Instrumentation

Content

Introduction

Part 1: Passage of particles through matter

- Charges particles, Photons, Neutrons, Neutrinos
- Momentum measurements, Combining measurements.

Part 2: Particle Detection

- Ionisation detector
- Scintillation detectors
- Semiconductor detectors

Instrumentation

Experimental Particle Physics

Accelerators

Luminosity, energy, quantum numbers

Detectors

Efficiency, speed, granularity, resolution

Trigger/DAQ

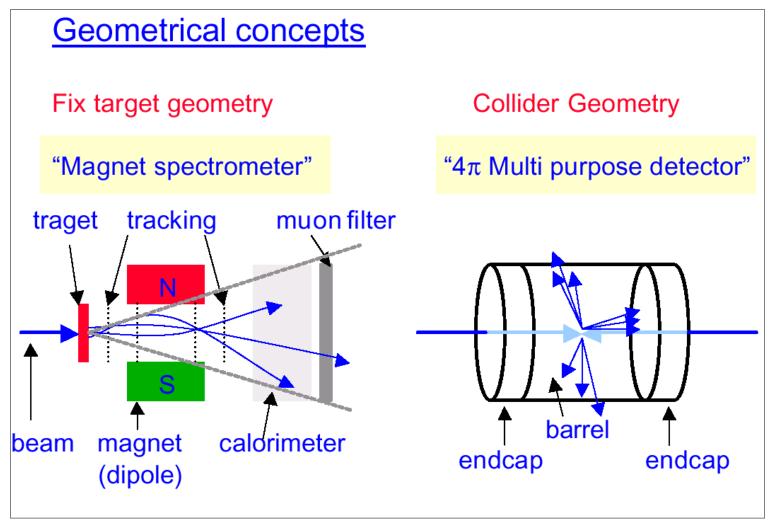
Efficiency, compression, through-put, physics models

Offline analysis

Signal and background, physics models.

The primary factors for a successful experiment are the accelerator and detector/trigger system, and losses there are not recoverable.

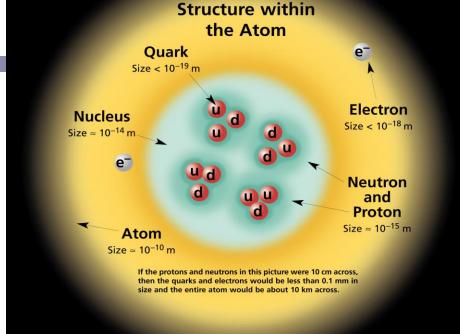
Detector systems



From C.Joram

Concentrate on electromagnetic forces since a combination of their strength and reach make them the primary responsible for energy loss in matter.

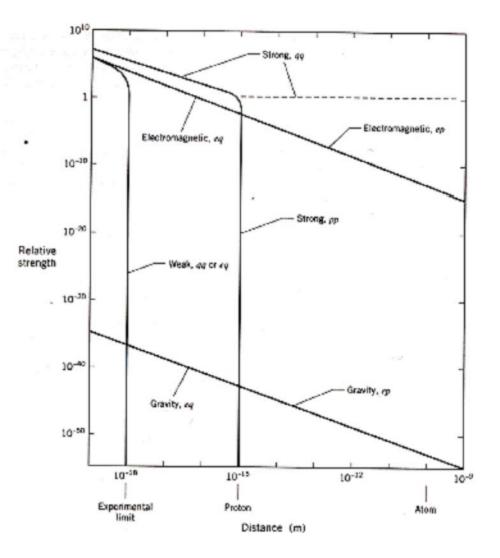
For neutrons, hadrons generally and neutrinos other effects obviously enter.



