The Long-Baseline Neutrino Experiment: Taking aim at the neutrino mass hierarchy, CP violation and South Dakota

> Christopher Mauger LANL 17 September 2013

#### Outline

- The Long-Baseline Neutrino Experiment (LBNE)
- Long-Baseline Neutrino Oscillation Physics
- The LBNE Detector Design
- More LBNE Physics
- Conclusions

#### The Long-Baseline Neutrino Experiment



- Intense neutrino beam at Fermilab
- Near detector systems at Fermilab
- 34 kt liquid argon time-projection chamber (TPC) at Sanford Laboratory at 4850 foot depth – 1300 km from Fermilab

## **Scientific Motivation**

- Neutrino oscillations requires physics beyond the standard model
- Detailed studies of neutrino oscillations will allow us to answer important scientific questions:
  - What is the neutrino mass hierarchy?
  - Do neutrinos violate CP symmetry?
- High precision studies of neutrino oscillation phenomena allow us to test the three-flavor paradigm
  - Do sterile neutrinos exist?
  - Are there non-standard interactions (NSI)
- Building an experiment to address these issues with accelerator neutrinos enables much more science

### Scientific Motivation II

- Physics enabled by intense neutrino source and high-precision near neutrino detector
  - Precision electroweak tests
  - Searches for high  $\Delta m^2$  neutrino oscillation physics
  - Searches for ``dark photons''
- Physics enabled by a large, underground far detector
  - Atmospheric neutrino studies
    - complementary oscillation physics studies
    - indirect WIMP searches
    - astrophysical neutrino searches
  - Burst supernova neutrino studies
    - complementary oscillation physics studies
    - supernova physics
  - Beyond the Standard Model nucleon decay studies
    - SUSY, Grand Unified Theories
- And many others, see arXiv:1307.7335 LBNE whitepaper

#### The LBNE Collaboration

Alabama Argonne Jakota Sta Davis lawan Indian Group Indiana lowa State -Tokvo erkelev Los Alamo Louisiana State Maryland Michigan State Minnesota

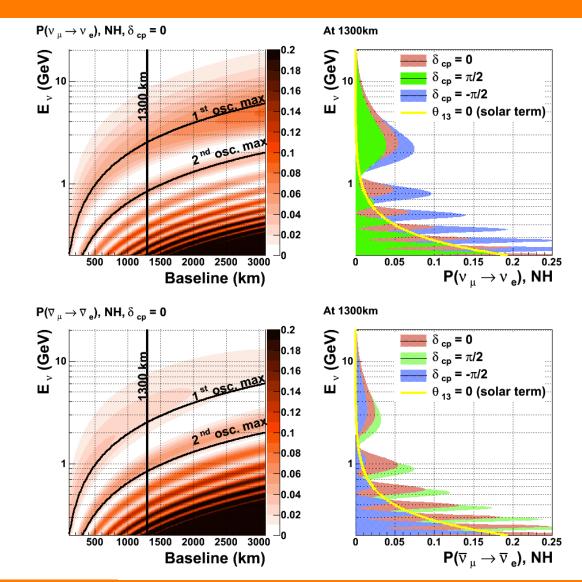
>400 members, 77 institutions, 6 countries
Co-spokespersons Milind Diwan (BNL), Bob Wilson (CSU)

NGA tern lotre Dame ford Rensselaer ochester Sanford Lab neffield South Carolina South Dakota ith Dako acuse rllington Austin Tufts **UCLA** Virginia Tech Washington William and Mary isconsin Yale

Fermilab, March 2013

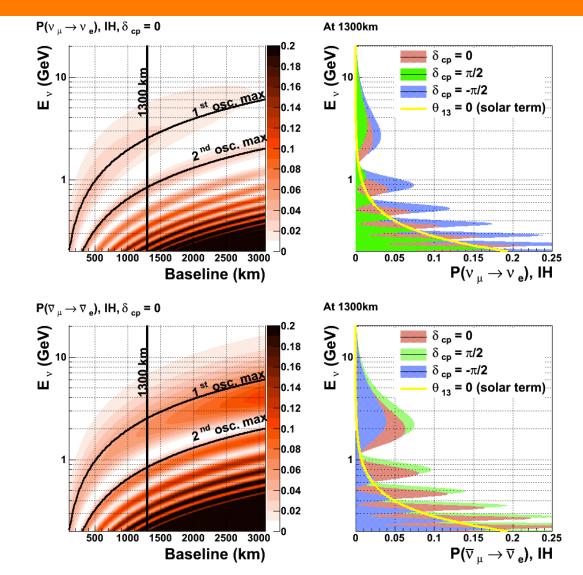
# Appearance Oscillograms (NH)

- Left plots: Neutrino oscillations vs energy and baseline for neutrinos (top) and antineutrinos (bottom) for  $\delta_{CP} = 0$
- Right: Neutrino oscillations as a function of neutrino energy for different values of  $\delta_{CP}$  for neutrinos (top) and antineutrinos (bottom) – solar term shown in yellow
- All plots assume Normal Hierarchy

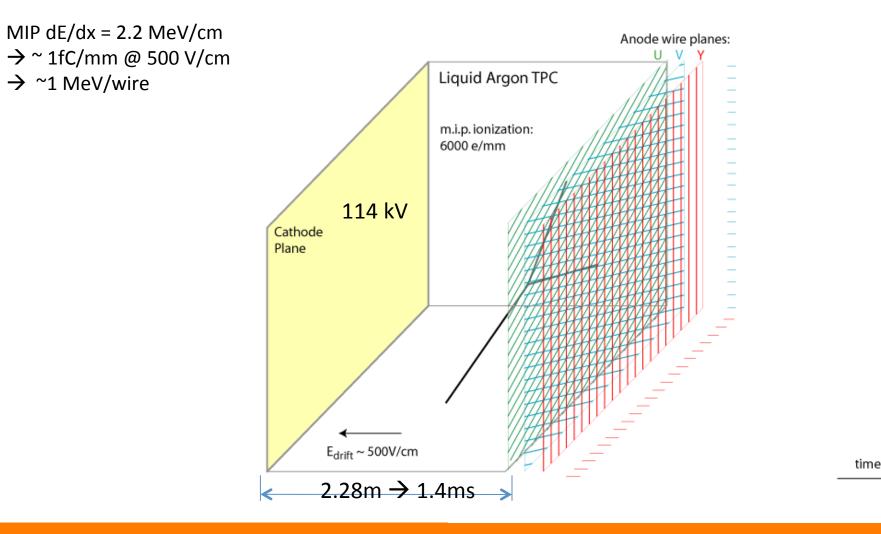


# Appearance Oscillograms (IH)

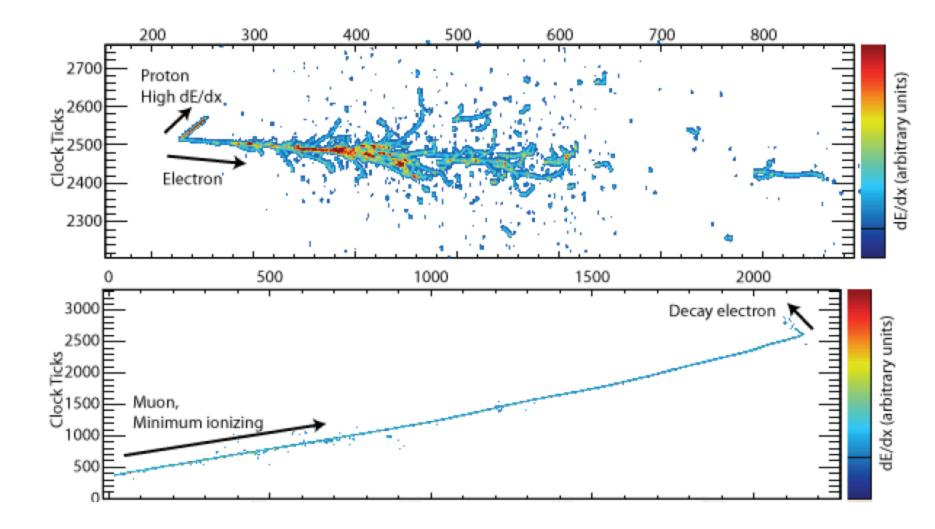
- Left plots: Neutrino oscillations vs energy and baseline for neutrinos (top) and antineutrinos (bottom) for  $\delta_{CP} = 0$
- Right: Neutrino oscillations as a function of neutrino energy for different values of  $\delta_{CP}$  for neutrinos (top) and antineutrinos (bottom) – solar term shown in yellow
- All plots assume Inverse Hierarchy



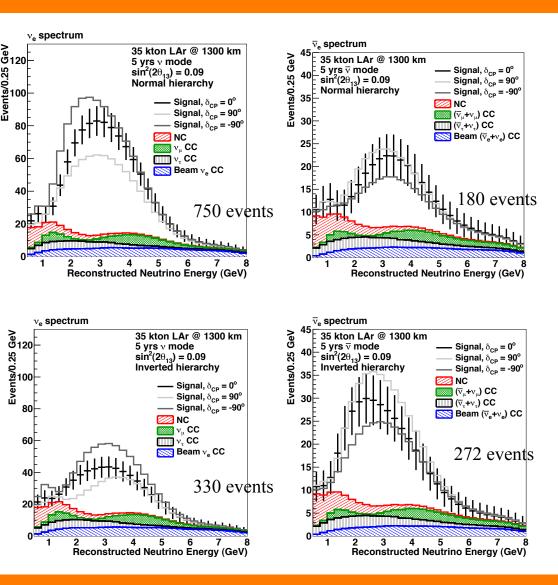
#### Liquid Argon Time-Projection Chambers (TPCs)



