



**ESI2013: 3rd EIROforum School on Instrumentation**

# **An Introduction to Synchrotron Radiation**

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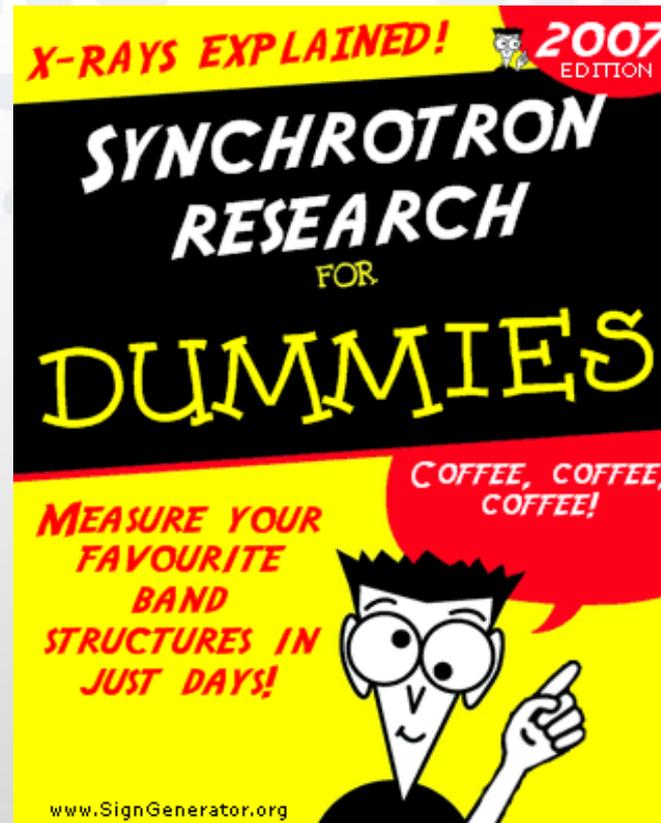
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## Part I

Introduction  
to  
synchrotron radiation  
(SR)



## Part II

SR Research

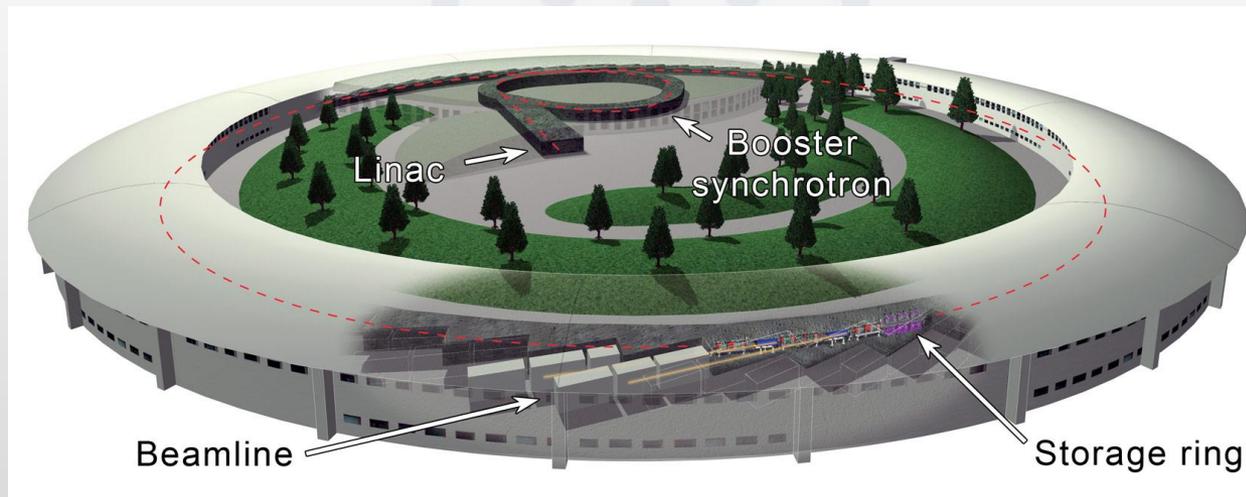
*Mostly Illustrated by examples from the ESRF*

## PART I

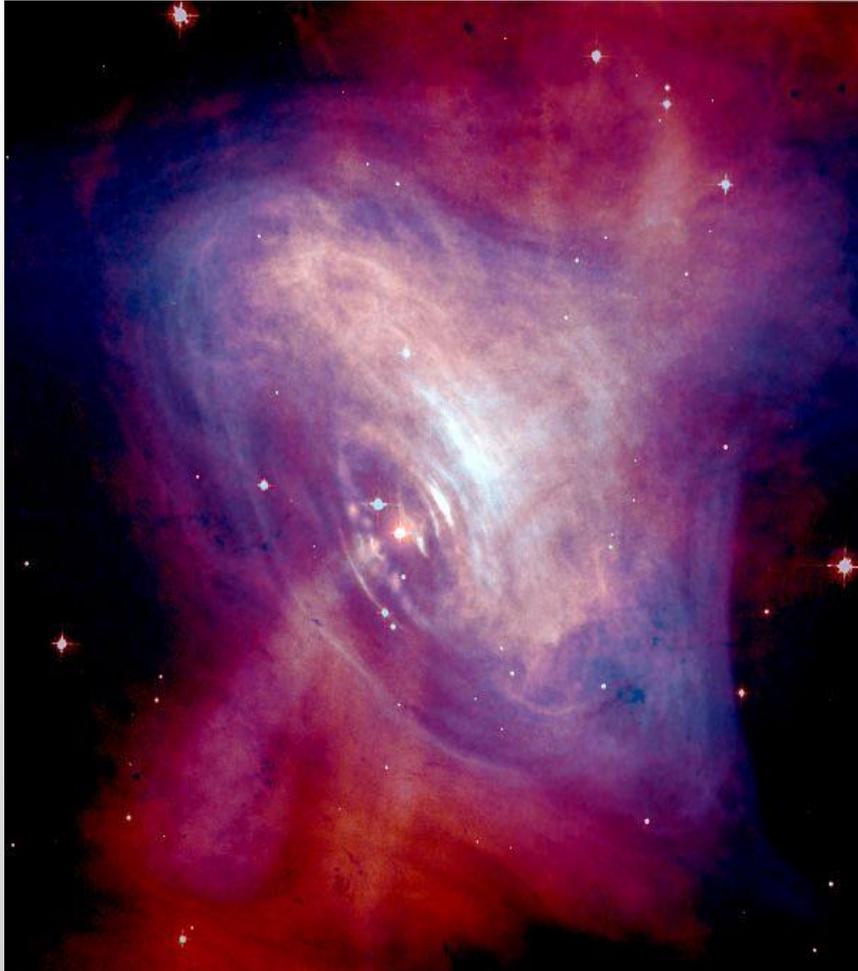
### Introduction to synchrotron radiation

1. Origin and physics
2. Historical evolution
3. Synchrotron radiation facilities
4. Three forms of synchrotron radiation
5. Sources and their operation

- An extremely intense source of X-rays (also IR & EUV)
  - Typically produced by highly energetic electrons (or positrons) moving in a large circle in the synchrotron (or more usually **Storage Ring**).



- Highly collimated X-rays are emitted tangentially to the storage ring and simultaneously serve multiple beamlines for scientific applications spanning physical, chemical and life sciences



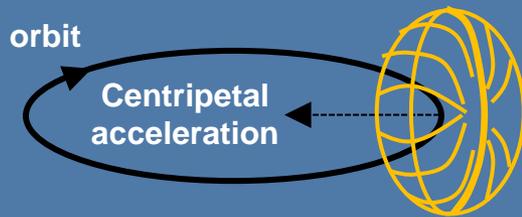
The Crab Nebula – a natural SR source  
CHANDRA (2001)

L. Shklovsky (1953)

- Synchrotron radiation is electromagnetic energy emitted by charged particles (e.g., electrons and ions) that are moving at speeds close to that of light when their paths are altered, e.g. by a magnetic field.
- It is thus termed because particles moving at such speeds in particle accelerators known as a synchrotrons produce electromagnetic radiation of this type.

## Classical case

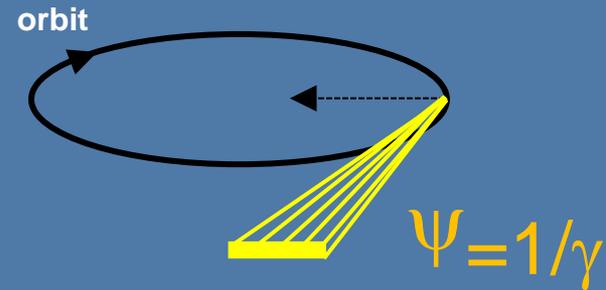
$$v \ll c$$



Isotropic emission

## Relativistic case

$$v \sim c$$



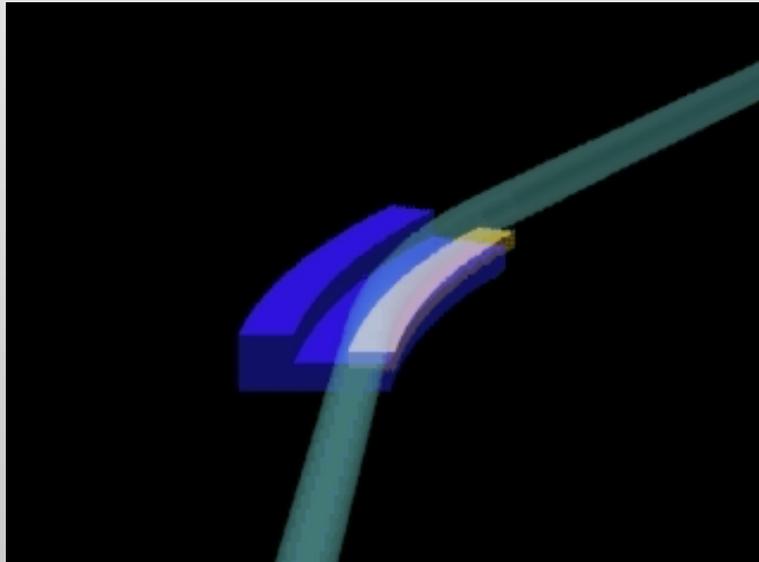
Emission concentrated within a narrow forward cone

$$\gamma = \frac{E_e}{mc^2}$$

$$E_e = 6 \text{ GeV (ESRF)}$$

$$mc^2 = 511 \text{ keV}$$

$$\Psi = 10^{-4} \text{ rad} = \text{a few } 10^{-3} \text{ deg}$$





1947

First observation of synchrotron radiation at General Electric (USA).

