

# Fine-tuning in the scale invariant NMSSM

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# Outline

- 1 The MSSM
  - Fine tuning in the MSSM
- 2 Scale Invariant NMSSM
  - Fine tuning in the scale invariant NMSSM
  - Phenomenology
- 3 Conclusions

# SUPERSYMMETRY (SUSY)

- The MSSM contains the **minimal supersymmetric particle content** compatible with the Standard Model (SM),  $\mathcal{N} = 1$  SUSY (one SUSY charge generator).
- **Gauge couplings unify** in the MSSM  $\implies$  strong hint towards GUT theories.
- If **R-parity** is conserved: avoid proton decay constraints, SUSY particles are produced in even numbers at colliders and the LSP is stable  $\implies$  **Dark Matter (DM)** candidate.

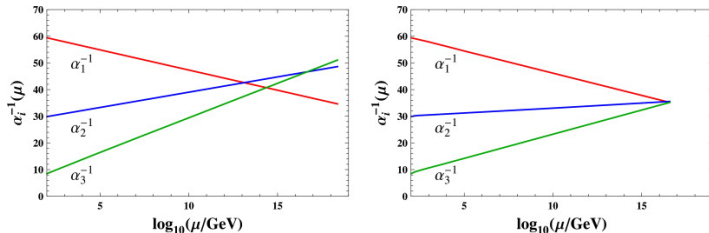


Figure: SM (left) vs MSSM (right).

# SUperSYmmetry (SUSY)

- SUSY avoids **quadratically** divergent quantum corrections to the Higgs mass involving a UV-cutoff  $\Lambda$ .

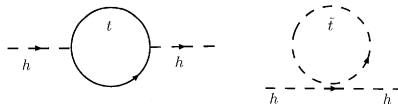


Figure: Top-stop contributions to Higgs mass.

- Quantum corrections still generate **large logarithms**  $\sim \log(\Lambda/m_{\text{soft}}) \Rightarrow$  Summed up via **Renormalization Group Equations** ( $\beta$ -functions).
- Finite quantum corrections** can be taken into account by **effective action** methods,

$$\mathcal{S}_{\text{eff}} = \int d^4x \left\{ \sum_{n=0}^{\infty} Z_i^n \partial_\mu \phi_i^\dagger \partial^\mu \phi_i - \sum_{n=0}^{\infty} V_n \right\}$$

- At one-loop we recover the Coleman-Weinberg formula which in the  $\overline{DR}$  scheme is

$$V_1 = \frac{1}{64\pi^2} \text{STr} M^4 \left[ \ln \left( \frac{M^2}{\mu_r^2} \right) - \frac{3}{2} \right]$$

- Energy scales involved:

$\Lambda$  \_\_\_\_\_ Scale at which SUSY breaking  
is transmitted  
from hidden sector to visible sector

$m_{soft}$  \_\_\_\_\_ Scale at which EWSB happens

$m_h \sim v_{EW}$  \_\_\_\_\_ Pole Higgs mass



# Fine tuning in the MSSM

- In the MSSM Higgs quartic coupling at tree-level  $\lambda_h \propto g^2$  with  $g \in SU(2)_W$ , (D-terms).
- However, finite loop corrections provide **additional** contributions to  $\lambda_h$ , the biggest coming from top/stop loops.
- **MSSM Higgs mass** in the decoupling limit ( $m_A^2 \gg m_h^2$ ) *Carena et al.*,

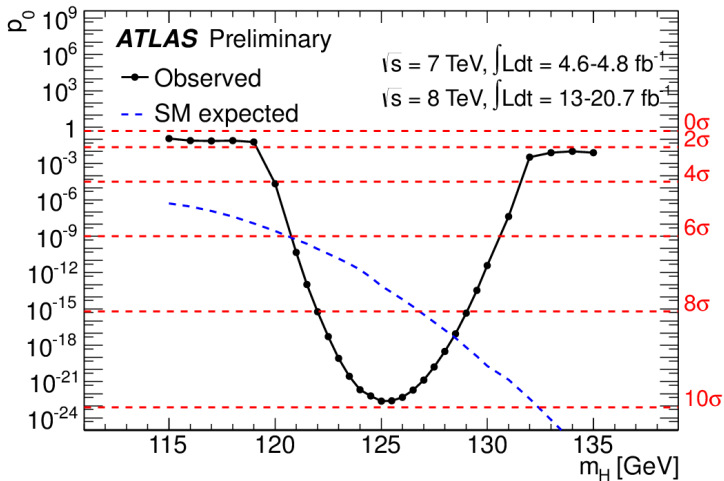
$$m_h^2 = m_Z^2 \cos^2 2\beta \left( 1 - \frac{3}{8\pi^2} y_t^2 t \right) + \frac{3}{4\pi^2} y_t^2 \left[ \frac{1}{2} X_t + t + \frac{1}{16\pi^2} \left( \frac{3}{2} \frac{m_t^2}{v^2} - 32\pi\alpha_3 \right) (X_t t + t^2) \right]$$

