

Hydrodynamic response to initial state fluctuations

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AA-collisions

- Initial particle/energy production, followed by
 - Hydrodynamic evolution, followed by
 - Freeze-out/Hadron cascade
-
- Goal is to determine QGP properties: EoS, transport coefficients . . .
 - Large(st) uncertainty in determining e.g. shear viscosity is initial state for hydrodynamic evolution

Complicated (and necessary) problem is to determine initial conditions simultaneously with QGP properties

Hydrodynamics

Conservation laws

$$\partial_\mu T^{\mu\nu} = 0$$

$$\partial_\mu n_i^\mu = 0$$

$$T^{\mu\nu} = e u^\mu u^\nu - (p + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu}$$

IS equations for viscous parts of $T^{\mu\nu}$

e.g. shear viscosity:

$$\tau_\pi \frac{d}{d\tau} \pi^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = 2\eta \nabla^{\langle\mu} u^{\nu\rangle} + \dots$$

To solve this set of equations we need

- Equation of state $p = p(e)$ and $T = T(e)$
- Initial condition $T^{\mu\nu}(\tau_0, \mathbf{x})$
- Transport coefficients, e.g. shear viscosity $\eta(T)$, relaxation time $\tau_\pi(T)$, ...

Hydrodynamics → particles (Freeze-out)

- Cooper-Frye freeze-out for particle i = calculate number of particles crossing the freeze-out hypersurface

$$E \frac{dN}{d^3\mathbf{p}} = \frac{g_i}{(2\pi)^3} \int d\sigma^\mu p_\mu f_i(\mathbf{p}, x)$$

+ decays

Hadron spectra

$$\rightarrow \frac{dN_h}{dy dp_T^2 d\phi}$$

~ 300 hadronic states: takes most of the computing time

Characterizing p_T -spectra

Fourier decomposition: w.r.t. event plane

$$\frac{dN}{dyd p_T^2 d\phi} = \frac{dN}{dyd p_T^2} [1 + 2v_1(p_T) \cos(\phi - \psi_1) + 2v_2(p_T) \cos[2(\phi - \psi_2)] + \dots]$$

$$\psi_n = (1/n) \arctan(\langle p_T \sin n\phi \rangle / \langle p_T \cos n\phi \rangle)$$

- $v_n(p_T)$, $\psi_n(p_T)$, dN/dy , ... characterize single event

Ensemble of events: Full characterization

- Averages: $\langle v_n \rangle$, $\langle \psi_n \rangle$, ...
- Probability distributions: $\mathcal{P}(v_n)$, $\mathcal{P}(\psi_n)$, ...
- Correlations: $\langle v_n, v_m \rangle$, $\langle \psi_n, \psi_m \rangle$, ...

Characterizing initial state

Eccentricity (for energy density ε)

$$\epsilon_{m,n} = -\frac{\int dx dy r^m \cos [n(\phi - \Psi_{m,n})] \varepsilon(x, y, \tau_0)}{\int dx dy r^m \varepsilon(x, y, \tau_0)}$$

$$\Psi_{m,n} = \frac{1}{n} \arctan \frac{\int dx dy r^m \sin(n\phi) \varepsilon(x, y, \tau_0)}{\int dx dy r^m \cos(n\phi) \varepsilon(x, y, \tau_0)} + \pi/n$$

- $e_{m,n}$, $\Psi_{m,n}$ characterize single event (initial energy density)

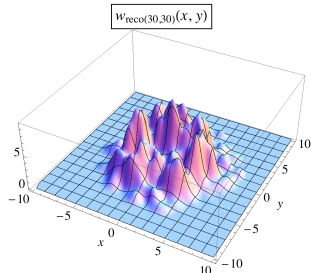
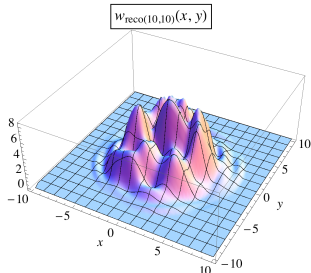
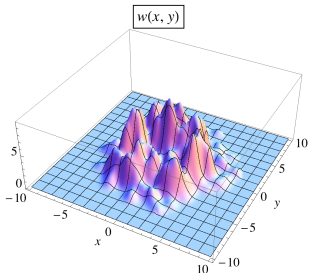
Ensemble of events (initial conditions): Full characterization

