





# P.S. Shternin, M. Baldo, P. Haensel

# Transport coefficients of nuclear matter in neutron star cores in the BHF framework

PRC88,065803 (2013)

The Structure and Signals of Neutron Stars, from Birth to Death Firenze, Italy, 24-28 March, 2014

# Introduction. Kinetic coefficients in NS cores



Non-superfluid beta-stable npeµ matter

Kinetic coefficients due to particle collisions

$$\kappa = \kappa_{e\mu}[ee, e\mu, ep] + \kappa_n[nn, np]$$
$$\kappa_p \ll \kappa_n$$

Electromagnetic part:  $\kappa_{e\mu}, \eta_{e\mu}$ 

Shternin, Yakovlev, 2007,2008

Present talk: nucleon sector

 $\kappa_{\rm n}, \eta_{\rm n}$ 

Strongly interacting multicomponent Fermi-liquid

Important: many-body effects in scattering

**Controversial results by different groups** 

- Thermal conductivity and shear viscosity in multicomponent Fermi-liquid
- Nucleon interaction in Brueckner-Hartree-Fock approximation
- Results
- Comparison with other works
- Conclusions

# Kinetic coefficients in multi-component Fermi-liquid: Formalism

$$\kappa = \sum_{c} \frac{\pi^2 k_B^2 T n_c \tau_c^{\kappa}}{3m_c^*}; \quad \eta = \sum_{c} \frac{n_c p_{\mathrm{F}c}^2 \tau_c^{\eta}}{5m_c^*}$$

**Perturbation**  $\nabla T, V_{\alpha\beta} = \frac{1}{2} \left( \frac{\partial V_{\alpha}}{\partial x_{\beta}} + \frac{\partial V_{\beta}}{\partial x_{\alpha}} \right) \implies$ 

 $\begin{array}{ll} \mbox{Deviation of the distribution} \\ \mbox{function} & F = f_{eq} - \Phi \frac{\partial f_{eq}}{\partial \epsilon} \end{array}$ 

$$\begin{array}{lll} \text{Kinetic equation} & \kappa : & (\epsilon_1 - \mu_1) \mathbf{v}_1 \frac{\nabla T}{T} \\ \text{(linearized)} & \eta : & \left( v_{1\alpha} p_{1\beta} - \frac{1}{3} \delta_{\alpha\beta} v_1 p_1 \right) V_{\alpha\beta} \end{array} \right\} \frac{\partial f_1}{\partial \epsilon_1} = \sum_i I_{ci}(12; 1'2'), \quad c, \ i = n, \ p = n$$

#### **Boltzmann collision integral**

$$I_{ci} = \frac{1}{(1+\delta_{ci})k_BT} \sum_{\sigma_{1'}\sigma_2\sigma_{2'}} \int \int \int \frac{\mathrm{d}\boldsymbol{p}_{1'}\mathrm{d}\boldsymbol{p}_2\mathrm{d}\boldsymbol{p}_{2'}}{(2\pi\hbar)^9} w_{ci}(12;1'2')f_1f_2(1-f_{1'})(1-f_{2'}) \left(\Phi_{1'}+\Phi_{2'}-\Phi_1-\Phi_2\right)$$
  
Transition probability 
$$\sum_{\mathrm{spins}} w_{ci}(12|1'2') = 4\frac{(2\pi)^4}{\hbar}\delta(\epsilon_1+\epsilon_2-\epsilon_1'-\epsilon_2')\delta(\boldsymbol{P}-\boldsymbol{P}')\mathcal{Q}_{ci}$$

Solution: 
$$\Phi^{\kappa,\eta} \to \tau_c^{\kappa,\eta} \to \kappa, \eta$$

Input: 
$$m^*, \ \mathcal{Q}$$
 on the Fermi surface

# Variational solution

2x2 algebraic system

$$\sum_{i=n,p} \nu_{ci} \tau_i =$$

 $\frown$ 

$$\begin{aligned} & \mathbf{2x2 \ algebraic \ system} \quad \sum_{i=n,p} \nu_{ci} \tau_i = 1 & \kappa = \sum_{c} - \nu_{ci} \tau_i = 1 \\ & \nu_{ci}^{(\kappa)} = \frac{64m_c^* m_i^{*2} (k_B T)^2}{5m_N^2 \hbar^3} S_{\kappa ci}, \quad \nu_{ci}^{(\eta)} = \frac{16m_c^* m_i^{*2} (k_B T)^2}{3m_N^2 \hbar^3} S_{\eta ci}. \end{aligned}$$

$$=\sum_{c}\frac{\pi^{2}k_{B}^{2}Tn_{c}\boldsymbol{\tau_{c}^{\kappa}}}{3m_{c}^{*}};\quad \eta=\sum_{c}\frac{n_{c}p_{\mathrm{F}c}^{2}\boldsymbol{\tau_{c}^{\eta}}}{5m_{c}^{*}}$$

