

New physics searches in ATLAS and relation to astroparticle physics

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Outline

► Introduction

- Evidence of Dark Matter
- DM candidates
- DM detection

► Search of DM at ATLAS

- LHC and the ATLAS experiment
- SUSY searches
- Other searches

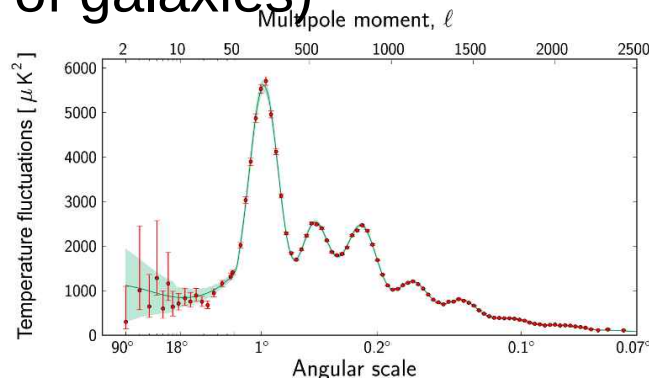
Evidence for Dark Matter

- From speed dispersion measurements in clusters of galaxies, and orbital speed of stars in galaxies
 - mass of luminous objects \ll total mass
- Weak gravitational lensing : get distribution of matter, in clusters

- most of it not accounted for by atoms

$$\Omega_{DM} \gtrsim 0.1 - 0.2$$

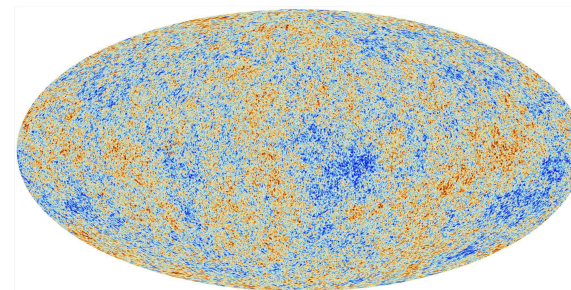
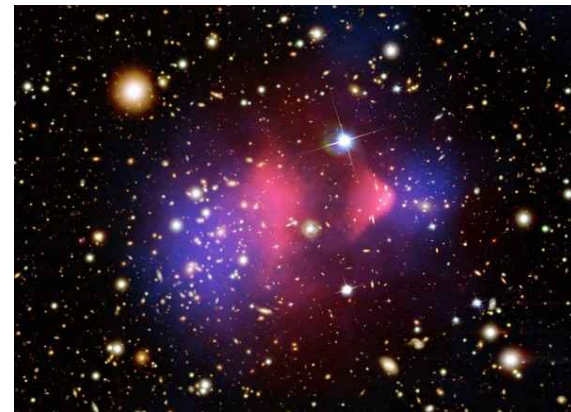
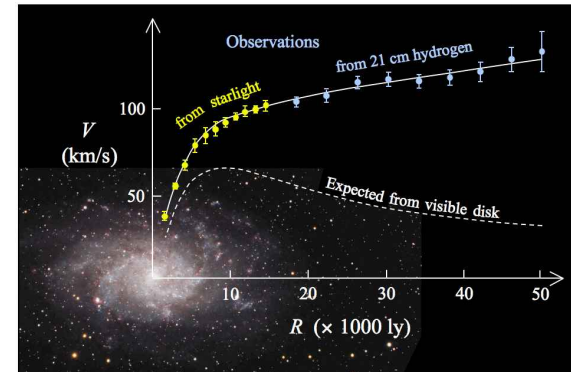
- From fit of cosmological parameters to measurements (CMB anisotropy, spacial distribution of galaxies)



$$\Omega_b = 0.0499(22)$$

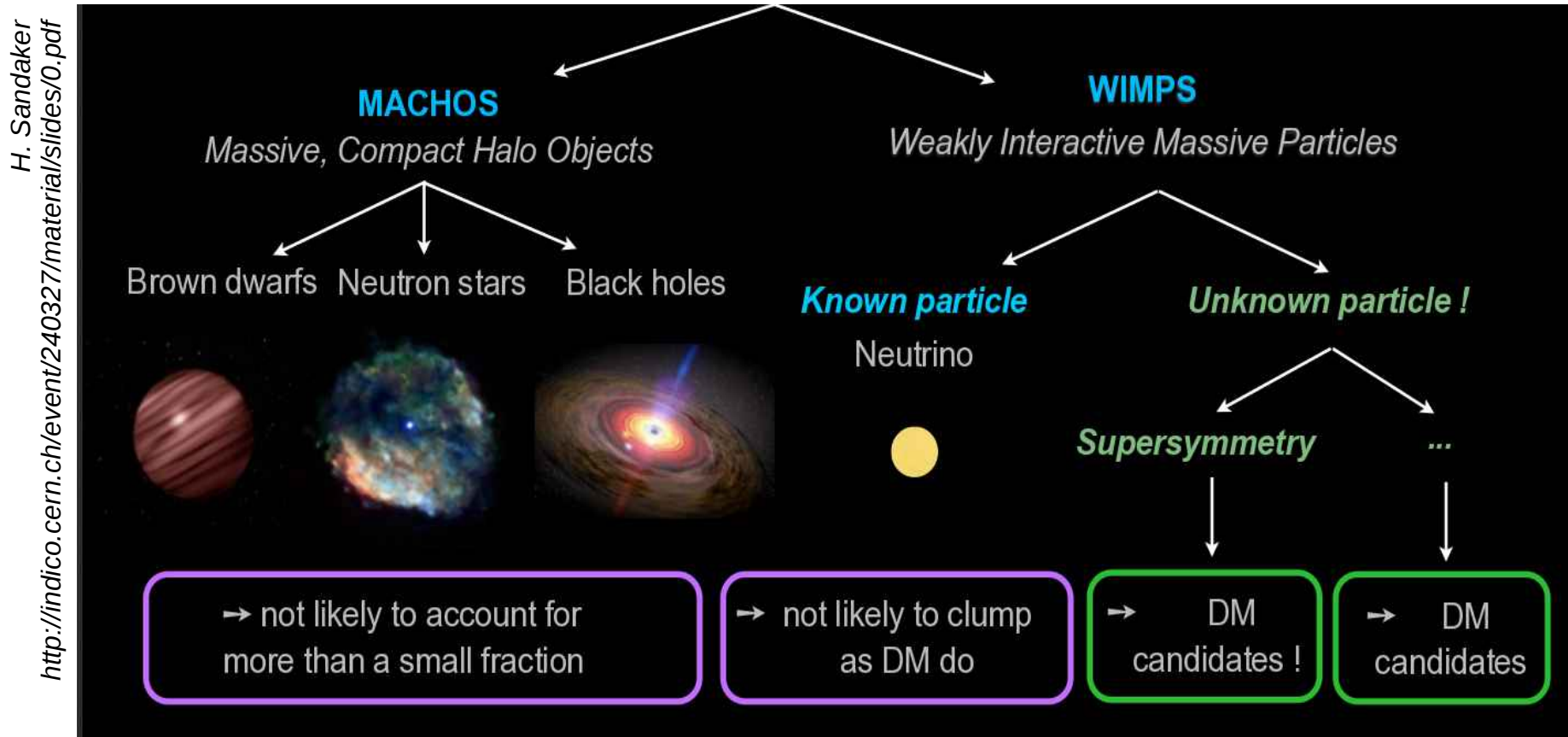
$$\Omega_{nbm} = 0.265(11)$$

- Predominance of non-baryonic matter in universe



Dark matter candidates

- ▶ Few points DM candidate should satisfy (arXiv:0711.4996)
 - Neutral, “cold”, consistent with relic density, ...

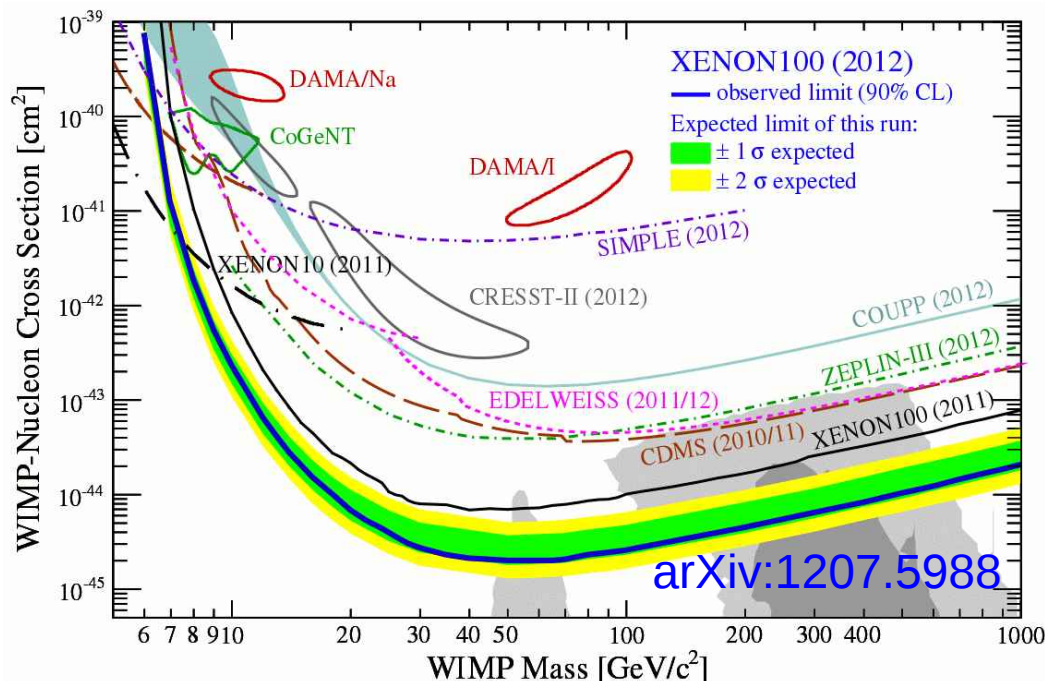
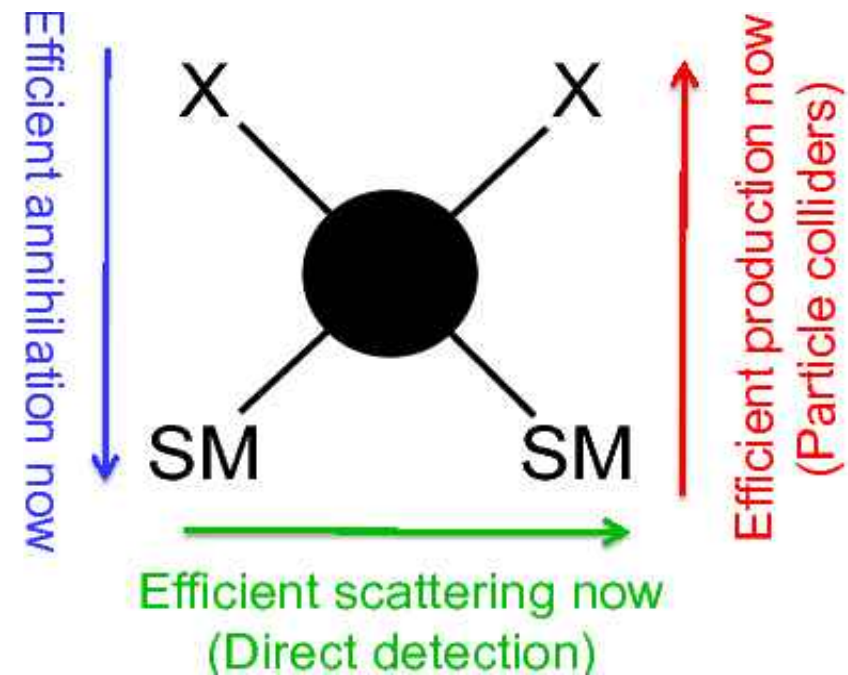


- ▶ Apparently none of the Standard Model particle has the right properties
- ▶ We need to look for new physics beyond the SM

Dark Matter detection

► Different types of experiments for DM search

- **Direct detection** : elastic scattering between nuclei (detector) and DM halo
- **Indirect detection** : look for decay products of DM annihilation
- **Particle colliders** : search for DM production from SM particle interaction



► Many results

- Some experiments found evidence of WIMP signature
- Some gave exclusion limits
- Sensitivity is improving
- LHC experiment can contribute with independent results

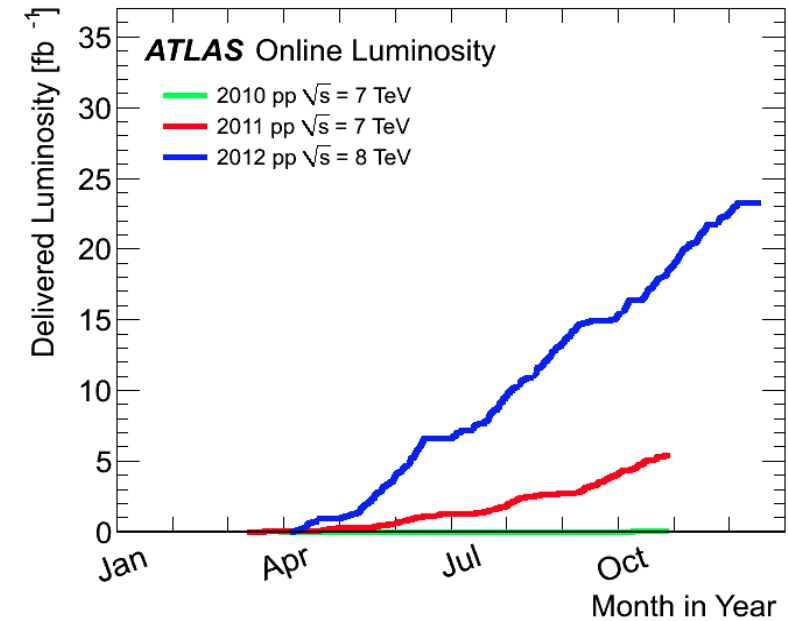
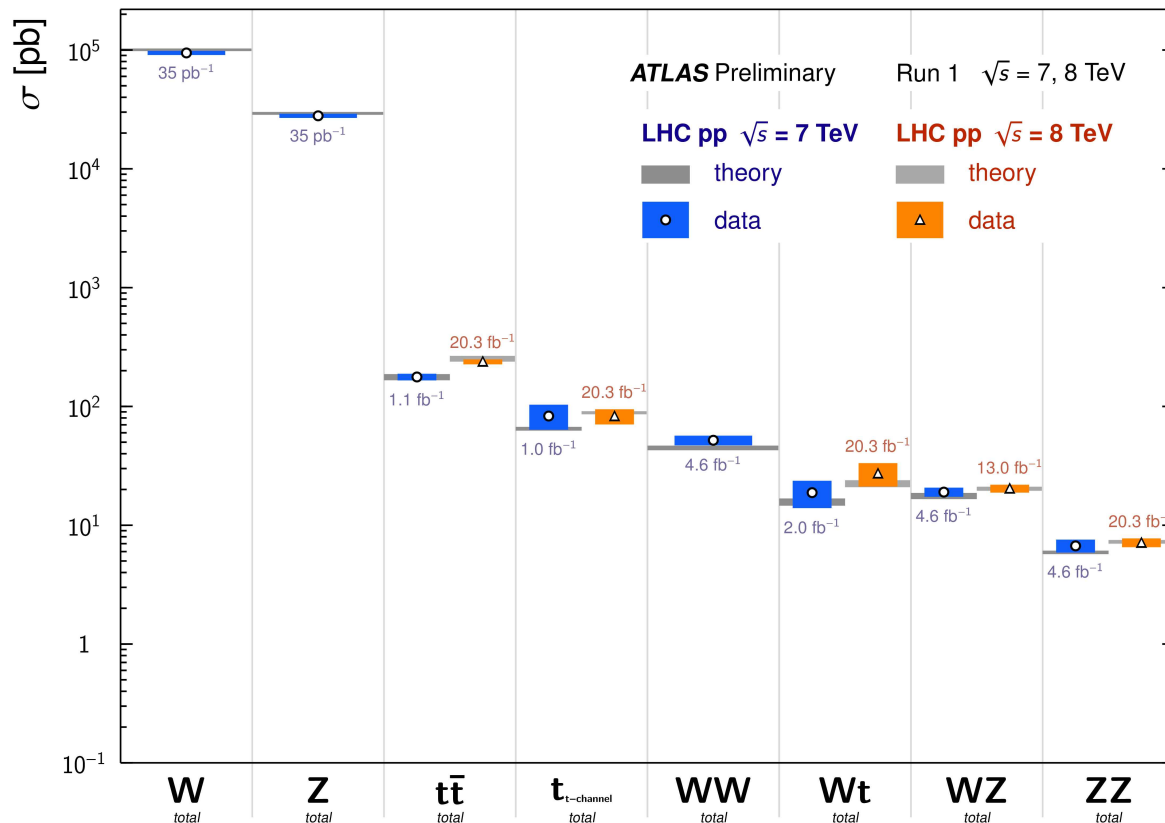
LHC and the ATLAS Experiment

