



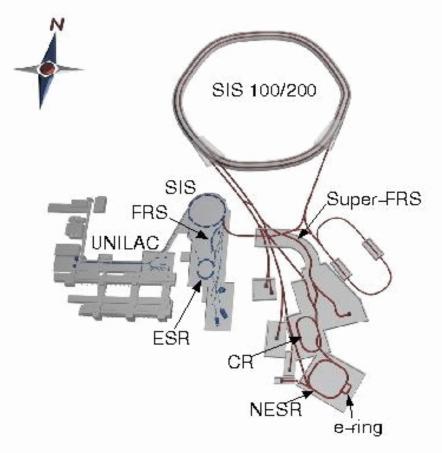
Analizing combined space charge and lattice nonlinearities based on an experiment at the CERN PS

G. Franchetti

Orsay, April 1st-2nd 2004 Lure, Building 209, room 110

Proposed GSI Upgrade

'An International Accelerator Facility for Research with lons and Antiprotons'



Predicted Gain Factors:

- primary beam intensities: 100 (10¹² uranium/s)
- → RIB intensities: up to 10⁴
- beam energy: 15

Special Properties:

- · fast cooled beams of exotic nuclei
- internal targets (p,e) for in-ring experiments

Technology Challenges:

- Fast cycling SC magnets for SIS 100/200
- Fast bunch compression in SIS 100
- High-power fragmentation target
- Fast stochastic cooling in the CR
- Medium energy electron cooling in the NESR

Accelerator Physics Challenges for SIS 100/200:

- Low tolerable relative beam loss
- Control of space charge effects
- Control of dynamic pressure effects



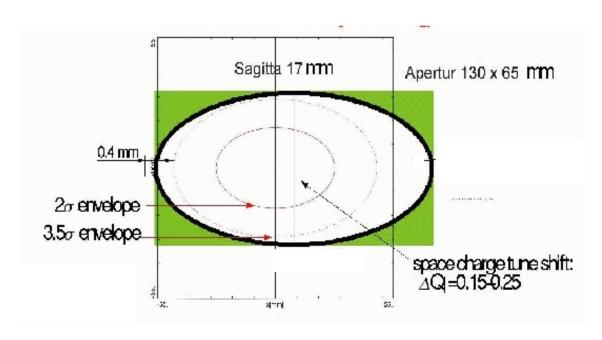


Of Particular Interest for GSI Project

- Accumulation of four SIS18 batches of U +28 at 96 MeV/u in the SIS100 in about 1 second.
- This takes hundreds thousand turns
- Maximum tuneshift of a bunched beam is 0.2

Challenges

- The beam fills a good fraction of the beam pipe
- Beam loss at the percent level





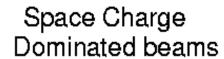


A New Area of Studies

Single Particle
Nonlinear Dynamics



- no self consistent effect
- nonlinearities intrinsic to the machine lattice
- hundred thousand turns





- self consistent effect
- nonlinearities from lattice+beam
- thousand turns





For SIS100 this distinction is no more justified.

1 s storage in presence of lattice nonlinearities + space charge





Work done:

- G.Franchetti et al. PRSTAB 6, 124201 (2003)
- G. Franchetti, I Hofmann, M. Giovannozzi, M. Martini, E. Metral 29th ICFA Workshop HALO03 693, 73
- G. Franchetti, I. Hofmann
 20th ICFA Workshop 642, 248

International effort on

Code Benchmarking on Space Charge and Nonlinear Dynamics

http://www-w2k.gsi.de/ihofmann/Benchmarking/Bench-marking.htm





In this talk

- Present our understanding of the dynamics of long term storage of space charge dominated bunched beams.
- We show that up to a certain extent this can be modeled in a single particle framework
- We present the comparison between experimental results and simulation

Our interests in this workshop

- Open a discussion for understanding better interplay of lattice nonlinearities and "collective nonlinearities".
- Role of FMA in disentangling in the theory or in experiments the dynamics of these complex system.