Considered first as a nuisance by particle physicists, synchrotron radiation was recognised in the 70s as having exceptional properties to explore matter.

## EVOLUTION

1930

First particle accelerators

1947

First observation of synchrotron radiation

More and more energetic particles,

Bigger and bigger machines

1980

Construction of the first dedicated machines

Particle physics

Synchrotron radiation

## Timeline

1994  
ESRF

1981  
SRS

1968  
Tantalus

1961  
Parasitic use

1950'  
Characterization

1947  
Observation

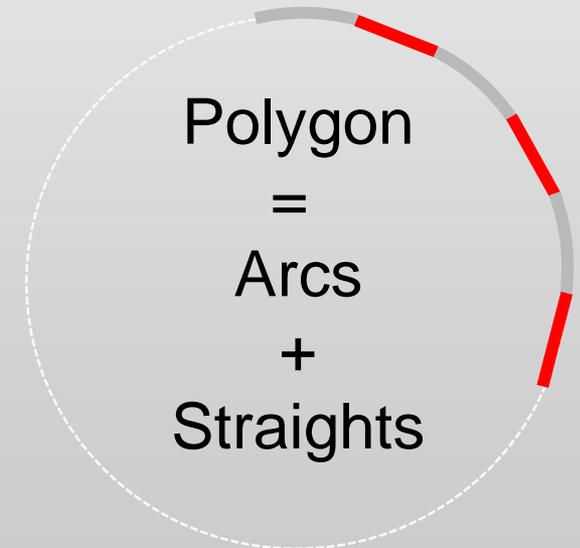
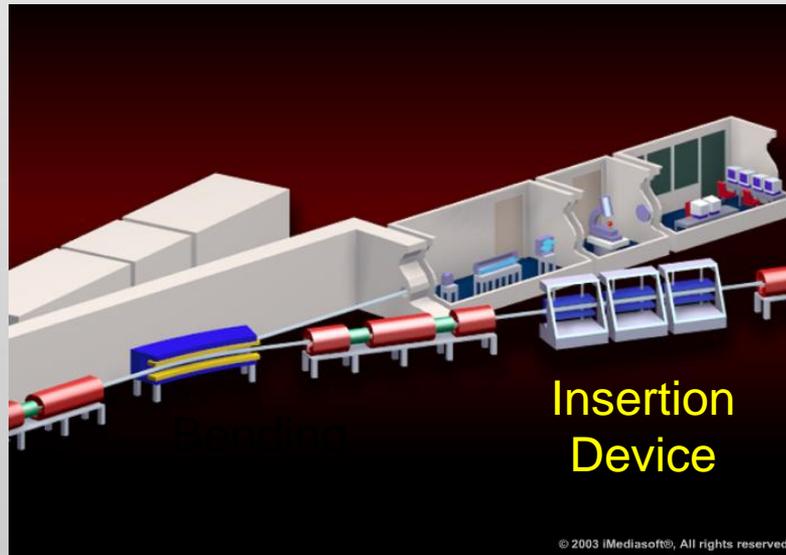
1897-1946  
Predictions



## ESRF

- Storage ring: 844 m circumf.
- 6 GeV electrons
- ~200 mA fill current
- Typical beam lifetime ~ 50 h
- Reinjection twice/day
- 24 hour operation 6 days/week

Fully optimized to produce bright X-ray beams



## Major X-Ray sources in the world



ESRF was the world's first 3<sup>rd</sup> generation hard X-ray source

Other hard X-ray sources: APS (USA) - SPring-8 (Japan) – Petra-III (D)

New national sources in Europe: Swiss Light Source (CH), Soleil (F) in 2006, Diamond (UK) in 2007, Petra-III (D) in 2009, ALBA (E) in 2010, Max IV (Sweden), Poland, Russia.....

New plans also in Brasil, China, India, .....

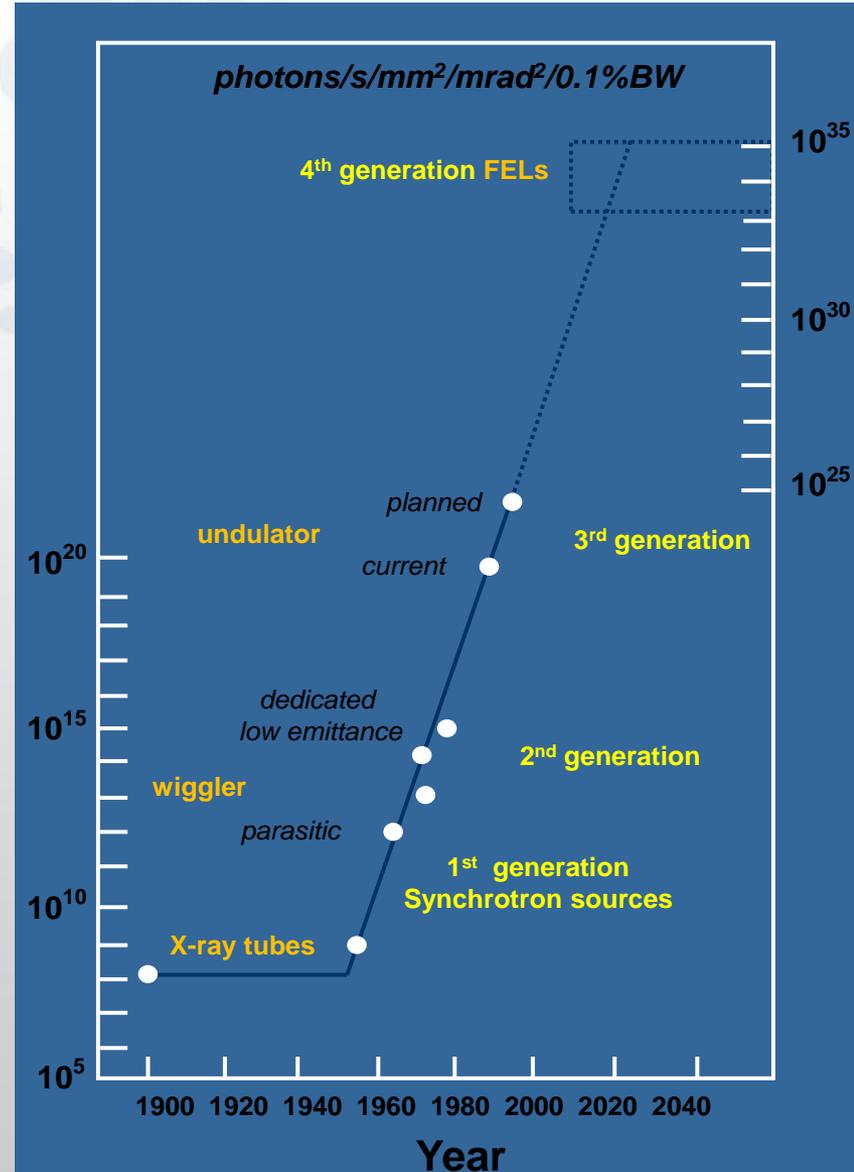
Brilliance or *Brightness* (flux density in phase space) is an invariant quantity in statistical mechanics, so that no optical technique can improve it.

