

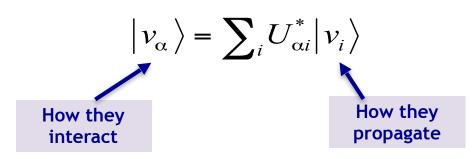
Latest Results from the Daya Bay Reactor Neutrino Experiment

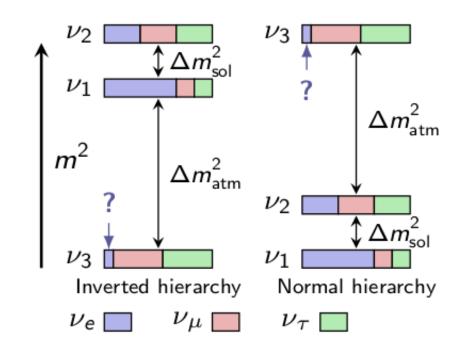




Three-Neutrino Framework: Current Status

Neutrino weak eigenstates are mixtures of mass eigenstates:





θ_{13} only recently well established by Daya Bay

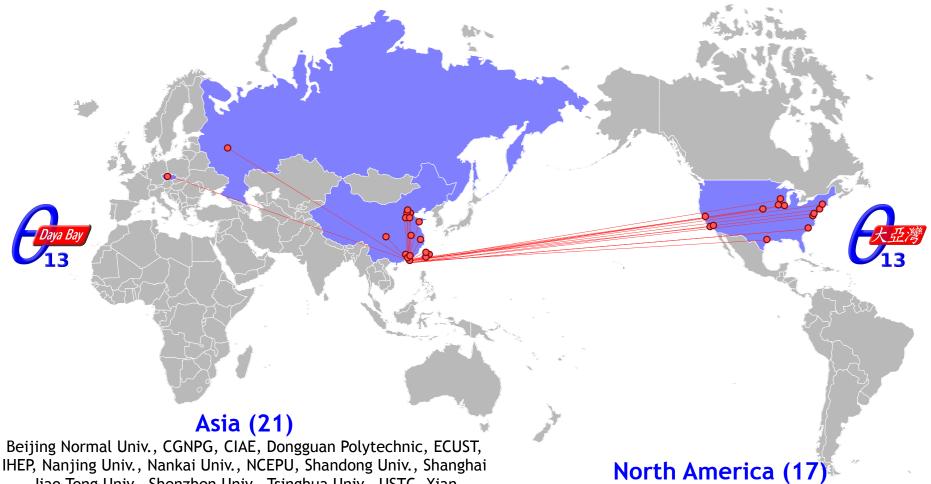
$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

θ₂₃ ~ 45° established through atmospheric and accelerator experiments: possibly maximal Gateway to CP violation and mass hierarchy

 θ_{12} ~ 34° established through solar experiments and KamLAND: large but not maximal

The Daya Bay Reactor Neutrino Experiment

The Daya Bay Collaboration:



Jiao Tong Univ., Shenzhen Univ., Tsinghua Univ., USTC, Xian Jiaotong Univ., Zhongshan Univ.,

Chinese Univ. of Hong Kong, Univ. of Hong Kong, National Chiao Tung Univ., National Taiwan Univ., National United Univ.

Europe (2)

Charles University, JINR Dubna

Brookhaven Natl Lab, CalTech, Illinois Institute of Technology, Iowa State, Lawrence Berkeley Natl Lab, Princeton, Rensselaer Polytechnic, Siena College, UC Berkeley, UCLA, Univ. of Cincinnati, Univ. of Houston,

UIUC, Univ. of Wisconsin, Virginia Tech, William & Mary, Yale

~230 Collaborators

Daya Bay Experimental Layout

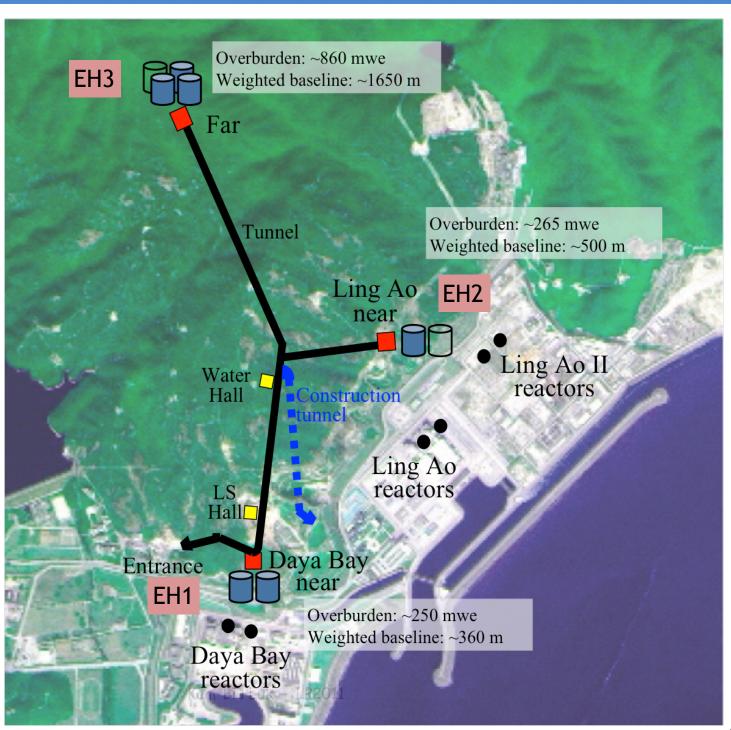
Electron antineutrinos are produced in copious amounts in nuclear reactors.