A 2D example

equation of motion

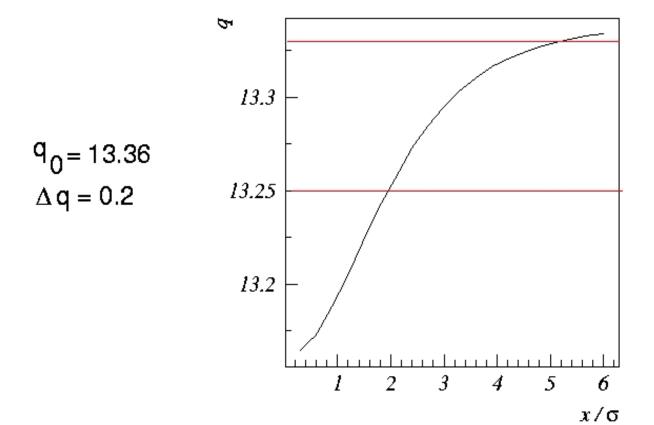
$$x'' + \left(\frac{q_0}{R}\right)^2 x = K\frac{x}{r^2}(1 - e^{-\frac{x^2}{2\sigma^2}})$$
 space charge nonlinear defocusing strength

Symplectic integration: space charge is included through kicks





Space charge induced nonlinear tune



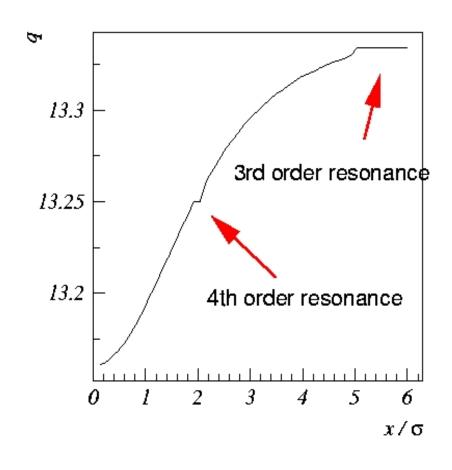
$$3 q = 40$$

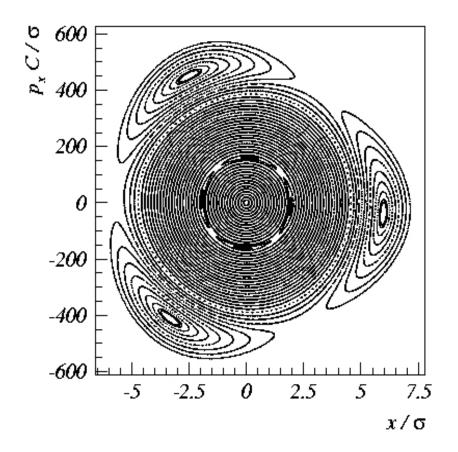
$$4 q = 53$$



When the lattice has nonlinearities







20 sext. error a2 = 0.01

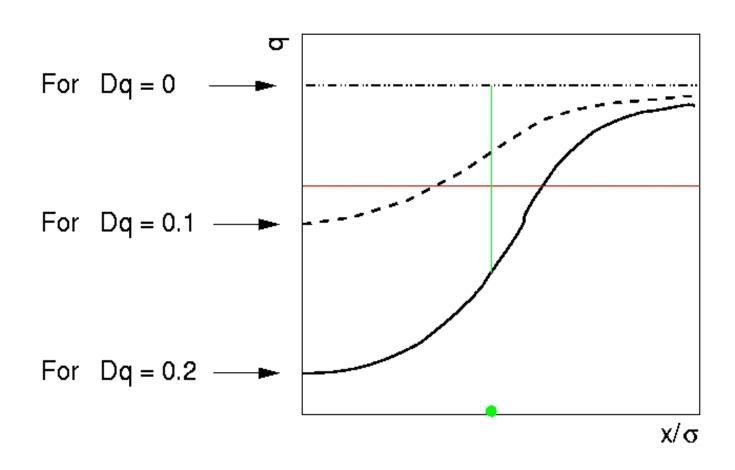
53 oct. errors a3 = 0.005

$$a3 = 0.005$$





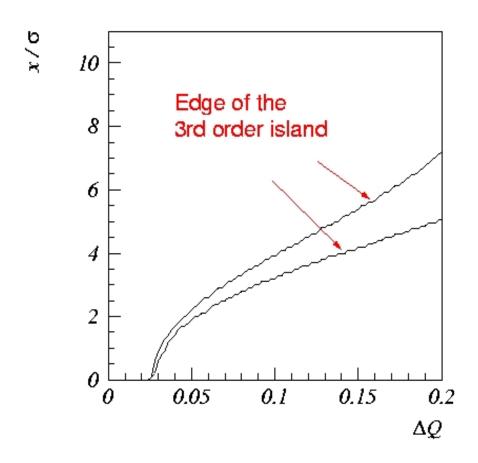
