

Froggatt-Nielsen Models with a Residual \mathbb{Z}_4^R Symmetry

Toby O. Opferkuch

Cargèse Summer School 2014 Student Talks

Based on arXiv:1308.0332
in collaboration with H. K. Dreiner and C. Luhn



July 17, 2014

Discrete symmetries

Use discrete symmetries to stabilise the proton

Exists many popular anomaly free discrete symmetries:

- R-parity R_p or Matter parity M_p

G. Farrar & P. Fayet (1978); S. Dimopoulos, S. Raby & F. Wilczek (1981)

- Baryon triality B_3

L. Ibáñez & G. Ross (1992)

- Proton hexality P_6

K.S. Babu, L. Gogoladze & K. Wang (2002); H.K. Dreiner, C. Luhn & M. Thormeier (2006)

⇒ These symmetries do not forbid the μ -term in the superpotential

Why a \mathbb{Z}_4^R symmetry?

One convenient parametrisation of the \mathbb{Z}_4^R symmetry is:

H. M. Lee, S. Raby, M. Ratz, G. G. Ross, et al. (2010)

	Q_i	\bar{U}_i	\bar{D}_i	L_i	\bar{E}_i	\bar{N}_i	H_u	H_d	θ
\mathbb{Z}_4^R	1	1	1	1	1	1	0	0	1

- Forbids μ -term in the superpotential
- Forbids dim-4 and 5 LNV and BNV terms
- Anomaly free
- Allows the Weinberg operator

The model

$$G_{\text{SM}} \times \text{U}(1)_R \xrightarrow{\text{family symmetry breaking}} G_{\text{SM}} \times \mathbb{Z}_4^R \xrightarrow{\text{SUSY breaking}} G_{\text{SM}} \times \mathbb{Z}_2$$

- ① Provides a gauge symmetry origin of the \mathbb{Z}_4^R symmetry
 \hookrightarrow the $\text{U}(1)_R$ must be gauged
- ② Obtain correct charged fermion masses and mixings
 \hookrightarrow Froggatt-Nielsen mechanism
- ③ Generate the μ -term at an acceptable scale
 \hookrightarrow Giudice-Masiero mechanism
- ④ Ensure anomaly freedom
- ⑤ Allow a mechanism to generate neutrino masses

Outline

- ❶ Motivation for BSM physics
- ❷ A six step procedure for building a \mathbb{Z}_4^R inspired flavour model
- ❸ Results
- ❹ Conclusion

Fermion Hierarchy

- Many free parameters in the SM unconstrained by theoretical arguments:
 - ↪ 9 fermion masses containing huge hierarchy

$$m_t \simeq 172.7 \text{ GeV}$$

⋮

$$m_e = 511 \text{ keV}$$

Further enhanced including neutrino masses $m_{\nu_i} \simeq 0.05 \text{ eV}$

Fermion Hierarchy

- Many free parameters in the SM unconstrained by theoretical arguments:
 - ↪ 9 fermion masses containing huge hierarchy

$$m_t \simeq 172.7 \text{ GeV}$$

⋮

$$m_e = 511 \text{ keV}$$

Further enhanced including neutrino masses $m_{\nu_i} \simeq 0.05 \text{ eV}$

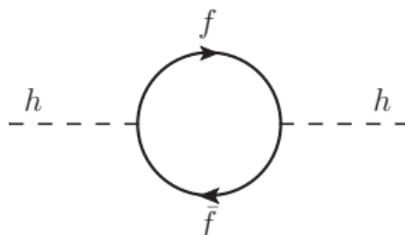
- Neutrino masses:

Assuming G_{SM} & accidental global symmetries

 No neutrino mass terms possible

Motivation for Supersymmetry

- Hierarchy Problem:

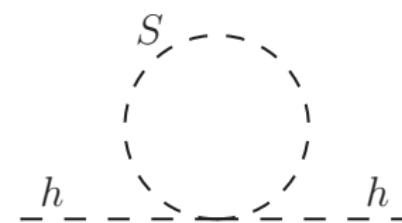
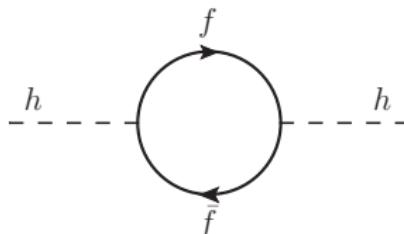


$$\Delta m_h^2 = -\frac{y_f^2}{8\pi^2} \Lambda^2 + \dots$$

- Gauge coupling unification: Does not occur in the SM
- Dark matter: No possibilities within the SM

Motivation for Supersymmetry

- Hierarchy Problem:



$$\Delta m_h^2 = -\frac{y_f^2}{8\pi^2} \Lambda^2 + \dots$$

$$\Delta m_h^2 = \frac{\lambda_S}{16\pi^2} \Lambda^2 + \dots$$

- Grand unification:** Does not occur in the SM
 - ↪ Rectified with extra field content from supersymmetry
- Dark matter:** No possibilities within the SM
 - ↪ Numerous possibilities in supersymmetry

