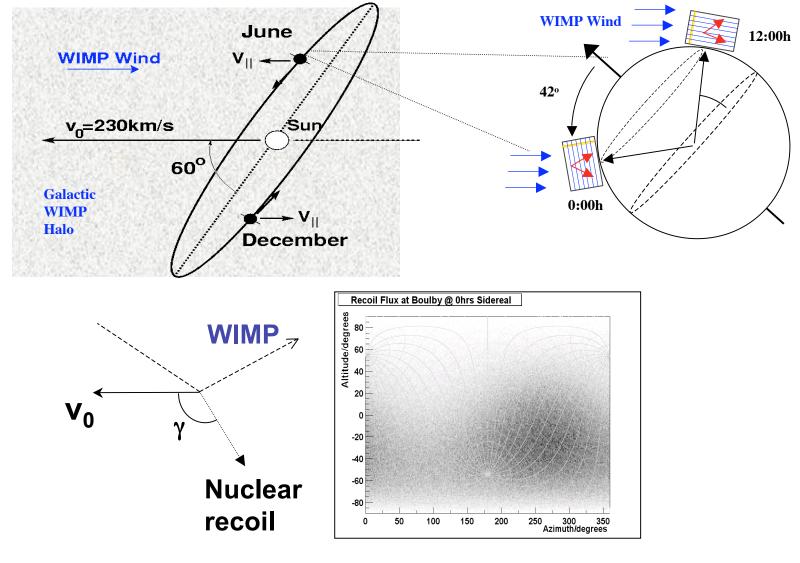
# A Review of the Directional Signature for Dark Matter Searches

# Dinesh Loomba TeVPA/IDM, Amsterdam June 22 2014

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## The WIMP directionality signature



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## **Theoretical Studies**

#### Motion of the Earth and the detection of weakly interacting massive particles

David N. Spergel\* Institute for Advanced Study, Princeton, New Jersey 08540 (Received 21 September 1987)

If the galactic halo is composed of weakly interacting massive particles (WIMP's), then cryogenic experiments may be capable of detecting the recoil of nuclei struck by the WIMP's. Earth's motion relative to the galactic halo produces a seasonal modulation in the expected event rate. The direction of nuclear recoil has a strong angular dependence that also can be used to confirm the detection of WIMP's. I calculate the angular dependence and the amplitude of the seasonal modulation for an isothermal halo model.

v <sub>th</sub> /v <sub>halo</sub>	Fraction of incident flux detected	Forward/back	July/January
0.00	1.00	4.00	1.04
0.20	0.97	4.17	1.04
0.40	0.90	4.66	1.05
0.60	0.78	5.44	1.07
0.80	0.65	6.56	1.08
1.00	0.50	8.10	1.11
1.20	0.37	10.18	1.13
1.40	0.25	12.98	1.16
1.60	0.16	16.73	1.20
1.80	0.10	21.77	1.24
2.00	0.06	28.54	1.28

• Sidereal modulation is about an order of magnitude larger than annual modulation

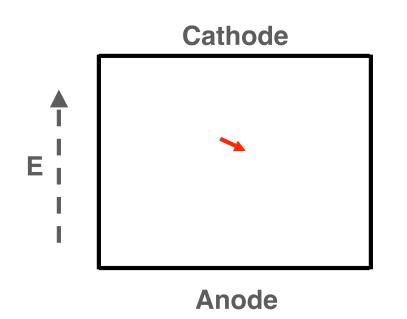
• Huge body of literature on directionality (non-standard halo models, real detector responses, various statistical tests, presence of backgrounds, etc, etc)

The result of this work is that of order 10's - 100's of events are needed to confirm a DM signal

(Lewin, Smith, Krauss, Copi, Green, Morgan, Gondolo, Billard, and many others)

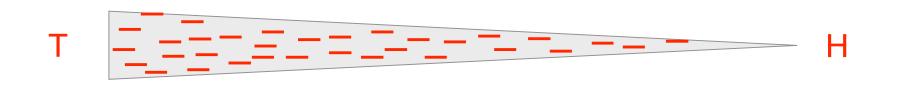
## The Experimental Challenges to Measuring Recoil Tracks

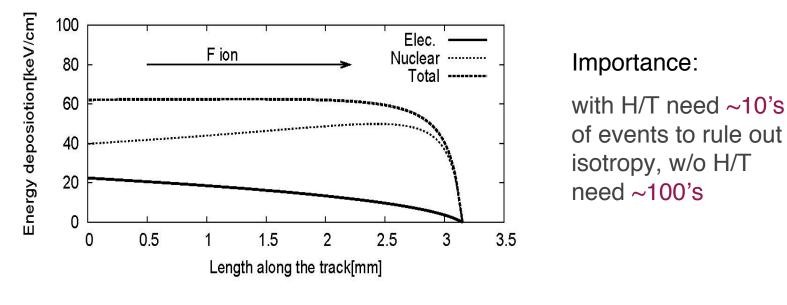
Most experiments use low pressure gas-based TPCs:



The challenge: detecting ~mm tracks in cubic meter volumes

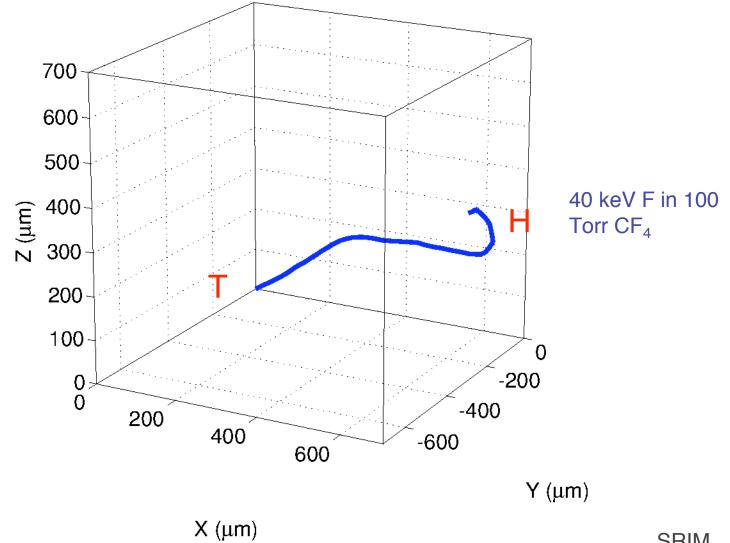
#### Zoom in on the recoil:



