Implications of Higgs boson discovery and other data for SUSY - A theorist's perspective

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L. Roszkowski, 22/3/2013

Outline

- \diamond Introduction: SUSY primer
- \diamond How to compare theory with data
- Implications of mh~126 GeV for favored SUSY mass scale
- \diamond Probe CMSSM with DM searches
- Implications of BR(B_s to mu mu)
- \diamond Beyond the CMSSM
- \diamond Comments on g-2

♦ Summary

Based on:

- Two ultimate tests of constrained SUSY, <u>1302.5956</u>
- The Constrained NMSSM with a 125 GeV Higgs boson -- A global analysis, <u>1211.1693</u>
- Constrained MSSM favoring new territories: The impact of new LHC limits and a 125 GeV Higgs boson, 1206.0264 ...with updates

BayesFITS group: A. Fowlie (UoS), M. Kazana, K. Kowalska, S. Munir, E. Sessolo, S. Tsai, S. Trojanowski, LR, plus external collaborators (M. Misiak, K. Turzyński, K. Jedamzik)



Many open questions in particle physics

- Origin of particle masses?
- Origin of EWSB?
- Origin and structure of flavor and CP X?
- New physics beyond the Standard Model?
- Dark matter in the Universe?
- Why is the Universe made up of matter and not antimatter?
- Unification of fundamental forces?
- Role of gravity?
- History of the early Universe?
 - LHC: chance to shed light on some of them

Many BSM ideas waiting to be tested...

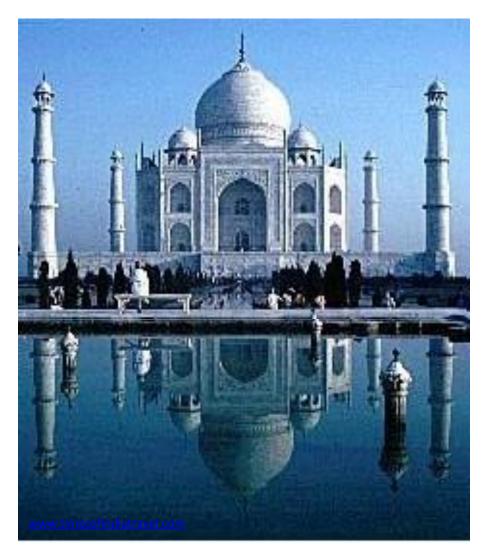
Supersymmetry of several sorts

... by far most attractive

- Large/warped extra dimensions
- Low-scale gravity, microscopic black holes
- Little Higgs framework
- Extra gauge bosons
- Extra fermions
- Extra interactions

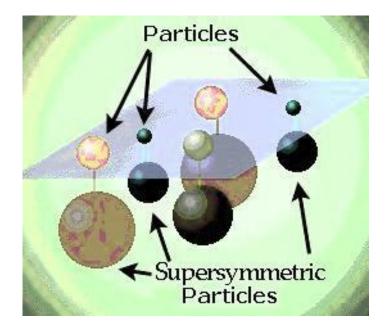


Supersymmetry

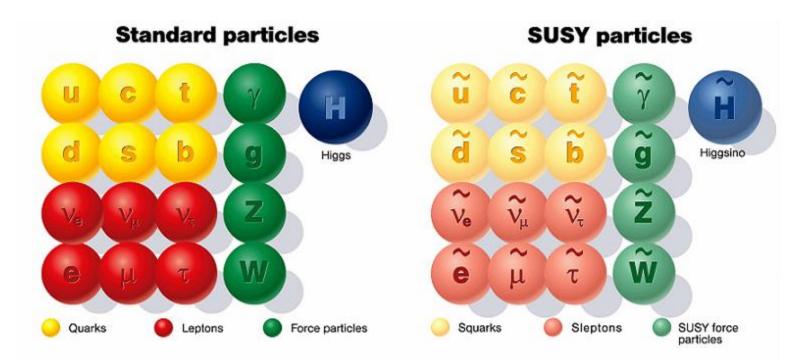


Symmetry among particles

bosons <-> fermions



Supersymmetry



SUSY: superpartners of known SM particles; mass expected at ~1 TeV

"neutralino" χ : lightest mass state of neutral gauginos and higgsinos stable WIMP: excellent DM candidate

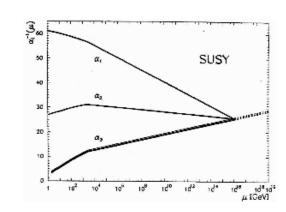
L. Roszkowski, 22/3/2013

Supersymmetry

particle physics

- grand unification,
- supergravity, superstrings,
- hierarchy/naturalness/fine-tuning,
- fermion masses and mixings,
- neutrino masses and mixings,
- CP/flavor violation,

. . .



astroparticle physics

- WIMP dark matter,
- E–WIMPs: gravitinos and axinos
- other relics

particle cosmology

- cosmic inflation,
- baryogenesis/leptogenesis,
- relic production and decay after BB,
- effect on and constrains from BBN,
- effect on and constrains from CMB,

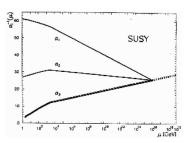
SUSY has dominated theoretical efforts beyond the SM for the last two-three decades...



SUSY: Constrained or Not?

• Constrained:

Low-energy SUSY models with grand-unification relations among gauge couplings and (soft) SUSY mass parameters



Virtues:

- Well-motivated
- Predictive (few parameters)
- Realistic

Many models:

- CMSSM (Constrained MSSM): 4+1 parameters
- NUHM (Non-Universal Higgs Model): 6+1
- CNMSSM (Constrained Next-to-MSSM) 5+1
- 22/03/2455M-NUHM: 7+1

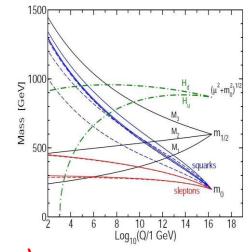


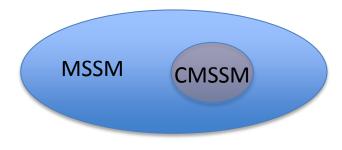
figure from hep-ph/9709356

<u>Phenomenological:</u>

Supersymmetrized SM...