Superfield formalism

Left-chiral superfields

$$\Phi(y, \theta) = \varphi(y) + \sqrt{2}\theta\xi(y) + \theta^2 F(y)$$

Vector superfields (Wess-Zumino gauge)

$$V_{\text{WZ}} = \theta\sigma^\mu\bar{\theta}A_\mu(x) + \theta\theta\bar{\theta}\bar{\lambda}(x) + \bar{\theta}\bar{\theta}\theta\lambda(x) + \frac{1}{2}\theta\theta\bar{\theta}\bar{\theta}D(x)$$

Invariance under both supersymmetry and gauge transformations

$$\mathcal{L}_{\text{SUSY}} = \int d^2\theta [\mathcal{W}^A \mathcal{W}_A + W(\Phi_i) + h.c.] + \int d^2\theta d^2\bar{\theta} K(\Phi_i, \tilde{\Phi}^{\dagger j})$$

Supersymmetric Standard Model

Superpotential terms invariant under G_{SM} :

$$\begin{aligned} W_{\text{ren}} = & Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\ & \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k \end{aligned}$$

Supersymmetric Standard Model

Superpotential terms invariant under G_{SM} :

$$\begin{aligned} W_{\text{ren}} = & Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\ & \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k \end{aligned}$$

$$W_{\text{dim } 5} = \kappa_{ij}^{(0)} L_i H_u L_j H_u + \kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots$$

Supersymmetric Standard Model

Superpotential terms invariant under G_{SM} :

$$\begin{aligned} W_{\text{ren}} = & Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\ & \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k \\ = & \kappa_{ij}^{(0)} L_i H_u L_j H_u + \kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots \end{aligned}$$

Violates only lepton number

Supersymmetric Standard Model

Superpotential terms invariant under G_{SM} :

$$W_{\text{ren}} = Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d$$

$$\frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

$$W_{\text{dim } 5} = \kappa_{ij}^{(0)} L_i H_u L_j H_u + \kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots$$

- Violates only lepton number
- Violates only baryon number

Supersymmetric Standard Model

Superpotential terms invariant under G_{SM} :

$$W_{\text{ren}} = Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\ \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

$$W_{\text{dim } 5} = \kappa_{ij}^{(0)} L_i H_u L_j H_u + \boxed{\kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots}$$

- Violates only lepton number
- Violates only baryon number
- Violates both lepton and baryon number

Supersymmetric Standard Model

Superpotential terms invariant under G_{SM} :

$$W_{\text{ren}} = Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\ \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

$$W_{\text{dim } 5} = \kappa_{ij}^{(0)} L_i H_u L_j H_u + \kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots$$

- Violates only lepton number
 - Violates only baryon number
 - Violates both lepton and baryon number
- μ -term has mass dimension

The μ -problem

J.E. Kim & H.P. Nilles (1984)

Consider an SU(5) GUT

$$\bar{\mathbf{5}}_H \longrightarrow \overbrace{(\mathbf{1}, \mathbf{2})_{-\frac{1}{2}}}^{H_d} \oplus \overbrace{(\bar{\mathbf{3}}, \mathbf{1})_{\frac{1}{3}}}^{H_d^T}$$

If SU(5) is broken at $M_{\text{GUT}} \sim 10^{16} \text{ GeV}$

$$\mu \bar{\mathbf{5}}_H \mathbf{5}_H = M_{\text{GUT}} \bar{\mathbf{5}}_H \mathbf{5}_H$$

