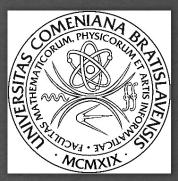
New Approach to an Effective Density Dependence within an Advanced Relativistic Mean-Field Theory

Kristian Petrík Štefan Gmuca











MOTIVATION



MASSIVE PULSARS OBSERVED

- masses around 2Msol
- precise measurements

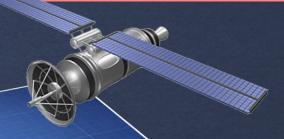
P. B. Demorest et al., **Nature** 467, 1081 (2010) J. Antoniadis et al., **Science** 340, 448 (2013)

DENSITY DEPENDENCE

- *not exactly known* for a nuclear interaction
- huge gap between compact stars and nuclei

HYPERON PUZZLE

- *hyperons* generally **soften** the equation of state
- many models with exotic matter (hyperons, kaon condensates, ...) are ruled out





EFFECTIVE MODELS

- simplicity vs fundamentality
- generally better results but with a "black box" like understanding

EFFECTIVE FIELD THEORY



- **Ab initio** approaches
 - microscopic treatment
 - parameter free EoS predictions
 - based on precise NN interaction

Equation of state

Phenomenological approaches

- relatively easy calculations
- very good results
- description of systems not reachable by ab initio models

EFFECTIVE FIELD THEORY



- **Ab** initio approaches
 - microscopic treatment
 - parameter free EoS predictions
 - based on precise NN interaction

Equation of state

Phenomenological approaches

- relatively *easy* calculations
- very good results
- description of systems not reachable by ab initio models

DIRAC **RELATIVISTIC** BRUECKNER MEDIUM EFFECTS

HARTREE DIRECT TERMS

FOCK **EXCHANGE TERMS** CORRELATIONS

EVERYTHING ELSE

EFFECTIVE Black Box MODEL Everything is hidden in parameters

Equation of state

EFFECTIVE FIELD THEORY



ADVANCED EFFECTIVE MODEL

DIRAC **RELATIVISTIC** BRUECKNER

MEDIUM EFFECTS

HARTREE

DIRECT TERMS

FOCK

EXCHANGE TERMS

CORRELATIONS

EVERYTHING ELSE

DENSITY DEPENDENCE

New form of a density dependence

EFFECTIVE MODEL

FOCK TERMS

Understanding exchange interaction

Equation of state

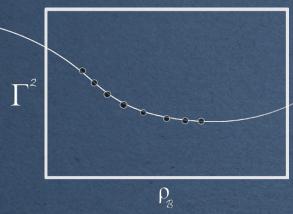
2 PARAMETRIC FUNCTIONS



■ Typical choice of a density dependence

$$\Gamma_i(\rho_B) = a_i \left[\frac{1 + b_i(\rho_B/\rho_{sat} + d_i)^2}{1 + c_i(\rho_B/\rho_{sat} + e_i)^2} \right],$$

Dirac-Brueckner-Hartree-Fock



- ambiguities
- no stable extrapolations
- too many free parameters

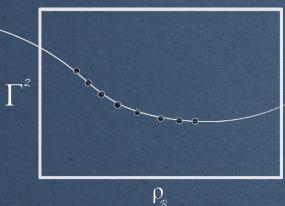
2 PARAMETRIC FUNCTIONS



Typical choice of a density dependence

$$\Gamma_i(\rho_{\mathcal{B}}) = \alpha_i \left[\frac{1 + b_i(\rho_{\mathcal{B}}/\rho_{sat} + d_i)^2}{1 + c_i(\rho_{\mathcal{B}}/\rho_{sat} + e_i)^2} \right],$$

Dirac-Brueckner-Hartree-Fock



- ambiguities
- no stable extrapolations
- too many free parameters

Other functional forms

4-parametric class

$${}^{A}\Gamma_{i}^{\bullet\bullet}(x) = a_{i} \left[\frac{c_{i} + (x + d_{i})^{2}}{e_{i} + (x)^{2}} \right],$$

$${}^{B}\Gamma_{i}^{\bullet\bullet}(x) = a_{i} \left[\frac{c_{i} + (x + d_{i})^{1/2}}{e_{i} + (x)^{1/2}} \right],$$