We position 8 detectors around the Daya Bay Power Plant in China, among the most powerful in the world.

Main principle:

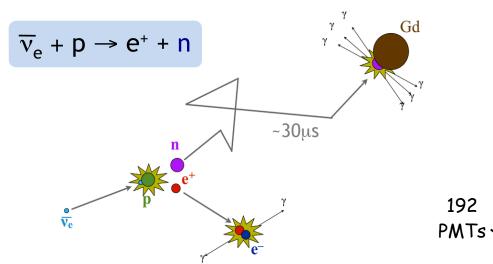
- (i) sample the reactor anti-neutrino flux in the near and far locations, and
- (ii) look for evidence of disappearance

Note: results shown here use data collected with 6 / 8 detectors



The Detectors

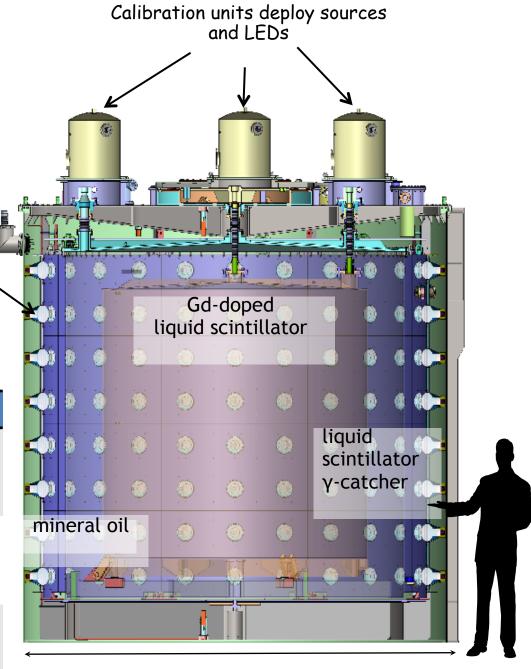
Anti-neutrinos are detected via the inverse beta decay (IBD) reaction:



❖ The detectors are ~100ton three-zone cylindrical modules:

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Zone	Mass	Liquid	Purpose
Inner acrylic vessel	20 t	Gd-doped liquid scintillator	Anti-neutrino target
Outer acrylic vessel	20 t	Liquid scintillator	Gamma catcher (from target zone)
Stainless steel vessel	40 t	Mineral Oil	Radiation shielding

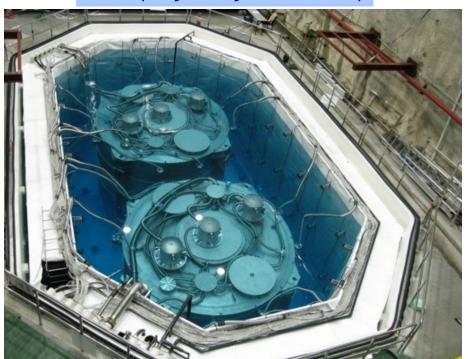


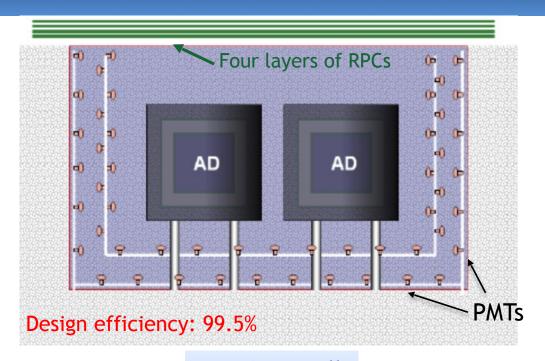
5 m

Muon Veto System

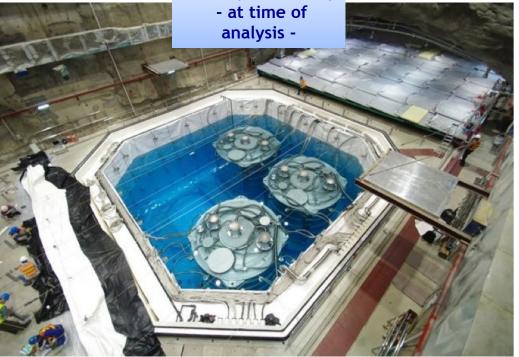
- The detectors are immersed in an instrumented water pool:
 - Double purpose:
 - ✓ Shields against gammas from ambient radioactivity and neutrons produced by cosmic rays
 - ✓ Serves as a Cerenkov detector to tag cosmic ray muons (thus reducing backgrounds)