- ✓ Stabilises proton namely H_d^T is heavy
- ✗ $\mu H_u H_d$ now too large

The μ -problem

In the MSSM $\mu \simeq M_{\text{EW}}$

↪ Necessary to avoid miraculous cancellation between μ and soft squared-mass

- Higgs obtains the correct VEV $v_u^2 + v_d^2 \simeq (174 \text{ GeV})^2$

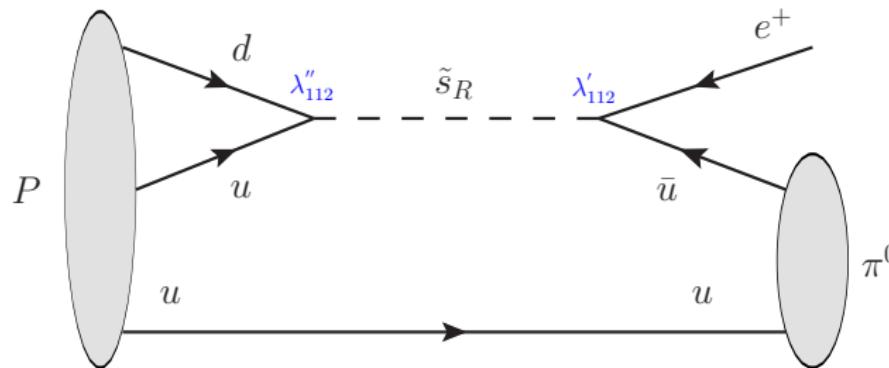
potential yields correct m_Z^2

⇒ μ -term must be forbidden in the high energy theory

Proton Decay

If both lepton and baryon number hazardous for proton decay

S. Dimopoulos, S. Raby & F. Wilczek (1982)



$$|\lambda'_{112} \lambda''_{112}| \leq 2 \times 10^{-27} \left(\frac{M_{\tilde{s}_R}}{100 \text{ GeV}} \right)^2$$

R. Barbier et al. (2005)

Forbid terms allowing proton decay in the Lagrangian!

Discrete symmetries

Use discrete symmetries to stabilise the proton

Exists many popular anomaly free discrete symmetries:

- R-parity R_p or Matter parity M_p

G. Farrar & P. Fayet (1978); S. Dimopoulos, S. Raby & F. Wilczek (1981)

- Baryon triality B_3

L. Ibáñez & G. Ross (1992)

- Proton hexality P_6

K.S. Babu, L. Gogoladze & K. Wang (2002); H.K. Dreiner, C. Luhn & M. Thormeier (2006)

⇒ These symmetries do not forbid the μ -term in the high energy theory

The model

$$G_{\text{SM}} \times \text{U}(1)_R \xrightarrow{\text{family symmetry breaking}} G_{\text{SM}} \times \mathbb{Z}_4^R \xrightarrow{\text{SUSY breaking}} G_{\text{SM}} \times \mathbb{Z}_2$$

- ➊ Forbid the μ -term
→ Possible with an R -symmetry
- ➋ Obtain correct charged fermion masses and hierarchy
→ Froggatt-Nielsen mechanism
- ➌ Generate the μ -term at an acceptable scale
→ Giudice-Masiero mechanism
- ➍ Ensure anomaly freedom
- ➎ Provides a gauge symmetry origin of the \mathbb{Z}_4^R symmetry
→ the $\text{U}(1)_R$ must be gauged
- ➏ Allow a mechanism to generate neutrino masses

Step 1: Forbid the μ -term

Continuous R -symmetries:

Superspace coordinates are charged

$$\theta \rightarrow \theta' = e^{iq_\theta^R \alpha(x)} \theta$$

$$\int d\theta \rightarrow \int d\theta' = e^{-iq_\theta^R \alpha(x)} \int d\theta$$

Step 1: Forbid the μ -term

Continuous R -symmetries:

Superspace coordinates are charged

$$\theta \rightarrow \theta' = e^{iq_\theta^R \alpha(x)} \theta$$

$$\int d\theta \rightarrow \int d\theta' = e^{-iq_\theta^R \alpha(x)} \int d\theta$$

Implications for model building:

- Superpotential must be charged under $U(1)_R$

$$\int d^2\theta W(\Phi) \rightarrow \int d^2\theta e^{2iq_\theta^R \alpha(x)} W'(\Phi)$$

- Kähler potential unchanged

$$\int d^2\theta d^2\bar{\theta} K(\Phi_i, \tilde{\Phi}^{\dagger j}) \rightarrow \int d^2\theta d^2\bar{\theta} e^{2iq_\theta^R \alpha(x)} e^{-2iq_\theta^R \alpha(x)} K(\Phi_i, \tilde{\Phi}^{\dagger j})$$

Step 1: Forbid the μ -term

A gauged $U(1)_R$ in supergravity includes:

- supergravity multiplet: (e_μ^a, ψ_μ)
- R -vector multiplet: $V^R = (A_\mu^R, \lambda^R, D^R)$
- usual MSSM vector multiplets: $V^k = (A_\mu^k, \lambda^k, D^k)$

Step 1: Forbid the μ -term

A gauged $U(1)_R$ in supergravity includes:

- supergravity multiplet: (e_μ^a, ψ_μ)
- R -vector multiplet: $V^R = (A_\mu^R, \lambda^R, D^R)$
- usual MSSM vector multiplets: $V^k = (A_\mu^k, \lambda^k, D^k)$

The spin-1/2 components transform under $U(1)_R$ as:

$$\lambda(x) \rightarrow e^{iq_\theta^R \alpha(x)} \lambda(x), \quad \xi(x) \rightarrow e^{i(q_\Phi^R - q_\theta^R) \alpha(x)} \xi(x)$$

- ↪ Expect a gaugino coupling to R -photon
- ↪ Contribution to anomalies from gauginos and gravitinos

Step 1: Forbid the μ -term

After the $U(1)_R$ symmetry is spontaneously broken we require a \mathbb{Z}_4^R symmetry of the form:

[H. M. Lee, S. Raby, M. Ratz, G. G. Ross, et al. \(2010\)](#)

