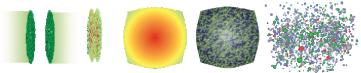
Aleksi Kurkela

Helsinki 02-08, ETH Zurich 08-10, McGill (Montreal) 10-13, CERN 13-



Heavy-ion collisions well described by relativistic hydrodynamics:

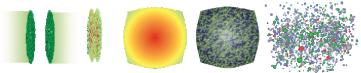
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 $T_{\mu\nu} = T^{\rm eq}_{\mu\nu} + {\rm small \ corrections}$

What are properties of (generic) quantum fields in equilibrium?

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• ... but initial state far from thermal equlibrium

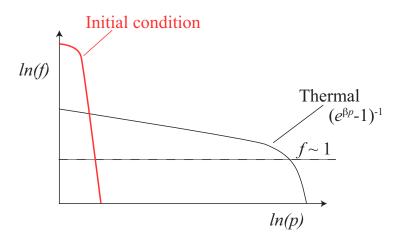
How do (generic) quantum fields relax towards equilibrium?

• Applications elsewhere: cosmological relics, reheating, cold atomic gases, neutron stars...

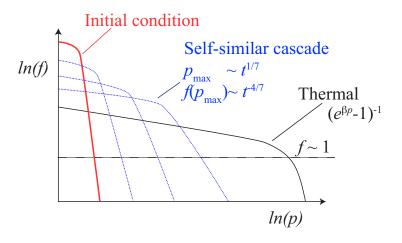
For generic theories, only weak coupling methods available:

- Mostly parametric estimates, not even LO results
- Even at weak coupling often non-perturbative: strong fields, secular divergences, instabilities...
- Weak coupling provides scale separations
- Case-by-case effective theories
 - Effective kinetic theory
 - Classical field theory
 - Hard loop effective theory/ Vlasov equations
 - . . .

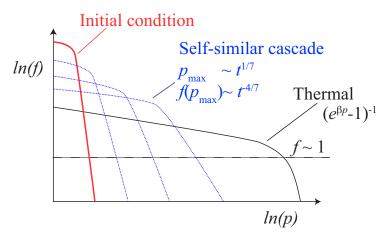
Simple example: what happens if you have too many soft gluons



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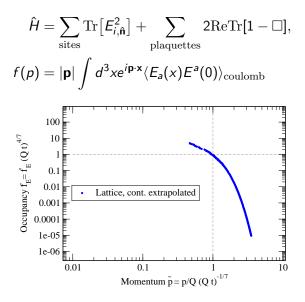


Quantitatively:

Strong fields (f ≫ 1): Classical (lat.) field theory
But not too (f ≪ 1/α): Effective kinetic theory

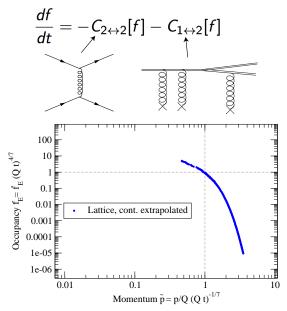
Shape of the self-similar cascade

On lattice, follow the evolution of gauge fields (A_i, E_i) :



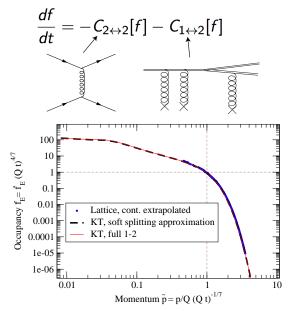
Shape of the self-similar cascade

In kinetic theory, interactions through medium-corrected matrix-elements



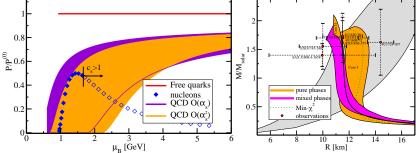
Shape of the self-similar cascade

In kinetic theory, interactions through medium-corrected matrix-elements



Neutron star radii from perturbation theory: NNLO Equation of state for T = 0, $\mu_B \neq 0$, $m_s \neq 0$:





Other stuff:

- Thermal photon production rate to NLO
- High-T eff. theories for thermodynamics:
- Overlap fermions with staggered kernel:
- Extra-dimensions on lattice:

1302.5970 0801.1566

1202.1867

1003.4643