

PINGU - Resolving the Neutrino Mass Hierarchy at the South Pole

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New Directions in Neutrino Physics
Aspen Center for Physics
February 2013



- Precision IceCube Next Generation Upgrade (PINGU)
- Physics goals
- Drive neutrino energy reach down to few GeV at multi-megaton scale size by infilling IceCube/DeepCore
- South Pole is an attractive option for a GeV-scale energy neutrino detector

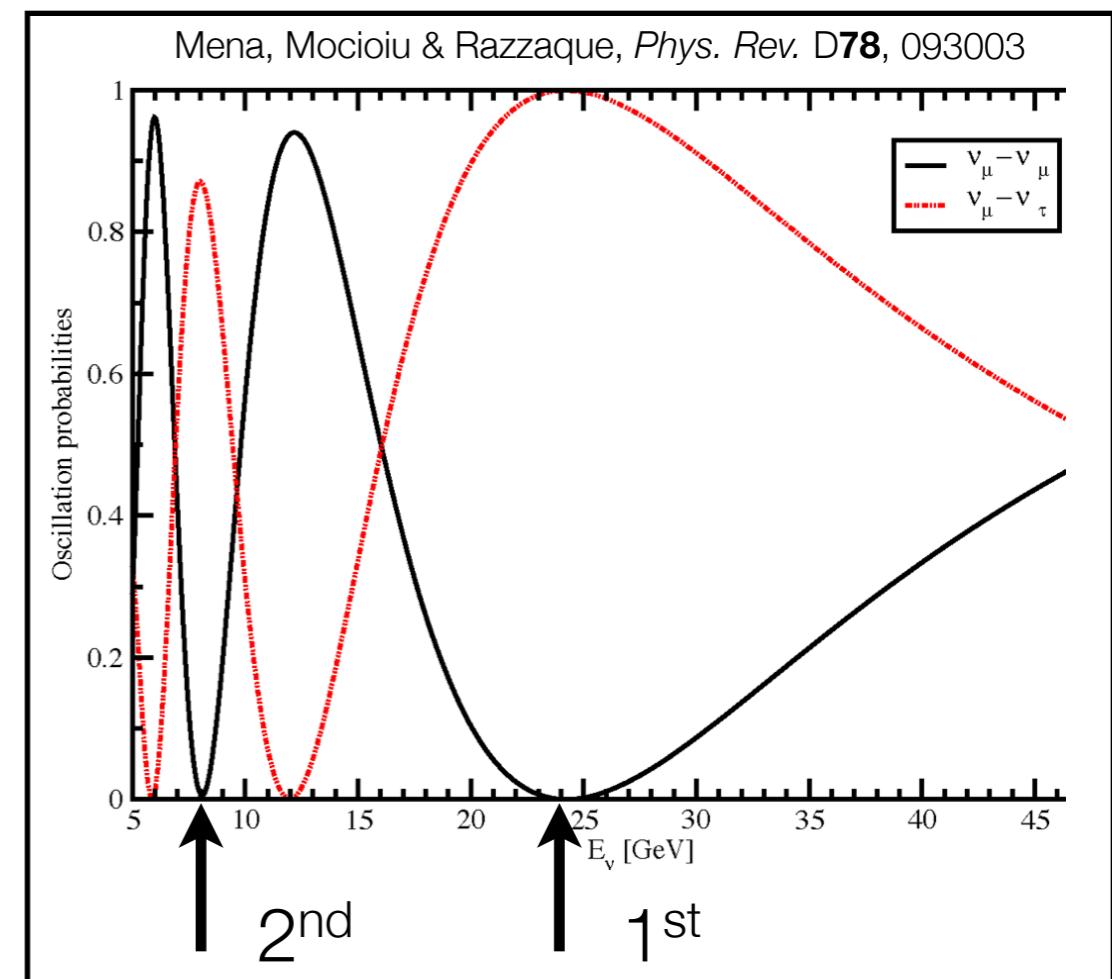
Physics Beyond DeepCore

- Physics/Detector
 - Hierarchy
 - Challenges
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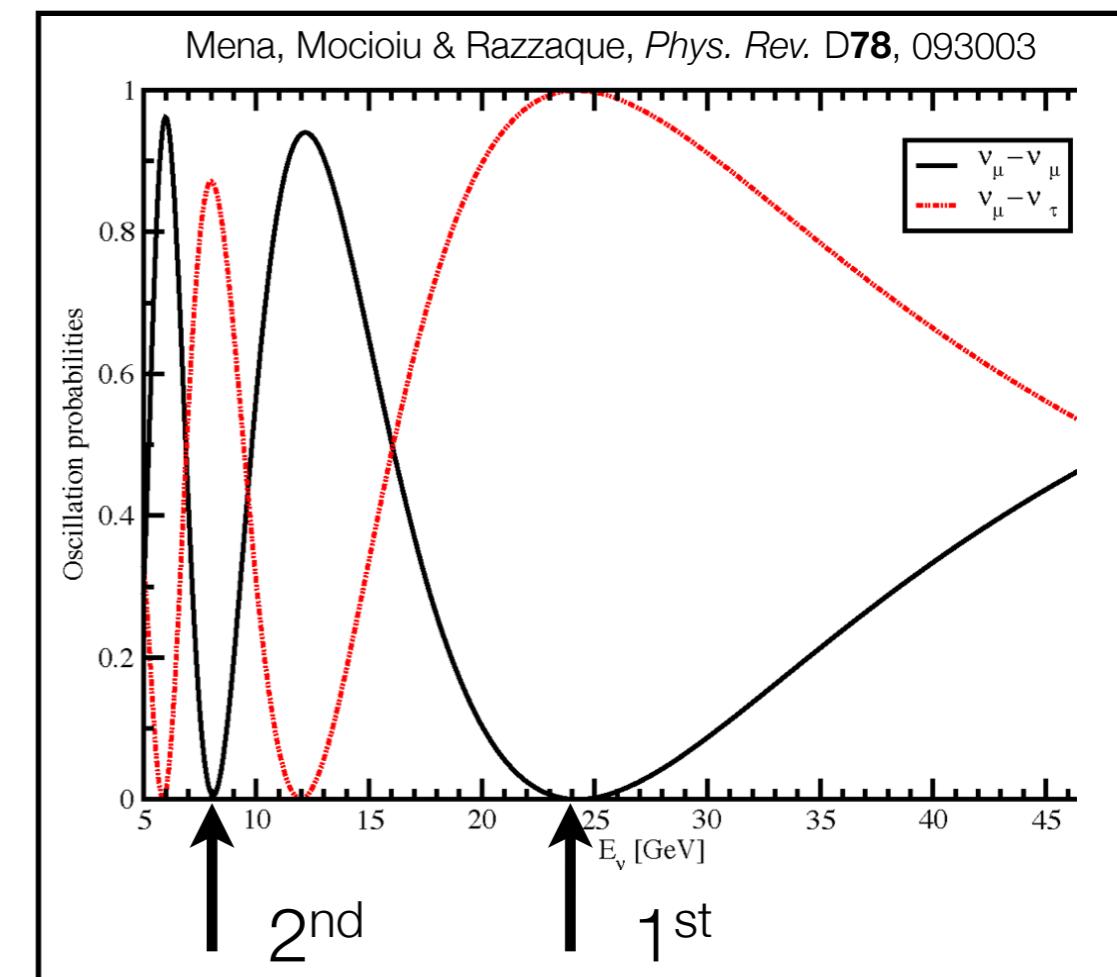
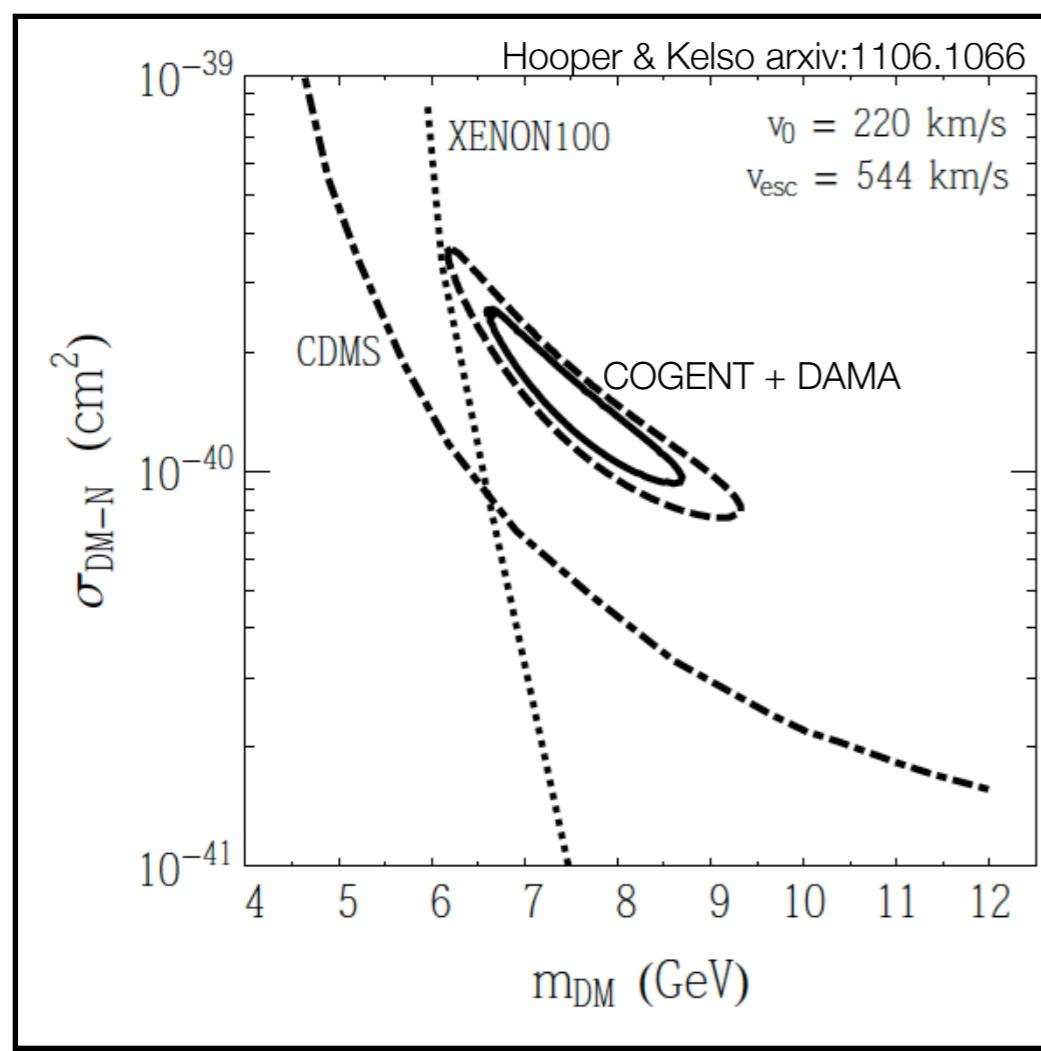
- Enhance/extend ongoing oscillation analyses - see E. Resconi



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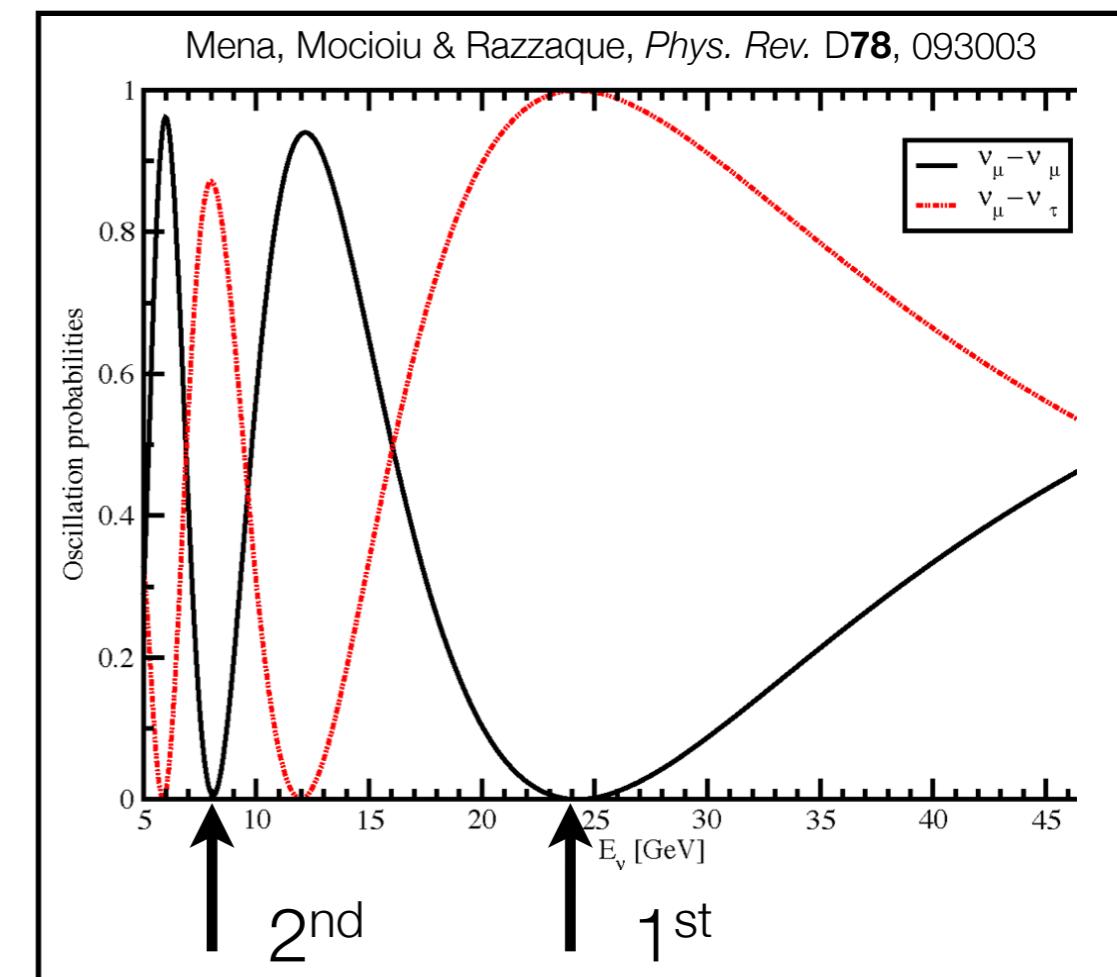
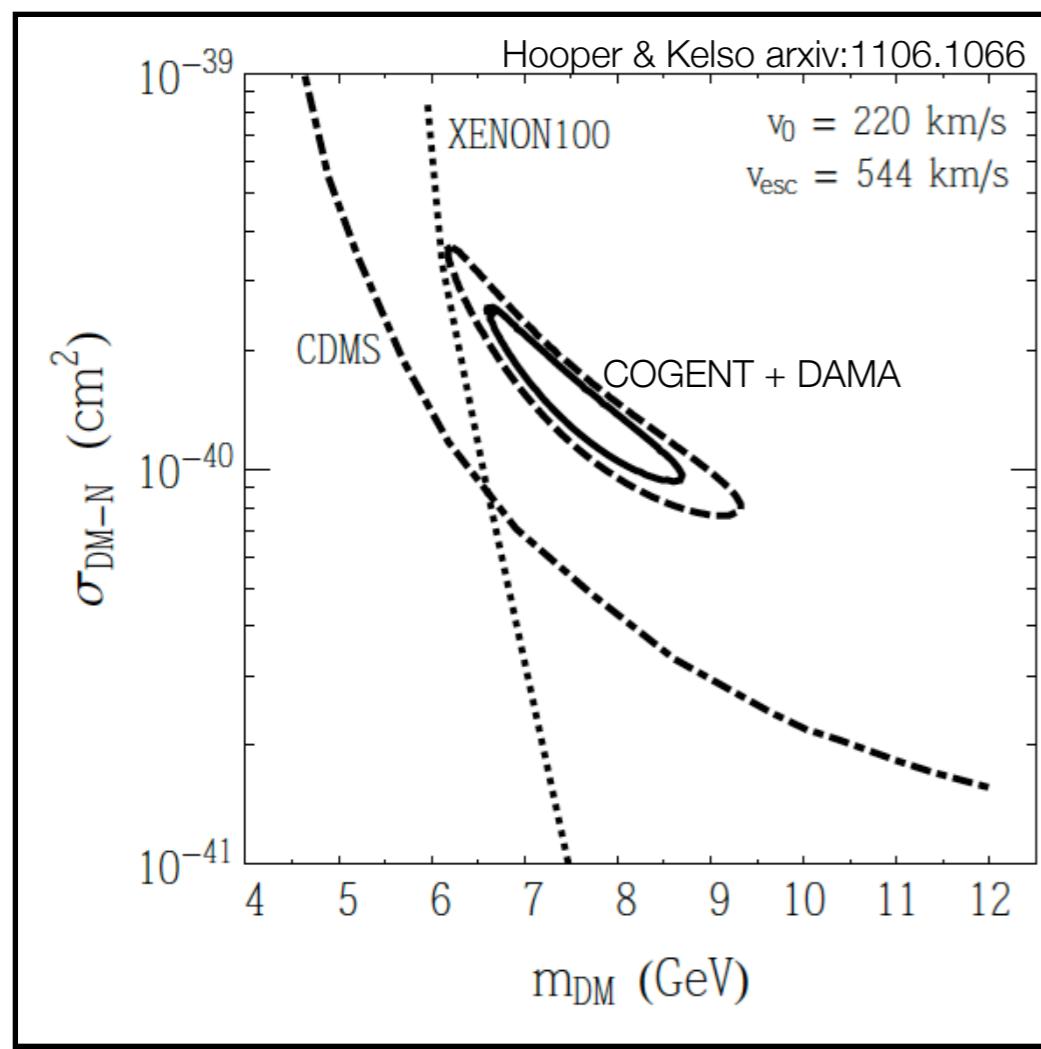
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- GeV Dark Matter - see C. Rott



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- θ_{23} maximal mixing (end of talk)