	Q_i	\bar{U}_i	\bar{D}_i	L_i	\bar{E}_i	\bar{N}_i	H_u	H_d	θ
\mathbb{Z}_4^R	1	1	1	1	1	1	0	0	1

- Forbids μ -term in the superpotential
- Forbids dim-4 and 5 LNV and BNV terms
- Anomaly free

Step 2: Froggatt-Nielsen mechanism

C. Froggatt & H. B. Nielsen (1979)

Result: Free matrix textures \rightarrow free charges under $U(1)_R$

Step 2: Froggatt-Nielsen mechanism

Provides a symmetry argument to fix Yukawa textures

C. Froggatt & H. B. Nielsen (1979)

Requires:

- additional gauge symmetry
- flavon superfield Φ (G_{SM} singlet)

Result: Free matrix textures \rightarrow free charges under $U(1)_R$

Charges under $U(1)_R$ constrained by:

- phenomenologically acceptable fermion masses
- anomaly cancellation
- requirement of a residual \mathbb{Z}_4^R symmetry

Step 2: Froggatt-Nielsen mechanism

Start with standard Yukawa term in superpotential

$$Y_d^{ij} Q_i H_d \bar{D}_j$$

Step 2: Froggatt-Nielsen mechanism

Start with standard Yukawa term in superpotential

$$Y_d^{ij} Q_i H_d \bar{D}_j \xrightarrow{\text{step 1}} y_d^{ij} \left(\frac{\Phi}{M_P} \right)^n \Theta[n] Q_i H_d \bar{D}_j$$

- ➊ multiply with powers of Φ introducing function $\Theta[n]$

$$\Theta[n] = \begin{cases} 1 & n \in \mathbb{N} \\ 0 & \text{otherwise} \end{cases}$$

Step 2: Froggatt-Nielsen mechanism

Start with standard Yukawa term in superpotential

$$\begin{aligned} Y_d^{ij} Q_i H_d \bar{D}_j &\xrightarrow{\text{step 1}} y_d^{ij} \left(\frac{\Phi}{M_P} \right)^n \Theta[n] Q_i H_d \bar{D}_j \\ &\xrightarrow{\text{step 2}} y_d^{ij} \left(\frac{\Phi}{M_P} \right)^{q_{Q_i}^R + q_{H_d}^R + q_{\bar{D}_j}^R - 2q_\theta^R} \Theta[\textcolor{red}{n}] Q_i H_d \bar{D}_j \end{aligned}$$

- ➊ multiply with powers of Φ introducing function $\Theta[n]$

- ➋

This choice implies q_θ^R must be a free parameter!

Step 2: Froggatt-Nielsen mechanism

Start with standard Yukawa term in superpotential

$$\begin{aligned}
 Y_d^{ij} Q_i H_d \bar{D}_j &\xrightarrow{\text{step 1}} y_d^{ij} \left(\frac{\Phi}{M_P} \right)^n \Theta[n] Q_i H_d \bar{D}_j \\
 &\xrightarrow{\text{step 2}} y_d^{ij} \left(\frac{\Phi}{M_P} \right)^{q_{Q_i}^R + q_{H_d}^R + q_{\bar{D}_j}^R - 2q_\theta^R} \Theta[n] Q_i H_d \bar{D}_j \\
 &\xrightarrow{\text{step 3}} y_d^{ij} \left(\frac{\langle \Phi \rangle}{M_P} \right)^{q_{Q_i}^R + q_{H_d}^R + q_{\bar{D}_j}^R - 2q_\theta^R} \Theta[n] Q_i H_d \bar{D}_j
 \end{aligned}$$

- ➊ multiply with powers of Φ introducing function $\Theta[n]$
- ➋ choose $q_\Phi^R = -1$ and fix n through $U(1)_R$ gauge invariance
- ➌ flavon field acquires a VEV breaking $U(1)_R$

Step 2: Froggatt-Nielsen mechanism

Define FN expansion parameter

$$\epsilon \equiv \left(\frac{\langle \Phi \rangle}{M_P} \right) \simeq \lambda_c = 0.23$$

Step 2: Froggatt-Nielsen mechanism

Define FN expansion parameter

$$\epsilon \equiv \left(\frac{\langle \Phi \rangle}{M_P} \right) \simeq \lambda_c = 0.23$$

