

Initial (FLUKA) calculations for synchrotron radiation at TLep

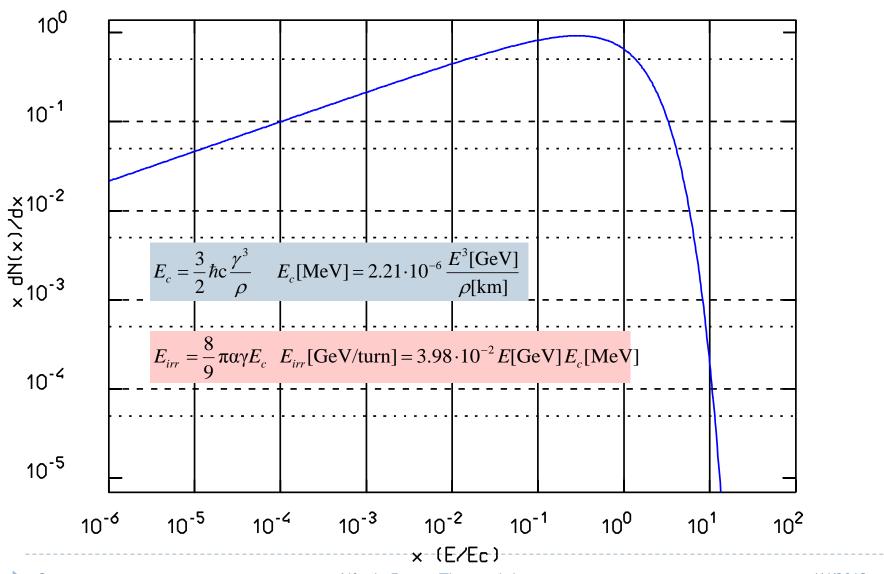
April 4th, 2013

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Outline

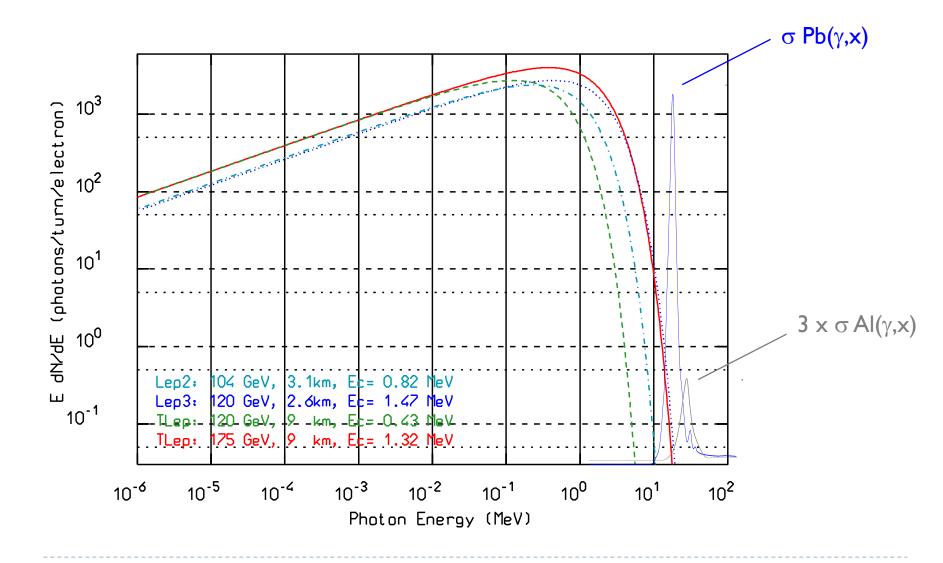
- 1. A few reminders about Synchrotron Radiation
- 2. FLUKA implementation
- 3. Some geometrical considerations
- 4. Photon attenuation and spectra for TLep
 - Results
 - 2. Physical interpretation
- Photoneutrons
- 6. Conclusions

