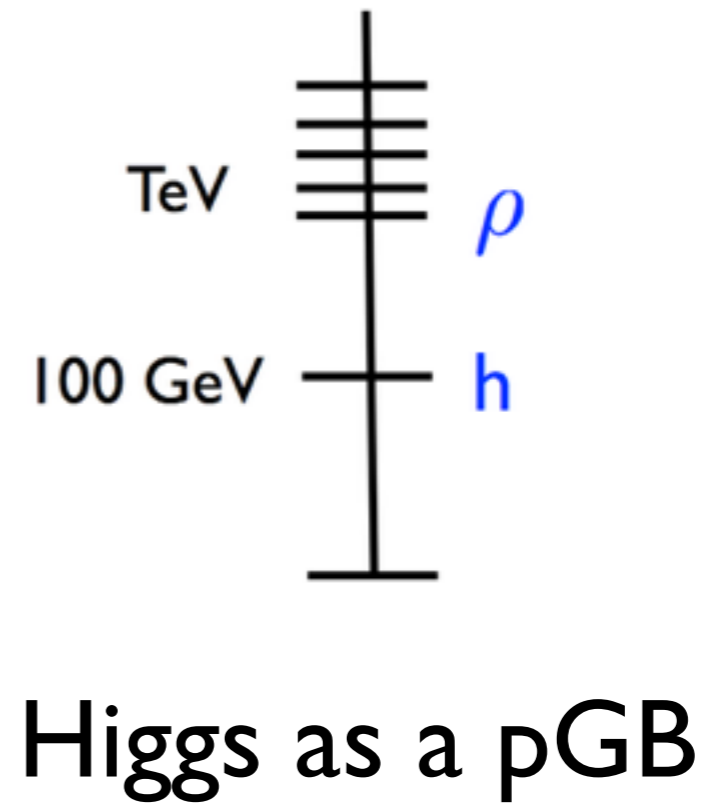
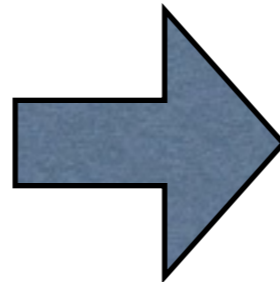
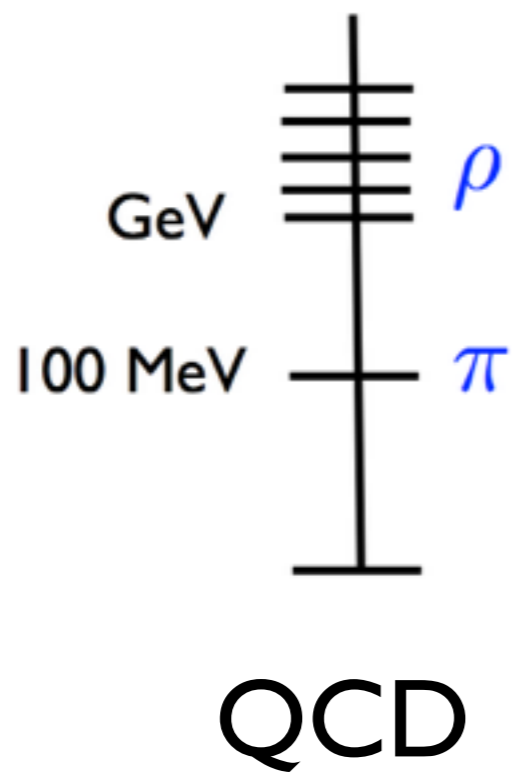


Beyond the Standard Model 2

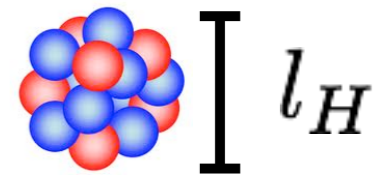
Andreas Weiler
(DESY&CERN)



CERN school
2014/6/29

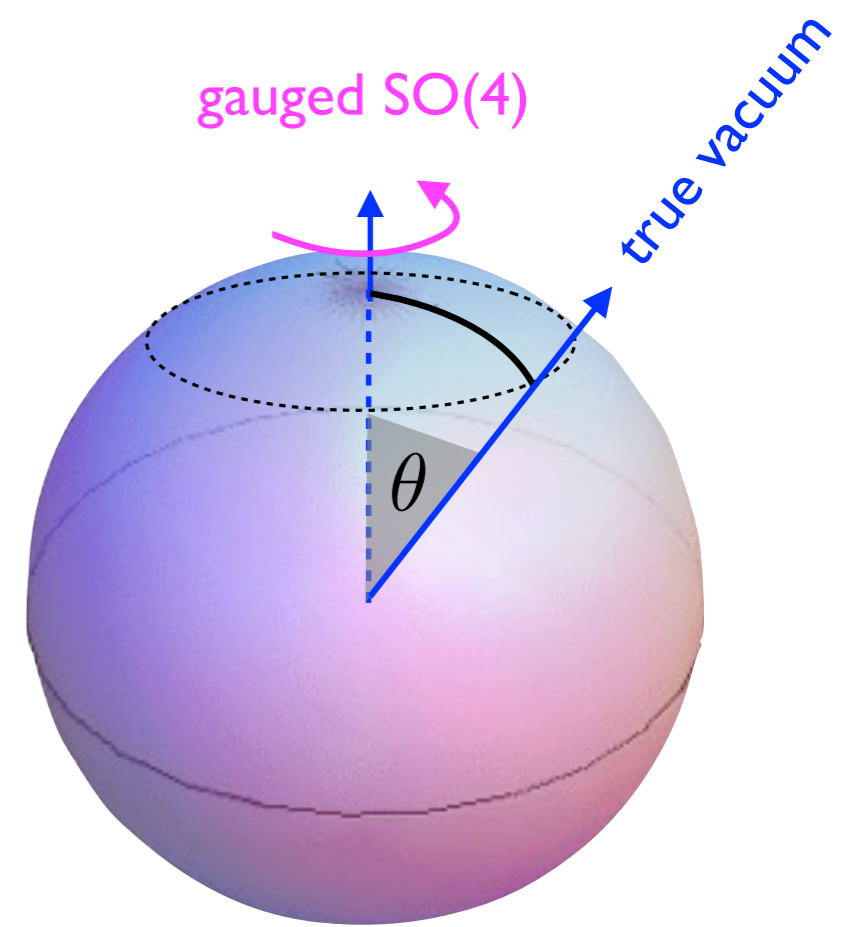


Composite Higgs



- Higgs is a hadron of a new strong force
- Solves the hierarchy problem (like QCD)
- Higgs is a pseudoGoldstone that's why it is lighter than the other resonances

$SO(5)/SO(4)$



Tree level: gauge $SO(4)$ aligned

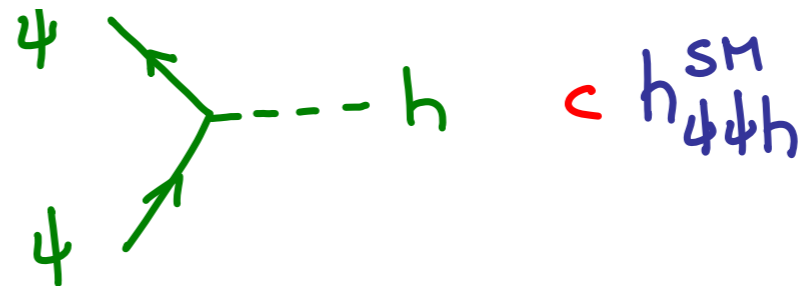
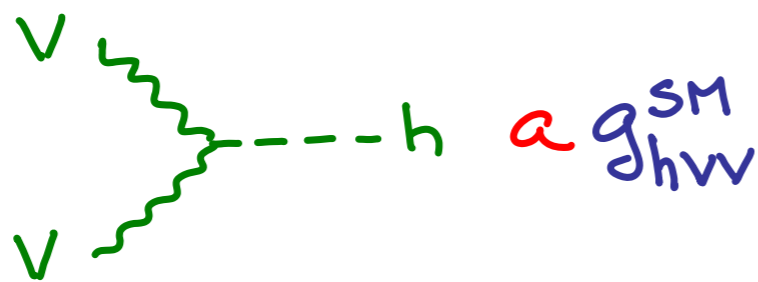
Higgs

$$\phi = e^{i\pi \hat{a} T^{\hat{a}} / f} \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \\ 1 \end{pmatrix} = \begin{pmatrix} \sin(\pi/f) \times \begin{pmatrix} \hat{\pi}^1 \\ \hat{\pi}^2 \\ \hat{\pi}^3 \\ \hat{\pi}^4 \end{pmatrix} \\ \cos(\pi/f) \end{pmatrix} \stackrel{\text{I-loop } \langle \phi(x) \rangle = \theta \cdot f}{=} \begin{pmatrix} \sin(\theta + h(x)/f) e^{i\chi^i(x) A^i / v} \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix} \\ \cos(\theta + h(x)/f) \end{pmatrix}$$

eaten by W_L, Z_L

Higgs couplings

Have been measured to 20-30% precision



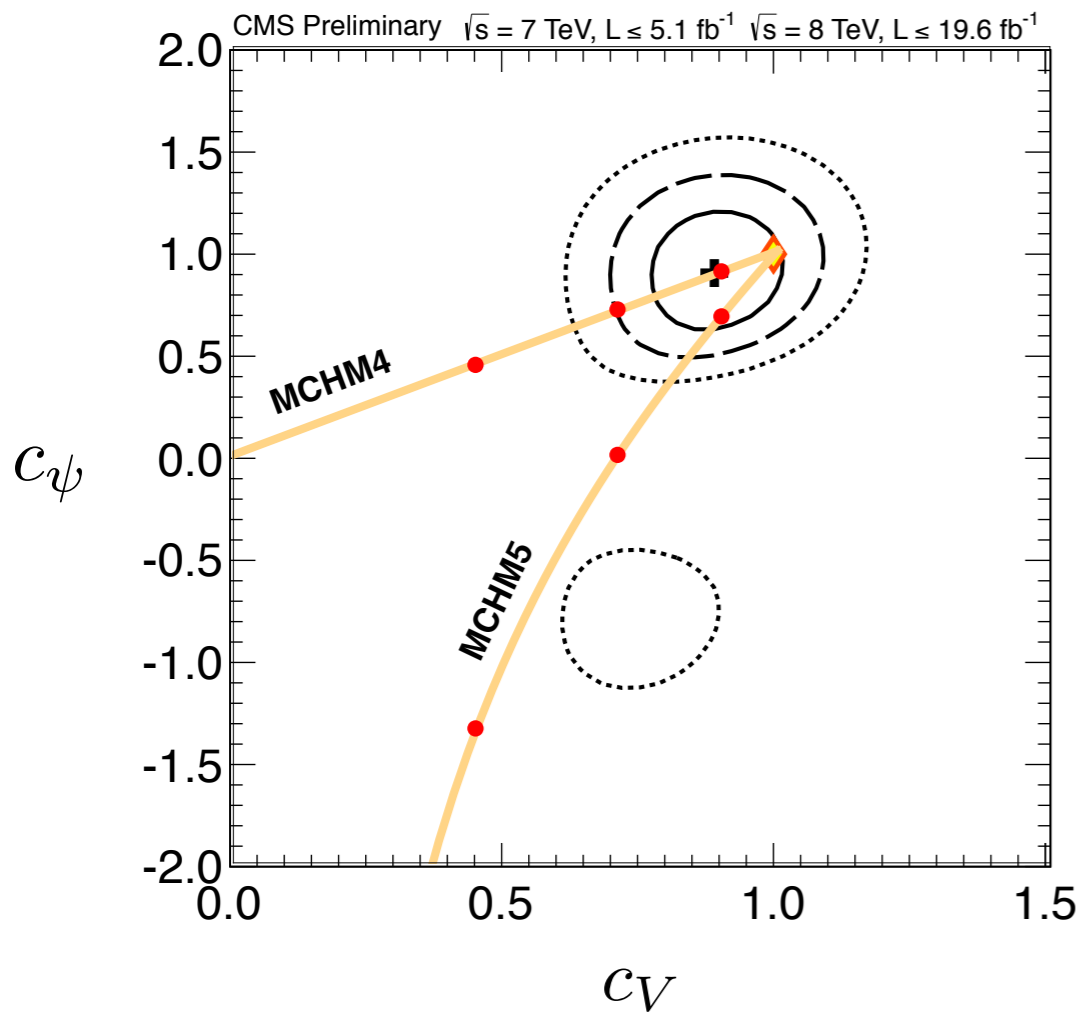
Expect deviations $\sim (v/f)^2$

$$\xi \equiv \frac{v^2}{f^2}$$

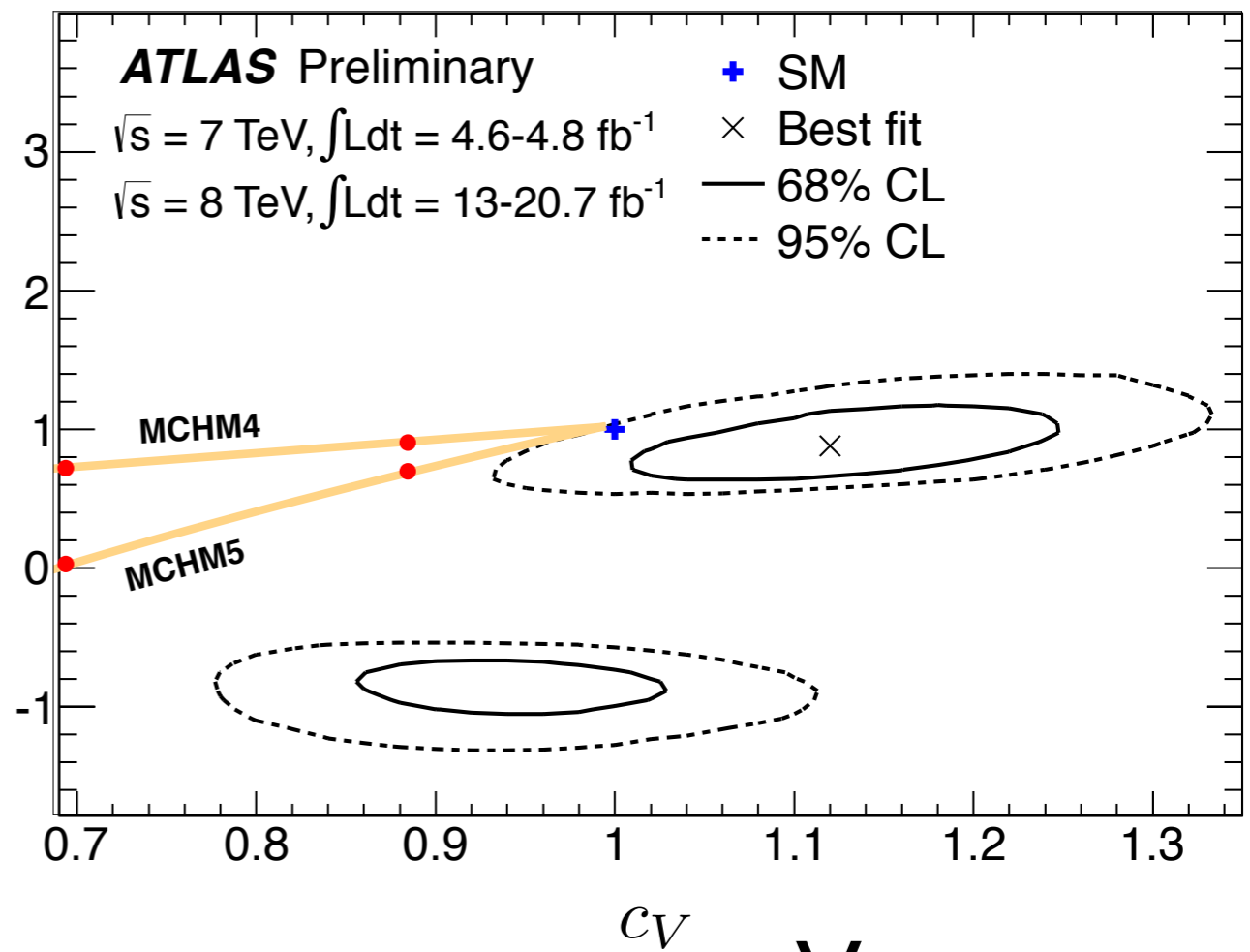
$$a = \sqrt{1 - \xi}$$

$$c_f = \frac{1 - (1 + n)\xi}{1 - \xi}$$

Higgs couplings



Fermion



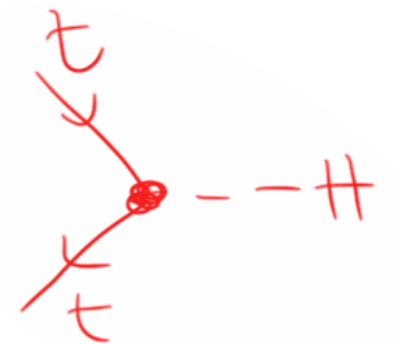
Red points at $\xi \equiv (v/f)^2 = 0.2, 0.5, 0.8$

Vector

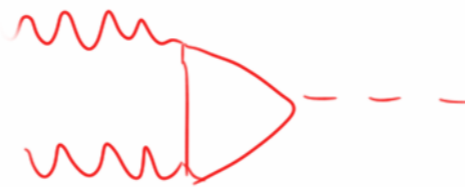
Higgs couplings

$$\mathbf{SM} + \mathcal{L} = \frac{\alpha_s c_g}{12\pi} |H|^2 G_{\mu\nu}^a{}^2 + \frac{\alpha c_\gamma}{2\pi} |H|^2 F_{\mu\nu} + y_t c_t \bar{q}_L \tilde{H} t_R |H|^2$$

$$\frac{\sigma(gg \rightarrow h)}{\text{SM}} = (1 + (c_g - c_t)v^2)^2$$



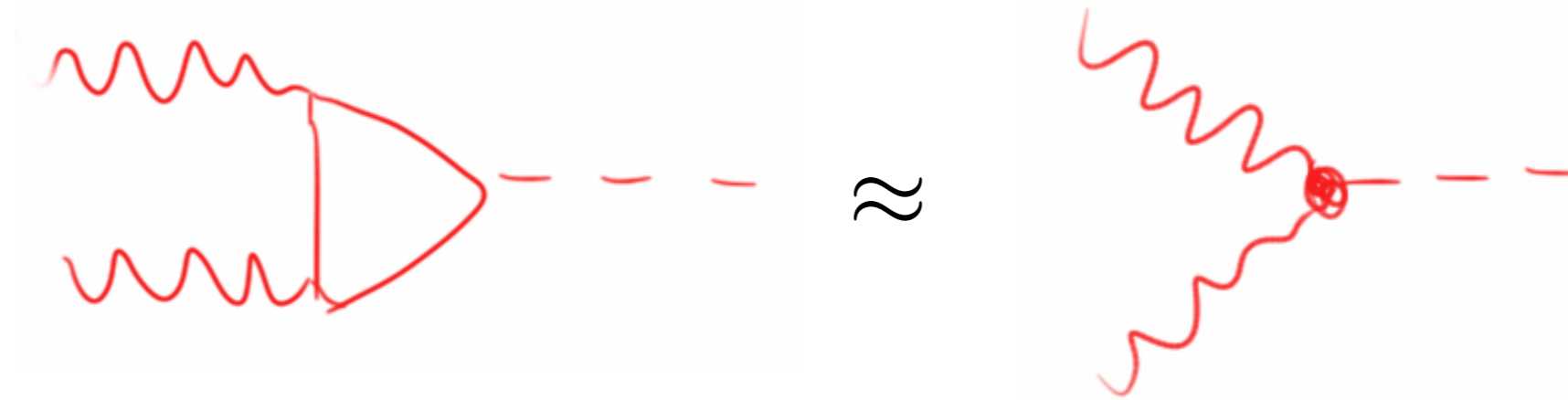
Degeneracy ‘short-distance’ vs ‘long-distance’



E.g. fermionic top partners MCHM: $\Delta c_t = \Delta c_g$

$$\sigma(pp \rightarrow H + X)_{\text{inclusive}}$$

Does not resolve short-distance physics

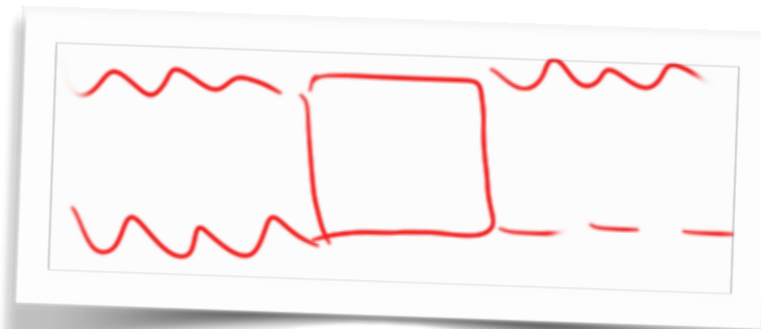
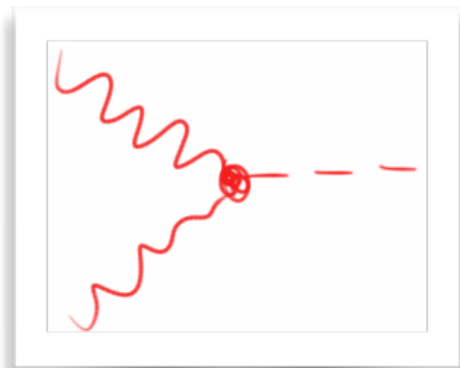


$m_H(\text{GeV})$	$\frac{\sigma_{NLO}(m_t)}{\sigma_{NLO}(m_t \rightarrow \infty)}$	$\frac{\sigma_{NLO}(m_t, m_b)}{\sigma_{NLO}(m_t \rightarrow \infty)}$
125	1.061	0.988
150	1.093	1.028
200	1.185	1.134

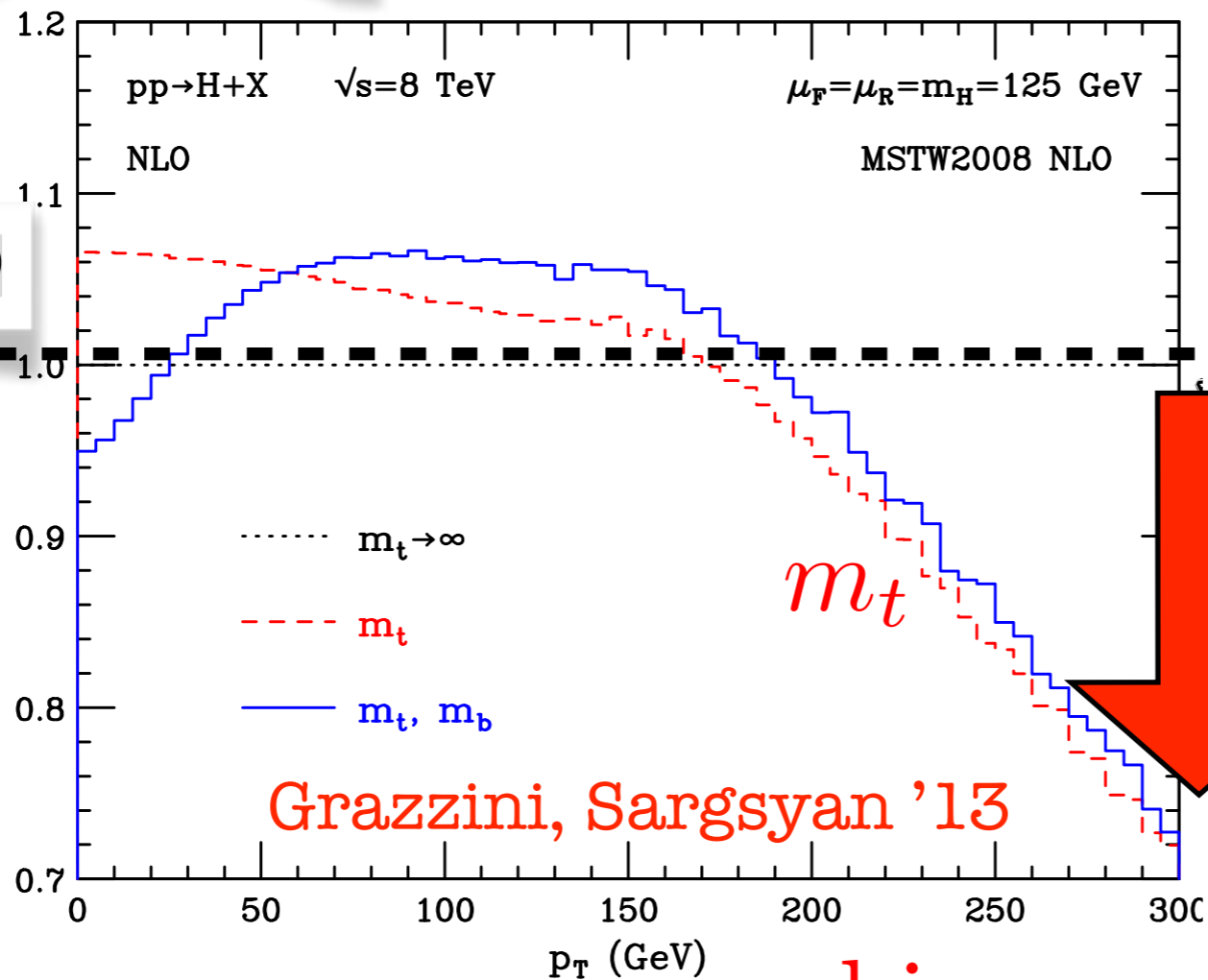
e.g. [1306.4581](#)