► LHC performances from 2010 to 2012

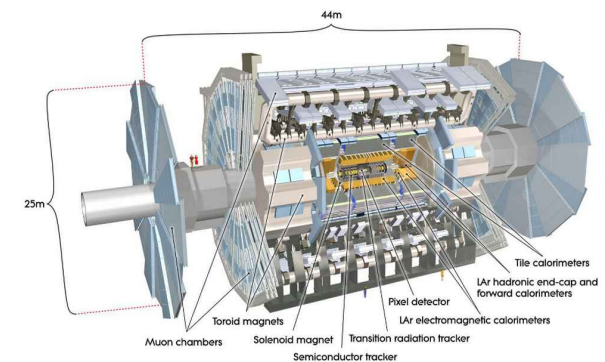
- pp run at 7 and 8 TeV
- Delivered 28 fb^{-1} ($>21 \text{ fb}^{-1}$ collected data in ATLAS)
- Restart in 2015 after technical stop with 13-14 TeV

Standard Model Total Production Cross Section Measurements

Status: March 2014



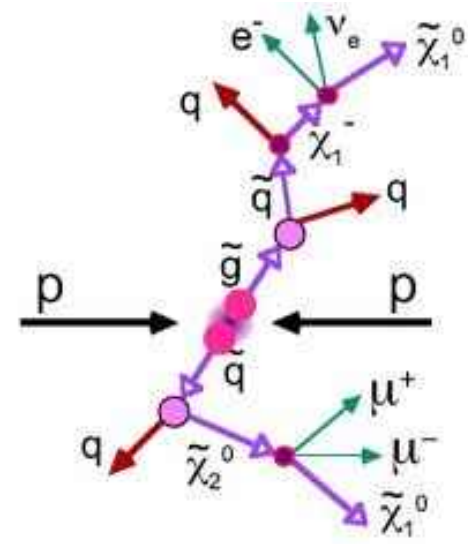
► Very good agreement between Standard Model theoretical and experimental cross sections



Search of Dark Matter in ATLAS

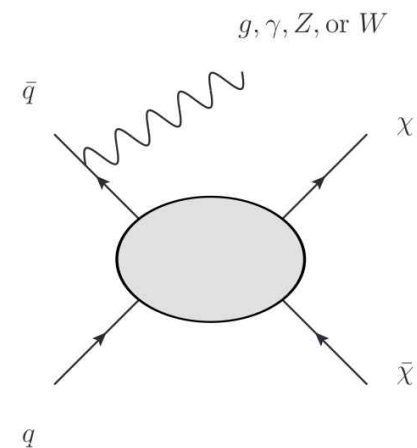
► SUSY searches

- Stability of DM ensured by imposing a symmetry that forbids the decay of DM into SM particle (R-parity conservation)
- Cascade ending with SM particles (jets, leptons) and LSP (missing E_t)
- Most widely studied candidates : lightest neutralino (spin 1/2), light gravitino (spin 3/2)



► Alternative searches

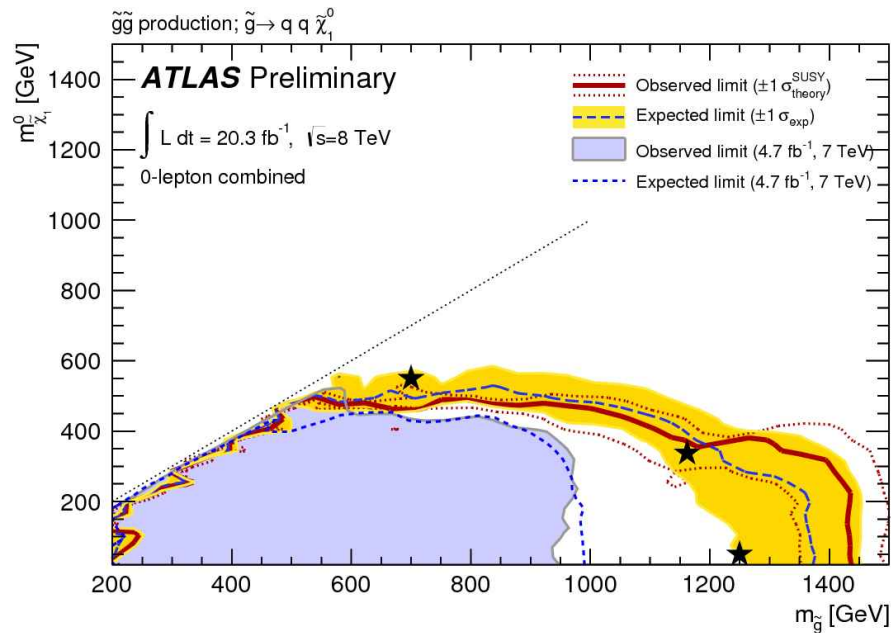
- Interaction SM-WIMP described by effective theory or simplified model
- WIMP pair production in association with an initial state radiation
 - Mono-jet : see talk from Valerio Rossetti
 - Mono-W/Z : see talk from Andy Nelson



Search for SUSY DM candidates in ATLAS

SUSY : strong production

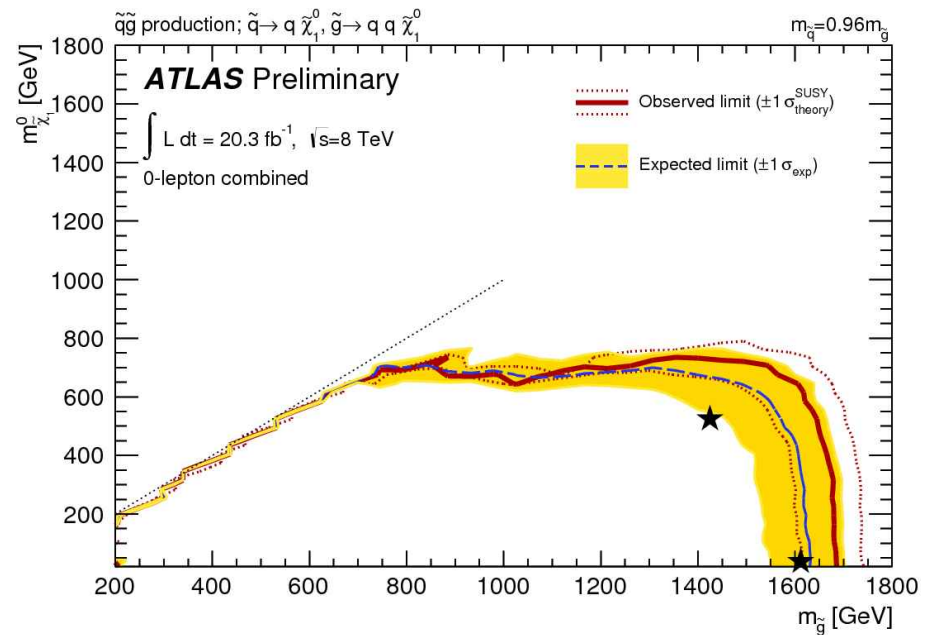
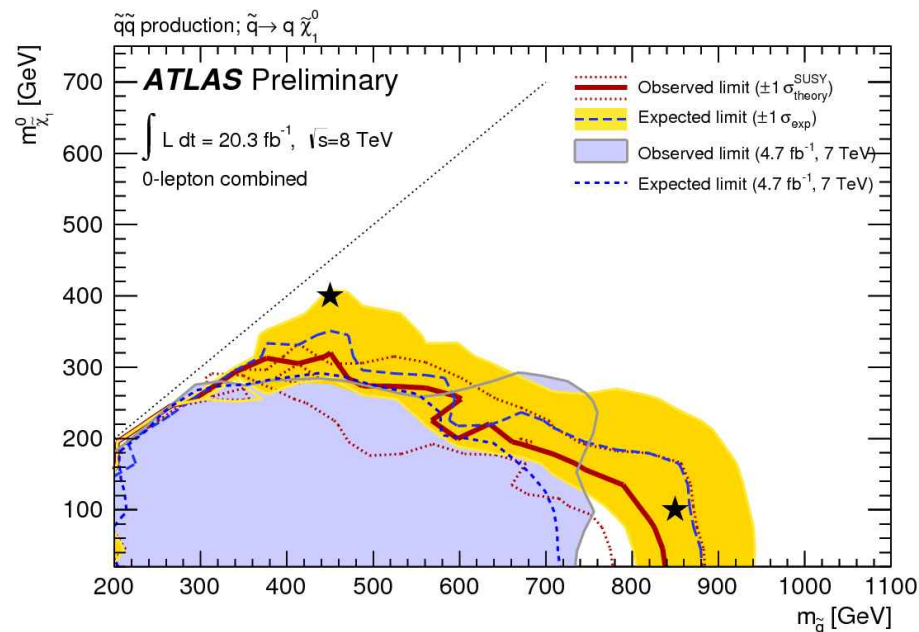
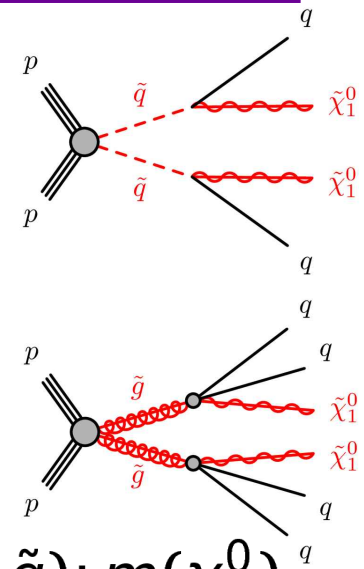
ATLAS-CONF-2012-147



► R-parity conservation, LSP lightest neutralino (DM candidate)

► Simplified model : 3rd squark generation, other neutralinos, sleptons at very high mass scale

