



# Results from T2K



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ND280

T2K

- T2K
  - Overview
  - Tokai
  - Kamioka
- Neutrino Oscillations
- Results
  - Muon neutrino disappearance
  - Electron neutrino appearance
  - Charged Current inclusive cross section
- Future Prospects
  - Current Run
  - Future Running



# T2K Overview

## *Physics Goals*

### **T2K - Long Baseline Accelerator based Neutrino Experiment**

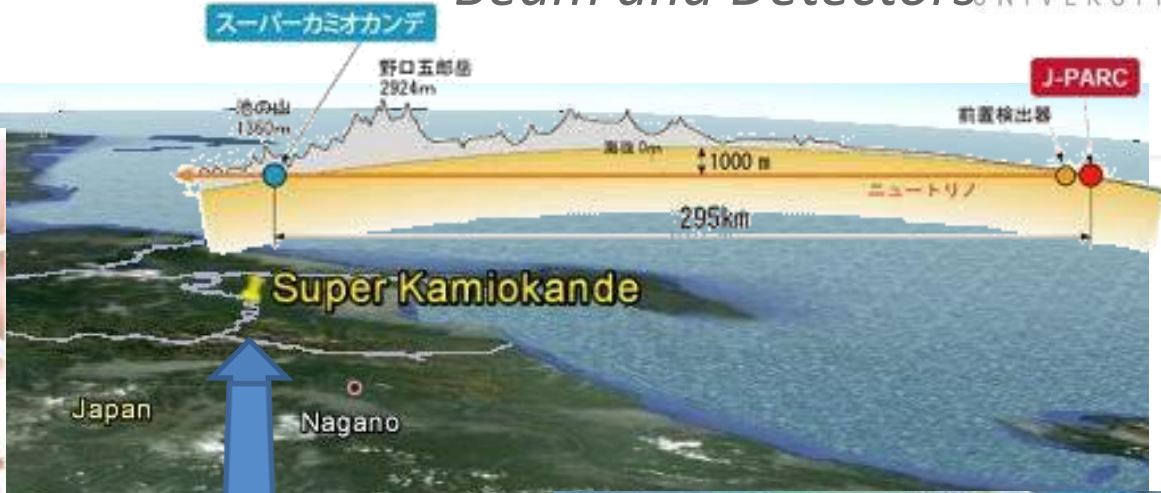
- Observe  $\nu_e$  appearance and measure  $\theta_{13}$
- Observe  $\nu_\mu$  disappearance and measure  $\theta_{23}$
- Measure  $\nu$  cross sections
- Search for exotic neutrinos

# T2K

# T2K Overview

## Beam and Detectors

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nizu

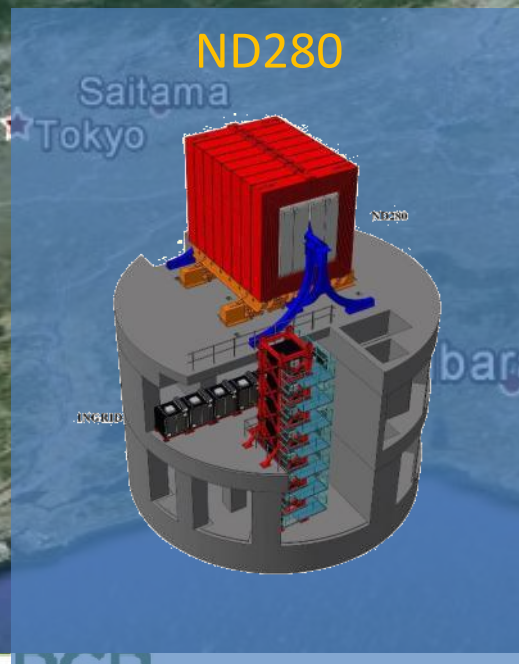
Japan

Nagano

Super Kamiokande

Chiba

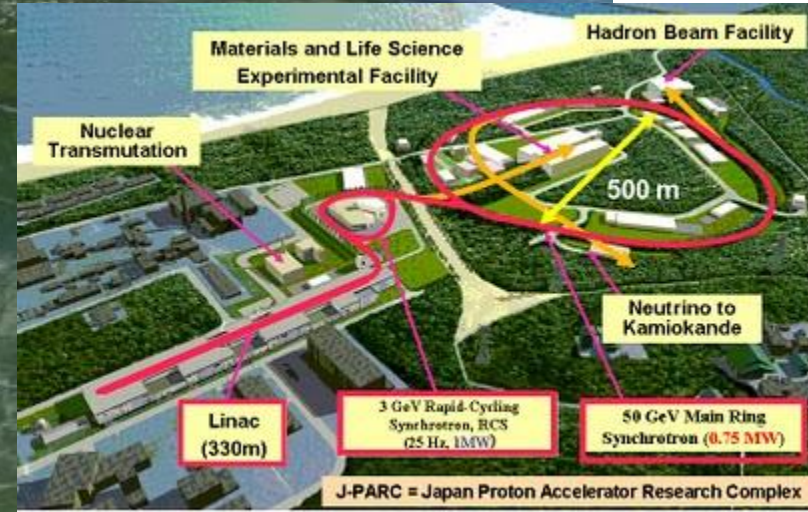
Asahi



Tochigi

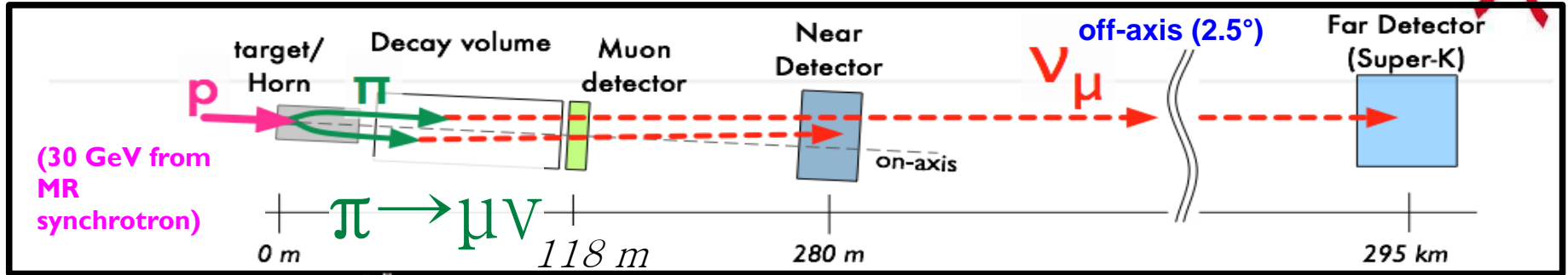
J-PARC

Tokai

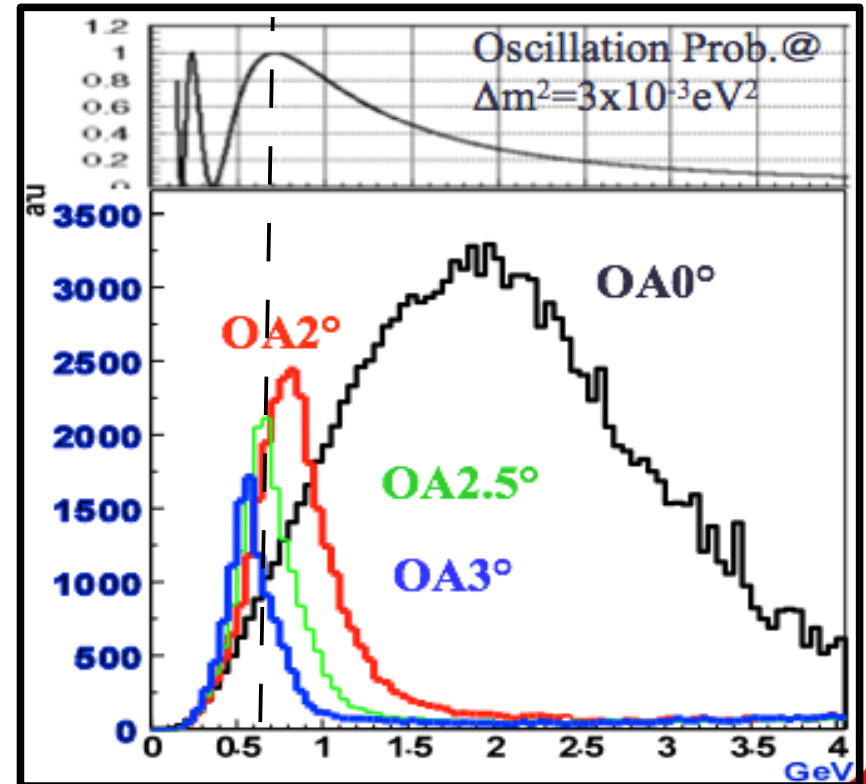
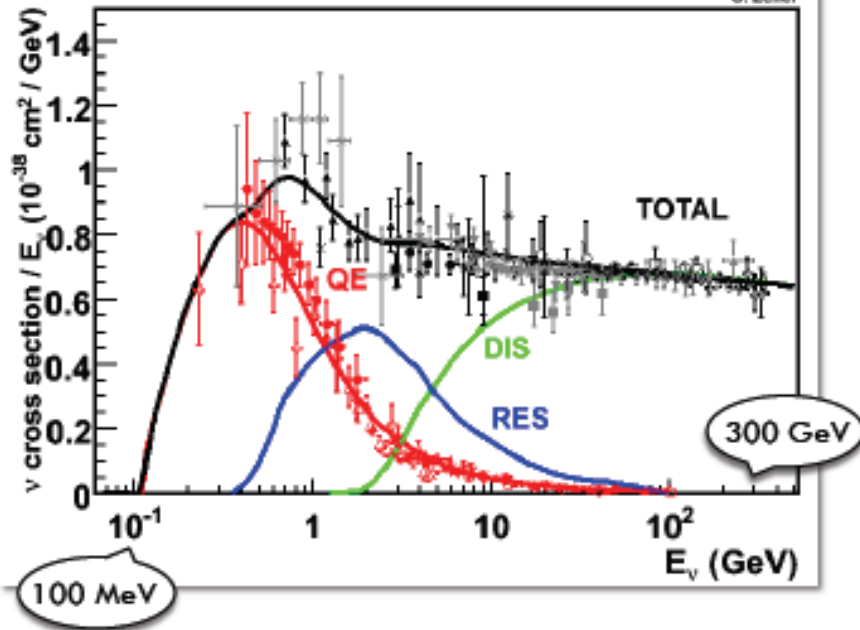


Iwaki

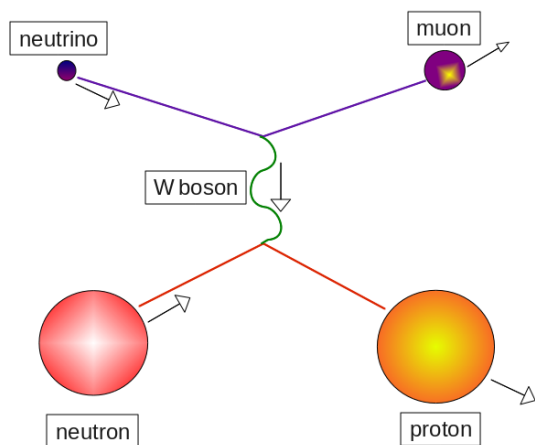




neutrino

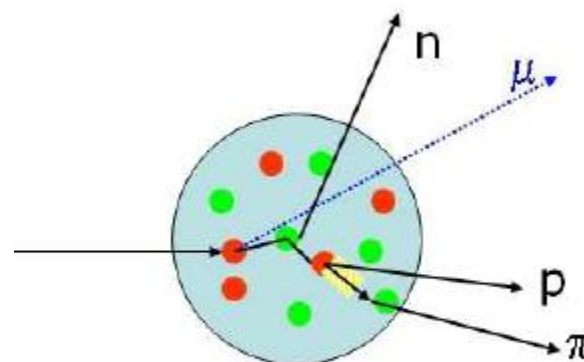


What we'd like:

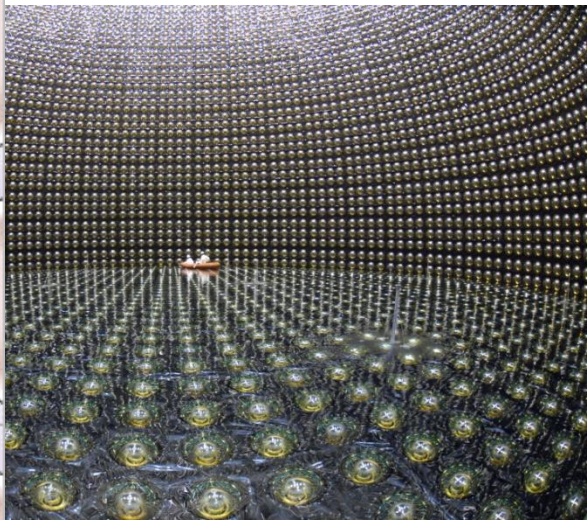
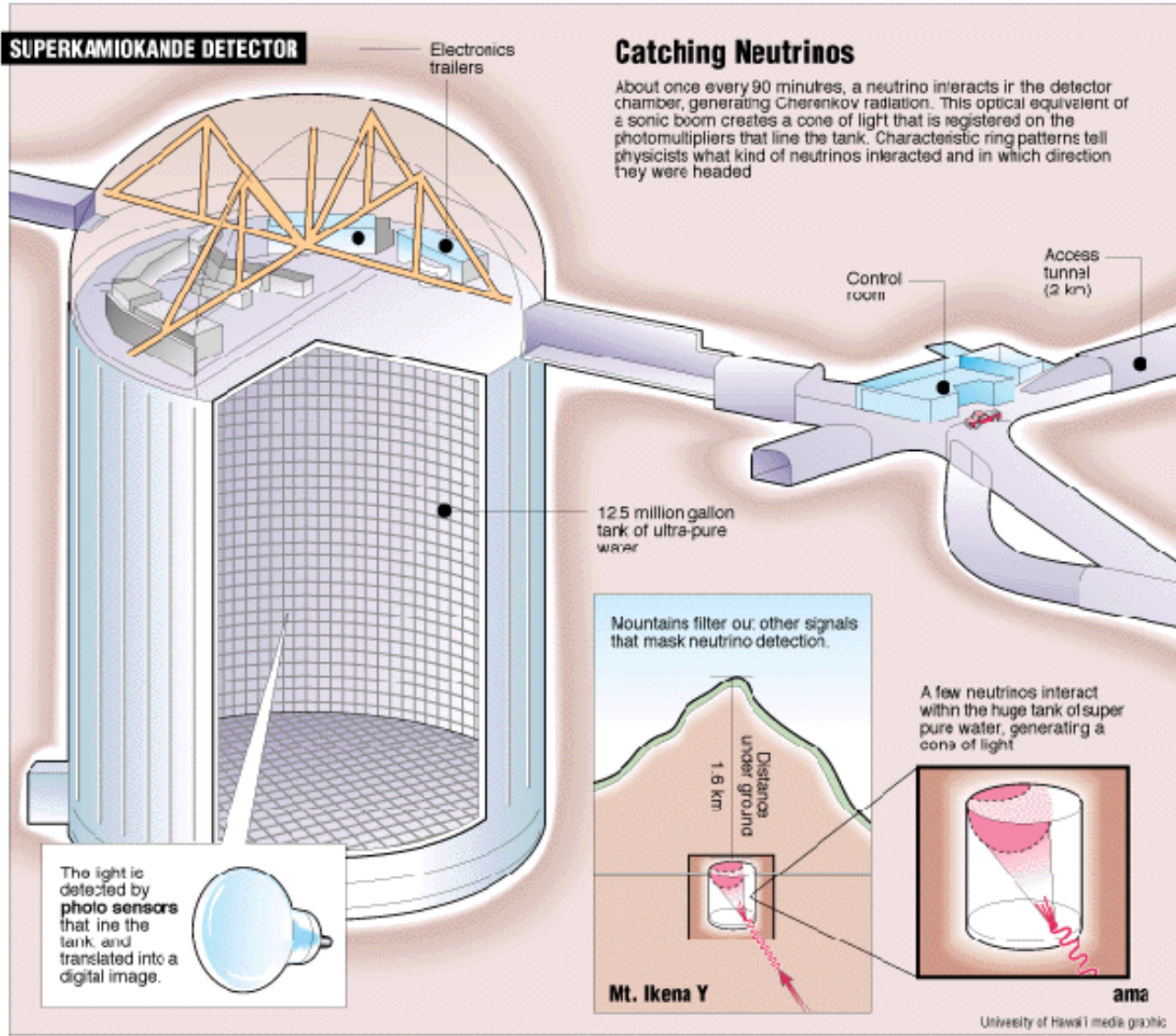


Nucleon changes but doesn't break up.

