Paradoxes of the Cosmological Physics in the beginning of the 21-st century

Yurij V. Baryshev

Astronomical Institute of the Saint-Petersburg State University

Protvino 2014

Contents

The Standard Cosmological Model (SCM)
 Observational and conceptual problems of the SCM in the beginning of the 21-st century
 Perspectives for physical testing the basis and predictions of the SCM

The Standard Cosmological Model in the beginning of the 21-st century

Initial assumptions (the basis):

- **Homogeneous** and isotropic matter distribution in the expanding Universe ($\rho = \rho(t)$; p = p(t); $g_{ik}(t)$)
- General Relativity (geometrical gravity theory) is applicable to the whole Universe
 (g_{ik}; R_{iklm}; T^{ik}_(m+de))
- Laboratory physics in the expanding Universe (QFT)
- Inflation in the early Universe (flatness, isotropy, initial conditions for structure formation)





Einstein-Friedmann expanding space paradigm

$$R^{ik} - \frac{1}{2}g^{ik}R = \frac{8\pi G}{c^4} (T^{ik}_{(m)} + T^{ik}_{(de)}) (T^{ik}_{(m)} + T^{ik}_{(de)}); k \equiv 0$$

$$\begin{split} R_{ik} &= \frac{\partial \Gamma^{l}{}_{ik}}{\partial x^{l}} - \frac{\partial \Gamma^{l}{}_{il}}{\partial x^{k}} + \Gamma^{l}{}_{ik} \Gamma^{m}{}_{lm} - \Gamma^{m}{}_{il} \Gamma^{l}{}_{km} \\ \Gamma^{i}{}_{kl} &= \frac{1}{2} g^{im} \left(\frac{\partial g_{mk}}{\partial x^{l}} + \frac{\partial g_{ml}}{\partial x^{k}} - \frac{\partial g_{kl}}{\partial x^{m}} \right) \\ T^{ik}{}_{(m+de)} &= diag \left(\rho c^{2}, p, p, p \right); \quad \rho = \rho(t); \quad p = p(t) \\ ds^{2} &= g_{ik} dx^{i} dx^{k} \quad \text{RW homogeneous metric:} \\ ds^{2} &= c^{2} dt^{2} - S(t)^{2} [d\chi^{2} + I_{k}(\chi)^{2} d\omega^{2}] \\ I_{k}(\chi) &= \sin \chi, \chi, \sinh \chi \quad for \quad k = 1, 0, -1 \\ r(t) &= S(t) \times \chi \quad - \text{expanding space} \\ \text{(increasing distance between galaxies)} \end{split}$$

Friedmann's equations: derivation from Einstein's equations

$$T^{i}_{k} = diag(\varepsilon, -p, -p, -p) = T_{(m)}^{i}_{k} + T_{(de)}^{i}_{k}$$

$$\varepsilon = \varepsilon(t) = \varrho(t)c^{2}, \quad p = p(t) \quad (\text{ no } T_{(g)}^{i}_{k})$$
Einstein's eq. (0, 0):
$$\frac{\dot{S}^{2}}{S^{2}} + \frac{kc^{2}}{S^{2}} = \frac{8\pi G}{3c^{2}}\varepsilon \qquad \varepsilon = \varepsilon_{m} + \varepsilon_{de}$$

$$p = p_{m} + p_{de}$$
Einstein's eq. (1, 1):
$$2\frac{\ddot{S}}{S} + \frac{\dot{S}^{2}}{S^{2}} + \frac{kc^{2}}{S^{2}} = -\frac{8\pi G}{c^{2}}p$$
Bianchi identity:
$$3\frac{\dot{S}}{z} = -\frac{\dot{\varepsilon}}{2}$$
Fluid EoS $p = \gamma\varepsilon$
Dust: $\gamma = 0$, Radiation: $\gamma = 1/3$