- Averages: $\langle e_{m,n} \rangle$, $\langle \Psi_{m,n} \rangle$, ...
- Probability distributions: $\mathcal{P}(e_{m,n})$, $\mathcal{P}(\Psi_{m,n})$
- Correlations: $\langle e_{m,n}, e_{m',n'} \rangle$, $\langle \Psi_{m,n}, \Psi_{m',n'} \rangle$, ...

Bessel-Fourier expansion: S. Floerchinger, U. Wiedemann: arXiv:1307.7611 [hep-ph]

$$w_l^{(m)} = \frac{2}{R^2 [J_{m+1}(k_l^{(m)} R)]^2} \int_0^R dr r w^{(m)}(r) J_m(k_l^{(m)} r)$$

$$w_{\text{reco}}(N_m, N_l)(r, \phi) = \sum_{l=1}^{N_l} w_l^{(m=0)} J_0(z_l^{(0)} r/R) + 2 \sum_{m=1}^{N_m} \sum_{l=1}^{N_l} |w_l^{(m)}| \cos[m(\phi - \varphi_l^{(m)})] J_m(z_l^{(m)} r/R)$$



Connecting initial condition to hadron spectra

Ensemble of events (initial conditions): Full characterization

- Averages: $\langle e_{m,n} \rangle, \langle \Psi_{m,n} \rangle, \dots$
- Probability distributions: $\mathcal{P}(e_{m,n}), \mathcal{P}(\Psi_{m,n})$
- Correlations: $\langle e_{m,n}, e_{m',n'} \rangle, \langle \Psi_{m,n}, \Psi_{m',n'} \rangle, \dots$



Hydrodynamic response (EoS, η/s , T_{dec} , ...)

Ensemble of events (spectra): Full characterization

- Averages: $\langle v_n \rangle, \langle \psi_n \rangle, \dots$
- Probability distributions: $\mathcal{P}(v_n), \mathcal{P}(\psi_n), \dots$
- Correlations: $\langle v_n, v_m \rangle, \langle \psi_n, \psi_m \rangle$

- Problem is that typically the connections depend on the full details of the space-time evolution. . .

$e_n - v_n$ correlation

How are v_n 's and ϵ_n 's related?

- $\langle v_n \rangle \propto \epsilon_n$
- $v_n \propto \epsilon_n$

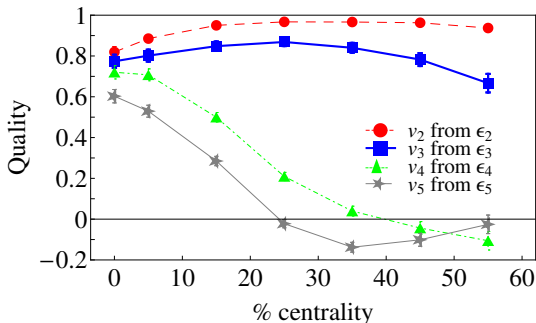
Measure by linear correlation coefficient:

$$c(a, b) = \left\langle \frac{(a - \langle a \rangle_{ev})(b - \langle b \rangle_{ev})}{\sigma_a \sigma_b} \right\rangle_{ev}$$

- $c = 0$ no (linear) correlation
- $c = 1(-1)$ fully (anti-)correlated

$e_n - v_n$ correlation

F. G. Gardim, F. Grassi, M. Luzum and J. -Y. Ollitrault, Nucl. Phys. A904-905 **2013**, 503c (2013)



- v_2 and v_3 strongly correlated to the corresponding eccentricities
- Higher harmonics: no correlation (except in central collisions)