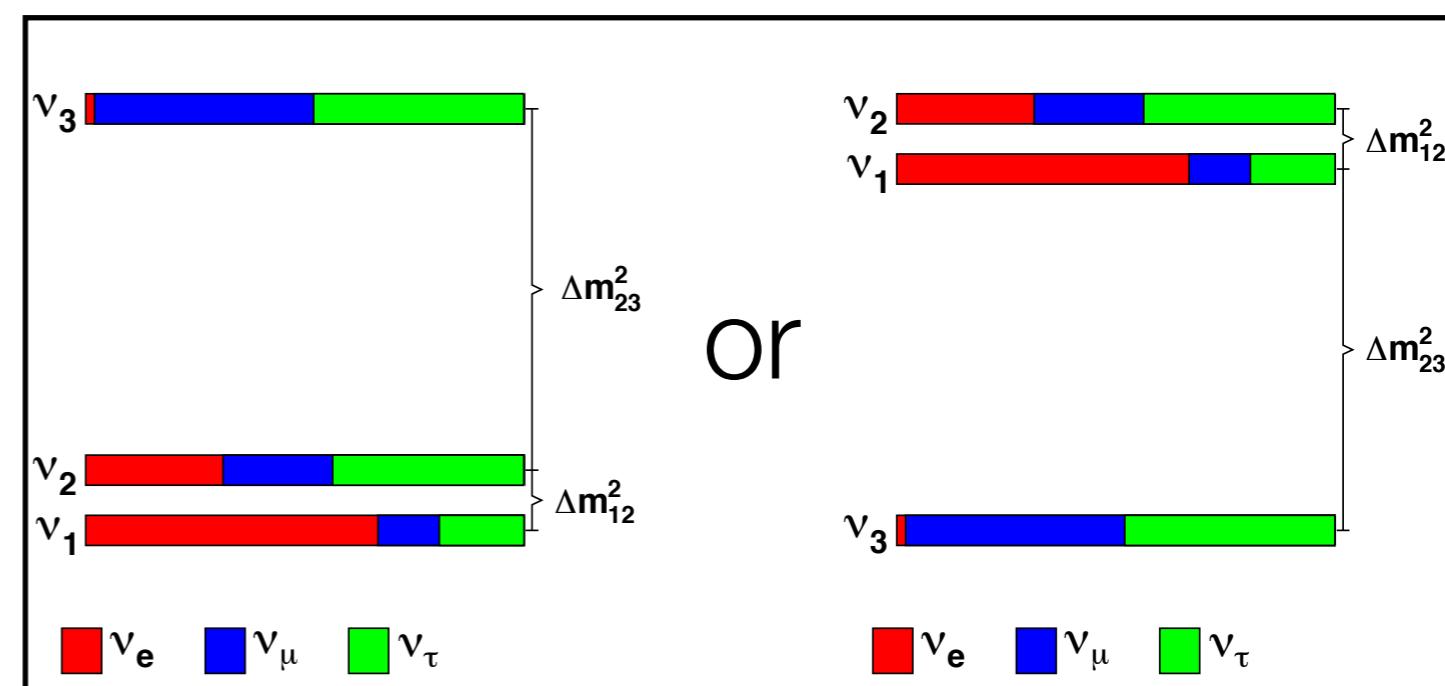
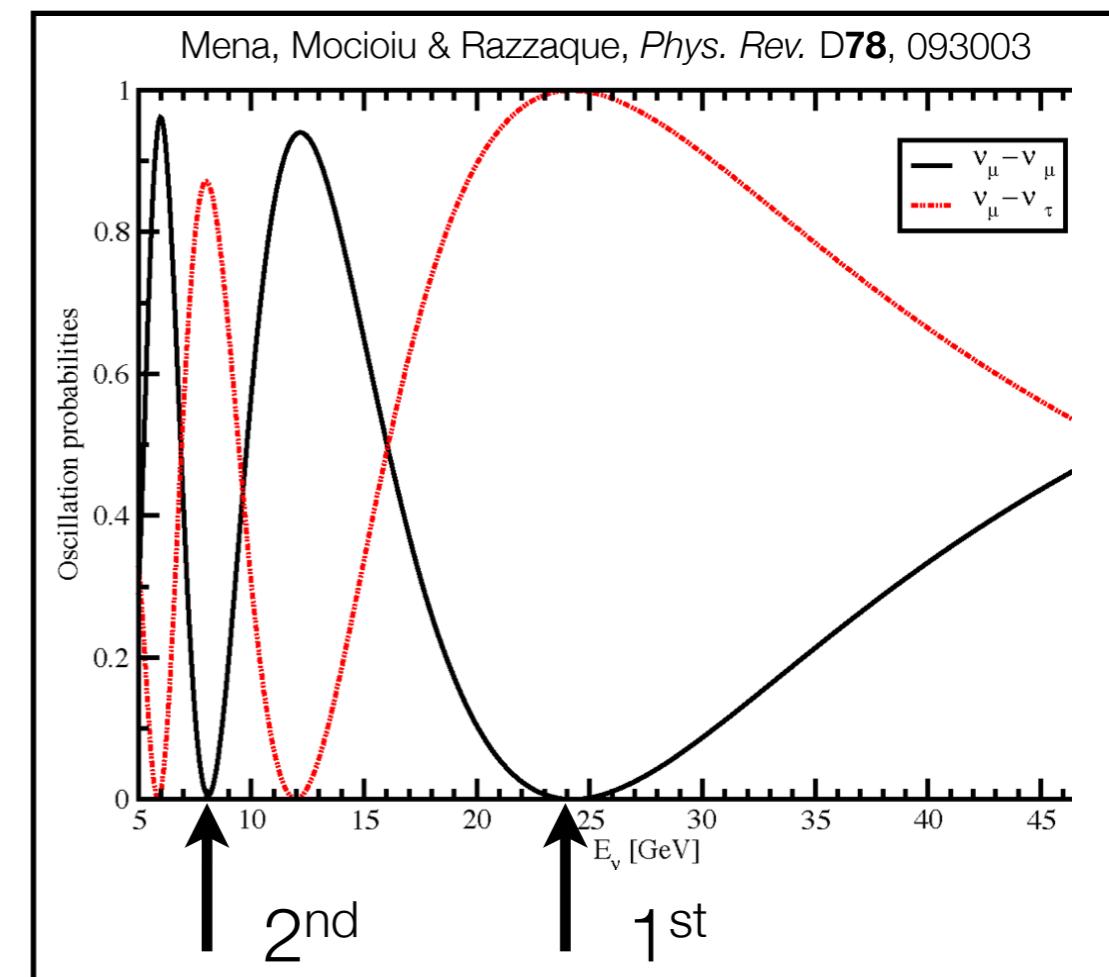
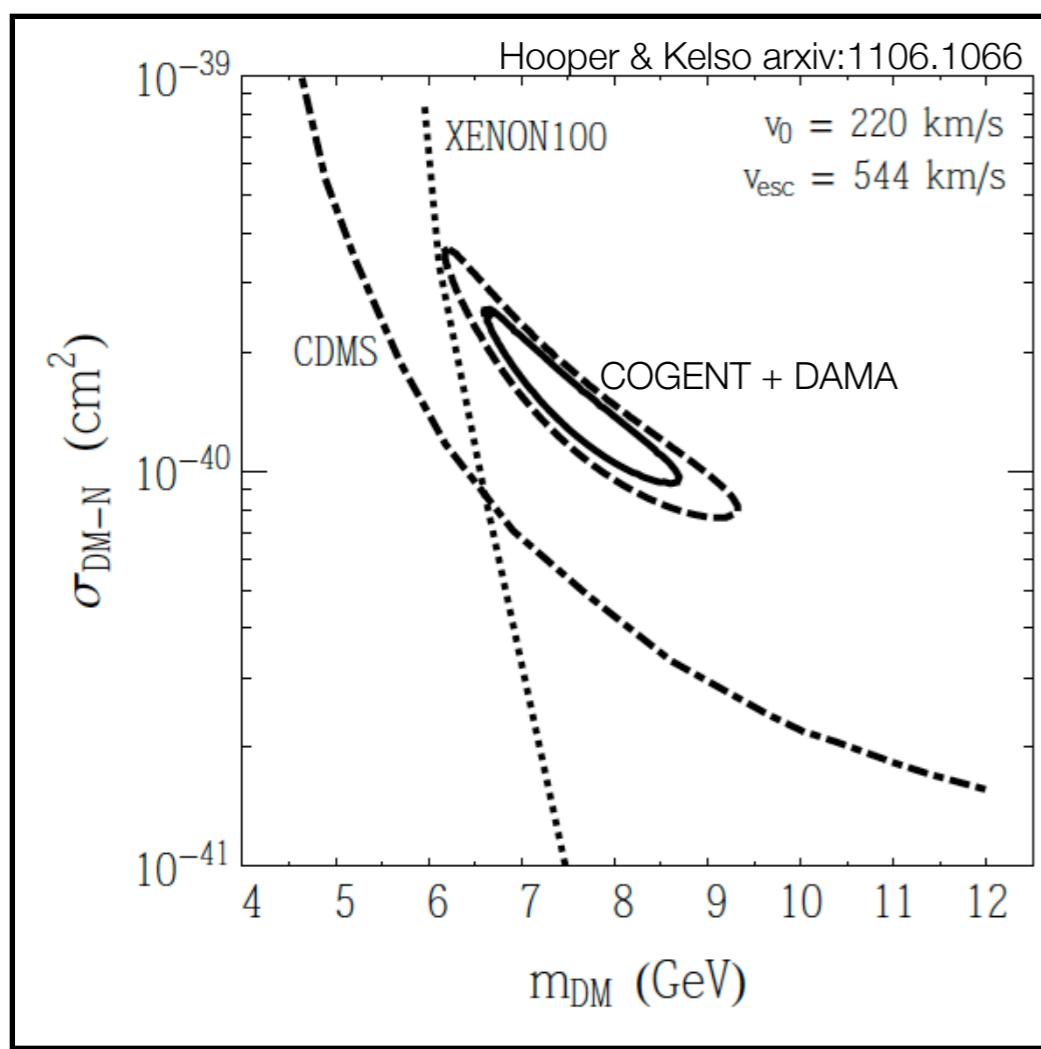
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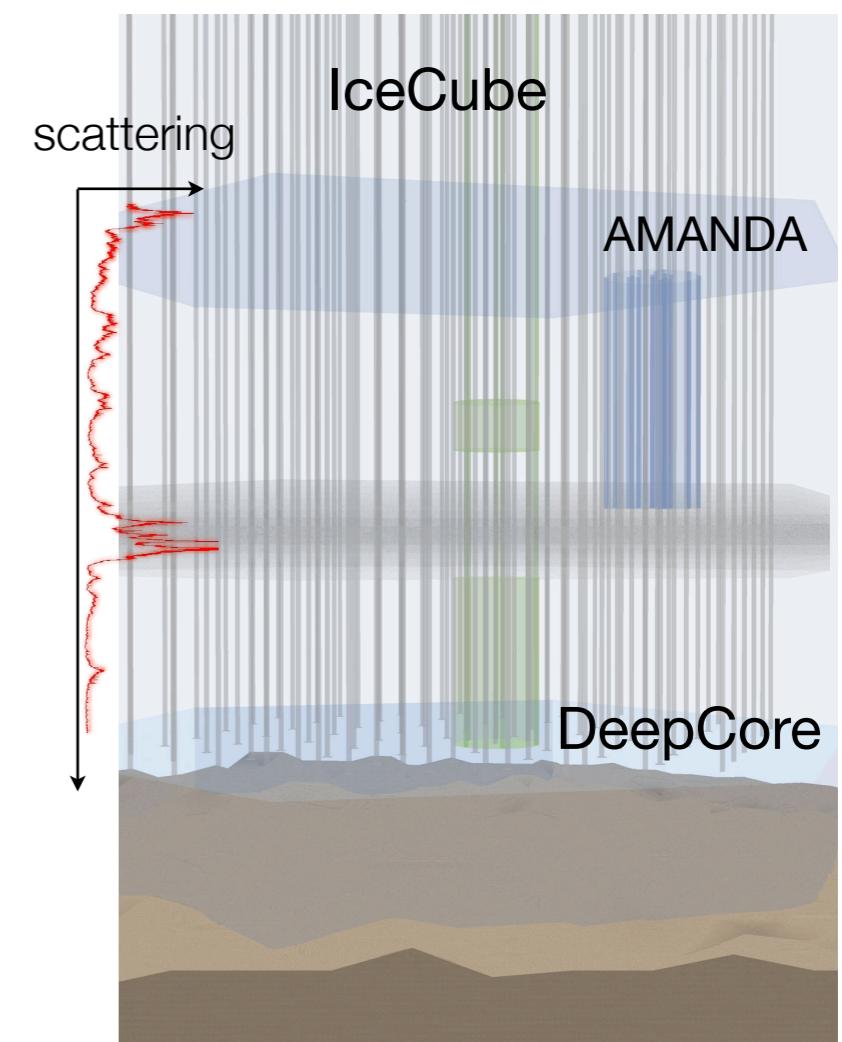
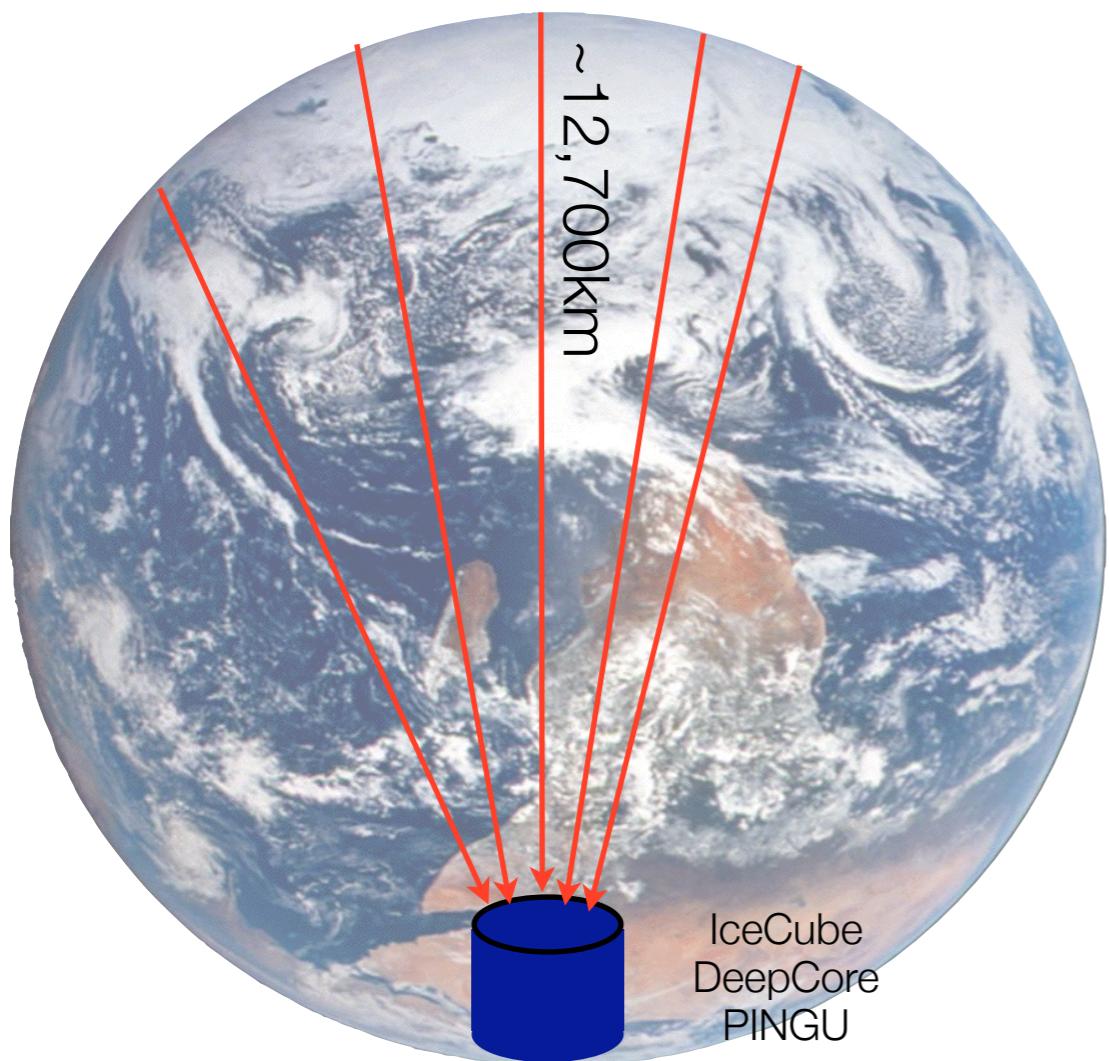
Neutrino Hierarchy



South Pole Neutrino Overview

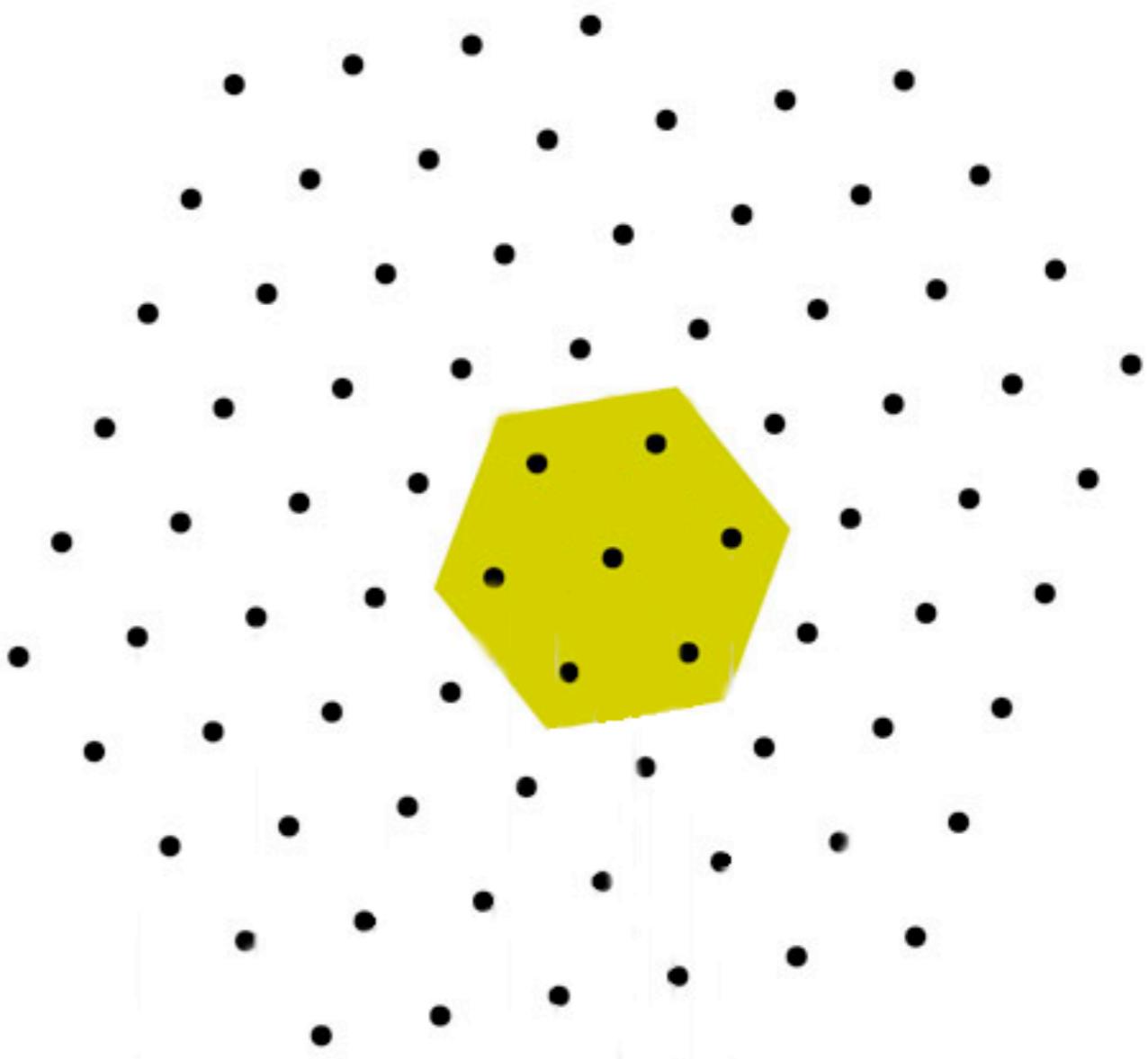
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- Free/cheap stuff is the best stuff
 - Atmospheric neutrino “beam” covers all baselines up to 12,700km and a large energy region (MeV - TeV+)
 - More clear ice than you can shake a stick at, just needs instrumentation
 - Gigaton veto (IceTop/IceCube/DeepCore) already built and operational



- Additional strings within IceCube/DeepCore volume
 - Number of strings, string-string spacing, DOM-DOM spacing, etc... under investigation
 - 10 different simulated geometries already
- Shorter DOM-DOM spacing than DeepCore
- R & D for future water/ice cerenkov detectors
- 1.5 year procurement/shipping + 2-3 year deployment

