

Complementarity in direct dark matter searches

Miguel Peiró

Based on arXiv: 1304.1758 & 1403.3539

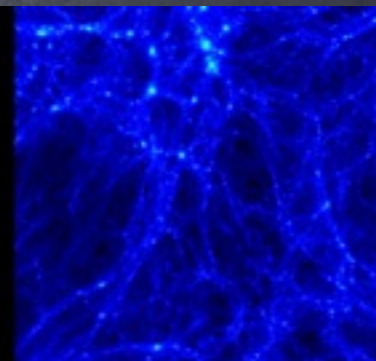
Work done in collab. with: D. G. Cerdeño, M. Fornasa, J-H. Huh, C. Marcos, and
the ROSEBUD collaboration.



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MultiDark

Multimessenger Approach
for Dark Matter Detection



Complementarity in direct dark matter searches

An update

Miguel Peiró

Based on arXiv: 1304.1758 & 1403.3539

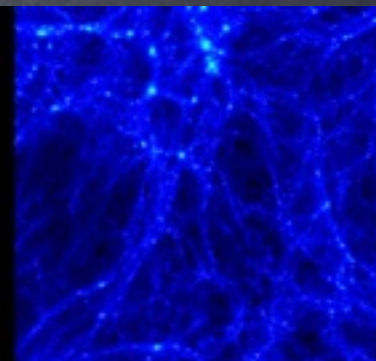
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MultiDark

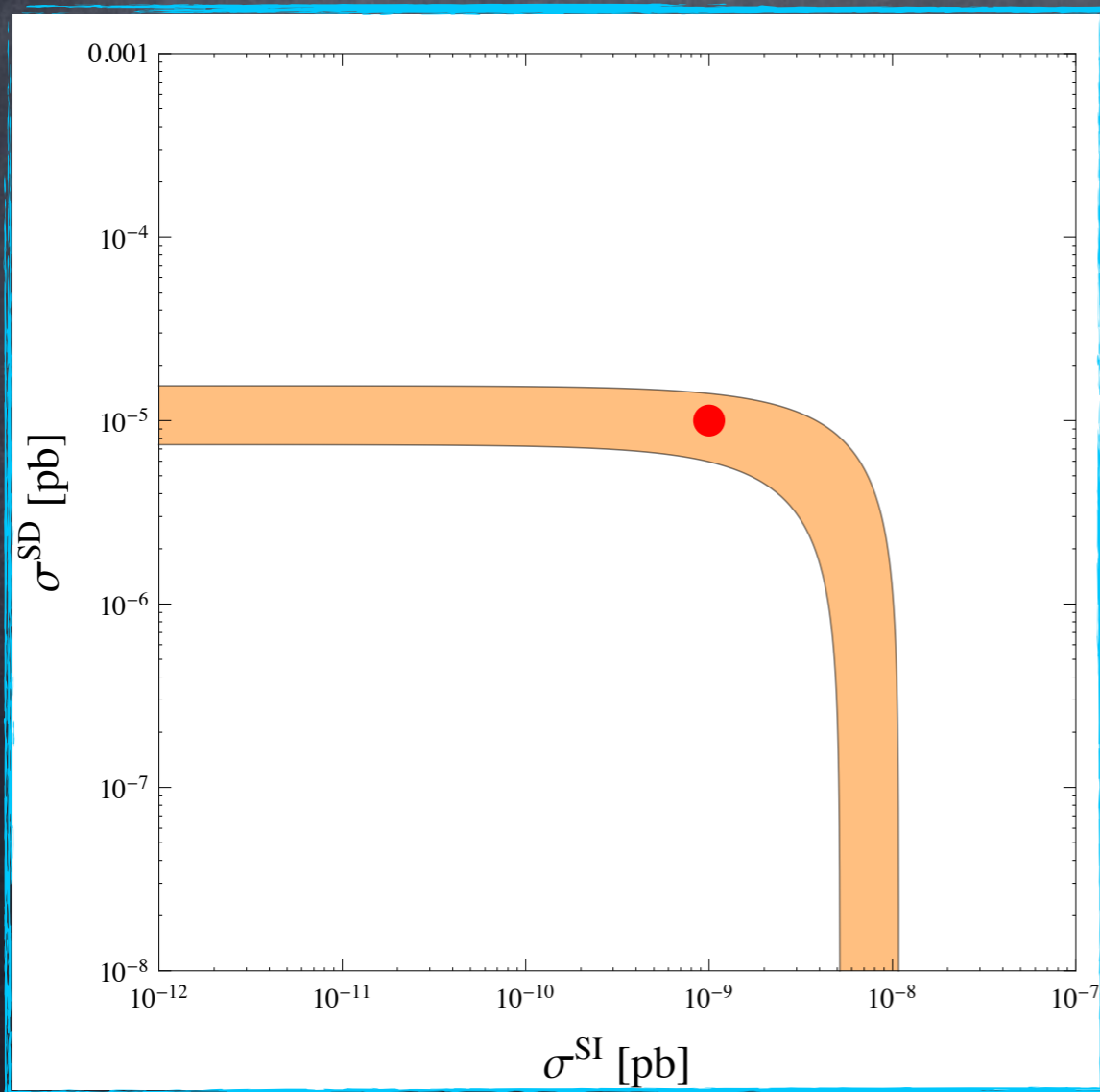
Multimessenger Approach
for Dark Matter Detection



Combining targets

Let's assume that an experiment 1 measures a signal, generated by the DM mass and the SI-SD cross sections, the number of events would be given by: $N_1 = A_1 \sigma^{\text{SI}} + B_1 \sigma^{\text{SD}}$ where the parameters A and B depends on the nucleus and DM mass.

If we assume that DM mass is well determined (for simplicity):

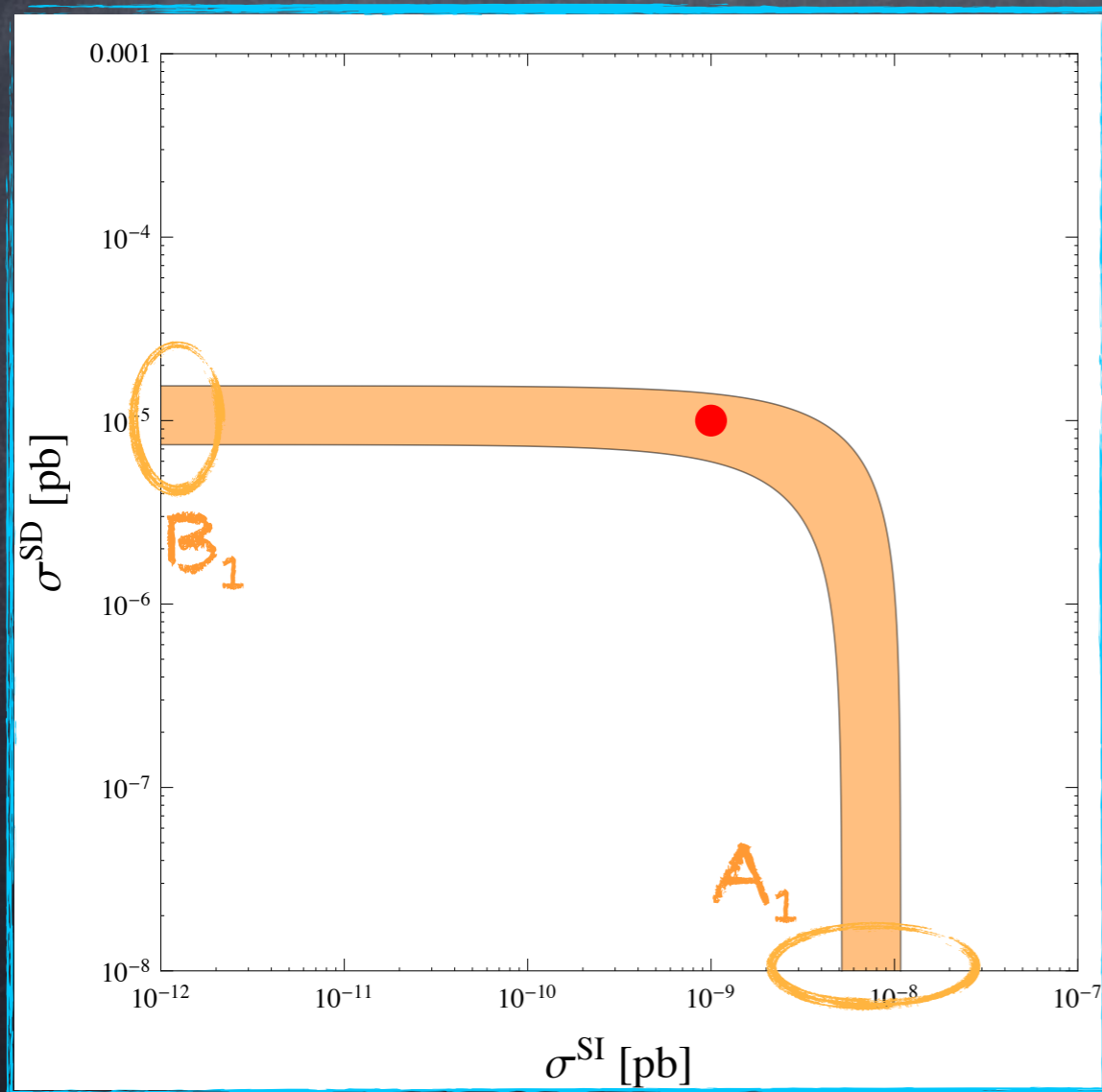


The red dot represents the values of the cross sections used to "simulate" the signal

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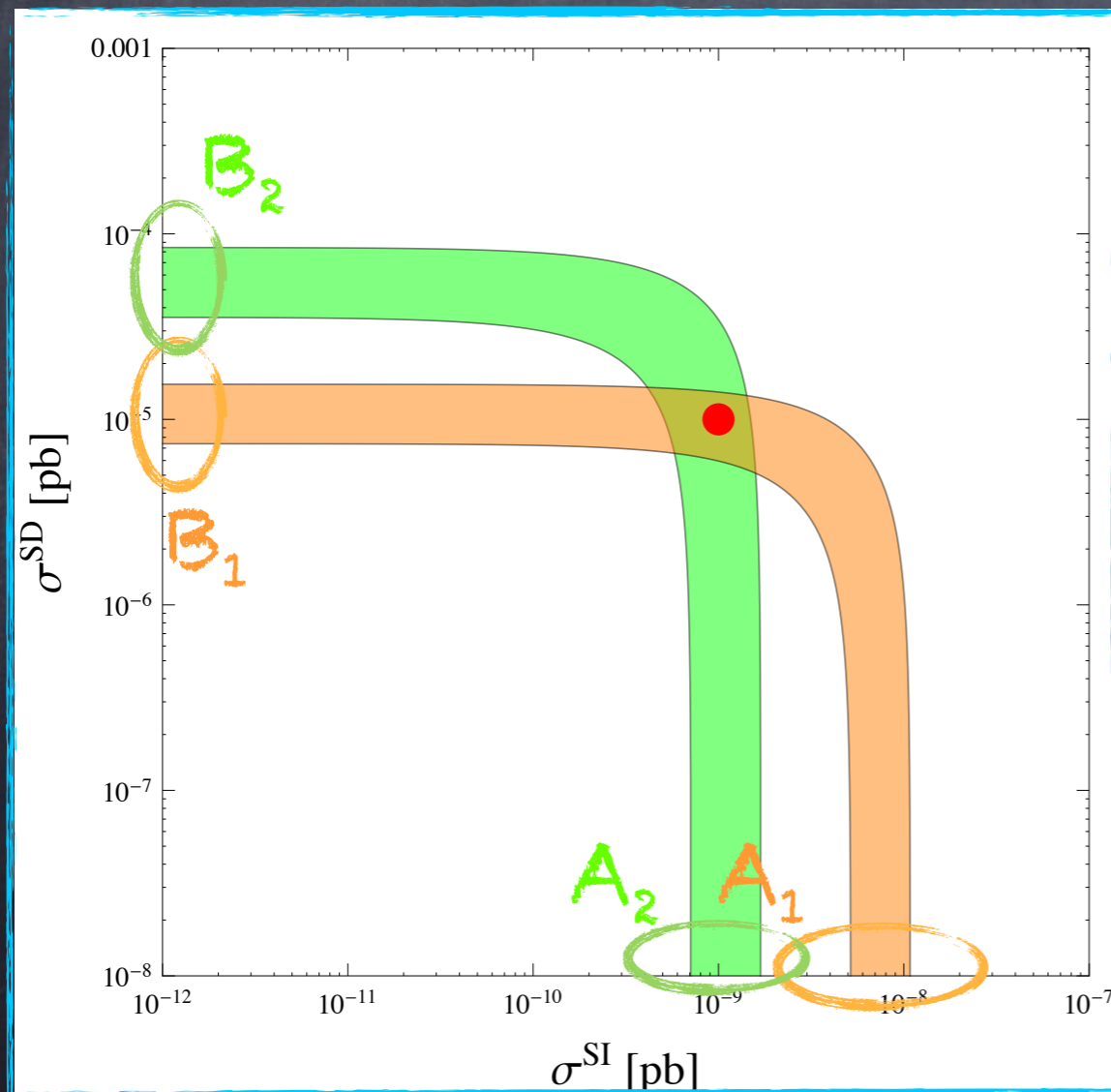
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There is a degeneracy between the cross sections (the signal can be explained by both alone or a combination)

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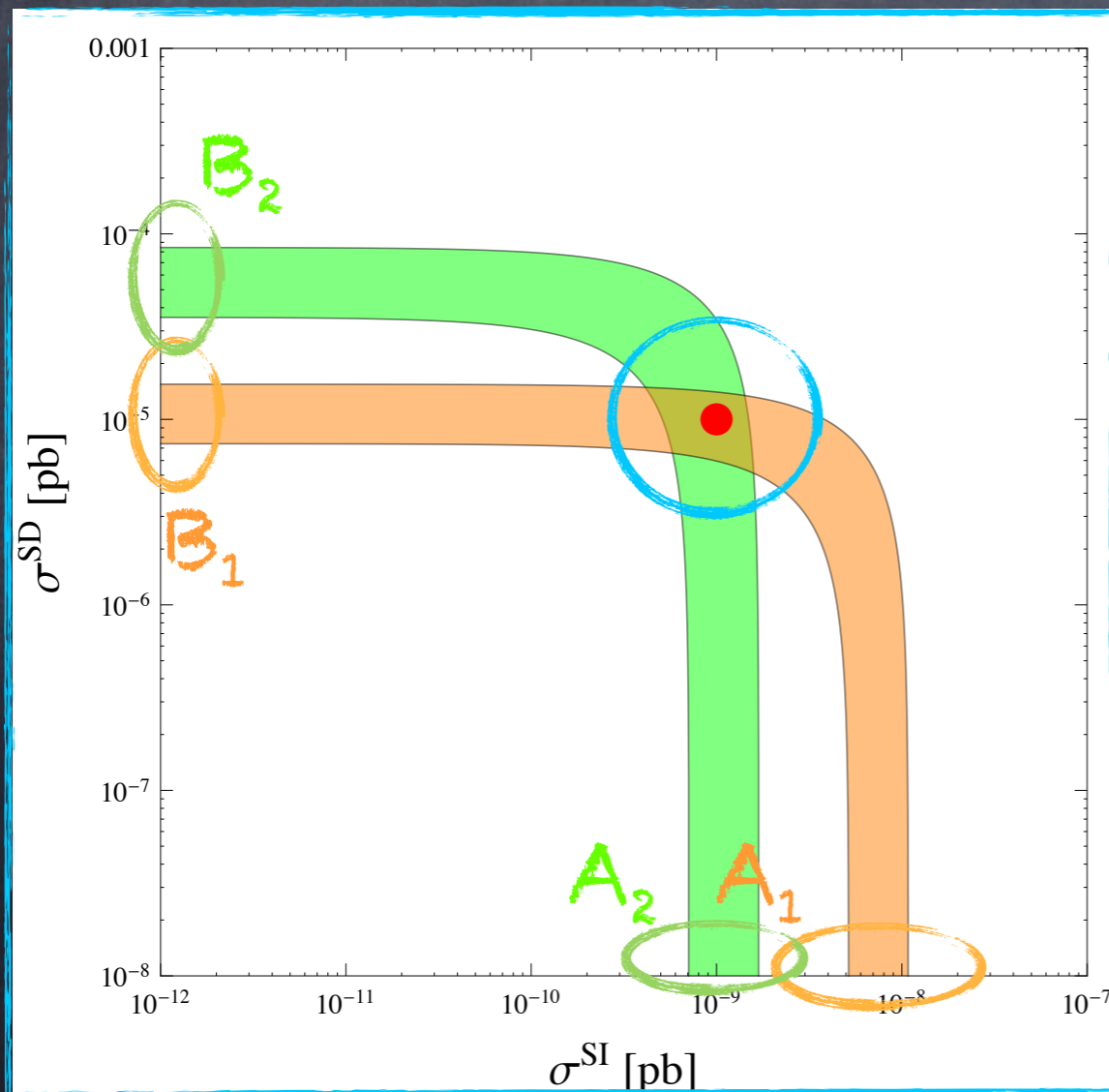
Including a second experiment

$$N_2 = A_2 \sigma^{SI} + B_2 \sigma^{SD}$$

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There is a degeneracy between the cross sections (the signal can be explained by both alone or a combination)

Including a second experiment

$$N_2 = A_2 \sigma^{SI} + B_2 \sigma^{SD}$$

If the parameters A and B from each of the experiments are different, then we end up with a