Features:

- Many free parameters
- Broader than constrained SUSY

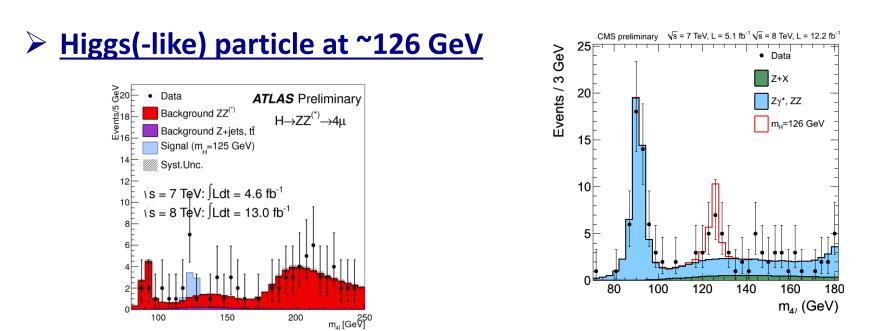


Many models:

- general MSSM over 120 params
- MSSM + simplifying assumptions
- pMSSM: MSSM with 19 params
- p9MSSM, p12MSSM, pnMSSM, ...
- L. Roszkowski, Moriond 9-16/3/13

• etc

Main news from the LHC so far...



No (convincing) deviations from the SM

$${
m BR}(\overline{B}_s
ightarrow \mu^+ \mu^-) = \left(3.2^{+1.5}_{-1.2}
ight) imes 10^{-9}$$

Stringent lower limits on superparner masses

SUSY masses close to 1 TeV scale...

...and from the media...

Is Supersymmetry Dead?

The grand scheme, a stepping-stone to string theory, is still high on physicists' wish lists. But if no solid evidence surfaces soon, it could begin to have a serious PR problem

SCIENTIFIC AMERICAN[™]

April 2012

Nothing new...



The negative result illustrates the risks of Big Science, and its often sparse pickings.

By MALCOLM W. BROWNE

Three hundred and fifteen physicists worked on the experiment.

Their apparatus included the Tevatron, the world's most powerful particle accelerator, as well as a \$65 million detector weighing as much as a warship, an advanced new computing system and a host of other innovative gadgets.

But despite this arsenal of brains and technological brawn assembled at the Fermilab accelerator laboratory, the participants have failed to find their quarry, a disagreeable reminder that as science gets harder, even Herculean efforts do not guarantee success.

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27 August 2011 Last updated at 06:41 GMT

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LHC results put supersymmetry theory 'on the spot'



By Pallab Ghosh Science correspondent, BBC News

Results from the Large Hadron Collider (LHC) have all but killed the simplest version of an enticing theory of sub-atomic physics.

Researchers failed to find evidence of so-called "supersymmetric" particles, which many physicists had hoped would plug holes in the current theory.

Theorists working in the field have told BBC News that they may have to come up with a completely new idea.

Data were presented at the Lepton Photon science meeting in Mumbai.

Supersymmetry predicts the existence of mysterious

super particles

Related Stories

Increasing energy, luminosity and the number of physicist failing to find SUSY have increased by factor of 10...

Constrained SUSY – still alive?

The constrained MSSM (CMSSM) paradigm is "hardly tenable"

At Open Symposium of the European Strategy Preparatory Group, Krakow, Poland, 10-12 Sept. 2012

Really?



Higgs discovery: ways to go

- SM confirmed end of the story (and collider physics?)
- Higgs is fundamental > SUSY
- Higgs is composite -> effective theory

In this talk: explore implications of Higgs properties for SUSY

Do not worry about theoretical and/or aesthetic arguments (finetuning, naturalness, etc).

See what the data says!

126 GeV Higgs boson and SUSY

- Is the discovered Higgs boson consistent with SUSY?
 Which SUSY model(s)?
- What does the mass of ~126 GeV tell us about superpartner masses?
- Is a SM-like Higgs boson a natural prediction of SUSY, or rather an oddity?
- Is the found Higgs consistent with all other relevant observational constraints on SUSY?

SUSY - most important constraints:

CMS: $m_h \sim 125.8 \text{ GeV}$ (in ZZ); $m_h = 124.9 \text{ GeV}$ (in $\gamma\gamma$)

ATLAS: $m_h = 124.3 \text{ GeV}$ (in ZZ); $m_h = 126.8 \text{ GeV}$ (in $\gamma\gamma$)

Direct search limits

Higgs mass

Lower limit...

Dark matter density

Positive measurement, inconsistent with SM

B_s -> mu mu

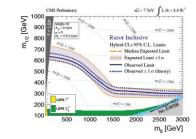
 ${
m BR}(\overline{B}_s o \mu^+ \mu^-) = \left(3.2^{+1.5}_{-1.2}
ight) imes 10^{-9}$

LHCb (Nov 2012)

- Other flavor (b to s gamma, etc)
- EW observables (M_W,...)

Baves

DITE



How to compare theory with experiment

Rigid step-function application of limits/allowed ranges (e.g. DM relic abundance, etc) Mahmoudi et al, Hewett et al, ... Frequentist (chi^2-based) MasterCode, Fittino, ... Bayesian

BayesFITS, Allanach, SuperBayes, Balazs,...

Frequentist: "probability is the number of times the event occurs over the total number of trials, in the limit of an infinite series of equiprobable repetitions"

Bayesian: "probability is a measure of the degree of belief about a proposition"

Both F and B are based on the likelihood function.



The Likelihood function

Central object: Likelihood function

Positive measurements:

Take a single observable $\xi(m)$ that has been measured

- c central value, σ standard exptal error
- define

$$\chi^2 = rac{[\xi(m)-c]^2}{\sigma^2}$$

assuming Gaussian distribution $(d \rightarrow (c, \sigma))$:

$$\mathcal{L} = p(\sigma, c | m{\xi}(m)) = rac{1}{\sqrt{2\pi}\sigma} \exp\left[-rac{\chi^2}{2}
ight]$$

• when include theoretical error estimate au (assumed Gaussian):

$$\sigma
ightarrow s = \sqrt{\sigma^2 + au^2}$$

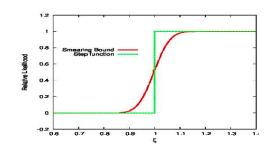


 $(e.g., M_W)$

for several uncorrelated observables (assumed Gaussian):

$$\mathcal{L} = \exp\left[-\sum_i rac{\chi_i^2}{2}
ight]$$

Limits:



- Smear out bounds.
- Add theory error.
- <u>LHC direct limits:</u>
 - Need careful treatment. Typically use Poisson.