Fermion mass ratios and CKM matrix expressed as

$$\begin{aligned} m_d : m_s : m_b &\sim \lambda_c^4 : \lambda_c^2 : 1 \\ m_u : m_c : m_t &\sim \lambda_c^8 : \lambda_c^4 : 1 \\ m_e : m_\mu : m_\tau &\sim \lambda_c^{4+z} : \lambda_c^2 : 1 \end{aligned}$$

$$U_{\text{CKM}} \sim \begin{pmatrix} 1 & \lambda_c^{1+y} & \lambda_c^{3+y} \\ \lambda_c^{1+y} & 1 & \lambda_c^2 \\ \lambda_c^{3+y} & \lambda_c^2 & 1 \end{pmatrix}$$

$$m_b : m_t \sim \lambda_c^x \cot \beta \quad \text{with parameters: } x = 0, 1, 2, 3$$

$$m_\tau : m_b \sim 1 \quad z = 0, 1 \quad y = -1, 0, -1$$

$$\Delta_{ij}^L = q_{L_i}^R - q_{L_j}^R$$

 \hookrightarrow left with $q_{Q_1}^R, q_{L_1}^R, q_{H_u}^R, q_{H_d}^R, q_\theta^R$ & introduced parameters

Step 3: Giudice-Masiero mechanism

Used to effectively generate terms from Kähler potential

[G.F. Giudice & A. Masiero \(1988\)](#)

$$K_{\text{Higgs}} \supset \left(\mu H_u H_d^\dagger + h.c. \right) + \left(\frac{K Z^\dagger}{M_P} H_u H_d + h.c. \right) + \dots$$

Step 3: Giudice-Masiero mechanism

Used to effectively generate terms from Kähler potential

[G.F. Giudice & A. Masiero \(1988\)](#)

$$K_{\text{Higgs}} \supset \left(\mu H_u H_d^\dagger + h.c. \right) + \left(\frac{K Z^\dagger}{M_P} H_u H_d + h.c. \right) + \dots$$

If the F -term acquires a VEV $\langle F_{Z^\dagger} \rangle \sim m_{3/2} M_P$

$$\int d^2\theta d^2\bar{\theta} \left(\frac{\langle F_{Z^\dagger} \rangle \bar{\theta}^2}{M_P} \right) K H_u H_d \longrightarrow \int d^2\theta m_{3/2} K H_u H_d$$

Step 3: Giudice-Masiero mechanism

Used to effectively generate terms from Kähler potential

G.F. Giudice & A. Masiero (1988)

$$K_{\text{Higgs}} \supset \left(\mu H_u H_d^\dagger + h.c. \right) + \left(\frac{K Z^\dagger}{M_P} H_u H_d + h.c. \right) + \dots$$

If the F -term acquires a VEV $\langle F_{Z^\dagger} \rangle \sim m_{3/2} M_P$

$$\int d^2\theta d^2\bar{\theta} \left(\frac{\langle F_{Z^\dagger} \rangle \bar{\theta}^2}{M_P} \right) K H_u H_d \longrightarrow \int d^2\theta m_{3/2} K H_u H_d$$

Applying FN mechanism

$$m_{3/2} \kappa \epsilon^{|q_{H_u}^R + q_{H_d}^R|} \tilde{\Theta} [q_{H_u}^R + q_{H_d}^R] H_u H_d$$

The important difference:

$$\tilde{\Theta}[n] = \begin{cases} 1 & n \in \mathbb{Z} \\ 0 & \text{otherwise} \end{cases}$$

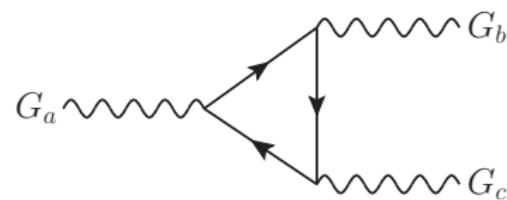
Step 4: Anomaly constraints

$$(\partial_\mu j^\mu(x))^a = \frac{-1}{32\pi^2} A^{abc} \epsilon^{\alpha\nu\beta\rho} F_{\alpha\nu}^b F_{\beta\rho}^c$$

$$A^{abc} = \frac{1}{2} \text{tr}\{T^a, T^b\} T^c$$

Anomaly freedom requires either:

- anomaly coefficients cancel $A^{abc} = 0 \quad \forall a, b, c$
- ↪ too restrictive: fermion mass hierarchy not possible



Step 4: Anomaly constraints

$$(\partial_\mu j^\mu(x))^a = \frac{-1}{2\omega^2} A^{abc} \epsilon^{\alpha\nu\beta\rho} F_{\alpha\nu}^b F_{\beta\rho}^c$$

$$T^b \} T^c$$

Ano

es either:

- cancel $A^{abc} = 0 \quad \forall a, b, c$

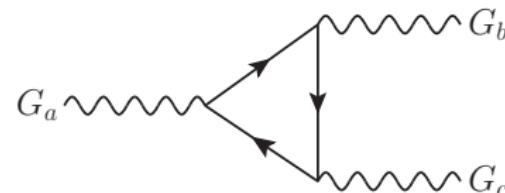
↪ too restrictive: fermion mass hierarchy not possible

- anomaly cancellation through Green-Schwarz mechanism

M. B. Green & J. H. Schwarz (1984)

$$\frac{A_{SU(3)_C^2 - U(1)_R}}{k_C} = \frac{A_{SU(2)_W^2 - U(1)_R}}{k_W} = \frac{A_{U(1)_Y^2 - U(1)_R}}{k_Y} = 2\pi^2 \delta_{GS}$$

↪ Can be used to fix $q_{Q_1}^R$ and $q_{H_u}^R$



Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

Why should discrete symmetries have gauge origins?

