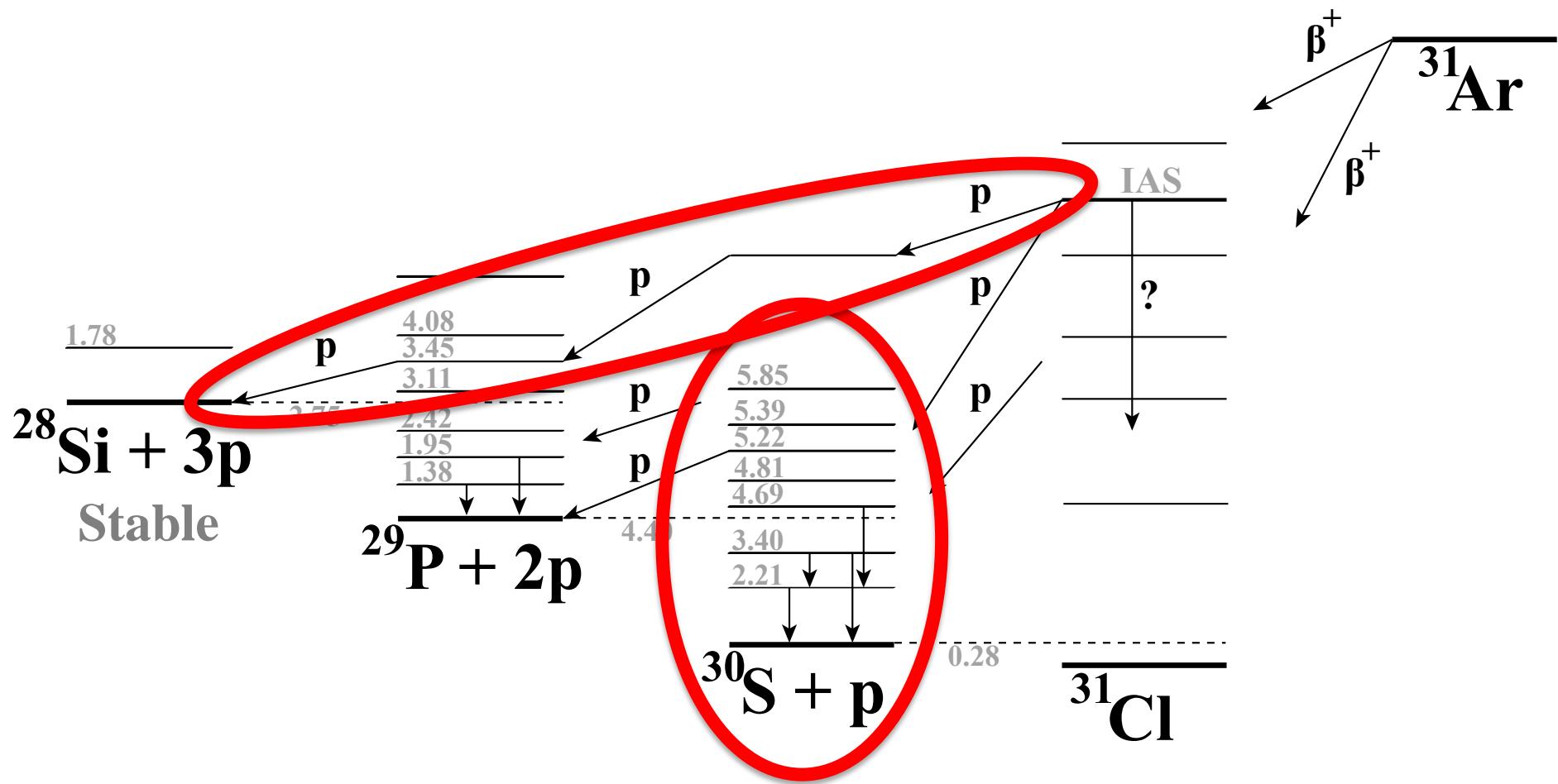


β^3 p spectroscopy and proton- γ width determination in the decay of ^{31}Ar

Presented by
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The decay of ^{31}Ar