#### All particles on Fermi surface

$$\begin{aligned} \mathbf{Effective transport cross-sections} \\ S_{\kappa cc} &= \frac{m_N^2}{128\pi^2\hbar^4 p_c^3} \int_0^{2p_c} \mathrm{d}P \int_0^{q_m(P)} \mathrm{d}q \frac{(4p_c^2 - P^2)}{\sqrt{q_m^2 - q^2}} \mathcal{Q}_{cc}, \\ S_{\kappa ci} &= \frac{m_N^2}{128\pi^2\hbar^4 p_c^3} \int_{|p_c - p_i|}^{p_c + p_i} \mathrm{d}P \int_0^{q_m(P)} \mathrm{d}q \frac{(4p_c^2 + q^2)}{\sqrt{q_m^2 - q^2}} \mathcal{Q}_{ci}, \quad c \neq i, \\ S_{\eta cc} &= \frac{3m_N^2}{128\pi^2\hbar^4 p_c^5} \int_0^{2p_c} \mathrm{d}P \int_0^{q_m(P)} \mathrm{d}q \frac{q^2(4p_c^2 - P^2 - q^2)}{\sqrt{q_m^2 - q^2}} \mathcal{Q}_{cc}, \\ S_{\eta ci} &= \frac{3m_N^2}{128\pi^2\hbar^4 p_c^5} \int_{|p_c - p_i|}^{p_c + p_i} \mathrm{d}P \int_0^{q_m(P)} \mathrm{d}q \frac{q^2(4p_c^2 - q^2)}{\sqrt{q_m^2 - q^2}} \mathcal{Q}_{ci}, \quad c \neq i, \end{aligned}$$

#### **Compare:** Scattering cross-sections:

$$\frac{d\sigma_{\alpha\beta}}{d\Omega} = \frac{m_{\alpha\beta}^{*2}}{16\pi^2\hbar^4} \mathcal{Q}_{\alpha\beta}$$

Accuracy (aposteriori)

 $\frac{\kappa_{exact}}{\kappa_{var}} \approx 1.2$  $\frac{\eta_{exact}}{\eta_{var}} \approx 1.05$ 

$$p_{1}$$
  $p_{2}$   $p_{2}$   $p_{2}$   $p_{2}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{1}$   $p_{1}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{1}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{1}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{1}$   $p_{2}$   $p_{1}$   $p_{2}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{2}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{1}$   $p_{2}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{1}$   $p_{2}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{3}$   $p_{1}$   $p_{2}$   $p_{3}$   $p_{3$ 

Two angles fix all momenta

 $\frac{1}{\tau_i} = \frac{1}{\tau_i} \left[ \langle w(12|1'2')\beta(\theta,\phi) \rangle \right]$ 

5

## **Approximate estimates**



# **Effects of the proton fraction**

PNM result is inaccurate even at small proton fraction



$$\kappa = \kappa_{\rm n} + \kappa_{\rm p}$$
$$\eta = \eta_{\rm n} + \eta_{\rm p}$$



# **Interactions in-medium**

$$\kappa = \sum_{c} \frac{\pi^{2} k_{B}^{2} T n_{c} \tau_{c}^{\kappa}}{3m_{c}^{*}}; \quad \eta = \sum_{c} \frac{n_{c} p_{F_{c}}^{2} \tau_{c}^{\eta}}{5m_{c}^{*}}$$
$$\sum_{i=n,p} \nu_{ci} \tau_{i} = 1$$
$$\nu_{ci}^{(\kappa)} = \frac{64m_{c}^{*} m_{i}^{*2} (k_{B}T)^{2}}{5m_{N}^{2} \hbar^{3}} S_{\kappa ci}, \qquad \nu_{ci}^{(\eta)} = \frac{16m_{c}^{*} m_{i}^{*2} (k_{B}T)^{2}}{3m_{N}^{2} \hbar^{3}} S_{\eta ci}.$$
From many-body theory we need
$$m^{*}, \quad Q$$