with  $t = \ln \frac{m_{\text{soft}}^2}{m_t^2}$ ,  $X_t = \frac{2(A_t - \mu \cot \beta)^2}{m_{\text{soft}}^2} \left( 1 - \frac{(A_t - \mu \cot \beta)^2}{12m_{\text{soft}}^2} \right)$ ,  $m_{\text{soft}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = \mu_r$ .

- If  $m_h \approx 126$  GeV  $\Rightarrow$  **large**  $m_{Q_3}^2$ ,  $m_{U_3}^2$  (enter logarithmically) or **large**  $A_t$  (enters as a power)  $\Rightarrow$  lead to large (UV-logarithmically) sensitivity on the Higgs potential at **quadratic** order.

## Higgs resonance at 126 GeV

Impressive ATLAS combined result for local probability  $p_0$  of background-only to be more signal-like than the observation



# Fine tuning in the MSSM

- In the MSSM the **minimization** conditions (including finite loop-corrections) take the form,

$$m_Z^2 = \frac{\hat{m}_{H_u}^2 - m_{H_d}^2}{\cos 2\beta} - \hat{m}_{H_u}^2 - \hat{m}_{H_d}^2 - 2\mu^2, \quad \frac{2b}{\sin 2\beta} = \hat{m}_{H_u}^2 + \hat{m}_{H_d}^2 + 2\mu^2$$

where  $\hat{m}_{H_u}^2 \equiv m_{H_u}^2 + \partial V_{\text{eff}} / \partial v_u^2$  and  $\hat{m}_{H_d}^2 \equiv m_{H_d}^2 + \partial V_{\text{eff}} / \partial v_d^2$ .

- From RG evolution in the SUSY theory (in the **leading-log** approximation)

$$m_{H_u}^2(m_{\text{soft}}) = m_{H_u}^2(\Lambda) - \frac{3y_t^2}{8\pi^2} \left[ m_{Q_3}^2(\Lambda) + m_{U_3}^2(\Lambda) + A_t^2(\Lambda) \right] \ln \left[ \frac{\Lambda}{m_{\text{soft}}} \right]$$

$\therefore$  We must **tune**  $m_{H_u}^2(\Lambda)$  in order to compensate **large stop corrections**  $\Rightarrow$  known as the **"little hierarchy problem"**.



# Fine tuning in the MSSM

One way to quantify the tuning is with the **measure** [Barbieri, Giudice](#)

$$\Sigma^v \equiv \max_{\xi_i} \left| \frac{\partial \ln v^2}{\partial \ln \xi(\Lambda)} \right|$$

- How much does  $v$  **change** when infinitesimally moving the **independent** parameters at the high scale  $\Lambda$ .
- In the MSSM the relevant parameters are  $\xi_i = (m_{H_u}^2, m_{H_d}^2, \mu, b, m_{Q_3}^2, m_{U_3}^2, m_{D_3}^2, A_t, A_b, M_1, M_2, M_3)$ .
- Using the **chain rule**,

$$\Sigma^v = \max_i \left| \sum_j \frac{\xi_j(\Lambda)}{v^2} \frac{dv^2}{d\xi_j(m_{\text{soft}})} \frac{d\xi_j(m_{\text{soft}})}{d\xi_j(\Lambda)} \right|$$

$\Rightarrow$

$$\left| \frac{d \log v^2}{d \log m_{H_u}^2(\Lambda)} \right| \simeq \left| \frac{3y_t^2}{8\pi^2 v^2} (m_{Q_3}^2(\Lambda) + m_{U_3}^2(\Lambda) + A_t^2(\Lambda)) \log \left[ \frac{\Lambda}{m_{\text{soft}}} \right] \times \frac{dv^2}{dm_{H_u}^2(m_{\text{soft}})} \right|$$

# Fine tuning in the MSSM

- Neglecting the Coleman-Weinberg corrections in the MSSM and in the regime  $\tan \beta \gg 1$

$$\frac{dv^2}{dm_{H_u}^2(m_{\text{soft}})} \simeq -\frac{2v^2}{m_Z^2} + \mathcal{O}\left(\frac{1}{\tan \beta}\right)$$

∴ **NO FREEDOM TO SUPPRESS DERIVATIVE AND REDUCE TUNING!**

- In the MSSM with current experimental constraints and  $m_h \approx 126$  GeV,  $\Sigma^v \gtrsim 500$ .