PROPERTIES OF TH

Property Interaction	Gravitational	Weak	Electromagnetic	Strong	
		(Electroweak)		Fundamental	Residual
Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge	See Residual Strong Interaction Note
Particles experiencing:	All	Quarks, Leptons	Electrically charged	Quarks, Gluons	Hadrons
Particles mediating:	Graviton (not yet observed)	W+ W- Z ⁰	γ	Gluons	Mesons
Strength relative to electromag 10 ⁻¹⁸ m	10 ⁻⁴¹	0.8	1	25	Not applicable
for two u quarks at: 3×10^{-17} m	10 ⁻⁴¹	10 ⁻⁴	1	60	to quarks
for two protons in nucleus	10 ⁻³⁶	10 ⁻⁷	1	Not applicable to hadrons	20

Strength versus distance



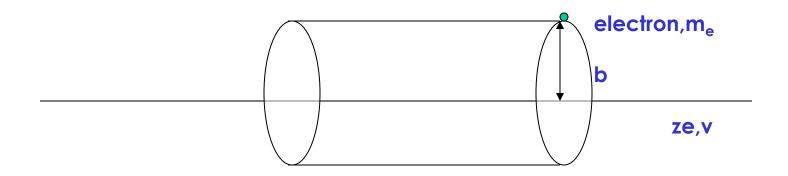
- At atomic distances only EM & gravity have sizable strengths
- EM is ~40 orders of magnitude stronger than gravity
- If quarks could be separated force would be enormous (see dashed line)
- At proton size distances strong force turns on & becomes ~ 100 times stronger than EM force
- At distances ~ 1/1000 of proton size weak force turns on abruptly

Heavy charged particles

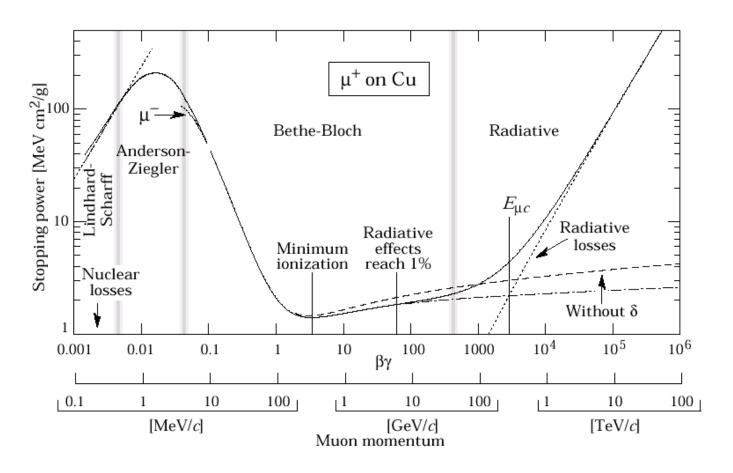
Heavy charged particles transfer energy mostly to the atomic electrons, ionising them. We will later come back to not so heavy particles, in particular electrons/positrons.

Usually the Bethe Bloch formally is used to describe this - and most of features of the Bethe Bloch formula can be understood from a very simple model :

- 1) Let us look at energy transfer to a single electron from heavy charged particle passing at a distance b
- 2) Let us multiply with the number of electrons passed
- 3) Let us integrate over all reasonable distances b



Heavy charges particles

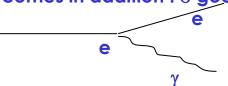


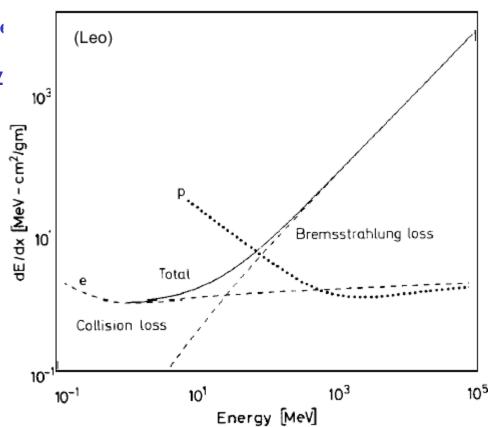
$$\left\langle \frac{dE}{dx} \right\rangle = -4\pi N_A r_e^2 m_e c^2 z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[\frac{1}{2} \ln \frac{2m_e c^2 \gamma^2 \beta^2}{I^2} T^{\text{max}} - \beta^2 - \frac{\delta}{2} \right]$$

Electrons and Positrons

Electrons/positrons; modify Bethe Bloch to take into account that incoming particle has same mass as the atomic electrons

Bremsstrahlung in the electrical field of a charge Z comes in addition : σ goes as $1/m^2$





