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Compact stars in minimal dilatonic gravity



The Structure and Signals of Neutron Stars, from Birth to Death
Florence, 28. 03. 2014

The basic lesson from cosmology:

GR and the Standard Particles Model are not enough !

- One may add some new content: Dark matter, Dark energy
- One may modify GR: simplest modifications are $F(R)$ and MDG
- Some combination of the above possibilities may work ?

What about the Star Physics ???

Do we have the same problem :

EOS versus gravity modification???

Maybe we need both possibilities
simultaneously ???

Plan of the talk:

- Minimal dilatonic gravity (MDG)
- The basic equations of SSSS in MDG
- The boundary conditions for SSSS in MDG
- Neutron SSSS with simplest EOS in MDG

SSSS = **S**tatic **S**pherically **S**ymmetric **S**tars

Minimal dilatonic gravity (MDG)

$$\mathcal{A}_{g,\Phi} = \frac{c}{2\kappa} \int d^4x \sqrt{|g|} (\Phi R - 2\Lambda U(\Phi))$$

O'Hanlon: PRL, 1972,

PPF: Mod. Phys. Lett. 2000; gr-qc/0202074; PRD 2013.

Brans-Dicke with $\omega = 0$

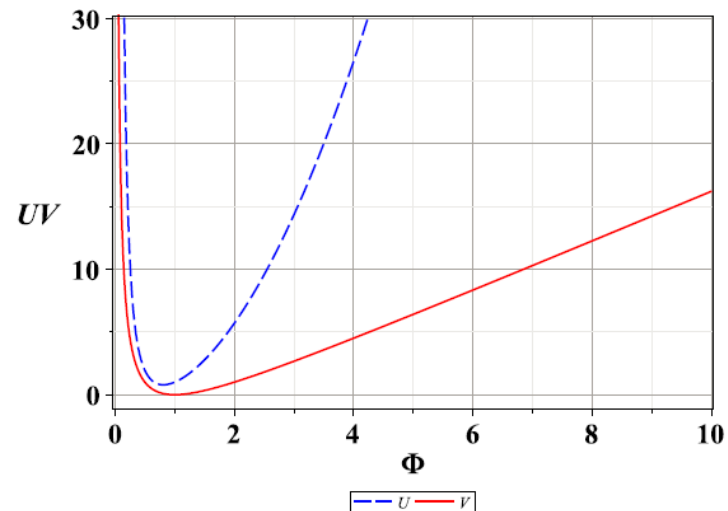
Locally equivalent to $f(R)$: $\mathcal{A}_{f(R)} = \frac{c}{2\kappa} \int d^4x \sqrt{|g|} f(R)$

Withholding potentials:

PPF: PRD 2013:

No ghosts! No tachions!

$$V'(\Phi) = \frac{2}{3} (\Phi U'(\Phi) - 2U(\Phi))$$



$$m' = 4\pi r^2 \epsilon_{eff} / \Phi,$$

Generalized TOV

$$\epsilon_{eff} = \epsilon + \epsilon_\Lambda + \epsilon_\Phi,$$

$$\Phi' = -4\pi r^2 p_\Phi / \Delta,$$

$$p_{eff} = p + p_\Lambda + p_\Phi$$

$$p'_\Phi = -\frac{p_\Phi}{r\Delta} \left(3r - 7m - \frac{2}{3}\Lambda r^3 + 4\pi r^3 \frac{\epsilon_{eff}}{\Phi} \right) - \frac{2}{r}\epsilon_\Phi,$$

$$p' = -\frac{p + \epsilon}{r} \frac{m + 4\pi r^3 p_{eff} / \Phi}{\Delta - 2\pi r^3 p_\Phi / \Phi},$$