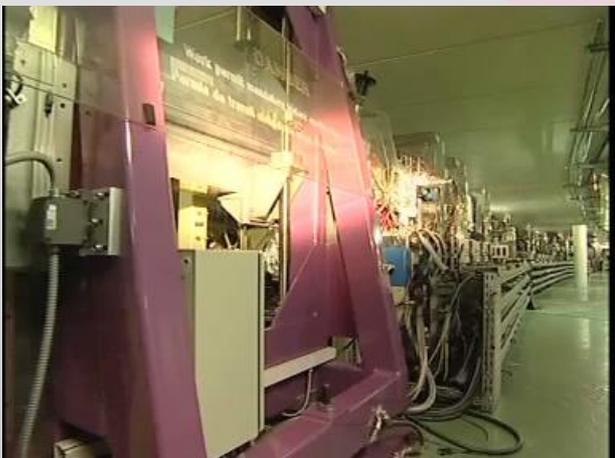
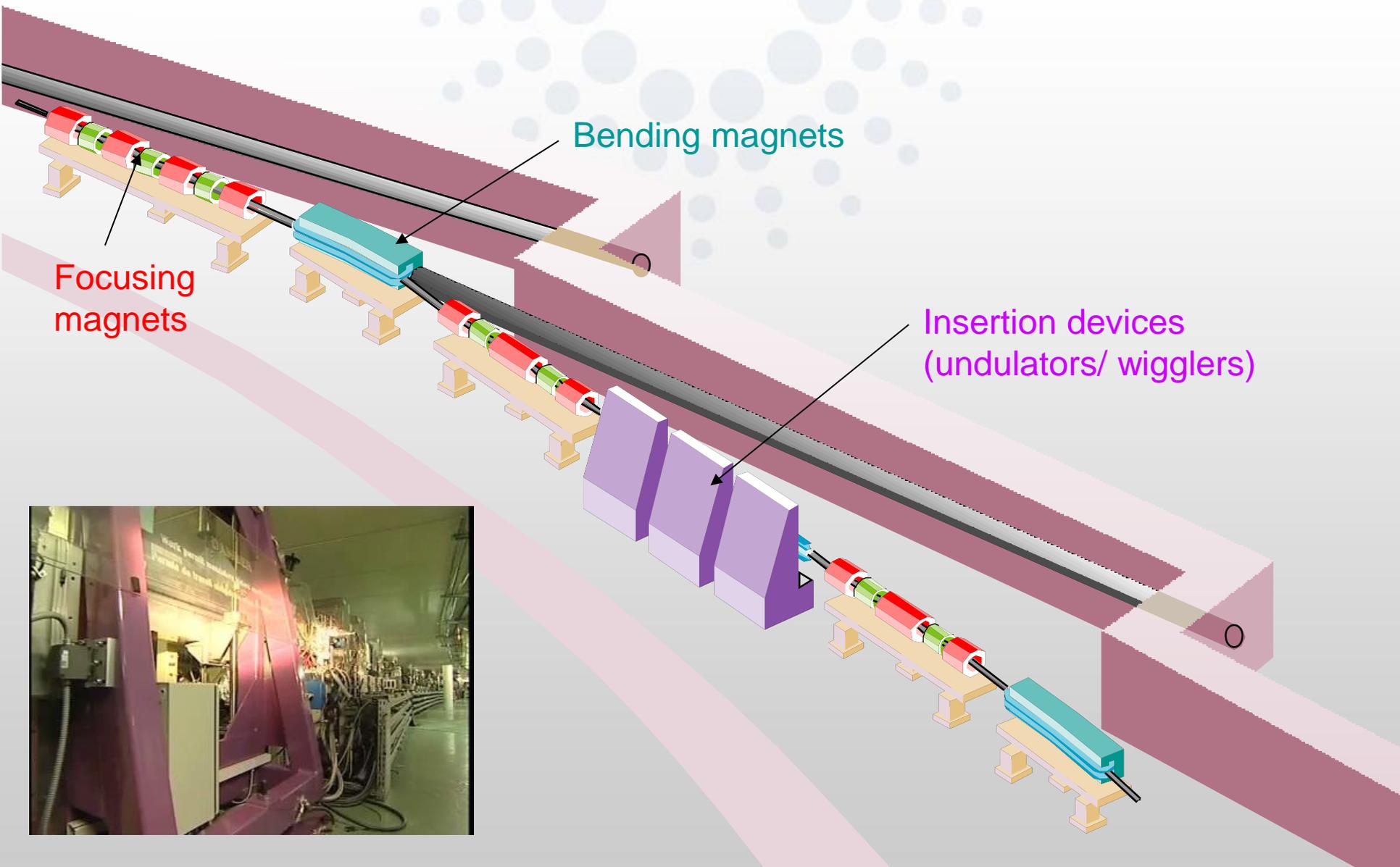
$$\text{Brightness} = \frac{\text{photon flux}}{(\Delta A) (\Delta \Omega)}$$

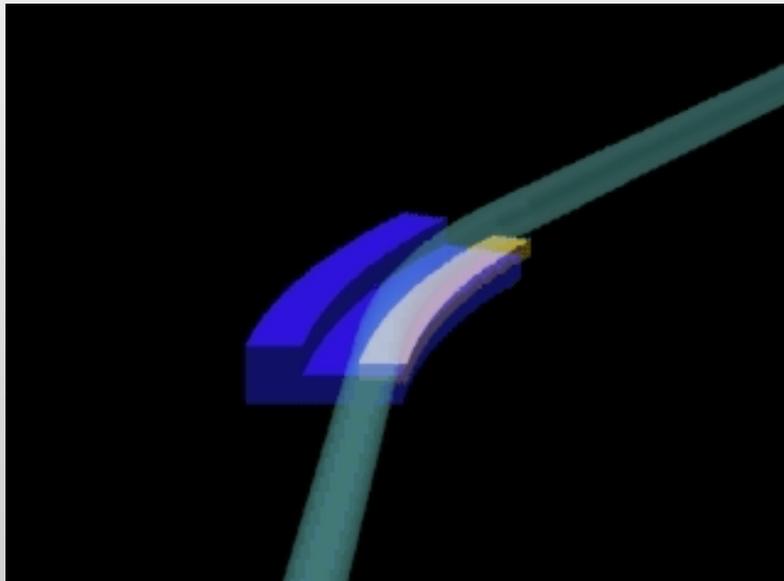
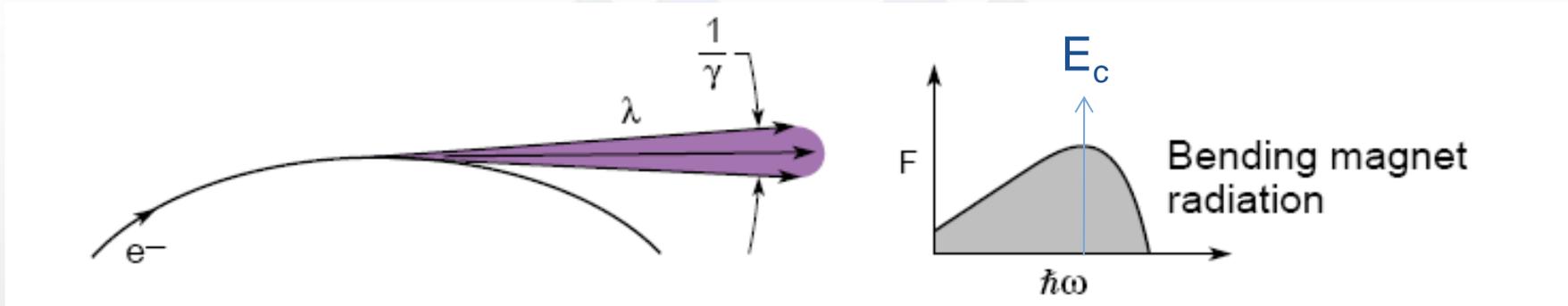
$$\text{Spectral Brightness} = \frac{\text{photon flux}}{(\Delta A) (\Delta \Omega) (\Delta \lambda / \lambda)}$$

[Photons/sec]

[mm]<sup>2</sup> [mrad]<sup>2</sup> [0.1% bandwidth]



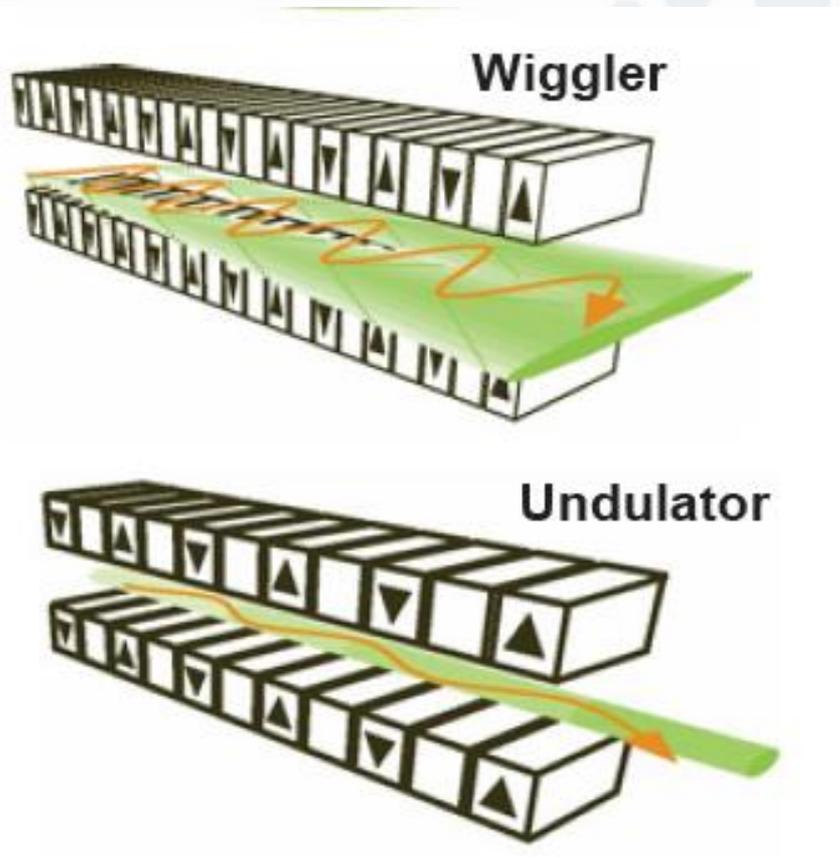




$$\gamma = \frac{E_e}{mc^2} = 1957 E_e (\text{GeV})$$

$$E_c = \frac{3e\hbar B \gamma^2}{2m}$$

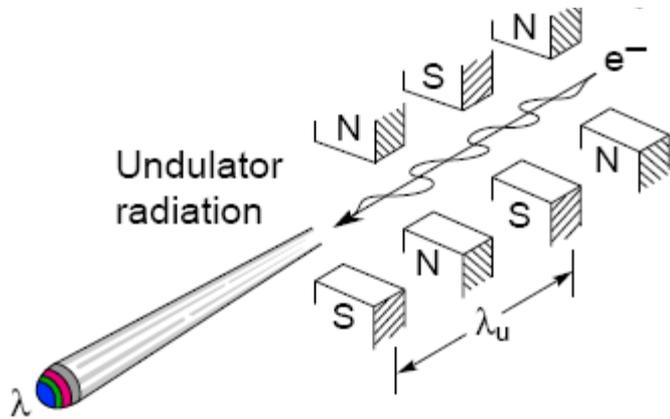
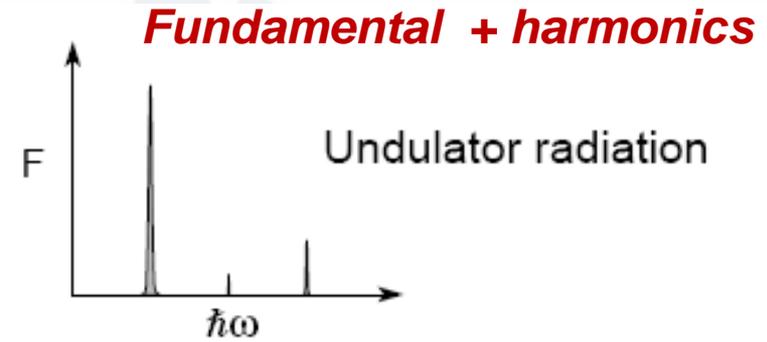
$$E_c (\text{keV}) = 0.665 E_e^2 (\text{GeV}) B (\text{T})$$