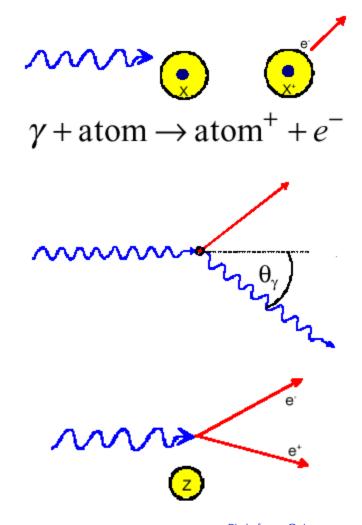
Photons

Three processes:

Photoelectric effect (Z⁵); absorption of a photon by an atom ejecting an electron. The cross-section shows the typical shell structures in an atom.

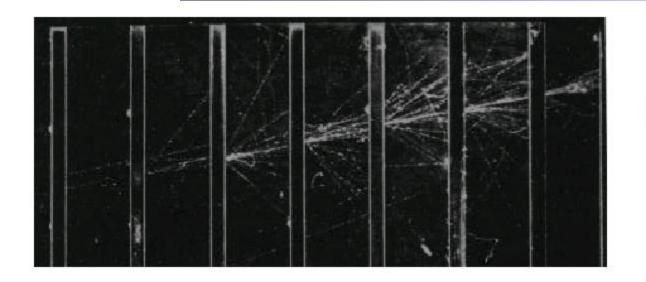
Compton scattering (Z); scattering of a photon again a free electron (Klein Nishina formula). This process has well defined kinematic constraints (giving the so called Compton Edge for the energy transfer to the electron etc) and for energies above a few MeV 90% of the energy is transferred (in most cases).

Pair-production (Z²+Z); essentially bremsstrahlung again with the same machinery as used earlier; threshold at 2 m_e = 1.022 MeV. Dominates at a high energy.



Plots from C.Joram

Electromagnetic calorimeters



Electron shower in a cloud chamber with lead absorbers

From C.Joram

Considering only Bremsstrahlung and Pair Production with one splitting per radiation length (either Brems or Pair) we can extract a good model for EM showers.

Text from C.Joram

Longitudinal shower development For Iron: a = 9.4, b=39 $t_{\text{max}}(\lambda_I) \approx 0.2 \ln E[GeV] + 0.7$ E =100 GeV $t_{95\%} \approx a \ln E + b$ $\rightarrow t_{95\%} \approx 80 \text{ cm}$ 1000 (C. Fabjan, T. Ludlam, CERN-EP/82-37) Longitudinal energy deposit (arbitrary units) 100 GeV 15 GeV Longitudinal depth (λ)

• Additional strong interactions for hadrons (p,n, etc); hadronic absorption/interaction length and hadronic showers

Neutrinos interact only weakly → tiny cross-sections For their detection we need again first a charged particle.

Possible detection reactions:

- v_{ℓ} + n $\rightarrow \ell^{-}$ + p ℓ = e, μ , τ
- \overline{v}_{ℓ} + p $\rightarrow \ell^+$ + n ℓ = e, μ , τ

The cross-section for the reaction $v_e^- + n \rightarrow e^- + p$ is of the order of 10⁻⁴³ cm² (per nucleon, $E_n^- \approx$ few MeV).