Vacuum: $\gamma = -1$

 $\mathcal{E} + p$

S

Friedmann's equations: standard form for scale factor S(t)

$$H^{2} - \frac{8\pi G}{3} \rho = -\frac{kc^{2}}{S^{2}}$$

$$\ddot{S} = -\frac{4\pi G}{3} \left(\rho + \frac{3p}{c^2}\right)$$

S

A

$$q = \frac{1}{2} \Omega \left(1 + \frac{3p}{\rho c^2} \right)$$

Main parameters of the Friedmann model:

 $1 - \Omega = -\Omega_k$

$$H(t) = \dot{S} / S$$

$$q = -\ddot{S} S / \dot{S}^{2}$$

$$\rho_{crit} = 3H^{2} / 8\pi G$$

$$\Omega = \rho / \rho_{crit}$$

$$\Omega_{k} = kc^{2} / S^{2}H^{2}$$

Solution for
$$p(t) = \gamma \varepsilon(t)$$
 :

$$S(t) \propto t^{2/(3+3\gamma)}, \ k = 0$$

$$S(t) \propto \exp(\alpha t), \ \gamma = -1$$

Fundamental conclusions of the LCDM SCM

- Cosmological redshift $z = d\lambda/\lambda$ and the linear Hubble Law $v_{exp} = Hr$ is the consequence of the homogeneous space expansion $r(t) = S(t) \times \chi$
- Cosmic microwave background radiation is the result of the photon gas cooling in the expanding space $T(z) = T_0(1+z)$
- Small anisotropy $\Delta T/T(\theta)$ of the CMBR is determined by the initial spectrum of density fluctuations which are the source of the large scale structure of the Universe
- The expanding Universe is made of unobservable in lab dark energy (70%), nonbaryonic dark matter (25%) plus ordinary matter (5%). Visible galaxies contribution is less than 0.5%.

Main mass-energy components in the LC(nonbaryonic)DM SCM



Planck 2013 (arXiv:1303.5076): Minimal (flat) 6 parameters LCDM $\Omega_i = \rho_i / \rho_{crit}$ $\rho_{crit} = \frac{3H^2}{8\pi G} = 0.853 \times 10^{-29} \, g/cm^3$ $\rho_{vac} \sim \rho_{Pl} = 10^{94} \, g/cm^3$
$$\begin{split} H_0 &= \ 67.4 \ \pm 1.4 \ km \ s^{-1} \ Mpc^{-1} \\ \Omega_m &= 0.314 \ \pm 0.020 \\ \Omega_c &= 0.263 \ \pm 0.0031 \\ \Omega_b &= 0.048 \ \pm 0.0003 \\ \Omega_{de} &= 0.686 \ \pm 0.020 \end{split}$$



Cosmology is almost finished !



Examples of observational problems in the L(nonbaryonic)CDM SCM

- Tasitsiomi A., The Cold Dark Matter crisis on galactic and subgalactic scales, Int.J.Mod.Phys. D12 (2003) 11572002
- Perivolaropoulos L., LCDM: Triumphs, Puzzles and Remedies, (2011), Proc. Int. Conf. (arXiv:1104.0539)
- Kroupa P. et al., The Failures of the Standard Model of Cosmology Require a New Paradigm, Int.J.Mod.Phys. D21 (2012), 1230003
- Sylos Labini F., Inhomogeneous Universe, Class.Quant.Grav., 28, iss.16, 4003 (2011)

Problem of galaxy structure – dark matter relations

Problem of cusps and satellite galaxies

Problem of large scale correlations in 3d galaxy surveys

Problem of CMBR – ecliptic correlation

Is the SCM ultimate physical picture of the World? Is the basis of SCM firmly established?