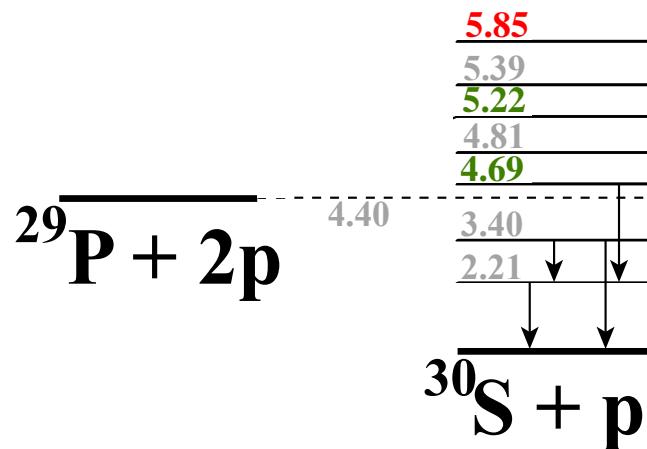
Levels of ^{30}S important for astrophysics

IS476 in 2009:

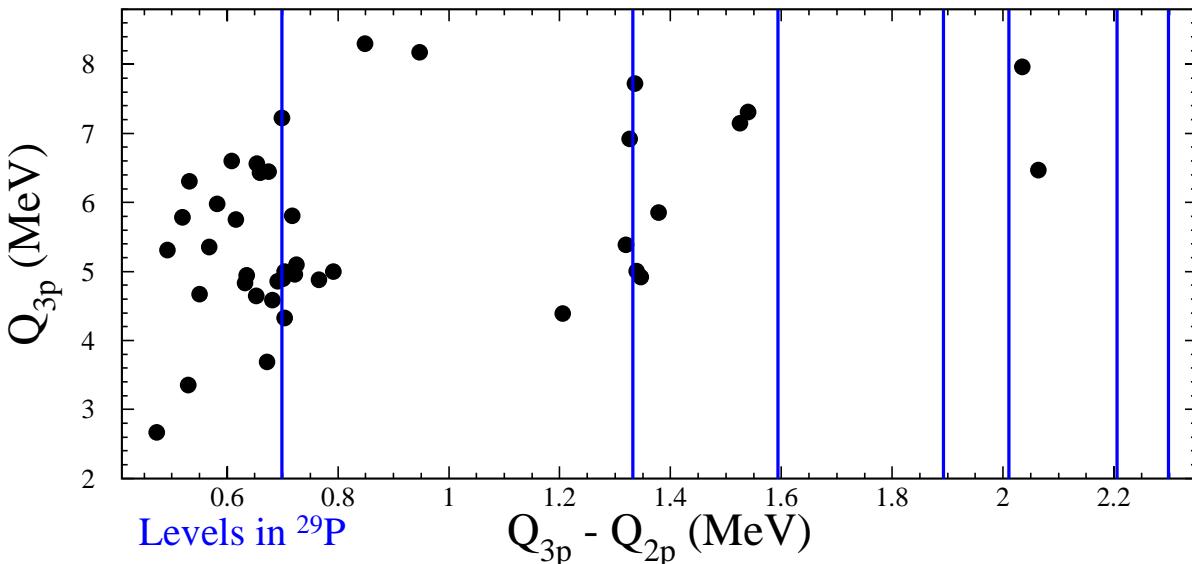
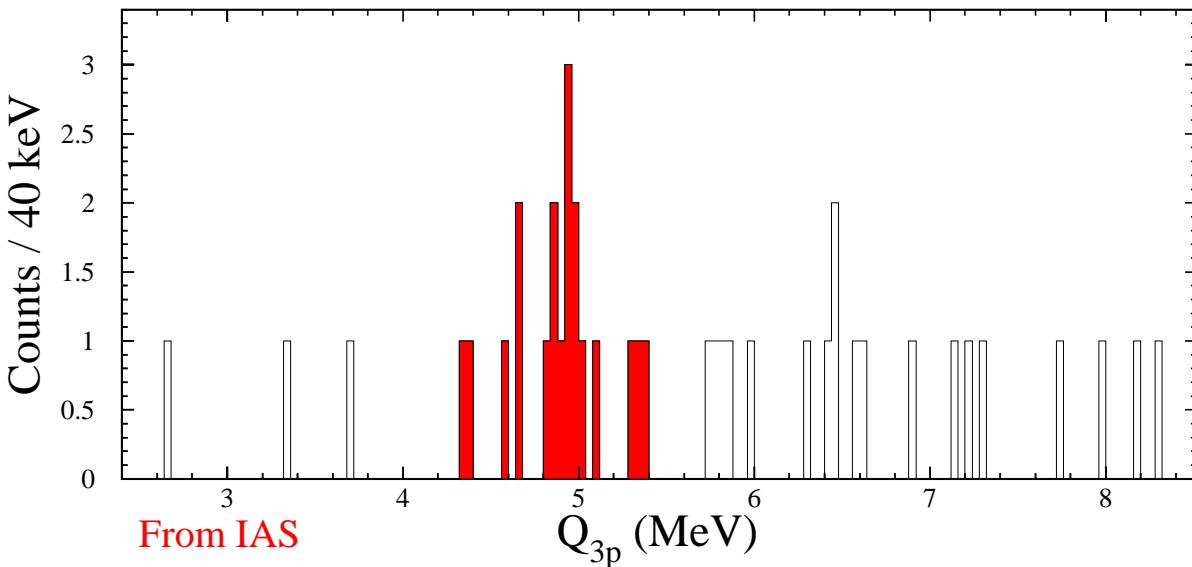
- Limits on the Γ_p/Γ_γ ratio for 3 levels
 - 1 contradict theory
 - 2 two orders of magnitude higher
- Phys. Rev. C 87 (2013) 055808
- Spin from p-p correlations
- Statistics limited...