FINITE CONTOUR
Both experiments are

COMPLEMENTARY

The real situation

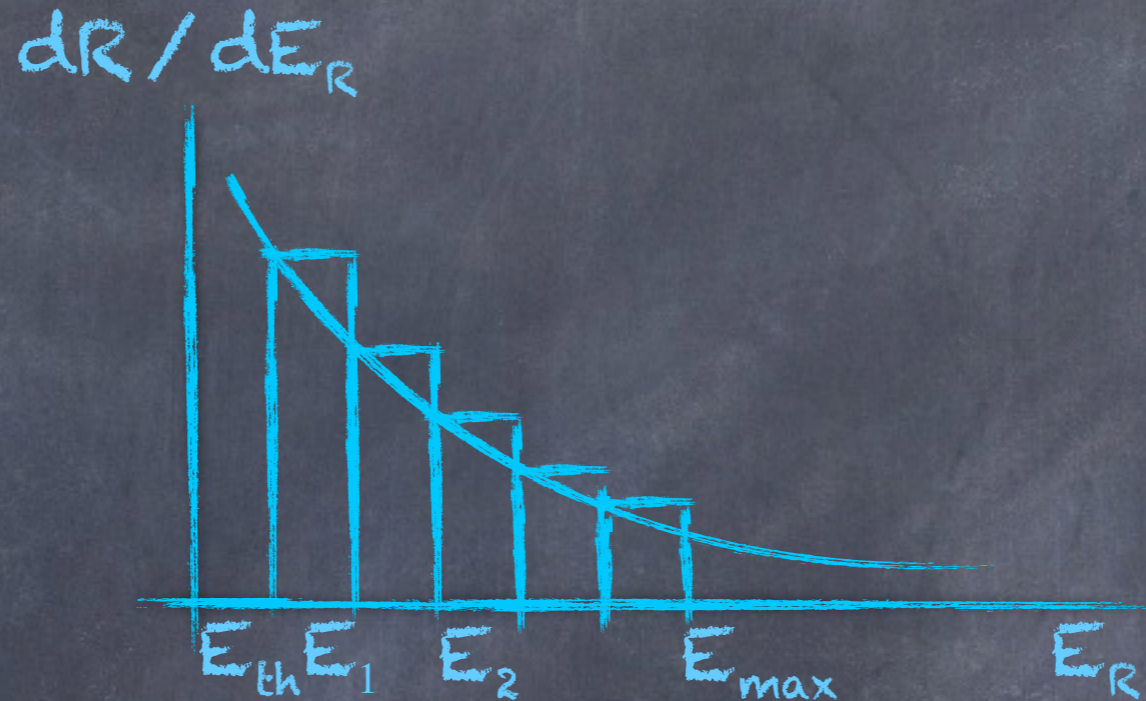
Ingredients of the analysis:

- We generate a "signal" by choosing a BM point defined by $m_{DM}, \sigma^{SI}, \sigma^{SD}$

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The total number of events in a series of bins must be calculated $N = R t$ ($t = \text{live time}$)

$$N_{12} = \int_{E_1}^{E_2} dE_R \frac{\epsilon \rho_0}{m_N m_{DM}} \int_{v_{\min}}^{v_{\text{esc}}} v f(v) \frac{d\sigma}{dE_R}(v, E_R) dv$$

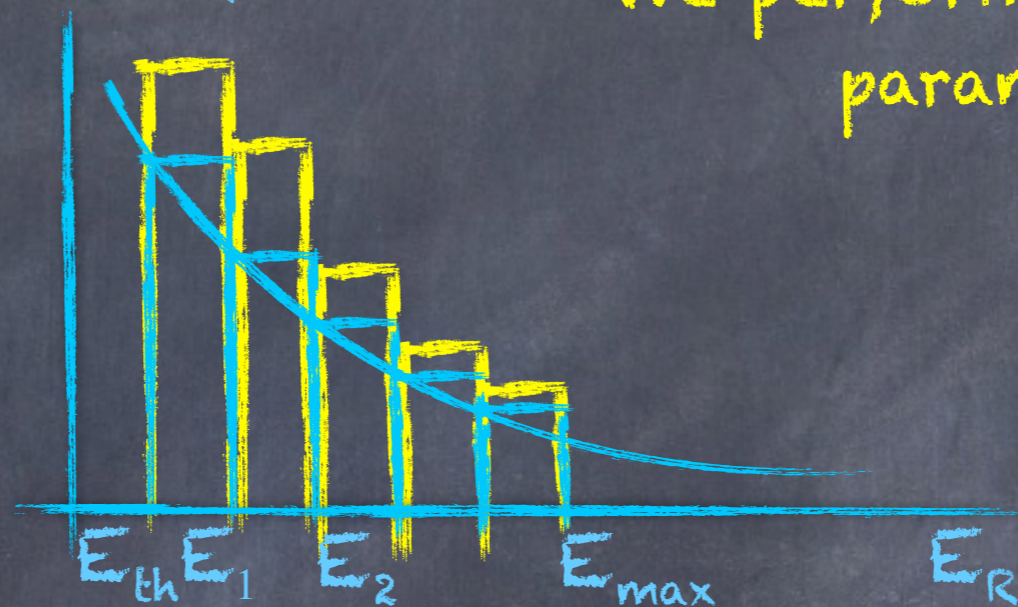
And extend from threshold energy up to the maximum (usually defined as energy window)

The real situation

Ingredients of the analysis:

- We generate a "signal" by choosing a BM point defined by $m_{DM}, \sigma^{SI}, \sigma^{SD}$

dR/dE_R



We perform a scan over the parameter space

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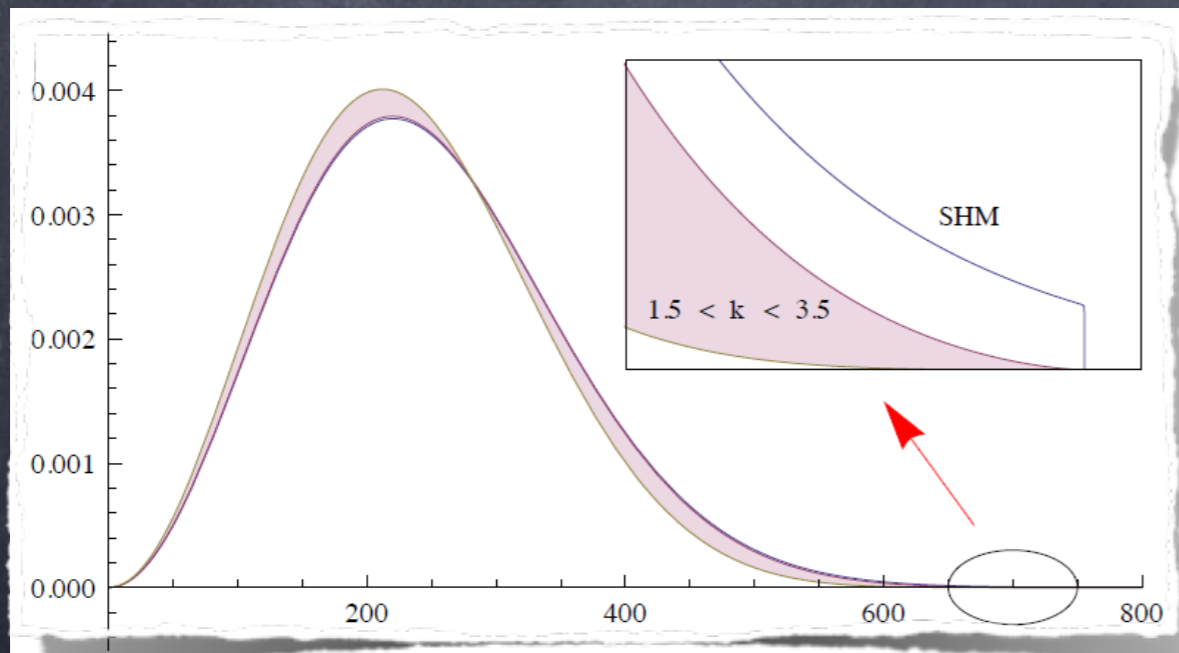
And extend from threshold energy up to the maximum (usually defined as energy window)

We compare both sets of data using a Poissonian likelihood function

The real situation

Ingredients of the analysis:

- We generate a "signal" by choosing a BM point defined by $m_{DM}, \sigma^{SI}, \sigma^{SD}$
- We perform a scan over the parameter space
- We include astrophysical uncertainties



Nuisance parameter	Range	Prior distribution
$\rho_{WIMP, \odot}$	[0.2, 0.6] GeV cm^{-3}	normal
v_{esc}	[478, 610] km s^{-1}	normal
v_{\odot}	[170, 290] km s^{-1}	normal
k	[0.5, 3.5]	flat

Lisante et al '10

$$F_k(v) = N_k^{-1} v^2 \left[e^{-v^2/k\Delta v^2} - e^{-v_{esc}^2/k\Delta v^2} \right]^k \theta(v_{esc} - v) \quad N_k = \Delta v^3 e^{-y_e^2} \int_0^{y_e} y^2 \left(e^{-(y^2 - y_e^2)/k} - 1 \right)^k dy$$

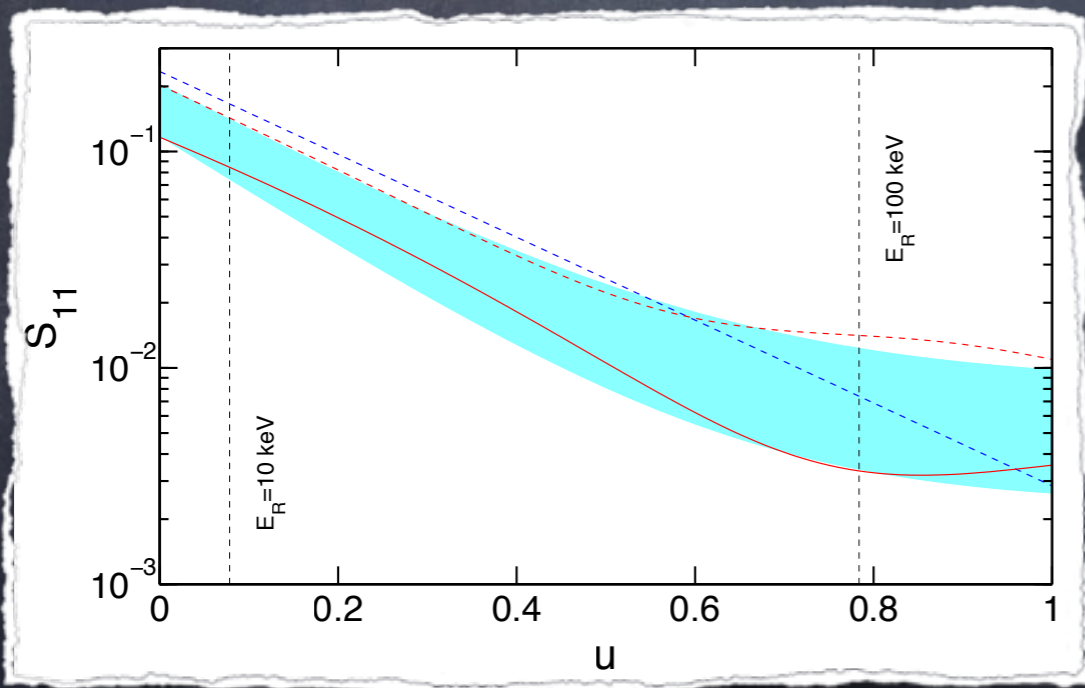
Binney & Tremaine

$$y_e = v_{esc} / \Delta v$$

The real situation

Ingredients of the analysis:

- We generate a "signal" by choosing a BM point defined by $m_{DM}, \sigma^{SI}, \sigma^{SD}$
- We perform a scan over the parameter space
- We include astrophysical uncertainties
- We include nuclear uncertainties



Isotope	N	α	β
^{73}Ge	0.0749 - 0.2071	5.0 - 6.0	0.0304 - 0.0442
^{129}Xe	0.0225 - 0.0524	4.0625 - 4.3159	0.001 - 0.0093
^{131}Xe	0.0169 - 0.0274	3.9913 - 4.7075	0.05 - 0.105
^{127}I	0.0297 - 0.0568	4.0050 - 4.4674	0.05 - 0.057
^{23}Na	0.0098 - 0.0277	2.0 - 3.5287	0 - 0.1250
^{19}F	0.0505 - 0.1103	2.9679 - 3.0302	0 - 0.0094

$$\left(\frac{d\sigma}{dE_R} \right)_{SD} = \frac{16m_N G_F^2 (J+1)}{\pi v^2 J} \left(a_p \langle S_p \rangle + a_n \langle S_n \rangle \right)^2 F_{SD}^2(E_R) \quad S_{ij}(u) = N \left[(1-\beta) e^{-\alpha u} + \beta \right] \quad E_R = \frac{ub^{-2}}{m_N}$$

$$S(q) = a_0^2 S_{00}(q) + a_0 a_1 S_{01}(q) + a_1^2 S_{11}(q) \quad F_{SD}^2(q) = S(q) / S(0)$$

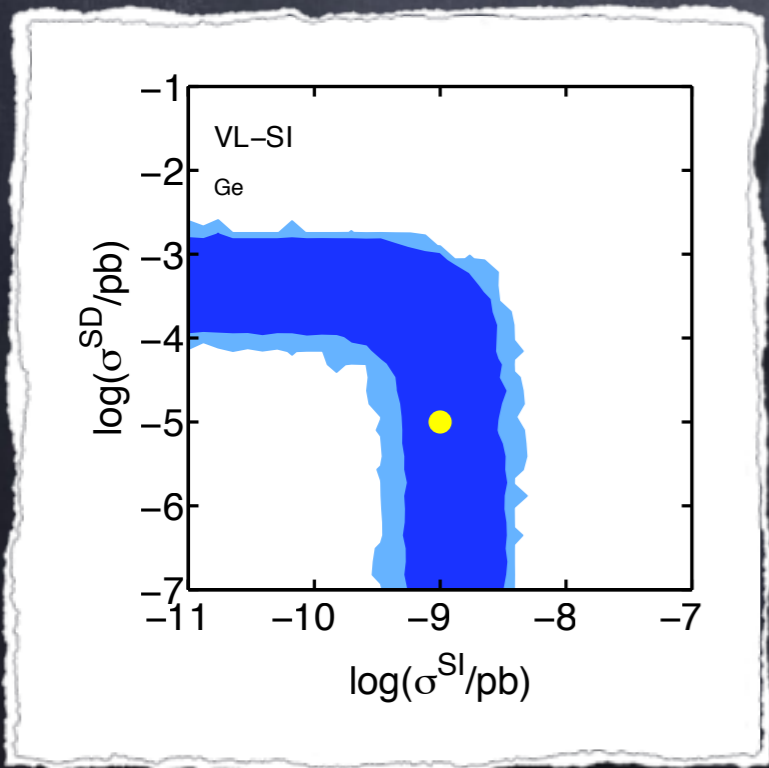
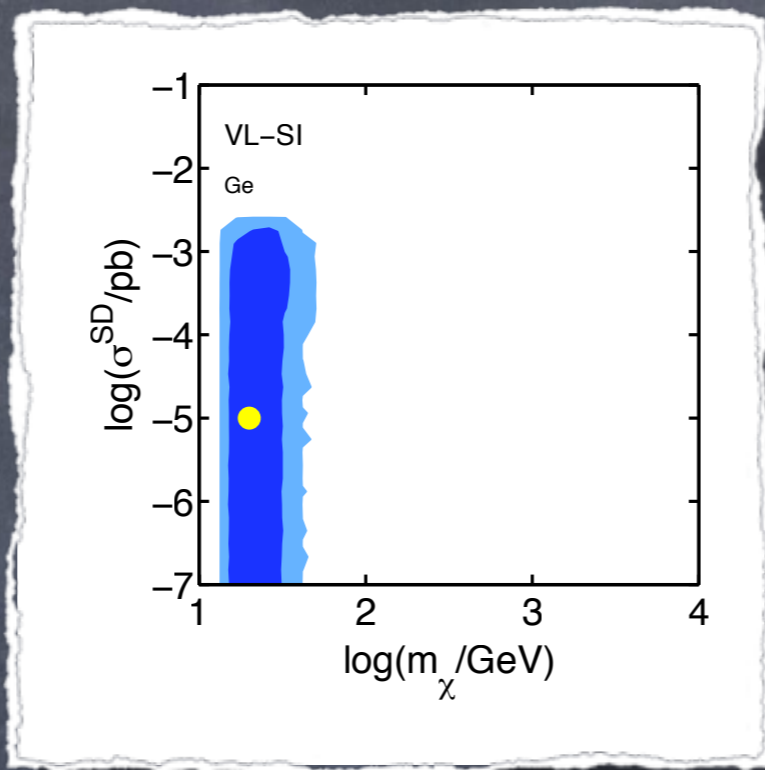
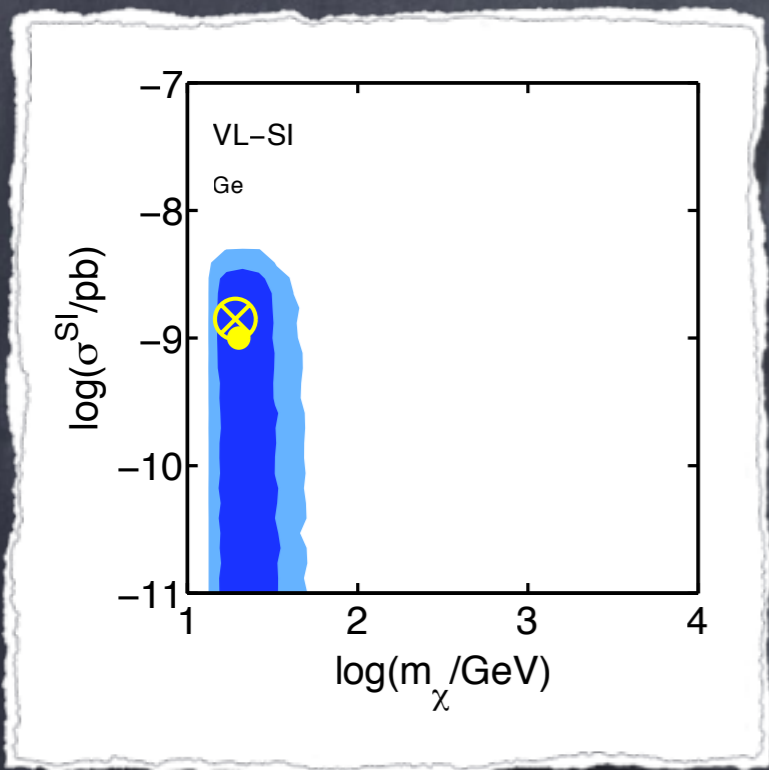
Cerdeño et al '12