#### LAr TPC Performance



# **Electron Neutrino Appearance**

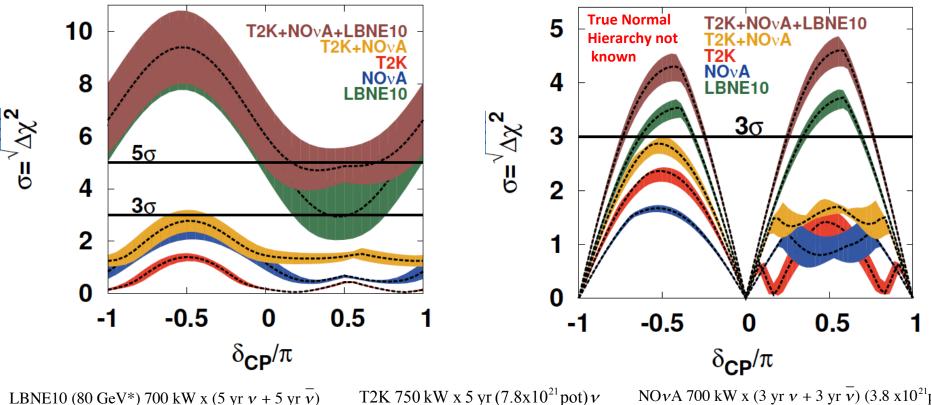


- Electron (anti)neutrino appearance spectra
- Tau neutrino background included
- Assume  $\sin^2 2\theta_{13} = 0.09, 80$ GeV proton energy beam, 700 kW, 5 years running in each mode
- Left upper: neutrino, normal hierarchy
- Left lower: neutrino, inverted hierarchy
- Right upper: antineutrino, normal hierarchy
- Right lower: antineutrino, inverted hierarchy

# 10 kt of LAr is compelling

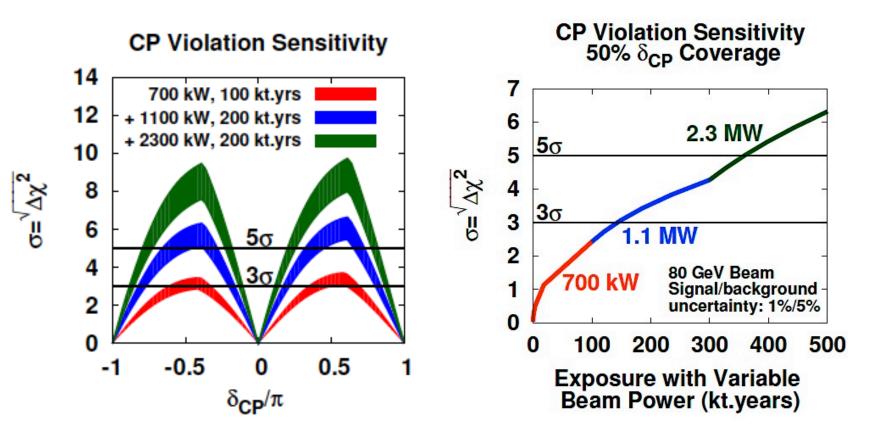
#### **Mass Hierarchy Sensitivity**

#### **CP Violation Sensitivity**



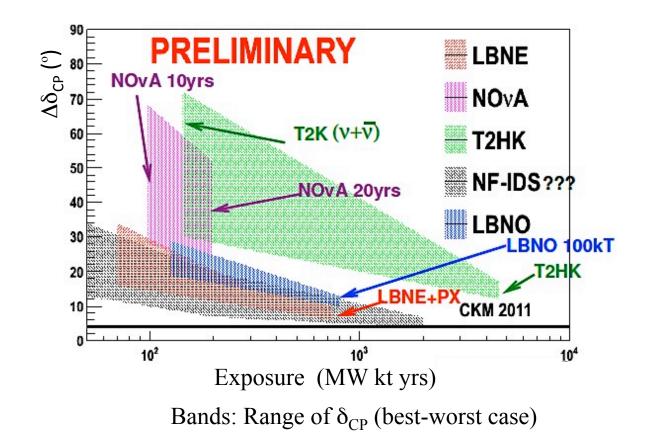
T2K 750 kW x 5 yr  $(7.8 \times 10^{21} \text{ pot})\nu$  NOvA 700 kW x  $(3 \text{ yr } \nu + 3 \text{ yr } \overline{\nu})$   $(3.8 \times 10^{21} \text{ pot})$ Bands: 1 $\sigma$  variations of  $\theta_{13}$ ,  $\theta_{23}$ ,  $\Delta m_{31}^2$  (Fogli et al. arXiv:1205.5254v3)

# LBNE + Project X (1.1-2.3 MW)



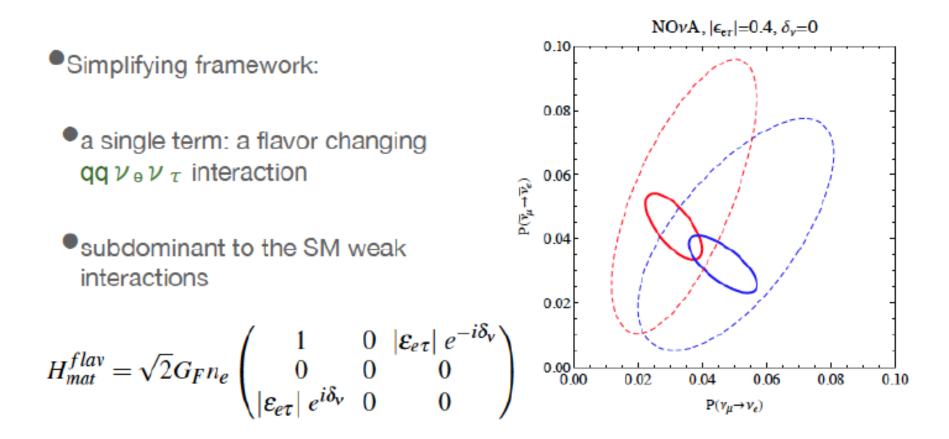
Comprehensive Global Science Program

#### **Global Context**



• LBNE + Project X approach level of precision of CKM matrix

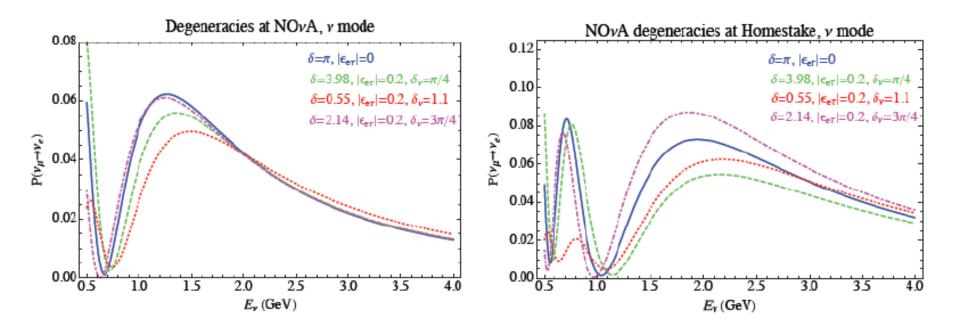
### Non-Standard Interactions (NSI)



Alex Friedland

#### NSI: Breaking degeneracy

• LBNE spectral measurements can break the degeneracy

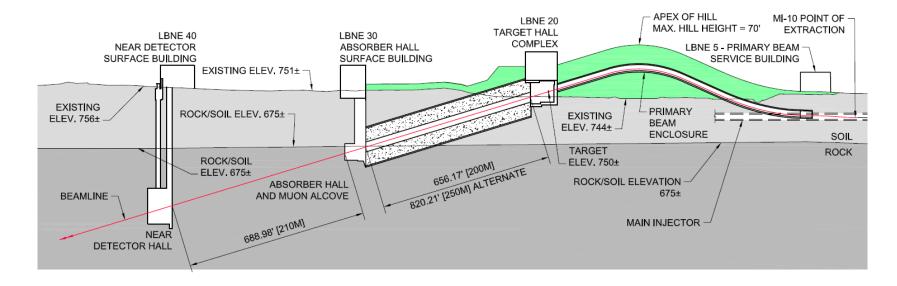


Requires a high precision near neutrino detector

From A. Friedland

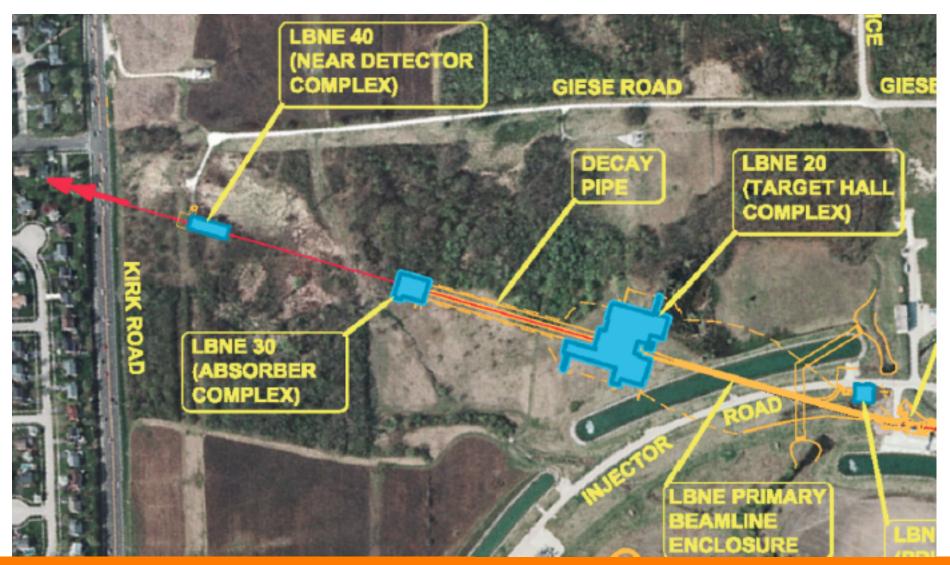
#### Near Detector Systems (NDS)