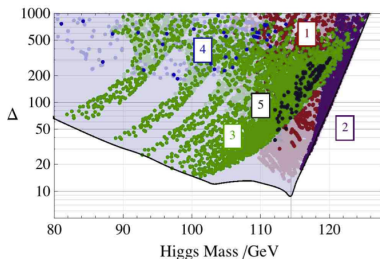


Figure: Cassel et al. 1101.4664 [hep-ph].

# Saving low energy SUSY

- 1 Give up on small tuning: landscape, anthropic principle, split SUSY, . . .
- 2 Keep naturalness as guiding principle and add additional non-decoupling D-term contributions to the Higgs potential to raise Higgs mass at tree-level (enlarge weak gauge group)
- 3 Add additional F-term contributions: e.g. NMSSM; introduce a gauge singlet chiral superfield S.

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# Scale Invariant NMSSM

- Introduce a gauge singlet chiral superfield  $S$  in addition to the MSSM superfield content.
- **No mass scale** in the new **superpotential** piece ( $\mathbb{Z}_3$  symmetry):

$$W_{NMSSM} = \lambda S H_d H_u + \frac{\kappa}{3} S^3$$

- Soft breaking terms:

$$V_{soft} = m_{H_u}^2 |H_u|^2 + m_{H_d}^2 |H_d|^2 + m_S^2 |S|^2 + \left( a_\lambda S H_d H_u + \frac{a_\kappa}{3} S^3 + h.c \right)$$

Notice from  $W_{NMSSM}$  that:

- 1 If  $\langle s \rangle \sim v$  for  $\lambda \sim 1 \Rightarrow$  **solve** the  $\mu$ -problem found in the MSSM.
- 2 Contributions to the Higgs potential of the form  $F_S F_S^*$  generate a **quartic coupling at tree-level** proportional to  $\lambda^2$ .

# Scale Invariant NMSSM

- Assume CP conservation on the Higgs-singlet sector  $\Rightarrow$  scalar Higgs-singlet sector decomposes in 3 CP-even states, 2 CP-odd states and 1 charged Higgs.
- Neutral components:  $H_u^0 = v_u + (h_u + ih_{u,I})/\sqrt{2}$ ,  $H_d^0 = v_d + (h_d + ih_{d,I})/\sqrt{2}$  and  $S = v_s + (s + is_I)/\sqrt{2}$ .
- For neutral CP-even sector is useful to rotate to the basis

$$\begin{pmatrix} h \\ H \\ s \end{pmatrix} = \begin{pmatrix} \sin \beta & \cos \beta & 0 \\ -\cos \beta & \sin \beta & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} h_u \\ h_d \\ s \end{pmatrix}$$

where only  $\langle h \rangle = v \neq 0$  (besides  $\langle s \rangle = v_s \neq 0$ ) and  $m_{hh}^2 = m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta$ .

- We want  $\tan \beta \approx 1$  to increase the tree-level Higgs mass (upper bound on the tree-level Higgs mass).
- If lightest eigenstate is mostly  $h$ , mixing with either  $H$  or  $s$  pulls mass down (level repulsion).

# Scale Invariant NMSSM

- We take the **best case** scenario for tuning:
  - ① Keep only 3<sup>rd</sup> generation squarks and gauginos **light**. All other sparticles have masses  $\tilde{m} \sim \Lambda$ .
  - ② Take a **low** "messenger" scale (effective cutoff)  $\Lambda = 20 \text{ TeV}, 100 \text{ TeV}, 1000 \text{ TeV}$ .
- In this way we obtain the **largest regions** of parameter space consistent with low fine tuning and whose collider and flavor constraints are ameliorated due to family splitting.
- Furthermore, low cutoff allows for a possible **large value** of  $\lambda(m_{\text{soft}}) \gtrsim 1$  (if  $\Lambda = M_{GUT} \Rightarrow \lambda(m_{\text{soft}}) \lesssim 0.65$ ).
- Models of  $\lambda$ -SUSY [Barbieri, Hall, Nomura, Ruderman, . . .](#)  
∴ **Loop corrections** not only from the top/stop sector but also from the **Higgs-singlet sector** become **very important** and should be included.

