



**High
Luminosity
LHC**

Tolerances on powering, alignment and ground motion

M. Fitterer

Acknowledgments: R. De Maria, M. Giovannozzi, S. Fartoukh,
J. Pfingstner

Outline

- Ground motion and alignment tolerances for the IT/D1 area:
 - short term (<10 s) effects on ground motion
 - BPM tolerances for orbit correction
 - review of corrector strength for (long term) misalignment correction
- Tolerances on powering of the IT

Ground motion and alignment tolerances

Short term effect of ground motion (1)

Ground motion model used for simulations:

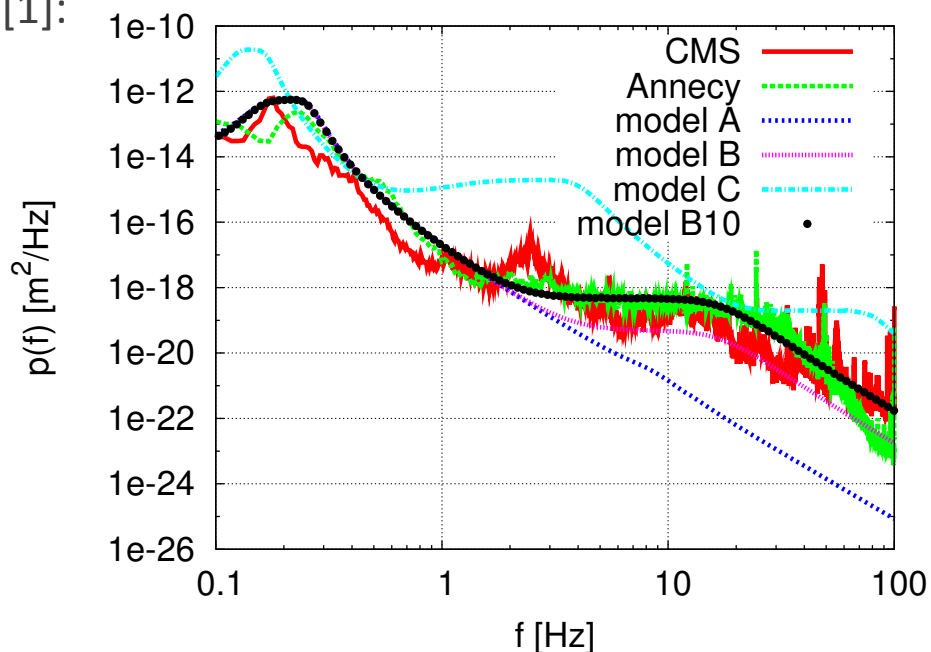
Real spectrum for short time scales (<1min) [1]:

CMS: measurement [1]

Annecy: measurement [2]

model **A/B/C**: ground motion models [3]

model **B10** (black): model B with 10x more noise for the range >2 Hz (D. Schulte)



ATL law for long time scales (>1min) [4]:

$$\langle dY^2 \rangle = ATL, A = 10^{-5 \pm 1} \mu\text{m}^2 / (\text{sm})$$

with T=time interval between measurements, L=distance between measurement points

[1] A. Kuzmin, Technical Report EDMS Nr. 1027459, CERN, 2009

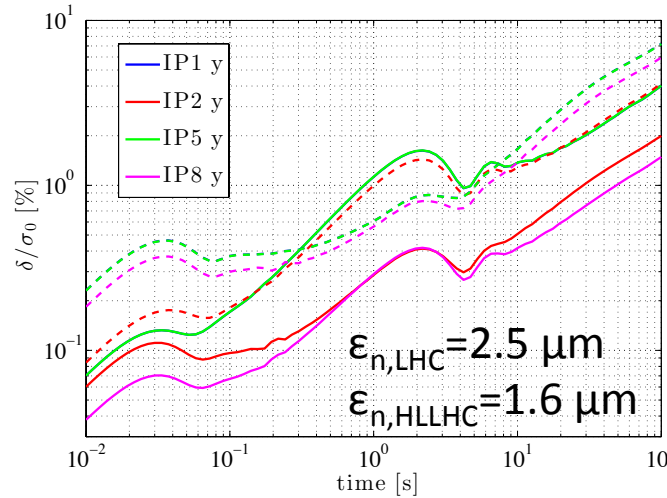
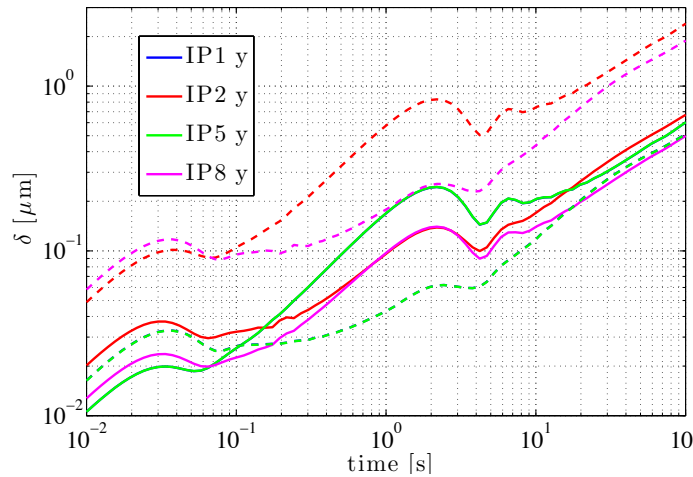
[2] B. Bolzon, PhD Thesis, Université de Savoie, 2007

[3] International Linear Collider Technical Review Committee: Second Report, SLAC Report-606 (2003)

[4] V. Shiltsev, PRSTAB 13, 094801 (2010)

Short term effect of ground motion (2)

Simulation results: — nom. LHC, 4 TeV, $\beta^*=60$ cm — HLLHCV1.0, 7 TeV, $\beta^*=15$ cm



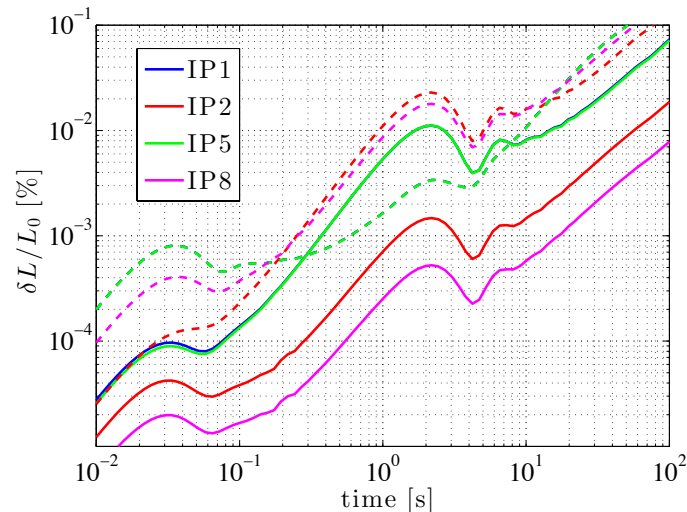
offset at IP

$$\sim (k \cdot l)$$

and

$$(k \cdot l)_{\text{nom.}} \gtrsim (k \cdot l)_{\text{HLLHC}}$$

➔ similar effects for nominal LHC than HL-LHC



luminosity loss

$$\sim \exp \left[-\frac{(\delta x/2)^2}{\sigma_x^{*2} \cos^2 \phi + \sigma_z^2 \sin^2 \phi} - \left(\frac{\delta y}{2\sigma_y^*} \right)^2 \right]$$

and

$$\beta_{\text{nom.}}^* > \beta_{\text{HLLHC}}^*$$

➔ larger effect for HL-LHC than nominal LHC

➔ effect negligible ($\sim 10^{-2}$) for time range < 10 s,

afterwards assumed correction by the orbit feedback

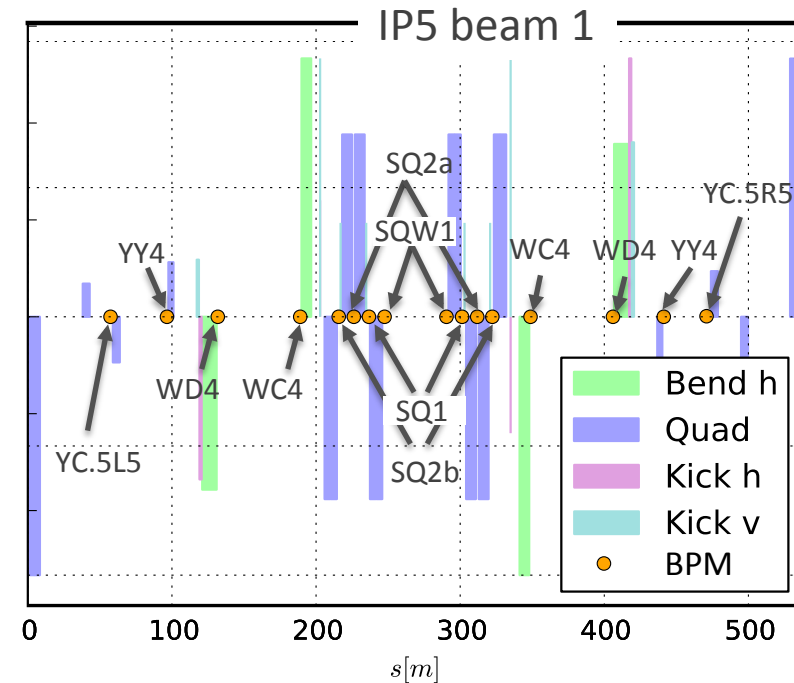
J. Pfingstner

BPM tolerances for orbit correction (1)

