

Dark Matter versus $h \rightarrow \gamma\gamma$ and $h \rightarrow \gamma Z$ with supersymmetric triplets

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(In collaboration with Chiara Arina and Germano Nardini, based on arXiv:1403.6434)

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Motivation

- Higgs found at the LHC with a mass of 126 GeV is SM-like.
- The minimal SUSY model requires heavy stops or large stop mixing (fine-tuning).
- Non-minimal SUSY models can raise the Higgs mass at tree level.
- Minimal extensions: Singlets (NMSSM) or Triplets ($Y = 0, \pm 1$).
- New EW sector \rightarrow Dark Matter/Collider phenomenology.

Addition of a $Y = 0$ $SU(2)_L$ triplet superfield.

$$\Sigma = \begin{pmatrix} \xi^0/\sqrt{2} & -\xi_2^+ \\ \xi_1^- & -\xi^0/\sqrt{2} \end{pmatrix}$$

The total superpotential reads

$$W_{\text{TMSSM}} = W_{\text{MSSM}} + \lambda H_1 \cdot \Sigma H_2 + \frac{1}{2} \mu_\Sigma \text{Tr} \Sigma^2,$$

and the soft-breaking Lagrangian

$$\mathcal{L}_{\text{TMSSM}_{\text{SB}}} = \mathcal{L}_{\text{MSSM}_{\text{SB}}} + m_4^2 \text{Tr}(\Sigma^\dagger \Sigma) + [B_\Sigma \text{Tr}(\Sigma^2) + \lambda A_\lambda H_1 \cdot \Sigma H_2 + h.c].$$

Electroweak precision observables impose $\langle \xi^0 \rangle \lesssim 4$ GeV which implies

$$|A_\lambda|, |\mu|, |\mu_\Sigma| \lesssim 10^{-2} \frac{m_\Sigma^2 + \lambda^2 v^2 / 2}{\lambda v}$$

If $A_\lambda, \mu, \mu_\Sigma$ at EW scale, then $m_\Sigma \gtrsim 2$ TeV

Decoupled scalar triplet ξ^0 ($m_\Sigma \gtrsim 5$ TeV)

$$\mathcal{M}_{h,H}^2 = \begin{pmatrix} m_A^2 \cos^2 \beta + m_Z^2 \sin^2 \beta & (\lambda^2 v^2 - m_A^2 - m_Z^2) \sin 2\beta/2 \\ (\lambda^2 v^2 - m_A^2 - m_Z^2) \sin 2\beta/2 & m_A^2 \sin^2 \beta + m_Z^2 \cos^2 \beta \end{pmatrix}$$

In the decoupling limit, $m_A \rightarrow \infty$

- Mass of the lightest Higgs state

$$m_{h,trees}^2 = m_Z^2 \cos^2 \beta + \frac{\lambda^2}{2} v^2 \sin^2 2\beta$$

- Couplings of the lightest Higgs state \rightarrow SM-like (except loop-induced processes)

Loop-induced couplings

- $h \rightarrow gg$

Heavy Squarks \rightarrow SM-like ggh coupling. The gluon fusion Higgs production is SM-like.

- $h \rightarrow \gamma\gamma / h \rightarrow Z\gamma$

New fields in the EW sector \rightarrow Deviation from the SM couplings.

$$M_{\tilde{\chi}^{\pm}}^{tree} = \begin{pmatrix} M_2 & gv \sin \beta & 0 \\ gv \cos \beta & \mu & -\lambda v \sin \beta \\ 0 & \lambda v \cos \beta & \mu_{\Sigma} \end{pmatrix}$$

$$h \rightarrow \gamma\gamma / h \rightarrow Z\gamma$$

$$R_{\gamma\gamma} = \left| 1 + \frac{A_{\tilde{\chi}_{1,2,3}^{\pm}}^{\gamma\gamma}}{A_W^{\gamma\gamma} + A_t^{\gamma\gamma}} \right|^2 \quad R_{Z\gamma} = \left| 1 + \frac{A_{\tilde{\chi}_{1,2,3}^{\pm}}^{Z\gamma}}{A_W^{Z\gamma} + A_t^{Z\gamma}} \right|^2$$

