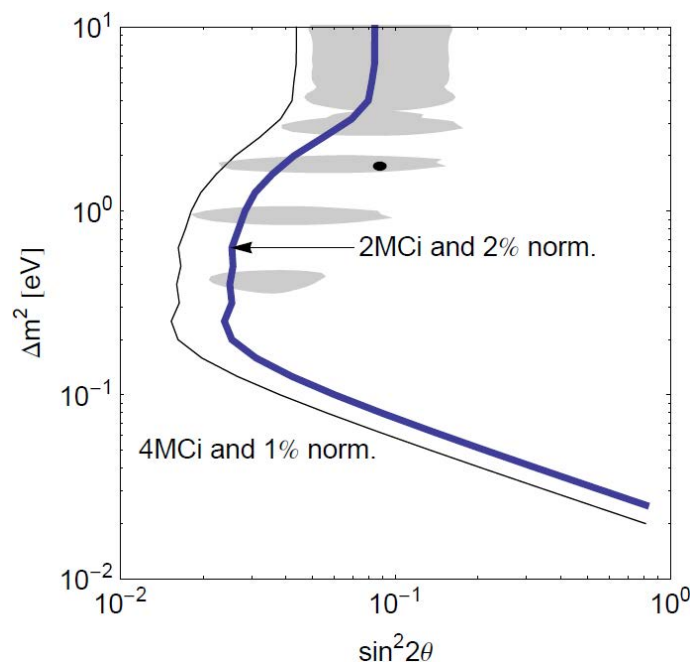
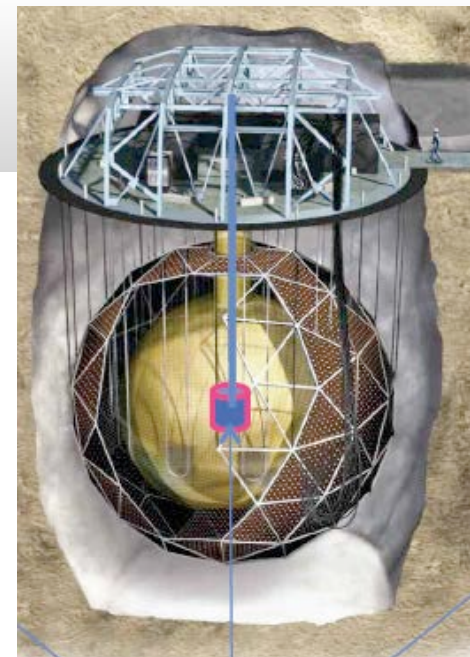


Signal/Noise vs Radius



SNO has widest neck/ chimney of all liquid scintillator detectors

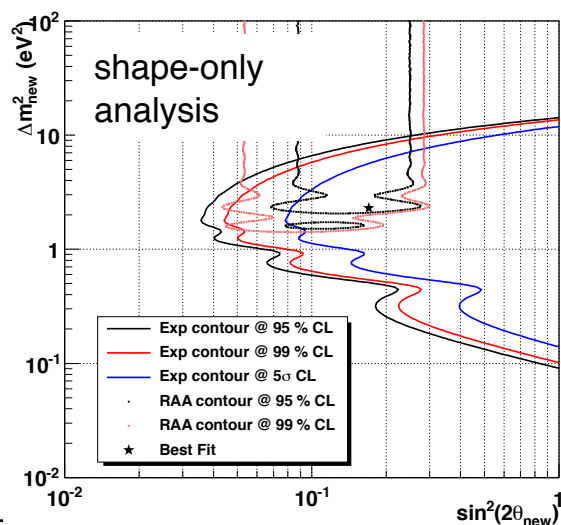
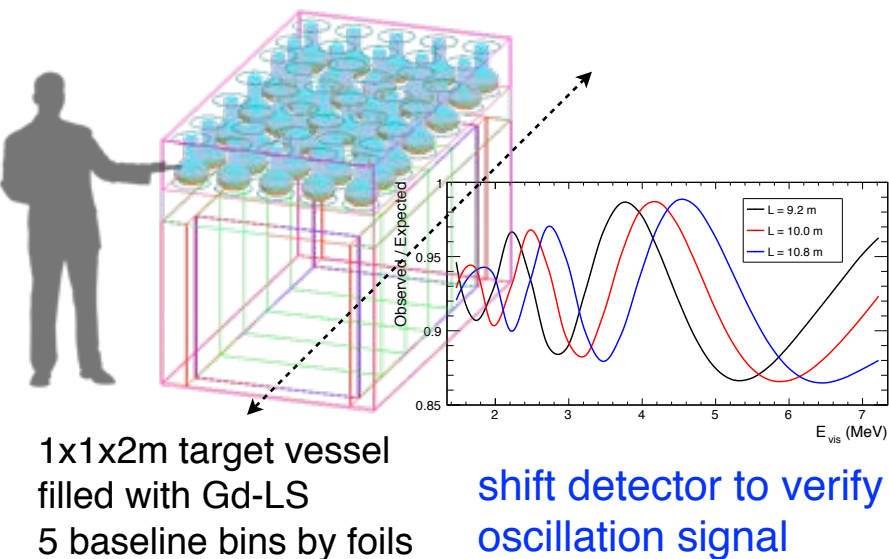
may be able to produce ^{51}Cr source in US at High Flux isotope reactor at ORNL