3-parametric class

$${}^{C}\Gamma_{i}^{\circ\bullet}(x) = a_{i} \left[\frac{1 + (x + d_{i})^{2}}{(x + c_{i})^{2}} \right],$$

$${}^{D}\Gamma_{i}^{\circ\bullet}(x) = a_{i} \left[\frac{1 + (x + d_{i})^{2}}{c_{i} + (x)^{2}} \right],$$

$${}^{E}\Gamma_{i}^{\circ\bullet}(x) = a_{i} \left[\frac{1 + (x + d_{i})^{2}}{c_{i} + (x + c_{i})^{2}} \right],$$

P

compact stars

nuclei

2 parametric class

2-parametric class

$$|\Gamma_i^{\circ\circ}(x)| = \alpha_i \left[\frac{\beta_i + x^2}{b_i + x^2} \right],$$

$$|\Gamma_i^{\circ\circ}(x)| = \alpha_i \left[\frac{\beta_i + x^{3/2}}{b_i + x^{3/2}} \right],$$

$$|\beta_i^{\circ\circ}(x)| = \beta_i + x^2 \Big]^{-1}$$

$$|\tilde{\Gamma}_{i}^{\circ\circ}(x)| = a_{i} \left[\frac{\beta_{i} + x^{2}}{b_{i} + x^{2}} \right]^{-1},$$

$$|\tilde{\Gamma}_{i}^{\circ\circ}(x)| = a_{i} \left[\frac{\beta_{i} + x^{3/2}}{b_{i} + x^{3/2}} \right]^{-1},$$

$$x = \frac{\rho_{\mathcal{B}}}{\rho_{sat}} \quad |\Gamma_{\delta}^{\circ \bullet}(x)| = a_{\delta} \left[\frac{\beta_{\delta} + (x + c_{\delta})^{2}}{b_{\delta} + x^{2}} \right],$$

$$||\Gamma_{\delta}^{\circ \bullet}(x)| = a_{\delta} \left[\frac{\beta_{\delta} + (x + c_{\delta})^{3/2}}{b_{\delta} + x^{3/2}} \right].$$

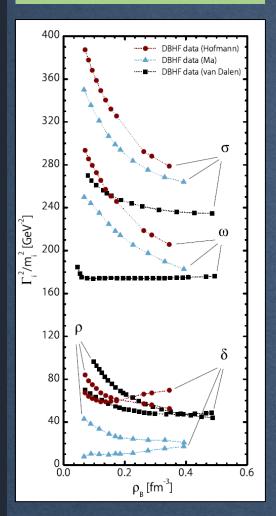
- 2+ parameters
- stable extrapolations
 - great universality