### Cross-section of NDS Layout

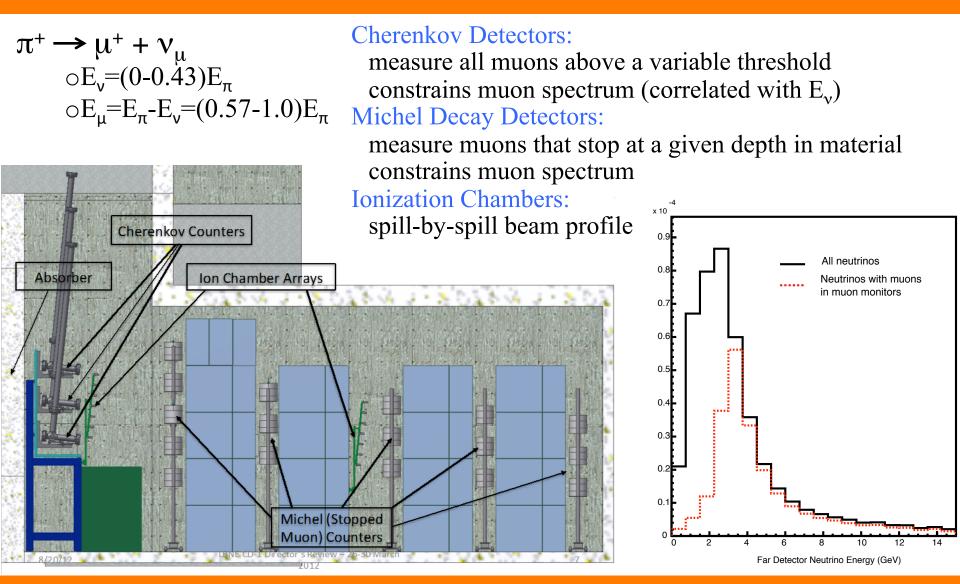


- Two sets of detector systems:
  - Measure muons after the absorber
  - Measure neutrinos

#### Plan View of NDS Layout

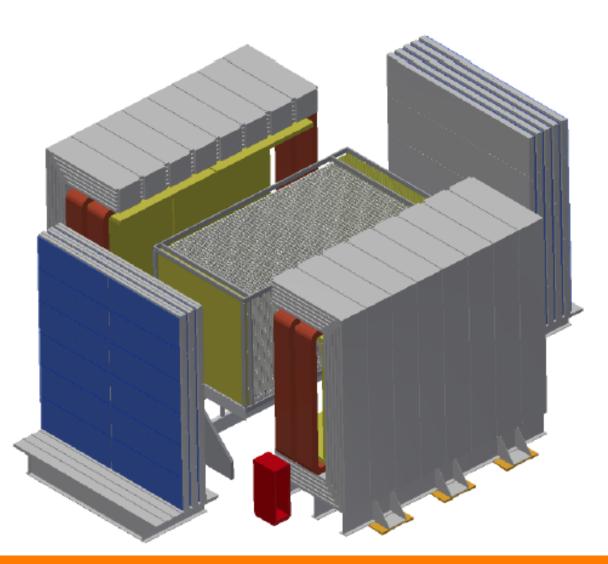


### Measurements of muons post-absorber



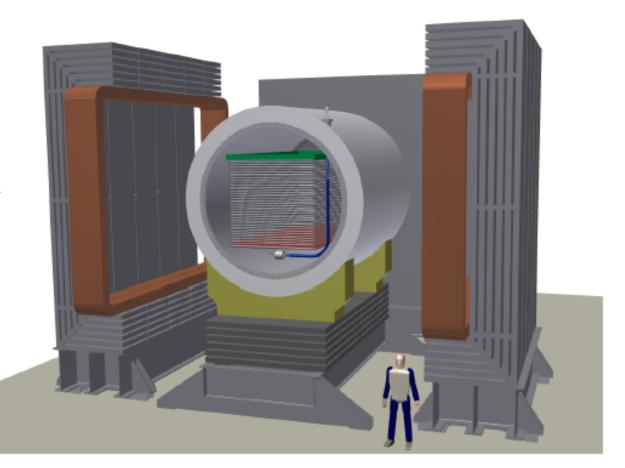
#### Near Neutrino Detector

- High precision straw-tube tracker with embedded high-pressure argon-gas targets
- Philosophy
  - make high-precision, highstatistics measurements of neutrino interactions with argon (far detector target nucleus)
  - measure inclusive and exclusive cross-sections to build and constrain models to predict the event signatures at the far site and correlate them with true neutrino energy
  - make detailed studies of electron (and muon) neutrino and anti-neutrinos separately



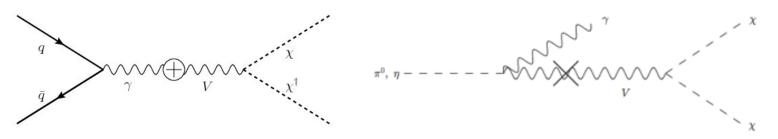
### Alternate Design

- Smaller (than the far detector) liquid argon TPC
- Philosophy
  - make high statistics measurements of event signatures in the same technology of the far site
  - try to minimize detector/reconstructionrelated systematic uncertainties by using the same technology as the far site
  - magnetization allows for charged muon separation for separate neutrino and antineutrino studies



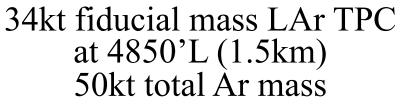
## Light Dark Matter Search

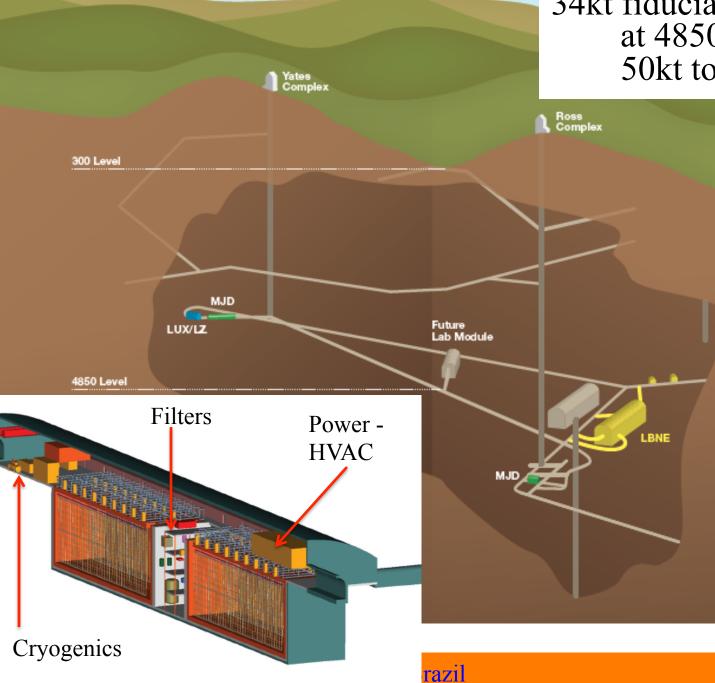
- Motivated by non-observation of SUSY, non-observation of direct detection experiments
- Simplest model, U(1) gauge field mixes with SM U(1) gauge field (dark photon)
- Can be produced when protons strike the LBNE target and decay into light (mass) dark-sector particles



- Detectable in a neutrino detector neutrino-like NC-like interactions
- Very forward electrons with late timing (mass much heavier than neutrinos)
- Special runs with no focussing would reduce the neutrino backround







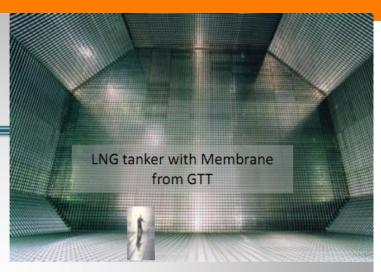
Location suitable for 10 to 100 kt of LAr

### Cryostat Design



The LNGC "Tembek", one of the thirty-one 216,000 m<sup>3</sup> LNG carriers ordered by Nakilat and delivered in 2008

#### Vendors: Gaz Transport & Technigaz (GTT) Ishikawajima-Harima Heavy Industries Co., Ltd. (IHI)



#### GTT Membrane Technology **IHI Membrane Technology** Stainless Steel Primary Plywood Membrane Membrane Sheet Reinforced GRE **Fire Proof Board** Polyurethan GRU Foam Secondary Moisture Barrier Membrane Reinforced Polyurethane Foam Plywood Concrete Bearing Mastic Vapor **GRE: Glass Cloth Reinforced Epoxy** Barrier GRU: Glass Cloth Reinforced urethane Nominal Panel Size Nominal Panel Size 3m x 8m x 2 mm thickness 1m x 3m x 1.2 mm



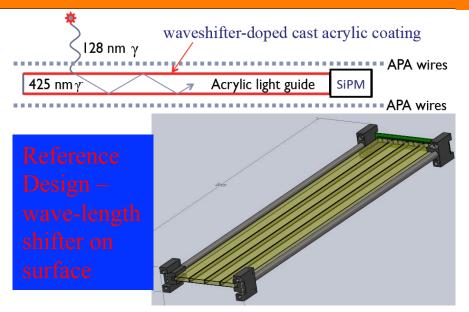
#### Christopher Mauger – WIN 2013, Natal, Brazil

