

Observables and initial conditions from exact rotational hydro solutions

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Continuation: observables

Two different family of equations of state

Single particle spectra

Elliptic and higher order flows

Oscillations of HBT radii

What have we learnt?

Summary

Manuscript in preparation

[arXiv:1309.4390v2](https://arxiv.org/abs/1309.4390v2)

Rotated Gaussian solutions

$$n(t, \mathbf{r}') = n_0 \frac{V_0}{V} \exp\left(-\frac{r'_x{}^2}{2X^2} - \frac{r'_y{}^2}{2Y^2} - \frac{r'_z{}^2}{2Z^2}\right),$$

$$\mathbf{v}'(t, \mathbf{r}') = \left(\frac{\dot{X}}{X} r'_x, \frac{\dot{Y}}{Y} r'_y, \frac{\dot{Z}}{Z} r'_z\right),$$

Note: (r', k') in fireball frame
rotated wrt lab frame (r, k)

$$\ddot{X}X - X^2\omega^2 = \ddot{Y}Y = \ddot{Z}Z - Z^2\omega^2 = \frac{T}{m},$$

$$\dot{T} \frac{d}{dT}(\kappa T) + T \left(\frac{\dot{X}}{X} + \frac{\dot{Y}}{Y} + \frac{\dot{Z}}{Z}\right) = 0,,$$

$$X = Z \equiv R$$

$$\dot{X} = \dot{Z} \equiv \dot{R}$$

Note: in this talk, rotation in the
 (X, Z) impact parameter plane !

Role of EoS on acceleration

$$\ddot{X}X - X^2\omega^2 = \ddot{Y}Y = \ddot{Z}Z - Z^2\omega^2 = \frac{1}{1 + \kappa}.$$

$$\ddot{X}X - X^2\omega^2 = \ddot{Y}Y = \ddot{Z}Z - Z^2\omega^2 = \frac{T}{m},$$

Lattice QCD type EoS is explosive

$$\theta(t) = \theta_0 + \int dt \omega(t),$$

$$\omega = \omega_0 \frac{R_0^2}{R^2},$$

$$R = X = Z \neq Y,$$

$$s = s_T + s_Z = \frac{r_x^2 + r_z^2}{R^2} + \frac{r_y^2}{Y^2},$$

$$V = XYZ = R^2Y.$$

Tilt angle θ integrates rotation in (X,Z) plane: sensitive to EoS!

Single particle spectra

$$E \frac{d^3n}{d\mathbf{k}'} \propto E \exp \left(-\frac{k'_x{}^2}{2mT'_x} - \frac{k'_y{}^2}{2mT'_y} - \frac{k'_z{}^2}{2mT'_z} \right),$$

$$T'_x = T_f + m(\dot{X}_f^2 + \omega^2 Z^2)$$

$$T'_y = T_f + m\dot{Y}_f^2$$

$$T'_z = T_f + m(\dot{Z}_f^2 + \omega^2 X^2),$$

In the rest frame of the fireball:

Rotation increases effective temperatures both in the longitudinal and impact parameter direction

in addition to Hubble flows

Directed, elliptic and other flows

From the single particle spectra
→ flow coefficients v_n

$$\frac{d^3n}{dk_z k_t dk_t d\phi} = \frac{d^2n}{2\pi dk_z k_t dk_t} \left[1 + 2 \sum_{n=1}^{\infty} v_n \cos(n\phi) \right].$$

$$\begin{aligned}v_1 &= 0, \\v_2 &= \frac{I_1(w)}{I_0(w)}, \\v_3 &= 0, \dots \\v_{2n} &= \frac{I_n(w)}{I_0(w)},\end{aligned}$$