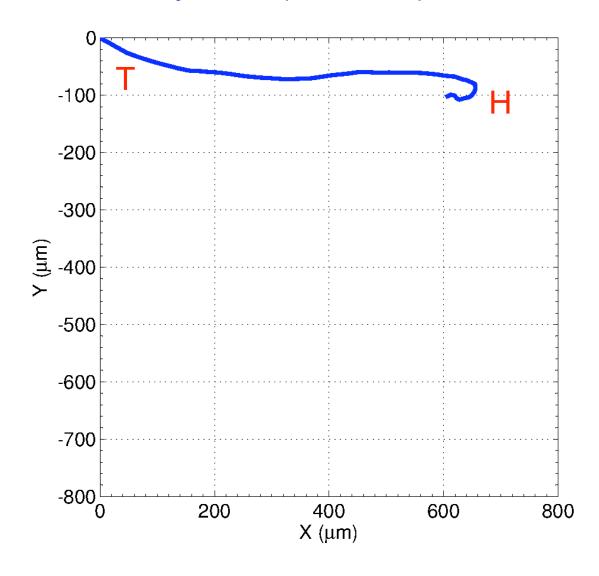


From Tanimori, et al Phys.Lett. B578 (2004) Hitachi's work

#### A real recoil has straggling:



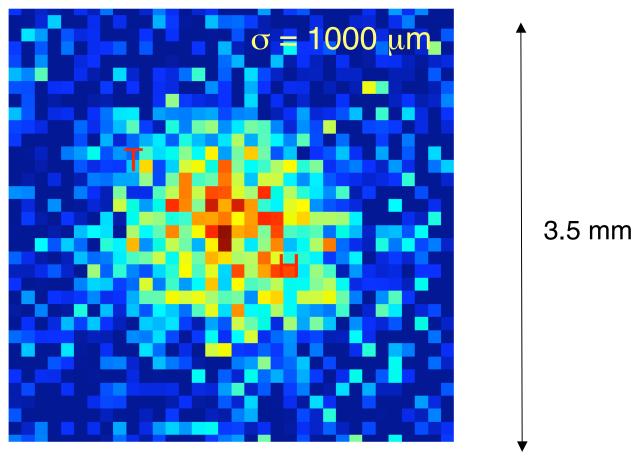
#### Projection (2D or 3D):



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#### Diffusion: you need to keep it low!!



100  $\mu$ m pixel readout

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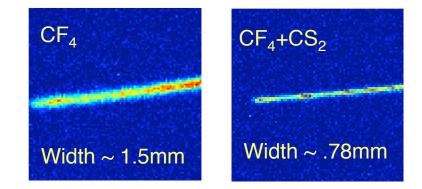
## Flexible technology

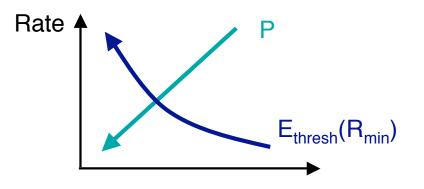
• Flexibility in choice of target A: light targets (He, C, O) for low mass WIMPs, F for spin-dependent, etc.

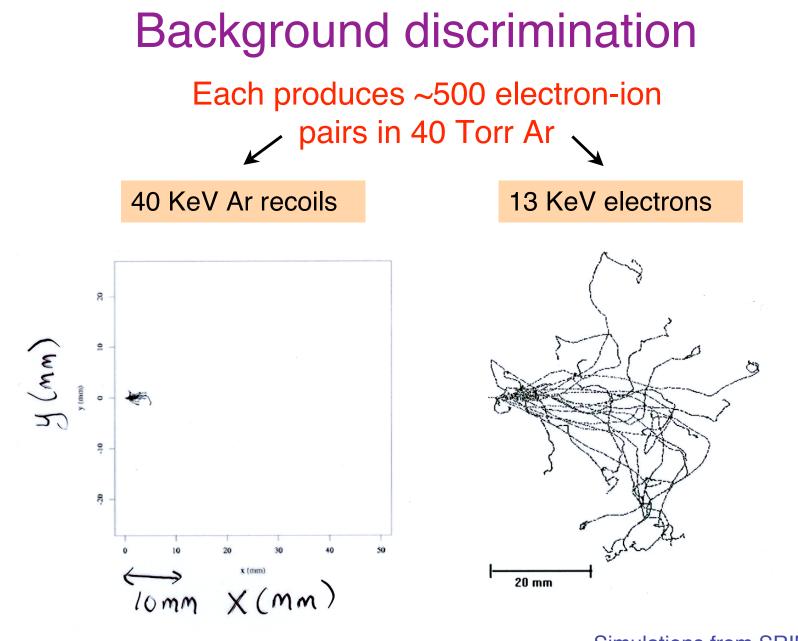
• Negative ion drift: target +CS<sub>2</sub> mixtures enable drift with thermal diffusion (Martoff).

vs Shorter drift distance

Pressure is tunable:
given a minimum
resolvable track-size,
R<sub>min</sub>, one can vary the
directionality E<sub>th</sub> by
lowering pressure:





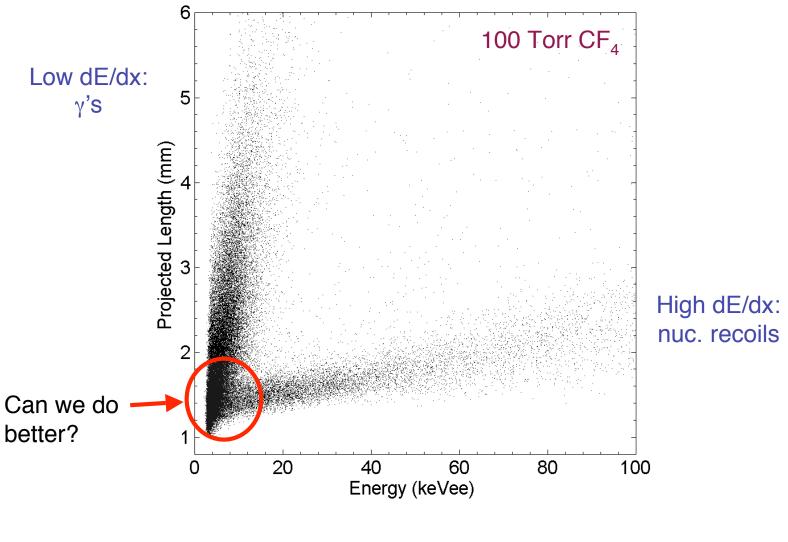


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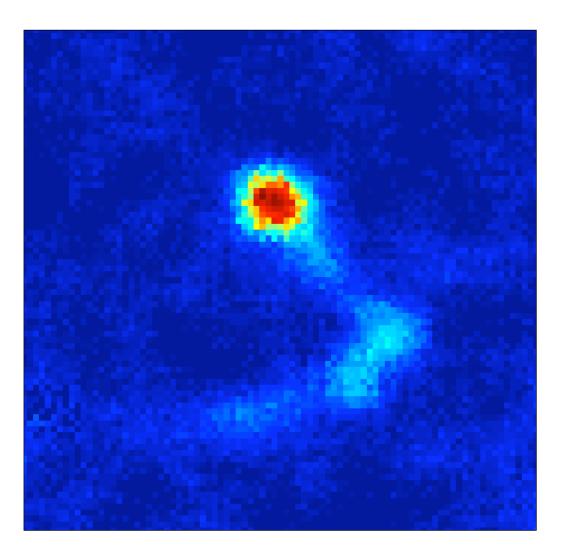
Simulations from SRIM97, EGS4/Presta

### Range vs Energy



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#### Gamma backgrounds have small dE/dx, large fluctuations: high signal-to-noise is important



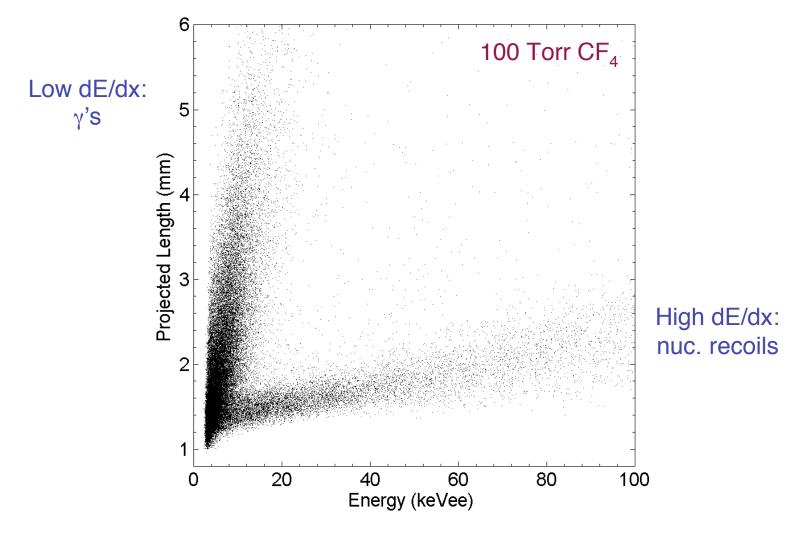
And 3D should help

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An example: discrimination and directionality from an optimized detector

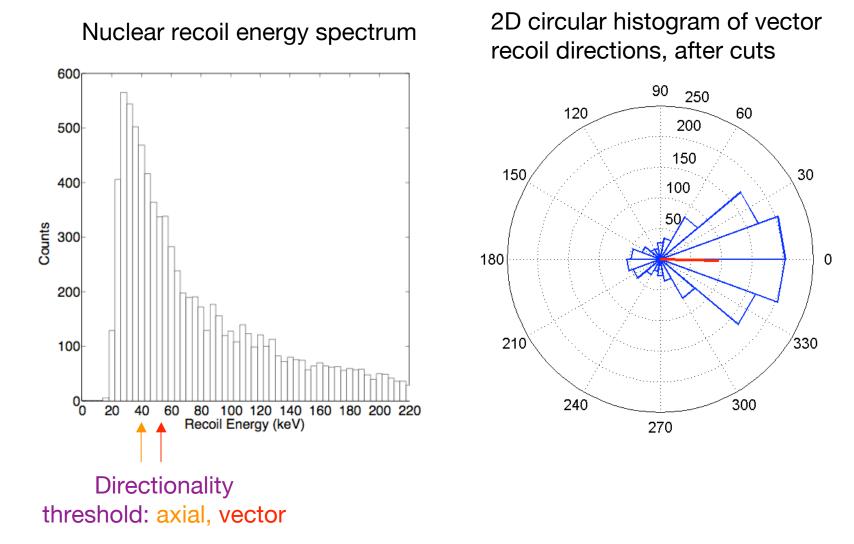
- 100 Torr CF<sub>4</sub>
- High resolution 2D readout with ~160  $\mu m$  pixels
- High signal-to-noise, gas gains ~100,000
- Low diffusion,  $\sigma$ ~0.4mm

#### Range vs Energy from Cf-252

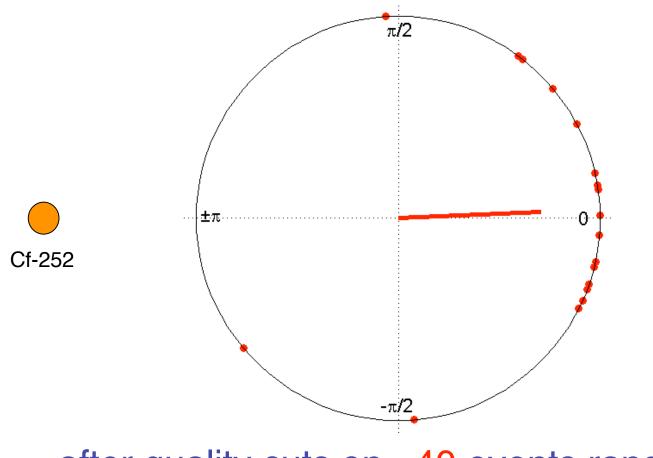


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#### Results from Cf-252 exposure:



#### ~18 events needed to point back to the source...



...after quality cuts on ~40 events randomly chosen from dataset with vector directionality

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### Comments:

- **40** events needed, from which ~18 with high track quality are selected
- These ~18 sit in the high energy tail of the spectrum
- About 1/4 1/3 of the recoils in the dataset are produced by scattered neutrons. So, these results are conservative

# Using Cf-252 results to simulate directionality for 100 GeV WIMP ~54 events needed to locate source (90% CL)

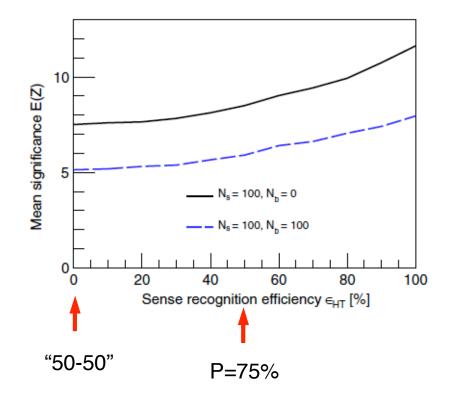
## Pretty good...BUT why ~50 events and not ~10, as expected from vector directionality??

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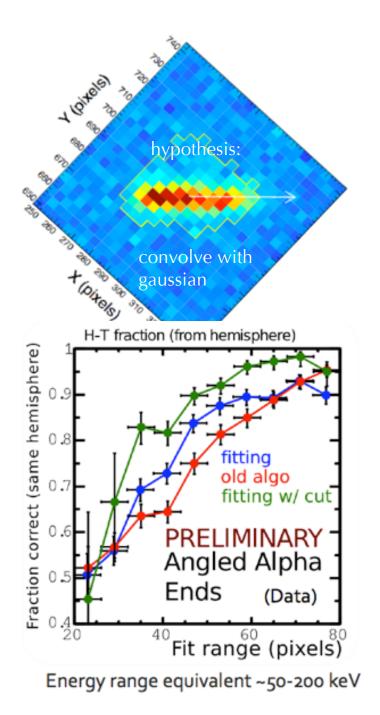
#### Quality of head-tail:

Green, Morgan (PRD 77, 2008), and Billard et al (PRD 85, 2012) point out importance of a high probability of getting head-tail correct:

"The linearly increasing sense determination probability with  $P(E_{th}) = 0.75 (0.50) \dots$  requires ~3 times (10 times) more events than the constant P = 1."



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#### Comments:

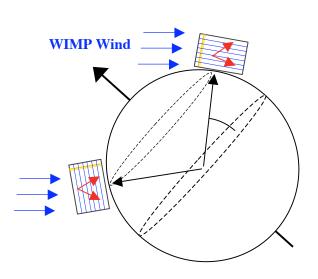
• 3D should help and nature provides a population of less straggled events.

• Maybe its better to pick a higher directionality threshold, e.g., one where P>0.7? If minimizing the directional exposure is the goal, this data says: NO, its better to pick the lowest possible threshold where there's directionality. Consistent with Green, Billard.

• Gas-based TPCs have a unique knob, P, plus the ability to vary target A. So one can lower the directionality threshold, e,g, for a low mass WIMP search. (See Nguyen Phan's talk).

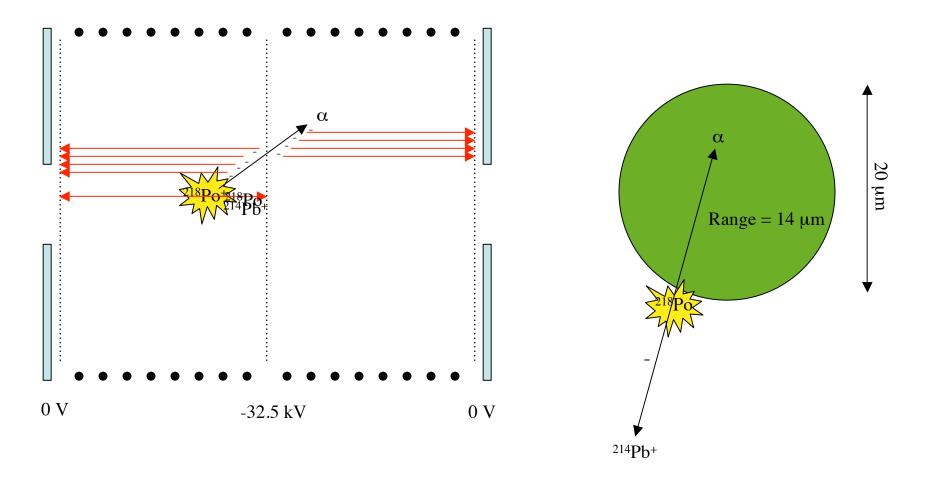
• Better sense-sensing algorithms. From *C. Deaconu, TAUP 2013 Proceedings* -See Jocelyn Monroe's talk)