What happen when the perveance is dynamically changed when in the lattice nonlinearities are present?







When the perveance K is changed





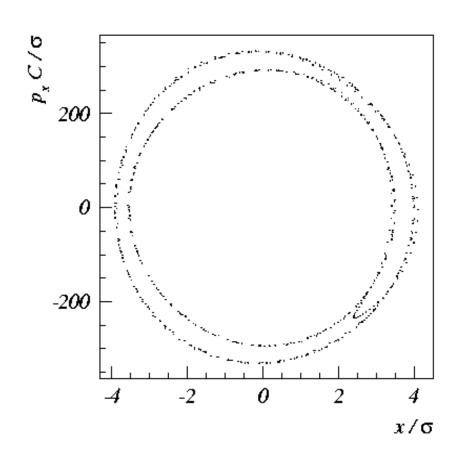
island coming from inside particle orbit

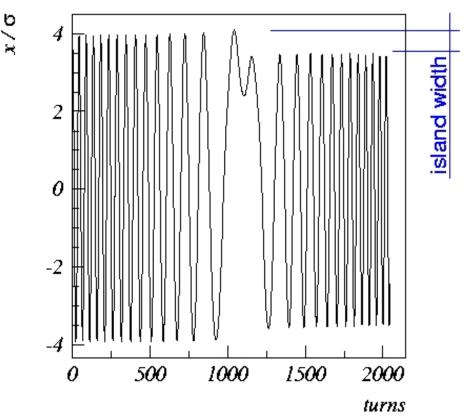


$$\begin{array}{l} x=4 \\ px=0 \end{array}$$

$$q = 13.36$$

current ramp 0.2 / 2048 [Δq /turns]







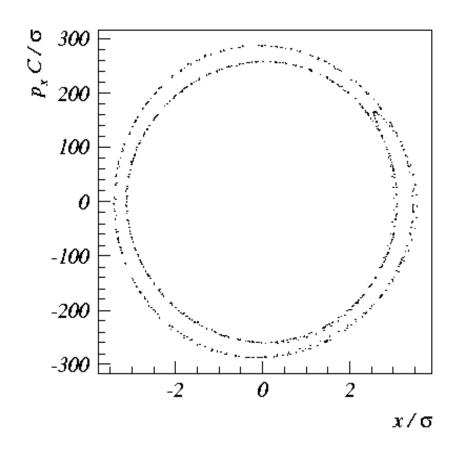
Island coming from outside the particle orbit

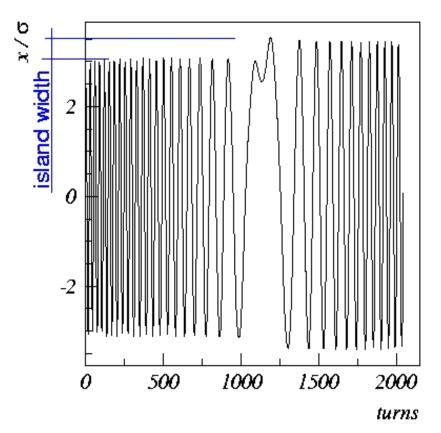


$$\begin{array}{l} x=3\\ px=0 \end{array}$$

$$q = 13.36$$

current ramp =0.2 / 2048 [∆ q/turns]