Large  $\lambda$  helps

- **Minimization** conditions for the scale invariant NMSSM

$$m_Z^2 = \frac{m_{H_u}^2 - m_{H_d}^2}{\cos 2\beta} - m_{H_u}^2 - m_{H_d}^2 - 2\lambda^2 v_s^2.$$

$$\lambda^2 v^2 = \frac{2(a_\lambda v_s + \lambda \kappa v_s^2)}{\sin 2\beta} - m_{H_u}^2 - m_{H_d}^2 - 2\lambda^2 v_s^2$$

$$m_s^2 = \lambda \kappa v^2 \sin 2\beta - 2\kappa^2 v_s^2 - \lambda^2 v^2 - \frac{a_\lambda v^2}{2v_s} \sin 2\beta - a_\kappa v_s$$

- Parameters  $\xi_i = (m_{H_u}^2, m_{H_d}^2, m_s^2, \lambda, \kappa, a_\lambda, a_\kappa, m_{Q_3}^2, m_{u_3}^2, m_{d_3}^2, A_t, A_b, M_1, M_2, M_3)$ .
- In this case we find

$$\frac{dv^2}{dm_{H_u}^2(m_{\text{soft}})} = \frac{\kappa}{\lambda^3} \cot 2\beta + \mathcal{O}\left(\frac{1}{\lambda^4}\right)$$

**Suppressed** for large values of  $\lambda$ .

$\therefore$  Large  $\lambda$  **seems** to help allowing for smaller tuning and/or larger stop masses.



# Large $\lambda$ helps

- All plots have 5 % vev tuning or better.

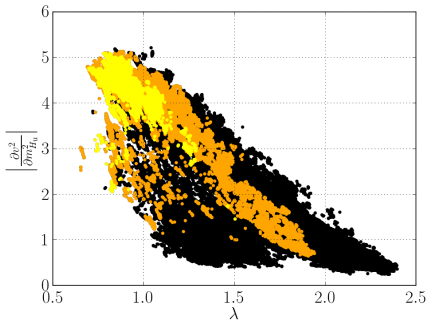


Figure:  $\left| dv^2 / dm_{F_u}^2 \right|$  vs  $\lambda$ .

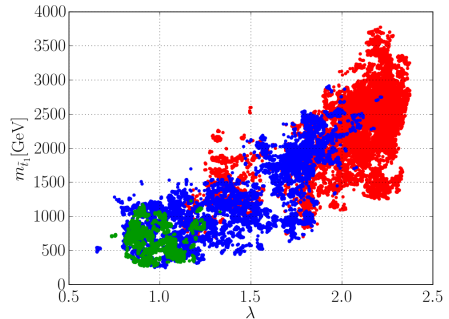


Figure:  $m_{\tilde{t}_1}$  [GeV] vs  $\lambda$ .

# Large $\lambda$ hurts

We can write the **Higgs mass** as,

$$m_h^2 = m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta + \delta m_{h,mix}^2 + \delta m_{h,stop}^2 + \delta m_{h,S}^2$$

- For  $\tan \beta \approx 1$  and  $\lambda \gtrsim 1$  we "overshoot" the Higgs mass at tree-level,  $m_{h,tree} \gg 126$  GeV (for  $\lambda \simeq 2.4$ ,  $m_{h,tree}^2 \approx 10(126 \text{ GeV})^2$ ).
- For most of our points **admixture** modifies mass as most by 40 % . Thus pull-down effect is typically **not large**.
- $\delta m_{h,stop}^2$  generically provides a **positive** contribution to  $m_h^2$ .
- The finite loop corrections from the **Higgs-singlet** sector can provide **large negative** contributions to  $m_h^2$ .

This leads us to define a **new tuning measure** in analogy with the EW vev tuning:

$$\Sigma^h \equiv \max_{\xi_i} \left| \frac{d \log m_h^2}{d \log \xi_i} \right|$$

Think of it as a **tuning in the quartic Higgs coupling** (contrary to the quadratic tuning associated with the  $v$ ).  $\xi_i$  are the same as before except  $m_{H_u}^2$ ,  $m_{H_d}^2$  and  $m_S^2$  have been replaced by  $v_u$ ,  $v_d$  and  $v_s$  which are kept fixed.

# Large $\lambda$ hurts

- $\Sigma^h \propto \lambda^2$
- At  $\tan \beta \gtrsim 3$ ,  $m_h^2$  decreases **two-folded** from  $\sin 2\beta$  small and because larger  $\tan \beta \Rightarrow$  larger higgsino contributions to the T-parameter  $\Rightarrow \lambda$  must be smaller.