Monte-Carlo Glauber

$$s(x, y) = W \sum_{i=1}^{N_{\text{part, bin}}} \exp \left\{ - \left[(x - x_i)^2 + (y - y_i)^2 \right] / (2\sigma^2) \right\}$$

- (x_i, y_i) position of wounded nucleon or binary collision
- W normalization constant
- Centrality selection according to N_{bin} or N_{part}

p_T -spectra for each event $\left. \frac{dN}{dy dp_T^2 d\phi} \right|_{ev}$

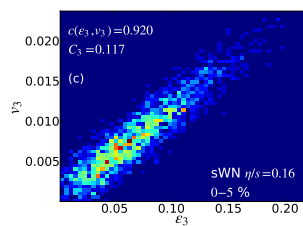
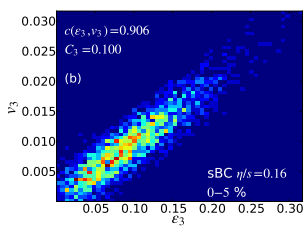
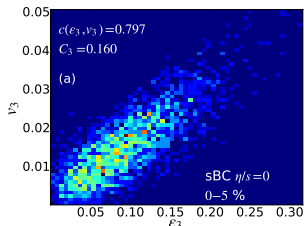
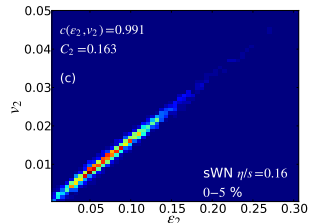
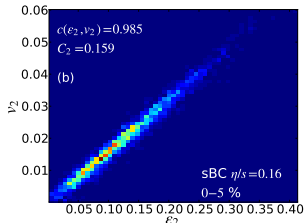
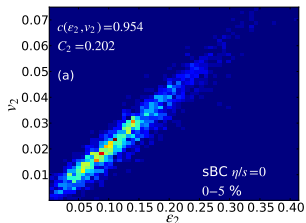
$v_2/3$ vs $\epsilon_{2/3}$ (0 – 5% centrality class)

2d histogram ~ 2000 hydro events per case

sBC $\eta/s = 0$

sBC $\eta/s = 0.16$

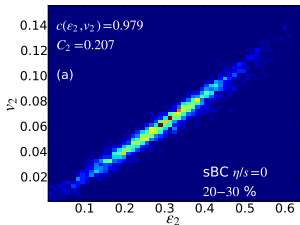
sWN $\eta/s = 0.16$



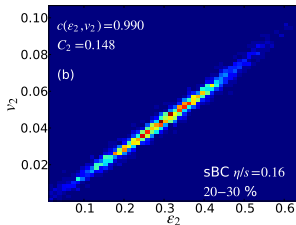
Both v_2 and v_3 correlated to corresponding eccentricities (v_2 event-by-event)

$v_{2/3}$ vs $\epsilon_{2/3}$ (20 – 30% centrality class)

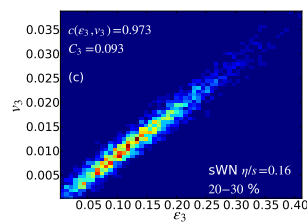
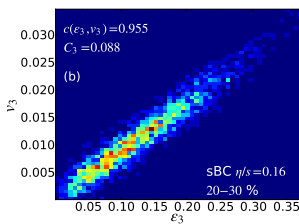
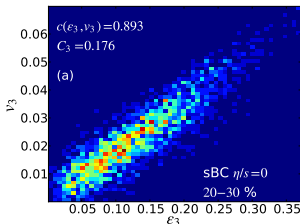
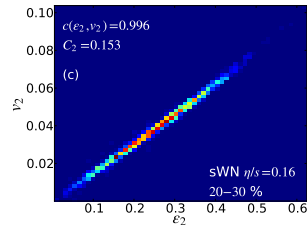
sBC $\eta/s = 0$



