

# *Neutrino mass measurement with the KATRIN experiment*

F. Glück\*

(for the KATRIN collaboration)

KIT: Karlsruhe Institute of Technology, Germany

Prague v13, 24 May 2013

\*e-mail: [Ferenc.Glueck@kit.edu](mailto:Ferenc.Glueck@kit.edu)



bmb+f - Förderschwerpunkt

Astroteilchenphysik

Großgeräte der physikalischen  
Grundlagenforschung

## Neutrino mass:

**Particle physics:** mass generation mechanism  
(fermion mass theory ?)

**Cosmology:** contribution to matter density, model input

## SM:

left-handed and massless neutrinos

**Neutrino oscillation experiments (solar, atmospheric, reactor, accelerator):**

non-zero mass-squared differences → at least two neutrinos have non-zero masses

**No information about absolute mass scale  
from neutrino oscillation experiments**

# Indirect neutrino mass determination

**Cosmology, neutrinoless double beta decay:**

**Sensitive to absolute neutrino mass scale (even below 0.1 eV possible in future)**

**But: model dependent**

**Example for  $00\nu\beta\beta$ :**

**It could be that the light massive neutrino mechanism is subdominant. Dominant mechanism could be f.e.: R-parity violating supersymmetry, right-handed couplings, leptoquark exchange, etc.**

**In that case neutrinoless double beta decay experiments give only upper limits for the absolute neutrino mass scale.**

**Also: sensitive to Dirac and Majorana phases  $\rightarrow$  cancellation possibility**

# Direct neutrino mass experiments

Electron energy spectrum close to endpoint of single beta decay:

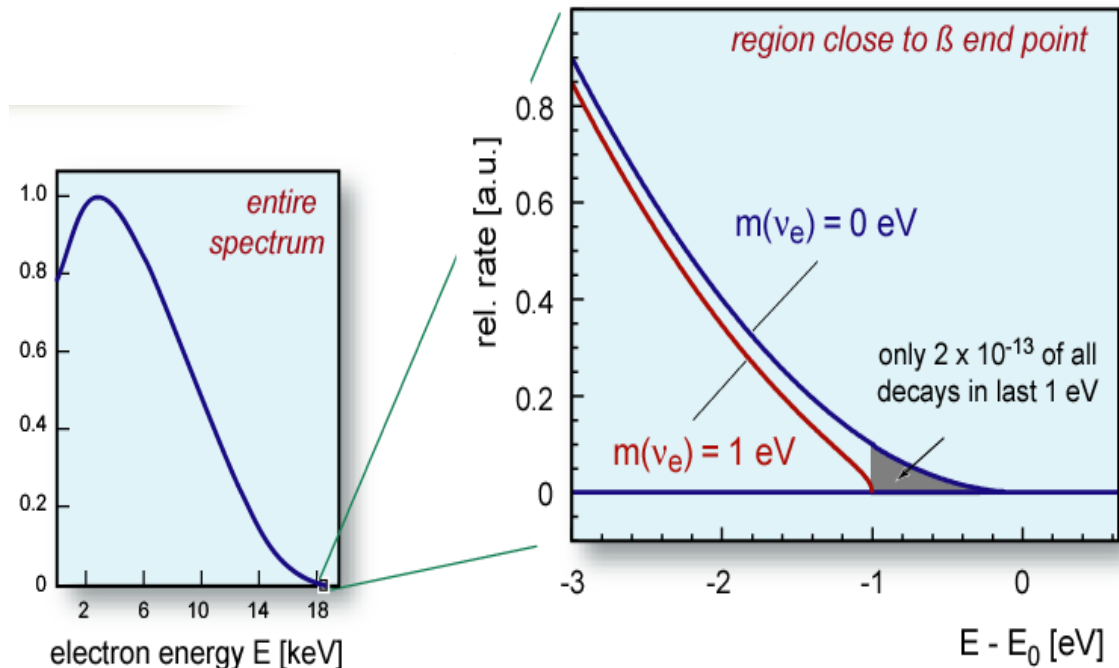
$$w(E_e) \sim E_e p_e E_\nu p_\nu$$

$$\sim E_e p_e (E_0 - E_e) \sqrt{(E_0 - E_e)^2 - m_\nu^2}$$



$T_2$  ( $E_0=18.6$  keV):

KATRIN



# Electron energy spectrum with general neutrino mixing:

$$\nu_e = \sum_i U_{ei} \nu_i$$

mass eigenstates

$$w(E_e) \sim E_e p_e (E_0 - E_e) \sum_i |U_{ei}|^2 \sqrt{(E_0 - E_e)^2 - m_{\nu i}^2} \Theta(E_0 - E_e - m_{\nu i})$$

lepton mixing matrix elements

**3 active neutrinos: small  $\Delta m_\nu^2$**

**→ effective electron neutrino mass definition:**

$$m_\beta^2 = \sum_{i=1}^3 |U_{ei}|^2 m_{\nu i}^2$$

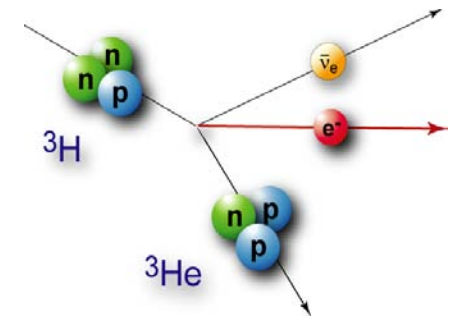
**Not sensitive to phases !**

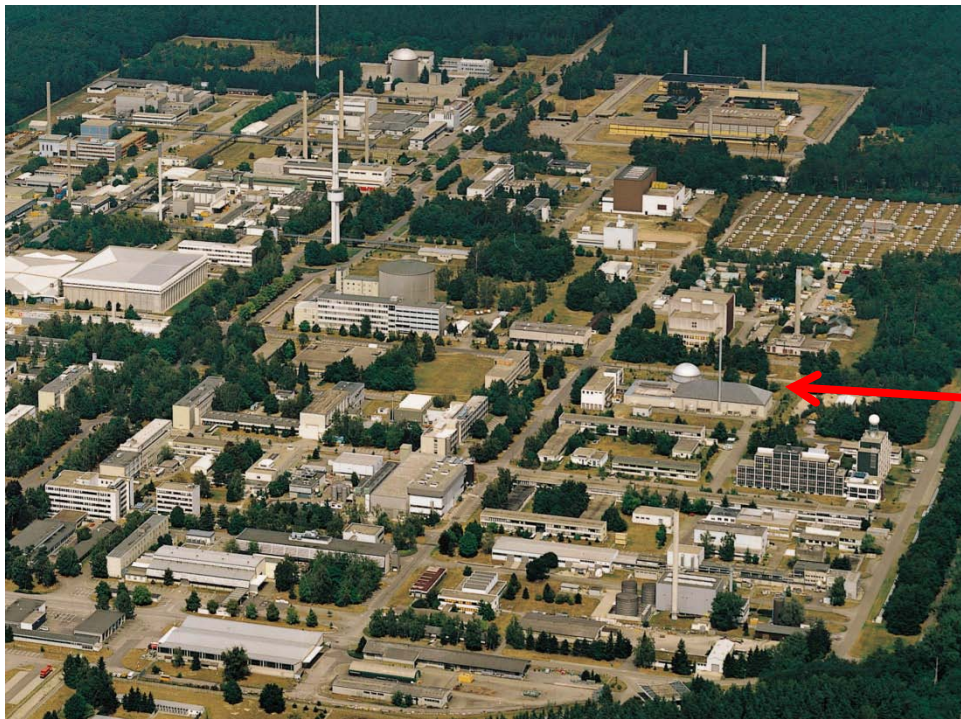
# KATRIN experiment

Absolute neutrino mass scale determination down to 0.2 eV, with small model dependence.

Integral electron energy spectrum measurement (MAC-E filter method), close to endpoint of molecular tritium beta decay.

$^3\text{H}$ : super-allowed	
$E_0$	18.6 keV
$t_{1/2}$	12.3 y





**KIT (Karlsruhe Institute of Technology) campus north, Germany**

**KATRIN: at TLK  
(Tritium Laboratory Karlsruhe)**

**Collaboration:  
Germany, USA, Russia,  
Czech Republic, Great Britain**



*University of Washington*



**Fachhochschule Fulda**  
University of Applied Sciences



## Advantages of tritium:

- **Superaligned transition:** → matrix element  $M$  is not energy dependent
- **Low endpoint energy:** → relative decay fraction at the endpoint is comparatively high
- **Short half life:**
  - specific activity is high
  - low amount of source material
  - low fraction of inelastic scattered electrons
- **Hydrogen isotope:**
  - simple atomic shell
  - final states precisely calculable



# Status of previous tritium experiments

*Mainz & Troitsk have reached their intrinsic limit of sensitivity*



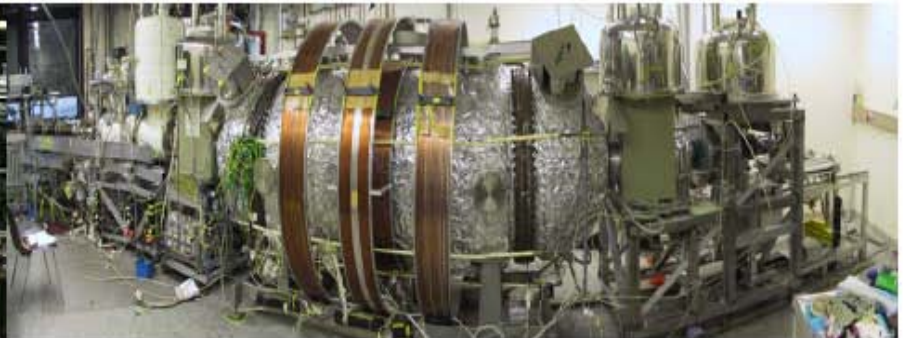
Troitsk

windowless gaseous  $T_2$  source

analysis 1994 to 1999, 2001

$$m_\nu^2 = -2.3 \pm 2.5 \pm 2.0 \text{ eV}^2$$

$$m_\nu \leq 2.2 \text{ eV (95\% CL.)}$$



Mainz

quench condensed solid  $T_2$  source

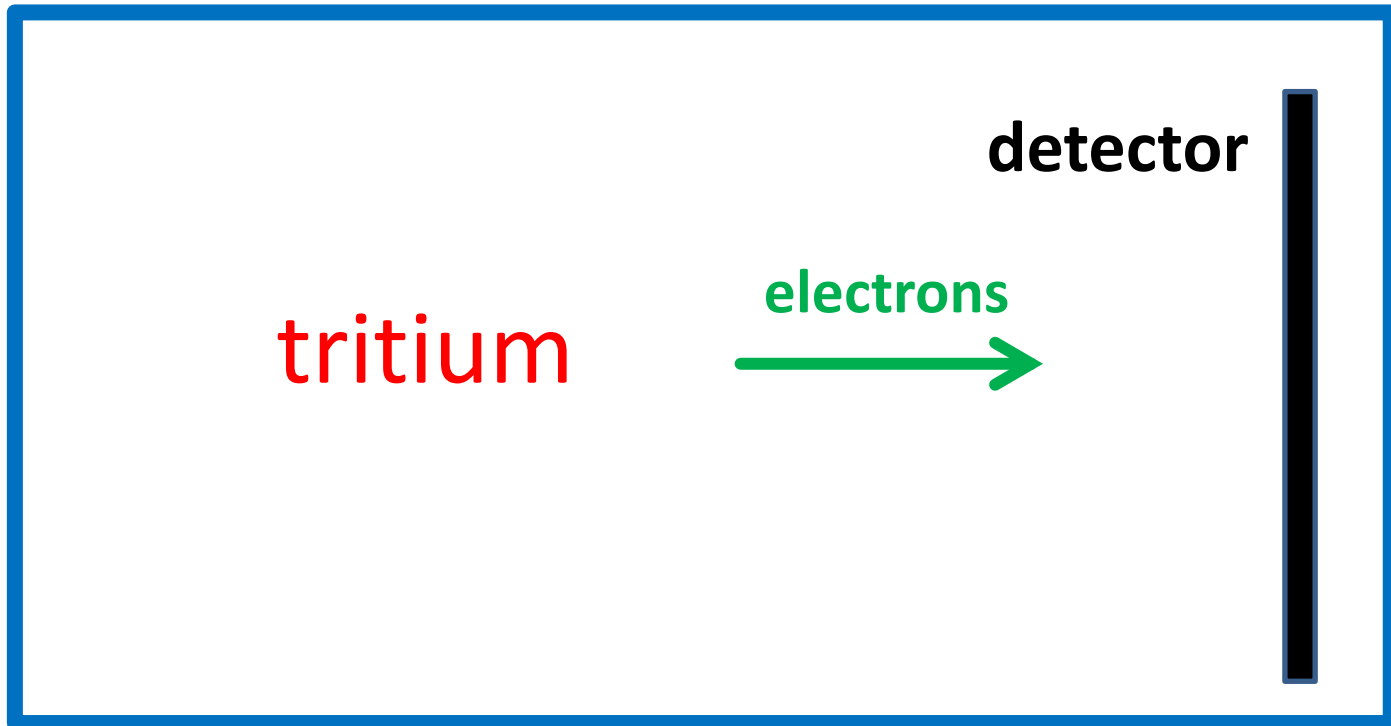
analysis 1998/99, 2001/02

$$m_\nu^2 = -1.2 \pm 2.2 \pm 2.1 \text{ eV}^2$$

$$m_\nu \leq 2.2 \text{ eV (95\% CL.)}$$

**Goal of KATRIN: factor 10 improvement**

# Simple energy spectrum measurement method



## Problems:

- no electron detector with 1 eV resolution
- extremely large rate

# The $\beta$ -Spectrum of $H^3$

G. C. HANNA AND B. PONTECORVO

*Chalk River Laboratory, National Research Council of Canada,  
Chalk River, Ontario, Canada*

January 28, 1949

**T**HE proportional counter technique previously described<sup>1,2</sup> has been used to study the  $\beta$ -spectrum of  $H^3$  an investi-

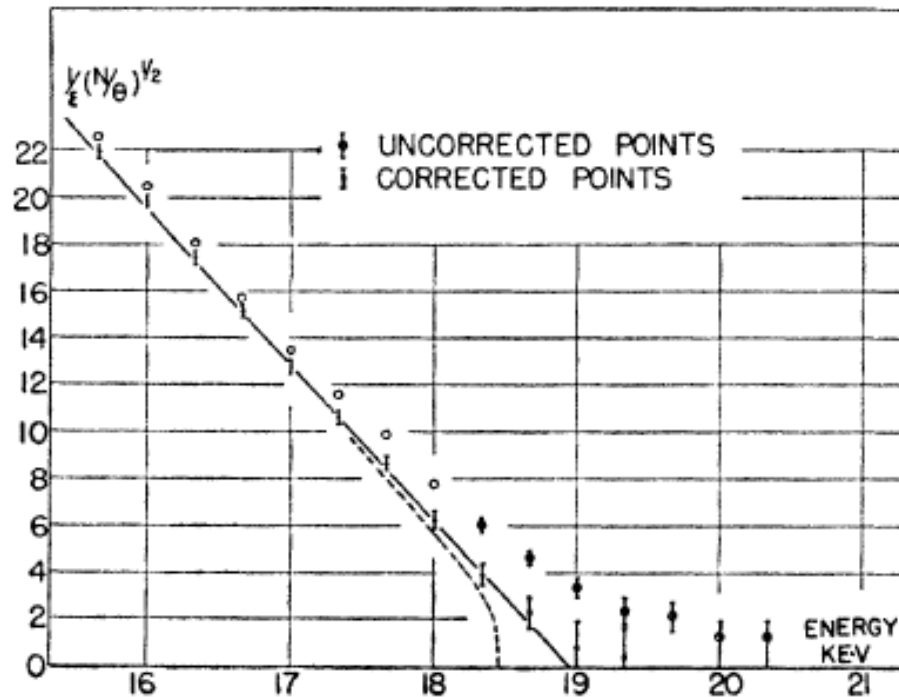
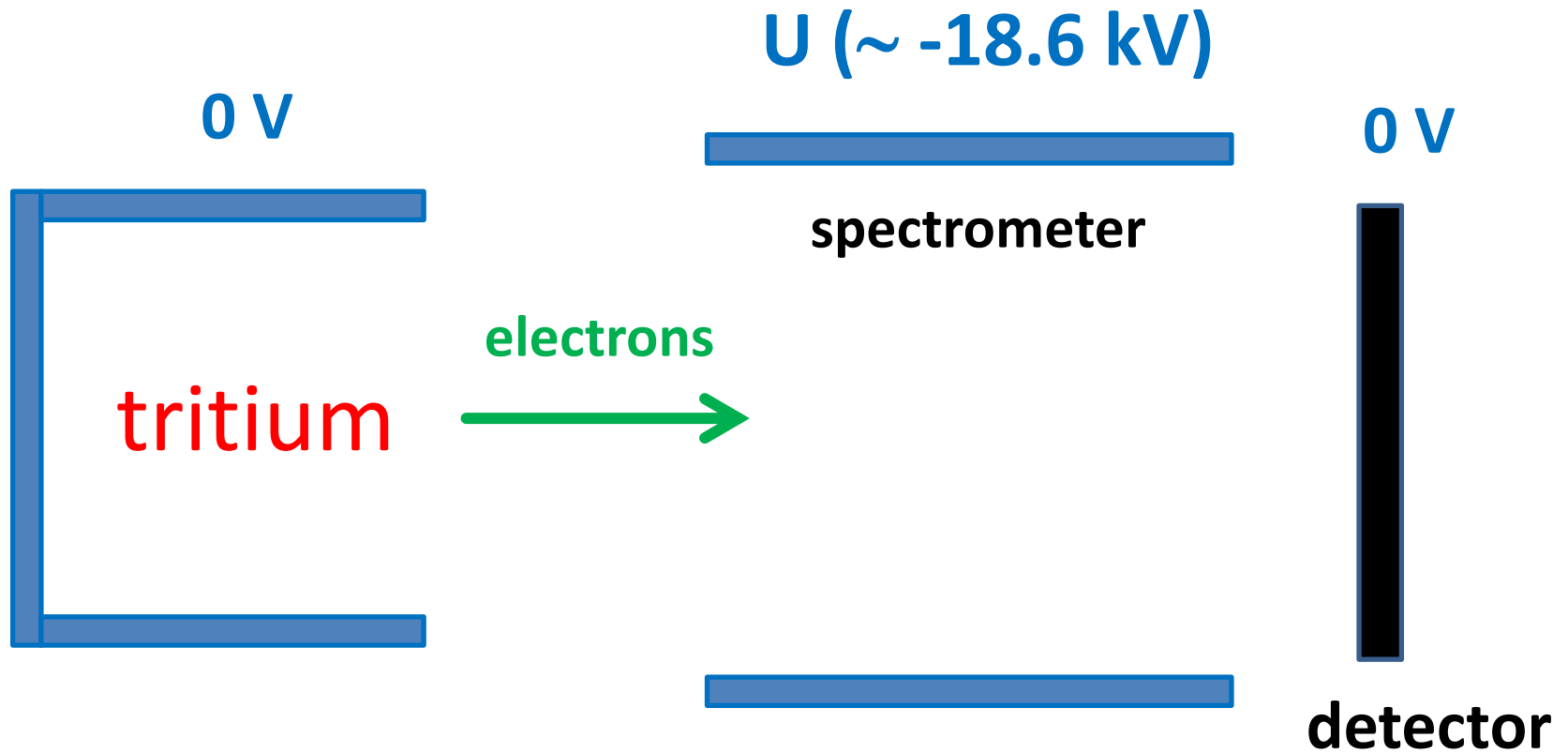
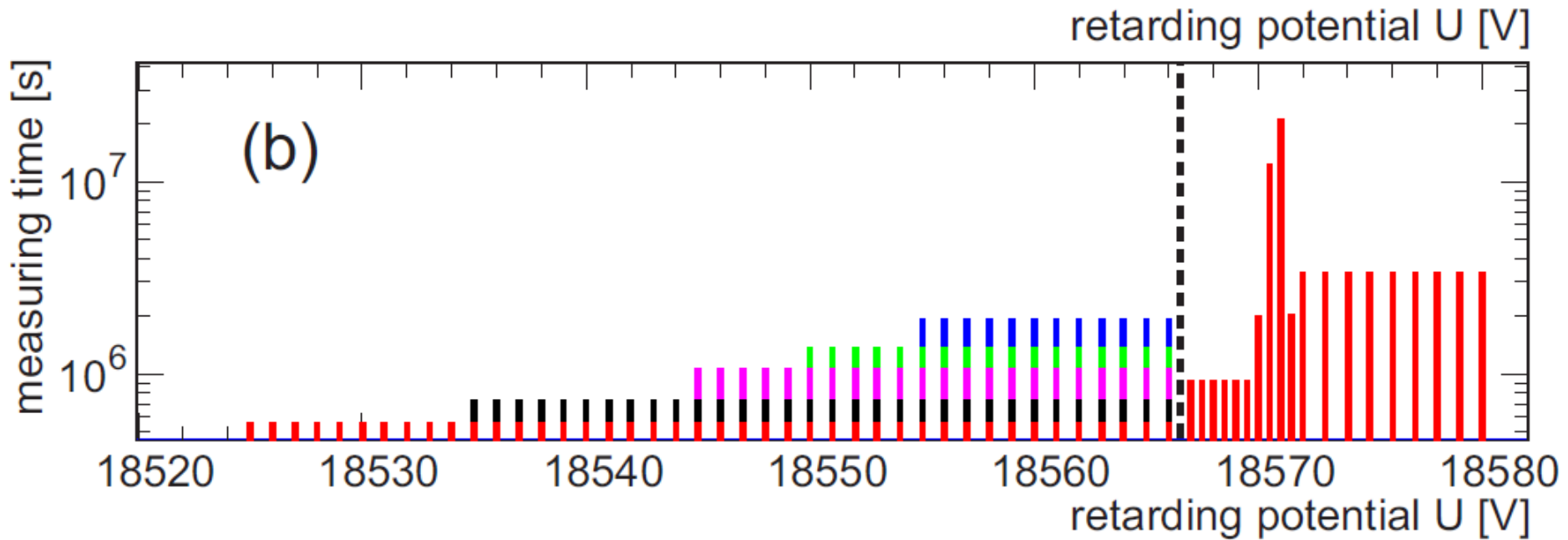
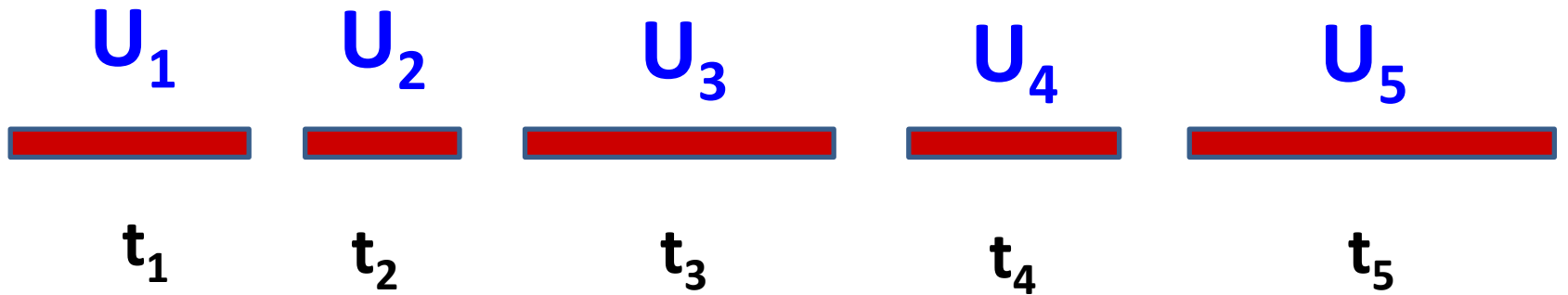


FIG. 2. "Kurie" plot of the end of the  $H^3$  spectrum. The theoretical curve (shown dotted) corresponding to a finite neutrino mass of 500 eV (or 1 keV —see text) has been included for comparison.