Playing with Ge and Xe

Experimental setups for Ge and Xe

- Background free experiments (in previous works we have shown that the expected levels of background in these detectors do not affect the complementary conclusions)
- We use natural abundances for each of the nuclei
- Exposure of 300 kg yr (1 ton yr with 30% of efficiency)
- 3 keV threshold for each
- Maximum energy: 100 keV for Ge and 43 keV for Xe
- Gaussian energy resolution

Current experiments (Ge, Xe)



● BM point

$$m_\chi = 20 \text{ GeV}$$

$$\sigma^{SI} = 10^{-9} \text{ pb}$$

$$\sigma^{SD} = 10^{-5} \text{ pb}$$

⊗ Best fit point

68% and 99%CL profile likelihood contours

Ge

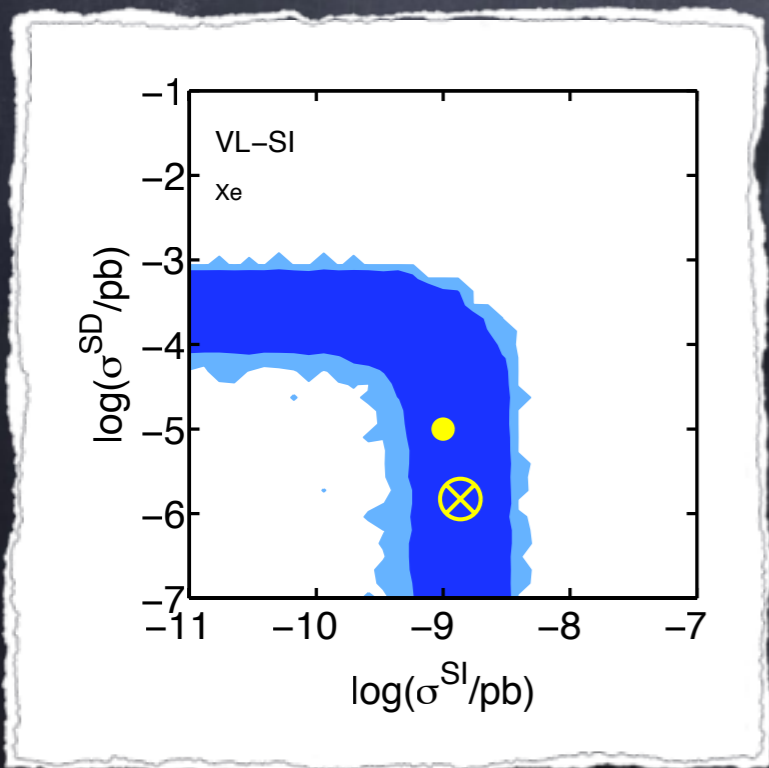
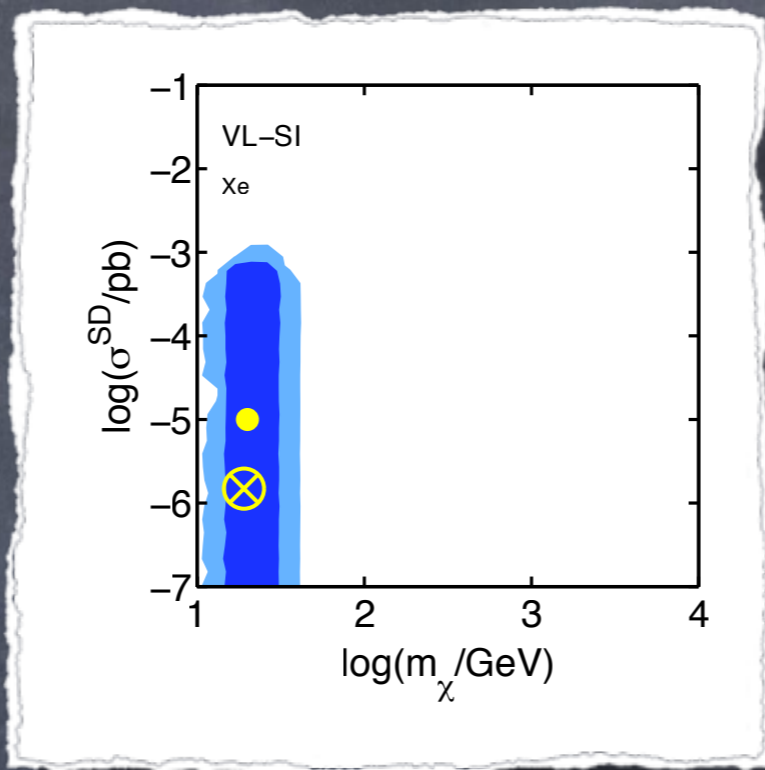
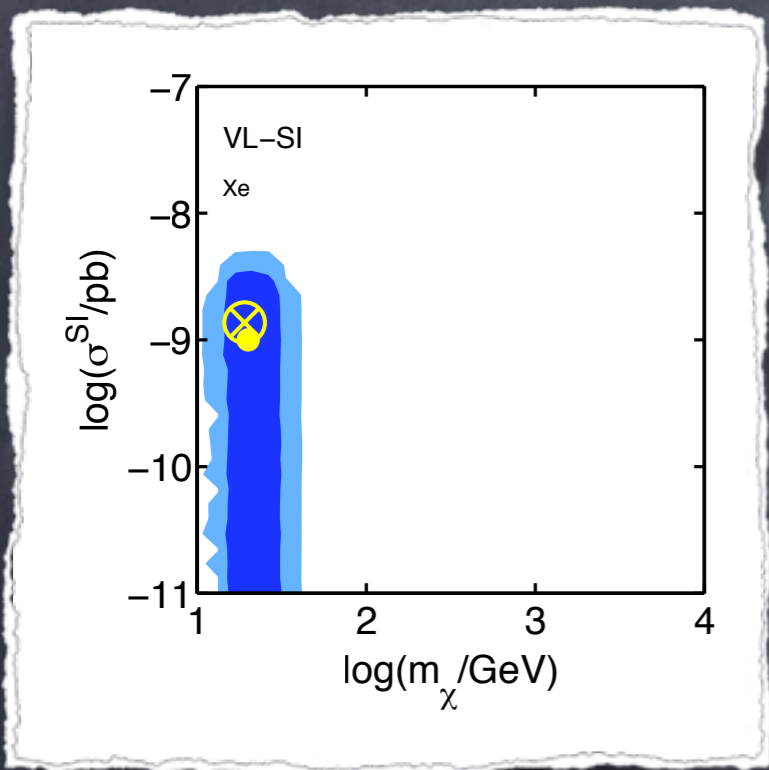
NOT ABLE TO RECONSTRUCT

THE CROSS SECTIONS

THE MASS OF
THE WIMP IS
WELL

RECONSTRUCTED

Current experiments (Ge, Xe)



● BM point

$$m_\chi = 20 \text{ GeV}$$

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⊗ Best fit point

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Xe

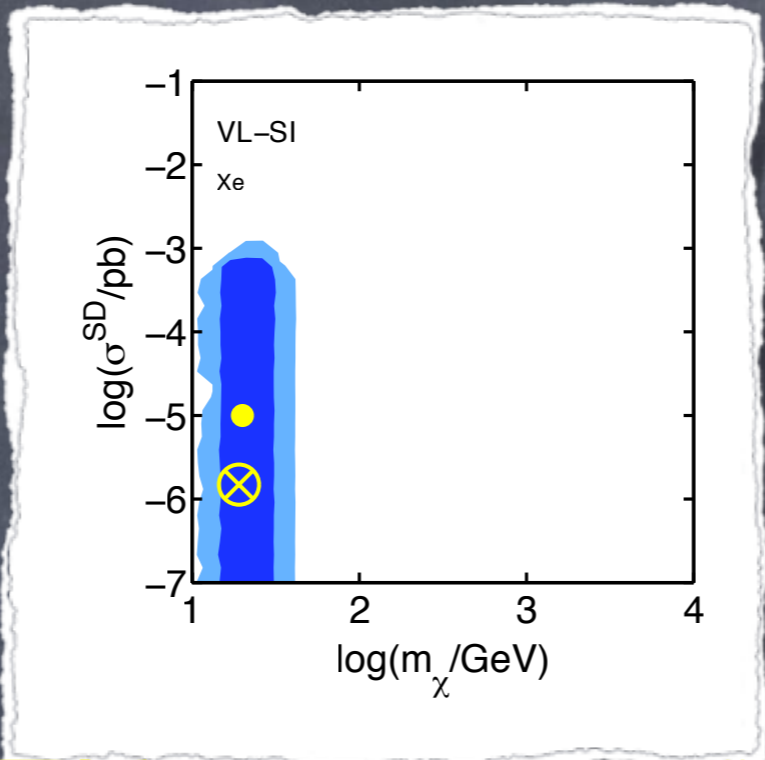
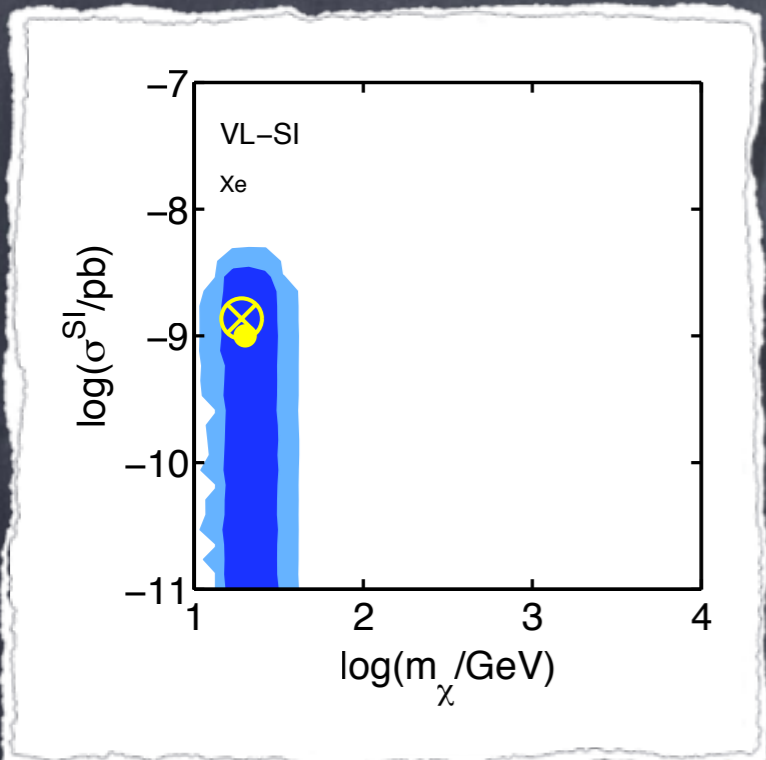
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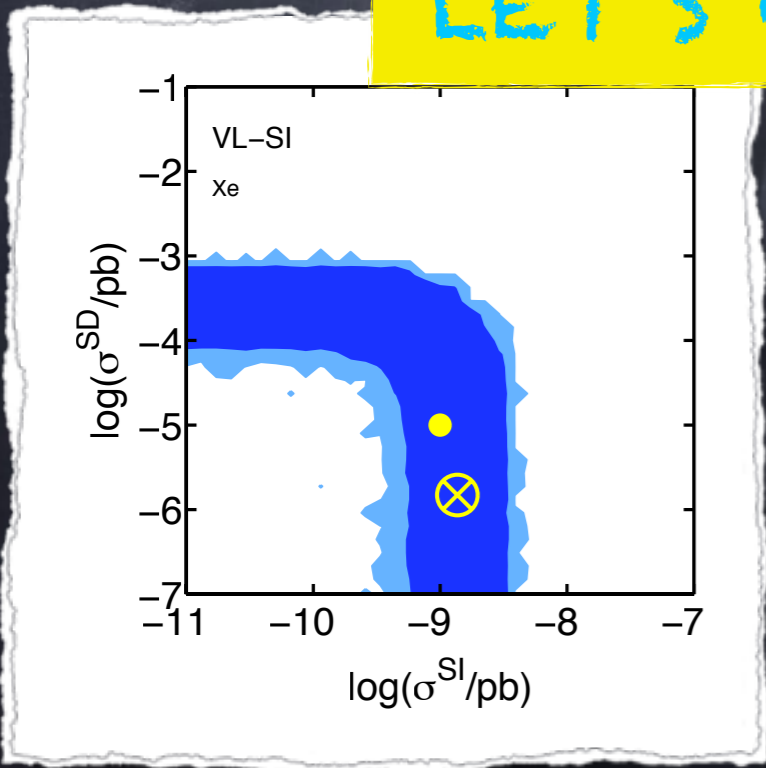
Current experiments (Ge, Xe)



Xe

NOT ABLE TO RECONSTRUCT

LET'S COMBINE BOTH TARGETS



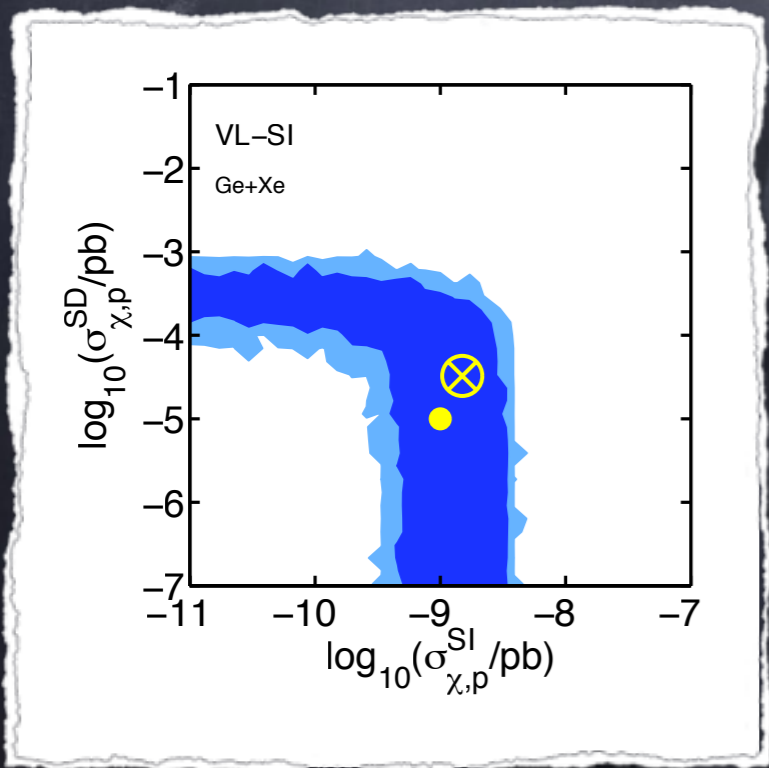
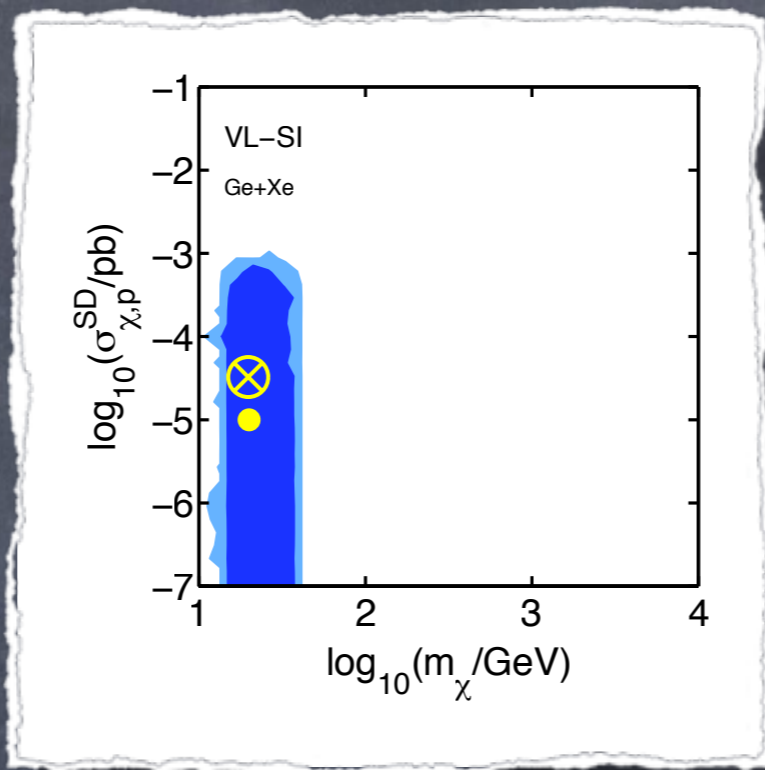
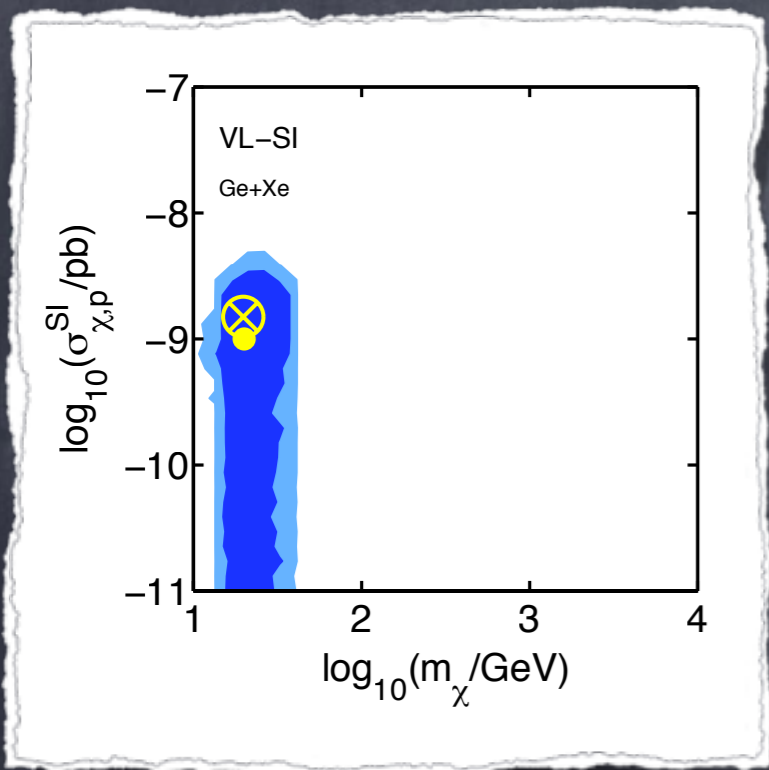
$m_\chi = 20 \text{ GeV}$
 $\sigma^{SI} = 10^{-9} \text{ pb}$
 $\sigma^{SD} = 10^{-5} \text{ pb}$

