

## **Motivation (TLEP Design study Draft)**

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**C**Running TLEP as a Giga- (Tera-)Z Factory (45 GeV/beam):

Polarization gives access to certain physics (e.g. SLC)

Want > 50%, maybe at lower luminosity

Want pol. e+ as well

**G**Running TLEP @ WW threshold (80...90 GeV/beam)

want energy calibration (*P* ≥ 5%)

**G**Running TLEP at the Higgs (120 GeV/beam):

Polarization not required.

*<u>OEnergy calibration?</u>* 

Running TLEP at tt (175 GeV/beam) <u>-</u><br>+

*<u>OEnergy calibration?</u>* 

## **Introduction: LEP Observations & Data**

**OLEP has had the highest-energy polarized electron beams** Energy spread reduces polarization at highest energy



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### **Polarization time constants**

**OSokolov-Ternov polarization:** 

$$
\tau_{\rm st}^{-1} = \frac{5\sqrt{3}}{8} \frac{r_{\rm e}\gamma^5 \hbar}{m_{\rm e}|\rho|^3} \ = \frac{2\pi}{99} \frac{E^5}{C\rho^2}
$$



## **A simple model to describe the energy limit**

Spin resonances every 0.441 GeV (for *e–*)

**Otails in a beam may extend beyond these** 

q.e. causes instantaneous jumps beyond the resonant energy

radiative damping causes crossing on the way back => some depolarization.

 $\bullet$  spin tune modulated by Qs => reduces space for spin tune.



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### **"Phenomenological Description"**

**O**lf the resonance-crossing causes depolarization *D*, we can make the *ansatz* for the polarization



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**O**To finish this, we need to know D

Crossing a depolarizing resonance => we can estimate D using the *Froissart-Stora* formula:

$$
1 - D = 2 e^{-\frac{1}{2} \frac{\pi w k^2}{\alpha}} - 1
$$

For our cases,  $wk \approx 0.001$  and  $\alpha \le 0.01$  so D turns out to be 1..4 x 10<sup>-4</sup> per crossing.

**This effect competes with Sokolov-Ternov to reduce the eq.** 

polarization:

\n
$$
Pol := \frac{P0}{1 + \frac{tau\_ST e^{-\frac{H}{H\_ave}} D}{tau} }
$$

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# **P**Putting it all together:



 $\Theta(dE = (0.5 - Q_s(E))^* 0.441$  spin-tune-space expressed in GeV)

 $\triangle Q_s$  comes in as the spin tune of each particle is modulated by  $Q_s$ 

**G** For a realistic estimate of  $Q_s(E)$  we need to have an estimate of  $V_r(E)$ .

We do this such that the momentum acceptance stays reasonable

 $\Theta$ =>  $V_{rf}$   $\propto$   $E^2$ , roughly hitting the nominal values per the parameter table.

#### **LEP2**

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## $\mathbf{V}_{\text{rf}} = 1/3 * E^2$  (*E* in GeV,  $V_{\text{rf}}$  in GV [3.64 GV @ 104.5 GeV])





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## **Comparison of model to LEP2 data**

Assmann (1999) estimates *wk* ≈ 0.0014

**Othis model would favour**  $wk = 0.002$ 



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#### **TLEP Polarization estimate**





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## **Optimizing TLEP**

**On this model, we can gain by** 

reducing the energy spread using *Js* –> 2  $\Theta$ being optimistic reducing  $w_k \rightarrow 0.001$ 

✦but note: LEP2 already used quite elaborate algorithm for spin matching **e**reducing the synchrotron tune using  $\alpha_p \rightarrow 5 \times 10^{-6}$ 



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## **Remarks to the proposed model**

**The inspiration to this model came from a paper by Derbenev,** Kondratenko and Skrinsky.

Their condition for correlated multiple resonance crossings is *violated*.

- On the other hand; the resonance-crossing model used here has also issues due to small # of synchrotron oscillation periods
- **O**The interesting difference between these descriptions:
	- D-K-S: a higher *Qs helps* polarization (for correlated crossings) this model: a higher *Qs hurts* polarization
- Also, at very high energy, D-K-S allow for an increase in polarization

high polarization rate trumps depolarization rate

 $\Theta$ this model allows for that as well

**O** Neither has higher-order or spin-betatron resonances.

significant e.g. in LHeC tracking

#### **Linear vs nonlinear Spin Tracking (SLICKTRACK, LHeC,** *Qs***=0.15) SLAC**



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#### **TLEP Polarization time**





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## **TLEP Polarizing Wiggler @ 45 GeV**

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Note:  $100 s \Rightarrow \sqrt{0} \approx 3 GeV$ ≈ 3E34 lumi for 50 MW sr power, most of power goes into wiggler(s)  $\frac{1}{26}$ 



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### **Where does this leave us?**



Sokolov-Ternov polarization in TLEP appears to be difficult to achieve

- $\odot$   $\odot$  Z energy: good polarization, but excessive build-up time (150 h)
- ✦Alain proposes wigglers power density [MW/small divergence] manageable??
- $\odot$  g0 GeV: enough polarization for an energy calibration may be possible. ✦here wigglers could help
- $\odot$   $\odot$  H energy: even under optimistic assumptions not much left
	- ✦enough for an energy calibration??? 1 hour build-up time would be enough for this.
- @ tt energy: very fast build-up (10 min), maybe some polarization ں<br>-F
	- ✦enough for energy calibration?.
	- $\triangle$ it would be very interesting as a test case for the theory!

**O** Can snakes come to the rescue??

### **Siberian Snakes (180° Spin Rotators)**

A pair of snakes with longitudinal/radial axis can suppress depolarizing resonances & stabilize  $\frac{on}{\lambda s}$  (up to a point). ∂*n*ˆ ∂δ

n

Spin direction opposite in the two halves of the ring => no radiative pol. One of the snakes can double-up as IR spin rotators.

Derbenev-Grote proposal

for  $I$  FP

 $\blacksquare$ For TLEP:

- ✦inject polarized
- ✦no "compacted" bends (but wigglers could work)
- ✦Similar arrangement of snakes in the injector ring
- ✦18.5…37 Tm of dipole required for each snake

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## **A conceivable scenario for TLEP ?**

Use Siberian Snakes to accelerate a polarized beam to 45 GeV

#### would need pairs of longitudinal & radial snakes, depending on resonance strength

 $\Theta$ snake pair in the collider to maintain polarization (for ≥ 1/2 h)

- $\triangle$ Allows polarized physics running  $\omega$  the Z.
- ✦Snakes will prevent energy calibration with polarization

At 90 GeV, *τpol* ≈ 5 h and *P* ≈ 0.2…0.4 (no snakes)

**Cenough to get an energy-calibration point.** 

At 120…175 GeV, *τpol* ≈ 1 h…10 min

might get enough polarization for energy calibration

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## **Summary**



# **C**Achieving polarization in TLEP will be challenging

**Caught between Scylla (long polarization time)** and Charybdis (depolarization due to unavoidable energy spread) Polarizing wiggler(s) present power-handling challenge **Siberian Snakes may come to the rescue** 

would need to accelerate polarized beam & maintain polarization  $\blacktriangleright$  vertical bends => potential to blow up vertical emittance needs a polarization-capable injector chain.

 $\Theta$ no polarized posi's  $\odot$ 

**A hybrid scenario may be conceivable** 

 $\Theta$ Snakes for physics running  $\omega$  Z energy

**■** self polarization with snakes off  $@ \approx 80...90$  GeV for energy calibration

**O** The theoretical situation is not particularly clear & may hold surprises.