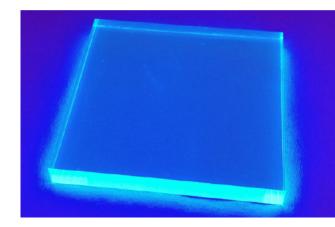
thickness

### **Photon Detection**

- The 1.4 ms drift time of the electrons is very slow.
  - The beam spill is  $10 \ \mu s \ long gives \ t0$
  - For non-beam related physics the interaction time is unknown. Use the light to determine the event time.
- A lot of light, but complicated structure
  - Scintillation and Cherenkov radiation 5 times more scintillation light
    - 23% of the scintillation light is prompt (~6ns)
    - 77% of the light is late (~1.6  $\mu$ sec).
  - Prompt yield 33,000 128nm photons per MeV for a MIP
  - Scattering length is ~95 cm
  - Present coverage is 0.4%

#### **Photon Detection System is critical for low energy events**



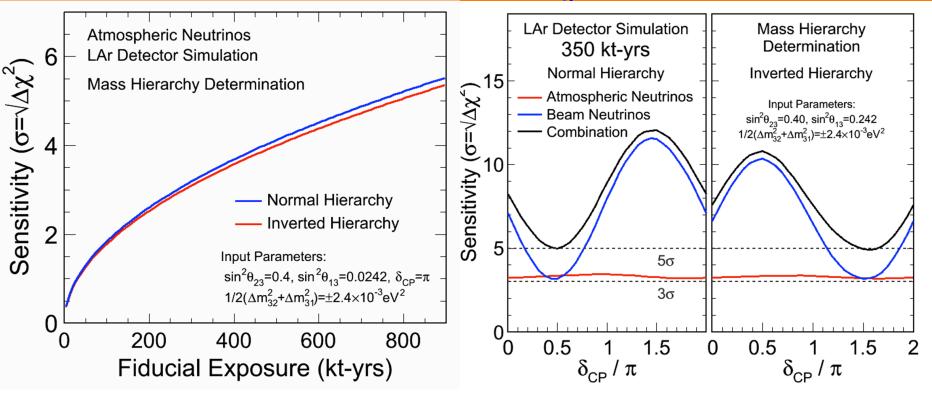


Design – doping of the bulk

### **Atmospheric Neutrinos**

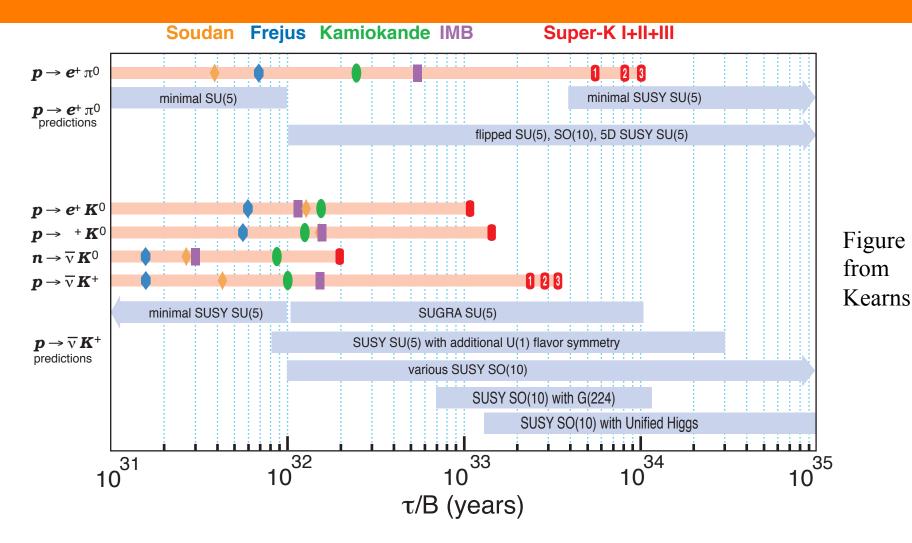
- Largest L/E range of any neutrino source above muon production threshold
- Protons and low-energy charged pions are visible in liquid argon TPCs better pointing resolution than water Cherenkov detectors at low-to-moderate neutrino energies
- Can separate muon neutrino and anti-neutrino events by searching for muon capture (75% probability for μ<sup>-</sup>), final state proton vs. neutron for quasi-elastic neutrino interactions
- These features give the LBNE far detector sensitivity to the neutrino mass hierarchy

#### Atmospheric Neutrinos Neutrino Mass Hierarchy

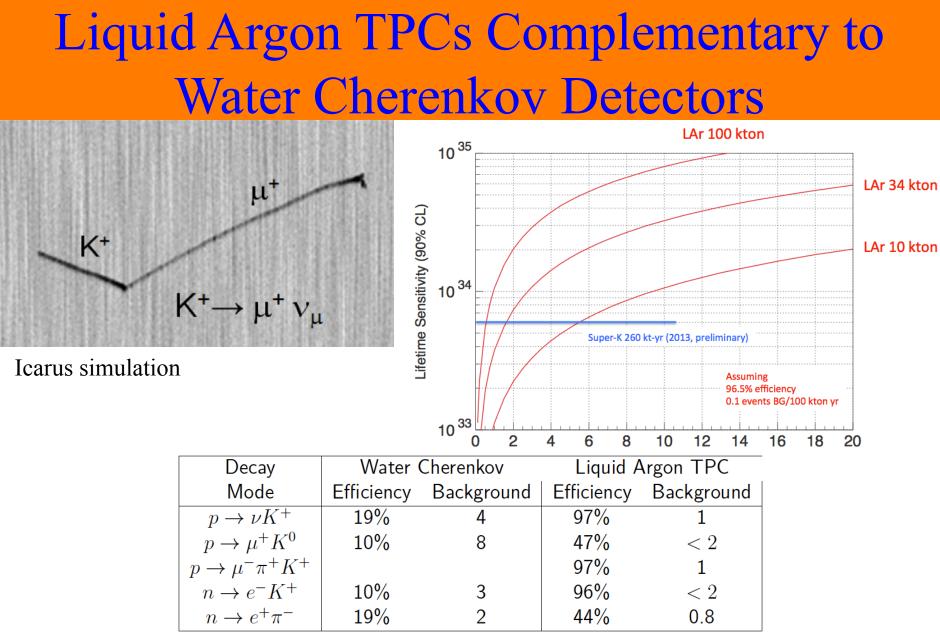


- Atmospheric neutrinos provide independent test in a different system with the same detector systematics
- Beam and atmospheric neutrinos combined give excellent sensitivity

### Nucleon Decay



Predicted by models beyond the Standard Model – Grand Unified Theories, Supersymmetry



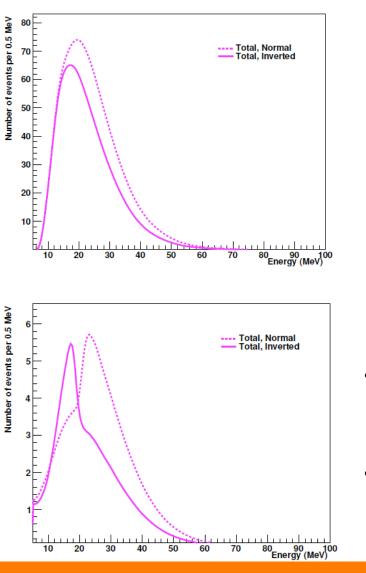
• Smaller exposure, so LAr shines with high efficiency modes

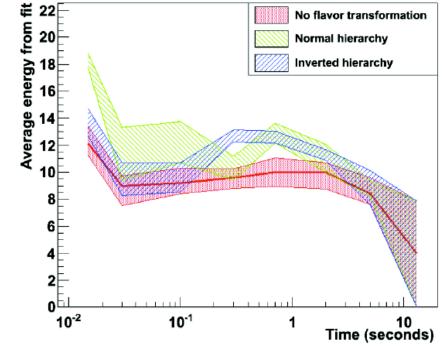
### Supernova Neutrinos

- Supernova bursts in our galaxy are a fantastic source of neutrinos
- Significant fluxes in < 10 seconds
- Matter effects unachievable from other sources
- Argon uniquely sensitive to CC electron neutrino interactions complementary to water Cherenkov detectors sensitive to CC electron anti-neutrino interactions

Reaction Type	Events / 10 kton	(at 10 kpc)
(CC) $v_e + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$	~700 [1]	
(CC) $\overline{\nu}_e + {}^{40}\text{Ar} \rightarrow e^+ + {}^{40}\text{Cl}^*$	~ <mark>60</mark> [1]	[1] K. Scholberg
(ES) $v_{\mathbf{x}} + e^{-} \rightarrow v_{\mathbf{x}} + e^{-}$	~85 [1]	[2] A. Hayes
(NC) $v_{\mathbf{x}} + {}^{40}\mathrm{Ar} \rightarrow v_{\mathbf{x}} + {}^{40}\mathrm{Ar}^*$	~90 [2]	

### Neutrino Mass Hierarchy Information





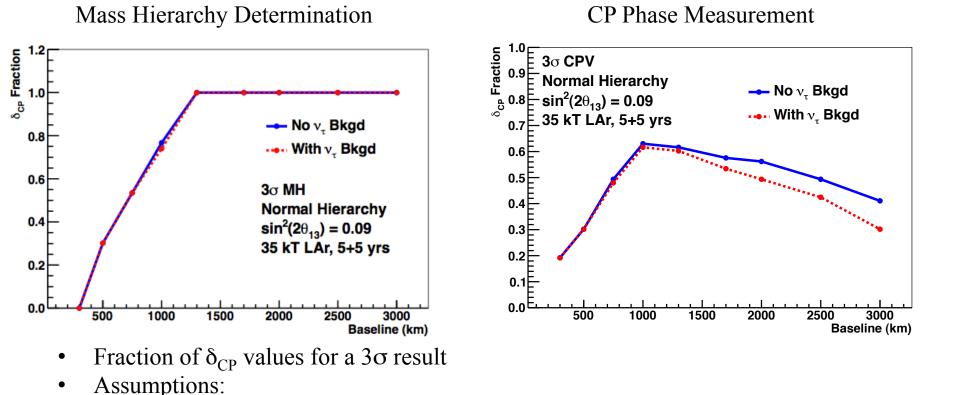
- Left: Event rates for a 100kt water Cherenkov detector (upper) and 17kt (sorry) liquid argon TPC (lower) (model from Duan and Friedland: Phys. Rev. Lett., 106:091101, 2011)
- Upper: Average electron neutrino energy as a function of time for different mass hierarchy assumptions with 34 kt (model from Keil, Raffelt, and Janka: Astrophys. J., 590:971-991, 2003)