$$A_{\tilde{\chi}_{1,2,3}^{\pm}}^{\gamma\gamma} = \sum_{i=1}^3 \frac{2m_W}{\sqrt{2}m_{\tilde{\chi}_i^{\pm}}} (g_{h\tilde{\chi}_i^+ \tilde{\chi}_i^-}^L + g_{h\tilde{\chi}_i^+ \tilde{\chi}_i^-}^R) A_{1/2}(\tau_{\tilde{\chi}_i^{\pm}})$$

$$A_{\tilde{\chi}_{1,2,3}^{\pm}}^{Z\gamma} = \sum_{j,k=1}^3 \frac{g_2 m_{\tilde{\chi}_j^{\pm}}}{g_1 m_Z} f(m_{\tilde{\chi}_j^{\pm}}, m_{\tilde{\chi}_k^{\pm}}, m_{\tilde{\chi}_k^{\pm}}) (g_{h\tilde{\chi}_i^+ \tilde{\chi}_i^-}^L + g_{h\tilde{\chi}_i^+ \tilde{\chi}_i^-}^R) (g_{Z\tilde{\chi}_i^+ \tilde{\chi}_i^-}^L + g_{Z\tilde{\chi}_i^+ \tilde{\chi}_i^-}^R)$$

$A_W^{\gamma\gamma} = -8.3$	$A_t^{\gamma\gamma} = 1.9$	$A_W^{Z\gamma} = -12$	$A_t^{Z\gamma} = 0.6$
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Invisible Higgs Decays

Invisible Decay Width

$$\Gamma(h \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = \frac{G_F m_W^2}{2\sqrt{2}\pi} m_h \left(1 - \frac{4m_{\tilde{\chi}_1^0}^2}{m_h^2}\right)^{3/2} g_{h\tilde{\chi}_1^0\tilde{\chi}_1^0}^2$$

$$g_{h\tilde{\chi}_1^0\tilde{\chi}_1^0} = (N_{12} - \frac{g_1}{g_2} N_{11})(\sin\beta N_{14} - \cos\beta N_{13}) + \frac{\lambda}{g_2} N_{15}(N_{14} \sin\beta + N_{13} \cos\beta)$$

$$R_{XY} \equiv [\text{BR}(h \rightarrow XY)/\text{BR}_{\text{SM}}(h \rightarrow XY)] \times [1 - \text{BR}(h \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0)]$$

$$\mathcal{M}_{\tilde{\chi}_1^0}^{\text{tree}} = \begin{pmatrix} M_1 & 0 & -\frac{1}{2}g_1 v_1 & \frac{1}{2}g_1 v_2 & 0 \\ 0 & M_2 & \frac{1}{2}g_2 v_1 & -\frac{1}{2}g_2 v_2 & 0 \\ -\frac{1}{2}g_1 v_1 & \frac{1}{2}g_2 v_1 & 0 & -\mu & -\frac{1}{2}v_2 \lambda \\ \frac{1}{2}g_1 v_1 & -\frac{1}{2}g_2 v_2 & -\mu & 0 & -\frac{1}{2}v_1 \lambda \\ 0 & 0 & -\frac{1}{2}v_2 \lambda & -\frac{1}{2}v_1 \lambda & \mu \Sigma \end{pmatrix}$$

SARAH

⇒ Lagrangian TMSSM. Diagonalization, RGE's ...

SPheno

⇒ SUSY Spectrum (Full 1 loop, 2 loop corrections for the Higgs, low energy observables, ...)