- A Periodic magnetic field causes local oscillations of electron trajectory. The influence of the field on the electron motion is characterized by *the deflection parameter  $K$*

$$K=0.934 \lambda_u [\text{cm}] B_0 [\text{T}]$$

- The maximum angular deflection of the orbit is  $K/\gamma$
- $K \gg 1$  Wiggler regime  
radiation from different parts of the electron trajectory adds incoherently  
→ *interference effects are less important*
- $K < 1$  Undulator regime  
radiation from different periods interferes coherently  
→ *strong interference phenomena*



On-axis

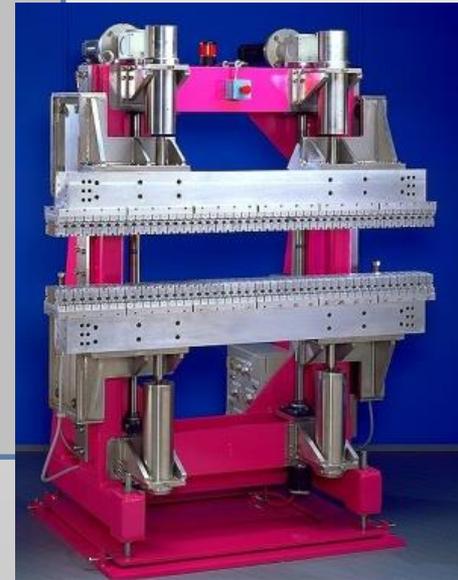
$$\lambda_1 = \frac{\lambda_u}{2\gamma^2} \left(1 + \frac{K^2}{2}\right)$$

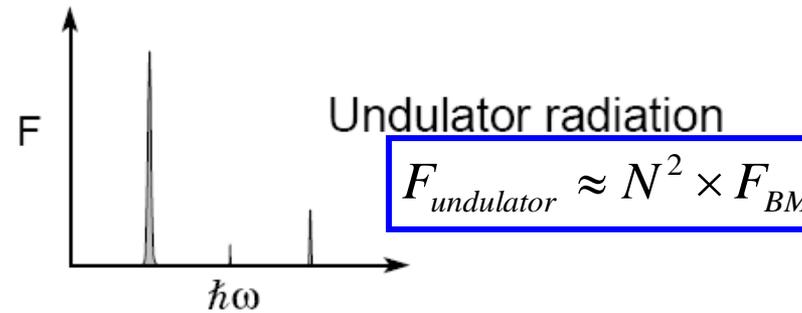
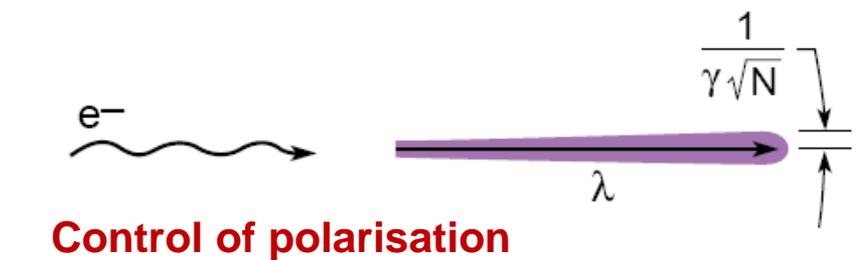
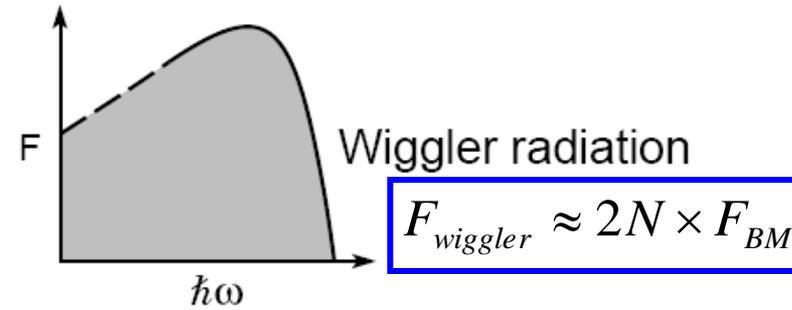
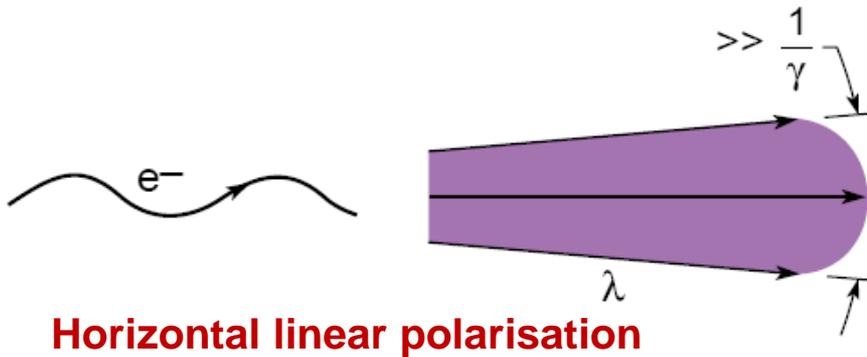
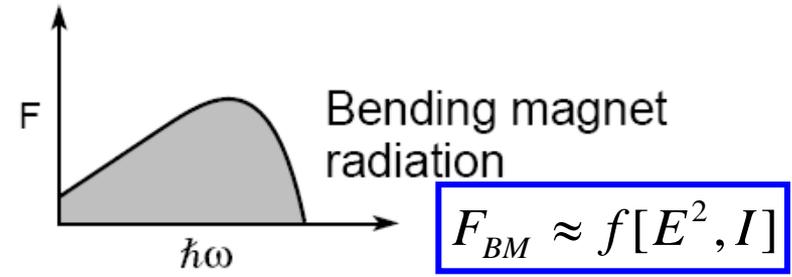
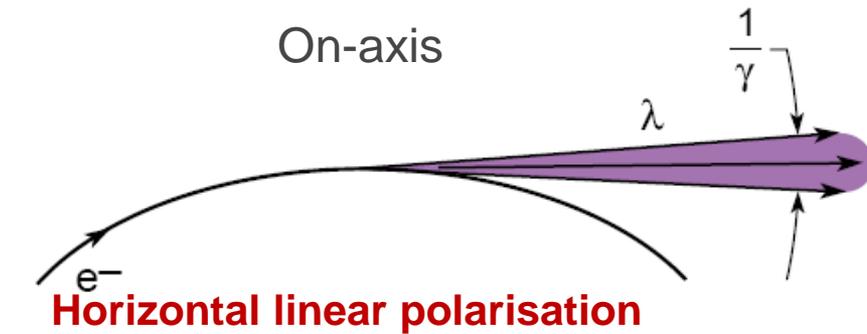
$$E(\text{keV}) = \frac{0.949 E_e^2 (\text{GeV})}{\lambda_u (\text{cm}) \left(1 + \frac{K^2}{2}\right)}$$

$$\frac{\Delta\lambda}{\lambda} = \frac{1}{N}$$

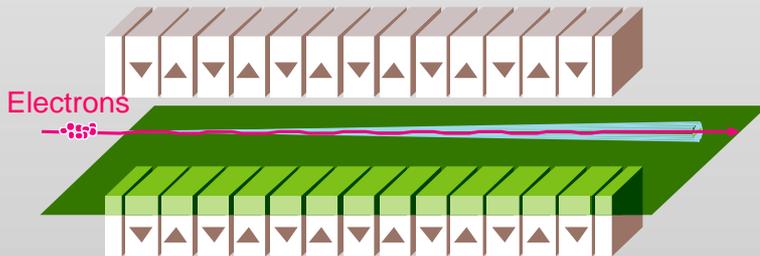
To tweak (or to scan) energy:

- $E = f[K]$
- $K = f[B(z)]$
- $B(z) = f[\text{gap}]$

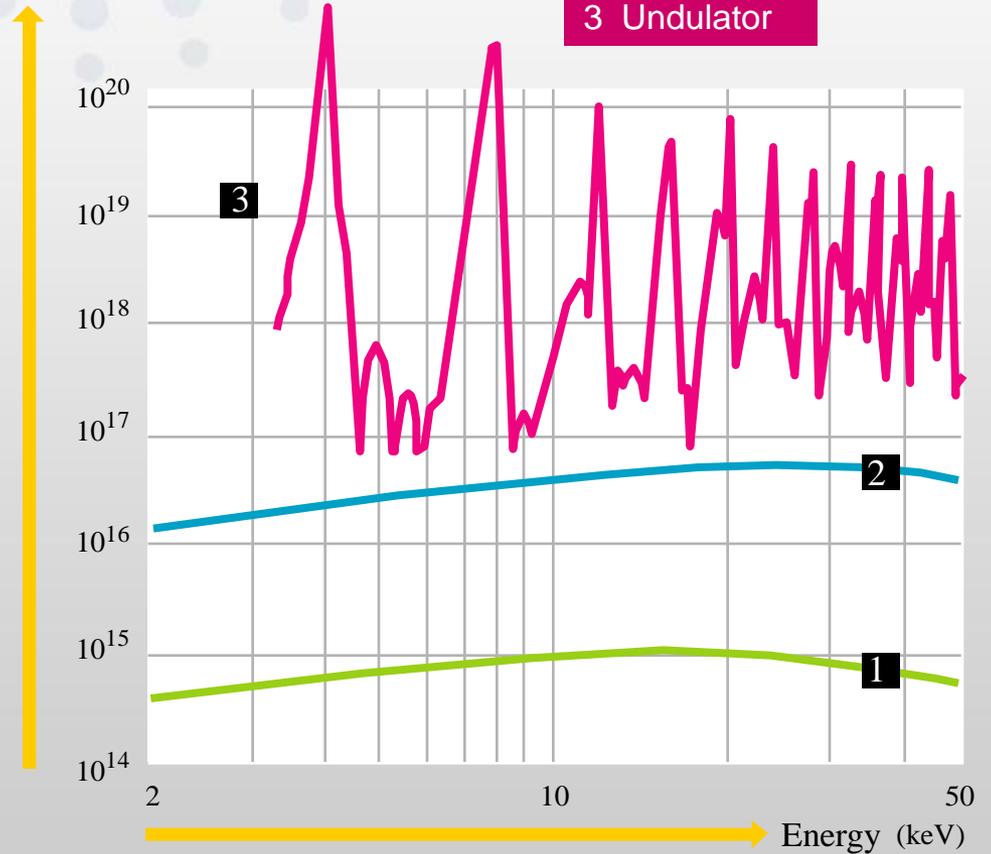




# Insertion devices



Brilliance  
(photons/s/mm<sup>2</sup>/mrad<sup>2</sup>/0.1% BW)

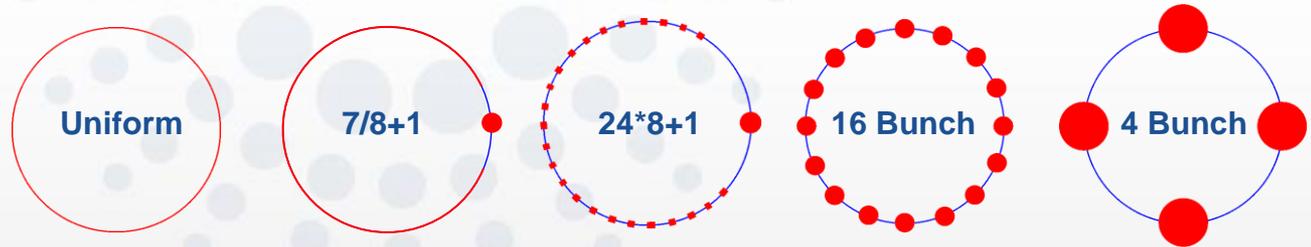


1 Bending magnet

2 Wiggler

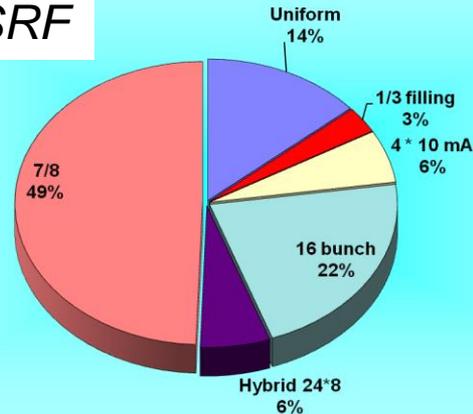
3 Undulator

## Filling patterns



Number of bunches	992	870	24x8+1	16	4
Maximum current [mA]	200	200	200	90	40
Rms bunch length [ps]	20	20	25	48	55

## Typical filling modes at the ESRF



- **Brightness**
  - $10^{22}$  ph/sec/mrad<sup>2</sup>/mm<sup>2</sup>/0.1% b.w. ( $10^{11}$  higher than conventional sources)
- **Small source and highly collimated beam**
  - $\sim 10\mu\text{m}$  and  $10\mu\text{rad}$
- **Broad emission spectrum: wavelength tunability**
  - $0.1\text{eV} < E < 100\text{keV}$  or  $1.2\mu\text{m} < \lambda < 0.01\text{\AA}$
- **Polarized radiation**
  - 100% linear or circular or elliptical
- **Pulsed radiation**
  - 50 ps pulses every ns
- **Power**
  - several kW total power, several  $100\text{ W/mm}^2$  power density.
- **High degree of coherence**