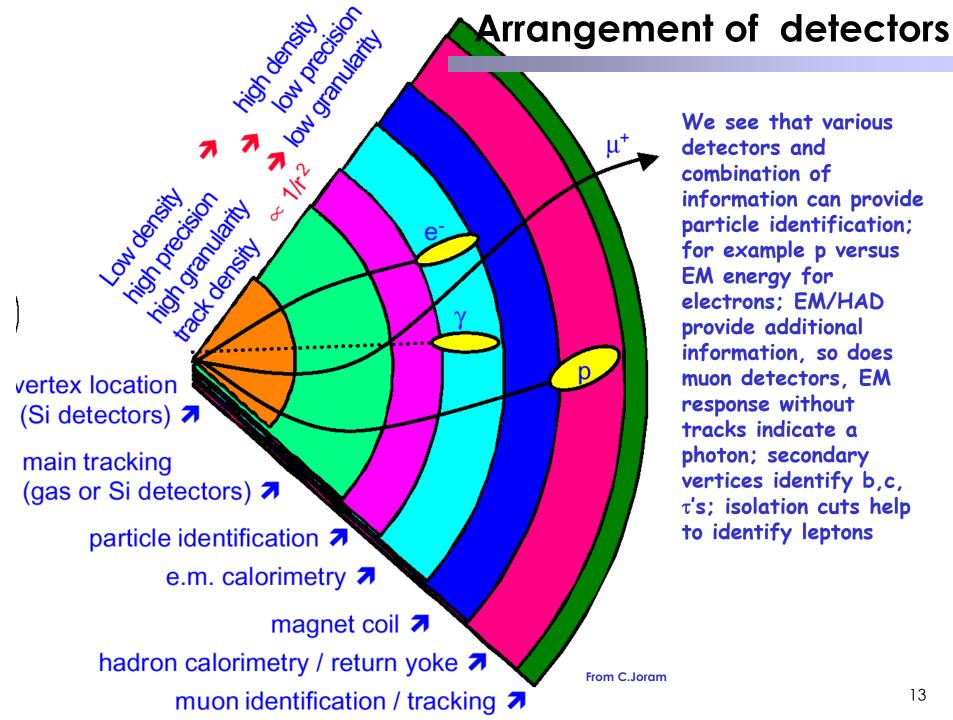
$$ightarrow$$
 detection efficiency $\varepsilon_{\rm det} = \sigma \cdot N^{surf} = \sigma \cdot \rho \frac{N_A}{A} d$
1 m Iron: $\varepsilon_{\rm det} \approx 5 \cdot 10^{-17}$

Neutrino detection requires big and massive detectors (ktons) and high neutrino fluxes.

In collider experiments fully hermetic detectors allow to detect neutrinos indirectly:

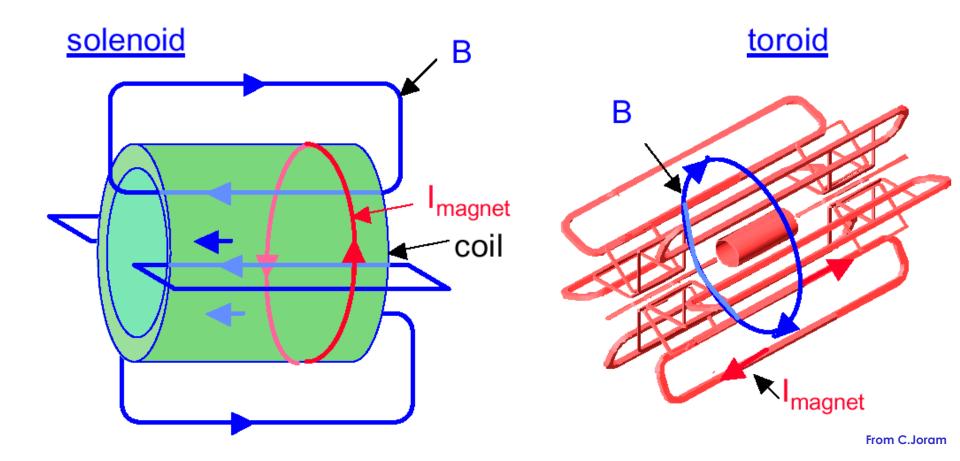
- Sum up all visible energy and momentum.
- ◆ Attribute missing energy and momentum to neutrino.

Neutrinos



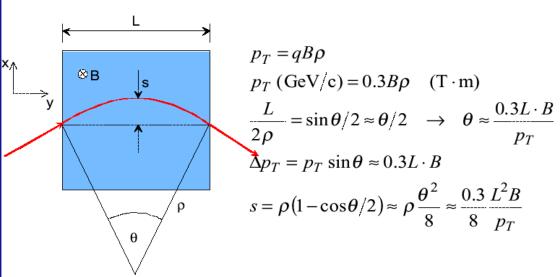
Magnetic fields

Magnetic field configurations:



See the Particle Data Book for a discussion of magnets, stored energy, fields and costs.