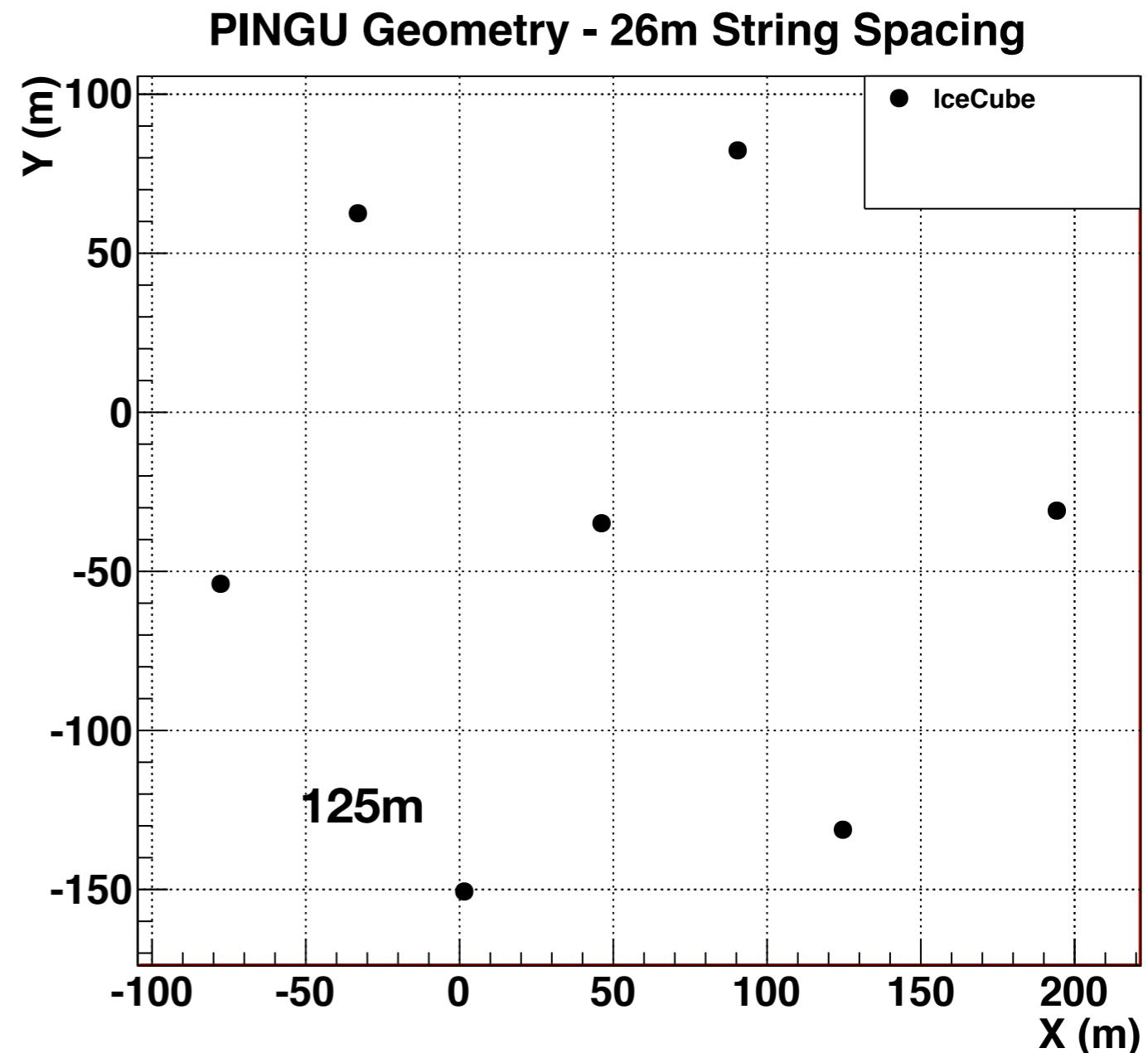
Top View



PINGU Detector

- Physics/Detector
- Hierarchy
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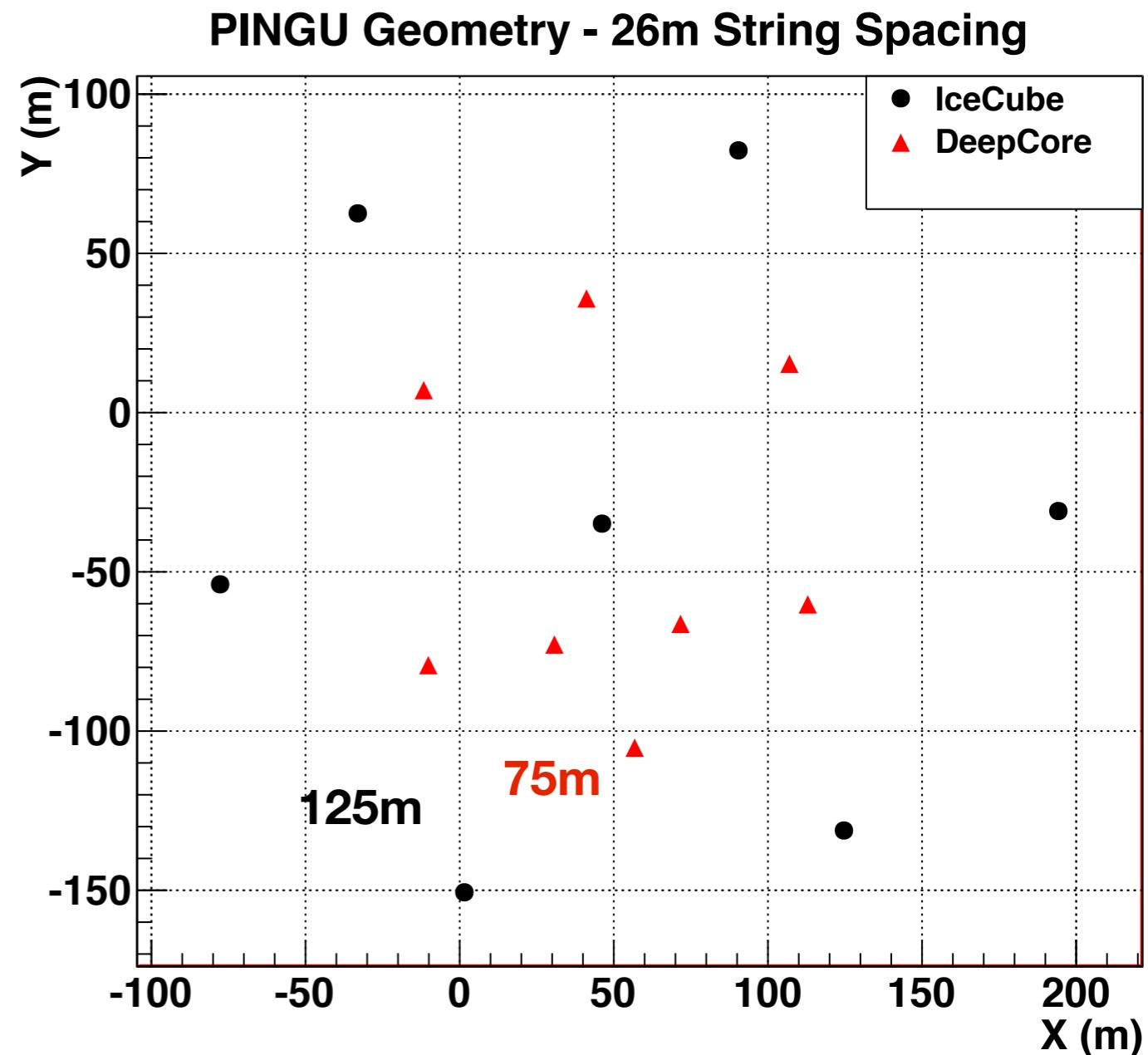
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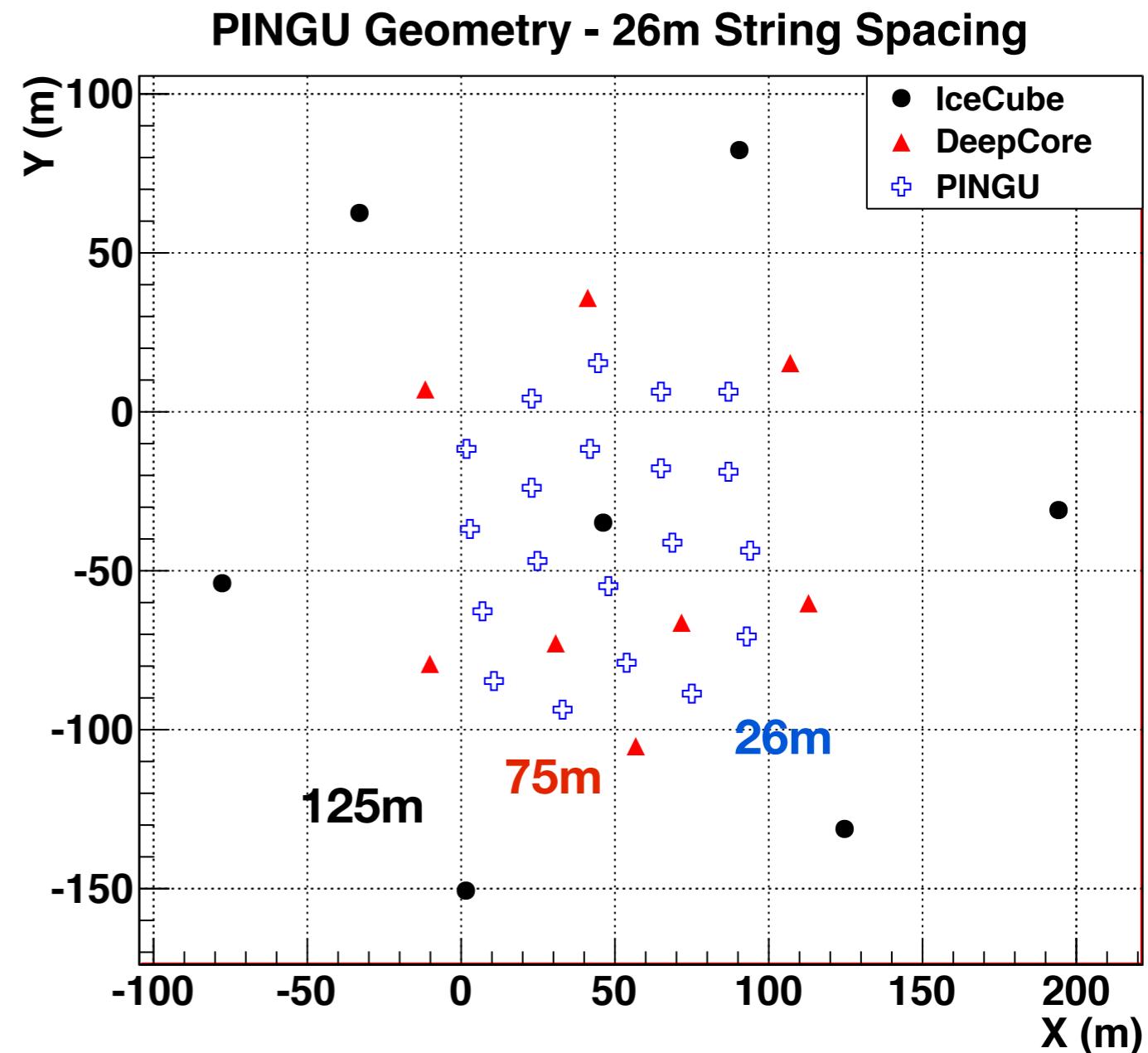
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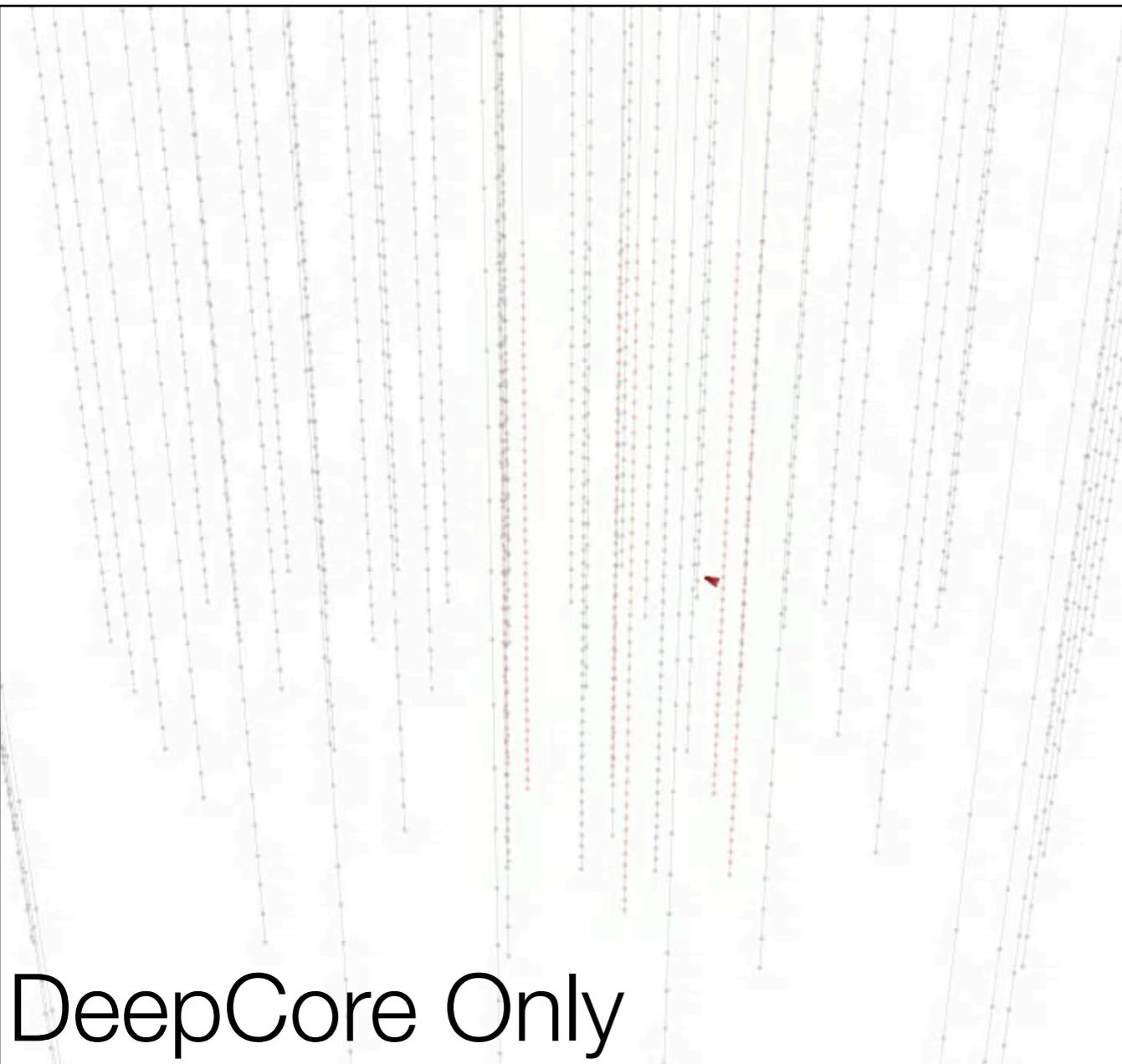


All following plots use V6 Geometry

V6 = 20 strings w/ 60 DOMs/string @ 26m string-string spacing, 5m DOM-DOM

Event Movies

- Physics/Detector
- Hierarchy
- Challenges



- 9.28 GeV Neutrino, 4.9 GeV muon, 4.5 GeV cascade

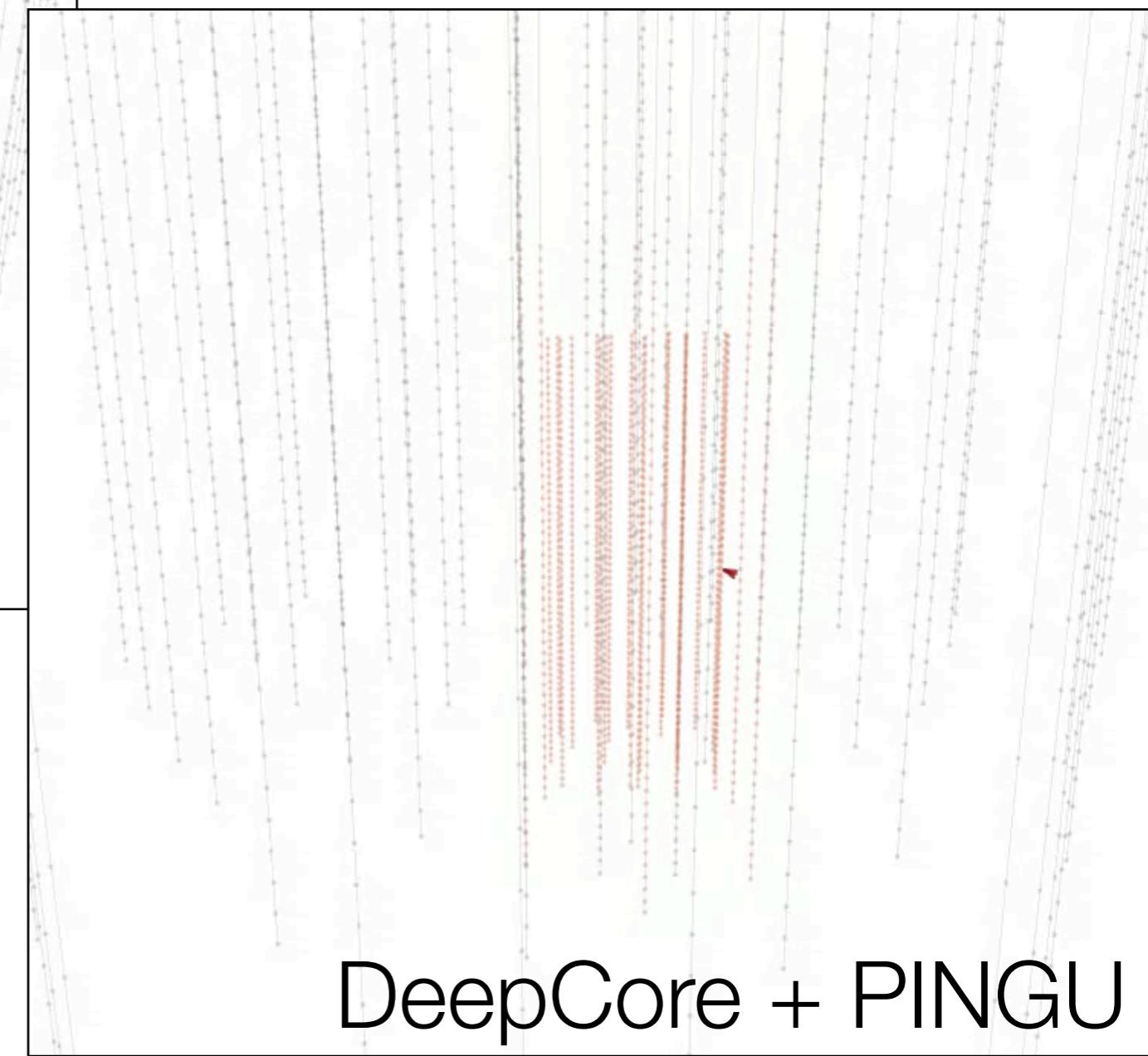
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DeepCore Only