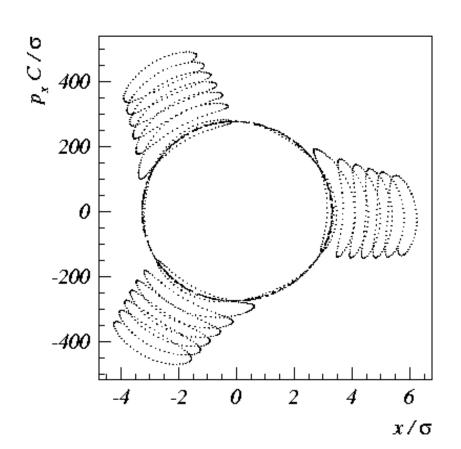


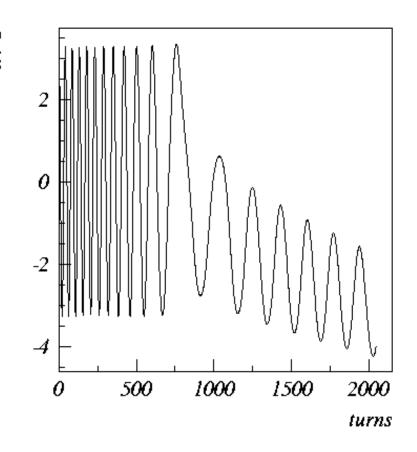


$$\begin{array}{l} x = 3.3 \\ px = 0 \end{array}$$

$$q=13.36\,$$

current ramp 0.2 / 2048 [\(\Delta q / turns \)]





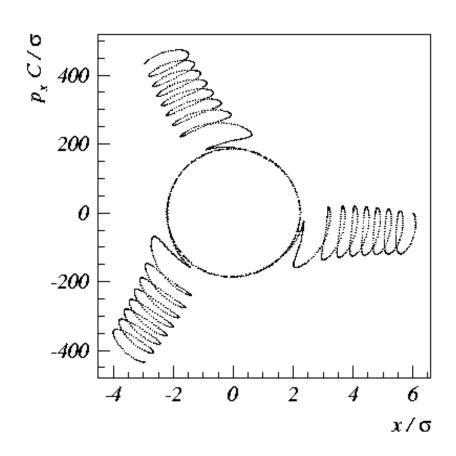


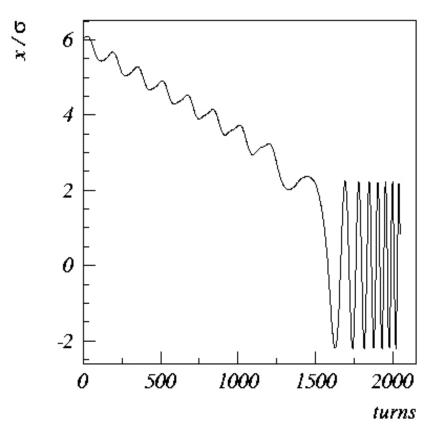


For a negative rate the capture process inverts the effect

initial coordintes x=6, px = 0

from
$$Dq = 0.02$$
 to $Dq = 0$









Change of space charge intensity causes resonance trapping

Resonance trapping is not new in accelerators:

Studies of trapping induced by chromaticity

- A.W. Chao et al. Nucl. Instrum. Methods 133, 405 (1976).
- T. Satogata et al. PRL 68, 1838 (1992)

BUT

Space charge induces a nonlinear tune, whereas the previous studies considered a detuning induced by chromaticity which is amplitude independent





Consequencies

- Particles may be transported out of a beam: HALO formation in rings
- Maximum distance of the particles depends on bare tunes and resonance position

Open questions

- What is the probability of trapping/detrapping with space charge?
- What is the fraction of particles in a bunch which will be brought up to a certain amplitude?
- What is the role of dynamic aperture?





