

## 126Gev Higgs in Next-to-Minimal-UED

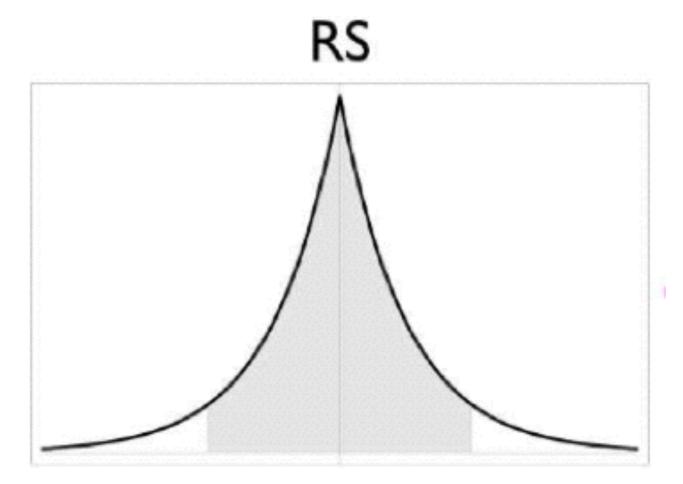
s.park (skku)

PASCOS, Taipei Nov 21, 2013

Flacke, SCP, Kong arXiv:1309.7077 to appear in PLB

## VED AS AN EFFECTIVE THEORY OF RS

Csaki, Heinonen, Hubisz, SCP, Shu (2011)



## IR UV IR

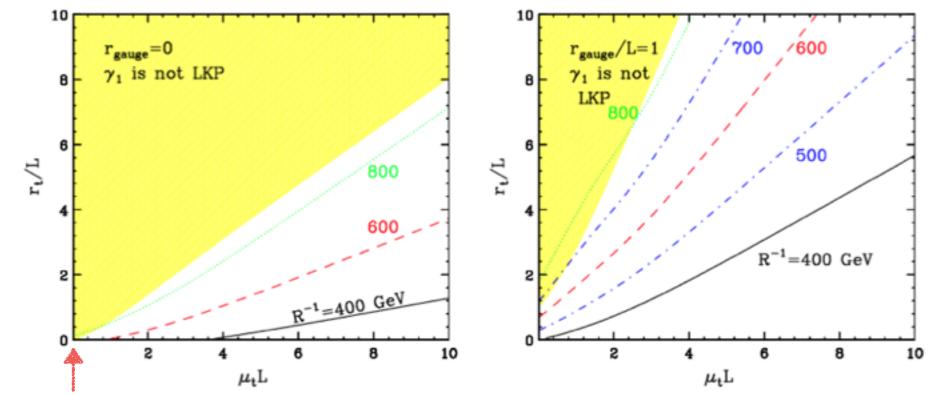
- -KK Dark matter with KK-parity
- -Big hierarchy with warping
- -Rich phenomenology~ bosonic SUSY
- -Still UED is an attractive model for the LHC!



Flacke, <u>SCP</u>, Kong arXiv:1309.7077 to appear in PLB

### In NMUED, I/R~O(100)GeV is still okay with H126

\*\* bulk mass makes KK-fermion heavier, BLT makes gauge boson lighter



I/R>700 GeV in MUED

$$\begin{split} S_{NMUED} &= S_{MUED} + \int d^4x \int_{-L}^{L} dy \left\{ -M_t \xi_{L/R} \overline{\Psi}_3 \Psi_3 \right. \\ &\left. + \left[ \delta(y-L) + \delta(y+L) \right] \left[ r_g \mathcal{L}_{MUED} + (r_t - r_g) i \overline{\Psi}_{3,L/R} \not\!\!\!D \Psi_{3,L/R} \right] \right\} \end{split}$$

