

Density correlations in the Glauber model

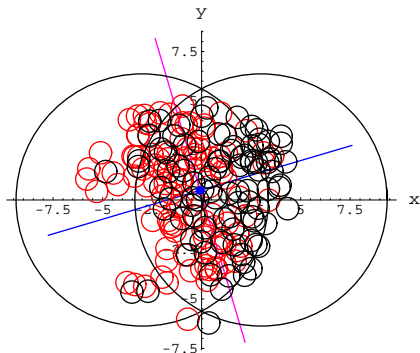
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[based on research with J.-Y. Ollitrault and J.-P. Blaizot]

Fluctuations in the initial state



Goal: understand $S(x, y)$ microscopically

Density-density correlator:

$$S(x, y) = \langle \rho(x)\rho(y) \rangle - \langle \rho(x) \rangle \langle \rho(y) \rangle$$

$\langle . \rangle$ - e-by-e average

[different approach from Coleman-Smith, Petersen, Wolpert 2012 or
Floerchinger-Wiedemann 2013]

Density-density correlator

$S(x, y)$ carries information on e-by-e fluctuations of observables

(2-body effects): $\text{var}(\mathcal{O}) = \int dx dy \mathcal{O}(x) S(x, y) \mathcal{O}(y)$

Embodies **short-range** correlations

- autocorrelations
- NN repulsion in colliding nuclei
- correlation **formed** in the production mechanism

and **long-range** correlations

- conservation laws
- constraints (e.g., choice of centrality class)
- technical issues (recentering)

(for simplicity all for head-on collisions, $b = 0$)

Glauber sources

In each collision n (point-like) sources (wounded nucleon, binary collisions) are created in the transverse plane with a distribution $f_n(x_1, x_2, \dots, x_n)$. Marginal distributions are

$$f_n^{(2)}(x_1, x_2) = \int dx_3 \dots dx_n f_n(x_1, \dots, x_n), \quad f_n^{(1)}(x_1) = \int dx_2 f_n^{(2)}(x_1, x_2)$$

The density is $\rho(x) = \sum_{i=1}^n \delta(x - x_i)$. Then

$$\langle \rho(x) \rangle = \left\langle \int dx_1 \dots dx_n f_n(x_1, \dots, x_n) \sum_i \delta(x - x_i) \right\rangle = \langle n f_n^{(1)}(x) \rangle$$

$$S(x, y) = \langle n f_n^{(1)}(x) \rangle \delta(x - y) + \langle n(n-1) f_n^{(2)}(x, y) \rangle - \langle n f_n^{(1)}(x) \rangle \langle n f_n^{(1)}(y) \rangle$$

Introducing the *pair distribution function*

$$g(x, y) = \frac{\langle n(n-1) f_n^{(2)}(x, y) \rangle}{\langle \rho(x) \rangle \langle \rho(y) \rangle}$$

we may write

$$S(x, y) = \langle \rho(x) \rangle \delta(x - y) + \langle \rho(x) \rangle \langle \rho(y) \rangle [g(x, y) - 1]$$

Some properties

Sum rules:

$$\int dx \rho(x) = n, \quad \int dx dy S(x, y) = \text{var}(n)$$

(sensitivity to constraining n)

The pair distribution function is normalized as

$$\int dx dy \langle \rho(x) \rangle \langle \rho(y) \rangle g(x, y) = \langle n(n-1) \rangle = \text{var}(n) + \langle n \rangle (\langle n \rangle - 1).$$

(increases with the $\text{var}(n)$ as expected)

No correlations, fixed n :

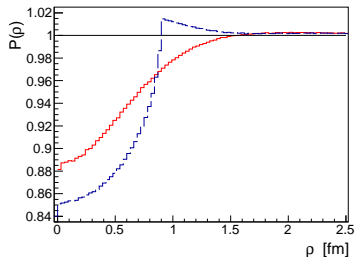
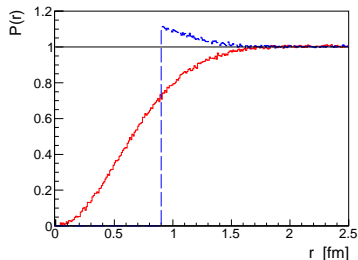
$$g(x, y) = 1 - \frac{1}{n}$$

GLISSANDO 2 is out!