Momentum measurement



the sagitta s is determined by 3 measurements with error $\sigma(x)$: $s = x_2 - \frac{x_1 + x_3}{x_1 + x_2}$

error
$$\sigma(x)$$
: $s = x_2 - \frac{x_1 + x_3}{2}$

$$\frac{\sigma(p_T)}{p_T}\Big|_{meas.} = \frac{\sigma(s)}{s} = \frac{\sqrt{\frac{3}{2}}\sigma(x)}{s} = \frac{\sqrt{\frac{3}{2}}\sigma(x) \cdot 8p_T}{0.3 \cdot BL^2}$$

for N equidistant measurements, one obtains

(R.L. Gluckstern, NIM 24 (1963) 381)

$$\frac{\sigma(p_T)}{p_T}\bigg|_{p_T}^{meas.} = \frac{\sigma(x) \cdot p_T}{0.3 \cdot BL^2} \sqrt{720/(N+4)} \qquad \text{(for N $\geq \approx 10$)}$$

ex: $p_T=1$ GeV/c, L=1m, B=1T, $\sigma(x)=200\mu m$, N=10

$$\left. \frac{\sigma(p_T)}{p_T} \right|^{meas.} \approx 0.5\%$$
 (s ≈ 3.75 cm)

Magnetic fields

Instrumentation

Lecture set 2: How are reactions of the various particles with detectors turned into electrical signals. We would like to extract position and energy information channel by channel from our detectors.

Three effects are usually used:

- 1 Ionisation
- 2 Scintillation
- 3 Semi Conductors

and these are used in either for tracking, energy measurements, photon detectors for Cherenkov or TRT, etc

and from then on it is all online (trigger, DAQ) and offline treatment and analysis

Ionisation Detectors

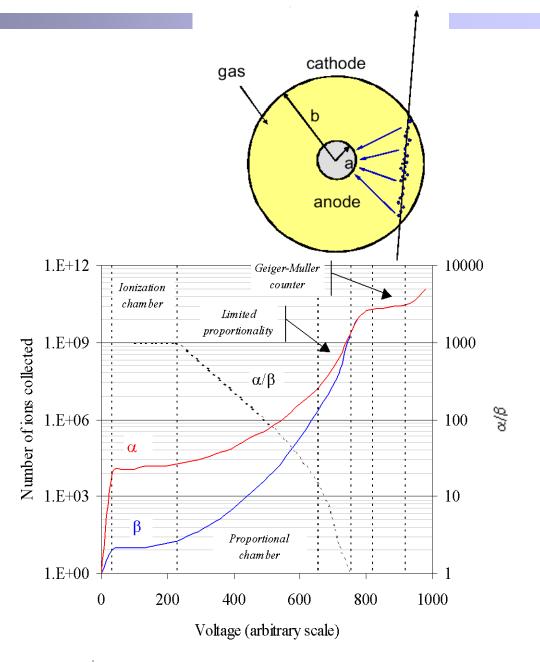
The different regions:

Recombination before collection.

lonisation chamber; collect all primary charge. Flat area.

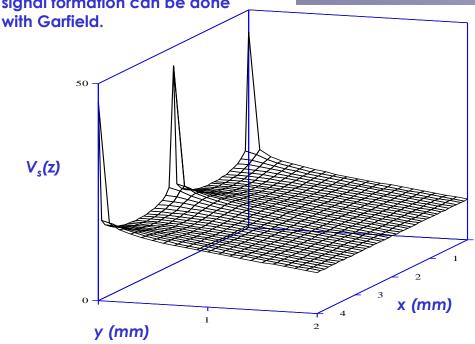
Proportional counter (gain to 10⁶); secondary avalanches need to be quenched. Limited proportionality (secondary avalanches distorts field, more quenching needed).

Geiger Muller mode, avalanches all over wire, strong photoemission, breakdown avoided by cutting HV.