1

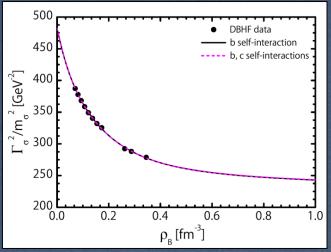
2 PARAMETRIC FUNCTIONS



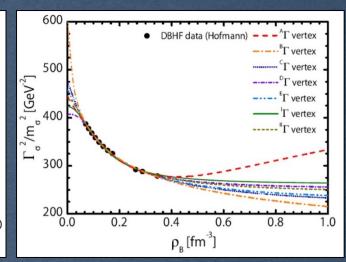
DBHF vertices



• *Single* density dependence



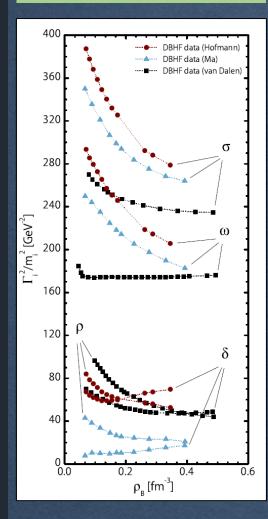
• *Several* density dependences



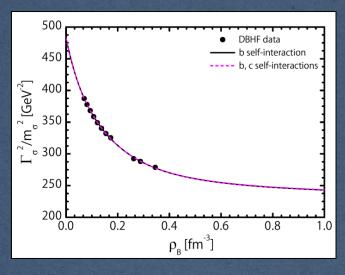
1 2 PARAMETRIC FUNCTIONS



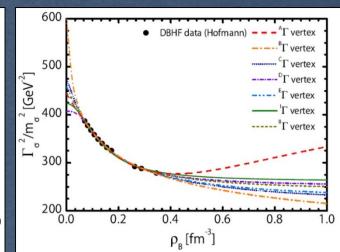




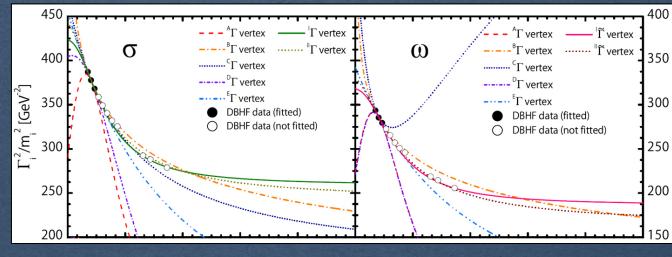
• Single density dependence



• Several density dependences



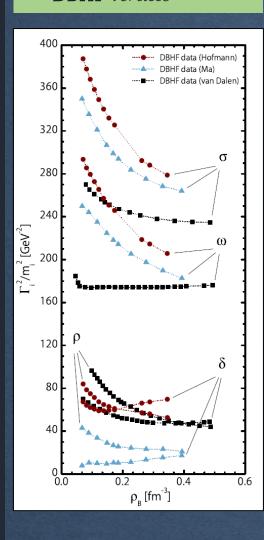
partial fits

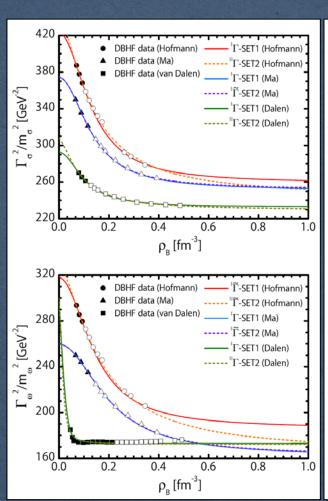


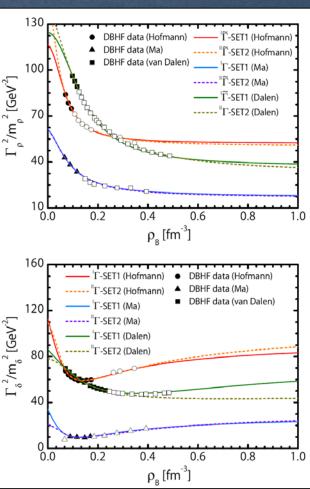


DBHF vertices

partial fits

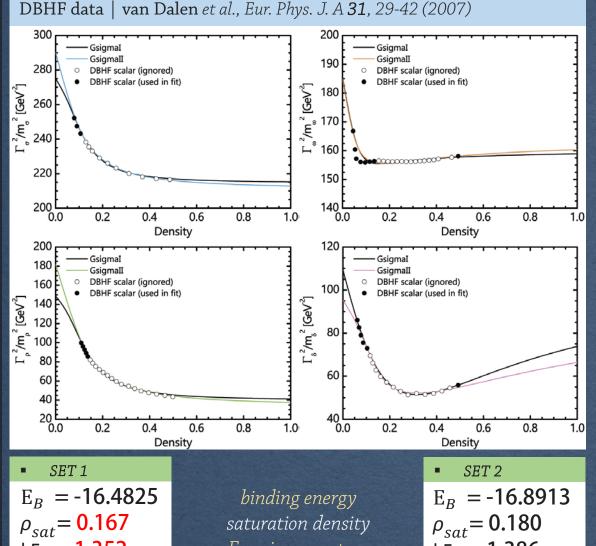


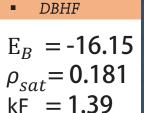


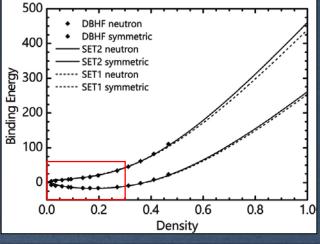


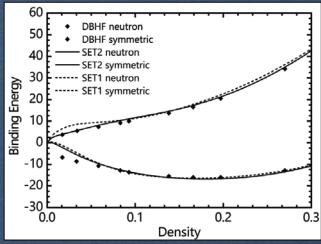
2 PARAMETRIC FUNCTIONS - RESULTS











kF = 1.352= 205

= 33.4

Fermi momentum compression modulus symmetry energy

kF = 1.386= 236 $a_s = 35.0$

MAPPING OF FOCK TERMS



• Lagrangian density for DHF model

$$\mathcal{L}_{HF} = \mathcal{L}^{fermi} + \mathcal{L}^{bose} + \mathcal{L}^{int},$$

$$\begin{split} \mathcal{L}^{\text{fermi}} &= \quad \bar{\psi} \left[i \gamma^{\mu} \partial_{\mu} - m_{N} \right] \psi, \\ \mathcal{L}^{\text{bose}} &= \quad \frac{1}{2} \partial_{\mu} \hat{\sigma} \partial^{\mu} \hat{\sigma} - \frac{1}{2} m_{\sigma}^{2} \hat{\sigma}^{2} - \frac{1}{4} \hat{\omega}_{\mu\nu} \hat{\omega}^{\mu\nu} + \frac{1}{2} m_{\omega}^{2} \hat{\omega}_{\mu} \hat{\omega}^{\mu} \\ &\quad + \frac{1}{2} \partial_{\mu} \hat{\pi} \partial^{\mu} \hat{\pi} - \frac{1}{2} m_{\pi}^{2} \hat{\pi}^{2} \\ \mathcal{L}^{\text{int}} &= \quad \bar{\psi} \left[g_{\sigma} \hat{\sigma} - g_{\omega} \hat{\omega}_{\mu} \gamma^{\mu} - \frac{f_{\pi}}{m_{\pi}} \gamma^{5} \gamma^{\mu} \partial_{\nu} \boldsymbol{\pi} \cdot \boldsymbol{\tau} \right] \psi, \end{split}$$

Nuclear self energy

$$\Sigma(k, k_F) = \Sigma^{s}(k, k_F) - \gamma^{0} \Sigma^{0}(k, k_F) + \boldsymbol{\gamma} \cdot \boldsymbol{k} \Sigma^{\vee}(k, k_F)$$



- exchange **Fock** terms taken into account
- pion contributions do not vanish
- significant enhancement of the maximum mass
- each coupling depends on a particular density



- density dependence of Fock part is not so strong
- in medium effects still **not covered**

DIRAC RELATIVISTIC

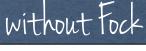
HARTREE DIRECT TERMS

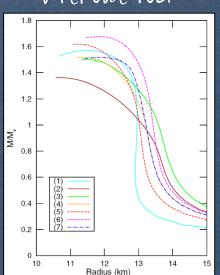
FOCK

EXCHANGE TERMS

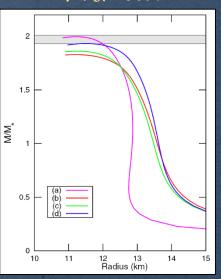
EFFECTIVE MODEL

Dark Gray Box





with Fock



MAPPING OF FOCK TERMS

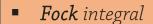


Nuclear self energy

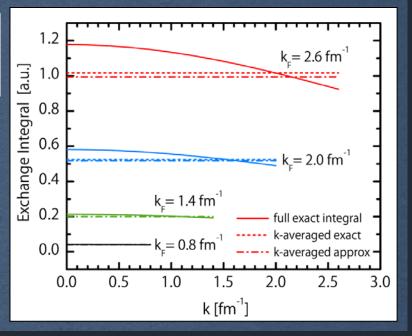
$$\Sigma(k, k_F) = \Sigma^{S}(k, k_F) - \gamma^{0} \Sigma^{0}(k, k_F) + \boldsymbol{\gamma} \cdot \boldsymbol{k} \Sigma^{\vee}(k, k_F)$$