sBC $\eta/s = 0.16$



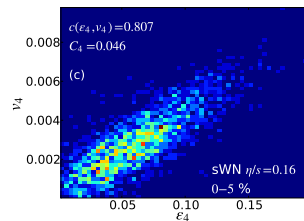
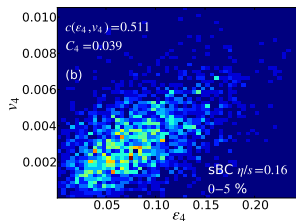
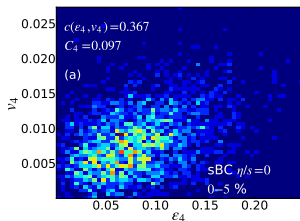
sWN $\eta/s = 0.16$



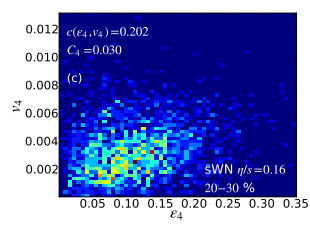
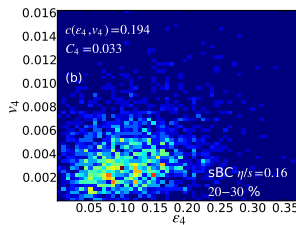
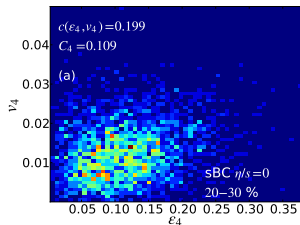
Even better correlation

V₄ VS ε₄

0 – 5 %



20 – 30 %



Weak correlation in central collision, no correlation in peripheral collisions

$e_{m,n} - v_n(p_T)$ correlation

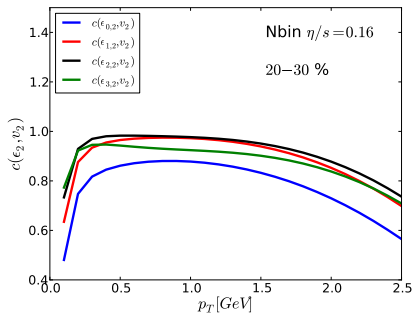
Do $v_n(p_T)$ correlate with different $\epsilon_{m,n}$ at different p_T 's ($\epsilon_{m,n}$ with different r^m weight)

- $v_2(p_T)$: not really.
- $v_3(p_T)$: kind of, but rather weak difference

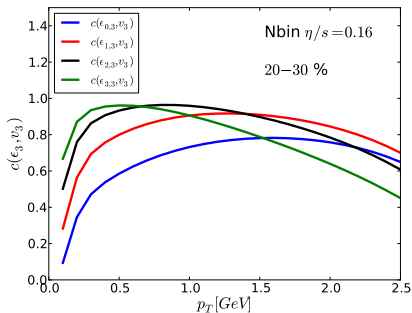
→ No straightforward access to r -dependence of initial density profiles

$c(v_2, e_2)$:

$c(v_3, e_3)$:



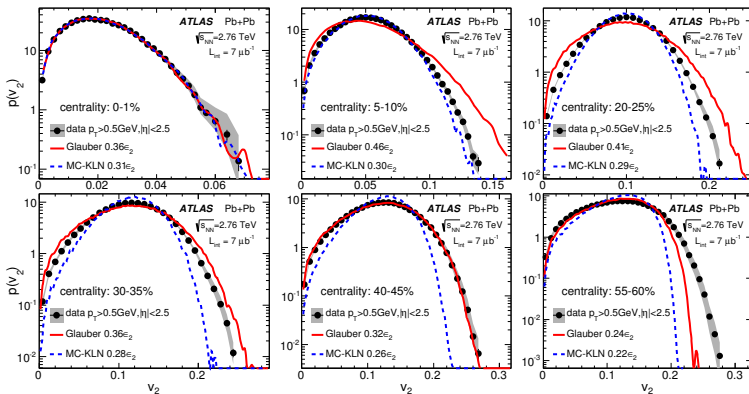
Bessel-Fourier expansion?