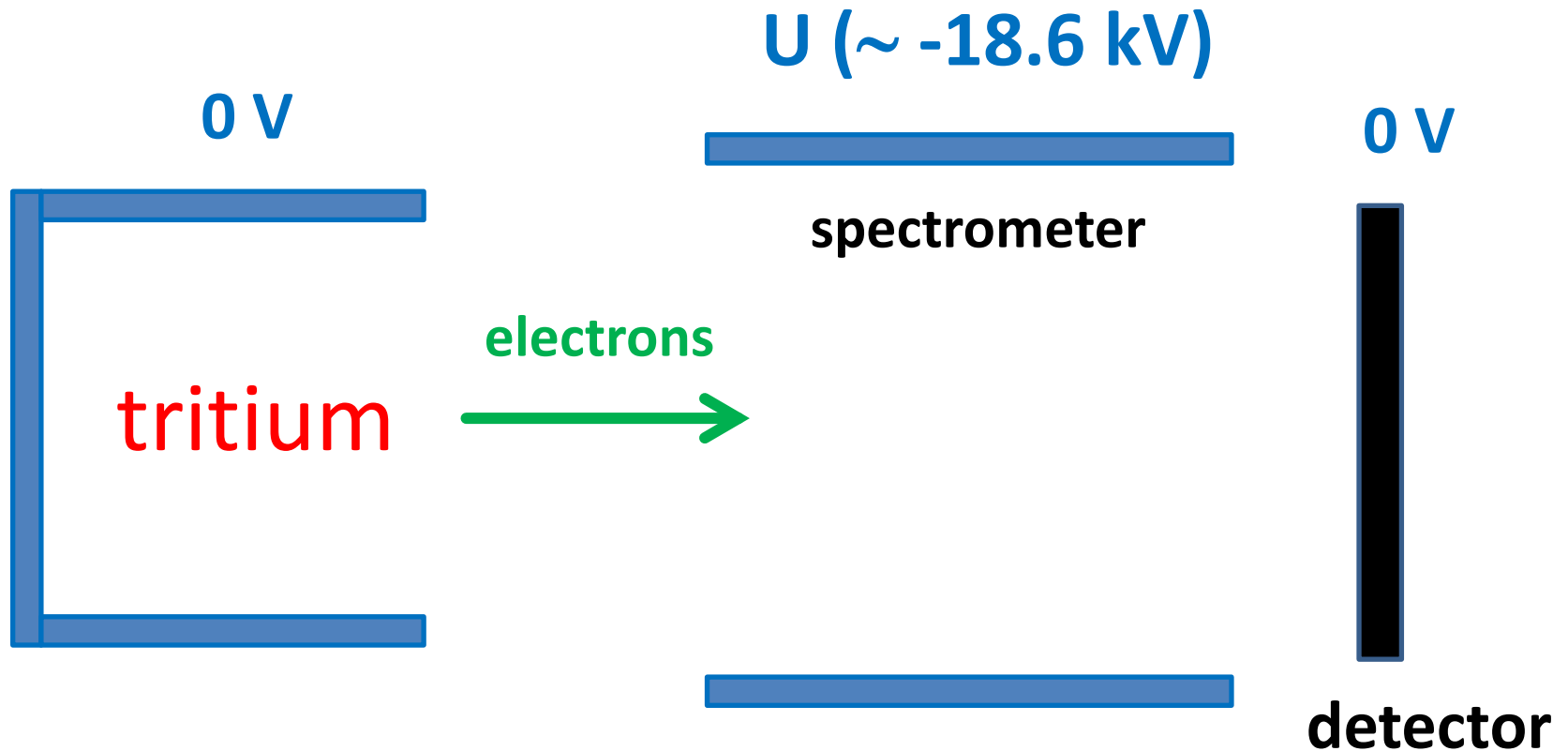
# Integral energy spectrum measurement with spectrometer at negative potential between source and detector





Measured number of events:  $N_1, N_2, N_3, \dots \rightarrow$  integral spectrum;  
 neutrino mass determination by fit

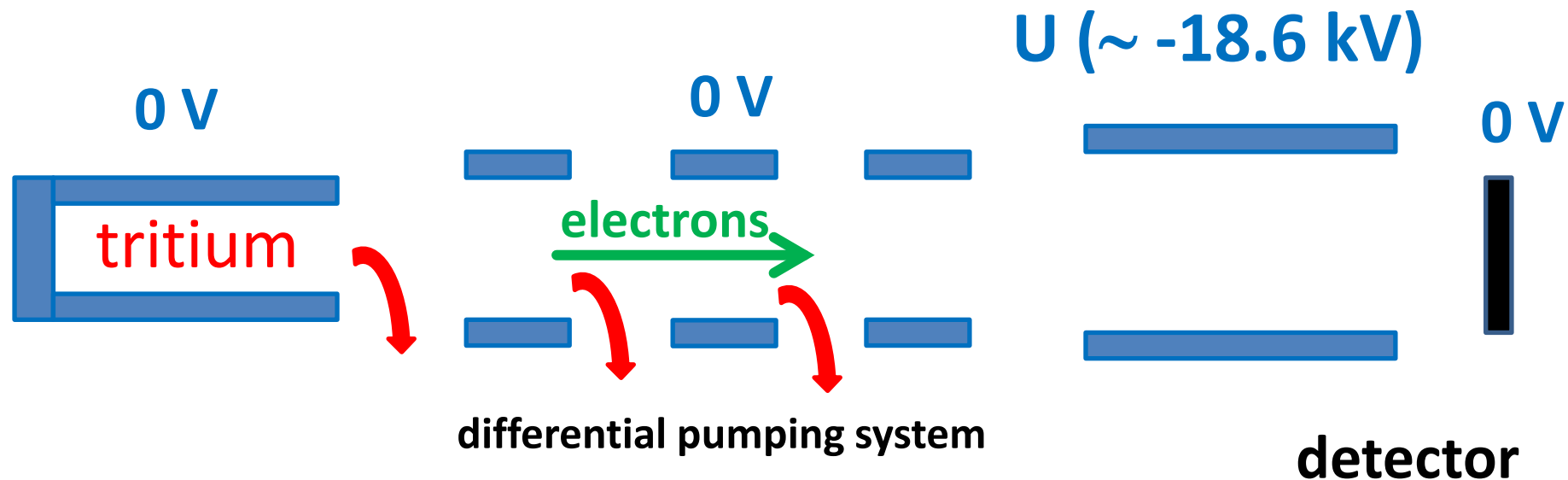
# Integral energy spectrum measurement with spectrometer at negative potential between source and detector



## Problem:

tritium beta decays inside and beyond the negative potential region (no material window allowed between source and spectrometer: energy loss !)

## Reducing the tritium flow into the spectrometer by differential pumping system (large distance between source and spectrometer)

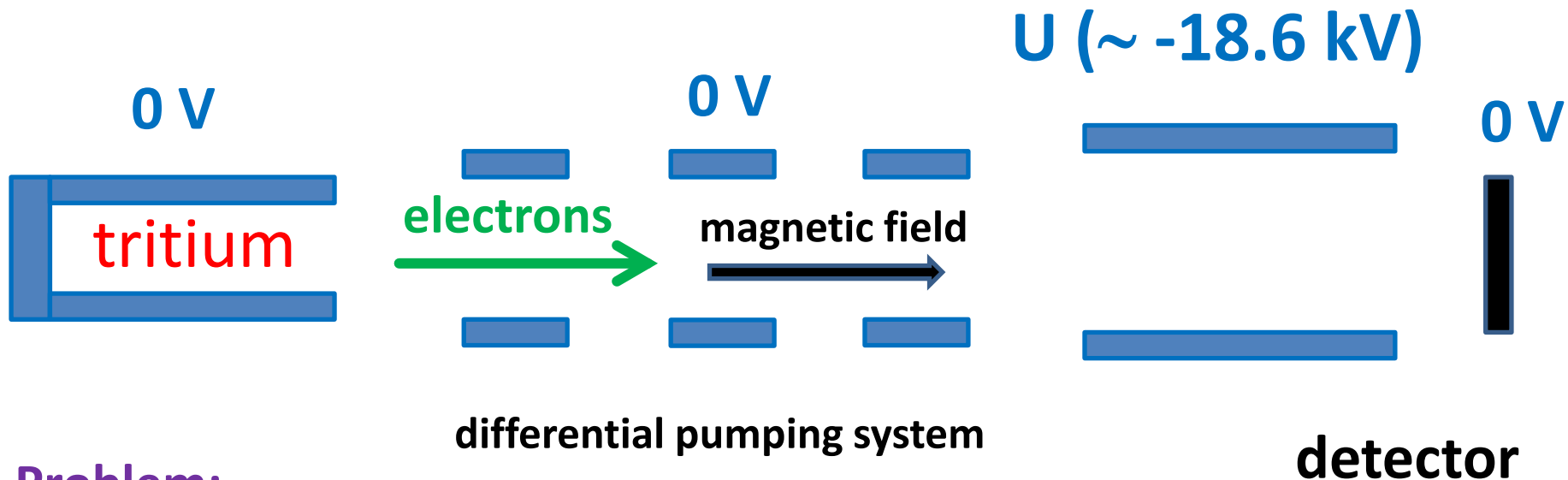


### Problem:

Most of the beta decay electrons do not get to detector (small solid angle)

# Guiding the beta electrons by magnetic field

Electrons follow the magnetic field lines  
(in adiabatic approximation)



## Problem:

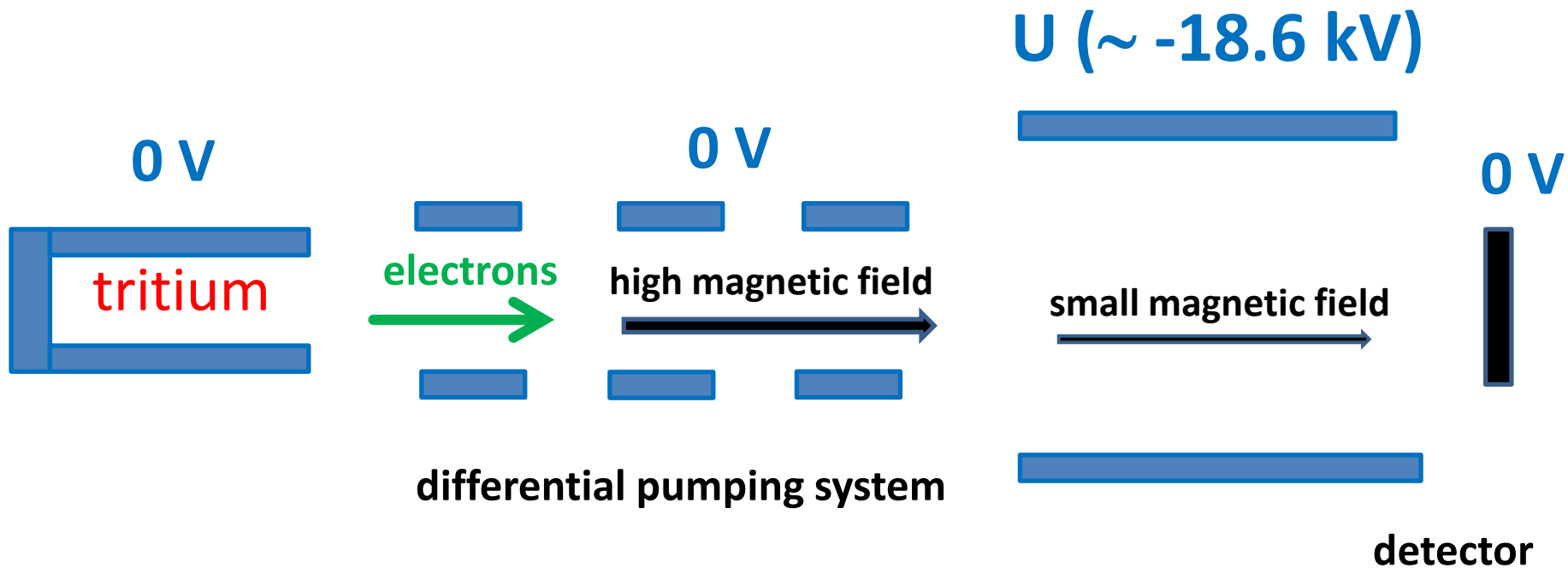
Electric field can change only the longitudinal energy  
→ most of the electrons with larger transversal energy do not get to detector



# Converting the transversal energy into longitudinal energy by inverse magnetic mirror effect

Adiabatic approx.:

orbital magnetic moment  $E_{trans} / B$  is constant



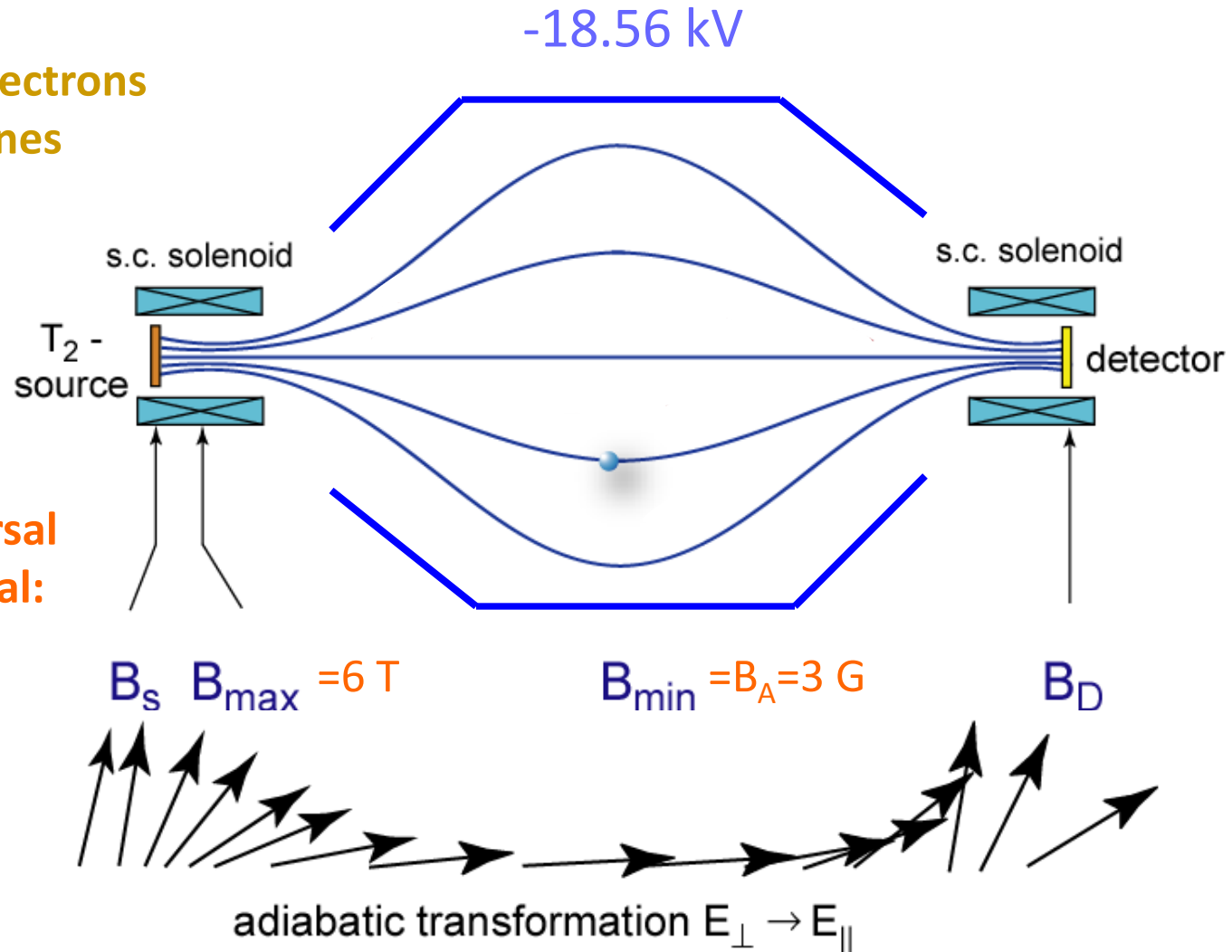
# principle of an electrostatic filter with magnetic adiabatic collimation (MAC-E)

Adiabatic guiding of electrons along magnetic field lines

Energy analysis: electric field can change longitudinal energy

Conversion of transversal energy into longitudinal: by inverse magnetic mirror effect

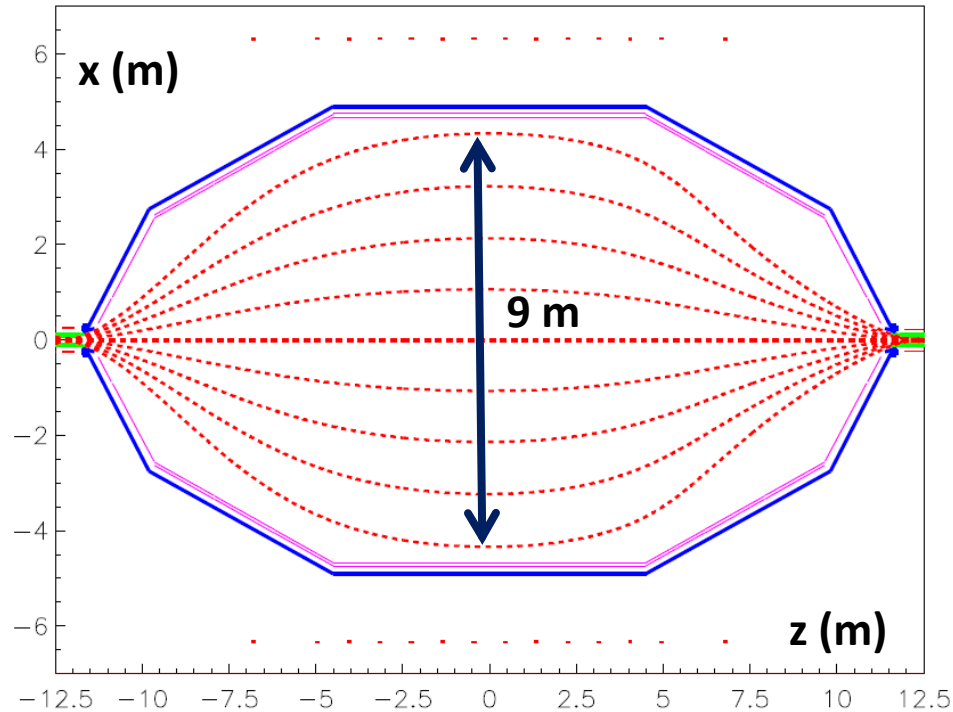
Constant magnetic flux,  $B_{\min}$  small: large spectrometer diameter (10m)