EH1 (Daya Bay Near Hall)









Analyzed Datasets

Two detector comparison [1202.6181]

- 90 days of data, Daya Bay near only
- NIM A 685 (2012), 78-97

First oscillation analysis [1203:1669]

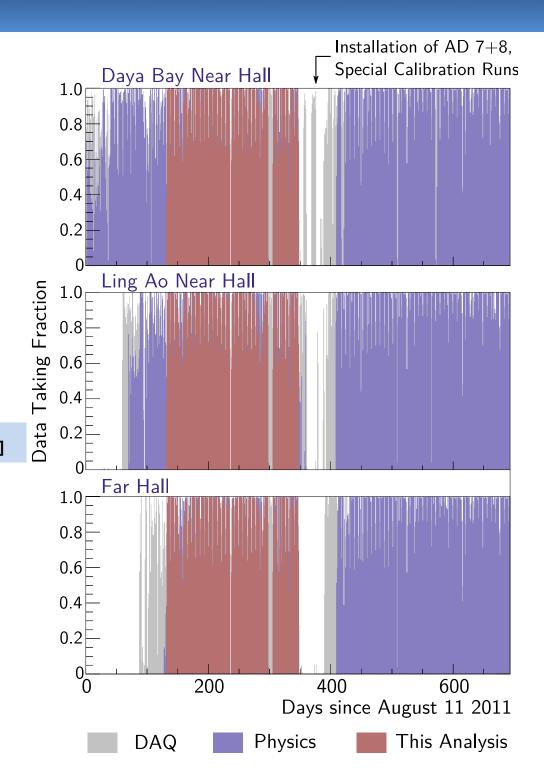
- 55 days of data, 6 ADs near+far
- PRL **108** (2012), 171803
- Top 10 breakthrough of 2012 by Science Magazine

Improved oscillation analysis [1210.6327]

- 139 days of data, 6 ADs near+far
- CP C 37 (2013), 011001

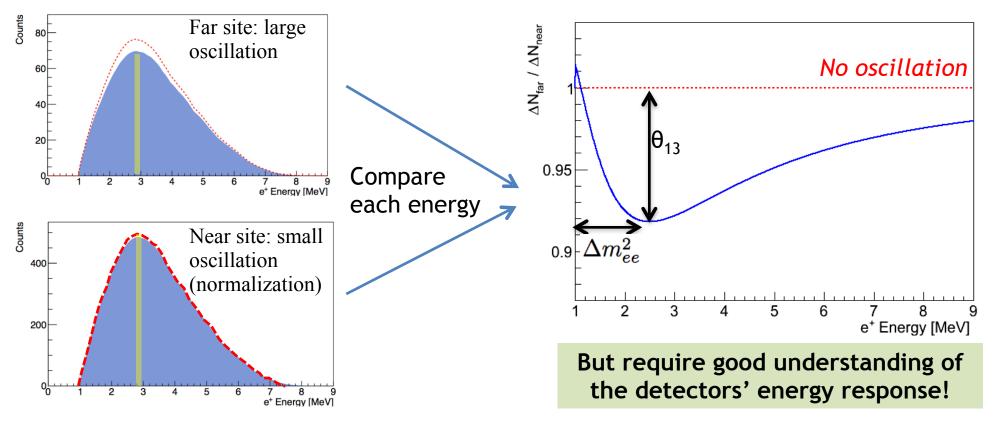
Spectral Analysis [1310.6732]

- 217 days complete 6 AD period
- 55% more statistics than CPC result



Doing a Spectral Measurement

With a spectral measurement can measure the mass splitting:



lacktriangle Which mass splitting do we measure? Define an effective mass splitting Δm_{ee}^2 :

$$P_{\bar{\nu_e} \rightarrow \bar{\nu_e}} = 1 - \sin^2 2\theta_{13} \sin^2 \left(\Delta m_{ee}^2 \frac{L}{4E} \right) - \sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \left(\Delta m_{21}^2 \frac{L}{4E} \right)$$

$$\sin^2 (\Delta m_{ee}^2 \frac{L}{4E}) \equiv \cos^2 \theta_{12} \sin^2 (\Delta m_{31}^2 \frac{L}{4E})$$