BPMs in IT/D1 area used for x-scheme/alignment of the beam

First estimate of required BPM resolution and number/placement by using an **idealized system** (results **scale linearly** with BPM resolution, assumed reference value is $\pm 0.5 \mu\text{m}$)

- no field or misalignment errors
- treat interaction region as line (from BPMYC.5L5 to BPMYC.5R5)
- assume that the orbit is corrected up to BPM resolution (here $\pm 0.5 \mu\text{m}$) at BPMYC.5L5 [*] (=initial condition) and BPMYC.5R5 (=matching constraint end of line)
- use perturbed BPM readings to match x-scheme and record resulting orbit



[*] assume thus $z = \pm 0.5 \mu\text{m} \Rightarrow J_z = z^2 / \beta \Rightarrow z = (\beta J_z)^{1/2} \cos(\phi_z)$, $z' = (J_z / \beta) (\sin(\phi_z) + \alpha \cos(\phi_z))$, with $J_z \in [0, J_z]$ and $\phi_z \in [0, 2\pi]$

BPM tolerances for orbit correction (2)

Observables: z, z' at IP and z at crab cavities

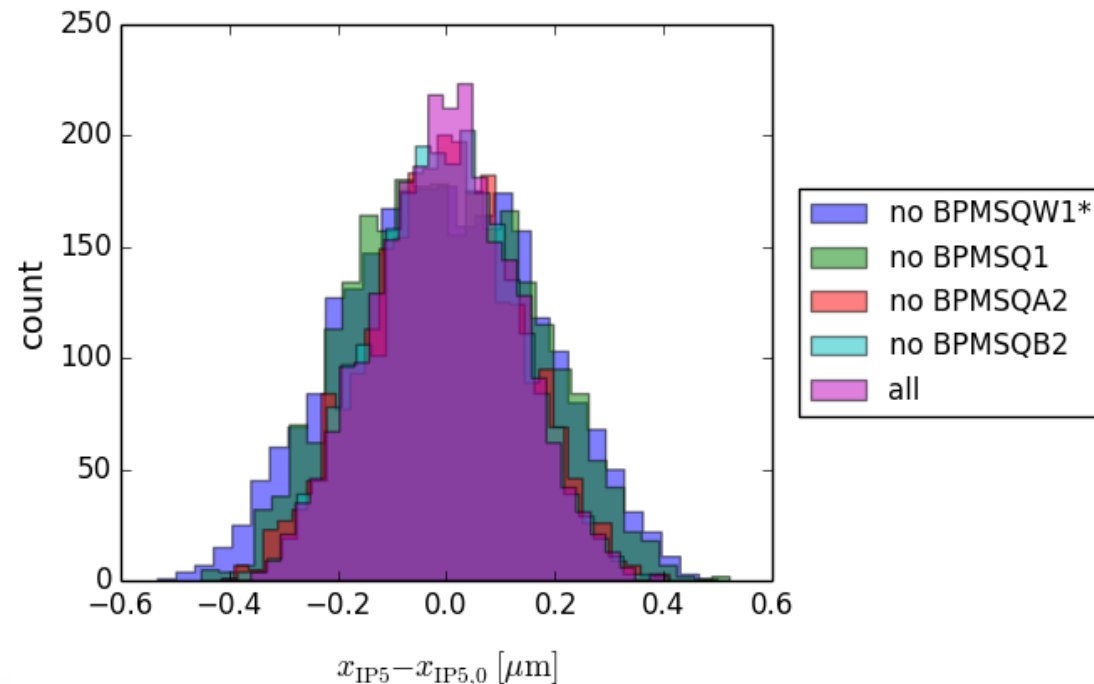
Comments/Studies:

a) **linear system** $\Rightarrow z/z'$ at IP/crab cavities scales linearly with BPM precision

b) use all BPMs except one to match orbit

\rightarrow BPMs closest to IP most effective to constraint orbit at IP

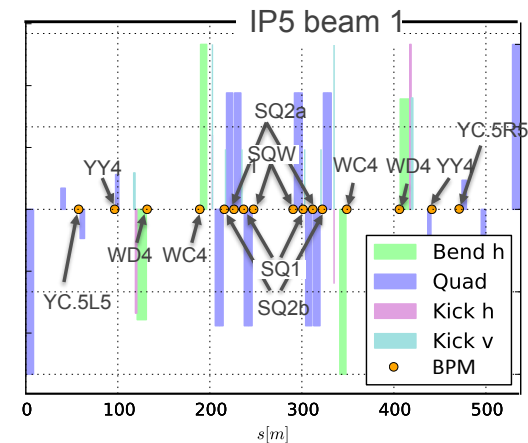
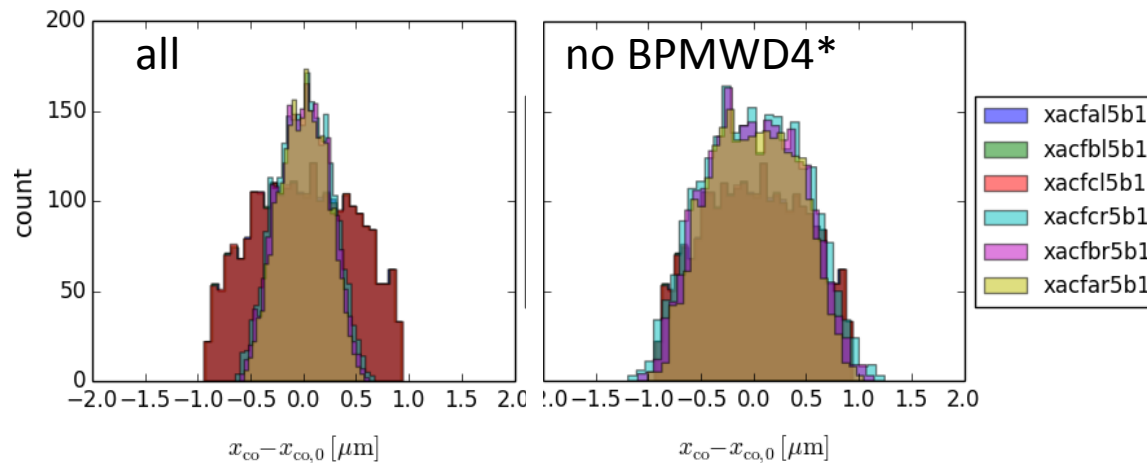
	$\text{rms}(x_{\text{IP5}} - x_{\text{IP5,0}})$
no BPMSQW1	0.180 μm
no BPMSQ1	0.165 μm
no BPMSQA2	0.138 μm
no BPMSQB2	0.131 μm
all	0.128 μm



BPM tolerances for orbit correction (3)

b) use all BPMs except one to match orbit (continued)

➔ BPMs closest to crab cavities most effective to constraint orbit at location of crab cavities



c) use selection of BPMs

➔ same performance using only **3x4 BPMs around IP** (BPMSQW1*, BPMSQ1, BPMSQA2) and **2x4 around crab cavities** (BPMWD4*, BPMYY4), explicitly **no BPMSQB2* (LR bb encounter)** and BPMWC4*