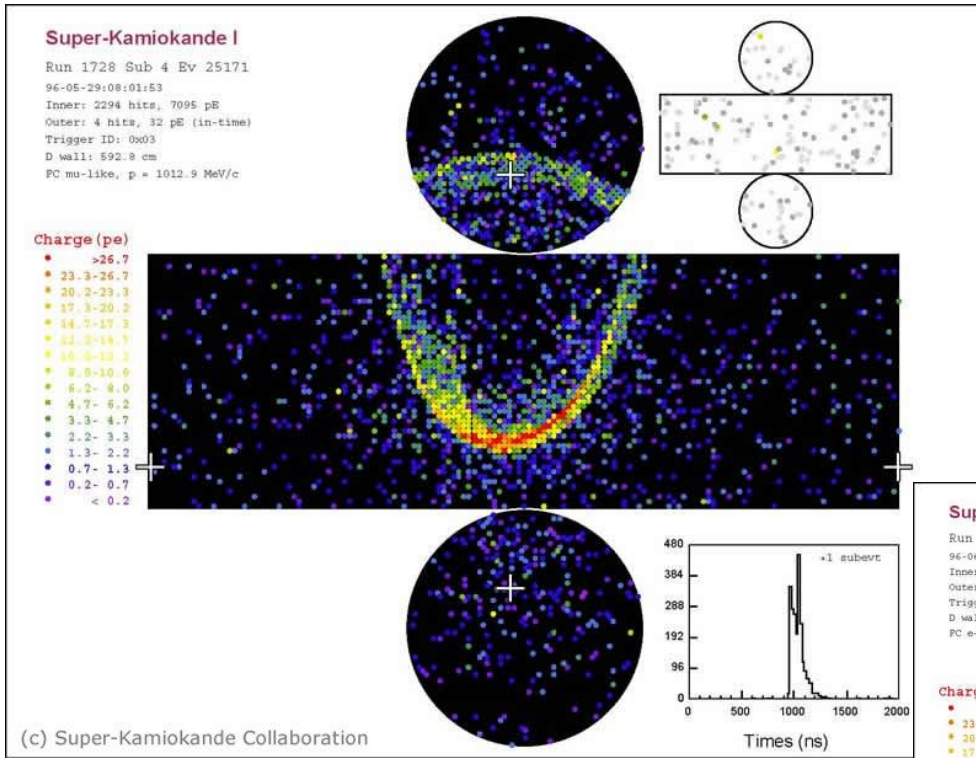
What we have:



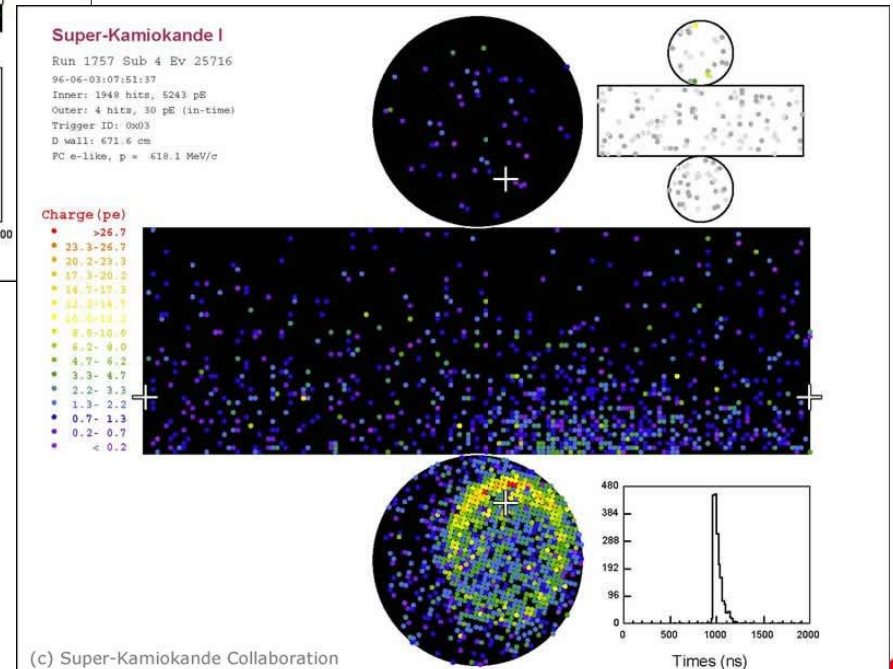
“Final State Interactions”  
confuse the picture







Muon event



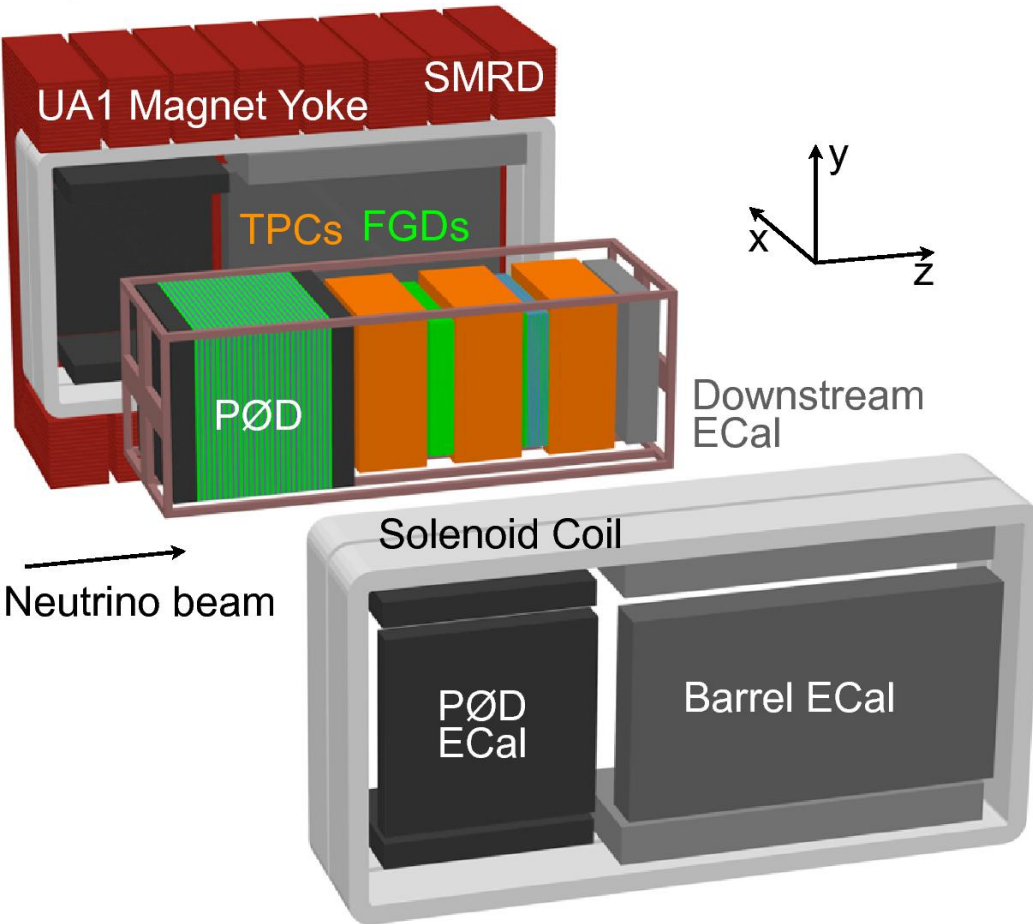
Electron event



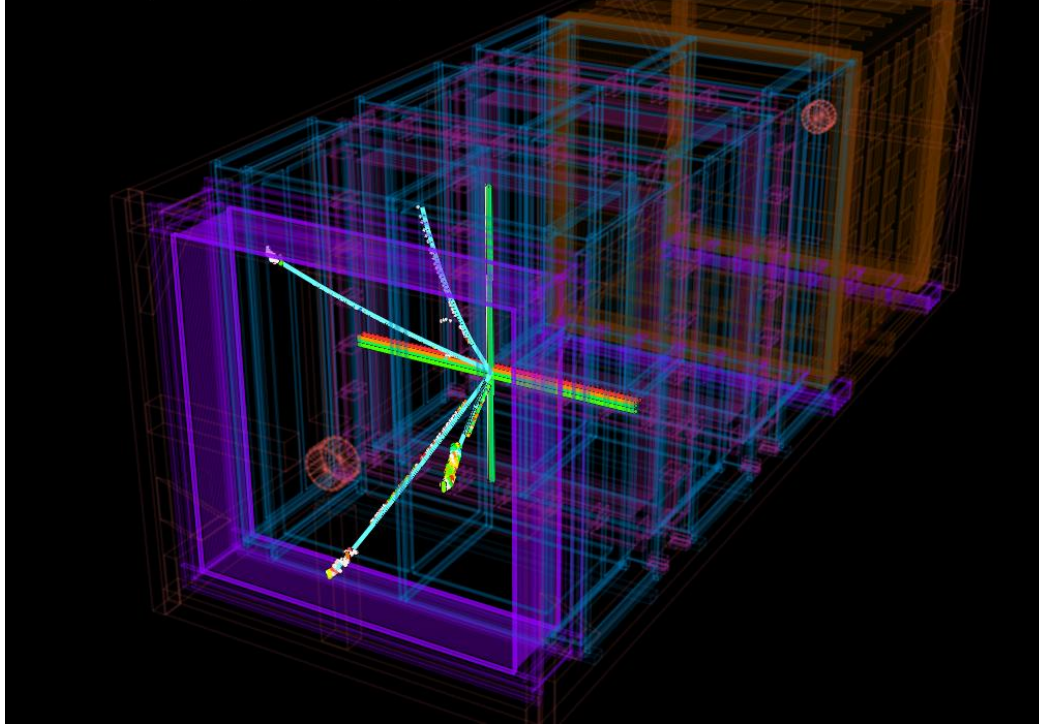
# T2K

# ND280

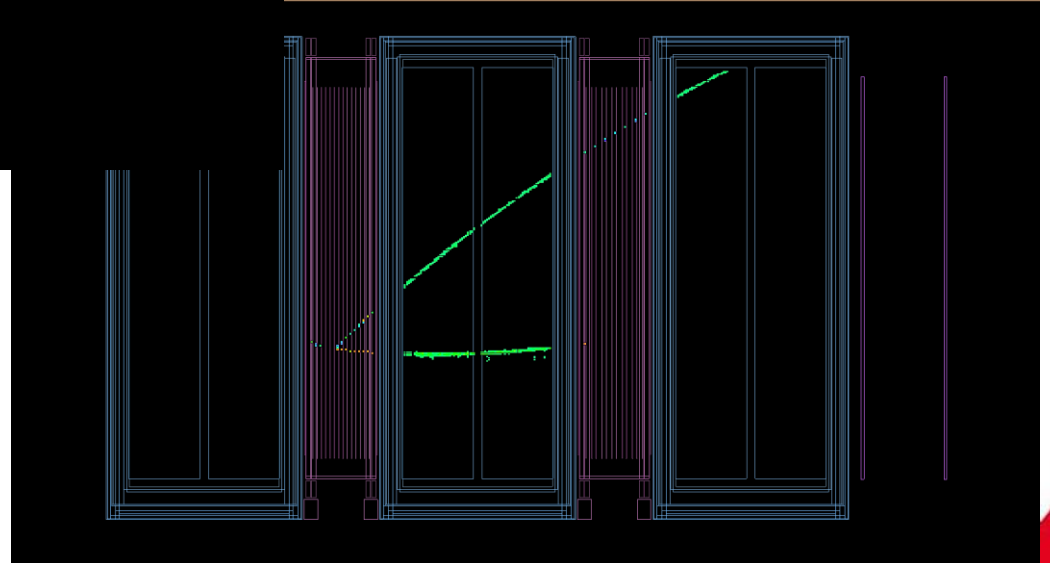
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Event number : 1184 | Partition : 63 | Run number : 4173 | Spill : 54560 | SubRun number : 0 | Time : Sat 2010-03-20 02:43:09 JST | Trigger: Beam Spill

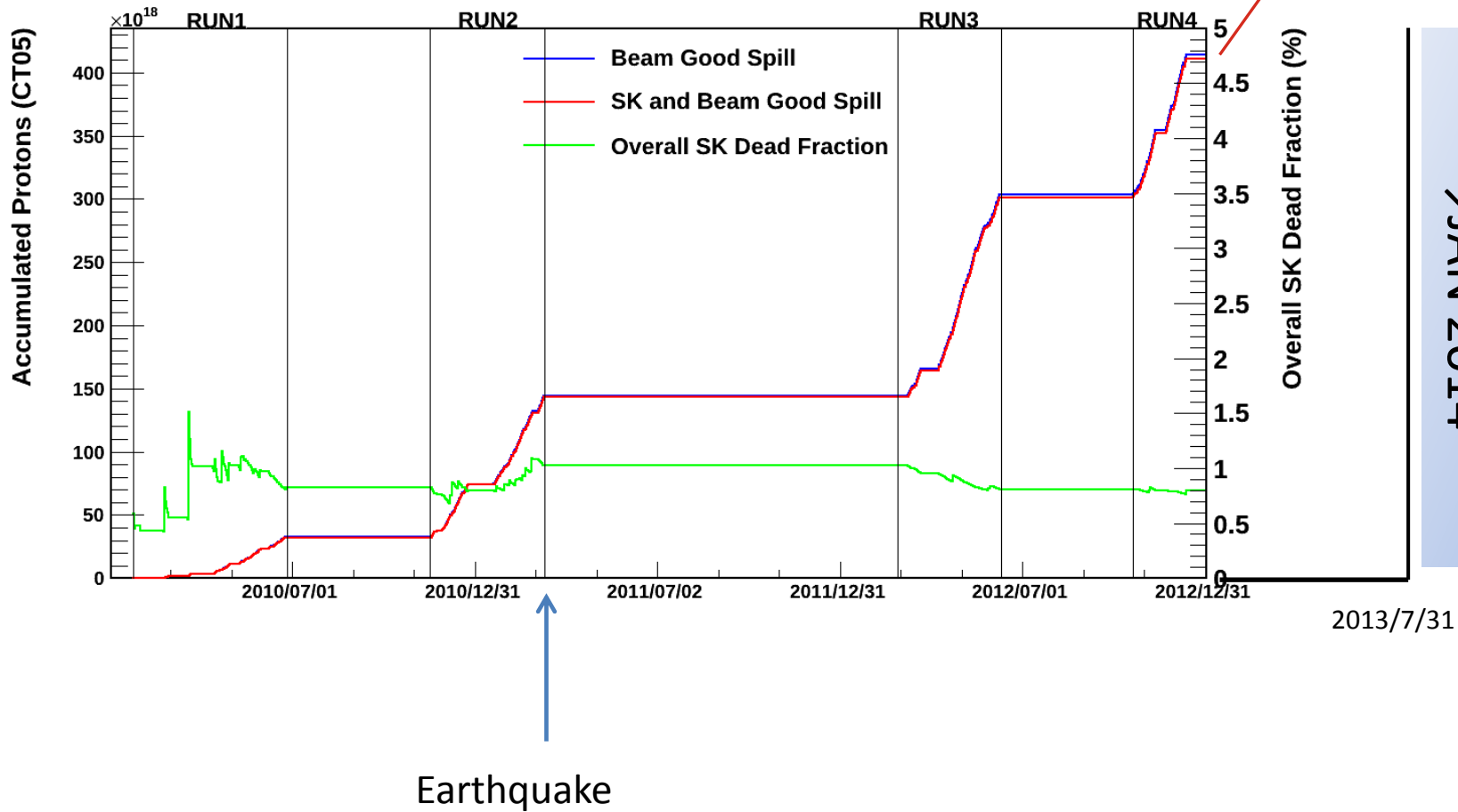


Event number : 4200 | Partition : 63 | Run number : 4173 | Spill : 54560 | SubRun number : 6 | Time : Sun 2010-03-21 22:33:25 JST | Trigger: Beam Spill





?



Long Shutdown  
→ JAN 2014



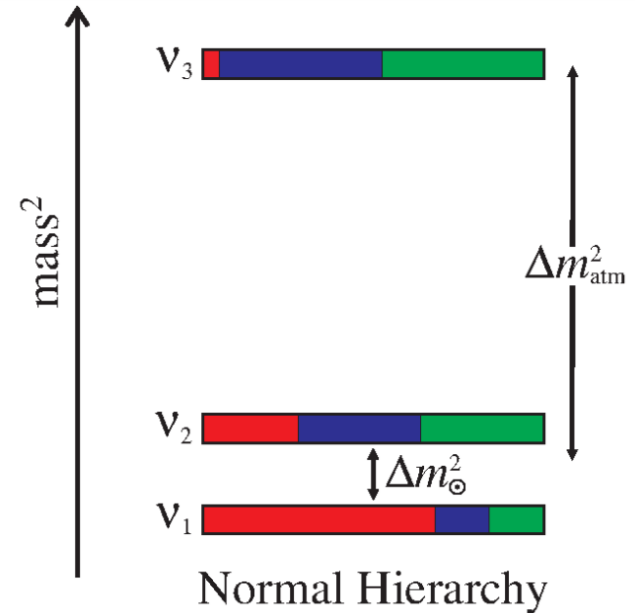
PMNS=>Pontecorvo-Maki-Nakagawa-Sakata

Neutrino mixing (PMNS) matrix is:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



\*Plotted for  $U_{e3} \pm \sqrt{0.05}i$



$$\Delta m^2_{21} \equiv \Delta m^2_{\text{sol}} = 7.6 \times 10^{-5} \text{ eV}^2$$

$$|\Delta m^2_{31}| \approx |\Delta m^2_{32}| \equiv \Delta m^2_{\text{atm}} = 2.4 \times 10^{-3} \text{ eV}^2$$

$$\Theta_{13} = 9^\circ$$

$$\Theta_{12} = 34^\circ$$

$$\Theta_{23} = 45^\circ$$



The mixing matrix is commonly parameterised as the product of two rotations and a unitary transformation. Writing  $s_{ij} = \sin\theta_{ij}$ , and  $c_{ij} = \cos\theta_{ij}$ :

$$\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}$$

The choice of parameterisation is convenient as the **solar** and **atmospheric** disappearance amplitudes can be approximated as functions of  $\theta_{12}$  and  $\theta_{23}$ , respectively. This approximation only works to the extent that the third angle  $\theta_{13}$  is small.

In the standard parameterisation, it turns out that

$$U_{e3} = \sin \theta_{13} e^{-i\delta}, \quad \text{and therefore} \quad \sin \theta_{13} = |U_{e3}|.$$

The value of  $\sin \theta_{13}$  particularly significant because a zero element in the mixing matrix would have eliminated the possibility of (KM-mechanism) leptonic CP violation.