27 May 2013

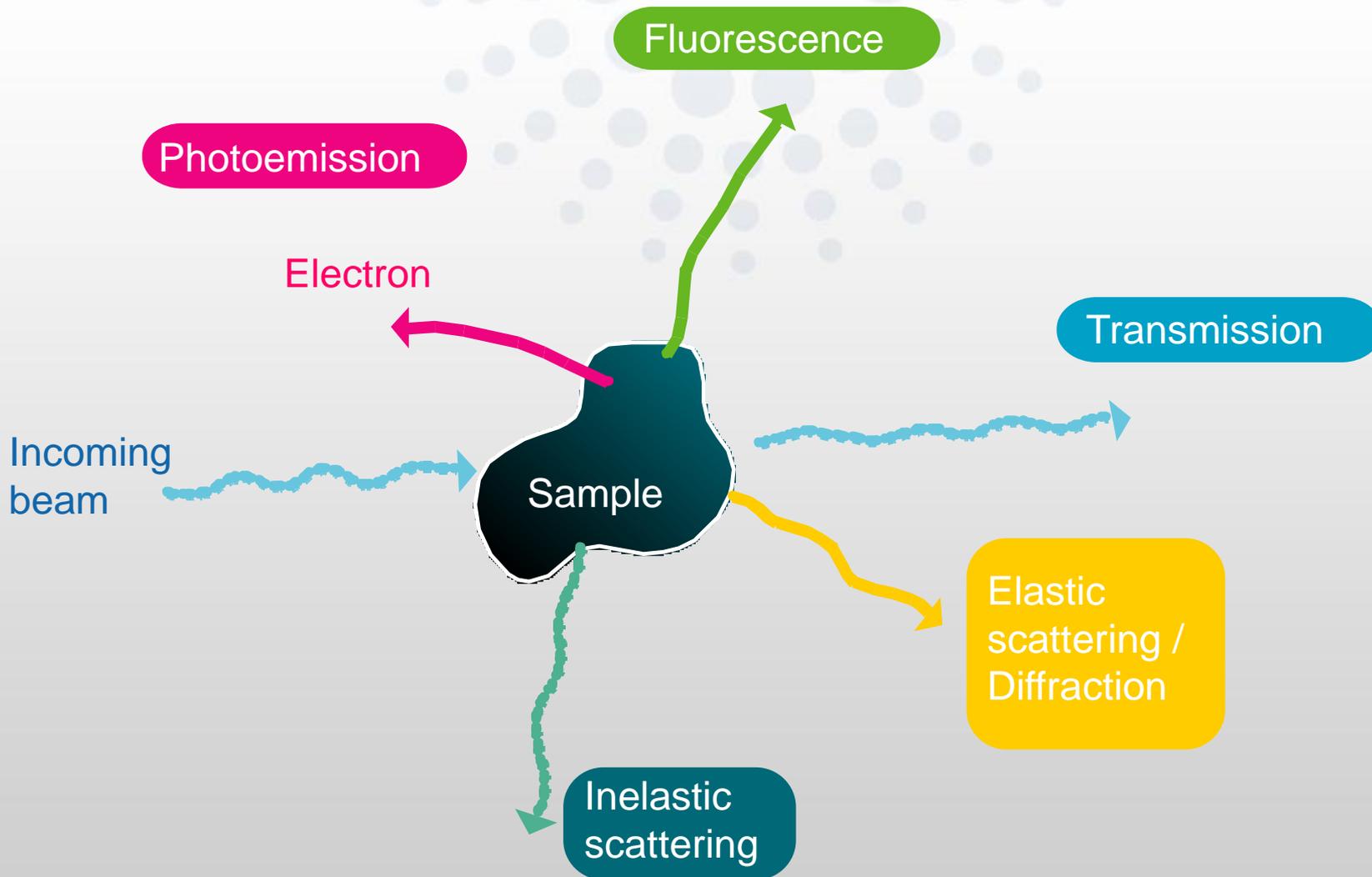
ESI 2013: Synchrotron Radiation

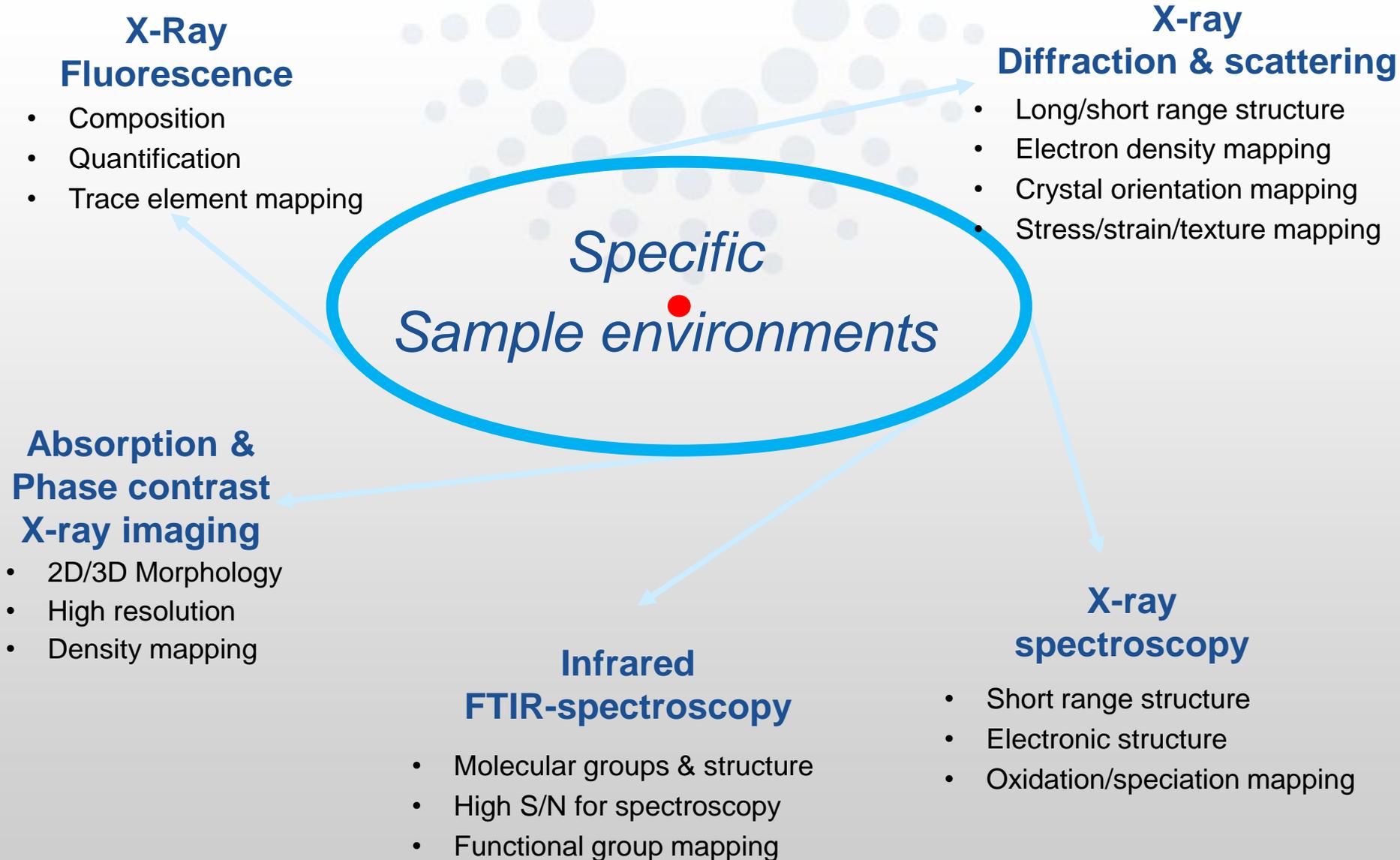
## PART II

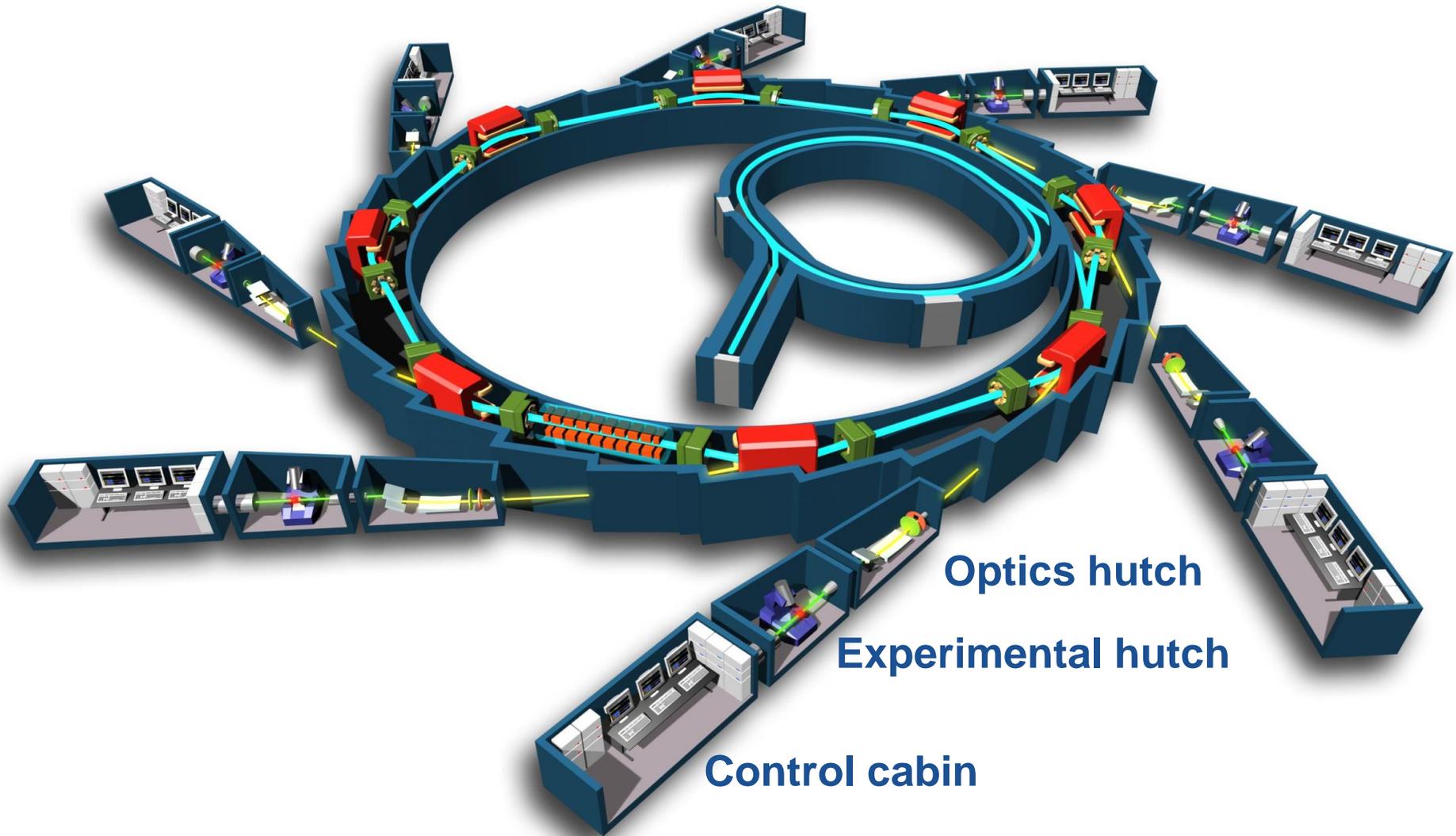
# SR Research

1. SR techniques
2. Some examples



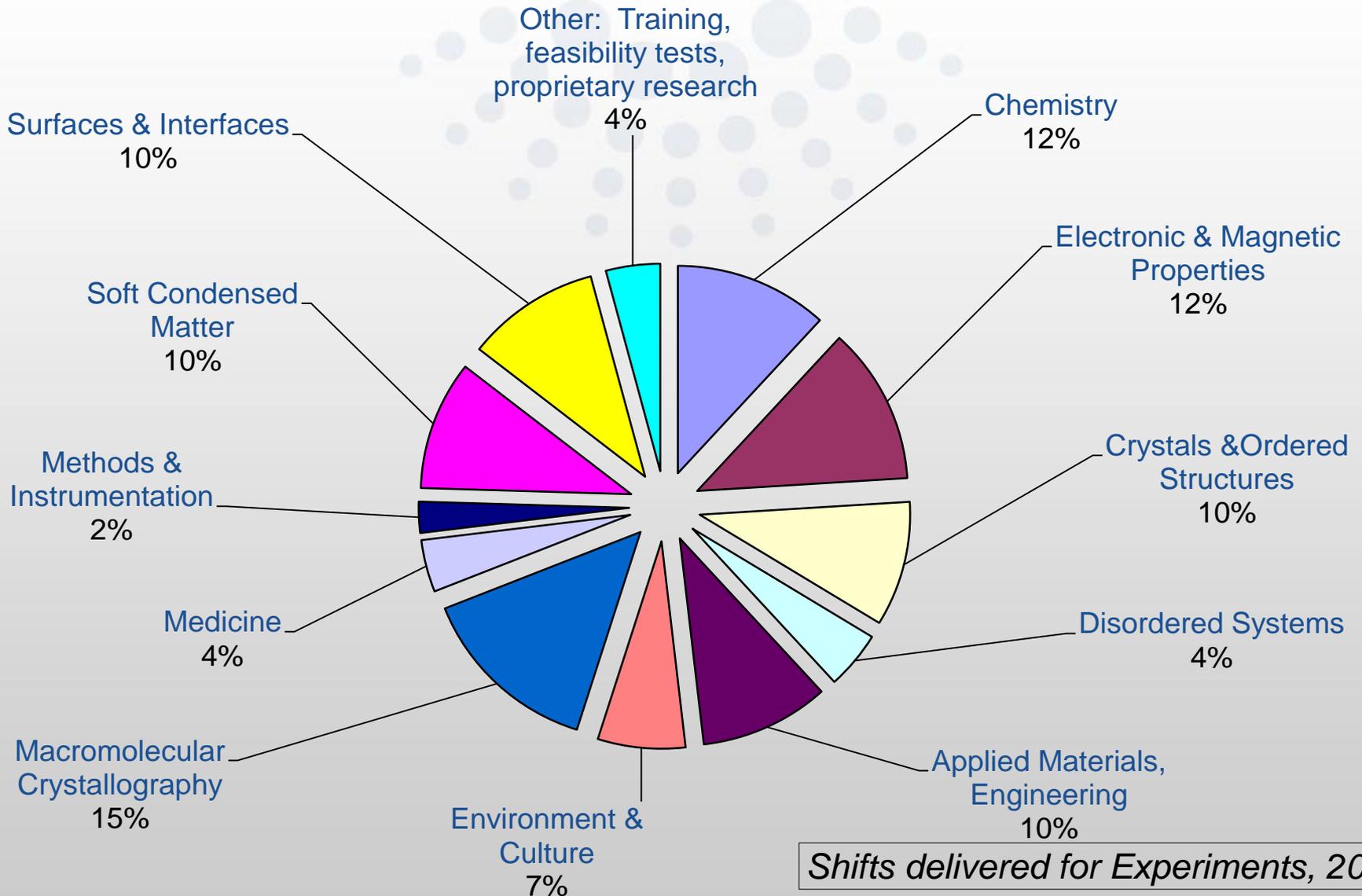






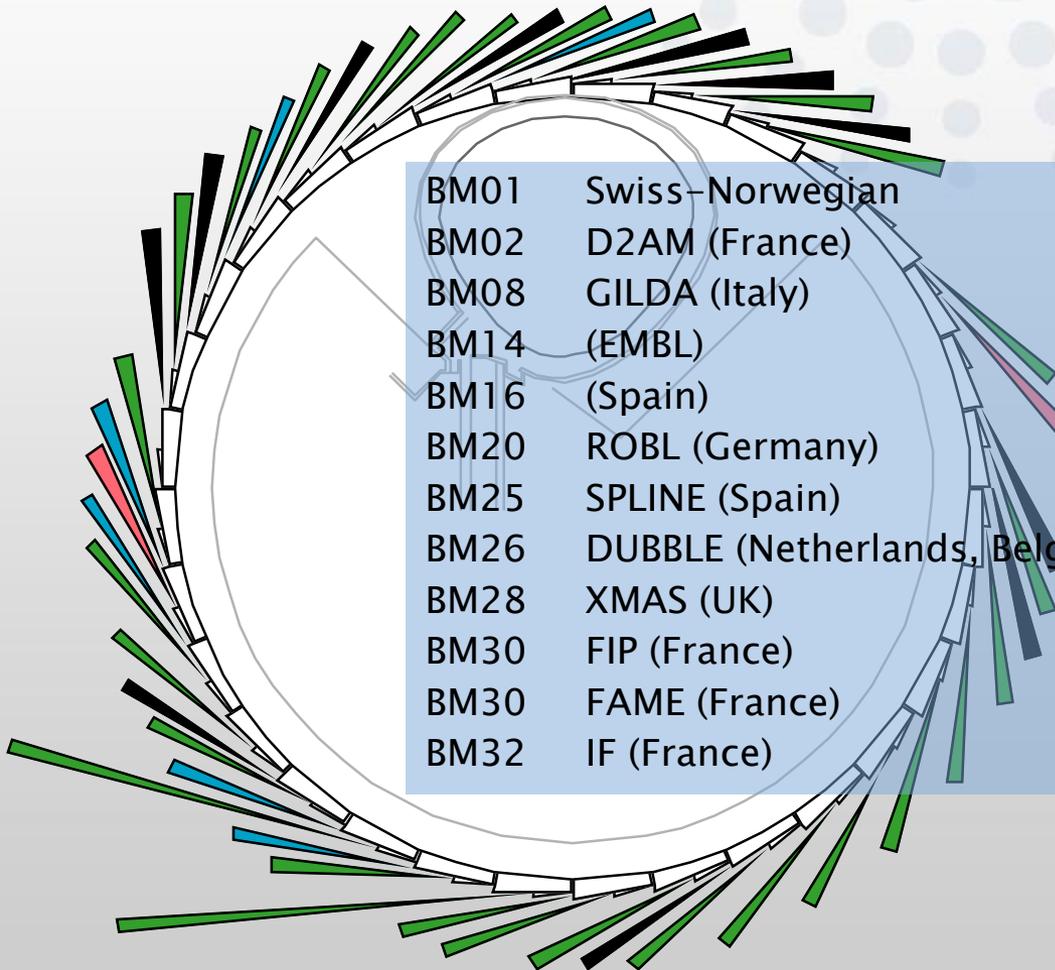
<http://www.synchrotron-soleil.fr/>

## Beamtime use at the ESRF



*Shifts delivered for Experiments, 2010*

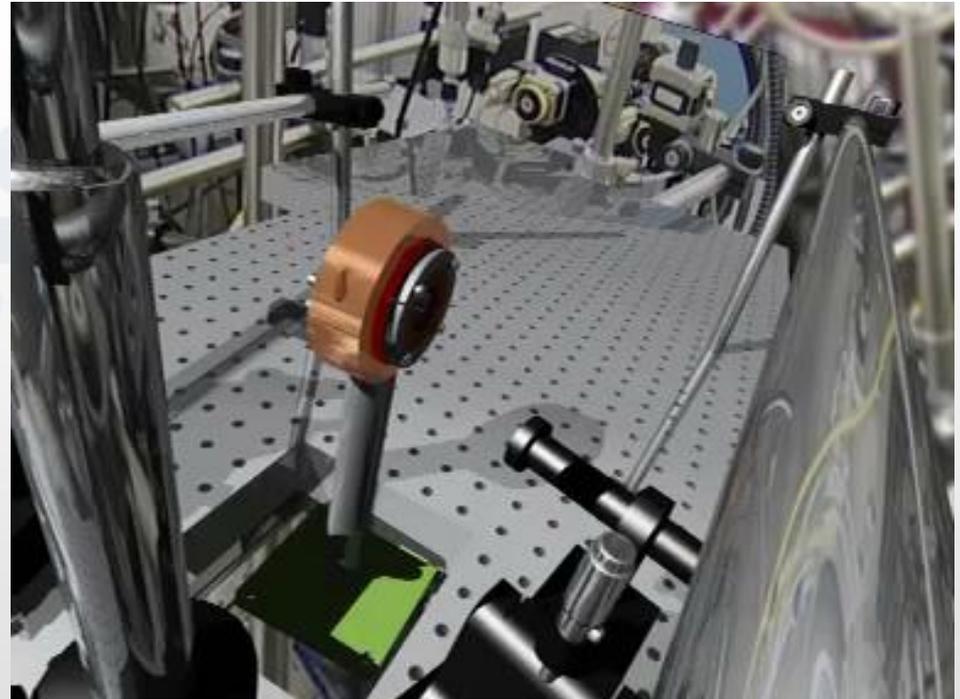
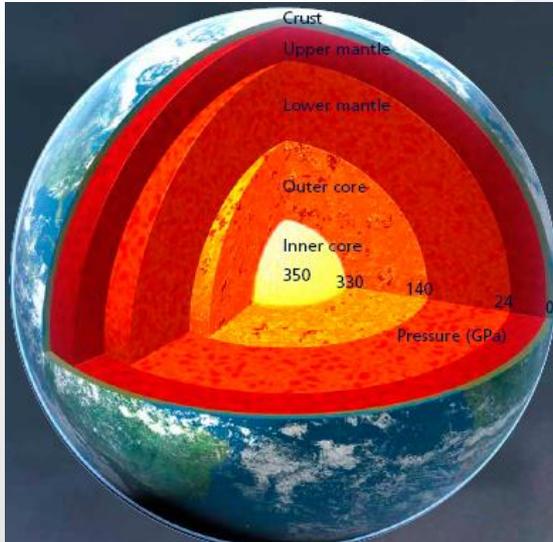
Over 40 different experiments can be conducted simultaneously



- BM01 Swiss-Norwegian
- BM02 D2AM (France)
- BM08 GILDA (Italy)
- BM14 (EMBL)
- BM16 (Spain)
- BM20 ROBL (Germany)
- BM25 SPLINE (Spain)