Advanced calculations of electric field, drift, diffusion and signal formation can be done

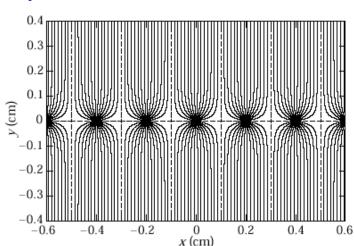
Ionisation Detectors

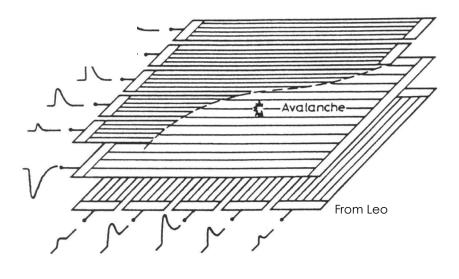


Two dimensional readout can be obtained by; crossed wires, charge division with resistive wires, measurement of timing differences or segmented cathode planes with analogue readout

Resolution given by (binary readout): $\sigma = d/\sqrt{12}$

Analogue readout and charge sharing can improve this significantly when the left/right signal size provide more information about the hit position.



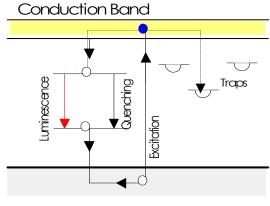


$$y = \frac{\sum (Q_i - b) y_i}{\sum (Q_i - b)} ,$$

Scintillators

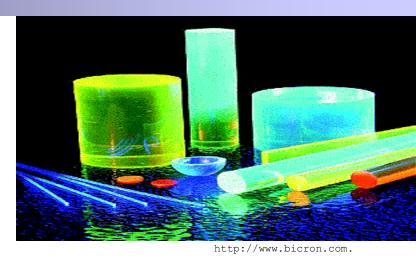
Inorganic Crystalline Scintillators
The most common inorganic scintillator is
sodium iodide activated with a trace
amount of thallium [NaI(TI)],

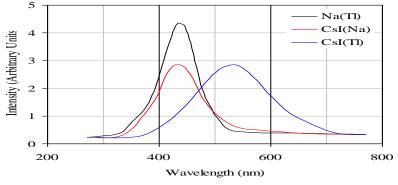
Energy bands in impurity activated crystal

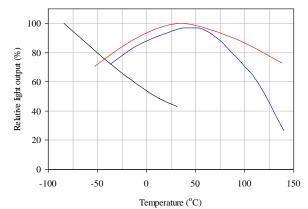


Valence Band

Strong dependence of the light output and the decay time with temperature.

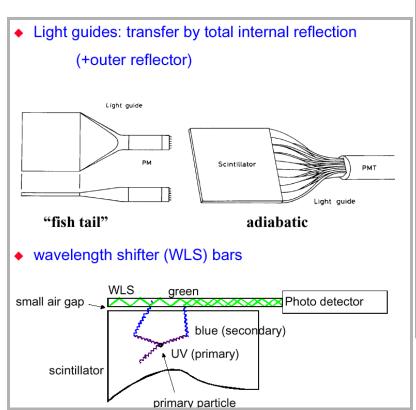


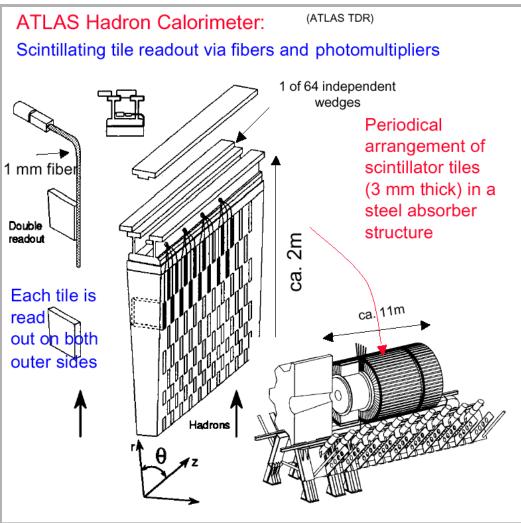




Scintillators

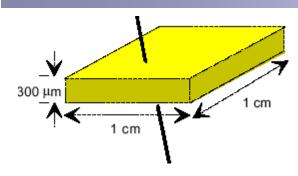
External wavelength shifters and light guides are used to aid light collection in complicated geometries; must be insensitive to ionising radiation and Cherenkov light. See examples.