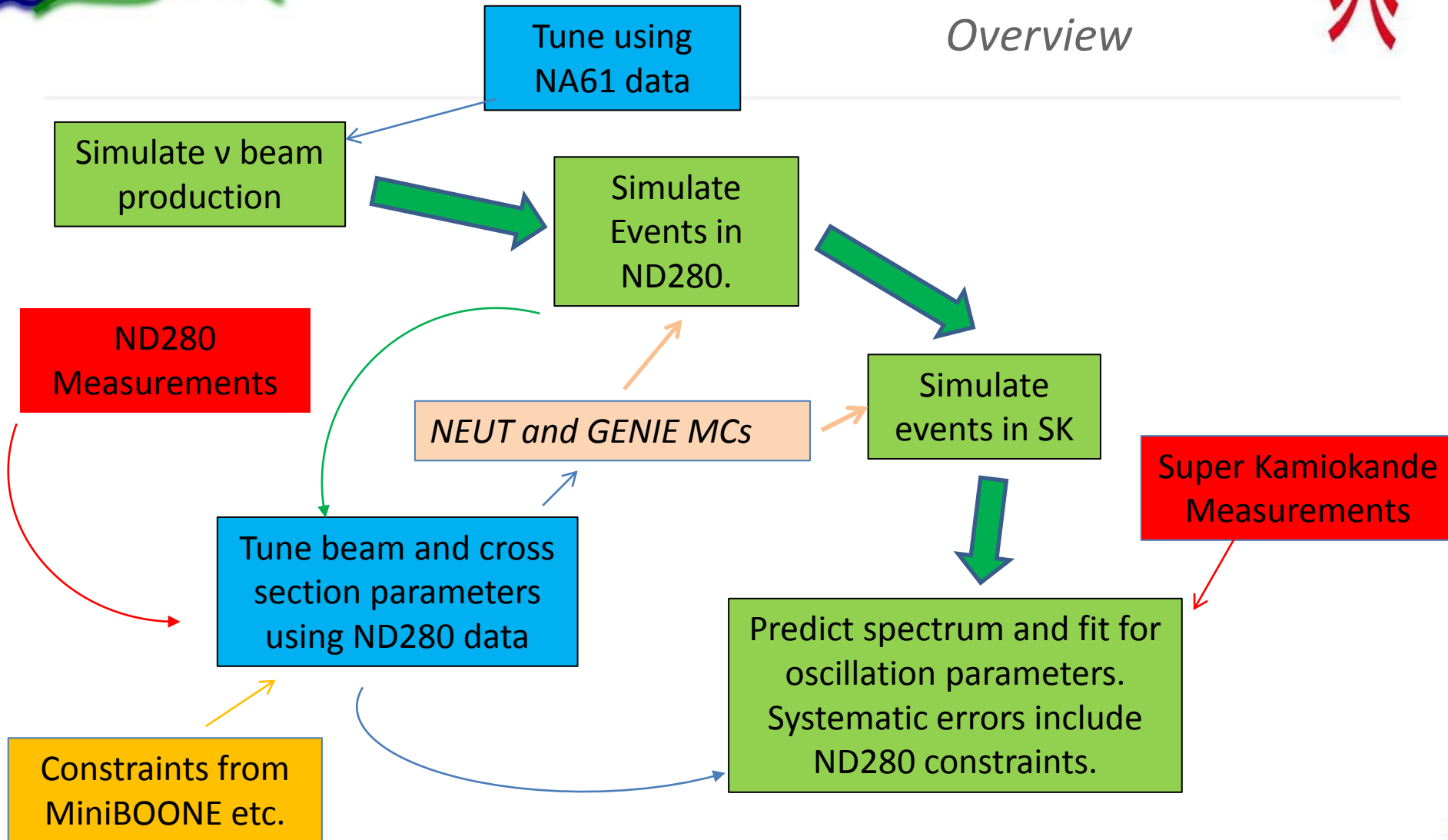
**The future program of neutrino physics is strongly dependent on the size of  $\theta_{13}$ .**

To study, need channels involving  $\langle \nu_e | \nu_3 \rangle$ . The most accessible are  $\bar{\nu}_e \rightarrow \bar{\nu}_e$  (reactor) and  $\nu_\mu \rightarrow \nu_e$  (**accelerator**) at first 'atmospheric' maximum ( $L/\text{km} \sim 0.5 \times E/\text{MeV}$ )

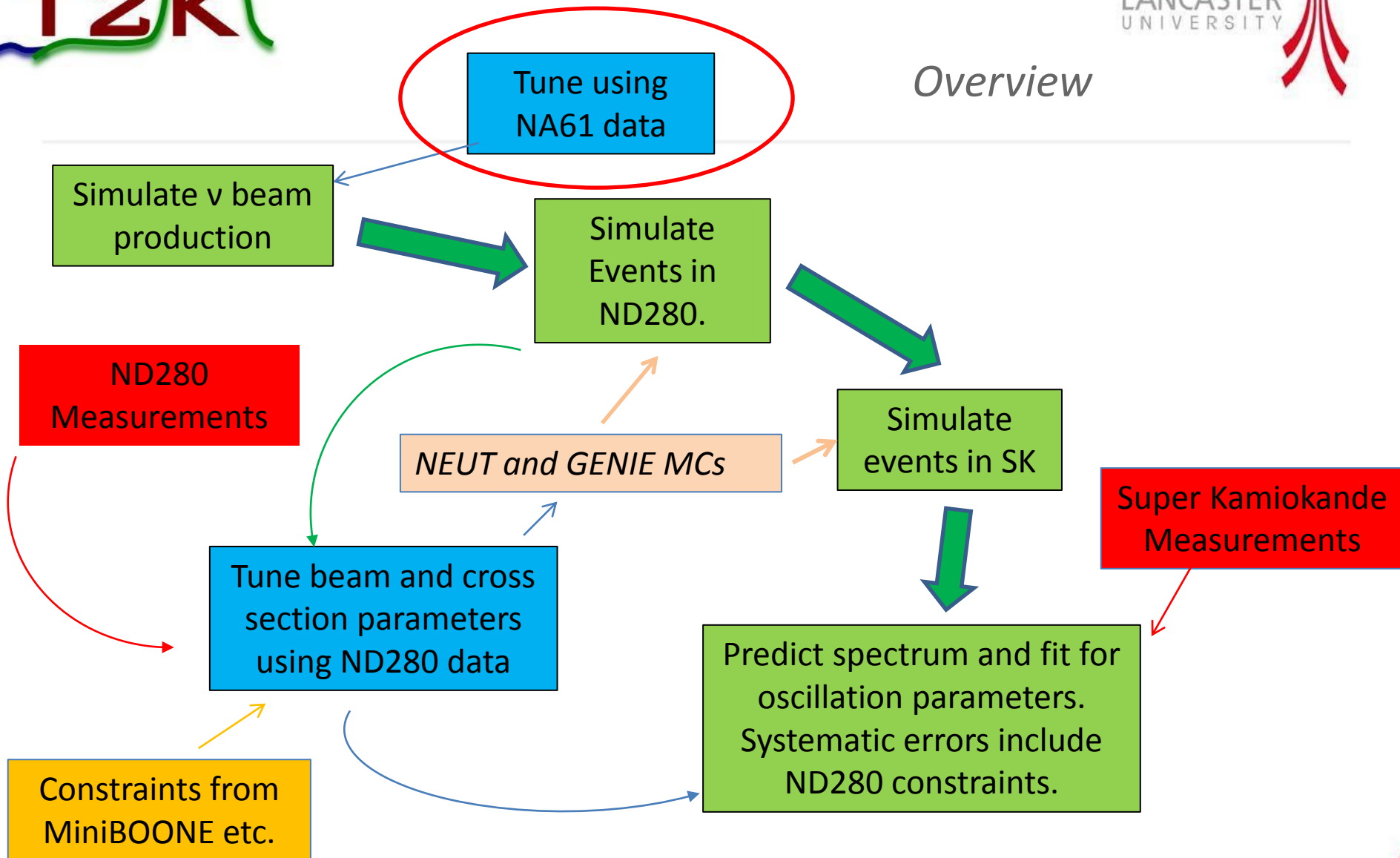




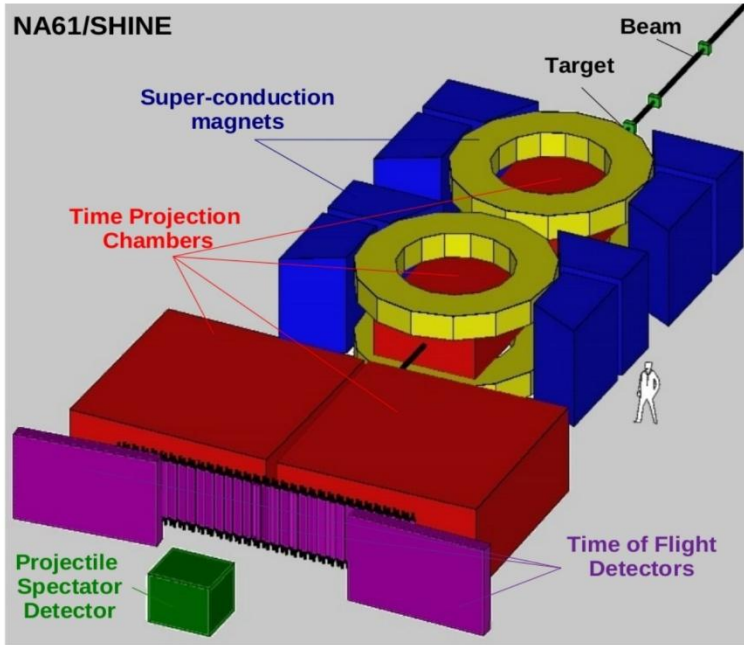
## Overview



Use event reweighting to model effect of varying parameters.



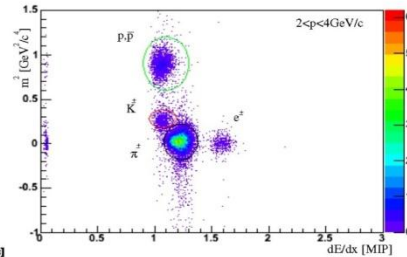
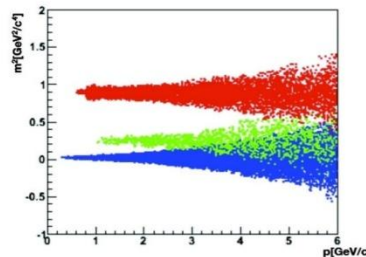
Use event reweighting to model effect of varying parameters.



NA61 is a large acceptance hadron spectrometer with excellent capabilities for momentum, charge and mass measurements. The experimental facility consists of Time Projection Chambers, Time of Flight and Projectile Spectator Detectors.

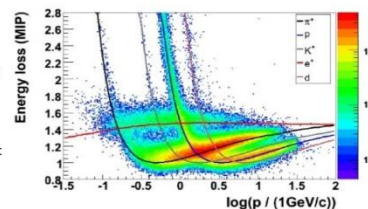
### Physics goals

... Hadron production reference measurements for neutrino (T2K) and cosmic-ray (Pierre Auger Observatory, KASCADE-Grande and KASCADE) experiments. ...



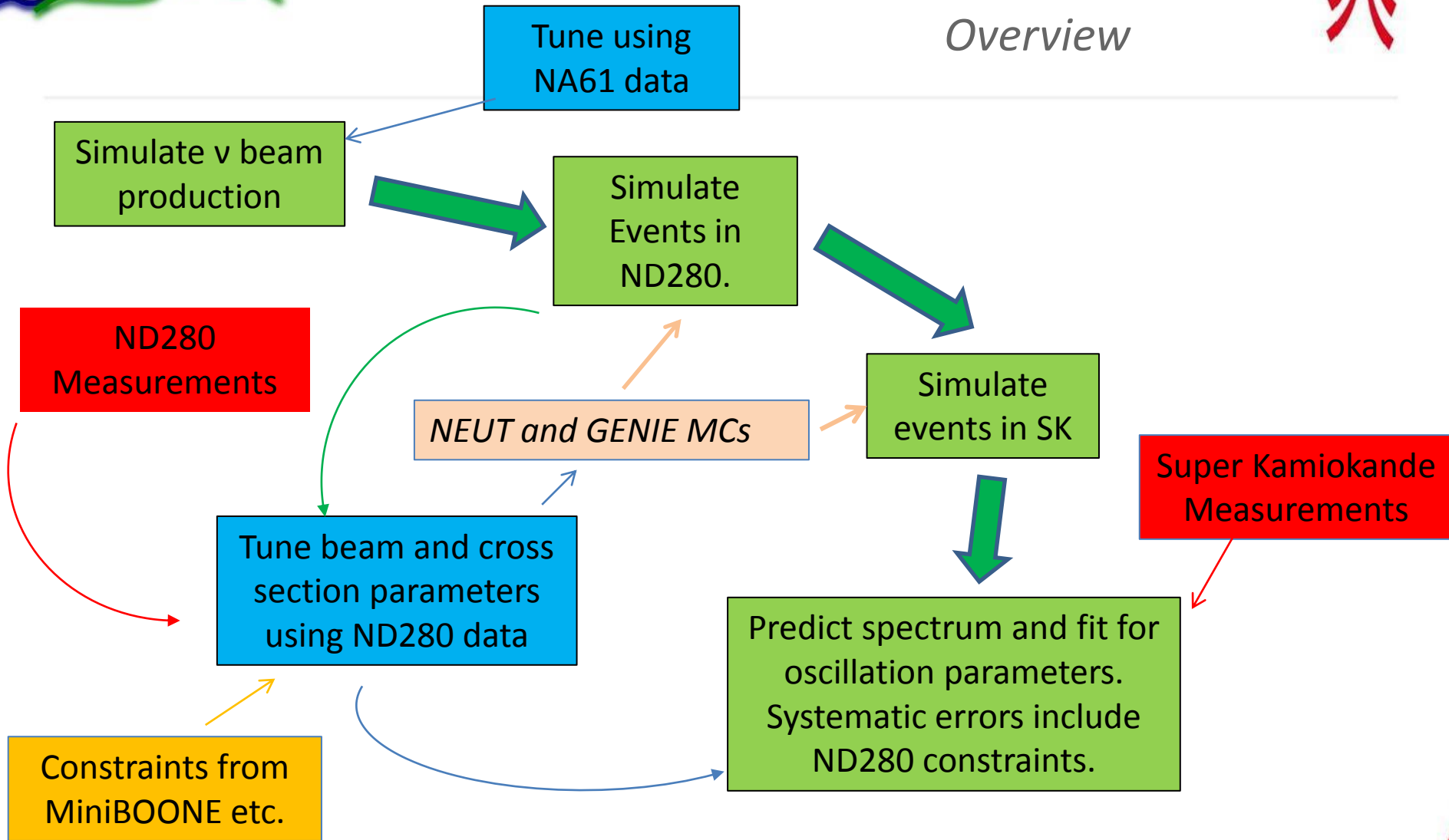
Particle identification capabilities for protons, kaons, pions and electrons:

- Specific energy loss vs. momentum (Bethe-Bloch plot) - lower right
- Reconstructed mass vs. specific energy loss - upper right
- Reconstructed mass vs. momentum - upper left