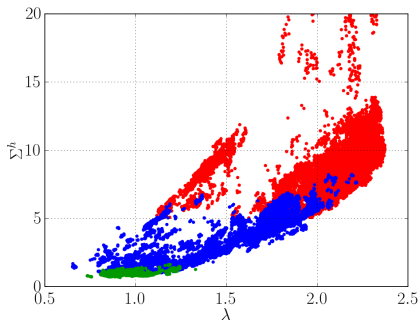


Figure:  $\Sigma^h$  vs  $\lambda$ .

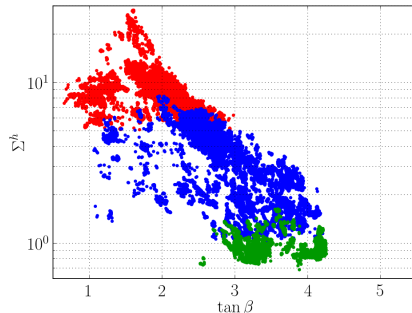


Figure:  $\Sigma^h$  vs  $\tan \beta$ .

# Combined tuning

Define a **combined tuning measure**  $\Sigma^h \times \Sigma^v$  (if two quantities are not correlated, the probability involving both is  $P(A \cap B) = P(A) \cdot P(B)$ ).

- Total tuning dominated at small  $\lambda$  by vev tuning and at large  $\lambda$  by Higgs mass tuning. Minimum at  $\lambda \approx 1$ .
- Brown band corresponds to a characteristic point in the MSSM with  $\mu = 200$  GeV,  $\tan \beta = 20$ ,  $\Lambda = 20$  TeV,  $m_a = 1$  TeV and  $A_t$  such that  $m_h \in [124, 127]$  GeV. Always **better** than the MSSM.

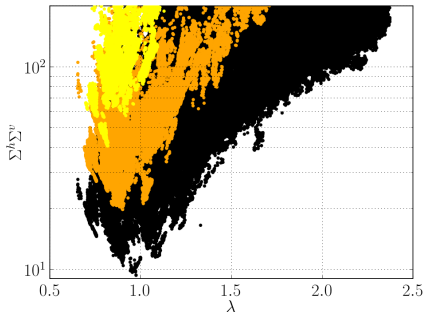


Figure:  $\Sigma^v \times \Sigma^h$  vs  $\lambda$

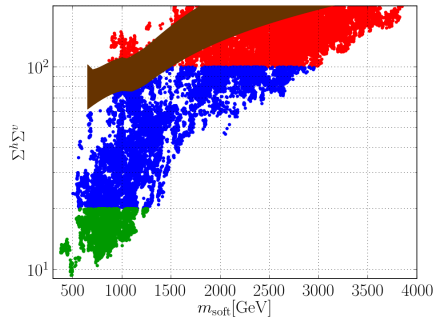


Figure:  $\Sigma^v \times \Sigma^h$  vs  $m_{\text{soft}}$  [GeV]

## Numerical Scan

- Markov Chain Monte Carlo
- We used a modified version of NMHDECAY, which is part of NMSSMTools 3.2.1, as well as MicrOMEGAs 2.4.5 for DM
- We scanned linearly in 14-dimensional parameter space

$\tan \beta$	$\tan \beta > 0.08$	$m_{Q_3}$	$\Delta_{\tilde{g}} m_{Q_3} < m_{Q_3} < 5$	$M_1$	$0 < M_1 < 8$
$\mu$	$ \mu  < 1$	$m_{u_3}$	$\Delta_{\tilde{g}} m_{u_3} < m_{u_3} < 5$	$M_2$	$0 < M_2 < 8$
$\lambda$	$0 < \lambda < 3$	$m_{d_3}$	$0 < m_{d_3} < 8$	$M_3$	$0.5 < M_3 < 8$
$\kappa$	$ \kappa  < 2.75$	$A_t$	$ \Delta_{\tilde{g}} A_t  <  A_t  < 5$	$\nu$	174
$A_\lambda$	$ A_\lambda  < 2$	$A_b$	$ A_b  < 8$	$\Lambda$	20, 100, 1000
$A_\kappa$	$ A_\kappa  < 1$				

⇒ We did not sample the full parameter space. Therefore, there is **no statistical interpretation** of the scatter plots.