### Conclusions

- The Long-Baseline Neutrino Experiment consists of an exciting and diverse physics program enabled by a strong and developing international partnership
- The beam neutrino physics probes the neutrino mass hierarchy, leptonic CP violation, non-standard neutrino interactions among others
- The highly capable near neutrino detector enables the long-baseline neutrino oscillation program as well as a significant high precision neutrino interaction program
- The far detector besides its duties as a far detector for the beam physics program enables searches for baryon number violation, measurements of neutrinos from supernova bursts, and measurements of atmospheric neutrinos
- LBNE has received approval to begin this program



## **Baseline Optimization**



120 GeV, 700kW beam, 5+5 years of running (5 neutrino, 5 anti-neutrino), 35-kt LAr TPC

1300km baseline economical solution for a comprehensive neutrino oscillation

#### program

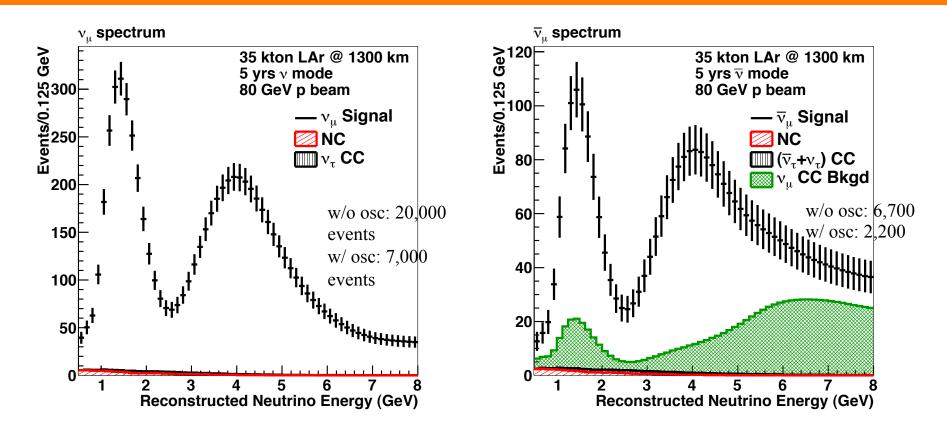
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Normal hierarchy,  $\sin^2 2\theta_{13} = 0.09$ 

## Muon Neutrino Disappearance

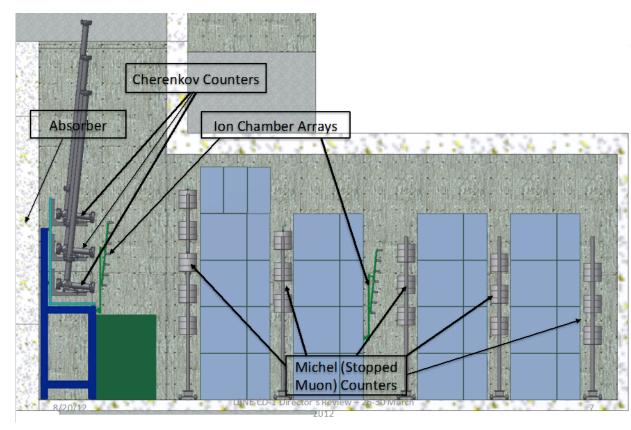


- 35-kt LAr TPC, 80 GeV, 700kW proton beam
- Tau neutrino background accounted for

# Measurements of muons post-absorber

### Ionization Chambers: spill-by-spill beam profile Cherenkov Detectors:

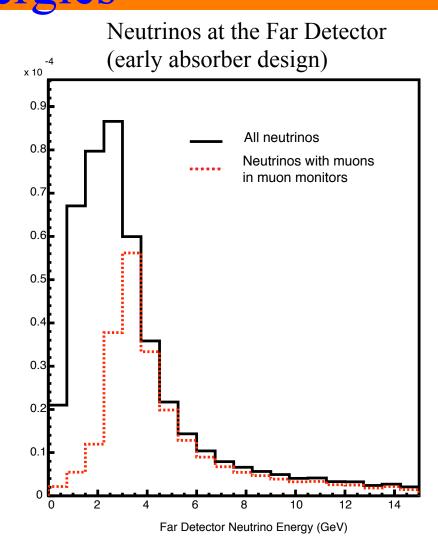
measure all muons above a variable threshold constrains muon spectrum (correlated with  $E_v$ ) Michel Decay Detectors: measure muons that stop at a given depth in material constrains muon spectrum may give absolute flux constraint



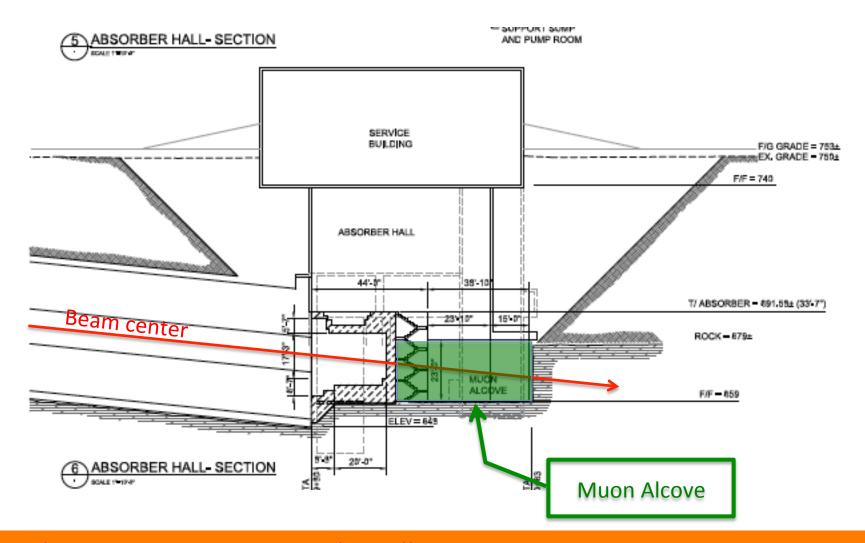
Includes planning measurements of hadron production in external beamlines on materials from which the target and horns are composed

### Correlation between muon and neutrino energies

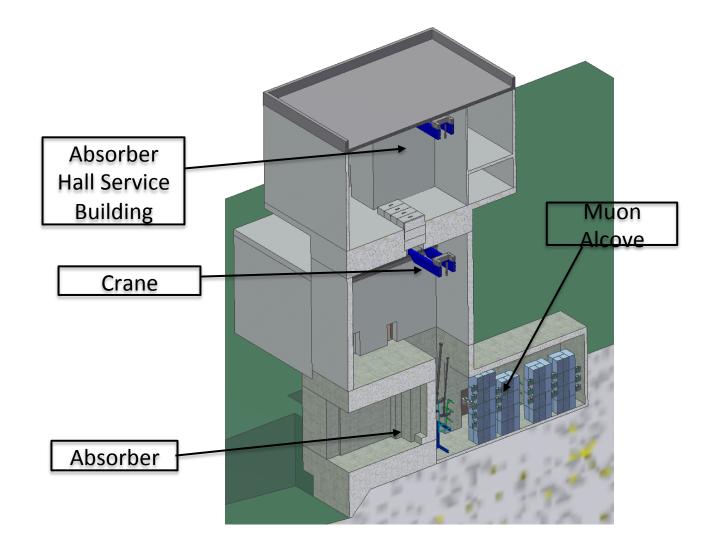
- Muons and neutrinos are anti-correlated in the two-body decays of pions and kaons
- Muons take most of the momentum in the decay
- For pions:
  - $\circ$  E<sub>v</sub>=(0-0.43)E<sub>n</sub>
  - $\circ E_{\mu} = E_{\pi} E_{\nu} = (0.57 1.0)E_{\pi}$