- BM26 DUBBLE (Netherlands, Belgium)
- BM28 XMAS (UK)
- BM30 FIP (France)
- BM30 FAME (France)
- BM32 IF (France)

- ID01 Anomalous scattering
- ID02 High brilliance
- ID03 Surface diffraction
- ID06 Instrumentation development
- ID08 Dragon / Spectroscopy using polarised soft X-rays
- ID23 Structural biology (MAD)
- ID09 Biology / High pressure
- ID10 Troika / Multipurpose
- ID11 Materials science
- ID12 Circular polarisation
- ID13 Microfocus
- ID14 Protein crystallography
- ID15 High energy diffraction
- ID16 Inelastic scattering
- ID17 Medical absorption spectroscopy

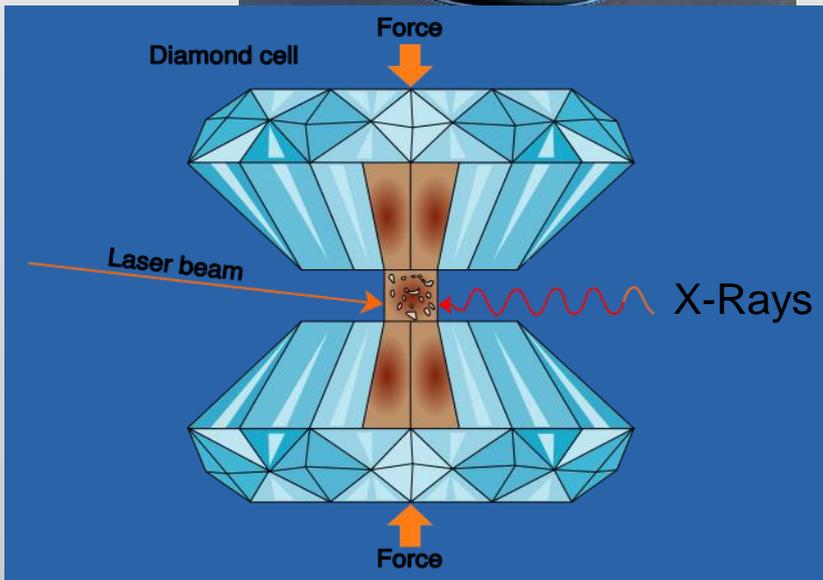
One can better understand the structure of matter at the center of the Earth ...