Turner M.S., Absurd universe, Astronomy, Vol. 31, No. 11, p. 44 - 47 (2003) (visible matter <0.5%)

- **Steinhardt P., Inflation debate: Is the theory at the heart of modern cosmology deeply flawed ?**, Sci. Am., April 2011 (+ arXiv: 1402.6980)
- Francis M. et al., Expanding Space: the Root of all Evil?, Publ. Astron. Soc. Austr., 24, 95 (2007)
- **Clifton T.** et al., **Modified gravity and cosmology**, Physics Reports, Volume 513, Issue 1, p. 1-189 (2012)
- Hawking S., arXiv: 1401.5761 + Nature, 24 January 2014, "there are no black holes no event horizons and no firewalls".

Conceptual problems of the SCM

Gravitation theory

(GR is not a quantum theory, absence of the EMT of gravitational field in GR, **Hawking 2014:** need for unite gravity with the other fundamental forces of nature – "no black holes")

Physics of space expansion

(Continuous creation of vacuum, violation of energy conservation) The nature of cosmological redshift

(It is not the Doppler effect, the global gravitational cosmological redshift should be taken into account)

Inhomogeneity of matter distribution (Discreteness and **fractality** of spatial galaxy distribution)

Inhomogeneity of matter distribution

Homogeneous point distribution with D = 3

Poisson, N=10 000



Inhomogeneous fractal point distribution with D = 2(fractal dimension $N(r) \sim r^D$)

Random Cantor, D=2, N=10 938



Conditional density $< n(r) >_p$: estimation

Integrated conditional density (in spheres)

$$\Gamma^*(r) = \langle n(r' < r) \rangle_p$$



Inhomogeneous spatial galaxy distribution in SDSS sample (~100000 galaxies)



There are structures with sizes 400 Mpc , while LCDM predicts homogeneity scale 5 – 10 Mpc . Fractal dimension $D \sim N(r)^D$, $D \sim 2$.

Conditional density of SDSS galaxy samples: Sylos Labini F. et al., Astron.Astrophys.508:17-43,2009, fractal dimension D ~ 2



The nature of cosmological redshift

Exact velocity – distance relation in the Friedmann model

$$V_{exp} = \frac{dr}{dt} = \frac{dS}{dt} \chi = \frac{dS}{dt} \frac{r}{S} = H(t) r = c \frac{r}{r_{H}}$$

$$r(t) = S(t) \chi \quad H(t) = \dot{S}/S \quad r_{H} = c/H(t) - Hubble radius$$
Homogeneity implies
$$V_{exp} = H r$$

Exact GR relation between galaxy expansion velocity and distance to the galaxy (no relativistic effects)

Linear Hubble Law

Peebles P.J.E., Schramm D.N., Turner E.L., Kron R., *"The case for the relativistic hot Big Bang cosmology",* Nature, 352, 769 (1991):

"The connection between homogeneity and Hubble's law was the first success of the expanding world model."

$\rho(r,t) = \rho(t) \iff V_{exp} = H r$





Hubble – de Vaucouleurs (HdeV) paradox

The linear Hubble Law cz = Hrcoexists with strong inhomogeneous Fractal Law – de Vaucouleurs Law of spatial galaxy distribution $\rho(r) = \rho_0 \left(\frac{r_0}{r}\right)^{\gamma}$

 $\gamma = 3 - D$

Global gravitational cosmological redshift

Fractal matter distribution: $M(r) \propto r^D$

Global gravitational cosmological redshift $z = d\lambda/\lambda$:

$$z_{\text{grav}} = \frac{\delta\phi(r)}{c^2} = \frac{1}{2} \frac{GM(r)}{c^2 r}$$

$$z_{\text{grav}}(r) = \frac{4\pi G \rho_0 r_0^2}{c^2 D (D-1)} \left(\frac{r}{r_0}\right)^{D-1} = \frac{2\pi G \rho_0 r_0}{c^2} r = \frac{H_g}{c} r \quad (D=2)$$
$$H_g = 2\pi \rho_0 r_0 \frac{G}{c} \approx 69 \text{ km s}^{-1}/\text{Mpc}$$
$$\rho_0 = 5 \times 10^{-24} \text{ g/cm}^3, r_0 = 10 \text{ kpc}$$

Physics of space expansion



What is space expansion?