- Likelihood function: product of Gaussians for the Higgs mass centered at 126 and for the VEV finetuning centered at 0.
- We impose the hard cuts summarised in the table as well as

$$|\xi(\Lambda) - \xi(m_{\text{soft}})| < |\xi(\Lambda)| \quad \text{for } \xi = m_{Q_3}^2, m_{u_3}^2, m_{d_3}^2, A_t, A_b$$

similar to "gluino sucks the stop mass up" [Arvanitaki, Craig, Dimopoulos, Villadoro\(2012\)](#)

## Electroweak Precision Tests (EWPT)

- impose EWPT at  $2\sigma$  with  $m_{h,ref} = 117$  GeV [PDG\(2012\)](#).

$$S_0 = -0.04 \pm 0.09, T_0 = 0.07 \pm 0.08 \text{ and correlation of } 88\% \text{ at } 95\% \text{ C.L.}$$

- Consistent with previous analyses [Barbieri et. al.\(2006\)](#); [Franceschini, Gori \(2010\)](#), we find and singlet scalars do not contribute much to S,T
- (3rd gen) squarks:  $\tan \beta$  can not be too large unless cancellation  $\tan \beta \neq 1$  breaks custodial SU(2)  $\rightarrow$  contribution to T
- Neutralino/chargino sector imposes strong constraint on  $\lambda$  as well as  $\tan \beta$

$$M_{\psi 0} = \begin{pmatrix} M_1 & 0 & -\cos \beta \sin \theta_W m_Z & \sin \beta \sin \theta_W m_Z & 0 \\ \cdot & M_2 & \cos \beta \cos \theta_W m_Z & -\sin \beta \cos \theta_W m_Z & 0 \\ \cdot & \cdot & 0 & -\mu & -\lambda v \sin \beta \\ \cdot & \cdot & -\mu & 0 & -\lambda v \cos \beta \\ \cdot & \cdot & \cdot & \cdot & -2\frac{\kappa}{\lambda} \mu \end{pmatrix}$$

$$M_{\psi \pm} = \begin{pmatrix} 0 & X^T \\ X & 0 \end{pmatrix} \quad \text{with} \quad X = \begin{pmatrix} M_2 & \sqrt{2} m_W \sin \beta \\ \sqrt{2} m_W \cos \beta & 2\mu \end{pmatrix}$$

in gauge-basis  $\psi^0 = (\tilde{B}, \tilde{W}^3, 1\tilde{H}_u^0, 1\tilde{H}_d^0, 1\tilde{S})$  and  $\psi^\pm = (\tilde{W}^\pm, 1\tilde{H}_u^\pm, \tilde{W}^\mp, 1\tilde{H}_d^\mp)$

- We find  $\tan \beta \lesssim 5$  depending on  $\lambda$

## SUSY searches

Assuming a neutralino LSP, we conservatively exclude the following regions:

- gluino search:  $\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$  [ATLAS-CONF-2012-151](#)

$$m_{\tilde{g}} < 1310 \quad \text{if } m_{\tilde{\chi}_1^0} < 650$$

- sbottom search [PDG \(2012\)](#); [CMS-PAS-SUS-12-028](#); [ATLAS-CONF-2012-106](#)

$$m_{\tilde{b}} < 89$$
$$150 < m_{\tilde{b}} < 650 \quad \text{if } m_{\tilde{\chi}_1^0} < 230$$

- stop search [PDG\(2012\)](#); [ATLAS \(1208.1447, 1208.2590\)](#)

$$m_{\tilde{t}} < 95.7$$
$$220 < m_{\tilde{t}} < 500 \quad \text{if } m_{\tilde{\chi}_1^0} < 160$$

- chargino search  $m_{\tilde{\chi}_{\pm}} < 94$  [PDG \(2012\)](#) The neutralino/chargino searches by ATLAS/CMS did not lead to further constraints.

# Higgs searches

$$R_X \equiv \frac{\sigma(h) \times BR(h \rightarrow X)}{\sigma(h_{\text{SM}}) \times BR(h_{\text{SM}} \rightarrow X)}$$

- Higgs resonance at 126

ATLAS (1207.7214), ATLAS-CONF-2012-162; ATLAS-CONF-2012-170; CMS (1207.7235), CMS-HIG-12-045

$$\begin{aligned} 0.81 < R_{ZZ} < 1.32, & \quad 0.74 < R_{WW} < 1.40, \\ 0 < R_{b\bar{b}} < 1.10, & \quad 0.27 < R_{\tau\tau} < 1.15, \end{aligned}$$

- Heavy Higgs searches [CMS-PAS-Higgs-11-024](#), [CMS-PAS-Higgs-11-041](#)

$$\begin{aligned} \frac{\sigma(s_i) \times BR(s_i \rightarrow ZZ)}{\sigma(h_{\text{SM}}) \times BR(h_{\text{SM}} \rightarrow ZZ)} &< 0.09 \\ \frac{\sigma(s_i) \times BR(s_i \rightarrow WW)}{\sigma(h_{\text{SM}}) \times BR(h_{\text{SM}} \rightarrow WW)} &< 0.2 \end{aligned}$$