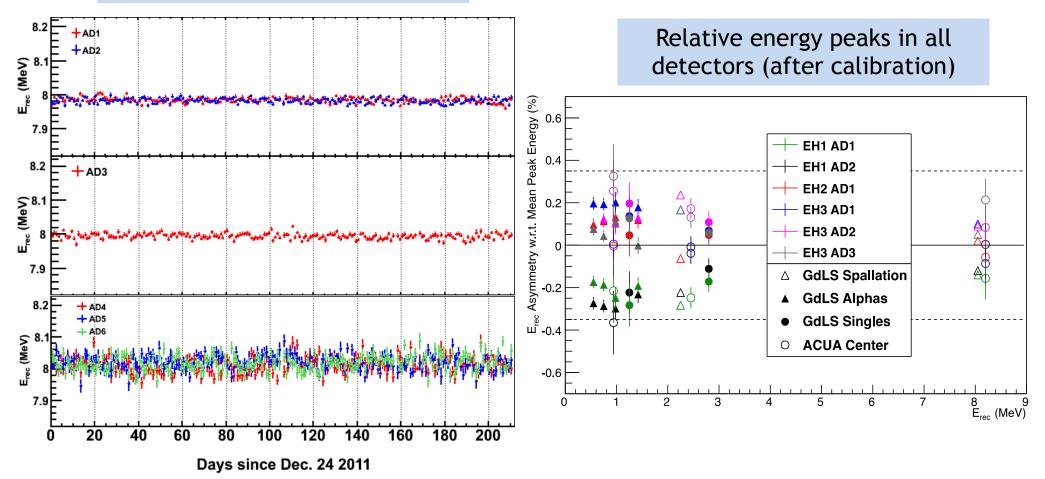
$$+ \sin^2 \theta_{12} \sin^2 (\Delta m_{32}^2 \frac{L}{4E})$$
 so that:
$$\left| \Delta m_{ee}^2 \right| \simeq \left| \Delta m_{32}^2 \right| \pm 5.21 \times 10^{-5} \text{eV}^2 \quad \begin{array}{l} +: \text{ Normal Hierarchy} \\ -: \text{ Inverted Hierarchy} \end{array}$$

Ingredient #1: Calibration

One key is achieving a <u>stable and consistent energy response</u> between detectors:

After calibration, achieve energy response that is **stable to ~0.1%** in all detectors, with a **total relative uncertainty of 0.35%** between detectors.

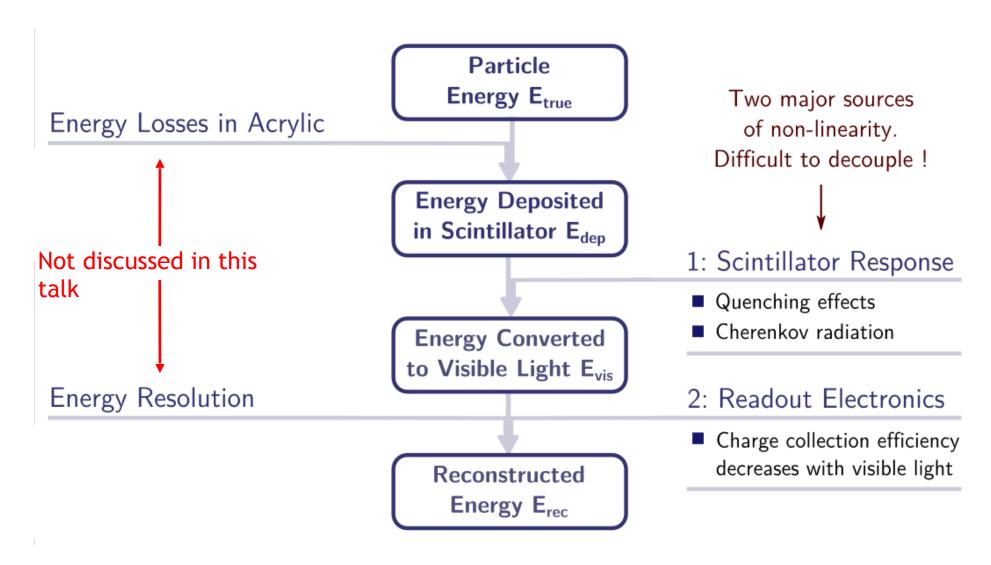
Spallation *n*Gd capture peak vs. time (after calibration)