⊗ Best fit point

68% and 99%CL profile likelihood contours

THESE SECTIONS
 THE MASS OF
 THE WIMP IS
 WELL
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Current experiments (Ge,Xe)



● BM point

$$m_{\chi} = 20 \text{ GeV}$$

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Ge+Xe

NOT ABLE TO RECONSTRUCT THE CROSS SECTIONS EVEN

IN COMBINATION

THE MASS OF THE WIMP IS WELL RECONSTRUCTED

WELL RECONSTRUCTED

Current experiments (Ge, Xe)

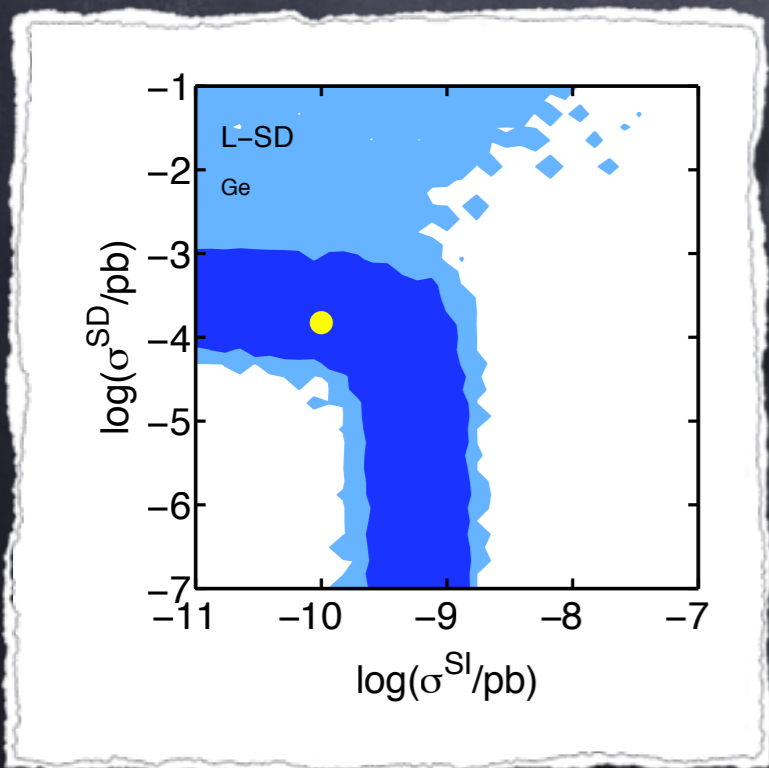
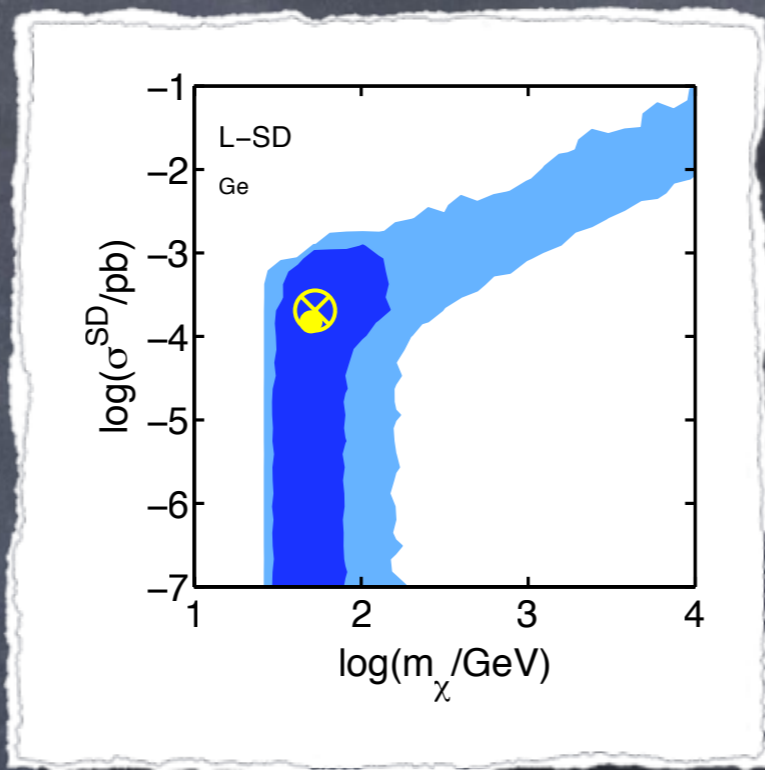
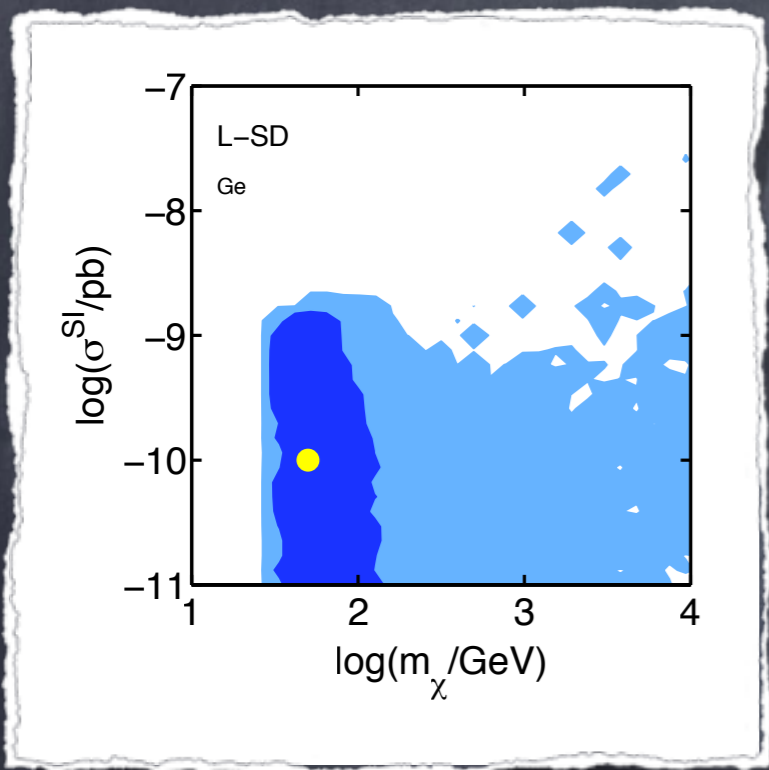
For light DM (20 GeV) and SI-dominated signals in Ge and Xe

the combination of Ge and Xe like experiments is

NOT COMPLEMENTARY

LET'S SEE A BENCHMARK POINT IN WHICH THE RATE IS SD-DOMINATED IN GE AND XE

Current experiments (Ge, Xe)



● BM point

$$m_\chi = 50 \text{ GeV}$$

$$\sigma^{SI} = 10^{-10} \text{ pb}$$

$$\sigma^{SD} = 1.5 \times 10^{-4} \text{ pb}$$

⊗ Best fit point

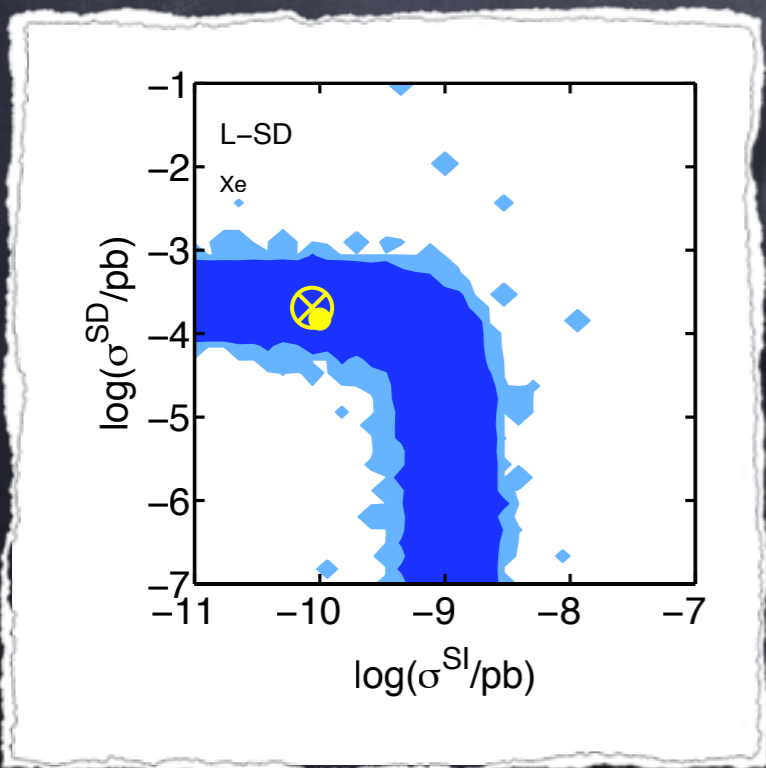
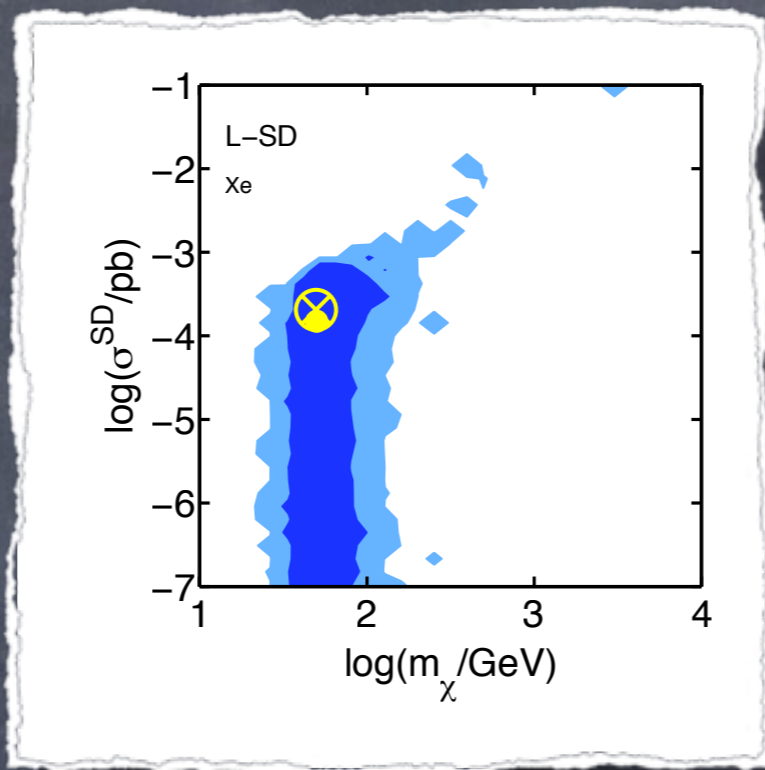
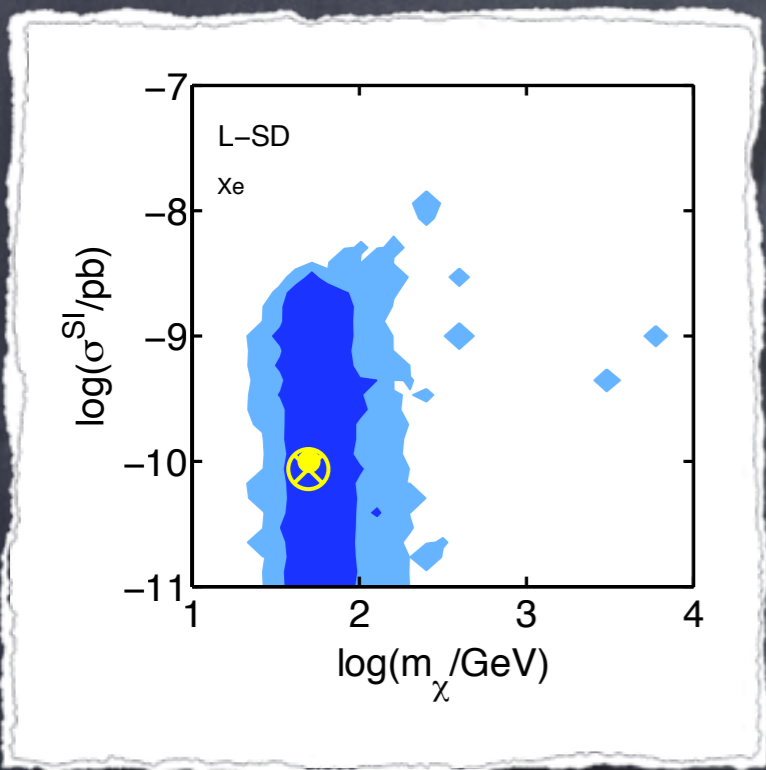
68% and 99%CL profile likelihood contours

Ge

NOT ABLE TO RECONSTRUCT THE CROSS SECTIONS

THE MASS OF THE WIMP IS WELL RECONSTRUCTED ONLY AT 68% CL

Current experiments (Ge, Xe)



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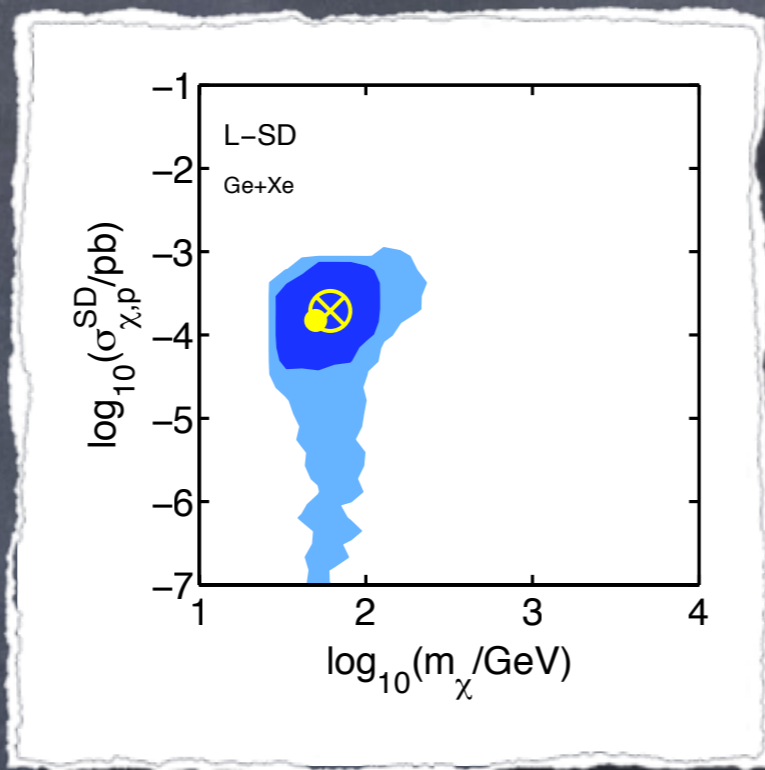
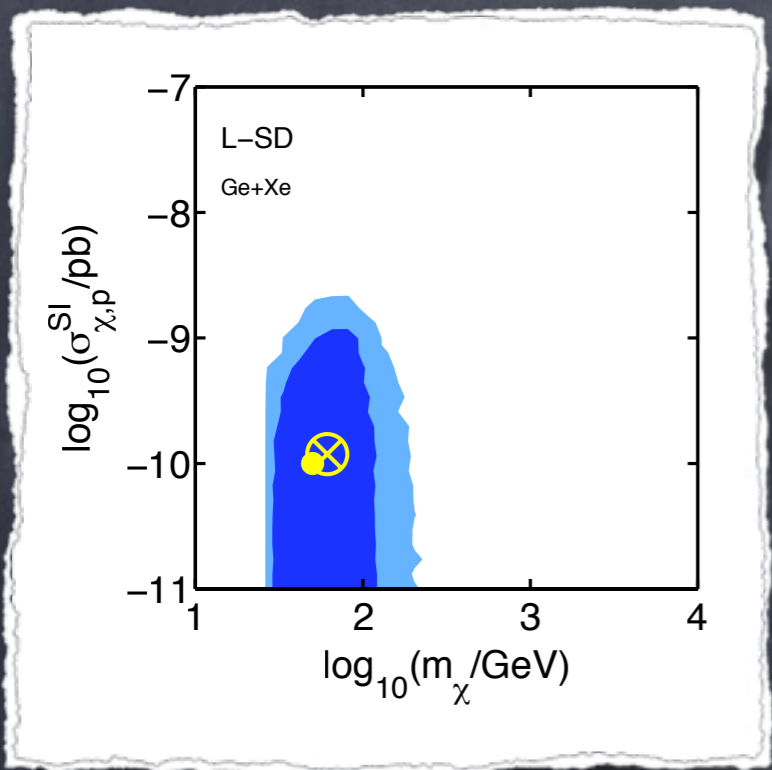
68% and 99%CL profile likelihood contours