► Exclusion in the plane $m(\tilde{q}, \tilde{g}); m(\chi_1^0)$

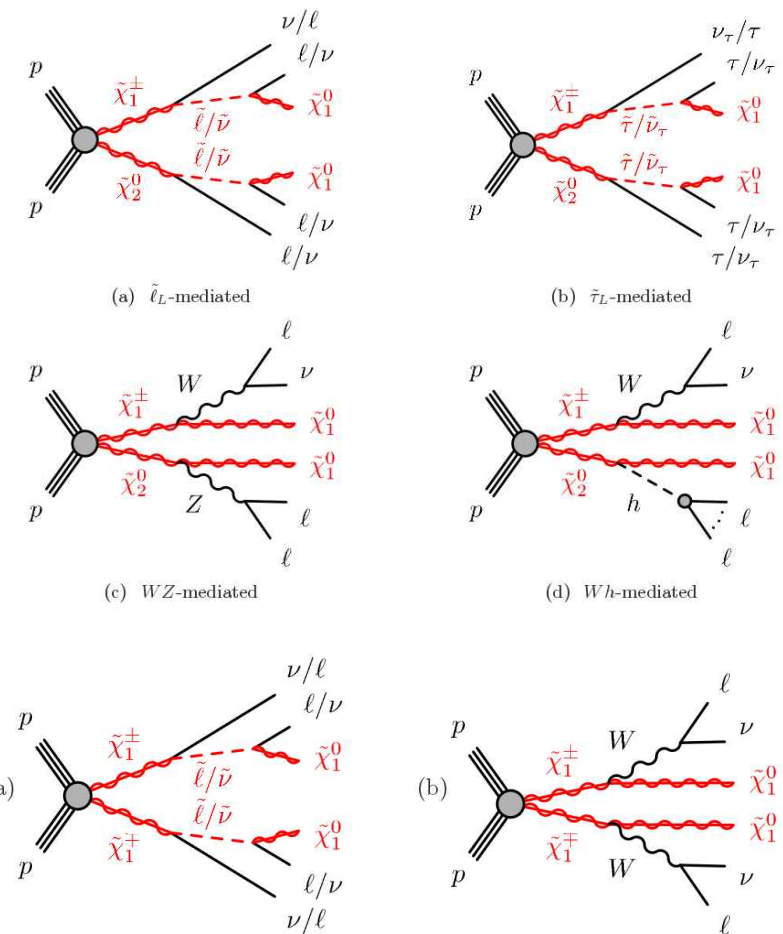
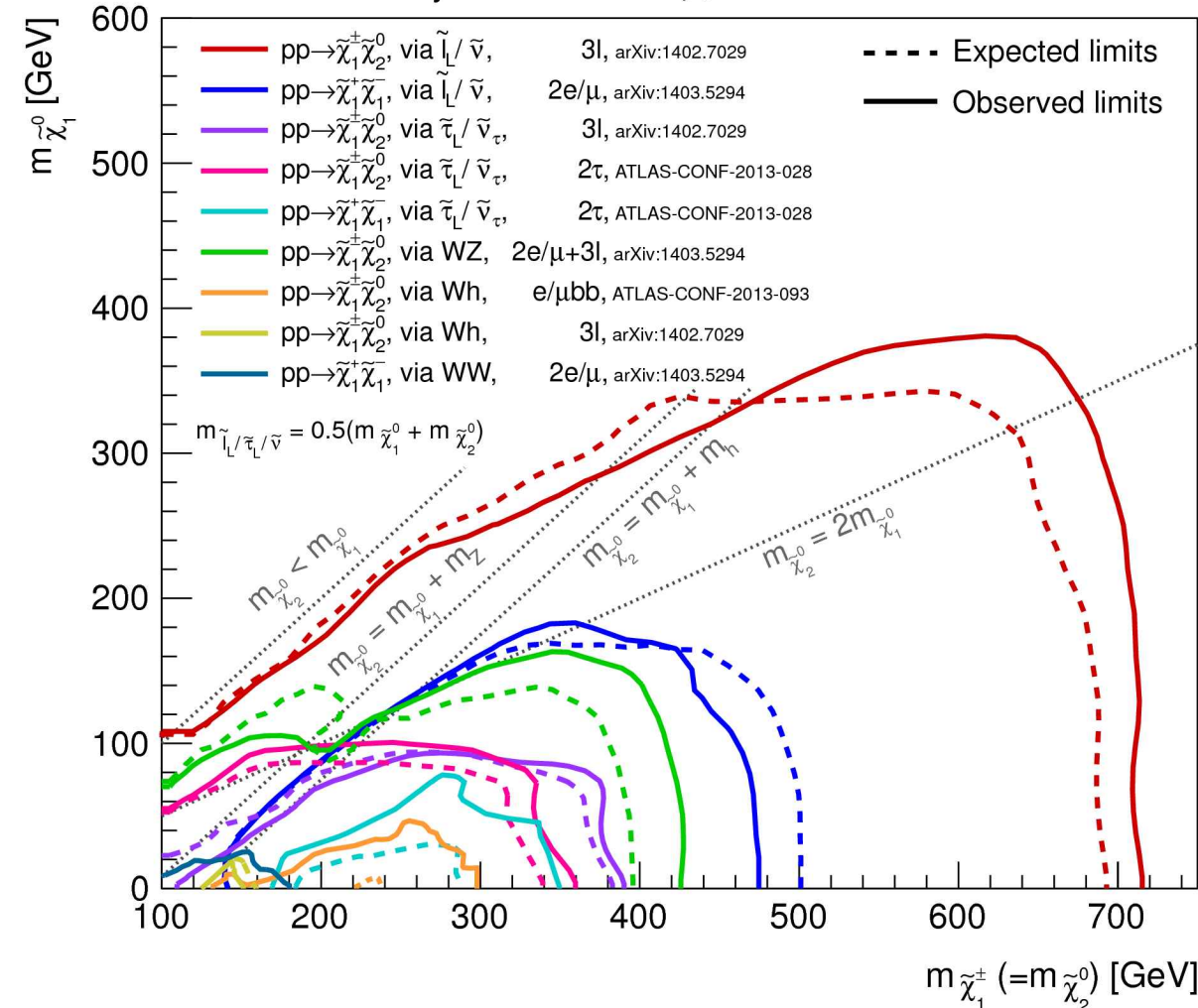


SUSY : electroweak production

- ▶ R-parity conservation, LSP lightest neutralino
- ▶ pMSSM, simplified model

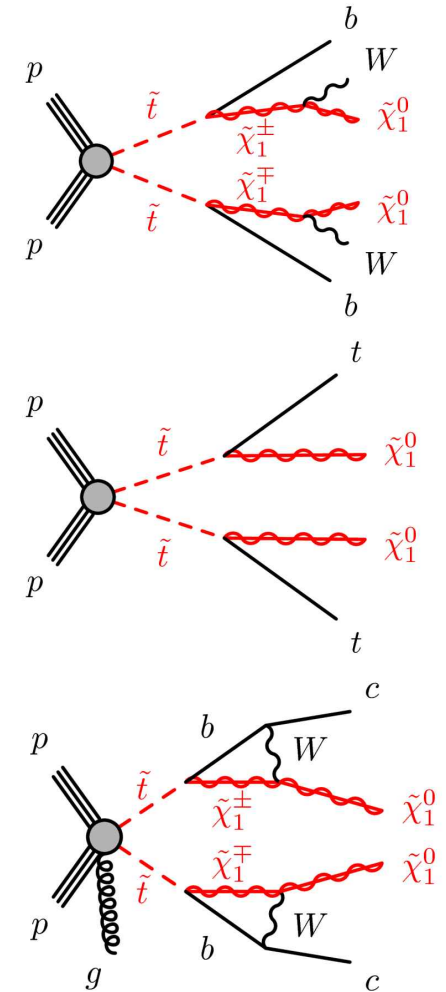
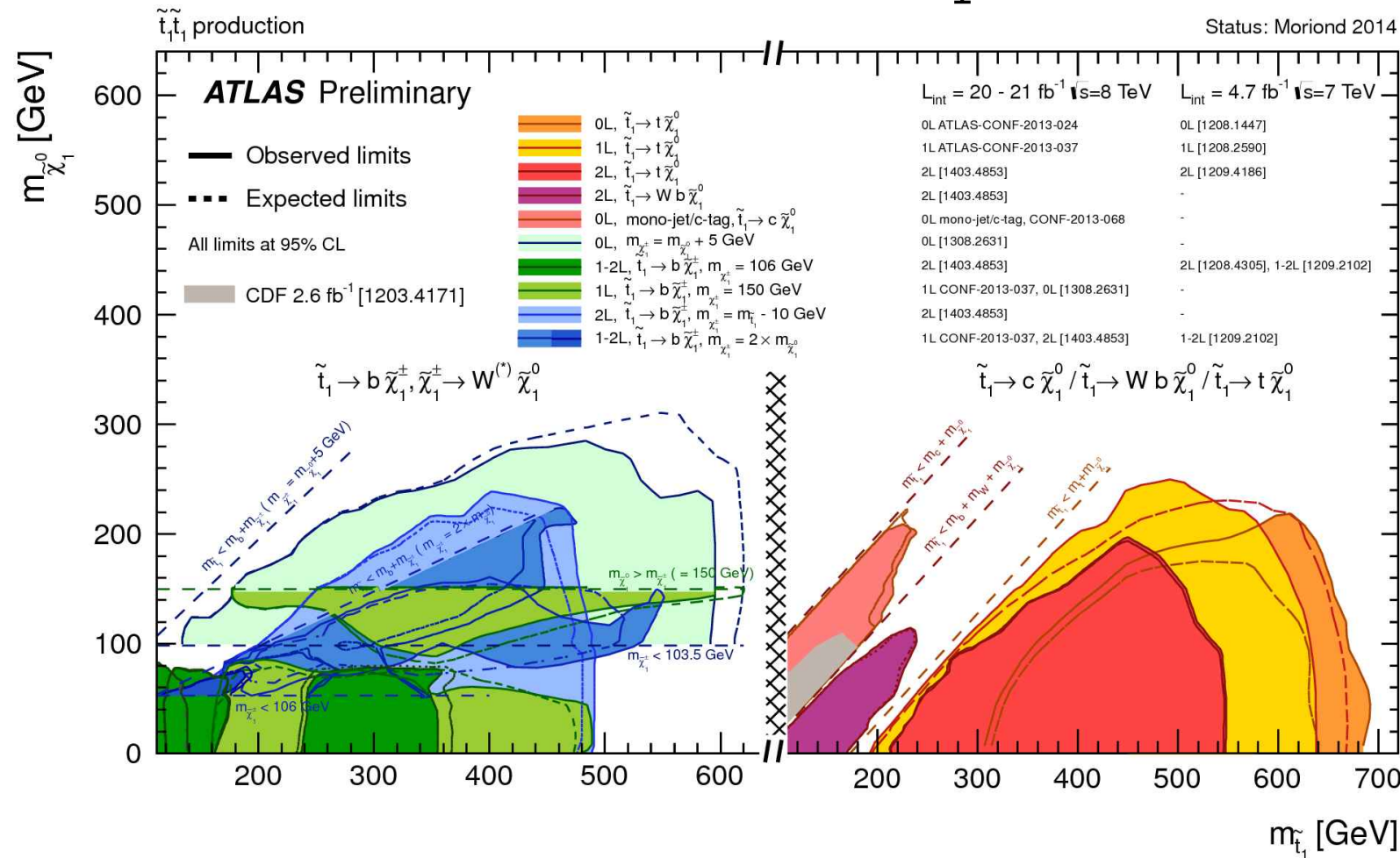
$$pp \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_1^\mp, \tilde{\chi}_2^0 \tilde{\chi}_1^\pm$$

ATLAS Preliminary 20.3-20.7 fb⁻¹, $\sqrt{s}=8$ TeV Status: Moriond 2014



SUSY : third generation production

- ▶ R-parity conservation, LSP = lightest neutralino
- ▶ Simplified models
- ▶ Exclusions in the plane $m(\tilde{t}); m(\chi_1^0)$

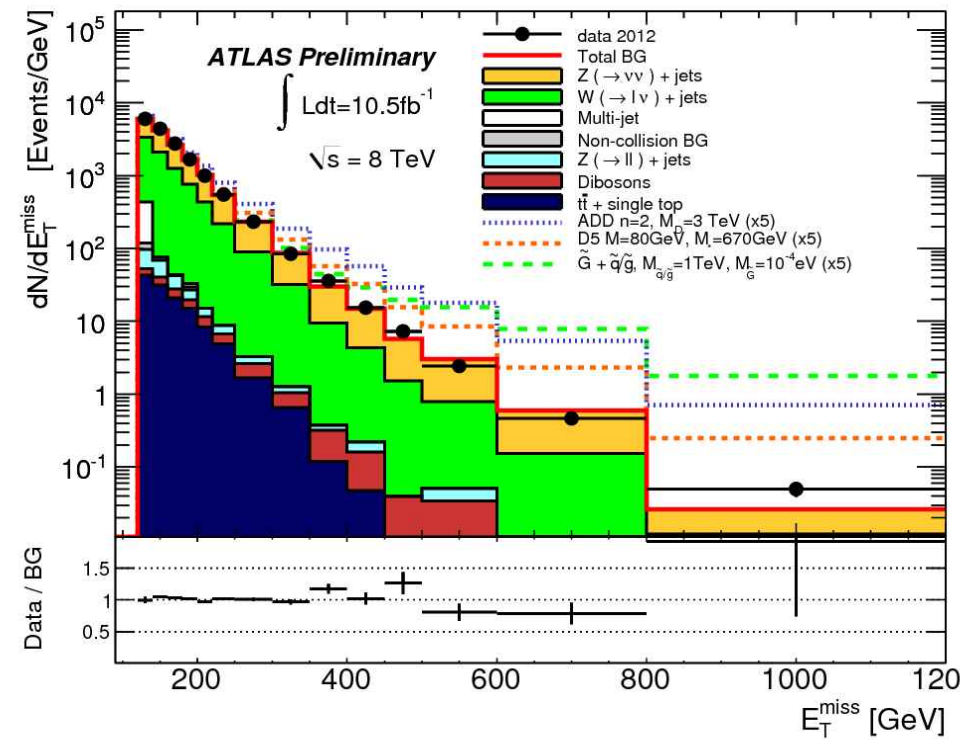


SUSY : summary of neutralino search

- ▶ Strong production
 - Neutralino (LSP) below 600 GeV excluded for squark/gravitino masses up to 1.6 TeV
- ▶ Electroweak and 3rd generation production
 - Neutralino (LSP) below 300 GeV excluded for chargino/stop masses up to ~700 GeV
- ▶ Simplified models (assume SUSY decays with 100% BR)
- ▶ Comparison with Direct DM detection experiments is non trivial
- ▶ In case an excess is seen by DM experiments, ATLAS SUSY results may help discriminating between models