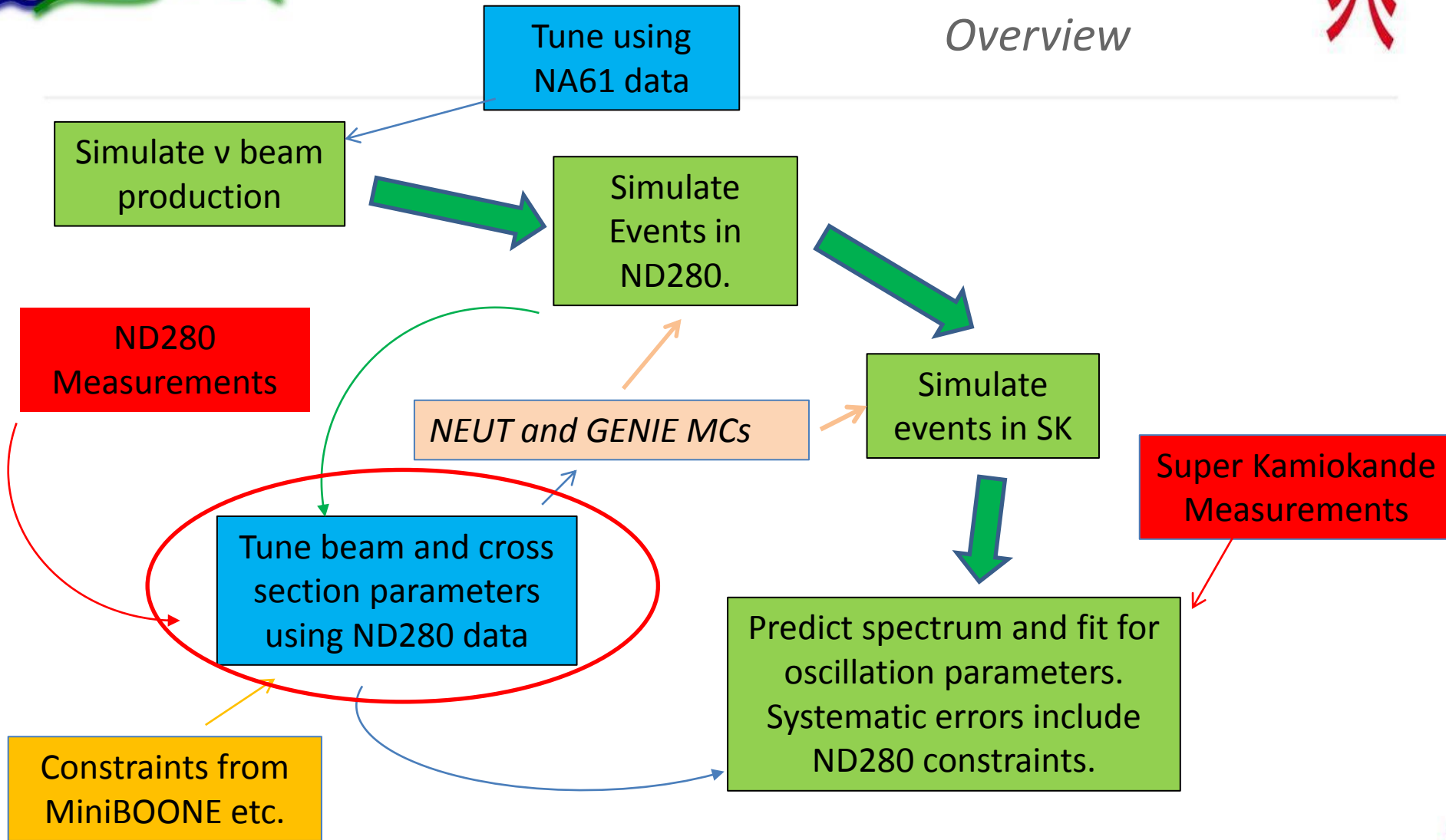


Phys.Rev. C84 (2011) 034604

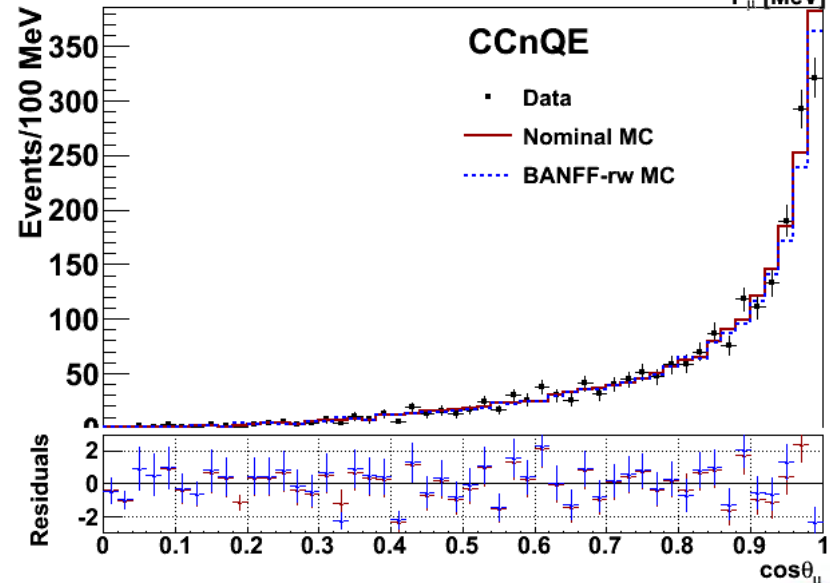
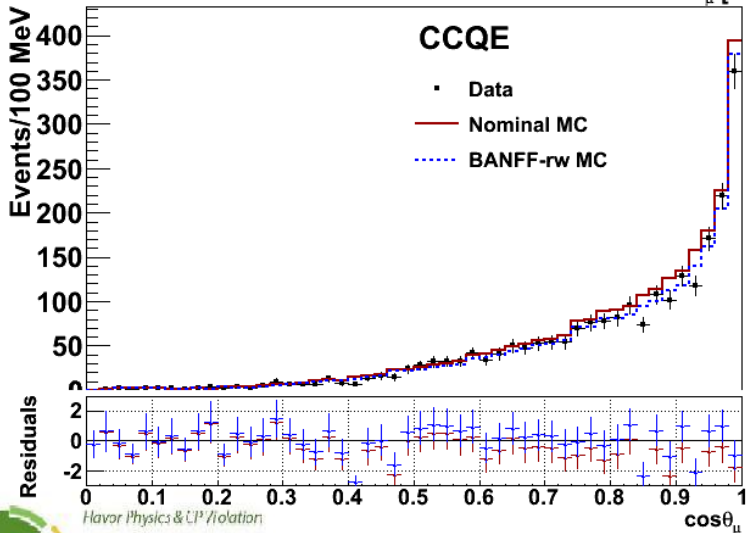
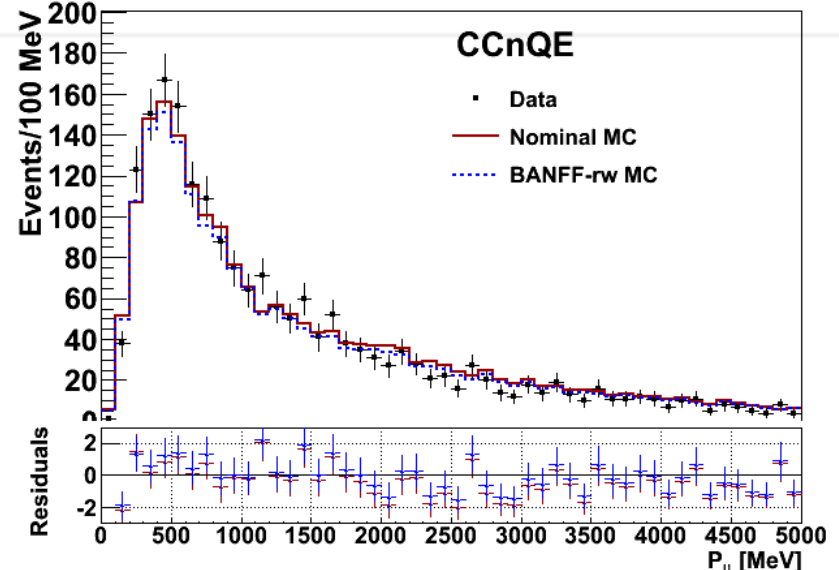
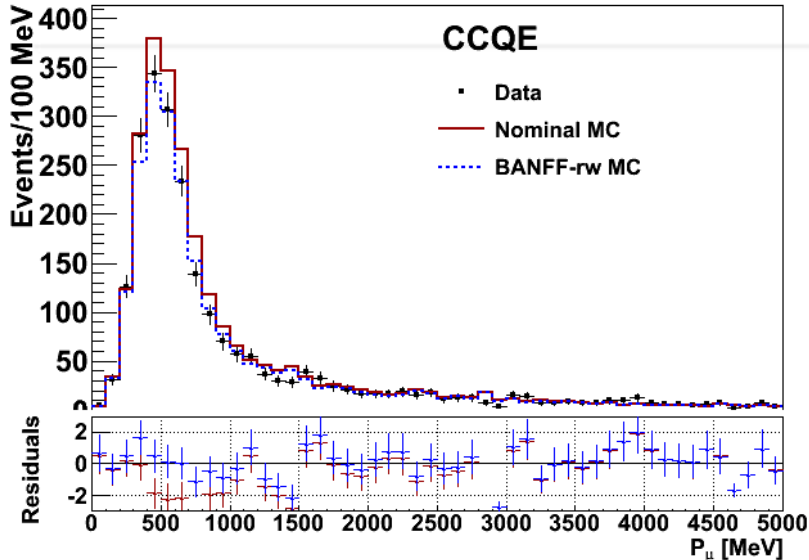


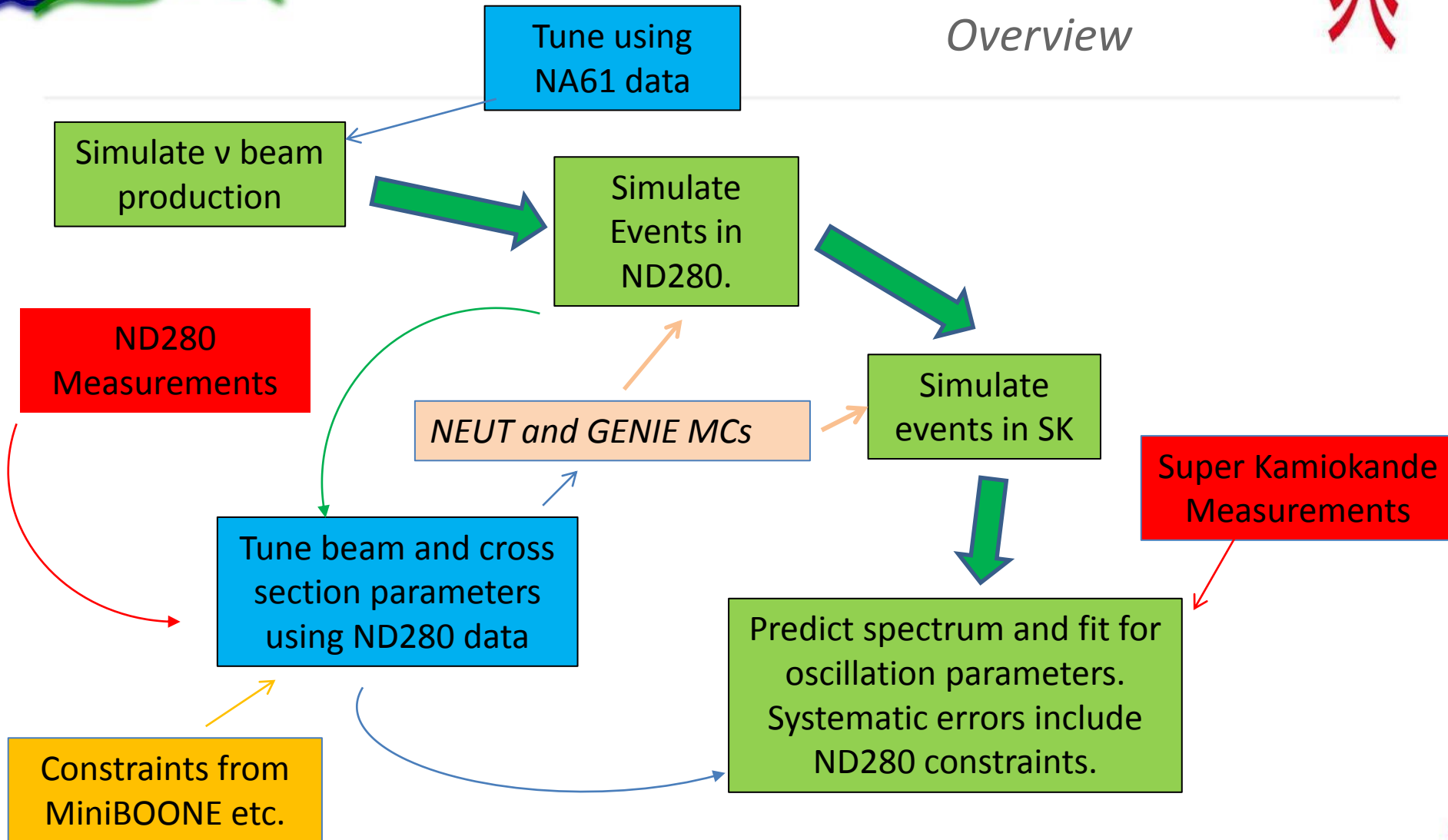


*Use event reweighting to model effect of varying parameters.*



Use event reweighting to model effect of varying parameters.



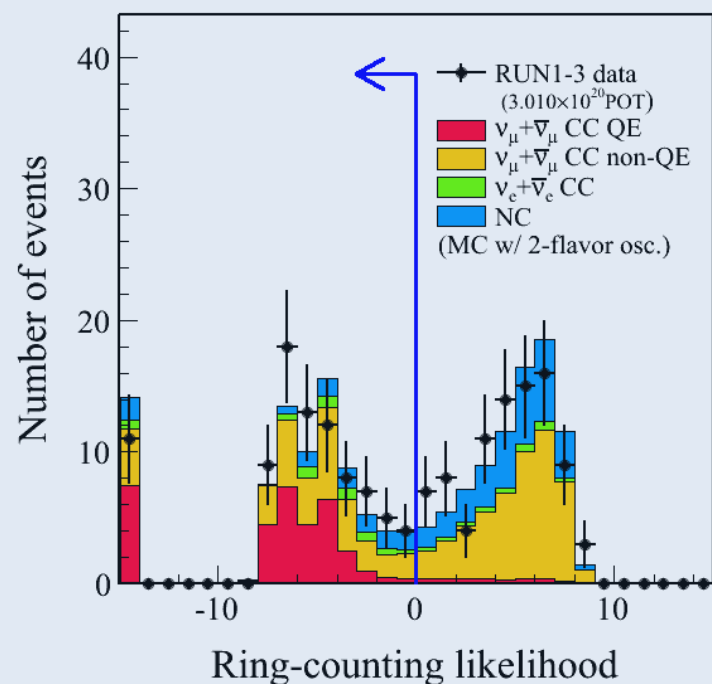
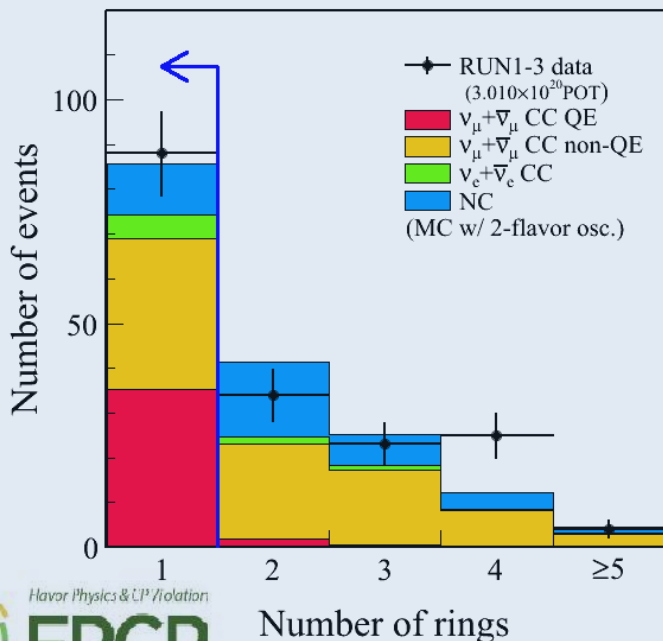
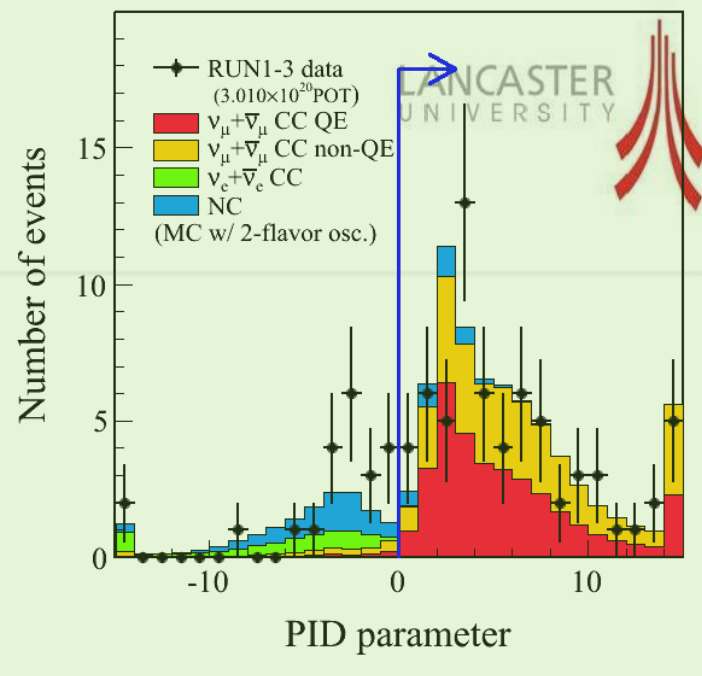
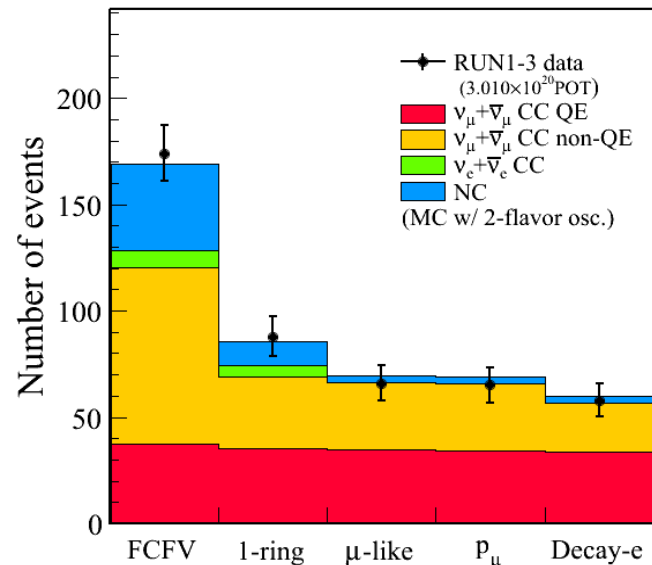