Xe

NOT ABLE TO RECONSTRUCT THE CROSS SECTIONS

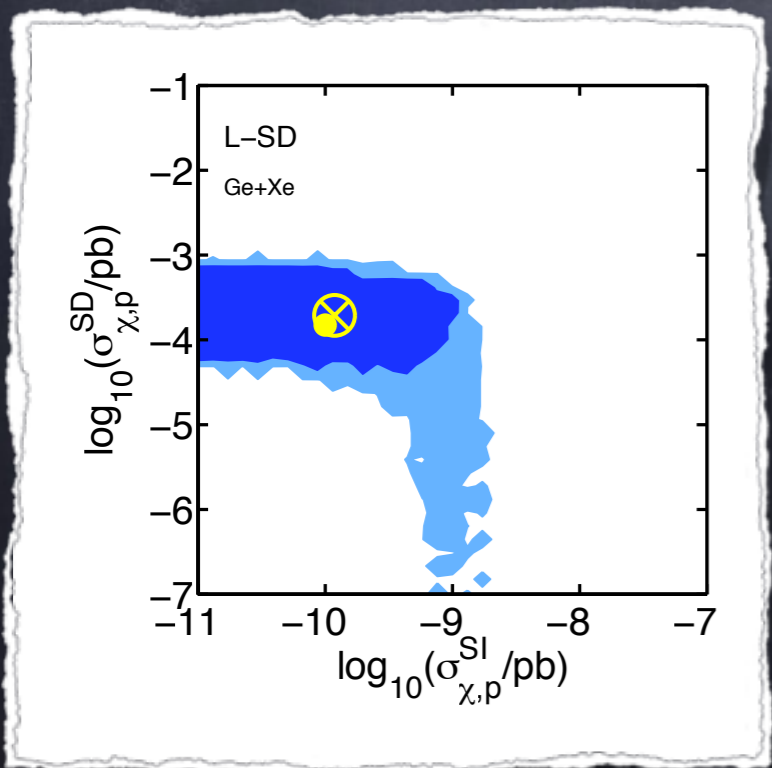
THE MASS OF THE WIMP IS BETTER RECONSTRUCTED

Current experiments (Ge,Xe)



Ge+Xe

NOT ABLE TO RECONSTRUCT
THE SI CROSS SECTION



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⊗ Best fit point

68% and 99%CL profile
likelihood contours

THE MASS OF
THE WIMP AND
THE SD CROSS
SECTION IS
RECONSTRUCTED

Including the bolometers

LiF - CaWO_4 - CaF_2 - NaI

Experimental setups for the bolometers

- Background free experiments (in previous works we have shown that the expected levels of background in these detectors do not affect the complementary conclusions)
- We use natural abundances for each of the nuclei
- Exposure of 300 kg yr (1 ton yr with 30% of efficiency)
- 10 keV threshold for each
- Maximum energy: 100 keV
- Energy resolution: 5% FWHM
- For NaI: thermal quenching of 0.85, 1.0 and 1.15

Including the bolometers

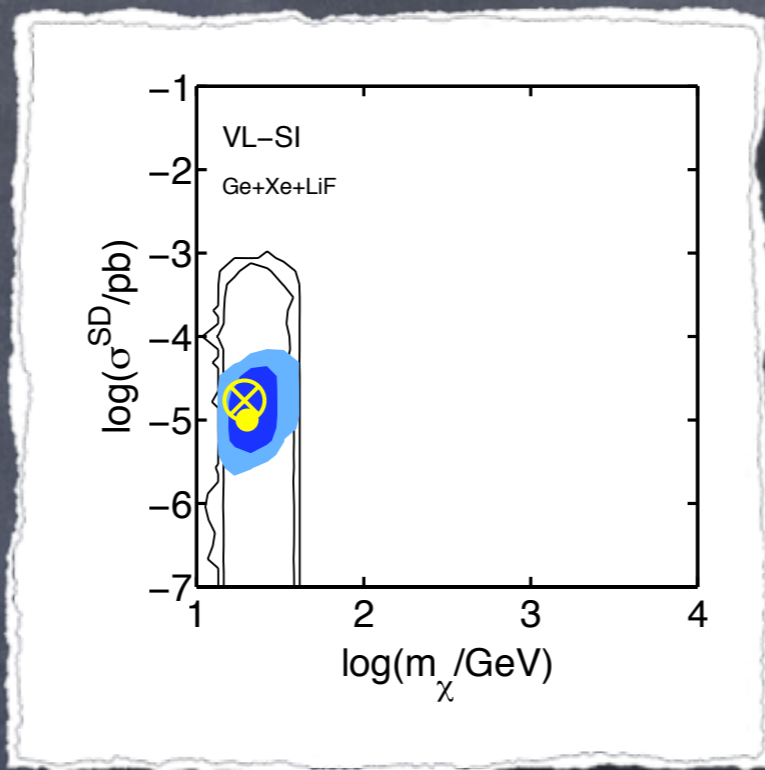
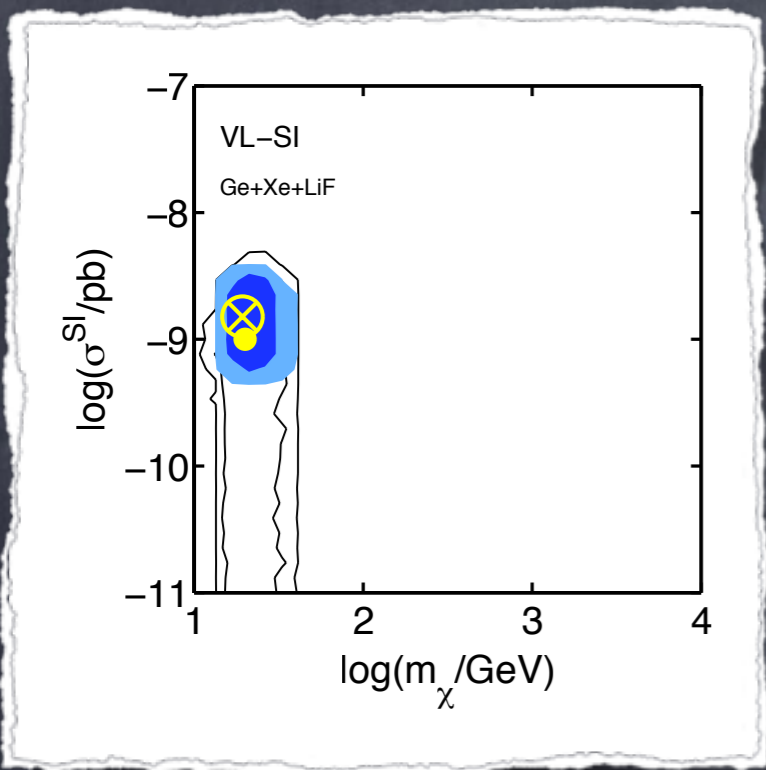
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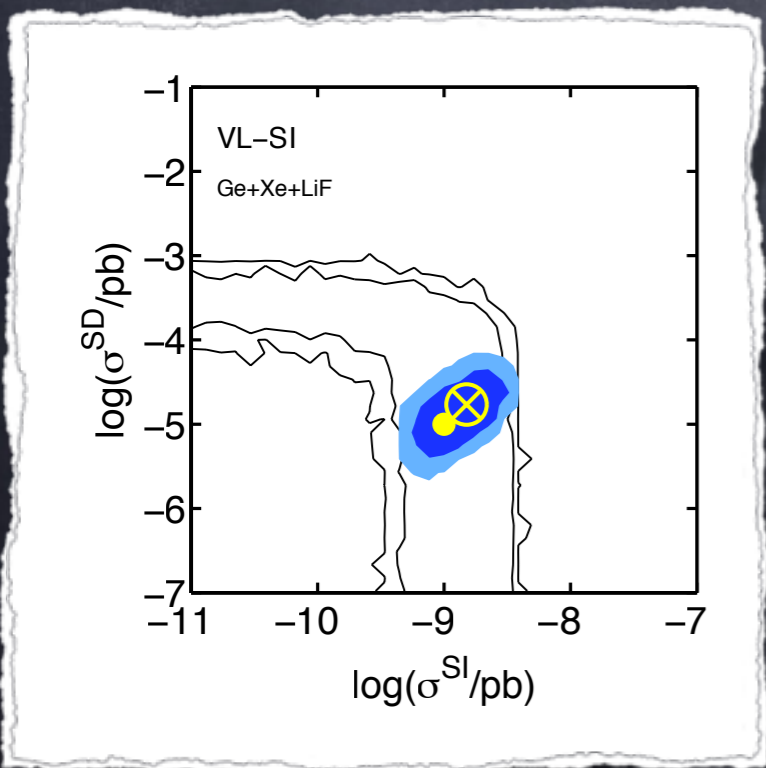
**WE START WITH CASE OF SI
DOMINATED RATE IN GE AND XE**

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+LiF



● BM point

$$m_\chi = 20 \text{ GeV}$$

$$\sigma^{SI} = 10^{-9} \text{ pb}$$

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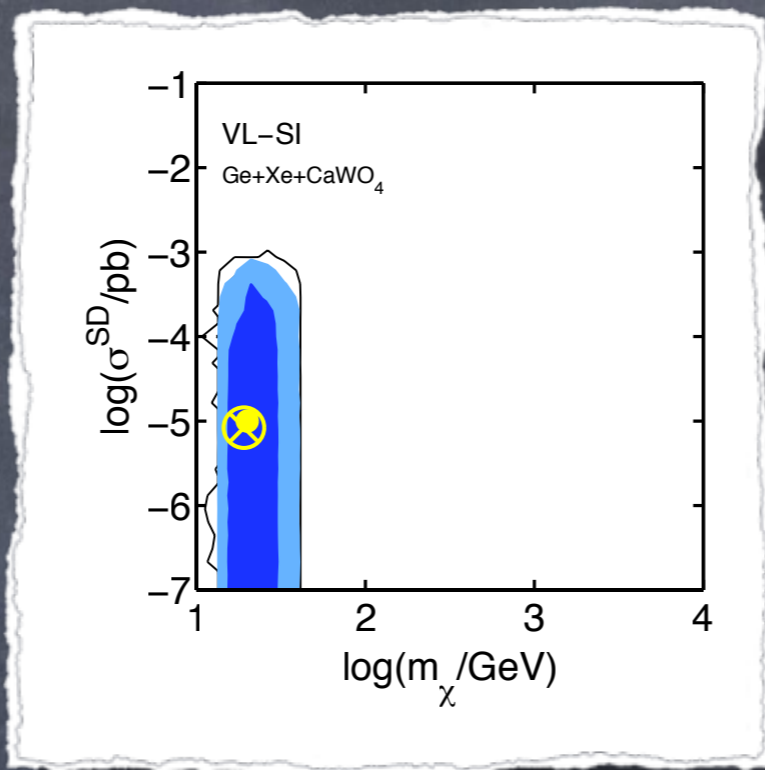
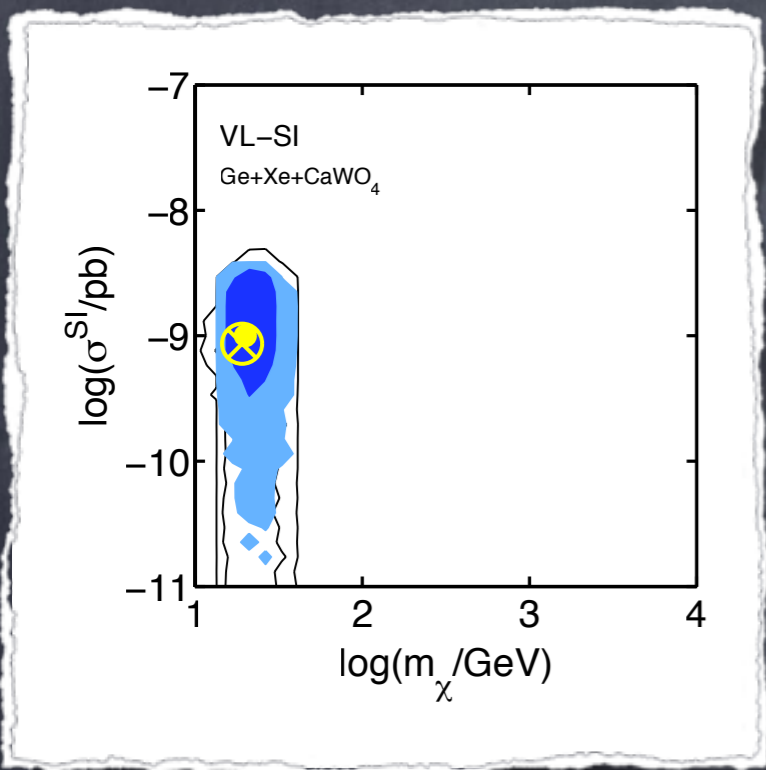
⊗ Best fit point

68% and 99%CL profile likelihood contours

COMPLEMENTARY!!!

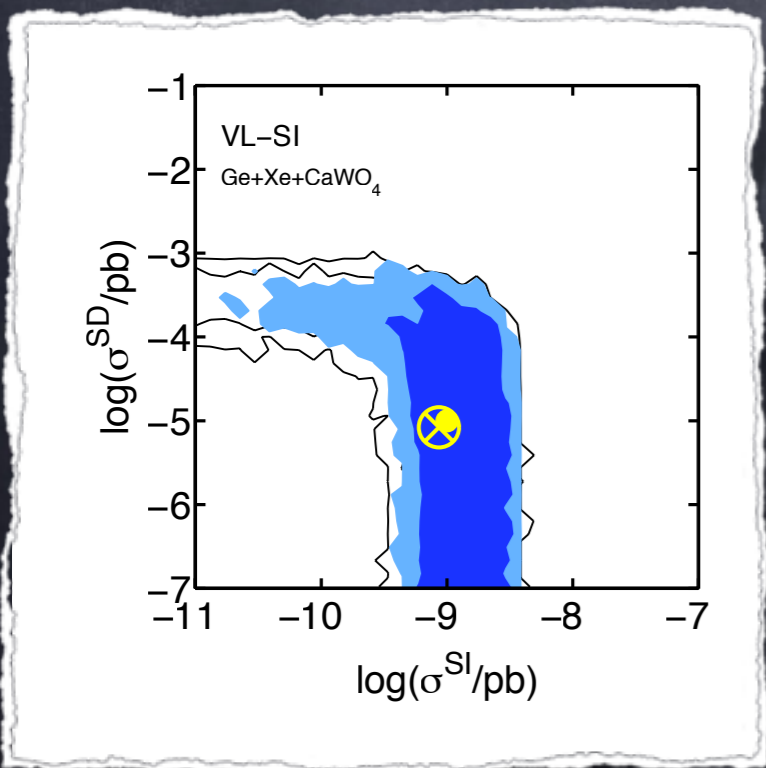
Very high SD sensitivity in both Li and F