Beyond current observables

Cut the loop open, recoil against hard jet



$$p_T \gg m_t$$



Baur, Glover '90,
Langenegger et. al '06,
1308.4771

high p_T tail resolves
loop dynamics

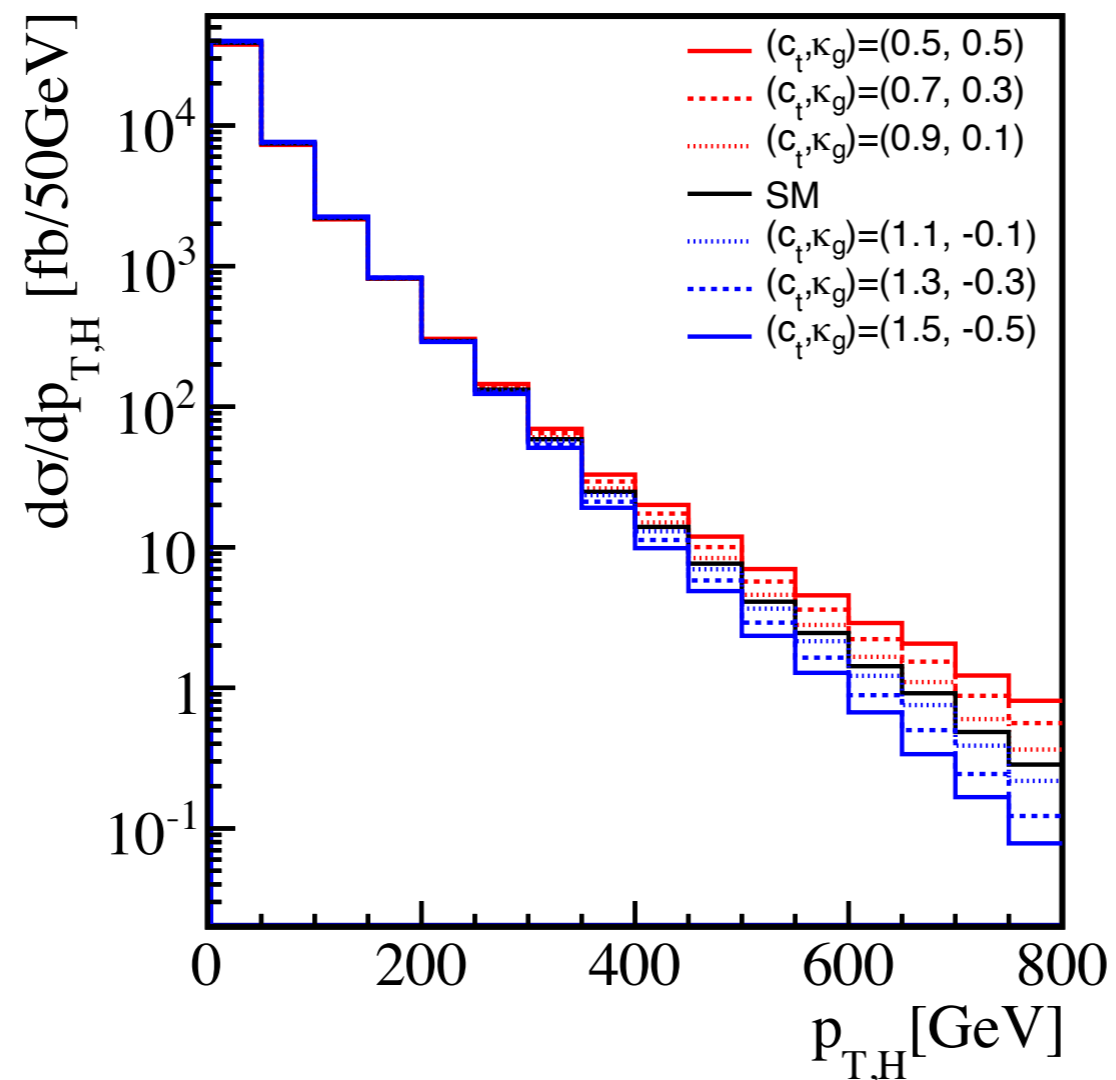
Grazzini, Sargsyan '13

$higgs - p_T$

$$m_t \rightarrow \infty$$

Measurement how-to

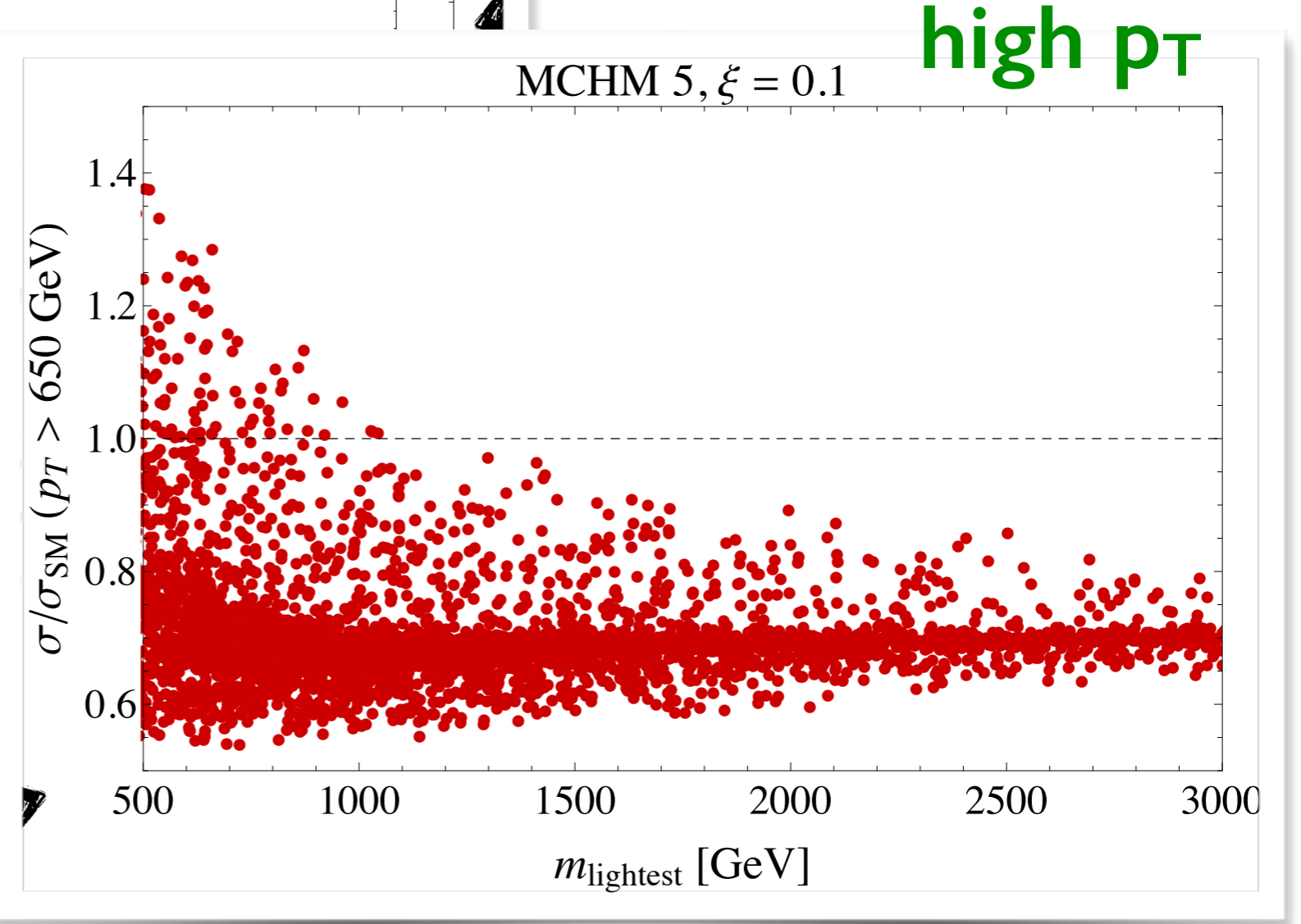
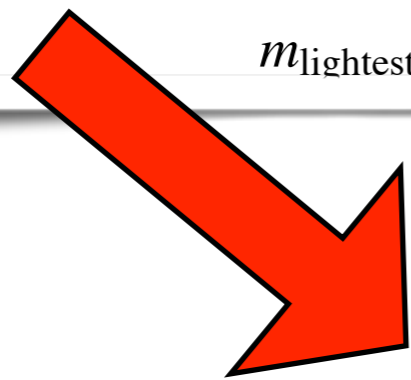
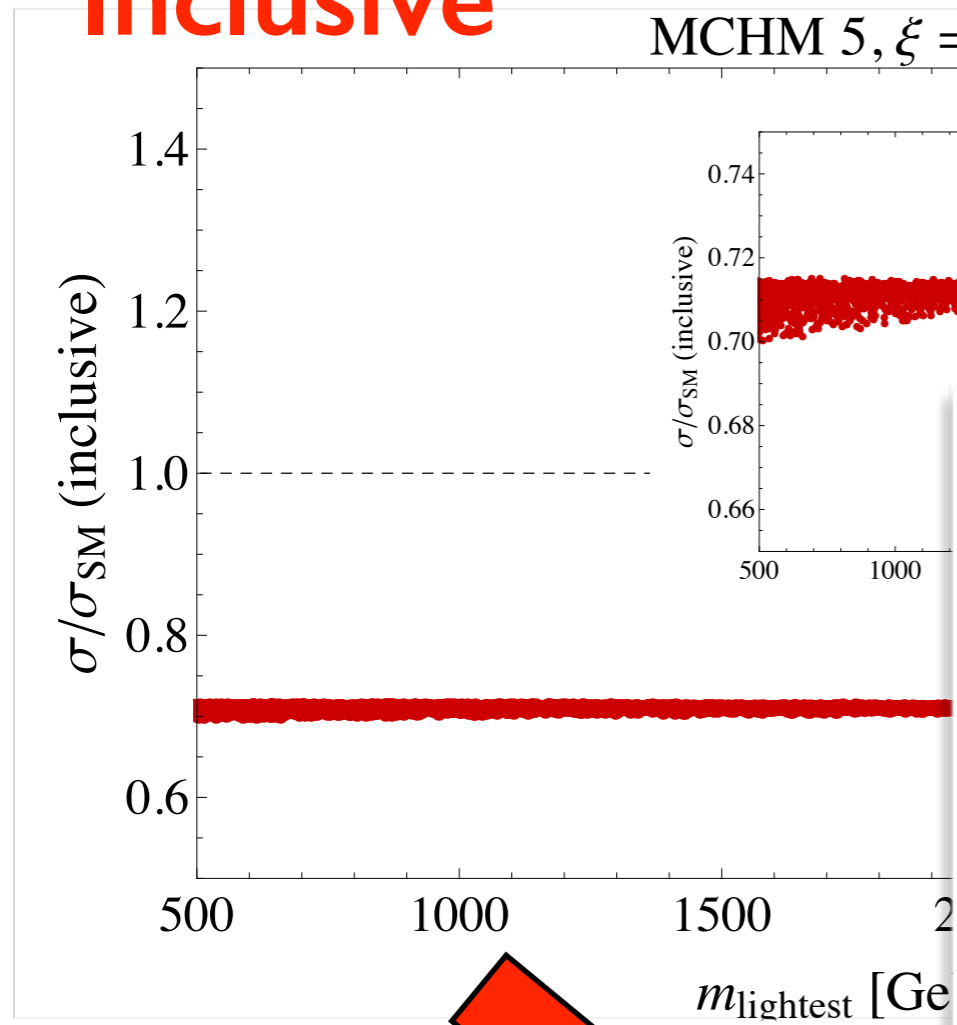
worst case: inclusive cross-section = SM



Top partner example

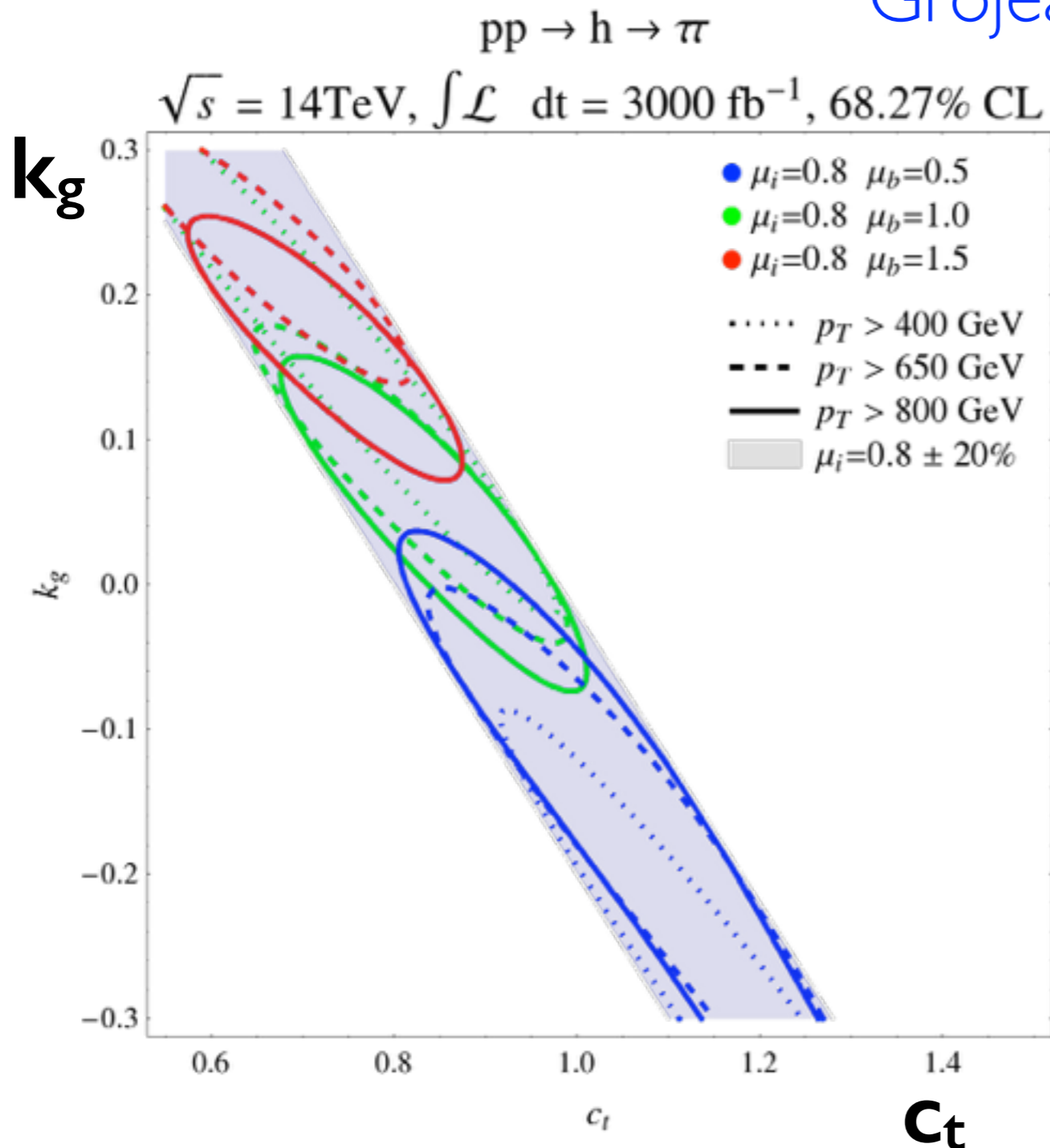
Grojean, Salvioni, Schlaffer, AW

Inclusive



Complementary to $h\bar{t}t$

Grojean, Salvioni, Schlaffer, AW, in progress



Competitive/complement to notoriously difficult $h\bar{t}t$ channel

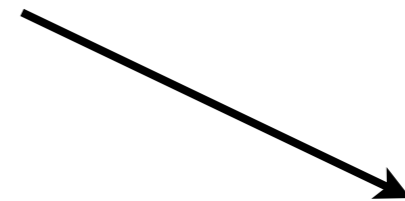
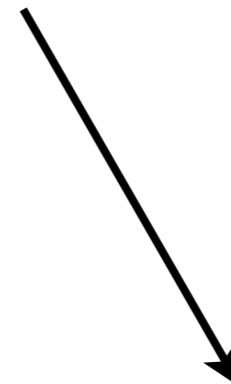
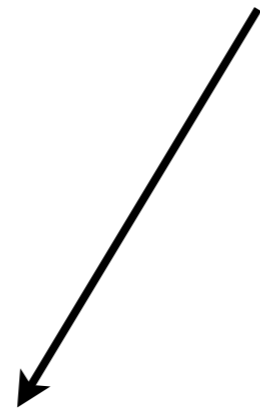
Theory frontier:

NLO $_{m_t}$ not yet calculated,
 $1/m_t$ known to $\mathcal{O}(\alpha_S^4)$:
few % up to $p_T \sim 150 \text{ GeV}$

Harlander et al '12

New physics & naturalness

Light Higgs



?

light stops_{1,2}, sbottom_L,
higgsinos, gluinos, ...

supersymmetry

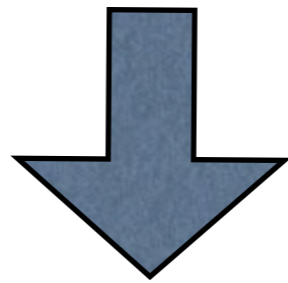
light top partners
($Q=5/3, 2/3, 1/3$),
anything else ?

composite Higgs

Flavor used to be a show-stopper

CPV in Kaon mixing

$$|\epsilon| = 2.3 \times 10^{-3} \implies \frac{M_{ETC}}{g_{ETC} \sqrt{\text{Im}(V_{sd}^2)}} \gtrsim 16,000 \text{ TeV}$$



$$m_{q,\ell,T}(M_{ETC}) \simeq \frac{g_{ETC}^2}{2M_{ETC}^2} \langle \bar{T}T \rangle_{ETC} \lesssim \frac{0.1 \text{ MeV}}{|V_{sd}|^2 N^{3/2}} \quad \text{vs. } m_{\text{top}}$$

Partial compositeness

Fermionic operators can excite composite fermions at low energy:

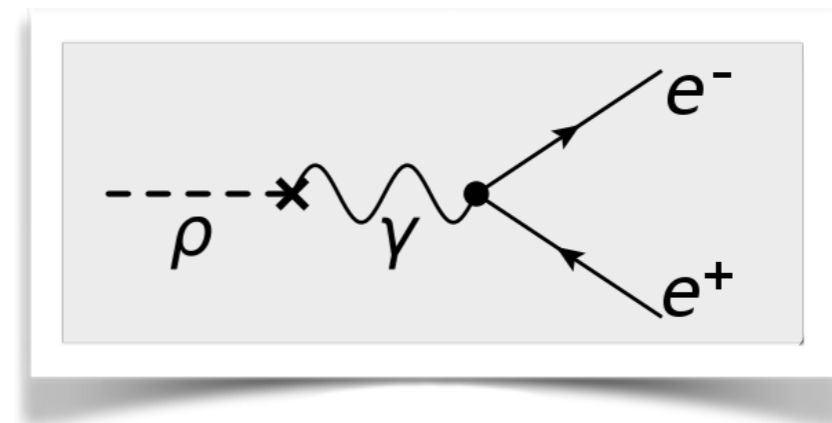
vector-like
composite fermion

↓

$$\langle 0 | O | \chi \rangle = \lambda f$$

Analogous to photon-rho mixing

$$\text{Br}(\rho \rightarrow e^+e^-) \sim 10^{-5}$$

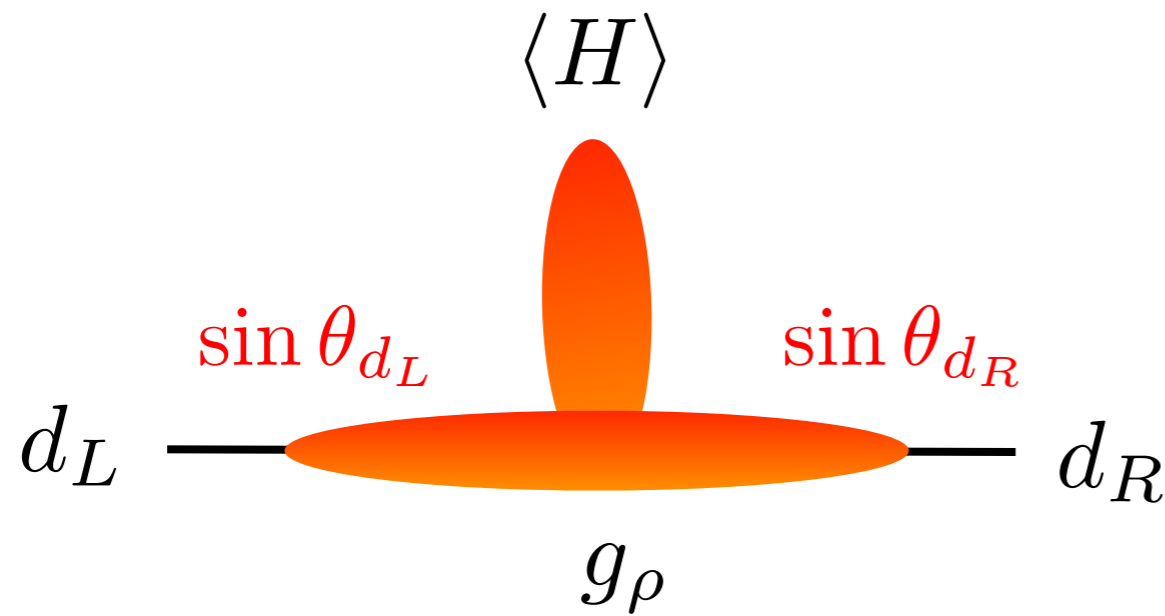


Linear couplings imply mass mixings:

$$\mathcal{L} = \bar{\psi} i \not{\partial} \psi + \bar{\chi} (i \not{\partial} - m_*) \chi + \lambda f \bar{\psi} \chi + h.c.$$