*Use event reweighting to model effect of varying parameters.*

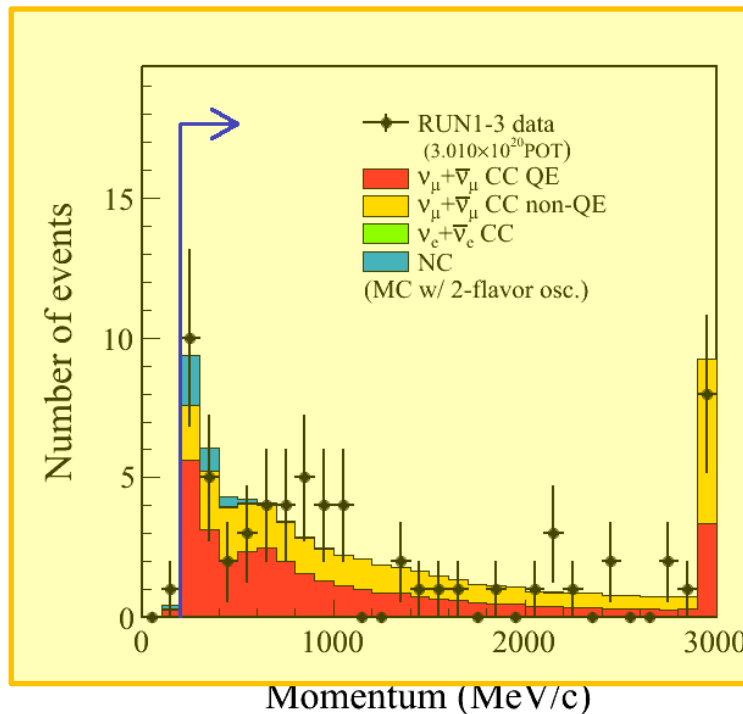
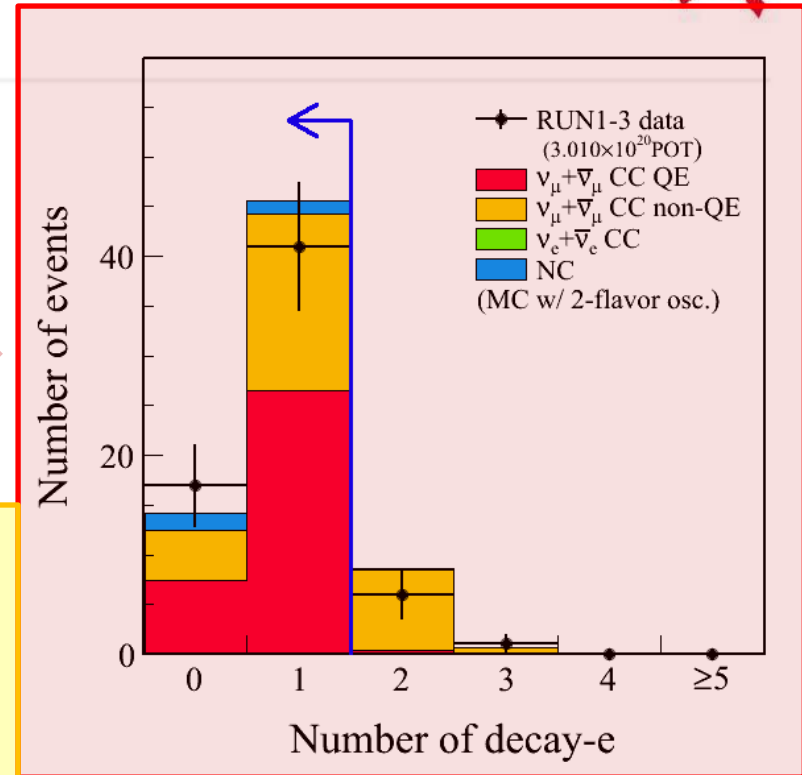
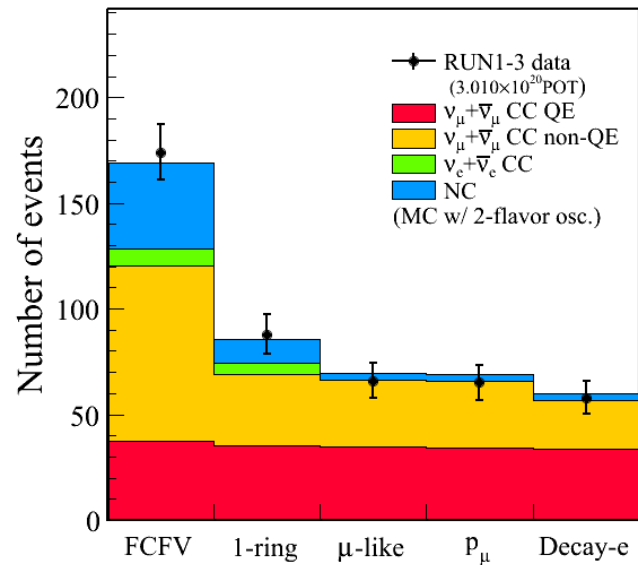


- Observe muon neutrinos from J-PARC in Super Kamiokande
- Predict number using beam MC constrained by ND280 measurements
- Fit reconstructed neutrino energy spectrum and thus
- Measure  $\sin^2 2\theta_{23}$  and  $\Delta m_{32}^2$

# $\nu_\mu$ Selection at Far Detector (SK)



# $\nu_\mu$ Event Selection contd.



> 200 MeV/c

*Electron observed later in time than  $\mu$  consistent with being from decay of  $\mu$ .*

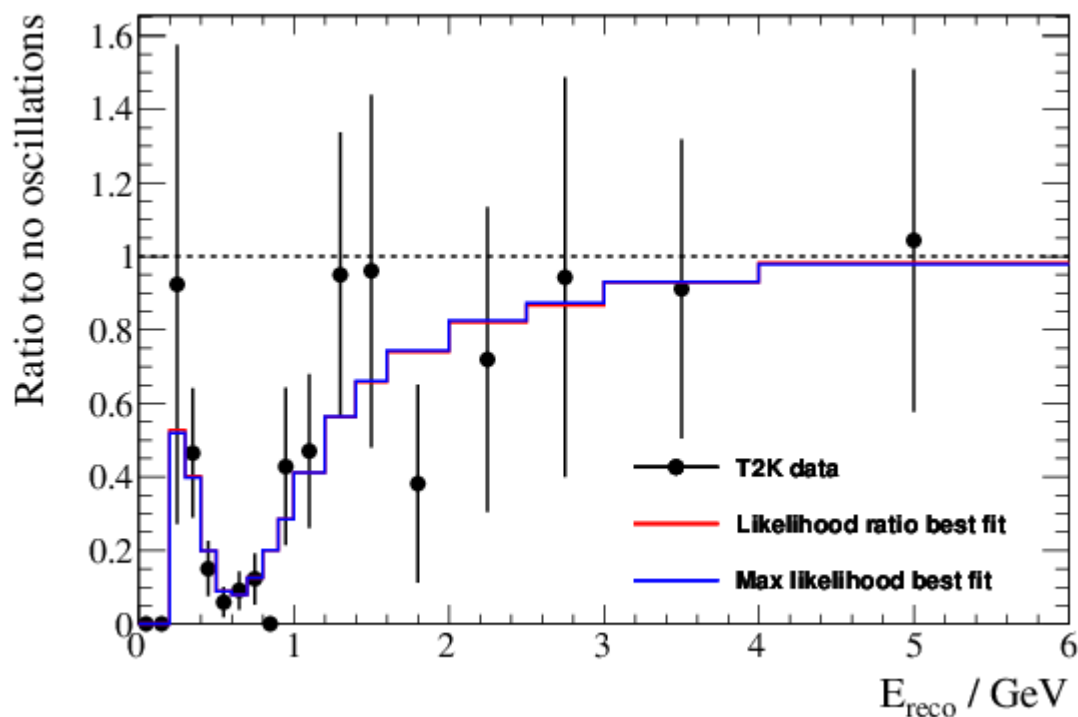
Count number of muon neutrino events in Super Kamiokande

RUN1+2+3 3.010x10 <sup>20</sup> POT	Data	MC Expectations w/ oscillation				
		MC total	$\nu_\mu + \bar{\nu}_\mu$ CCQE	$\nu_\mu + \bar{\nu}_\mu$ CC non-QE	$\nu_e + \bar{\nu}_e$ CC	NC
True FV	-	299.35	49.67	109.50	8.62	131.56
FCFV	174	168.86	37.60	82.80	8.24	40.23
One-ring	88	85.65	35.27	33.67	5.28	11.43
$\mu$ -like	66	69.67	34.58	31.61	0.04	3.43
$p_\mu > 200 \text{ MeV}/c$	65	69.25	34.34	31.54	0.04	3.33
$N_{\text{dcy-e}} \leq 1$	58	59.86	33.90	22.73	0.04	3.19
Efficiency [%]	-	20.0	68.2	20.8	0.4	2.4

95% CC – of which 64% CCQE



Plot spectrum as a function of reconstructed  $\nu$  energy.  
Fit for oscillation parameters.



90% Confidence Level allowed region:

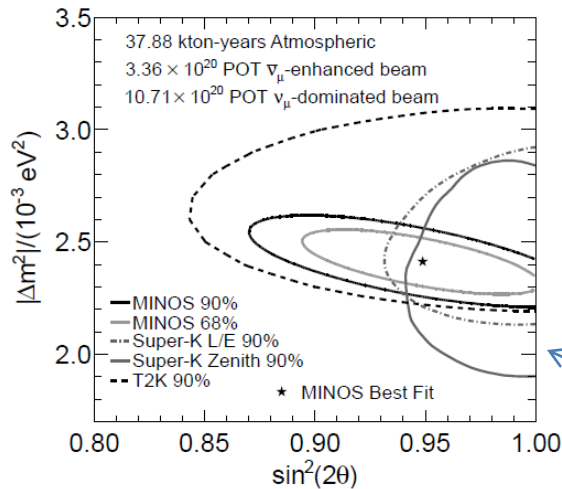
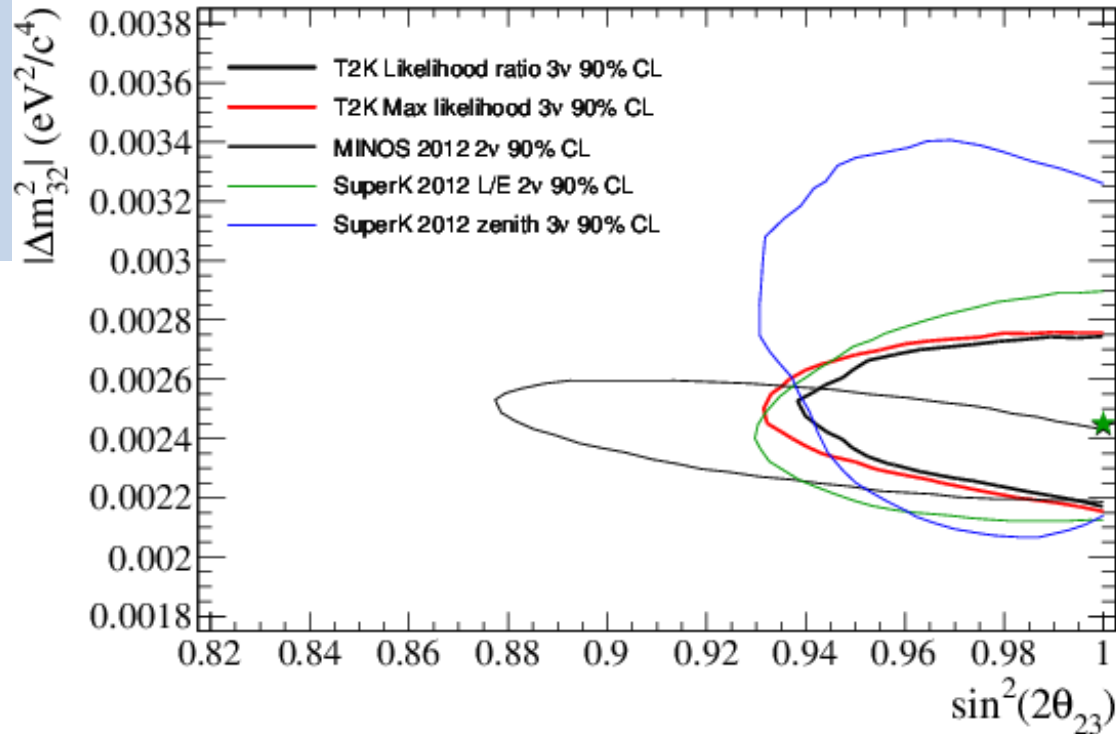
$$2.14 \times 10^{-3} \text{ eV}^2/\text{c}^4 < |\Delta m_{32}| < 2.76 \times 10^{-3} \text{ eV}^2/\text{c}^4$$

$$\sin^2 2\theta_{23} > 0.957$$

Best fit values:

$$|\Delta m_{32}| = 2.44 \times 10^{-3} \text{ eV}^2/\text{c}^4$$

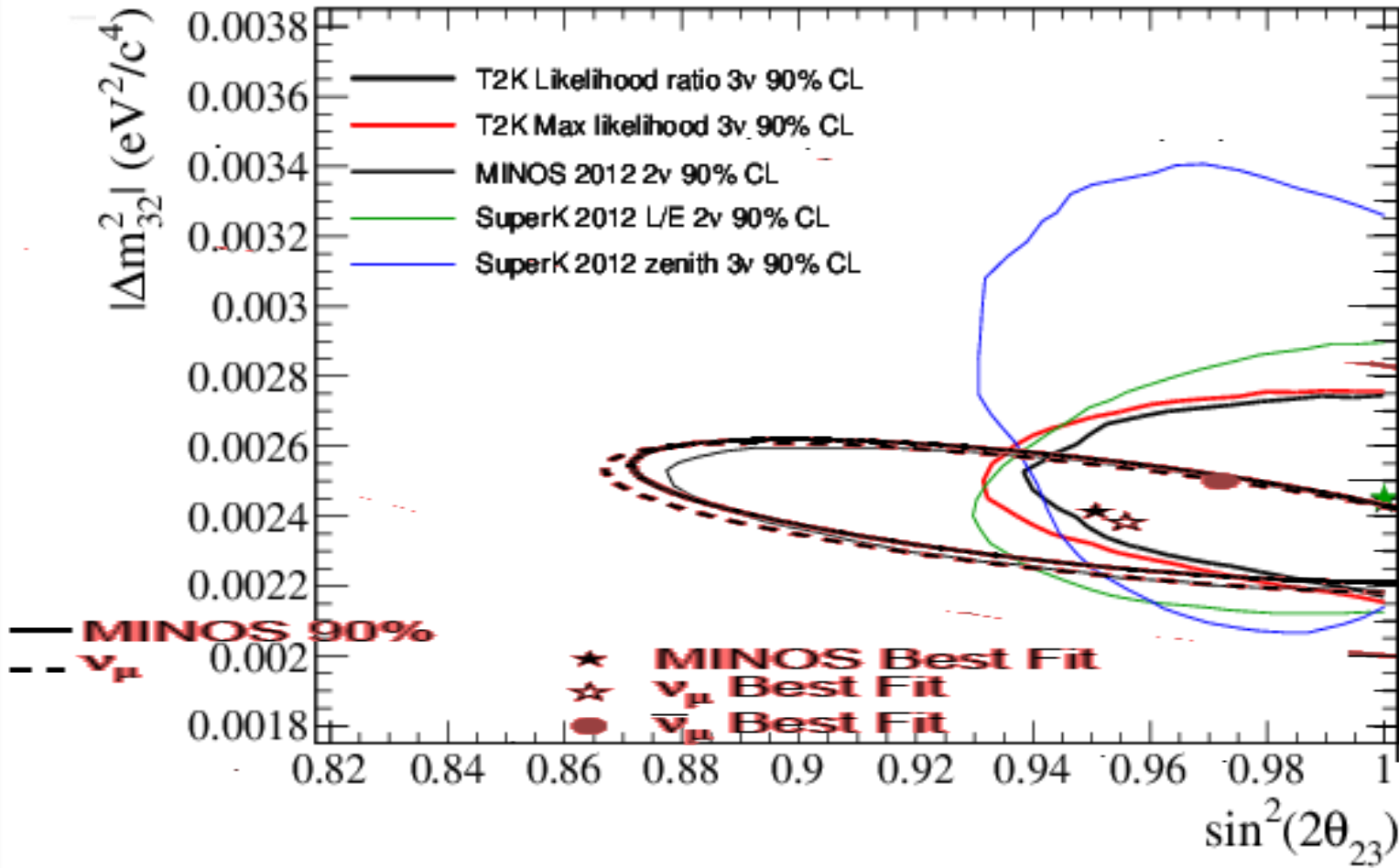
$$\sin^2 2\theta_{23} = 1.0$$



New MINOS result: combined fit of beam and atmospheric results

<http://uk.arxiv.org/abs/1304.6335>

8<sup>th</sup> May 2013

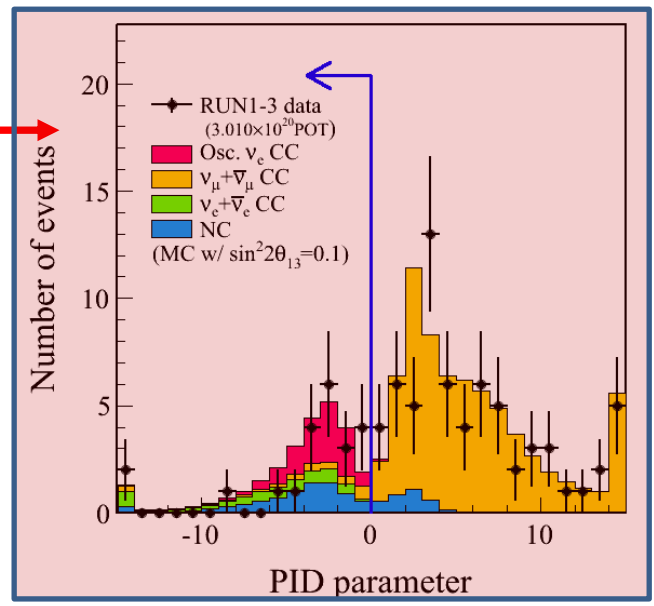
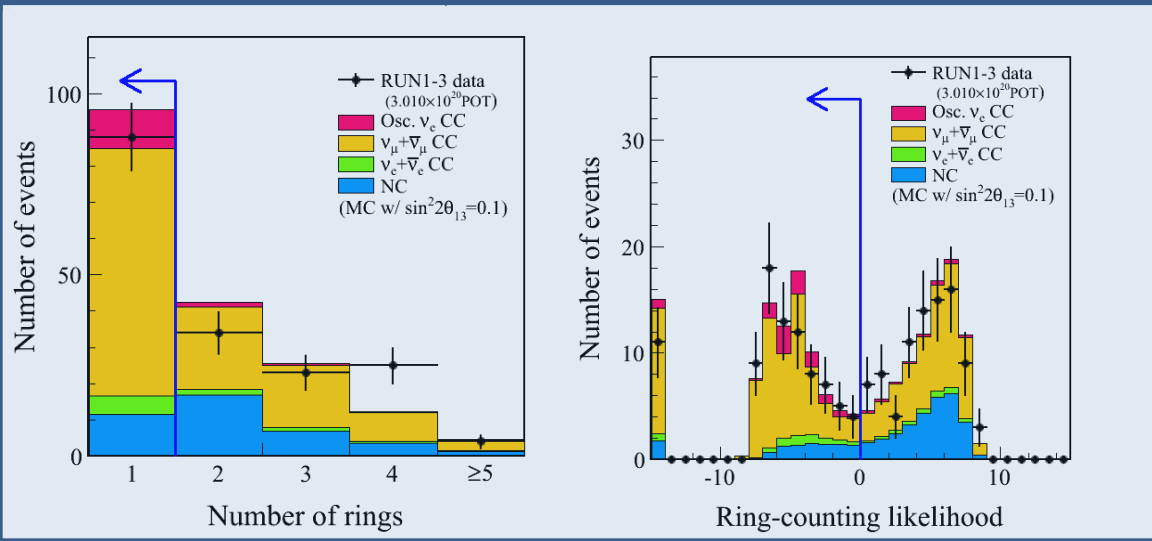
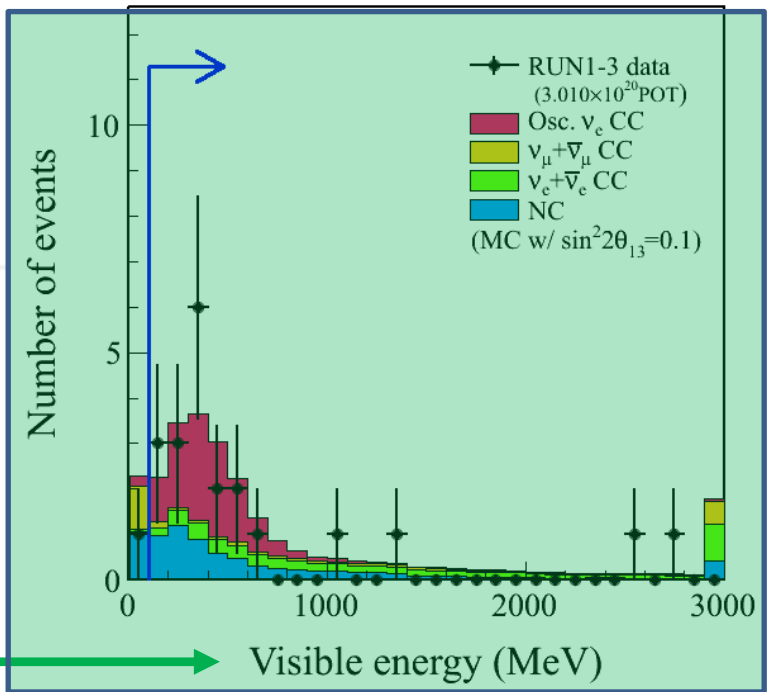
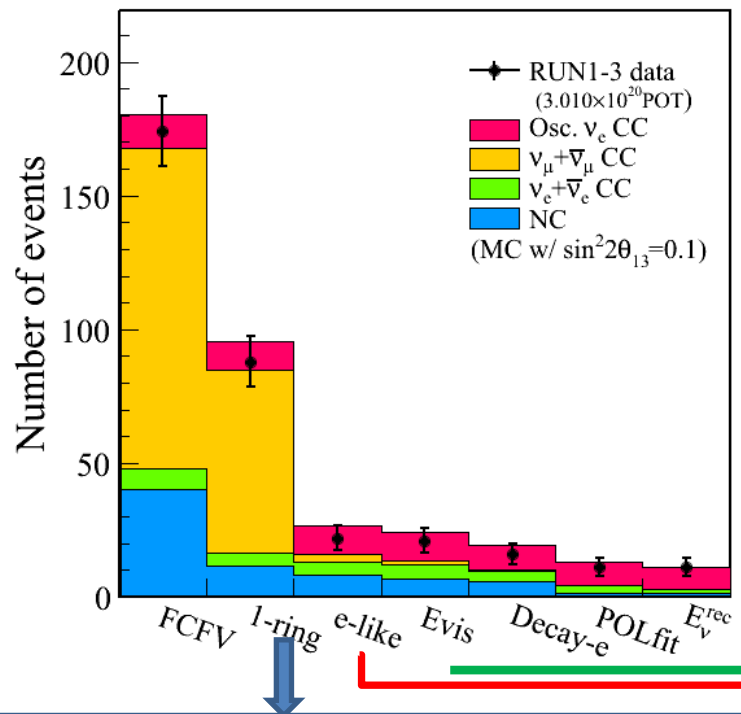


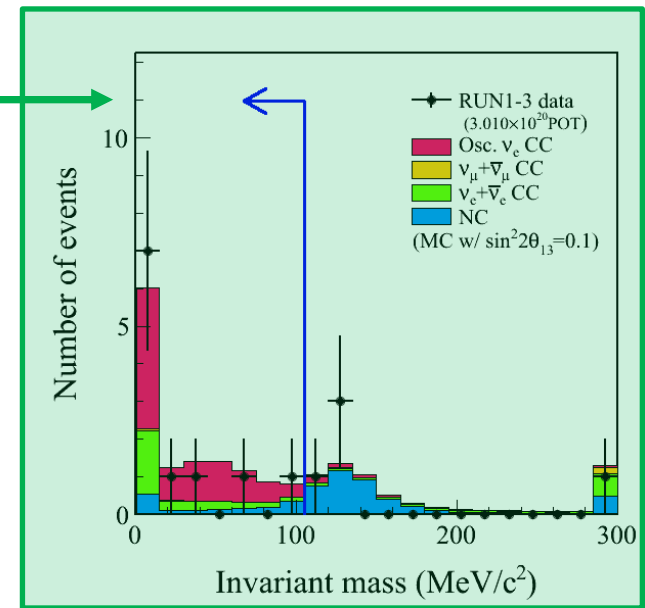
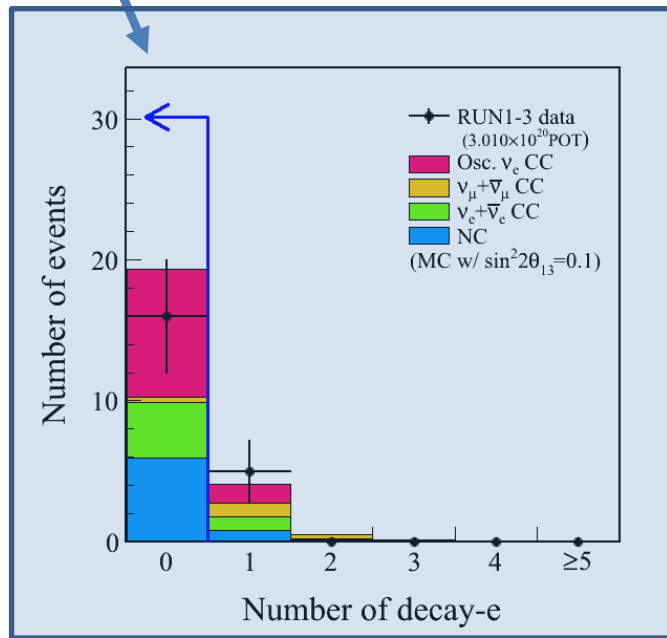
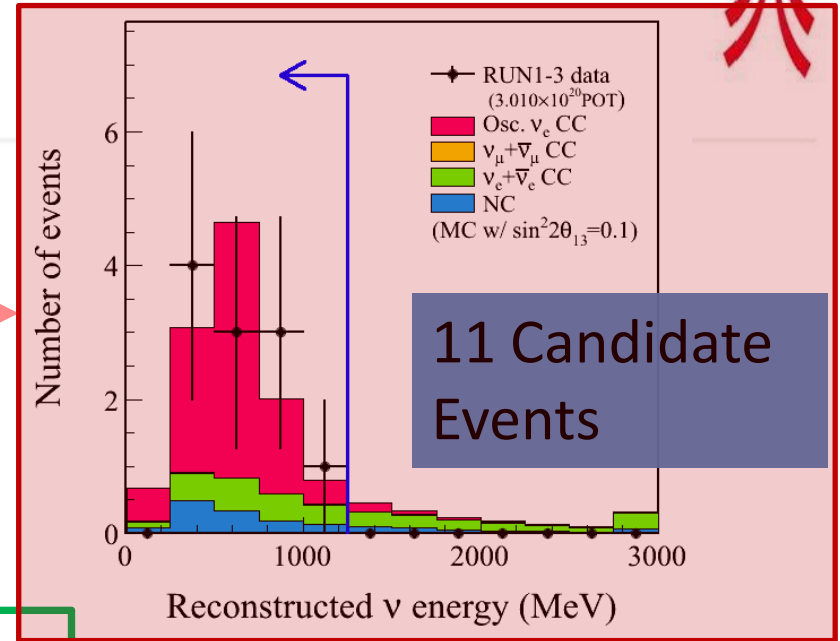
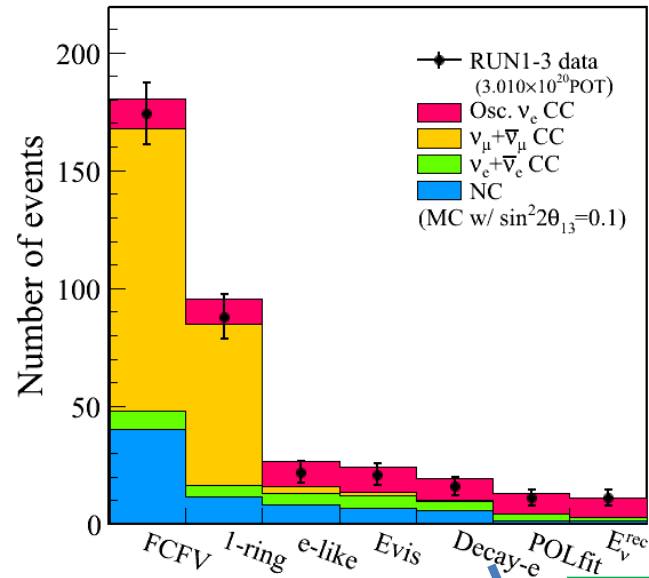
## Recent history

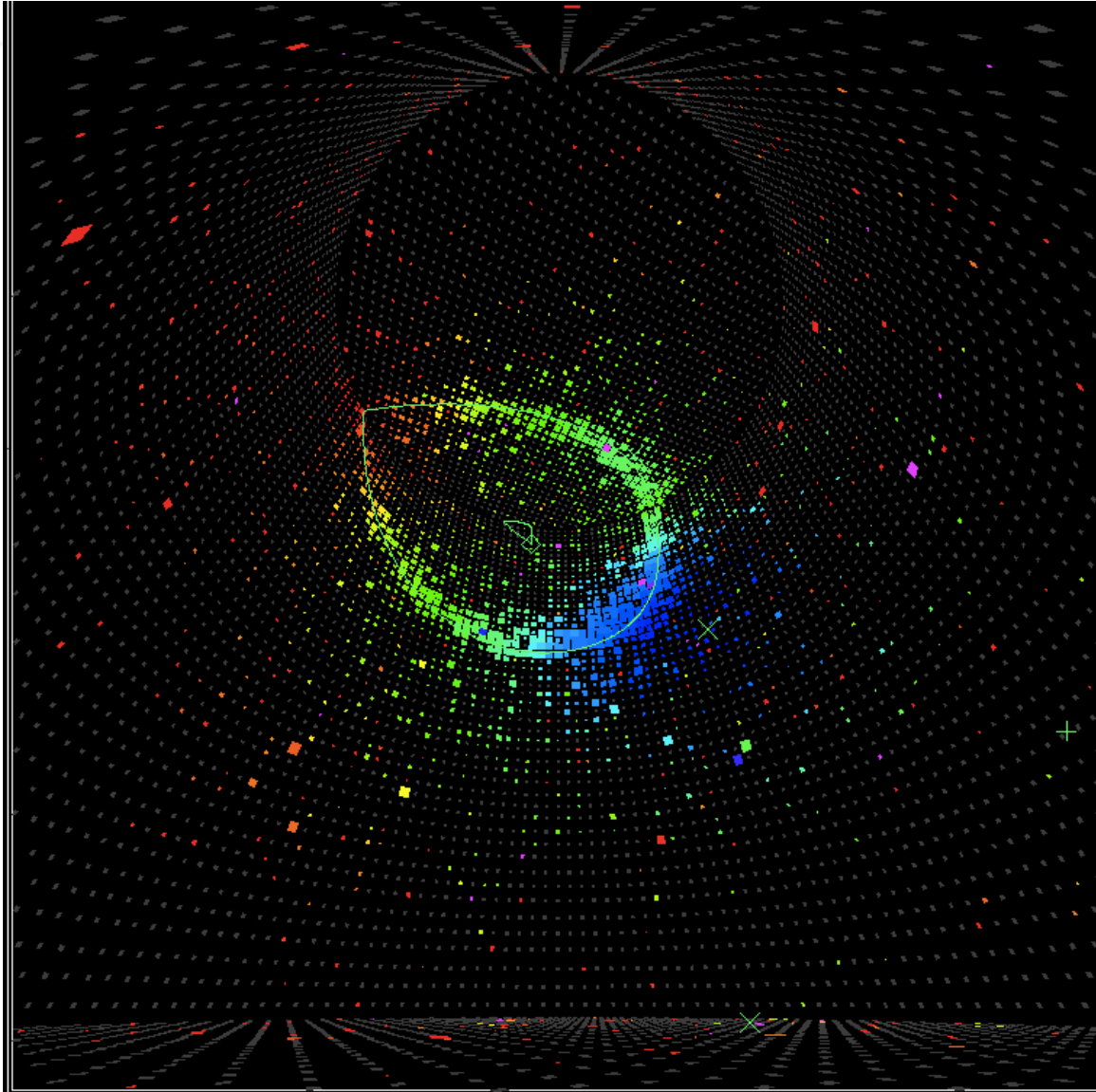
- 2011 result
  - Observed 6 events
    - (Run 1 and 2 data – before earthquake 11 March 2011)
  - $0.03 < \sin^2 2\theta_{13} < 0.28$ 
    - for  $\delta CP=0$  and normal hierarchy.
- Reactor results
  - 29 Dec 2011 Double Chooz  $0.017 < \sin^2 2\theta_{13} < 0.16$ . ( 90% CL)
  - 8 March 2012 Daya Bay announces  $5.2\sigma$  measurement of  $\theta_{13}$   
 $\sin^2 2\theta_{13} = 0.092 \pm 0.016$  (stat)  $\pm 0.005$  (syst)  
Confirmed 1 month later by RENO  
See talk by Kwong Lau
- Our new result
  - Run1 + Run2 + Run 3

