

# Parametrization of $\gamma$ -ray production cross-sections for $pp$ interactions in a broad proton energy range from the kinematic threshold to PeV energies

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ISVHECRI 2014

# Outline

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- 3 Parametrization  $pp \rightarrow \pi^0$  cross section
- 4 Parametrization of  $pp \rightarrow \gamma$  differential cross section
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# Why a new parametrization?

We need an accurate parametrization that:

- 1 Can fit all available Experimental Data (especially at low energies)
- 2 Is able to fit the High Energy Hadronic models results (MC codes)
- 3 Connects smoothly High and Low energy regions

Motivation

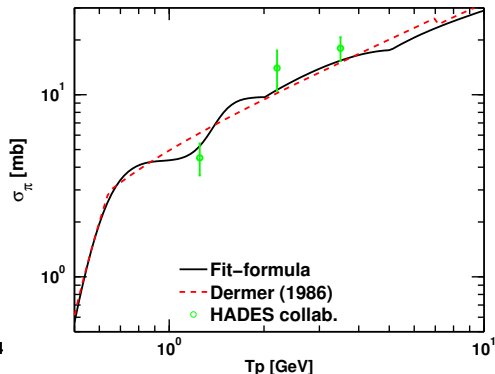
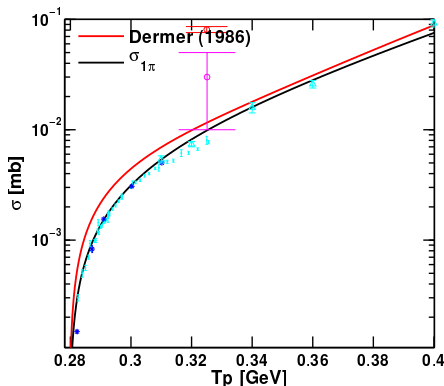
- Space-based  $\gamma$ -ray instruments probe with high accuracy  
 $E_\gamma \sim 100\text{MeV}$  (e.g. Fermi-LAT)
- Very hot astrophysical plasma applications (BH accretion disk)
- Astrophysical sources have wide range of proton energy and none of the current parametrizations can cover this energy interval accurately

## Widely used $pp \rightarrow \gamma$ models/parametrisation

- **Isobar + scaling model Dermer (1986)**  
(connects  $\Delta(1232)$  *Isobar + scaling model* calculations)
- **Parametrization from Kamae et.al. (2006),  $488\text{MeV} < T_p < 512\text{ TeV}$**   
(Uses  $\Delta(1232)$  and *res(1600) group of resonances + Blattnig et.al. (2000) parametrizations + PYTHIA 6*)
- **Parametrization from Kelner et.al. (2006),  $0.1 \leq T_p \leq 10^5\text{ TeV}$**   
(Parametrization of *QGSJET and SIBYLL* results)

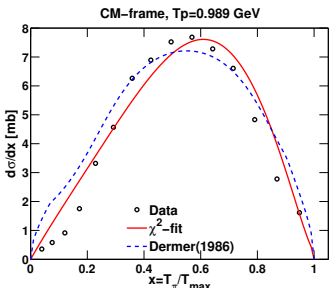
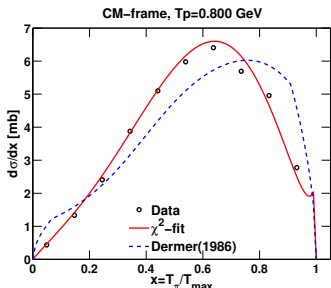
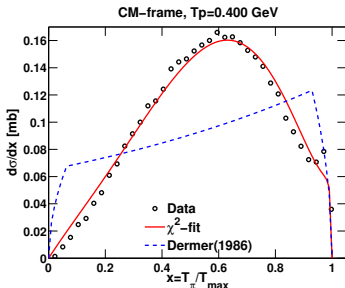
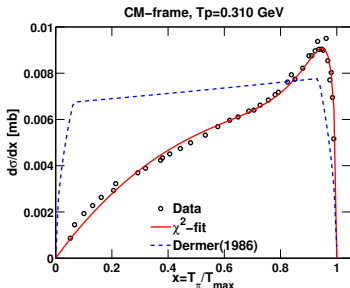
Exp. DATA  
V.S.  
Present Parametrizations

