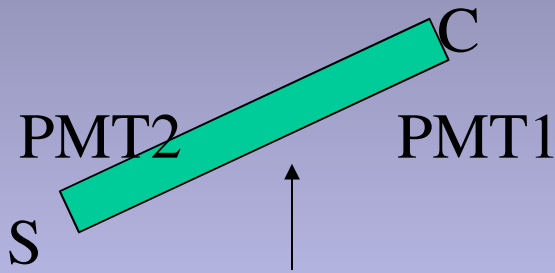


BGO Pions

Cosenza Group

To Measure Cerenkov light percentage



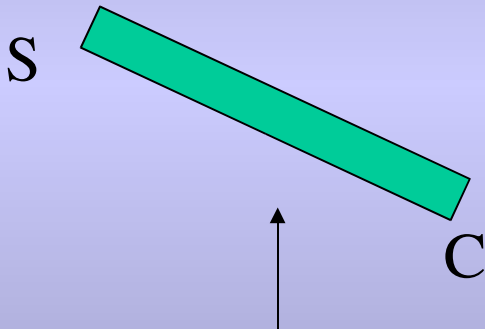
$$\theta > 0$$

$$\text{Counts } C1 = \eta \text{ cere1} + \eta \alpha \text{ sci1}$$

$$\text{Counts } S1 = \gamma \text{ sci1}$$

filter efficiency $\rightarrow \alpha$

geometrical factors $\rightarrow \eta, \gamma$



$$\theta < 0$$

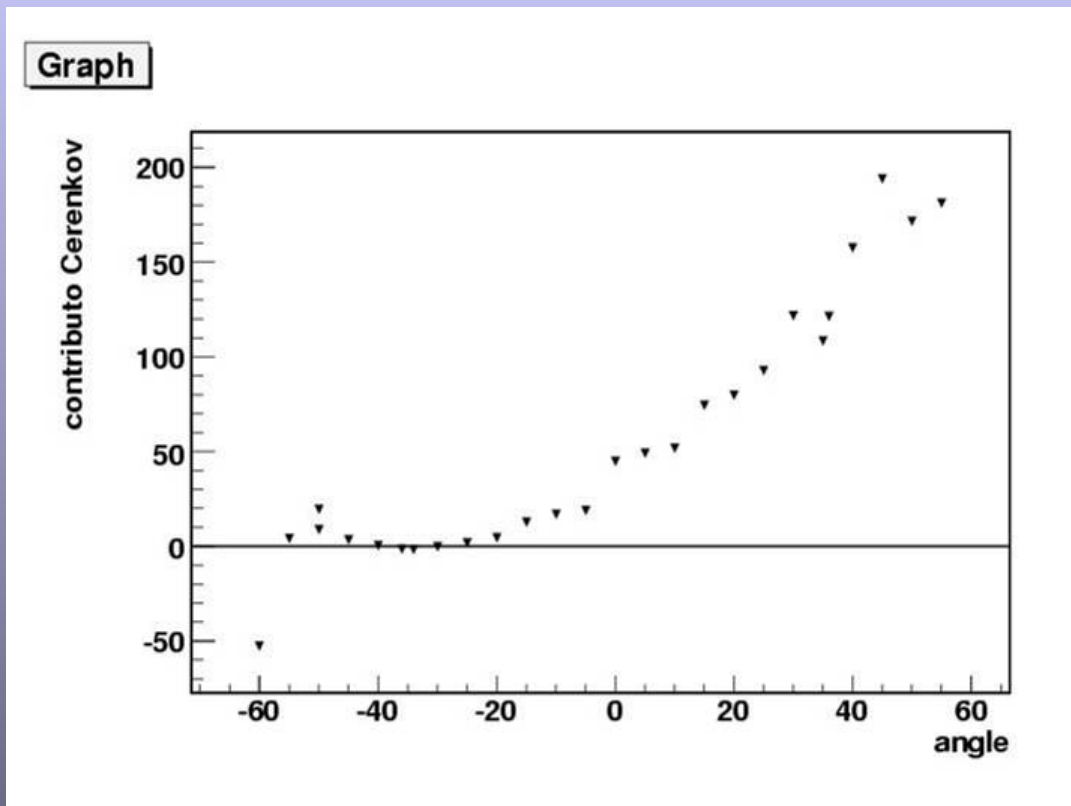
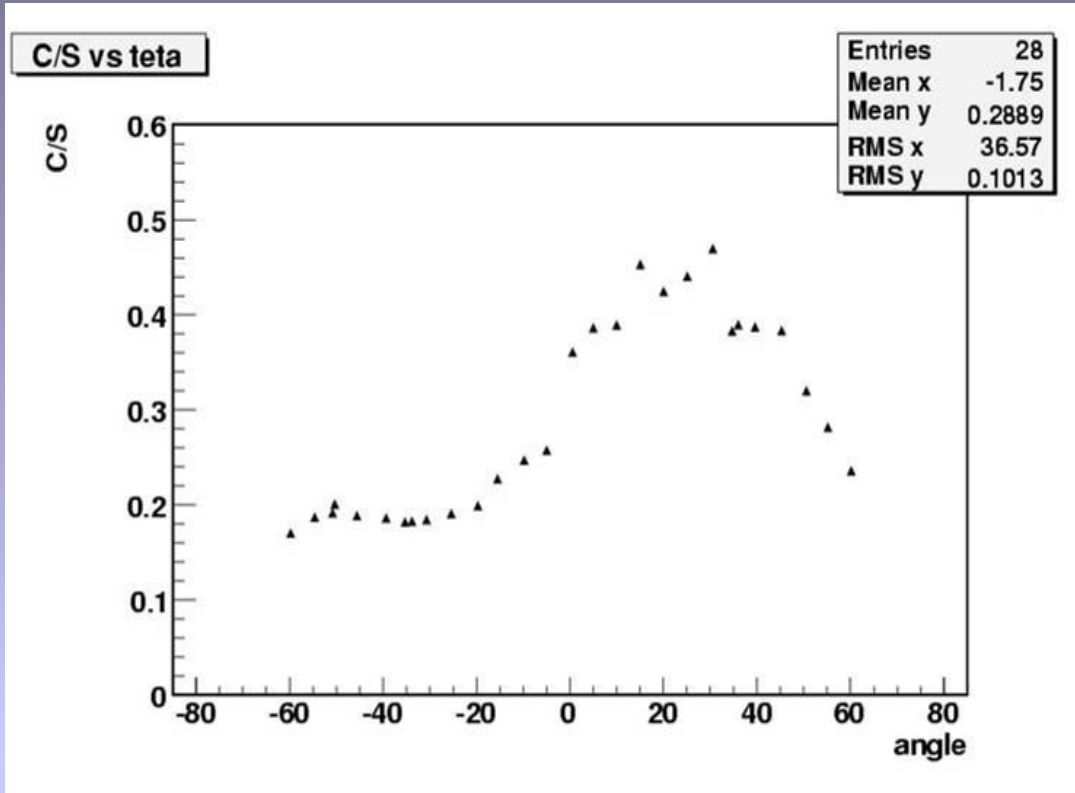
$$\text{Counts } C2 = \eta \alpha \text{ sci2}$$