CPSuperH

⇒ Passarino-Veltman functions

micrOMEGAS

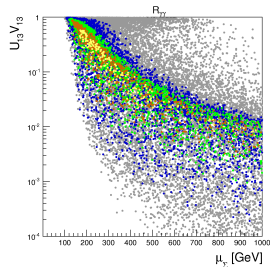
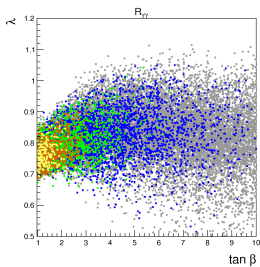
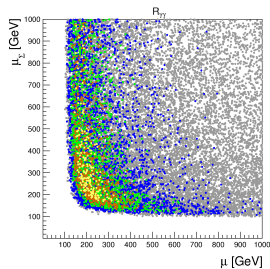
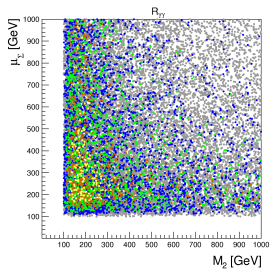
⇒ Relic Abundance, direct detection cross sections

Multinest

$$p(\theta_i|d) \propto \mathcal{L}(d|\theta_i)\pi(\theta_i), \quad \{\theta_i\} = \{M_1, M_2, \mu, \mu_\Sigma, \tilde{m}, \tan \beta, \lambda\}$$

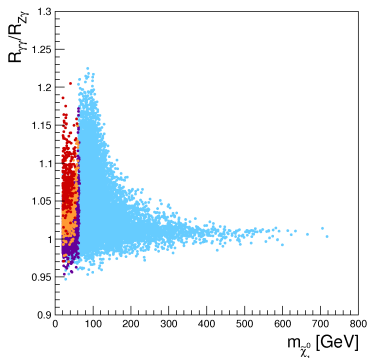
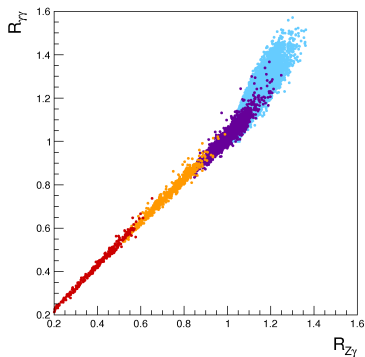
NS parameters	Prior range
$\log_{10}(M_1/\text{GeV}), \log_{10}(\mu_\Sigma/\text{GeV})$	$1 \rightarrow 3$
$\log_{10}(\mu/\text{GeV}), \log_{10}(M_2/\text{GeV})$	$2 \rightarrow 3$
\tilde{m}/TeV	$0.63 \rightarrow 2$
$\log_{10}(\tan \beta)$	$0 \rightarrow 1$
λ	$0.5 \rightarrow 1.2$

Type	Observable	Measurement/Limit
<u>Collider data</u>	m_h	$125.85 \pm 0.4 \text{ GeV (exp)} \pm 3 \text{ GeV (theo)}$
	$\Gamma(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0)$	$< 2 \text{ MeV}$
	$m_{\tilde{t}_1}$	$> 650 \text{ GeV (LHC 90\% CL)}$
	$m_{\tilde{\chi}_1^+}$	$> 101 \text{ GeV (LEP 95\% CL)}$
<u>DM data</u>	$\Omega_{\text{DM}} h^2$	$0.1186 \pm 0.0031 \text{ (exp)} \pm 20\% \text{ (theo)}$
	$\sigma_{\text{Xe}}^{\text{SI}}$	LUX (90\% CL)