From C.Joram

Semi-Conductors

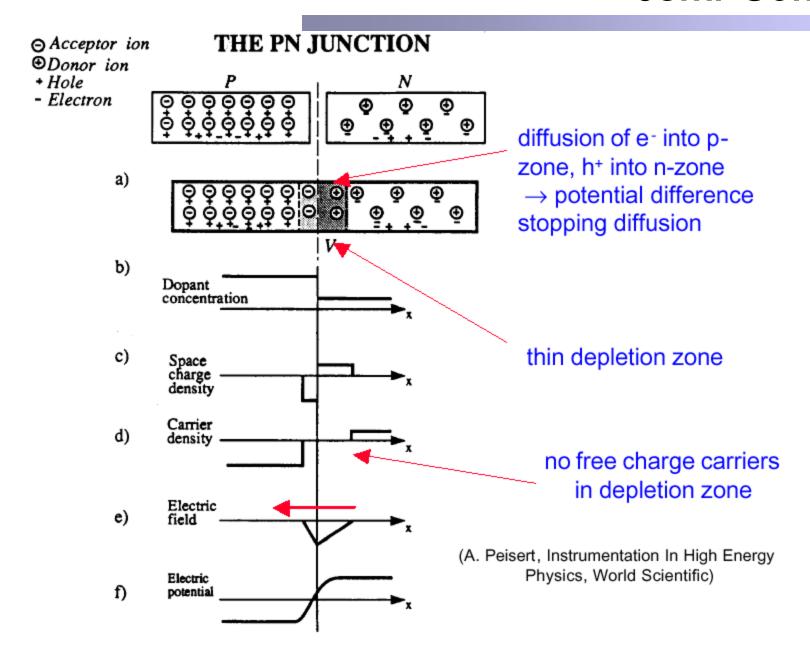


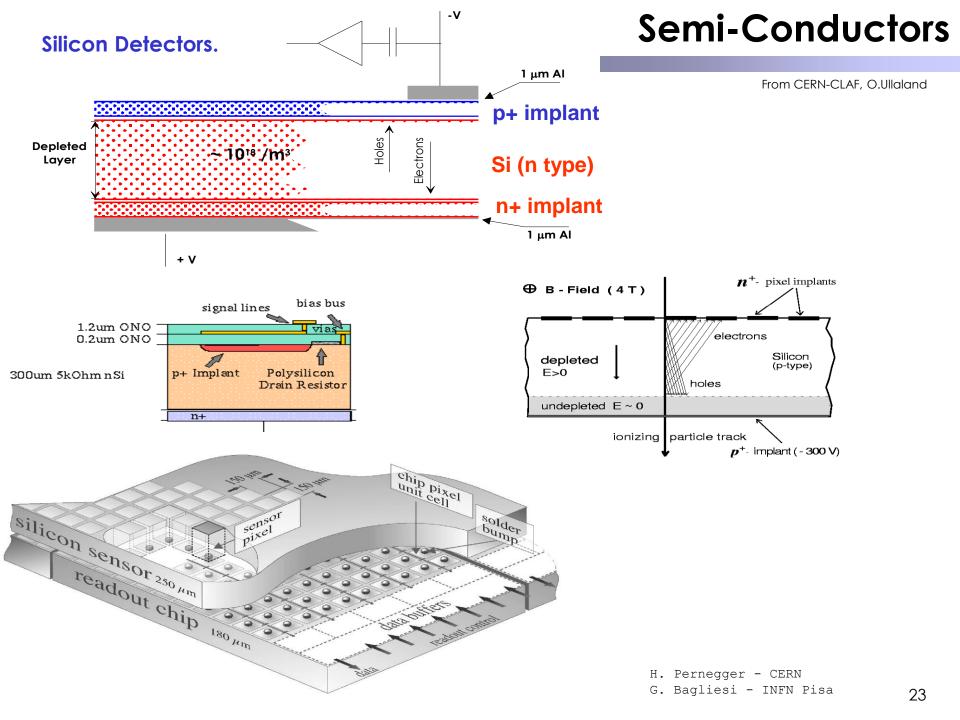
Intrinsic silicon will have electron density = hole density; 1.45 10¹⁰ cm⁻³ (from basic semiconductor theory).

In the volume above this would correspond to 4.5 10⁸ free charge carriers; compared to around 3.2 10⁴ produces by MIP (Bethe Bloch loss in 300um Si divided by 3.6 eV).