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+CaWO4



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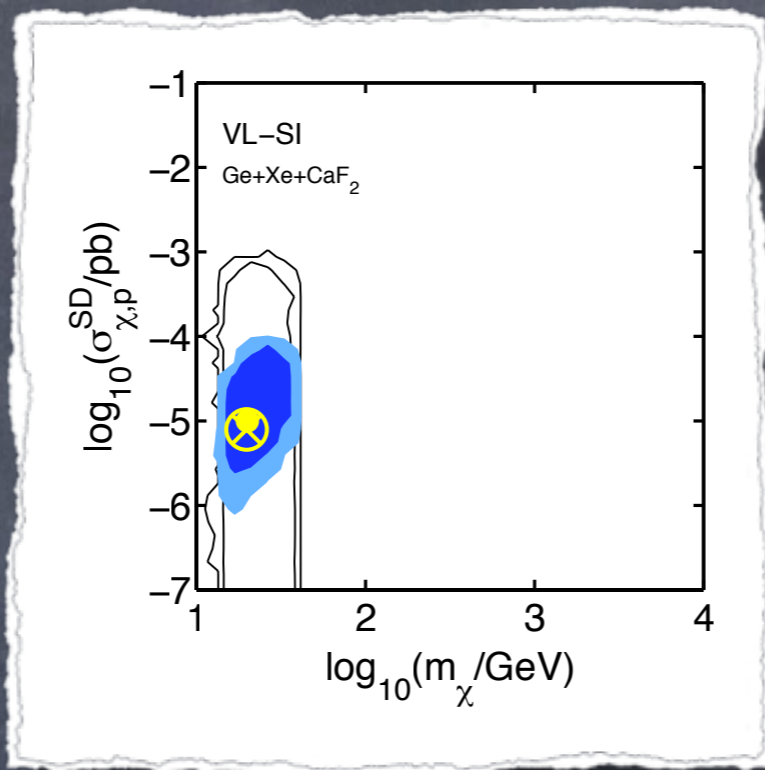
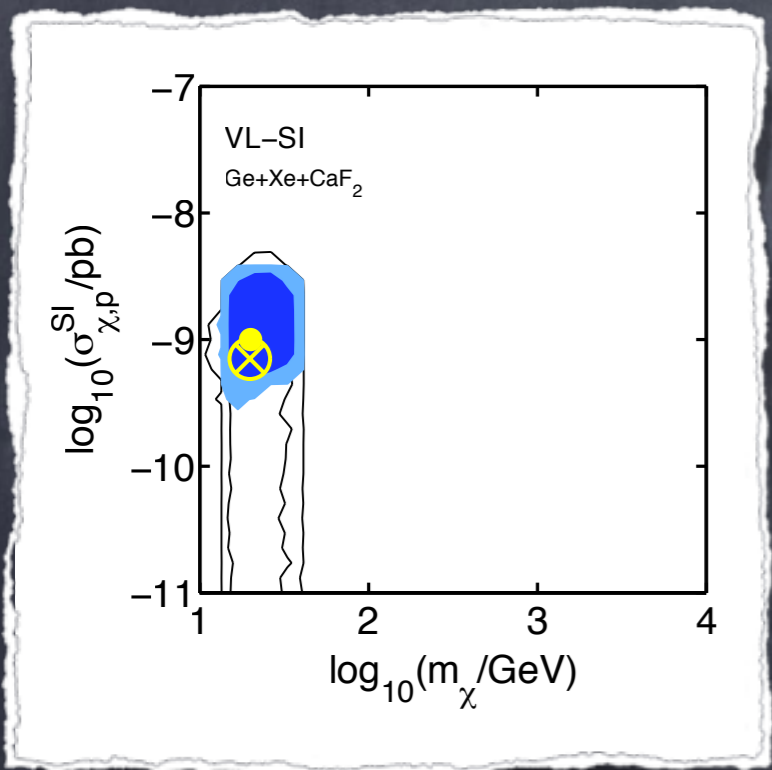
⊗ Best fit point

68% and 99%CL profile likelihood contours

SI and mass well reconstructed

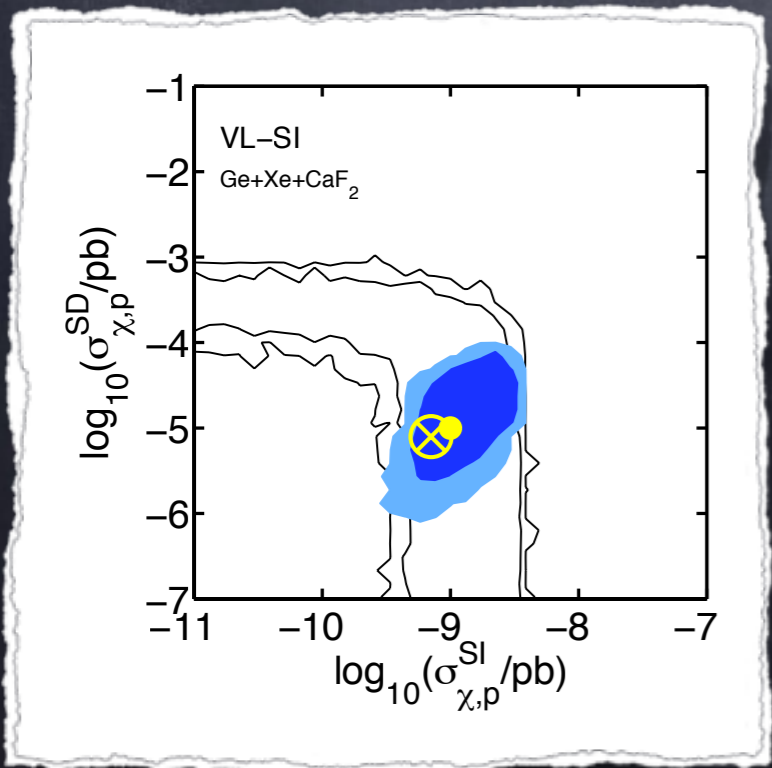
SD cross section sensitivity of CaWO4 similar to those of Ge and Xe

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+CaF₂



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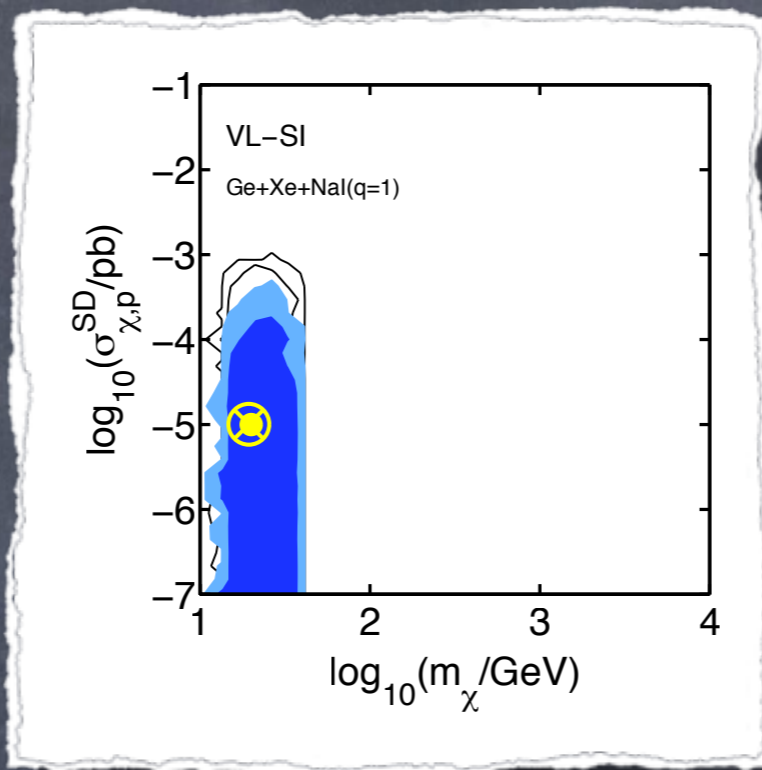
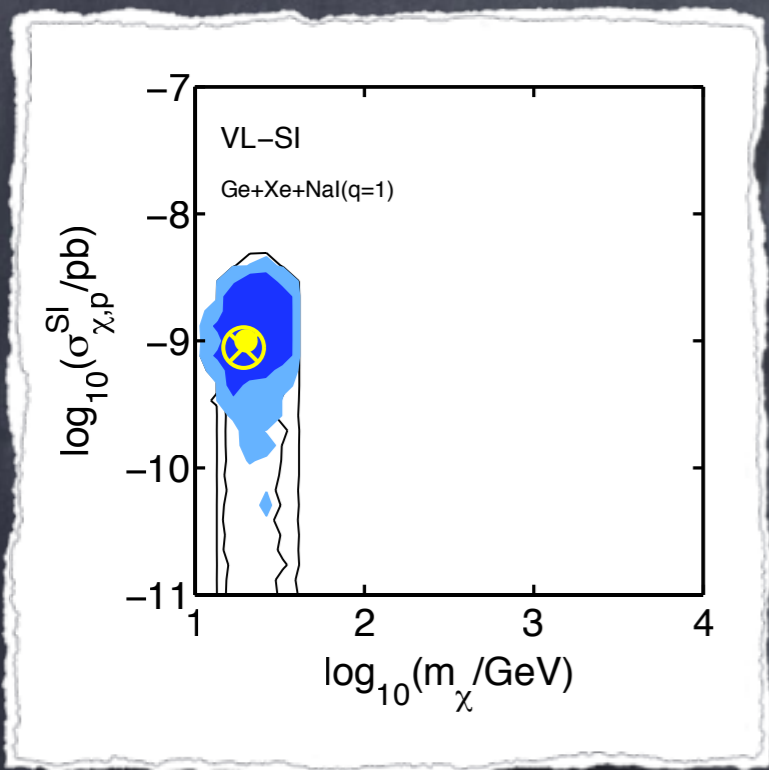
⊗ Best fit point

68% and 99%CL profile likelihood contours

COMPLEMENTARY!!!

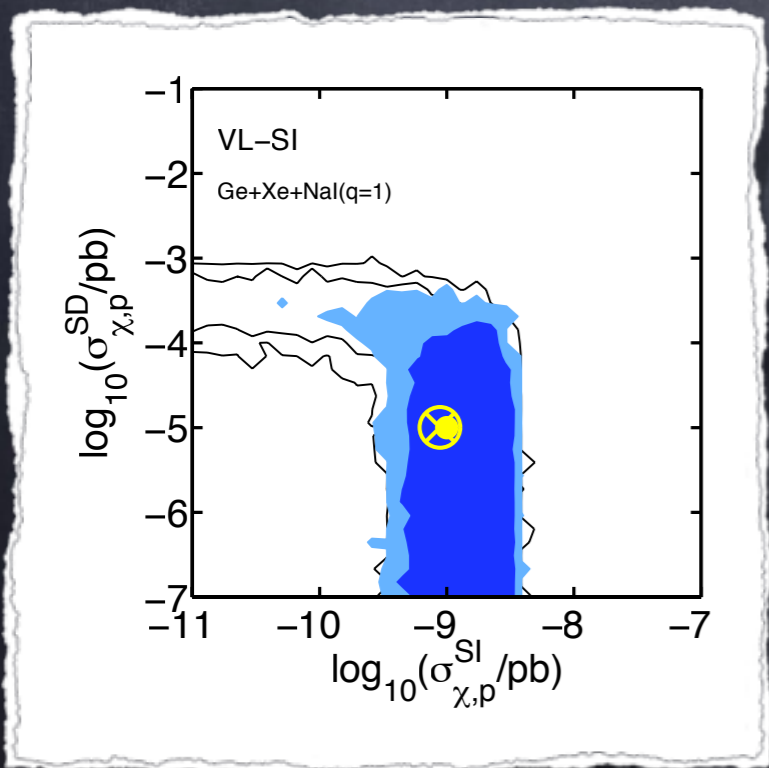
Very high SD sensitivity in F (similar to LiF)

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+NaI



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⊗ Best fit point

68% and 99%CL profile likelihood contours

SI and mass well reconstructed

SD cross section sensitivity of NaI similar to those of Ge and Xe

Current experiments (Ge, Xe)

Ge + Xe + LiF

Ge + Xe + CaF₂

COMPLEMENTARY

Ge + Xe + CaWO₄

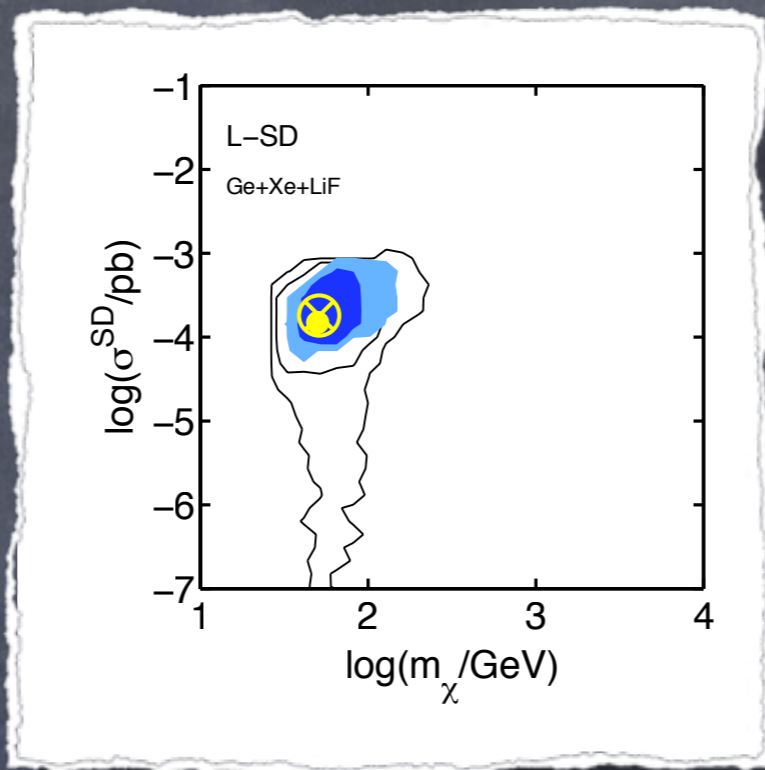
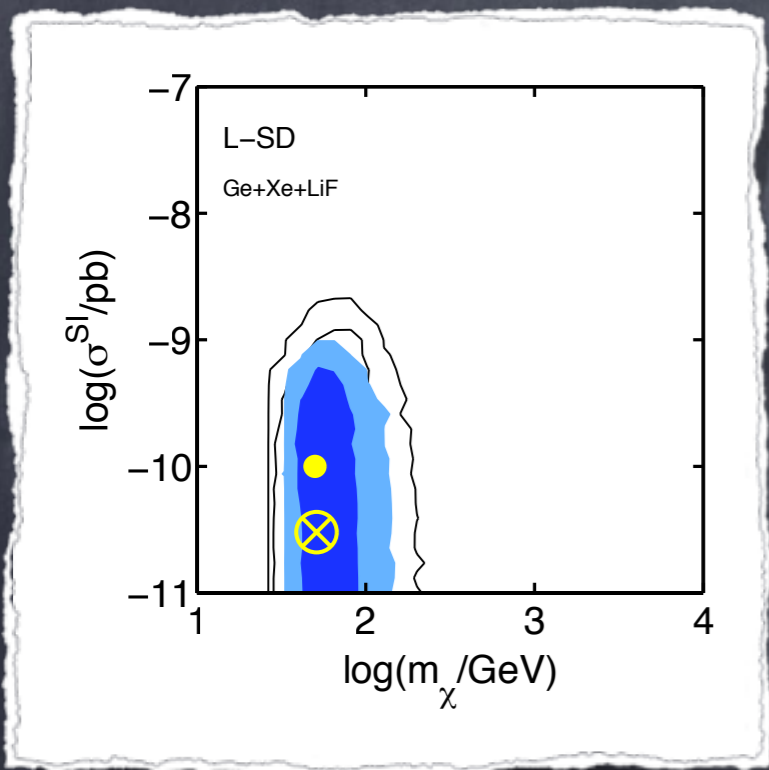
Ge + Xe + NaI

NOT COMPLEMENTARY

(but a considerable improvement)

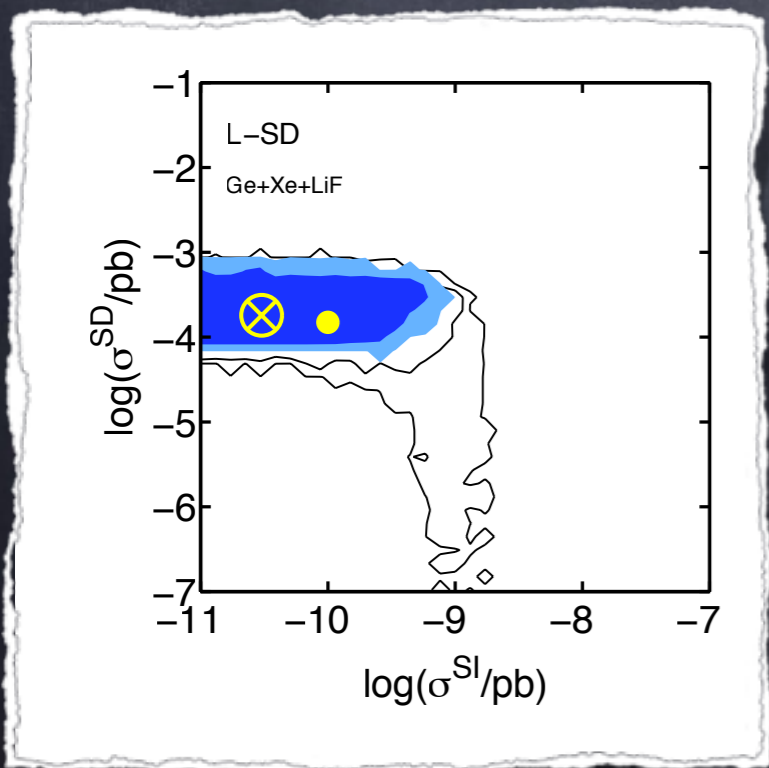
AGAIN LET'S SEE A BENCHMARK POINT IN WHICH
THE RATE IS SD-DOMINATED IN GE AND XE