# Good directionality and discrimination.. but where's the scaleup??



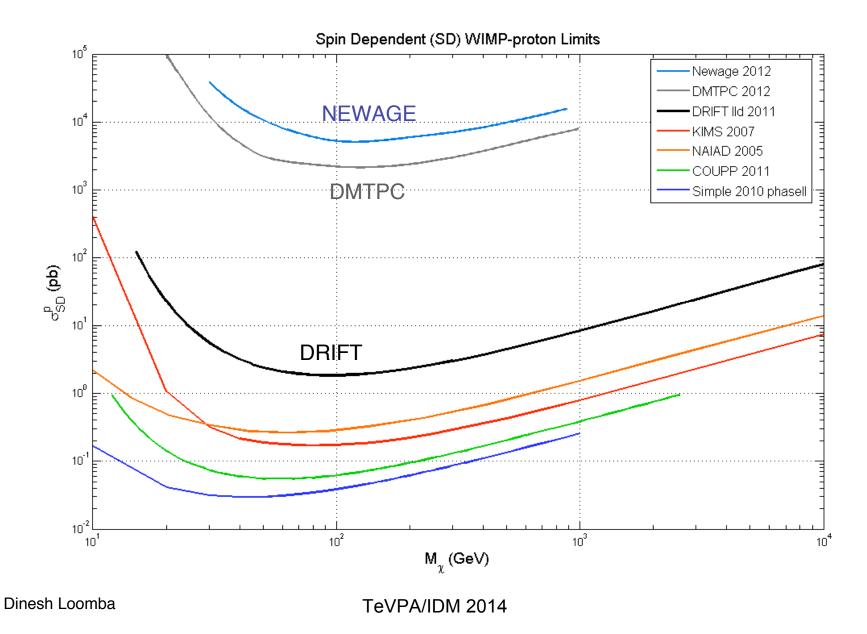
WIMP Wind

## A detour: Radon Progeny Recoils (RPRs)



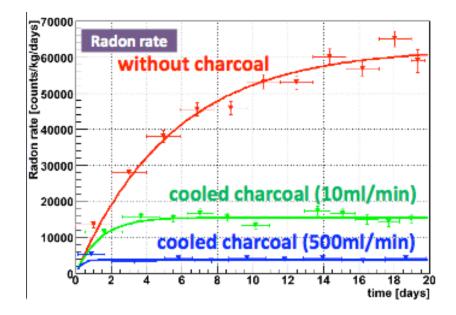
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## Limits nonetheless



## Radon/RPR reduction

- Identifying and eliminating radon emanating materials
- Recirculation thru cooled charcoal (NEWAGE)
- Development of thin film cathodes to improve efficiency of vetoing RPRs (DRIFT)

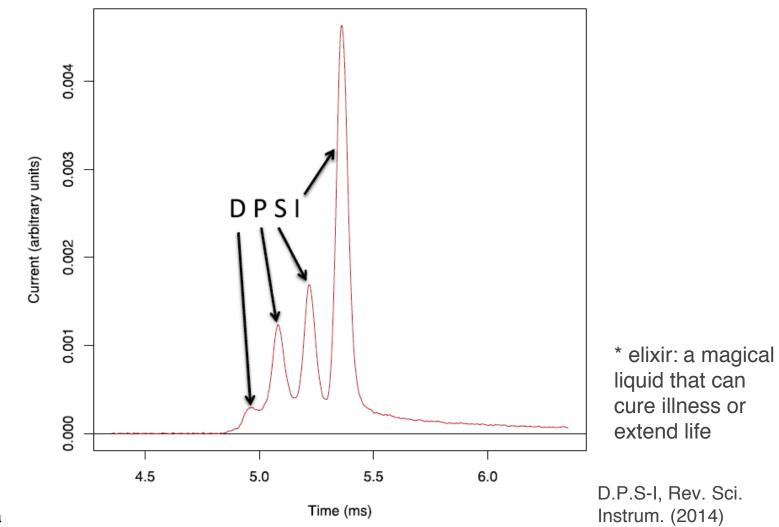




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## Z Fiducialization: the "holy grail"

Discovery of "minority peaks" in  $CS_2^* + O_2$  mixtures:

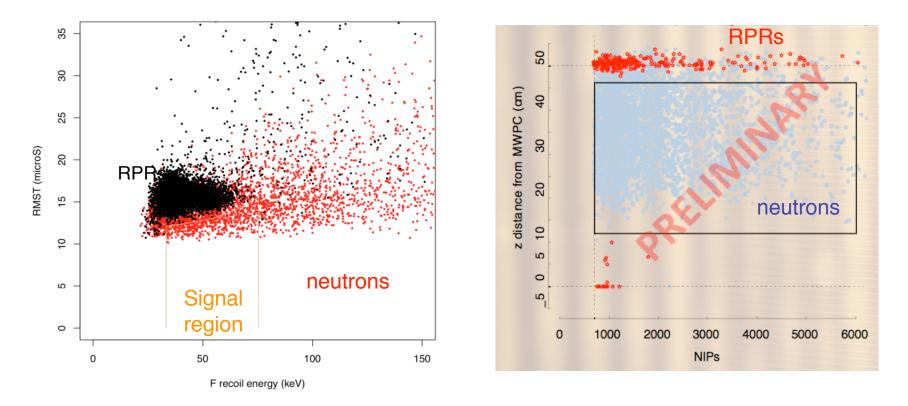


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#### For DRIFT: an expanded signal region and **zero** backgrounds

Before





Both are from ~50 day dark matter runs at Boulby

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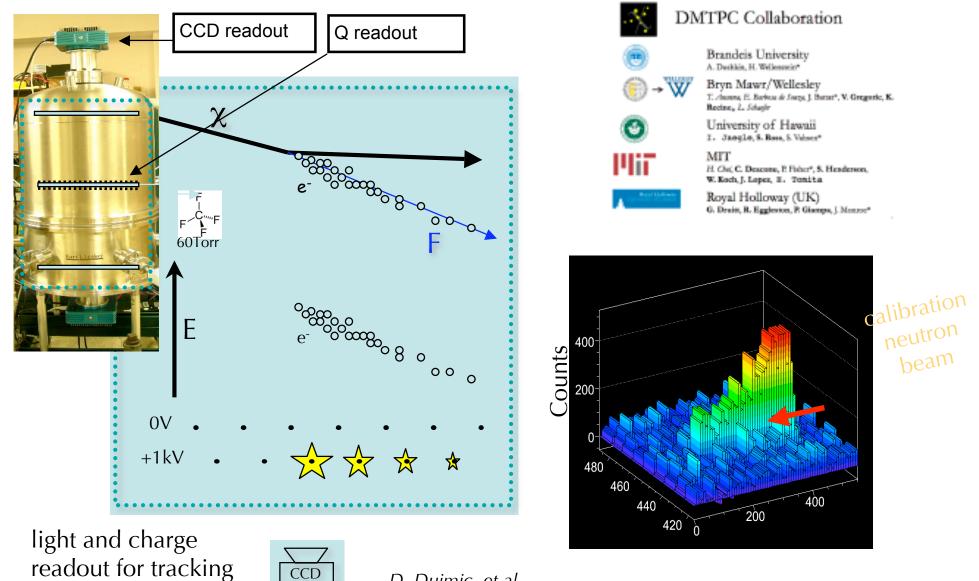
(See Neil Spooner's talk)



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## **Directional Experiments**