L.M. Krauss & F. Wilczek (1989)

- global discrete symmetries violated by quantum effects
- ensure domain walls are not present
- multiple U(1) symmetries present in many UV completions

Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

Why should discrete symmetries have gauge origins?

L.M. Krauss & F. Wilczek (1989)

- global discrete symmetries violated by quantum effects
- ensure domain walls are not present
- multiple U(1) symmetries present in many UV completions

Simple $U(1)_X$ example:

$$\phi_i \longrightarrow \phi'_i = e^{i\alpha(x)q_{\phi_i}^X} \phi_i$$

If $q_{\phi_1}^X = N$ obtains a VEV

$$\sum_{i'} q_{\phi_{i'}}^X = 0 \xrightarrow{\langle \phi_1 \rangle \neq 0} \sum_{i'} q_{\phi_{i'}}^X = 0 \mod N$$

Remaining fields transform as

$$\phi_{i'} \longrightarrow \phi'_{i'} = e^{\frac{2\pi i}{N} q_{\phi_{i'}}^X} \phi_{i'}$$

Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

Ensure the \mathbb{Z}_4^R symmetry results after Φ obtains a VEV
⇒ reverse engineer remaining unfixed $U(1)_R$ charges
⇒ not transparent as $q_\Phi^R = -1$

Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

Ensure the \mathbb{Z}_4^R symmetry results after Φ obtains a VEV

\implies reverse engineer remaining unfixed $U(1)_R$ charges

\implies not transparent as $q_\Phi^R = -1$

Method:

- start with a completely general term

$$q_{\text{total}}^R = n_{H_u} q_{H_u}^R + n_{H_d} q_{H_d}^R + \sum_i (n_{Q_i} q_{Q_i}^R + n_{\bar{U}_i} q_{\bar{U}_i}^R + \dots)$$

Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

Ensure the \mathbb{Z}_4^R symmetry results after Φ obtains a VEV

\implies reverse engineer remaining unfixed $U(1)_R$ charges

\implies not transparent as $q_\Phi^R = -1$

Method:

- start with a completely general term

$$q_{\text{total}}^R = n_{H_u} q_{H_u}^R + n_{H_d} q_{H_d}^R + \sum_i (n_{Q_i} q_{Q_i}^R + n_{\bar{U}_i} q_{\bar{U}_i}^R + \dots)$$

- eliminate parameters in above equation by:

- enforcing SM gauge invariance
- ensuring the minimum superpotential

Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

- The charge under the \mathbb{Z}_4^R symmetry is expressed as

$$\sum_i (n_{Q_i} + n_{L_i} + n_{\bar{U}_i} + n_{\bar{D}_i} + n_{\bar{E}_i} + n_{\bar{N}_i}) = 4\mathcal{N}_{\mathbb{Z}_4^R} + 2S + C_4$$

$$S = \begin{cases} 1 & \text{for superpotential terms} \\ 0 & \text{for Kähler potential terms} \end{cases}$$

$\mathcal{N}_{\mathbb{Z}_4^R}$ implements mod 4 nature of \mathbb{Z}_4^R

Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

- The charge under the \mathbb{Z}_4^R symmetry is expressed as

$$\sum_i (n_{Q_i} + n_{L_i} + n_{\bar{U}_i} + n_{\bar{D}_i} + n_{\bar{E}_i} + n_{\bar{N}_i}) = 4\mathcal{N}_{\mathbb{Z}_4^R} + 2S + C_4$$

$$S = \begin{cases} 1 & \text{for superpotential terms} \\ 0 & \text{for Kähler potential terms} \end{cases}$$

$\mathcal{N}_{\mathbb{Z}_4^R}$ implements mod 4 nature of \mathbb{Z}_4^R

- Main idea: fix $U(1)_R$ charges such that:

If $C_4 = 0$ then Eq. (1) must be fulfilled

If $C_4 \neq 0$ then Eq. (1) must not be fulfilled

$$q_{\text{total}}^R = \begin{cases} \mathbb{N} + 2q_\theta^R, & \text{for superpotential terms} \\ \mathbb{Z}, & \text{for Kähler potential terms} \end{cases} \quad (1)$$

Step 5: Residual discrete gauge origin of \mathbb{Z}_4^R

Performing above steps:

Fixes remaining charges and constraints q_θ^R

$$q_\theta^R = \frac{2\mathbb{Z} + 1}{4}$$

Remaining integer parameters are $x, y, z, \Delta_{31}^L, \Delta^H$ and ζ plus q_θ^R