# Inclusive $\sigma_\pi$ , Dermer(1986) v.s. Exp. Data

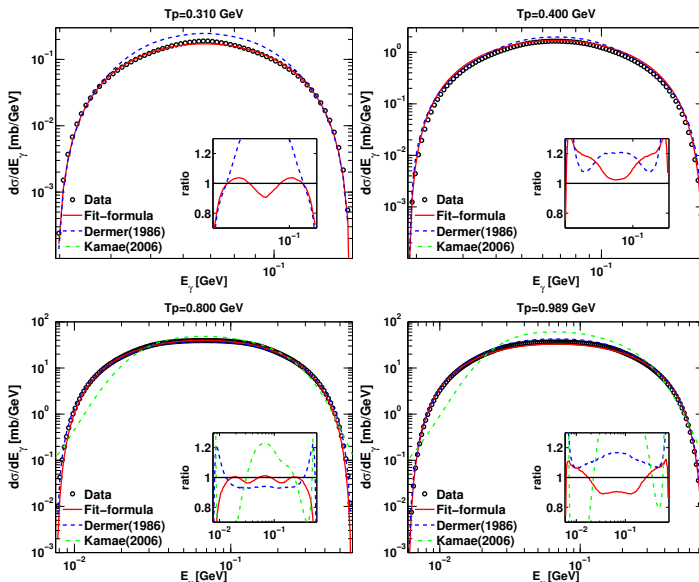


- 1)  $T_p \sim 0.3$  GeV, is 70-80 % higher (important for very hot plasma)
- 2)  $1 < T_p < 2$  GeV, is at least 20 % different and doesn't mimic data

# C.M. frame $\pi^0$ spectra, Dermer(1986) v.s. Exp. Data



# $\gamma$ -ray spectra, Exp.Data, Dermer(1986), Kamae(2006)

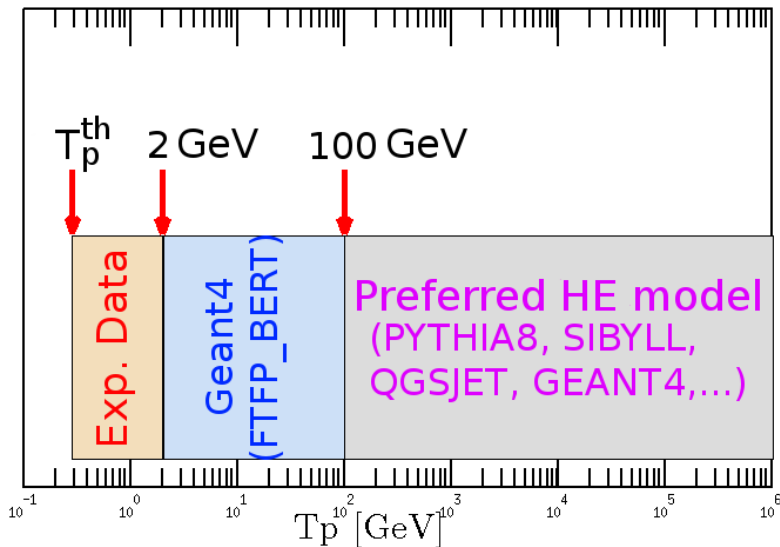




Thus, a parametrization is needed that:

- Revise the low energy  $\sigma(pp \rightarrow \gamma)$
- Expands to very high energies

## Models considered for the parametrization



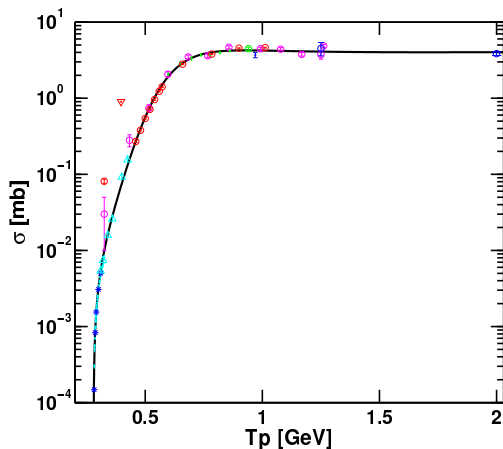
# Pion production Cross-Section $\sigma_\pi$