### **Dark Matter Time Projection Chamber**



D. Dujmic, et al., NIM A 584:337 (2008)

camera

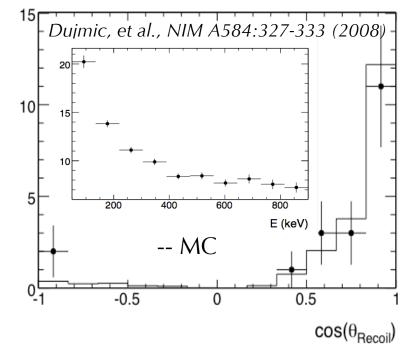
& backgrounds

#### Directionality

- 2D angle + head-tail from fitting light asymmetry (measure skewness)
- Require range/width > 3 for ID
- Diffusion has a big impact! Therefore working to lower pressure
- R&D for 3D proceeding



cosine (E>200 keVr), 5 cm drift

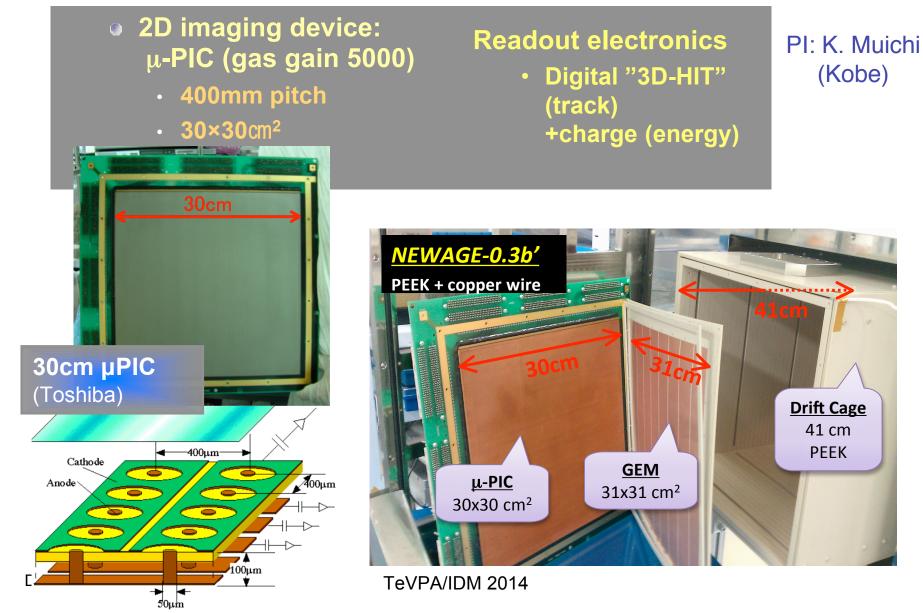


Future: DMTPCino 1m<sup>3</sup> Detector prototype for very large detector: build many 1m<sup>3</sup> modules, because of diffusion limit.

Detector under construction now, vacuum vessel acceptance test May 2014, commissioning Fall 2014

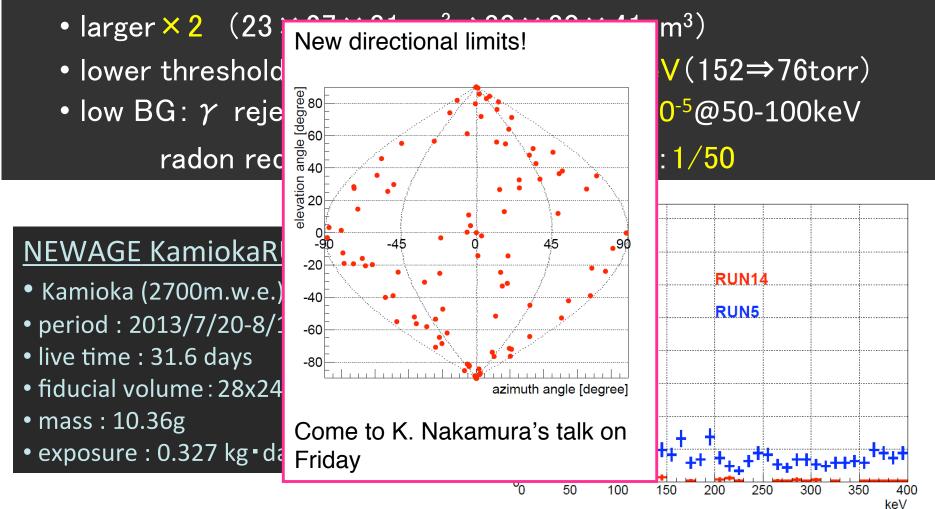
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# NEWAGE - New generation WIMP search with an advanced gaseous tracker experiment)



# NEWAGE 2013 results

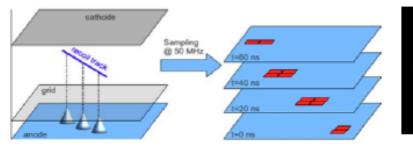
#### • NEWAGE-0.3b



## MIMAC

## **Concept:** low pressure CF<sub>4</sub>, CHF<sub>3</sub> and H with charge readout via Micromegas + pixel technology

- X and Y coordinates are measured on the pixelated anode
- Z direction by anode sampling at 50 MHz
- The anode is read every 20 ns. The 3D track is reconstructed, from the consecutive number of images defining the event Bi-chamber module 2 x (10.8x10.8x25 cm<sup>3</sup>)



Pixel micromegas from IRFU (Saclay) - 200 µm

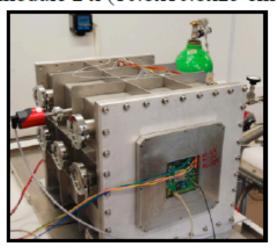


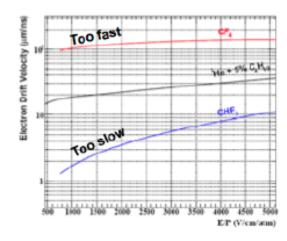
LPSC (Grenshle) : J. Lamblin, F. Mayet, D. Santas J. Billard (Ph.D.) (left in July 2012), Q. Riffard (Ph.D) (started in October 2012)