Bayesian statistics

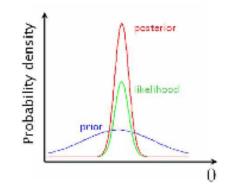


Bayes theorem: $Posterior = \frac{Prior \times Likelihood}{Evidence}$

- **Prior**: what we know about hypothesis BEFORE seeing the data.
- Likelihood: the probability of obtaining data if hypothesis is true.
- **Posterior**: the probability about hypothesis AFTER seeing the data.
- Evidence: normalization constant, crucial for model comparison.

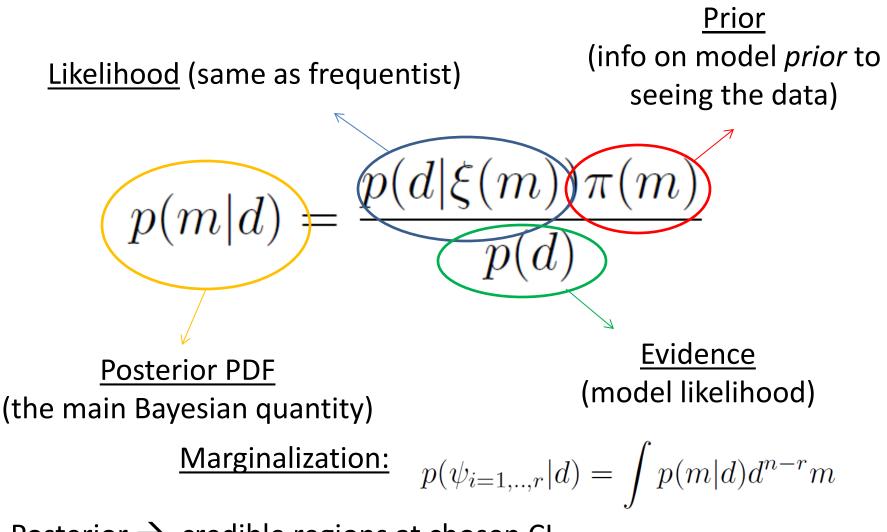
If hypothesis is a function of parameters, then posterior becomes posterior probability function (pdf).

Posterior \rightarrow credible regions at chosen CL



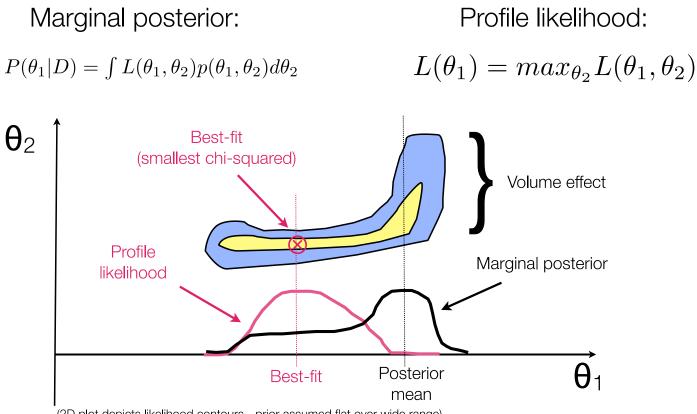


Bayesian Approach



Posterior \rightarrow credible regions at chosen CL

Bayesian \neq Frequentist



(2D plot depicts likelihood contours - prior assumed flat over wide range)

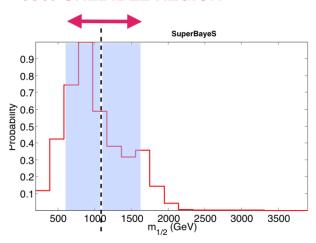


Bayesian \neq Frequentist

Bayesian

Best fit plays no special role.

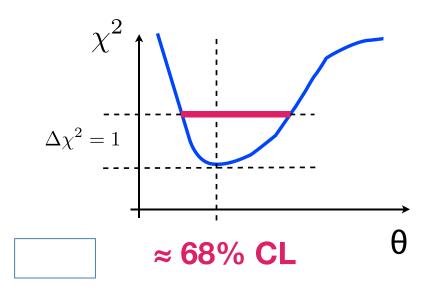
Central object: posterior probability



68% CREDIBLE REGION

Frequentist

Determine best-fit parameters: find minimum of -2Log(Likelihood)=chi^2



Determine posterior credible regions: e.g. symmetric interval around the mean containing 68% of samples Determine approximate confidence intervals: Local Δ (chi-squared) method

The Likelihood function

Central object: Likelihood function

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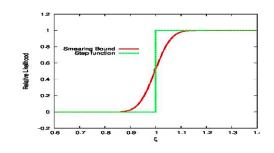


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- <u>LHC direct limits:</u>
 - Need careful treatment. Typically use Poisson.



Constrained Minimal Supersymmetric Standard Model (CMSSM)

G. L. Kane, C. F. Kolda, L. Roszkowski and J. D. Wells, Phys. Rev. D 49 (1994) 6173

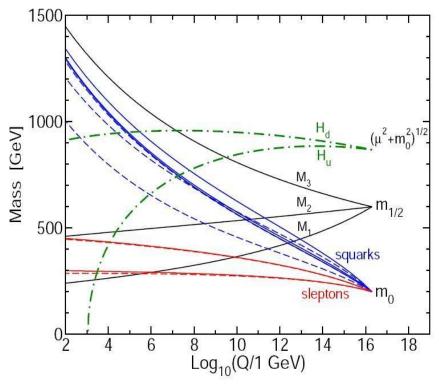


figure from hep-ph/9709356

Bayes

FITS22/03/2013

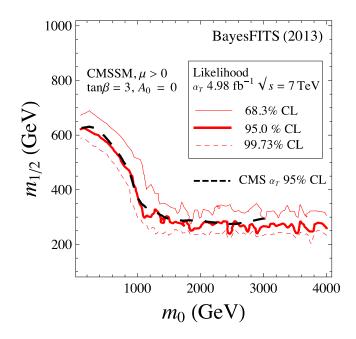
At $M_{\rm GUT} \simeq 2 \times 10^{16} \, {\rm GeV}$:

- gauginos M₁ = M₂ = m_ğ = m_{1/2}
 scalars m²_{q̃i} = m²_{l̃i} = m²_{Hb} = m²_{Ht} = m²₀
 3-linear soft terms A_b = A_t = A₀
 radiative EWSB μ² = m²_{Hb} - m²_{Ht} tan² β - m²_Z tan² β - 1 - m²_Z
 five independent parameters: m_{1/2}, m₀, A₀, tan β, sgn(μ)
- well developed machinery to compute masses and couplings

Reproducing CMS limits on SUSY

We approximate CMS limits by deriving likelihood maps

First, validate our method:

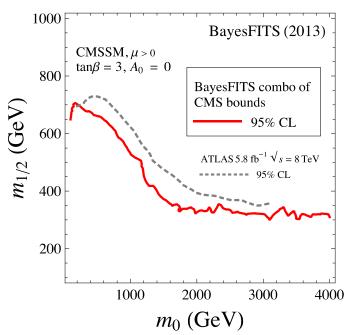


Excellent agreement

Next, derive combined CMS limit based on datasets:

 $\alpha_T \, 11.7/\mathrm{fb}, \, \sqrt{s} = 8 \, \mathrm{TeV}$

Razor 4.4/fb, $\sqrt{s} = 7 \,\mathrm{TeV}$



Applies to both signs of mu And to similar models: NUHM, CNMSSM,...



Below will use combined CMS fimit via likelihood function

Specialty 24 of BayesFITS

CMSSM: numerical scans

- Perform random scan over 4 CMSSM +4 SM (nuisance) parameters <u>simultaneously</u>
- Very wide ranges:

$$egin{aligned} 100\,{
m GeV} &\leq m_0 \leq 20\,{
m TeV} \ 100\,{
m GeV} &\leq m_{1/2} \leq 10\,{
m TeV} \ -20\,{
m TeV} \leq A_0 \leq 20\,{
m TeV} \ &3 \leq aneta \leq 62 \end{aligned}$$

 Use Nested Sampling algorithm to evaluate posterior

Γ	Nuisance	Description	Central value \pm std. dev.	Prior Distribution
	M_t	Top quark pole mass	$173.5 \pm 1.0 \mathrm{GeV}$	Gaussian
	$m_b(m_b)_{ m SM}^{\overline{MS}}$	Bottom quark mass	$4.18\pm0.03{\rm GeV}$	Gaussian
	$\alpha_s(M_Z)^{\overline{MS}}$	Strong coupling	0.1184 ± 0.0007	Gaussian
	$1/\alpha_{ m em}(M_Z)^{\overline{MS}}$	Inverse of em coupling	127.916 ± 0.015	Gaussian

Use 4 000 live points

Use Bayesian approach (posterior)



Hide and seek with SUSY

The experimental measurements that we apply to constrain the CMSSM's parameters. Masses are in GeV.

	Measurement	Mean or Range	Error: (Exp., Th.)	Distribution
	Combination of:			
	CMS razor 4.4/fb , $\sqrt{s} = 7 \text{TeV}$	See text	See text	Poisson
	CMS $\alpha_T \ 11.7/\text{fb}$, $\sqrt{s} = 8 \text{ TeV}$	See text	See text	Poisson
	m_h by CMS	$125.8{ m GeV}$	$0.6{ m GeV}, 3{ m GeV}$	Gaussian
	$\Omega_\chi h^2$	0.1120	0.0056,10%	Gaussian
>	$\delta \left(g-2 ight)^{ m SUSY}_{\mu} imes 10^{10}$	28.7	8.0, 1.0	Gaussian
	$\operatorname{BR}\left(\overline{B} \to X_s \gamma\right) \times 10^4$	3.43	0.22,0.21	Gaussian
	$BR(B_u \to \tau \nu) \times 10^4$	1.66	0.33, 0.38	Gaussian
	ΔM_{B_s}	$17.719{\rm ps}^{-1}$	$0.043 \mathrm{ps}^{-1}, \ 2.400 \mathrm{ps}^{-1}$	Gaussian
	$\sin^2 heta_{ m eff}$	0.23116	0.00012, 0.00015	Gaussian
	M_W	80.385	0.015, 0.015	Gaussian
	$BR\left(B_s \to \mu^+ \mu^-\right)_{current} \times 10^9$	3.2	+1.5 - 1.2, 10% (0.32)	Gaussian
	$\frac{\mathrm{BR} \left(B_s \to \mu^+ \mu^- \right)_{\mathrm{current}} \times 10^9}{\mathrm{BR} \left(B_s \to \mu^+ \mu^- \right)_{\mathrm{proj}} \times 10^9}$	$3.5 (3.2^*)$	$0.18~(0.16^*),~5\%~[0.18~(0.16^*)]$	Gaussian

SM value: $\simeq 3.5 \times 10^{-9}$

10 dof



SUSY - most important constraints:



CMS: $m_h \sim 125.8 \text{ GeV}$ (in ZZ); $m_h = 124.9 \text{ GeV}$ (in $\gamma\gamma$) ATLAS: $m_h = 124.3 \text{ GeV}$ (in ZZ); $m_h = 126.8 \text{ GeV}$ (in $\gamma\gamma$)

Direct search limits

Lower limit...

Dark matter density

Positive measurement, inconsistent with SM

➢ B_s → mu mu

 ${
m BR}(\overline{B}_s
ightarrow \mu^+ \mu^-) = \left(3.2^{+1.5}_{-1.2}
ight) imes 10^{-9}$

LHCb (Nov 2012)

1000

Razor Inclusive Hybrid CLs 95% C.L. Limits Median Expected Lim

1500 2000

2500 3000 m_n [GeV]

- Other flavor (b to s gamma, etc)
- EW observables (M_W,...)