SR: generalities





SR: generalities





FLUKA

implementation of SR

- Sophisticated low energy photon transport including polarization effects for Compton (see next slide), photoelectric and coherent scattering, and full account for bound electron effects: already available in FLUKA since several years
- New: dedicated "generic" source for SR radiation accounting for:
 - √ Spectrum sampling
 - √ Polarization as a function of emitted photon energy
 - ✓ Angular distribution
 - ✓ Arbitrary orientation emitting particle vs magnetic field
 - ✓ Photon emission along arcs/helical paths



Compton scattering: dynamics

Klein-Nishina cross section (see for example Heitler, "The Quantum Theory of Radiation"):

$$\frac{d\sigma_{KN}}{d\Omega} = \frac{1}{4} r_e^2 \frac{k'^2}{k^2} \left[\frac{k}{k'} + \frac{k'}{k} - 2 + 4\cos^2 \Theta \right]$$

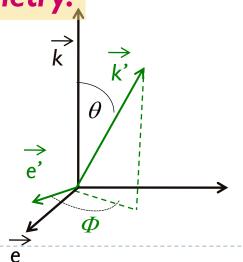
Let \vec{e} be the polarization vector of the incident photon, and $\vec{e'}$ that of the scattered one: $\cos\Theta = \vec{e} \cdot \vec{e'}$

Split σ into the two components, \bot and \bot to e respectively (actually with $e'\bot$ to the plane (e,k'), or contained in the plane (e,k')): breaking the azimuthal symmetry!

$$\frac{d\sigma_{\perp}}{d\Omega} = \frac{1}{4} r_e^2 \frac{k'^2}{k^2} \left[\frac{k}{k'} + \frac{k'}{k} - 2 \right]$$

$$\frac{d\sigma_{||}}{d\Omega} = \frac{1}{4} r_e^2 \frac{k'^2}{k^2} \left[\frac{k}{k'} + \frac{k'}{k} + 2 - 4\sin^2\theta \cos^2\Phi \right]$$

$$\cos^2 \Theta = 1 - \sin^2 \theta \cos^2 \Phi$$



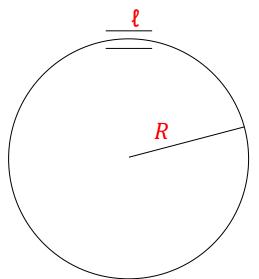
Tlep: parameters for the calculations

- E = 175 GeV, R = 9000 m
- \triangleright E_{crit}= 1.32 MeV, \triangle E=9.2 GeV/turn, dE/ds=1.63 keV/cm
- ▶ P = 9.2 I[mA] MW, dP/ds=1.6 I[mA] W/cm
- Simplified geometry, cylindrical Al beam pipe and (Pb) shielding
- ▶ SR photons generated and tracked above 100 eV (99.999% of the total power), average energy of the photons <E>=430 keV (E>100 eV)

A decrease of \sim 15% in R (eg Holzer talk) \rightarrow a corresponding increase in E_c , power, and likely a factor a few in photoneutron production



Synchrotron Radiation Interception



R accelerator bending radius

ℓ dipole length

r vacuum chamber radius

SR hitting inside the same dipole only if $\ell > \sqrt{2 r R}$

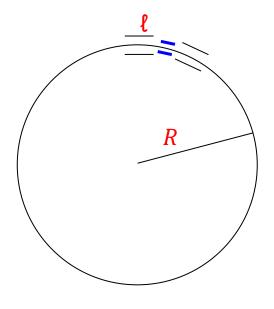
for R = 9 km and r = 4.5 cm $\ell > 28.5 \text{ m}$

for R = 3.1 km and r = 6.5 cm (LEP2) $\ell > 20$ m

totally escaping (-> hitting downstream elements) for shorter dipoles

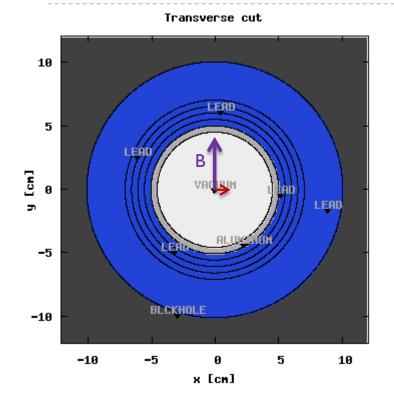
Pb shielding in the interconnects?