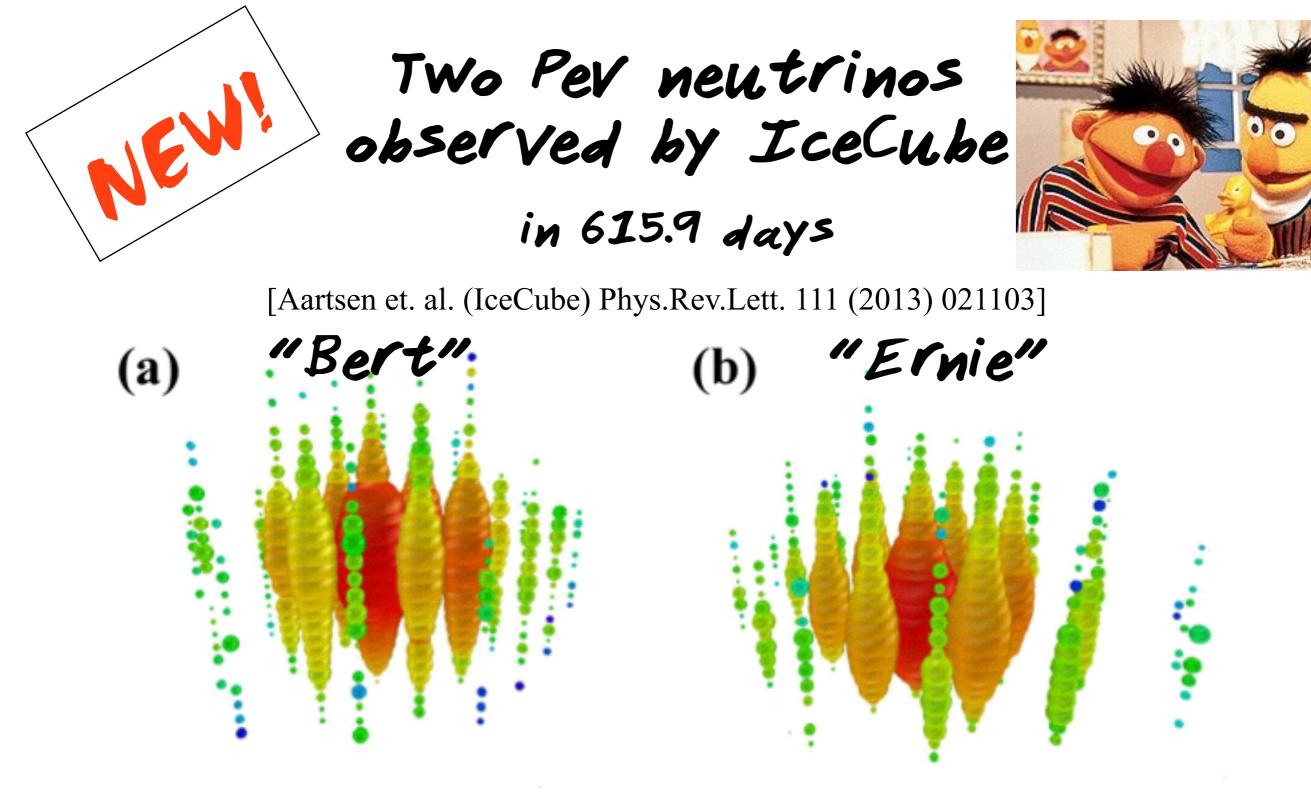


## 126Gev Higgs and Pev scale DM

#### s.park (skku)

PASCOS, Taipei Nov 21, 2013

work with K. Kohri (KEK), C. Rott (IceCube, SKKU)