 $\mathbf{r}(t) = \mathbf{S}(t) \boldsymbol{\chi}$

→ increasing distance between galaxies with violation of limiting velocity

→ continues space and vacuum creation

 \rightarrow violation of energy conservation

Cosmological physics of the expanding space is essentially different from the lab physics

Newtonian physics of the relativistic Friedmann equations

 $r(t) = S(t) \cdot \chi$ - distance to a galaxy

$$\ddot{S} = -\frac{4\pi G}{3} \left(\rho + \frac{3p}{c^2}\right) S \equiv \frac{d^2 r}{dt^2} = -\frac{G M_g(r)}{r^2}$$
 Newtonian ball r(t) expansion

$$H^{2} - \frac{8\pi G}{3} \rho = -\frac{k c^{2}}{S^{2}} \equiv \frac{V^{2}_{exp}}{2} - \frac{GM}{r} = const$$

Newtonian kinetic and potential energy

$$M_g(r) = -\frac{4\pi}{3} \left(\rho + \frac{3p}{c^2}\right) r^3$$

$$p = \gamma \rho c^2 \qquad M_g(r) = \frac{4\pi}{3} (1 + 3\gamma) \rho r^3 \quad \propto \quad S^{-3\gamma}(t)$$

Cosmological redshift as the Lemaitre effect in the expanding Universe $(1+z) = \lambda_0 / \lambda_1 = S_0 / S_1$

$$V_{\exp}(r) = \frac{dr}{dt} = H r$$
 Velocity – distance relation (for receding galaxy)

$$V_{
m exp}(z) = H \left| r(z) \right|$$
 Velocity – redshift relation (for receding galaxy)

$$r(t_0, z) = r(z) = \frac{c}{H_0} \int_0^z \frac{dz'}{h(z')}$$

Distance – redshift relation

 $h(z) = \left[\Omega^{0}_{dm}(1+z)^{3} + \Omega^{0}_{rad}(1+z)^{4} + \Omega^{0}_{de}(1+z)^{3(1+w)} + (1-\Omega^{0}_{tot})(1+z)^{2}\right]^{1/2}$

Comparison of the Lemaitre and the Doppler effects

$$V_{\exp}(z) = c \frac{r(z)}{r_H} \Rightarrow V_{\exp}(z) > c \quad if \quad z > 3$$

$$(q_0 = 0.5, \quad p = 0)$$

$$V_{Dop}(z) = c \frac{2z + z^2}{2 + 2z + z^2}$$

The expansion velocity of a galaxy in SCM Models: 1: $(\Omega m = 1, \Omega v = 0)$; 2: $(\Omega m = 0, \Omega v = 0)$; 3: $(\Omega m = 0, \Omega v = 1)$

Z,	$v_{\rm dop}/c$	$v_{\rm exp}/c, 1$	$v_{\rm exp}/c, 2$	$v_{\rm exp}/c, 3$
1	0.6	0.6	0.75	1
2	0.8	0.84	1.33	2
6	0.96	1.25	3.43	6
10	0.98	1.4	5.45	10
1000	$1 - 2 \cdot 10^{-6}$	1.94	500	1000
∞	1	2	∞	∞

Gravitation theory

Energy paradox in the SCM



T.Davis, Scientific American, July 2010, p.39

Is the Universe Leaking Energy?