Bayes

DITE

~126 Gev Higgs in SUSY

- In SUSY m_h is a calculated quantity.
- 1-loop corr: positive, up to ~45 GeV

$$\Delta m_h^2 = \frac{3m_t^4}{4\pi^2 v^2} \left[\ln\left(\frac{M_{\rm SUSY}^2}{m_t^2}\right) + \frac{X_t^2}{M_{\rm SUSY}^2} \left(1 - \frac{X_t^2}{12M_{\rm SUSY}^2}\right) \right]$$

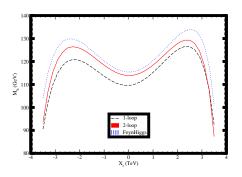
• 2-loop corr: negative, ~3 GeV

two most complete calculations differ by a 2-5 GeV

(DR-bar (Slavich,...) used in SoftSusy, Spheno, Suspect, and on-shell (Hollik,...) in FeynHiggs

Substantial theory error!





$$\begin{split} M_{\rm SUSY} &\equiv \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} \\ X_t &= A_t - \mu \cot \beta \end{split}$$

Two ways to obtain m_h~126 GeV:
1. increase M_SUSY -> heavy superparners! or
2. take large |X_t|~|A_t|-> stop_1 at ~1TeV

Applies to SUSY generally, not just constrained models.

~126 GeV Higgs in the CMSSM

 $m_{1/2}$ (TeV)

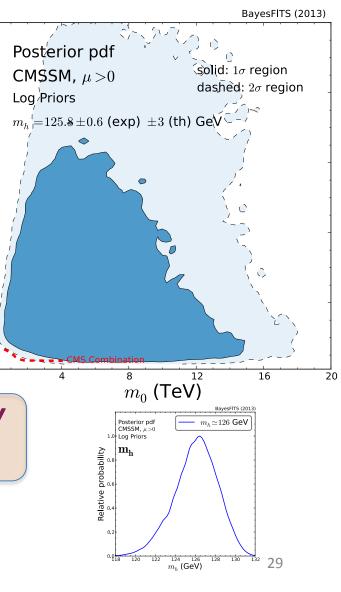
 Include <u>only</u> m_h~126 GeV and lower limits from direct SUSY searches

$$\mathcal{L} \sim e^{rac{(m_h-125.8\,{
m GeV})^2}{\sigma^2+ au^2}}$$

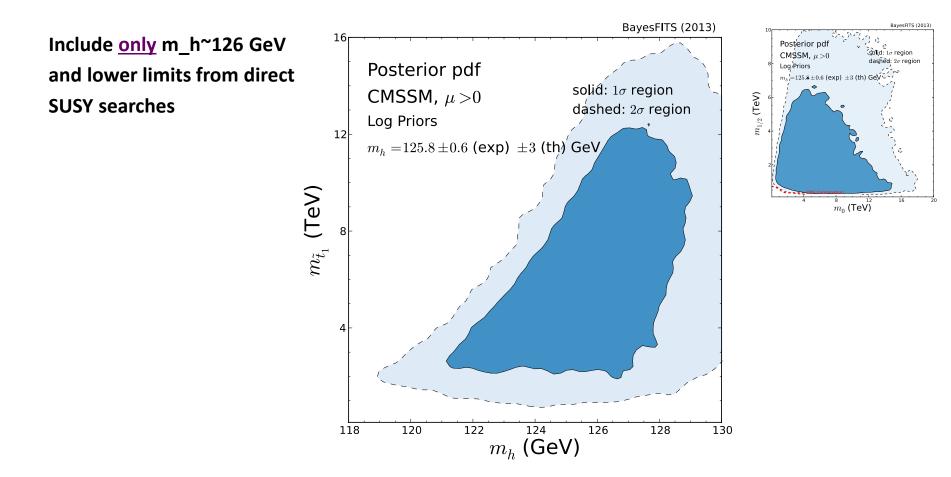
$$\sigma=0.6~{\rm GeV}, \tau=2~{\rm GeV}$$

~126 GeV Higgs mass implies multi-TeV SUSY masses

NO tension with LHC direct lower limits



~126 GeV Higgs in the CMSSM



~126 GeV Higgs mass implies multi-TeV SUSY

masses

SUSY - most important constraints:

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Direct search limits

Higgs mass

Lower limit...

Dark matter density

Positive measurement, inconsistent with SM

B_s -> mu mu

 $BR(\overline{B}_s \to \mu^+ \mu^-) = \left(3.2^{+1.5}_{-1.2}\right) \times 10^{-9}$ LHCb (Nov 2012)

- Other flavor (b to s gamma, etc)
- EW observables (M_W,...)

Bayes

DITE

Razor Inclusive Hybrid CLs 95% C.L. Limits Median Expected Lim

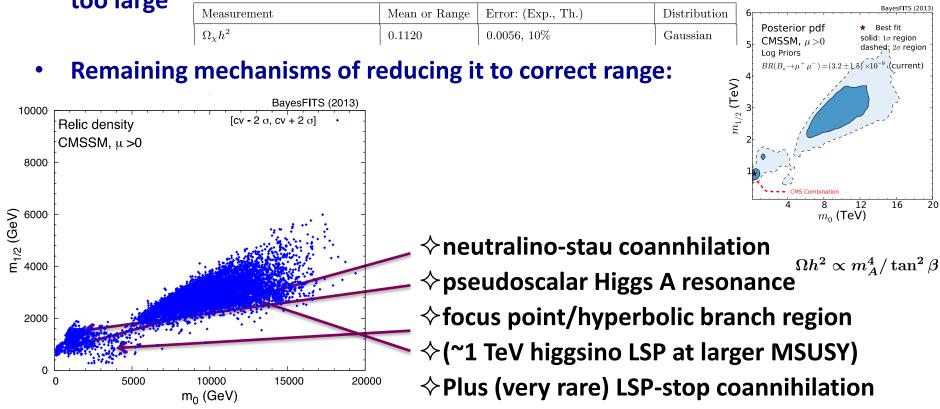
1500 2000

2500 3000 m_n [GeV]

1000

Dark matter density

• Unified SUSY: neutralino relic density is typically 1-2 orders of magnitude too large