v_n distributions

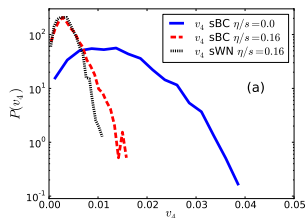
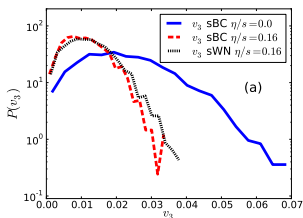
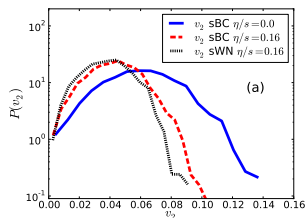
v_n distributions: ATLAS measurement. G. Aad *et al.* [ATLAS Collaboration], arXiv:1305.2942 [hep-ex]

G. Aad *et al.* [ATLAS Collaboration], arXiv:1305.2942 [hep-ex]



v_N distributions

- Full description of heavy-ion collisions should also get distributions of v_N correct (not only event averaged value)
- Each case gives different probability distribution

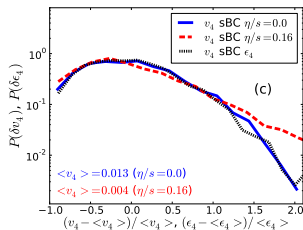
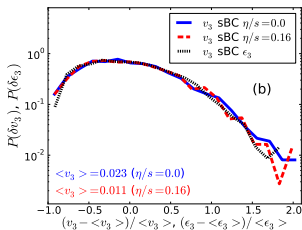
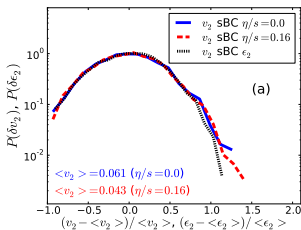


v_n distributions (20 – 30 % centrality)

Each type of initial state and η/s gives different distribution of v_n 's, but if we introduce scaled variable

$$\delta v_n = \frac{v_n - \langle v_n \rangle_{ev}}{\langle v_n \rangle_{ev}}, \quad \text{and} \quad \delta \epsilon_n = \frac{\epsilon_n - \langle \epsilon_n \rangle_{ev}}{\langle \epsilon_n \rangle_{ev}}.$$

Then...

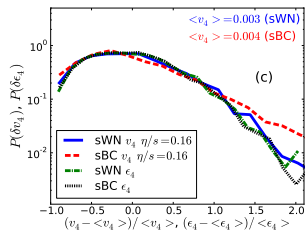
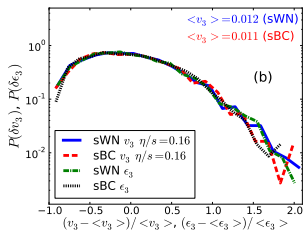
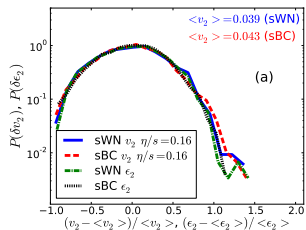