**Flux tube** : defined by magnetic field lines going through the source

Magnetic flux constant in flux tube:  $d_S^2 B_S = d_A^2 B_A$

**KATRIN:**  $B_S / B_A \approx 10000$        $d_S = 9 \text{ cm}$        $d_A = 9 \text{ m}$



**MAC-E filter:** high energy resolution (no tail at the low energy part),  
high luminosity, low background

gaseous T2, 30 K

Main

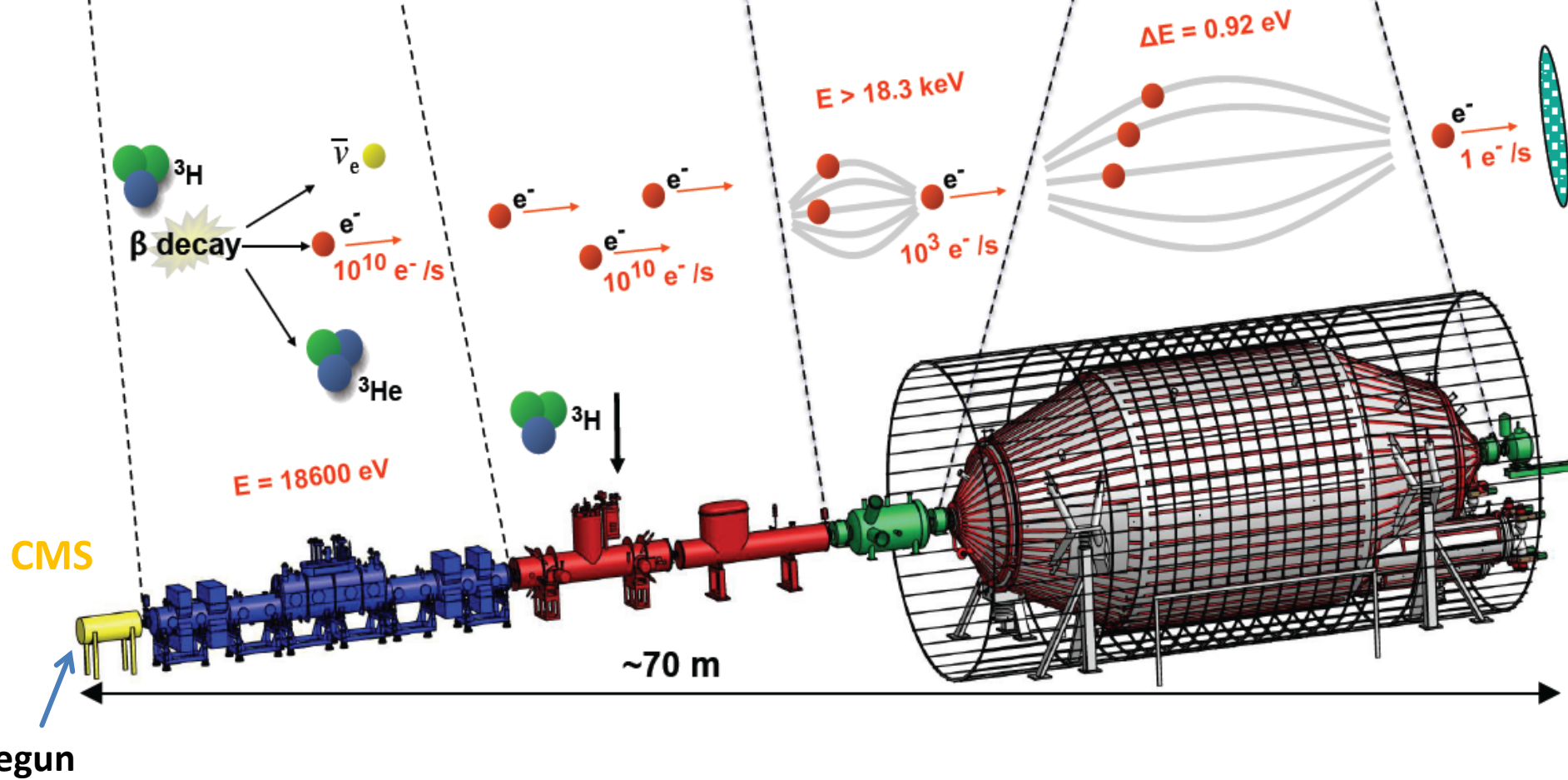
Tritium source

Transport section

Pre spectrometer

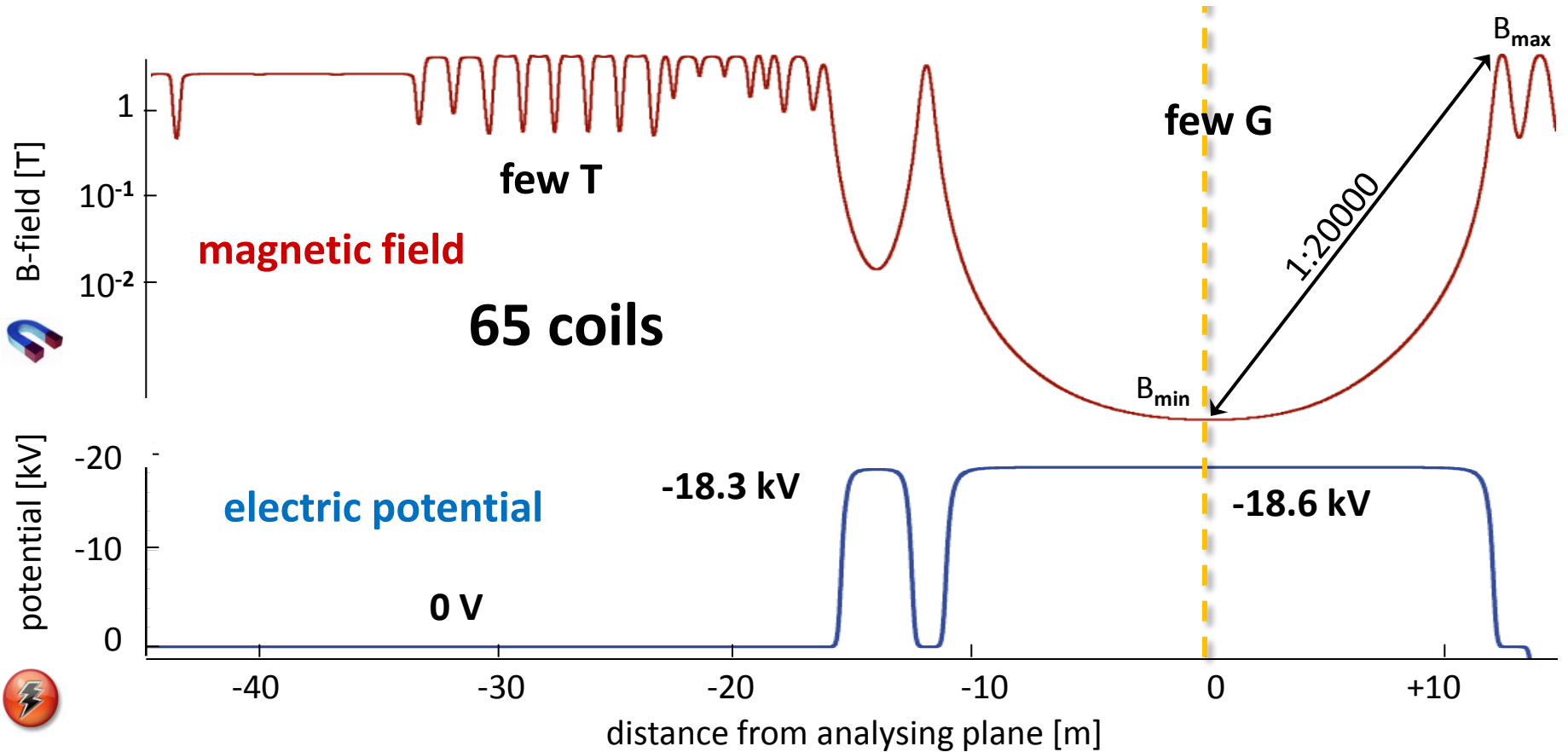
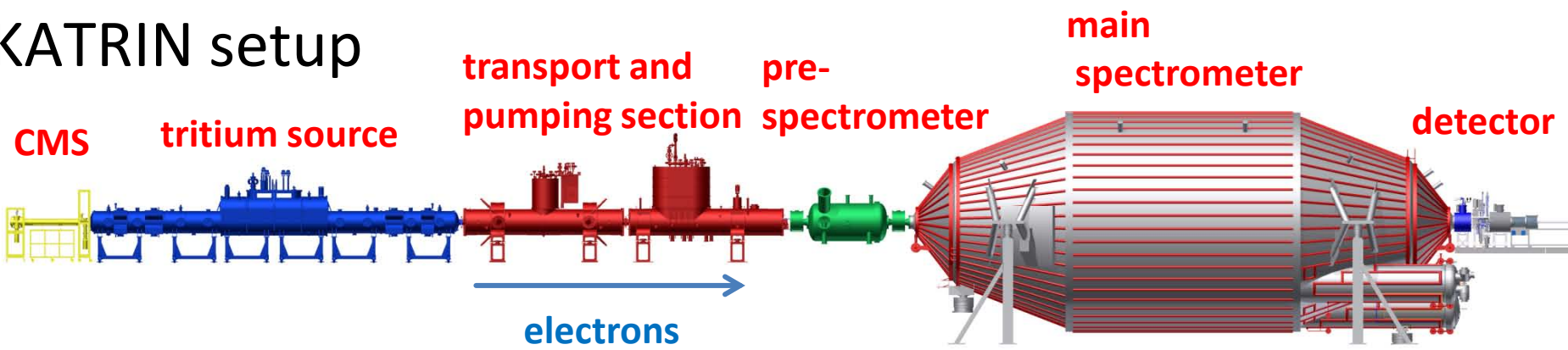
Spectrometer

Detector



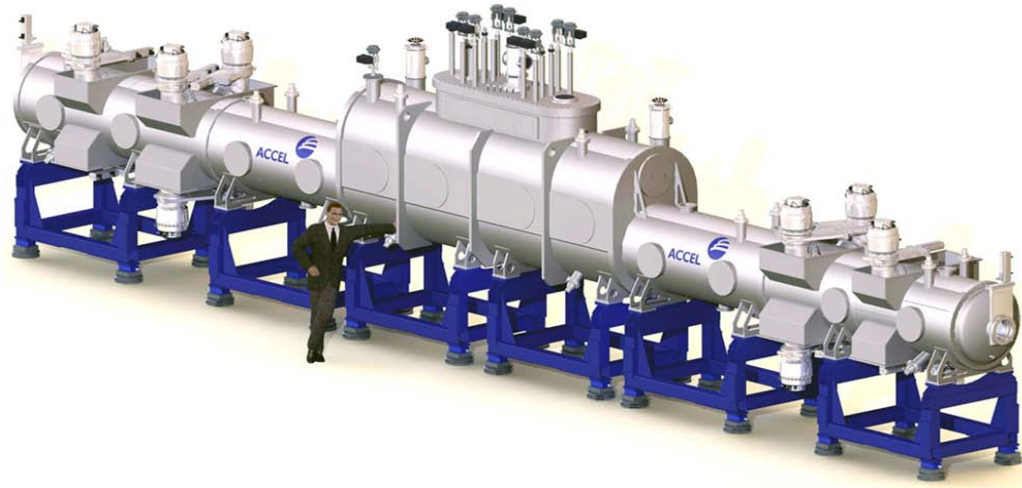
KATRIN system

# KATRIN setup



# WGTS

## Windowless Gaseous Tritium Source



**7 sc. magnet modules (3.6 T, 5.6 T); axisymmetric + dipole coils**

**12 turbomolecular pumps**

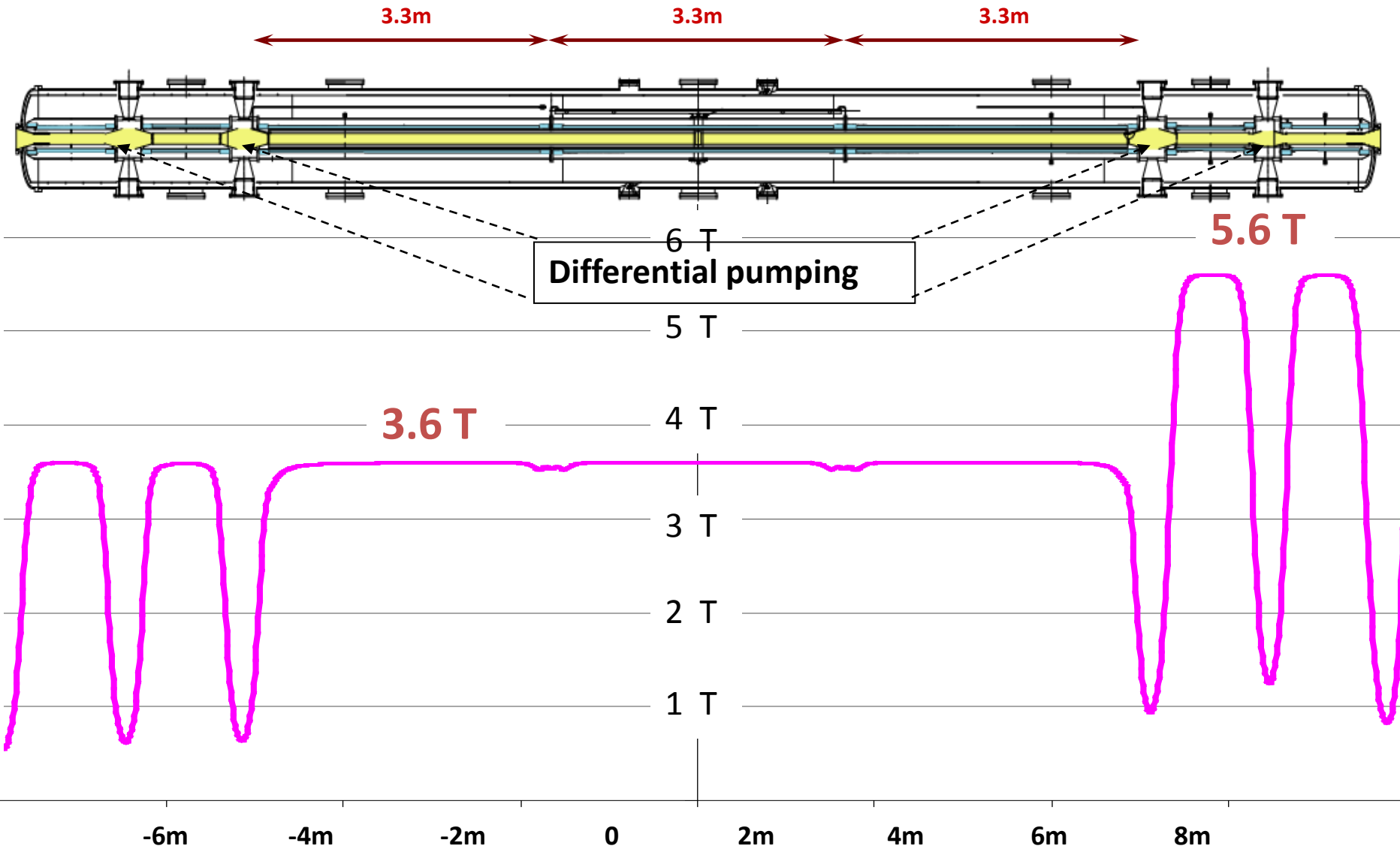
**Tritium gas temperature: 30 K**

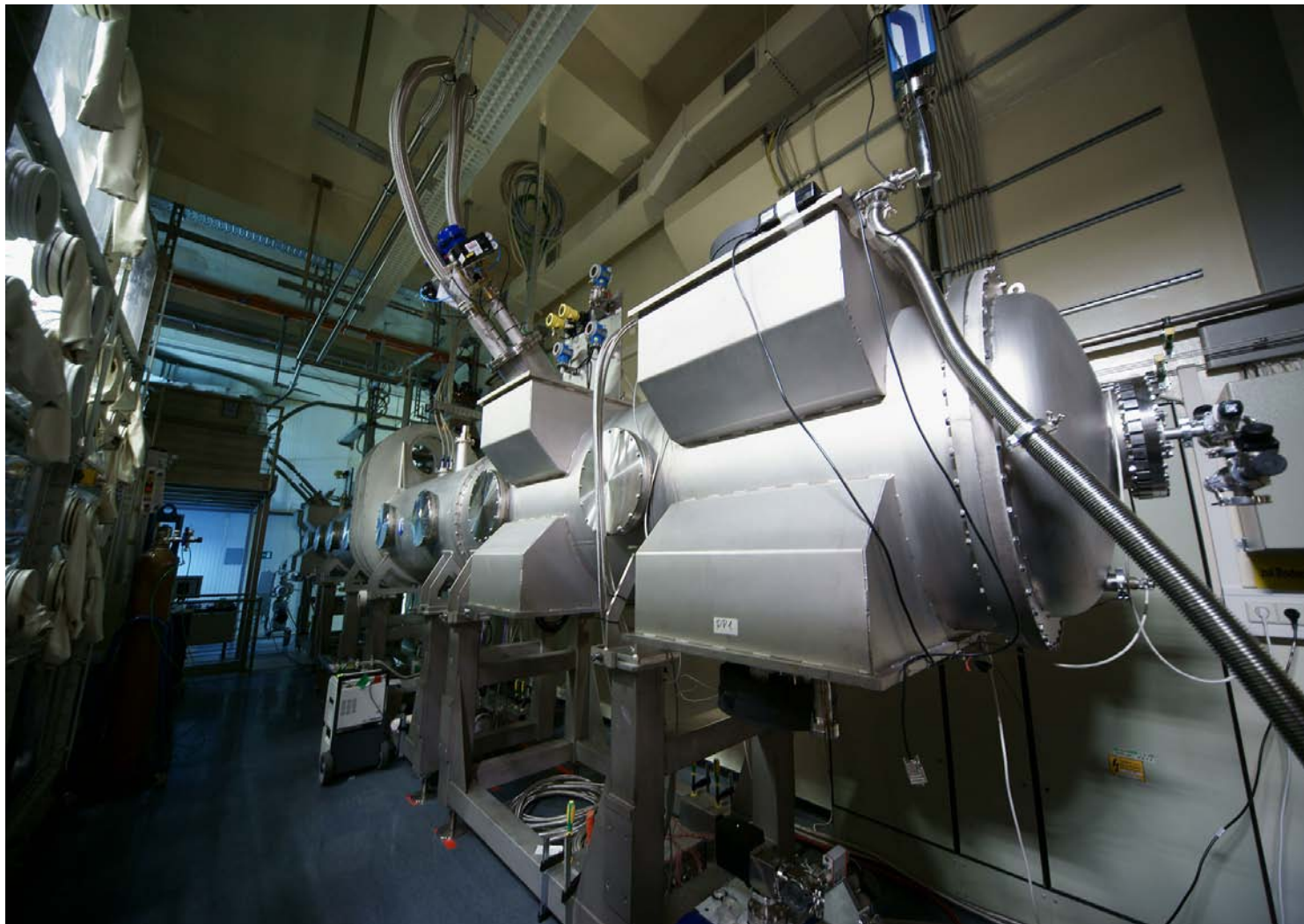
**Tritium purity: > 95 %**

**Column density:  $5 \times 10^{17}$  mol/cm<sup>2</sup>**

**Beta decay rate:  $10^{11}$ /s**

# WGTS – magnetic field





**WGTS Demonstrator at KIT-TLK: to test cooling concept  
(without sc magnets)**

**Temperature stability: 3 mK  
(requirement: 30 mK)**



# Plasma effects

## Beta decay and ionization:



a lot of electrons and  
ions in tritium source  
→ plasma



## From simulation:

secondary electrons cool down to gas  
temperature → average electron energy:  
few meV → space charge potential cannot  
be much larger than few mV → systematic effect  
to neutrino mass probably not significant

## Experimental information:

- ${}^{83m}Kr$  in tritium source (change of Kr lines due to plasma effects)
- egun electrons through tritium source (change of energy due to plasma instabilities)

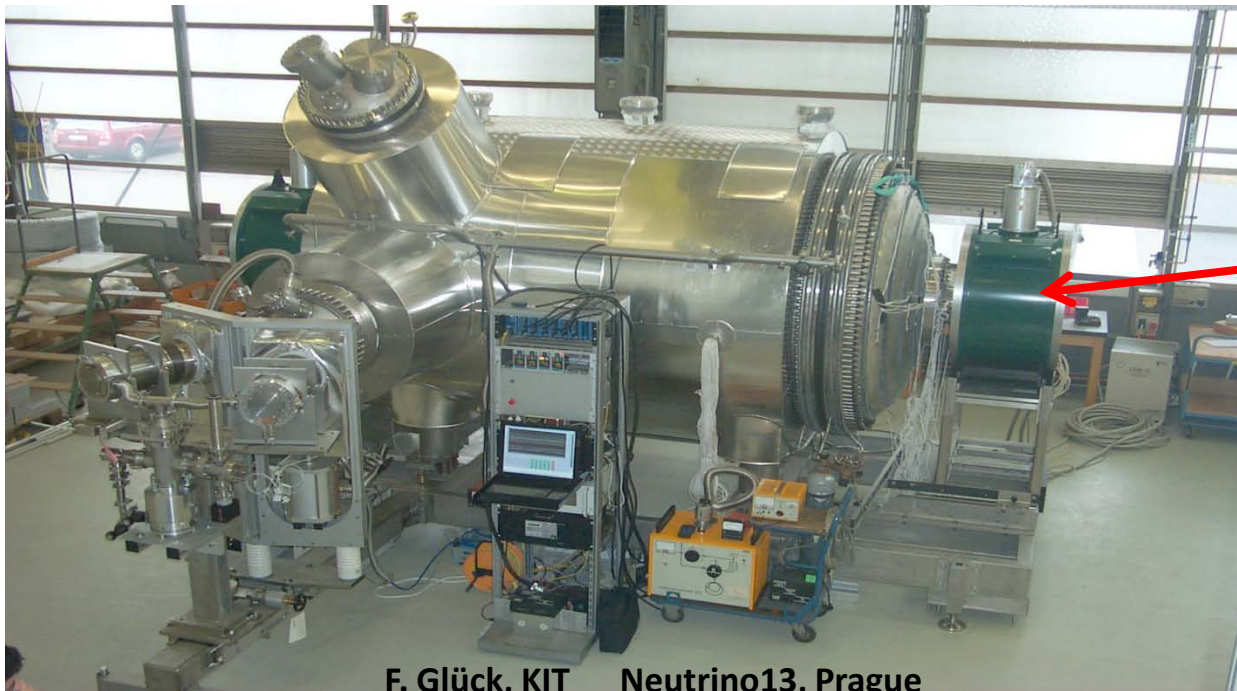
# Pre-spectrometer

- Optimize the background by its potential
- Additional tritium pump
- Prototype for main spectrometer (important !)

**Delivered : 2003**

**Length: 3.4 m, diameter: 1.5 m**

**First measurements: 2006**



**Sc. magnet  
max.  
4.5 T**

## First pre-spectrometer measurements:

**Strong Penning discharge:**

increasing magnetic field →

pressure and leakage current increase, electric breakdown

**Caused by deep (few kV) Penning traps**

**Eliminating the Penning traps (by a new shielding electrode)**

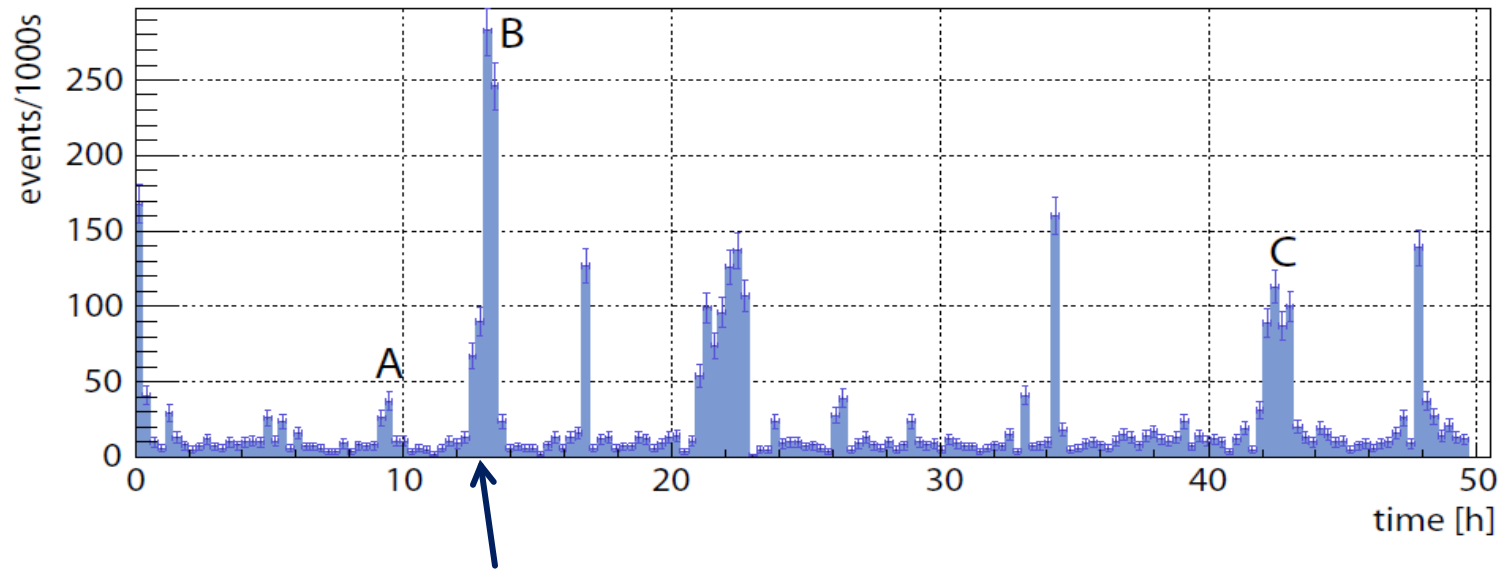
**→ strong Penning discharge disappeared !**