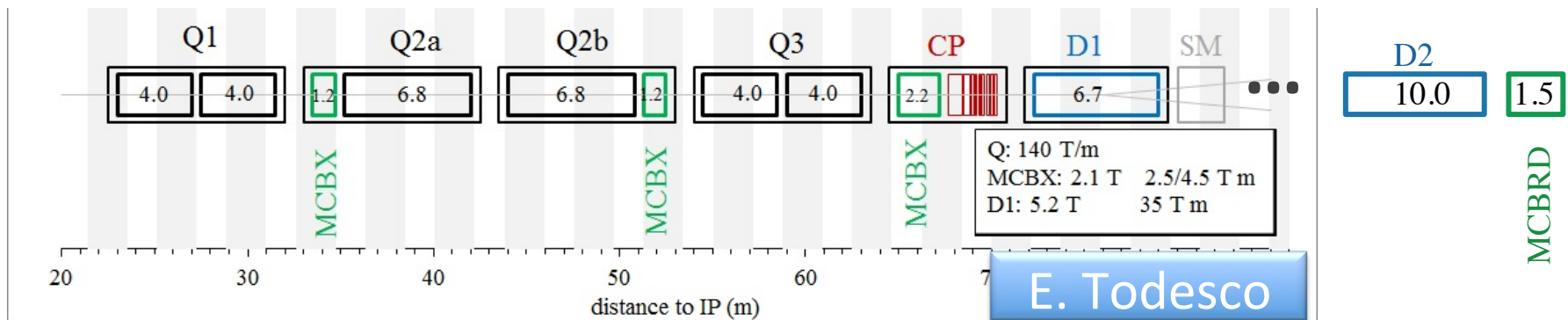
➔ x2 worse orbit at position of crab cavities using only BPMSQW1*, BPMSQ1* and BPMWD4*, BPMYY4* (no BPMSQA2*)

Misalignment correction (1)

- Observables:**
- corrector strength
 - maximum orbit in triplet (aperture)

Relative transverse movement of the triplet per year: +/-0.5 mm

→ misalignment of the triplet by +/-0.5 mm (uniform distribution) and correction with the MCBX* (2.5/4.5 Tm) and MCBRD* (8 Tm) correctors



Correction Schemes:

- MCBX* and two MCBRD
- all correctors, but limit strength in MCBRDs

For x-scheme:

crossing plane:

MCBX1: 0.4 Tm, MCBX2: 0 Tm

MCBX3: 2.1 Tm, MCBRD: 4.5 Tm

separation plane:

MCBX1: 0.15 Tm, MCBX2: 0 Tm

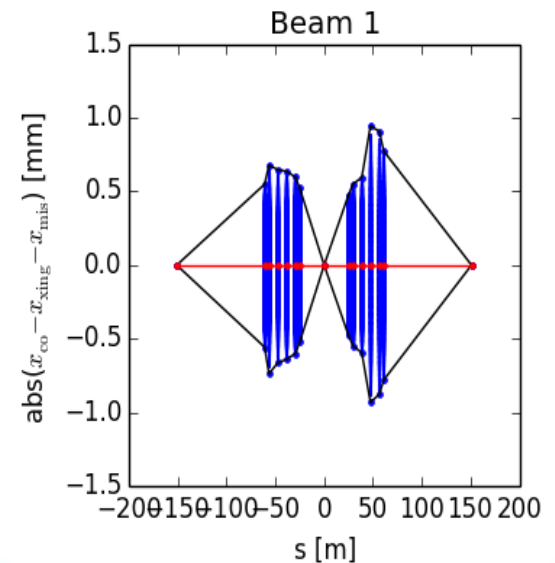
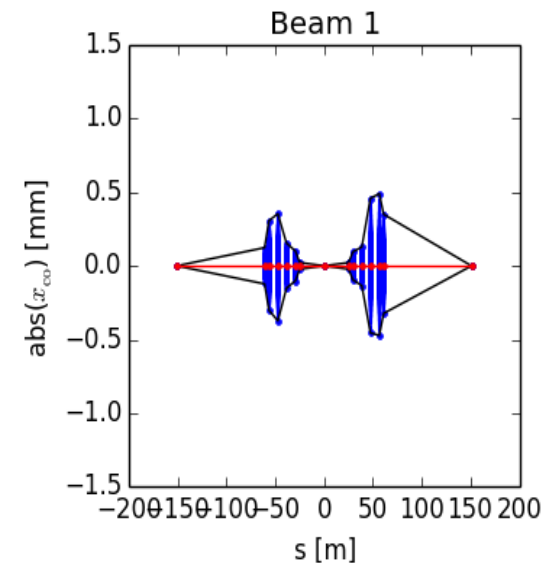
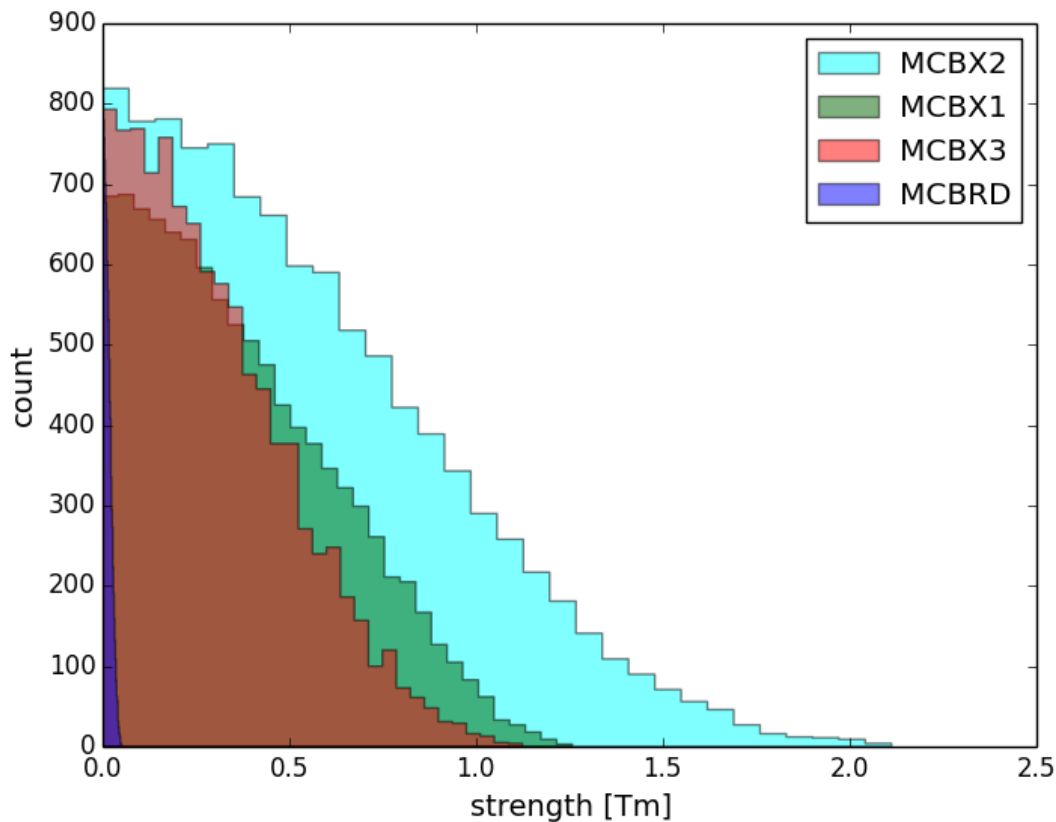
MCBX3: 0.24 Tm, MCBRD: 0.14 Tm

Misalignment correction (2)

a) MCBX* and two MCBRDs

➔ **small orbit deviation** in triplet (<1.0 mm)