This proposal:

- Γ_p/Γ_γ ratios:
 - 40 protons from 4.69 MeV
 - 4 γ 's from 5.22 MeV
 - γ 's and protons from 4.81 MeV
- Spin assignment of 5.22 MeV



β 3p-decay



IS476 in 2009:

- Approx. 30 3p-events
- Identified β 3p-decay
 - From IAS
 - Higher lying levels
- (partly) sequential
- Populate higher lying levels in ^{29}P
- Publication in preparation

This proposal:

- Approx. 500 3p-events
- Precise B.R. for IAS
 - both 3p and 2p γ
- Thorough investigation of decay mode

Setup

IS476 in 2009:

- 42 % of 4π particle detection
 - 6 DSSSD's
 $4 \times 300 \mu\text{m}$, $1 \times 69 \mu\text{m}$, $1 \times 500 \mu\text{m}$
 - Only backing on 4
- Approx. 6 % of $4\pi \gamma$ detection
 - 2 Miniball detectors
- CaO target
 - Yield: approx. 1 ^{31}Ar per sec.
 - $^{16,17}\text{N}$ contamination (N_2)
- 25 shifts used of 27

This proposal:

- 70 % of 4π particle detection
 - Compact setup of 6 DSSSD's with backing
- Approx. 25 % of $4\pi \gamma$ detection
 - **ISOLDE Decay Station**
- CaO nano-structured target:
 - Yield: $10 \, ^{31}\text{Ar}$ per sec.
 - Less contamination
- 24 shifts in total

Summary

24 shifts: CaO nano-structured target

^{30}S :

- Γ_p/Γ_γ ratios:
 - 40 protons from 4.69 MeV
 - 4 γ 's from 5.22 MeV
 - γ 's and protons from 4.81 MeV
- Spin assignment of 5.22 MeV

β 3p:

- Approx. 500 3p-events
- Thorough investigation of decay mode
- Which states in ^{31}Cl