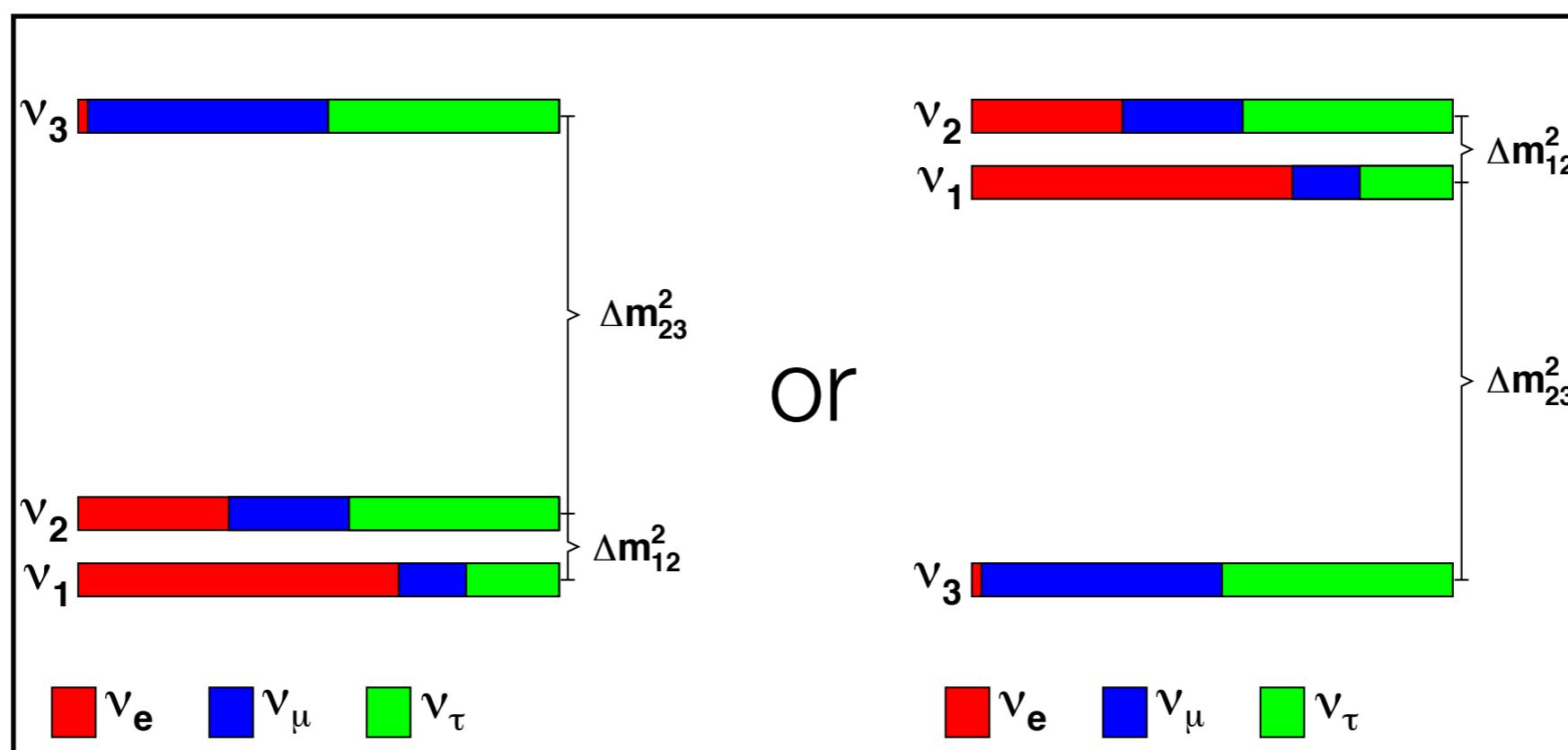
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- ~20 vs. ~50 Hit Modules



DeepCore + PINGU

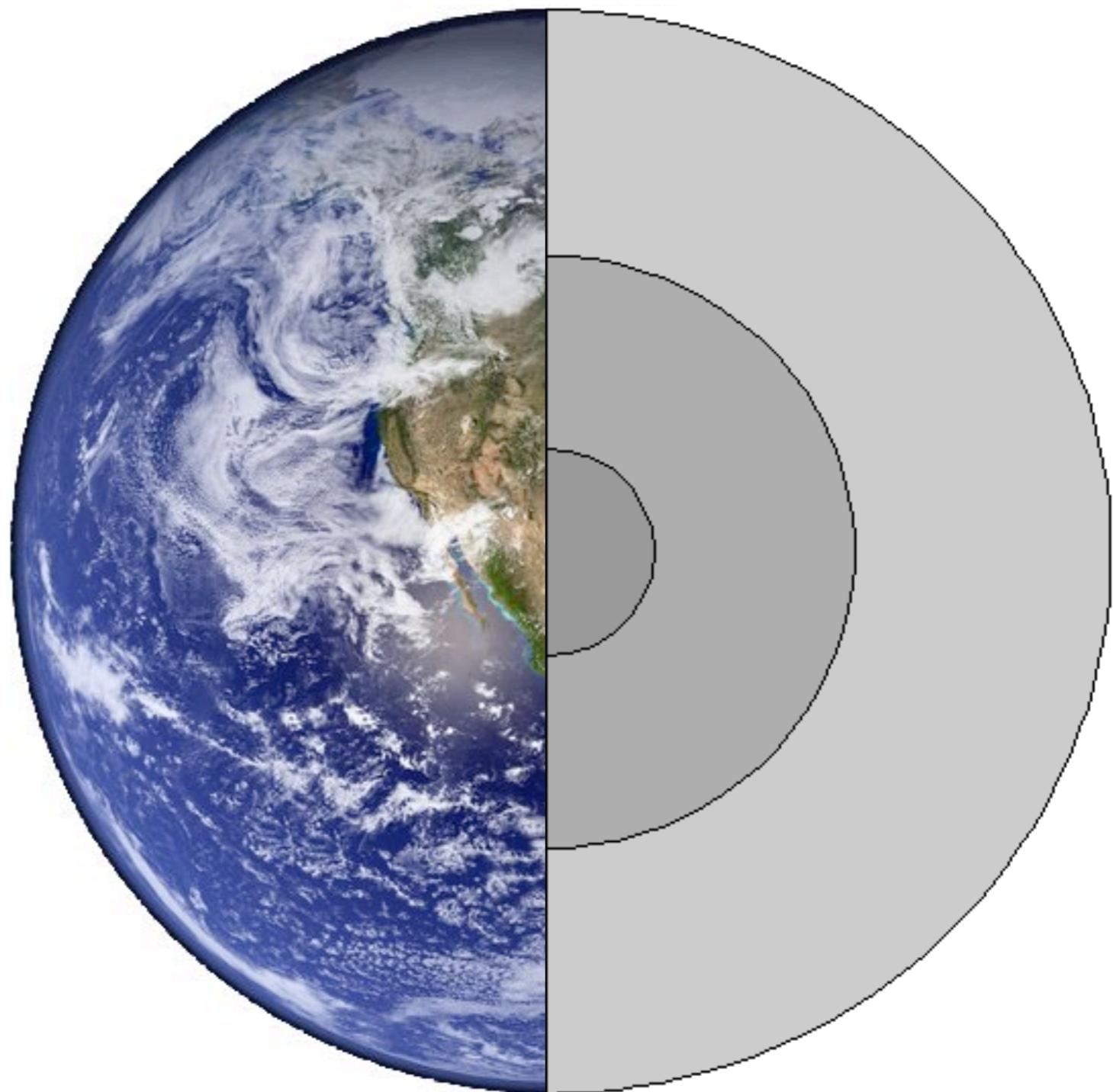
Resolving the Mass Hierarchy

Resolving the Mass Hierarchy



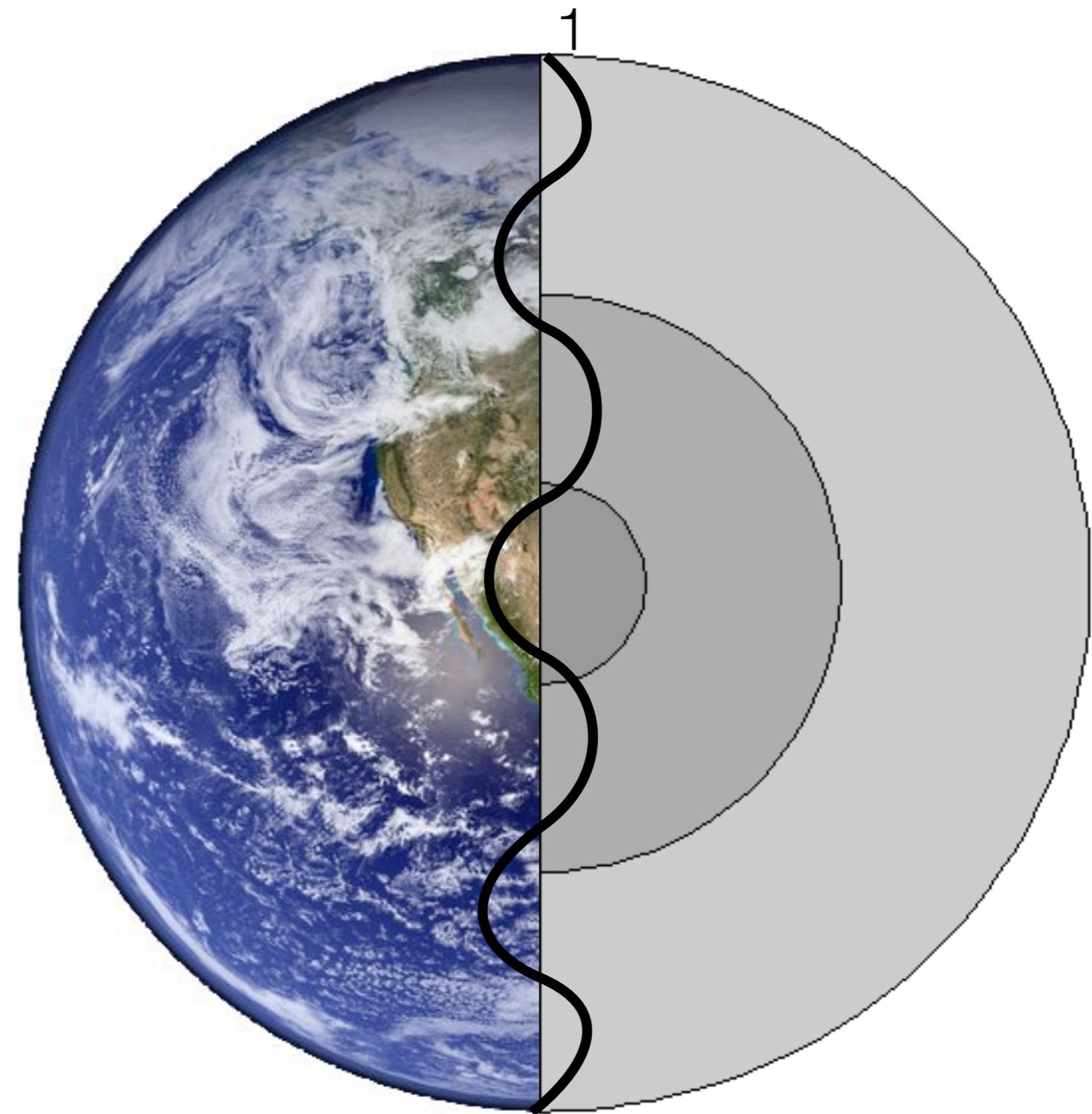
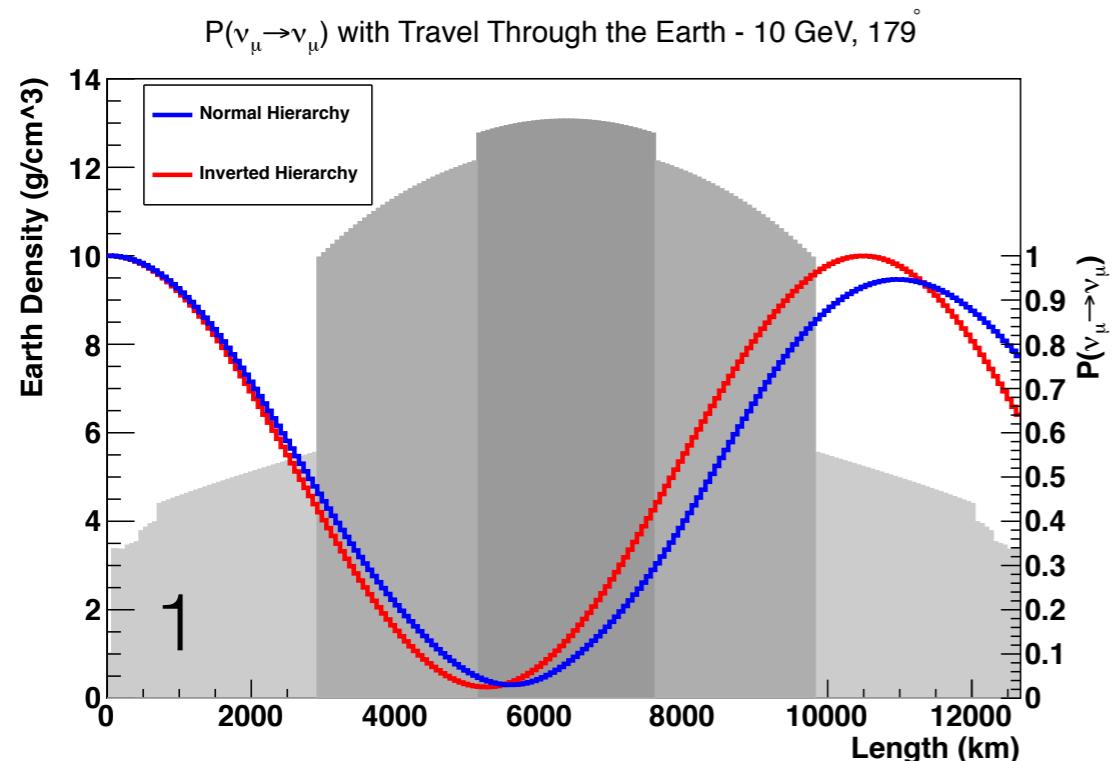
Matter Effects & Neutrino Hierarchy

- Physics/Detector
- Hierarchy
- Challenges



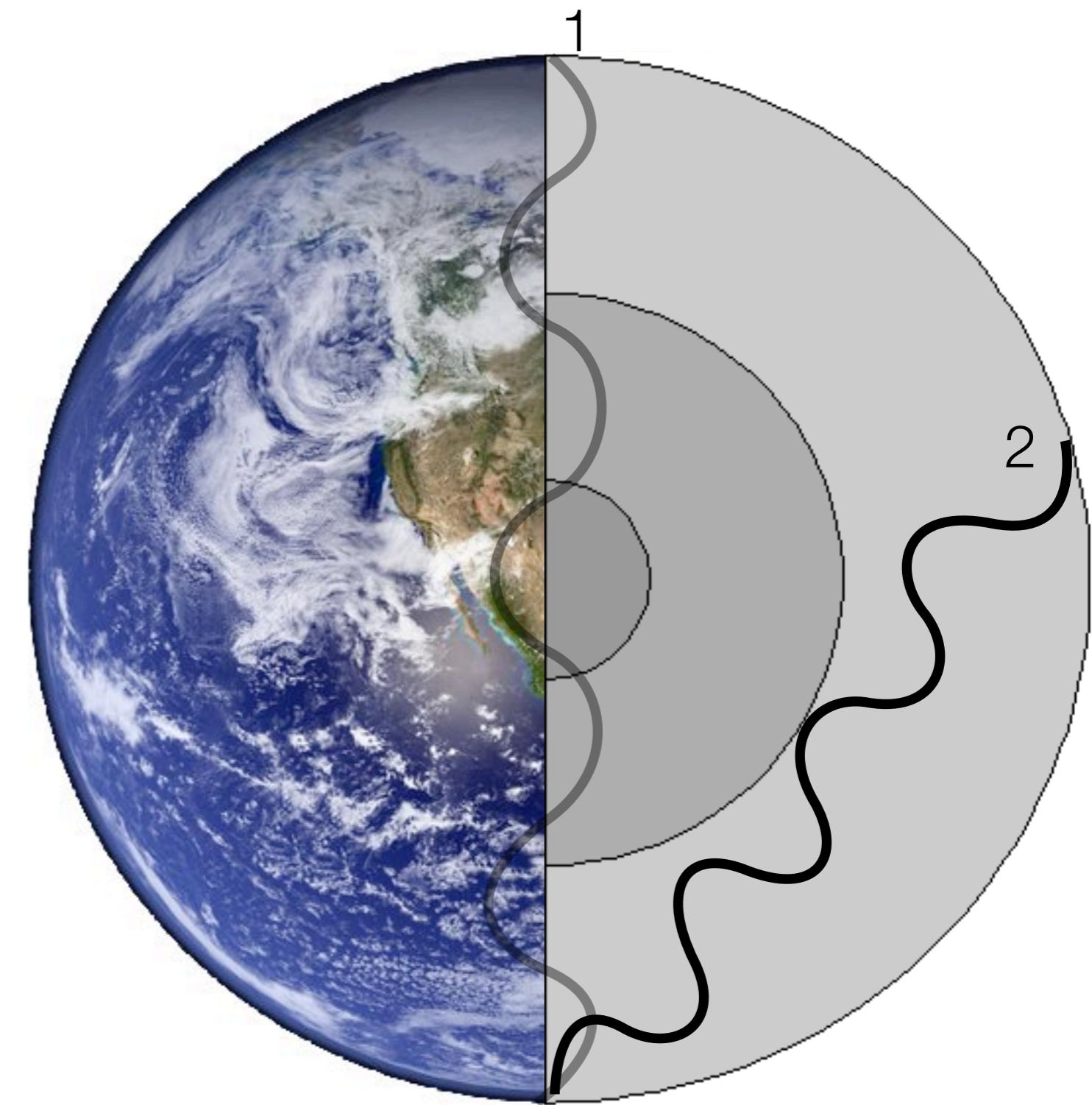
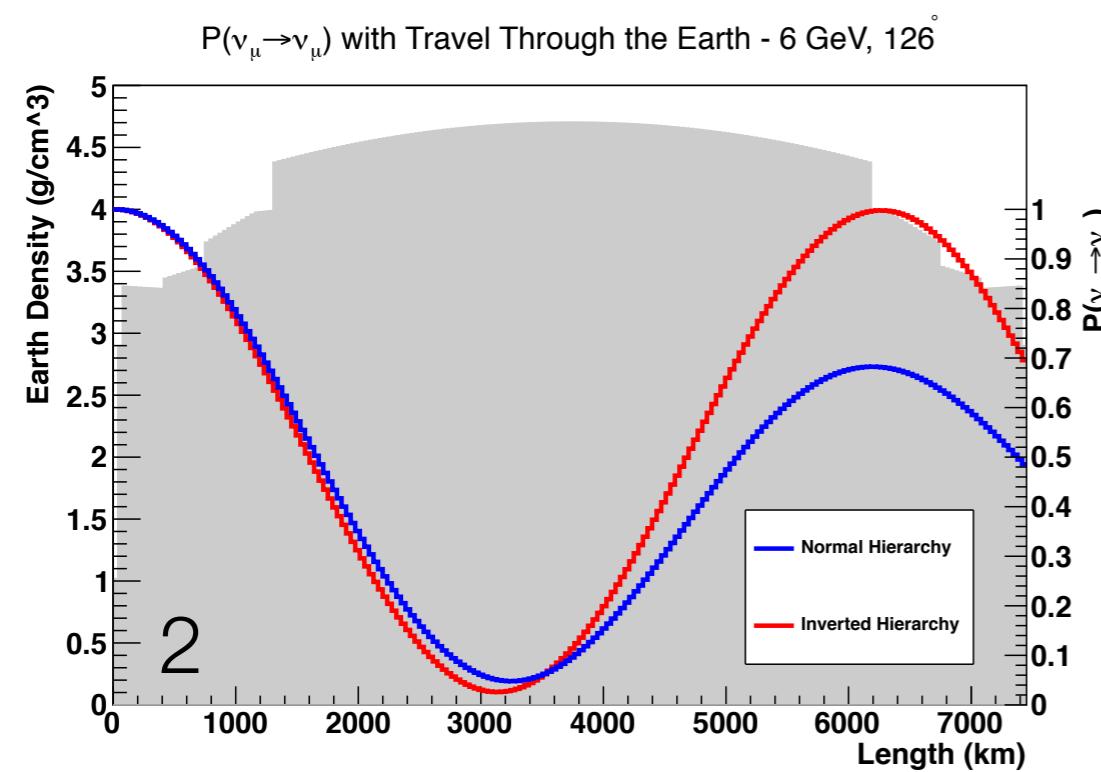
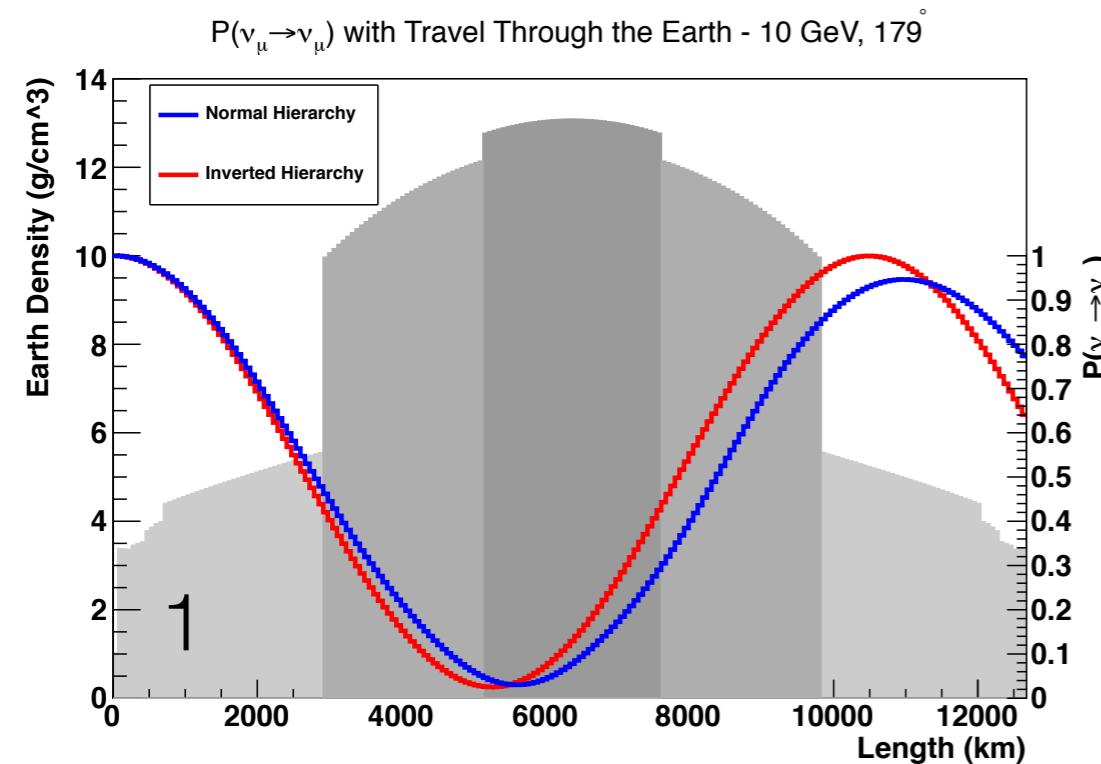
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Matter Effects & Neutrino Hierarchy