- For  $T_p \leq 2$  GeV; Dominate 1 & 2 $\pi$  production channels
  - $pp \rightarrow pp\pi^0$
  - $pp \rightarrow pp2\pi^0$
  - $pp \rightarrow \{pn/D\}\pi^+\pi^0$
- For  $T_p > 2$  GeV; Multi pion production (multiplicity from MC codes)

Cross section

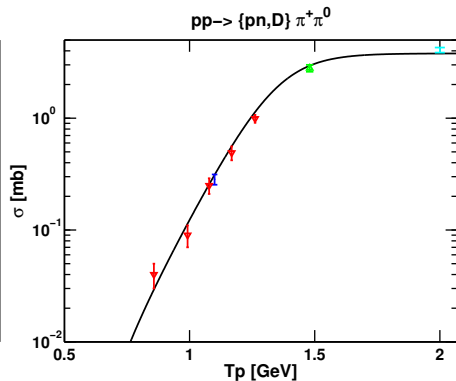
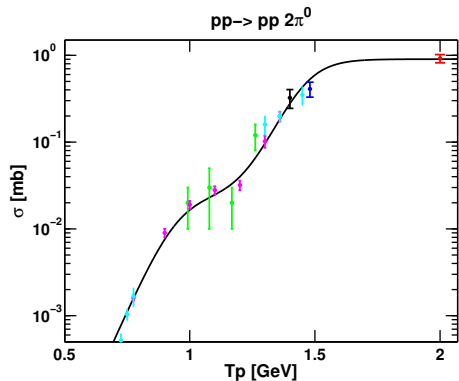
$$\sigma_\pi = \begin{cases} \sigma_{1\pi} + \sigma_{2\pi} & : (T_p^{\text{th}} < T_p \leq 2 \text{ GeV}) \\ \sigma_{\text{inel}} \times \langle n_{\pi^0} \rangle & : (T_p > 2 \text{ GeV}) \end{cases} \quad (1)$$

# Single pion production channel



$$\sigma_{1\pi} = \sigma_0 \times \eta^{1.95} (1 + \eta + \eta^5) \times [f_{BW}(\sqrt{s})]^{1.86} \quad (2)$$

## Two pion production channels



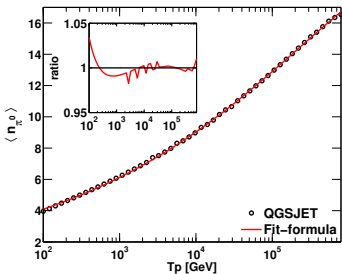
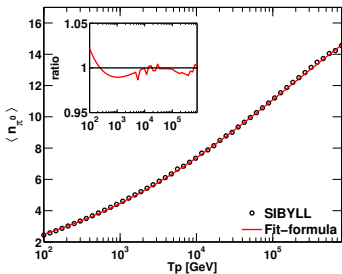
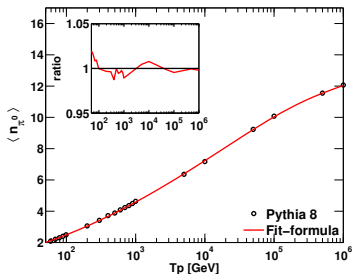
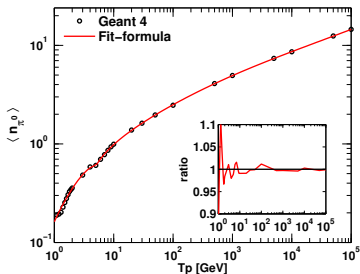
$$\sigma_{2\pi} = \frac{5.7 \text{ mb}}{1 + \exp(-9.3(T_p - 1.4))} \quad (3)$$