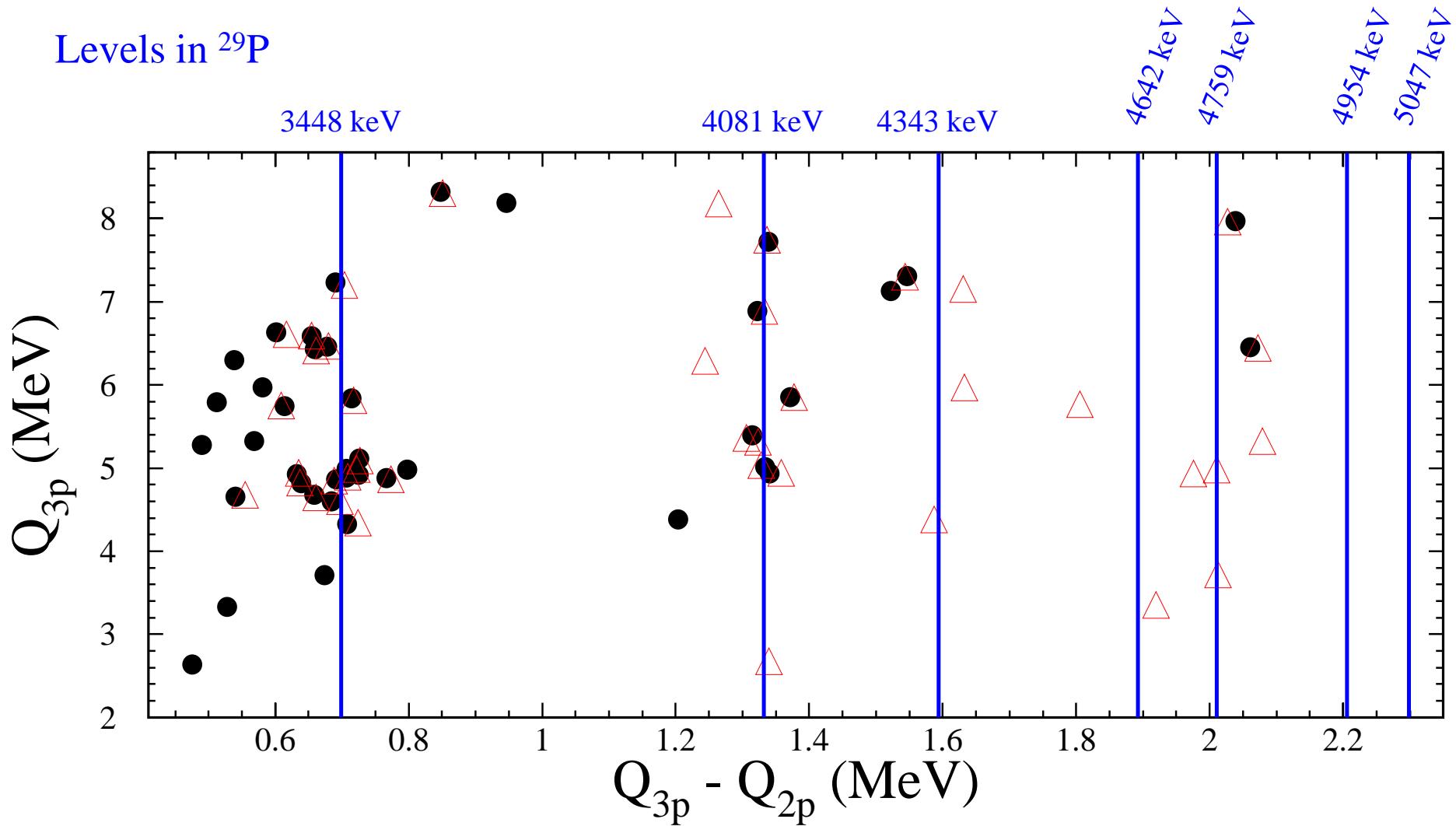
β -strength:

- Fermi strength
 - IAS 1p-, 2p-, 3p, γ -decay
- Gamow-Teller strength
 - Correctly assigned above IAS