After initial reconstruction, position non-uniformity is also corrected for

Ingredient #2: Energy Response Model

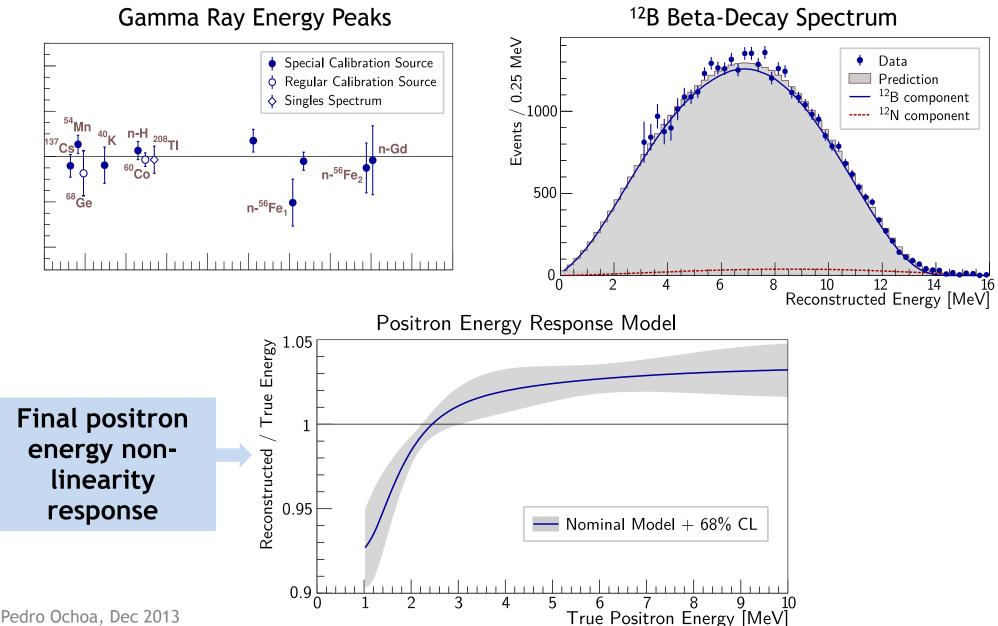
 \diamond Also need to relate reconstructed kinetic energy E_{rec} to true energy E_{true} :



- ✓ Minimal impact on oscillation measurement
- ✓ Crucial for measurement of reactor spectra (in progress)

Ingredient #2: Non-Linearity Response Model

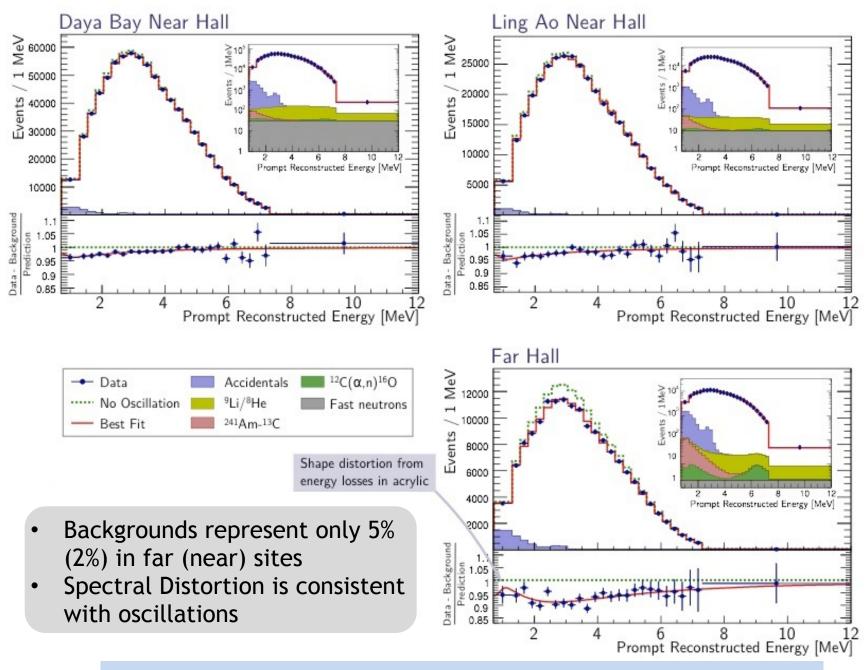
Model is constrained using monoenergetic gamma lines from various sources and continuous spectrum from ¹²B produced by muon spallation inside the scintillator:



J. Pedro Ochoa, Dec 2013

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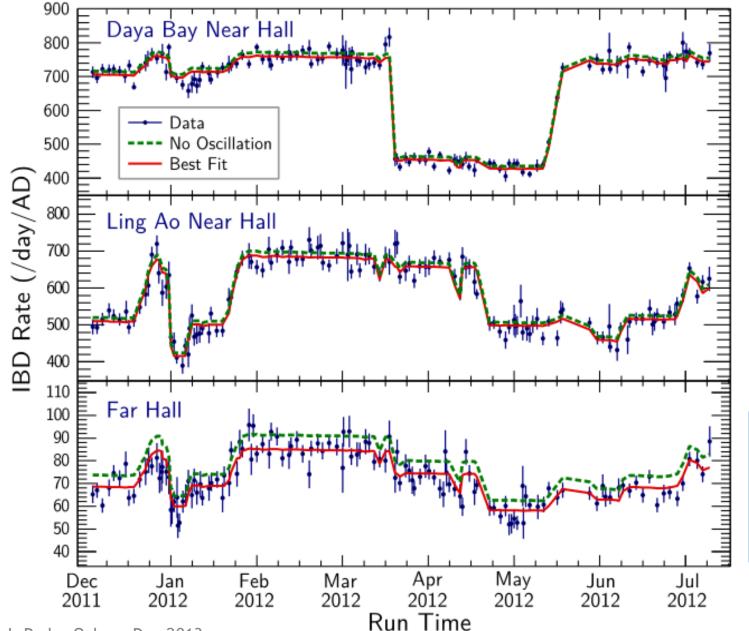
Dataset for Oscillation Analysis



This analysis uses more than 300k antineutrino interactions

Antineutrino Rates vs. Time

For main analysis we simultaneously fit all detectors <u>using reactor model</u>, with the absolute normalization as a free parameter:

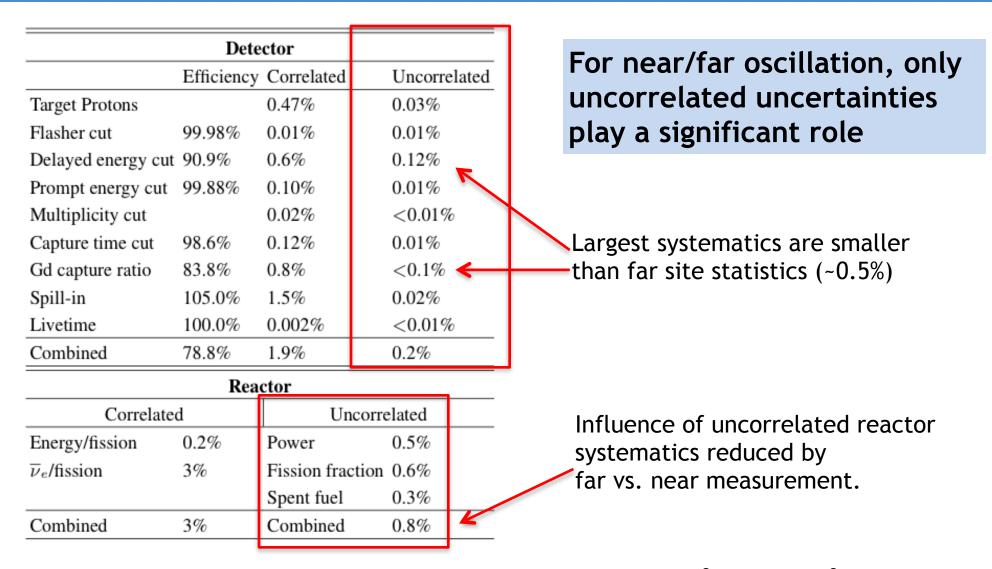


Note:

- Normalization is determined by fit to data. It is within a few percent of expectations.
- Paper on absolute reactor neutrino flux and shape is in preparation

Detected rate strongly correlated with reactor flux expectations

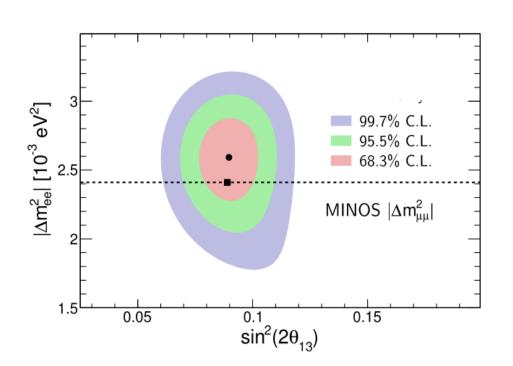
Systematic Uncertainties

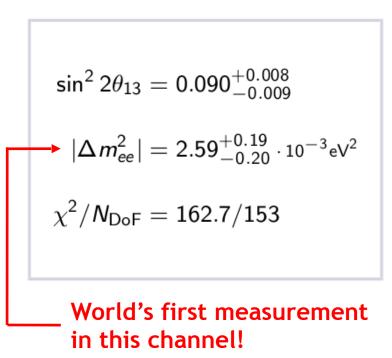


- Statistics contribute 73% (65%) to total uncertainty in $\sin^2 2\theta_{13}$ ($|\Delta m^2_{ee}|$)
- Major systematics:
 - θ_{13} : Reactor model, relative + absolute energy, and relative efficiencies
 - $|\Delta m^2_{ee}|$: Relative energy model, relative efficiencies, and backgrounds