Scan with all other relevant constraints imposed



CMSSM: these are the only DM-favored regions

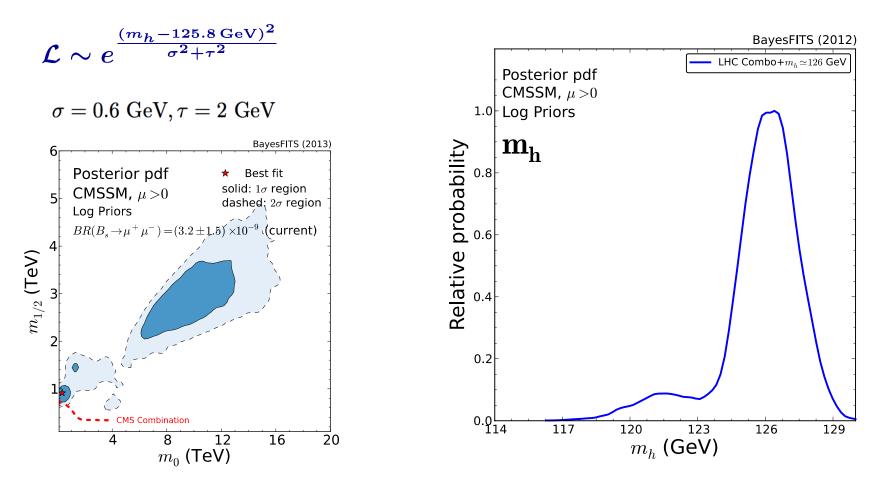
L. Roszkowski, 22/3/2013



Light Higgs in the CMSSM

Likelihood function

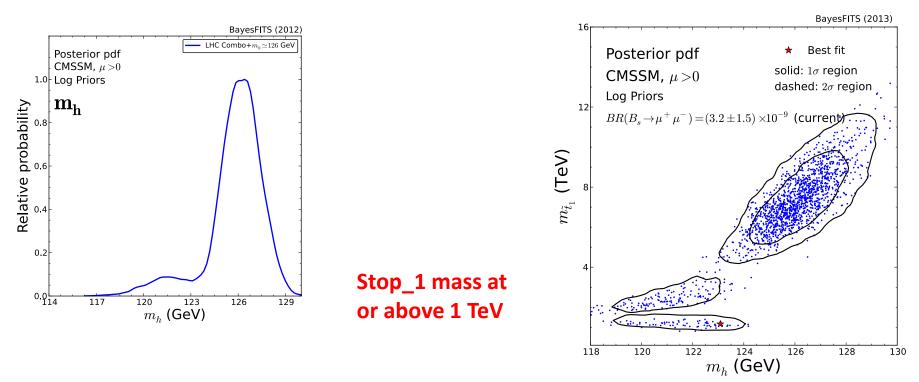
...with all relevant constraints imposed



~126 GeV Higgs at/near lowest chi2 (S.C./AF) and at X_SUSY>> 1TeV



Higgs vs stop mass



Best fit to ~126 GeV Higgs -> M_SUSY~ or >> 1 TeV $egin{aligned} ext{best-fit point } \chi^2_{ ext{min}}/ ext{dof} &= 18.26/10 \ & [\chi^2_{ ext{min}}/ ext{dof} &\simeq 4/9 ext{ when drop } (g-2)_\mu] \end{aligned}$

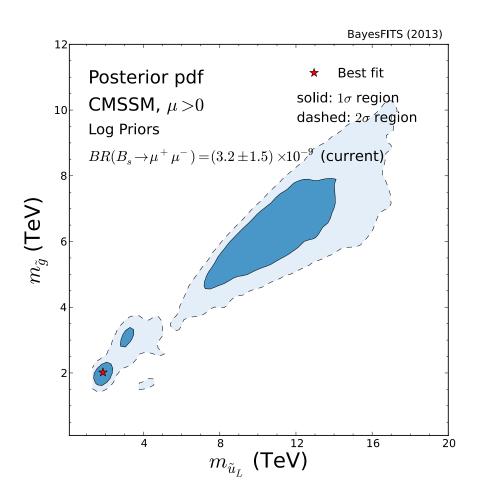
Dark matter relic density: selects some regions

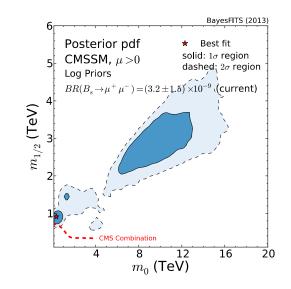
L. Roszkowski, 22/3/2013

Can mulit-TeV ranges of parameters be experimentally tested?



LHC?

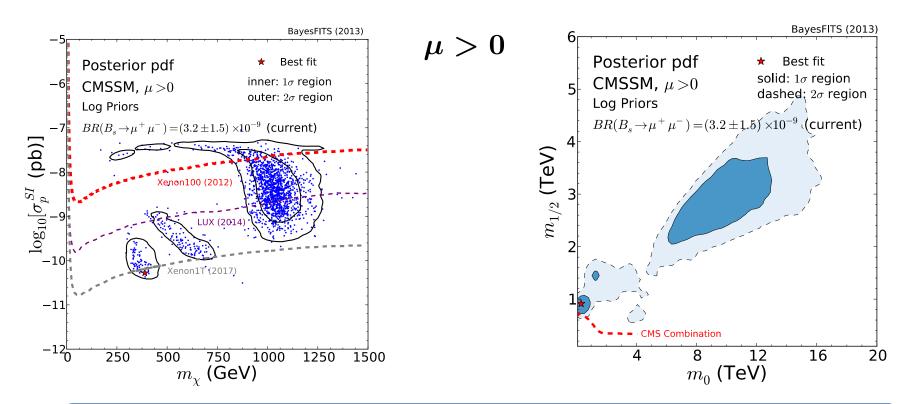




LHC reach: Gluino: ~2.7 GeV Squarks: ~3 TeV

...not guaranteed





1-tonne DM detectors to cover most of CMSSM predictions

... over ALL multi-TeV ranges of mass parameters

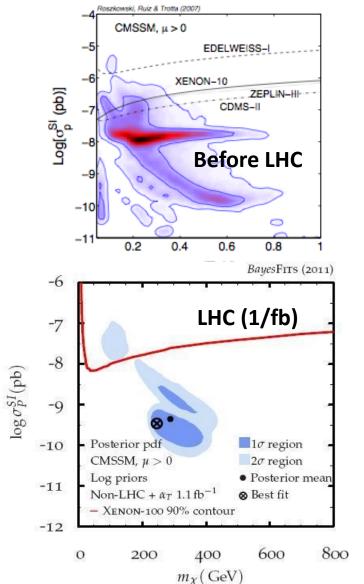
(Except for some cases at mu<0)