$$\Sigma^{0}(k,k_{F}) = -\frac{\gamma}{2\pi^{2}} \frac{g_{\omega}^{2}}{m_{\omega}^{2}} \int_{0}^{k_{F}} q^{2}dq \quad \text{Hartree}$$

$$-\frac{1}{(4\pi)^{2}k} \int_{0}^{k_{F}} qdq \left[g_{\sigma}^{2}\Theta_{\sigma}(k,q) + 2g_{\omega}^{2}\Theta_{\omega}(k,q) - (\gamma - 1) \left(\frac{f_{\pi}}{m_{\pi}} \right)^{2} m_{\pi}^{2}\Theta_{\pi}(k,q) \right],$$



$$I_{i}(k, k_{F}) = \frac{1}{k} \int_{0}^{k_{F}} q dq \ln \frac{(k+q)^{2} + m_{i}^{2}}{(k-q)^{2} + m_{i}^{2}}.$$



- **Expand** the exact integral in k and keep the terms up to k^2
- Average it over the Fermi sphere

Approximate k-averaged exchange integral

$$\bar{I}_i(k_F) = 4\left(k_F - m_i \arctan\left(\frac{k_F}{m_i}\right) - \frac{k_F^5}{5(k_F^2 + m_i^2)^2}\right).$$

$$X(x) = 3\left(x - \arctan(x) - \frac{x^5}{5(1+x^2)^2}\right)/x^3$$

MAPPING OF FOCK TERMS



New density dependent couplings

• Scalar channel

$$\frac{\Gamma_{\sigma}^2(k_F)}{m_{\sigma}^2} = \frac{g_{\sigma}^2}{m_{\sigma}^2} + \sum_{i=\{\sigma,\omega,\pi\}} b_i \frac{g_i^2}{m_i^2} X(k_F/m_i).$$

direct part

exchange part

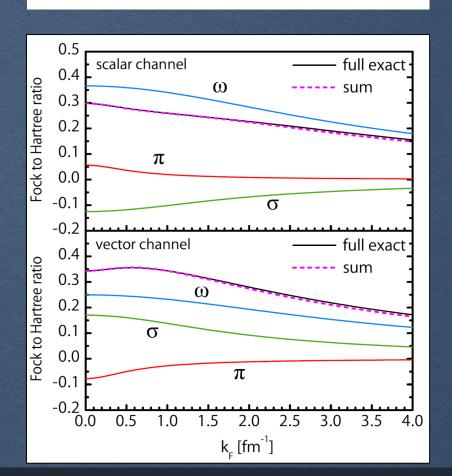
mixing

Dependence on a **specific** nuclear **density**

$$\Sigma^{j}(k_F) = \frac{\Gamma_j^2(k_F)}{m_j^2} \rho_j(k_F)$$

• **Vector** channel

$$\frac{\Gamma_\omega^2(k_F)}{m_\omega^2} = \frac{g_\omega^2}{m_\omega^2} + \sum_{i=\{\sigma,\omega,\pi\}} \alpha_i \frac{g_i^2}{m_i^2} X(k_F/m_i),$$



PERSPECTIVE



1 DENSITY DEPENDENCE

2 parametric density functions

- less ambiguous DBHF data
- **2 parametric functions** with a dependence on **all** nuclear densities?
- **hyperons** scaling?

FOCK TERMS

exchange interaction

- in medium correlations still play an important role
- **an improvement** through **non-linear** self interactions?