# Parametrization of $\langle n_{\pi^0} \rangle$ for Geant4, Pythia8, SIBYLL and QGSJET

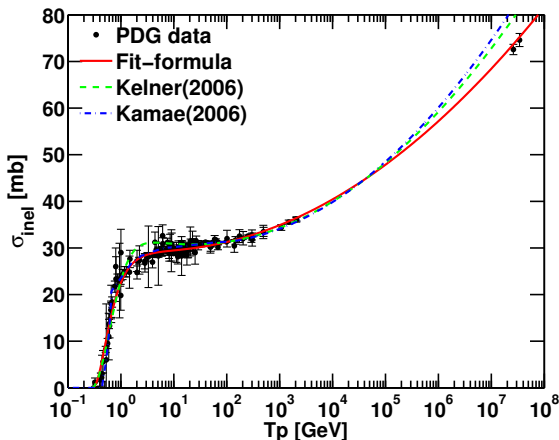
If  $Q_p = (T_p - T_p^{\text{th}})/m_p$  and  $\xi_p = (T_p - 3 \text{ GeV})/m_p$

$$\langle n_{\pi^0} \rangle = \begin{cases} a_1 + a_2 Q_p + a_3 Q_p^2 & : 2 < T_p < 5 \text{ GeV} \\ b_1 \xi_p^{b_4} [1 + \exp(-b_2 \xi_p^{b_5})] [1 - \exp(-b_3 \xi_p^{1/4})] & : T_p \geq 5 \text{ GeV} \end{cases} \quad (4)$$

# $\langle n_{\pi^0} \rangle$ for Geant4, Pythia8, SIBYLL and QGSJET



# p-p inelastic cross section



$$\sigma_{\text{inel}} = \left( 30.7 - 0.96 \log \left( \frac{T_p}{T_p^{\text{th}}} \right) + 0.18 \log^2 \left( \frac{T_p}{T_p^{\text{th}}} \right) \right) \times \left[ 1 - \left( \frac{T_p^{\text{th}}}{T_p} \right)^{1.9} \right]^3 \quad (5)$$



# Parametrization of the $\gamma$ -ray spectrum

$$\frac{d\sigma}{dE_\gamma}(T_p, E_\gamma) = A_{\max}(T_p) \times F(T_p, E_\gamma) \quad (6)$$

$$A_{\max}(T_p) = b_1 \times \theta_p^{-b_2} \times \exp(-b_3^2 \times \log^2(\theta_p)) \times \frac{\sigma_\pi(T_p)}{m_p} \quad (7)$$

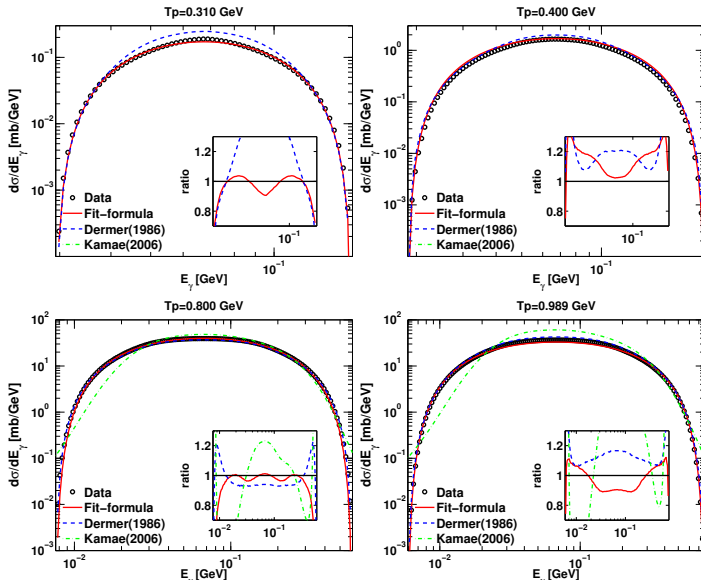
$$Y_\gamma = E_\gamma + \frac{m_\pi^2}{4E_\gamma}, \quad X_\gamma = \frac{Y_\gamma - m_\pi}{Y_\gamma^{\max} - m_\pi}$$

$$F(T_p, E_\gamma) = (1 - X_\gamma)^{\kappa(T_p)}, \quad \text{for } T_p^{\text{th}} \leq T_p < 1 \text{ GeV},$$