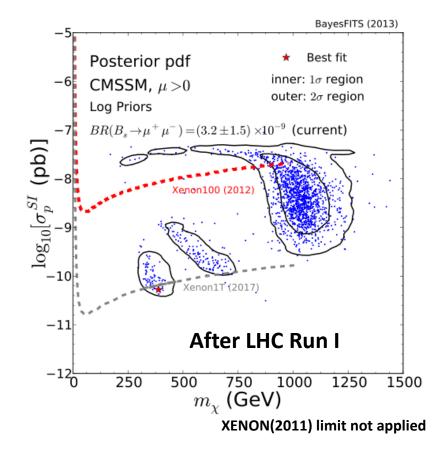
LUX (2014) to improve sensitivity by ~1 decade

Generic prediction of multi-TeV SUSY: ~1TeV LSP (higgsino)
L. Roszkowski, 22/3/2013



LHC: Impact on DM Searches





extended ranges of SUSY parameters
no other SUSY regions exist

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Direct search limits

Higgs mass

Bayes

DITE

Lower limit...

Dark matter density

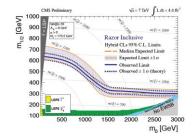
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LHCb (Nov 2012)

- Other flavor (b to s gamma, etc)
- EW observables (M_W,...)





BR(Bs->mu mu)

$${
m BR}(\overline{B}_s o \mu^+ \mu^-) = \left(3.2^{+1.5}_{-1.2}
ight) imes 10^{-9}$$

M. Palutan (LHCb), 13 Nov 2012

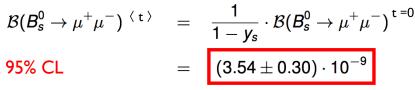
.1"
$$10^{!9} \le B(B^{0}_{s} \rightarrow \mu^{+}\mu^{-}) \le 6.4$$
" $10^{!9}$ at 95% CI

Note this gives weaker upper bound than before.

LHC combination (June 2012): B(B⁰s! µ⁺µ[!])<4.2" 10⁻⁹ at 95% CL

We approximate the signal with a Gaussian





ปี้ ¹ 0.8

0.6 Cobse

expected

0.4

0.2

SM value

De Bruyn et al., PRL 109, 041801 (2012) uses LHCb-CONF-2012-002

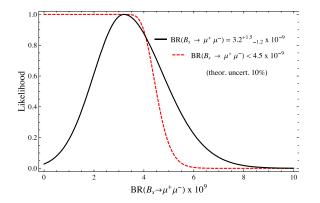
expected

bkg+SM

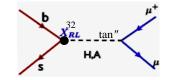
LHCb

 $\begin{array}{ccc} 6 & 0 \\ B(B_s^0 ! & m^+ m^-) & [10^{-9}] \end{array}$

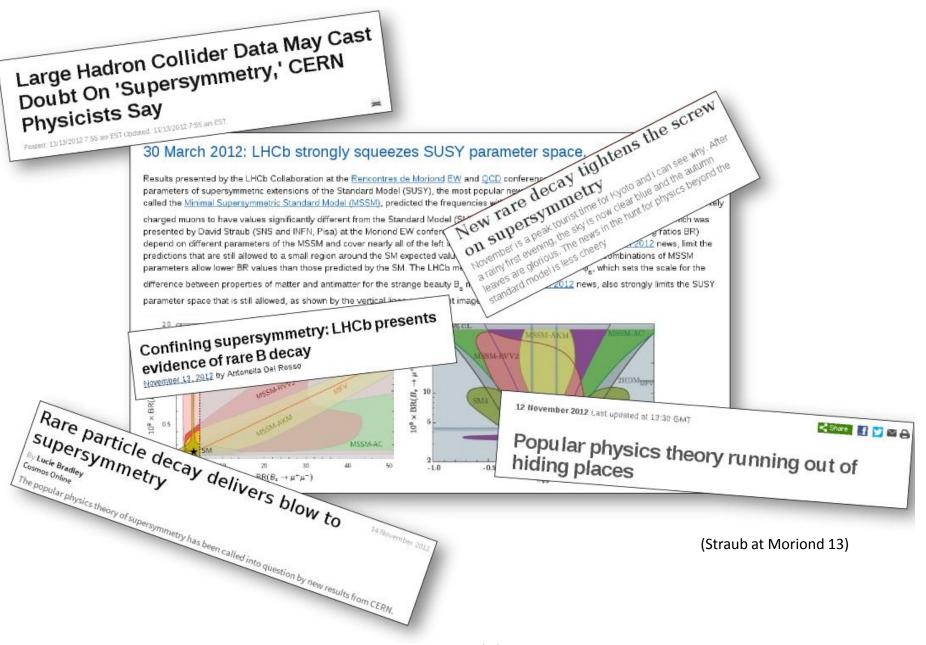
Note the Gaussian Like allows larger BR than 4.2 bound before.



- sensitive probe of new physics ${
m BR}(\overline{B}_s o \mu^+ \mu^-) \propto an^6 \, eta/m_A^4$