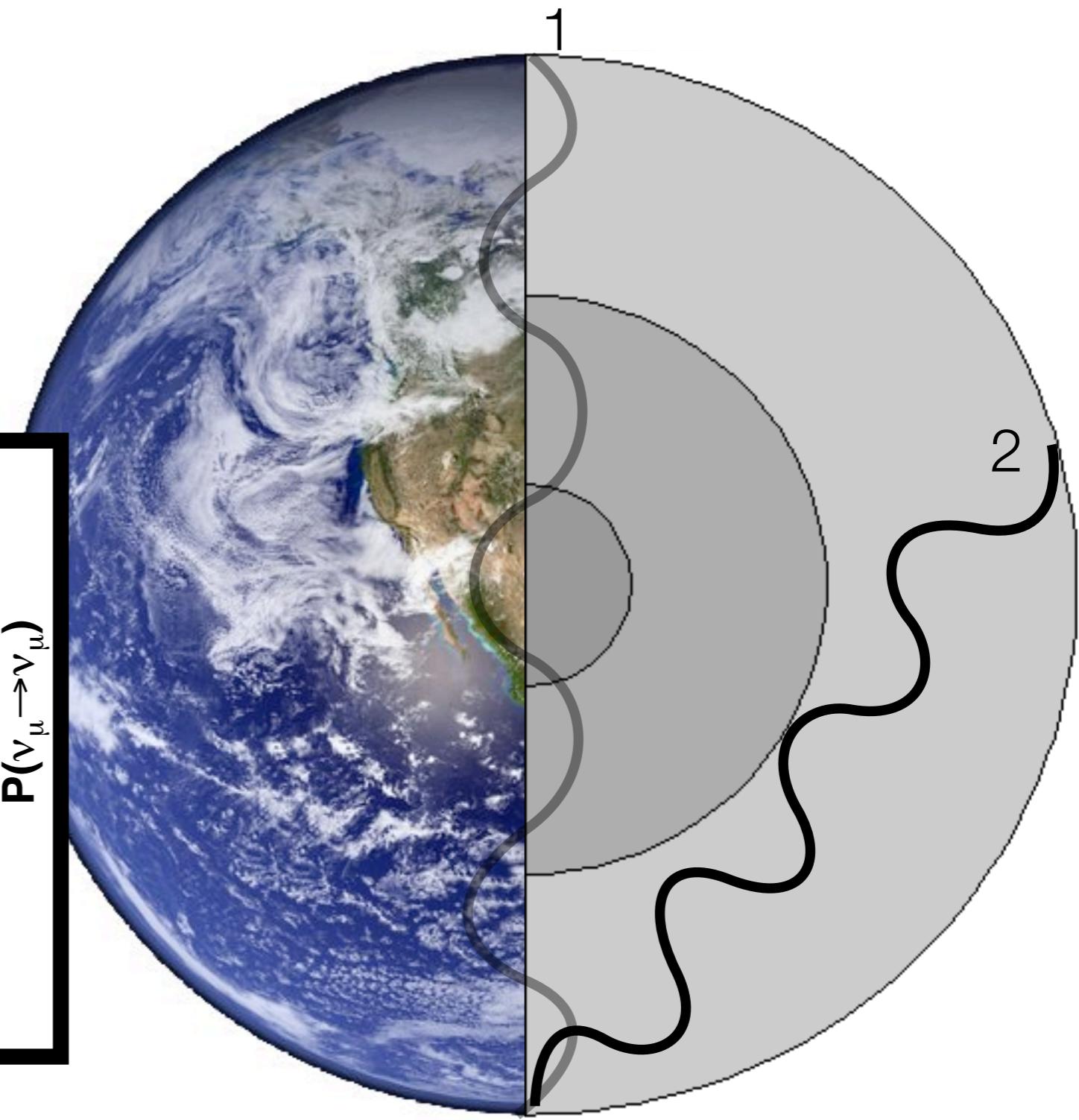
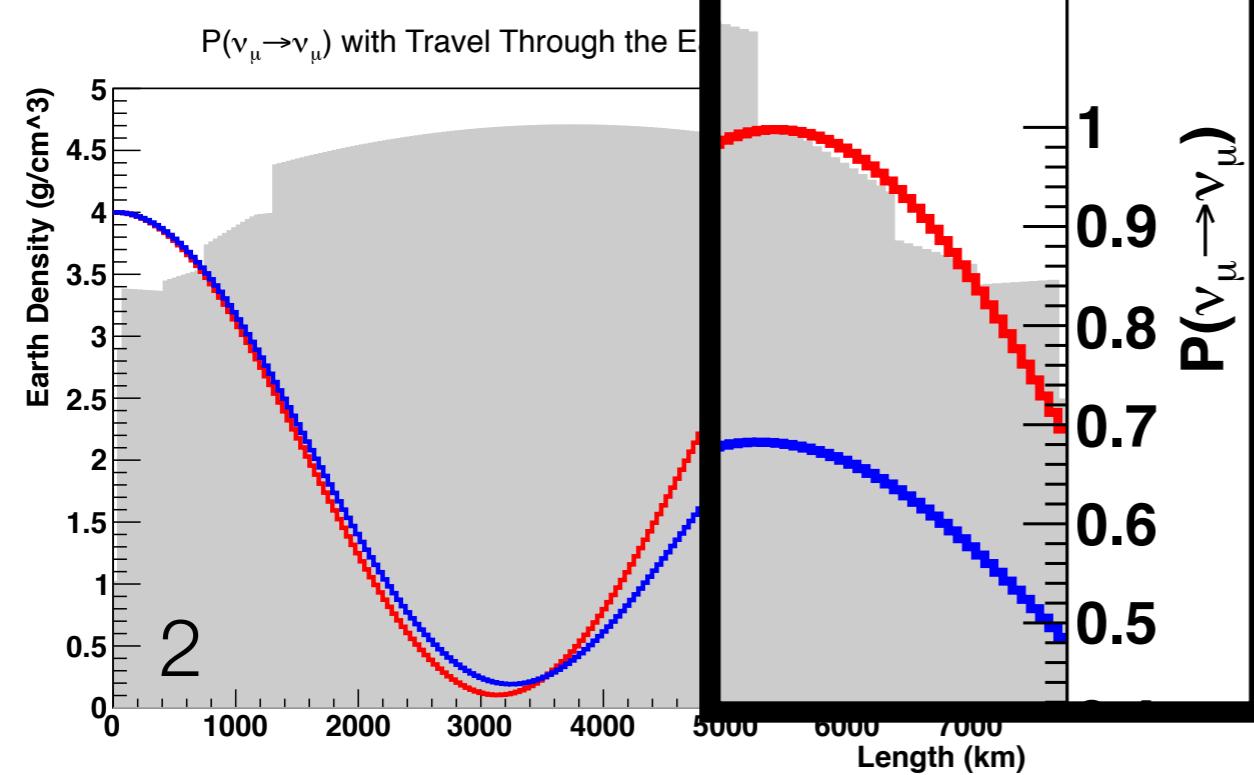
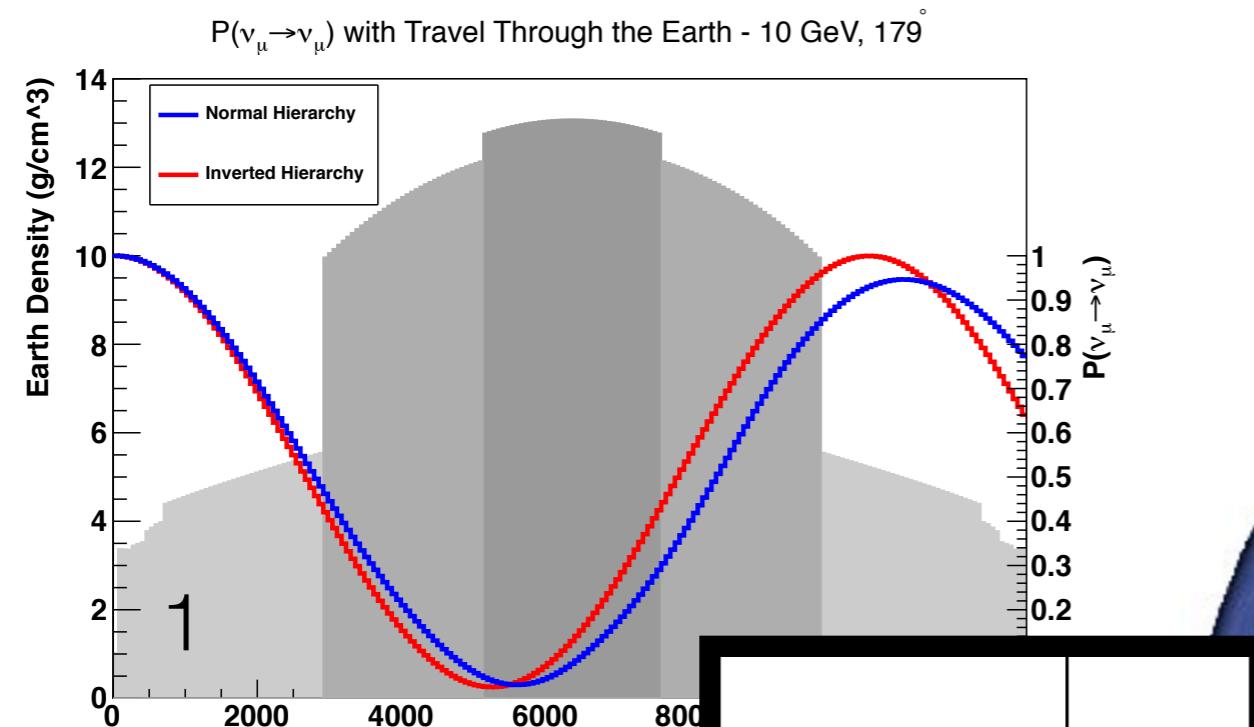
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- Inverted/Normal hierarchy has up to a 20% difference in ν_μ oscillation probability for specific energies and zenith angles (baselines)

Matter Effects & Neutrino Hierarchy

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- Inverted/Normal hierarchy has up to a 20% difference in ν_μ oscillation probability for specific energies and zenith angles (baselines)

- Use method outlined in Akhmedov, Razzaque, Smirnov - [arXiv:1205.7071](https://arxiv.org/abs/1205.7071)

$$S_{tot} = \sqrt{\sum_{ij} \frac{(N_{ij}^{IH} - N_{ij}^{NH})^2}{N_{ij}^{NH}}} \quad \begin{aligned} i &= \cos(\text{zenith}) \\ j &= \text{energy} \\ V^{eff} &= \text{effective volume} \end{aligned}$$

$$N_{i,j}^{NH} = P(\nu_\mu)_{i,j}^{NH} * \Phi(\nu_\mu)_{i,j} * \sigma(\nu_\mu)_j * V_{i,j}^{eff} + P^{NH}(\bar{\nu}_\mu)_{ij} * \Phi(\bar{\nu}_\mu)_{i,j} * \sigma(\bar{\nu}_\mu)_j * V_{i,j}^{eff}$$

- Essentially bin, sum, and subtract one hierarchy from the other.
It works because:

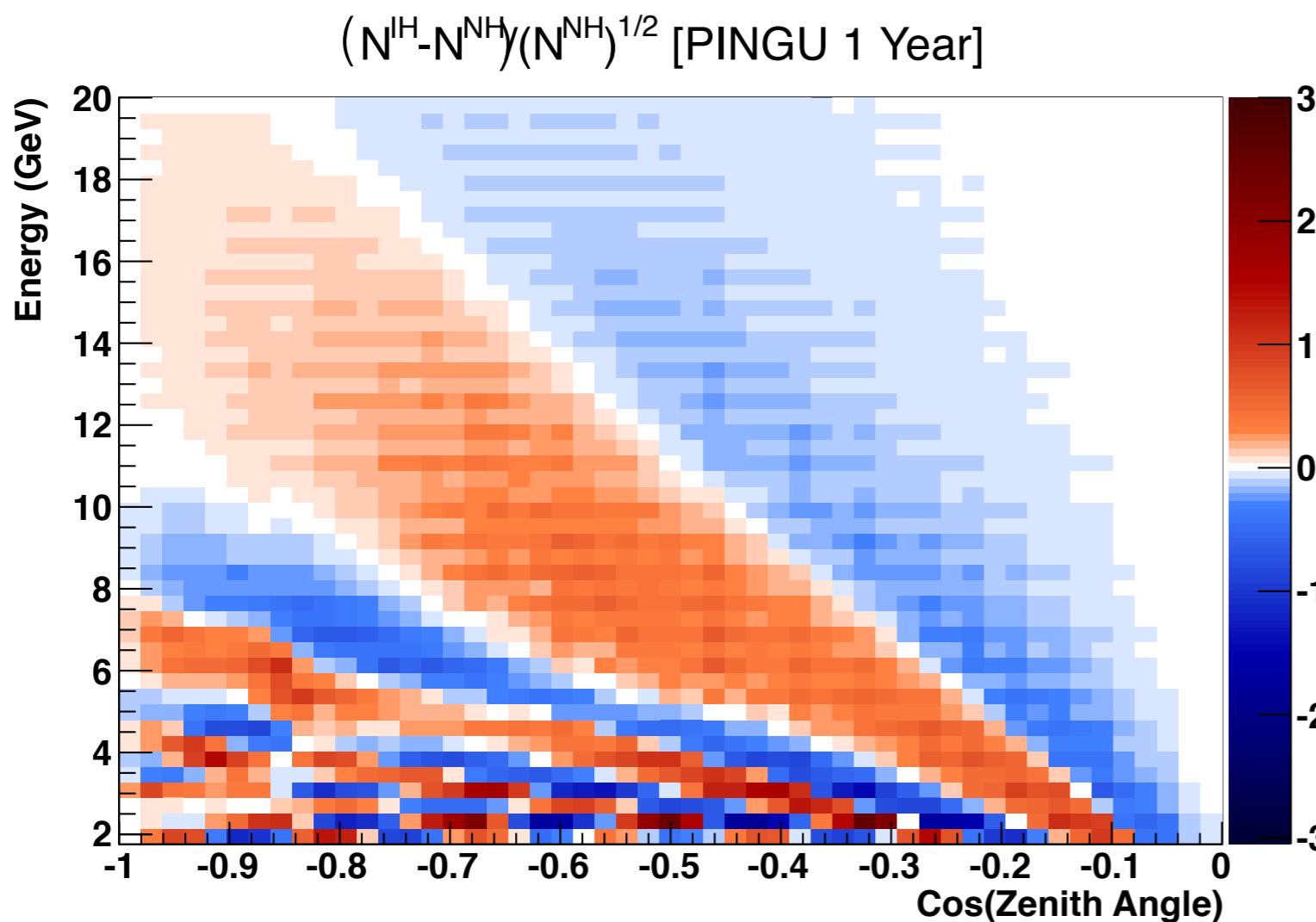
Probability : $P(\nu_\mu)^{IH} + P(\bar{\nu}_\mu)^{IH} \neq P(\nu_\mu)^{NH} + P(\bar{\nu}_\mu)^{NH}$