Measurements at the CERN-PS

(October, 15-18, 2002)

R. Cappi, G. Franchetti, M. Giovannozzi, I. Hofmann, M. Martini, E. Metral

excite the resonance 4 Qx = 25 by using a single octupole

working point range: qx = 6.23 - 6.28, qy = 6.08 - 6.4

octupole strength: $K_3 = 1.215 \ \text{I} \ (\text{m}^{-3} \) \ \text{I} = 0 \sim 400 \ \text{A}$

momentum spread dp/p₀ = $2.6 ext{ } 10^{-3}$

beam emittances (2σ unnormalized): ex = 9 mm mrad

ey = 4.5 mm mrad

bunch length: 200 ns

beam energy: 1.4 GeV

flat-top: of 1.2 seconds

Emittance measurements: flying wire (< 1ms)





momentum spread $dp/p_0 = 10^{-3}$

beam emittances (2σ unnormalized): ex = 9 mm mrad

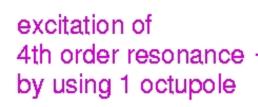
ey = 4.5 mm mrad

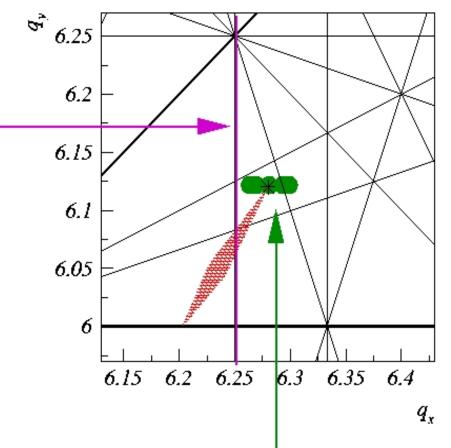
bunch length: 200 ns





The PS experiment



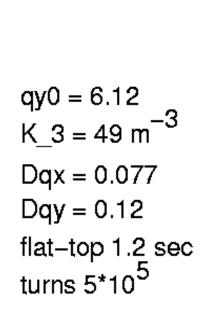


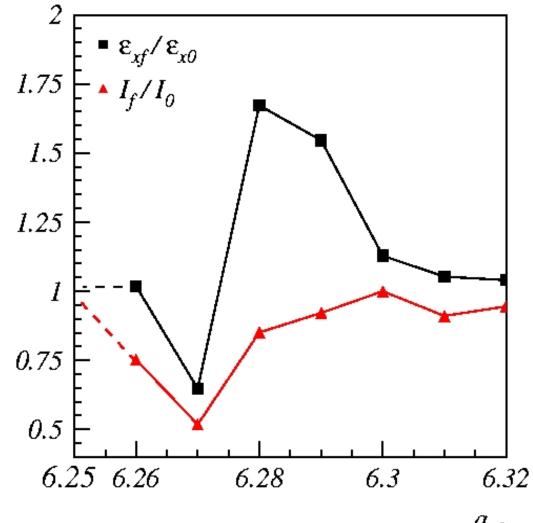
scan of qx0 from 6.32 to 6.25 qy0 = 6.12





PS experiment on Resonance Crossing: Summary

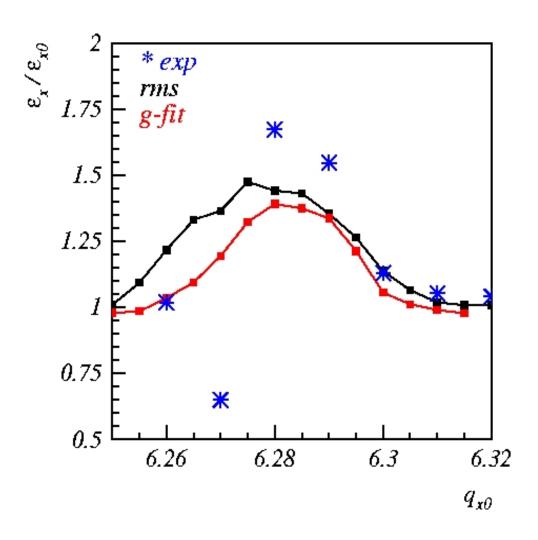








Benchmarking: experiment vs simulation



3D long term simulation



$$n(x,y,z) = \hat{n}(T)/(4\pi abc), \qquad \int_0^1 \hat{n}(t^2)t^2dt = 1,$$

with

$$T = \frac{x^2}{a^2} + \frac{y^2}{b^2} + \frac{z^2}{c^2}.$$

The charge distribution is given by $\rho(x, y, z) = Q n(x, y, z)$, with Q the total charge in the bunch. The electric field is given by