- Charged Higgs [PDG\(2012\)](#)

$$m_{H^\pm} \gtrsim 79.3$$



# Flavour Constraints

We use the following flavour physics constraints

- $B$ -meson mixing [HFAG \(1207.1158\)](#)

$$\Delta M_S = (17.719 \pm 0.086) \text{ ps}^{-1}$$

$$\Delta M_D = (0.507 \pm 0.008) \text{ ps}^{-1}$$

- rare  $B$ -decays [HFAG \(1207.1158\)](#)

$$\text{Br}(B^+ \rightarrow \tau^+ \nu_\tau) = (1.67 \pm 0.60) \times 10^{-4}$$

$$\text{Br}(B \rightarrow X_S \gamma) = (3.55 \pm 0.48 \pm 0.18) \times 10^{-4}$$

- recently measured rare decay  $B_S^0 \rightarrow \mu^+ \mu^-$  [LHCb \(1211.2674\)](#)

$$\text{Br}(B_S^0 \rightarrow \mu^+ \mu^-) = 3.2_{-2.4}^{+3.0+1.0}_{-0.6} \times 10^{-9}$$

# Particle Spectrum

- $\mu \in [105, 850]$  GeV.
- EW sector is light (possible light charginos) while coloured sector can be heavy.

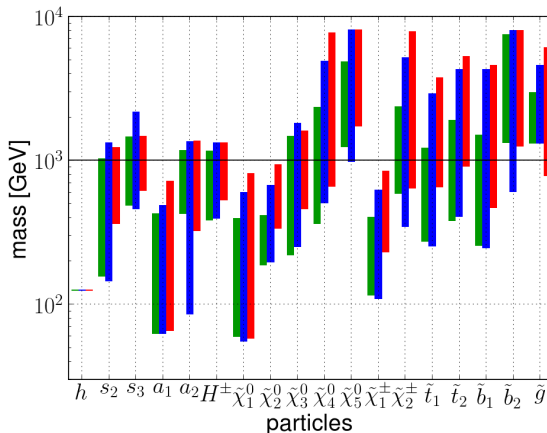


Figure: Particle spectrum

# Higgs signal strengths

- Signal strengths **always** deviate from complete SM-likeness.
- Decrease in  $R_{b\bar{b}}$   $\Rightarrow$  increase in  $R_{\gamma\gamma} \approx R_{ZZ}$ .
- Red points enhancement in  $R_{\gamma\gamma}$  due to large  $\lambda$  but also somewhat heavy  $\tilde{\chi}_1^+$  ( $g_{h\tilde{\chi}_1^+\tilde{\chi}_1^-} \approx \sqrt{2}\lambda V_{s,h}$  and  $V_{s,h} < 0$ ).
- **Interesting points** with small tuning and enhanced  $R_{\gamma\gamma}$  due to **light charginos** and  $\lambda \sim 1.1$  and  $\tan\beta \gtrsim 3$ ,  $m_{\tilde{\chi}_1^+} \approx 110$  GeV.

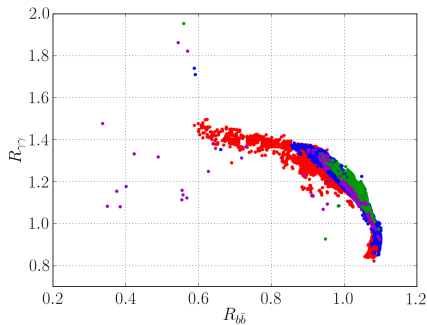
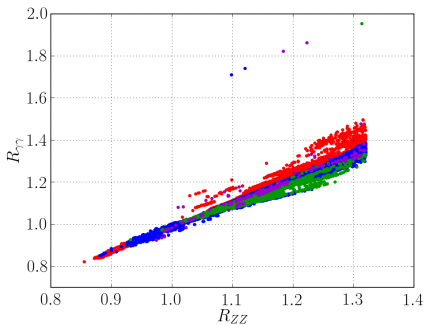


Figure:  $R_{\gamma\gamma}$  vs  $R_{ZZ}$

Figure:  $R_{\gamma\gamma}$  vs  $R_{b\bar{b}}$

# SUSY searches

- For total tuning better than 5 %,  $m_{\tilde{t}_1} \lesssim 1.3$  TeV,  $m_{\tilde{g}} \lesssim 2.5$  TeV,  $m_{\tilde{\chi}_1^0} \lesssim 400$  GeV.