SUSY searches : gravitino

ATLAS-CONF-2012-147



Events selection

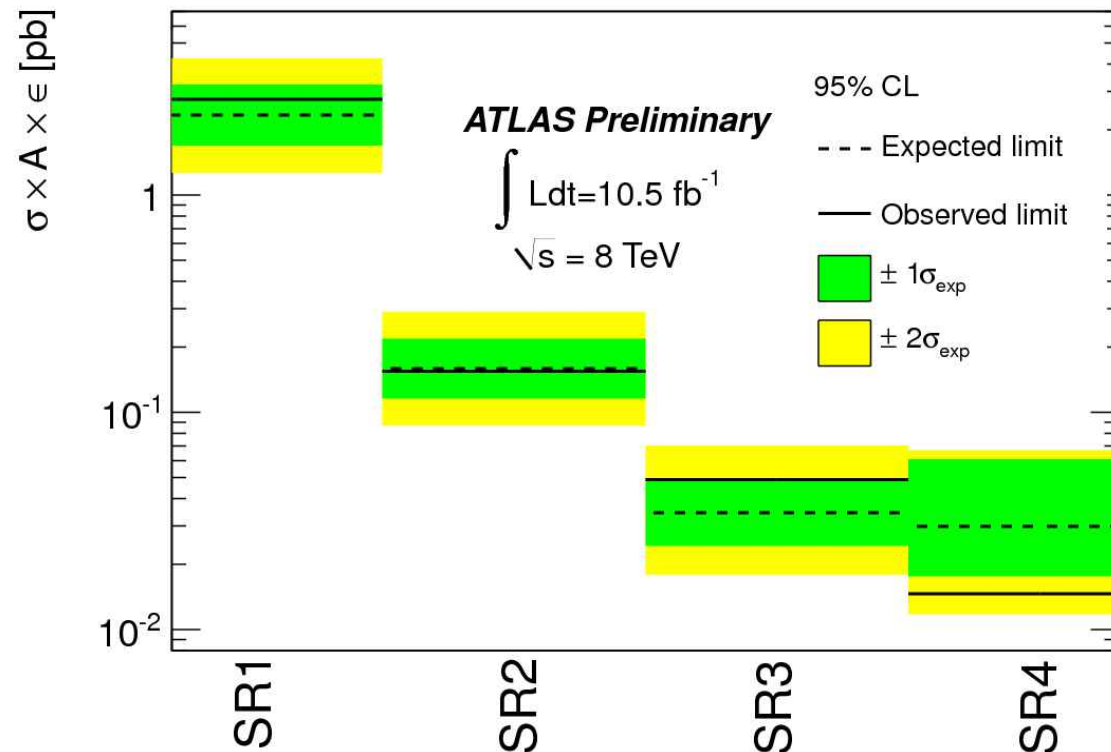
Lepton veto

$N_{\text{jets}} (p_T > 30 \text{ GeV}) < 3$

$E_T^{\text{miss}} > 120, 220, 350, 500 \text{ GeV}$

Leading jet $p_T > 120, 220, 350, 500 \text{ GeV}$

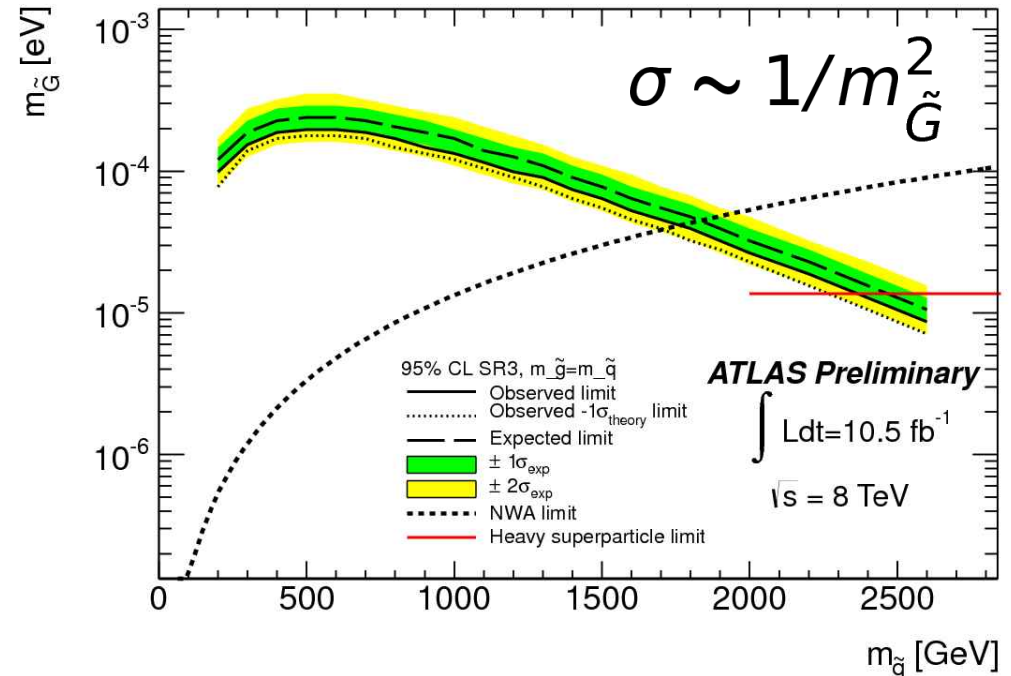
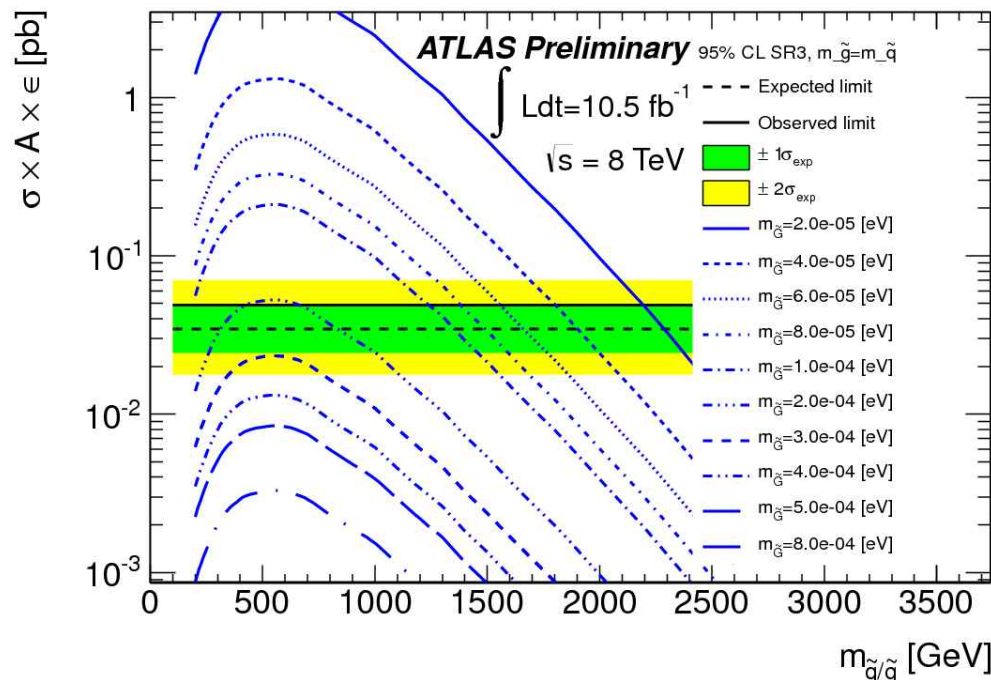
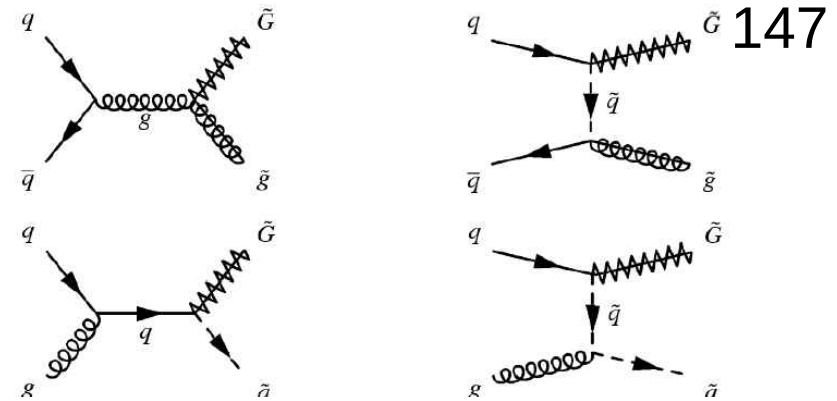
- Data driven determination of the dominant backgrounds W/Z+jets
- Good agreement with SM
 - Limits set on cross-section of physics Beyond the SM



SUSY searches : gravitino

ATLAS-CONF-2012-

- Interpretation in GMSB scenario
 - Gravitino LSP
 - Extremely heavy squark and gluino



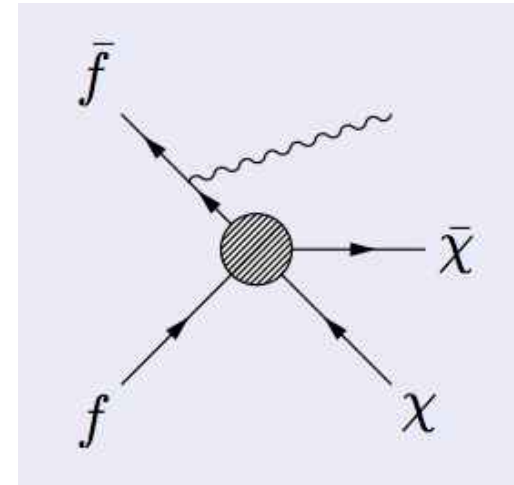
- Gravitino with $m < 10^{-4}$ eV (low $m(\tilde{q}, \tilde{g})$) to $m < 10^{-5}$ eV (high $m(\tilde{q}, \tilde{g})$) excluded at 95% CL

Alternative DM searches in ATLAS

Effective Field Theory (EFT) Phys.Rev.D82:116010,2010

► Effective Lagrangian

- Contact interaction
- Interaction mediated by a heavy particle of mass M with coupling g_1 to SM and g_2 to DM
- 2 parameters : m_χ , $M^* = M^2/g_1g_2$



► Limits : effective theory only valid for momentum transfer of the interaction $Q^2 \ll M^*$

► Different couplings SM-DM, with DM as Dirac fermion

- D1, D11, D5 spin-independent
- D8, D9 spin-dependent

Name	Initial state	Type	Operator
D1	qq	scalar	$\frac{m_q}{M_\star^3} \bar{\chi} \chi \bar{q} q$
D5	qq	vector	$\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \chi \bar{q} \gamma_\mu q$
D8	qq	axial-vector	$\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{q} \gamma_\mu \gamma^5 q$
D9	qq	tensor	$\frac{1}{M_\star^2} \bar{\chi} \sigma^{\mu\nu} \chi \bar{q} \sigma_{\mu\nu} q$
D11	gg	scalar	$\frac{1}{4M_\star^3} \bar{\chi} \chi \alpha_s (G_{\mu\nu}^a)^2$

Validity of the EFT

PhysRevD.85.056011

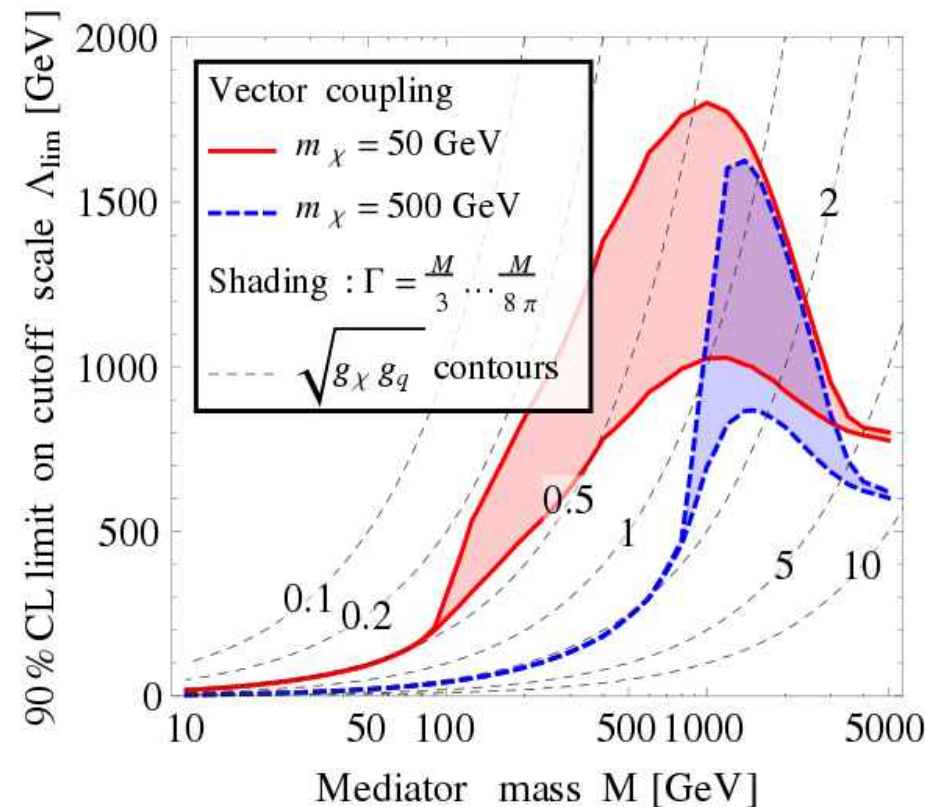
- ▶ LHC operating at high energy scales, where the validity of the Effective Field Theory may not be satisfied
- ▶ Investigation on how the predictions of the effective theory are modified once a propagating particle of mass M is introduced to mediate the interaction of SM and DM :

$$\sigma(pp \rightarrow \bar{\chi}\chi + X) \sim \frac{g_q^2 g_\chi^2}{(q^2 - M^2)^2 + \Gamma^2/4} E^2$$

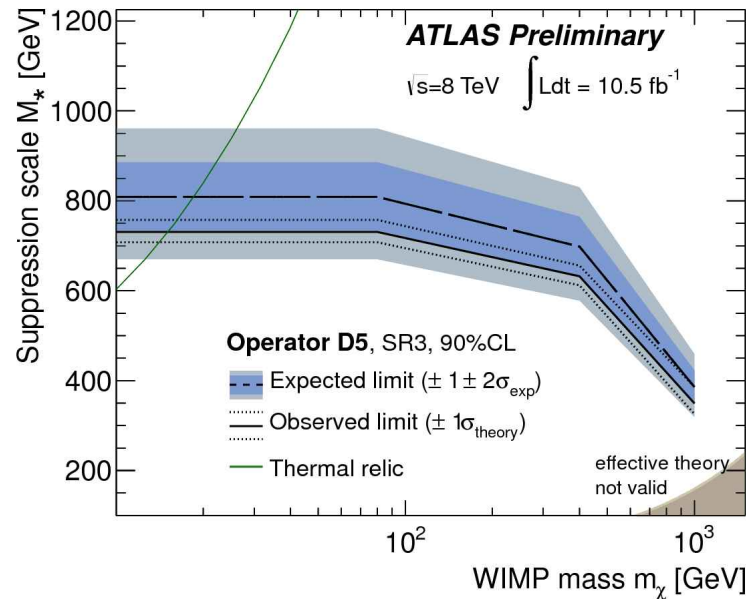
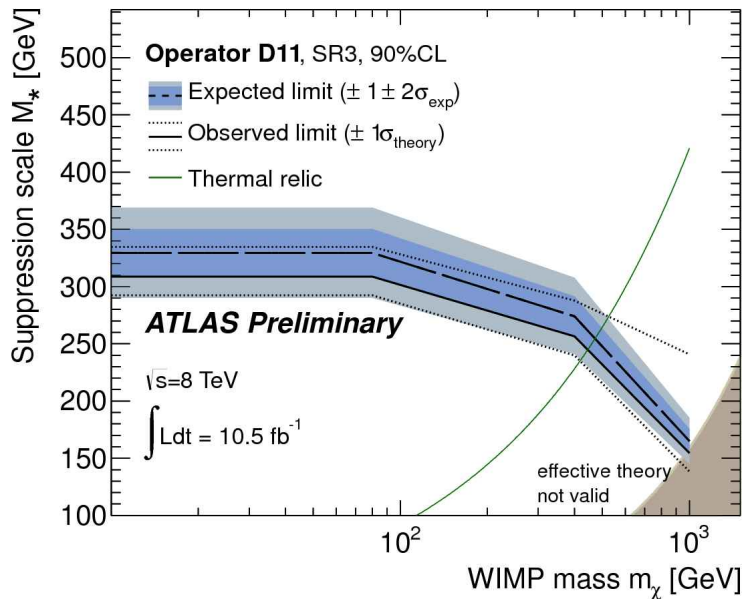
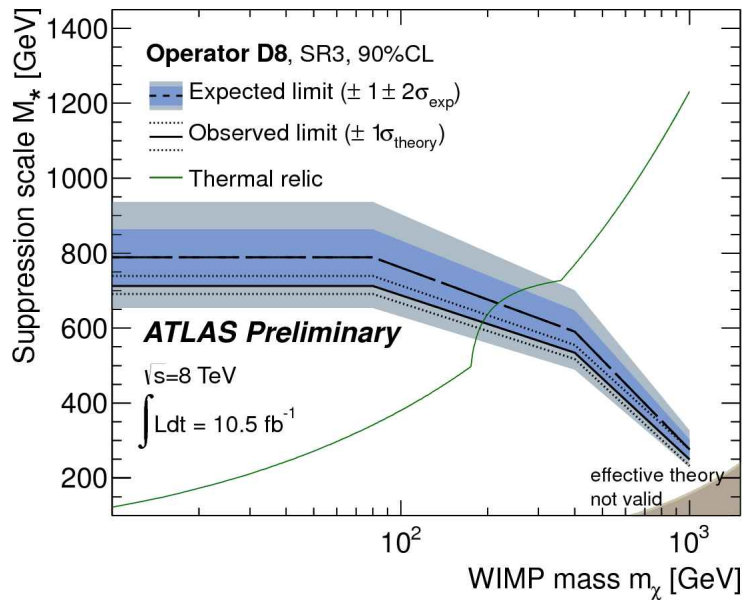
- ▶ To compare with EFT we define :

$$\Lambda = M / \sqrt{g_q g_\chi}$$

- ▶ At very large M (>5 TeV), limits on Λ asymptote to those of EFT
- ▶ For lighter mediator, ($M \sim \text{few} \times 100$ GeV) EFT limits are weaker due to resonant behavior of mediator
- ▶ For very light mediators, ($M < 100$ GeV), limits on direct detection cross sections are considerably weakened