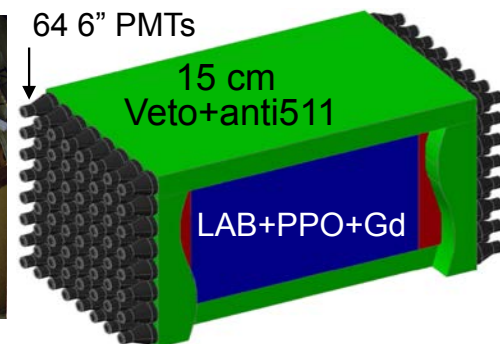
Ref: Link, Huber

Worldwide Effort Towards Sterile ν Search

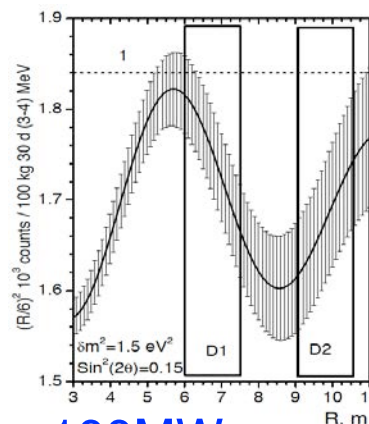
Stereo at ILL, France



POSEIDON at Reactor PIK, Russia



Gd-LS Detector: $2.1 \times 1.3 \times 1.3 \text{ m}^3$
 Energy resolution: $\sigma = 7\%$ at 1 MeV
 Spatial resolution: $\sigma_x = 15 \text{ cm}$ at 1 MeV

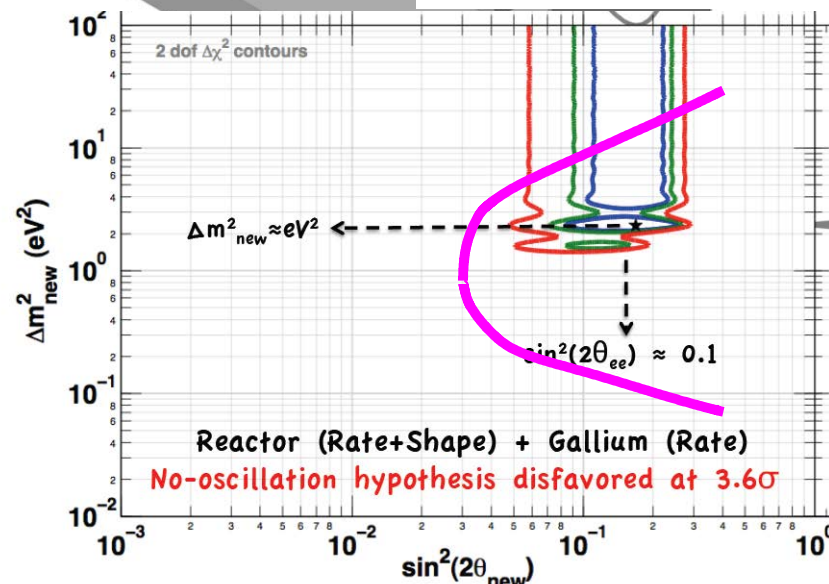
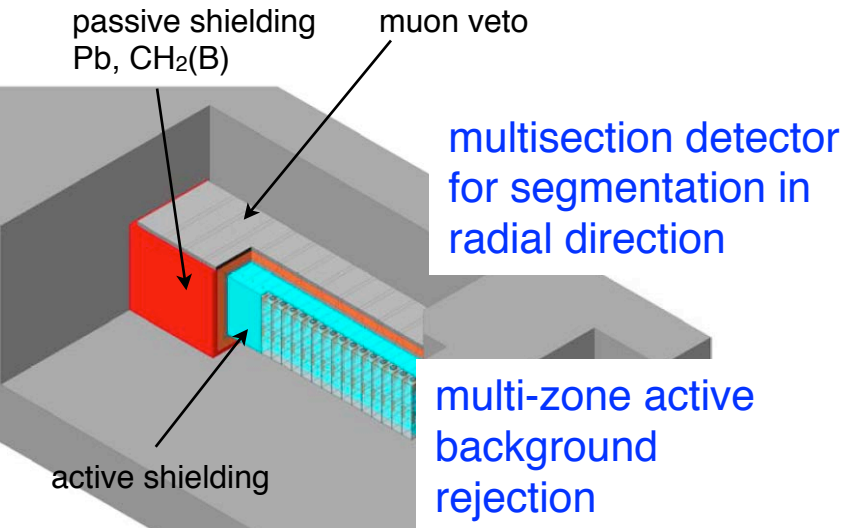