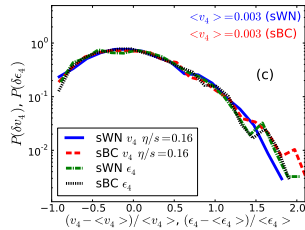
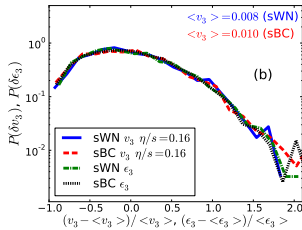
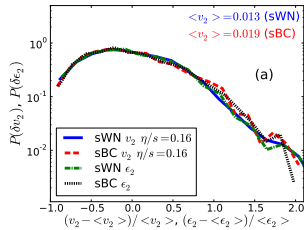
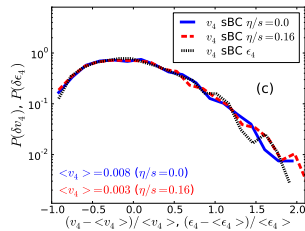
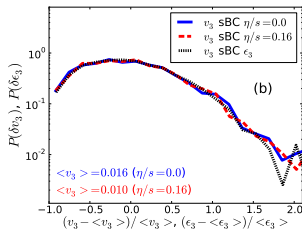
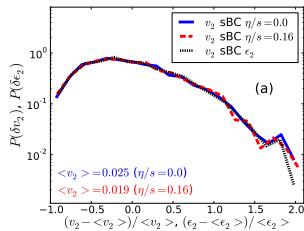
- Probability distributions of δv_2 and δv_3 independent of hydrodynamic evolution, and follow the corresponding $\delta \epsilon_n$ distribution.

HN, Denicol, Holopainen and Huovinen, Phys. Rev. C 87, 054901 (2013), arXiv:1212.1008 [nucl-th]

v_n distributions (20 – 30 % centrality)

Same result if we change between the Glauber variants (sBC or sWN):



v_n distributions (central collisions)

v_2 distributions

$$\mathcal{P}(\delta v_2) = \mathcal{P}(\delta \epsilon_2)$$

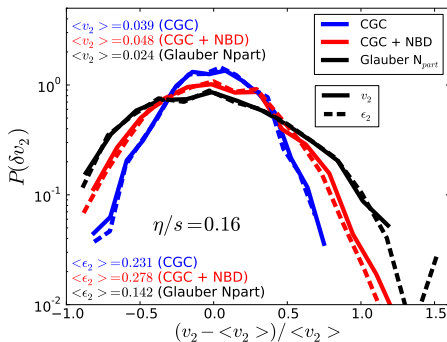
$$\mathcal{P}(\delta v_3) = \mathcal{P}(\delta \epsilon_3)$$

(probes initial condition directly)

and what is this good for...

v_n distributions: MC-KLN (CGC) vs MC-Glauber

Two different Glauber limits give the same distribution, but ...



- For example different variations of MC-KLN model give clearly different δv_2 distributions

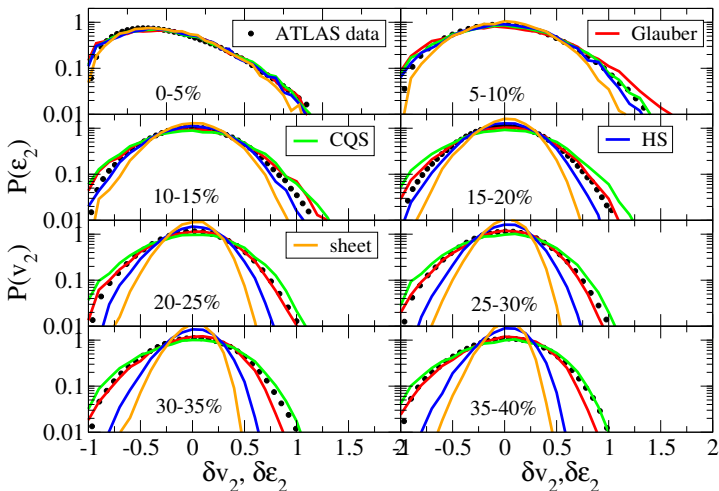
- δv_2 distributions can distinguish between the initial state models (without doing a single hydrodynamic calculation)

Initial conditions from A. Dumitru

Can we save the Glauber model...