**Measurements with detector: large background (few 100 Hz)  
at high magnetic field**

**Caused again by Penning traps (at the ground electrode, few mm  
dimensions)**

**New ground electrode: no Penning traps → background reduction  
by factor 100000 !**

## Background from radon decays



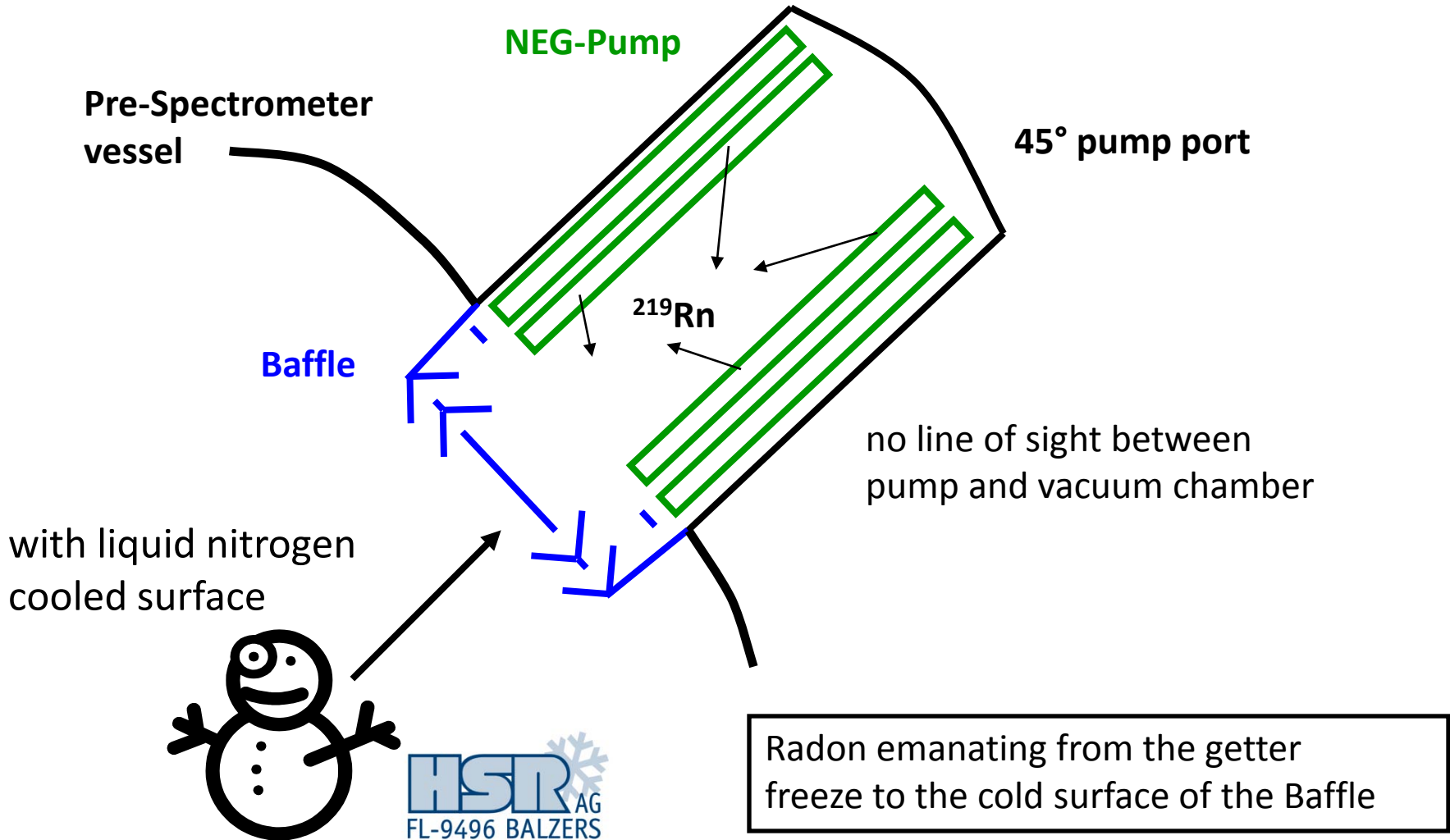
many secondary electrons  
from 1 high energy primary electron

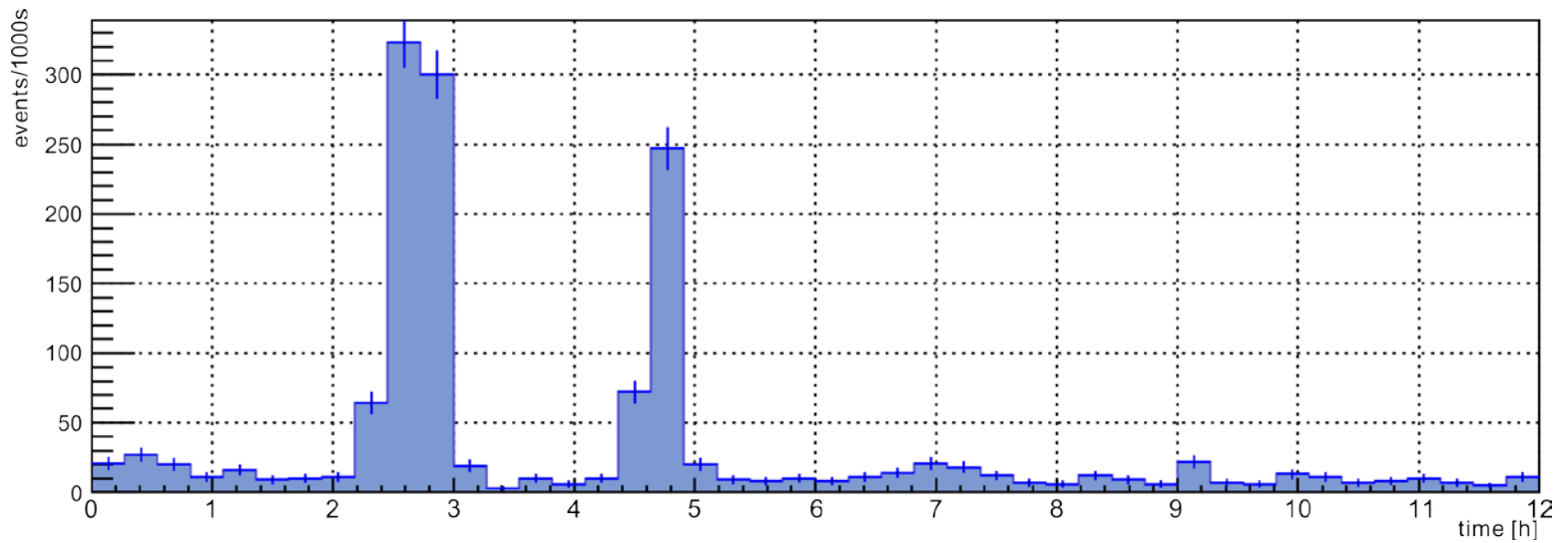
- background from high energy stored electrons
- high energy electrons come from Rn decays
- Rn comes from getter

Background reduction  
(from few 10 mHz to few mHz):

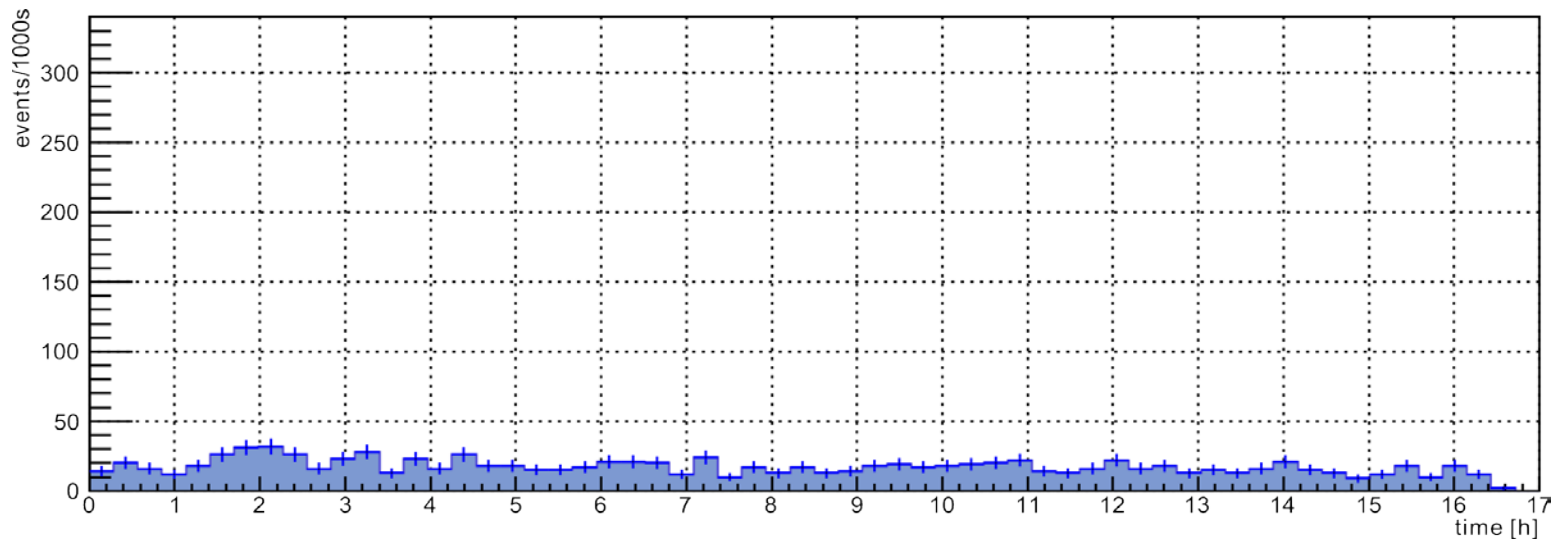
- removing getter from pump ports
- cooled baffle, preventing radon transport from getter into spectrometer

# Baffle Setup at Pre-Spectrometer





**Background in pre-spectrometer with warm baffle**



**Background in pre-spectrometer with cold baffle**

# KATRIN main spectrometer





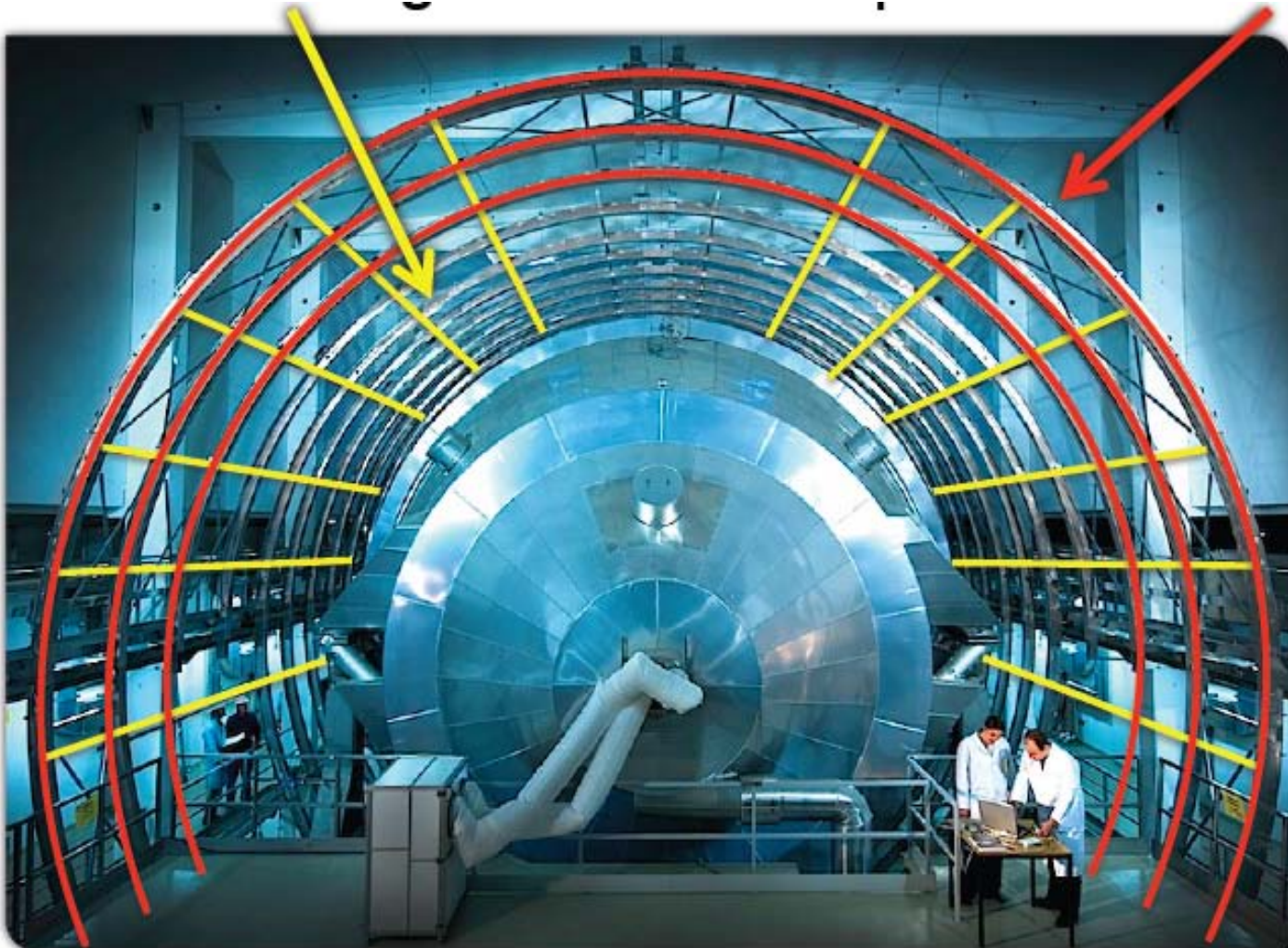
**24 m long, 10 m diameter, stainless steel: 3 cm thick, 200 t.  
Vacuum inside:  $10^{-11}$  mbar.**



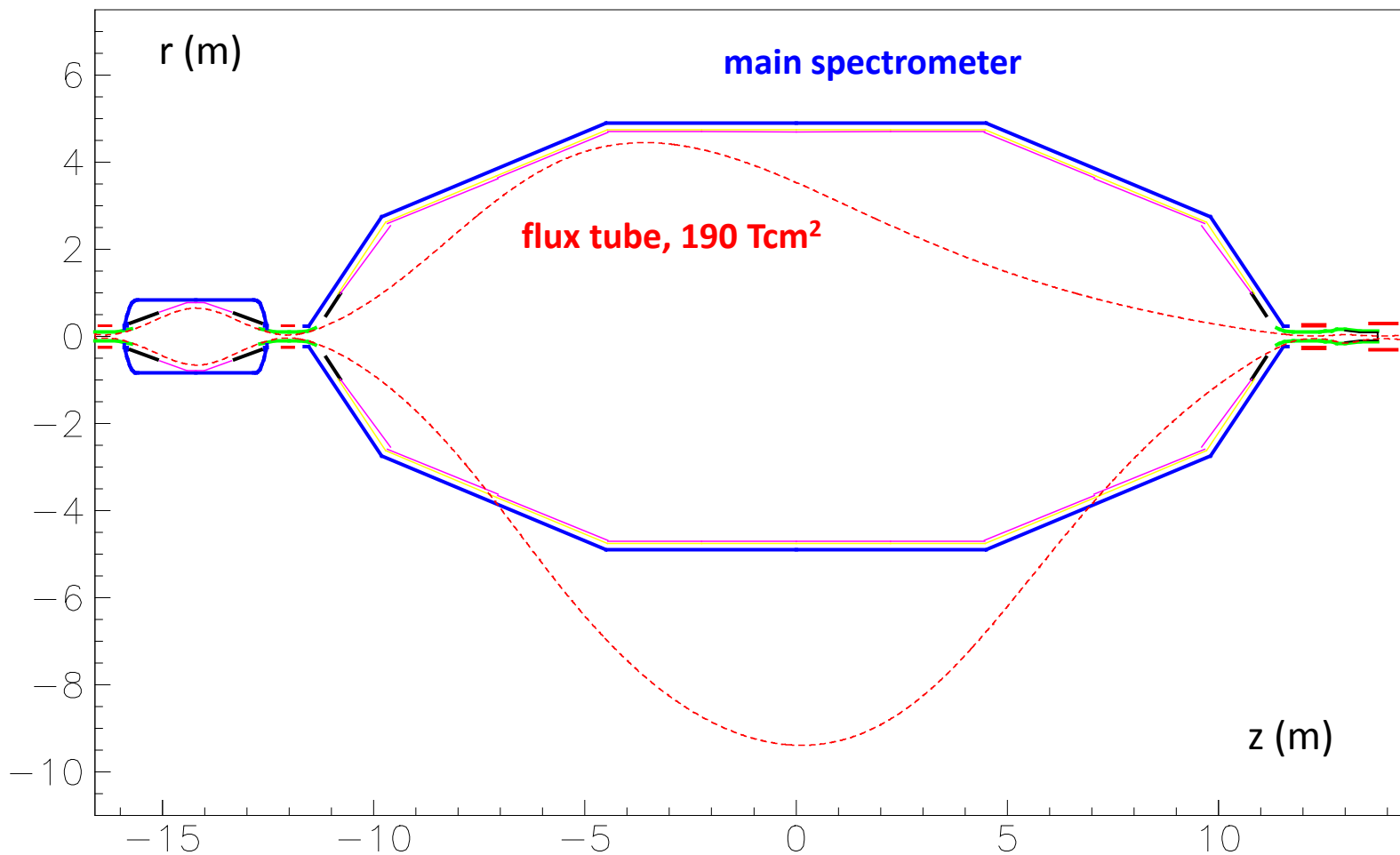
# Air coils

earth magnetic field compensation coils

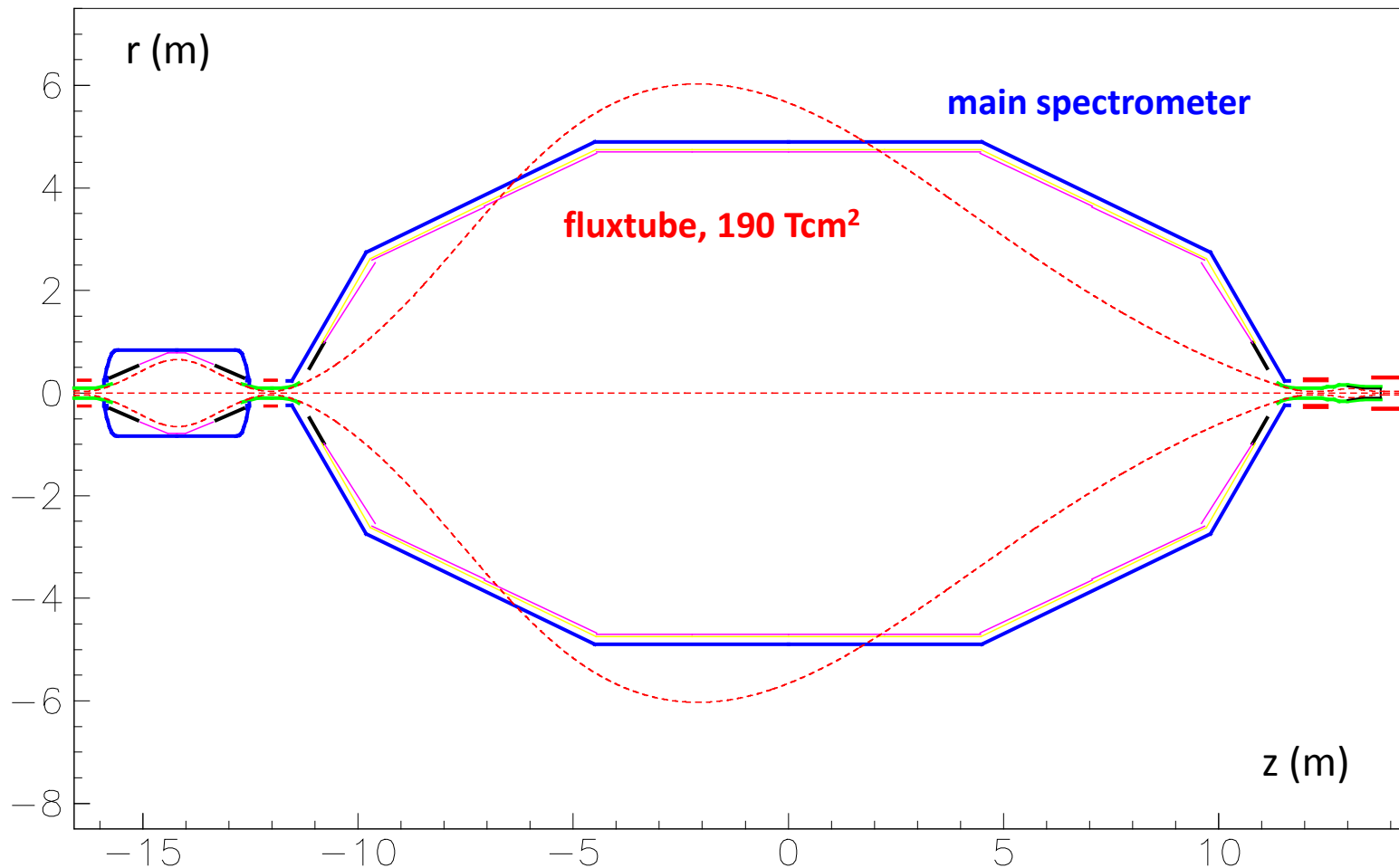
axisymmetric air coils



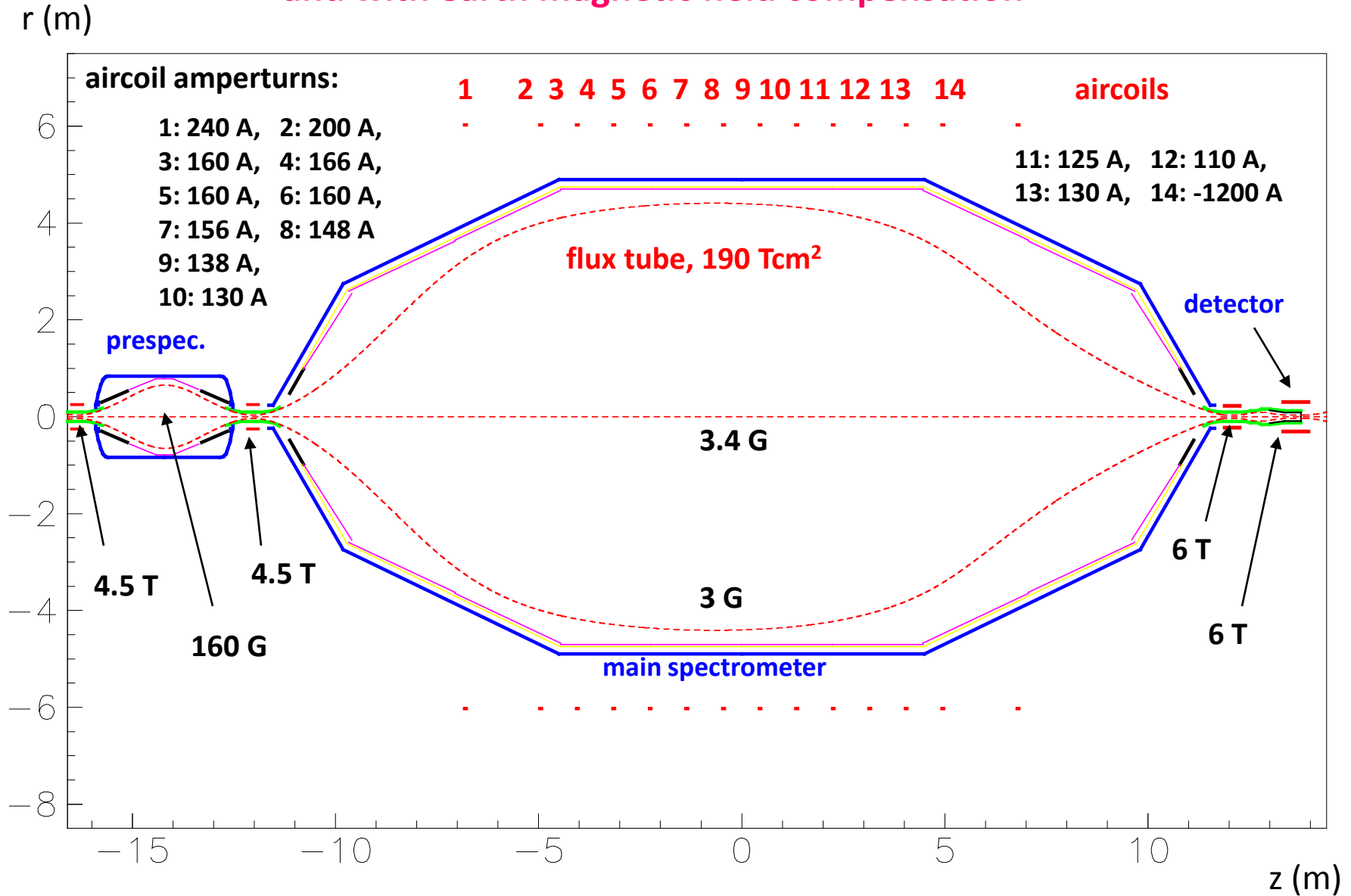
# Flux tube magnetic field lines without air coils: shift of field lines due to earth magnetic field