$$\text{Counts } S2 = \gamma \text{ sci2}$$

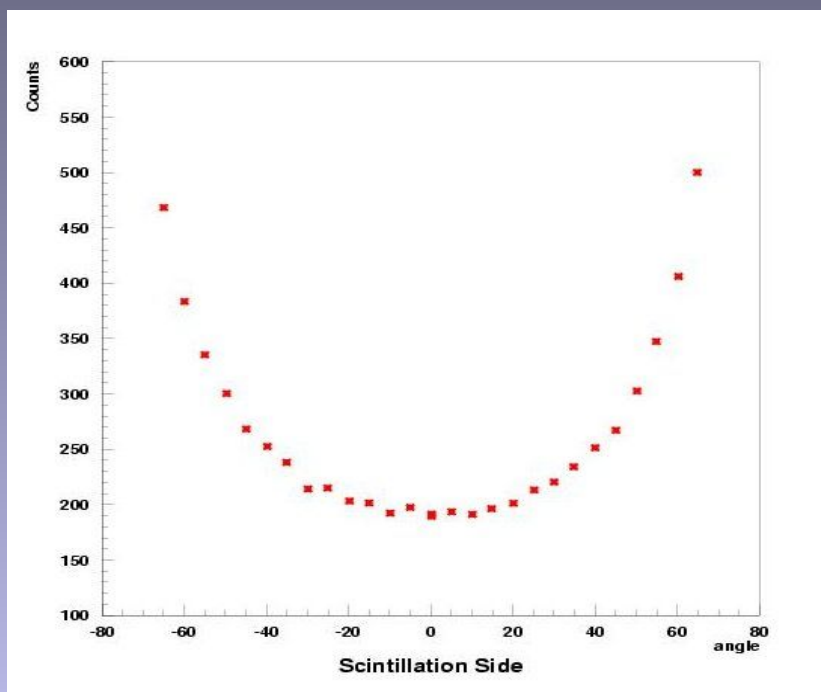
$$\text{Counts } C2 / \text{Counts } S2 = \eta \alpha / \gamma = K$$