Diamond Anvil Cell (DAC)

... by studying samples put under extreme conditions of pressure and temperature.

- 200GPa ( 2Mbar)
- T=3600K

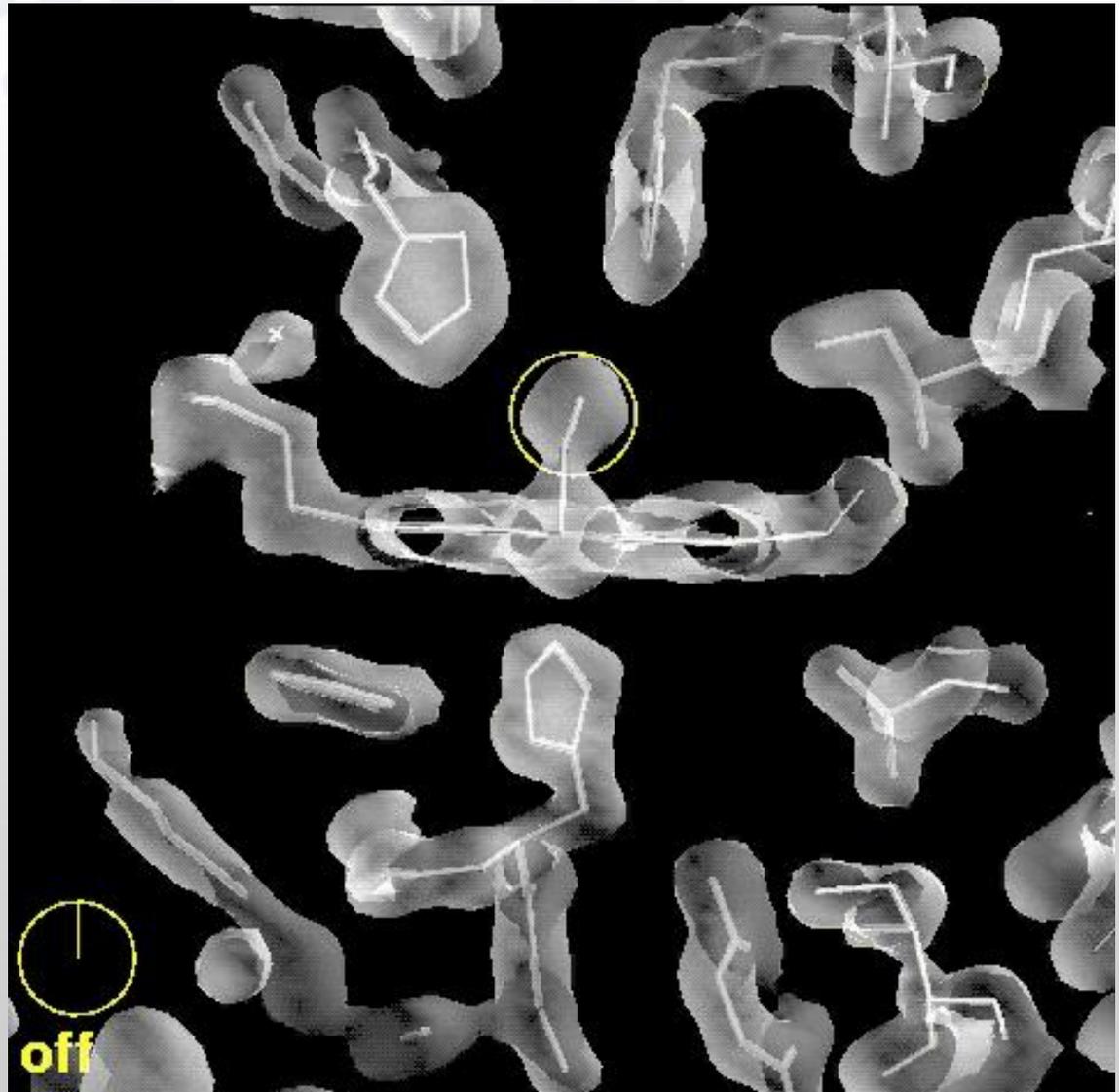


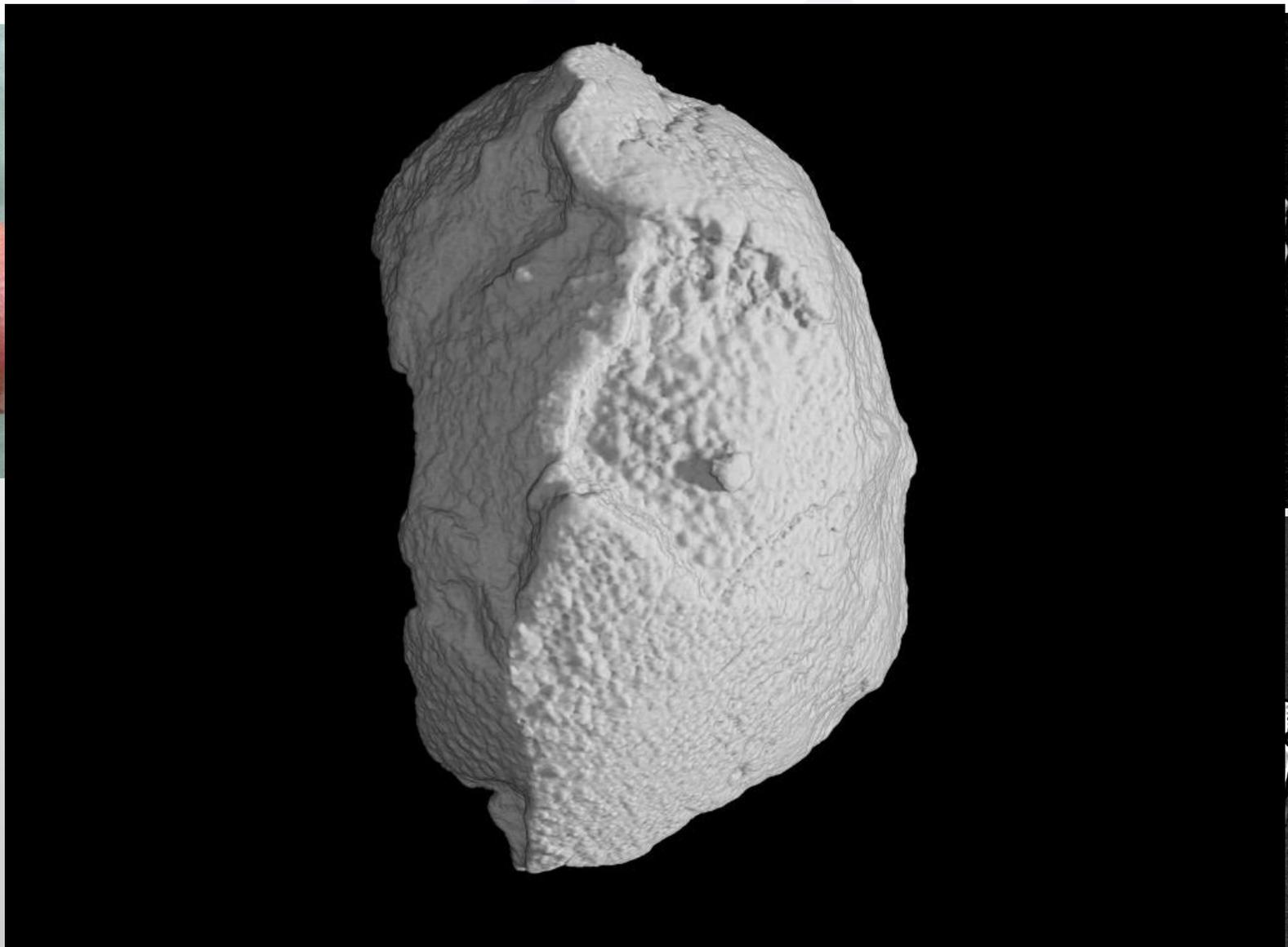
Molecular movies obtained by using the pulsed time structure of the SR

Migration of photodissociated CO in myoglobin (iron- and oxygen-binding protein in muscle tissue) using time-resolved single crystal diffraction in pump-probe experiment

Courtesy: M. Wulff, ESRF ID09

Srajer V, et al., Photolysis of the carbon monoxide complex of myoglobin: nanosecond time-resolved crystallography. *Science*. 1996;**274**:1726–1729





Courtesy P. Tafforeau et al. ESRF-ID19