**Flux tube magnetic field lines without axisymmetric air coils,  
but with earth magnetic field compensation:  
axisymmetric air coils are also needed**



# Flux tube magnetic field lines with axisymmetric air coils and with earth magnetic field compensation

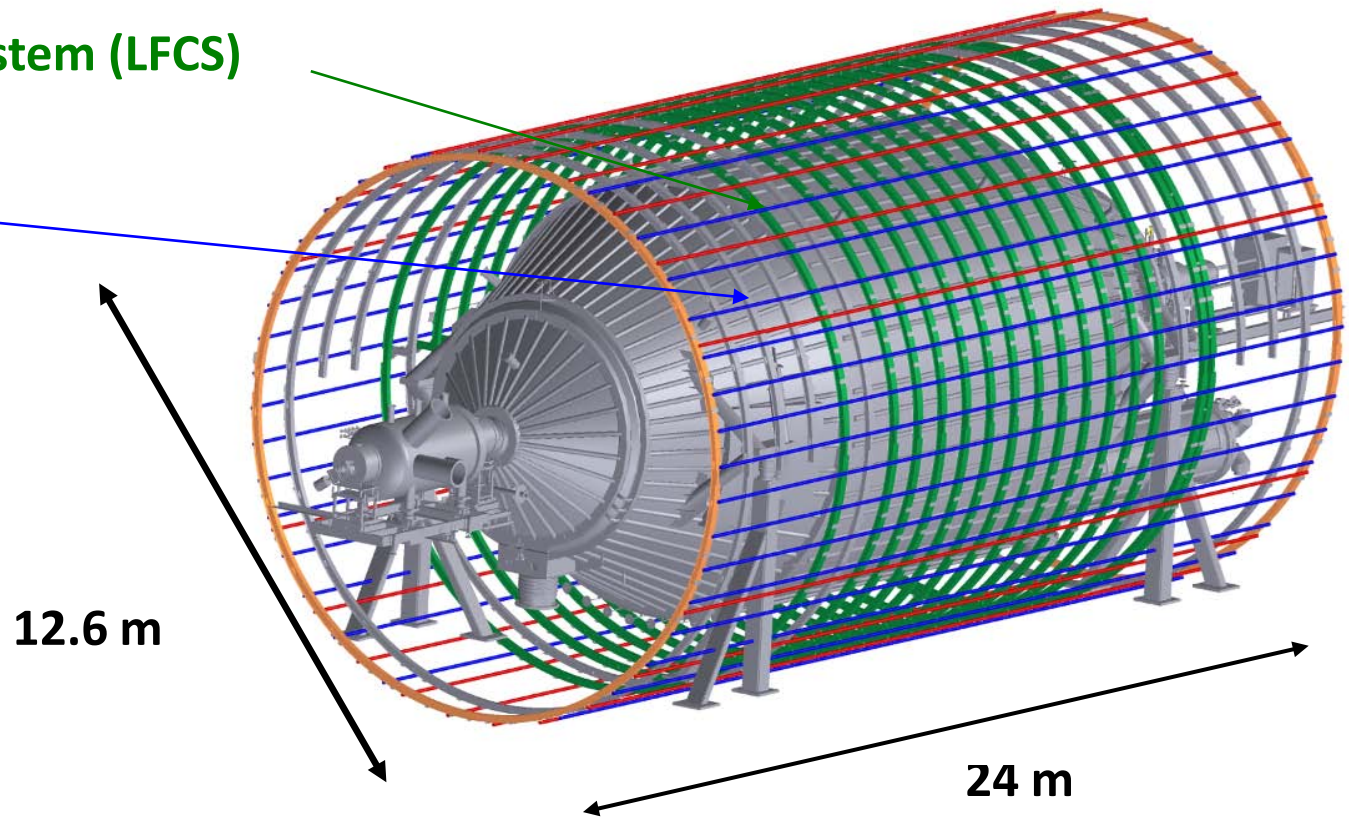


## Aim of air coils:

- to optimize the axisymmetric magnetic field inside the tank (field of supercond. coils alone is not optimal for electron transmission)
- to compensate the earth magnetic field

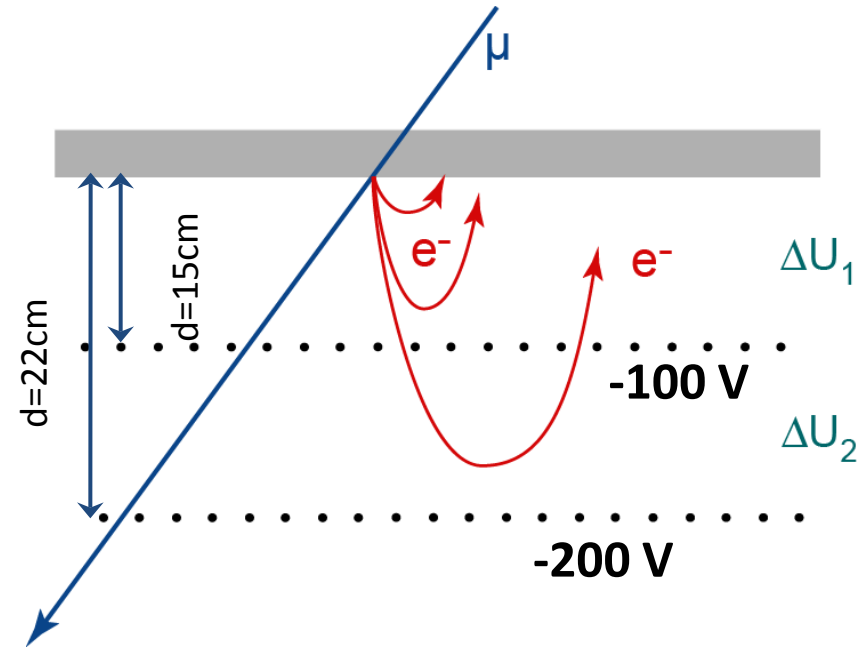
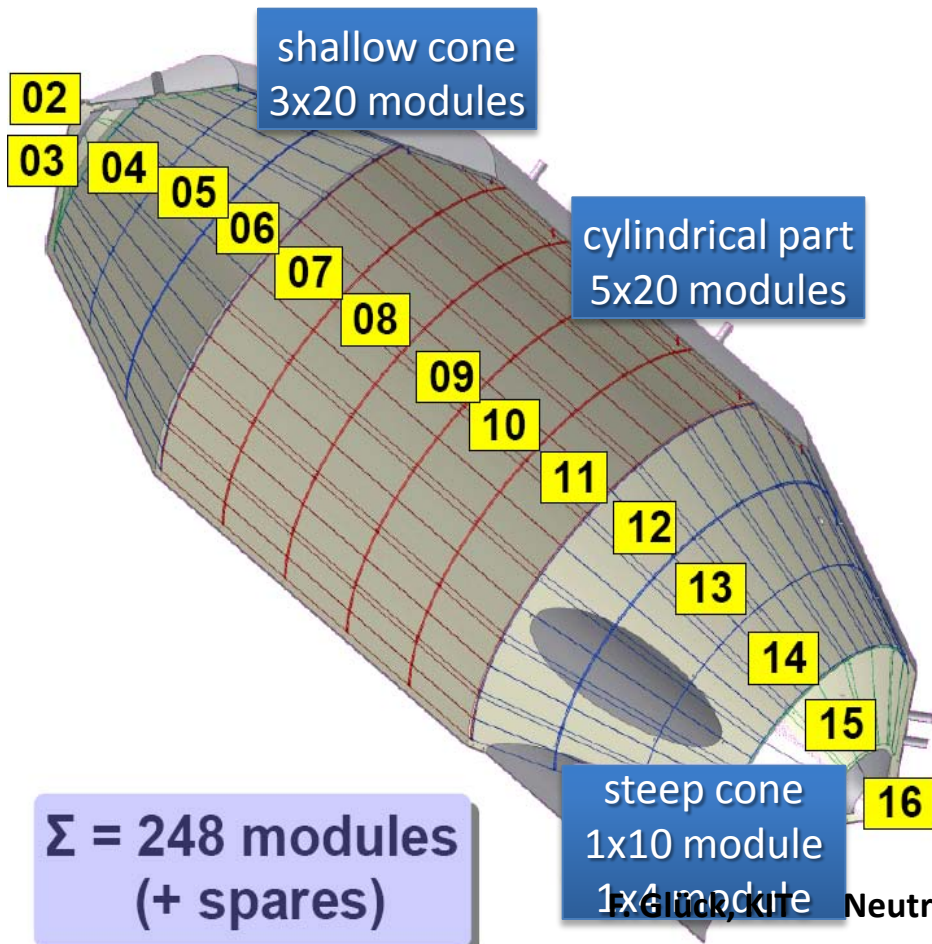
### low field coil system (LFCS)

earth magnetic field  
compensation  
system (EMCS)  
→ cosine coil system



# Electric shielding with wire electrodes: background reduction

Wires more negative than tank  $\rightarrow$   
 small energy secondary electrons  
 are reflected by the wires



## Wires:

0.2 and 0.3 mm diameter

total length: 42 km

tank surface: 700 m<sup>2</sup>

**Wire modules:  
University of Münster**

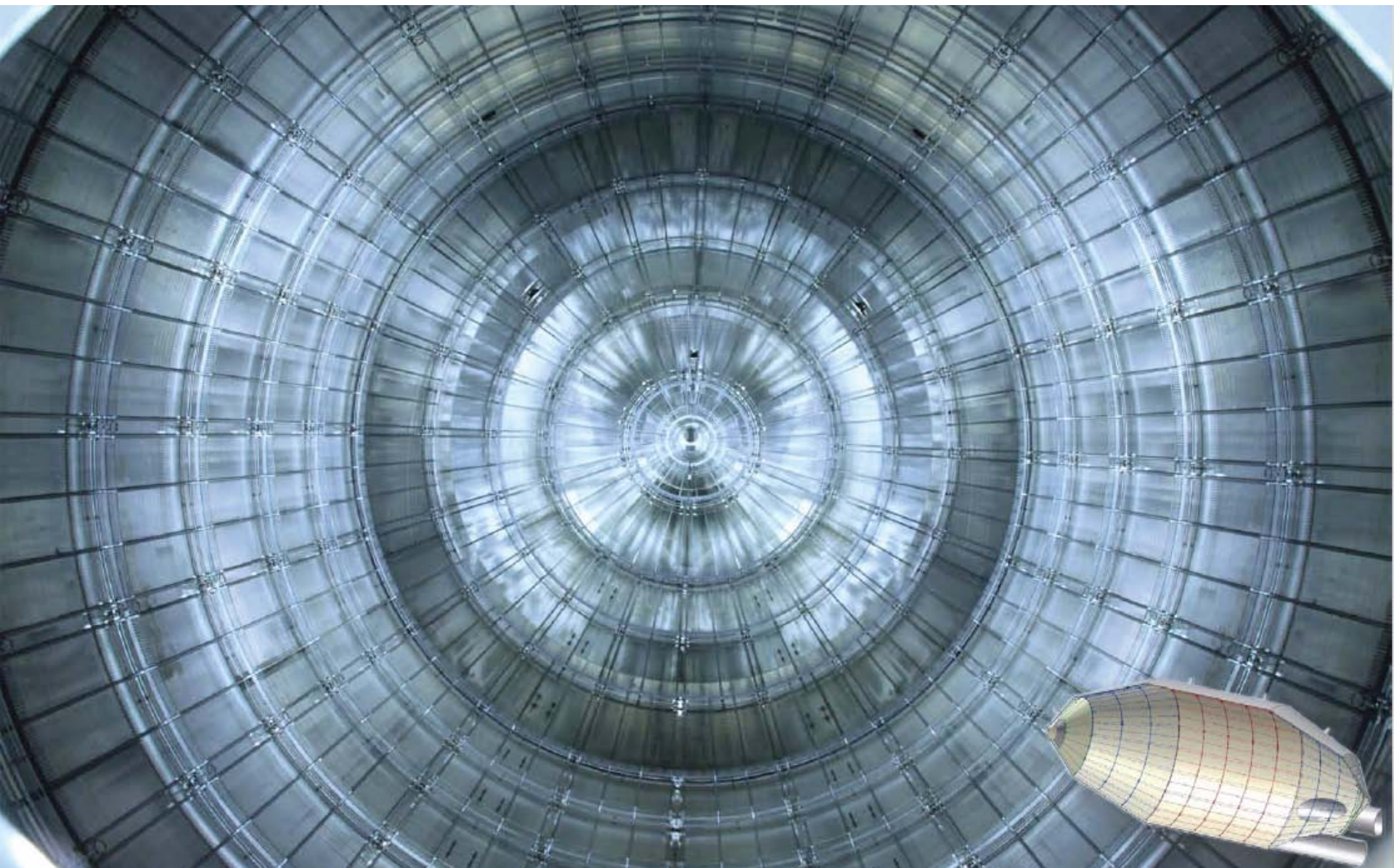


**Scaffolding:  
KIT**







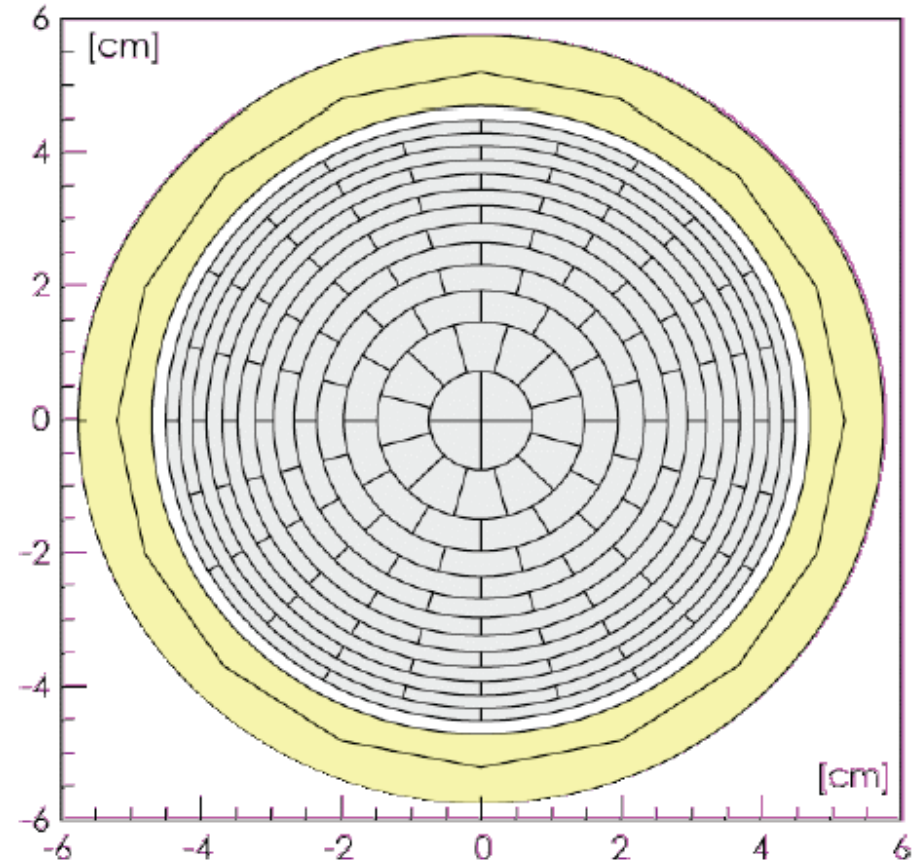


# Detector system

Assembled and commissioned:  
University of Washington, Seattle

## Focal-Plane detector:

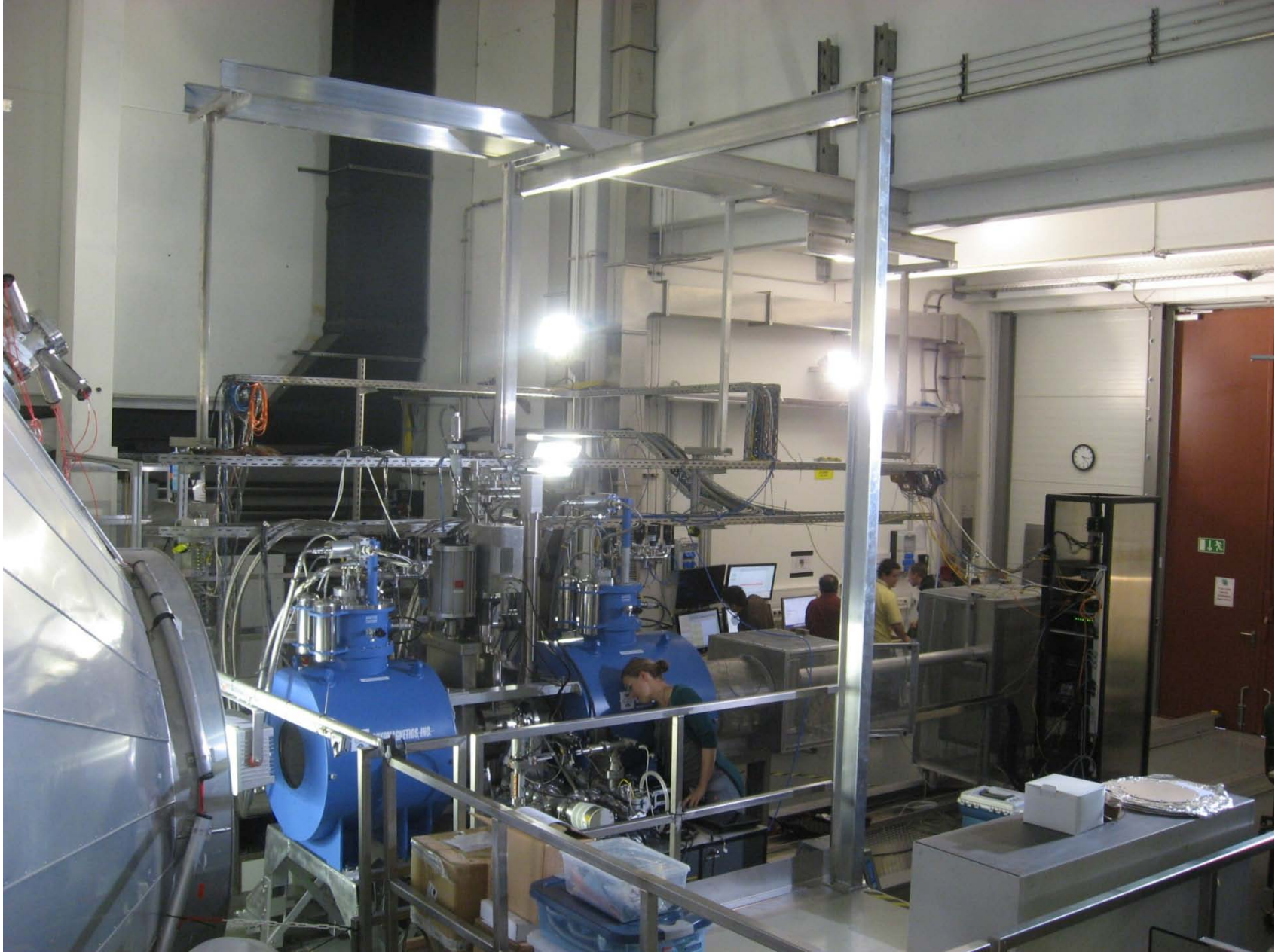
- segmented Si-PIN diode array (148 pixels: to record radial and azimuthal profile of signal and background electrons)
- energy resolution: 1.5 keV
- minimal electron energy: 5 keV
- detector region background: few mHz possible (with muon veto: MIT)



pinch magnet

detector magnet





**Detector system installed near main spectrometer**

F. Glück, KIT Neutrino13, Prague

**Wire electrode installation ready, main spectrometer vessel was baked up to 300 °C (Jan. 2013).**

**Main spectrometer commissioning experiments start next week**

- **first background experiments with low potential**
- **electron transmission experiments (with egun)**
- **radon background with warm and cold baffle; removing stored electrons by electric dipole field and magnetic pulse (fast decrease of magnetic field)**
- **high voltage experiments (field emission, Penning discharge, background from secondary electron emission)**

# Variation of the retarding potential in main spectrometer

Unaccounted variation of the analyzing plane potential in middle of the main spectrometer → shift of measured neutrino mass squared

$$\Delta m_\nu^2 = -2\sigma_{\Delta U}^2 \quad \Delta U = U_A - U_S$$

$U_A, U_S$  : analyzing plane and source potentials

→ **better than 60 mV stability of analyzing plane potential is required during spectrum scanning**

**Absolute value of the analyzing plane potential is not critical (since endpoint is free parameter)**

## Retarding potential monitoring:

- direct voltage measurement (18 kV  $\rightarrow$  10 V by high-precision voltage divider)
- direct calibration of analyzing plane potential of main spectrometer, with conversion electrons from  $^{83\text{m}}\text{Kr}$  or photoelectrons from  $^{241}\text{Am}/\text{Co}$  source; not possible during data taking
- main spectrometer potential connected to monitor spectrometer; continuous calibration of monitor spectrometer potential by photoelectron or conversion electron sources

Quench condensed Kr source, high-precision voltage divider:

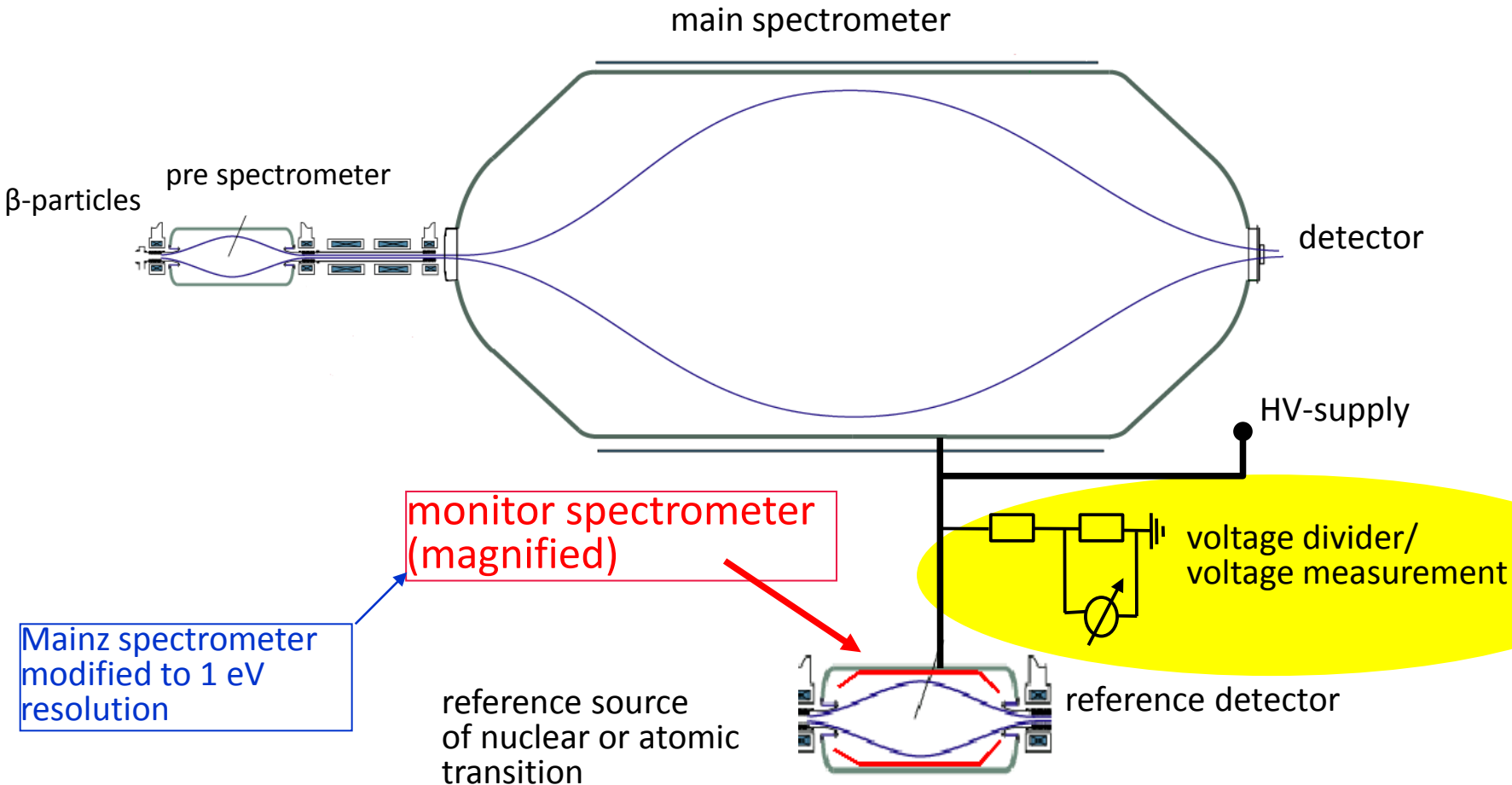
Münster (group of Ch. Weinheimer)