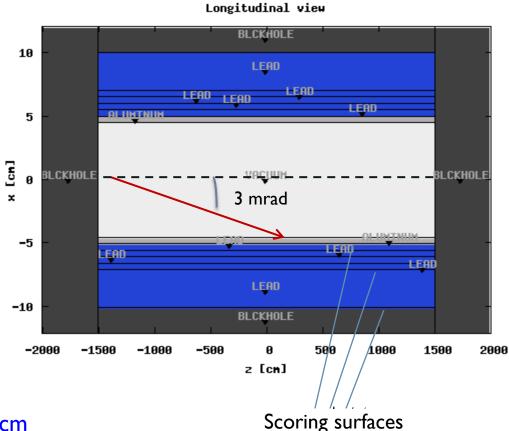
For the time being: impact angle as for "curved" geometry (eg, very short magnets, or very long curved ones)





Tlep Idealized geometry

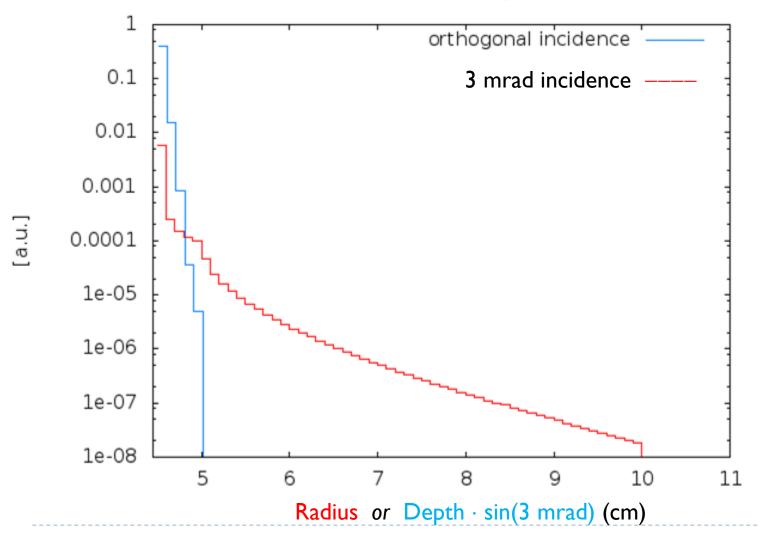




- □ Vacuum pipe: round R = 4.5 cm
- ☐ Aluminum pipe: thickness = 0.5 cm
- ☐ Lead shielding: thickness = 5.0 cm

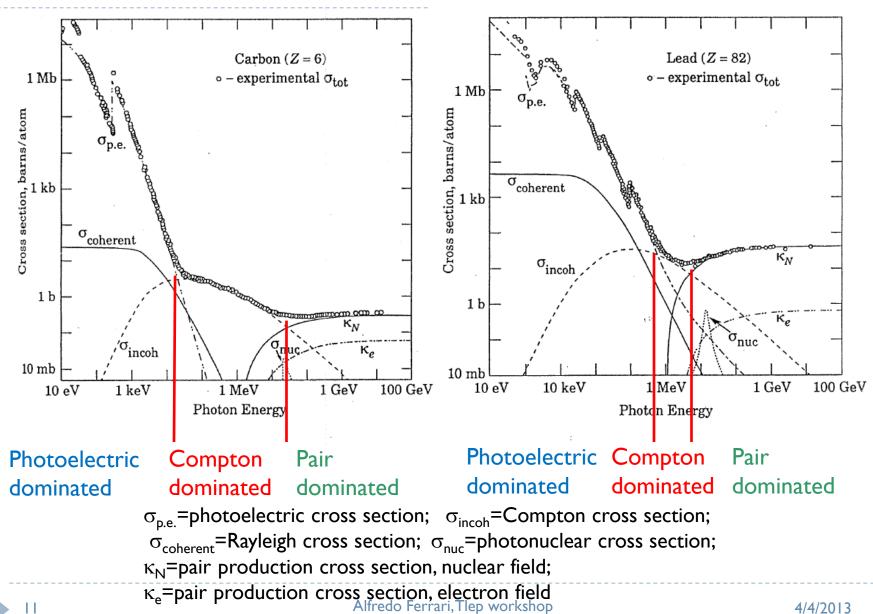
How different is the attenuation vs the equivalent line-of-sight?