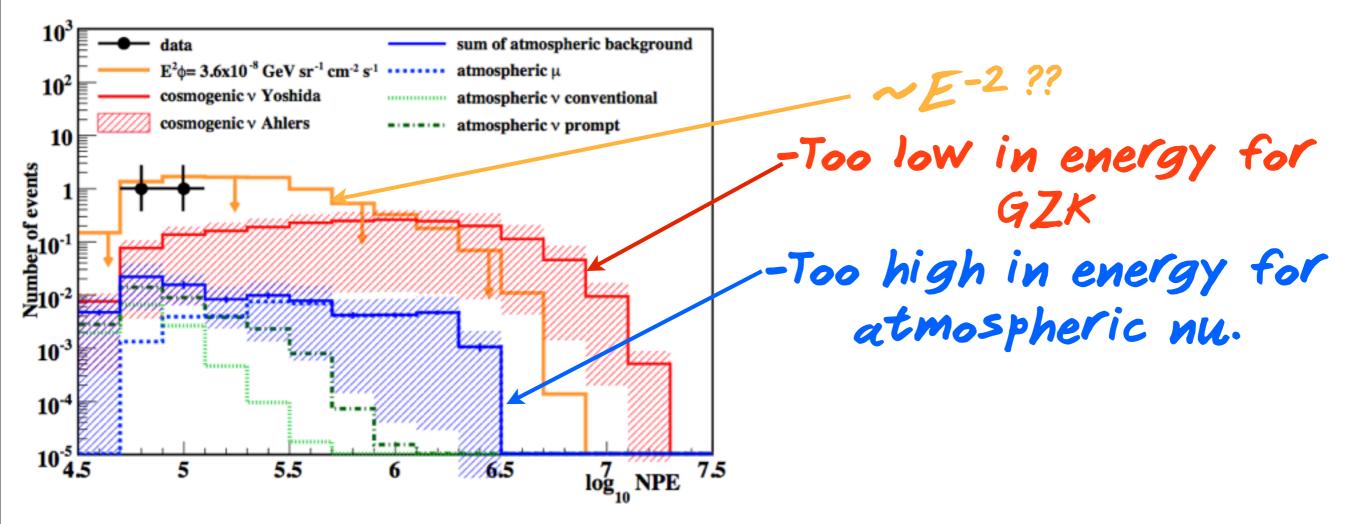
#### $1.04 \pm 0.16 \mathrm{PeV}$

 $1.14 \pm 0.17 \mathrm{PeV}$ 

~consistent with fully contained simulated particle showers induced by neutral-current  $v_{e,\mu,\tau}$  or charged-current  $v_e$  interactions within the IceCube detector.

The observational result looks odd ..

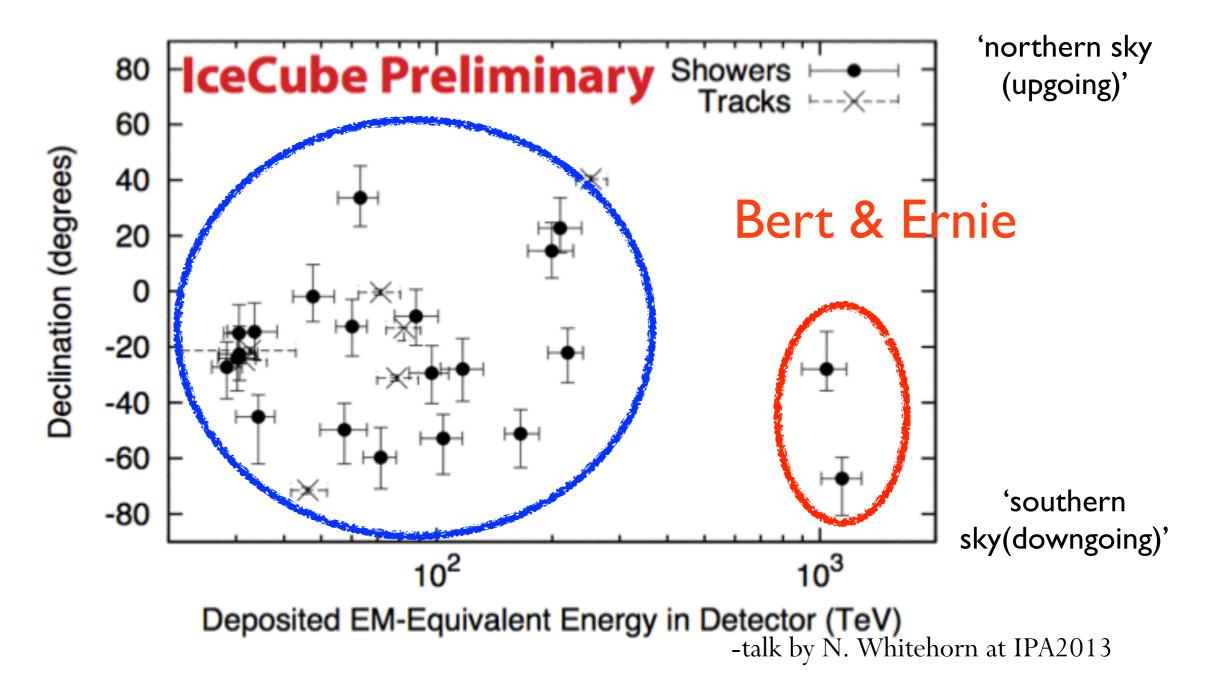
**\*\****Expected*:  $0.082 \pm 0.0024^{+0.041}_{-0.057}$ 