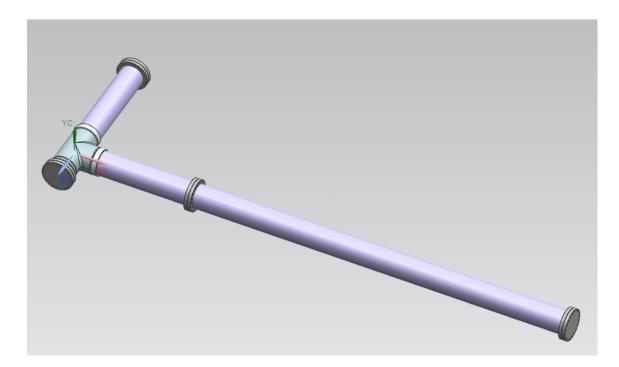
### Absorber an Muon Alcove



### Absorber and Alcove Region

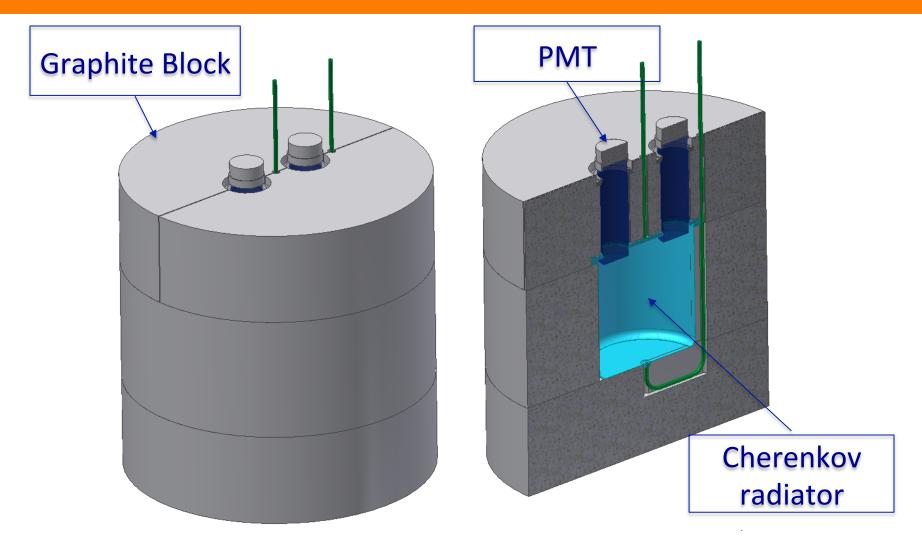


## **Threshold Cherenkov Detector**



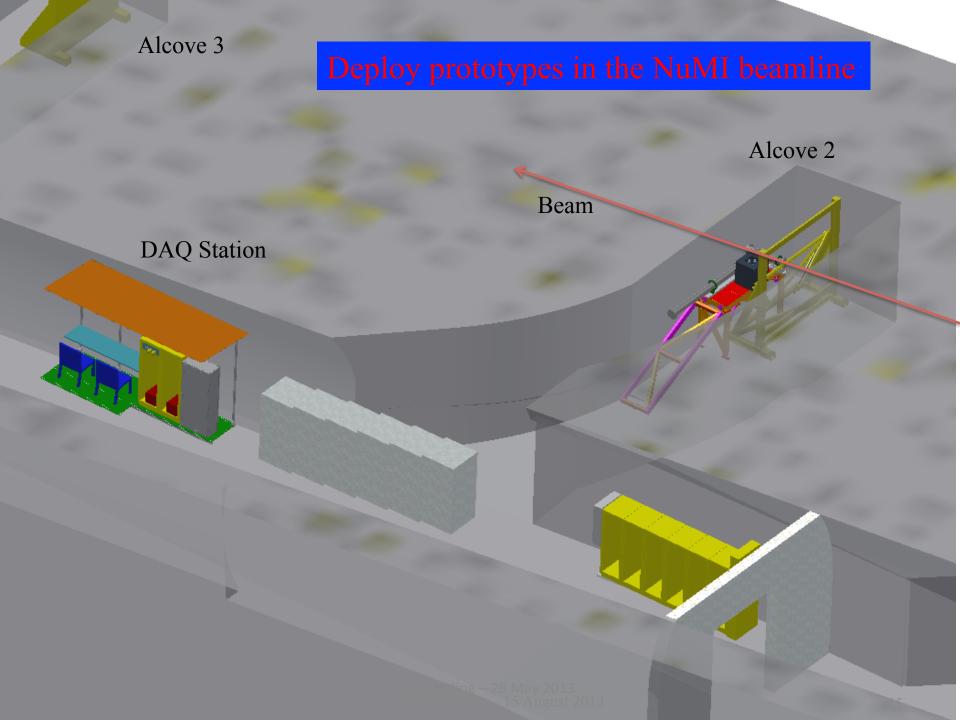
- Pressure varies from vacuum to ~20 atm
- Collect forward light from muons near Cherenkov threshold
- Flat mirror optics

## Stopped muon detector



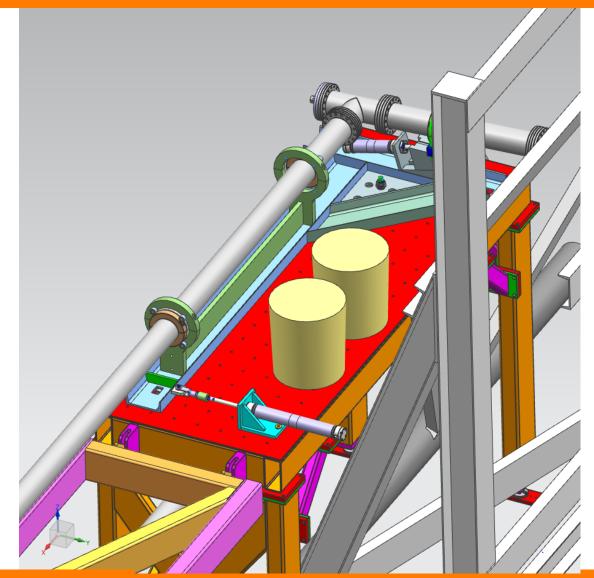
Measures decay of muons stopped in the radiator and decay of boron from muon capture

# **Prototyping Activities**



# Table for All Prototypes

- During NuMI shutdown - install table for easy deployment of prototypes during brief subsequent shutdowns
- Install all required services
- NuMI beamline running imminent
- Some prototypes are ready for installation



### Threshold Cherenkov Detector Prototype

Mirror Alignment in MI8

Laser

A 20 P

Laser alignment lines

### Mirror Installed in Tee

3.5 inch mirror

Jan

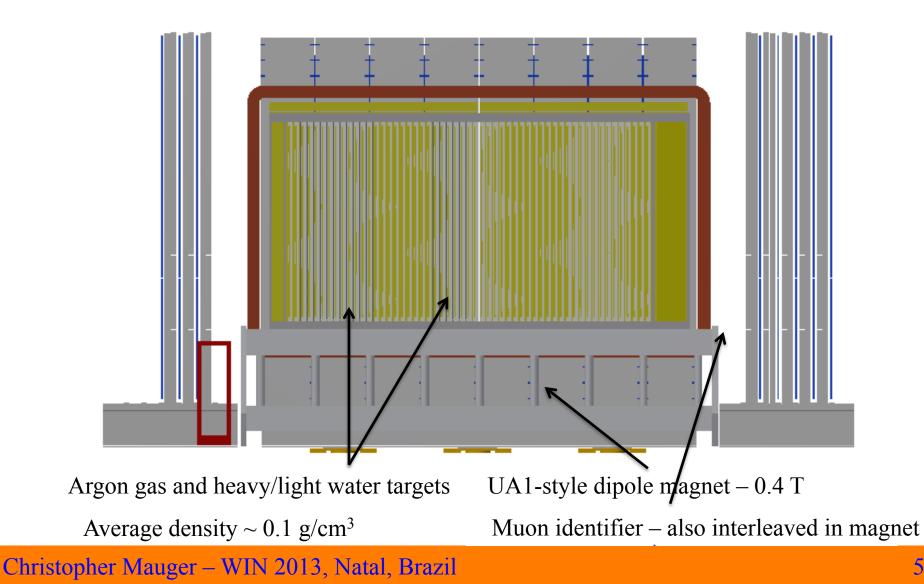
### Sealed system being pumped down



## Strawtube Tracker

- Low density allows for electron sign discrimination
- Detailed absolute and relative flux measurements
  - low  $v_0$  method
  - neutrino-electron elastic scattering
  - inverse muon decay
  - low-momentum transfer quasi-elastic events on hydrogenic targets
- High-precision measurements of all important interaction modes for LBNE on argon
  - Quasi-elastic
  - Resonant production including strange production
  - Deep Inelastic Scattering
- Detailed measurements of various nuclear effects especially those that impact neutrino energy reconstruction in the far detector
- Enables high-precision prediction of events in the far site for any neutrino oscillation paradigm

### Strawtube Tracker



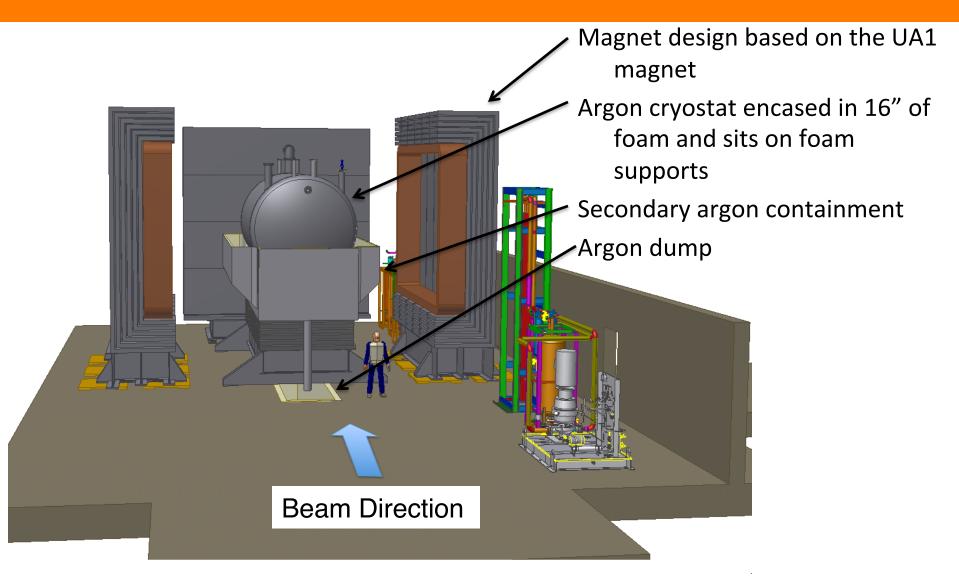
Barrel ECal Side Module 3.96m Y direction, 3.71m Z direction 8 layers of 7mm thick lead, 10mm thick scintillator (4 total modules)

 $4\pi$  coverage gives excellent neutral pion identification and containment Barrel ECal Top/Bottom Module 2.06m X direction, 3.71m Z direction 8 layers of 7mm thick lead, 10mm thick scintillator (8 total modules)