Results

Rate + shape results are consistent with previous results:





Strong confirmation of oscillation-interpretation of observed $\overline{v}_{\!_{e}}$ deficit

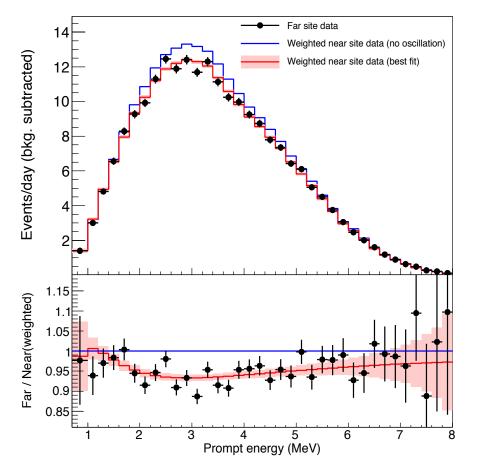
	Normal MH Δm_{32}^2 [10 ⁻³ eV ²]	Inverted MH Δm_{32}^2 [10 ⁻³ eV ²]	
From Daya Bay Δm_{ee}^2	$2.54_{-0.20}^{+0.19}$	$-2.64^{+0.19}_{-0.20}$	-
From MINOS $\Delta m_{\mu\mu}^2$	$2.37^{+0.09}_{-0.09}$	$-2.41^{+0.11}_{-0.09}$	A. Radovic, DPF2013

Independent Cross-Check

Independent crosscheck with minimal reactor assumptions

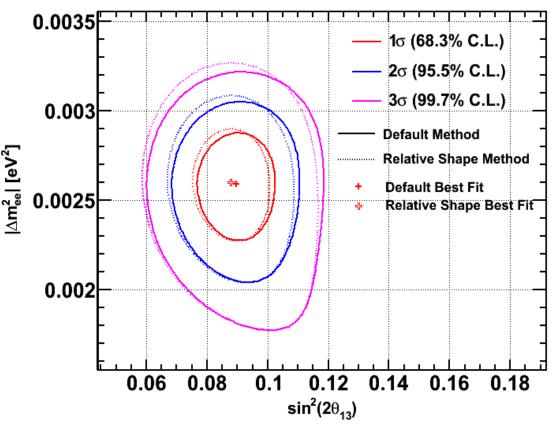
Predict far spectra directly from measured near site spectra

→ Minimizes impact of absolute flux and spectra prediction.

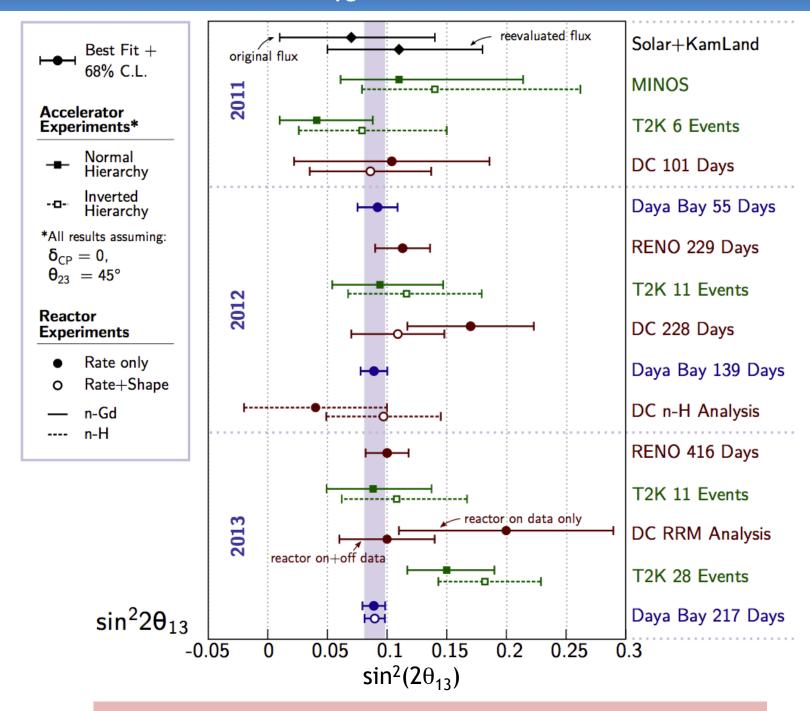


Use covariance matrices to account for systematic errors

→ Alternate method finds consistent uncertainties for neutrino parameters.