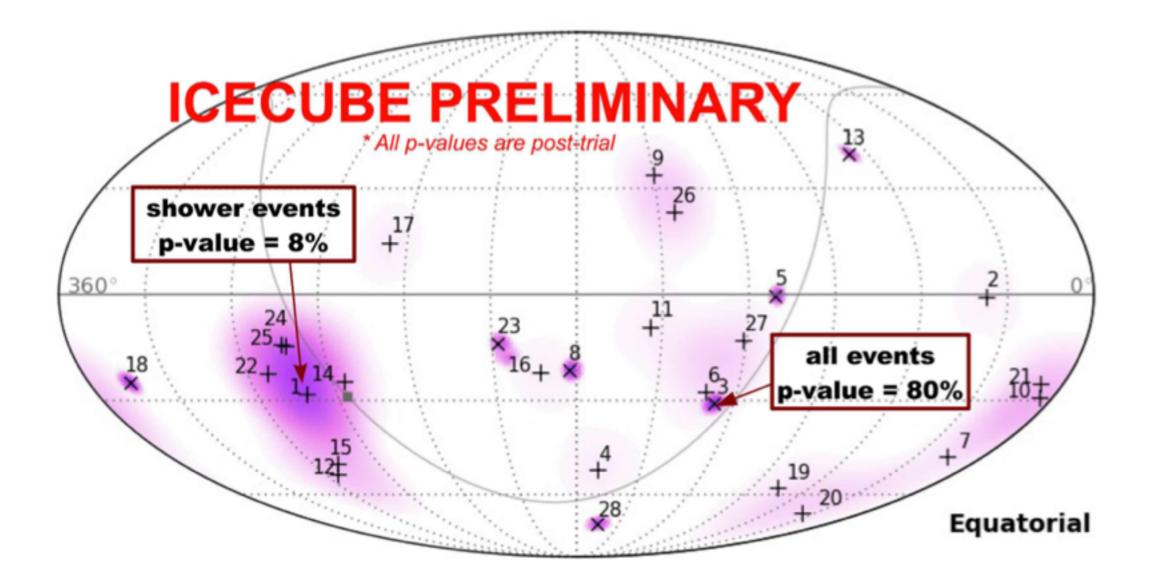
upshot:

These events cannot be understood by Known sources!

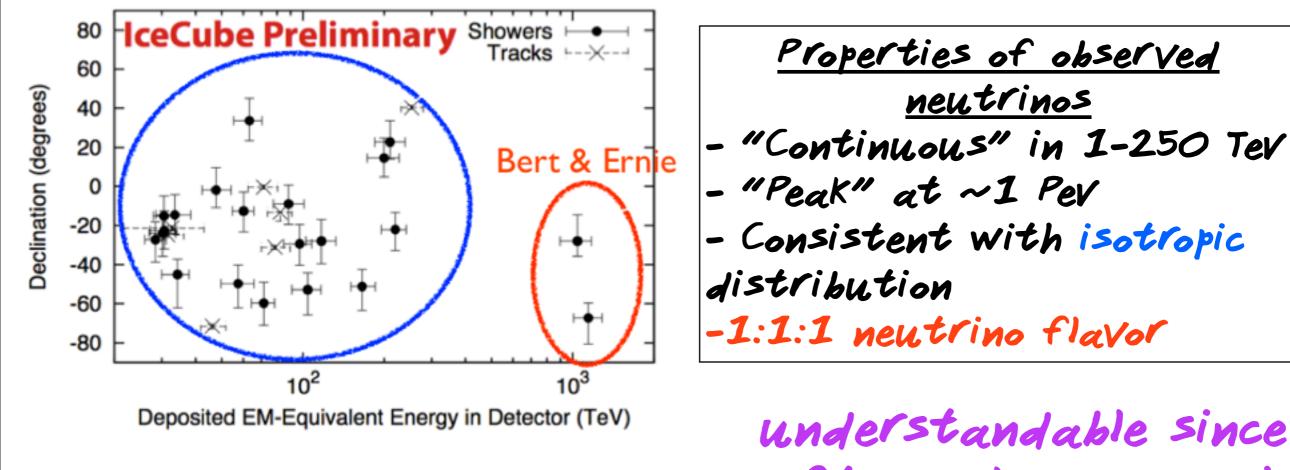
### In addition, 26 more neutrinos observed in ITev-250Tev Window, (cf) background is 10.6+-4.5