photon tracklength



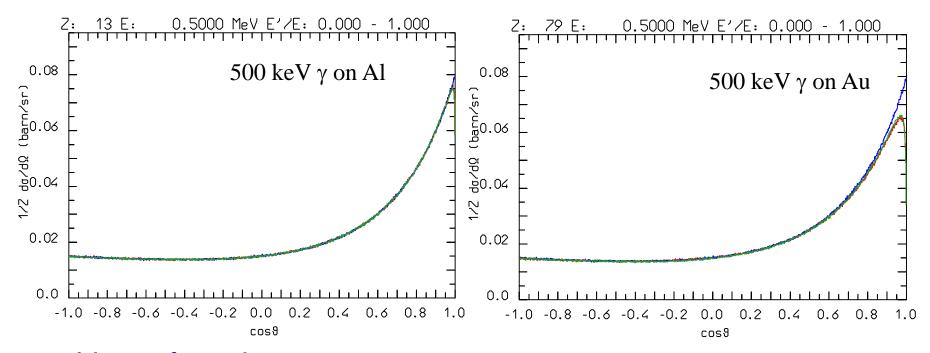
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Photon cross sections:





Compton ang. distr.: examples



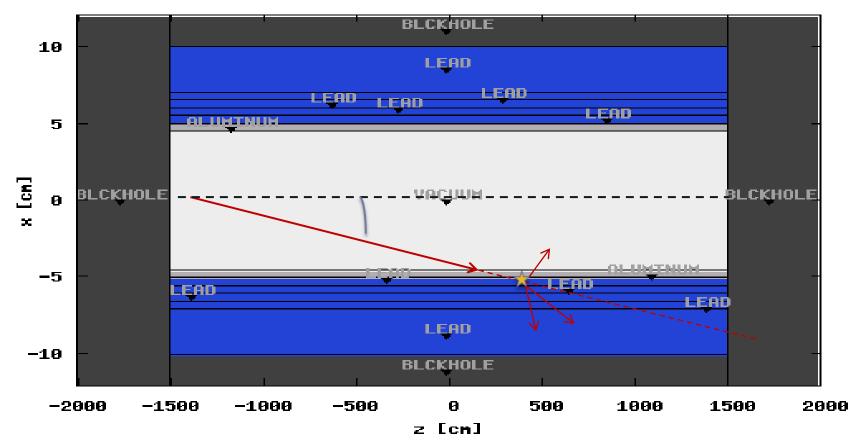
blue = free electron
green = binding with form factors
red = binding with shells and orbital motion