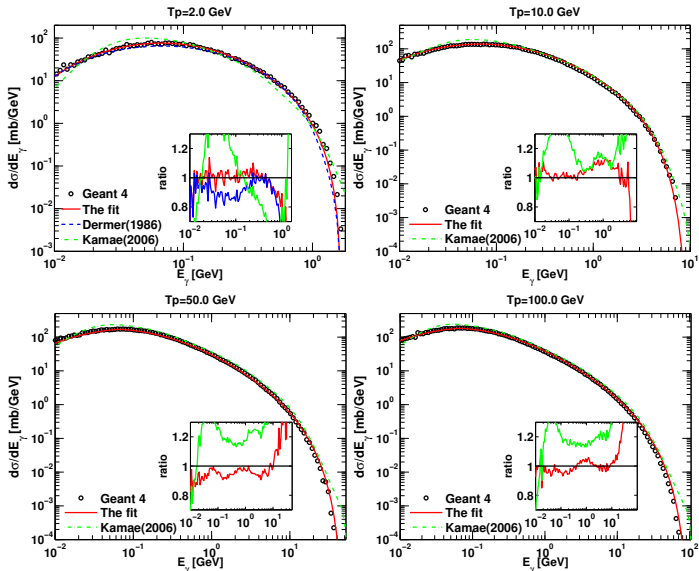
$$F(T_p, E_\gamma) = \frac{(1 - X_\gamma)^{\beta(T_p)}}{\left(1 + \frac{X_\gamma}{C}\right)^{\alpha(T_p)}}, \quad \text{for } 1 \leq T_p < 20 \text{ GeV}, \quad (8)$$

$$F(T_p, E_\gamma) = \frac{(1 - \sqrt{X_\gamma})^\delta}{\left(1 + \frac{X_\gamma}{C}\right)}, \quad \text{for } T_p \geq 20 \text{ GeV}$$

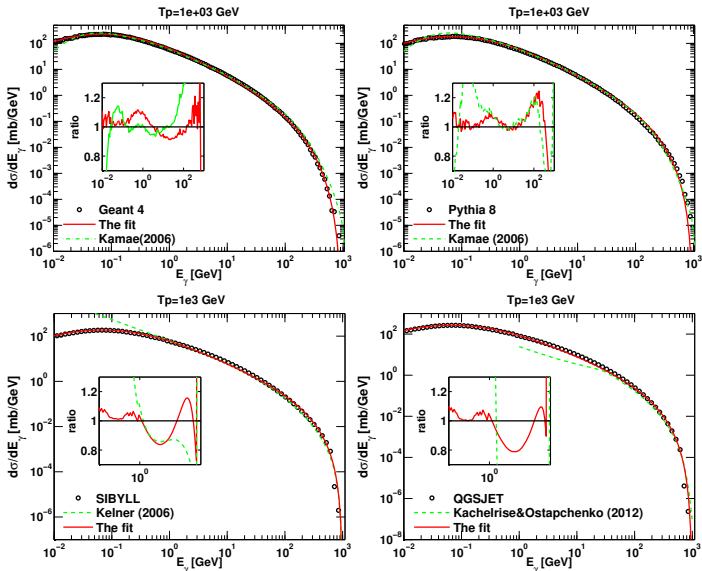
# Gamma-ray spectra for $T_p < 1$ GeV



# Geant4 $\gamma$ -ray spectra for $1 \leq T_p \leq 100$ GeV



# Gamma-ray spectra fits at $T_p = 1$ TeV



# Summary

We have a new parametrization for  $pp \rightarrow \gamma$  production cross section

- It covers the energy  $T_p^{\text{th}} \leq T_p \leq 1 \text{ PeV}$
- Simple & Accurate formulas ( $< 20 \%$ )
- Flexible to switch between different Hadronic models

For more details  $\Rightarrow$  [arXiv:1406.7369](https://arxiv.org/abs/1406.7369) [astro-ph].