Skymap: No Significant Clustering ~not from a local source



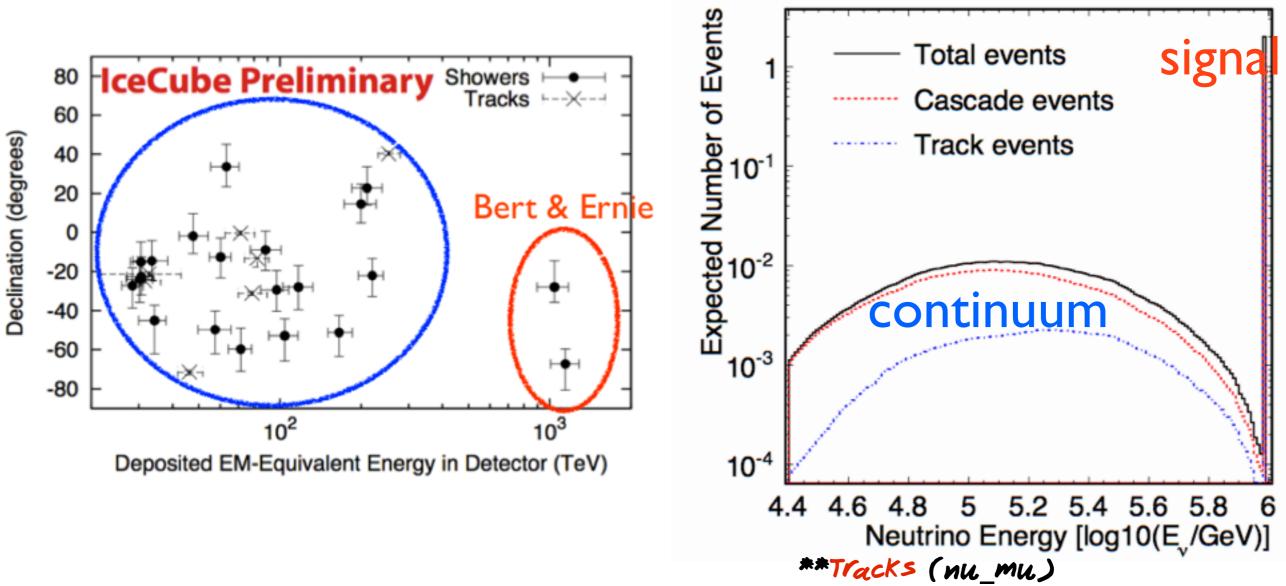
## Closer look at the DATA



$$P(\nu_e \leftrightarrow \nu_e) = 0.56,$$
  
$$P(\nu_e \leftrightarrow \nu_{\mu}) = P(\nu_e \leftrightarrow \nu_{\tau}) = 0.22,$$
  
$$P(\nu_{\mu} \leftrightarrow \nu_{\mu}) = P(\nu_{\mu} \leftrightarrow \nu_{\tau}) = P(\nu_{\tau} \leftrightarrow \nu_{\tau}) = 0.39.$$

understandable since after a long enough propagation, neutrino flavor info. Would disappear

## The "continuum+peak" may imply particle DM! $\chi \rightarrow \nu_L + X(\nu + \cdots)$



\*\*Cascades (nu\_e+nu\_tau)

Annihilation

[Feldman, Kusenko, Matsumoto, Yanagida]

$$\begin{split} \Gamma_{Events} \sim VL_{MW} n_N \sigma_N \times \left(\frac{\rho_{\rm DM}}{m_{\rm DM}}\right)^2 \langle \sigma_A v \rangle \\ \text{with} & \begin{array}{c} \sigma_N \sim 9 \times 10^{-34} {\rm cm}^2 & M_{\rm DM} \sim 1 \ {\rm PeV} \\ n_N \simeq n_{\rm Ice} \simeq 5 \times 10^{23} / {\rm cm}^3 \\ \rho_{\rm DM} \simeq 0.4 {\rm GeV} / {\rm cm}^3, \\ v \sim 10^{-3}, \\ L_{\rm MW} \simeq 10 {\rm kpc} \\ V \simeq 1 {\rm km}^3 \\ \sigma_A \leq 4\pi / (m_{\rm DM}^2 v^2) \\ \Gamma \lesssim 1 \ {\rm per \ few \ hundred \ years} \end{split}$$
(essentially impossible to explain IceCube data)