Technical Coordination :	O. Guillandin	
- Electronics :	G. Bosson, O.Bourrion, J-P. Richer	
- Gas detector :	O. Guillaudin, A. Pellisier	
- Data Acquisition:	O. Bourrion	
- Mechanical Structure :	Ch. Fourel, S. Roudier, M. Marton	
- Ion source (quenching) :	J-F. Muraz, J. Médard (CDD-1year)	

CCPM (Marselle): J. Busto, Ch. Tao, D. Fouchez, J. Brunner (Radon filtering)

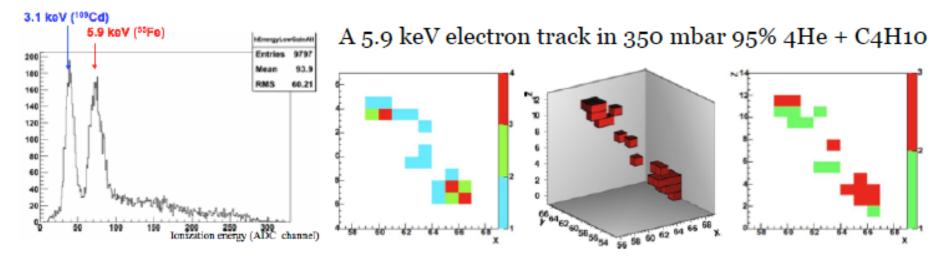
Neutron facility (AMANDE) : IRSN (Cadarache): L. Lebreton, D. Maire (Ph. D.)



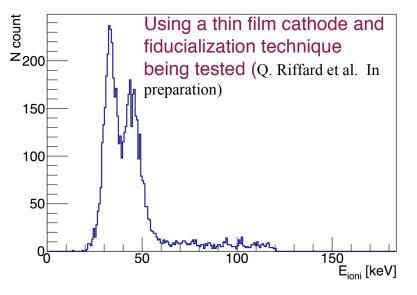


New mixed gas MIMAC target needed to slow drift velocity to match speed of electronics time slicing : CF4 + 30% CHF3

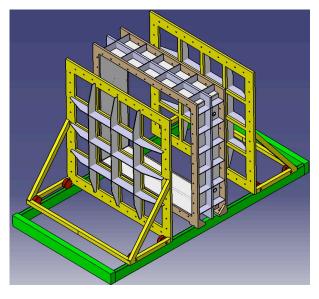
#### Underground in Modane since 2013



#### RPRs! Rn-222 daughter spectra:



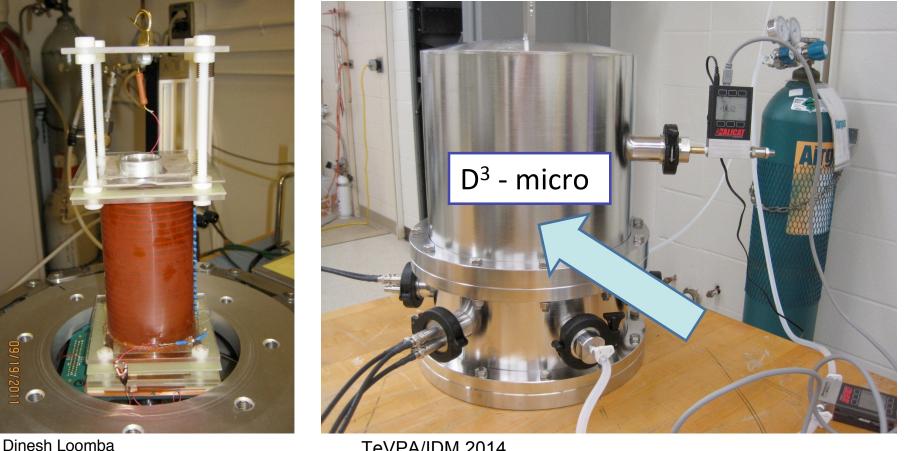
Future: MIMAC – 1m<sup>3</sup> = 16 bi-chamber modules (2x 35x35x25.5 cm<sup>3</sup>)



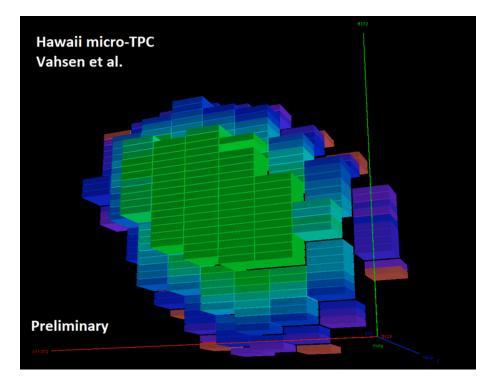
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## D<sup>3</sup> - Directional Dark Matter Detector

- Investigating feasibility of directional DM search w/ micro-pattern gas detectors
- Technology also of interest for detecting neutrons and charged particles
- Small (1-60 cm<sup>3</sup>) prototypes built at LBNL (PI: JA Kadyk) and U. Hawaii (PI: S. Vahsen)
- Ongoing since ~Fall 2010



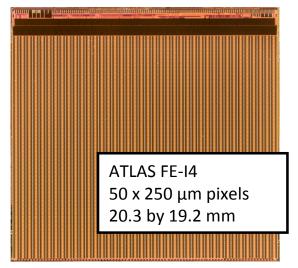
# 3D tracks: ~25 keVee recoiling He nucleus

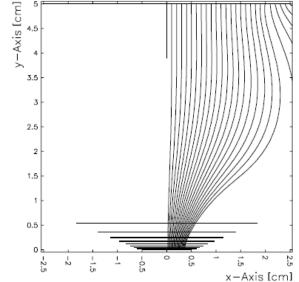


- each block: 50x250x250 μm<sup>3</sup>
- Color: ionization density

**S. Ross** et al., "Charge-Focusing Readout of Time Projection Chambers", proceedings of IEEE NSS 2012

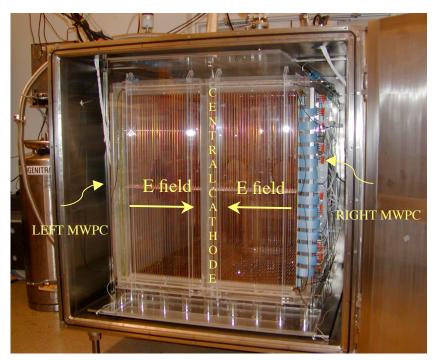
How to reduce # chips: larger ATLAS chips + focusing