$^{241}\text{Am}/\text{Co}$  source: Rez, group of O. Dragoun

# Monitor Spectrometer

## Precise monitoring of the main spectrometer energy scale:

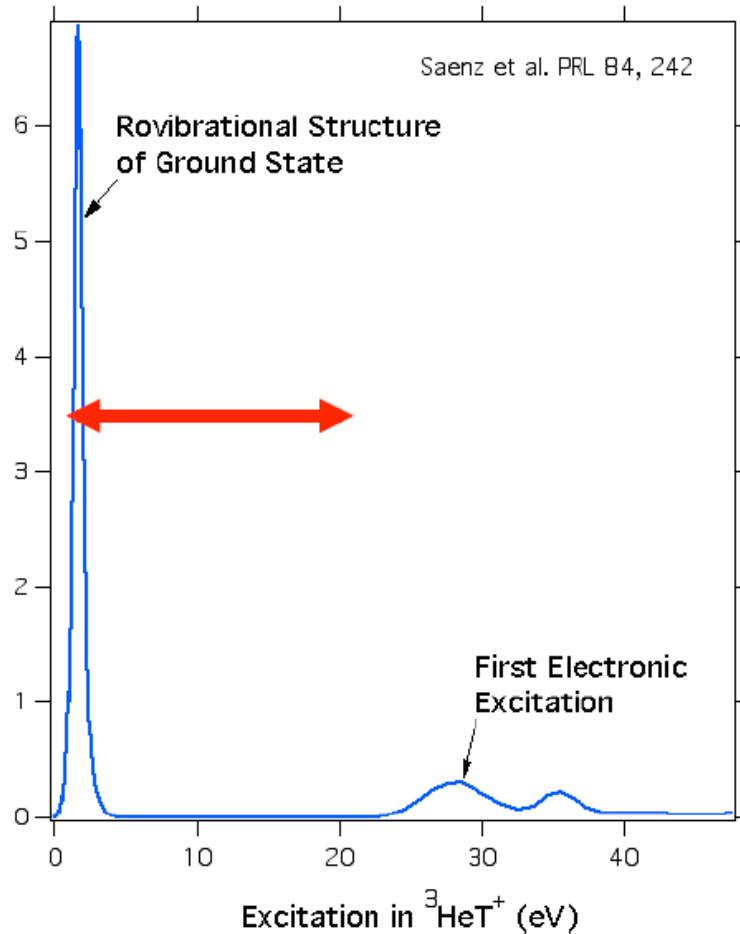
precise measurement of retarding potential  
+ comparison to reference energy



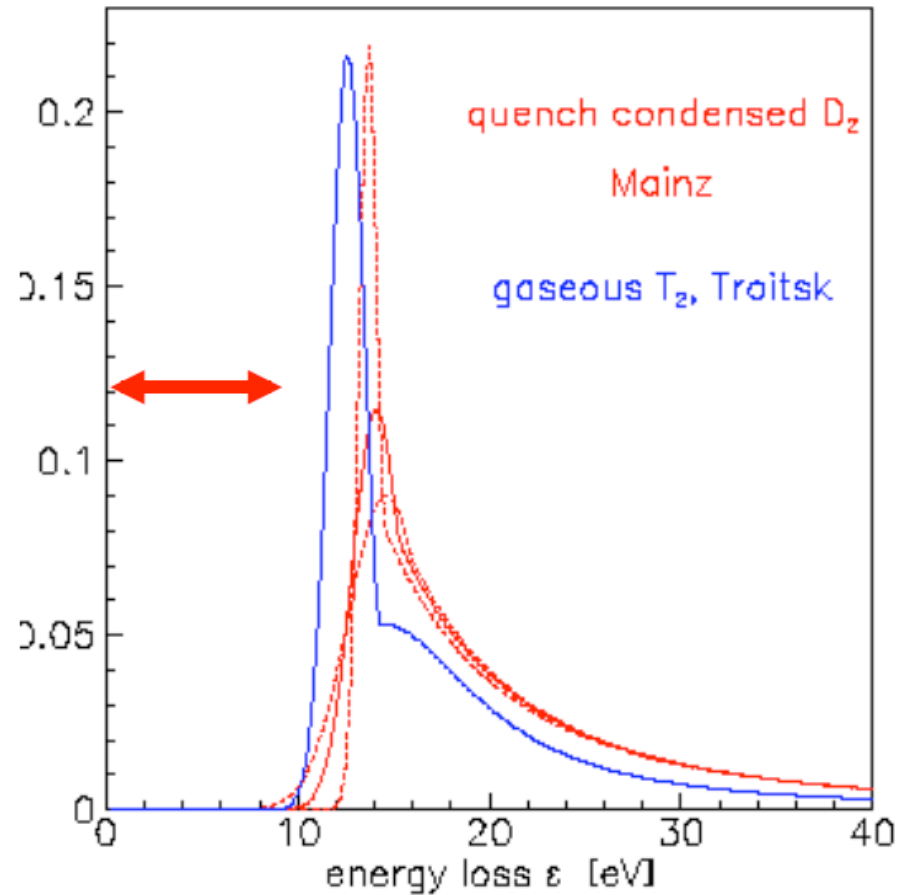


# A window to work in

## Molecular Excitations



## Energy loss function



# Simulation codes

## Various C and C++ codes

Standard C++ simulation package of KATRIN: **KASSIOPEIA**

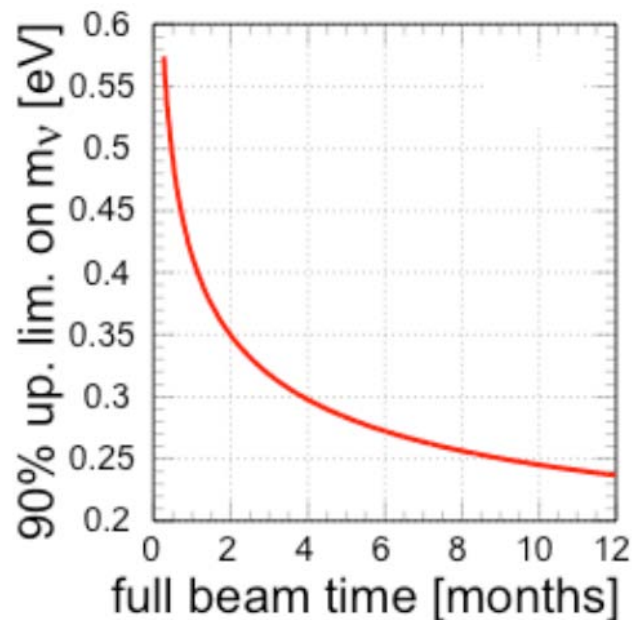
- relativistic charged particle tracking: 8th order Runge-Kutta
- electric field calculations: BEM, direct and iterative solvers, zonal harmonic expansion
- magnetic field: zonal harmonic expansion, elliptic integrals, integrated Biot-Savart
- e-H<sub>2</sub> scattering : elastic, excitation, ionization, total and differential cross sections. Electron scattering in silicon.
- gas flow
- various statistical simulation methods

## After 3 years data (5y realtime):

**discovery potential**  
 $m(\nu) = 0.35 \text{ eV} (5\sigma)$

**sensitivity (90% CL)**  
 $m(\nu) < 0.2 \text{ eV}$

**Planned start of data taking with tritium: 2015**



# KATRIN and sterile neutrinos

**Sterile neutrinos:** chiral right-handed singlets, no weak interaction

But they can mix with left-handed active neutrinos (if massive)

LSND anomaly: short-baseline  $\bar{\nu}_e$  appearance from  $\bar{\nu}_\mu$  beam

Reactor antineutrino anomaly: short-baseline detected  $\bar{\nu}_e$  rate smaller than calculated rate

Gallium anomaly: detected  $\nu_e + {}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge} + e^-$  rate from artificial  $\nu_e$  EC sources ( ${}^{51}\text{Cr}$ ,  ${}^{37}\text{Ar}$ ) smaller than calculated rate

*large mixing:  $\sin^2(2\theta) \approx 0.1$*

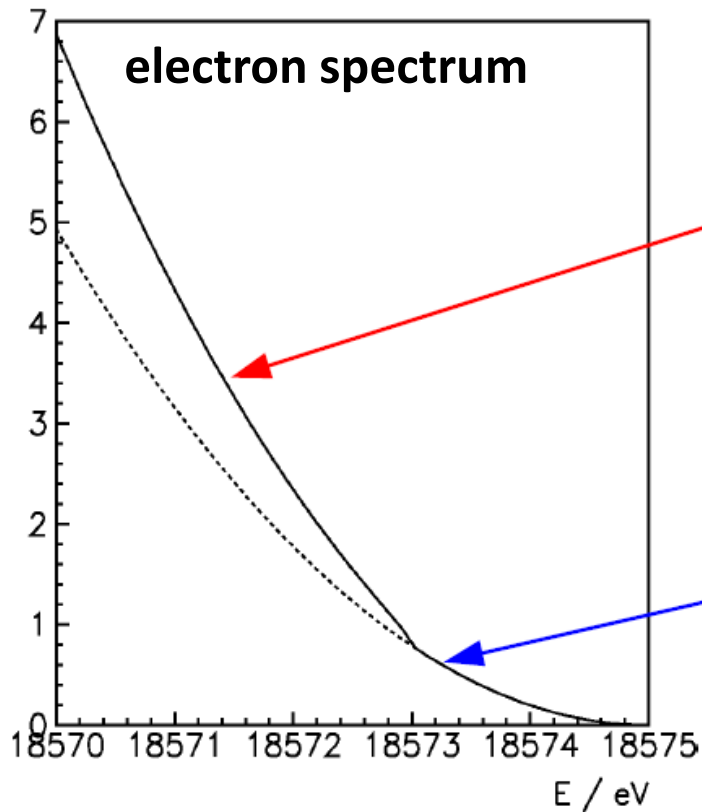
*large masses:  $\Delta m^2 > 1 \text{ eV}^2$*

$$\nu_e = \sum_{i=1}^3 U_{ei} \nu_i + U_{e4} \nu_4$$

$$\sum_{i=1}^3 |U_{ei}|^2 = \cos^2(\theta)$$

$$|U_{e4}|^2 = \sin^2(\theta)$$

$$dN/dE = K F(E,Z) p_{E_{tot}} (E_0 - E_e) \left( \cos^2(\theta) \sqrt{(E_0 - E_e)^2 - m(\nu_{1,2,3})^2} + \sin^2(\theta) \sqrt{(E_0 - E_e)^2 - m(\nu_4)^2} \right)$$



e.g. **sterile neutrino**

$$m(\nu_4) = 2 \text{ eV}$$

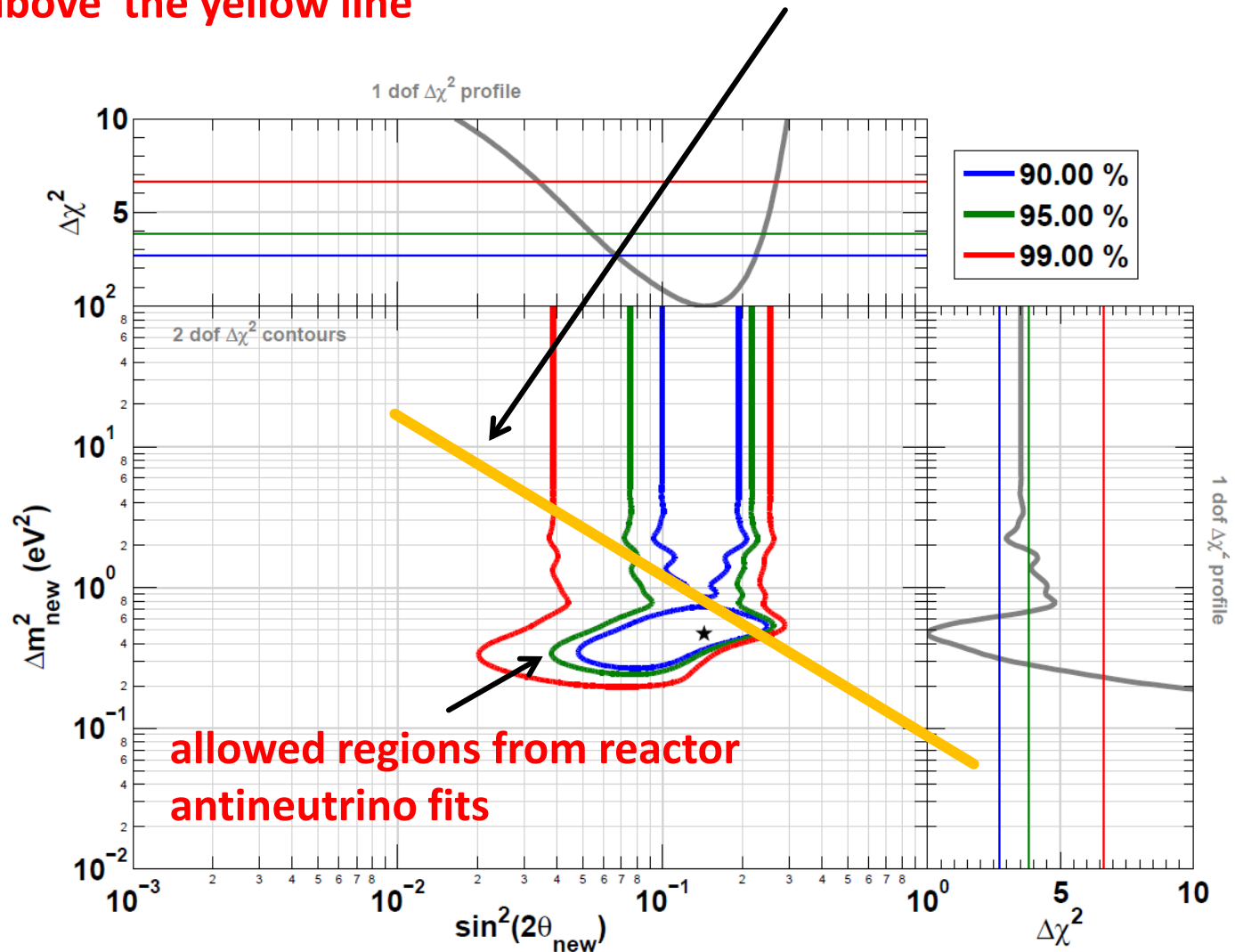
$$\sin^2(\theta) = 0.3$$

e.g. **active neutrinos**

$$m(\nu_{123}) \approx 0 \text{ eV}$$

$$\cos^2(\theta) = 0.7$$

# KATRIN 90 % exclusion possibility for light sterile neutrino above the yellow line

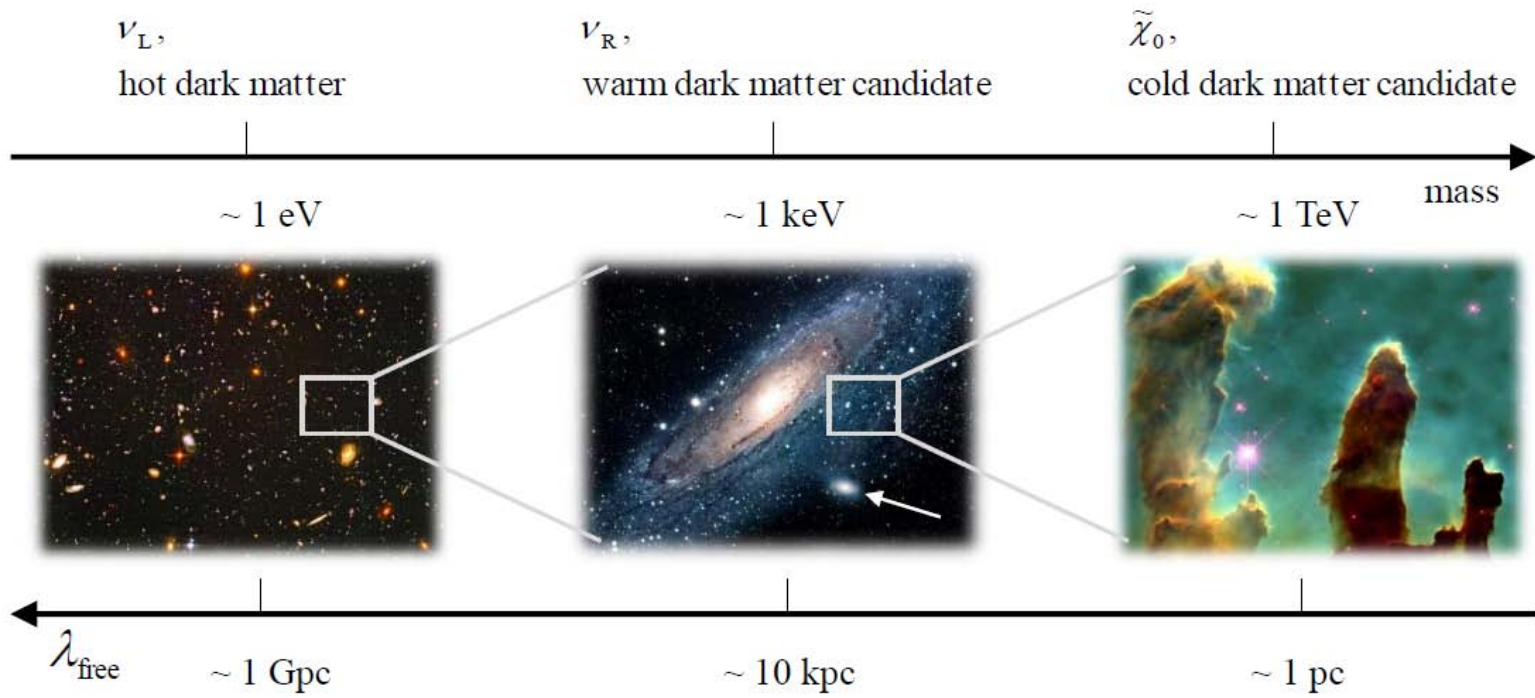


Riis, Hannerstad, JCAP 1102 (2011) 011

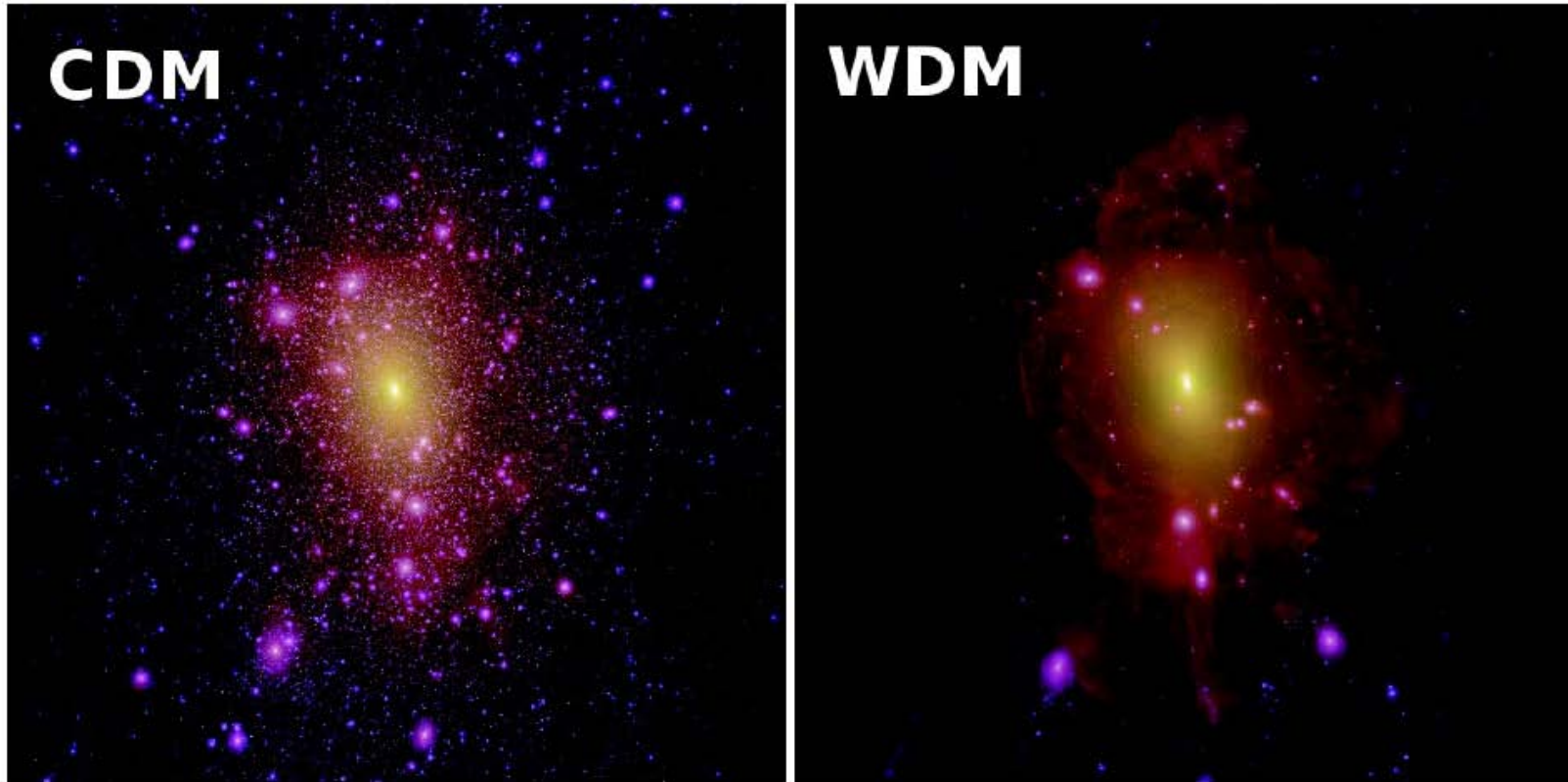
Formaggio, Barrett, PLB 706 (2011) 68

Esmaili, Peres, arXiv 1203.2632

# Warm dark matter (WDM) and keV sterile neutrinos



Some problems with CDM (e.g.: too many satellite dwarf galaxies predicted)