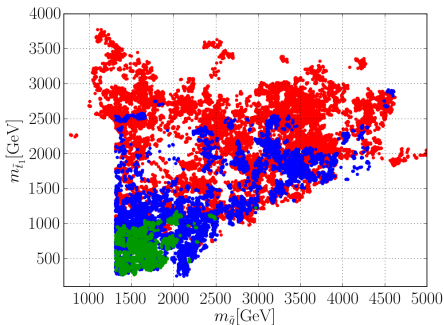


Figure:  $m_{\tilde{t}_1}$  [GeV] vs  $m_{\tilde{g}}$  [GeV].

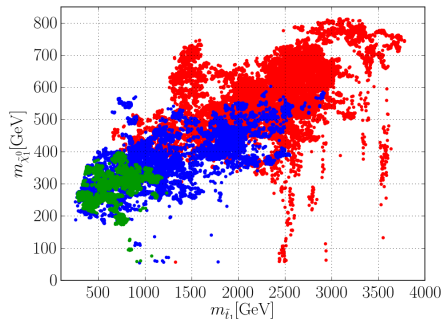


Figure:  $m_{\tilde{\chi}_1^0}$  [GeV] vs  $m_{\tilde{t}_1}$  [GeV].

# Cosmological Constraints and Dark Matter

- Most of the time  $\tilde{\chi}_1^0$  is **mostly singlino or higgsino**  $\Rightarrow$  **underproduced**.
- Colors: green (singlino), blue (higgsino), orange (wino), red (bino), purple (gravitino,  $m_{3/2} \simeq 0.01$  eV  $\Rightarrow \Omega_{3/2} h^2 \ll \Omega_{WMAP-9} h^2$ ).
- For points with  $\Omega_{DM} h^2 \approx 0.1$ , dark matter is **mostly Bino with small singlino component** (compensates with large  $\lambda$  and  $\kappa$ ).

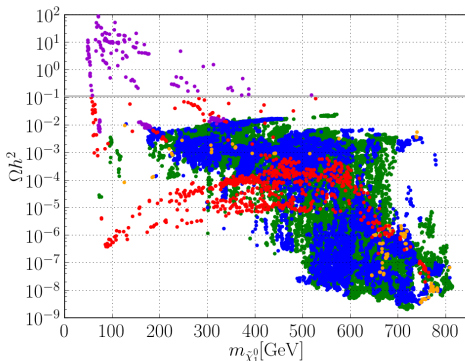


Figure:  $\Omega_{\tilde{\chi}_1^0} h^2$  vs  $m_{\tilde{\chi}_1^0}$  [GeV].

Dominant annihilation channels to get correct dark matter relic density,

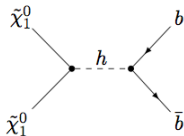


Figure: Resonant p-wave annihilation via  $h$  to bottom quarks

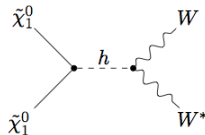


Figure: Resonant p-wave annihilation via  $h$  to  $W$  bosons

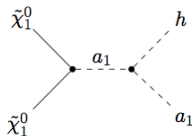


Figure: s-wave annihilation via lightest CP-odd  $a_1$

# Outline

- 1 The MSSM
  - Fine tuning in the MSSM
- 2 Scale Invariant NMSSM
  - Fine tuning in the scale invariant NMSSM
  - Phenomenology
- 3 Conclusions

# Conclusions

- A **natural Higgs** with a mass  $m_h \approx 126$  GeV in accordance with the latest LHC measurements can still be accomplished in SUSY by considering the second simplest extension of the SM, the **NMSSM**.
- Though large values of the parameter  $\lambda$  seem to help with the usual EW scale tuning, an **additional tuning** is induced in the effective Higgs quartic coupling which grows as  $\lambda^2$ .
- The **total tuning measure** defined as the product  $\Sigma^v \times \Sigma^h$  is minimized at  $\lambda \approx 1$  and can always be better than the tuning in the MSSM.
- For total tuning better than 5 %,  $m_{\tilde{t}_1} \lesssim 1.3$  TeV,  $m_{\tilde{g}} \lesssim 2.5$  TeV,  $m_{\tilde{\chi}_1^0} \lesssim 400$  GeV.
- **Enhancement in the two photon** Higgs discovery channel can be accomplished in regions of parameter space that have a and  $\tan \beta \gtrsim 3$  with moderate tuning.
- A **dark matter** candidate can be obtained with a dominant **Bino component and a small but non-vanishing singlino** component.