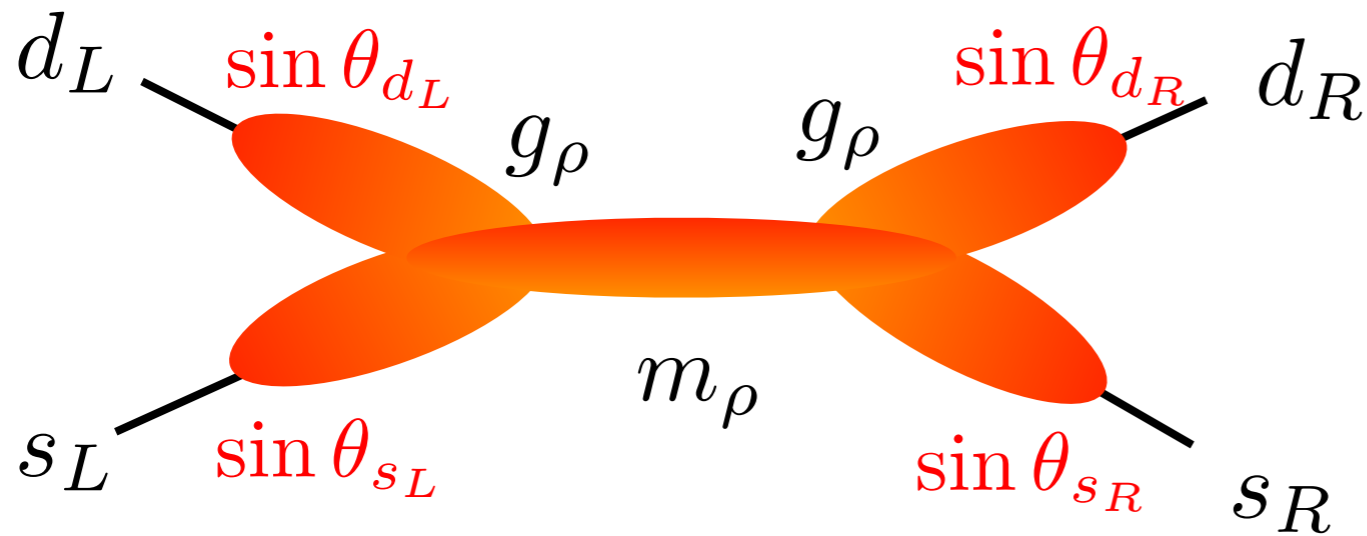
Rotate to mass eigenbasis:

$$\begin{pmatrix} \psi \\ \chi \end{pmatrix} \rightarrow \begin{pmatrix} \cos \varphi & \sin \varphi \\ \sin \varphi & \cos \varphi \end{pmatrix} \begin{pmatrix} \psi \\ \chi \end{pmatrix} \quad \tan \varphi = \frac{\lambda f}{m_*}$$



Yukawa

$$g_\rho \sin \theta_{d_L} \langle H \rangle \sin \theta_{d_R} \sim m_d$$



FCNC

$$\sim s_{d_L} s_{d_L} s_{d_R} s_{s_R} \sim \frac{m_d m_s}{v^2}$$

GIM-like protection

... almost works $\Lambda_{\epsilon_K} = 10^5 \text{ TeV} \rightarrow m_\rho \gtrsim 10 \text{ TeV}$

“Into the Extra-dimension
and back”

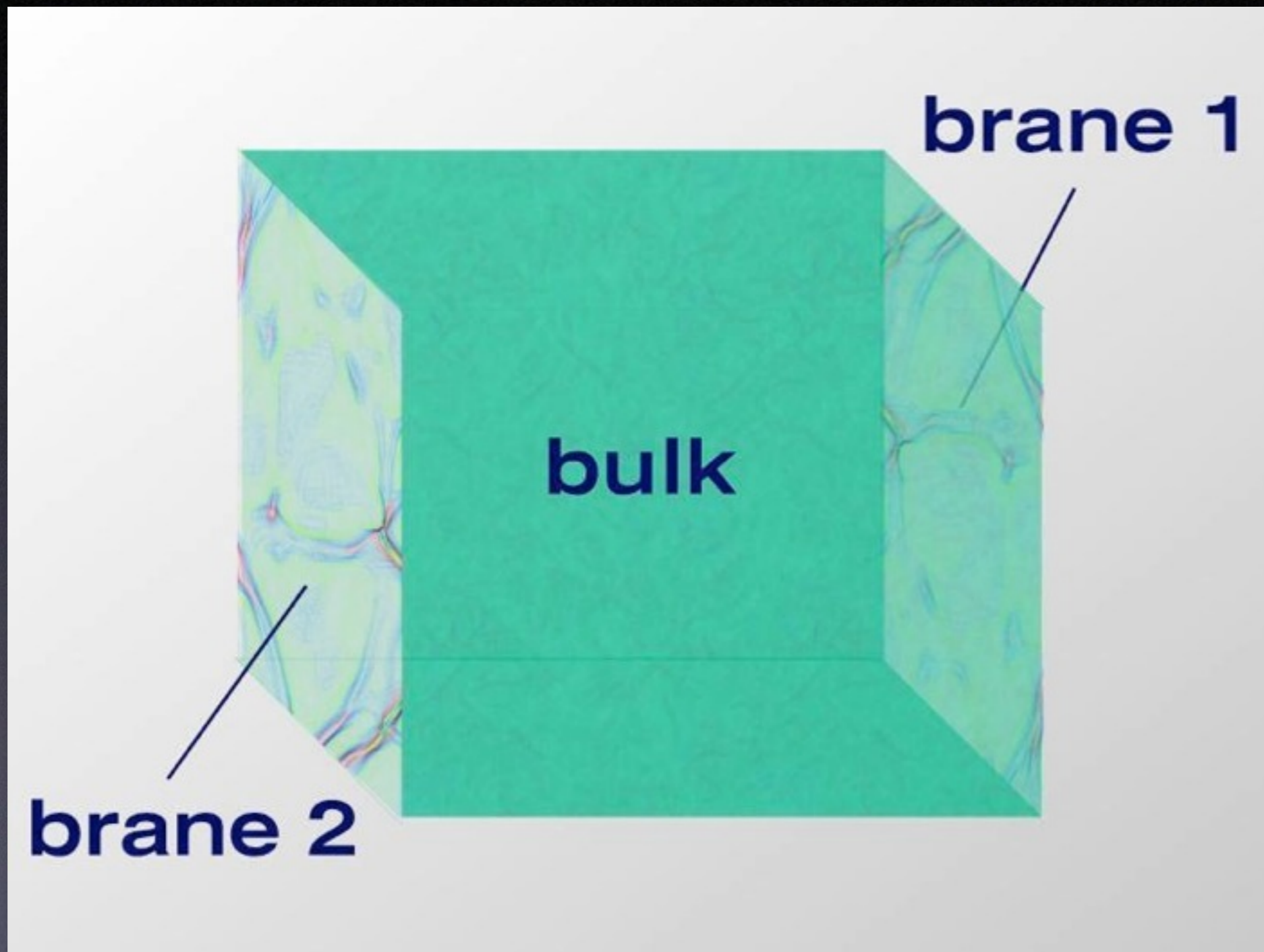
Exciting journey...



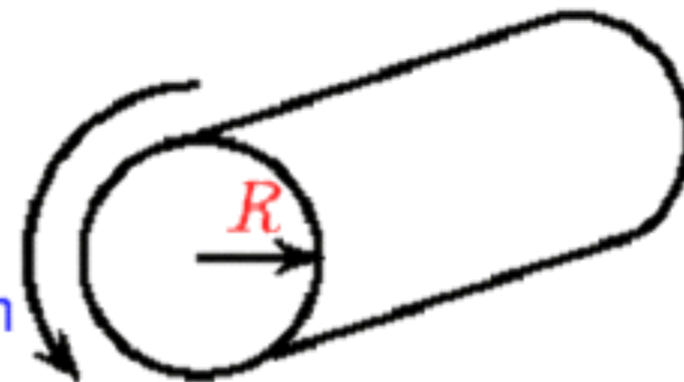
Depends on the perspective...



Extra-dimensions



Compact
Dimension



$$\phi(x) = \phi(x + k2\pi R)$$

$$(k = 0, 1, 2, \dots)$$

$$p = k/R$$

Compact extra dimensions

Compact Extra-dimension => momentum in extra-dim' direction is quantized: $p_{ED} = n/(\text{size of ED})$

$$p^2 = m^2 \quad \longrightarrow \quad p_{5D}^2 = p^2 - (n/R)^2 = m^2$$

4D 5D

Two pictures (n/R on LHS or RHS):

- 1) 5D field with quantized momentum and mass m^2
- 2) **infinite** tower of 4D fields labeled by 5 momentum n/R with masses

$$M_n^2 = m^2 + (n/R)^2$$

new particles: Kaluza Klein (KK) modes



Kaluza Klein states

Free scalar field, massless

$$S = \int d^5x \frac{1}{2} \partial_M \Phi \partial^M \Phi$$

Expand in fourier modes

$$\Phi(x, y) = \frac{1}{\sqrt{2\pi R}} \sum_n \Phi^{(n)}(x) e^{i \frac{n}{R} y}$$

with $(\Phi^{(n)})^\dagger = \Phi^{(-n)}$ to 'keep it real'



$$\partial_\mu \phi \partial^\mu \phi = \partial_\mu \phi \partial^\mu \phi - (\partial_y \phi)^2$$

orthogonal
 $m = -n$

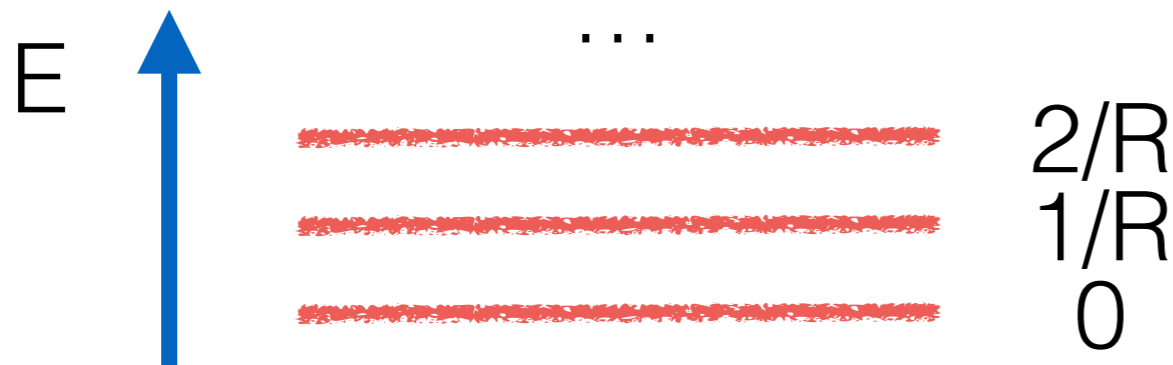
$$S = \int d^4x \sum_{m,n} \int dy \frac{1}{2\pi R} e^{i \frac{(m+n)}{R} y}$$

$$\cdot \frac{1}{2} \left[\partial_\mu \phi^{(m)}(x) \partial^\mu \phi^{(n)}(x) + \frac{m \cdot n}{R^2} \phi^{(m)} \cdot \phi^{(n)} \right]$$

$$= \frac{1}{2} \sum_n \int d^4x \left[\partial_\mu \phi^{(n)} \partial^\mu \phi^{(n)} - \frac{n^2}{R^2} \phi^{(-n)} \phi^{(n)} \right]$$

Infinite tower of massive 4D fields

$$m^{(n)} = \frac{n}{R}$$



The SM flavor puzzle

$$Y_D \approx \text{diag} (2 \cdot 10^{-5} \quad 0.0005 \quad 0.02)$$

$$Y_U \approx \begin{pmatrix} 6 \cdot 10^{-6} & -0.001 & 0.008 + 0.004i \\ 1 \cdot 10^{-6} & 0.004 & -0.04 + 0.001i \\ 8 \cdot 10^{-9} + 2 \cdot 10^{-8}i & 0.0002 & 0.98 \end{pmatrix}$$

Why this structure?

Other dimensionless parameters of the SM:

$$g_s \sim 1, \quad g \sim 0.6, \quad g' \sim 0.3, \quad \lambda_{\text{Higgs}} \sim 1, \quad |\theta| < 10^{-9}$$

Log(SM flavor puzzle)

$$-\log |Y_D| \approx \text{diag} (11 \quad 8 \quad 4)$$

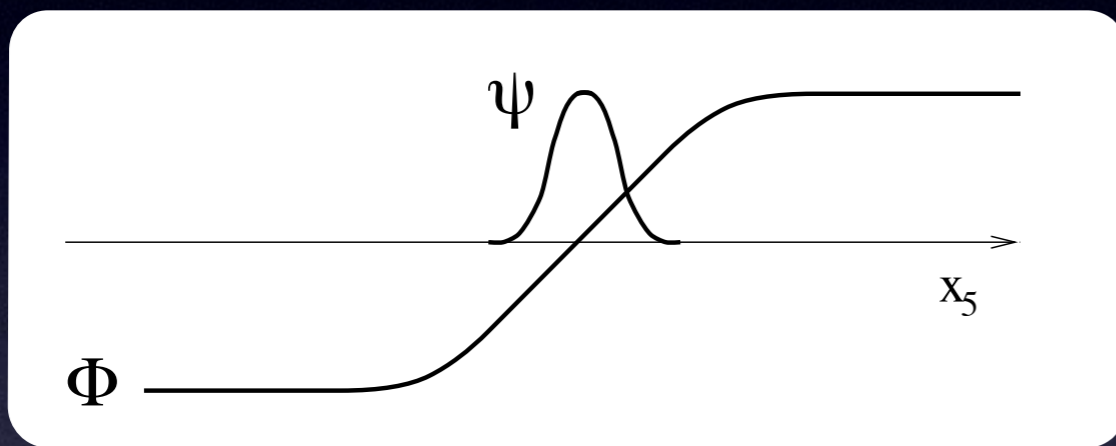
$$-\log |Y_U| \approx \begin{pmatrix} 12 & 7 & 5 \\ 14 & 6 & 3 \\ 18 & 9 & 0 \end{pmatrix}$$

If $Y = e^{-\Delta}$, then the Δ don't look crazy.

Hierarchies w/o Symmetries

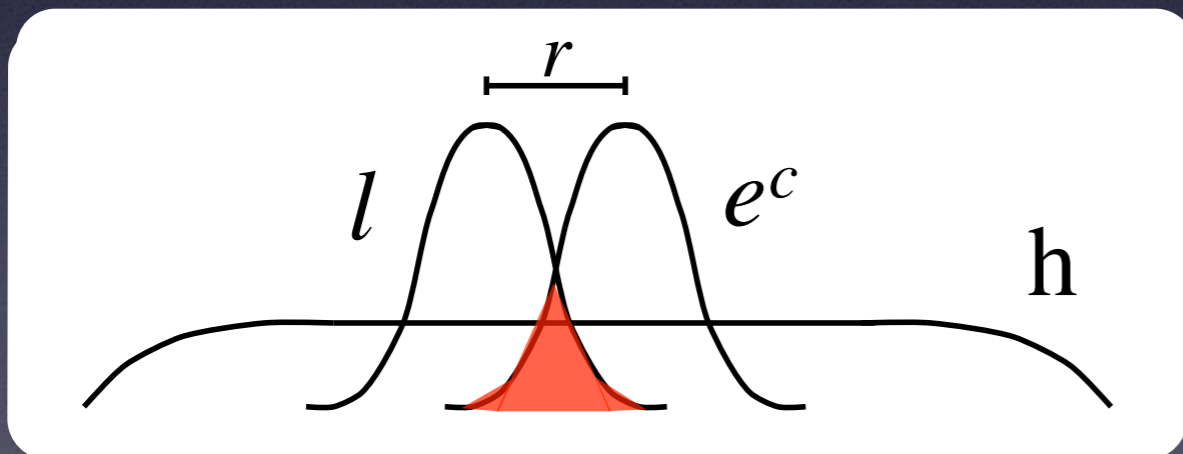
Arkani-Hamed, Schmaltz

SM on thick brane & domain wall \Rightarrow chiral localization



$$\mathcal{S} = \int d^5x \sum_{i,j} \bar{\Psi}_i [i \not{\partial}_5 + \lambda \Phi(x_5) - m]_{ij} \Psi_j$$

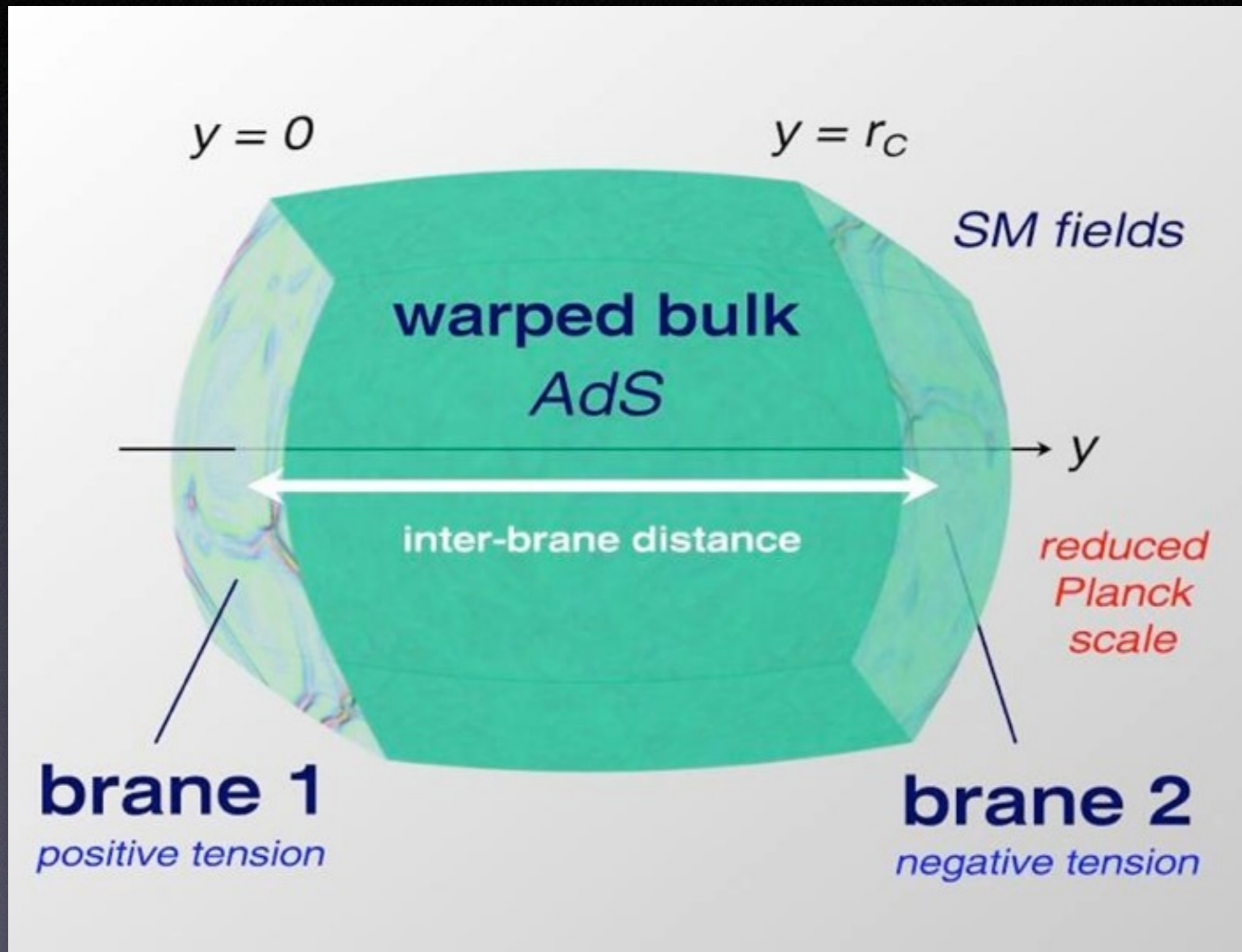
$$\Psi = \begin{pmatrix} \Psi_L \\ \Psi_R \end{pmatrix} = \begin{pmatrix} \psi_L^0 \\ 0 \end{pmatrix} + \text{KK modes}$$



Log(flavor hierarchy)!

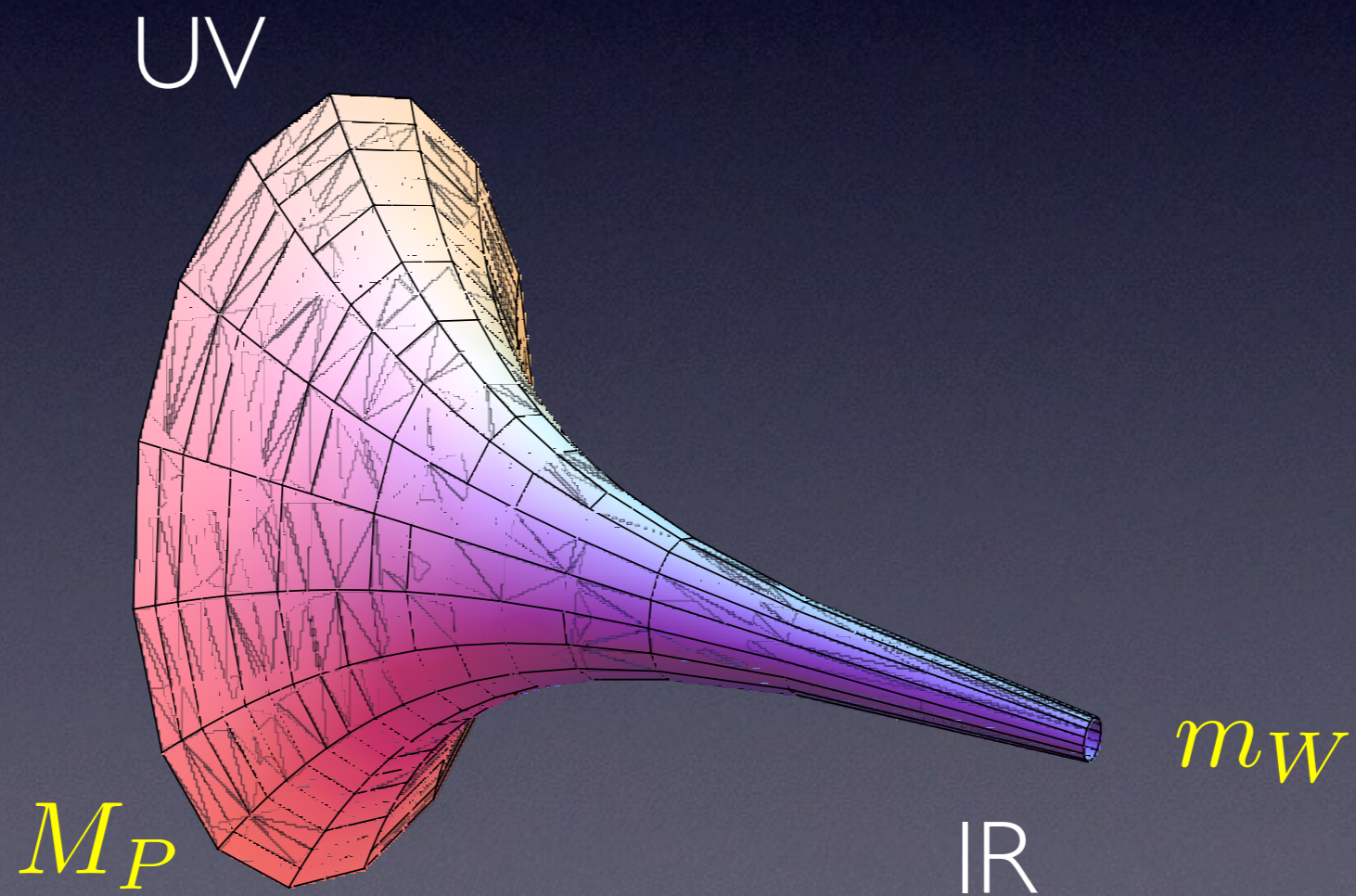
$$\int dx_5 \phi_l(x_5) \phi_{e^c}(x_5) = \frac{\sqrt{2}\mu}{\sqrt{\pi}} \int dx_5 e^{-\mu^2 x_5^2} e^{-\mu^2 (x_5 - r)^2} = e^{-\mu^2 r^2 / 2}$$

Warped Extra Dimensions



How to do calculations in a strongly coupled theory?