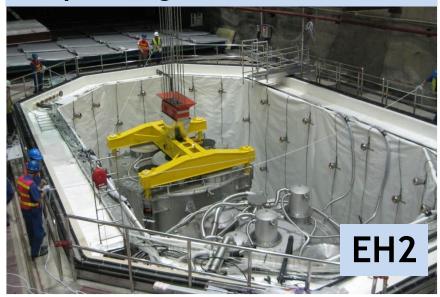
Global Landscape of θ_{13} Measurements



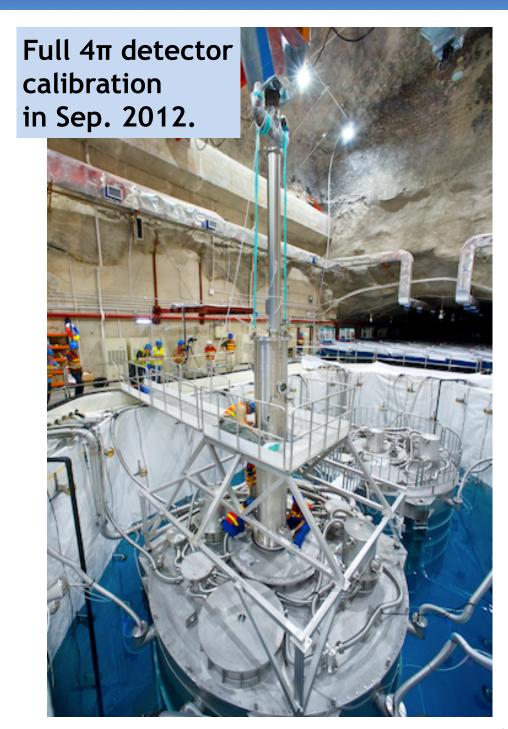
World's most precise measurement of θ_{13} to date.

Daya Bay Onsite Progress

Final two detectors installed, operating since Oct. 2012.

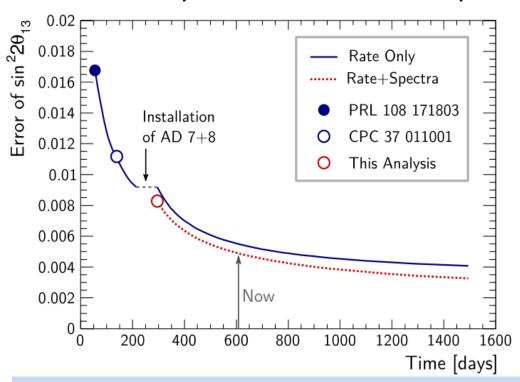


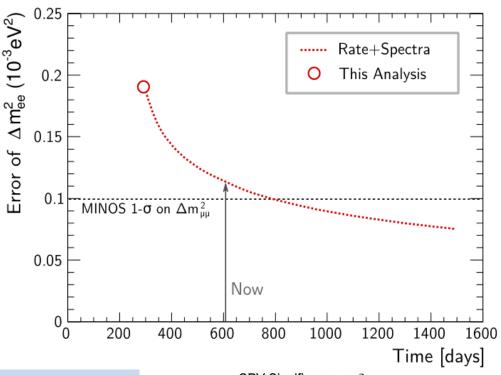




Daya Bay's Future

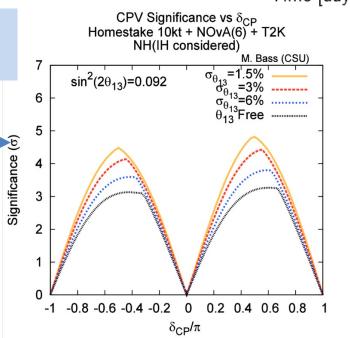
Increased precision in oscillation parameters:





World's most precise measurement of θ_{13} for a long time to come, and very precise estimate of Δm^2

- ✓ Constrains non-standard oscillation models
- ✓ Improves reach of next-generation experiments -
- Absolute reactor neutrino spectrum flux and shape measurement:
 - ✓ Probe reactor models and explore reactor antineutrino 'anomaly'
- Others (cosmogenic production, supernovae... etc)



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Summary & Conclusions

First direct measurement of the short-distance electron antineutrino oscillation frequency:

$$|\Delta m_{ee}^2| = 2.59_{-0.20}^{+0.19} \times 10^{-3} eV^2$$

• Most precise estimate of the θ_{13} mixing angle:

$$\sin^2(2\theta_{13}) = 0.090^{+0.008}_{-0.009}$$

Stay tuned for more exciting results from Daya Bay!







Thank you for your attention!