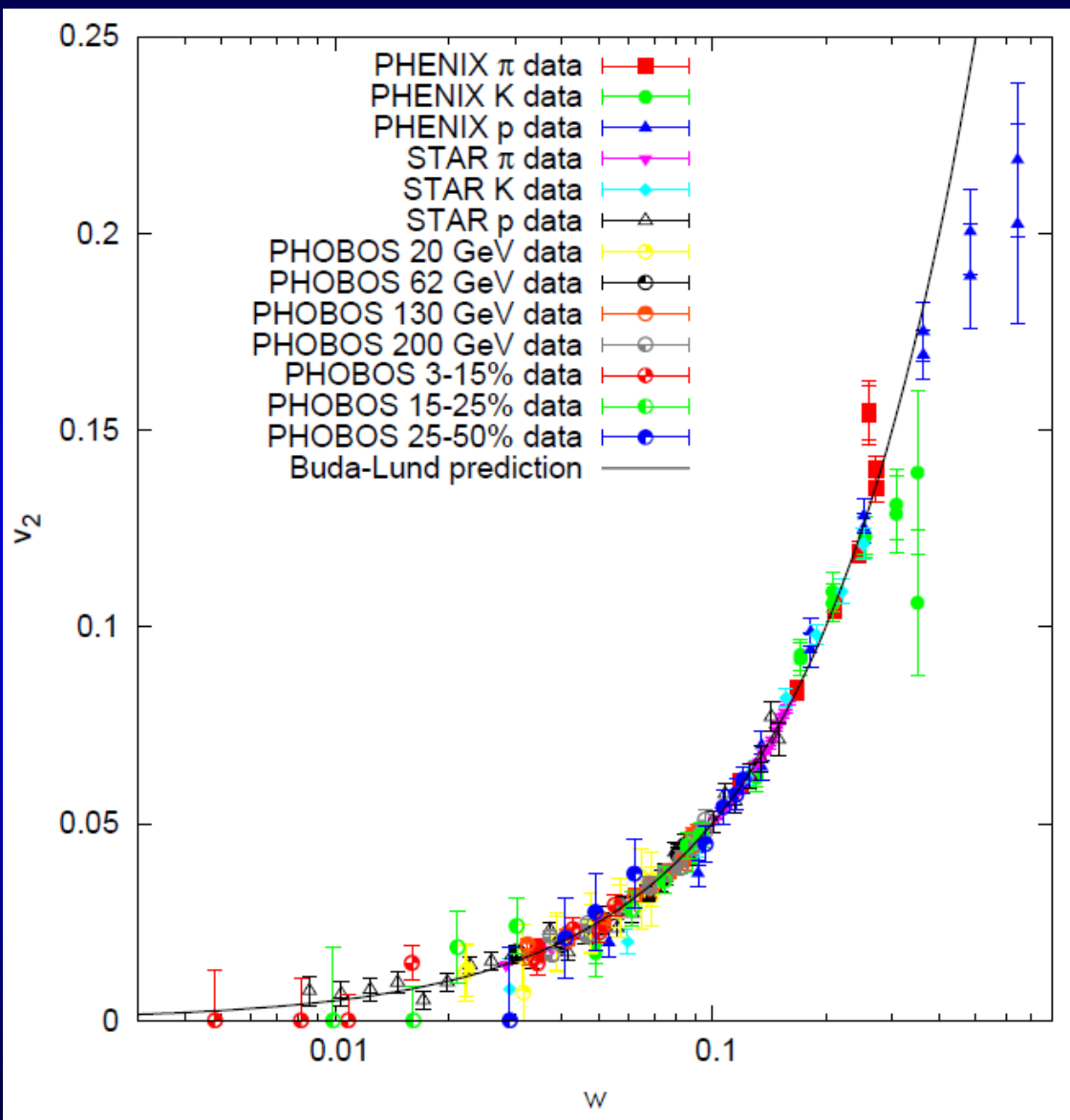
Note:
model is fully analytic

As of now, only $X = Z = R(t)$
spheroidal solutions are found

→ vanishing odd order flows

See next talk for fluctuations

Reminder: Universal w scaling of v_2



$$v_2 = \frac{I_1(w)}{I_0(w)},$$

Notes:

Rotation does not
change this

Just modifies
Radial flow parameters

See
+for more details

HBT radii for rotating spheroids

Diagonal Gaussians in natural frame

$$C(\mathbf{K}', \mathbf{q}') = 1 + \lambda \exp \left(-q_x'^2 R_x'^2 - q_y'^2 R_y'^2 - q_z'^2 R_z'^2 \right),$$
$$\mathbf{K}' = \mathbf{K}'_{12} = 0.5(\mathbf{k}'_1 + \mathbf{k}'_2),$$

$$\mathbf{q}' = \mathbf{q}'_{12} = \mathbf{k}'_1 - \mathbf{k}'_2 = (q_x', q_y', q_z'),$$
$$R_x'^{-2} = X_f^{-2} \left(1 + \frac{m}{T_f} (\dot{X}_f^2 + Y_f^2 \omega^2) \right),$$
$$R_y'^{-2} = Y_f^{-2} \left(1 + \frac{m}{T_f} \dot{Y}_f^2 \right),$$
$$R_z'^{-2} = Z_f^{-2} \left(1 + \left(\frac{m}{T_f} \dot{Z}_f^2 + X_f^2 \omega^2 \right) \right).$$

New terms with blue

→ Rotation decreases HBT radii similarly to Hubble flow.

HBT radii in the lab frame

$$C_2(\mathbf{K}, \mathbf{q}) = 1 + \lambda \exp \left(- \sum_{i,j=s,o,l} q_i q_j R_{ij}^2 \right),$$

$$R_s^2 = R_y'^2 \cos^2 \phi + R_x^2 \sin^2 \phi,$$

$$R_o^2 = R_x^2 \cos^2 \phi + R_y'^2 \sin^2 \phi + \beta_t^2 \Delta t^2,$$

HBT radii without flow

$$R_l^2 = R_z'^2 \cos^2 \theta + R_x'^2 \sin^2 \theta + \beta_1^2 \Delta t^2,$$

$$R_{ol}^2 = (R_x'^2 - R_z'^2) \cos \theta \sin \theta \cos \phi + \beta_t \beta_1 \Delta t^2,$$

$$R_{os}^2 = (R_x^2 - R_y'^2) \cos \phi \sin \phi,$$

$$R_{sl}^2 = (R_x'^2 - R_z'^2) \cos \theta \sin \theta \sin \phi,$$

$$R_x^2 = R_x'^2 \cos^2 \theta + R_z'^2 \sin^2 \theta$$

But spheroidal symmetry of the solution

Makes several cross terms vanish → need for ellipsoidal solutions

HBT radii in the lab frame

$$C_2(\mathbf{K}, \mathbf{q}) = 1 + \lambda \exp \left(- \sum_{i,j=s,o,l} q_i q_j R_{ij}^2 \right),$$

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$$R_o^2 = R_x^2 \cos^2 \phi + R_y'^2 \sin^2 \phi + \beta_t^2 \Delta t^2,$$

$$R_l^2 = R_z'^2 \cos^2 \theta + R_x^2 \sin^2 \theta + \beta_l^2 \Delta t^2,$$

$$R_{ol}^2 = (R_x^2 - R_z'^2) \cos \phi \sin \phi,$$

$$R_{os}^2 = (R_x^2 - R_y'^2) \cos \phi \sin \phi,$$

$$R_{sl}^2 = (R_x^2 - R_z'^2) \cos \theta \sin \theta,$$

$$R_x^2 = R_x'^2 \cos^2 \theta + R_z'^2 \sin^2 \theta + \beta_x^2 \Delta t^2,$$

HBT radii without flow

$$C_2(\mathbf{K}, \mathbf{q}) = 1 + \lambda \exp \left(- \sum_{i,j=s,o,l} q_i q_j R_{ij}^2 \right),$$