v_n distributions: Tuning MC-Glauber model

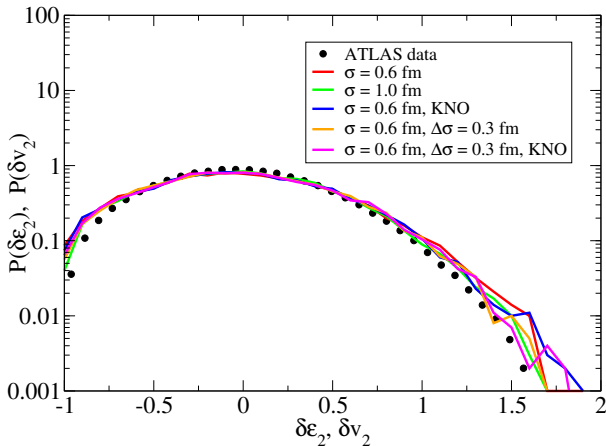
Different distributions of nucleons (or quarks): CQS = constituent quark scaling, HS = hard sphere



v_n distributions: Tuning MC-Glauber model

Change the width of the Gaussian peaks, multiplicity fluctuations (normalization),...

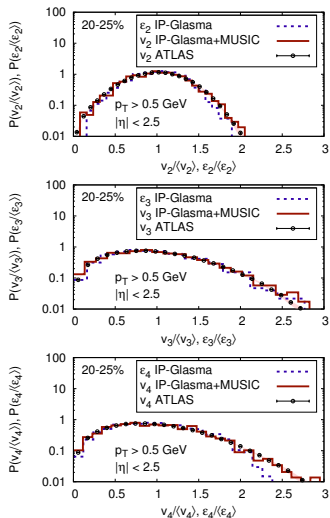
5-10% centrality

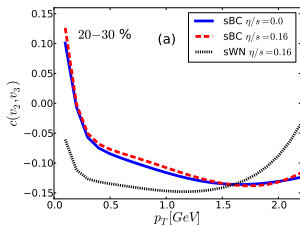
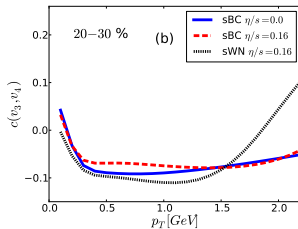
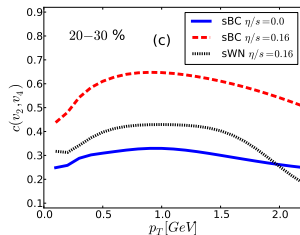


Thorsten Renk and HN

v_n distributions: IP-Glasma

C. Gale, S. Jeon, B. Schenke, P. Tribedy and R. Venugopalan, Phys. Rev. Lett. **110**, 012302 (2013)



(v_i, v_j) correlations $c(v_2, v_3)$  $c(v_3, v_4)$  $c(v_2, v_4)$ 

	$c(\epsilon_2, \epsilon_3)$	$c(v_2, v_3)$	$c(\epsilon_2, \epsilon_4)$	$c(v_2, v_4)$	$c(\epsilon_3, \epsilon_4)$	$c(v_3, v_4)$
sBC $\eta/s = 0.0$	-0.09	-0.11	0.26	0.32	-0.03	-0.11
sBC $\eta/s = 0.16$	-0.09	-0.11	0.25	0.63	-0.03	-0.09
sWN $\eta/s = 0.16$	-0.15	-0.14	-0.04	0.42	0.03	-0.11

Conclusions

- v_2 and v_3 linearly correlated to corresponding eccentricities (v_2 event-by-event)
- Correlation between v_4 and e_4 depends strongly on the details of fluid dynamics
- Distributions $P(\delta v_2)$ follow directly from the initial conditions (details of the fluid dynamical evolution do not matter)
- Clear constraints for the initial condition (without doing a single hydro run)
- (v_i, v_j) correlations provide additional information, but not as easy as the distributions