Flux : $\Phi(\nu_\mu) > \Phi(\bar{\nu}_\mu)$

Cross – Section : $\sigma(\nu_\mu) > \sigma(\bar{\nu}_\mu)$

Hierarchy

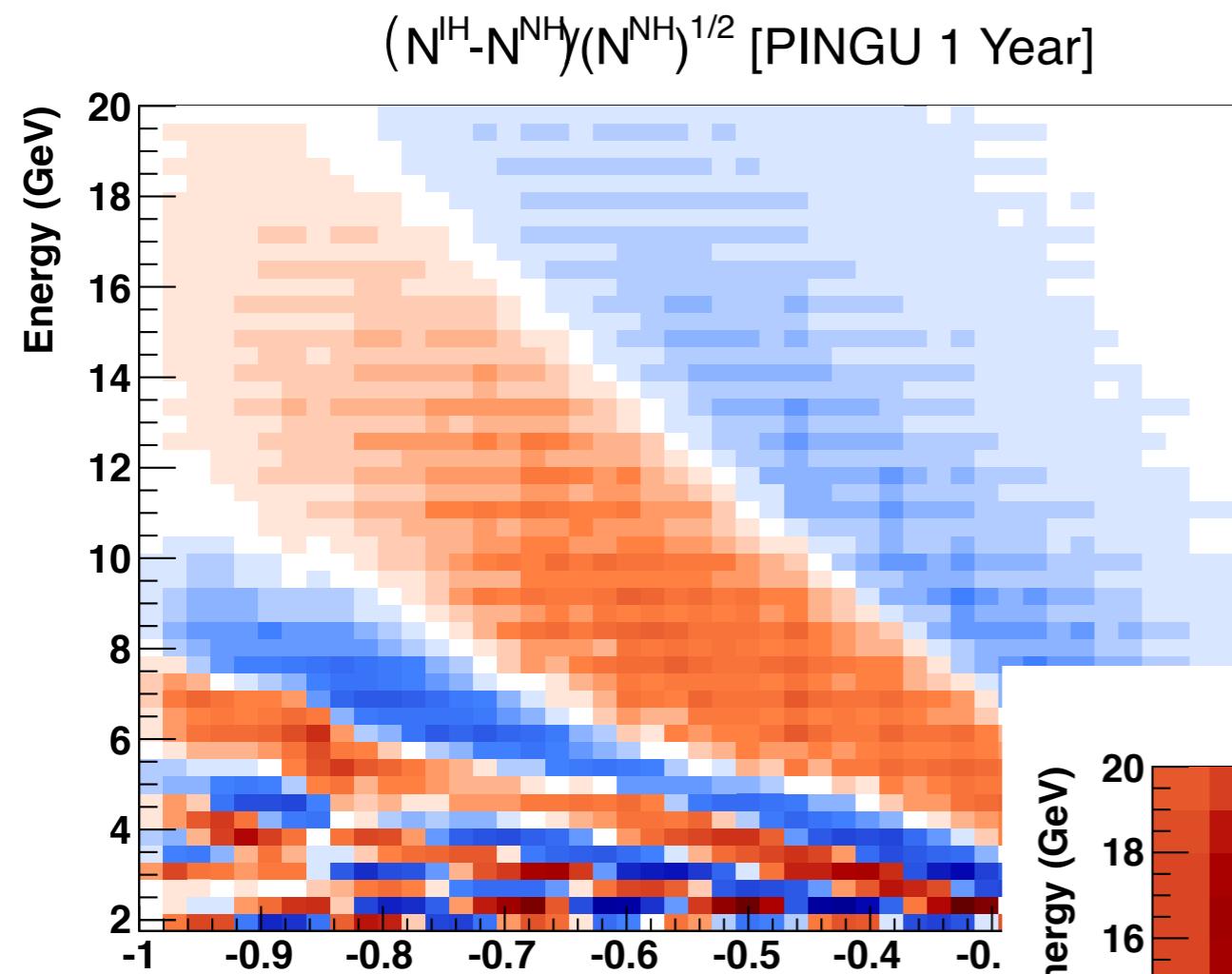
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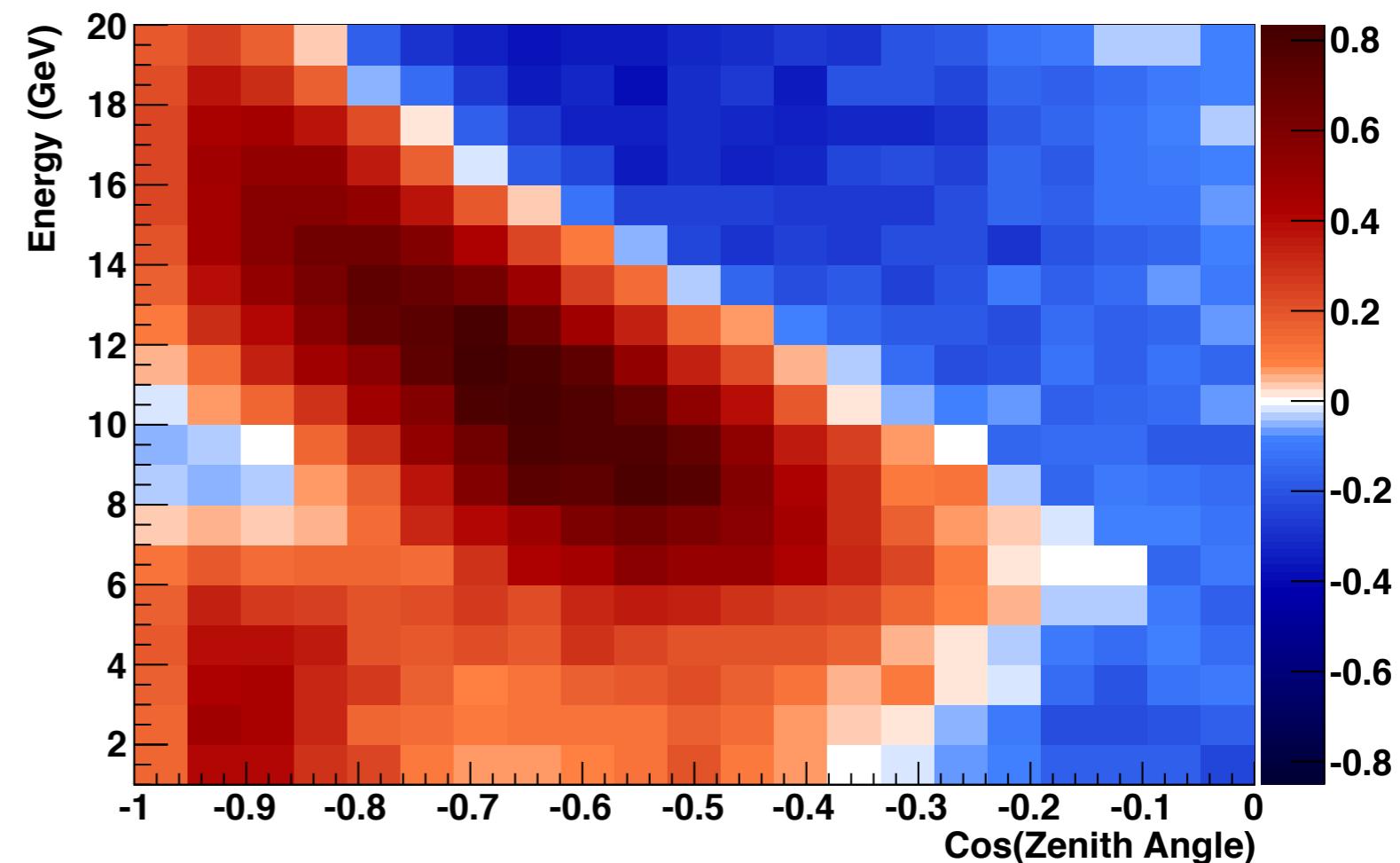
- **Idealized case** w/ perfect event ID, 100% event selection efficiency, no quality cuts and no background
- Detector response (ice modeling, DOM efficiency) may play a major role in final sensitivity
- Evaluations of angular and energy resolution are ongoing

Hierarchy

- Physics/Detector
- Hierarchy
- Challenges



smeared: 3 GeV in ν_μ energy and
11.25° in μ zenith resolution

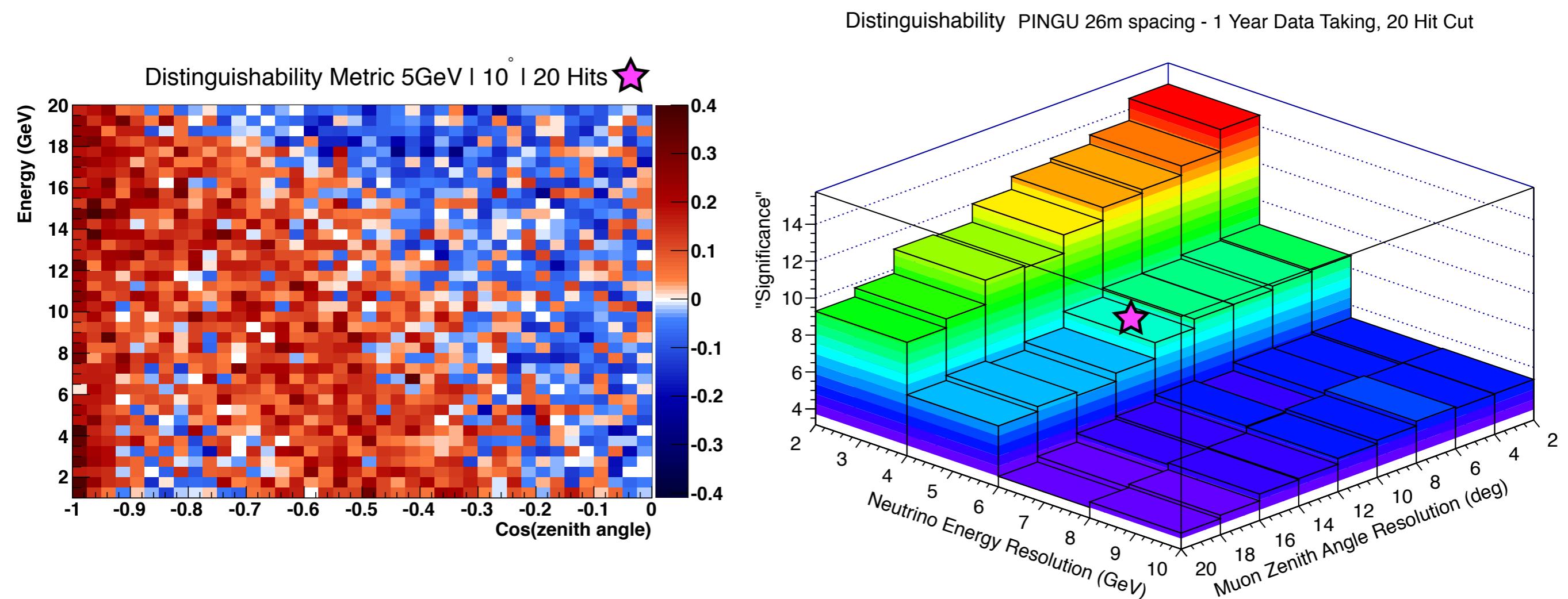


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Detector/Reconstruction Issues

- Physics/Detector
- Hierarchy
- Challenges

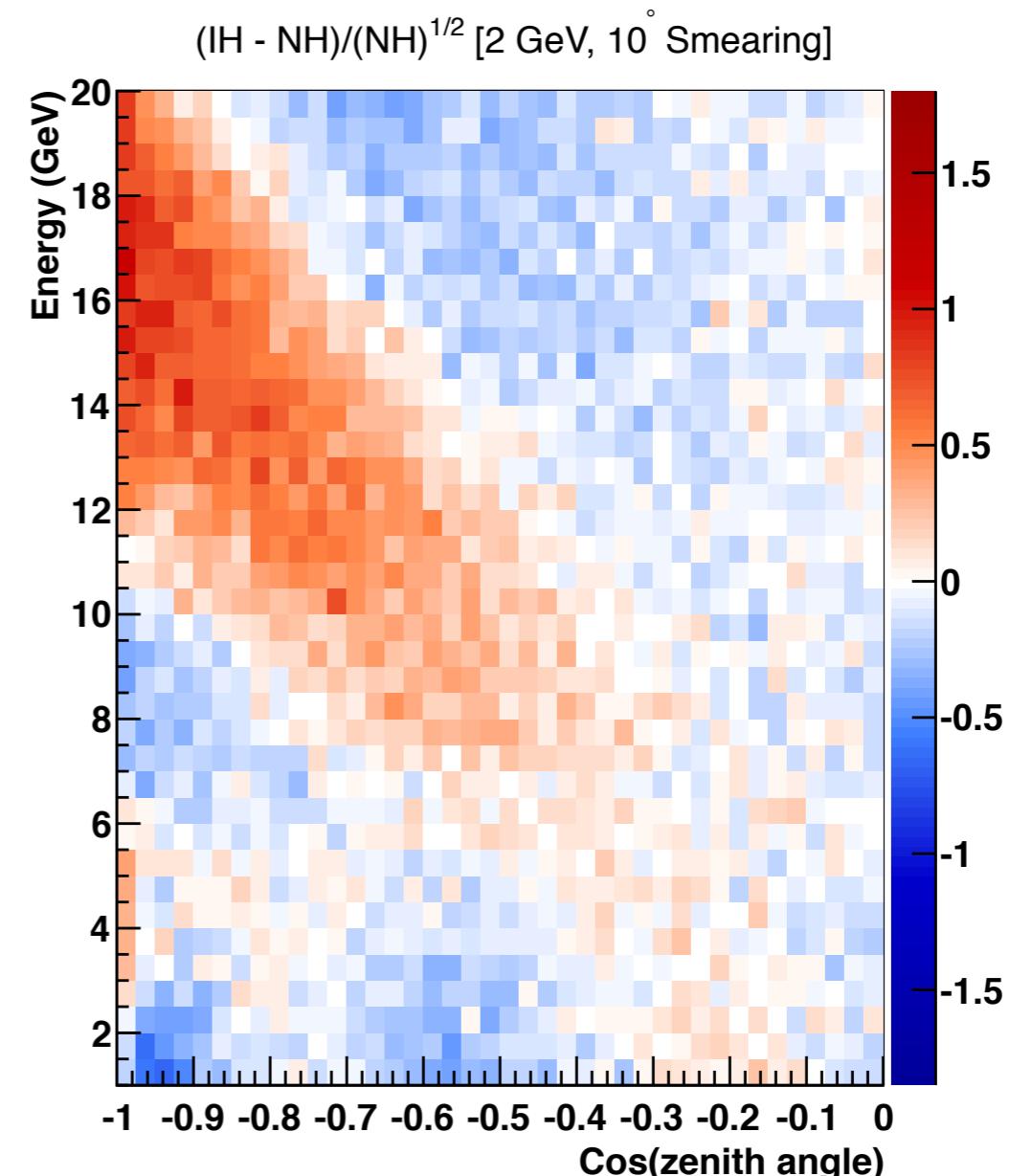
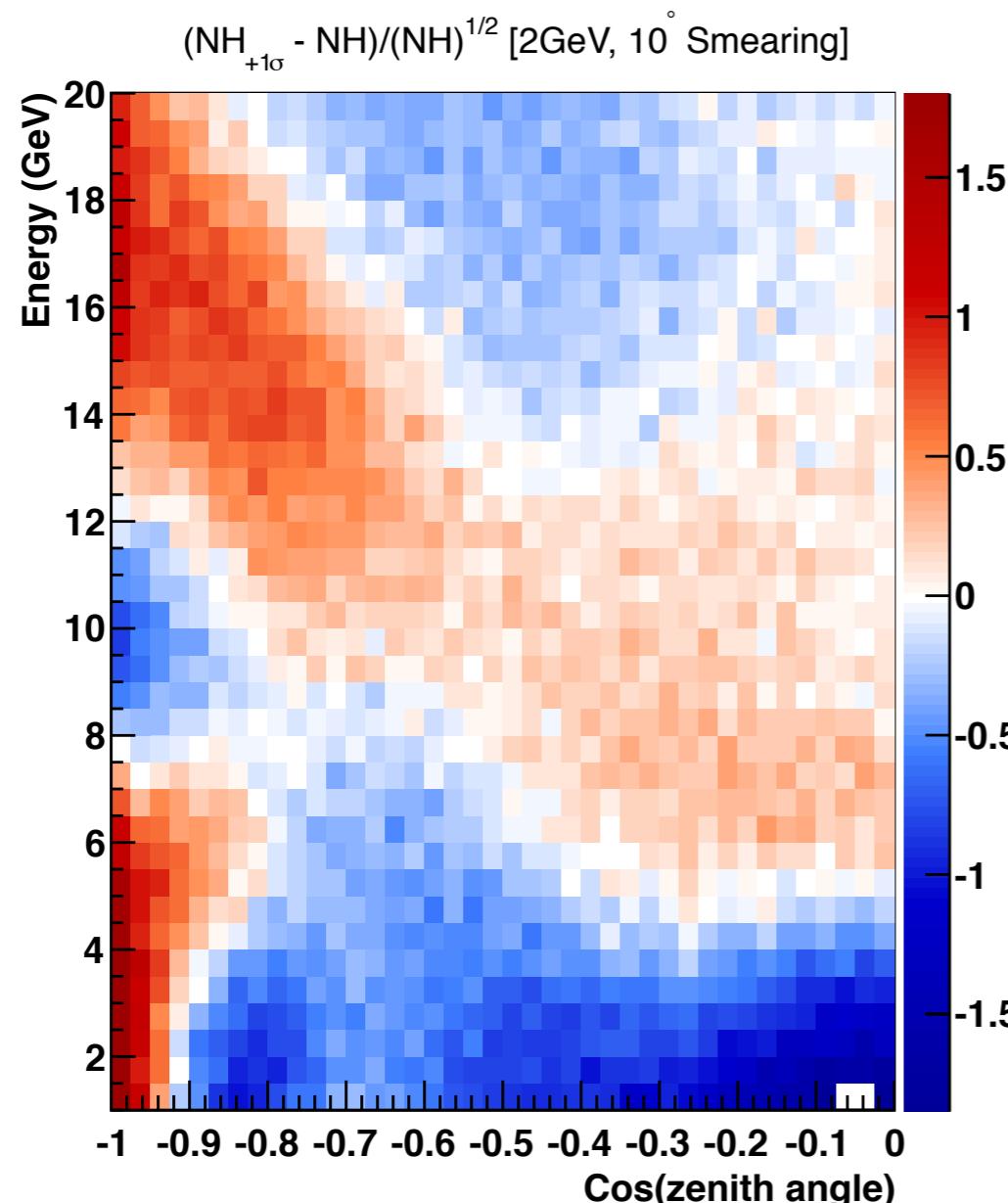
- There are benchmarks that relate hierarchy distinguishability to reconstruction uncertainties
 - A cut on 20 hits represents a reconstruction efficiency
 - Provide targets for ongoing reconstruction effort



Physics Challenges

- Physics/Detector
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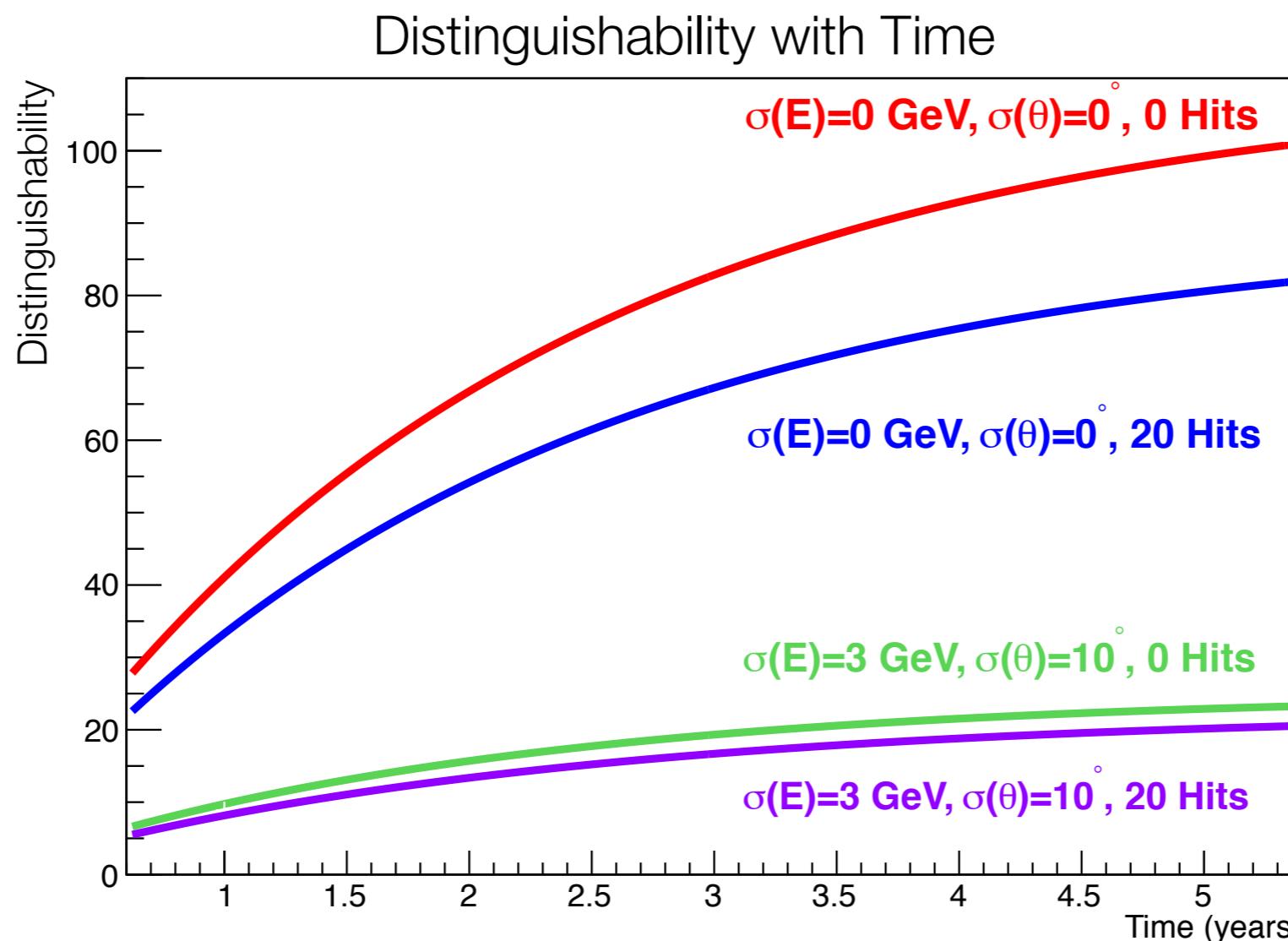
- Uncertainty in global Δm_{31}^2 creates some degeneracy in the distinguishability metric
 - Plots have perfect event ID and 100% selection efficiency, but include energy and angle smearing



Time to Distinguishing

- Physics/Detector
- Hierarchy
- Challenges

- The statistical power of PINGU makes systematics a critical factor sooner rather than later for hierarchy
 - PINGU specific - angular reco, energy reco, ice modeling...
 - Neutrino field at large - MC neutrino generators, cross-sections, atmospheric flux...



Leaving Feasible and Moving on to Ambitious

θ_{23} Maximal? Octant?