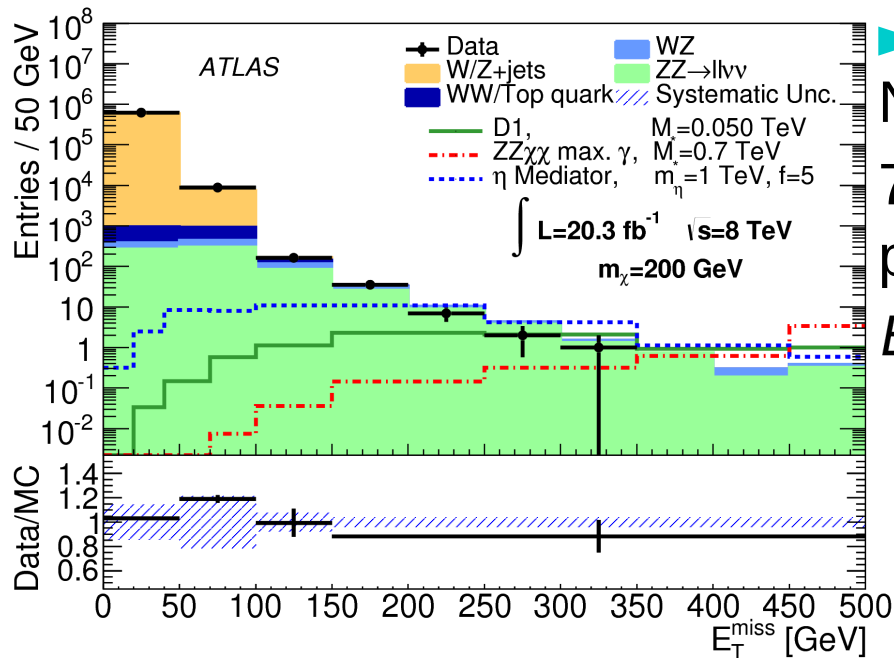
- ▶ Monojet selection : see slide #13
- ▶ 90% CL on the visible cross section for new physics translated into limits on M^* for Effective Theory (assuming its validity) with operators :
 - Operator D8 (axial vector) spin-dependent
 - Operator D5 (vector) spin-dependent
 - Operator D11 (scalar) spin-independent



▶ Green line : M^* and m_χ compatible with thermal relic abundance

Mono-Z($\rightarrow \ell\ell$)

CERN-PH-EP-2013-231



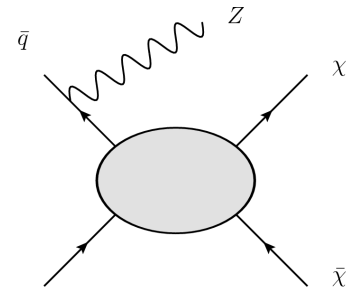
Events selection

No central jet with $p_T > 25 \text{ GeV}$

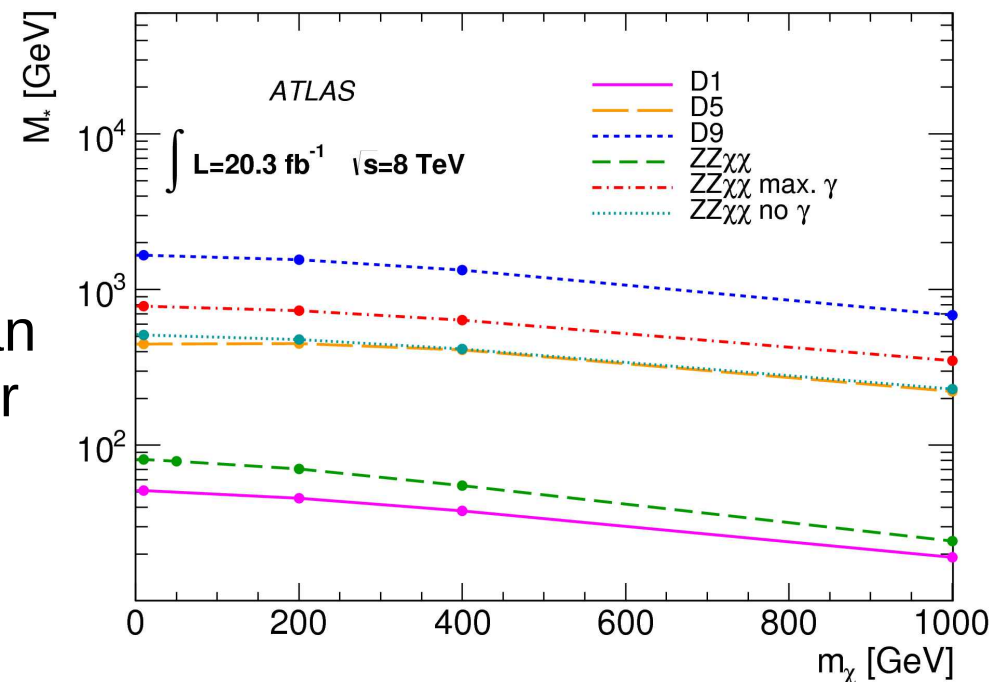
$76 < m(\ell\ell) < 106 \text{ GeV}$

$p_T(\text{lepton}) > 20 \text{ GeV}$

$E_T^{\text{miss}} > 150, 250, 350, 450 \text{ GeV}$



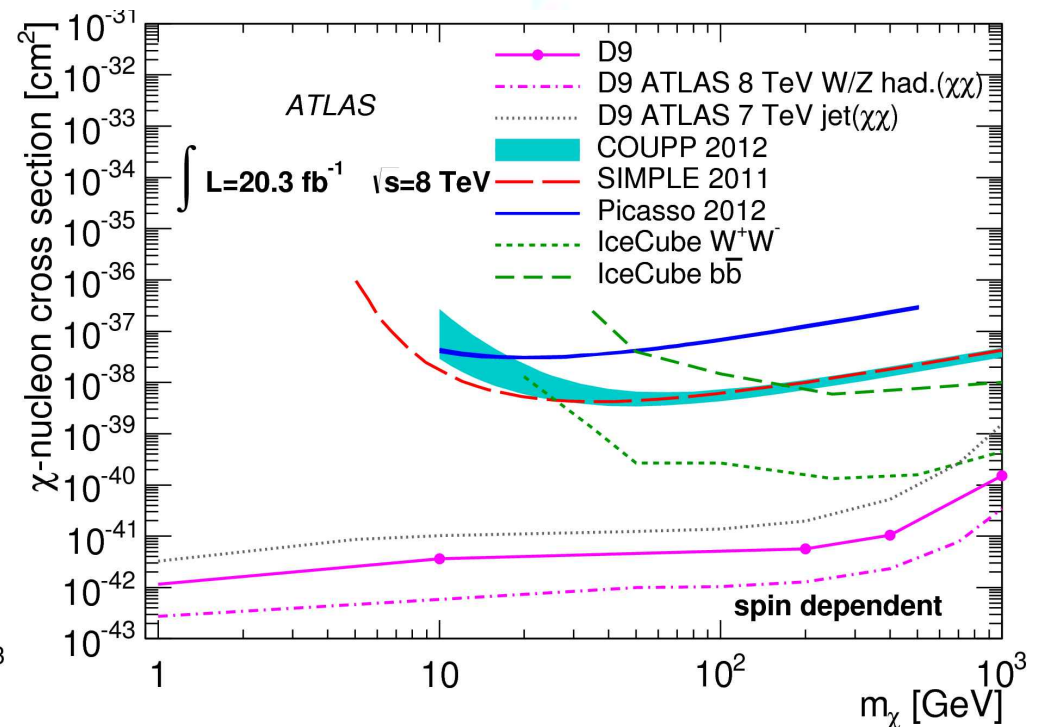
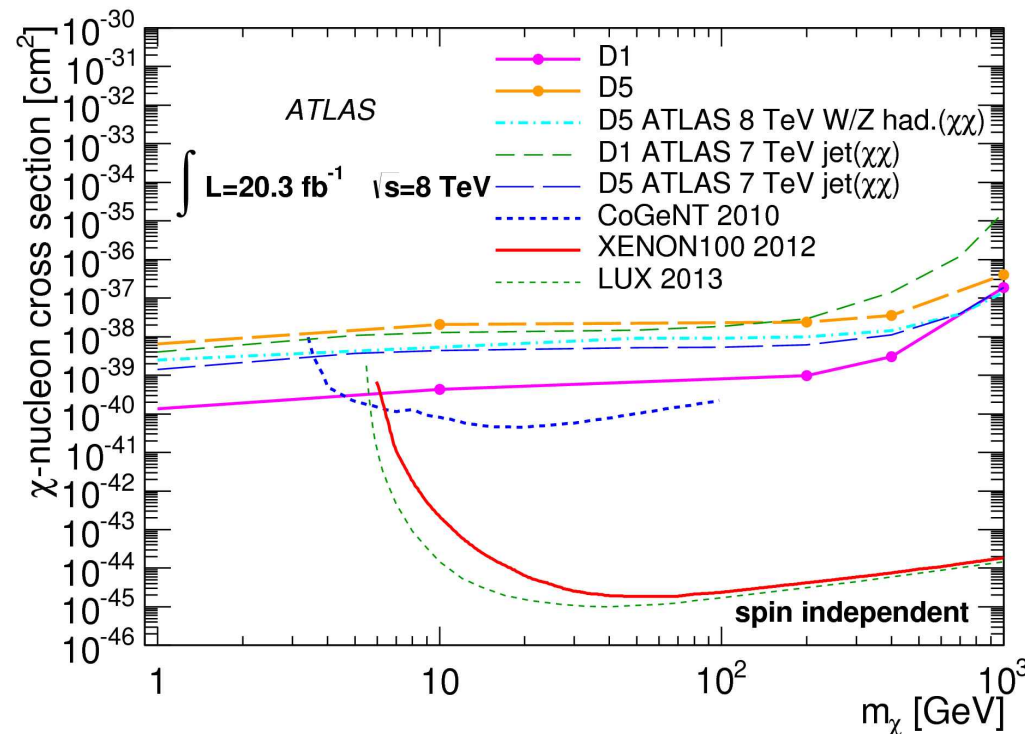
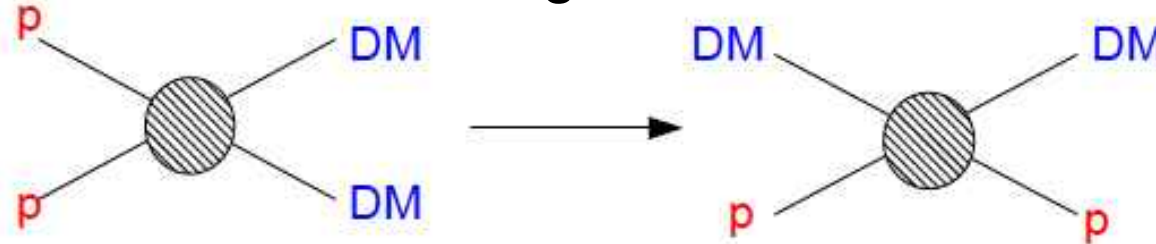
- ▶ No excess over the background is observed
- ▶ Interpretation DM effective lagrangian
- ▶ 95% CL limit on M^* for each operator



Mono-Z($\rightarrow \ell\ell$)

CERN-PH-EP-2013-231

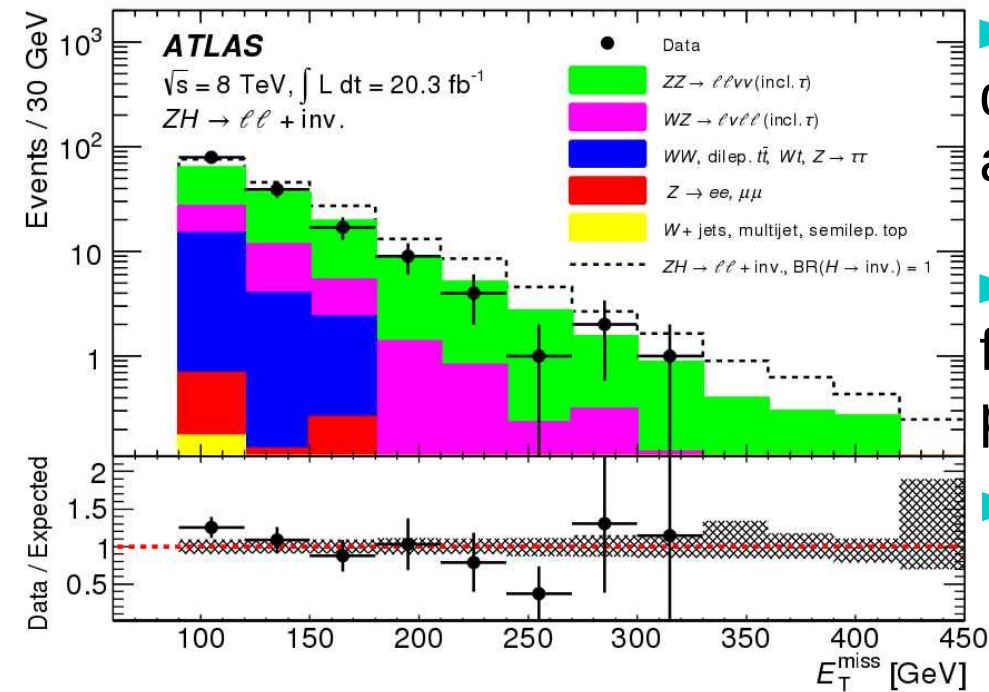
- 95% CL on nucleon scattering cross section



- Assuming the validity of the effective theory the results are competitive to direct detector experiments (particularly relevant at $m_\chi < 10$ GeV)

Mono-Z($\rightarrow \ell\ell$)

CERN-PH-EP-2013-210



► Search for invisible decay of Higgs boson in association with a Z

► Search for enhancements in the decay fraction to invisible particle due to new physics