Decay

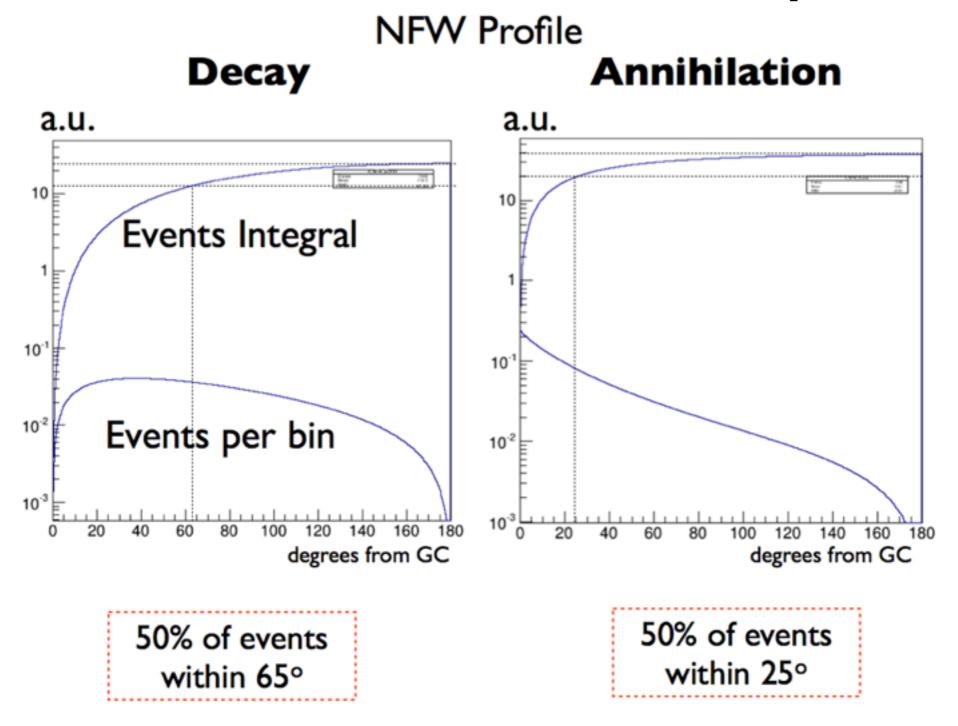
[Kohri, SCP, Rott (2013)]

$$\begin{split} \Gamma_{\rm Events} &\sim V L_{\rm MW} n_N \sigma_N \frac{\rho_{\rm DM}}{m_{\rm DM}} \Gamma_{\rm DM} \\ \sigma_N &\sim 9 \times 10^{-34} {\rm cm}^2 & M_{\rm DM} \sim 2 \ {\rm PeV} \\ n_N &\simeq n_{\rm Ice} &\simeq 5 \times 10^{23} / {\rm cm}^3 \\ \rho_{\rm DM} &\simeq 0.4 {\rm GeV} / {\rm cm}^3, & V \simeq 1 {\rm km}^3 \\ v &\sim 10^{-3}, & L_{\rm MW} &\simeq 10 {\rm kpc} \end{split}$$

 $\tau_{\rm DM}\simeq 1.9 N_{\nu}\times 10^{28} {\rm sec}$ 

## Directional information

[Kohri, SCP, Rott (2013)]



Ann VS Decay

[Kohri, SCP, Rott (2013)]