Total energy must be conserved. Every student of physics learns this fundamental law. The trouble is, it does not apply to the universe as a whole By Tamara M. Davis

The problem of energy conservation in the expanding Universe

Bianchi identity:

$$(T^{ik})_{;k} = (T^{ik}_m + T^{ik}_{de})_{;k} = 0$$

$$3\frac{\dot{S}}{S} = -\frac{\dot{\varepsilon}}{\varepsilon + p} \equiv dE + p \, dV = 0$$

$$dE = d(\varepsilon V) = d(\rho c^2 V) \quad dE = -p \ dV$$

 $V = const \times S^3$

Change of energy in each local comoving volume during space expansion $r(t) = S(t) \cdot \chi$ (without performing work). Continuity equation – no energy conservation (no gravity energy)

L.D.Landau & E.M.Lifshitz $(T_{(m)}^{ik} + t_{(g)}^{ik})$ "The Classical Theory of Fields", Oxford, 1971

§ 101. The energy-momentum pseudotensor

In the absence of a gravitational field, the law of conservation of energy and momentum of the material (and electromagnetic field) is expressed by the equation $\partial T^{ik}/\partial x^k = 0$. The generalization of this equation to the case where a gravitational field is present is equation (94.7):

$$\underline{T_{i;k}^{k}} = \frac{1}{\sqrt{-g}} \frac{\partial (T_{i}^{k}\sqrt{-g})}{\partial x^{k}} - \frac{1}{2} \frac{\partial g_{kl}}{\partial x^{i}} T^{kl} = 0.$$
(101.1)

In this form, however, this equation does not generally express any conservation law whatever.[†] This is related to the fact that in a gravitational field the four-momentum of the matter alone must not be conserved, but rather the four-momentum of matter plus gravitational field; the latter is not included in the expression for T_i^k .

To determine the conserved total four-momentum for a gravitational field plus the matter located in it, we proceed as follows.[‡] We choose a system of coordinates of such form that

[†] Because the integral $\int T^k \sqrt{-g} \, dS_k$ is conserved only if the condition

$$\frac{\partial(\sqrt{-g}T_i^k)}{\partial x^k} = 0$$

Edward Harrison, "Cosmology: the science of the Universe", Cambridge Univ. Press, 2000

Lab Physics

Cosmological Physics





"The conclusion, whether we like it or not, is obvious: energy in the Universe is not conserved."

dE = -p dV

Exact law for disappearance/creation of energy for each local comoving volume

Disappearance/creation of energy in each local comoving volume of a ball with radius

$$r(t) = S(t) \chi \text{ for } \mathbf{p} = \gamma \varepsilon$$

$$E(r,t) = \int_{0}^{r} T_{0}^{0} dV = \frac{4\pi}{3} \varepsilon(t) S^{3}(t) \chi^{3} \sigma_{k}(\chi)$$

$$E(r,t) = \frac{4\pi}{3} \rho(t) c^{2} r^{3}(t) \sigma_{k}(\chi) \propto S^{-3\gamma}(t)$$

$$E_{dust}(t) \propto const$$

=> the only Friedmann model ($\gamma = 0$), where the energy conserved (Lemaitre 1933; Mitra 2011)

$$E_{rad}(t) \propto S^{-1}(t)$$

=> cooling photon gas (CMBR) in the Friedmann model ($\gamma = 1/3$) is the result of continue disappearance of energy in a comoving volume

 $E_{vac}(t) \propto S^{+3}(t) = vacuum energy in the Freidmann model$ $(\gamma = -1) is continuously increases (created) in a comoving volume$

Perspective for solution of paradoxes

- Developing the theory of gravitational interaction, Ghctheory, laboratory and astrophysical testing gravity theories predictions (Einstein's geometrical approach and Feynman's field approach, nearby SN explosions, RCO, neutrino detection)
- Galaxy properties and large scale structure at very high redshifts (wide-angle and pencil-beam very deep surveys of faint galaxies and QSO)
- Observational tests of the nature of the cosmological redshift (Сэндидж m(z, SN Ia) u dZ/dt; Kopeikin dλ/λ<0 in the Solar system)



Joseph Hooton Taylor, (born March 29, 1941)

Nobel Prize in Physics (1993) with Russell Alan Hulse

"for the discovery of a <u>new</u> type of <u>pulsar</u>, a discovery that has opened up new possibilities for the study of <u>gravitation</u>"

PSR 1913+16 : Decreasing orbital energy via radiation of *positive energy of the gravitational waves*



E E Y M A M LECTURES ON GRAVITATION

Richard P. Feynman Fernando B. Morinigo = William G. Wagne

1995

Edited by Brian Hatfield

With a Foreword by John Preskill and Kip S. Thorne

From Feynman to his wife:

Ì

"Remind me not to come to any more gravity conferences!"