EN Engineering Department The physical explanation

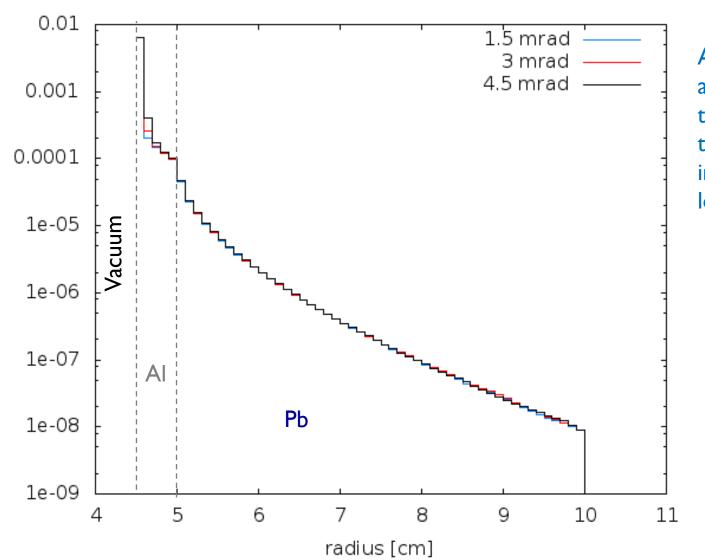
The first scattering effect: after a Compton interaction the photon loses "memory" of the initial, grazing, incidence because of the much larger scattering angle

Longitudinal view



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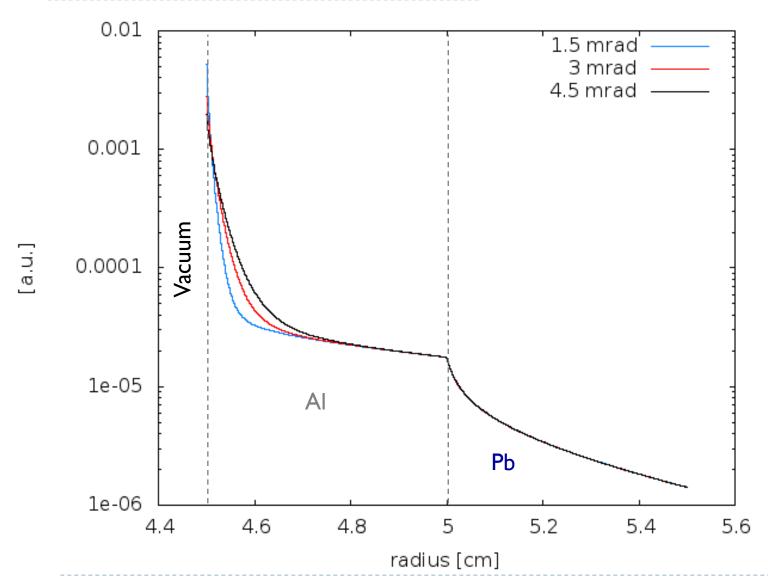
Photon fluence attenuation for 3 angles