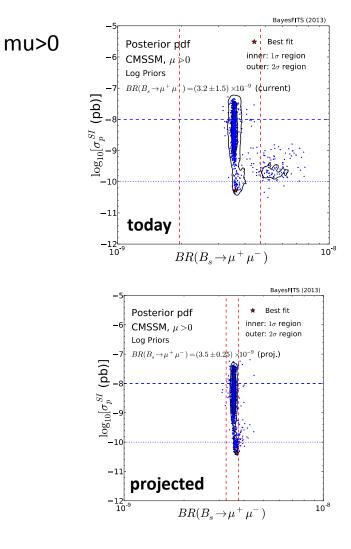


LHCb result agrees with SM value => limits on SUSY 40

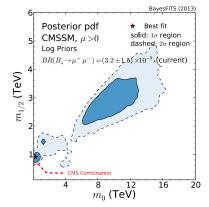


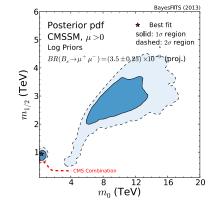
L. Roszkowski, 22/3/2013

Effect of precise $BR(\overline{B}_s \to \mu^+ \mu^-)$



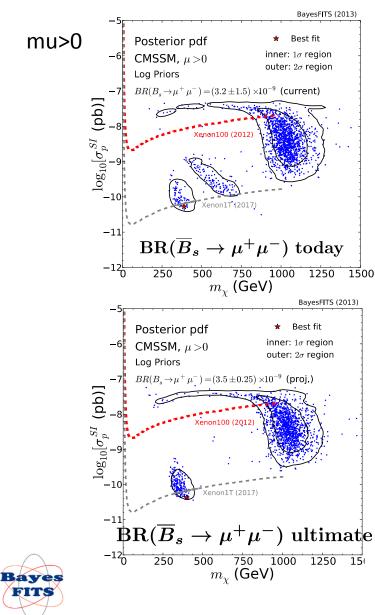
If $BR(\overline{B}_s \to \mu^+ \mu^-) \simeq SM$ value with 5-10% precision (both TH and EXPT) \Rightarrow A funnel region gone







Effect of precise $BR(\overline{B}_s \to \mu^+ \mu^-)$



If $BR(\overline{B}_s \to \mu^+ \mu^-) \simeq SM$ value with 5-10% precision \Rightarrow A funnel region gone

Ways to rule out the CMSSM:

- No DM signal in 1-tonne detectors
- DM signal at ~500 to 750
 GeV

SC: for $\mu < 0 \ \sigma_p^{\text{SI}}$ lower (cancellations)

NUHM, CNMSSM: similar ranges of sigma_p but DM-favored regions overlap

L. Roszkowski, 22/3/2013

Even the simplest unified SUSY model (CMSSM) is consistent with all data (Higgs mass, DM relic density, direct limits, flavor-violating processes, ...)

...except for g-2, R(gamma gamma)

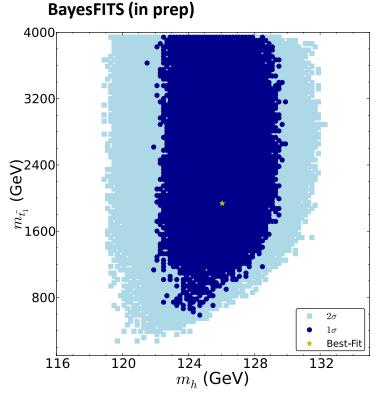
- M_SUSY >~ (or even >>) 1 TeV favored by ~126 GeV Higgs
- In less unified models somewhat lower SUSY masses are allowed (but not by much)

... except for very fine tuned corners

~126 GeV Higgs in general MSSM

• More free parameters, more freedom

...here 9 parameters



~126 GeV Higgs still implies heavy superpartners

...except for very fine tuned corners which allow much lighter staus, stops, charginos

SUSY - most important constraints:

CMS: $m_h \sim 125.8 \text{ GeV}$ (in ZZ); $m_h = 124.9 \text{ GeV}$ (in $\gamma\gamma$)

ATLAS: $m_h = 124.3 \text{ GeV}$ (in ZZ); $m_h = 126.8 \text{ GeV}$ (in $\gamma\gamma$)

Direct search limits

Higgs mass

Lower limit...

Dark matter density

Positive measurement, inconsistent with SM

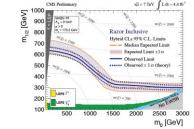
B_s -> mu mu

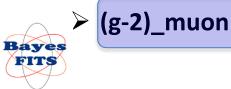
 $\mathrm{BR}(\overline{B}_s
ightarrow \mu^+ \mu^-) = \left(3.2^{+1.5}_{-1.2}
ight) imes 10^{-9}$

LHCb (Nov 2012)

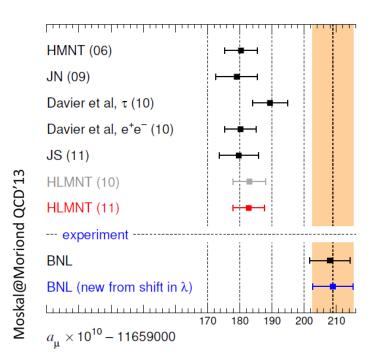
- Other flavor (b to s gamma, etc)
- EW observables (M_W,...)

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(g-2)_muon

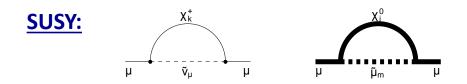


The anomalous magnetic moment of the muon

>3 sigma deviation

Now more believable with recent results on hadr. contribution from Kloe and Kloe-2

New physics?



This is the only result pointing towards low superparner masses!

Need sneutrino/chargino and/or smuon/neutralino in ~ few hundred GeV range

Unified SUSY: sleptons are unified with squarks and are too heavy

General MSSM: if (g-2)_muon anomaly is true: expect light sleptons/chargino/neutralino

... a question on many people's mind...

But what about fine-tuning/naturalness?!

- I prefer to follow what the data implies, rather than theoretical prejudice
- Naturalness: fundamental Higgs -> SUSY
- 126 GeV -> M_SUSY ~1TeV or >> 1TeV
- Fine-tuning is needed at any scale above the EW scale!

1 TeV is not a magic number

- If SUSY is discovered, the FT issue will have to be understood
- If SUSY is not discovered, the issue will become irrelevant
- There are ideas around of how to live comfortably with high fine-tuning



To take home:

 CMSSM: consistent with all experimental constraints. except (g-2)_muon, R(gamma gamma)

(Other simple constrained SUSY models: similar story.)

• Higgs of 126 GeV --> typically M_SUSY at multi-TeV scale.

Plus a window of light stop_1 (~1TeV) – best fit region (stau coann.)

• 1-tonne DM detectors to probe most CMSSM parameters.

Far beyond direct LHC reach.

Other simple constrained SUSY models: similar story.

- 1TeV (higgsino) LSP DM generic prediction of constrained SUSY models (and also MSSM) – look for it!
- precise determination of BR(B_s to mu mu) can be very helpful in CMSSM (but not beyond)



Lighter superpartners allowed in general MSSM