
Naturally light uncolored and heavy colored superparticles

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Motivation

● Squarks and gluino are heavy

- $m_h = 125 \text{ GeV} \Rightarrow$ multi-TeV stops with small mixing, or TeV stops with large mixing
- Meson oscillation and decays \Rightarrow first two generation squarks heavy and degenerate
- Non-observation of squarks and gluino in 7 and 8 TeV LHC

● Sleptons and weak gauginos may be light

- Muon $(g - 2)$ has 3.5σ discrepancy \Rightarrow light smuons
- Neutralino as DM expected in 100 GeV range
- Light staus may slightly alter Higgs diphoton rate
- Collider bounds on them are not so strong

How to reconcile this splitting between colored and uncolored superparticles?

Fusion of three issues

- Gauge coupling unification even with incomplete multiplets at string scale $>$ GUT scale (Bachas, Fabre, Yanagida '96; Bastero-Gil, Brahmachari '97).
 - Adjoint octet (Σ_8) of color SU(3), adjoint triplet (Σ_3) of weak SU(2)
 - Origin of these states can be traced to the non-Goldstone modes of scalar adjoint 24-plet of SU(5)
- Presence of intermediate states characterizing GMSB.

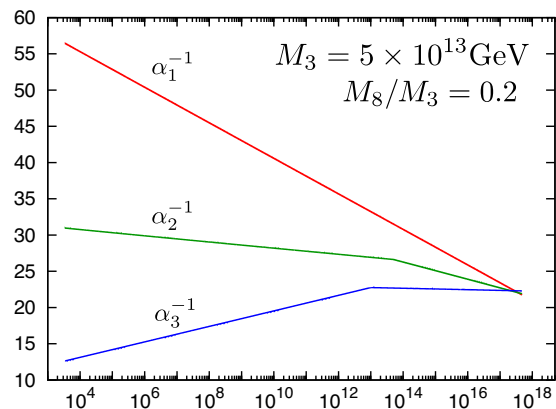
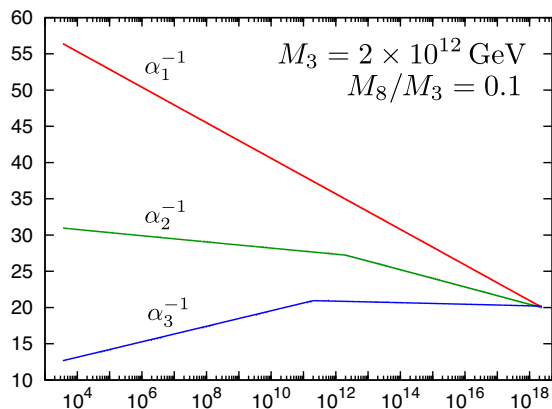
$$W_{\text{mess}} = (M_8 + \lambda_8 X) \text{Tr}(\Sigma_8^2) + (M_3 + \lambda_3 X) \text{Tr}(\Sigma_3^2)$$

F -term vev of hidden sector field X transmits SUSY breaking to visible sector via messenger multiplets.

- Dynamically ensure $\tilde{m}_{\text{color}} \gg \tilde{m}_{\text{uncolor}}$ by delinking the sources of mass generation for colored and uncolored super-particles (Han, Yanagida, Zhang '98).

Aim is to reproduce m_h , $(g - 2)_\mu$, and other data

Unification with Σ_3 and Σ_8



$$\alpha_1^{-1}(M_{\text{str}}) = \alpha_1^{-1}(m_{\text{SUSY}}) - \frac{(33/5)}{2\pi} \ln \frac{M_{\text{str}}}{m_{\text{SUSY}}}$$

$$\alpha_2^{-1}(M_{\text{str}}) = \alpha_2^{-1}(m_{\text{SUSY}}) - \frac{1}{2\pi} \ln \frac{M_{\text{str}}}{m_{\text{SUSY}}} - \frac{2}{2\pi} \ln \frac{M_{\text{str}}}{M_3}$$

$$\alpha_3^{-1}(M_{\text{str}}) = \alpha_3^{-1}(m_{\text{SUSY}}) - \frac{(-3)}{2\pi} \ln \frac{M_{\text{str}}}{m_{\text{SUSY}}} - \frac{3}{2\pi} \ln \frac{M_{\text{str}}}{M_8}$$

$$M_{\text{str}}^2 \sqrt{M_3 M_8} = M_{\text{GUT}}^3 \quad \text{with} \quad M_3 > M_8 \quad (m_{\text{SUSY}} \simeq 3.5 \text{ TeV})$$