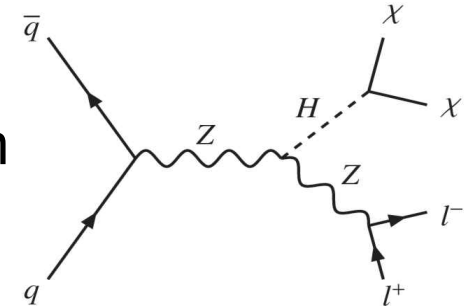
► Events selection

2 leptons

$p_T(\ell) > 20$ GeV

$76 < m(\ell\ell) < 106$ GeV

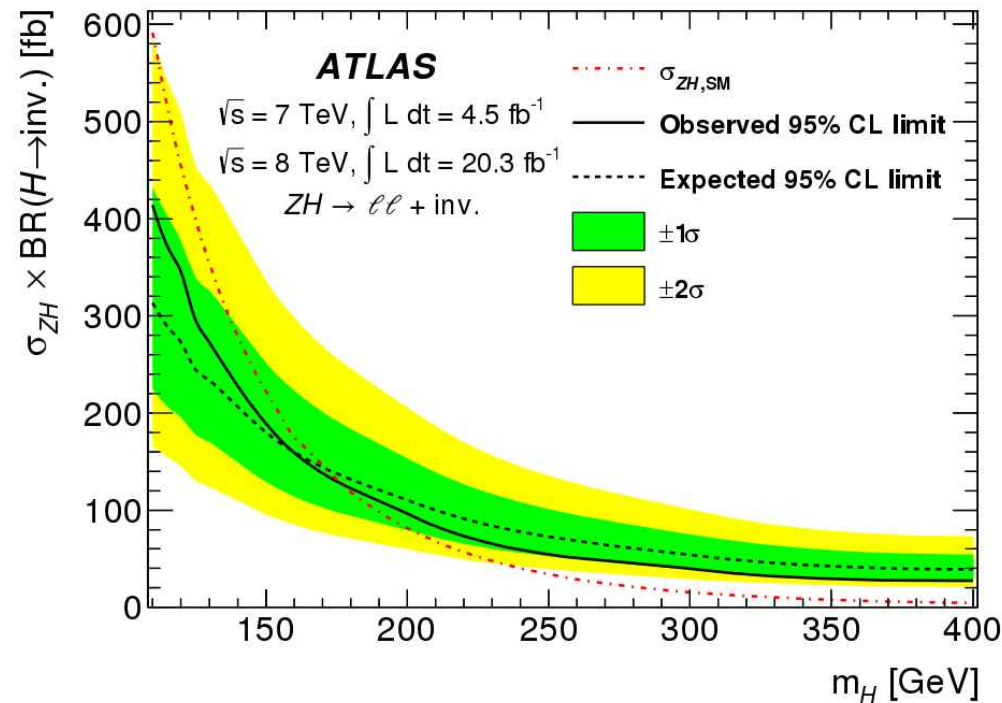
$E_T^{\text{miss}} > 90$ GeV



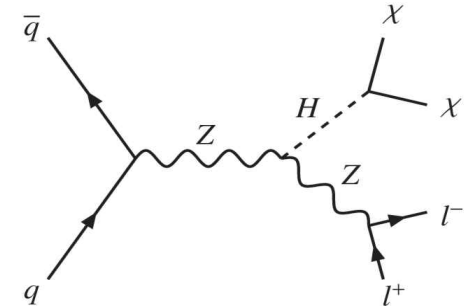
Data Period	2011 (7 TeV)	2012 (8 TeV)
$ZZ \rightarrow \ell\ell\nu\nu$	$20.0 \pm 0.7 \pm 1.6$	$91 \pm 1 \pm 7$
$WZ \rightarrow \ell\nu\ell\ell$	$4.8 \pm 0.3 \pm 0.5$	$26 \pm 1 \pm 3$
Dileptonic $t\bar{t}$, Wt , WW , $Z \rightarrow \tau\tau$	$0.5 \pm 0.4 \pm 0.1$	$20 \pm 3 \pm 5$
$Z \rightarrow ee$, $Z \rightarrow \mu\mu$	$0.13 \pm 0.12 \pm 0.07$	$0.9 \pm 0.3 \pm 0.5$
W + jets, multijet, semileptonic top	$0.020 \pm 0.005 \pm 0.008$	$0.29 \pm 0.02 \pm 0.06$
Total background	$25.4 \pm 0.8 \pm 1.7$	$138 \pm 4 \pm 9$
Signal ($m_H = 125.5$ GeV, $\sigma_{\text{SM}}(ZH)$, $\text{BR}(H \rightarrow \text{inv.}) = 1$)	$8.9 \pm 0.1 \pm 0.5$	$44 \pm 1 \pm 3$
Observed	28	152

Mono-Z($\rightarrow \ell\ell$)

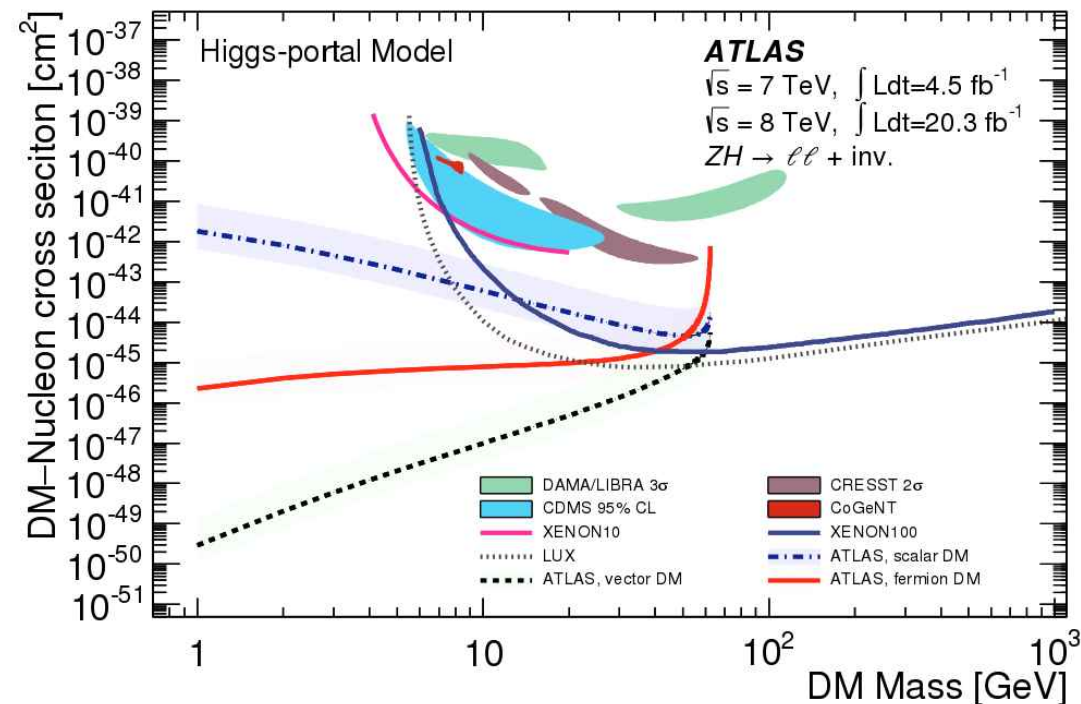
CERN-PH-EP-2013-210



► No significant excess above SM



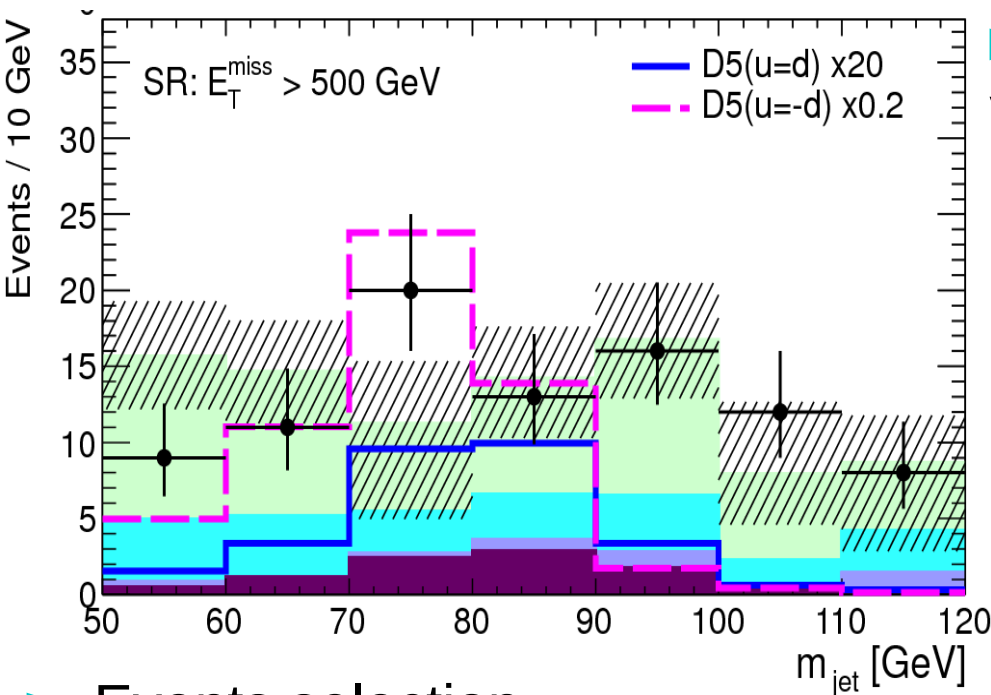
► Limit on DM-Nucleon cross section in Higgs-portal scenario [arXiv:hep-ph/0605188]



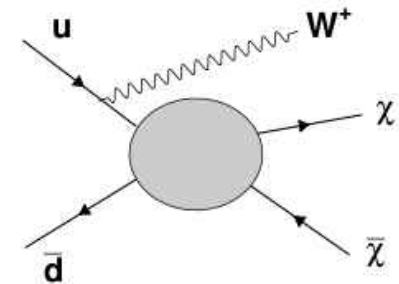
- For 125.5 GeV Higgs, Expected (without new physics)
 - $\text{BR}(H \rightarrow \text{inv.}) = 62\%$ (at 95% CL)
- Observed
 - $\text{BR}(H \rightarrow \text{inv.}) = 75\%$ (at 95% CL)

Mono-W($\rightarrow jj$)/Z($\rightarrow jj$)

Phys. Rev. Lett. 112, 041802



► WIMPs pair production via an unknown intermediate state, with initial-state radiation of W or Z, **hadronic decay**



► Use large-radius jets to capture the hadronic products of quarks from W or Z

► Events selection

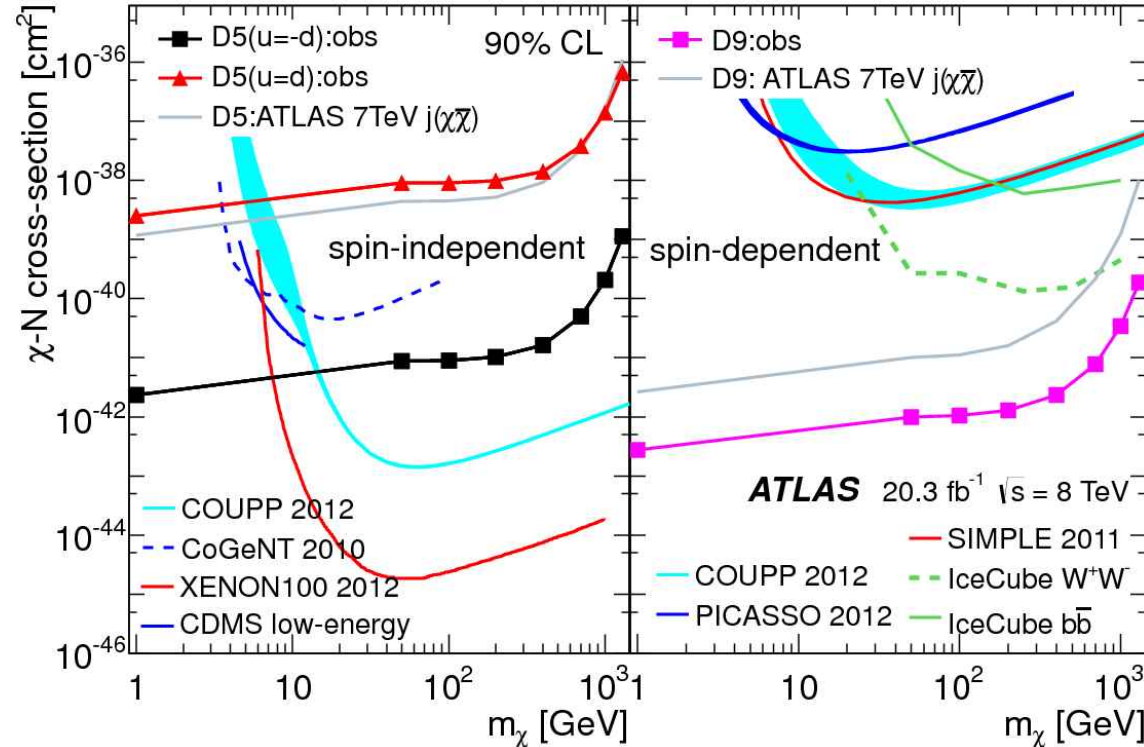
- 1 large-radius jet with $p_T > 250$ GeV
- $E_T^{\text{miss}} > 350, 500$ GeV
- No more than 1 narrow jet $p_T > 40$ GeV
- No isolated leptons

Process	$E_T^{\text{miss}} > 350$ GeV	$E_T^{\text{miss}} > 500$ GeV
$Z \rightarrow \nu\bar{\nu}$	402^{+39}_{-34}	54^{+8}_{-10}
$W \rightarrow \ell^\pm \nu, Z \rightarrow \ell^\pm \ell^\mp$	210^{+20}_{-18}	22^{+4}_{-5}
WW, WZ, ZZ	57^{+11}_{-8}	$9.1^{+1.3}_{-1.1}$
$t\bar{t}$, single t	39^{+10}_{-4}	$3.7^{+1.7}_{-1.3}$
Total	707^{+48}_{-38}	89^{+9}_{-12}
Data	705	89

► Good agreement with SM expectation. Limits set on physics BSM

Mono-W($\rightarrow jj$)/Z($\rightarrow jj$)

Phys. Rev. Lett. 112, 041802



► Limit on WIMP-Nucleon cross section with effective field theory

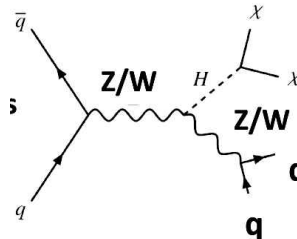
► Interpretation with invisible $H \rightarrow \chi\chi$

- For 125 GeV Higgs :