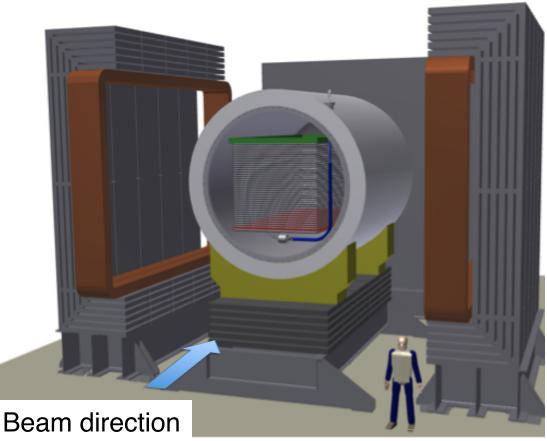
## Near Detector Physics List

- Weak mixing angle
- Strangeness content of the nucleon
- Nucleon structure
- Search for heavy neutrinos
- Search for high  $\Delta m^2$  neutrino oscillations
- Light dark matter searches
- Many others, see the whitepaper

# Alternative Design – Liquid Argon TPC



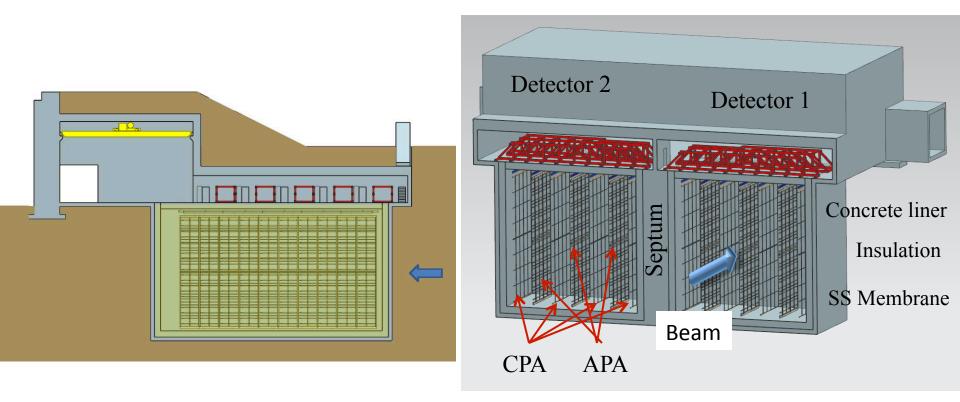
## Near Liquid Argon TPC



Based on MicroBooNE design 4.0m x 1.8m x 1.8m active volume – 18 tons of liquid argon Tracks charged particles Electromagnetic calorimetry 0.4 Tesla field Sign of muons, momentum for long tracks Downstream/side detectors in steel layers for muon/pion separation

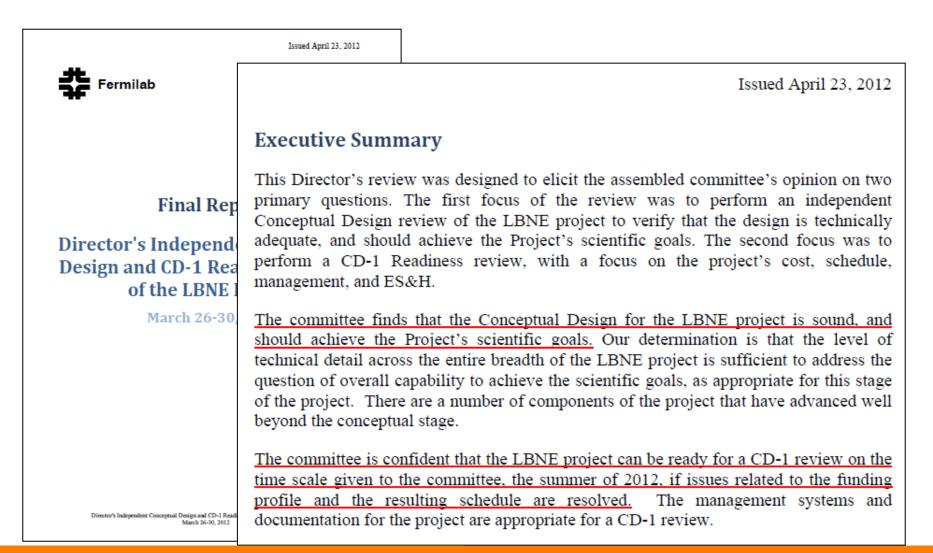
### Far Detector Layout

- 10 kt fiducial mass Liquid Argon (LAr) detector located on the surface in Lead, SD (two 5 (9.4) kt Fiducial Mass (Total Mass) modules)
- Detector designed to detect accelerator neutrinos
- Anode and Cathode Plan Assemblies (APAs, CPAs)



### **LBNE Reconfiguration**

### Complete Design Independently Reviewed and Found to be Sound





- Last year US funding agency (DOE) asked us to stage LBNE construction and gave us a budget of \$867M for the first phase
  - They also encouraged us to develop new partnerships to maximize the scope of the first stage.
- We chose to proceed with emphasis on the most important aspects of the experiment: 1300 km baseline and the full capability beam
  - With just the DOE budget, the far detector would be 10 kt LAr TPC at the surface and there would be no near neutrino measurements.
- An external review panel recommended this phase 1 configuration.
- DOE approved "CD-1" in December 2012 for this phase-1 scope.
- Our plan continues to be to build the full scope originally planned, and are working with domestic and international partners to make the first phase as close as possible to the original goal.

### **DOE CD-1** Approval Document

Ibne-doc-6681

Critical Decision 1 Approve Alternative Selection and Cost Range of the Long Baseline Neutrino Experiment (LBNE) Project (Line Item Project 11-SC-40) at the Fermi National Accelerator Laboratory and Sanford Underground Research Facility Office of High Energy Physics Office of Science

#### Purpose

The purpose of this paper is to document the review and approval by the DOE Office of Science Energy Systems Acquisition Advisory Board-equivalent for Critical Decision 1 (CD-1) "Approve Alternative Selection and Cost Range" for the Long Baseline Neutrino Experiment (LBNE) Project at the Fermi National Accelerator Laboratory (Fermilab) and Homestake Mine Critical Decision 1, Approve Alternative Selection and Cost Range for the LBNE Project

#### Approval

Based on the information presente din this document and at the ESAAB review, I approve Critical Decision 1, Approve Alternative Selection and Cost Range for the Long Baseline Neutrino (LBNE) Project.

William Brinkman, Acquisition Executive Director, Office of Science 12/10/12 Date

Tailoring of the scope definition prior to CD-2 to enhance scientific capabilities may also be considered. The physics opportunities offered by the beam from Fermilab and the long baseline may attract the support of other agencies both domestic and international. Contributions from such other agencies offer alternative funding scenarios that could enhance the science capabilities of the Project. If additional domestic or international funding commitments are secured sufficiently prior to CD-2, the DOE LBNE Project baseline scope could be refined before CD-2 to include scope opportunities such as a Near Neutrino Detector complex at Fermilab or an underground location at SURF for the far detector.

the neutrino mass states, would not be obtained, compromising the ability to understand the matter-antimatter asymmetry and resulting dominance of matter in the universe.

To meet the scientific and technical objectives for the LBNE experiment, the following draft key performance parameters have been developed.

http://lbne2-docdb.fnal.gov/cgi-bin/RetrieveFile?docid=6681;filename=LBNE%20CD-1%20appr.pdf

# LBNE Phase I Goal

- Together with additional partners, build:
  - Neutrino beam for 700 kW, upgradeable to 2.3 MW
  - Highly-capable near neutrino detector
  - >10 kt fiducial mass LAr far detector at A baseline of 1300 km A depth of 4300 m.w.e.
- The world-wide community can build upon the substantial investment planned by the US to make LBNE a world facility for neutrino physics, astrophysics, and searches for non-conservation of baryon number.

# **Engaging International Partners**

- We are in discussion with a number of potential non-US partners, both physics groups and funding agencies, in:

   Brazil
   India
   Italy
   UK
- LBNE and LAGUNA-LBNO have established a working group to explore joining forces
- Italian ICARUS groups joined LBNE
- Brazilian groups joined LBNE
- We have initiated preliminary discussions with:
   -CERN Dubna
- Engaging others potential partners:

   Japan
   Additional countries in the Americas, Asia and Europe
- Also exploring how to engage domestic US funding agencies beyond the DOE

## Near Neutrino Detector

- Indian Institutions have proposed to build a near neutrino detector based on a strawtube tracker
- Moving forward in the context of a broader Indian Institutions-Fermilab Collaboration
- Nascent prototyping effort at Panjab University and IIT-Guwahati
- Significant LBNE intellectual partner
- This is one example giving us hope for an exciting phase I physics program

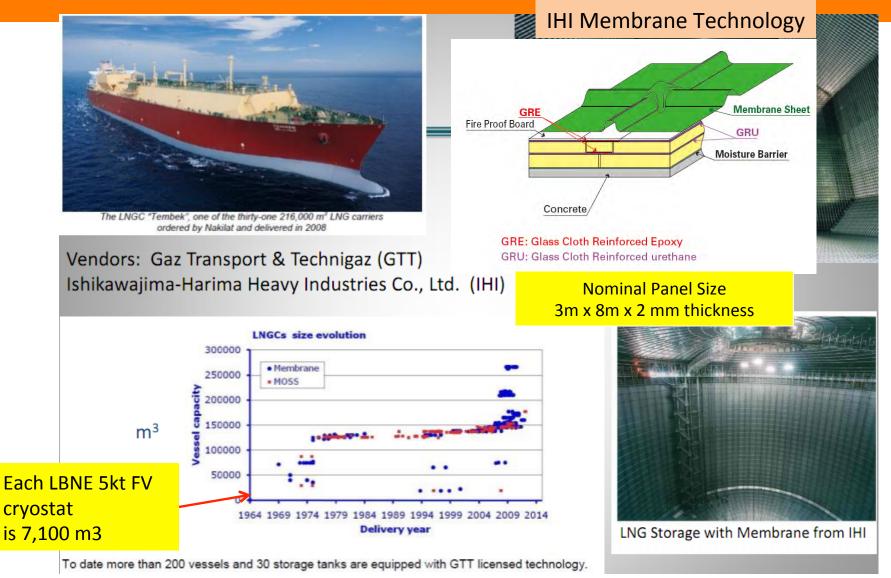
### LBNE and IIFC-nuP Working Meeting 6-7 June 2013