Gravitation is just a next fundamental physical interaction: symmetric tensor field ψ^{ik}

"The geometric Interpretation is not really necessary or essential to physics." (p. 113)

"the situation is exactly analogous to electrodynamics - and in the quantum interpretation, every radiated graviton carries away an amount of energy ħw." (p.220)

Field Gravity and Geometry

Field Gravity Geometry $g^{ik} \approx \eta^{ik} + \phi^{ik}$ $\eta^{ik}, \psi^{ik}(\vec{r},t), \ \vec{F}_{q}$ $g^{lk}(\vec{r},t), \mathcal{R}_{iklm},$ $g_{ik} g^{ik} = 4$ Spin 2 $\psi(\vec{r},t) = \eta_{ik}\psi^{ik}$ Spin 2 \oplus Spin 0 Einstein(1915) Bronstein(1936) Hilbert Poincare(1905) Fierz, Pauli Gupta Fock Thirring Deser Landau, Lifshitz Kalman Weinberg MTW, ... Feynman Π \parallel Zakharov Straumann (2000) ╢ Clifton et al. (2012) Padmanabhan(2008) Sokolov, Baryshev(1980) "no way to geometry" "modifications "New Relativistic of GR" Astrophysics" BH, noEMT, GW(T)RCO (noBH), EMT, GW(T+S)

Theory of gravitational interaction

Field Gravity

General Relativity

- The inertial reference frames The flat Minkowski space η^{ik} – conservation laws The concept of potential $\psi^{ik}(\vec{r},t), \ \psi(\vec{r},t) = \eta_{ik}\psi^{ik}$ scalar part, force, gravitons The Energy-Momentum **Tensor** of the gravity field $T_{(a)}^{ik}$ * The universality of gravitational interaction $\Lambda_{(int)} = \psi_{ik} T^{ik}, \ m_0$
- The non-inertial reference frames
- * The **curved Riemannian** space-time $g_{ik} g^{ik} = 4$
- * The metric tensor $g^{ik}(\vec{r}, t)$, the curvature tensor \mathcal{R}_{iklm}
- * The **Pseudo-Tensor** of the gravity field $t_{(g)}^{ik}$
- The Equivalence Principle

 $m_{inert} = m_{grav}$ free falling frames

Relativistic Compact Objects, SN explosion, neutrino detection

Pencil-beam and wide-angle galaxy surveys for high redshift galaxies and large scale structures



Observational test: stellar magnitude – redshift i.e. flux – distance relation

SN la



m(z, SN Ia)



SNIa Hubble Diagram

Hicken M. et al., Ap.J., 2009, 700, 1097

Residual

Observational test of the nature of the cosmological redshift



Sandage (1962) ApJ, 136, 319 "the change of redshift with time"

$$\frac{dz}{dt}(z) \sim 1 cm/s$$

Liske et al. (2008), MNRAS, 386, 1192 Astrophysics and Space Science Library 383

Yurij Baryshev Pekka Teerikorpi



Fundamental Questions of Practical Cosmology

Fundamental Questions of Practical Cosmology

Exploring the Realm of Galaxies

Yurij Baryshev, Pekka Teerikorpi

"Fundamental Questions of Practical Cosmology", Springer, Dordrecht Heidelberg London New York, 2012, 332p.

(Introduction to foundations and problems of modern cosmology) BT2012





Thanks!