As expected, after a few mm the memory of the initial incidence angle is lost

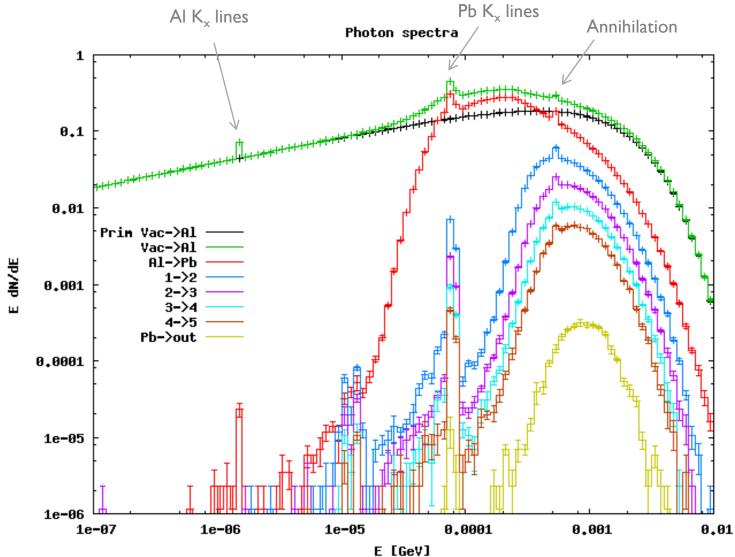


Photon fluence attenuation curves for 3 angles



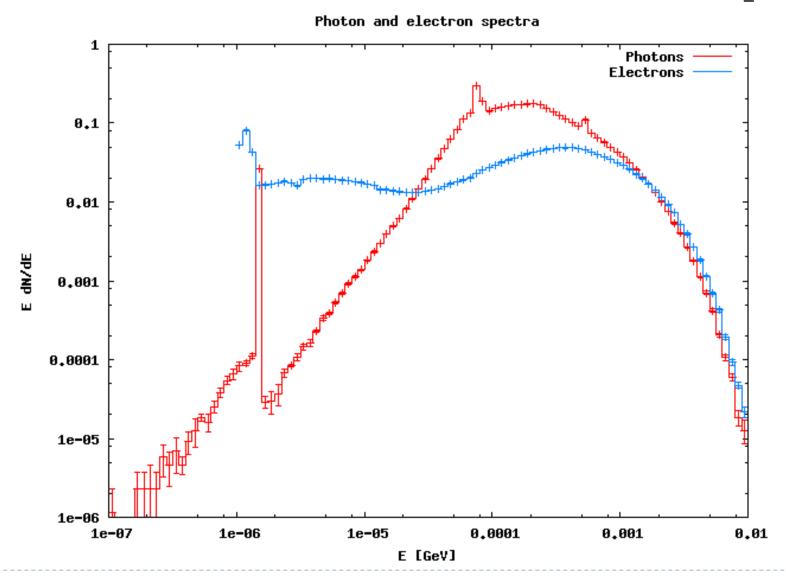


Photon spectra at various depths

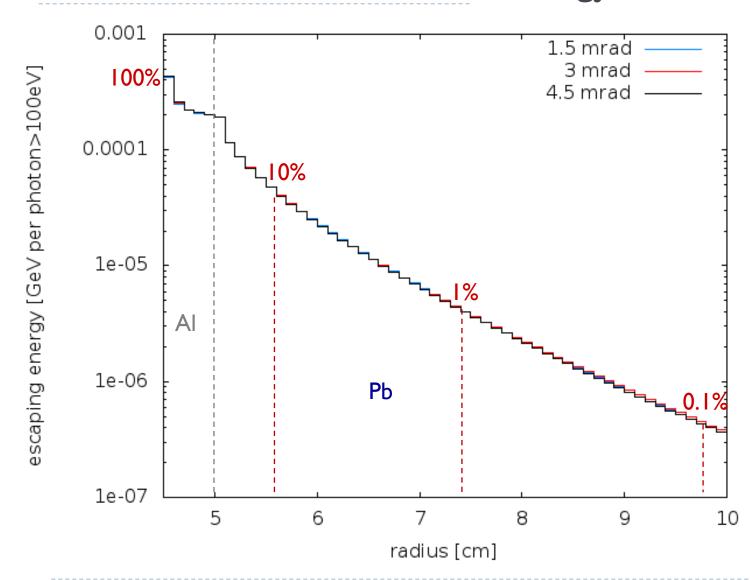




Backscattered photon and electron spectra

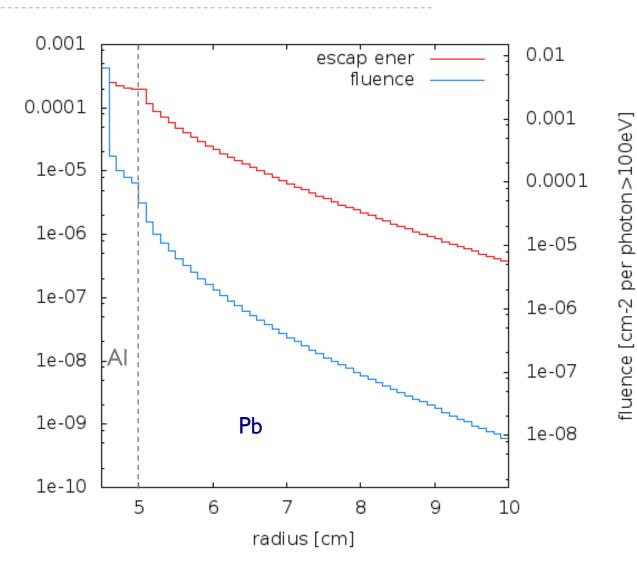


Power attenuation: escaping energy as a function of radius





Escaping energy vs fluence as a function of radius



Remarks, after 5 cm of Pb:

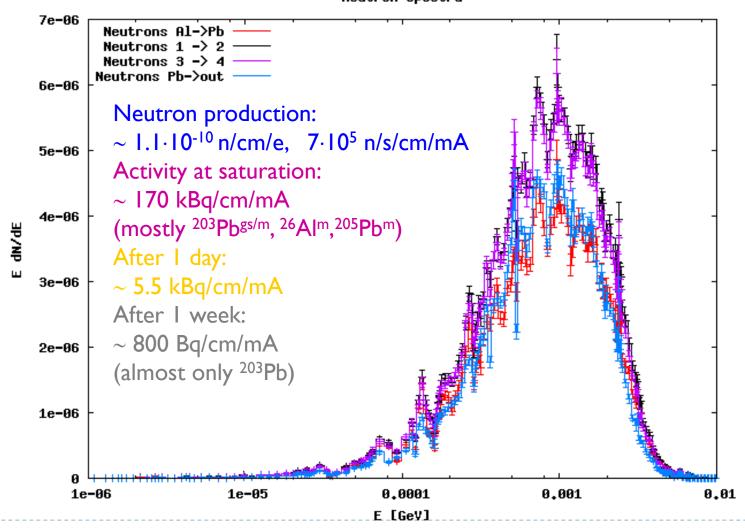
- Power and fluence are not yet exponentially attenuated
- The escaping power is decreasing slower than the photon fluence

escaping energy [GeV per photon>100eV]



Neutron spectra at various depths

Neutron spectra





Conclusions

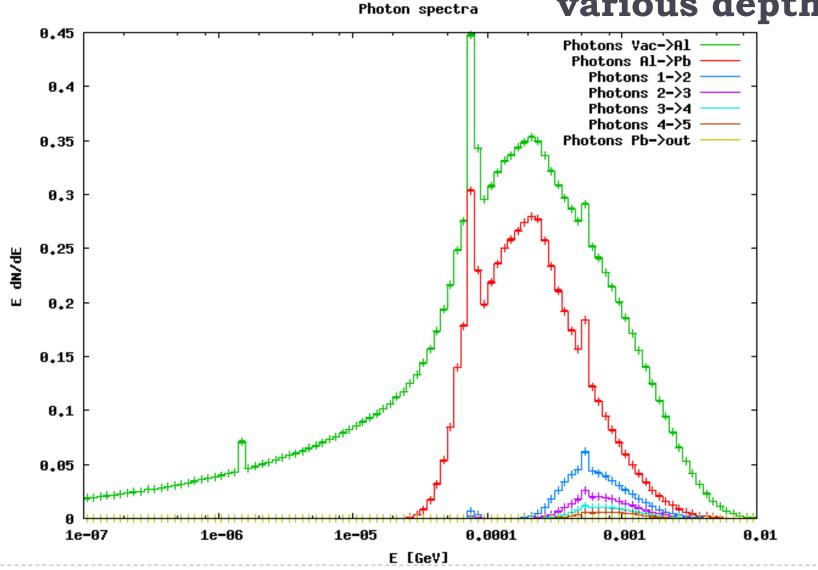
- ▶ **SR calculations possible** with full generality
- Some minimal layout specs (dipole length, curved or straight, beam pipe radius) required in order to start devising a shielding strategy, maybe possible to intercept most of SR at interconnections?
- Specs about the "tolerable" escaping power levels required as well
- As expected the attenuation curve is insensitive to the incidence angle and (unfortunately) far from naïve line-of-sight approximations
- Photoneutron production and associated activation (relatively)
 minor, it will change steeply with E_c