- Maximal Mixing
- Beam?

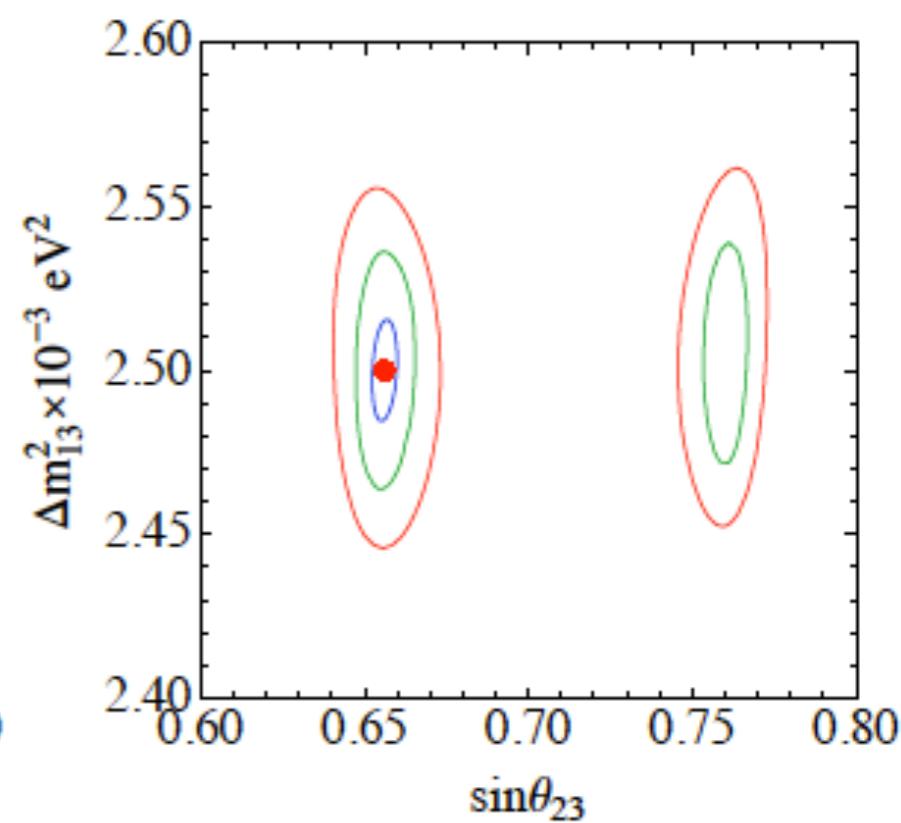
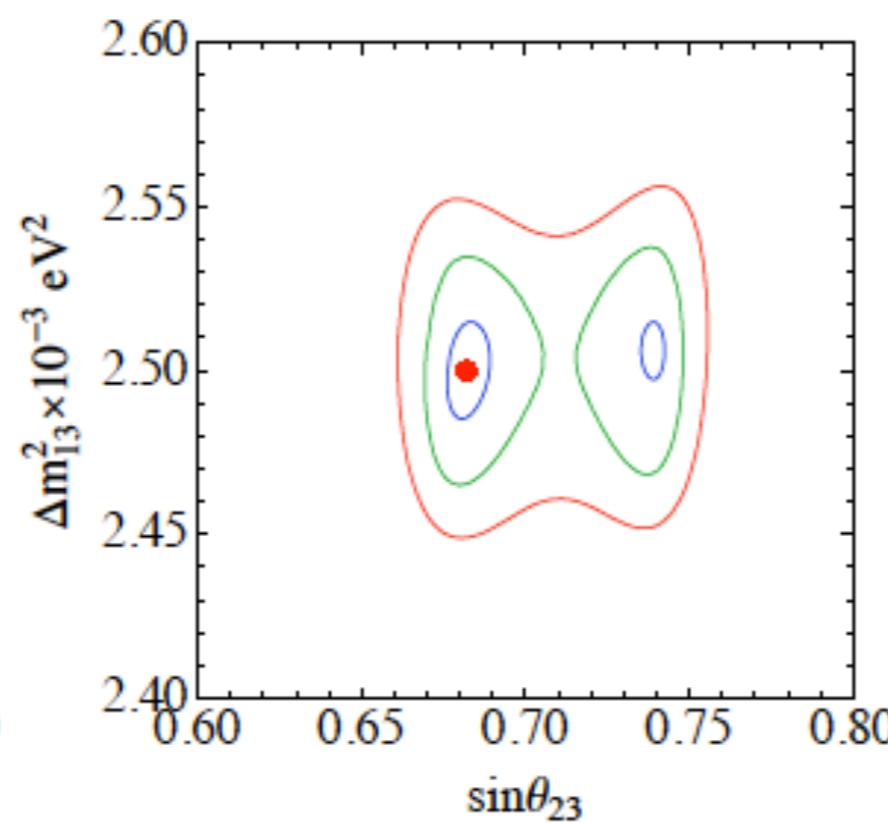
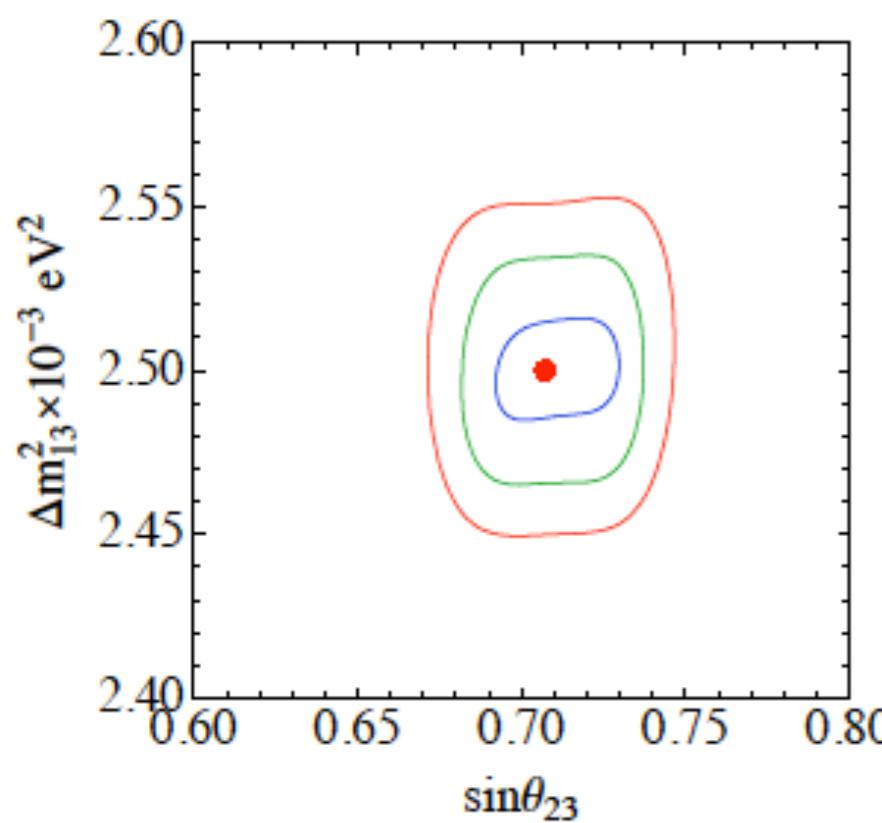
- Instead of fitting $\sin^2 2\theta_{23}$ fit $\sin \theta_{23}$
- Requires lots and lots of events
 - Plots below are 10 years of DeepCore exposure w/ ambitious reco assumptions
 - But, the requirements are similar to what is necessary for resolving the hierarchy in PINGU

Observable energies of 5 to 50 GeV
10 energy bins, 4 angular bins
vs.

1st energy bin, 1 angular bin +
9 energy bins, 4 angular bins

vs.

Exclude first 2 energy bins:
8 energy bins, 4 angular bins



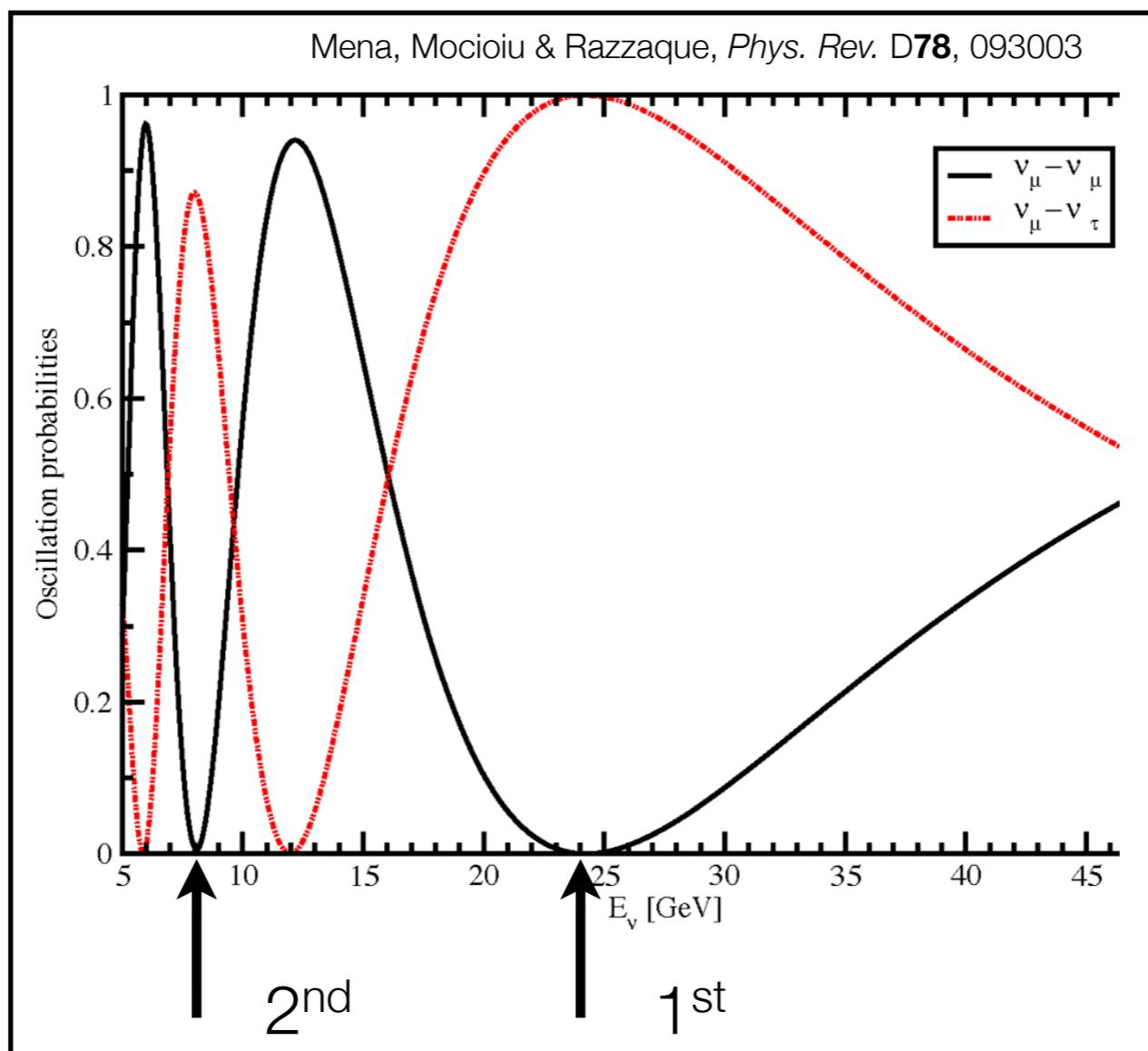
- Precision of atm. oscillation fit parameters will improve drastically the more minima/maxima can be resolved
- Trigger efficiency is much higher at lower energies for PINGU versus DeepCore
- PINGU/DeepCore covers all zeniths (baselines), while the oscillation minimum is at ~25 GeV @12700km, the minimum shifts to lower energies for decreasing baselines

$$\sin^2(2\theta_{13})=0.1$$

PINGU Maximal Mixing

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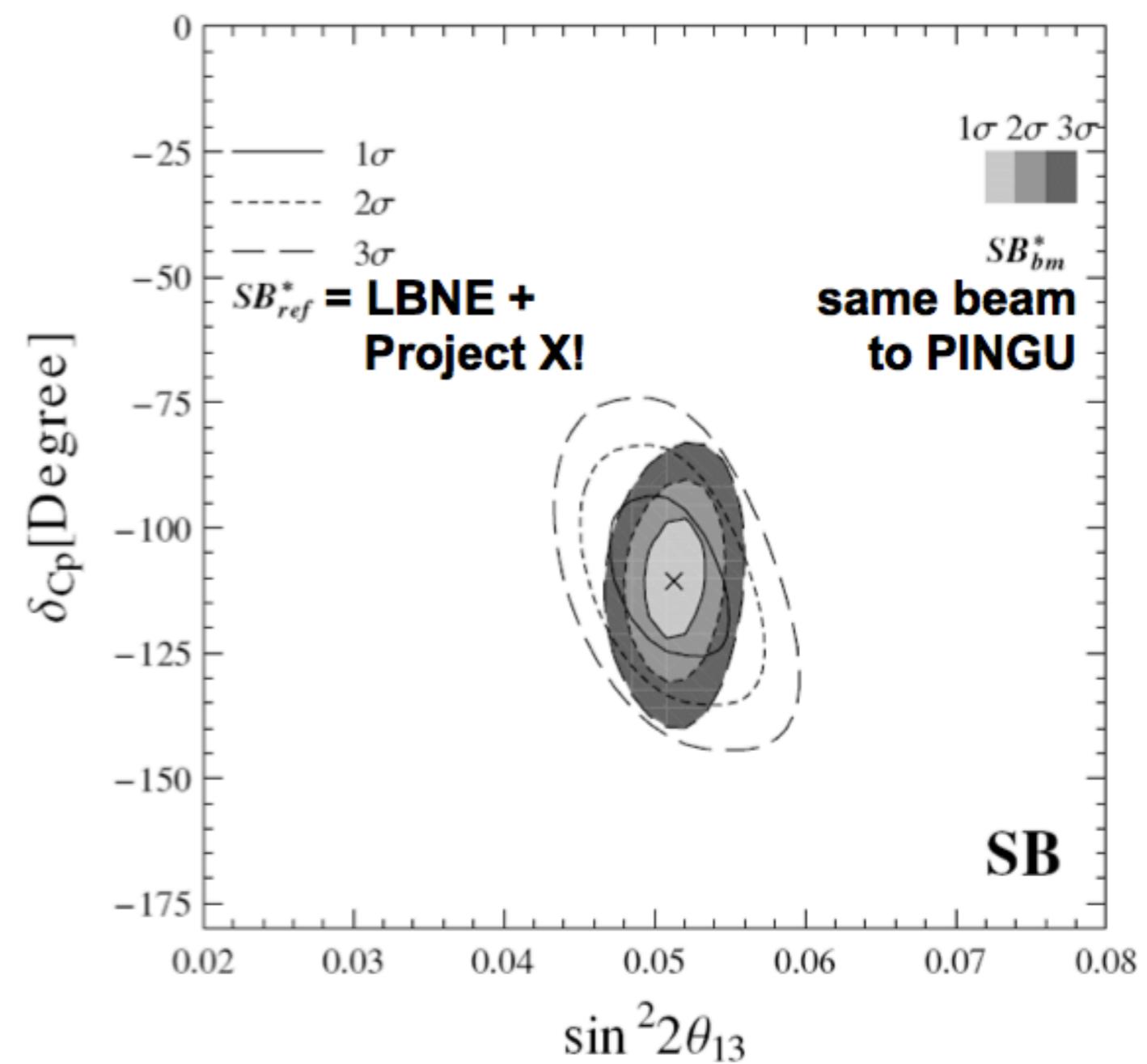
Leaving Ambitious
and Moving onto...

Leaving Ambitious and Moving onto...



Upgrade path towards δ_{CP} ?

- Measurement of δ_{CP} in principle possible, but challenging
- Requires:
 - Electromagnetic shower ID (here: 1% mis-ID)
 - Energy resolution (here: 20% $\times E$)
 - Maybe: volume upgrade (here: ~ factor two)
 - Project X
- Performance and optimization of PINGU, and possible upgrades (MICA, ...) require further study



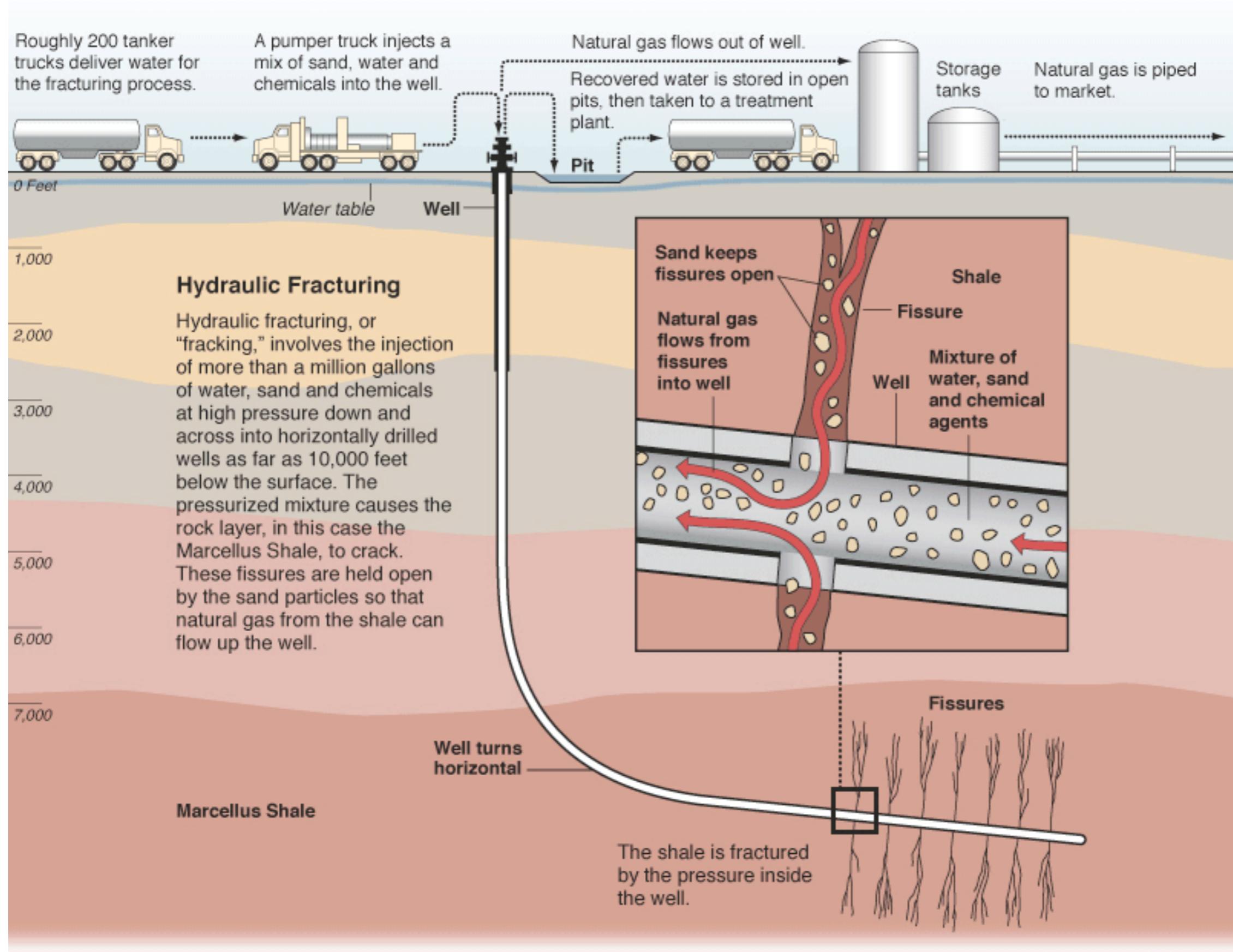
(Tang, Winter, JHEP 1202 (2012) 028)

15

- PINGU and MICA physics portfolio makes us of natural neutrino sources. Adding a beam will strengthen the diversity.
- Beam construction more of a headache than detector
 - 11620 baseline has a tilt angle of 65.8° from FNAL (similar for CERN)
 - Hydraulic fracturing drills may provide help