Excursion into AdS/CFT



AdS/CFT

Maldacena

$$ds^2 = \left(\frac{R}{z}\right)^2 (dx_\mu dx_\nu - dz^2)$$

Anti-de-Sitter (AdS)

Compactification

Red-shifting of scales



Conformal (CFT)

Mass gap

Dimensional trans-
mutation

$$m_W = \sqrt{\frac{g(IR)}{g(UV)}} M_P \ll M_P$$

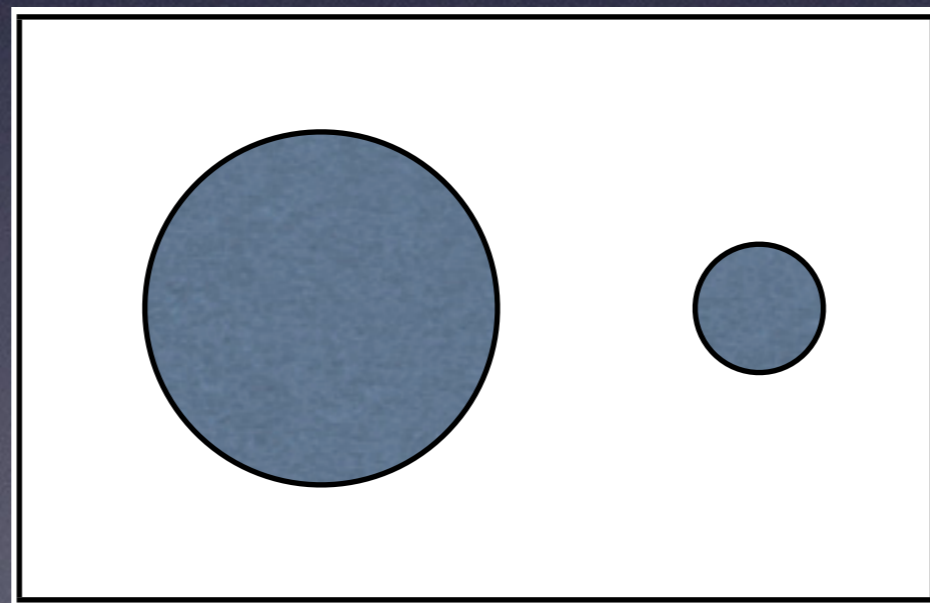
$$m_W \sim e^{-4\pi/\alpha} M_P$$



CFT & Extra-dimensions

Question

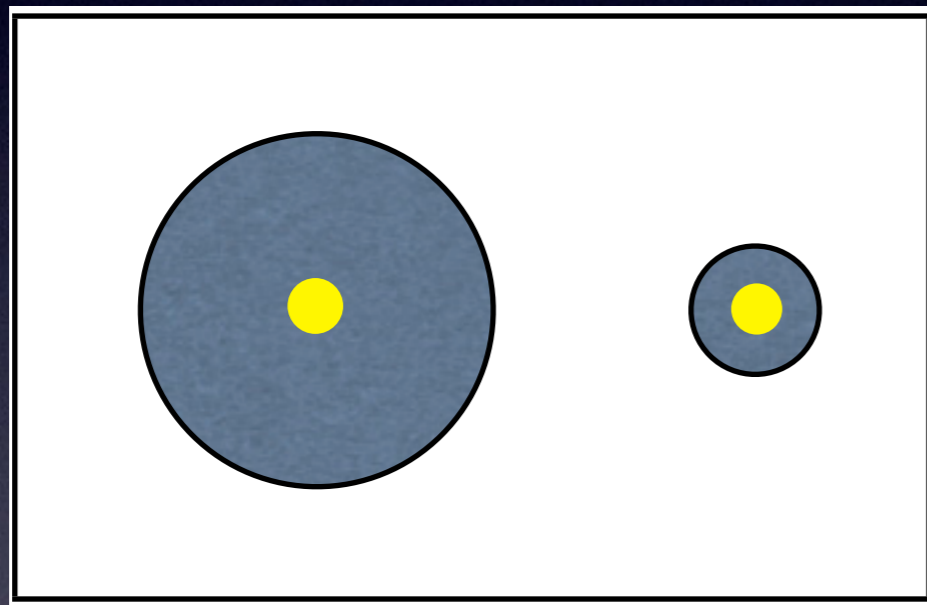
Is this a picture of a big ball and a small ball side-by-side in 2D or two identical balls at different distances in 3D?



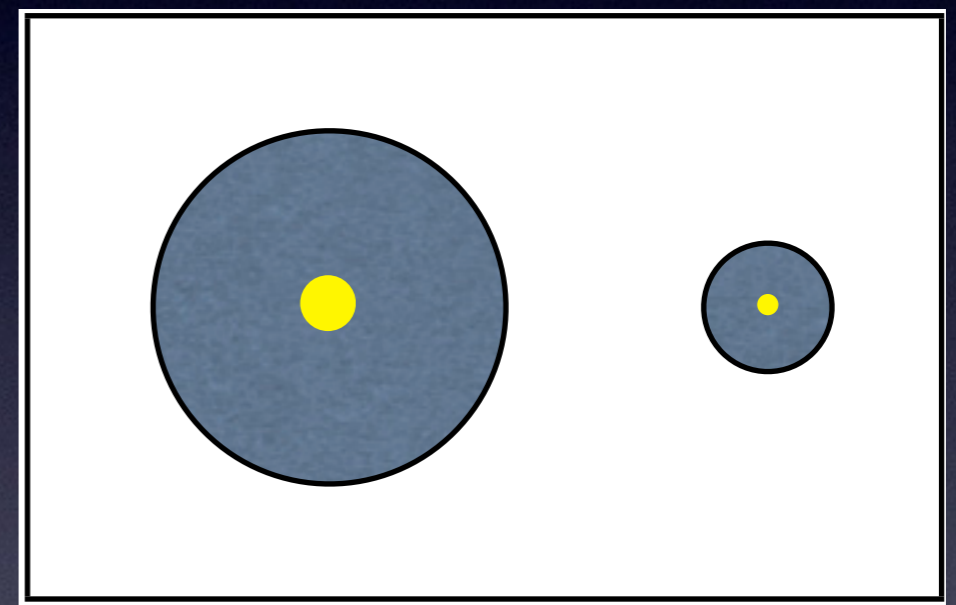
analogy by T. Okui

In the real world, we can tell the difference because atomic size is fixed.

(● Atom)

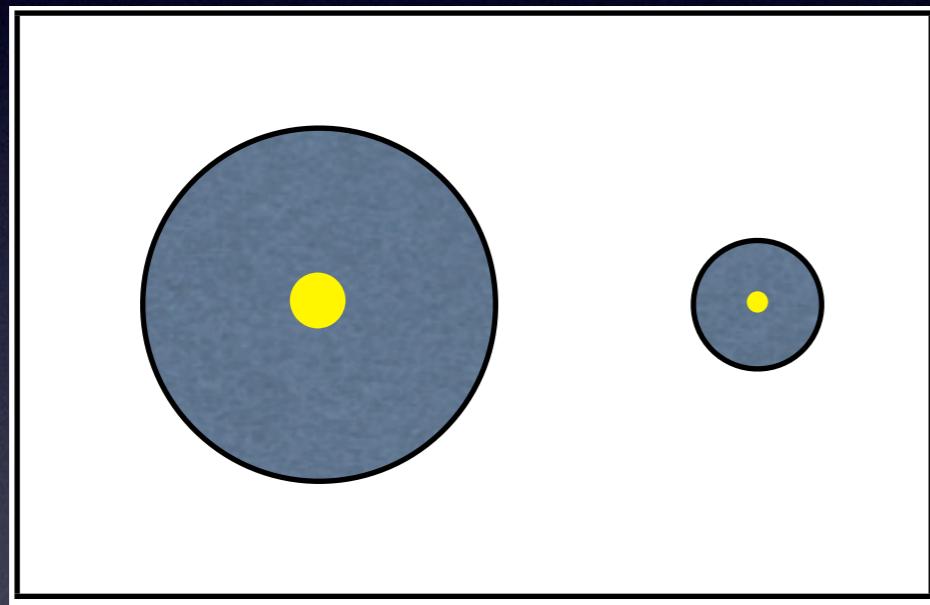


2D: Big & small balls

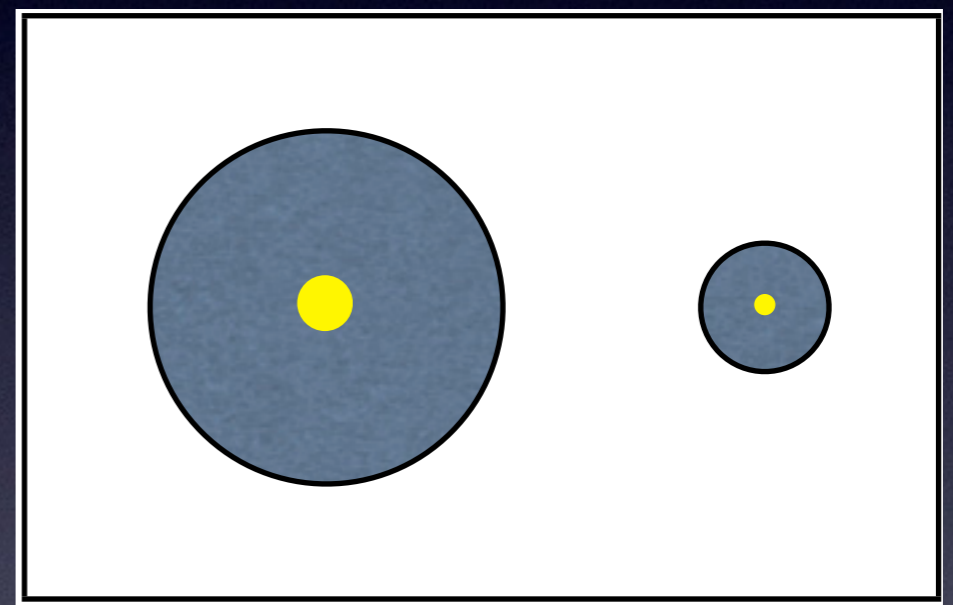


3D: Near & Far

But in a scale invariant world, atomic size ● would also **scale**.



2D + scale invariance



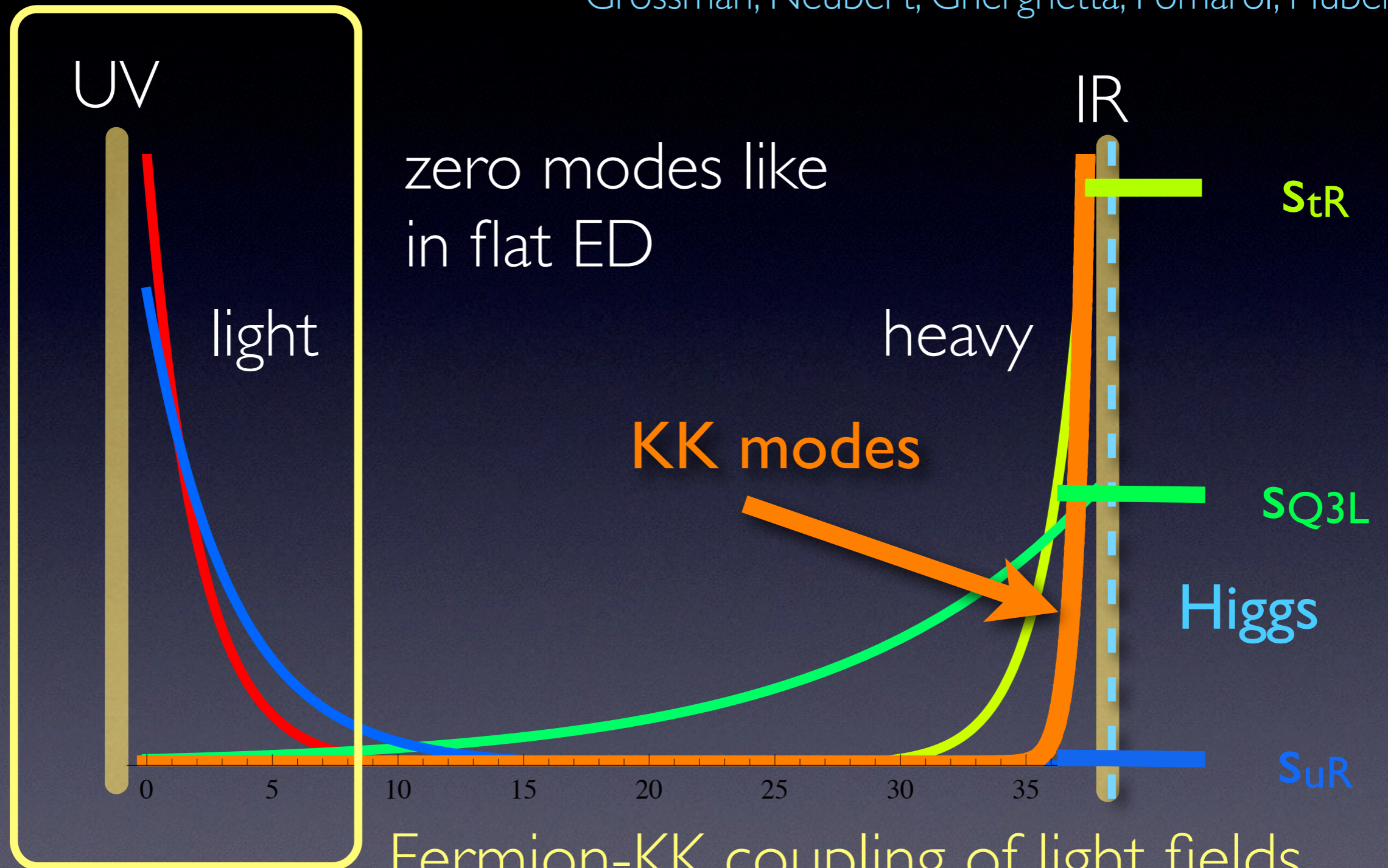
3D

scale invariance = extra-dimension !



Fermion location in AdS

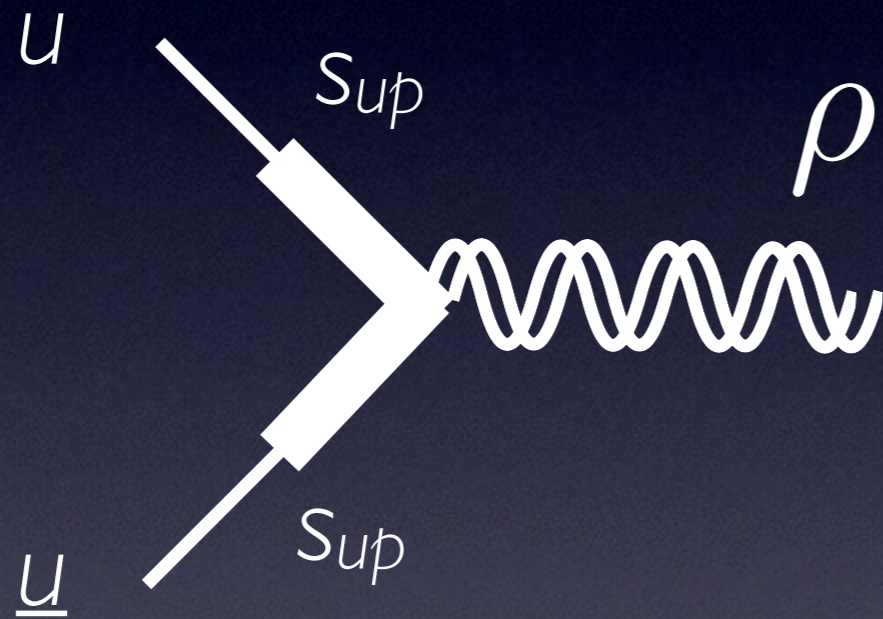
Grossman, Neubert; Gherghetta, Pomarol; Huber;



Fermion-KK coupling of light fields almost universal!

high p_T

Resonance production (option 1)

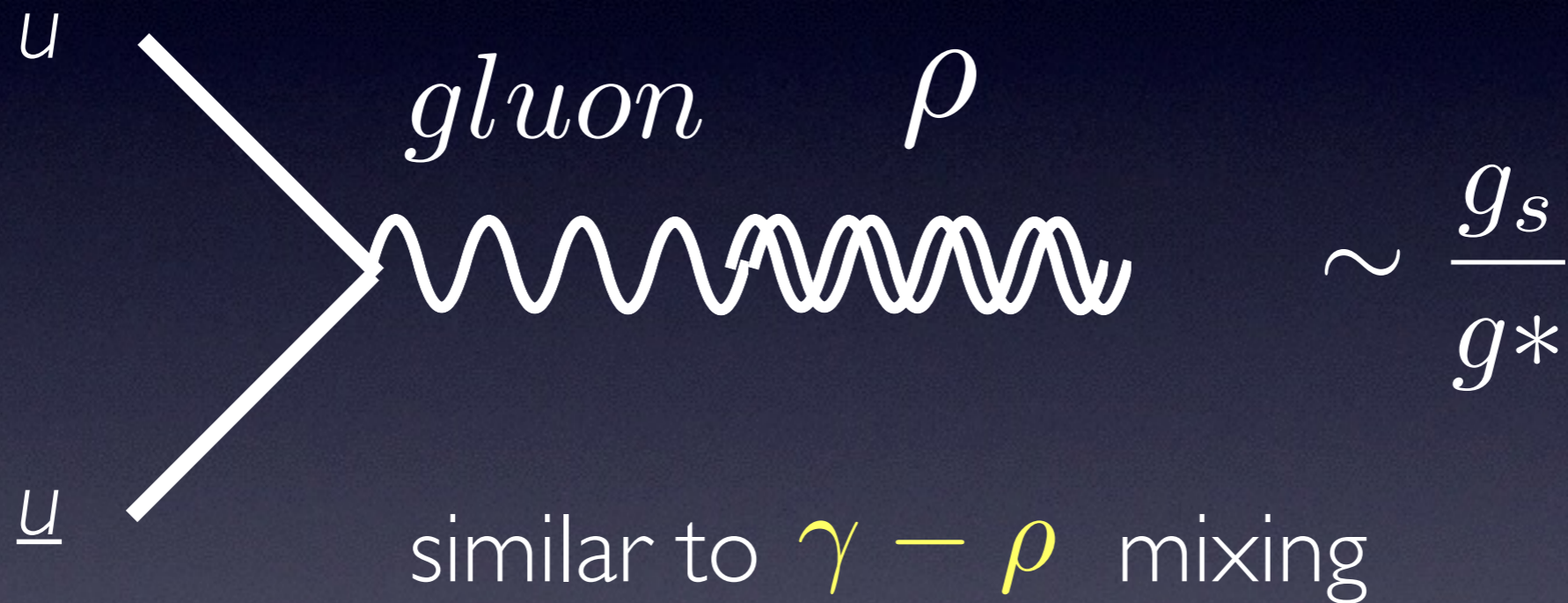


$$\sim g_*^2 \sin^2 \theta_{u_R}$$

strongly suppressed for
light quarks!

high p_T

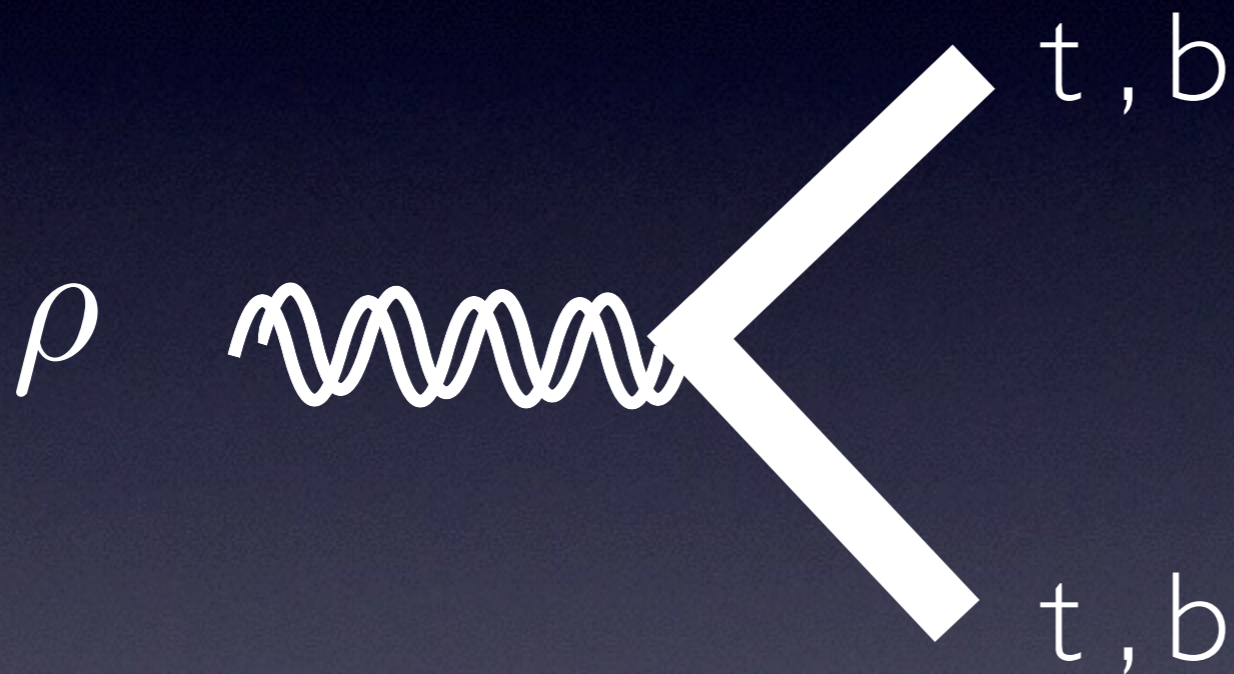
Resonance production (option 2)



NB, gluon-rho-rho = 0

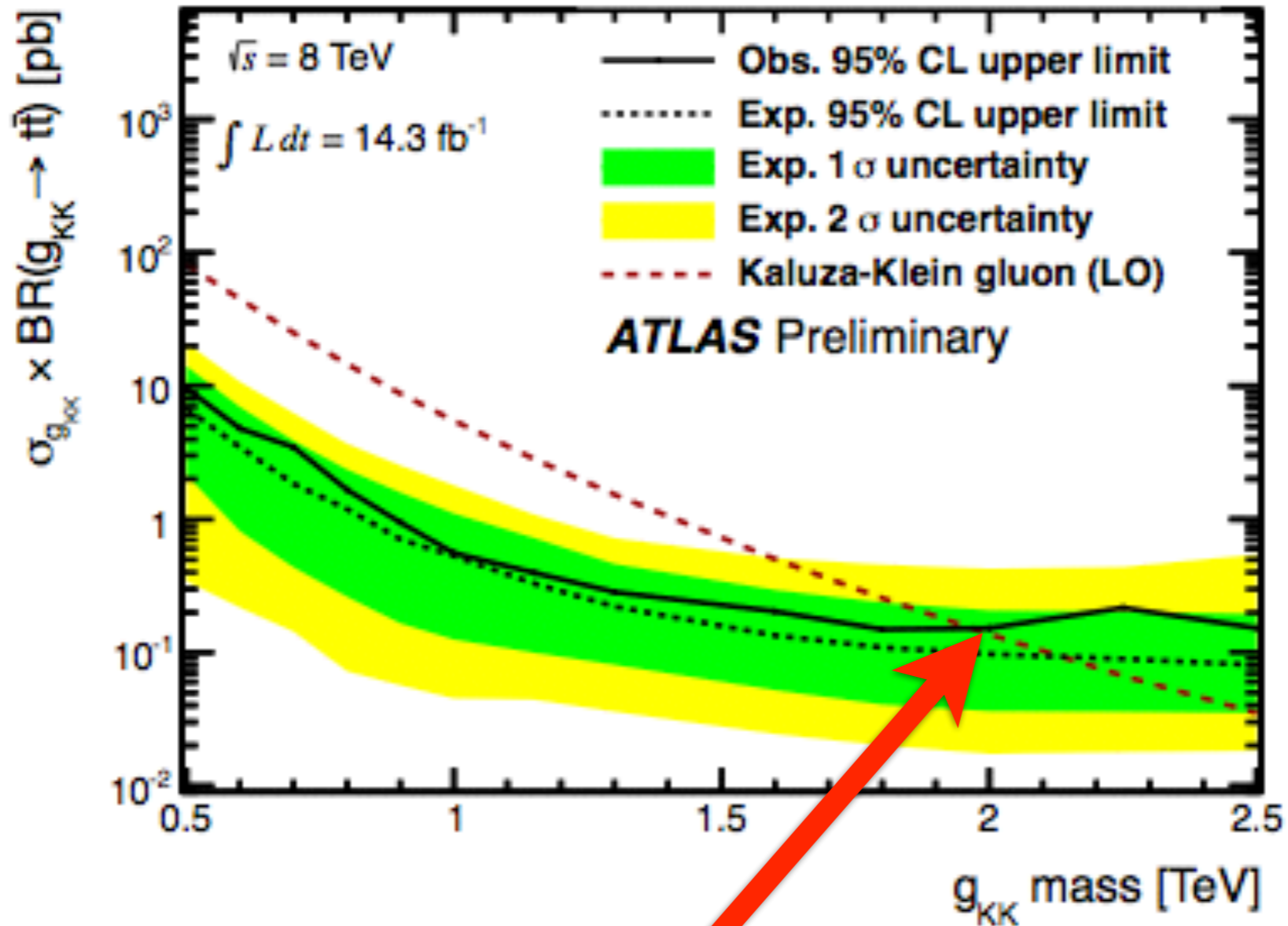
high p_T

Resonance decay



decays dominantly
into 3rd generation!
(tt, bt, bb)