## Annihilating $\chi \chi \rightarrow \nu_L + X(\rightarrow \nu + \cdots)$

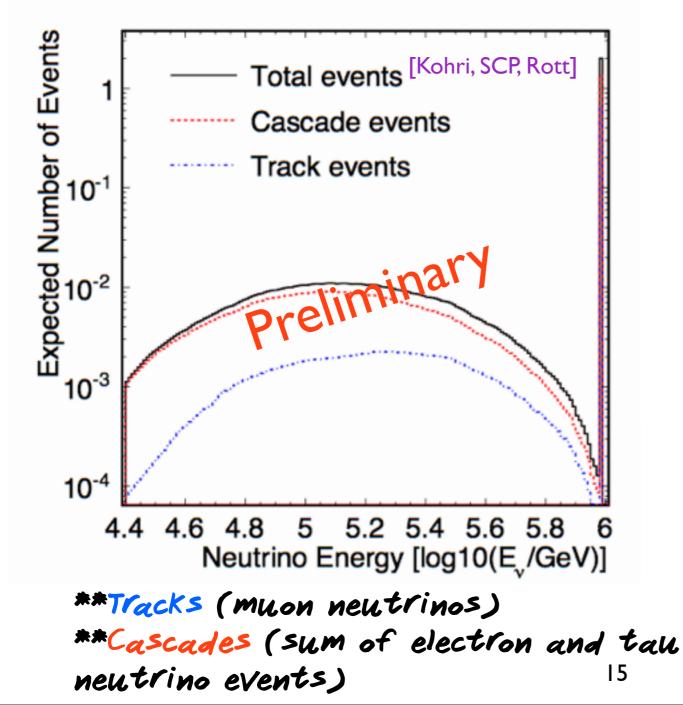
-less than one event/100 years with PeV DMs -centered (50% within 25°)

## Decaying (preferred)

 $au_{\chi} \sim 10^{28-29} {
m sec}$  would fit the "peak" -broadly distributed (50% within 65°)

A simple case

## We consider a simple decay $\chi \rightarrow \nu_L + H$ and found it can fit the observation pretty well!



-peak by VL -continuum by nu from Higgs decay

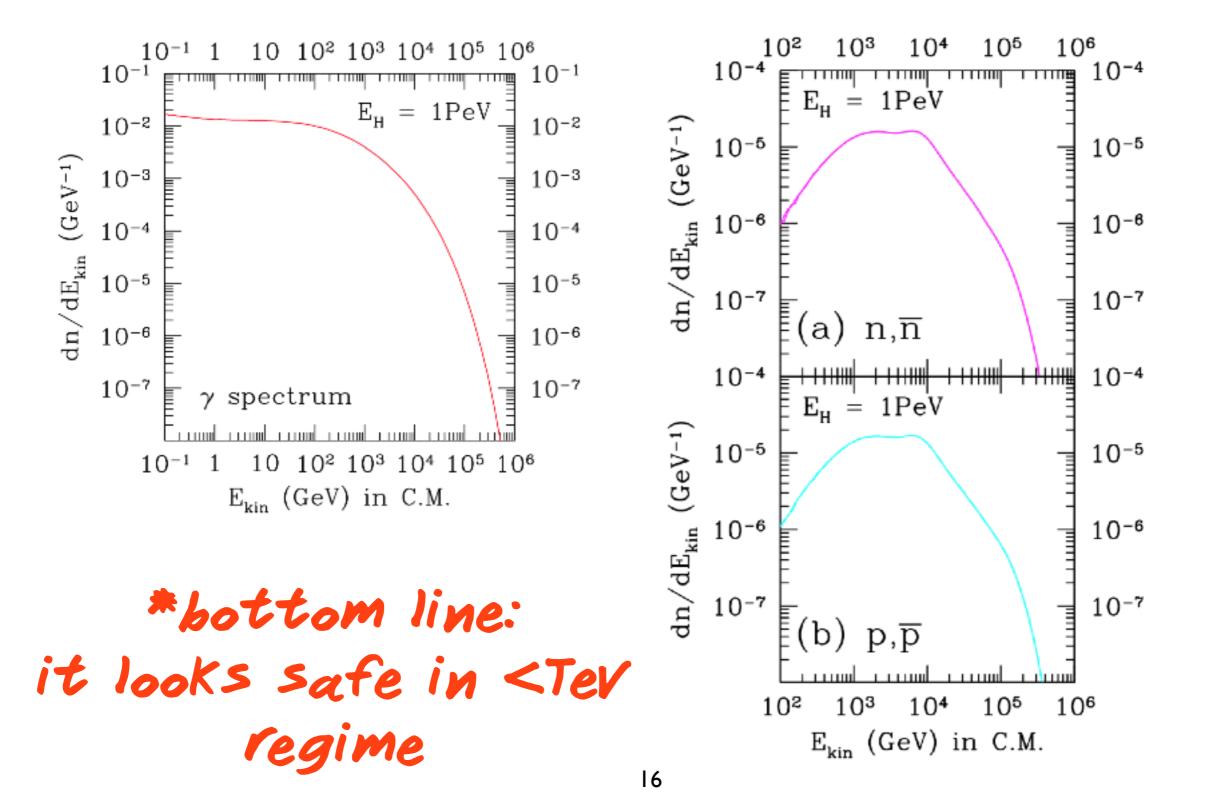
> $m_{\chi} = 2 \mathrm{PeV}$  $au_{\chi} = 9.7 \times 10^{28} \mathrm{sec}$ gives