Backup Slides

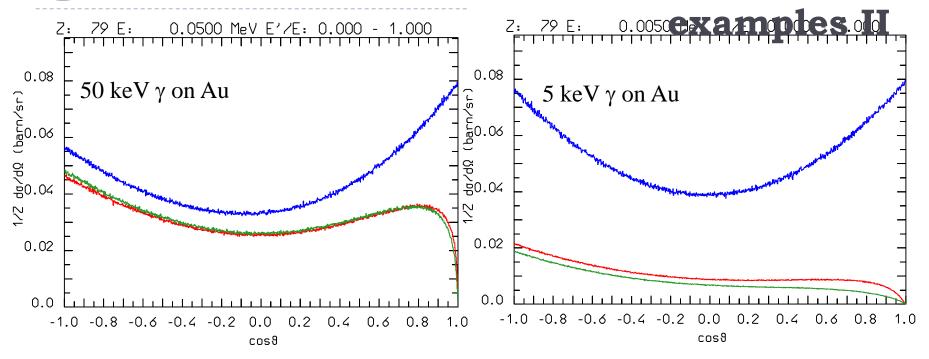


Photon spectra at various depths





Compton ang. distr.:



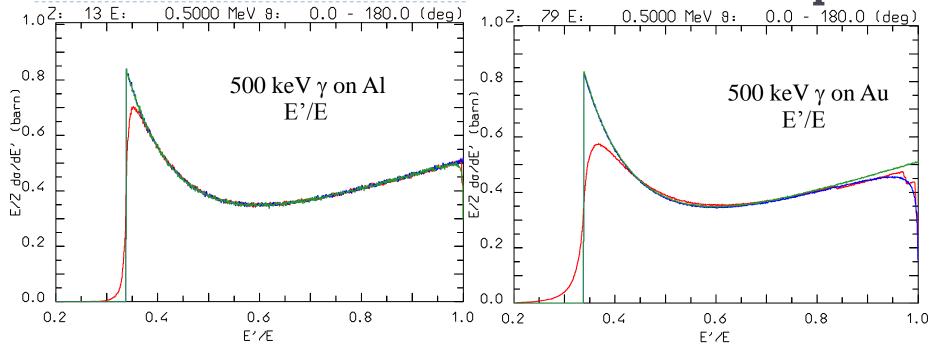
blue = free electron
green = binding with form factors
red = binding with shells and orbital motion

Effects visible only at $\cos\theta$ close to 1. The S(q,Z) approximation is still very good at 50 keV,

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Compton profile:

examples

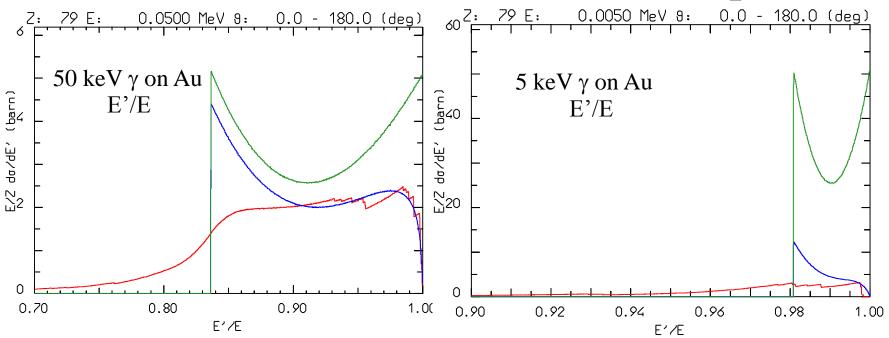


green = free electron
blue = binding with form factors
red = binding with shells and orbital motion

Larger effect at very low energies, where, however, the dominant process is photoelectric. Visible: shell structure near E'=E, smearing from motion at low E'



Compton profile: examples II



green = free electron
blue = binding with form factors
red = binding with shells and orbital motion

Larger effect at very low energies, where, however, the dominant process is photoelectric. Please note that the actual cross section goes down again at low energies!! Visible: shell structure near E'=E, smearing from motion at low E'