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+LiF



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$$m_\chi = 50 \text{ GeV}$$

$$\sigma^{SI} = 10^{-10} \text{ pb}$$

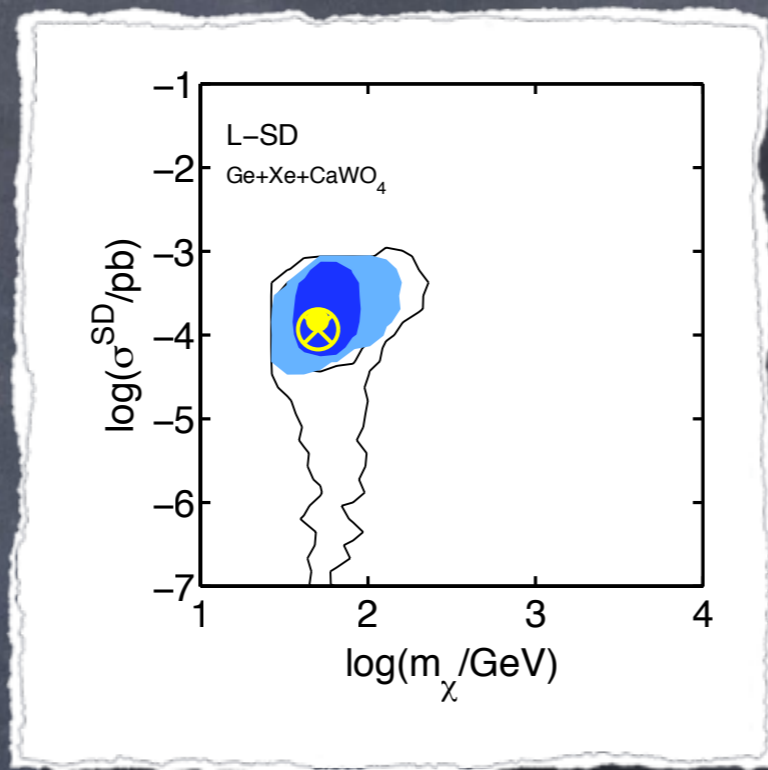
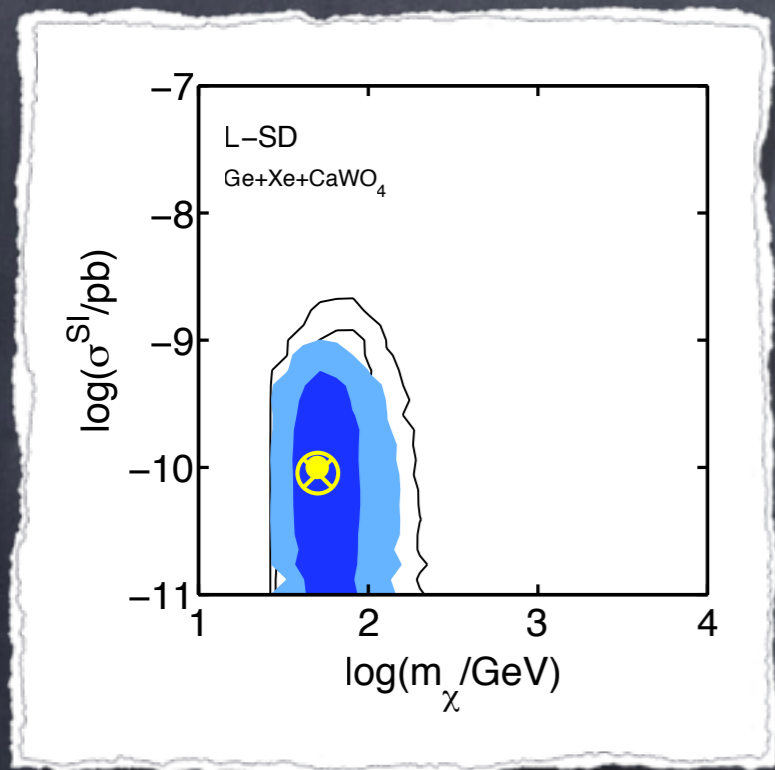
$$\sigma^{SD} = 1.5 \times 10^{-4} \text{ pb}$$

⊗ Best fit point

Mass and SD cross section reconstructed

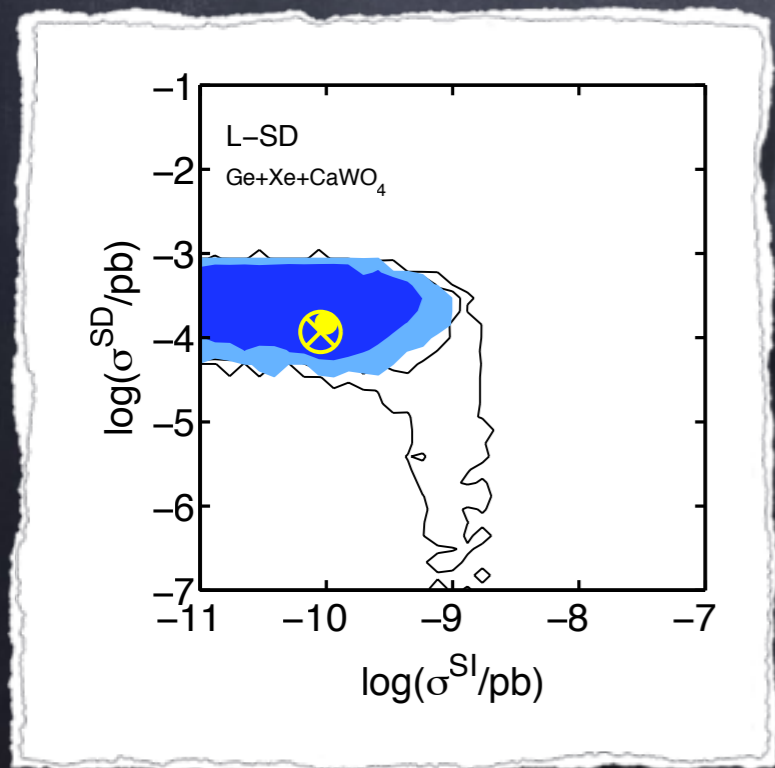
68% and 99%CL profile likelihood contours

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+CaWO₄



● BFM point

$$m_\chi = 50 \text{ GeV}$$

$$\sigma^{SI} = 10^{-10} \text{ pb}$$

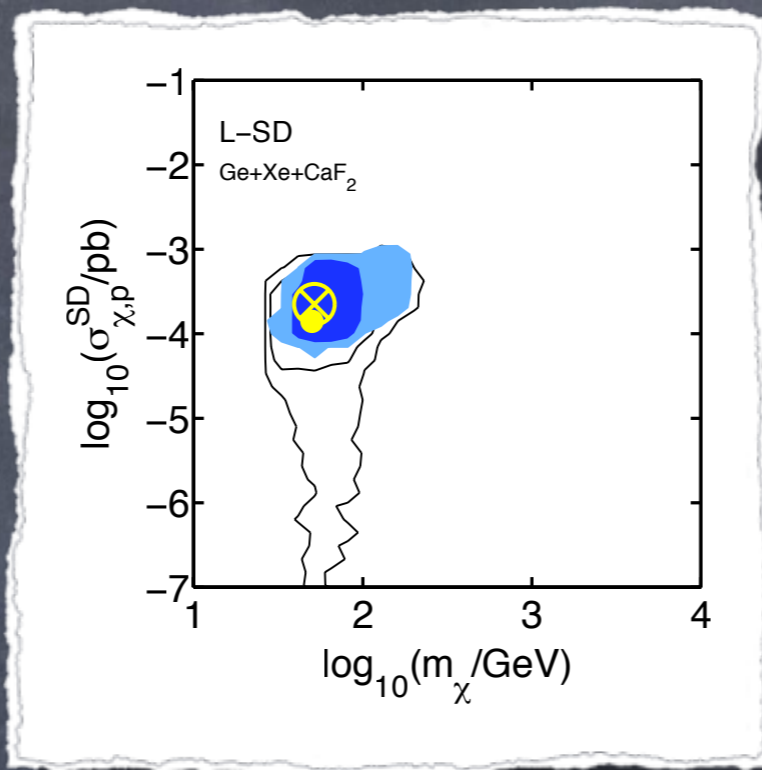
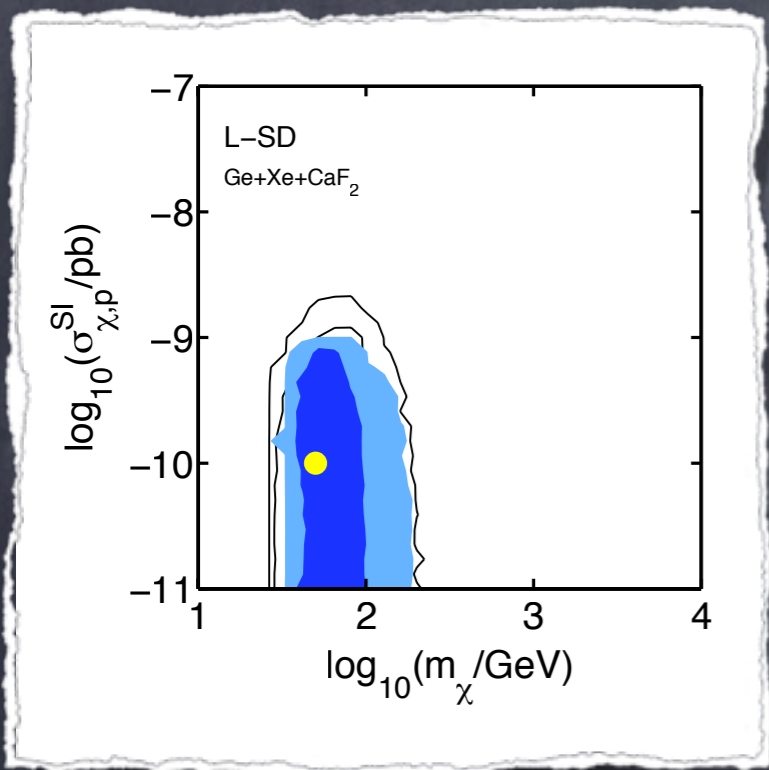
$$\sigma^{SD} = 1.5 \times 10^{-4} \text{ pb}$$

⊗ Best fit point

SD and mass well reconstructed

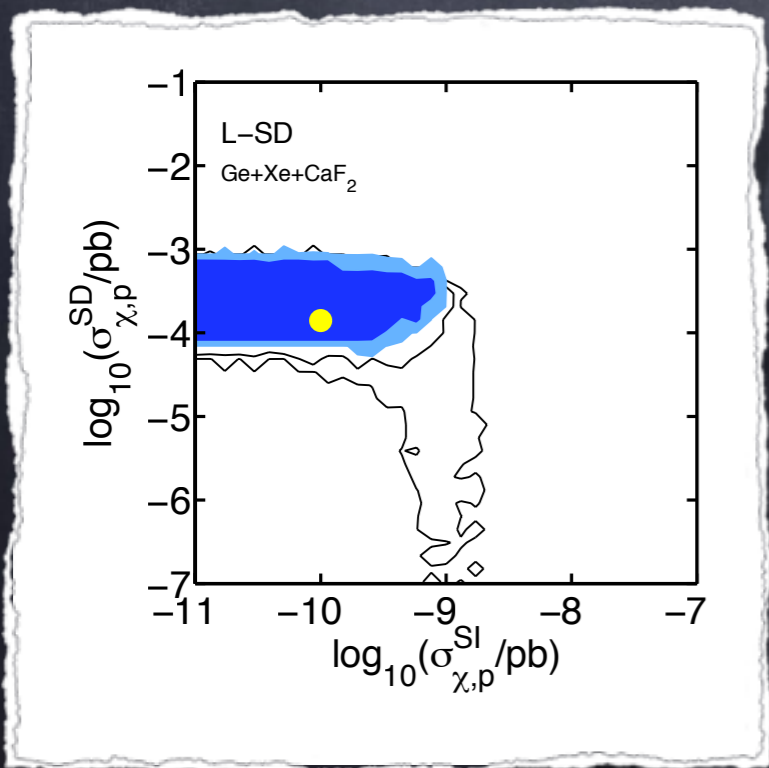
68% and 99%CL profile likelihood contours

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+CaF₂



● BM point

$$m_\chi = 50 \text{ GeV}$$

$$\sigma^{SI} = 10^{-10} \text{ pb}$$

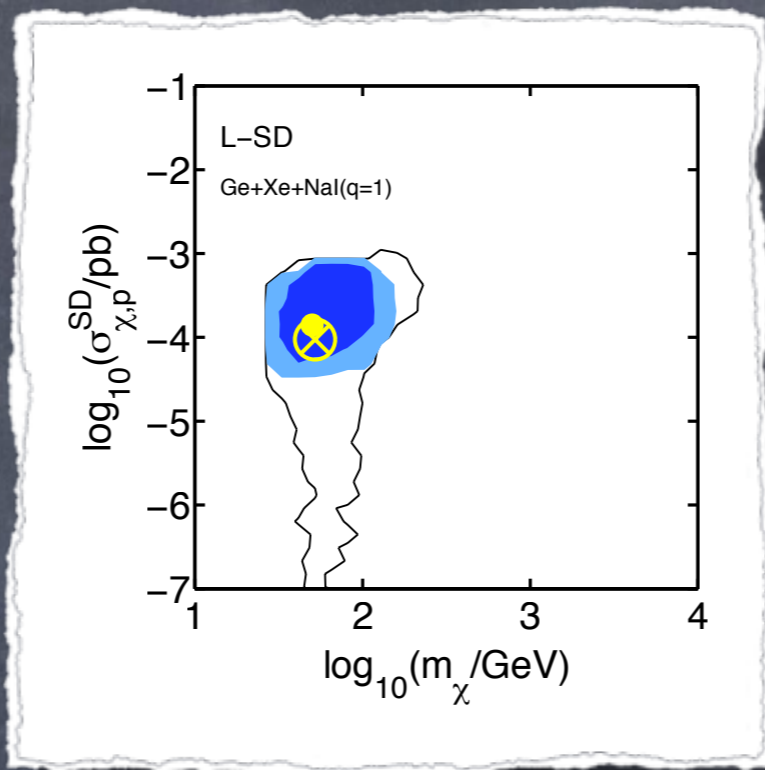
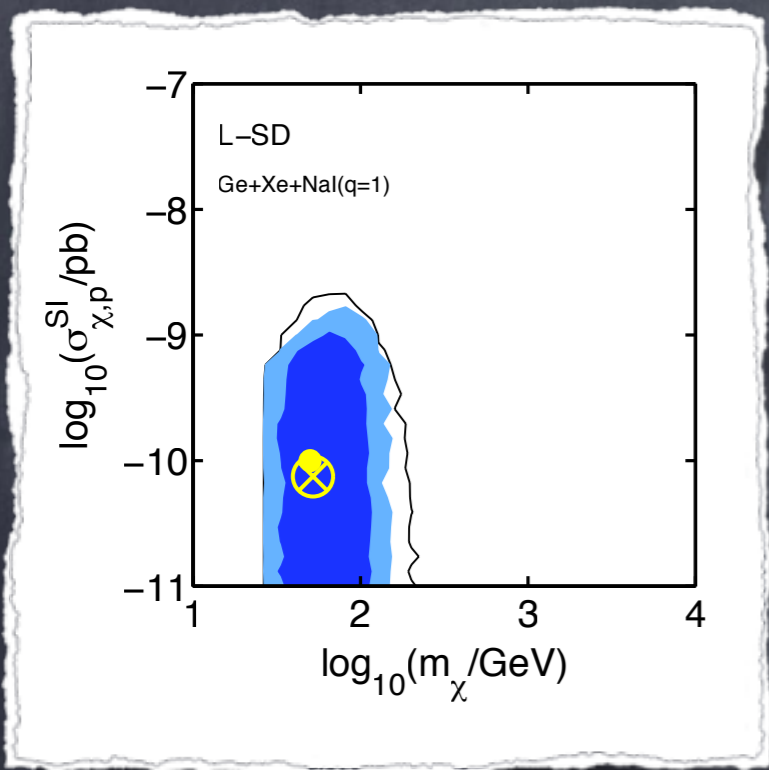
$$\sigma^{SD} = 1.5 \times 10^{-4} \text{ pb}$$

⊗ Best fit point

SD and mass well reconstructed

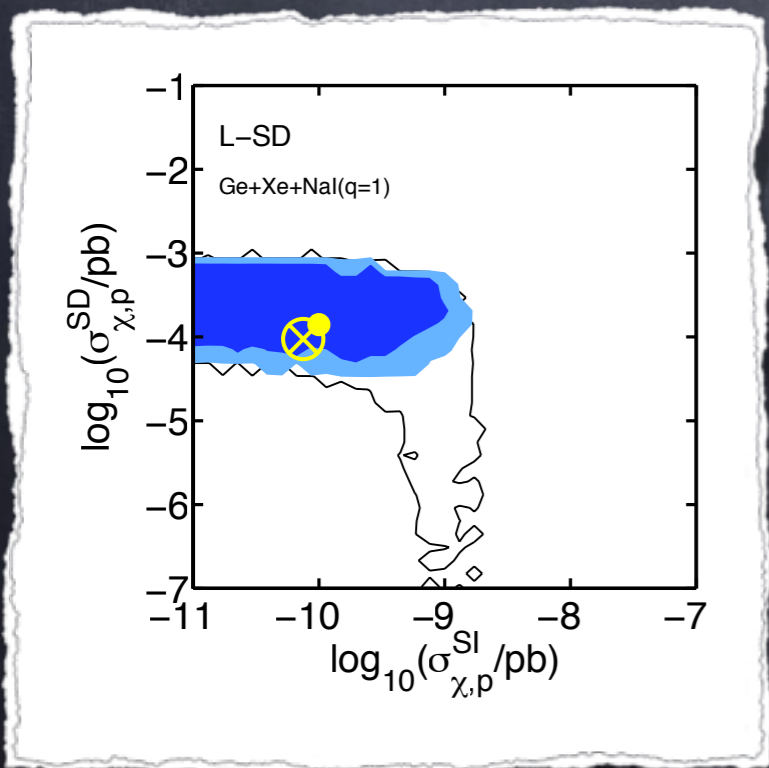
68% and 99%CL profile likelihood contours

Ge + Xe + Bolometer



In black: 68% and 99%CL contours for Ge+Xe

Ge+Xe+NaI



● BM point

$m_{\chi} = 50 \text{ GeV}$

$\sigma^{SI} = 10^{-10} \text{ pb}$

$\sigma^{SD} = 1.5 \times 10^{-4} \text{ pb}$

⊗ Best fit point

SD and mass well reconstructed

68% and 99%CL profile likelihood contours

Current experiments (Ge, Xe)