All simulations are carried out with

GLISSANDO 2: GLauber Initial-State Simulation AND mOre..., ver. 2
Maciej Rybczynski, Grzegorz Stefanek, Wojciech Broniowski, Piotr Bozek
e-Print: [arXiv:1310.5475](https://arxiv.org/abs/1310.5475)

NN repulsion



Left: The pair distribution function in the relative distance for the Pb nucleus for the hard-sphere expulsion (dashed line) and Gaussian correlation (solid line)

Right: projection on the transverse plane → “geometric quenching”

These correlations sneak into the fireball! ($\sim 10\%$ effect)

Analytic model for p+Pb

The probability that p incident at an impact parameter b interacts with N participants in a given configuration and does not interact with the remaining $A - N$ nucleons

$$f(s_1, \dots, s_A; b) = c_1 \theta(b - s_1) \dots \theta(b - s_N) (1 - \theta(b - s_{N+1})) \dots (1 - \theta(b - s_A)) T(s_1, \dots, s_A)$$

c_1 - normalization, T - thickness function, $\theta(u)$ - *wounding profile*

$$\left(\int du \theta(u) = \sigma_{\text{inel}}^{NN} \equiv \sigma_w \right)$$

Since d is small, we can include the 2-body correlations perturbatively

$$T(s_1, \dots, s_A) \simeq c_2 T_0(s_1) \dots T_0(s_A) \prod_{\substack{i,j=1 \\ i \neq j}}^A (1 - d(s_i, s_j)) \simeq c_2 T_0(s_1) \dots T_0(s_A) \sum_{i \neq j} (1 - d(s_i, s_j))$$

Then

$$f^{(2)}(s_1, s_2; b) = \frac{\theta(b - s_1) \theta(b - s_2) T_0(s_1) T_0(s_2) (1 - d(s_1, s_2))}{\int ds'_1 ds'_2 \theta(b - s'_1) \theta(b - s'_2) T_0(s'_1) T_0(s'_2) (1 - d(s'_1, s'_2))}$$

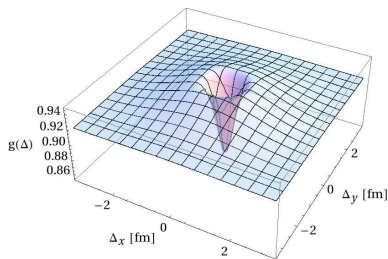
$$f^{(1)}(s_1; b) = \int ds_2 f^{(2)}(s_1, s_2; b)$$

Results of the analytic model for p+Pb

For simplicity, we take $b = 0$ and use the Gaussian parameterizations

$$\theta(u) = A \exp\left(-\frac{u^2}{2\sigma_w^2}\right), \quad A = 0.92, \quad \sigma_w = 1.08 \text{ fm} \quad (\text{LHC})$$

$$d(s_1, s_2) = B \exp\left(-\frac{(s_1 - s_2)^2}{2\sigma_d^2}\right), \quad B = 0.11, \quad \sigma_d = 0.56 \text{ fm}$$



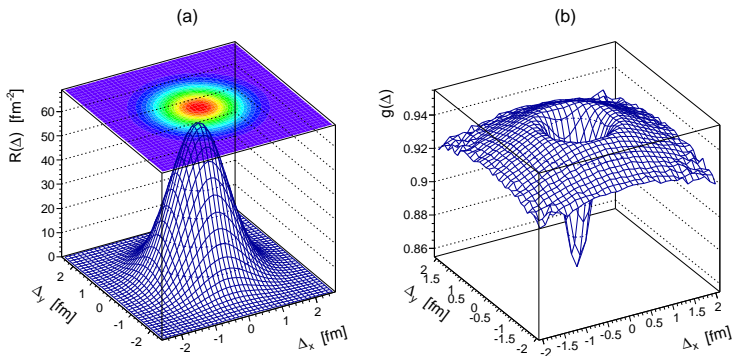
(“Aetna plot”)