Superparticle spectrum

- Define $\Lambda_8 \equiv \frac{\lambda_8 F_X}{M_8}$, $\Lambda_3 \equiv \frac{\lambda_3 F_X}{M_3}$
- Recall $M_3 > M_8$ (unification), tune λ_8 and λ_3 to ensure $\Lambda_8 \gg \Lambda_3$
- Messenger scale spectrum

$$m_{\tilde{B}} \simeq 0, \quad m_{\tilde{W}} \simeq \frac{g_2^2}{16\pi^2} (2\Lambda_3), \quad m_{\tilde{g}} \simeq \frac{g_3^2}{16\pi^2} (3\Lambda_8)$$

$$m_{\tilde{Q}}^2 \simeq \frac{2}{(16\pi^2)^2} \left[\frac{4}{3} g_3^4 (3\Lambda_8^2) + \frac{3}{4} g_2^4 (2\Lambda_3^2) \right], \quad m_{\tilde{D}}^2 = m_{\tilde{U}}^2 \simeq \frac{2}{(16\pi^2)^2} \frac{4}{3} g_3^4 (3\Lambda_8^2),$$

$$m_{\tilde{L}}^2 \simeq \frac{2}{(16\pi^2)^2} \frac{3}{4} g_2^4 (2\Lambda_3^2), \quad m_{\tilde{E}}^2 \simeq 0$$

- Bino mass is generated by Planck scale suppressed gravitational interaction (which also generates wino and gluino masses).

Introduce $M_{1/2}$ as universal gaugino mass at M_{str} .

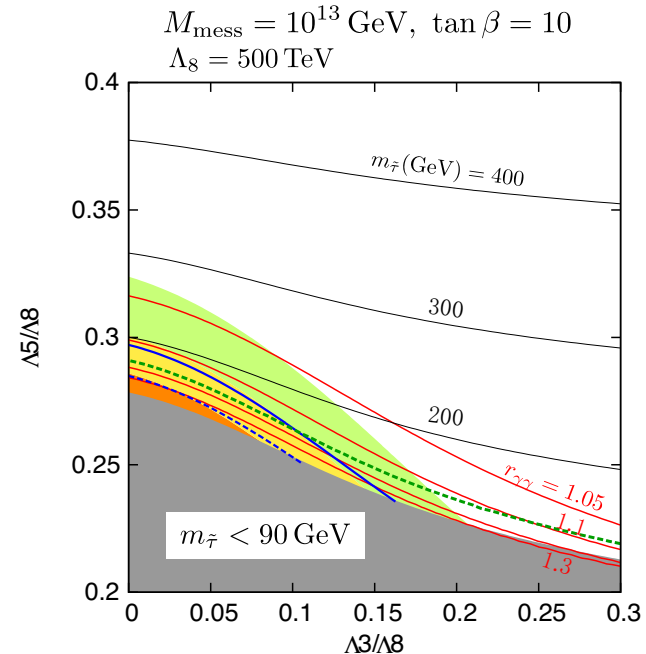
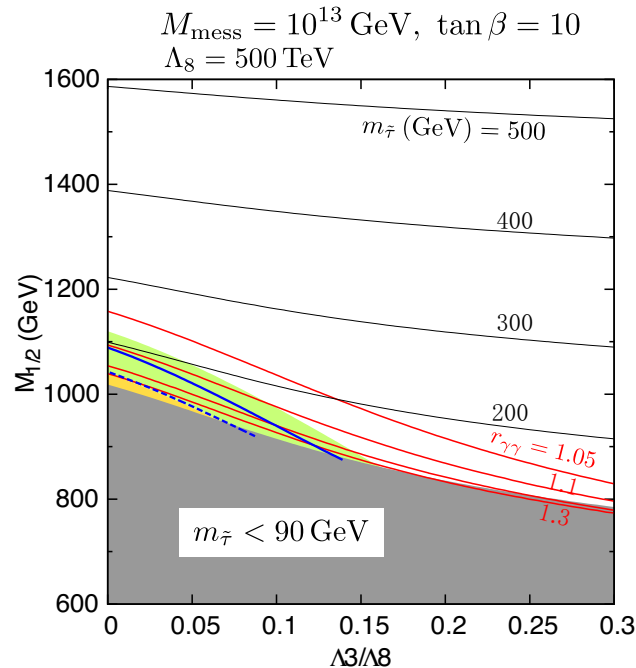
$$\int d^2\theta c_X \frac{X}{M_P} W_\alpha W^\alpha + \text{h.c.} \Rightarrow M_{1/2} = 2c_X F_X / M_P$$