 $N_{\nu}(\text{PeV}) = 2.04$ 

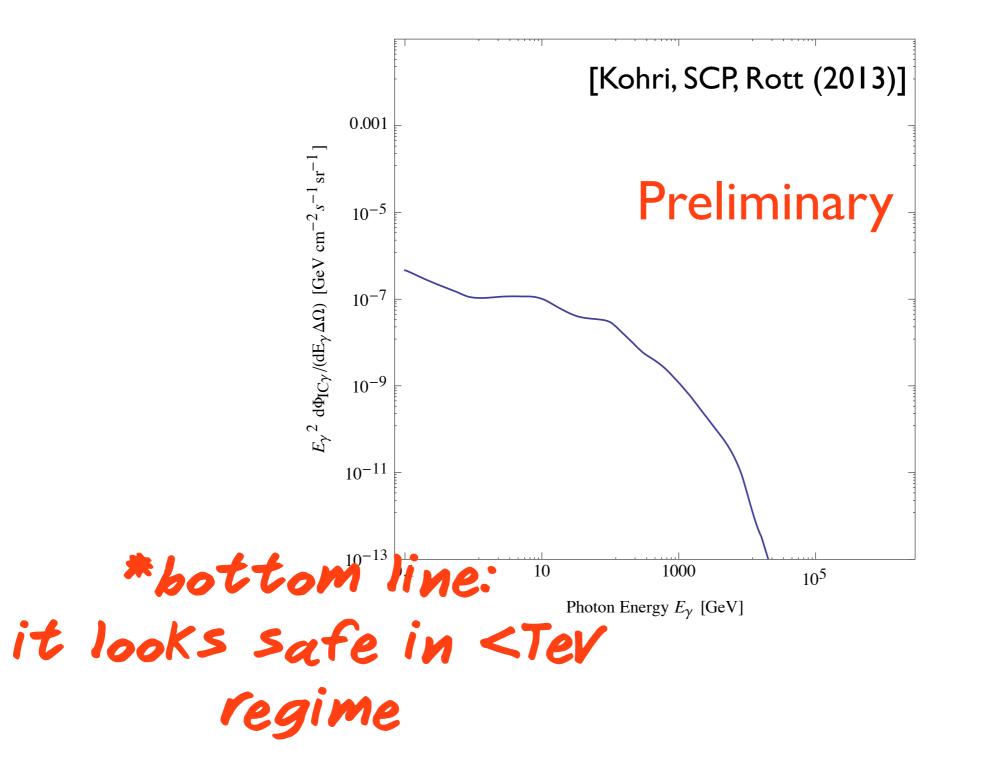
### with continuum

[Kohri, SCP, Rott (2013)]

contributions to CR



# gamma by ICS



Model building  

$$\mathcal{L} = y\bar{\nu}Hn + \overline{(n^{c},\chi)} \begin{pmatrix} M_{n} & \sigma \\ \sigma & M_{\chi} \end{pmatrix} \begin{pmatrix} n \\ \chi \end{pmatrix}$$

We can arrange <u>seesaw mechanism + small mixing in n & DM</u> such that DM can decay to neutrino + Higgs with a suppressed rate

$$\Gamma_{\chi \to \nu_L + H} = \frac{(y\epsilon)^2}{8\pi} M_-$$
  

$$\epsilon \approx -\frac{\sigma}{M_n - M_{\chi}} \ll 1$$
  

$$M_- \approx \frac{1}{2} (M_n + M_{\chi}) - \sqrt{\delta^2 + \sigma^2}, \delta = \frac{1}{2} (M_n - M_{\chi})$$

# Summary

- 1/R~O(100)GeV is fine in NMUED
   framework!
- Pev neutrinos at IceCube can be explained by a decaying DM model based on seesaw + small mixing.