Half-integrated distributions
(for visualization in the relative coordinate)

$$R(\Delta) = \int dr \langle \rho(r + \Delta/2) \rangle \langle \rho(r - \Delta/2) \rangle$$

$$g(\Delta) = \frac{1}{R(\Delta)} \int dr \langle n(n-1) f_n^{(2)}(r + \Delta/2, r - \Delta/2) \rangle$$

GLISSANDO for p+Pb



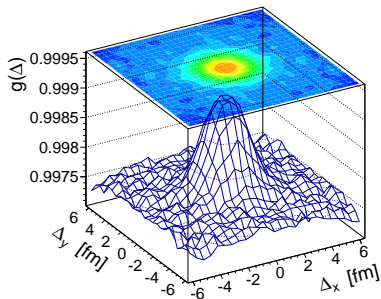
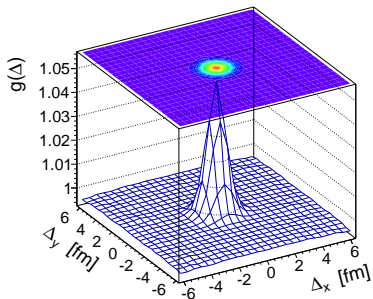
(a) $R(\Delta)$ (b) and $g(\Delta)$ for p+Pb at $b = 0$ and $N = 15$.

(full agreement with the analytic model)

GLISSANDO for Pb+Pb

(for the moment no NN repulsion in the nuclear distributions)

Why do we see the peaks?



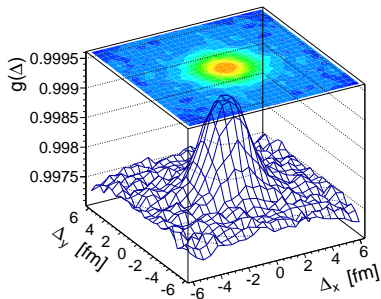
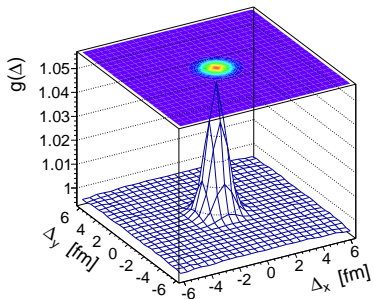
Left: $g(\Delta)$ for the Pb+Pb collisions at the impact parameter $b = 0$ with $\sigma_w = 20$ mb and $N_w = 371$

Right: Same for with $\sigma_w = 68$ mb and $N_w = 410$

GLISSANDO for Pb+Pb

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Peaks from the **twin-production!**

Twin production

Wounded nucleons are created in **partnership**: one nucleon from nucleus A and one from nucleus B. The range of the correlation is $\sim \sqrt{\sigma_w/\pi}$. For very **small** σ_w they come in isolated pairs, hence

$$f_n^{(2)}(x_i, x_j) = \delta(x_i - x_j) f_n^{(1)}(x_i), \quad i \in A, j \in B \quad (\sigma_w \rightarrow 0)$$

There are $N_w/2$ such tightly correlated pairs, while the remaining pairs are uncorrelated: $f_n^{(2)}(x_i, x_j) = f_n^{(1)}(x_i) f_n^{(1)}(x_j)$, $i, j \in$ different nuclei. This leads to