$$\eta \text{ cere1} = \text{Counts } C1 - K \text{ Counts } S1$$

Runs 1681 – 1722 e^- results in agree with Rome

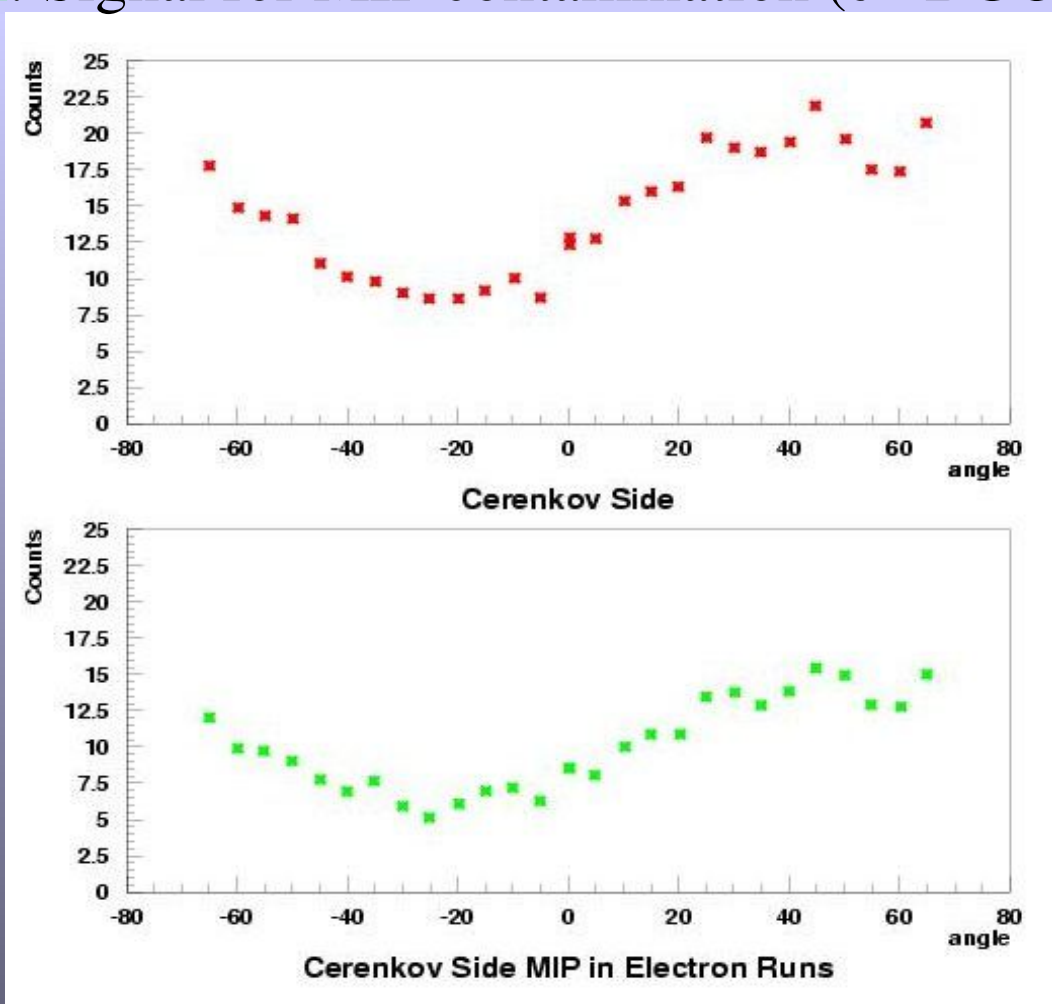


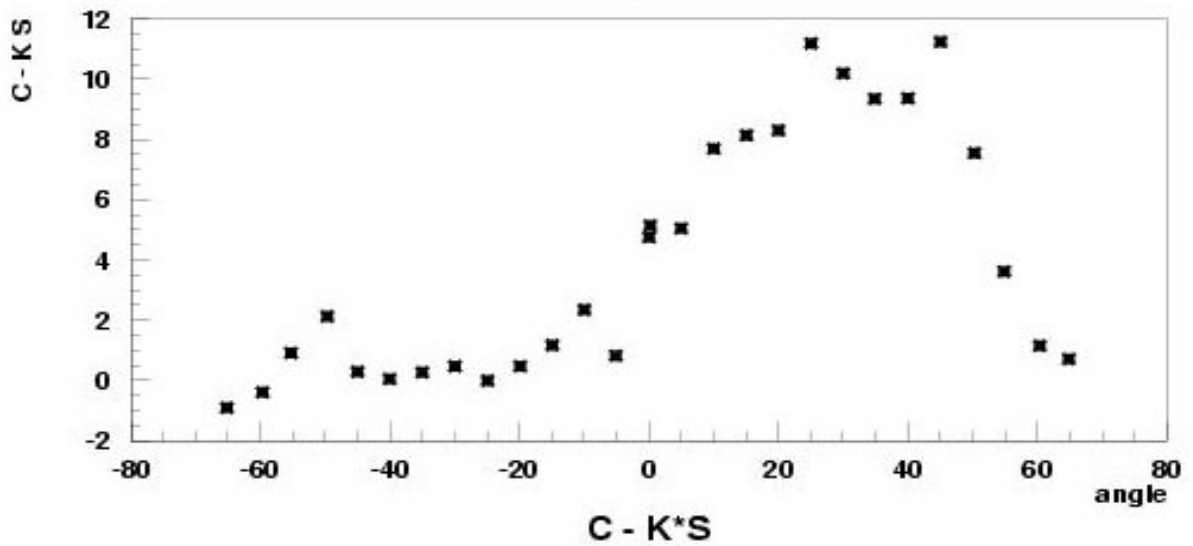
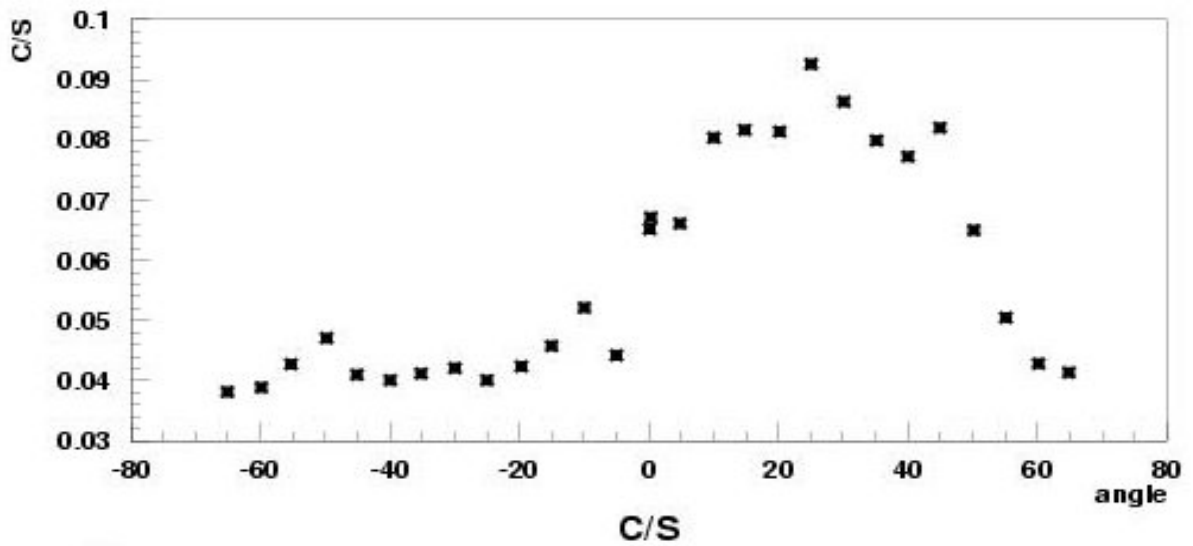
Same analysis for 200 GeV Pions; runs 1613-1647



Red: Signal for Cerenkov side

Green: Signal for MIP contamination (e^- -BGO)





Cerenkov contribution shape in agree with simulation & Cecilia analysis.