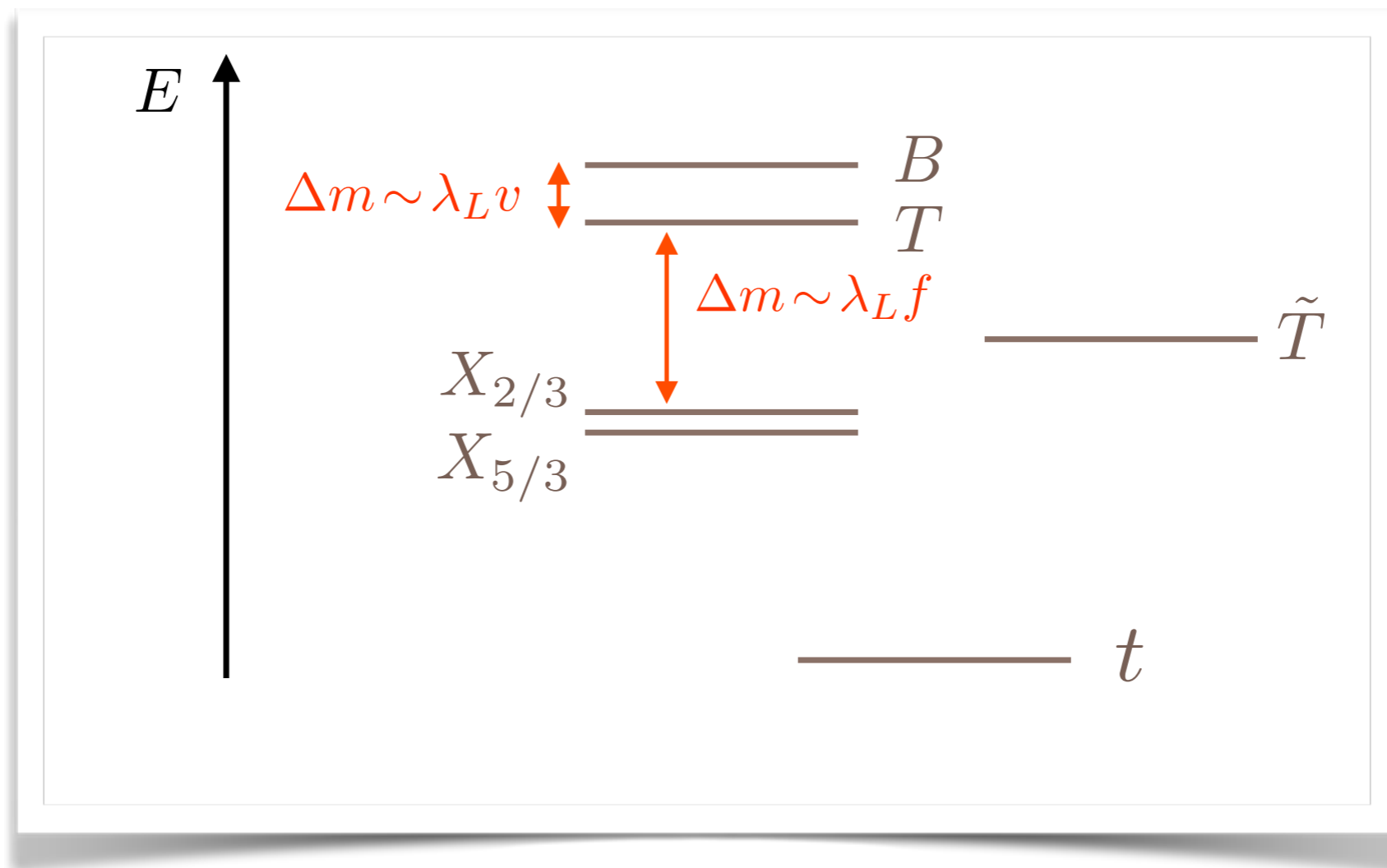
Agashe et al, Lillie et al



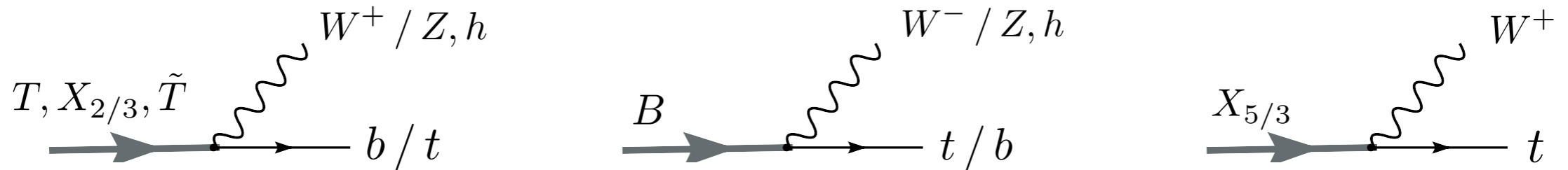
(b) g_{KK} upper cross section limits.

$$M_{KK} > 2 \text{ TeV @ } 95\text{CL}$$

Top partners



Decay modes



Current limits
 $> 700 - 800 \text{ GeV}$

CMS preliminary $\sqrt{s} = 8 \text{ TeV}$ 19.6 fb^{-1}

$BR(hW)$

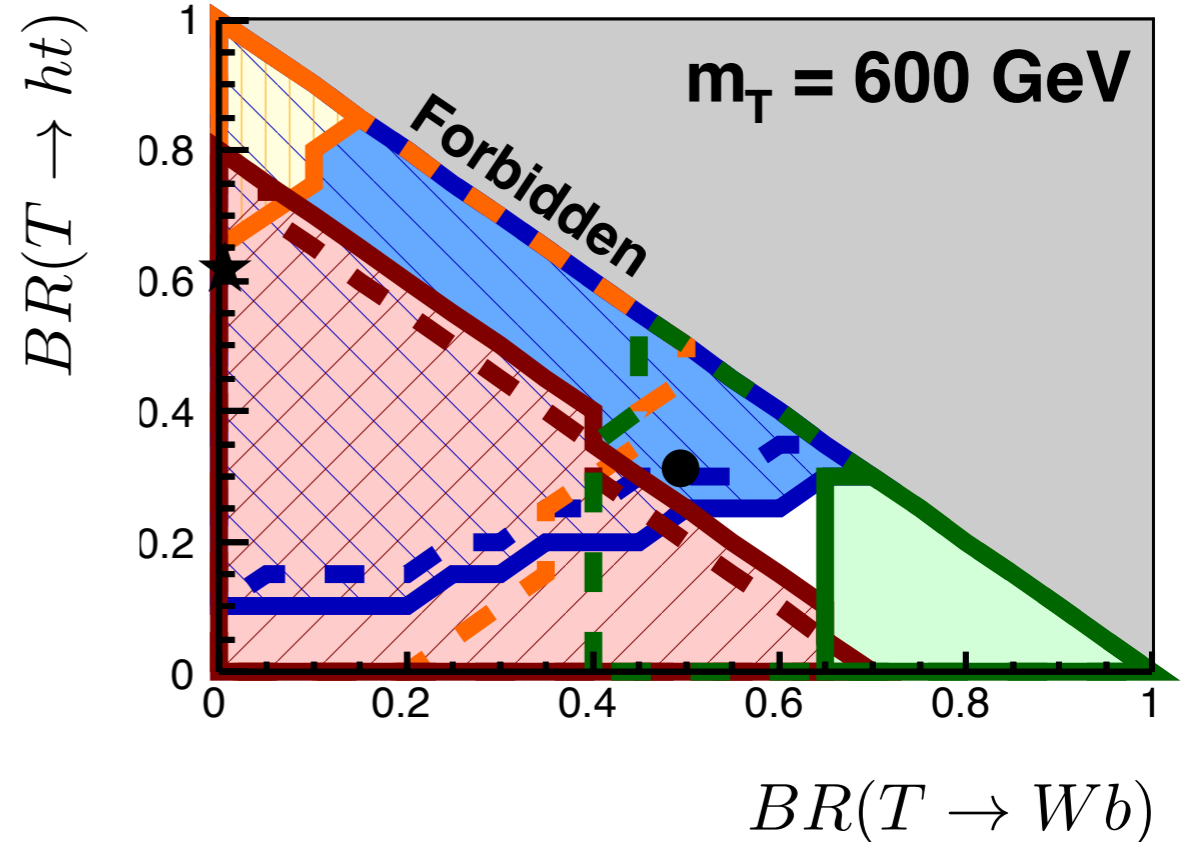
ATLAS Preliminary

Status: Lepton-Photon 2013

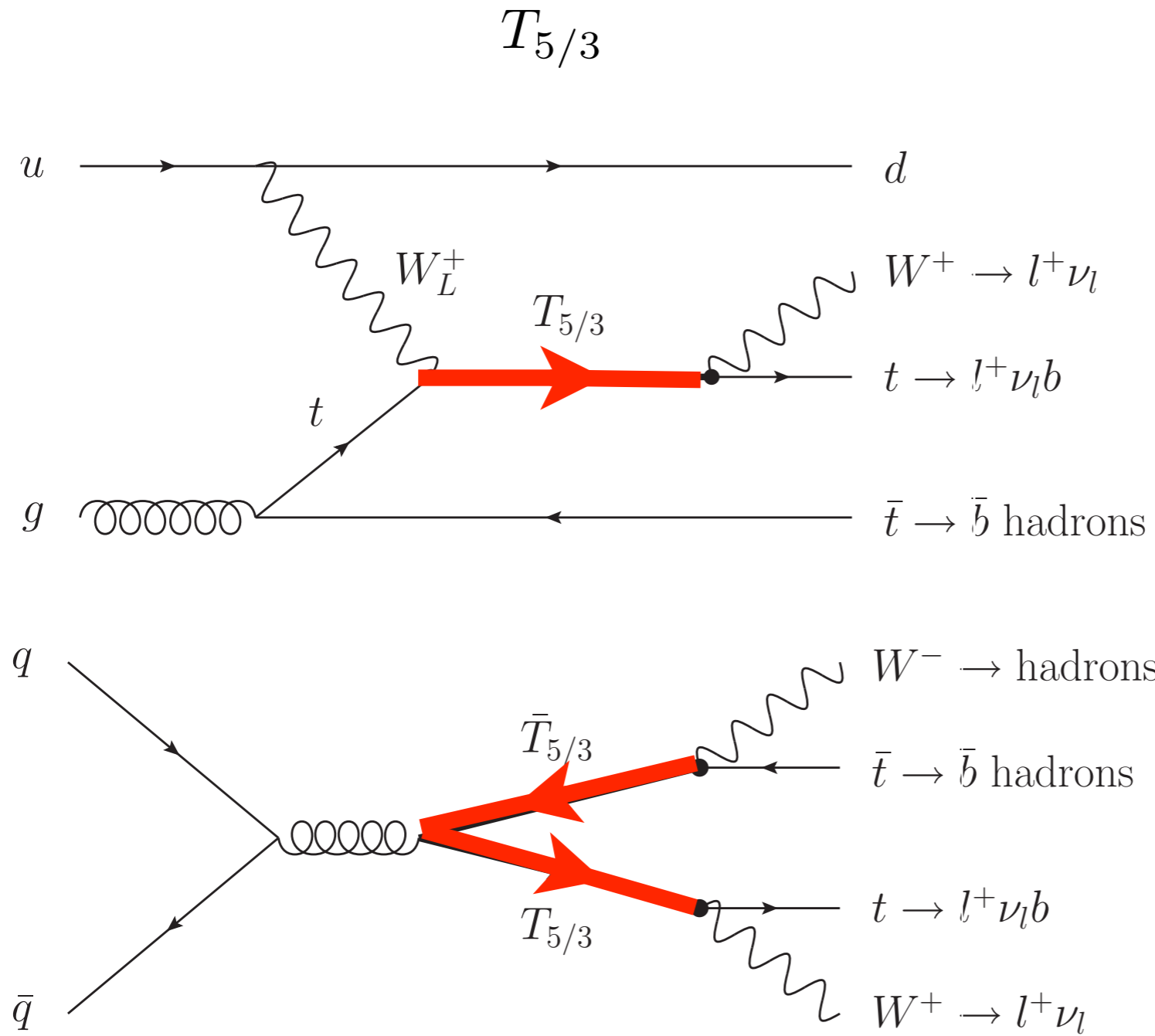
$\sqrt{s} = 8 \text{ TeV}$, $\int L dt = 14.3 \text{ fb}^{-1}$

- 95% CL exp. excl. 95% CL obs. excl.
- Ht+X [ATLAS-CONF-2013-018]
- Same-Sign [ATLAS-CONF-2013-051]
- Zb/t+X [ATLAS-CONF-2013-056]
- Wb+X [ATLAS-CONF-2013-060]
- ★ SU(2) (T,B) doub. ● SU(2) singlet

$BR(t)$

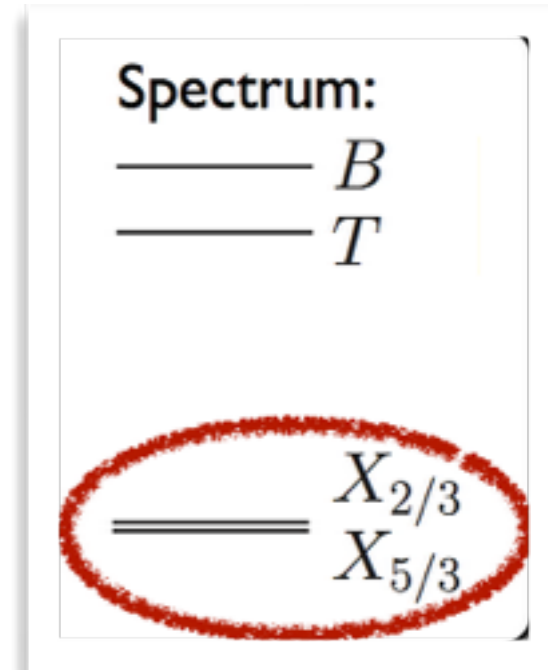


e.g. Perelstein, Pierce, Peskin
 Contino, Servant; Mrazek, Wulzer;
 De Simone, Matsedonkyi, Rattazzi, Wulzer



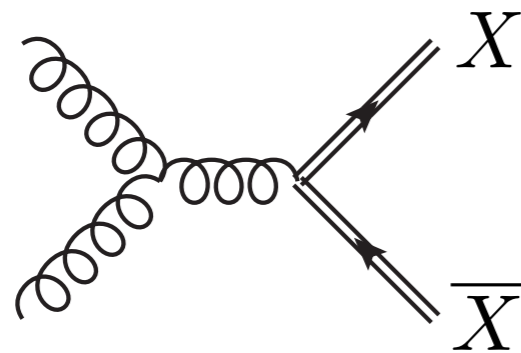
Single

Double

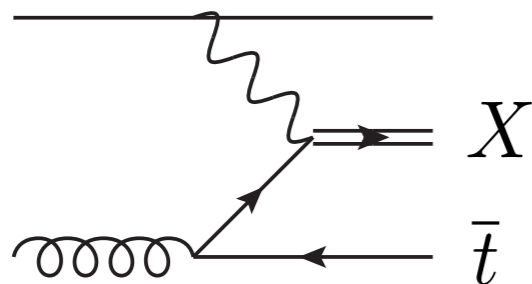


Phenomenology

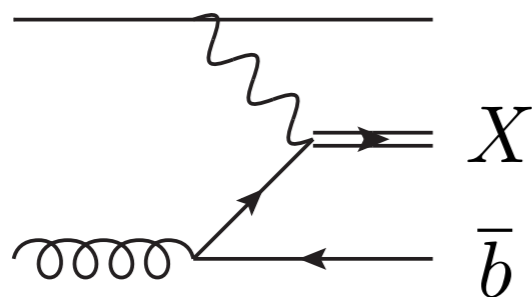
Three possible production mechanisms



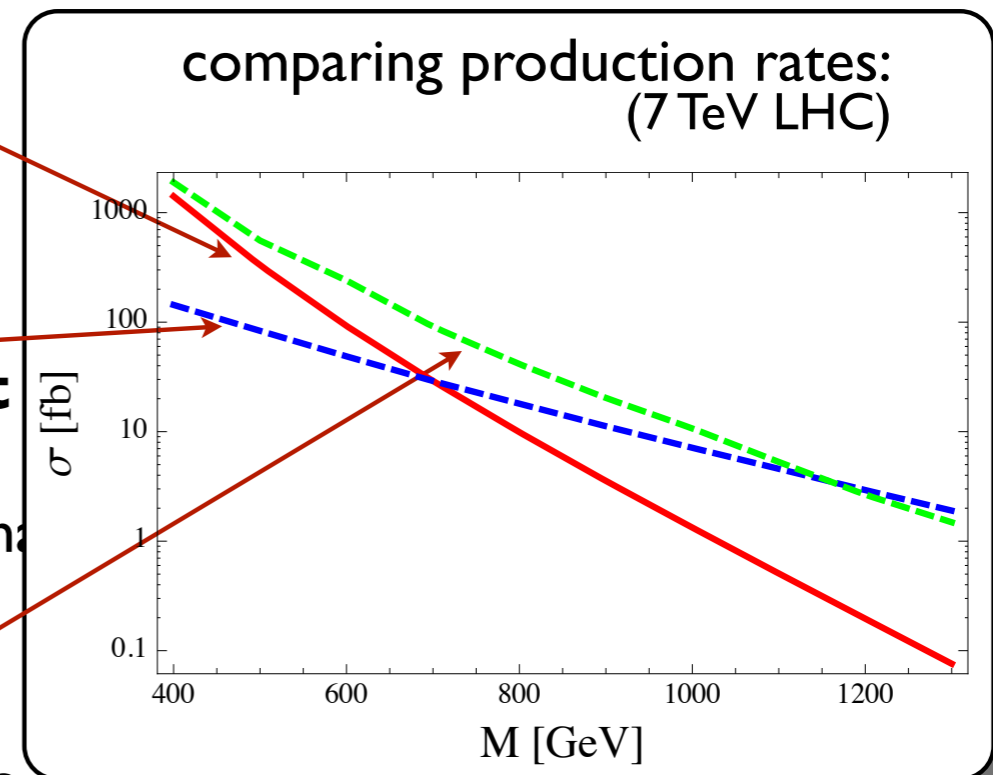
QCD pair prod.
model indep.,
relevant at low mass



single prod. with t
model dep. coupling
pdf-favored at high mass



single prod. with b
favored by small b mass
dominant when allowed



Exotics

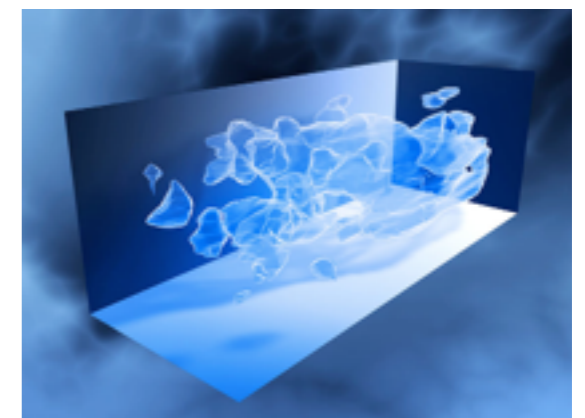
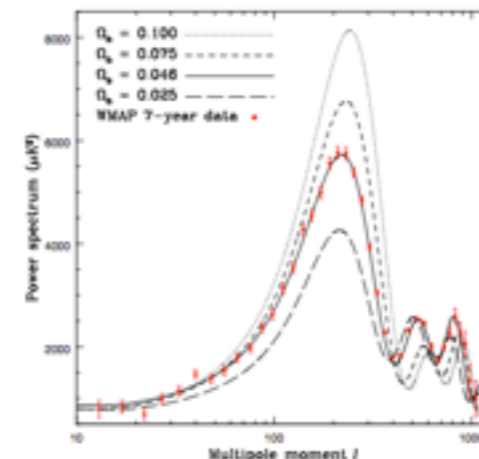
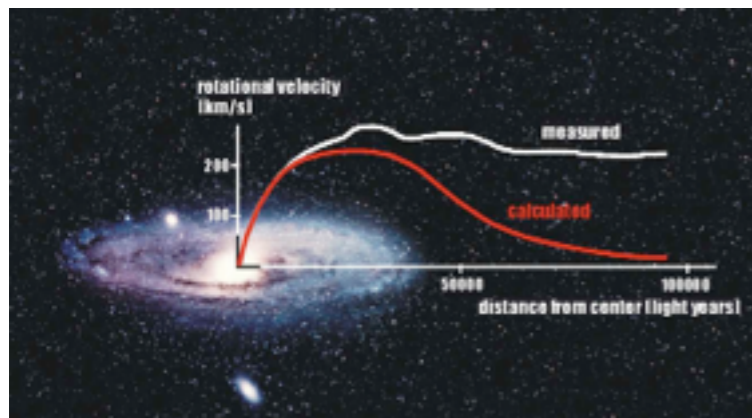
Have we thought hard enough
about non-standard options?

DM emerging jets

with D. Stolarski and P. Schwaller

Maybe DM is just part of a larger dark sector

- Example: Proton is massive, stable, composite state
- DM self interactions solve structure formation problems
- New signals, new search strategies!



Coincidence?

$$\Omega_{DM} \simeq 5\Omega_B$$

QCD like?