Step 5: Residual

Performing above step

Fixes remaining charge

gauge origin of \mathbb{Z}_4^R strains q_θ^R

$$q_\theta^R = \frac{2\mathbb{Z} + 1}{4}$$

Remaining integer parameters are $x, y, z, \Delta_{31}^L, \Delta^H$ and ζ plus q_θ^R Previously $\epsilon \stackrel{!}{\simeq} \lambda_c = 0.23$

$$-2 \lesssim q_\theta^R \lesssim 1 \implies q_\theta^R = \left\{ -\frac{7}{4}, -\frac{5}{4}, -\frac{3}{4}, -\frac{1}{4}, \frac{1}{4}, \frac{3}{4} \right\}$$

Only two values lead to acceptable μ -term ($|n| \leq 5$)

$$m_{3/2} \kappa \epsilon^{|n|} H_u H_d \implies q_\theta^R = \pm \frac{1}{4}$$

Step 6: Neutrino masses and mixing

Applying FN mechanism to generate masses:

- Non-renormalisable Weinberg term
 - ✗ maximum obtainable mass scale 10^{-5} eV
- See-saw
 - ✓ phenomenologically acceptable mass scale possible
 - ✓ requires RH-neutrino fields \bar{N}_i
 - ✗ $M_R^{ij} \bar{N}_i \bar{N}_j$ must be in superpotential

Via see-saw formula light neutrino mass matrix

$$M_\nu^{ij} \simeq \frac{\langle H_u \rangle^2}{M_P} \epsilon^{q_{L_i}^R + q_{L_j}^R + 2q_{H_u}^R - 2q_\theta^R}$$

$$\hookrightarrow m_{\text{abs}}^\nu \simeq \frac{\langle H_u \rangle^2}{M_P} \epsilon^{\Delta^H - 2z + \frac{1}{2} + 2\Delta_{31}^L + 14q_\theta^R}$$

Step 6: Neutrino masses and mixing

Enforce suitable PMNS mixing matrix

$$U_{\text{PMNS}} = V_{e_L} \tilde{V}_{\nu_L}^\dagger T \simeq \begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix} \simeq \begin{pmatrix} \epsilon^{0,1} & \epsilon^{0,1} & \epsilon^{0,1} \\ \epsilon^{0,1} & \epsilon^{0,1} & \epsilon^{0,1} \\ \epsilon^{0,1} & \epsilon^{0,1} & \epsilon^{0,1} \end{pmatrix}$$

FN structure $V_{\nu_L}^{ij} \simeq \epsilon^{|q_{L_i}^R - q_{L_j}^R|}$

$$\epsilon^{0,1} \simeq \epsilon^{|q_{L_i}^R - q_{L_j}^R|}$$

This sets the values of $(\Delta_{21}^L, \Delta_{31}^L)$ where $\Delta_{ij}^L = q_{L_i}^R - q_{L_j}^R$:

- $(0, 0)$ allows for normal, inverted hierarchy and degenerate
- $(0, -1)$ allows for normal hierarchy
- $(-1, -1)$ allows for normal hierarchy and inverted hierarchy

Step 6: Neutrino masses and mixing

Fixing m_{abs}^ν depending on neutrino spectrum

$m_1 < m_2 \ll m_3 \sim 0.05 \text{ eV}$ (normal hierarchy)

$m_3 \ll m_1 < m_2 \sim 0.05 \text{ eV}$ (inverted hierarchy)

$0.2 \text{ eV} < m_1 \sim m_2 \sim m_3 < 2.2 \text{ eV}$ (degenerate)

This determines all remaining charges and parameters except RH-neutrino charges

102 sets of unique charges and parameter values

Anomaly constraints

Three sets of promising charges:

Set 1

i	$q_{Q_i}^R$	$q_{U_i}^R$	$q_{D_i}^R$	$q_{L_i}^R$	$q_{E_i}^R$
1	$\frac{69}{20}$	$\frac{109}{20}$	$\frac{13}{20}$	$-\frac{7}{20}$	$\frac{89}{20}$
2	$\frac{69}{20}$	$\frac{29}{20}$	$-\frac{27}{20}$	$-\frac{27}{20}$	$\frac{69}{20}$
3	$\frac{29}{20}$	$-\frac{11}{20}$	$-\frac{27}{20}$	$-\frac{27}{20}$	$\frac{29}{20}$

$q_{H_u}^R$	$q_{H_d}^R$	q_θ^R	q_Φ^R
$-\frac{7}{5}$	$-\frac{3}{5}$	$-\frac{1}{4}$	-1

- ✓ Accommodates both an inverted and normal neutrino hierarchy
- ✓ $\epsilon = 0.192$
- ✓ absolute neutrino mass scale $m_{\text{abs}}^\nu = 0.049 \text{ eV}$
- ✗ μ -term has the size $m_{3/2}\epsilon^2$
- ✗ $y = 1 \implies \text{CKM matrix not optimal}$