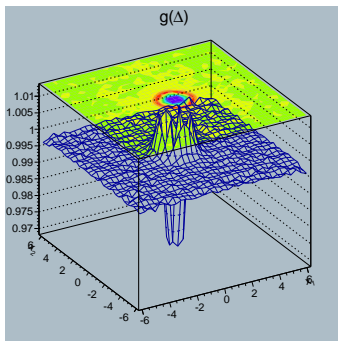
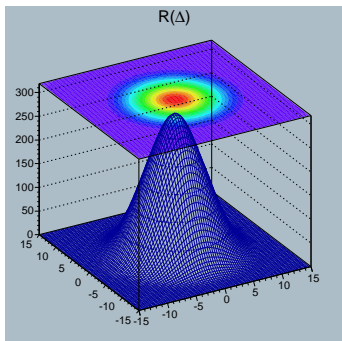
$$\begin{aligned} S(x, y) &= 2\langle n \rangle \left(\langle f^{(1)}(x) \rangle \delta(x - y) - f^{(1)}(x) f^{(1)}(y) \right) + \text{var}(n) f^{(1)}(x) f^{(1)}(y). \\ &= 4 \left[\langle n_{\text{pairs}} \rangle f^{(1)}(x) \delta(x - y) + \langle n_{\text{pair}}(n_{\text{pair}} - 1) \rangle f^{(1)}(x) f^{(1)}(y) \right] \end{aligned}$$

(pairs are the basic objects)

In the limit $\sigma_w \rightarrow \infty$ all nucleons are wounded, hence there is no correlation!

For intermediate values of σ_w clusters form and things are complicated

Same with NN repulsion present



(two effects: positive peak from twin production and negative peak from NN repulsion in projectiles)

Observables

Radiusⁿ:

$$\langle r^n \rangle_{\text{incl}} = \langle \sum_j r_j^n \rangle = \int dx r^n \langle \rho(x) \rangle, \quad \text{var}(r^n)_{\text{incl}} = \int dx dy r^n r'^n S(x, y)$$

r and r' – transverse radii corresponding x and y

Eccentricities:

$$\langle \epsilon_n e^{i\Psi_n} \rangle_{\text{incl}} = \frac{\int d^2x r^n e^{in\phi} \rho(x)}{\langle r^n \rangle_{\text{incl}}}, \quad \text{var}(|\epsilon_n|)_{\text{incl}} = \frac{\int dx dy r^n e^{in\phi} r'^n e^{-in\phi'} S(x, y)}{\langle r^n \rangle_{\text{incl}}^2}$$

Ψ_n – event-plane angles, ϕ and ϕ' – azimuths corresponding to x and y

No correlations, $b = 0$:

$$\begin{aligned} \langle r^n \rangle_{\text{incl}}^{\text{no corr.}} &= \langle n \rangle \langle r^n \rangle \\ \text{var}(r^n)_{\text{incl}}^{\text{no corr.}} &= \langle n \rangle \langle r^{2n} \rangle + (\text{var}(n) - \langle n \rangle) \langle r^n \rangle^2 \\ \langle \epsilon_n e^{i\Psi_n} \rangle_{\text{incl}}^{\text{no corr., central}} &= 0 \quad (\text{symmetry}) \\ \text{var}(|\epsilon_n|)_{\text{incl}}^{\text{no corr., central}} &= \frac{\langle r^{2n} \rangle}{\langle n \rangle \langle r^n \rangle^2} \end{aligned}$$

To focus on the effects of correlations in the fluctuation measures, we introduce

$$R(O) = \frac{\text{var}(O)_{\text{incl}}}{\text{var}(O)_{\text{incl}}^{\text{no corr.}}}$$

Observables in p+Pb

Analytic model:

($B = 0.11$ – depth of the “soft-core”, $\sigma_d = 0.56$ fm – its width, $\sigma_w = 1.08$ fm)

$$R(r^2) = 1 - 2B \frac{\omega(n) + \langle n \rangle}{\omega(n) + 1} \frac{\sigma_d^2 \sigma_w^4}{(\sigma_d^2 + 2\sigma_w^2)^3} + \mathcal{O}(B^2) = 1 - 0.69B + \mathcal{O}(B^2)$$

$$R(\epsilon_2) = 1 - B(\omega(n) + \langle n \rangle) \frac{\sigma_d^2 \sigma_w^4}{(\sigma_d^2 + 2\sigma_w^2)^3} + \mathcal{O}(B^2) = 1 - 0.35B + \mathcal{O}(B^2)$$