➔ corrector strength within limits

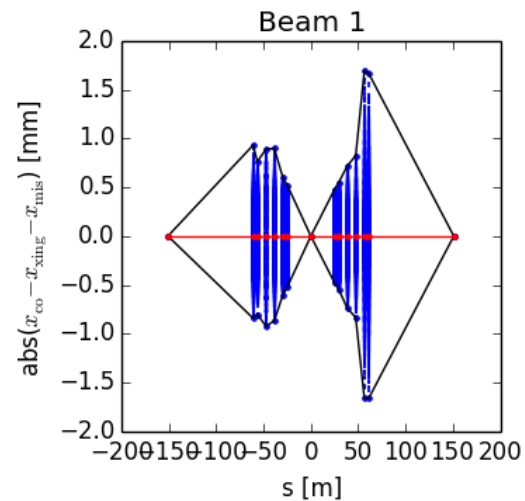
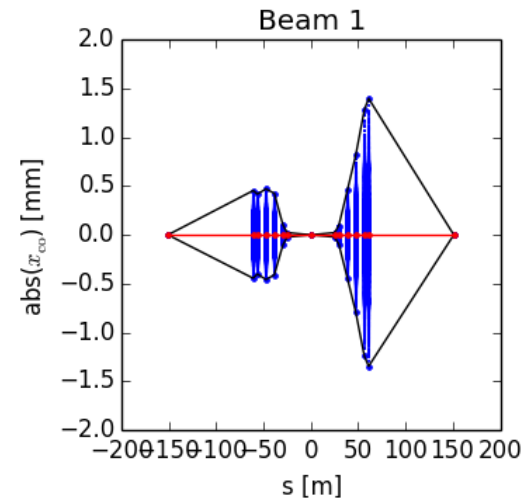
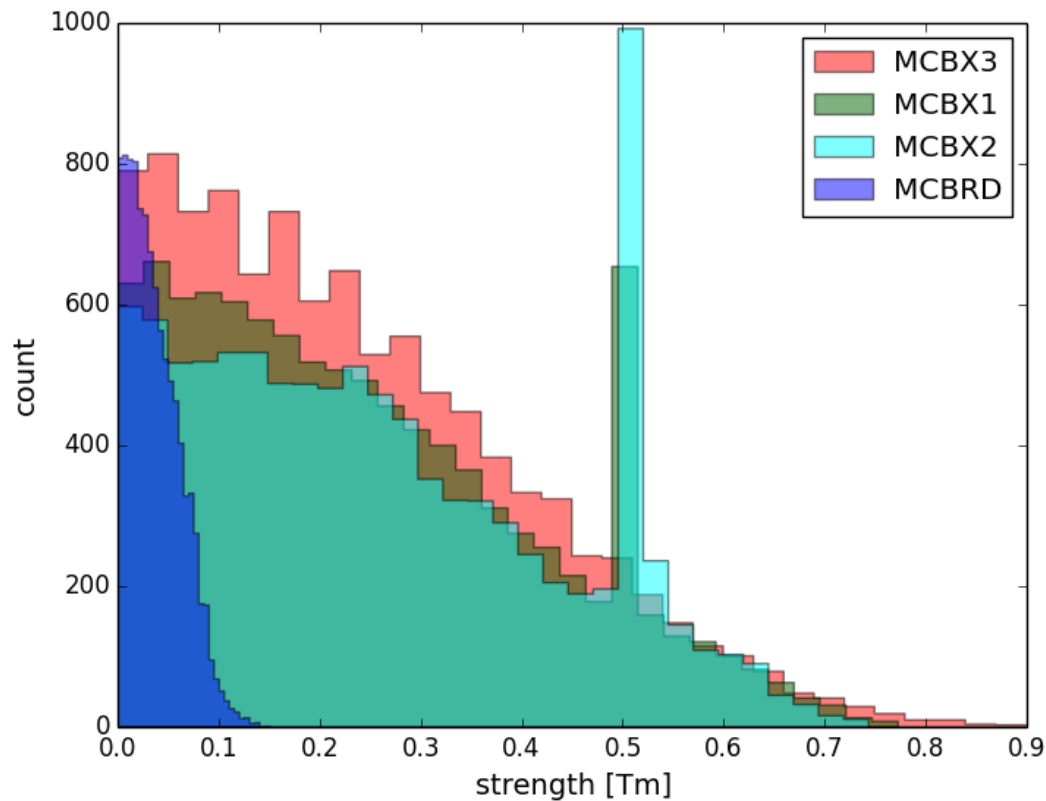


Misalignment correction (3)

a) all correctors, but limit strength in MCBRD (in order to limit orbit in triplet)

➔ larger orbit deviation in triplet (<1.5 mm) than for case a)

➔ by x2 smaller corrector strength than for case a)



Tolerances on powering of the IT

Tolerances on powering of the IT (1)

Summary of results of previous studies:

- A ripple on the current/voltage induces a change in tune, beta-beating, orbit, chromaticity ... In general the changes in beta-beating, orbit and chromaticity are negligible, but the **induced tune ripple** can be **non-negligible**.
- **Experiments** at the **SPS** [1,2] suggest that a tune ripple of 10^{-4} is acceptable while experiences at **HERA** [3] show that for low frequencies even a tune ripple of 10^{-5} and for high frequencies 10^{-4} can lead to significant particle diffusion.
- Experiment [1,2], theory and tracking studies [4,5] show that several ripple frequencies are much more harmful than a single one.
- Typical ripple frequencies lie between 5-1200 Hz [1,2,3]

[1] X. Altuna et al., CERN SL/91-43 (AP)

[2] W. Fischer, M. Giovannozzi, F. Schmidt, Phys. Rev. E 55, Nr. 3 (1996)

[2] O. S. Brüning, F. Willeke, Phys. Rev. Lett. 76, Nr. 20 (1995)

[3] O. S. Brüning, Part. Acc. 41, pp. 133-151 (1993)

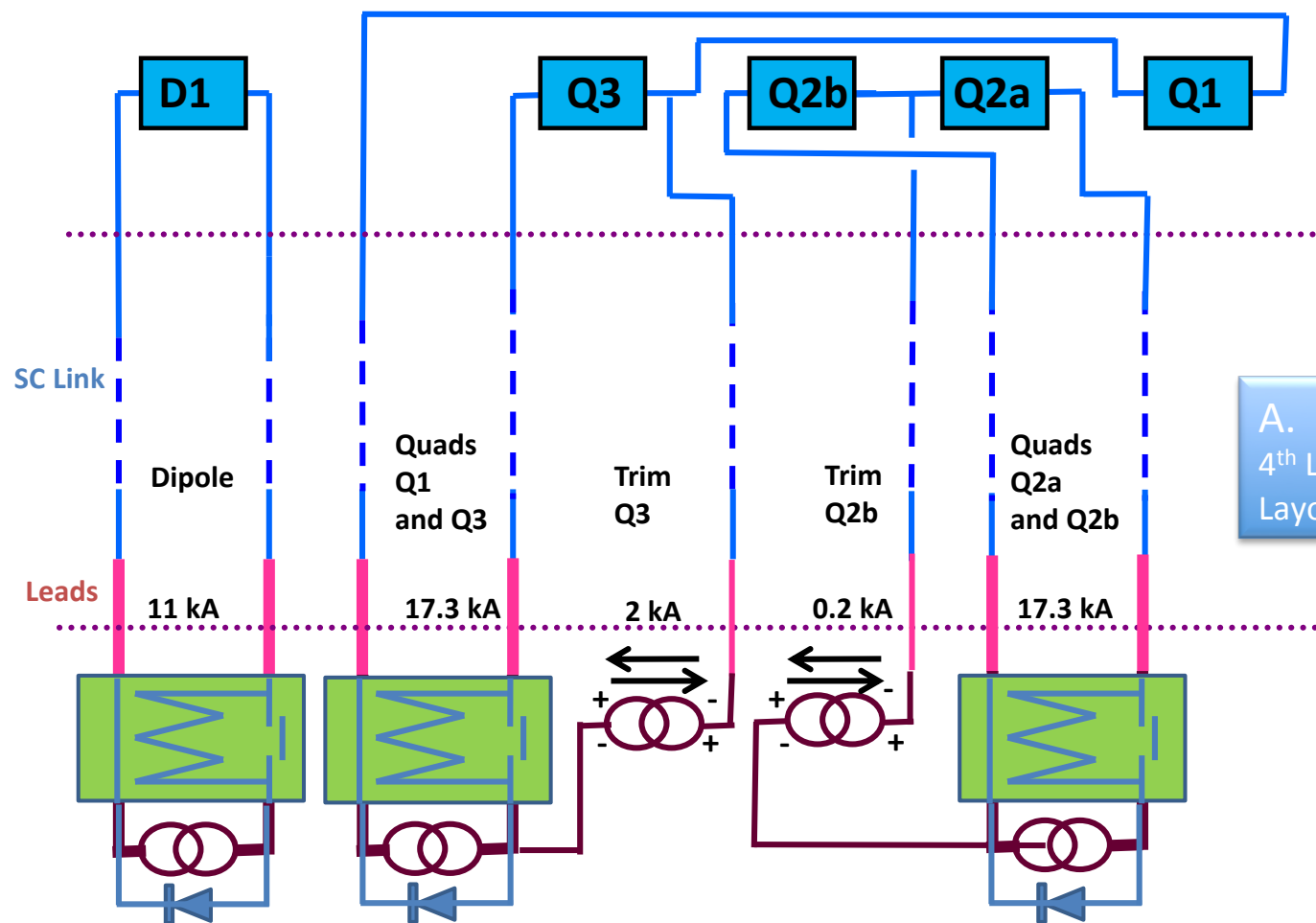
[4] M. Giovannozzi, W. Scandale, E. Todesco, Phys. Rev. E 57, Nr.3 (1998)



First estimate by calculating the tune ripple induced by a uniformly distributed error on the current, which should stay below 10^{-4}

Tolerances on powering of the IT (2)

Proposed powering scheme HL-LHC:



A. Ballarino,
4th LHC Parameter and
Layout Committee



Tolerances on powering of the IT (3)

Simulation using MAD-X, HLLHCV1.0 optics:

- uniformly distributed (independent) errors on current => gradient error:

$$itok = kmax/(17.3)$$

$$di = 1e-06$$

$$dk1l5 = 2*itok*17.3*di*(ranf()-0.5)$$

$$dkt3l5 = 2*itok*2.0*di*(ranf()-0.5);$$

...

$$kqx1.L5 := kqx10.L5 + dk1l5$$

$$kqx2a.L5 := kqx2a0.L5 + dk2l5 ;$$

$$kqx2b.L5 := kqx2b0.L5 + dk2l5 + dkt2bl5 ;$$

$$kqx3.L5 := kqx30.L5 + dk1l5 + dkt3l5 ;$$

...

$$kmax=0.599599999902e-02$$

+/-1 ppm ripple on current

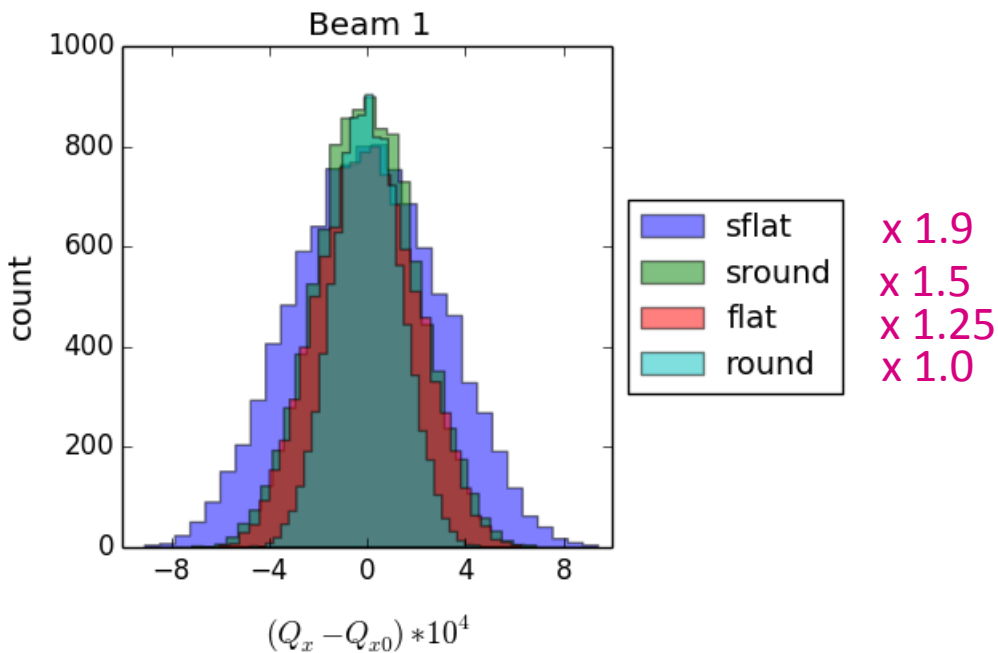
+/-di uniformly distributed error

powering
scheme

- optics: round ($\beta^*=15$ cm), flat ($\beta^*=7.5/30$ cm), sround ($\beta^*=10$ cm), sflat ($\beta^*=5.0/20$ cm)

Tolerances on powering of the IT (4)

- **linear** dependence on relative current error
(note: $\Delta Q = \frac{1}{4\pi} \oint \beta(s) \Delta k(s) ds$)
- almost **no effect from trims** for Q3 and Q2b
- dependence on β^* : apply +/- 1.0 ppm current ripple



$$\Delta Q = \frac{1}{4\pi} \oint \beta(s) \Delta k(s) ds \quad \text{and}$$

$$\beta(s) = \beta^* + \frac{s^2}{\beta^*} \approx \frac{s^2}{\beta^*} \quad \text{and}$$

$$\Delta Q_{\text{tot}} = \Delta Q_{\text{IR1}} + \Delta Q_{\text{IR5}} \quad \text{and}$$

alternate xing + anti-sym. triplet

$$\Rightarrow \Delta Q_z \sim \left(\frac{1}{\beta_{1,z}^*} + \frac{1}{\beta_{2,z}^*} \right)$$

Conclusion and open questions

Conclusion

- Ground motion in IT/D1 area (<10s):
 - effect on orbit similar to the LHC (but smaller beam spot size for the HLLHC) and negligible in the range <10 s
- BPM tolerances:
 - orbit at the location of the crab cavities negligible (in the range of 2 μm for a resolution of +/-0.5 μm)
 - the simplified model used shows that for a luminosity loss <1%, a BPM resolution of +/-1.25 μm would be needed assuming $\epsilon_n=2.0 \mu\text{m}$, $\beta^*=0.15 \text{ m}$ [*]
 - a matching of the IP is also possible with only a subset of BPMs with an eventual loss in precision (note: same precision for 1x4 BPM less around the IP and 1x4BPM less around crab cavities)

[*] linear system, thus $\delta_{\text{IP,max}} = \text{const} * (\text{BPM resolution})$

Conclusion

- Misalignment Correction:

- Two feasible correction schemes:

- a) MCBX*+2 MCBRD correctors: minimization of the orbit in the IT

- b) all correctors (MCBX*+4 MCBRD, limit on MCBRD strength): minimization of corrector strength

- Tolerances on powering of the IT:

- uniformly distributed relative current error of +/-1.0 ppm leads to a tune ripple of $4.0-8.0 \times 10^{-4}$ depending on the optics

- scaling with beta function according to $(1/\beta_1^*+1/\beta_2^*)$ and linear scaling with relative error

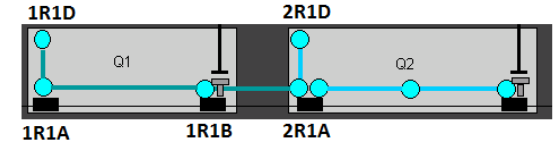
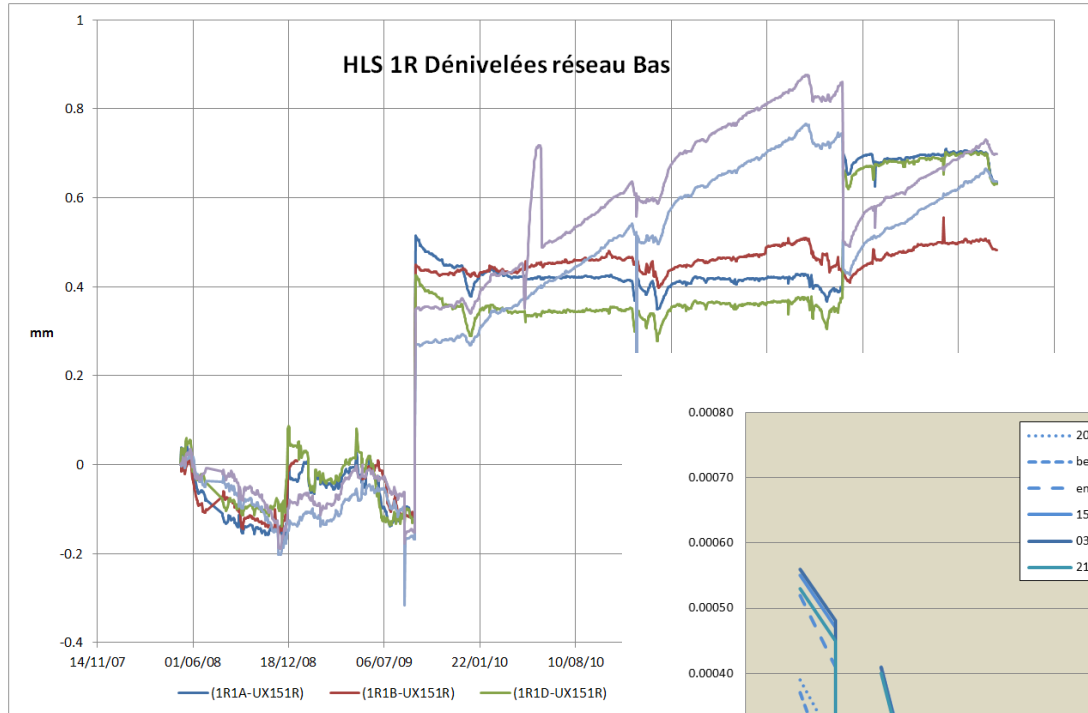
Open questions?

- Ground motion in the IT/D1 area:
 - here misalignment = misalignment of the cryostat.
Movement of the magnet inside cryostat?
- BPM tolerances:
 - precision (reproducibility of measurements) during one fill?
 - effect of gradient errors?
 - effect of longitudinal misalignment of the BPMs (high z' , thus small longitudinal misalignment results in relatively large difference orbit) and also the triplet?
- Tolerances on powering of the IT:
 - acceptable tune ripple?
 - frequencies of ripple?
 - if tolerances on ripple turn out to be too tight, possible compensation like in HERA?