🛟 Fermilab

by Fermi Research Aldernice As Science/U.S. Department of Unit

Xinchun Tian, K Naga Depthi, Christopher Mauger, Bindu Bambah, Roberto Petti, Amandeep Singh, Bipul Bhuyan, Anjan Giri Brian Mercurio, Kuldeep Kaur, Rukmani Mohanta, Ashok Kumar, Ramesh Babu T, Venktesh Singh, Milind Diwan, Raj Gandhi Sanjib Mishra, Bob Wilson, Sonam Mahajan, Jim Strait, Shekhar Mishra, Brajesh Choudhary, Baba Potukuchi

## Cryostat Design



### Membrane Cryostats

#### **GTT Membrane Technology IHI Membrane Technology** Stainless Steel Primary Plywood Membrane Membrane Sheet GRE Reinforced **Fire Proof Board** Polyurethane GRU Foam Secondary **Moisture Barrier** Membrane Reinforced Polyurethane Foam Plywood Concrete Bearing Vapor **GRE: Glass Cloth Reinforced Epoxy** Mastic Barrier GRU: Glass Cloth Reinforced urethane Nominal Panel Size Nominal Panel Size 3m x 8m x 2 mm thickness 1m x 3m x 1.2 mm thickness

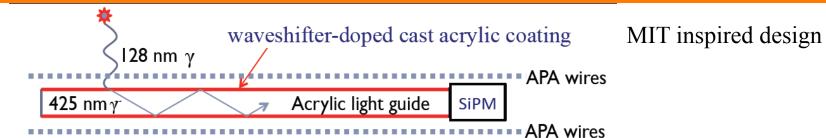
- Two 5-kton (fiducial mass) twin cryostats:
  - 18.8 kton total LAr mass
  - 13.5 kton total active mass
  - 10 kton total fiducial mass
- Estimated heat leak during steady state operations: 27.5 kW (per cryostat).
- No large vacuum hazard, no hazard from instantaneous heat leak change

## **Photon Detection**

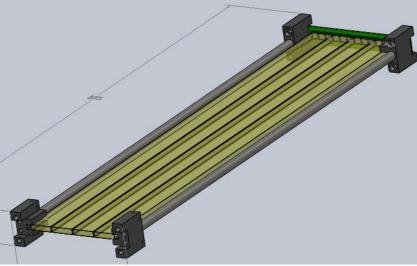
- The 1.4 ms drift time of the electrons is very slow.
  - The beam spill is  $10 \ \mu s \ long gives \ t0$
  - For non-beam related physics the interaction time is unknown. Use the light to determine the event time.
- A lot of light is produced by particles interacting in LAr
  - Both scintillation and Cherenkov radiation are generated but 5 times more scintillation light
    - 23% of the scintillation light is prompt (~6ns)
    - 77% of the light is late (~1.6  $\mu$ sec).
  - Prompt yield 33,000 128nm photons per MeV for a MIP
- Detecting the light and extracting the maximum information is complicated.
  - Scattering length is ~95 cm
  - Present coverage is 0.4%
- Photon Detection System critical for low-energy events

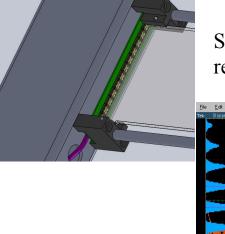
G.M. Seidel, R.E. Lanou, W. Yao, Rayleigh scattering in rare-gas liquids, NIM A489 (2002) 189.

## **Current Reference Design**

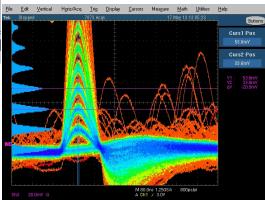


Bis-MSB coated cast acrylic 25 mm x 6 mm x 2.25 m bars 4 bars per paddle SiPM readout





SensL MicroFB-60035-SMT readout at LN2 temp

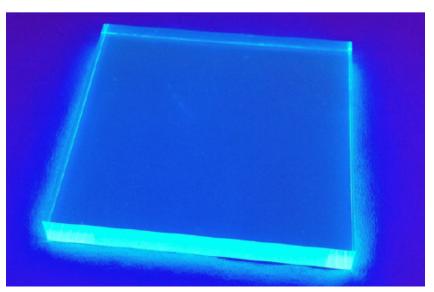


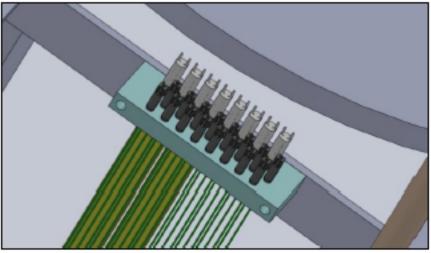
### Assemblies tested in Fall

# Outline

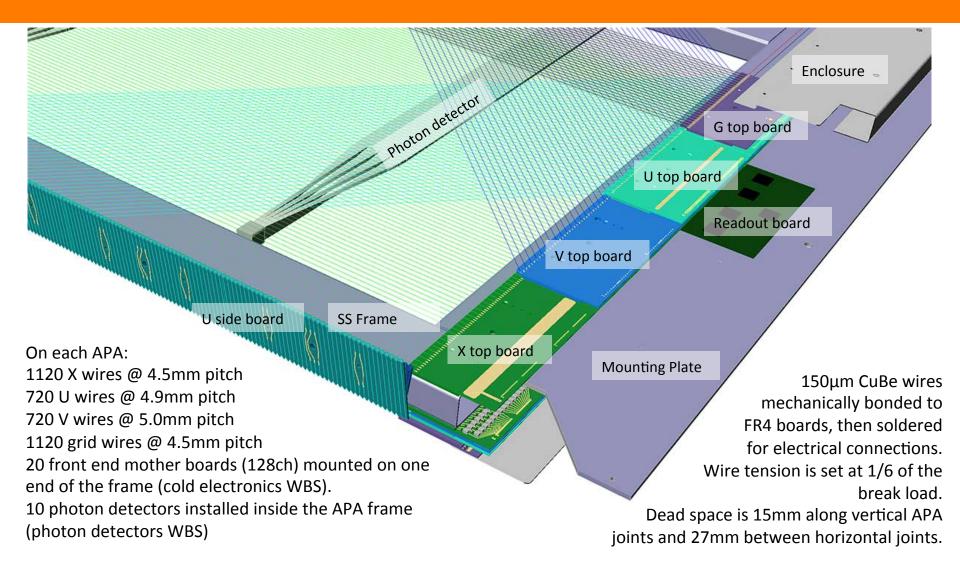
- Develop large area paddles with WLS doped in the bulk.
  - Only WLS near the surface interacts with 128 nm light

• Use bulk doped square fibers with small SiPM readout on each fiber.

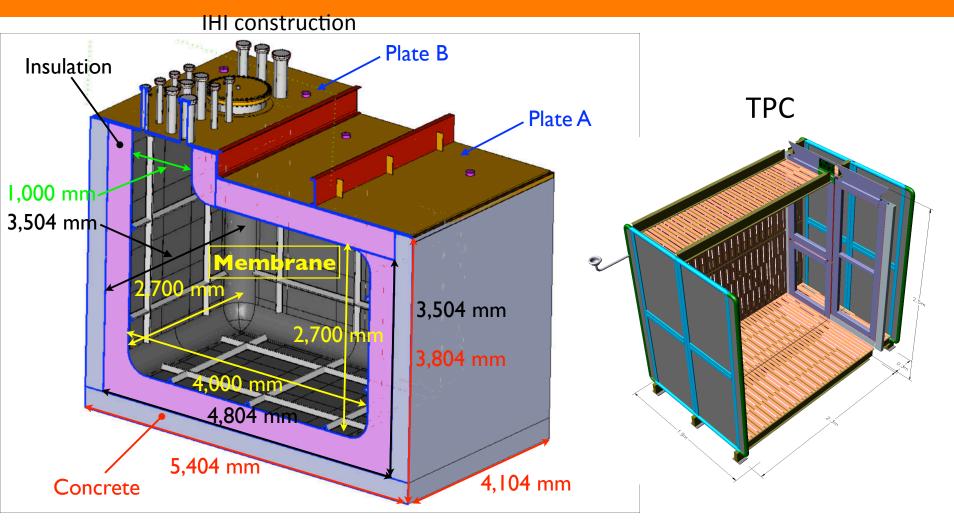




# Anode Plane Assembly Close-up View



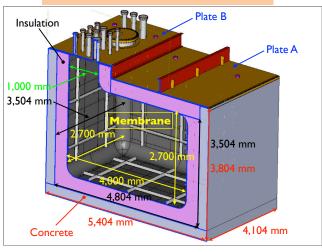
# 35 t Prototype Cryostat and TPC



• Test the cryostat this fall and install a TPC next year

## **Prototype Construction**

### 3 D Model of IHI Tank



### Concrete Structure @ PC4



### **Carbon Steel Vapor Barrier**







Two layers of foam (0.4 m)Top View of Two Layers Foam

SS membrane Insert Begins