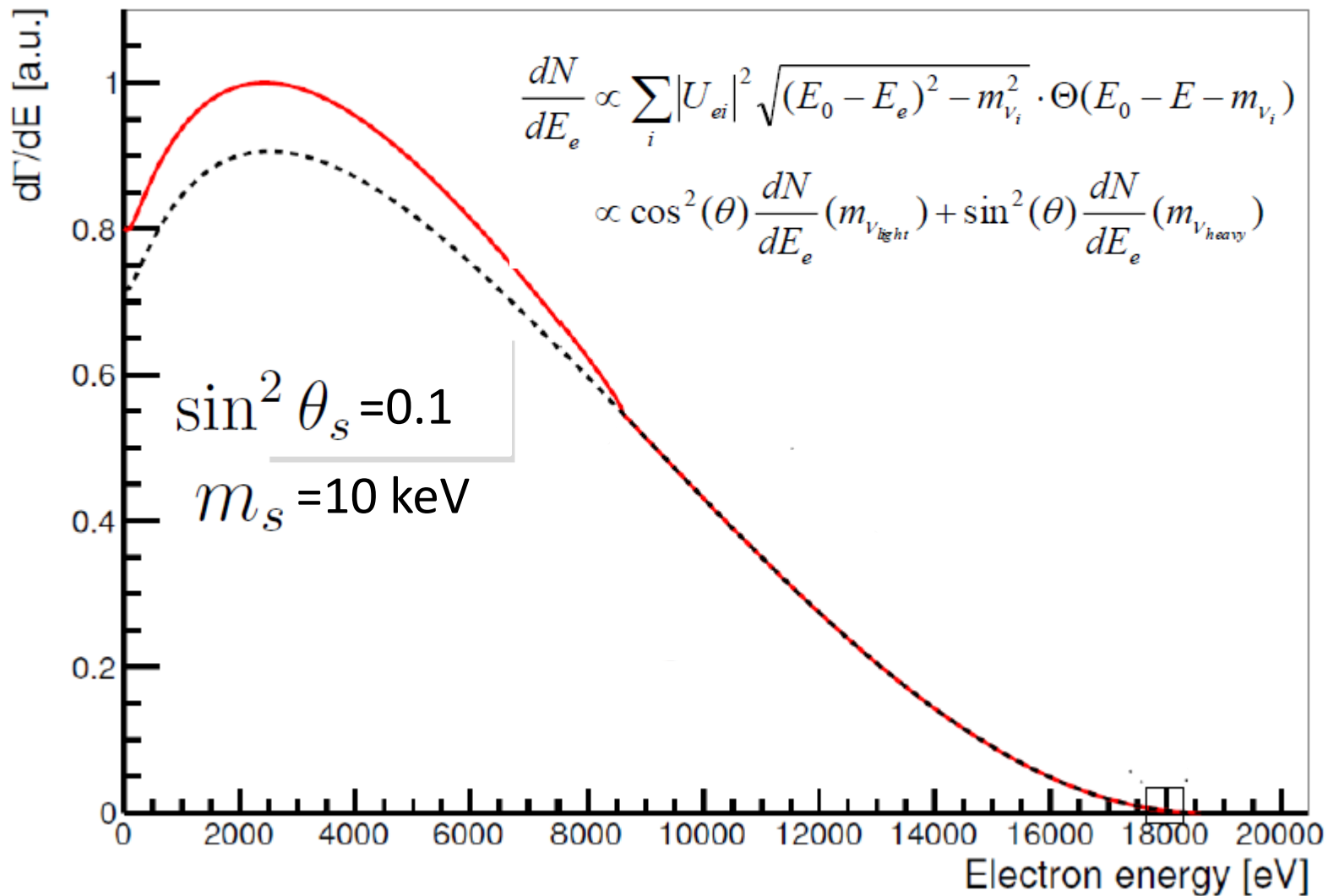


keV mass sterile neutrinos: possible WDM candidates

Predictions: mass: 1-15 keV, mixing angle:  $\sin^2\Theta_s < 10^{-7}$



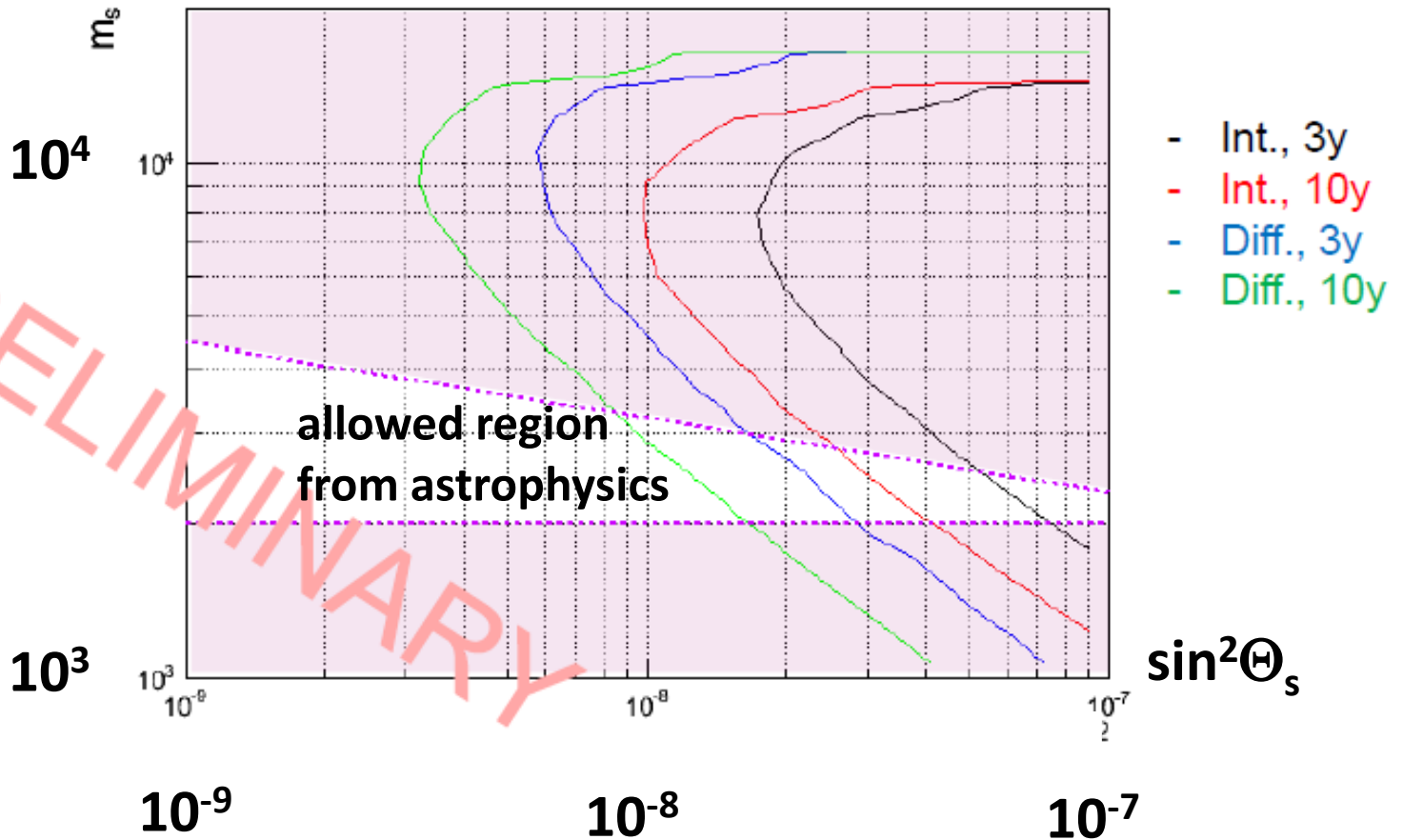
# Tritium beta decay spectrum with sterile neutrino (example)



# KATRIN statistical 90 % exclusion possibility for WDM sterile neutrino

Sterile neutrino mass (eV)

90% C.L.



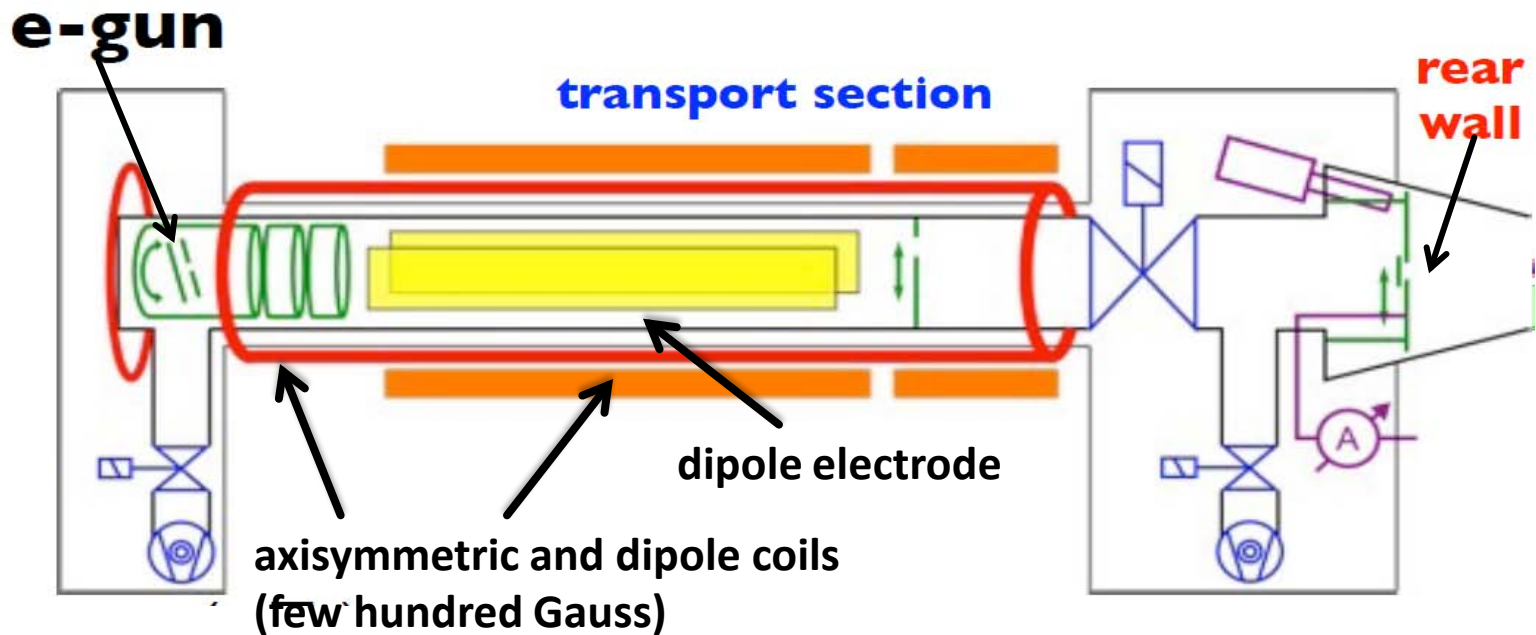
**Systematics ???**

# Summary

- **absolute neutrino mass scale determination: important task for particle physics and cosmology**
- **KATRIN: direct , model independent absolute neutrino mass scale determinatio by electron energy spectrum measurement of tritium  $\beta$  decay**
- **Spectrometer method; gaseous tritium source far from detector; magnetic guiding; integral spectrum (MAC-E filter); high statistics, high resolution, small background; complicated system; data taking start: not before 2015**
- **goal: 0.2 eV neutrino mass limit (90 % CL)**
- **possibility of light and WDM keV sterile neutrino detection**

# CMS

## Calibration and Monitoring Section



### Egun:

- electrons with sharp energy and angular distribution
- transmission function measurements
- source thickness monitoring
- plasma effects

### Rear wall:

- source potential definition and stabilization
- source activity monitoring
- gold-plated

## Advantages of tritium:

- **Superaligned transition:** → matrix element  $M$  is not energy dependent
- **Low endpoint energy:** → relative decay fraction at the endpoint is comparatively high
- **Short half life:**
  - specific activity is high
  - low amount of source material
  - low fraction of inelastic scattered electrons
- **Hydrogen isotope:**
  - simple atomic shell
  - final states precisely calculable

Final state distribution calculation is needed for differential spectrum of  $T_2$  decay:

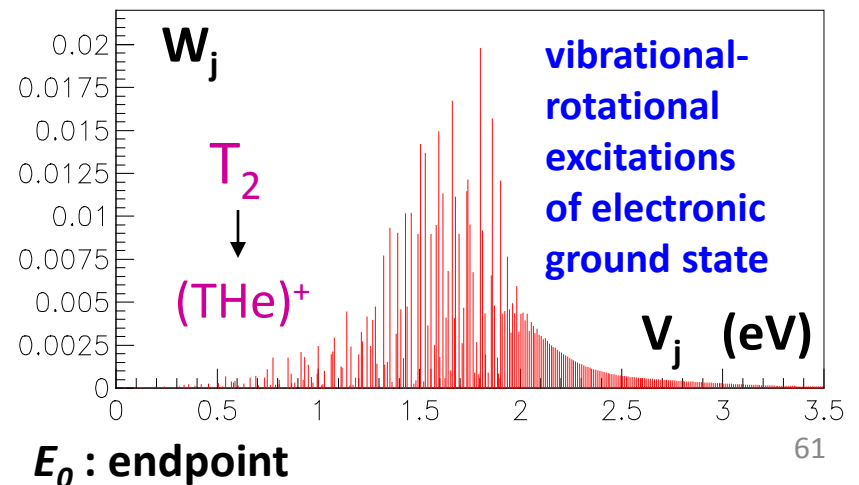
$$w_{diff}(E) = \sum_j W_j \cdot E_{\nu j} \sqrt{E_{\nu j}^2 - m_\nu^2}$$

$$E_{\nu j} = E_0 - V_j - E$$

(assuming degenerate neutrino masses)

Ground state  
probability: 0.57

mean: 1.7 eV  
 $\sigma = 0.4$  eV

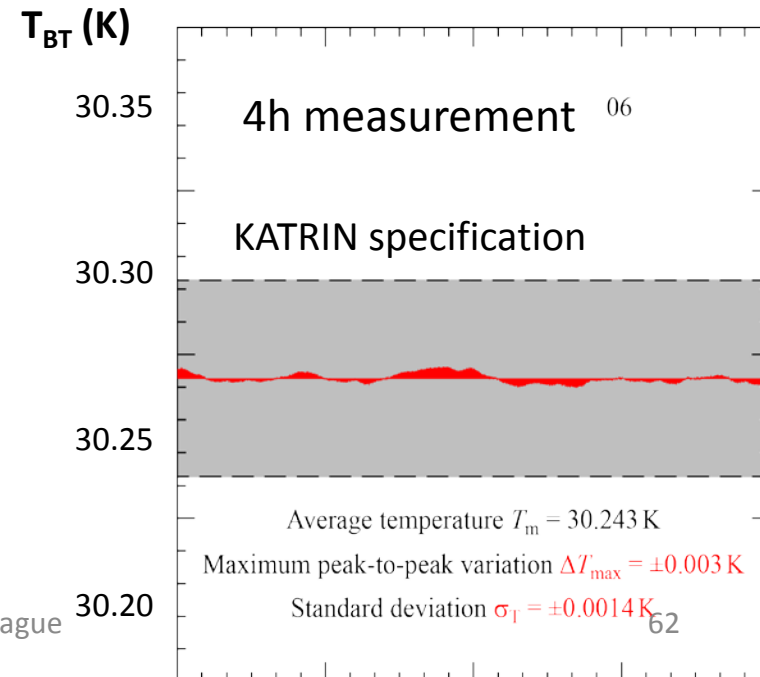
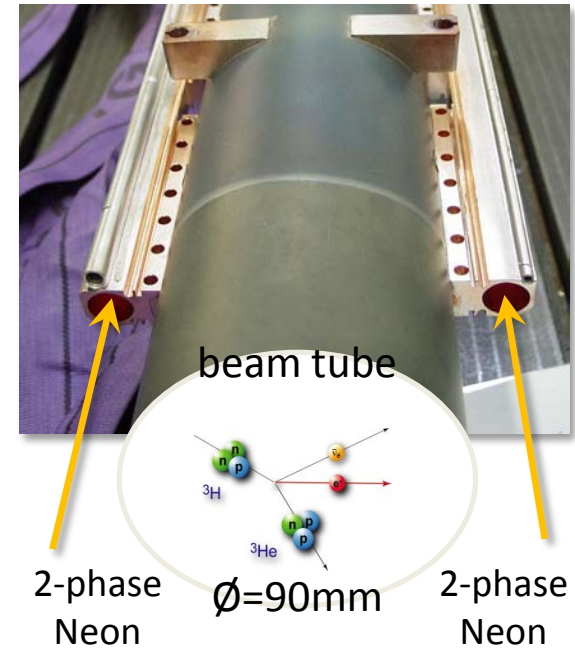


# WGTS demonstrator

- beam tube cooling system:  $T_{BT} = 28\text{-}32\text{ K}$   
 $\Delta T = \pm 30\text{ mK}$  - stability (1h) & homogeneity
- initial stability results: **proof-of-principle**

Parameter	$\Delta T$ (4h)	$\Delta T$ (24h)
peak-to-peak variation	3 mK	9 mK
standard deviation $\sigma_t$	<b>1.4 mK</b>	3.6 mK

- implications:  $\Delta T_{BT} \ll 10^{-3}$ ,  $\Delta p_{in} < 10^{-3}$ 
  - ↪ **super-stable  $\beta$ -emitting source**
  - ↪ **reduced systematics from source**
  - ↪ demonstrator  $\Rightarrow$  WGTS cryostat (2013)



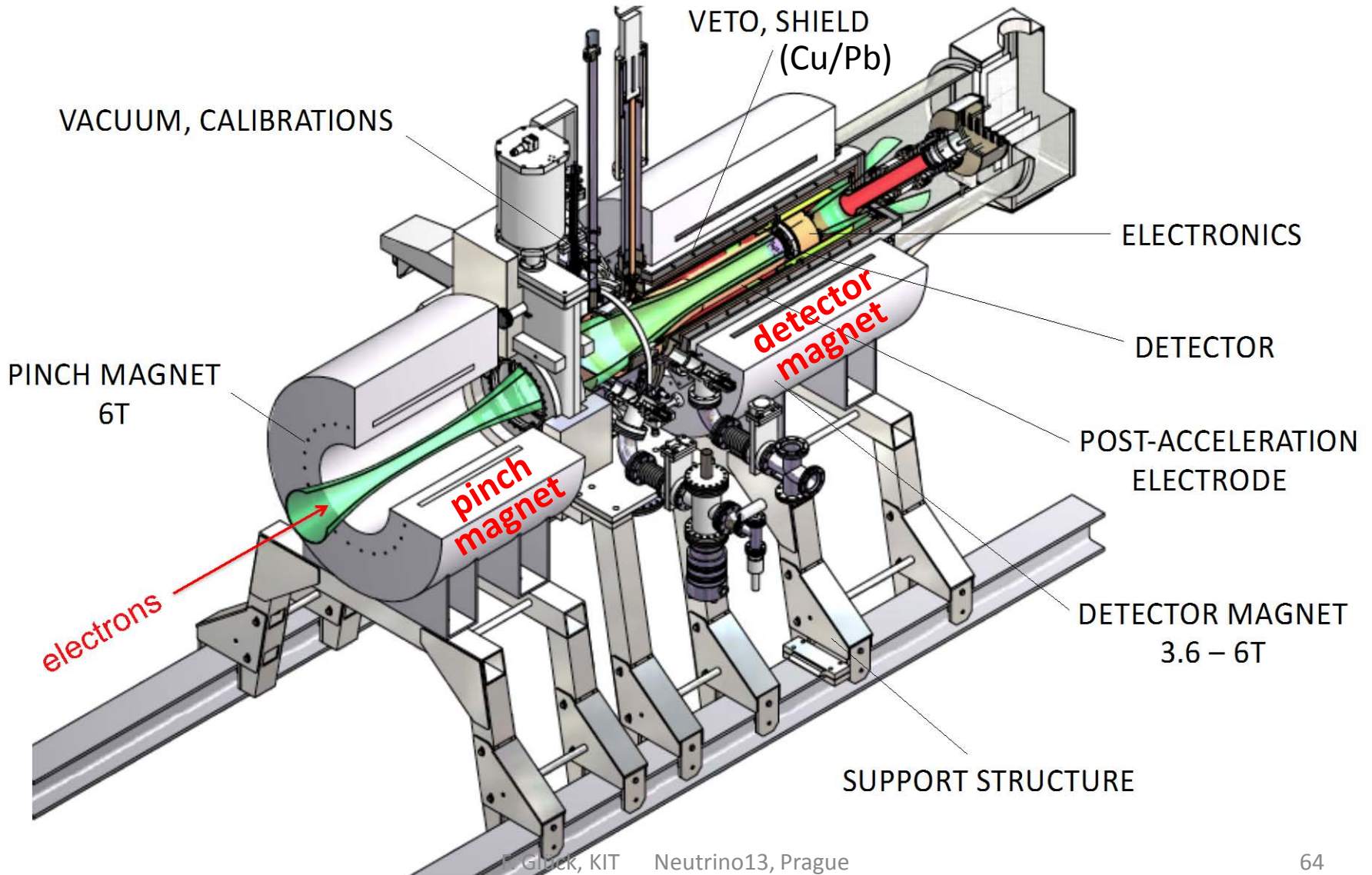
# WGTS

## Windowless Gaseous Tritium Source



- T<sub>2</sub> gas:
  - high luminosity and high stability
  - differential pumping (TMPs); closed tritium loop
  - high tritium purity; measurement of T<sub>2</sub>, DT etc. isotopic composition (with Laser Raman spectroscopy)
- Homogeneous magnetic field and adiabatic guiding of electrons
- Systematic effects for neutrino mass: as small as possible