$$E_x = \frac{Q}{2} x \int_0^\infty \frac{\hat{n}(\hat{T})}{(a^2 + t)^{3/2} (b^2 + t)^{1/2} (c^2 + t)^{1/2}} dt,$$

with

$$\hat{T} = \frac{x^2}{(a^2 + t)} + \frac{y^2}{(b^2 + t)} + \frac{z^2}{(c^2 + t)}.$$

if a = b then:

$$\hat{n}(t) = \sum_{l=0}^{\infty} c_l t^l$$

$$E_x = \frac{Q}{2} x \sum_{l=0}^{\infty} c_l \sum_{i+j=l}^{\infty} \binom{l}{i} r^{2i} z^{2j} I_{i+1,j}(0) \qquad I_{n,m}(t_1) = \int_{t_1}^{\infty} \frac{1}{(a^2 + t)^{1+n} (c^2 + t)^{1/2+m}} dt$$

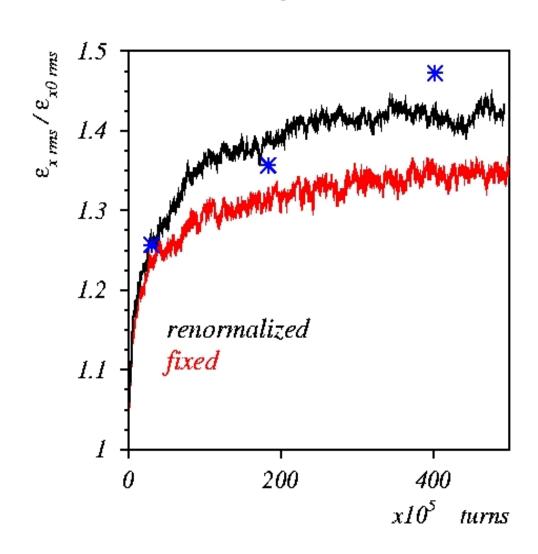
For the simulation

$$\hat{n}(t) = \begin{cases} \frac{105}{8}(1-t)^2 & if & t \le 1 \\ 0 & if & t > 1 \end{cases} \qquad n(x) = \frac{35}{32a}[1 - (x/a)^2]^3$$





renormalizing beam size by computing the rms emittances every 500 turns

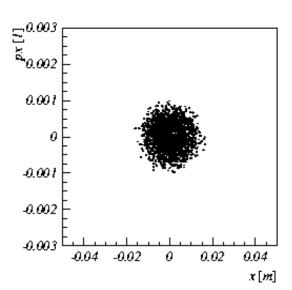


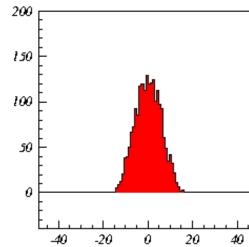


qx0 = 6.26

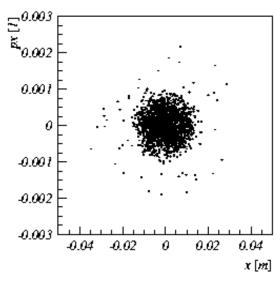


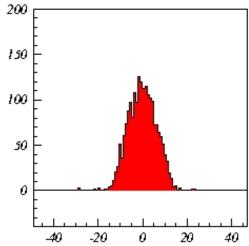
tums = 0





turns = 5000000









Conclusion

- We have described a mechanism for halo formation in rings
- Simulation prediction are in good agreement in the region of emittance growth
- We predict a halo radius which increases as Qx0 approaches the resonance
- Beam loss may occur when halo radius exceed dynamic aperture