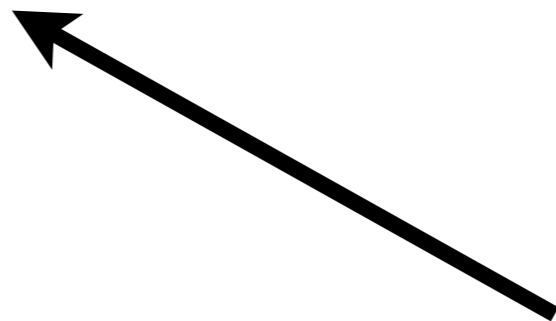


$$\Omega_{DM} = m_{DM} n_{DM}$$

Controlled by complicated
(known) QCD dynamics



$$\Omega_B = m_p n_B$$



Unknown dynamics
of baryogenesis

Dark QCD

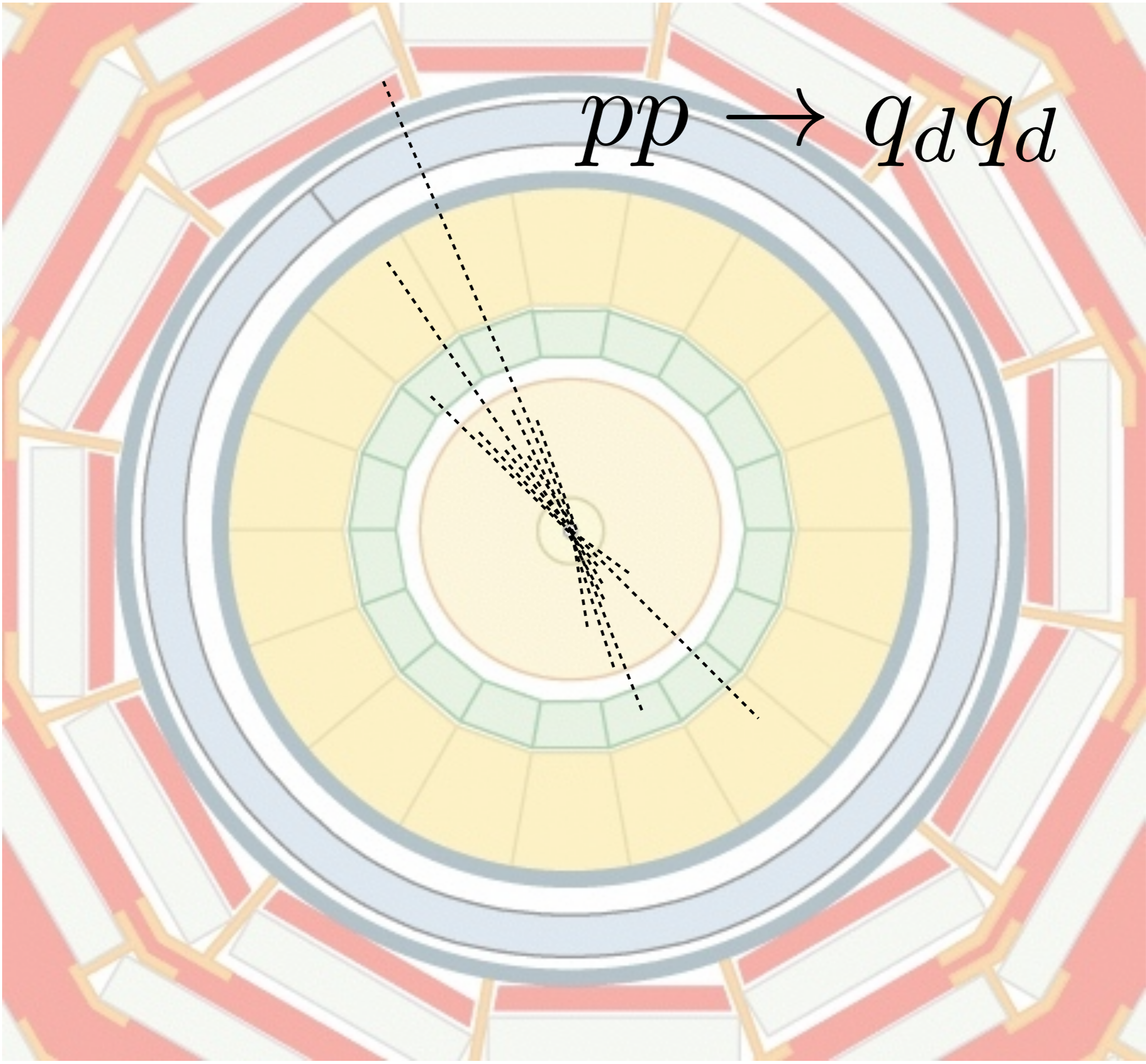
Imagine a QCD like “dark sector” with 1–10 GeV mass scale

$$p_d \quad \pi_d \quad Z_{00}d$$

Connected to SM in two ways:

- TeV scale mediator (hidden valley)
Strassler, Zurek, PLB 07.

$pp \rightarrow qdqd$



Dark QCD

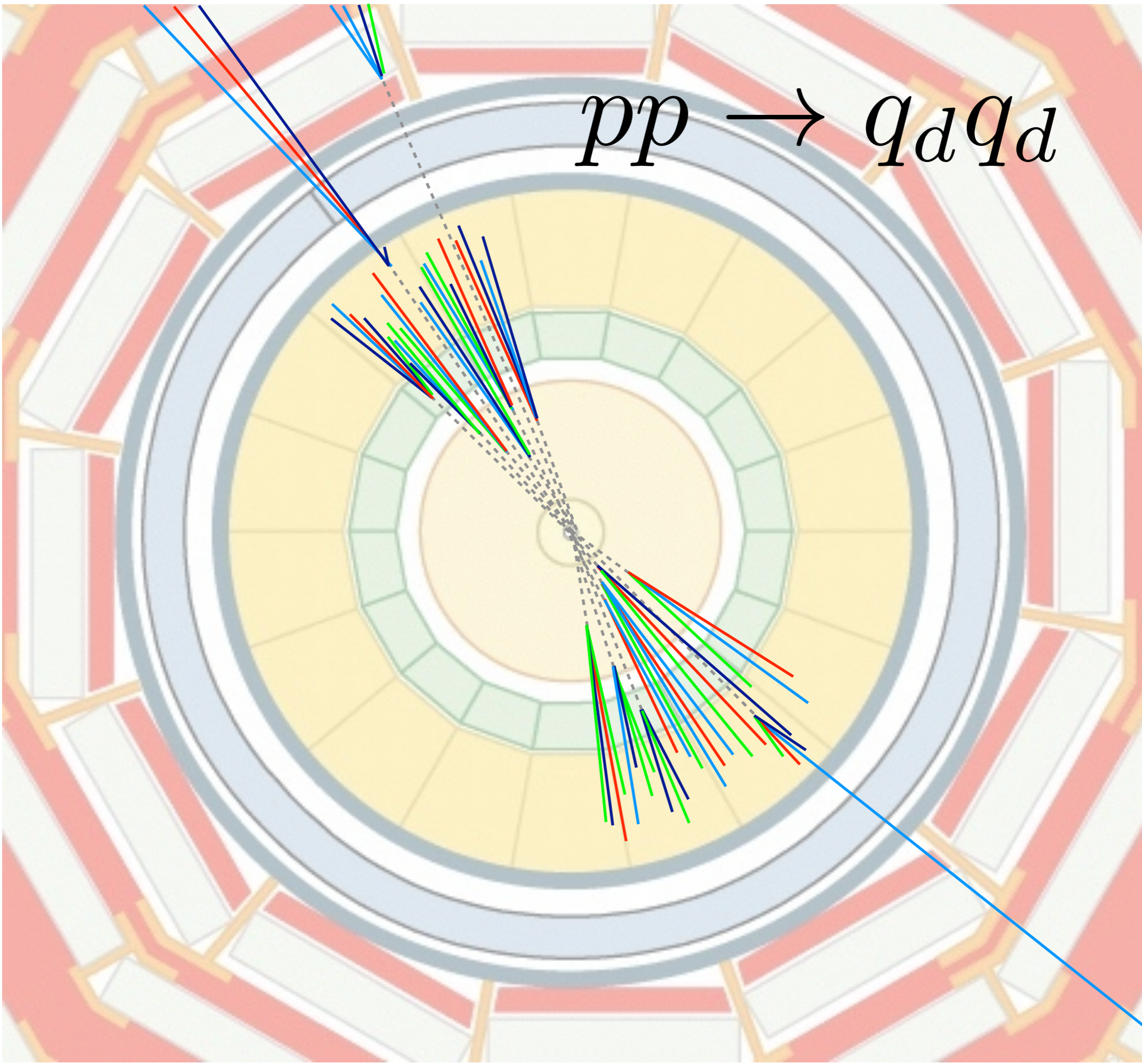
Imagine a QCD like “dark sector” with 1–10 GeV mass scale

$$p_d \quad \pi_d \quad Z_{00}d$$

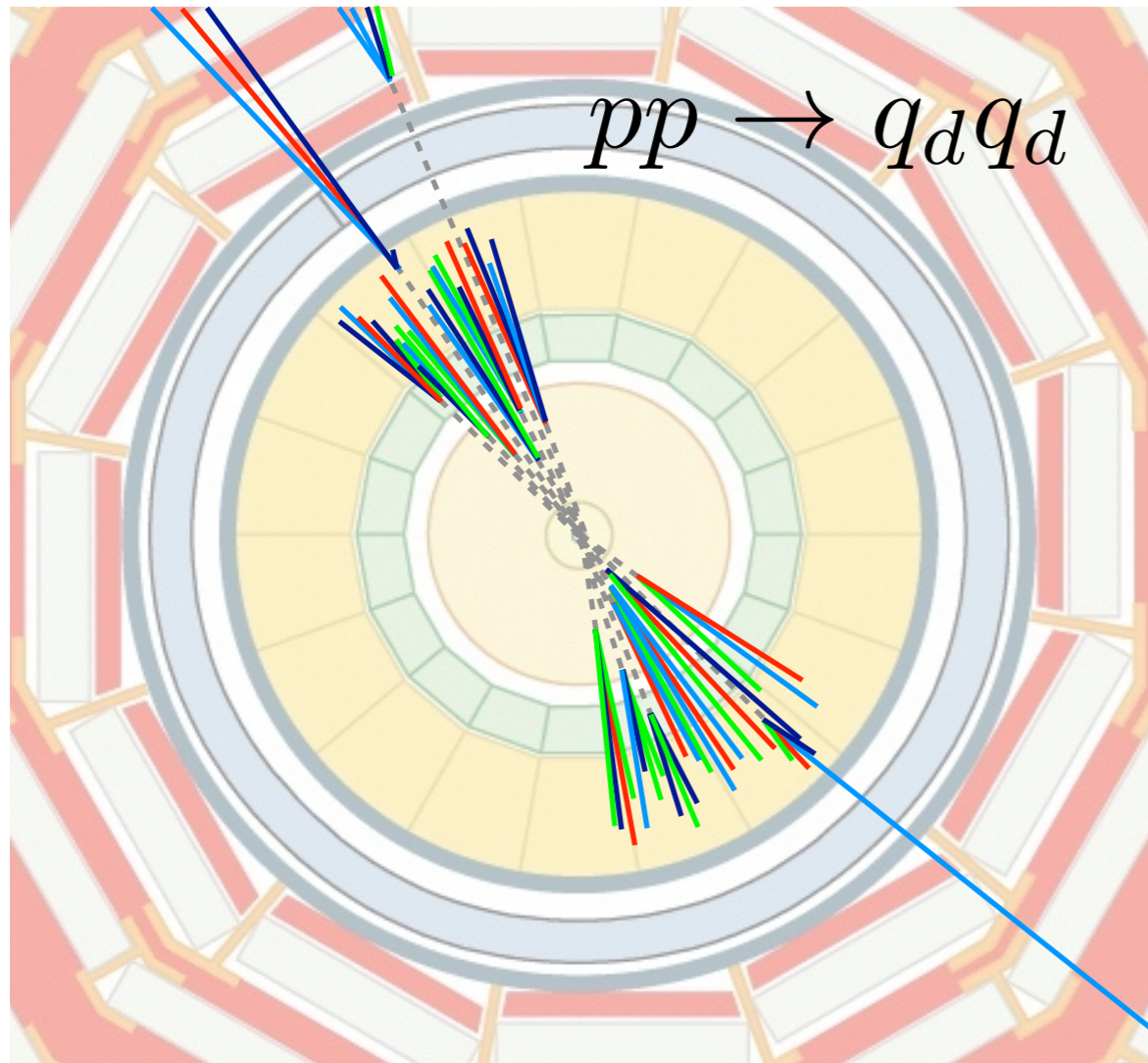
Connected to SM in two ways:

- TeV scale mediator (hidden valley)
[Strassler, Zurek, PLB 07.](#)
- Weak pion decay operator

$pp \rightarrow qdqd$



emerging jets



Decay lifetime of \sim cm

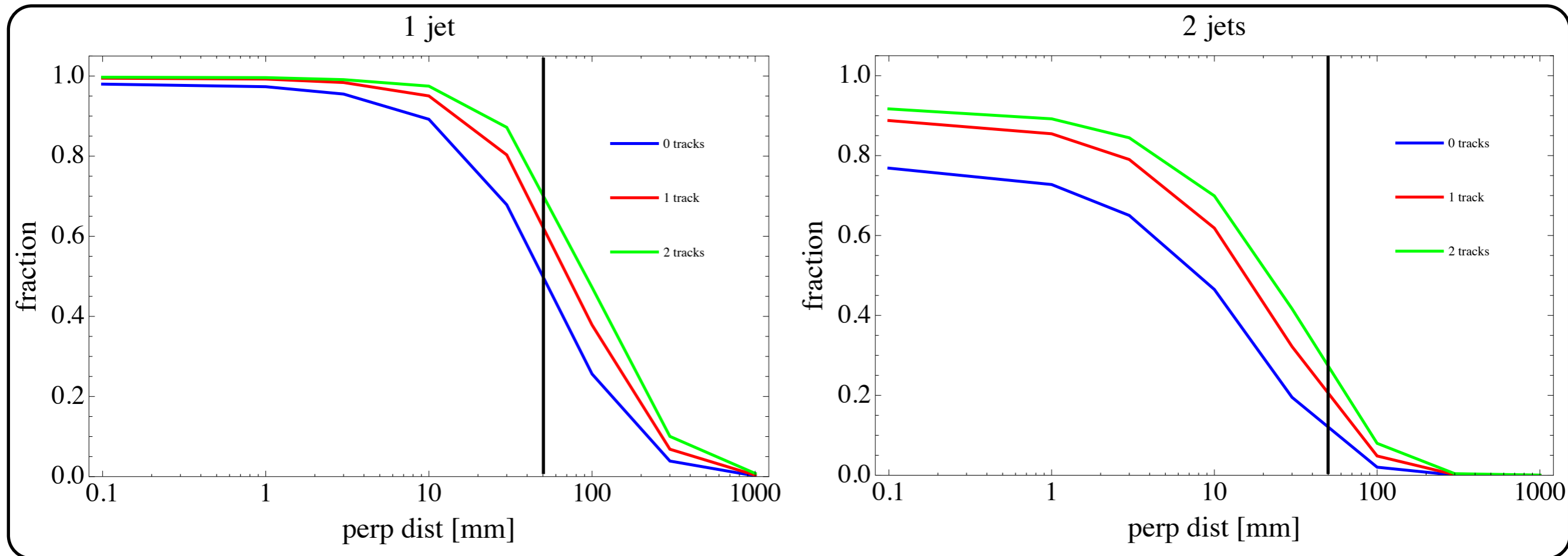
Exponential decay means jets emerge at different distances

No/few tracks originating from interaction point

$pp \rightarrow Q_d Q_d$

Look for events with
no/few tracks in the
circle

$$pp \rightarrow \Phi \Phi^\dagger \rightarrow \bar{q} Q_d \bar{Q}_d q$$

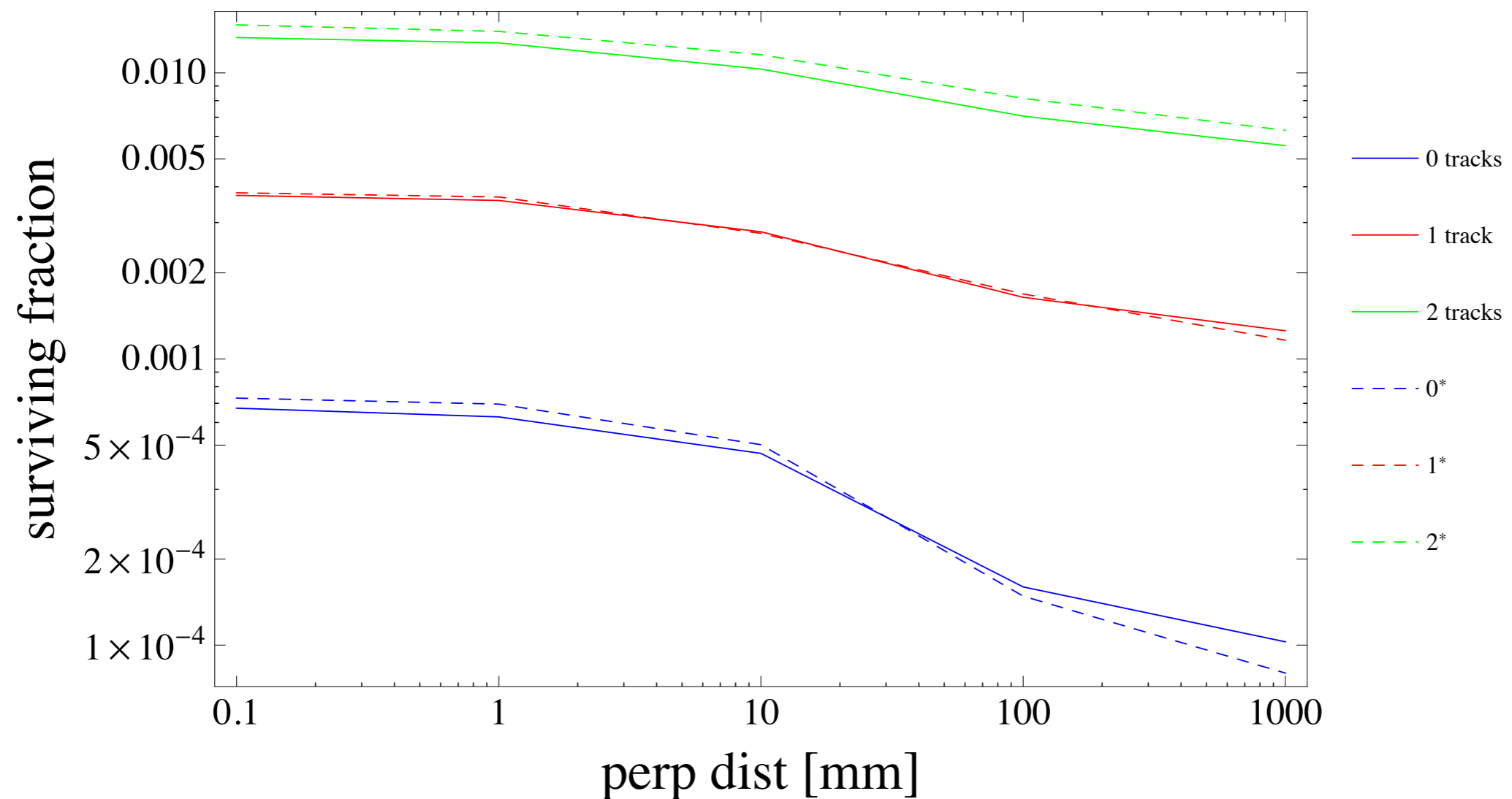


$$m_{\pi_d} = 5 \text{ GeV}$$

$$c\tau_{\pi_d} = 50 \text{ mm}$$

QCD bgd's

QCD 4-jet production in Pythia8



* – modified Pythia settings to increase QCD contribution

What will we learn
from run II?

Collider-reach

w/ Gavin Salam (CERN)

estimates of the reach of future colliders
based on existing limits

www.cern.ch/collider-reach

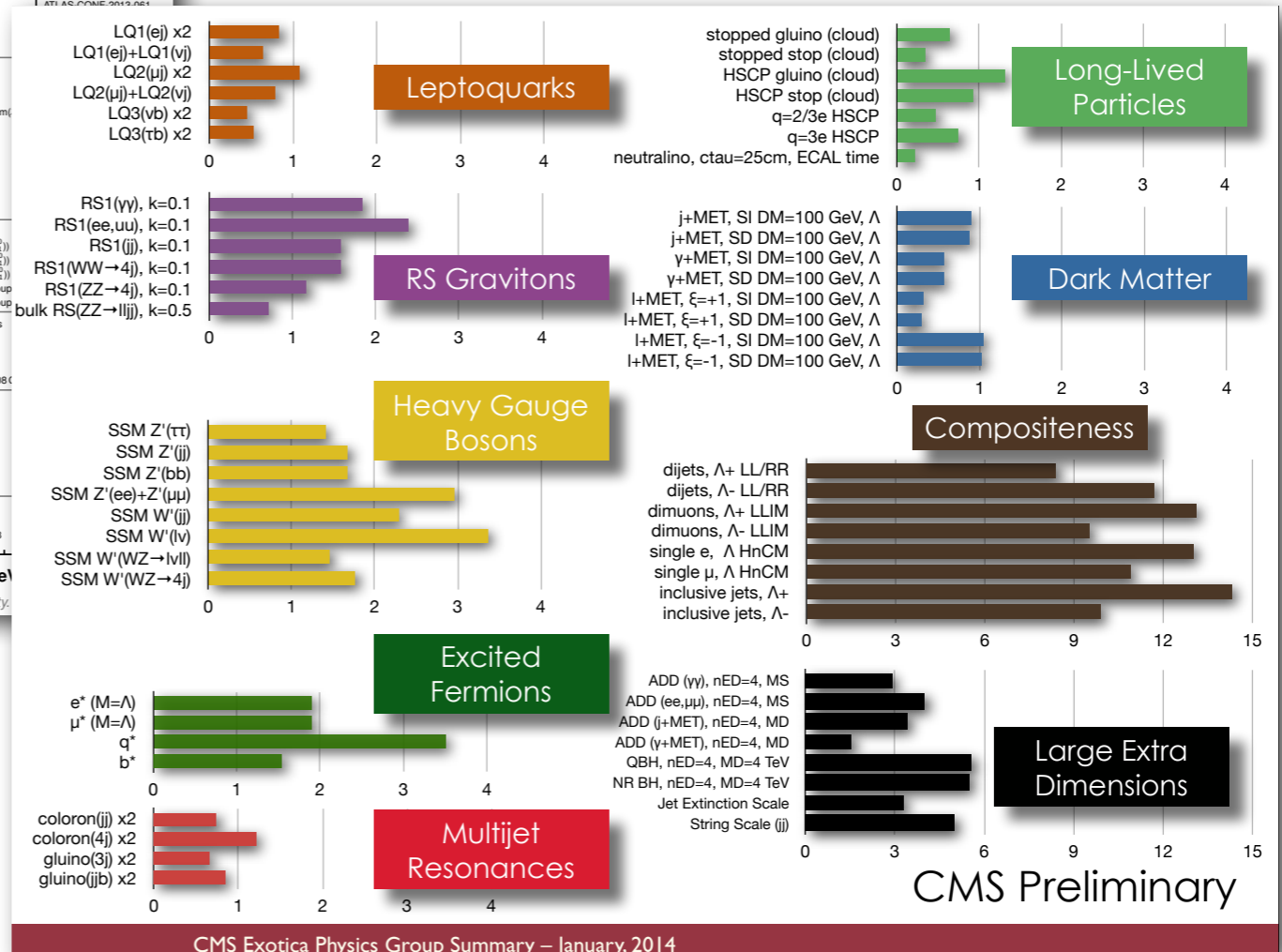
There are already many well-designed searches

ATLAS SUSY Searches* - 95% CL Lower Limits
 Status: SUSY 2013 ATLAS Preliminary
 $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$ $\sqrt{s} = 7, 8 \text{ TeV}$

Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	Reference	
Inclusive Searches	MSUGRA/CMSSM	0	2-6 jets	Yes	20.3	\tilde{g}, \tilde{g} 1.7 TeV	ATLAS-CONF-2013-047
	MSUGRA/CMSSM	1 e, μ	3-6 jets	Yes	20.3	\tilde{g} 1.2 TeV	ATLAS-CONF-2013-062
	MSUGRA/CMSSM	0	7-10 jets	Yes	20.3	\tilde{g} 1.1 TeV	1308.1841
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	\tilde{q} 740 GeV	ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	1 e, μ	2-6 jets	Yes	20.3	\tilde{g} 1.3 TeV	ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	1 e, μ	3-6 jets	Yes	20.3	\tilde{g} 1.18 TeV	ATLAS-CONF-2013-062
	GMSB (\tilde{L} NLSP)	2 e, μ	0-3 jets	-	20.3	\tilde{g} 1.12 TeV	ATLAS-CONF-2013-089
	GMSB (\tilde{L} NLSP)	2 e, μ	2-4 jets	Yes	4.7	\tilde{g} 1.24 TeV	1208.4688
	GMSB (\tilde{L} NLSP)	1-2 τ	0-2 jets	Yes	20.7	\tilde{g} 1.4 TeV	ATLAS-CONF-2013-026
	GGM (bino NLSP)	2 γ	-	Yes	4.8	\tilde{g} 1.07 TeV	1209.0753
3 rd gen. squarks	$\tilde{g} \rightarrow b\tilde{b}^0$	0	3 b	Yes	20.1	\tilde{g} 1.2 TeV	$m(\tilde{t}_1) < 600 \text{ GeV}$
	$\tilde{g} \rightarrow t\tilde{t}^0$	0	7-10 jets	Yes	20.3	\tilde{g} 1.1 TeV	$m(\tilde{t}_1) < 350 \text{ GeV}$
	$\tilde{g} \rightarrow t\tilde{t}^0$	0-1 e, μ	3 b	Yes	20.1	\tilde{g} 1.34 TeV	$m(\tilde{t}_1) < 400 \text{ GeV}$
	$\tilde{g} \rightarrow b\tilde{b}^0$	0-1 e, μ	3 b	Yes	20.1	\tilde{g} 1.3 TeV	$m(\tilde{t}_1) < 300 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{t}_1^0$	2 e, μ (SS)	0-3 b	Yes	20.1	\tilde{b}_1 100-620 GeV	$m(\tilde{t}_1) < 90 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{t}_1^0$	2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{b}_1 275-430 GeV	$m(\tilde{t}_1) = 2 m(\tilde{t}_2)$
	$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow b\tilde{t}_1^+$	1-2 e, μ	1-2 b	Yes	4.7	\tilde{t}_1 110-167 GeV	$m(\tilde{t}_1) = 55 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow W\tilde{b}_1^0$	2 e, μ	0-2 jets	Yes	20.3	\tilde{t}_1 130-220 GeV	$m(\tilde{t}_1) = m(\tilde{t}_2) - m(W) - 50 \text{ GeV}, m(\tilde{t}_1) < m(\tilde{t}_2)$
	$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t\tilde{t}_1^0$	2 e, μ	2 jets	Yes	20.3	\tilde{t}_1 225-525 GeV	$m(\tilde{t}_1) < 60 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow b\tilde{t}_1^+$	1 e, μ	1 b	Yes	20.1	\tilde{t}_1 150-580 GeV	$m(\tilde{t}_1) < 200 \text{ GeV}, m(\tilde{t}_1) - m(\tilde{t}_2) = 5 \text{ GeV}$
EW direct	$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{t}_1^0$	0	2 b	Yes	20.5	\tilde{t}_1 200-610 GeV	$m(\tilde{t}_1) = 0 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{t}_1^0$	0	2 b	Yes	20.5	\tilde{t}_1 320-660 GeV	$m(\tilde{t}_1) = 0 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.7	\tilde{t}_1 90-200 GeV	$m(\tilde{t}_1) + m(\tilde{t}_2) < 85 \text{ GeV}$
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	20.7	\tilde{t}_2 500 GeV	$m(\tilde{t}_1) > 150 \text{ GeV}$
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	20.7	\tilde{t}_2 271-520 GeV	$m(\tilde{t}_1) = m(\tilde{t}_2) + 180 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{t}_1^0$	2 e, μ	0	Yes	20.3	\tilde{t}_1 85-315 GeV	$m(\tilde{t}_1) = 0 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{t}_1^0$	2 e, μ	0	Yes	20.3	\tilde{t}_1 125-450 GeV	$m(\tilde{t}_1) = 0 \text{ GeV}, m(\tilde{t}_1) = 0.5(m(\tilde{t}_1) + m(\tilde{t}_2))$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{t}_1^0$	2 τ	-	Yes	20.7	\tilde{t}_1 180-330 GeV	$m(\tilde{t}_1) = 0 \text{ GeV}, m(\tilde{t}_1) = 0.5(m(\tilde{t}_1) + m(\tilde{t}_2))$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{t}_1^0$	3 e, μ	0	Yes	20.7	\tilde{t}_1 315 GeV	$m(\tilde{t}_1) = m(\tilde{t}_2), m(\tilde{t}_1) = 0, m(\tilde{t}_1) = 0.5(m(\tilde{t}_1) + m(\tilde{t}_2))$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{t}_1^0$	3 e, μ	0	Yes	20.7	\tilde{t}_1 285 GeV	$m(\tilde{t}_1) = m(\tilde{t}_2), m(\tilde{t}_1) = 0, \text{ sleptons decouple}$
Long-lived particles	Direct $\tilde{\chi}_1^0\tilde{\chi}_1^0$ prod., long-lived $\tilde{\chi}_1^0$	Disapp. trk	1 jet	Yes	20.3	$\tilde{\chi}_1^0$ 270 GeV	$m(\tilde{\chi}_1^0) - m(\tilde{\chi}_2^0) = 160 \text{ MeV}, \tau(\tilde{\chi}_1^0) = 0.2 \text{ ns}$
	Stable, stopped \tilde{R} -hadron	0	1-5 jets	Yes	22.9	\tilde{g} 832 GeV	$m(\tilde{\chi}_1^0) = 100 \text{ GeV}, 10 \mu\text{s} < \tau(\tilde{g}) < 1000 \text{ s}$
	GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tau(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$	1-2 μ	-	-	15.9	$\tilde{\tau}$ 475 GeV	$10 < \tau_{\text{stop}} < 50$
	GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma G$, long-lived $\tilde{\chi}_1^0$	2 γ	-	Yes	4.7	$\tilde{\chi}_1^0$ 230 GeV	$0.4 < \tau(\tilde{\chi}_1^0) < 2 \text{ ns}$
	$\tilde{q}\tilde{q}, \tilde{\chi}_1^0 \rightarrow q\tilde{q}\mu$ (RPV)	1 μ , displ. vtx	-	-	20.3	\tilde{q} 1.0 TeV	$1.5 < \tau < 156 \text{ mm}, \text{BR}(\mu) = 1, m(\tilde{\chi}_1^0) = 108 \text{ GeV}$
	LFV $pp \rightarrow \tilde{\nu}_e + X, \tilde{\nu}_e \rightarrow e + \mu$	2 e, μ	-	-	4.6	$\tilde{\nu}_e$ 1.61 TeV	$A_{131} = 0.10, A_{132} = 0.05$
	LFV $pp \rightarrow \tilde{\nu}_e + X, \tilde{\nu}_e \rightarrow e(\mu) + \tau$	1 $e, \mu + \tau$	-	-	4.6	$\tilde{\nu}_e$ 1.1 TeV	$A_{131} = 0.10, A_{132} = 0.05$
	Bilinear RPV CMSSM	1 e, μ	7 jets	Yes	4.7	\tilde{g}, \tilde{g} 1.2 TeV	$m(\tilde{q}) = m(\tilde{g}), c_{\tau LSP} < 1 \text{ mm}$
	$\tilde{\chi}_1^0\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow ee\tilde{\nu}_e, e\mu\tilde{\nu}_e$	4 e, μ	-	Yes	20.7	$\tilde{\chi}_1^0$ 760 GeV	$m(\tilde{\chi}_1^0) > 300 \text{ GeV}, A_{122} > 0$
	$\tilde{\chi}_1^0\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tilde{\nu}_e, e\tau\tilde{\nu}_e$	3 $e, \mu + \tau$	-	Yes	20.7	$\tilde{\chi}_1^0$ 350 GeV	$m(\tilde{\chi}_1^0) > 80 \text{ GeV}, A_{122} > 0$
$\tilde{g} \rightarrow q\tilde{q}$	0	6-7 jets	-	20.3	\tilde{g} 916 GeV	$\text{BR}(\tau) = \text{BR}(b) = \text{BR}(c) = 0\%$	
Other	$\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow b\tilde{s}$	2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{g} 880 GeV	
	Scalar gluon pair, sgluon $\rightarrow q\tilde{q}$	0	4 jets	-	4.6	sgluon 100-287 GeV	incl. limit from 1110.2693
	Scalar gluon pair, sgluon $\rightarrow t\tilde{t}$	2 e, μ (SS)	1 b	Yes	14.3	sgluon 800 GeV	
WIMP interaction (DS, Dirac χ)	0	mono-jet	Yes	10.5	M^* scale 704 GeV	$m(\chi) < 80 \text{ GeV}, \text{limit of } \sim 687 \text{ GeV for DS}$	

Legend: $\sqrt{s} = 7 \text{ TeV}$ full data (blue), $\sqrt{s} = 8 \text{ TeV}$ partial data (green), $\sqrt{s} = 8 \text{ TeV}$ full data (yellow)

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1σ theoretical signal cross section uncertainty.



CMS Exotica Physics Group Summary – January, 2014

How do we leverage that experience to estimate future reaches?

A rough way of doing it

Suppose ATLAS/CMS are currently sensitive to gluinos of 1250 GeV (95% CL_s , 8 TeV, 20 fb⁻¹)



Work out how many signal events that corresponds to



Find out for what gluino mass you would get the same number of signal events at 14 TeV with 300 fb⁻¹ (assume # of background events scales same way)

Too simplistic

Backgrounds may not scale in the same way as signal

New irreducible backgrounds may appear at higher scales

Reconstruction efficiencies may depend on mass scale

Detector effects (e.g. granularity), and run conditions (pileup) vary across energy scales and luminosities



It can't possibly work!

Too complicated

Calculating mass for constant # of signal events is pretty straightforward

But it still requires some work and setup

E.g. need cross section calculators for each new physics process (Prospino/Pythia/...), run them for a range of masses, etc.



an iPhone app?

$$\frac{N_{\text{signal-events}}(M_{\text{high}}^2, 14 \text{ TeV}, \text{Lumi})}{N_{\text{signal-events}}(M_{\text{low}}^2, 8 \text{ TeV}, 19 \text{ fb}^{-1})} = 1$$

Coupling constants & other prefactors mostly cancel in the ratio.

Dependence on M and on \sqrt{s} mostly comes about through parton distribution functions (PDFs) & simple dimensions.

Z' example

$$\hat{\sigma}_0(\hat{s}) = C \frac{\hat{s}}{(\hat{s} - M_{Z'}^2)^2 + \Gamma_{Z'}^2 M_{Z'}^2}$$

$$\frac{d\sigma}{dm^2} = \int dx_1 dx_2 [f_1(x_1) f_2(x_2)] \hat{\sigma}_0(\hat{s}) \delta(m^2 - \hat{s}^2),$$

$$= \sum_{ij} \left[\tau \int \frac{dx}{x} f_i(x) f_j(\tau/x) \right] \frac{C}{(m^2 - M_{Z'}^2)^2 + \Gamma_{Z'}^2 M_{Z'}^2}$$

$$\sigma \approx \int dm^2 \sum_{ij} \mathcal{L}_{ij}(m^2, s) C \frac{\pi}{\Gamma_{Z'} M_{Z'}} \delta(m^2 - M_{Z'}^2) \quad \Gamma_{Z'} \propto M_{Z'}$$

$$= \frac{1}{M_{Z'}^2} \sum_{ij} C' \mathcal{L}_{ij}(M_{Z'}^2, s)$$

“1/M² x parton-lumi”

$$= N(M_{Z'}, s)$$

Instead of cross section ratio, use **parton luminosity ratio**

Equation we solve to find M_{high} is then

$$\frac{\mathcal{L}_{ij}(M_{\text{high}}^2, s_{\text{high}})}{\mathcal{L}_{ij}(M_{\text{low}}^2, s_{\text{low}})} \times \frac{\text{lumi}_{\text{high}}}{\text{lumi}_{\text{low}}} = \frac{M_{\text{high}}^2}{M_{\text{low}}^2}$$

The tools we use for this are
LHAPDF and HOPPET
most plots with MSTW2008 NNLO PDFs

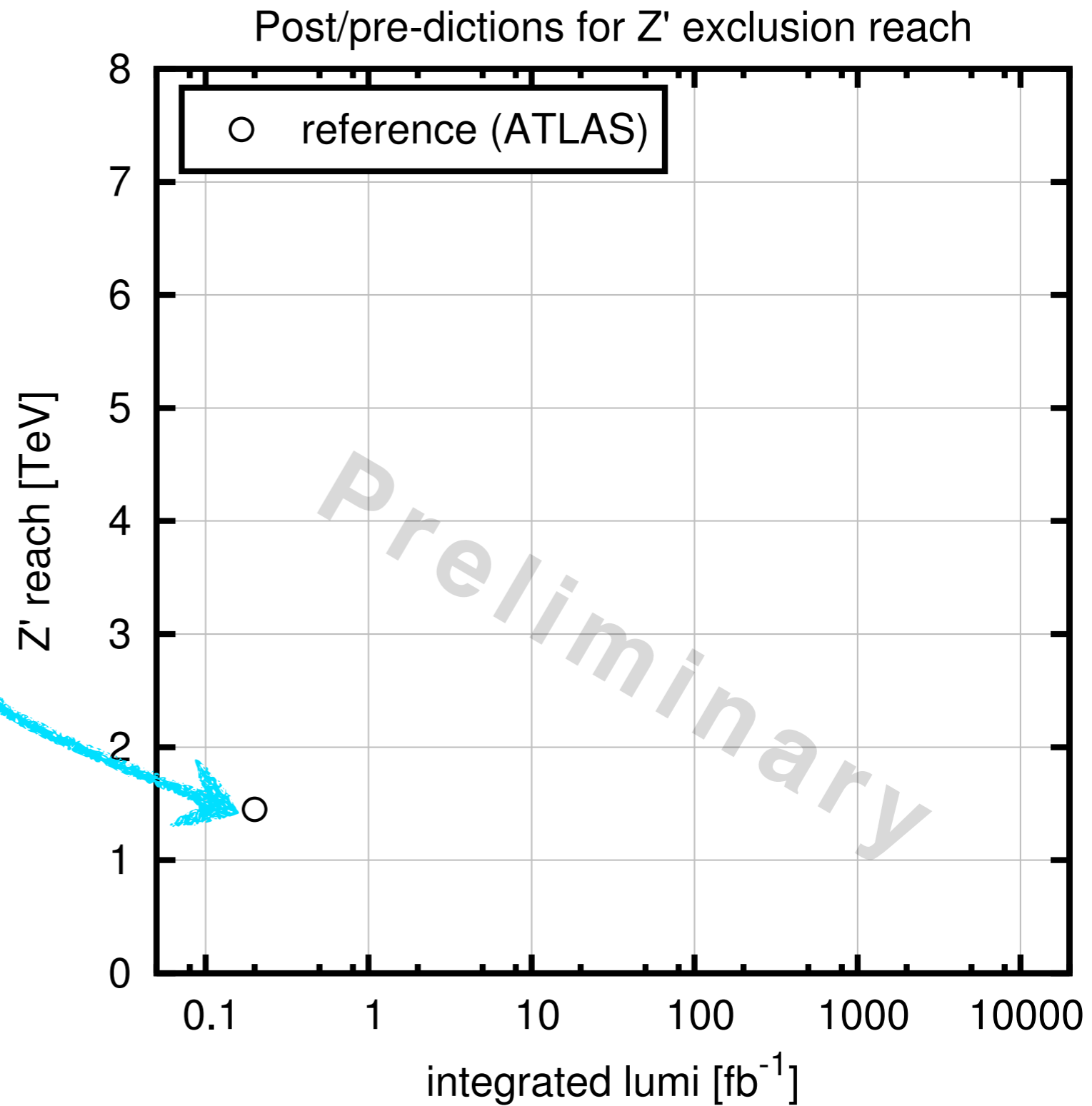
$$\mathcal{L}_{ij}(M^2, s) = \int_{\tau}^1 \frac{dx}{x} x f_i(x, M^2) \frac{\tau}{x} f_j\left(\frac{\tau}{x}, M^2\right) \quad \tau \equiv \frac{M^2}{s}$$

i & j parton

Does it work?

Try a Z' search. Take a baseline analysis:

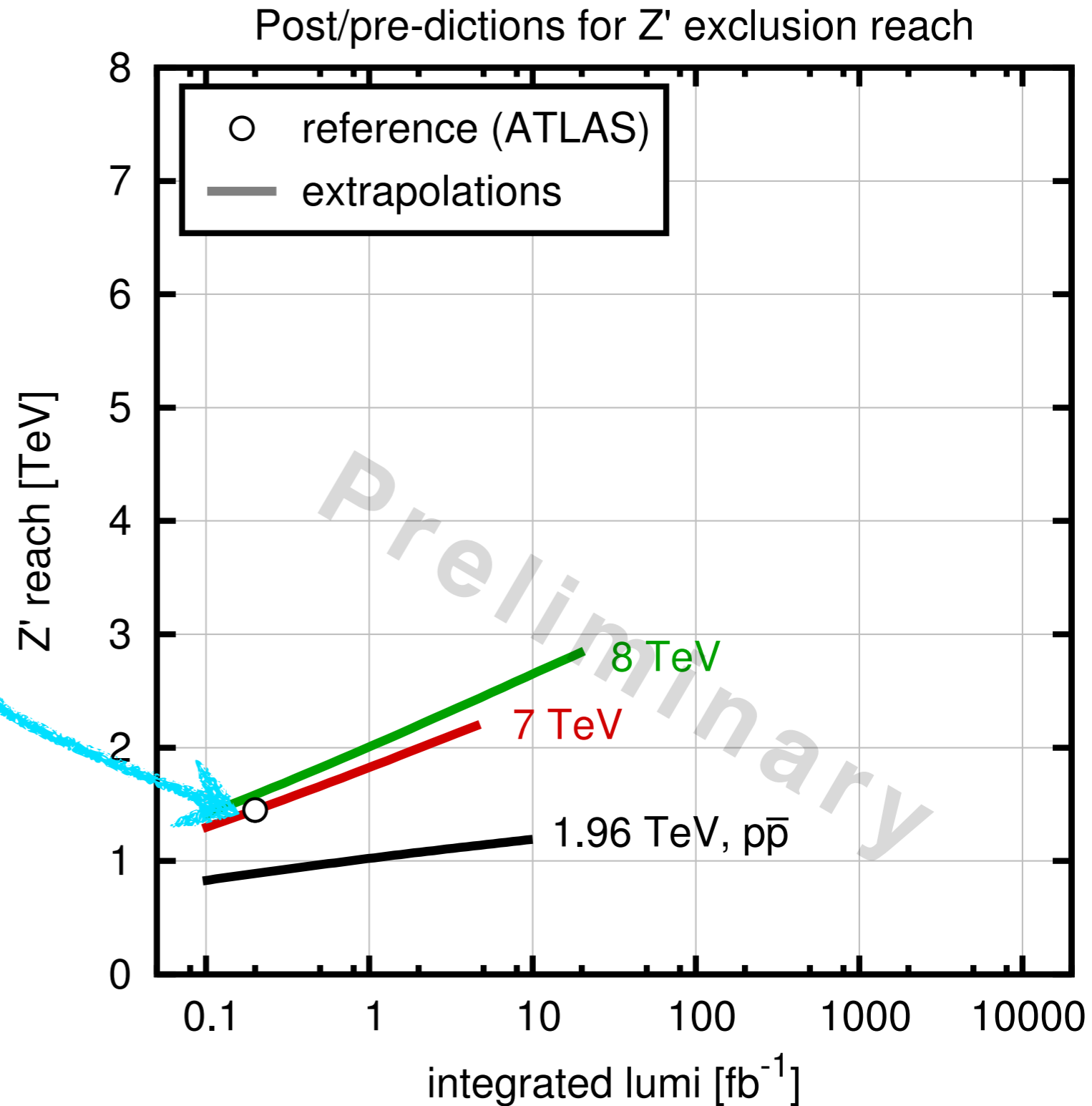
ATLAS,
 0.2 fb^{-1} @ 7 TeV
excludes $M < 1450 \text{ GeV}$



Try a Z' search. Take a baseline analysis:

ATLAS,
 0.2 fb^{-1} @ 7 TeV
excludes $M < 1450 \text{ GeV}$

“Predict” exclusions
at other lumis &
energies (assume $q\bar{q}$)

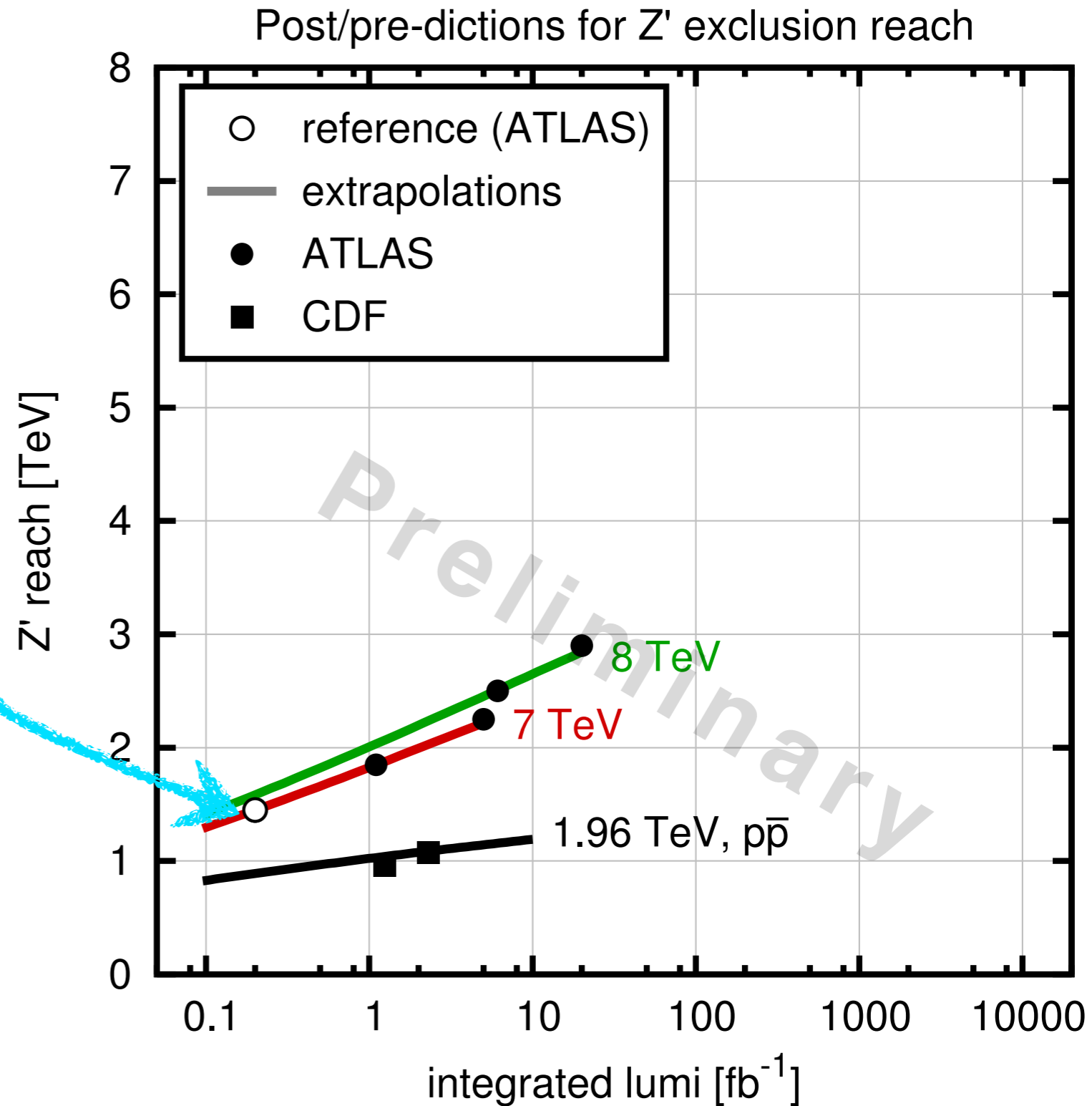


Try a Z' search. Take a baseline analysis:

ATLAS,
 0.2 fb^{-1} @ 7 TeV
excludes $M < 1450 \text{ GeV}$

“Predict” exclusions
at other lumis &
energies (assume $q\bar{q}$)