'Frakking'

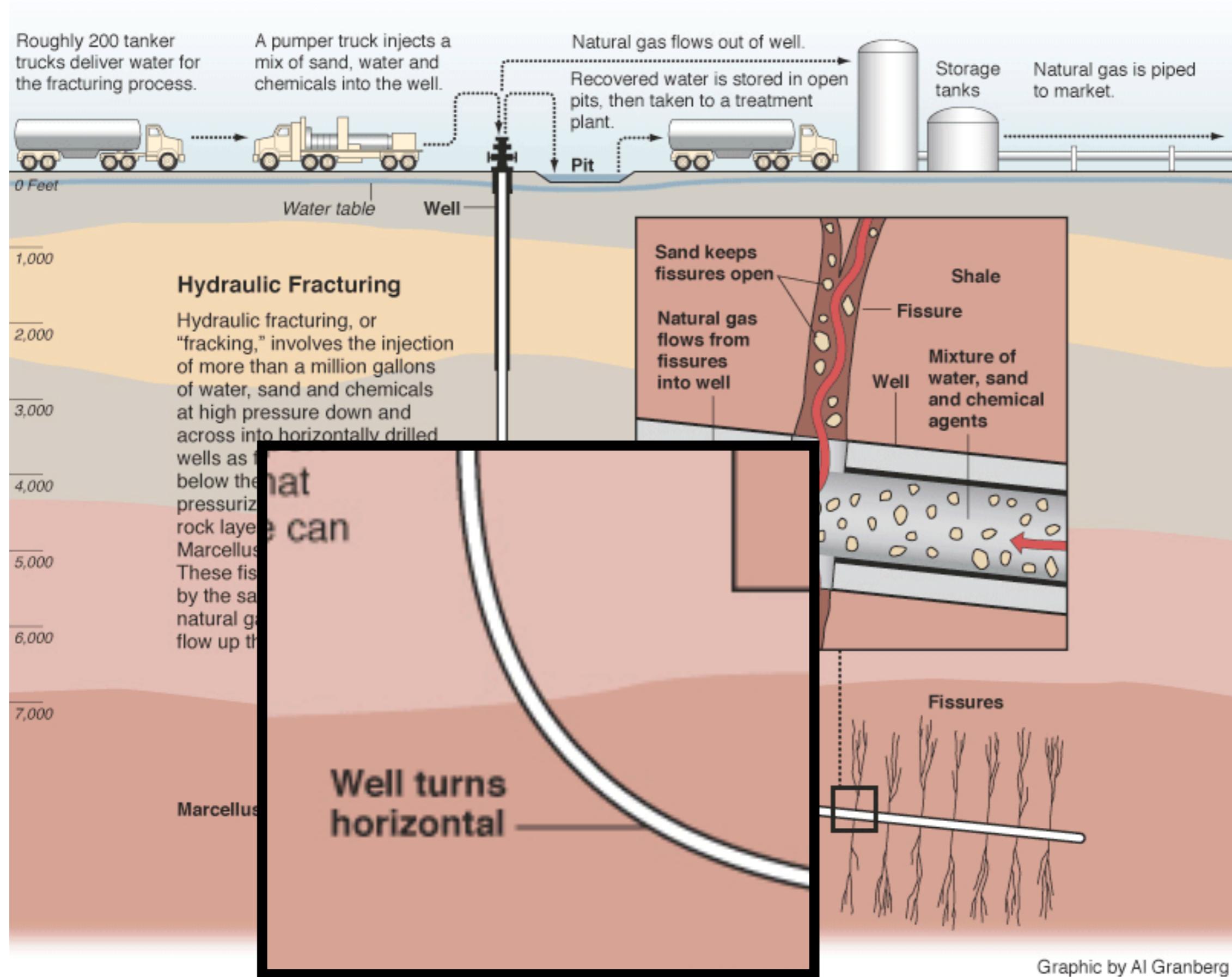
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Graphic by Al Granberg

'Frakking'

- Maximal Mixing
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 - Hydraulic fracturing drills may provide help
- Melting ice is cheaper than excavating rock
 - Going bigger underground gets non-linearly expensive for civil construction (in a bad way)
 - Whereas, the difficulty of in-ice deployment is going smaller

PINGU Advantages

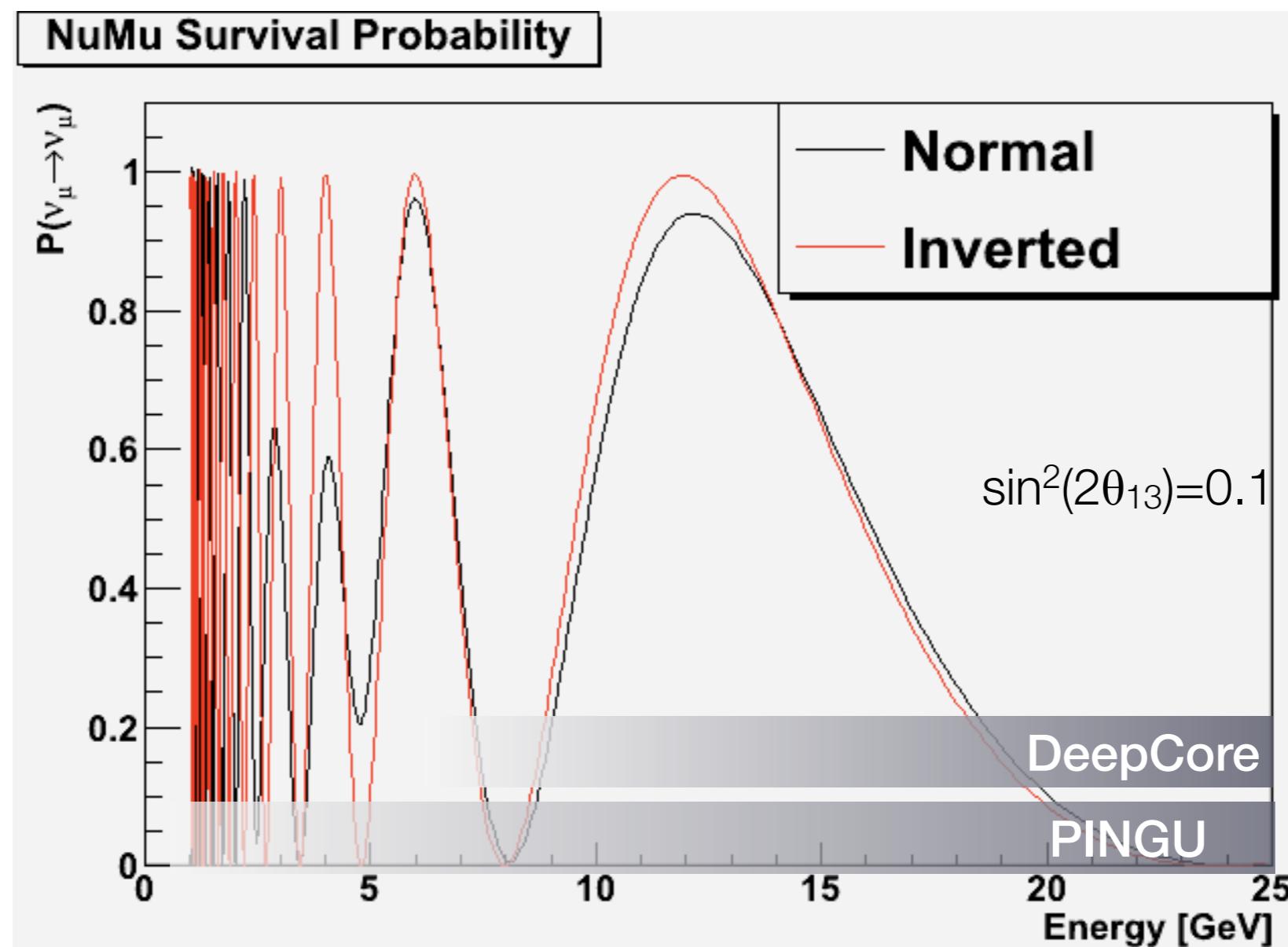
- Relatively quick, cost effective, huge and unique
 - 2-3 season deployment w/ additional ~1.5 year procurement/shipping
 - ~10M\$ start up and ~1.25M\$/string based on IceCube experience
 - Megaton size at GeV energies
 - Samples many angle, many baselines and crosses the earth core
- Enhance on-going DeepCore physics
 - muon disappearance
 - tau appearance
- Gains sensitivity to additional neutrino oscillation features
 - 2nd oscillation minima/maxima
 - Maximal θ_{23}
 - Neutrino hierarchy over all possible values of δ_{cp}
 - Maximal mixing and beam option

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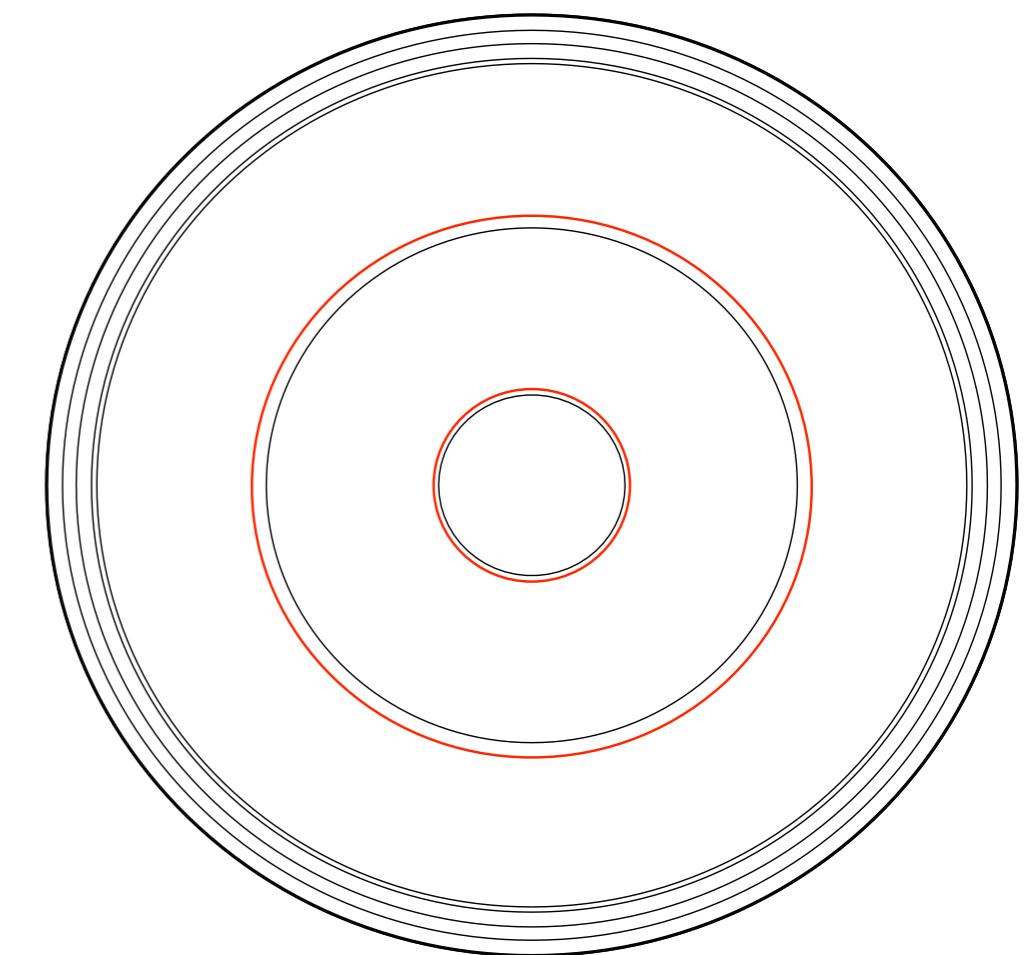
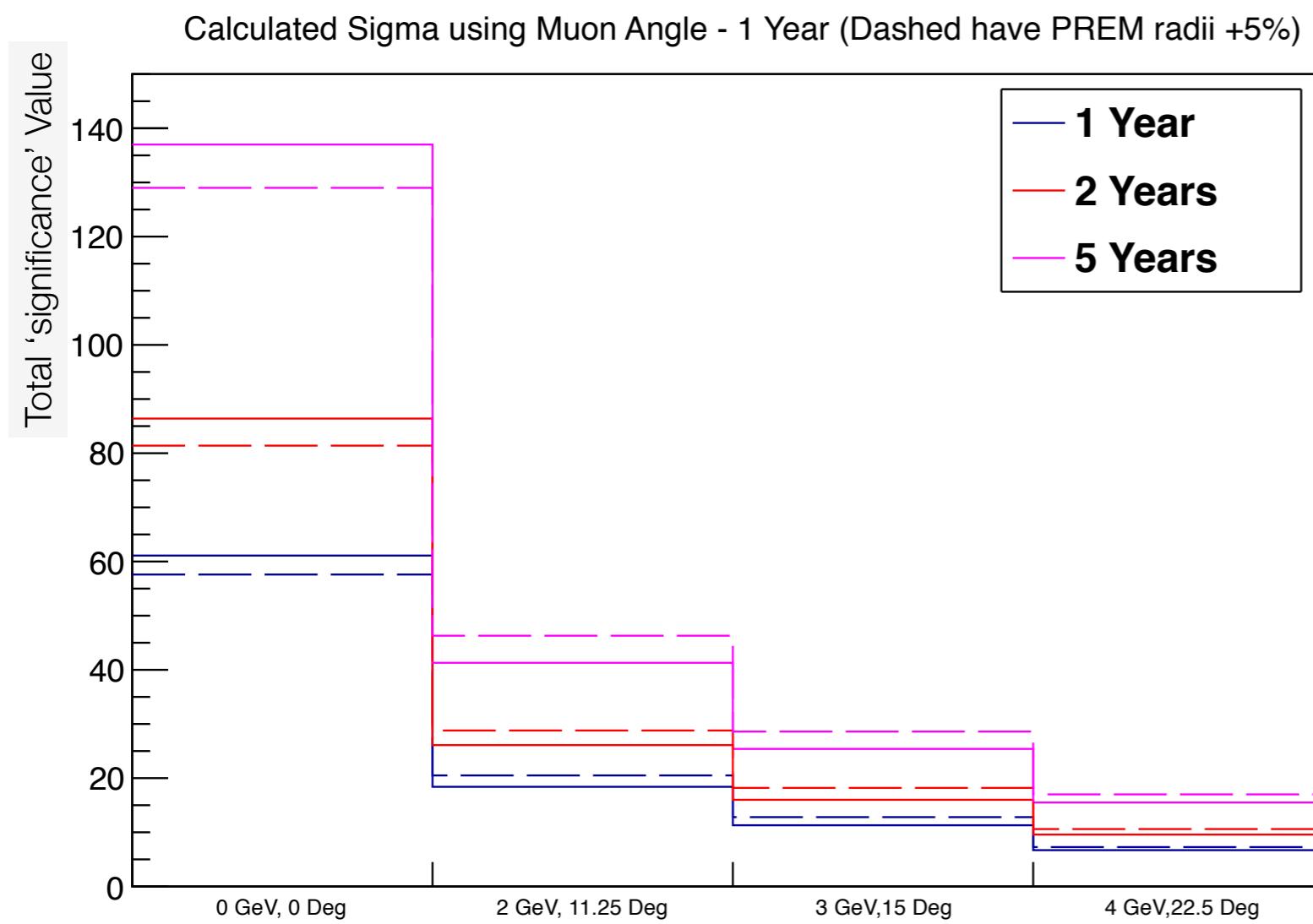


Backup

- Qualitative comparison of DeepCore to PINGU energy region



- Change the earth layers within the PREM model radii +5%
- Distinguishability is largely unaffected by PREM model uncertainty



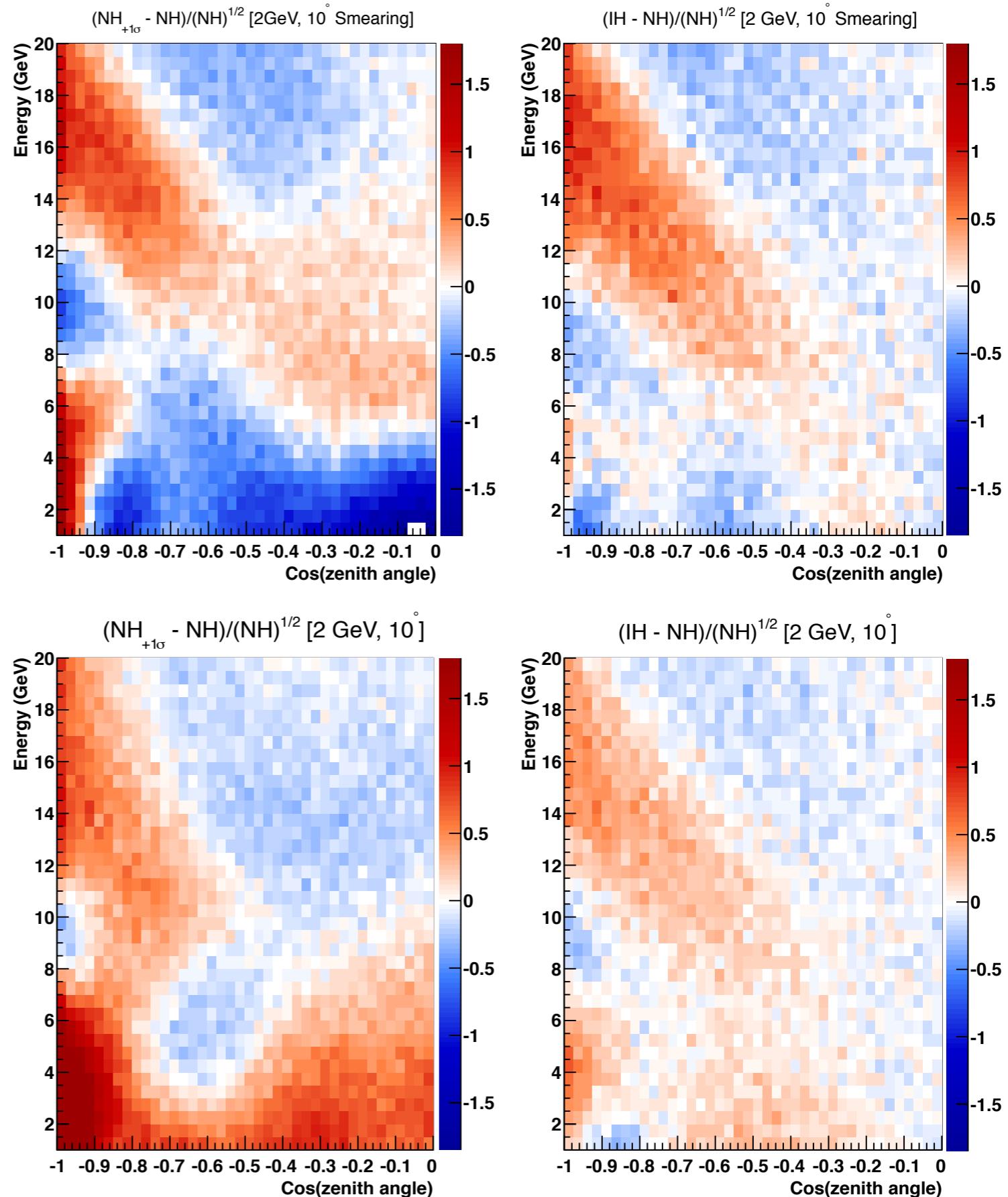
Physics/Detector Challenges

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- Including a $20 \geq$ hit cut (bottom row) still shows some degeneracy between hierarchy ordering and uncertainty in Δm_{31}^2

- Event selection and analysis techniques can be used to reduce the degeneracy

- Optimize for inelastic event selection
- Likelihood analysis instead of a chi-squared like test



Calibration

- Physics/Detector
- Hierarchy
- Challenges

- In-situ cameras show that refrozen hole ice has a central column of ‘cloudy’ ice
- Degassing filters can be included on the hot water drill to reduce the addition of bubbles