$$R(\epsilon_3) = 1 - 0.09B + \mathcal{O}(B^2)$$

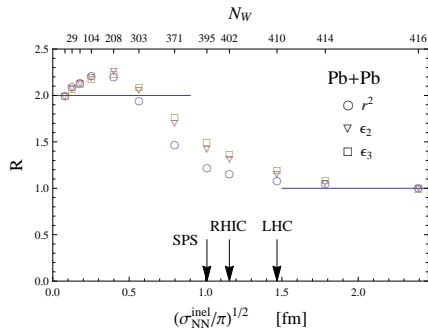
Small but non-negligible effect (a few %)

Maximum when $\sigma_w = \sigma_d$ – matching of the two scales: probing and internal

Observables in Pb+Pb

($R = \text{"correlated/uncorrelated"}$)

GLISSANDO:

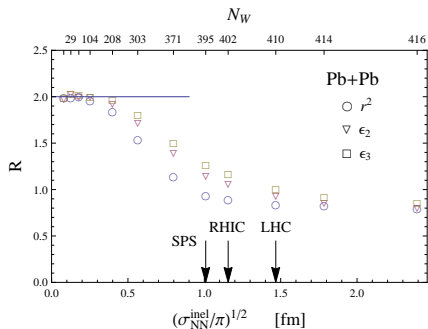
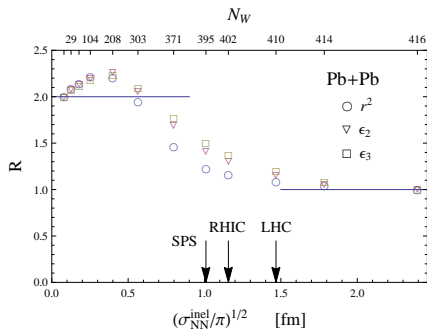


The ratios R at various values of σ_w . The corresponding fixed values of the number of the wounded nucleons is shown on the upper axis

Observables in Pb+Pb

($R = \text{“correlated/uncorrelated”}$)

GLISSANDO:



Left: The ratios R at various values of σ_w . The corresponding fixed values of the number of the wounded nucleons is shown on the upper axis

Right: Same with the **NN repulsion included**

Superposition model

$$\rho(x) = \sum_{i=1}^n w_i \delta(x - x_i) \quad \text{-- varying strength}$$

w_i – weights generated independently according to some suitable distribution. Then

$$(\omega(w) = \text{var}(w)/\bar{w})$$

$$\langle \rho(x) \rangle = \bar{w} \langle n f_n^{(1)}(x) \rangle$$

$$S(x, y) = (\omega(w) + \bar{w}) \langle \rho(x) \rangle \delta(x - y) + \langle \rho(x) \rangle \langle \rho(y) \rangle [g(x, y) - 1]$$

$$g(x, y) = \frac{\bar{w}^2 \langle n(n-1) f_n^{(2)}(x, y) \rangle}{\langle \rho(x) \rangle \langle \rho(y) \rangle}$$

Upon integration over x and y we find the superposition-model formula

$$\text{var}(\rho) = \text{var}(w) \langle n \rangle + \bar{w}^2 \text{var}(n)$$

$\omega(w)/\bar{w}$ enhances the relative contribution of the autocorrelation term

Conclusions

- Microscopic understanding achieved
- Smearing: $\delta(x - y) \rightarrow d(x - y)$
- Short-range correlations assume the form $S_{\text{short}}(x, y) = A(x)d(x - y)$
(short range dominance)
- Autocorrelations + NN repulsion in projectiles + twin production
- Long range term $\sim f^{(1)}(x)f^{(1)}(y)$ from constraining n (in general, global constraints lead to long-range correlations)
- Correlations affect fluctuation of observables (NN repulsion already studied in [WB & M. Rybczyński, PRC 81 (2010) 064909])
- Fluctuations of eccentricities dominated by $S_{\text{short}}(x, y)$