Need to decrease number of free carriers; use depletion zone (reduce temperature would also help but one would need to go to cryogenic temperatures)

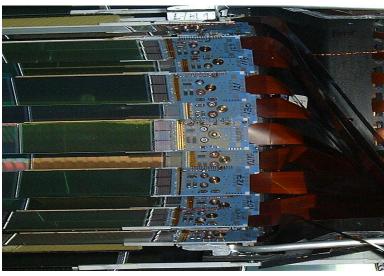
Semi-Conductors



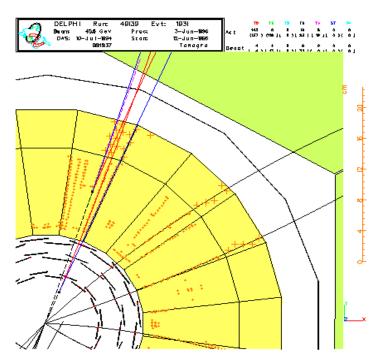


Semi-Conductors

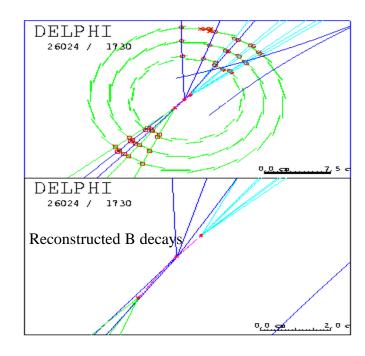
The DELPHI Vertex Detector



K0 and Lambda reconstruction

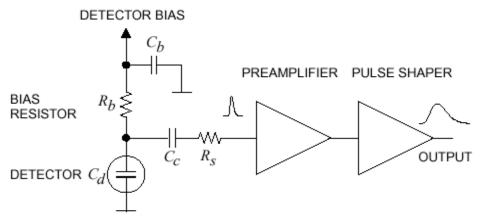


From CERN-CLAF, O.Ullaland



Front End electronics

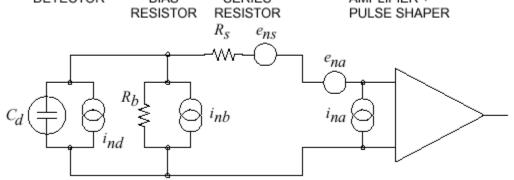
Most detectors rely critically on low noise electronics. A typical Front End is shown below:



where the detector is represented by the capasitance C_d , bias voltage is applied through R_b , and the signal is coupled to the amplifier though a capasitance C_c . The resistance R_s represent all the resistances in the input path. The preamplifier provides gain and feed a shaper which takes care of the frequency response and limits the duration of the signal.

The equivalent circuit for noise analysis includes both current and voltage noise sources labelled in and en respectively. Two important noise sources are the detector leakage current (fluctuating-some times called shot noise) and the electronic noise of the amplifier, both unavoidable and therefore important to control and reduce. The diagram below show the noise sources and their representation in the noise analysis:

DETECTOR BIAS SERIES AMPLIFIER +



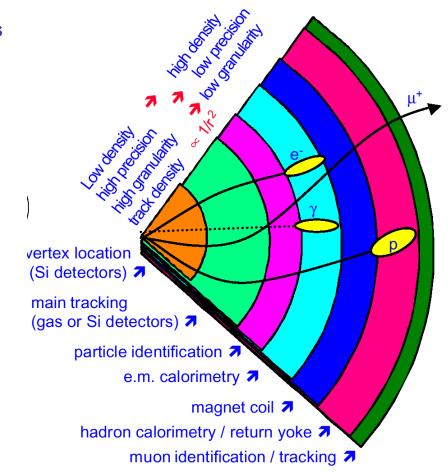
Instrumentation

We now know how most particles (i.e all particles that live long enough to reach a detector; e,u,p, π ,k,n, γ , neutrinos,etc) react with matter.

We now know how to identify particles to some extend, how to measure E and p, v, and how to measure lifetimes using secondary vertices, etc

Essential three detector types are used:

- 1 Ionisation detectors
- 2 Scintillators
- 3 Semi Conductors
- 4 Finally we have looked briefly at how electrical signals are treated in FE electronics



The detector-types mentioned are either for tracking, energy measurement, photon detectors, etc in various configurations.