Results

Three sets of promising charges:

Set 2

i	$q_{Q_i}^R$	$q_{U_i}^R$	$q_{D_i}^R$	$q_{L_i}^R$	$q_{E_i}^R$
1	$\frac{65}{12}$	$\frac{67}{12}$	$-\frac{17}{12}$	$-\frac{5}{4}$	$\frac{25}{4}$
2	$\frac{41}{12}$	$\frac{43}{12}$	$-\frac{17}{12}$	$-\frac{5}{4}$	$\frac{13}{4}$
3	$\frac{17}{12}$	$-\frac{19}{12}$	$-\frac{17}{12}$	$-\frac{5}{4}$	$\frac{5}{4}$

$q_{H_u}^R$	$q_{H_d}^R$	q_θ^R	q_Φ^R
$-\frac{5}{2}$	$\frac{7}{2}$	$\frac{1}{4}$	-1

- ✓ Accommodates only a degenerate neutrino mass spectrum
- ✓ $\epsilon = 0.205$
- ✗ absolute neutrino mass scale $m_{\text{abs}}^\nu = 4.158 \text{ eV}$
- ✓ μ -term has the size $m_{3/2}\epsilon$
- ✗ $y = -1 \implies \text{CKM matrix not optimal}$

Summary

Constructed a set of viable flavour models that using a gauged $U(1)_R$ family symmetry

- stabilise the proton
- generate correct μ -term
- result in a residual discrete \mathbb{Z}_4^R gauge symmetry
- generate the necessary hierarchies in the Yukawa matrices

Thank you for your attention!

Discrete Symmetries

$$= Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\ \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

$$W_{\text{dim } 5} = \kappa_{ij}^{(0)} L_i H_u L_j H_u + \kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots$$

Forbidden by matter parity / R-parity

G. Farrar & P. Fayet (1978)

S. Dimopoulos, S. Raby & F. Wilczek (1981)

- ✓ Allows for neutrino masses from the Weinberg term
- ✗ Allows dim 5 proton decay operators

Discrete Symmetries

$$W_{\text{ren}} = Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\ \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

$$W_{\text{dim}} = \kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots$$

Forbidden by baryon triality

L. Ibáñez & G. Ross (1992)

- ✓ Allows for neutrino masses from the Weinberg term
- ✓ Allows for neutrino masses from the LNV term LH_u
- ✓ Proton stability ensured as all BNV terms forbidden
- ✗ Other \mathcal{M}_p terms strongly constrained
- ✗ LSP decays \Rightarrow no dark matter candidate

Discrete Symmetries

$$\begin{aligned}
 W_{\text{ren}} = & Y_{ij}^e L_i H_d \bar{E}_j + Y_{ij}^d Q_i H_d \bar{D}_j + Y_{ij}^u Q_i H_u \bar{U}_j + \mu H_u H_d \\
 & \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \mu_i L_i H_u + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k \\
 W_{\text{dim}} = & \kappa_{ijkl}^{(1)} Q_i Q_j Q_k L_l + \kappa_{ijkl}^{(2)} \bar{U}_i \bar{U}_j \bar{D}_k \bar{E}_l + \dots
 \end{aligned}$$

Forbidden by proton hexality

K.S. Babu, L. Gogoladze & K. Wang (2002)

H.K. Dreiner, C. Luhn & M. Thormeier (2006)

- ✓ Allows for neutrino masses from the Weinberg terms
- ✓ Forbids all dim 4 and 5 proton decay inducing terms
- ✓ Forbids all LNV and BNV terms except the Weinberg term

Anomaly constraints

Green-Schwarz anomaly cancellation does not include $A_{U(1)_R^2 - U(1)_Y}$

$$\hookrightarrow A_{U(1)_R^2 - U(1)_Y} \stackrel{!}{=} 0$$

Can be used to fix $q_{H_d}^R$

Anomaly constraints

Green-Schwarz anomaly cancellation does not include $A_{U(1)_R^2 - U(1)_Y}$

$$\hookrightarrow A_{U(1)_R^2 - U(1)_Y} \stackrel{!}{=} 0$$

 Can be used to fix $q_{H_d}^R$

Aside: anomalous nature of $U(1)_R$ can be used to fix $\langle \Phi \rangle$

Dine-Seiberg-Wen-Witten mechanism induces Fayet-Iliopoulos term $\xi \sim \delta_{GS}$

M. Dine, N. Seiberg, X.G. Wen & E. Witten (1986,1987)

For SUSY to be unbroken:

$$\Rightarrow \epsilon = \frac{\langle \Phi \rangle}{M_P} = \frac{g_C}{4\pi} \sqrt{A_{SU(3)_C^2 - U(1)_R}} = \frac{g_C}{4\pi} \sqrt{\frac{3}{2} (x + z - 6q_\theta^R + 6)}$$