$$\Delta(r) = r - 2m(r) - \frac{1}{3}\Lambda r^3$$

ALL NOVEL

$$\epsilon_\Lambda = \frac{\Lambda}{8\pi} \left(U(\Phi) - \Phi \right)$$

$$p_\Lambda = -\frac{\Lambda}{8\pi} \left(U(\Phi) - \frac{1}{3}\Phi \right)$$

$$V'(\Phi) = \frac{2}{3} \left(\Phi U'(\Phi) - 2U(\Phi) \right)$$

$$\epsilon_\Lambda = -p_\Lambda - \frac{\Lambda}{12\pi} \Phi; \quad \leftarrow \text{CEOS}$$

$$\epsilon_\Phi = p - \frac{1}{3}\epsilon + \frac{\Lambda}{8\pi} V'(\Phi) \leftarrow \text{DEOS}$$

$$+ \frac{p_\Phi}{2} \frac{m + 4\pi r^3 p_{eff} / \Phi}{\Delta - 2\pi r^3 p_\Phi / \Phi};$$

$$\epsilon = \epsilon(p) \quad \leftarrow \text{MEOS}$$

Assuming:

$$r_c = 0$$

$$m(0) = m_c = 0, \quad \Phi(0) = \Phi_c, \quad p(0) = p_c,$$

$$p_\Phi(0) = p_{\Phi_c} = \frac{2}{3} \left(\frac{\epsilon(p_c)}{3} - p_c \right) - \frac{\Lambda}{12\pi} V'(\Phi_c).$$

SSSS edge:

$$p = 0 \Rightarrow r^*$$

$$m^* = m(r^*; p_c, \Phi_c), \quad \Phi^* = \Phi(r^*; p_c, \Phi_c),$$

$$p_\Phi^* = p_\Phi(r^*; p_c, \Phi_c).$$

Cosmological horizon:

$$r_U$$

$$r \in [r^*, r_U] \quad p \equiv 0 \text{ and } \epsilon \equiv 0$$

$$r_U: \Delta(r_U; p_c, \Phi_c) = 0,$$

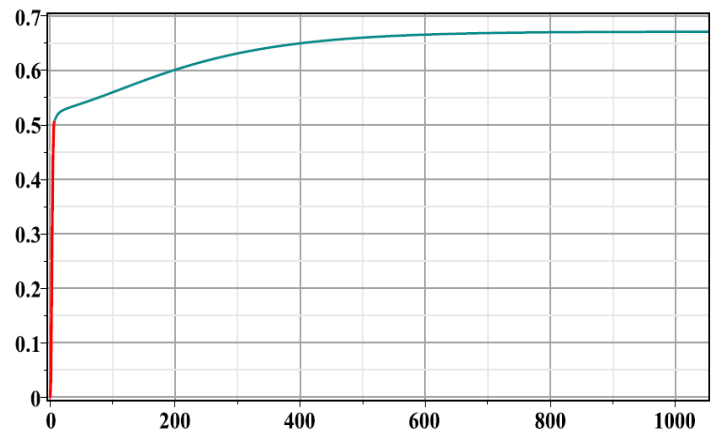
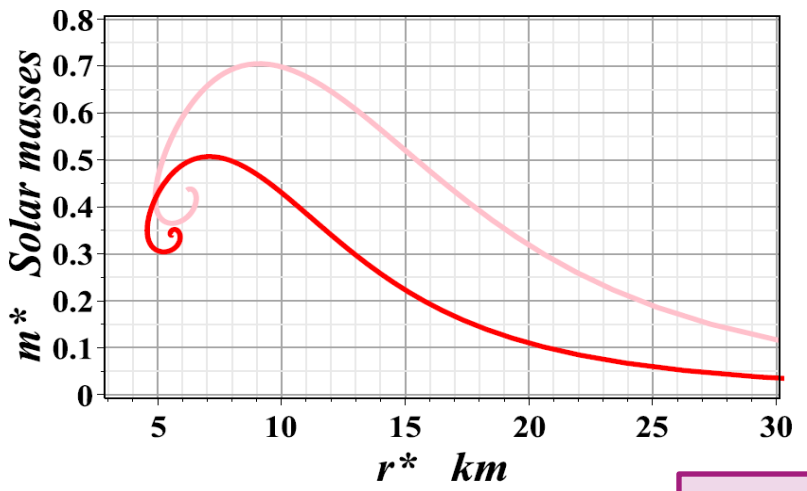
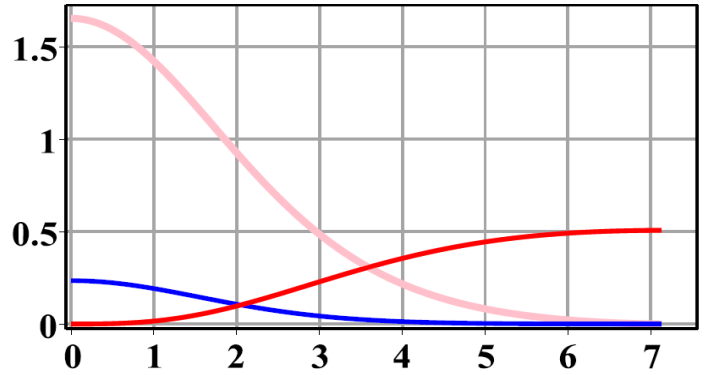
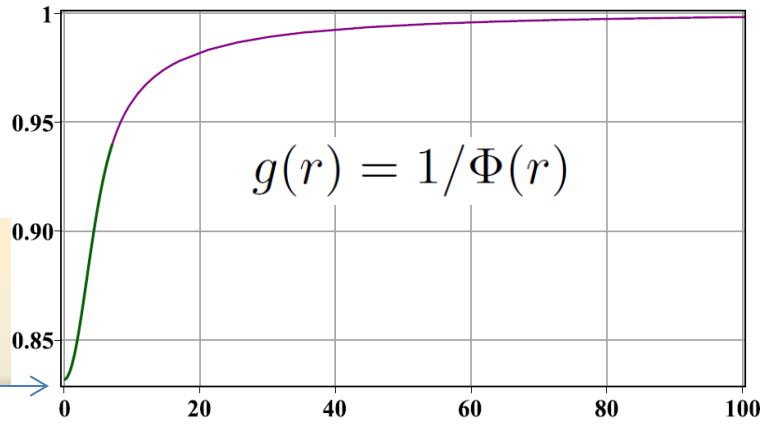
$$\Phi(r_U; p_c, \Phi_c) = 1 \quad \leftarrow \text{De Sitter vacuum}$$

$$F_\Phi(p_{\Phi_c}, p_c, \Phi_c) = 0, \quad F_\Lambda(p_c, \Phi_c) = 0, \quad \leftarrow \text{Two specific MDG relations}$$

One parametric (p_c) family of SSSS – as in GR and the Newton gravity !

$$\epsilon = \frac{1}{4\pi} (\sinh t - t), \quad p = \frac{1}{12\pi} (\sinh t - 8 \sinh(t/2) + 3t).$$