Bonus slides

β 3p-decay

Levels in ^{29}P

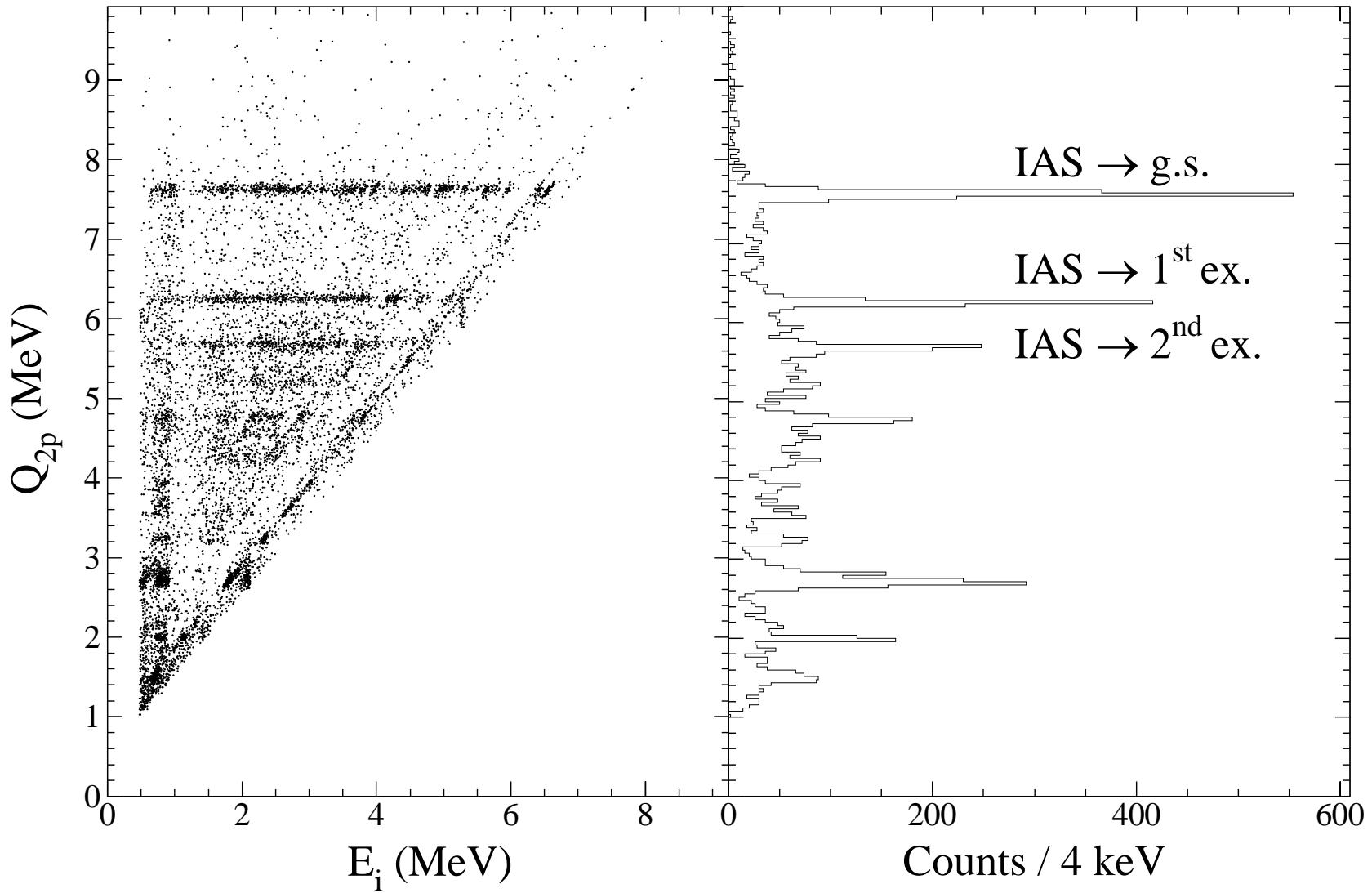


Spin determination from proton-proton angular correlations

	^{31}Cl	A_2	$\Delta\chi^2$
$W(\cos \theta) = \sum_{\nu=0}^{\nu_{\max}} A_\nu P_\nu(\cos \theta)$	7.7	0.42(16)	6.52
	7.0	0.11(22)	0.15
IAS	0.5(3)	2.95	

$\begin{array}{c} ^{31}\text{Cl} \\ \diagdown \\ ^{30}\text{S} \end{array}$	$\frac{3}{2}^+$	$\frac{5}{2}^+$	$\frac{7}{2}^+$
0 ⁺	0	0	0
1 ⁺	0	0	0
2 ⁺	0	0	[-0.7; -0.25]
3 ⁺	[0.15; 0.87]	0	0
4 ⁺	[0.76; 1.00]	[0.13; 0.95]	0
0 ⁻	0	0	0
1 ⁻	[-0.80; -0.64]	[-0.60; -0.75]	[-0.67; -0.53]