- Gravitino mass is generated roughly in the same order: $m_{3/2} = \frac{F_X}{\sqrt{3}M_P}$

Sparticle spectrum (continued)

- The right-slepton mass is generated by RG running, same order as bino mass, i.e. **bino, slepton and gravitino are of the same order.**
- Due to large left-right stau mixing, one stau can be very light. Light stau can modify diphoton BR of Higgs by 10-20%.
- **Stau is the NLSP, with gravitino LSP.** However, stau is long-lived, as its decay to gravitino of the same order mass is suppressed. CMS limit: $m_{\tilde{\tau}} > 340$ GeV. This is too heavy for diphoton anomaly.
- This implies smuon is too heavy to explain muon $(g - 2)$ anomaly.
- **Way out:** Allow mild ($\leq 10^{-7}$) RPV, so that stau can promptly decay to a lepton and neutrino. Then stau can be lighter than 340 GeV. Then muon $g-2$ can be explained at slightly better than 2σ level.
- **Can we avoid RPV?**

muon ($g - 2$)



Introduction of $(5 + \bar{5})$ messengers explains muon ($g - 2$) better (right panel).

Key point: Bino/stau and gravitino mass generation de-linked. Gravitino can be ultra-light, while bino/stau can weigh around 100 GeV (since 5-plets have non-zero Y).

Unification is not affected by complete multiplets.

Region below blue solid line is excluded by vac stability limit arising from large LR slepton mixing, which sets an upper limit on $\mu \tan \beta$.

New developments

- 3-loop corrections to m_h . **Stop mass in (3-5) TeV range even with minimal mixing can reproduce $m_h = 125$ GeV** (Feng et al '13).
- Since SUSY breaking scale comes down, μ gets smaller.