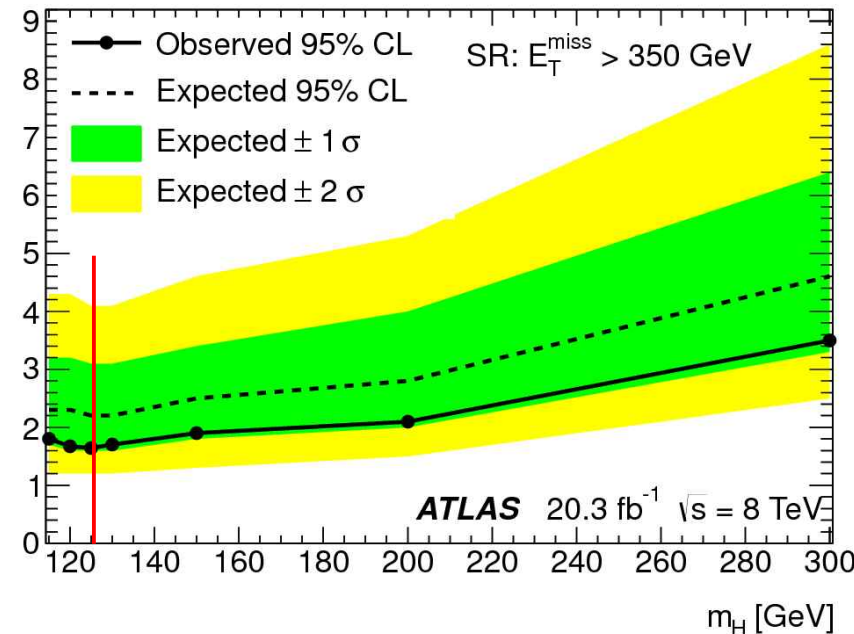
$$\sigma(WH, ZH, H \rightarrow \text{inv.})$$

$$< 1.6$$

$$\sigma(WH, ZH, \text{Standard model})$$

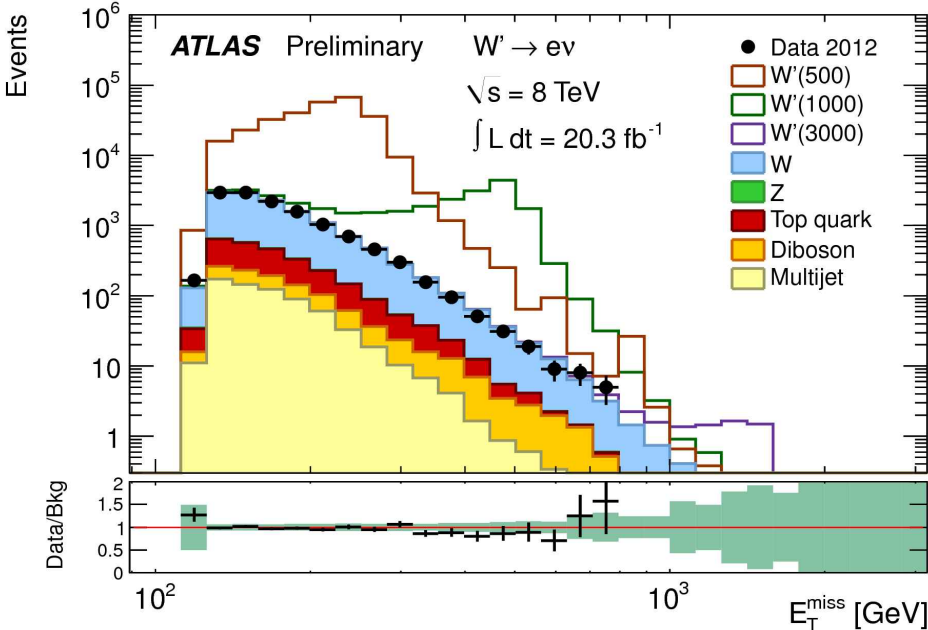


$$\sigma(W/Z \rightarrow W/Z \text{ inv}) / \sigma_{\text{total SM}}(W/Z H)$$



Mono $W(\rightarrow \ell \nu)/Z(\rightarrow \ell \ell)$

ATLAS-CONF-2014-017

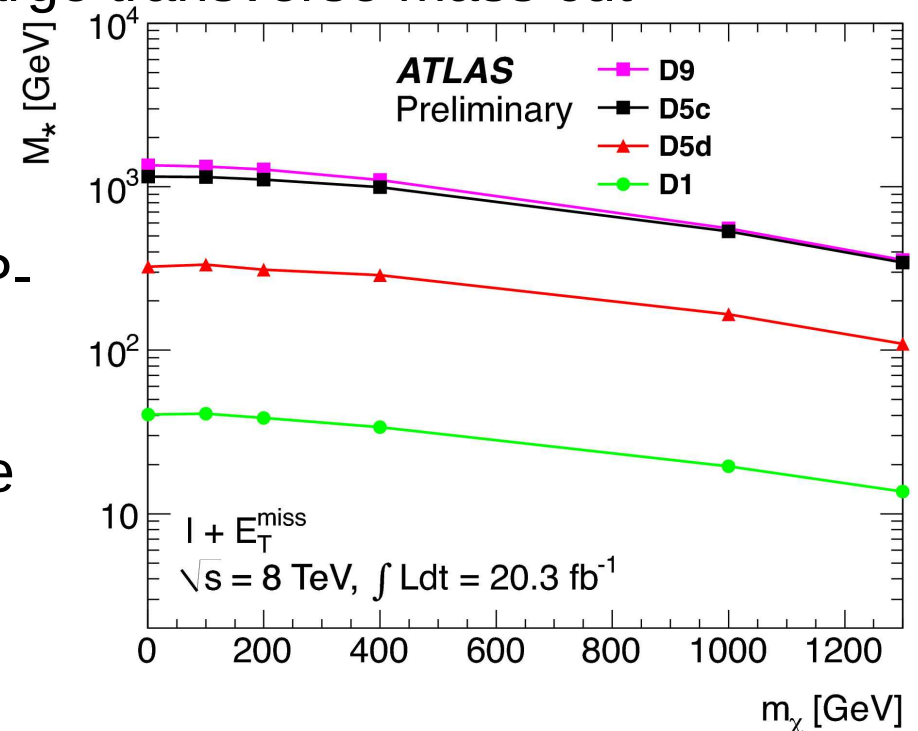


- ▶ No excess with respect to SM expectation
- ▶ Limits set on effective theory of WIMP-SM interaction
 - D9 tensor
 - D5 vector (constructive & destructive interference)
 - D1 scalar

▶ WIMPs pair production via an unknown intermediate state, with initial-state radiation of W or Z, **leptonic decay**

▶ Events selection

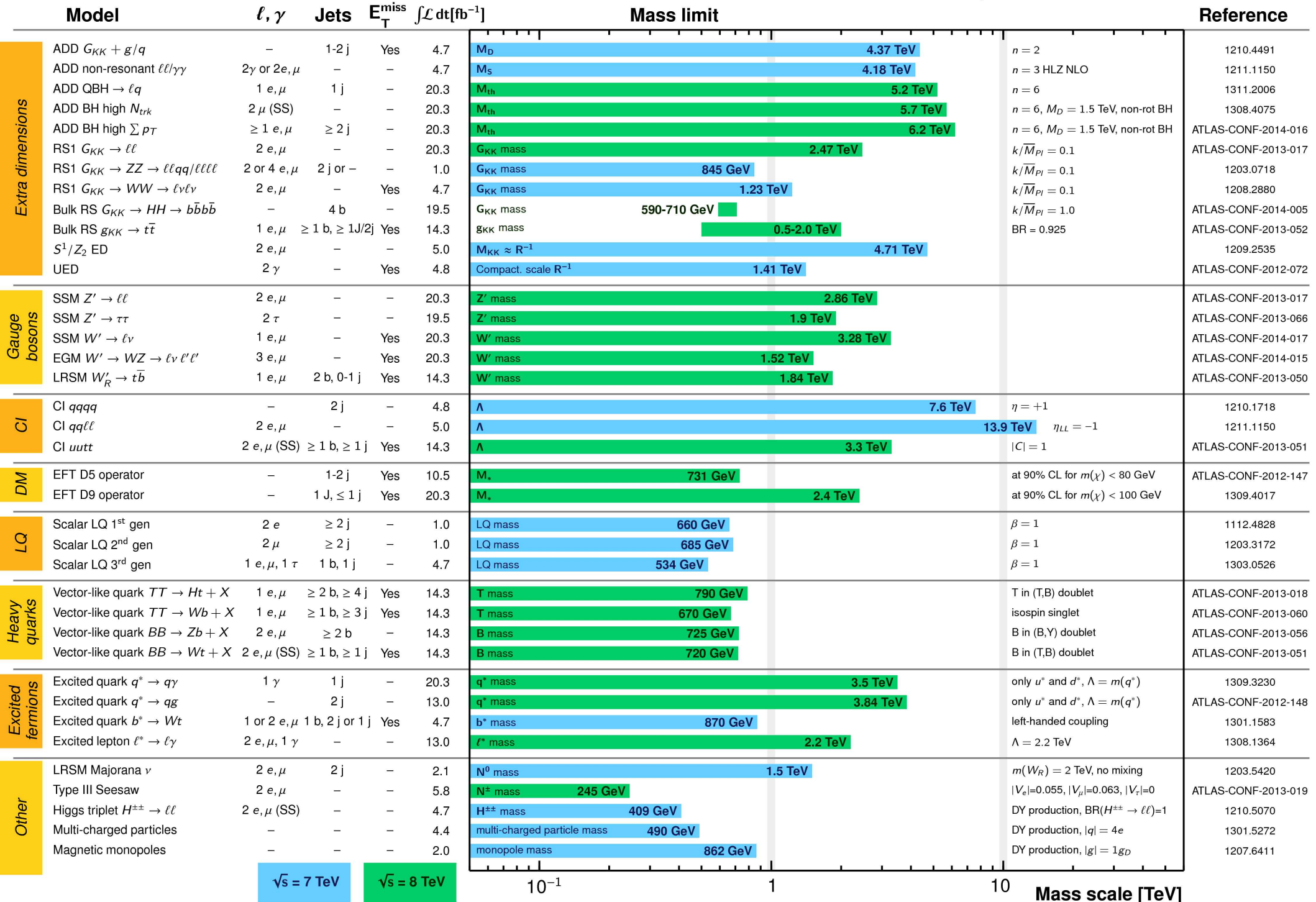
- At least 1 e ($p_T > 125$ GeV) or 1 μ ($p_T > 45$ GeV)
- $E_T^{\text{miss}} > 125$ GeV (e), 45 GeV (μ)
- Large transverse mass cut



Conclusion

- ▶ Very successful LHC operations during the last 3 years : more than 26 fb^{-1} of data on tape for ATLAS/CMS (7 TeV & 8 TeV)
- ▶ Large expected signal in ATLAS for many of the most interesting theories for Dark Matter.
- ▶ So far no significant excess indicating physics beyond the Standard Model
- ▶ ATLAS results put strong constraint on different SUSY models which may constrain/discriminate possible models for Dark Matter particles
- ▶ DM search within Effective Theory (model independent), simplified models, with comparison to direct DM search experiments is a very active topic
- ▶ LHC will restart soon, with higher energy. If DM is within reach of LHC
 - We should be able to detect it, and measure its mass and properties
 - We could be able to say something about the physics theory behind Dark Matter, possibly predict the right model
- ▶ Effort to share and discuss ATLAS results with astroparticle physics Communities : ATLAS Astro Forum (see backup)

Backup



*Only a selection of the available mass limits on new states or phenomena is shown.

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: Moriond 2014

ATLAS Preliminary

$$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \quad \sqrt{s} = 7, 8 \text{ TeV}$$