β^2 p-events



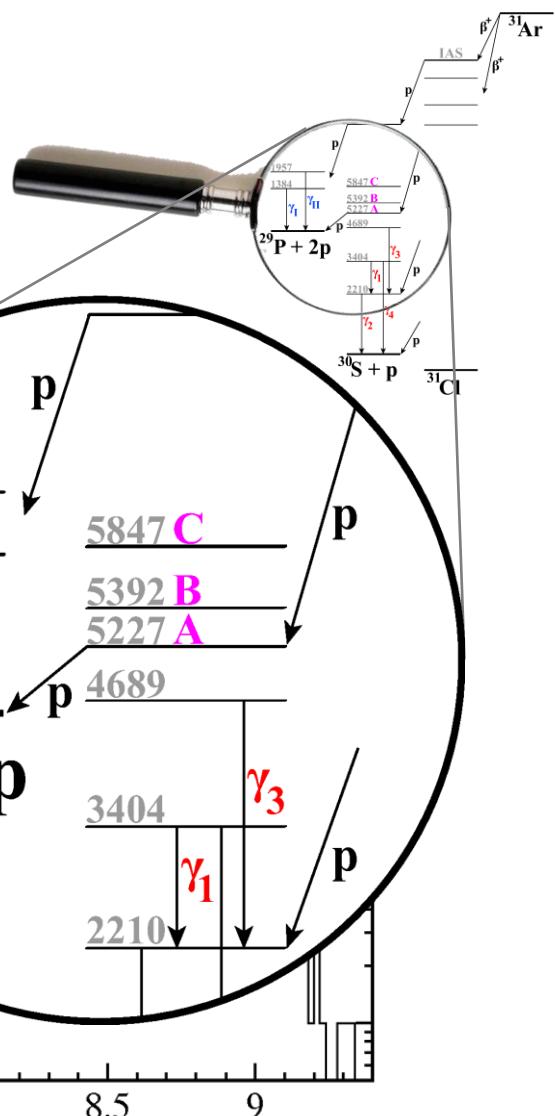
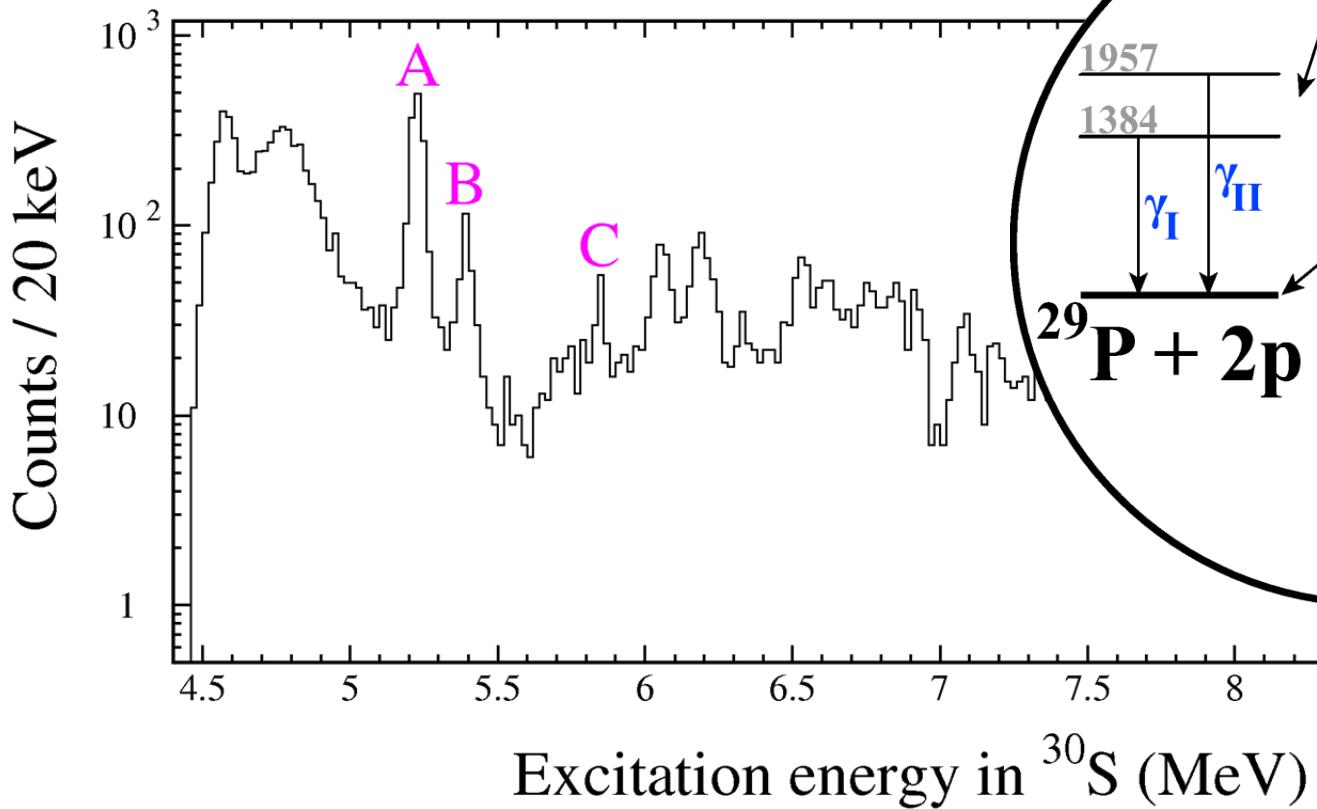
B.R. for the decay of IAS

TABLE I. Total theoretical branching ratio is 4.17(10) %.