detector can be placed
at 5-8m

100MW research reactor being built in
Gatchina (~2014yr)

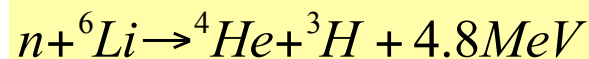
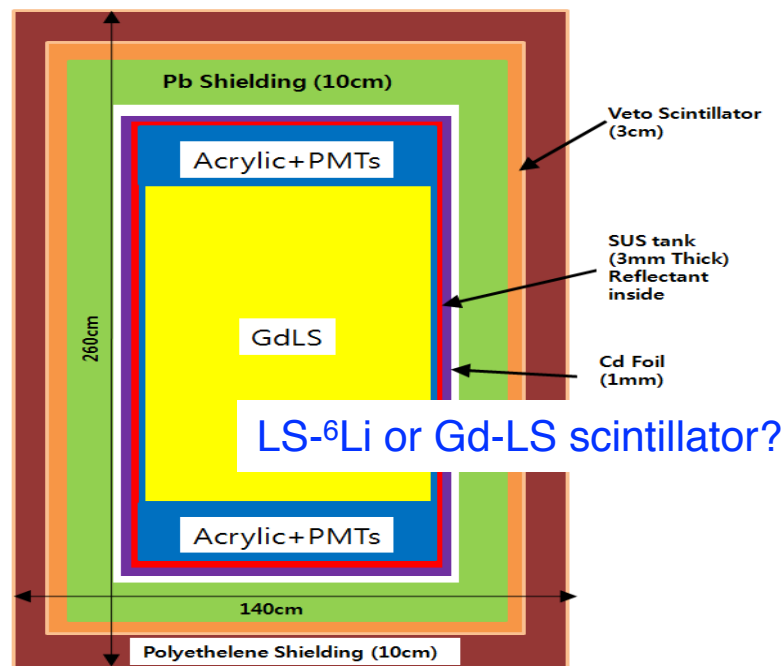
Worldwide Effort Towards Sterile ν Search

Neutrino4, Russia



Karsten Heeger, Univ. of Wisconsin

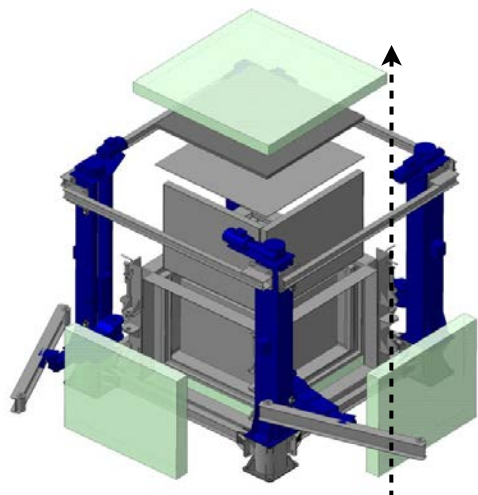
Hanaro-SBL, Korea



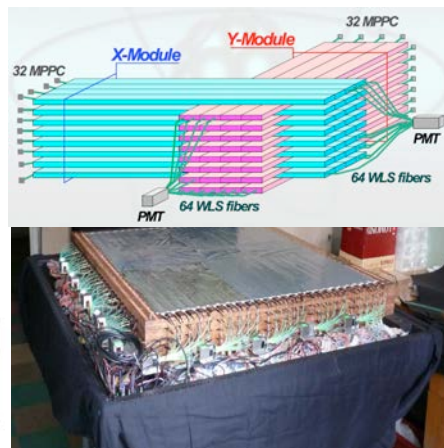
- γ - α coincidence can effectively reject backgrounds
- PSF with ⁶Li-loaded scintillator may enable on-surface detector with minimal overburden