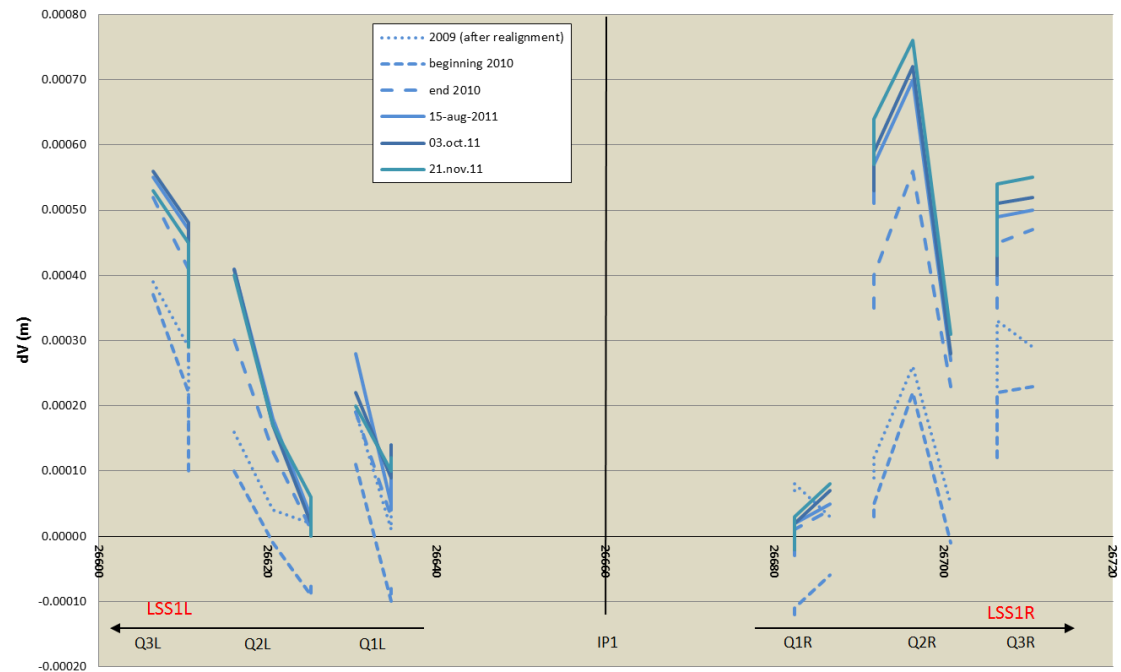
Ground motion measurements LHC IT



H. Mainaud Durand

4th HL-LHC PLC meeting

IR1 - Low beta Vertical



- Leveling measurements once per year during the winter shutdown
- Relative measurements every 10s

Simulation method ground motion

Simulation method [3], [4]:

- Ground motion and thus misalignment of the IT is described as a superposition of sine and cosine like oscillations with the frequency and wavelength given by the 2-d power spectrum $p(f, k)$
- The effect of the misalignment on the closed orbit is calculated by using the orbit response matrix M :

$$x_{\text{co}} = M \cdot x_{\text{mis}}$$

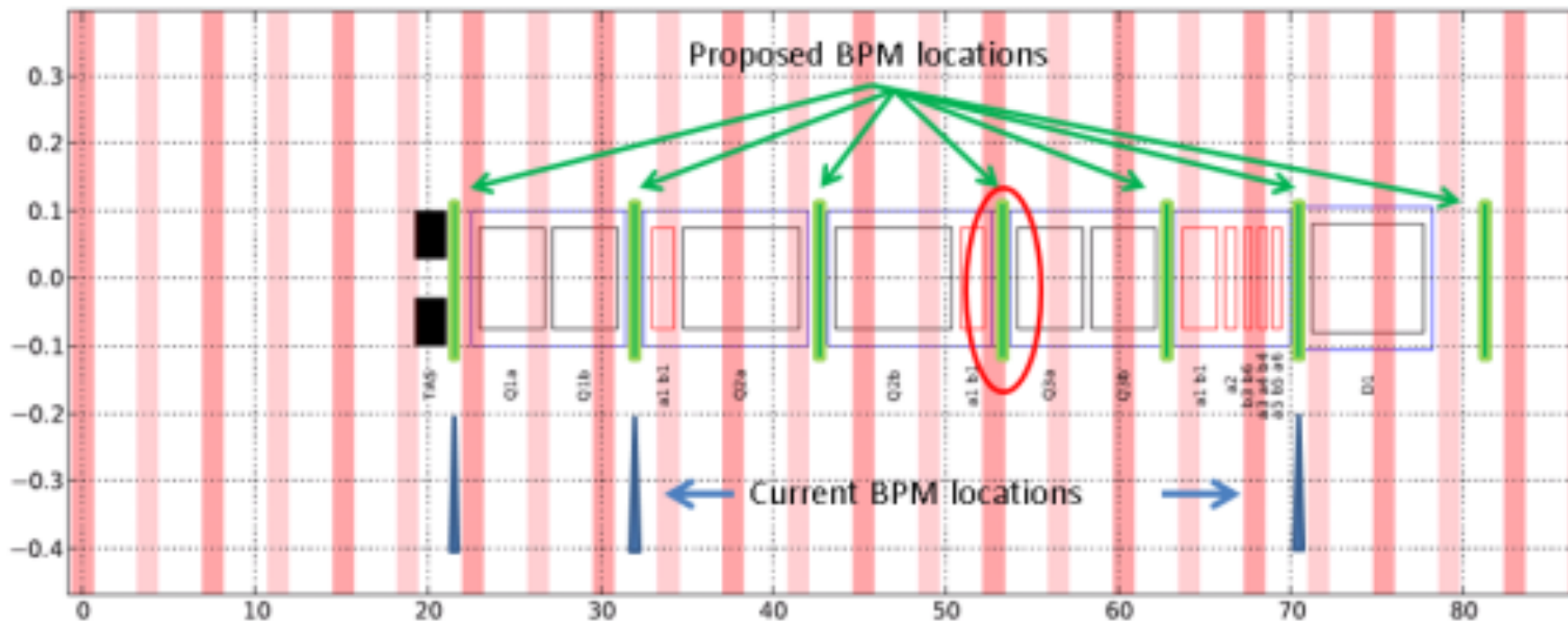
- The effect of the misalignment on the orbit (expressed by the sensitivity function $G(k)$) is first calculated for each frequency and wavelength individually
- All waves are assumed to be independent, thus the rms closed orbit offset is given by:

$$\sigma_{\text{co,tot}} = \sum_{f,k} (p(f, k) \cdot \Delta f \cdot \Delta k \cdot G(k)^2)$$

[3] A. Sery and O. Napoly, Influence of ground motion on the time evolution of beams in linear colliders, Phys. Rev. E, 53:5323 (1996)

[4] J. Pflugstner, Mitigation of ground motion effects via feedback systems for the Compact Linear Collider, PhD Thesis (2012)

BPM tolerances for orbit correction (1)

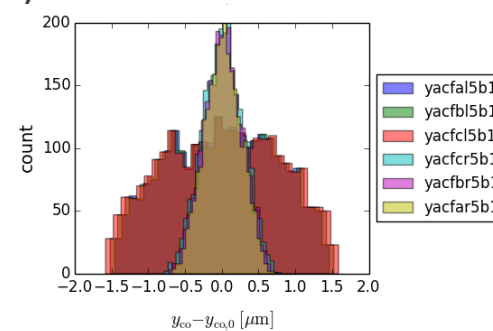
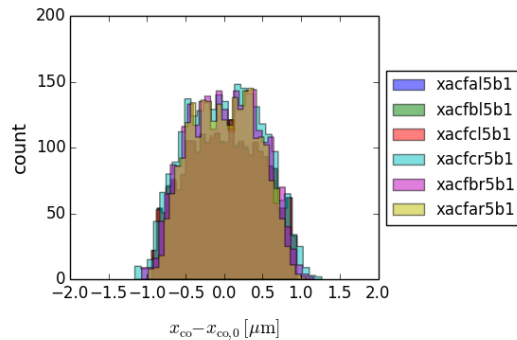