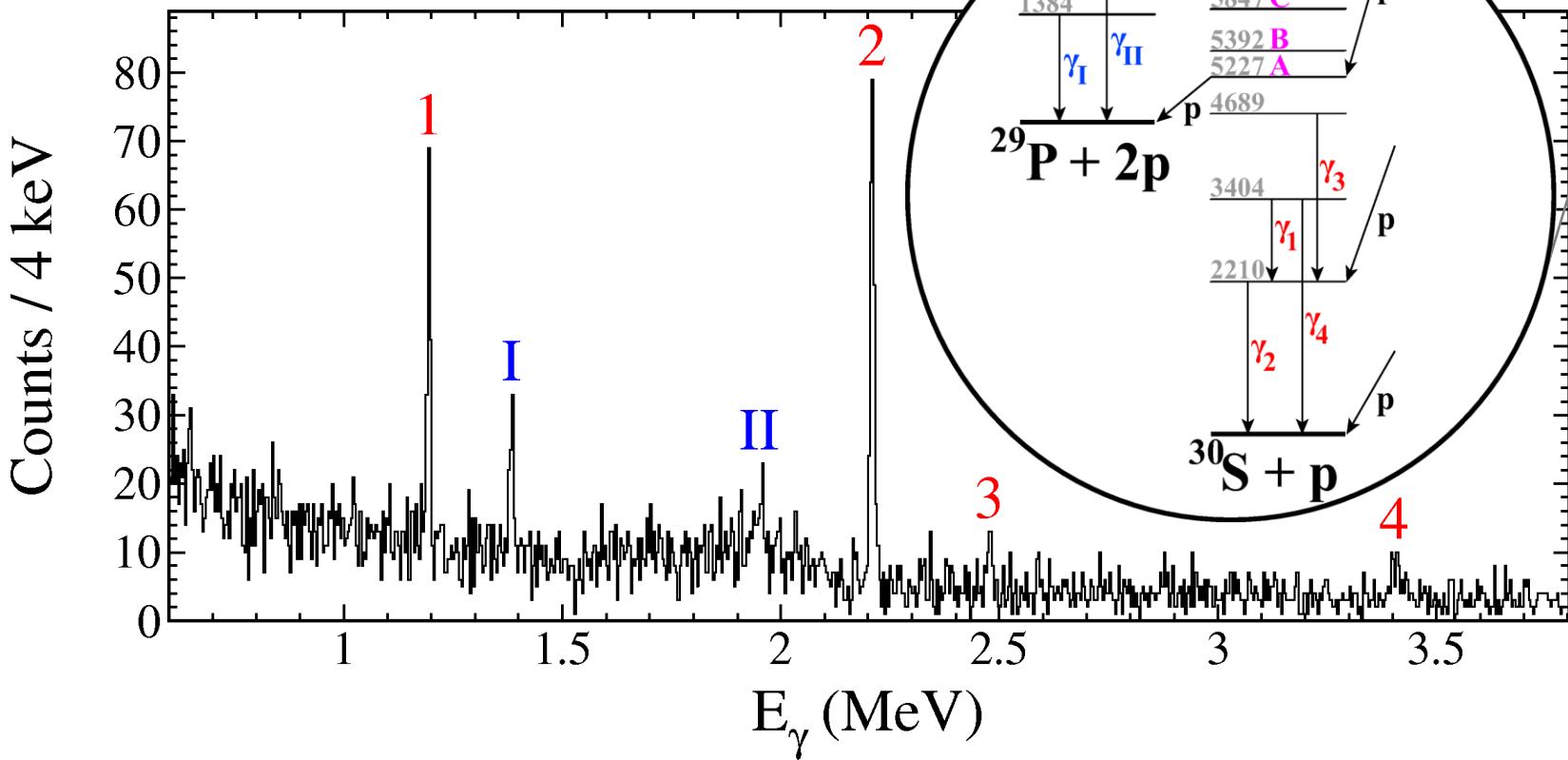
Three-proton branch				
Final state in ^{28}Si (keV)	J^π	Q_{3p} (MeV)	E_{IAS} (MeV)	B.R. (%)
0	0^+	4.89(29)	12.32(29)	0.04(2)
Two-proton branch				
Final state in ^{29}P (keV)	J^π	Q_{2p} (MeV)	E_{IAS} (MeV)	B.R. (%)
0	$\frac{1}{2}^+$	7.633(4)	12.311(6)	1.47(28)
1383.55(7)	$\frac{3}{2}^+$	6.251(4)	12.313(6)	0.88(18)
1953.91(17)	$\frac{5}{2}^+$	5.688(6)	12.320(8)	0.40(11)
2422.7(3)	$\frac{3}{2}^+$	5.22(8)	12.32(8)	0.07(5)
One-proton branch				
Final state in ^{30}S (keV)	J^π	E_p (MeV)	E_{IAS} (MeV)	B.R. (%)
0	0^+	11.57(8)	12.24(8)	0.049(12)
2210.2(1)	2^+	9.46(8)	12.27(8)	0.104(21)
3404.1(1)	2^+	8.33(8)	12.30(8)	0.108(21)
3667.7(3)	0^+	8.08(8)	12.30(8)	0.101(24)
Total			12.313(4)	3.22(36)

^{30}S spectrum

$$E(^{30}\text{S}) = Q_{2p} + E(^{29}\text{P}) - \frac{M(^{30}\text{S}) + m_p}{M(^{30}\text{S})} E_1 + S_{p2}$$



