### Thermalization at weak coupling

Aleksi Kurkela CERN (PH-TH)

... or equilibration in three acts:

- 1. Melting of the coherent fields
- 2. Competition between dynamics and expansion
- 3. Bottom-up thermalization

#### What?

Hydrodynamics provides a fantastic description of HIC, but:

- assumes local thermal equilibrium
  - needs thermalization time  $au_0$  and initial geometry  $T_{\mu
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### How?

- Extremely big and at extremely high energy
  - Saturation scale  $Q_s \gg \Lambda_{QCD}$
  - weak coupling:  $\alpha_s(Q_s) \ll 1$
  - Scale separation allows for effective theories: HTL, kinetic theory.
  - Strong fields, not perturbative.
- Provide description consistent with CGC assumptions
- Philosophy: Do the whole calculation at weak coupling, extrapolate to finite  $\alpha_s$  only at the end (Compare to McLerran's talk)

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Units  $\alpha^{\#}Q_s^{-1}$ , not fm/c in this talk

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# Initial condition: $t \sim Q_s^{-1}$

Blaizot Mueller '87, McLerran Venugopalan '94, Kovner McLerran, Weigert '95, Gelis Epelbaum '13

At weak coupling, well understood: Color Glass Condensate

- Strong, boost invariant fields
- Characteristic transverse coherence length  $\sim Q_s^{-1}$
- Characteristic energy density  $\epsilon \sim Q^4/lpha$



- Boost invariance broken only by quantum fluctuations
- Fields decohere and isotropize in a time scale  $au \sim Q_s^{-1} imes \log(lpha)^{-1}$ 
  - For details: Thomas Epelbaum's talk

The Map:

#### Kurkela, Moore '11

Once the fields have decohered, description in terms of gluons  $(\eta = 0, \tau = t)$ :

 $Q_{s}$ 

- Characteristic, *saturation* scale:
- Highly over-occupied:
- Highly anisotropic:

$$\begin{split} f(p < Q_s) &\sim \alpha^{-c}, \quad c \gtrsim 1 \\ \left\{ \begin{array}{ll} p_t \sim Q_s \\ p_z \sim \alpha^d Q_s \end{array} \right. \end{split}$$



Big question: How to reach the thermal point c = 0, d = 0

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# Competition between dynamics and expansion: c(t), d(t)

Subsequent evolution competition between

- Longitudinal expansion (anisotropizes)
- Momentum broadening due interactions (isotropizes, dilutes)



# Longitudinal expansion

Spatial expansion translates into redshift in  $p_z \sim lpha^d Q_s$ 

- Reduces  $p_z = \alpha^d Q$ ,  $Q_s$  stays constant
- Reduces energy:  $\varepsilon(t) \sim \alpha^{-1} Q_s^4/(Q_s t)$



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Baier, Mueller, Schiff, Son '00



 $(\delta \equiv \alpha^d)$ 



- Along the attractive solution scattering and expansion compete
- Eventually system becomes underoccupied
- Suggested other solution BGLMV, that might stay  $\mathcal{O}(1)$  isotropic.

Blaizot, Gelis, Liao, McLerran, Venugopalan 2011

# Plasma instabilities: Anisotropic screening

#### Isotropic distributions:

• Screening stabilizes soft E-fields

$$\omega_{pl}^2 \sim m^2 \sim \alpha \int d^3 p \frac{f(p)}{p}$$



• B-fields induces a rotation on  $f(p) \rightarrow$  No screening for static B-fields

#### Anisotropic distributions:

- B-field induces non-trivial rotation:
  - Some B-fields stabilized
  - ...others destabilized: Plasma-unstable modes



• Unstable modes grow exponentially...

### Plasma instabilites: Unstable modes

Which *B*-modes become unstable depend on the distribution of gluons Romatshcke, Strickland '03, Arnold, Moore '07
 Strong anisotropy: Weak anisotropy:



### Plasma instabilities: Growth and saturation

Saturation when fields become non-perturbatively strong:

$$D_{\mu} = k_{\mathrm{inst}} + igA_{\mu} \Rightarrow \left\{ egin{array}{l} A \sim k_{\mathrm{inst}}/g \ f(k_{\mathrm{inst}}) \sim 1/lpha \end{array} 
ight.$$