R. De Maria

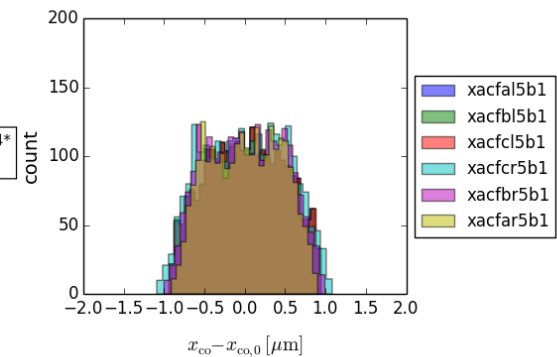
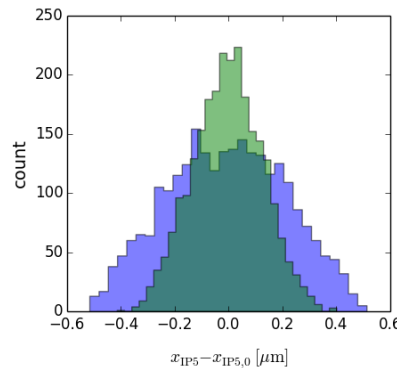
BPM tolerances for orbit correction (4)

c) use selection of BPMs

- same performance using only BPMSQW1*,BPMSQ1,BPMSQA2 and BPMWD4*,BPMYY4 (no BPMSQB2*, no BPMWC4*)
- x2 worse orbit at position of crab cavities using only BPMSQW1*,BPMSQ1* and BPMWD4*,BPMYY4* (no BPMSQA2*)



- possible to match x-scheme using only BPMSQ1* and BPMYY4*, but worse orbit at IP and location of crab cavities and a max. orbit excursion of up to 0.45 mm



Misalignment correction (4)

c) all correctors, but limit the maximum orbit in the triplet

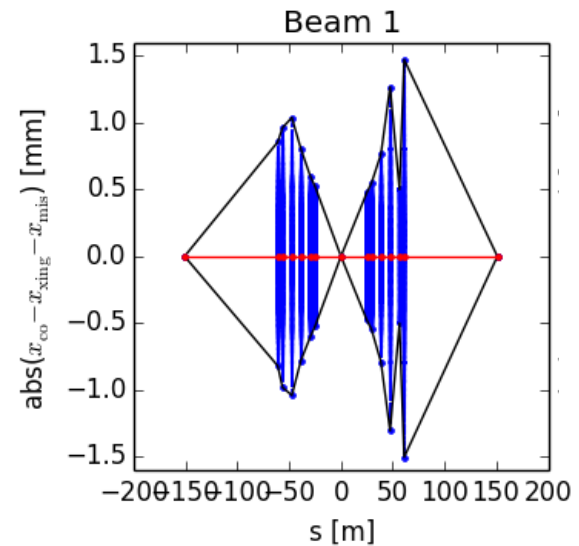
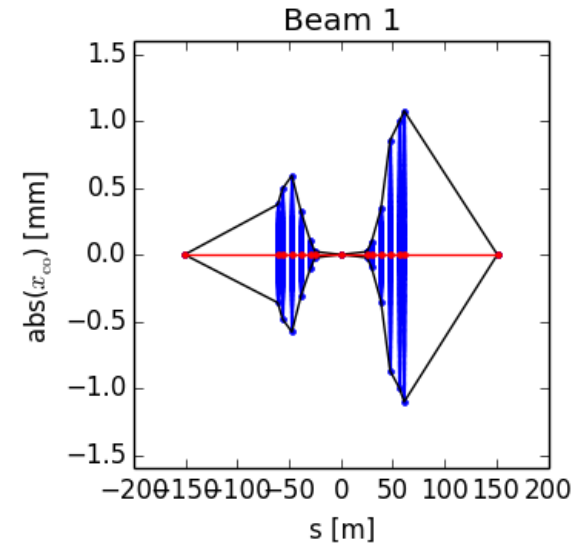
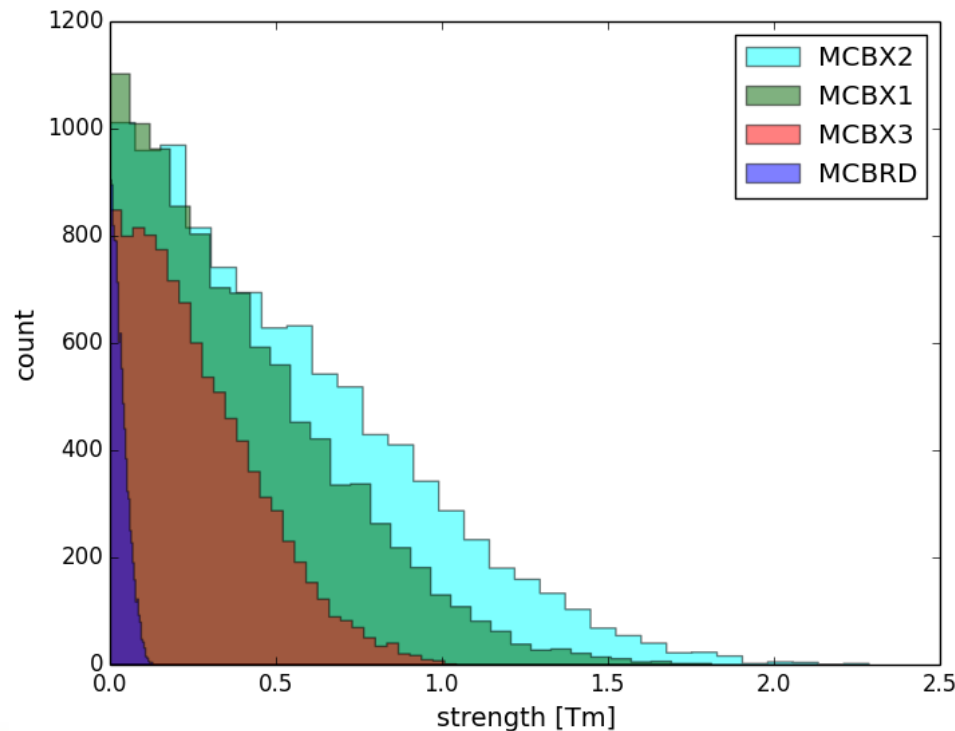
→ slightly smaller orbit deviation (<1.1 mm)

than for case b)

→ but x2 larger corrector strength than for case b)

(MCBX1 <1.5 Tm, MCBX2 <2 Tm, MCBX3 <1.0 Tm,

MCBRD <0.1 Tm)



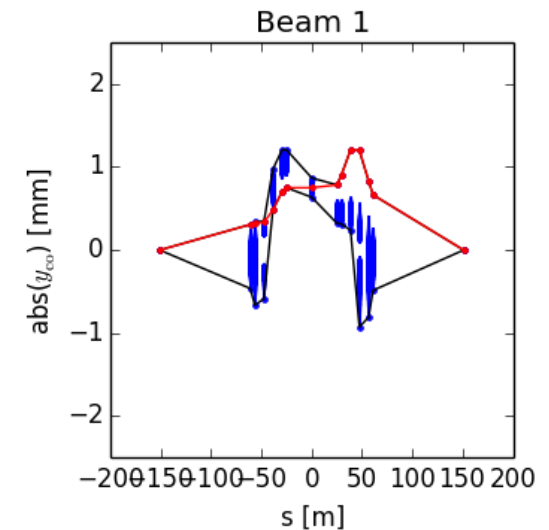
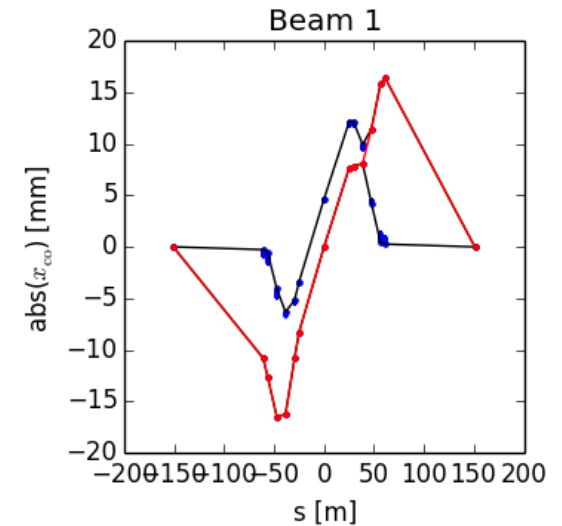
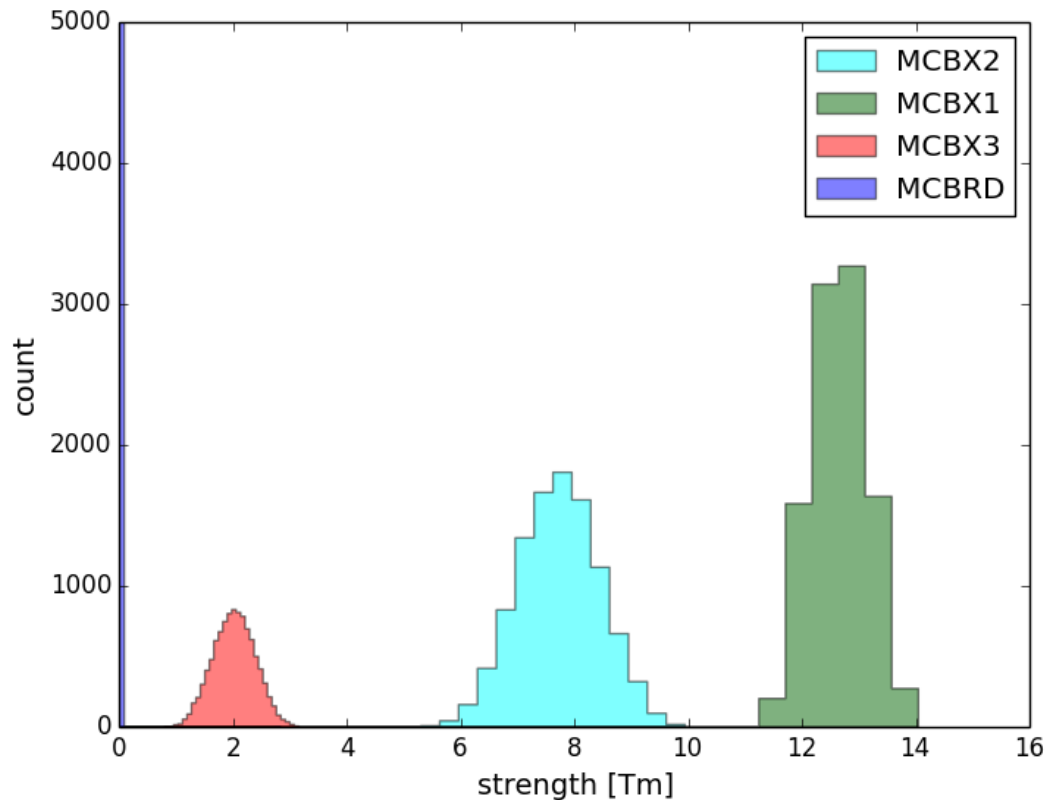
Misalignment correction (4)