Model		e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit		Reference	
Inclusive Searches	MSUGRA/CMSSM	0	2-6 jets	Yes	20.3	\tilde{q}, \tilde{g}	1.7 TeV	$m(\tilde{q})=m(\tilde{g})$	ATLAS-CONF-2013-047
	MSUGRA/CMSSM	1 e, μ	3-6 jets	Yes	20.3	\tilde{g}	1.2 TeV	any $m(\tilde{q})$	ATLAS-CONF-2013-062
	MSUGRA/CMSSM	0	7-10 jets	Yes	20.3	\tilde{g}	1.1 TeV	any $m(\tilde{q})$	1308.1841
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	\tilde{q}	740 GeV	$m(\tilde{\chi}_1^0)=0$ GeV	ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	\tilde{g}	1.3 TeV	$m(\tilde{\chi}_1^0)=0$ GeV	ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^\pm \rightarrow q\tilde{q}W^\pm\tilde{\chi}_1^0$	1 e, μ	3-6 jets	Yes	20.3	\tilde{g}	1.18 TeV	$m(\tilde{\chi}_1^0)<200$ GeV, $m(\tilde{\chi}^\pm)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$	ATLAS-CONF-2013-062
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$	2 e, μ	0-3 jets	-	20.3	\tilde{g}	1.12 TeV	$m(\tilde{\chi}_1^0)=0$ GeV	ATLAS-CONF-2013-089
	GMSB ($\tilde{\ell}$ NLSP)	2 e, μ	2-4 jets	Yes	4.7	\tilde{g}	1.24 TeV	$\tan\beta<15$	1208.4688
	GMSB ($\tilde{\ell}$ NLSP)	1-2 τ	0-2 jets	Yes	20.7	\tilde{g}	1.4 TeV	$\tan\beta>18$	ATLAS-CONF-2013-026
	GGM (bino NLSP)	2 γ	-	Yes	20.3	\tilde{g}	1.28 TeV	$m(\tilde{\chi}_1^0)>50$ GeV	ATLAS-CONF-2014-001
	GGM (wino NLSP)	1 $e, \mu + \gamma$	-	Yes	4.8	\tilde{g}	619 GeV	$m(\tilde{\chi}_1^0)>50$ GeV	ATLAS-CONF-2012-144
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	4.8	\tilde{g}	900 GeV	$m(\tilde{\chi}_1^0)>220$ GeV	1211.1167
GGM (higgsino NLSP)	2 e, μ (Z)	0-3 jets	Yes	5.8	\tilde{g}	690 GeV	$m(\tilde{H})>200$ GeV	ATLAS-CONF-2012-152	
Gravitino LSP	0	mono-jet	Yes	10.5	$F^{1/2}$ scale	645 GeV	$m(\tilde{g})>10^{-4}$ eV	ATLAS-CONF-2012-147	
3 rd gen. \tilde{g} med.	$\tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	20.1	\tilde{g}	1.2 TeV	$m(\tilde{\chi}_1^0)<600$ GeV	ATLAS-CONF-2013-061
	$\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0	7-10 jets	Yes	20.3	\tilde{g}	1.1 TeV	$m(\tilde{\chi}_1^0)<350$ GeV	1308.1841
	$\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^\pm$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	1.34 TeV	$m(\tilde{\chi}_1^0)<400$ GeV	ATLAS-CONF-2013-061
	$\tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^\pm$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	1.3 TeV	$m(\tilde{\chi}_1^0)<300$ GeV	ATLAS-CONF-2013-061
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	20.1	\tilde{b}_1	100-620 GeV	$m(\tilde{\chi}_1^0)<90$ GeV	1308.2631
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$	2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{b}_1	275-430 GeV	$m(\tilde{\chi}_1^\pm)=2 m(\tilde{\chi}_1^0)$	ATLAS-CONF-2013-007
	$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	1-2 e, μ	1-2 b	Yes	4.7	\tilde{t}_1	110-167 GeV	$m(\tilde{\chi}_1^0)=55$ GeV	1208.4305, 1209.2102
	$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$	2 e, μ	0-2 jets	Yes	20.3	\tilde{t}_1	130-210 GeV	$m(\tilde{\chi}_1^0)=m(\tilde{t}_1)-m(W)-50$ GeV, $m(\tilde{t}_1)<m(\tilde{\chi}_1^\pm)$	1403.4853
	$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$	2 e, μ	2 jets	Yes	20.3	\tilde{t}_1	215-530 GeV	$m(\tilde{\chi}_1^0)=1$ GeV	1403.4853
	$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	0	2 b	Yes	20.1	\tilde{t}_1	150-580 GeV	$m(\tilde{\chi}_1^0)<200$ GeV, $m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)=5$ GeV	1308.2631
	$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$	1 e, μ	1 b	Yes	20.7	\tilde{t}_1	200-610 GeV	$m(\tilde{\chi}_1^0)=0$ GeV	ATLAS-CONF-2013-037
	$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^\pm$	0	2 b	Yes	20.5	\tilde{t}_1	320-660 GeV	$m(\tilde{\chi}_1^0)=0$ GeV	ATLAS-CONF-2013-024
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet/c-tag	Yes	20.3	\tilde{t}_1	90-200 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)<85$ GeV	ATLAS-CONF-2013-068
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_1	150-580 GeV	$m(\tilde{\chi}_1^0)>150$ GeV	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_2	290-600 GeV	$m(\tilde{\chi}_1^0)<200$ GeV	1403.5222
	EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ	0	Yes	20.3	$\tilde{\ell}$	90-325 GeV	$m(\tilde{\chi}_1^0)=0$ GeV
$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\nu}(\ell\tilde{\nu})$		2 e, μ	0	Yes	20.3	$\tilde{\chi}_1^\pm$	140-465 GeV	$m(\tilde{\chi}_1^0)=0$ GeV, $m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1403.5294
$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\tau}(\nu\tilde{\tau})$		2 τ	-	Yes	20.7	$\tilde{\chi}_1^\pm$	180-330 GeV	$m(\tilde{\chi}_1^0)=0$ GeV, $m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	ATLAS-CONF-2013-028
$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_{L,R}\tilde{\ell}_{L,R}(\tilde{\nu}\tilde{\nu}), \tilde{\ell}\tilde{\nu}\tilde{\ell}_{L,R}(\tilde{\nu}\tilde{\nu})$		3 e, μ	0	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	700 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1402.7029
$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$		2-3 e, μ	0	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	420 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0$, sleptons decoupled	1403.5294, 1402.7029
$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0$		1 e, μ	2 b	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	285 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0$, sleptons decoupled	ATLAS-CONF-2013-093
Long-lived particles	Direct $\tilde{\chi}_1^+\tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	20.3	$\tilde{\chi}_1^\pm$	270 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)=160$ MeV, $\tau(\tilde{\chi}_1^\pm)=0.2$ ns	ATLAS-CONF-2013-069
	Stable, stopped \tilde{g} R-hadron	0	1-5 jets	Yes	22.9	\tilde{g}	832 GeV	$m(\tilde{\chi}_1^0)=100$ GeV, $10 \mu\text{s}<\tau(\tilde{g})<1000$ s	ATLAS-CONF-2013-057
	GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$	1-2 μ	-	-	15.9	$\tilde{\chi}_1^0$	475 GeV	$10<\tan\beta<50$	ATLAS-CONF-2013-058
	GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$	2 γ	-	Yes	4.7	$\tilde{\chi}_1^0$	230 GeV	$0.4<\tau(\tilde{\chi}_1^0)<2$ ns	1304.6310
	$\tilde{q}\tilde{q}, \tilde{\chi}_1^0 \rightarrow q\tilde{q}\mu$ (RPV)	1 μ , displ. vtx	-	-	20.3	\tilde{q}	1.0 TeV	$1.5<c\tau<156$ mm, $\text{BR}(\mu)=1, m(\tilde{\chi}_1^0)=108$ GeV	ATLAS-CONF-2013-092
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e + \mu$	2 e, μ	-	-	4.6	$\tilde{\nu}_\tau$	1.61 TeV	$\lambda'_{311}=0.10, \lambda_{132}=0.05$	1212.1272
	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e(\mu) + \tau$	1 $e, \mu + \tau$	-	-	4.6	$\tilde{\nu}_\tau$	1.1 TeV	$\lambda'_{311}=0.10, \lambda_{1(2)33}=0.05$	1212.1272
	Bilinear RPV CMSSM	1 e, μ	7 jets	Yes	4.7	\tilde{q}, \tilde{g}	1.2 TeV	$m(\tilde{q})=m(\tilde{g}), c\tau_{LSP}<1$ mm	ATLAS-CONF-2012-140
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow ee\tilde{\nu}_\mu, e\mu\tilde{\nu}_e$	4 e, μ	-	Yes	20.7	$\tilde{\chi}_1^\pm$	760 GeV	$m(\tilde{\chi}_1^0)>300$ GeV, $\lambda_{121}>0$	ATLAS-CONF-2013-036
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tilde{\nu}_e, e\tau\tilde{\nu}_\tau$	3 $e, \mu + \tau$	-	Yes	20.7	$\tilde{\chi}_1^\pm$	350 GeV	$m(\tilde{\chi}_1^0)>80$ GeV, $\lambda_{133}>0$	ATLAS-CONF-2013-036
	$\tilde{g} \rightarrow q\tilde{q}q$	0	6-7 jets	-	20.3	\tilde{g}	916 GeV	$\text{BR}(h)=\text{BR}(b)=\text{BR}(c)=0\%$	ATLAS-CONF-2013-091
	$\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$	2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{g}	880 GeV		ATLAS-CONF-2013-007
Other	Scalar gluon pair, sgluon $\rightarrow q\tilde{q}$	0	4 jets	-	4.6	sgluon	100-287 GeV	incl. limit from 1110.2693	1210.4826
	Scalar gluon pair, sgluon $\rightarrow t\tilde{t}$	2 e, μ (SS)	2 b	Yes	14.3	sgluon	350-800 GeV		ATLAS-CONF-2013-051
	WIMP interaction (D5, Dirac χ)	0	mono-jet	Yes	10.5	M^* scale	704 GeV	$m(\chi)<80$ GeV, limit of <687 GeV for D8	ATLAS-CONF-2012-147

$\sqrt{s} = 7$ TeV
full data

$\sqrt{s} = 8$ TeV
partial data

$\sqrt{s} = 8$ TeV
full data

10^{-1}

1

Mass scale [TeV]

$\sqrt{s} = 7$ TeV full data
 $\sqrt{s} = 8$ TeV partial data
 $\sqrt{s} = 8$ TeV full data

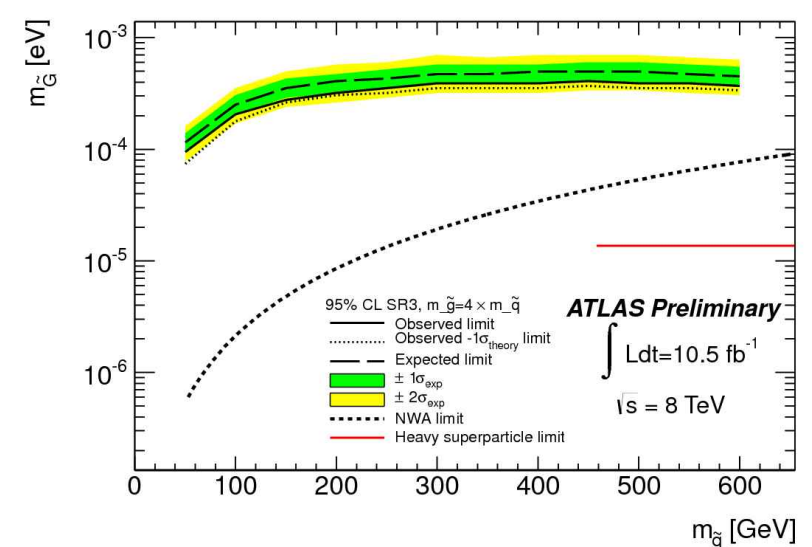
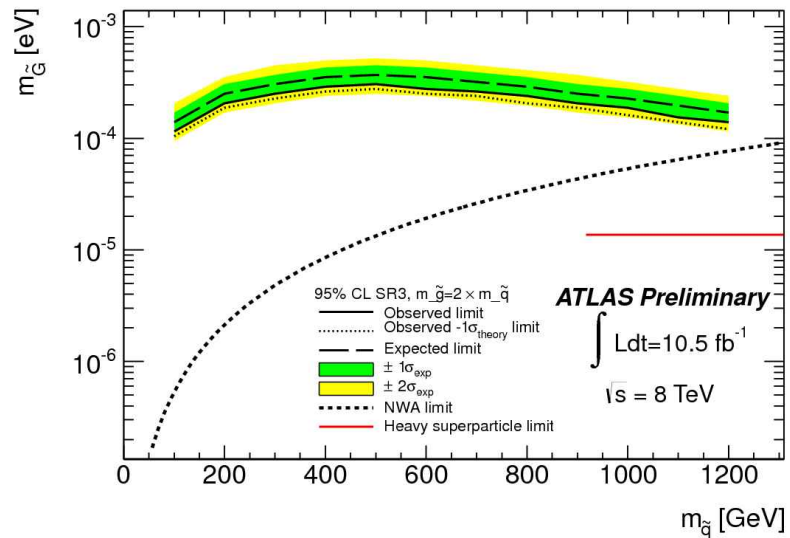
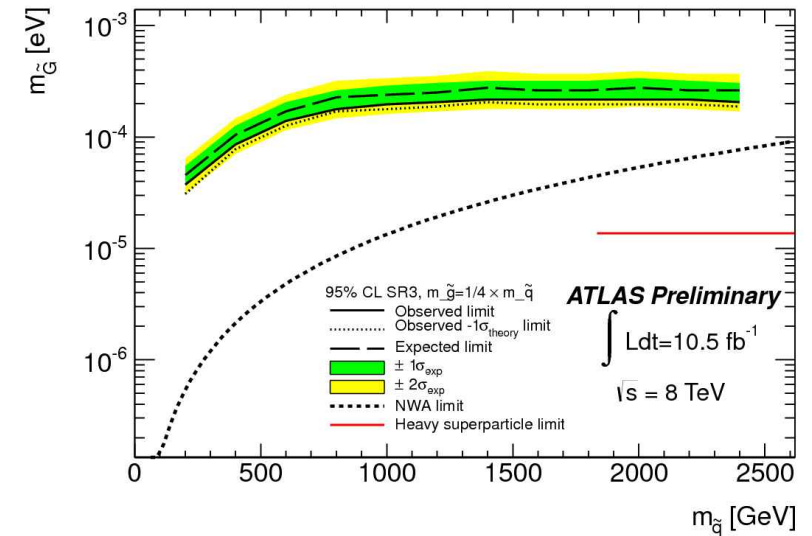
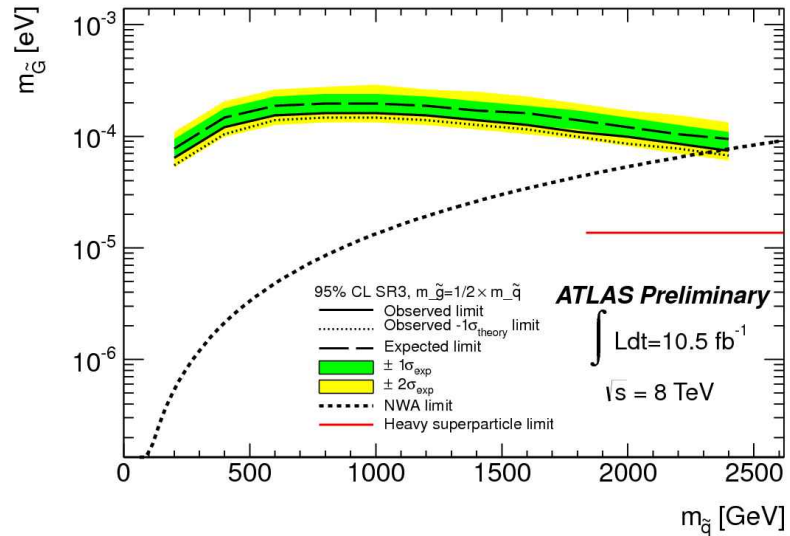
10⁻¹

1

Mass scale [TeV]

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 σ theoretical signal cross section uncertainty.

SUSY searches : gravitino



ATLAS Astroparticle Forum

Exchange ideas about opportunities of using the capability of ATLAS to contribute to the unsolved problems in cosmology and astrophysics.

- Invite astroparticle physics experts from other experiments or theorists to discuss recent developments or results and their possible cross feed with ATLAS analyses.

- Follow up and coordinate astroparticle physics aspects of analyses within existing ATLAS physics working groups and provide expertise of a broader astroparticle physics context.

- Propose astroparticle physics interpretation to an existing analysis and contribute to the work within the relevant physics group.

- Propose re-optimization of an existing analysis for an astroparticle physics interpretation and contribute to the work within the relevant physics group.

- Propose new analyses related to astroparticle physics and contribute to the work within the relevant physics group.

Work on publicizing ATLAS results with astroparticle relevance in the astroparticle community.

The AAF will work proactively on invitations for ATLAS talks (which would be assigned via the normal procedures) by contacting conference organisers and submitting abstracts (together with the physics group in which the analysis of interest is done).

LHC New Physics signals may be related to measurements of astroparticle experiments in terms of Dark Matter. And results from astroparticle experiments may serve to guide some of the ATLAS data analysis and interpretations.

► relevant analyses, compiled based on the ATLAS astro workshop

a) Monojets / photons

- Search for pair production of WIMPs, tagged by ISR photon or jet

b) Top-philic Dark Matter searches

- Motivation: if WIMP miracle is reality, WIMPs may be related to electroweak symmetry breaking and sizable couplings to top quarks would be expected. See e.g. <http://arxiv.org/abs/0912.0004> for a model where top couples via Z' to WIMPs

- Signatures to search for at the LHC: four top and $t\bar{t}$ plus MET final states

c) High tan beta analysis (contact: Anna Lipniacka, Heidi Sandaker)

- Motivation: High tan-beta studies are very interesting to the astroparticle physics community, both in terms of ATLAS reach but also to show overlap or complementarity between the particle and astroparticle experiments. See references and arguments in this talk:

d) Grid searches (Contact: Anna Lipniacka, Christophe Clement, Heidi Sandaker)

- Motivation: Often ATLAS experiments present models/results which are already partially excluded by astroparticle physics data or other experiments.