Ge + Xe + LiF

Ge + Xe + CaF₂

NOT COMPLEMENTARY

Ge + Xe + CaWO₄

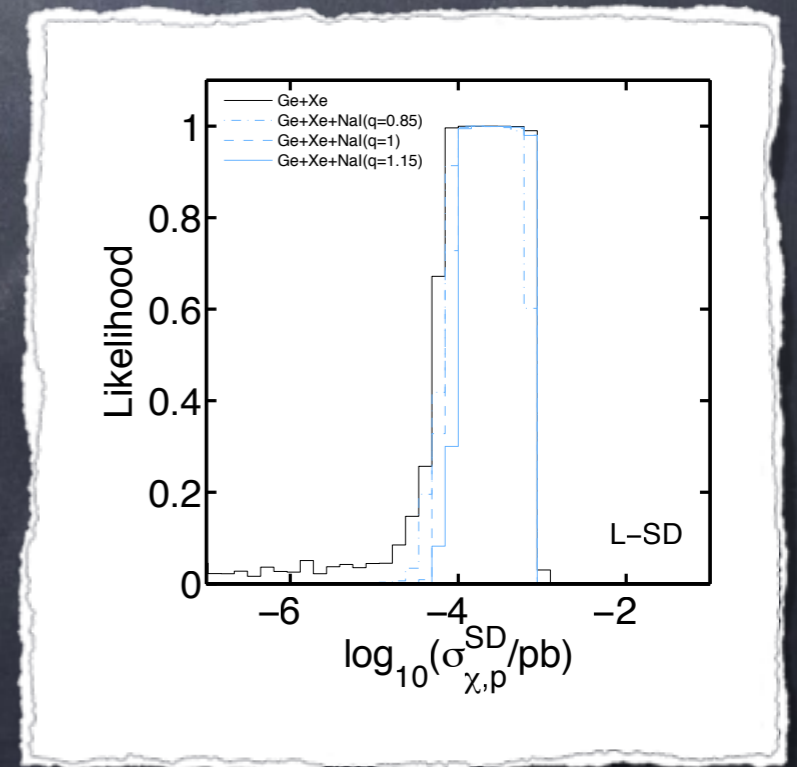
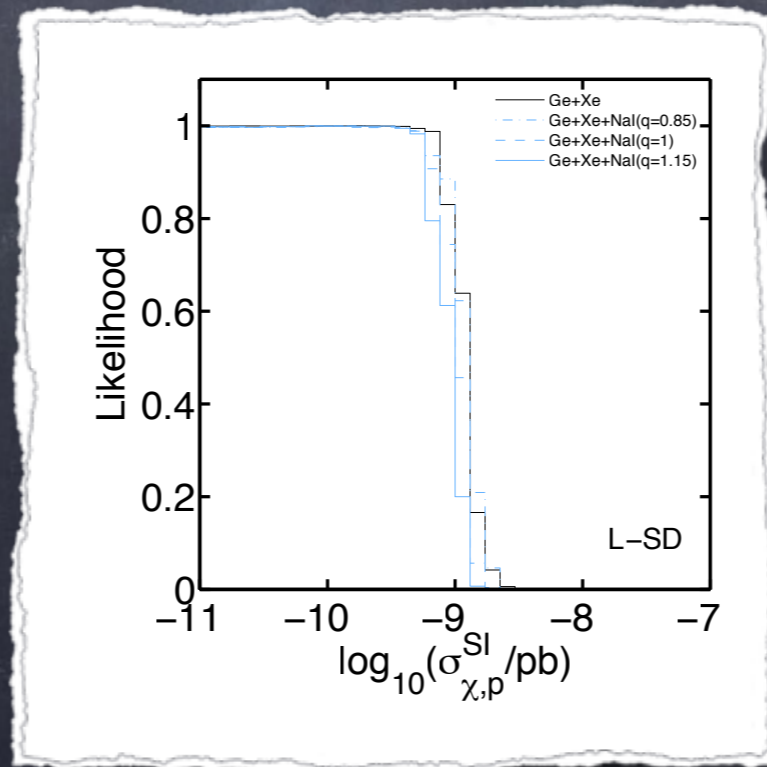
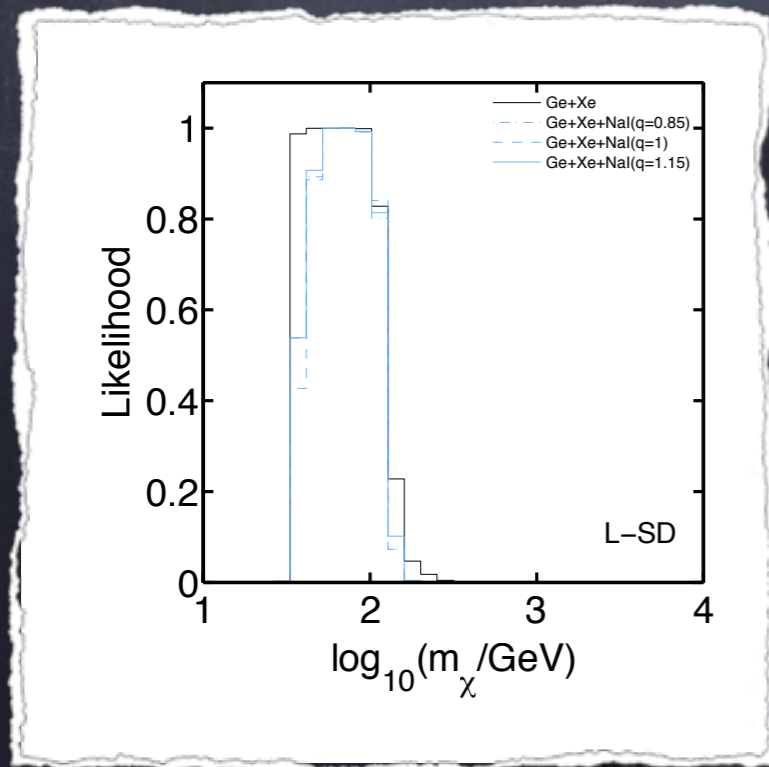
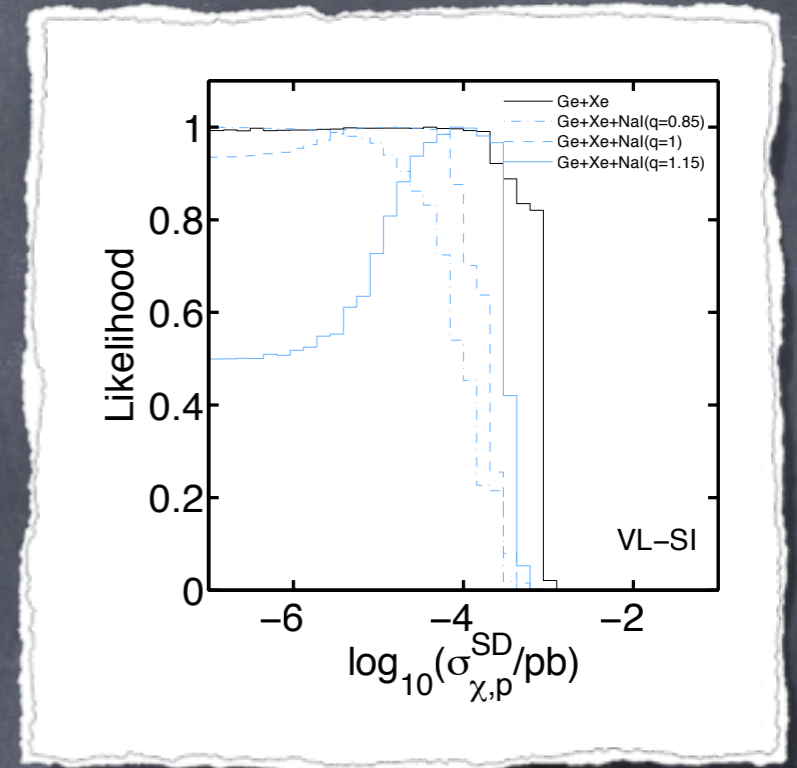
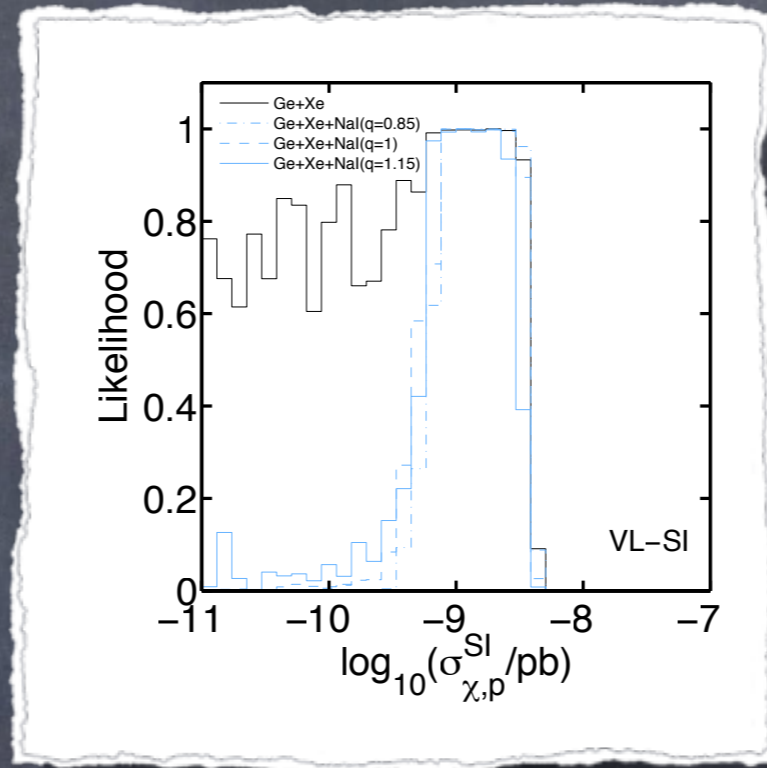
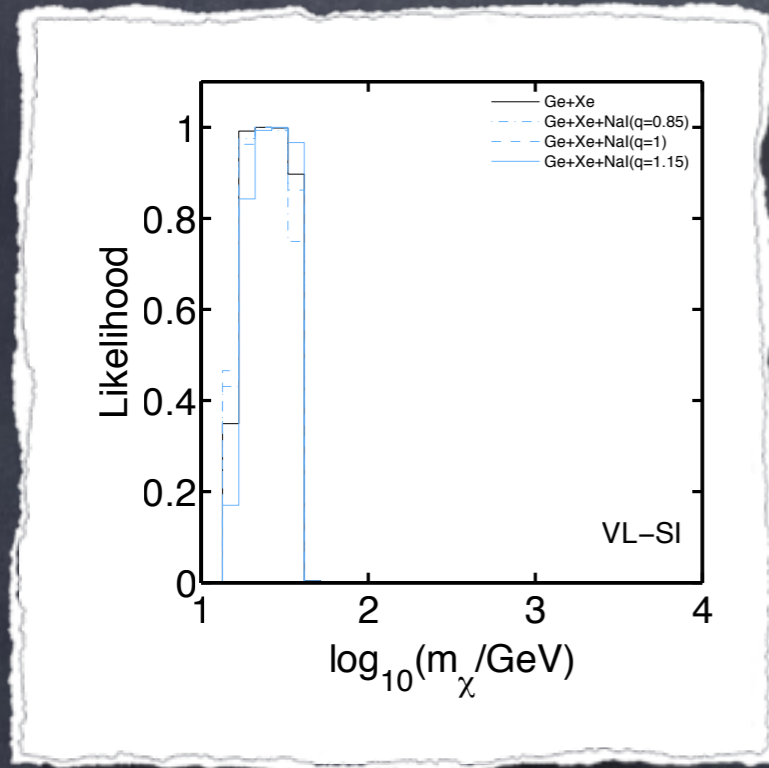
Ge + Xe + NaI

NOT COMPLEMENTARY

HOWEVER, IN ALL CASES THE MASS AND SD CROSS
SECTION CAN BE RECONSTRUCTED AT 99% CL

A GOOD IMPROVEMENT!!!

EFFECT OF QUENCHING (NaI)



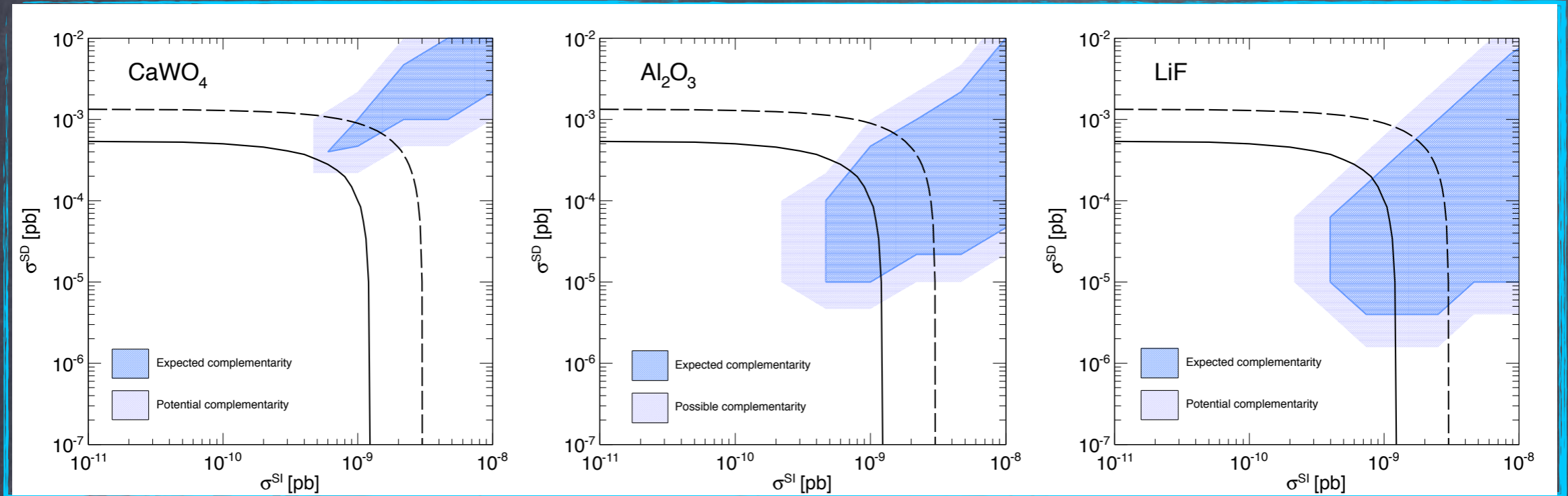
Conclusions

- We investigate how the combination of different targets in direct detection experiments helps in determining the DM properties
- Ge and Xe can be good for discovery but might not be able to measure all DM properties
- The situation seems to be more promising for the targets investigated here: LiF, CaF₂, NaI and CaWO₄
- We are generalizing the analysis for bigger regions of the parameter space and implementing the new experimental features for targets like CaWO₄ (see talk by R. Strauss)

Thanks!

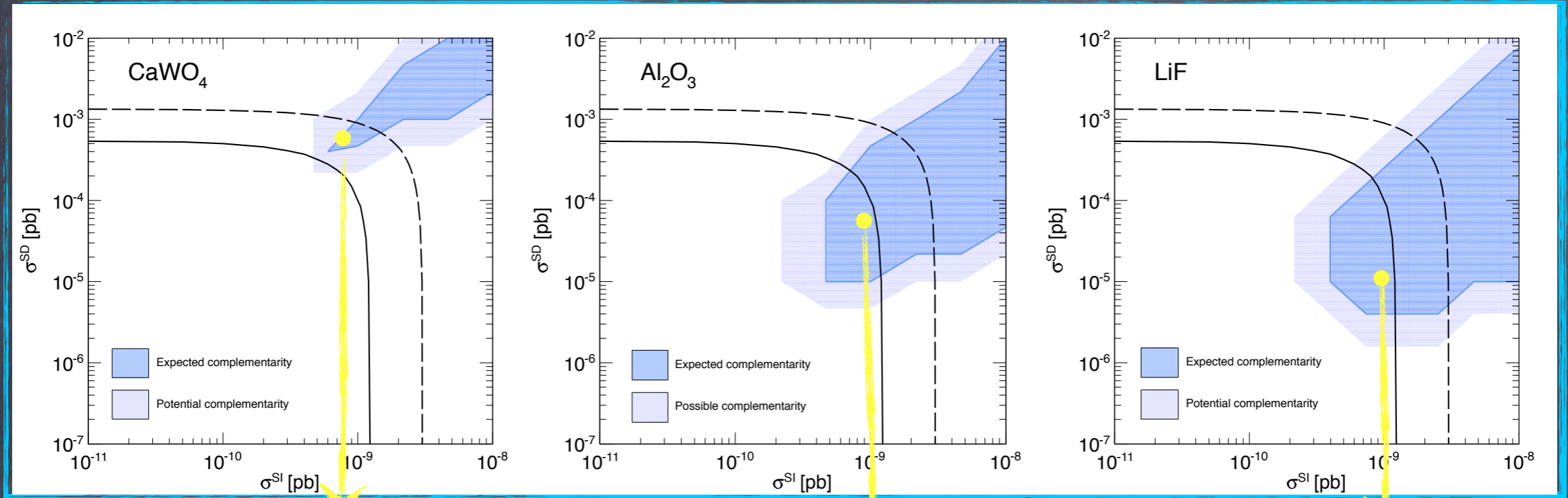
Backup slides

Full scan fixing the mass to 50 GeV



Backup slides

Full scan fixing the mass to 50 GeV



Scanning the background and exposure level to have complementarity