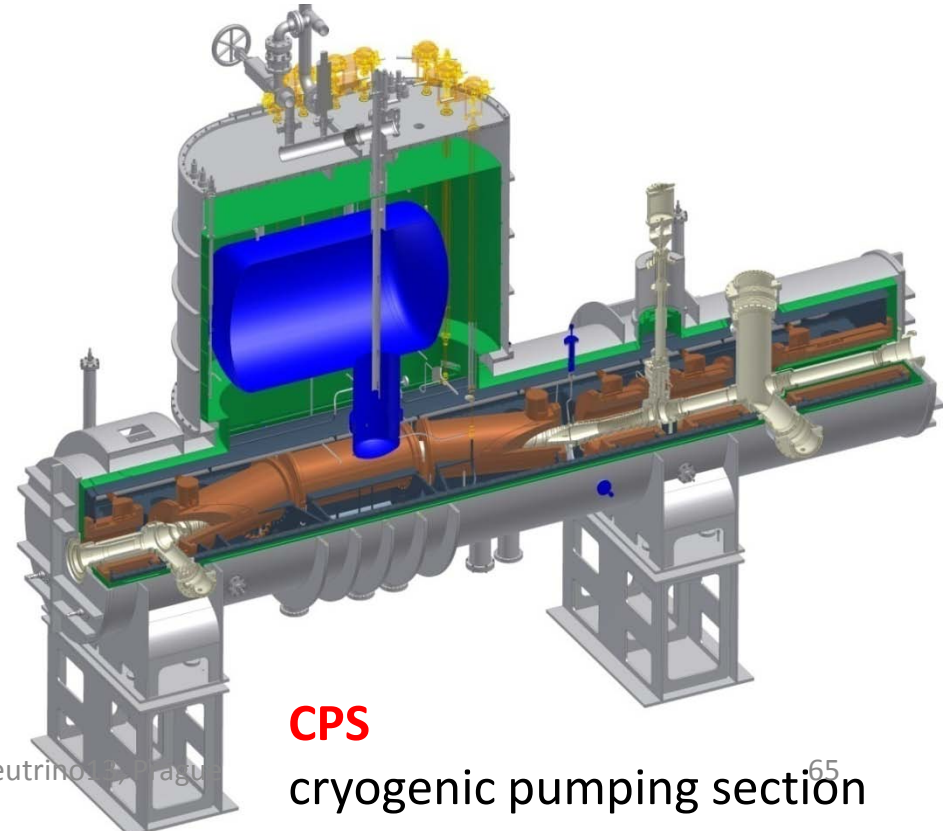
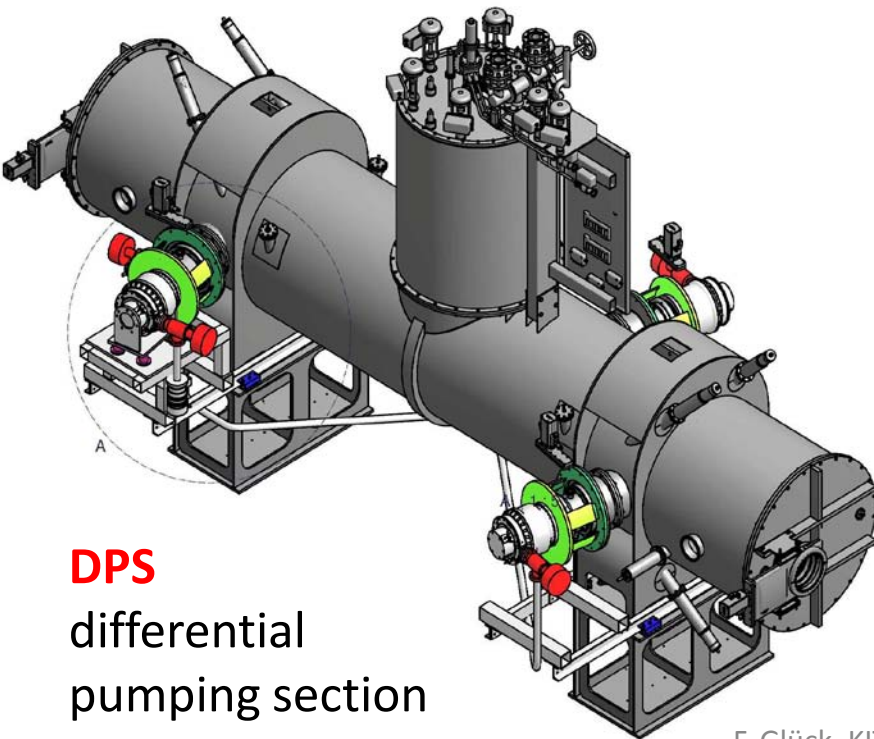
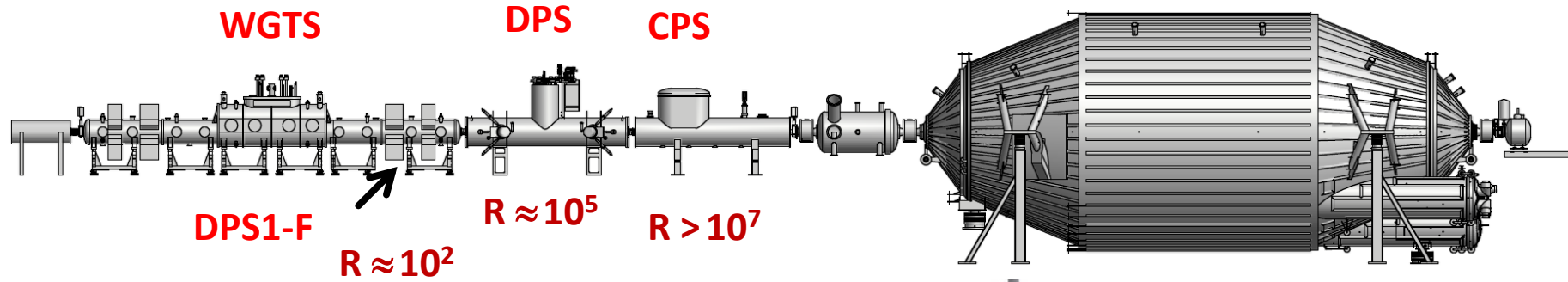
# Focal Plane Detector System





# Tritium retention: DPS1-F, DPS2-F and CPS

Only very small amount of tritium allowed to enter the main spectrometer (background !)

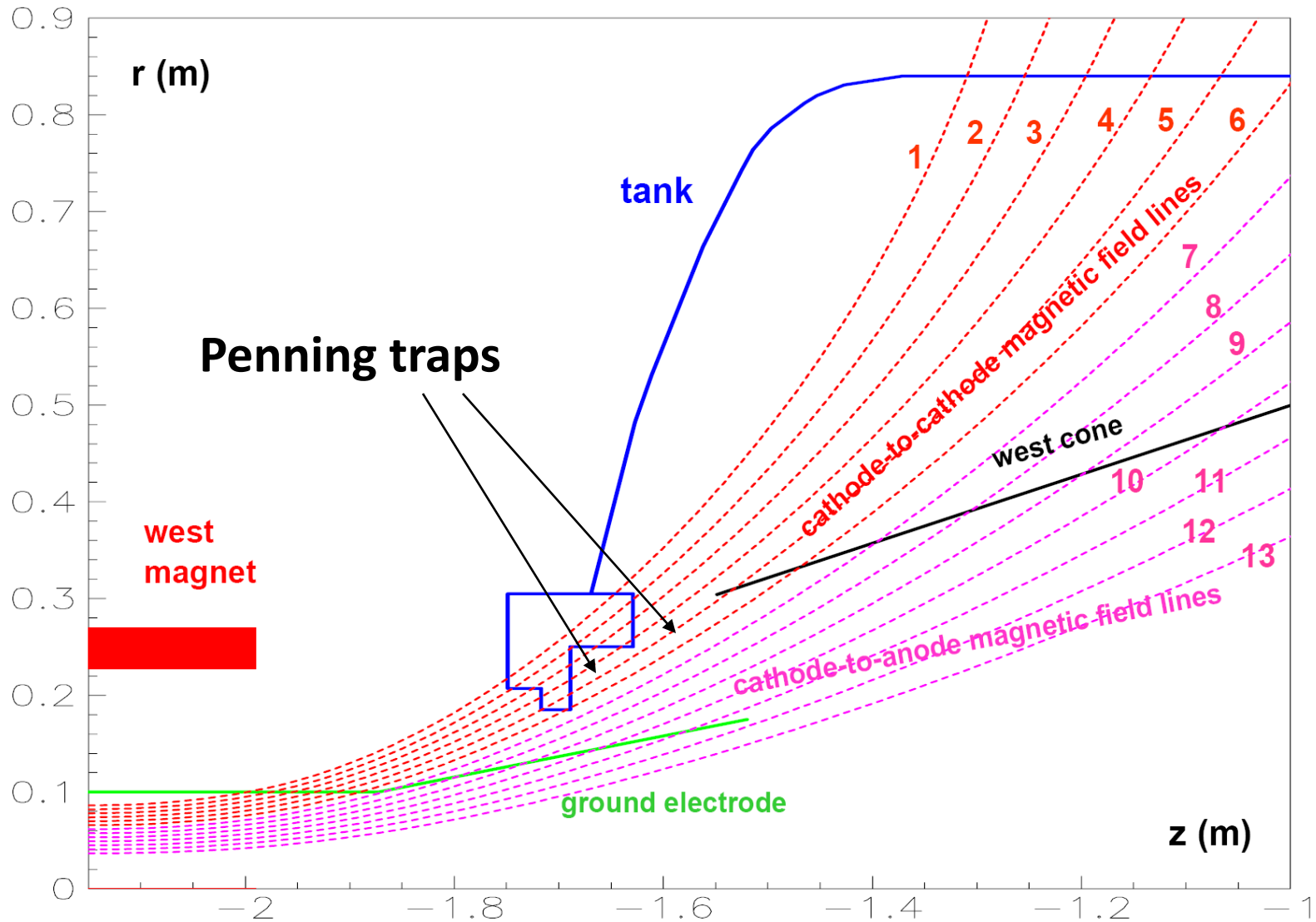


## First test experiments at KATRIN pre-spectrometer (end of 2006):

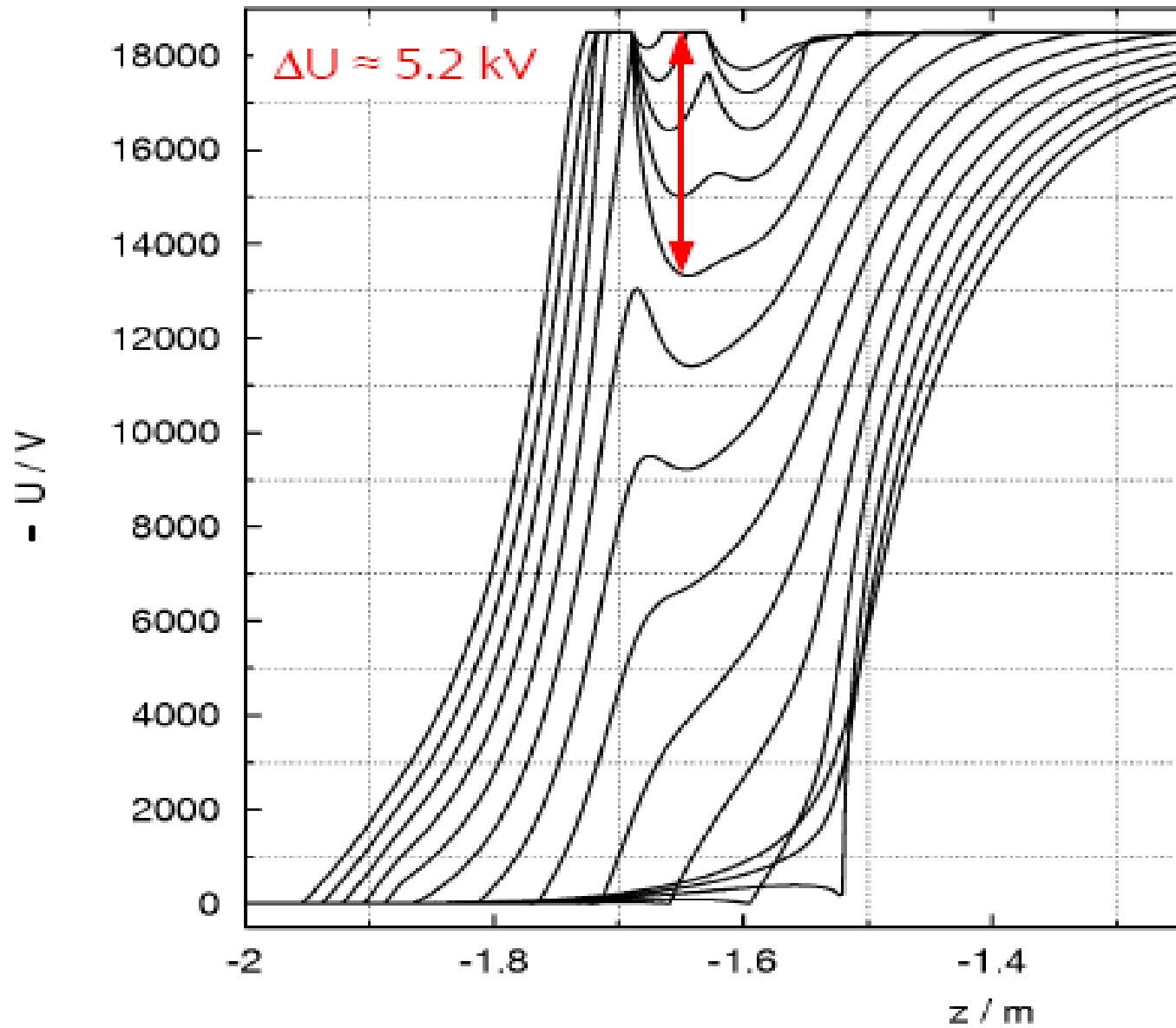
Tank and electrodes at -18 kV, increase of magnetic field;  
at 0.15 T

- increase of vacuum pressure from  $10^{-9}$  mbar to  $10^{-6}$  mbar
- increase of leakage current from  $0.2 \mu\text{A}$  to few mA (limit of power supply)
- drop of high voltage from -18 kV down to -3 kV (electric breakdown)

**Penning discharge !**



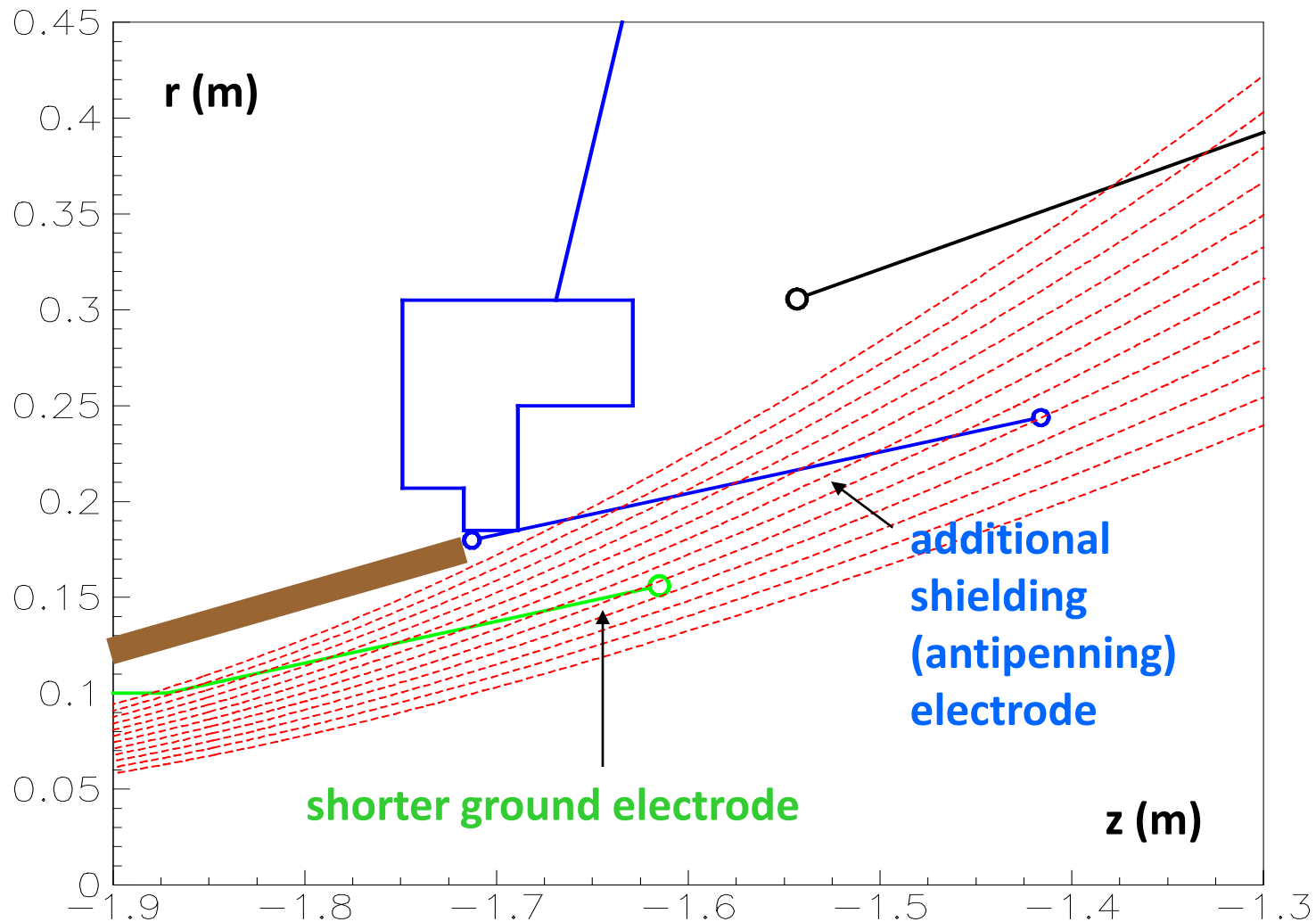
**Cathode-to-cathode magnetic field lines, potential penetration from anode (ground electrode)**



**Electric potential on various magnetic field lines at the entrance of pre-spectrometer**

# Solution of the Penning discharge problem in the pre-spectrometer

by shielding the Penning trap region against potential penetration from the anode by an additional electrode → no deep Penning traps



**After installation of the new electrodes (Sept 2007):**

**No pressure and leakage current increase, no electric breakdown (up to -30 kV, 4.5 T)**

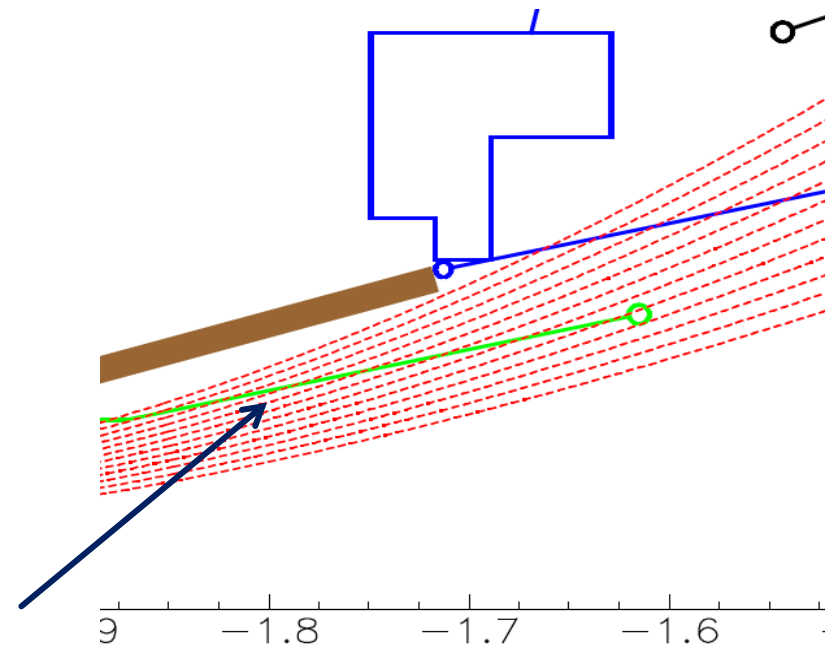
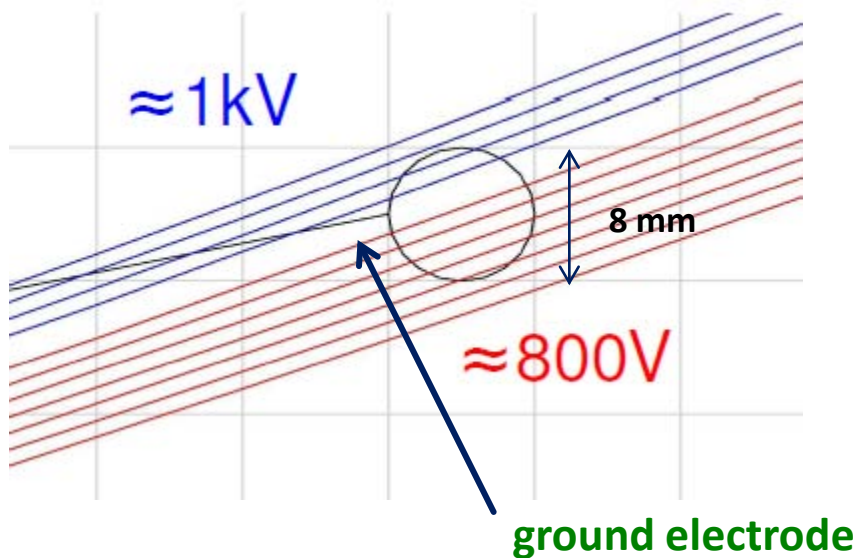
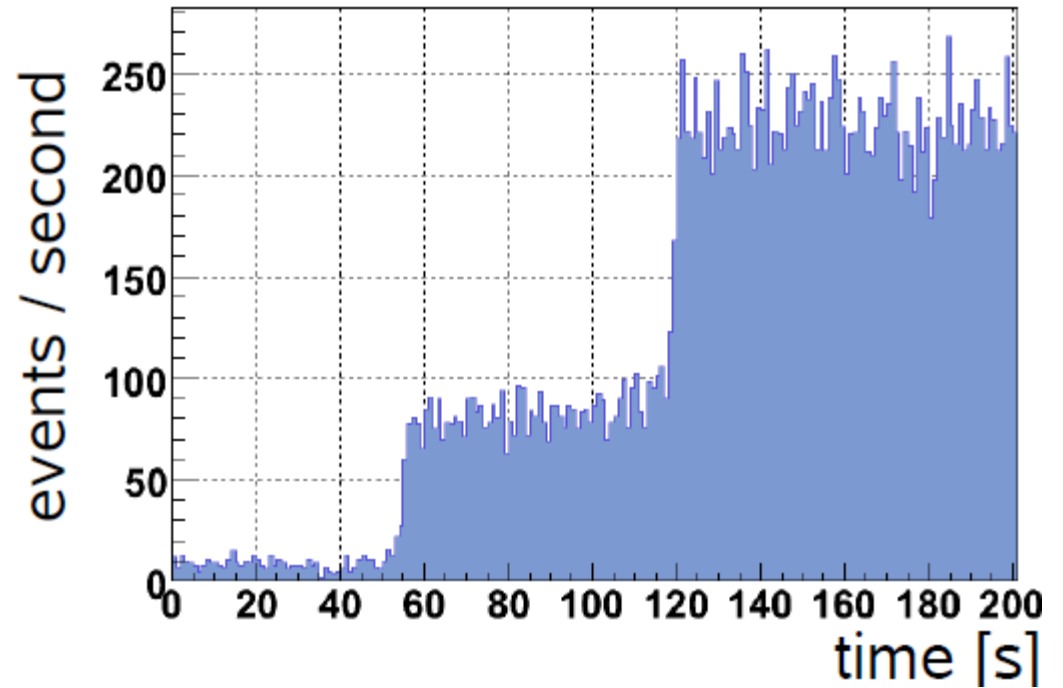
—————→ **large Penning discharge disappeared !**

**It was then possible to start background measurements with the segmented PIN-diode detector.**

Measurements at high (>2 T) magnetic field:

background:  
order of few 100 Hz

Again Penning traps:  
at the end ring of the  
ground electrode,  
small dimensions (few mm)



## Solution:

new ground electrode  
design, with detailed  
simulations,  
Penning traps eliminated

Using the new ground electrode:

Background decreased from  
few 100 Hz to few 10 mHz  
(factor 10000)

After reducing background  
from radon decays:

Background: few mHz  
(reduction factor of new ground electrode: 100000)