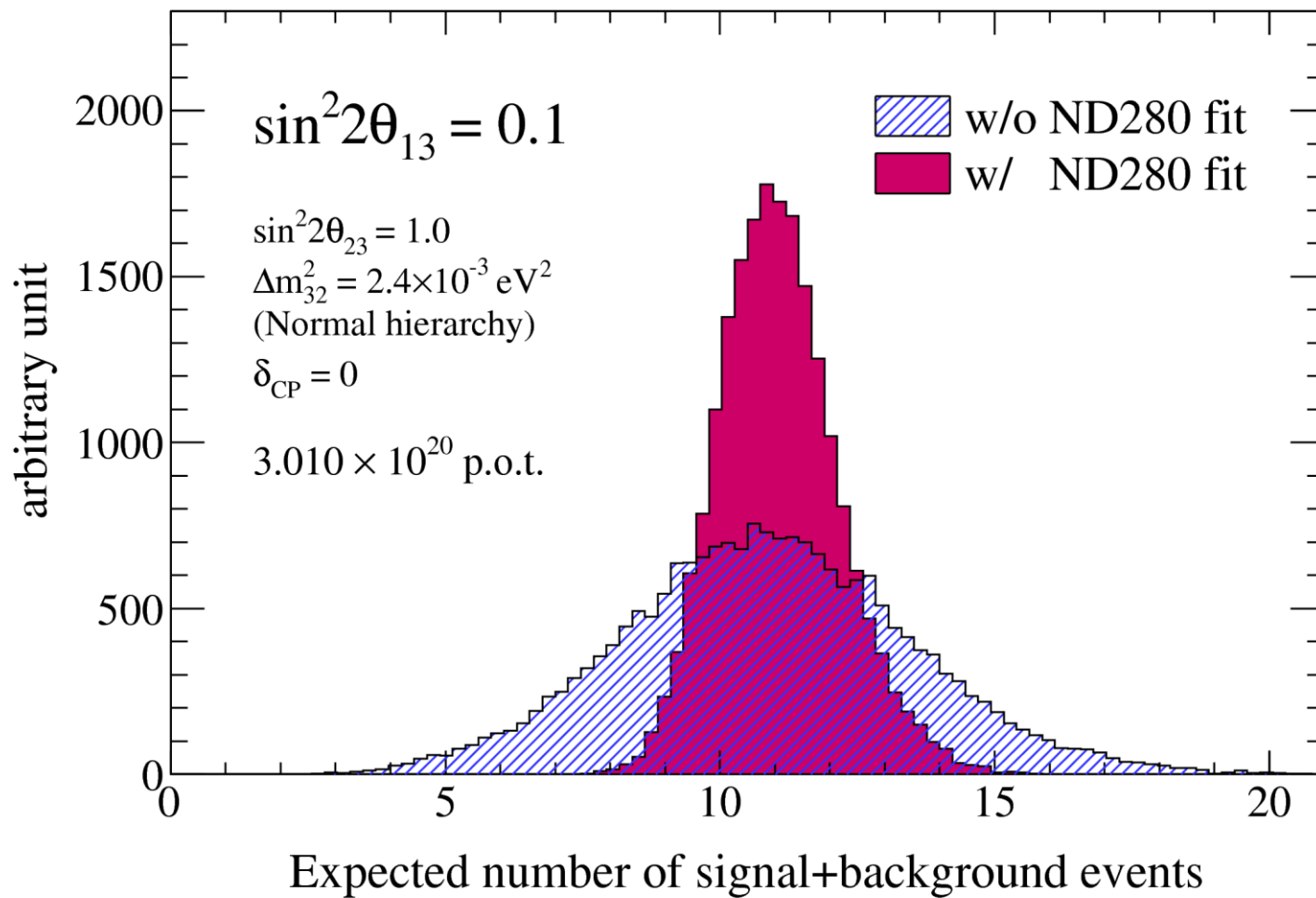


# T2K Electron Neutrino Candidate Selection (at SK)



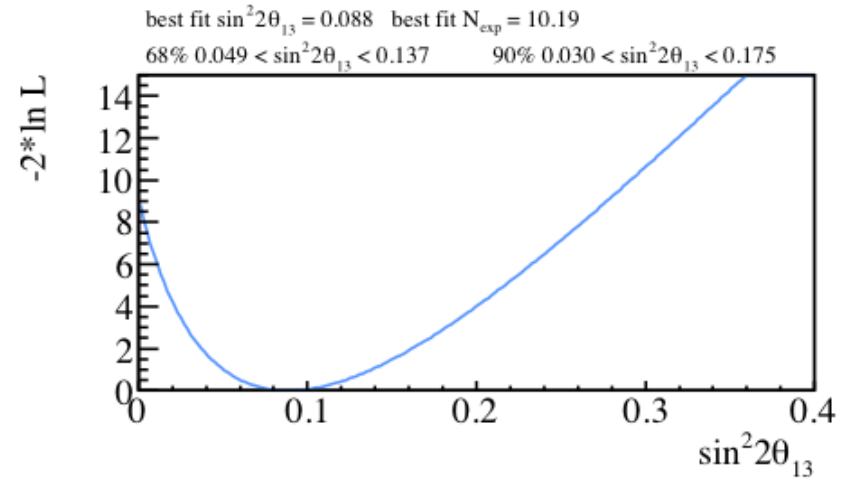
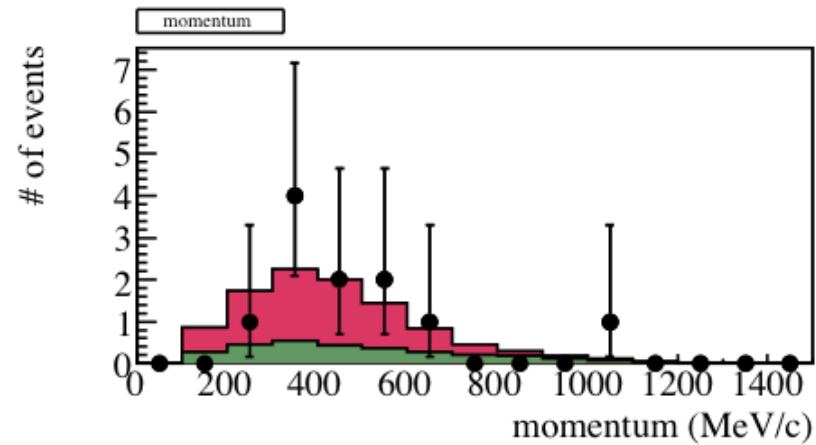
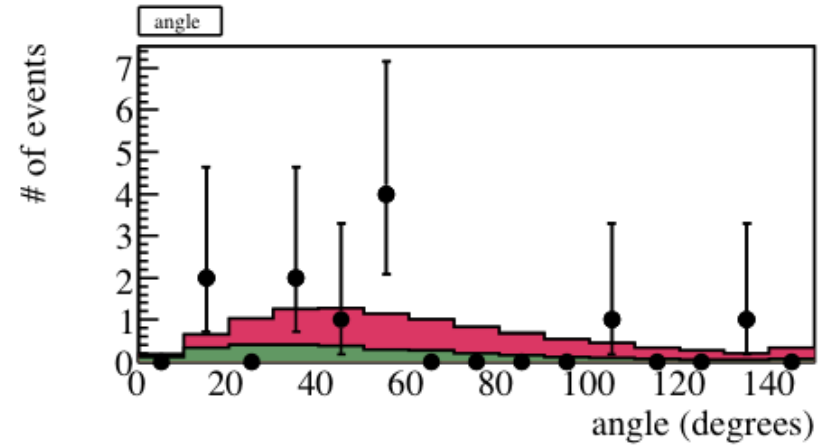
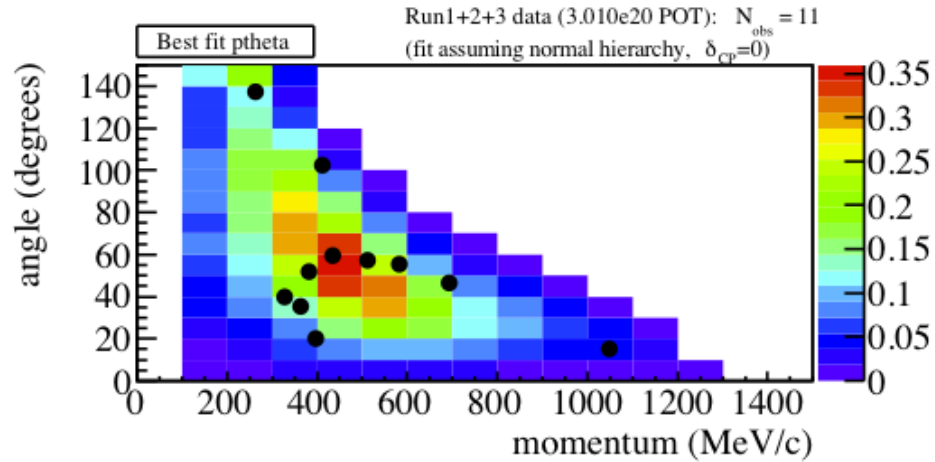




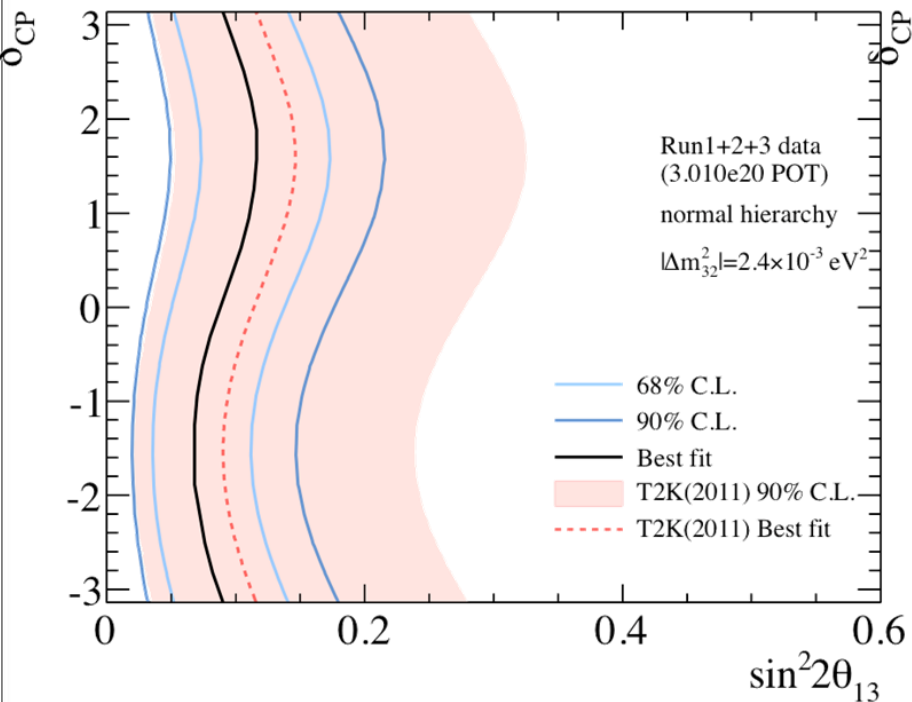




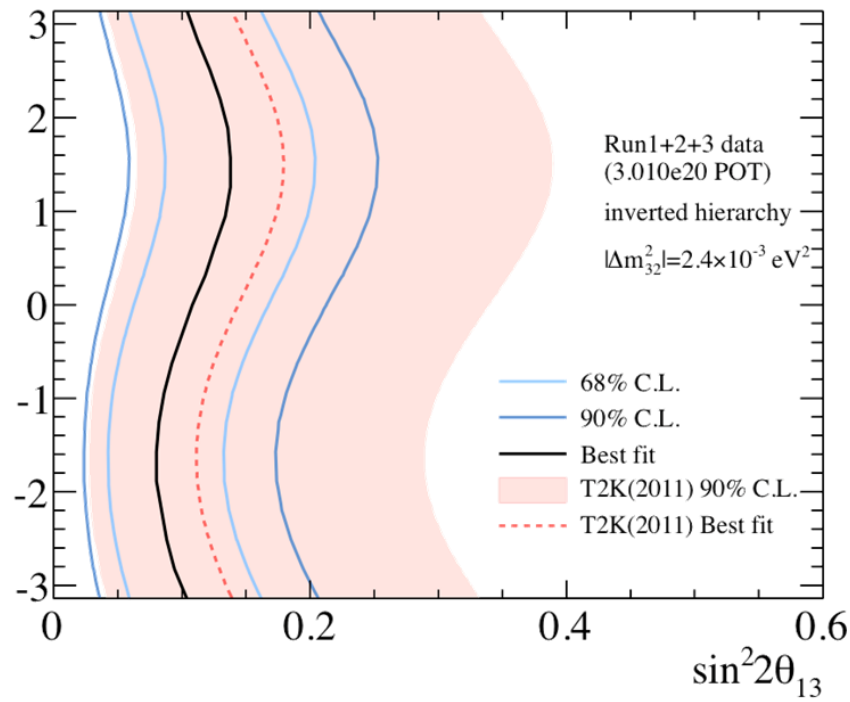
## Summary



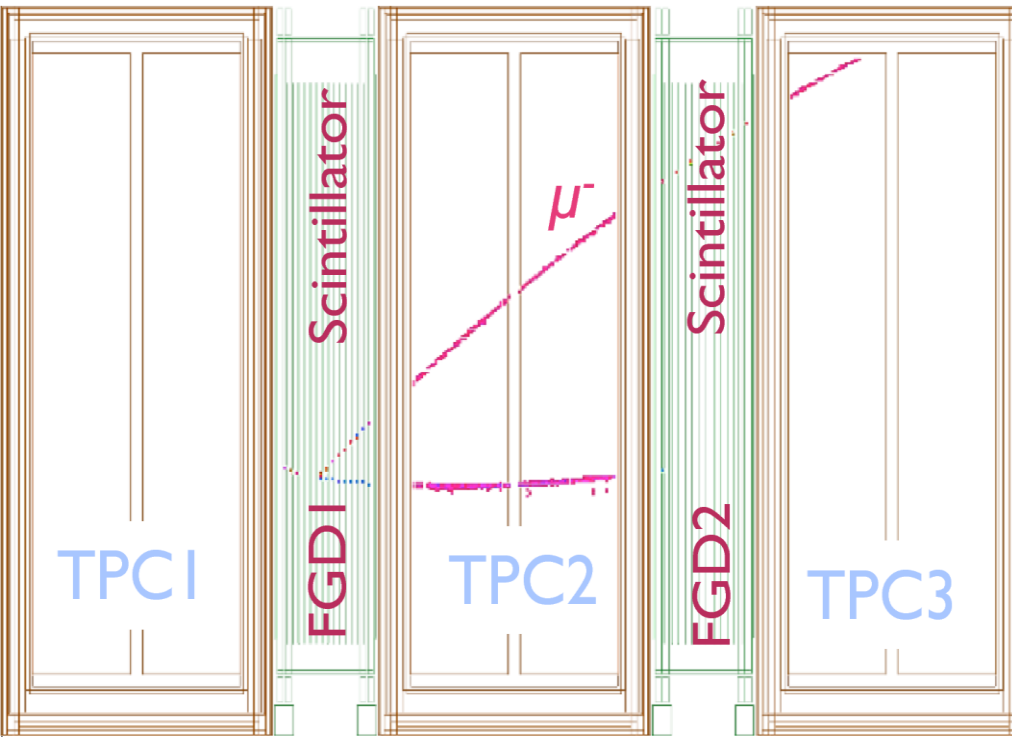
## Fit results



$$0.030 < \sin^2 2\theta_{13} < 0.175 \text{ (90\%CL)}$$



$$0.038 < \sin^2 2\theta_{13} < 0.212 \text{ (90\%CL)}$$



- ND280 Run1 and 2 data
- Detect  $\nu_\mu$  in FGD1 of ND280
  - Good Data Quality
  - $>0$  -ve track in TPC
  - Track starts in FGD1
  - $dE/dx$  of -ve track consistent with  $\mu$
  - No activity upstream of FGD
- Unfold to true  $\mu$   $p/\theta$  bins
- Convert to differential cross section

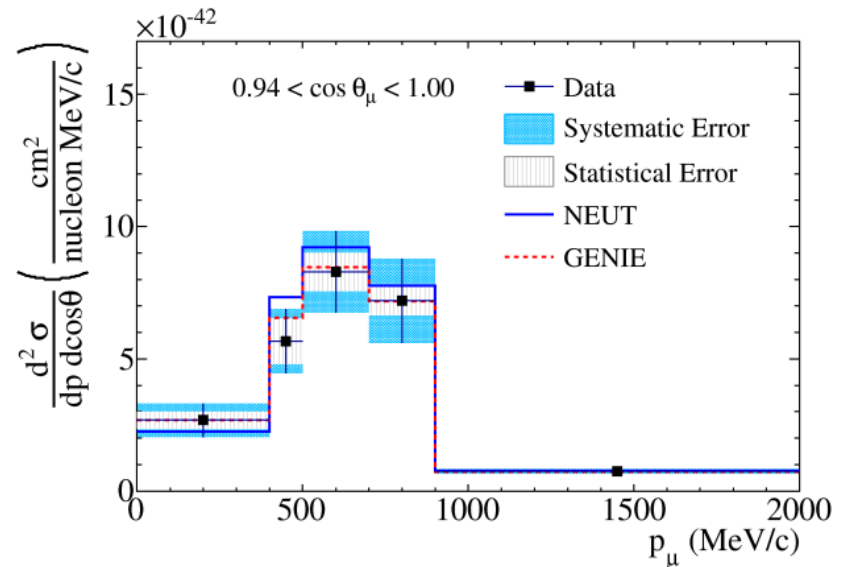
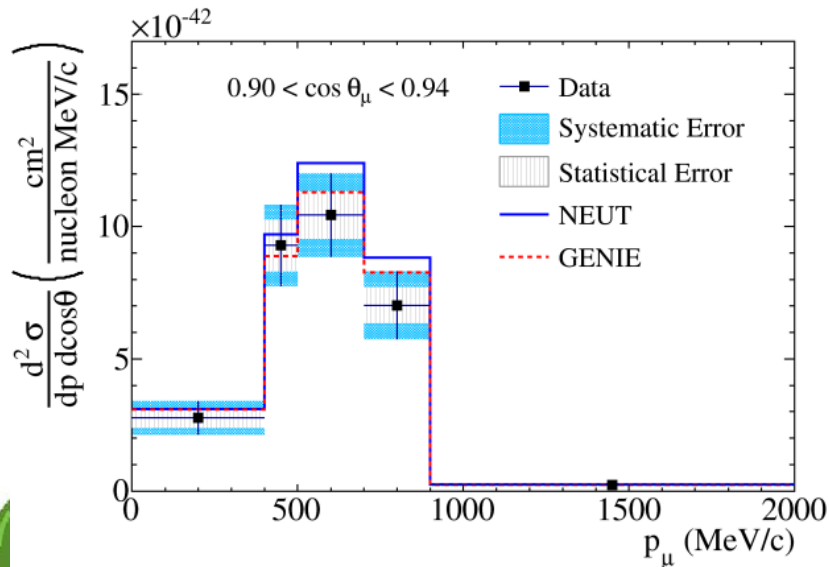
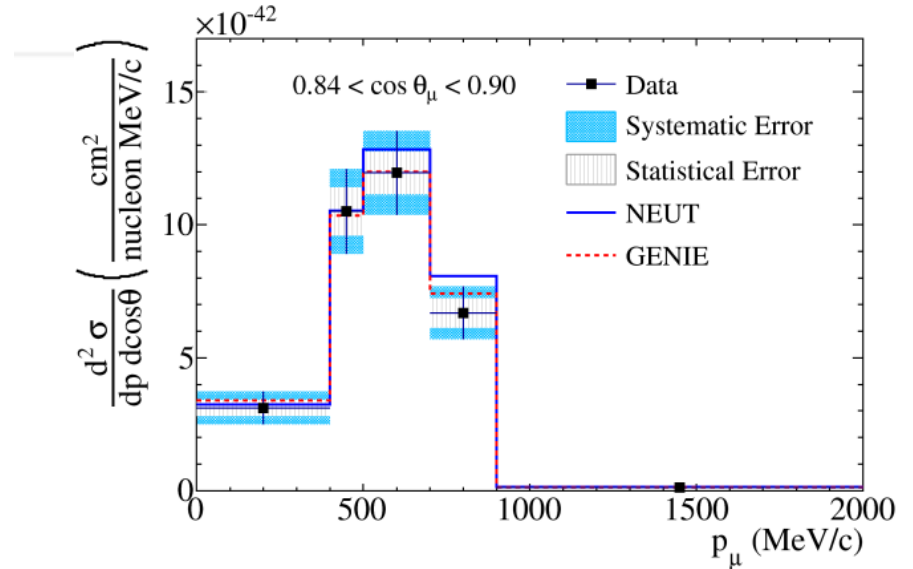
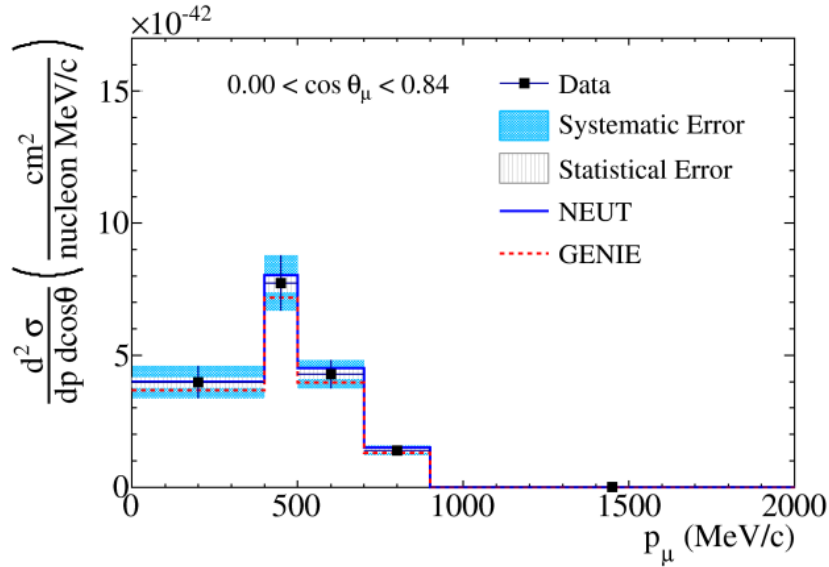
Efficiency = 50%