# 4<sup>th</sup> power of m<sup>\*</sup> – the main effect?

#### **Bethe-Brueckner-Goldstone equation**

 $V \to G$ 

All diagrams on two hole-line level are included

$$\langle p_1 p_2 | G^{\alpha\beta}(\omega) | p_3 p_4 \rangle = \langle p_1 p_2 | V^{\alpha\beta} | p_3 p_4 \rangle + \sum_{k_1, k_2} \langle p_1 p_2 | V^{\alpha\beta} | k_1 k_2 \rangle \frac{Q^{\alpha\beta}(k_1, k_2)}{\omega - \epsilon_{\alpha}(k_1) - \epsilon_{\beta}(k_2)} \langle k_1 k_2 | G^{\alpha\beta} | p_3 p_4 \rangle$$

$$Q^{\alpha\beta}(k_1, k_2) \quad \text{- Pauli operator} \qquad \epsilon_{\alpha}(p) = \frac{p^2}{2m_{\alpha}} + U_{\alpha}(p)$$

Continuous choice of auxiliary single-particle potential

$$U_{\alpha}(p_1) = \sum_{\beta; p_2 < p_{F\beta}} \langle p_1 p_2 | G^{\alpha\beta}(\epsilon_1(p_1) + \epsilon_2(p_2)) | p_1 p_2 \rangle_{\mathcal{A}}$$

*minimizes the contribution from the three-hole line diagrams* 

**BHF energy:** 
$$E = E_{kin} + \frac{1}{2} \sum_{\alpha\beta} \sum_{p_1 < p_{F\alpha}, p_2 < p_{F\beta}} \langle p_1 p_2 | G^{\alpha\beta}(\epsilon_1(p_1) + \epsilon_2(p_2)) | p_1 p_2 \rangle_A, \quad \alpha, \beta = n, p_1 \langle p_1 p_2 | G^{\alpha\beta}(\epsilon_1(p_1) + \epsilon_2(p_2)) | p_1 p_2 \rangle_A,$$

#### **Three-nucleon interactions**

$$V = V_{12} + V_{12}^{(3)}$$

Effective interaction (averaged over the third particle)

$$V_{12}^{(3)} = \rho \int \mathrm{d}\boldsymbol{r}_3 \, g^2(r_{13}) g^2(r_{23}) V_{123}$$

$$g(r)$$
 – defect function

# **Results. Equation of state**

#### Argonne v18 + Urbana IX



3bf produce correct s.p. of SNM EOS is stiffer

#### **Proton fraction in beta-stable matter**



# **Results. Effective mass**

# Fermi-sea of quasi-particles $\epsilon(p)$ interacting via G

$$m^* = \left(\frac{1}{p} \frac{\mathrm{d}\epsilon(p)}{\mathrm{d}p}\right)_{p=p_F}^{-1}$$

2bf decrease effective masses 3bf increase



NB: Fermi-liquid effects are not included

#### **Results. In-medium cross-sections**



# **Results. Kinetic coefficients.**

Exact solutions are shown



Urbana IX 3bf results are comparable with 'free-scattering'

m\* modification only is insufficient

# **Comparison with other works**

Carbone&Benhar, J.Phys.Conf.Ser. 336, 012015 (2011)

Zhang, Lombardo, Zuo, PRC 82, 015805 (2010)



#### Strong disagreement!

#### CB11: CBF approach, no 3bf

ZLZ10: BHF approach, different 3bf, rearrangement in m\*

• Kinetic coefficients of neutrons in neutron star cores are calculated in the BHF framework with account for the effective three-nucleon interactions

• Both effective mass and the intermediate state blocking by Pauli principle are important

Three-nucleon forces can lead to lower kinetic coefficients

• Many-body effects in the present model do not result in significant change of kinetic coefficients

#### Var0 – carriers are neutrons Var – solution of 2x2 system



#### In-vacuum cross-sections

#### **Comparison with experiments**



## pp cross-section with Coulomb interaction (Mott) included



Approximation for the pure neutron matter

$$\kappa_{\text{Var}} = \frac{5}{3} \frac{p_F^3}{m^2} \frac{1}{T} \left[ 32\sigma_{\text{tot}} \left( E_{\text{cm}} = \frac{p_F^2}{m} \right) \right]^{-1} \qquad \qquad \eta_{\text{Var}} = \frac{p_F^5}{16\pi^2 m^2} \frac{1}{(k_B T)^2} \left[ \sigma_{\text{tot}} \left( E_{\text{cm}} = \frac{p_F^2}{m} \right) \right]^{-1}$$





Baym, Pethick "Landau Fermi-Liquid Theory: concepts and applications, 1991"

Characteristic relaxation time  $\tau \equiv \frac{8\pi^4\hbar^6}{m^{*3}\langle W \rangle (k_B T)^2}$ Where  $W = \frac{\pi}{\hbar} \sum_{\sigma_2 \sigma_3 \sigma_4}' |\langle 34|t|12 \rangle|^2 = \pi Q$ Benhar, Valli:  $W = \frac{16\pi^2}{m^{*2}} \frac{d\sigma}{d\Omega}$ Zhang e

Factor of  $\pi$ 





# **Kinematics**