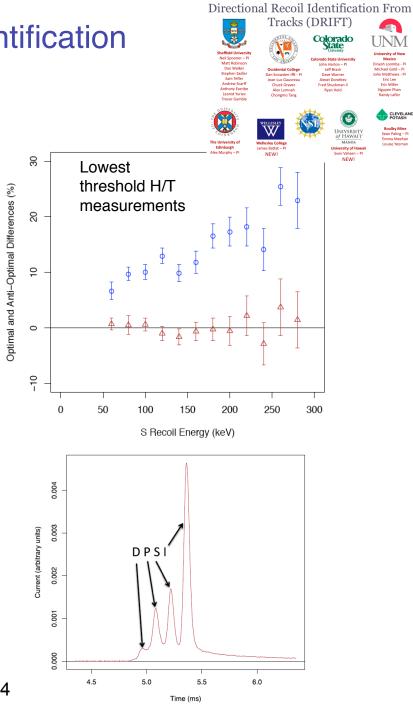


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### DRIFT - Directional Recoil Identification From Tracks



- + 1m<sup>3</sup> Negative Ion TPC with 30+10 Torr  $CS_2$ +CF<sub>4</sub>
- MWPC readouts
- Operating in Boulby for >10 years
- Operating with ZERO backgrounds for >50 days!

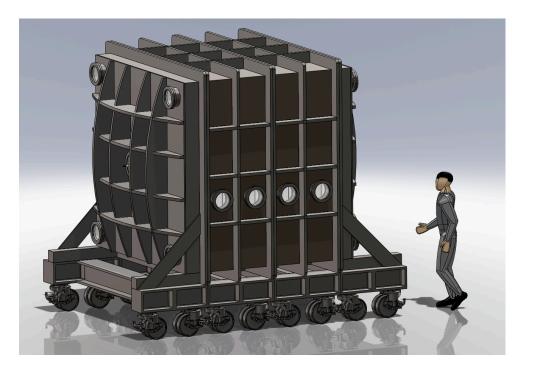


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#### Future goals

- Deploy 2nd DRIFT-II detector with transparent MWPCs capable of reading from both sides
- Higher signal-to-noise every-wire readout
- R&D on high resolution readout (GEM/Micromegas + strips/pixels)
- R&D on directionality for low-mass WIMPs

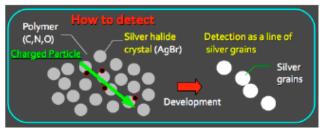
DRIFT has now entered a new phase, going from being background limited, to being volume limited. Ready to scale up by ~30 to the DRIFT-III detector:

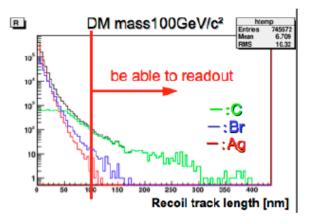


## **Emulsion Dark Matter Search Project**

Nagoya (T. Naka), Napoli, Padova, LNGS

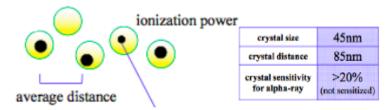
**Concept (1)**: Use of emulsion film to give 3D tracking - solid detector (3g/cc), high spatial resolution, low cost, target Ag(46%), Br(34%), C(N,O) (19%)





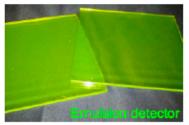
 Progress made to produce stable very fine crystals by using the PVA techniques

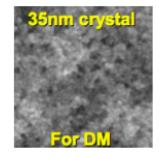
• Track produces line of silver grains



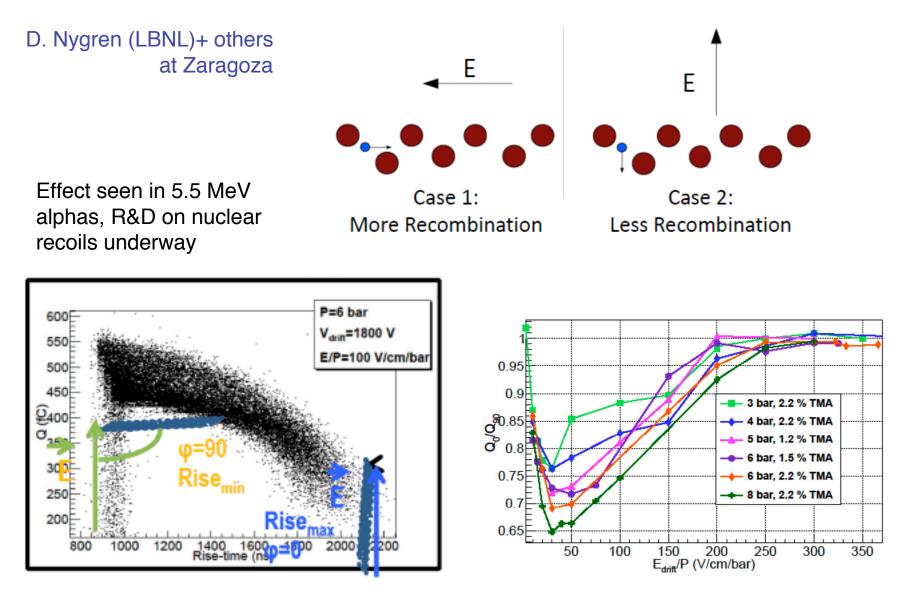
- Challenge is to get: (i) small grains <40nm (OPERA had 200 nm), (ii) closely packed, and (iii) sensitive to low ionisation
- Typical recoils are order 100nm Ag, Br likely produce tracks too short so need to use

C, N, O target





#### Columnar recombination in HPXe



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## Summary

- Experimental effort have demonstrated and improved their directionality. We are approaching the simulated predictions on strength of signature.
- Discrimination with range vs energy is excellent. But good signal-to-noise and resolution will be needed as we push to lower thresholds.
- Flexibility of technology enables numerous approaches for probing a wide range of WIMP parameters.
- The RPR problem has a solution that works.
- Experiments are all getting larger, but we are volume limited and need to scale up.
- Many new groups entering field, some with novel ideas in solids-state and high pressure gas