$R_{\gamma\gamma} < 1.1$, 1.2 , 1.3 , 1.4 and $R_{\gamma\gamma} > 1.4$

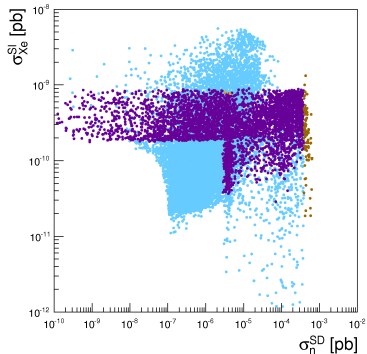
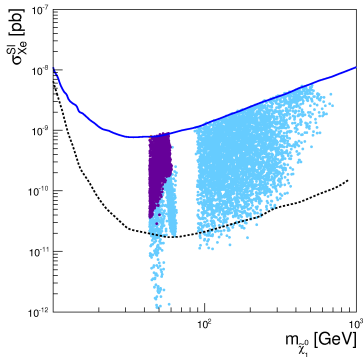
$BR_{inv} < 1\%$, 20% , 50% and $BR_{inv} > 50\%$.



Enhancement up to 60% for $R_{\gamma\gamma}$ and 40% for $R_{Z\gamma}$
Correlation in $h \rightarrow \gamma\gamma$ and $h \rightarrow Z\gamma$ channels!!

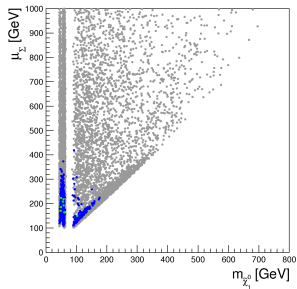
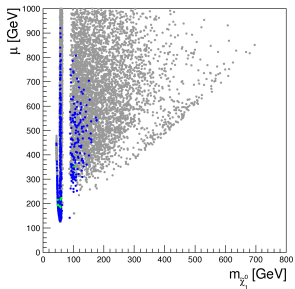
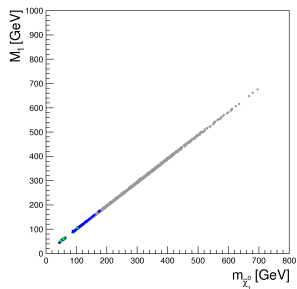
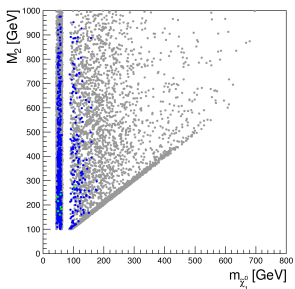
Dark Matter. Neutralino Phenomenology

Lightest Neutralino is a viable DM candidate in the Z, h resonances and in the well-tempered region.

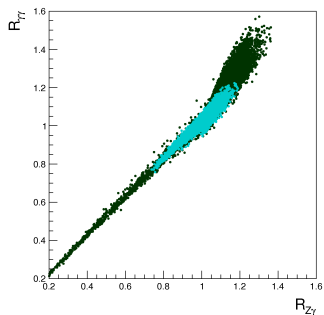
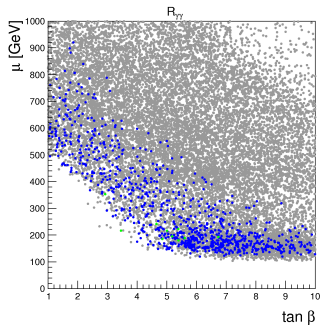


[Brown points at odds with the XENON100 exclusion bound for σ_n^{SD}]

Dark Matter. Neutralino Composition



Dark Matter.



DM constraints reduce the enhancement up to 20%.

Conclusions

- The triplet extension achieves a Higgs mass of 126 GeV reducing the fine-tuning.
- Loop induced processes in Higgs Physics may deviate from the SM ones.
- [No DM Pheno] Large enhancement in the decays $h \rightarrow \gamma\gamma$ (60%) and $h \rightarrow Z\gamma$ (40%). Both channels are correlated.
- [With DM Pheno] Lightest neutralino is a viable DM candidate in different regions.
- [With DM Pheno] Constraints on SI cross-section coming from LUX reduces $R_{\gamma\gamma}$ and $R_{Z\gamma}$

Thank you for your attention!