Worldwide Effort Towards Sterile ν Search

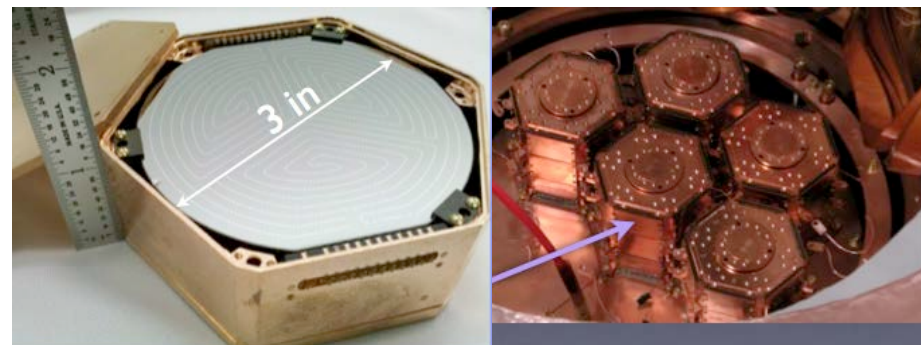
DANSS, Russia



movable distance

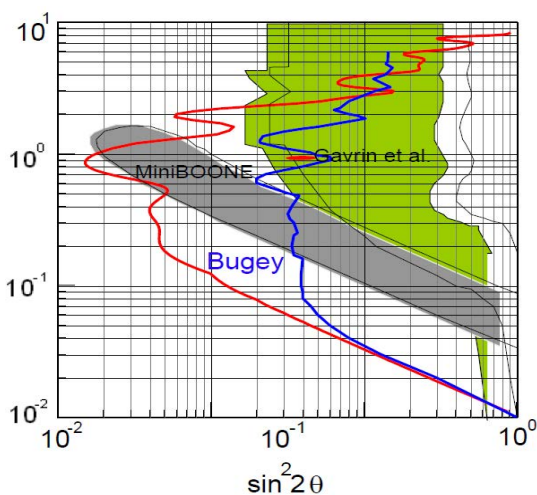
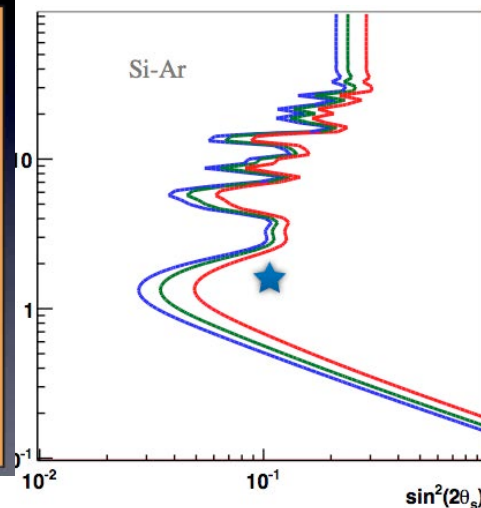
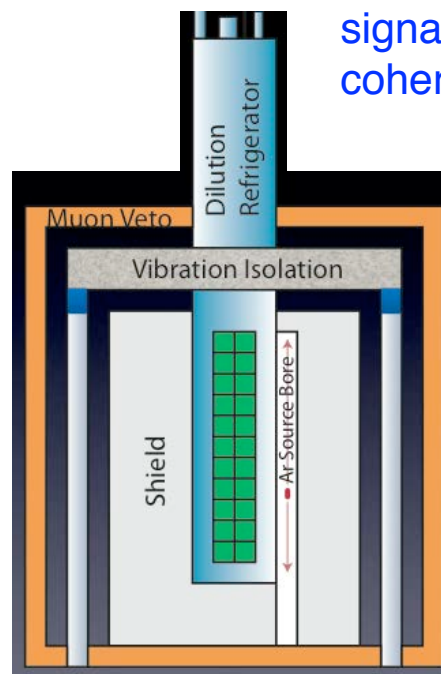


Ricochet, USA



signal detection through
coherent scattering

also used for
neutrino magnetic
moment searches
with Ge detectors



Worldwide Effort Towards Sterile v Search

A Submarine Base in Russia or Ukraine?