d) only MCBX* and $x_{IP,b1}-x_{IP,b2}=sep$, $px_{IP,b1}-px_{IP,b2}=x-angle$

➔ smaller orbit in triplet ($x_{co}<12$ mm, $y_{co}<1.2$ mm)

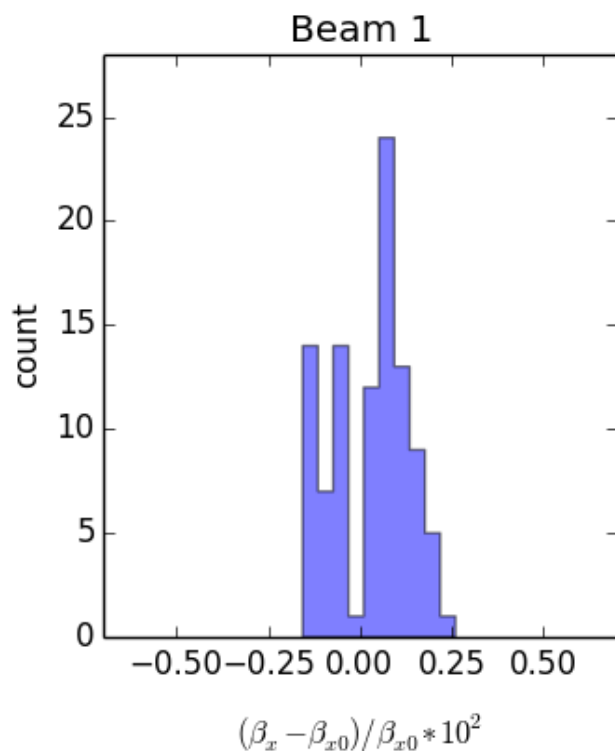
➔ **but too high corrector strength!**

(MCBX1<14 Tm, MCBX2<10 Tm, MCBX3<3.0 Tm)

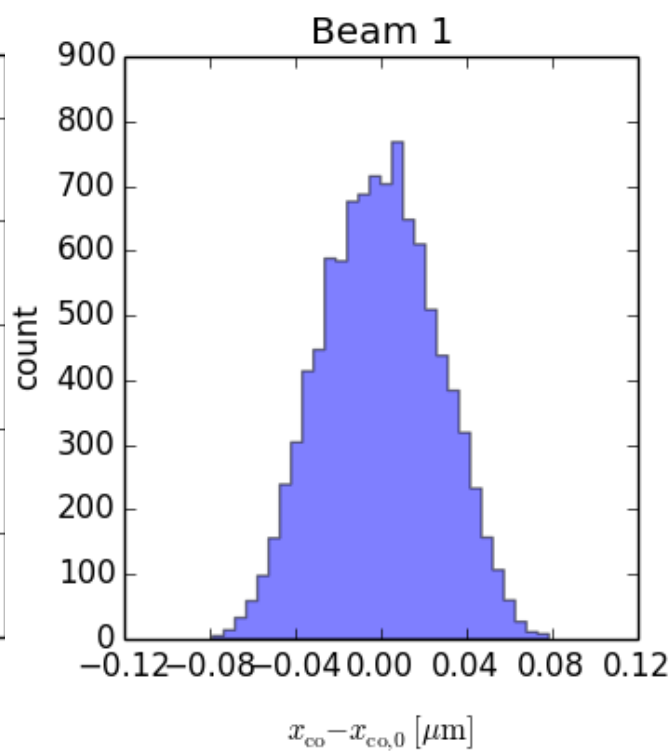


Tolerances on powering of the IT (1)

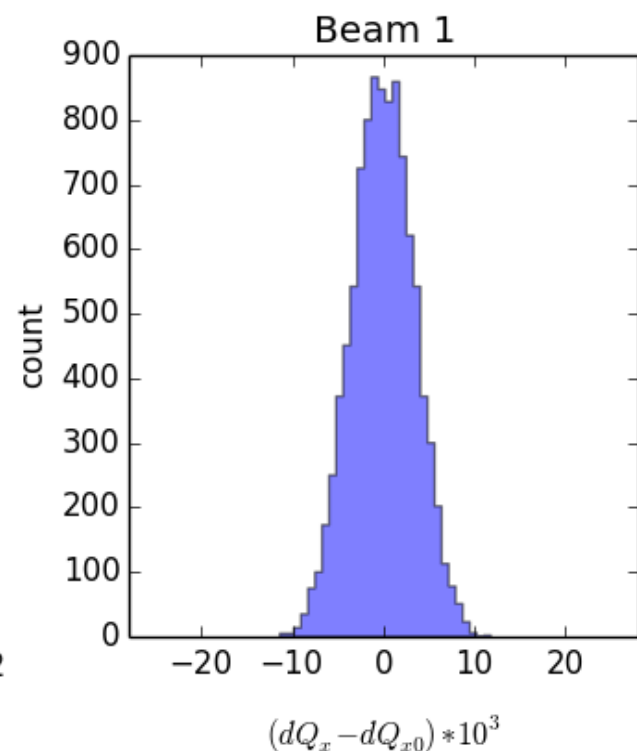
Effect on beta-beating, orbit and chromaticity for round optics
(15 cm β^*) and +/-1.0 ppm ripple on current :



max. beta-beating
(complete ring)



closed orbit IP5,
beam 1



chromaticity

Tolerances on powering of the IT (3)

Simulation setup for MAD-X:

- uniformly distributed ripple on current => gradient ripple:

itok = kmax/(17.3); (kmax=0.5995999999902E-02)

di = 1e-06; (ripple on current)

dk1l5 = 2*itok*17.3*di*(ranf()-0.5); (+/-di uniformly distributed error)

kqx1.L5 := kqx10.L5 + dk1l5 ;

kqx2a.L5 := kqx2a0.L5 + dk2l5 ;

kqx2b.L5 := kqx2b0.L5 + dk2l5 + dkt2bl5 ;

kqx3.L5 := kqx30.L5 + dk1l5 + dkt3l5 ;

kqx1.R5 := kqx10.R5 + dk1r5 ;

kqx2a.R5 := kqx2a0.R5 + dk2r5 ;

kqx2b.R5 := kqx2b0.R5 + dk2r5 + dkt2br5 ;

kqx3.R5 := kqx30.R5 + dk1r5 + dkt3r5 ;

powering scheme
(same for IR1)

- simulation for 10^4 different seeds