$$R_s^2 = R_y'^2 \cos^2 \phi + R_x^2 \sin^2 \phi,$$

$$R_o^2 = R_x^2 \cos^2 \phi + R_y'^2 \sin^2 \phi + \beta_t^2 \Delta t^2,$$

$$R_l^2 = 0 + \beta_l^2 \Delta t^2,$$

$$R_{ol}^2 = 0 + \beta_t \beta_l \Delta t^2,$$

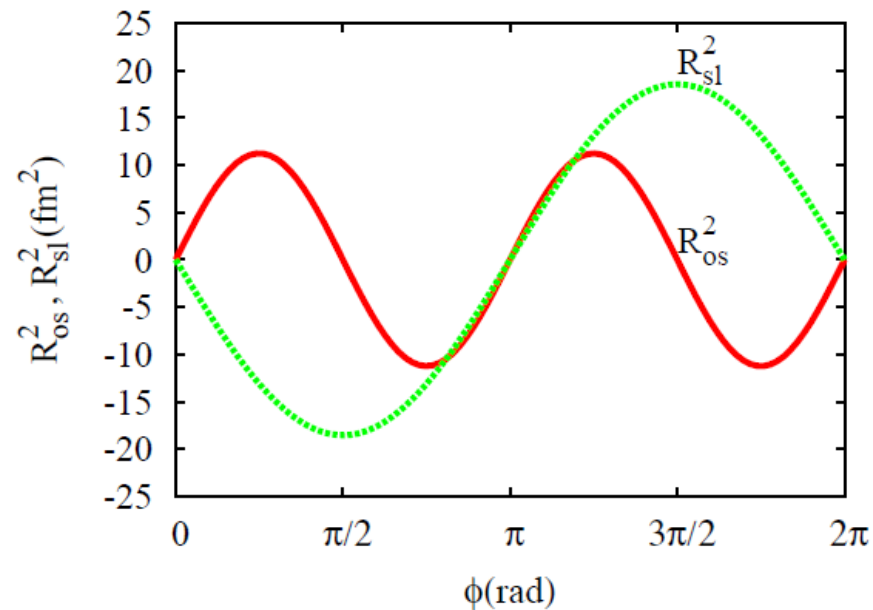
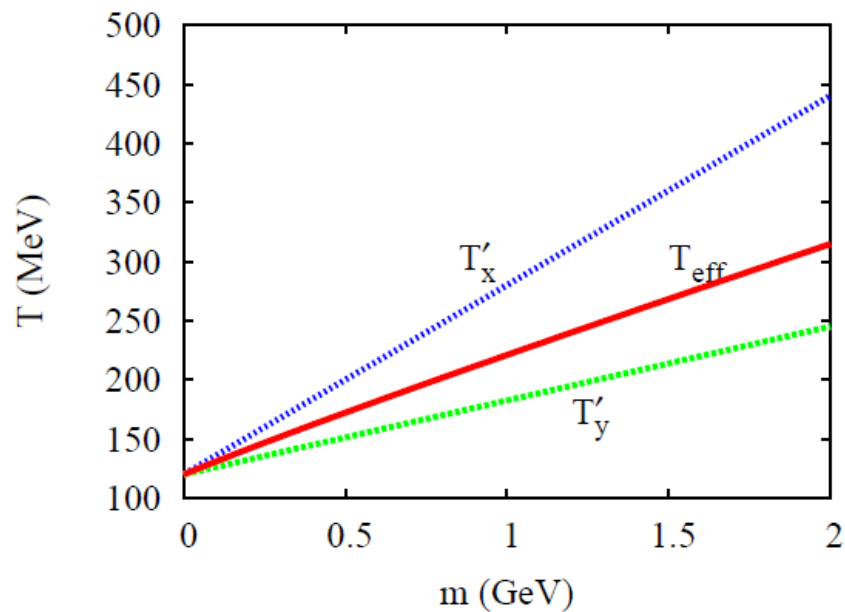
$$R_{os}^2 = (R_x^2 - R_y'^2) \cos \phi \sin \phi,$$

$$R_{sl}^2 = 0.$$

But spheroidal symmetry of the solution

Makes several cross terms vanish → need for ellipsoidal solutions

Qualitatively rotation and flow similar



But need more time
dependent calculations

and less academic studies
(relativistic solutions)

Summary

Observables calculated

Effects of rotation and flow

Combine and have same mass dependence

Spectra: slope increases

v_2 : universal w scaling remains valid

HBT radii:

Decrease with mass intensifies

Even for spherical expansions:

v_2 from rotation.

Picture: vulcano

How to detect the rotation?

Next step: penetrating probes

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Observables calculated

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How to detect the rotation?

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