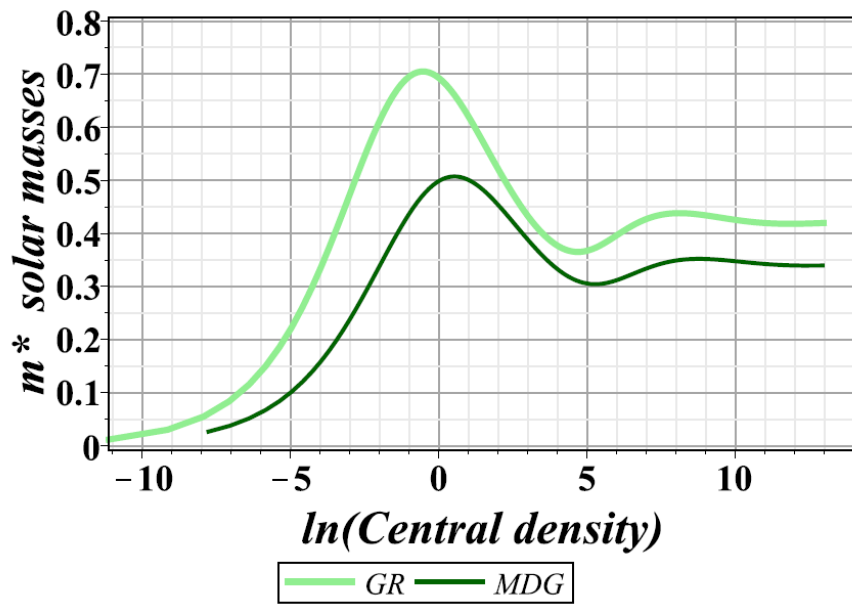
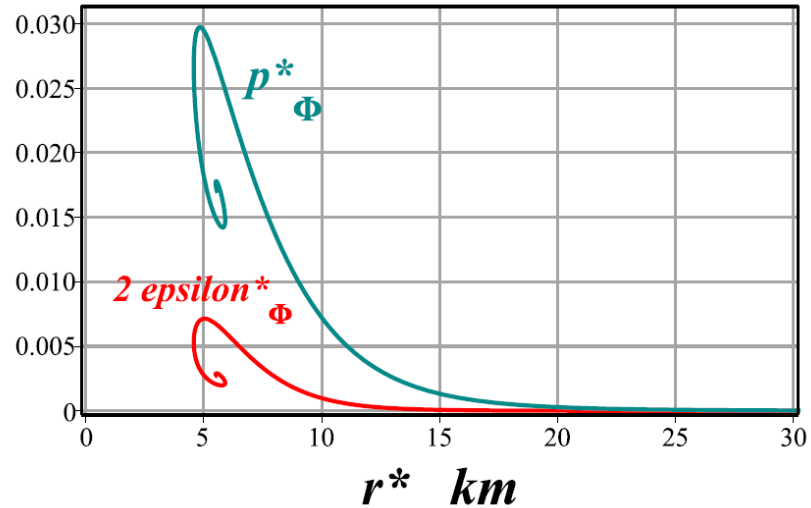
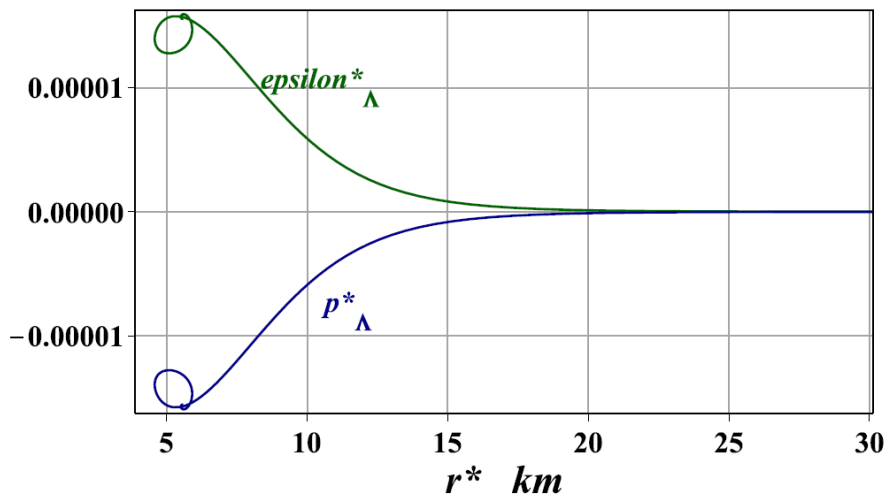
17% weaker gravity



$m_{total} \approx 0.6710 M_{\odot}$

GR MDG

m(r) m_disph(r)



In MDG we have the same stability properties of SSSS as in GR

Conclusion

Most probably we need to look
simultaneously and coherently

for **a realistic EOS**

and

for **a realistic withholding cosmological potential**

which are able to describe the variety of

**cosmological, astrophysical, gravitational and star
phenomena**

at different scales.

**MICHELANGELO Buonarroti, 1511:
Separation of Light from Darkness**



Thank you !