γ -spectrum



β^+ ^{31}Ar

β^+

JAS

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

p

Levels of ^{30}S important for astrophysics

Ritcher & Brown 2013			IS476		
E_x (keV)	J^π	Γ_γ/Γ_p	E_x (keV)	Γ_γ/Γ_p	
4688	3^+	372	4689.2(24)	> 3.8	
4809	2^+	2.2			
5130*	4^+	3.5×10^{-2}			
5218	0^+	9.3×10^{-4}			
5219	3^+	4.0×10^{-3}	5227(3)	< 0.5	
5312	3^-	5.2×10^{-3}			
5382	2^+	7.6×10^{-3}			
5836*	4^+	16	5847(4)	< 9	

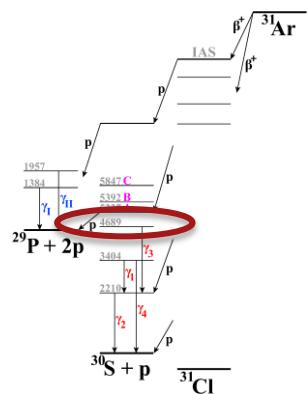
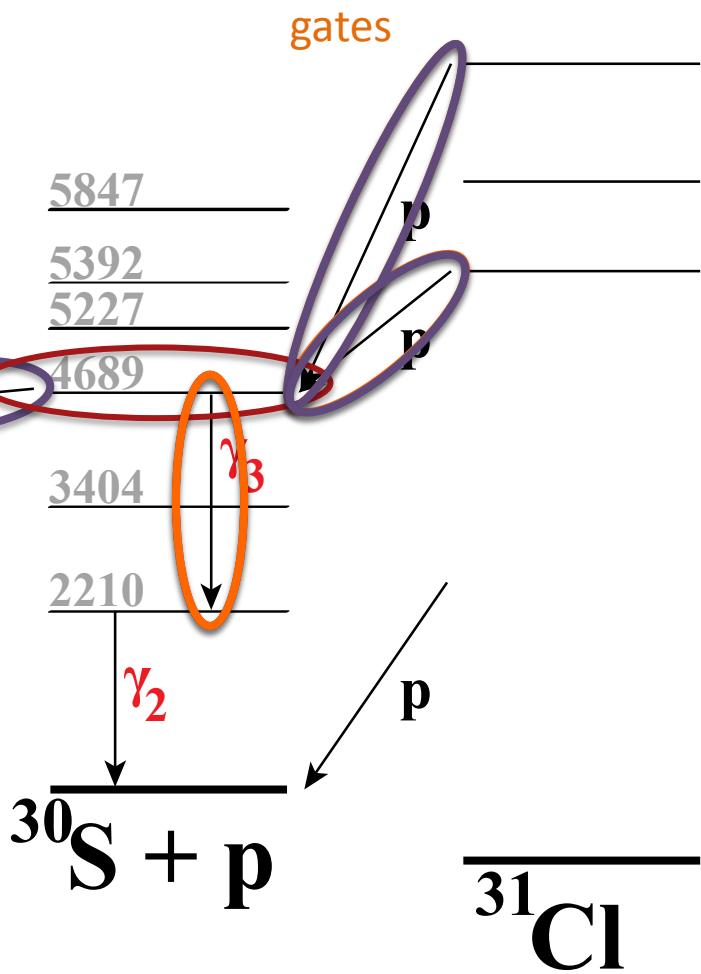
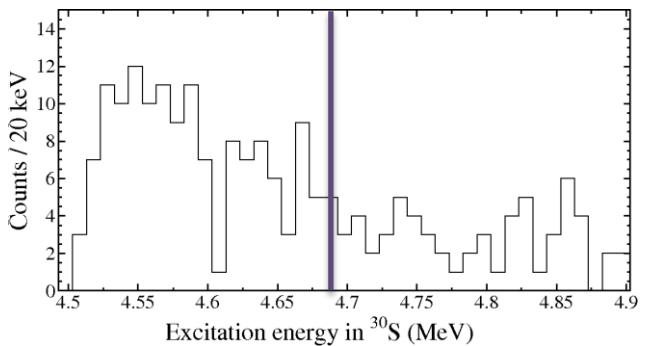
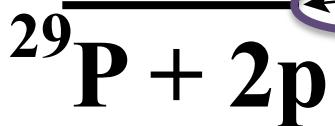
$^{30}\text{S} + \text{p}$

*From Almaraz-Calderon *et al.* 2012

Γ_p/Γ_γ for the 4689 keV level

$$\#p < 16 \pm 5$$

$$\#\gamma = 10 \pm 4$$



Levels of ^{30}S important for astrophysics

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