Purity = 88%

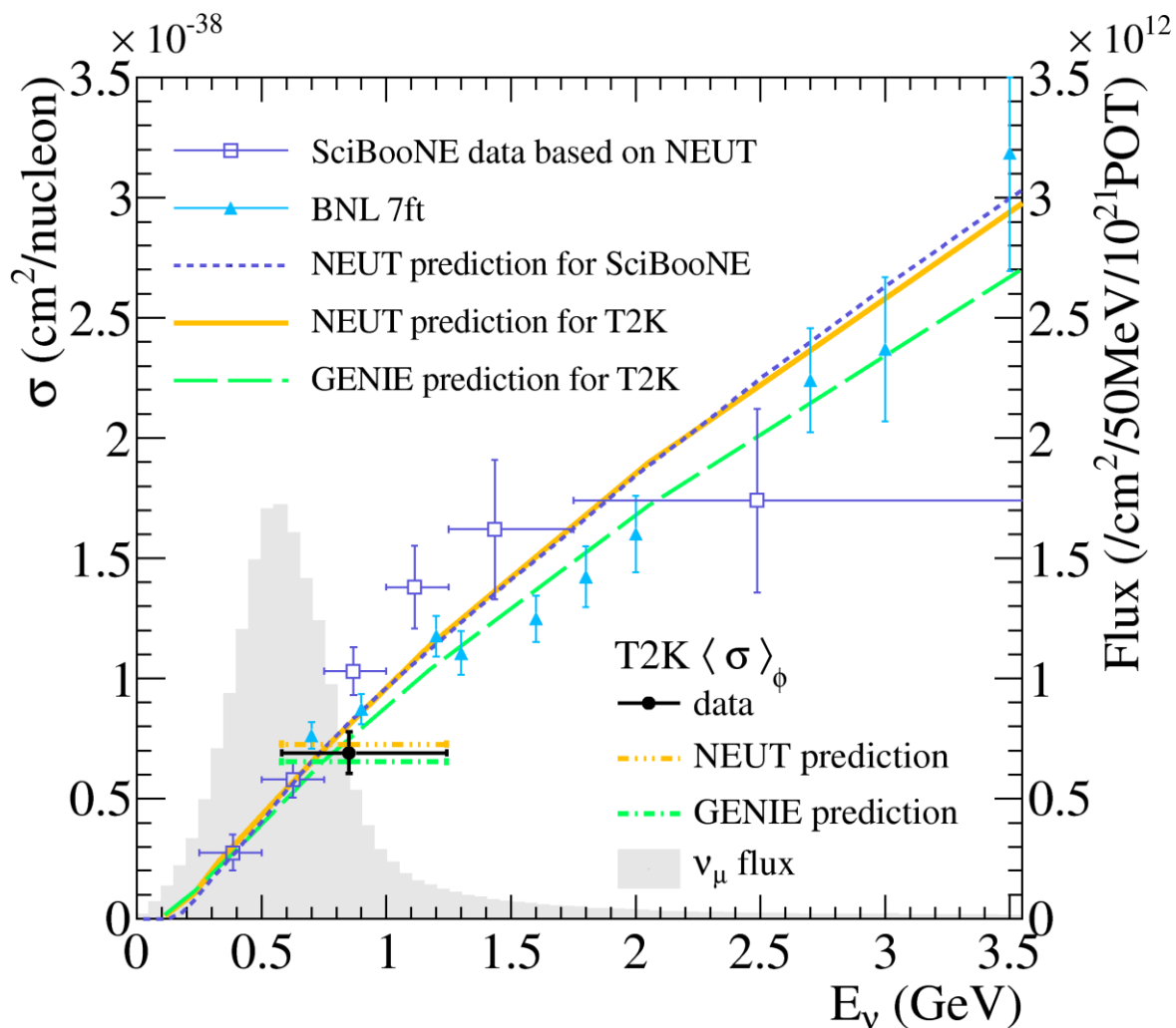
Main backgrounds:  
 Events not in FGD1  
 Neutral current



## Results (Binned)



$$\langle \sigma_{CC} \rangle_{\phi} = (6.93 \pm 0.13(\text{stat}) \pm 0.85(\text{syst}) \times 10^{-39} \text{cm}^2/\text{nucleon}$$





- More data
  - Plan is to collect sufficient data before long shutdown for a  $5\sigma$  measurement of  $\theta_{13}$
- Improved reconstruction
- More ND280 cross section measurements
- Which will all help to produce more accurate measurement.

## *Future Runs / Long Term*

- More of the above and..
- Anti neutrino running?
- Synergies with NOVA...
- Possible sensitivity to CP violation.

### NOVA:

Long baseline experiment Fermilab -> Ash River (810 km)  
=> Larger matter effects than T2K

Off Axis

E=2GeV

First beams now (Spring 2013)

Detector completed this time next year

Expect  $5\sigma$  on  $\theta_{13}$  by May 2014

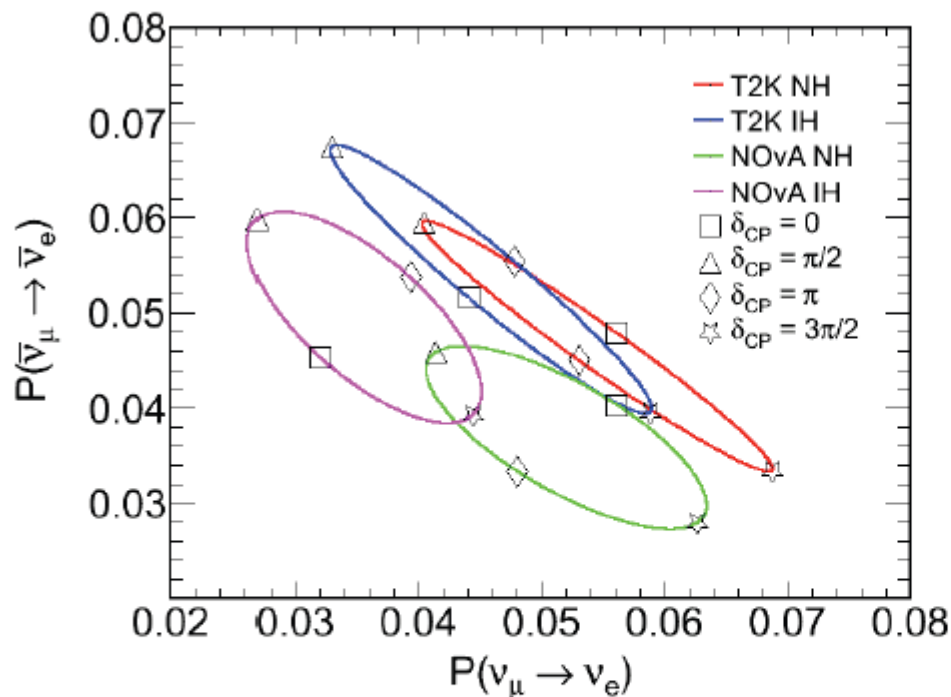
- T2K is well on the way to achieving its original Physics Goals
  - Electron neutrino appearance
  - Muon neutrino disappearance
  - Neutrino cross sections
- Lots more to come in the next few years



# BACKUP



## with matter effects



T2K might be able to measure  $\delta_{CP}$  if it is close to  $3\pi/2$  (normal hierarchy) or  $\pi/2$  (inverted hierarchy). However for most values of  $\delta_{CP}$ , its effects are entangled with those of the mass hierarchy.

NOvA might be able to measure  $\delta_{CP}$  if it is between  $\pi-2\pi$  (normal hierarchy) or  $0-\pi$  (inverted hierarchy).

$$\sin^2(2\theta_{12}) = 0.87$$

$$\sin^2(2\theta_{13}) = 0.1$$

$$\sin^2(2\theta_{23}) = 1.0$$

$$\Delta m_{12}^2 = 7.6 \times 10^{-5} \text{ eV}^2$$

$$|\Delta m_{32}^2| = 2.32 \times 10^{-3} \text{ eV}^2$$

$$\text{Earth crust density} = 2.6 \text{ g/cm}^3$$

NH = normal mass hierarchy

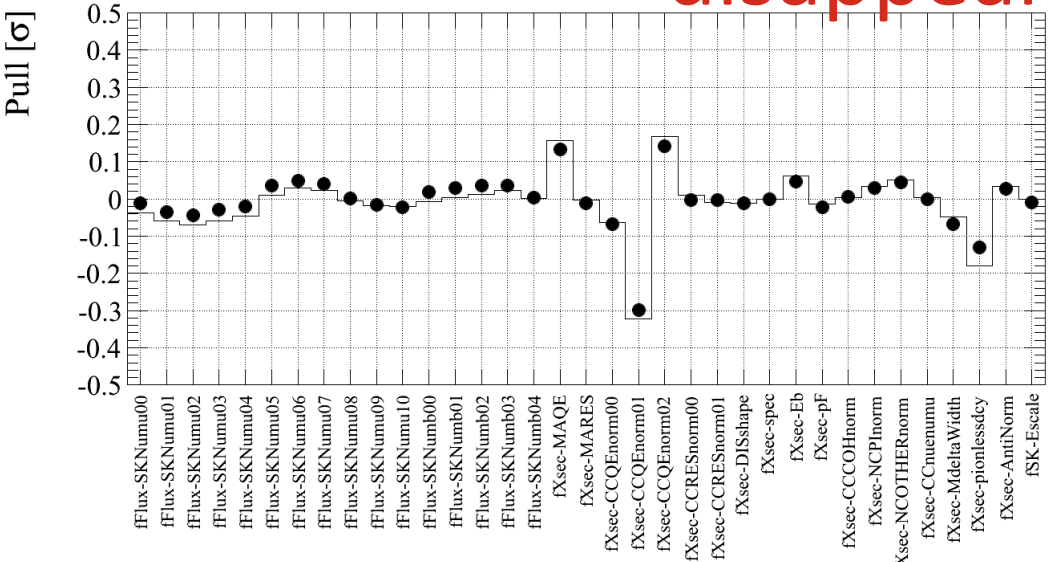
IH = inverted mass hierarchy

T2K energy = 0.6 GeV

NOvA energy = 2.0 GeV

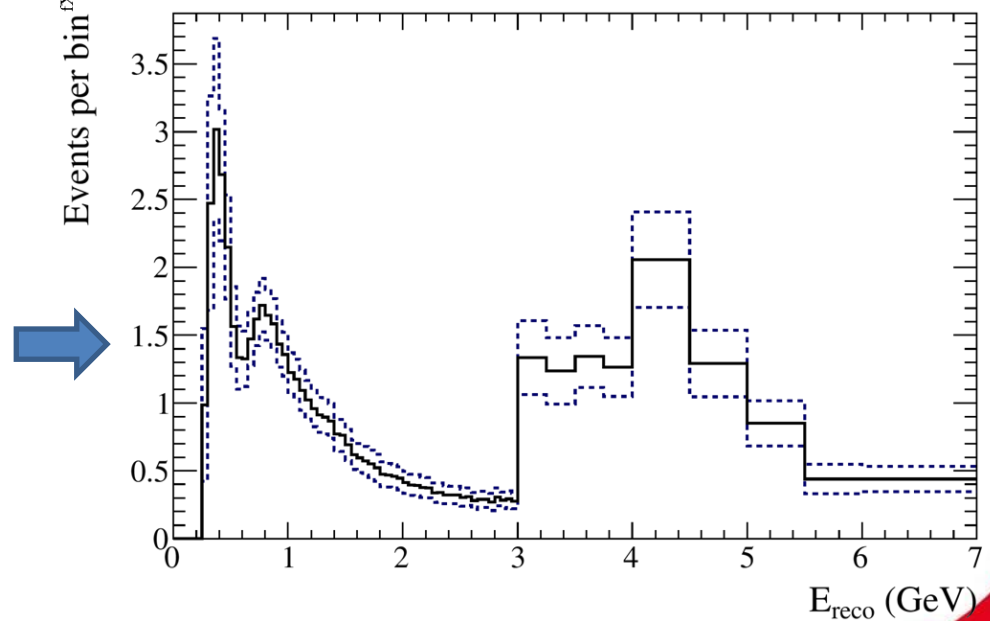
Note that, for  $\delta_{CP} = 0$ ,  $P(\nu_\mu \rightarrow \nu_e) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ . This is due to matter effects.

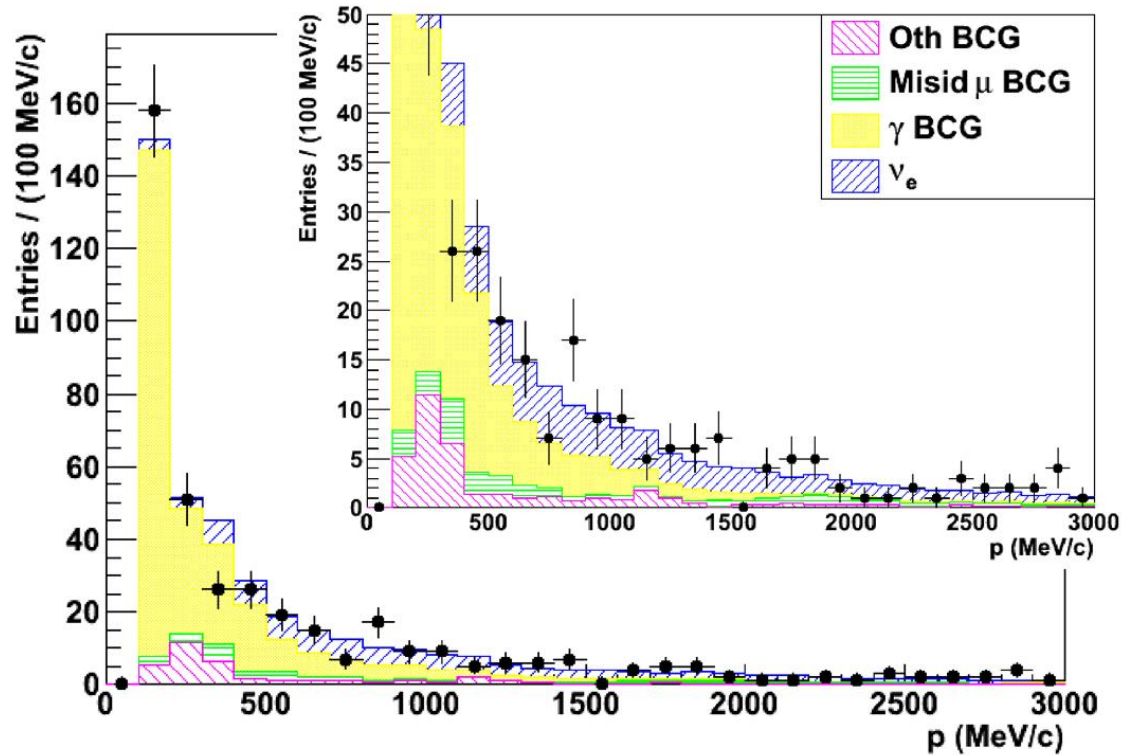




Pulls of systematic parameters at the best fit point using two different fitting methods (histogram/points)

Total error envelope calculated as  $\pm 1\sigma$  spread of bin contents using 100k toy MC experiments generated with randomized systematic parameters. All systematic parameters and their correlations were taken into account.





Unfold to find true bins of muon momentum and angle

Differential cross section definition:

$$\left\langle \frac{\partial^2 \sigma}{\partial p_\mu \partial \cos \theta_\mu} \right\rangle_{kl} = \frac{\overset{\text{\# of interactions in true bin}}{N_{kl}^{\text{int}}}}{\underset{\text{\# of target flux nucleons}}{T \phi \Delta p_{\mu,k} \Delta \cos \theta_{\mu,l}}}$$

2D binning: (k,l)

Method

**Unfolding**

$$N_k^{\text{int}} \approx \hat{N}_k = \frac{\overset{\text{un-smearing matrix}}{U_{kj}}}{\underset{\text{efficiency}}{\epsilon_k}} \left( \underset{\text{\# of sel. events}}{N_j^{\text{sel}}} - \underset{\text{background in rec. bin}}{B_j} \right)$$

ID binning: k

unfolding based on Bayes' theorem

$$U_{kj} = P(k|j) = \frac{P(j|k)P(k)}{\sum_{\alpha} P(j|\alpha)}$$

$U_{kj}$  = probability to have an interaction in bin k, when having reconstructed the event in bin j

recc  
↓  
true