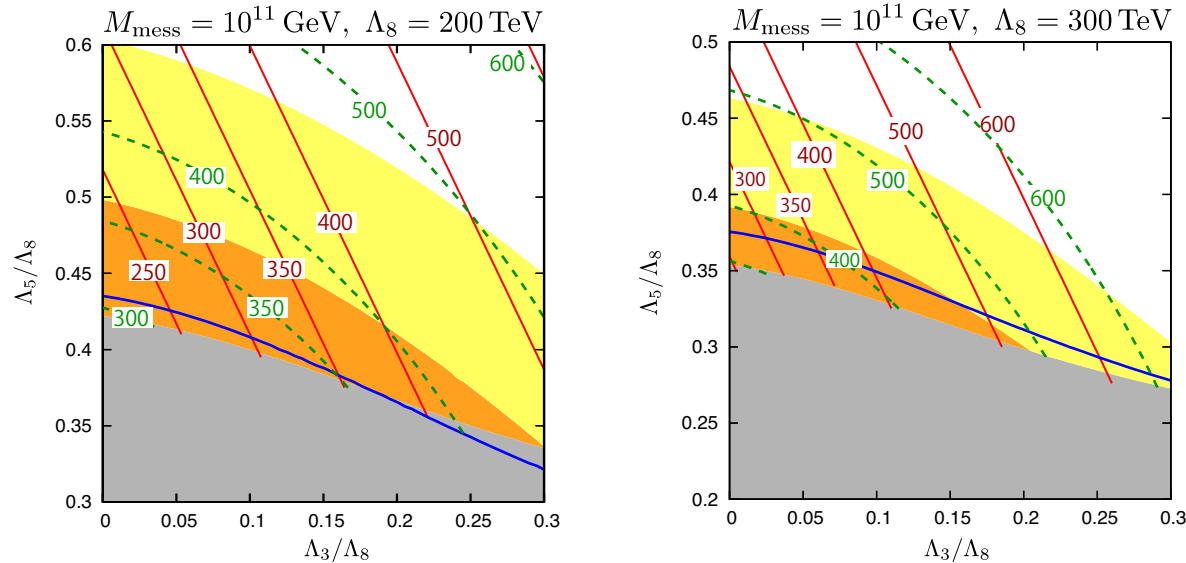
$$\mu^2 \sim (-m_{H_u}^2) \sim \frac{3}{4\pi^2} y_t^2 (m_{\text{stop}}^2) \ln \left(\frac{M_{\text{mess}}}{m_{\text{stop}}} \right)$$

- LR mixing also goes down. So $\tilde{\tau}_1$ need not be that light.
- **Bino/RH-slepton and gravitino masses de-correlated, thanks to 5-plets. Bino is NLSP.** At messenger scale $m_{\tilde{B}} = \frac{\alpha_1}{4\pi} \Lambda_5$, $m_{\tilde{E}}^2 = \frac{1}{8\pi^2} \left[\frac{3}{5} \alpha_1^2 \Lambda_5^2 \right]$.
- Gravitino can be made light

$$m_{3/2} \simeq 0.01 \text{ GeV} \left(\frac{\Lambda_8}{200 \text{ TeV}} \right) \left(\frac{(\Lambda_3/\Lambda_8)}{0.2} \right) \left(\frac{M_8}{10^{11} \text{ GeV}} \right) \left(\frac{(M_3/M_8)}{10} \right)$$

- 100 GeV Neutralino decays into 10 MeV gravitino in a BBN safe way (Kawasaki et al '08). No need of RPV.

Muon ($g - 2$) (updated)



- $(g - 2)_\mu$ is dominated by bino-smuon loop. In the orange (yellow) region it is explained at 1 (2)- σ level. In the gray region, stau is lighter than 90 GeV.

$$(\Delta a_\mu)_{\text{SUSY}} \simeq \frac{3}{5} \frac{g_1^2}{8\pi^2} \frac{m_\mu^2 \mu \tan \beta}{M_1^3} F_b \left(\frac{m_{\tilde{L}}^2}{M_1^2}, \frac{m_{\tilde{E}}^2}{M_1^2} \right)$$

- Viable regions are above the blue solid line where bino is NLSP. A stau NLSP is stable inside the detector (hence > 340 GeV (CMS '13)), which makes smuons too heavy!

Conclusions

- Does naturalness demand that super-particles all have to be simultaneously heavy? OR, sleptons/weak gauginos can remain significantly lighter than squarks/gluino by internal dynamics?
- Key observation: With unconventional choice of matter multiplets, a color SU(3) octet and a weak SU(2) triplet, GMSB works:
 - unification at string scale (between GUT and Planck scale).
 - colored mass \gg uncolored mass of sparticles by intrinsic dynamics.
 - Introducing in addition the SU(5) 5-plets, it is possible to explain Muon ($g - 2$) within 1σ . Consequences of 3-loop correction to Higgs mass have been crucial.
 - Lighter stau is in (100-200) GeV range which can be a target at ILC.

GMSB with incomplete matter multiplets is an attractive option.