Arnold, Moore, Yaffe '07, Bödeker, Rummukainen '07

Hard Loop:  $Q_s$  modes as currents, unstable modes classical YM

### Plasma instabilites: Momentum diffusion $\hat{q}$

Hard parton traveling through magnetic fields receives coherent kicks from patches of same-sign magnetic fields



### Attractor with plasma instabilities

Broadening of the hard particle distribution dominated by the plasma instabilities  $(\hat{q}_{el} \ll \hat{q}_{inst})$  originating from the scale  $Q_s$ 



BMSS: Baier, Mueller, Schiff, Son 2009, KM: Kurkela, Moore 2011, BGLMV: Blaizot, Gelis, Liao, McLerran, Venugopalan 2011

# Numerical approaches



- Hard Expanding Loops: Assumes free streaming for the particles, sees the growth of plasma instabilities.
- BBSV: Classical YM simulations in expanding background at (c = 1, d = 0)-point. Favours BMSS. Not yet evidence of plasma instabilities.

- 1. Melting of the coherent fields
- 2. Competition between dynamics and expansion
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  - ...or how very dilute and anisotropic systems equilibrate



# Soft radiation

Physics of under occupied qualitatively different from over occupied. Formation of soft thermal bath:



• Soft modes quick to emit

 $n_s \sim \alpha n_{col}$ 



- Low p: easy to bend ⇒thermalize quickly
- Can dominate dynamics!
  - (i.e. scattering, screening, ...)

 $\Rightarrow$  Right way to think: Few energetic "jets" propagating in thermal bath

# Hard splitting

Baier, Mueller, Schiff & Son '00

 $Q_s$  modes break before they bend!



- In vacuum: on-shell particles, no splitting
- In medium: Particles receive small kicks frequently
  - For stochastic uncorrelated kicks: Brownian motion in *p*-space

$$\Delta p_{\perp}^2 \sim \hat{\boldsymbol{q}} t, \qquad t_{\mathrm{split}}(k) \sim l_{\mathrm{stop}}(k) \sim \alpha^{-1} \underbrace{\sqrt{k/\hat{\boldsymbol{q}}}}_{t_{\mathrm{form}}} \qquad (\mathrm{LPM})$$

Physics of Landau-Pomeranchuk-Migdal suppression important

## Bottom-Up

Thermal bath eats the hard particles away:



• Scales below  $k_{\text{split}}$  have cascaded down to T-bath

 $t_{
m split}(k_{
m split}) \sim t \Rightarrow k_{
m split} \sim lpha^2 \hat{q} t^2$ 

• "Falling" particles heat up the thermal bath

$$T^4 \sim k_{
m split} \int d^3 p f(p)$$

• Thermalization when  $Q_s$  gets eaten

$$k_{
m split} \sim Q_s$$

Thermalization time determined by how strongly hard particle motion is affected by the thermal bath,  $\hat{q}$ 

### Bottom up:

If the  $\hat{q}$  is dominated by elastic scattering with the soft thermal bath

$$\begin{cases} \hat{q} \sim \alpha^2 T^3 \\ T \sim \alpha^3 Q_s(Q_s t) \\ k_{\text{split}} \sim \alpha^{13} (Q_s t)^5 Q \\ \tau_0 \sim \alpha^{-13/5} Q_s^{-1} \end{cases}$$



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However, the thermal bath is slightly anisotropic and can give rise to plasma instabilities and increase  $\hat{q}$  to faster thermalization Kurkela, Moore '11

$$\begin{cases} \hat{q} \sim \alpha^3 Q_s^3 \\ T \sim \alpha Q_s (Q_s t)^{1/4} \\ k_{\rm split} \sim \alpha^5 (Q t)^2 Q_s \\ \tau_0 \sim \alpha^{-5/2} Q_s^{-1} \end{cases}$$



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- Role of plasma instabilities under investigation
- Thermalization time  $\tau \sim \# \alpha^{-5/2} Q_s^{-1}$ .
  - Is this long or short? Depends on #. Numerics needed.



Momentum scales vs. time scales

### Comparison to other work



BGLMV: Blaizot, Gelis, Liao, McLerran, Venugopalan arXiv:1107.5296 BMSS: Baier, Mueller, Schiff, Son hep-ph/0009237 HEL: Rebhan, Stricland arXiv:1207.5795 CGC: Gelis, Iancu, Jalilian-Marian, Venugopalan arXiv:1002.0333

BS: Berges, Schlichting: arXiv:1209.0817