Compare to actual
exclusions

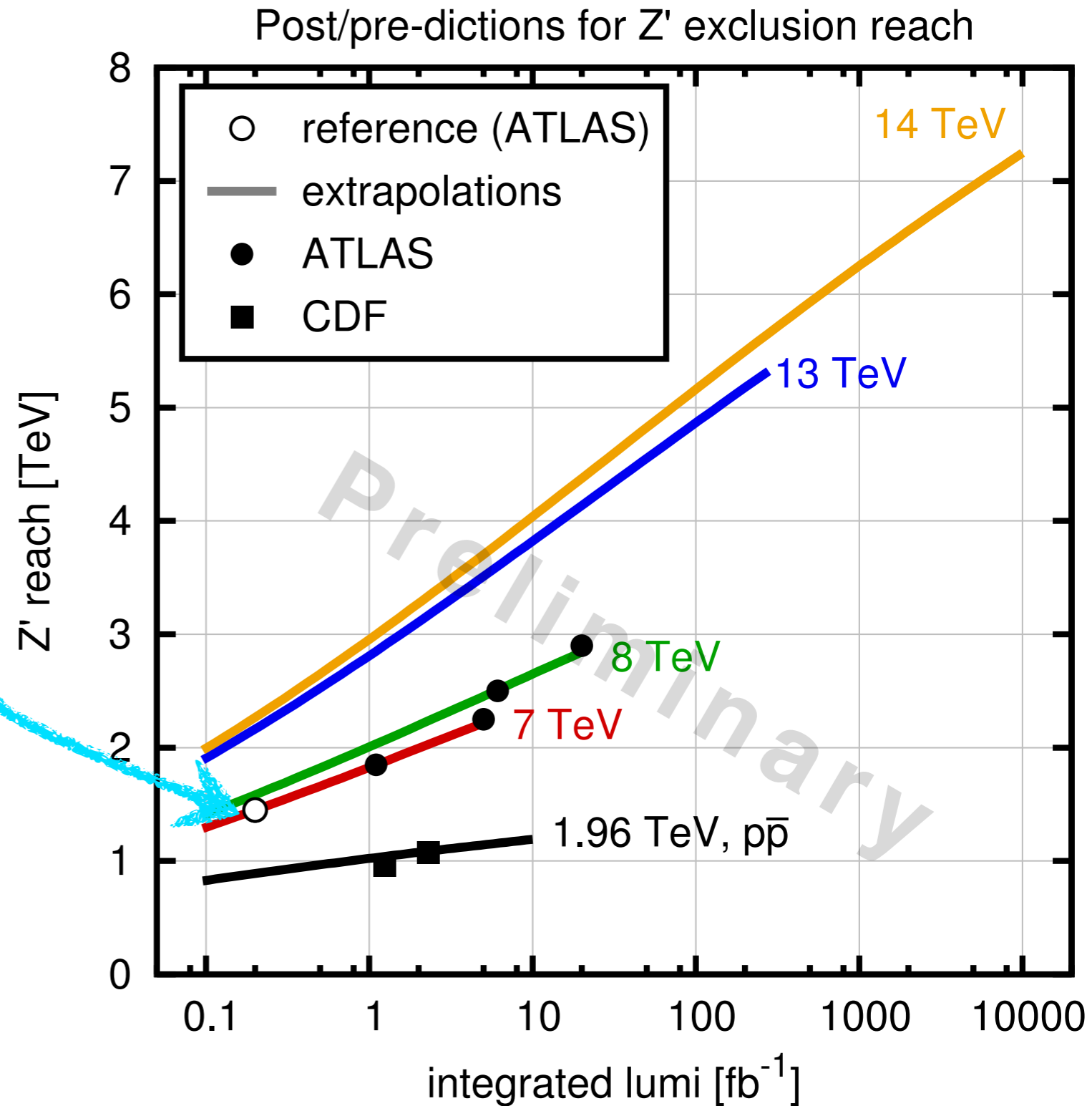


Try a Z' search. Take a baseline analysis:

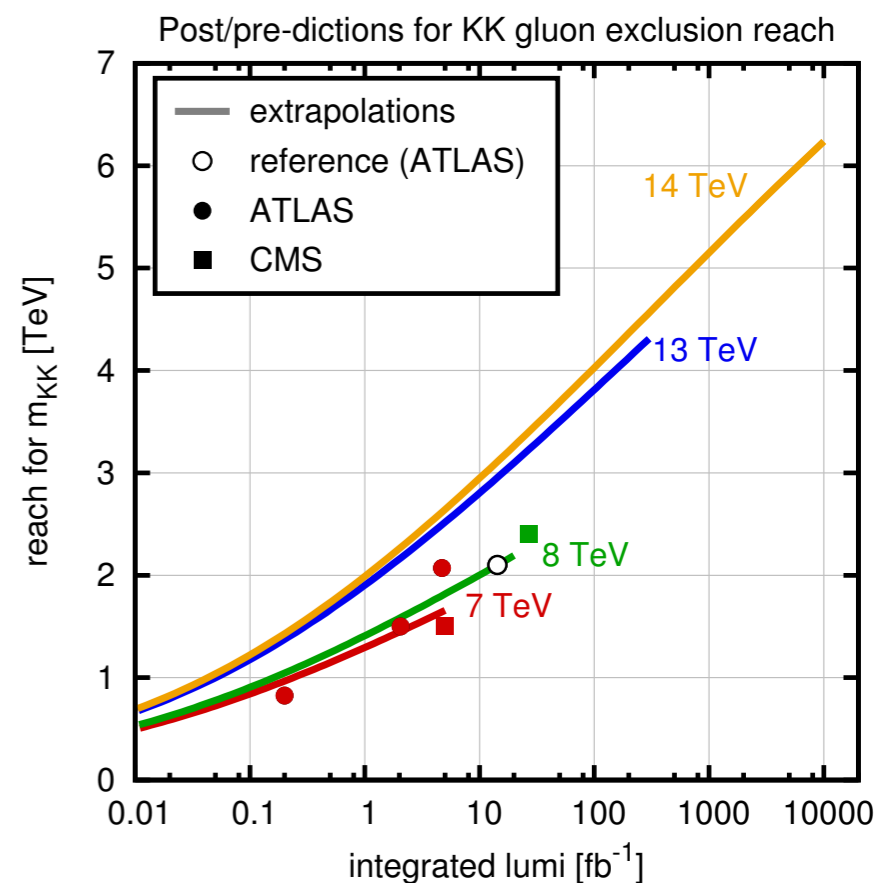
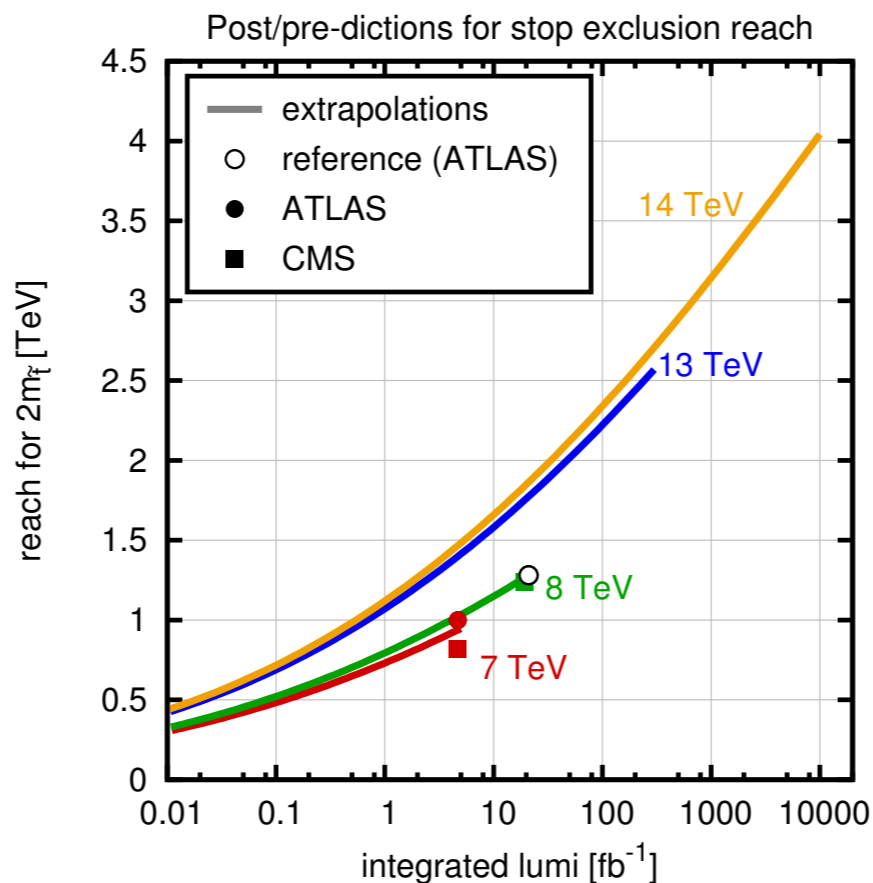
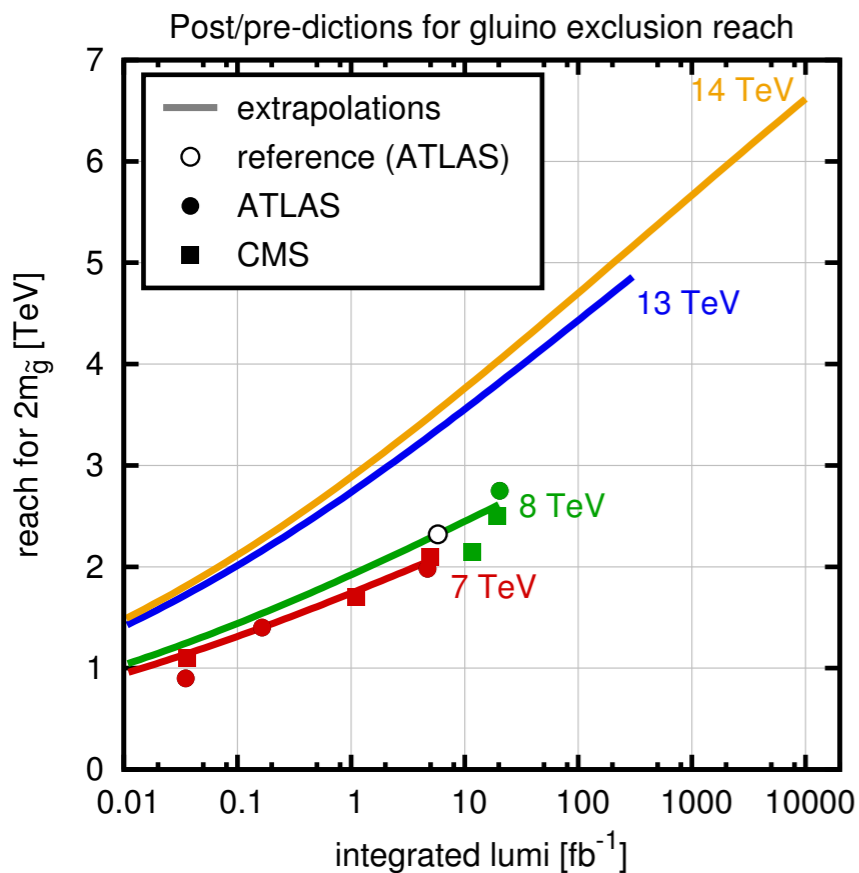
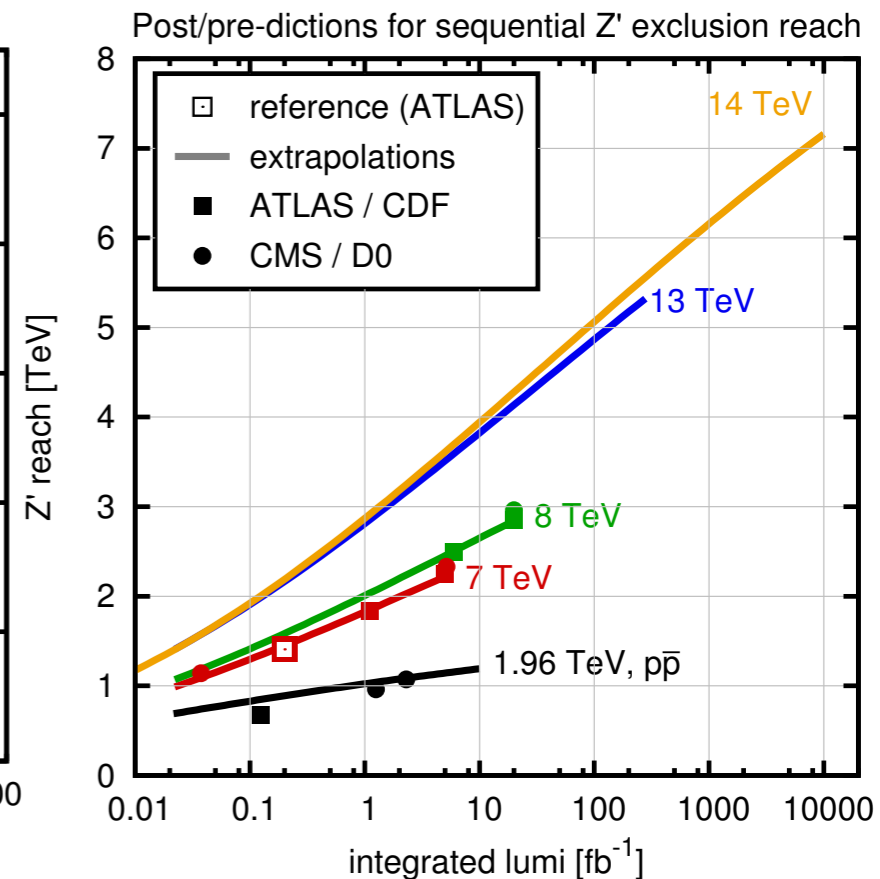
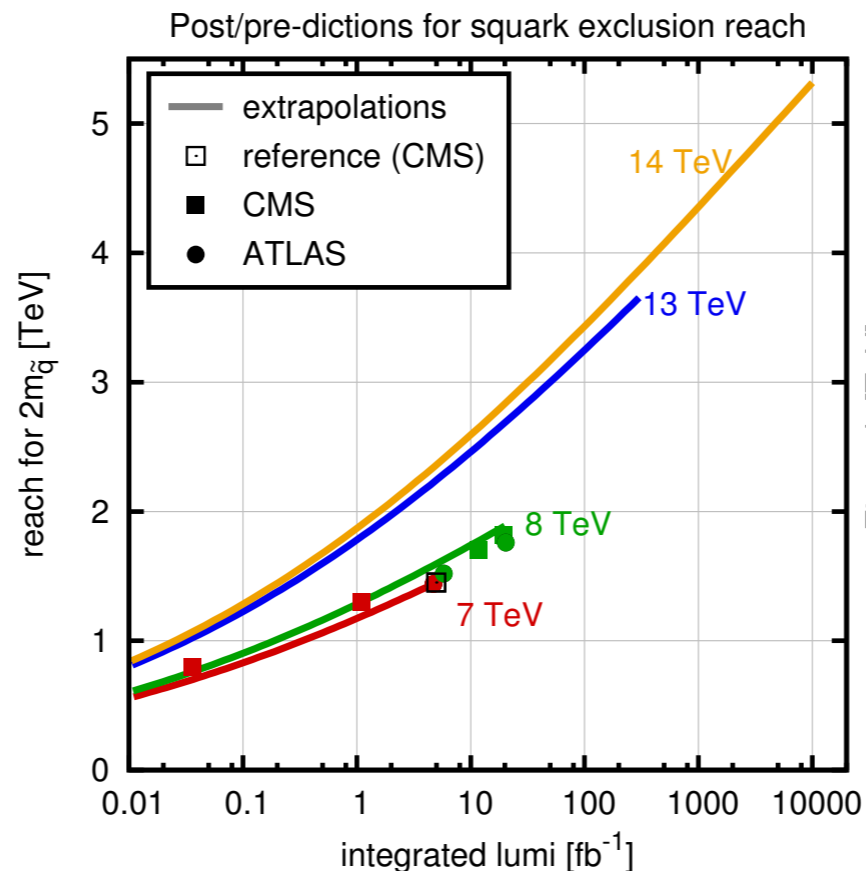
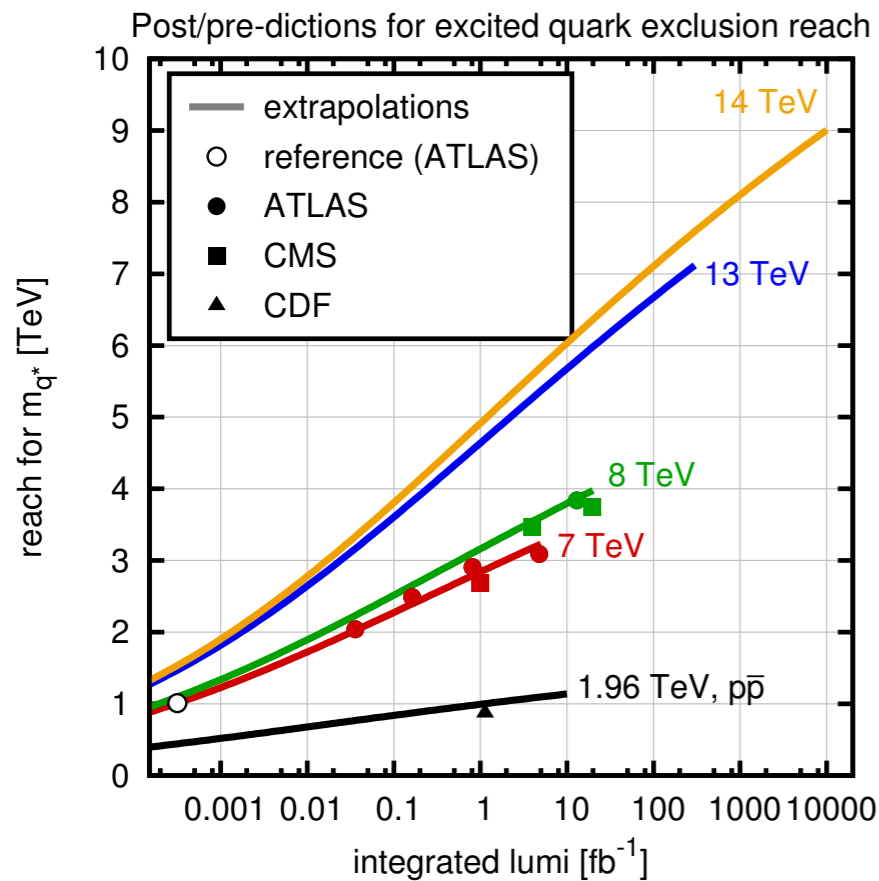
ATLAS,
 0.2 fb^{-1} @ 7 TeV
excludes $M < 1450 \text{ GeV}$

“Predict” exclusions
at other lumis &
energies (assume $q\bar{q}$)

Compare to actual
exclusions



Maybe it only works so well because it's a simple search?
(Signal & Bkgd are both $q\bar{q}$ driven)



From your iPhone/Android
(or a generic browser)
cern.ch/collider-reach

Collider 1: CoM energy

8

TeV, integrated luminosity

20

fb^{-1}

Collider 2: CoM energy

14

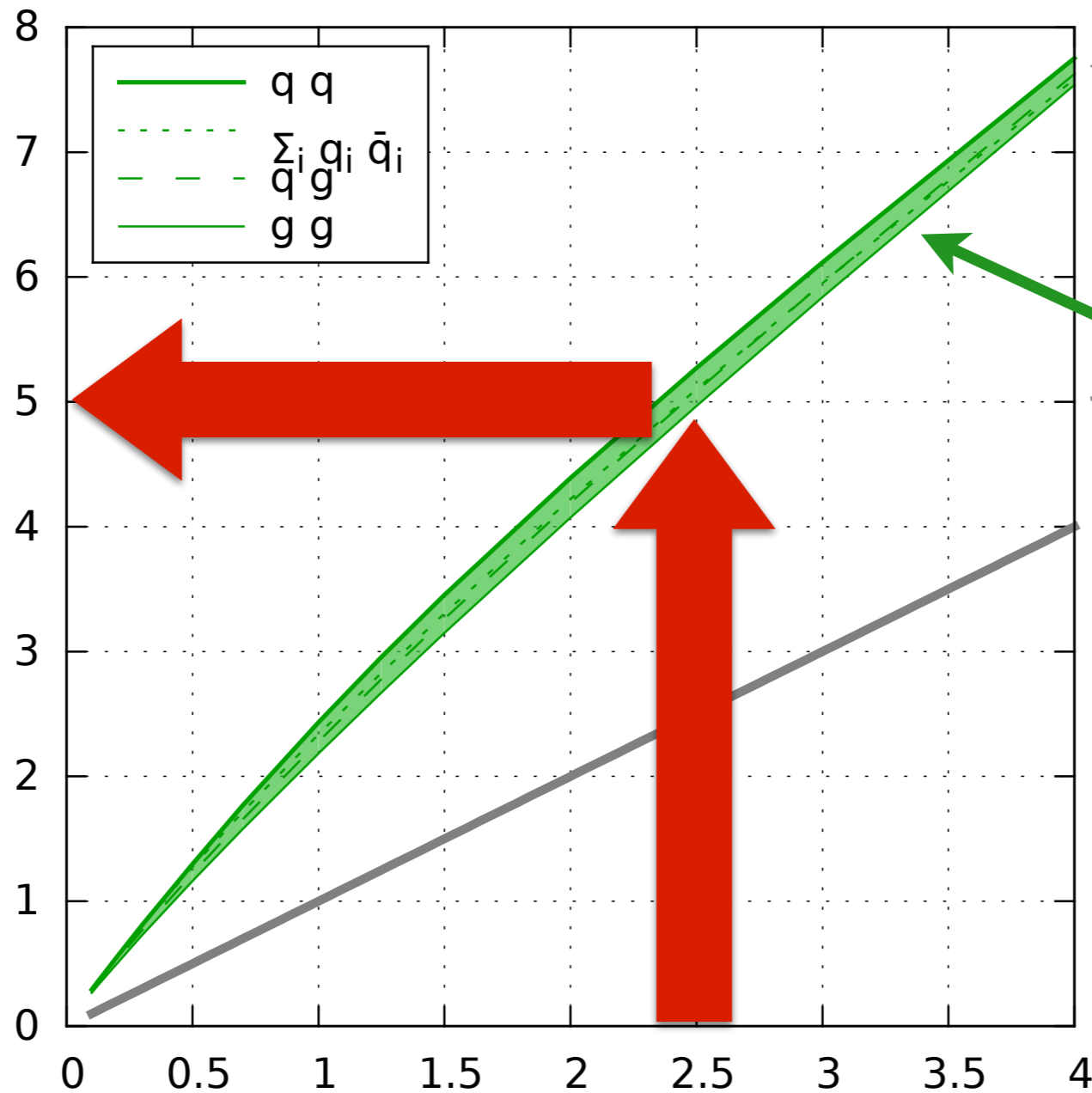
TeV, integrated luminosity

300

fb^{-1}

PDF:

MSTW2008nnlo68cl



Mass [TeV] at
collider #2

Spread of
partonic
channels
(assume same
channel for
S & B)

Mass [TeV] at collider #1

The Collider Reach tool gives you a quick (and dirty) estimate of the relation between the mass reaches of different proton-proton collider setups.

Collider 1: CoM energy TeV, integrated luminosity fb⁻¹

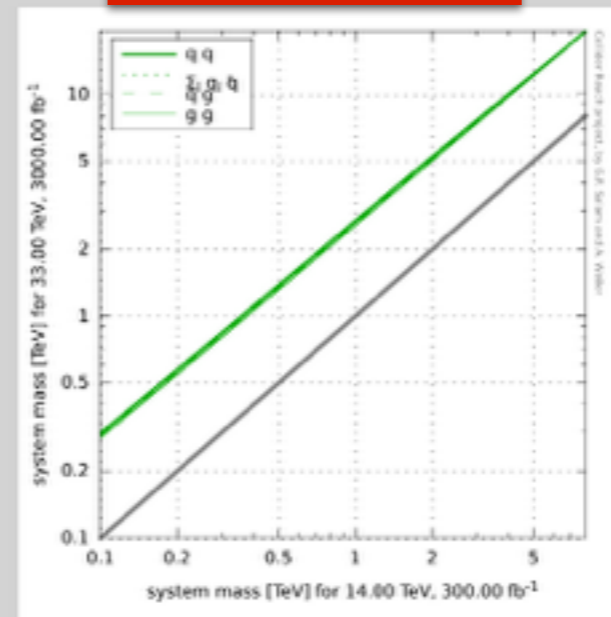
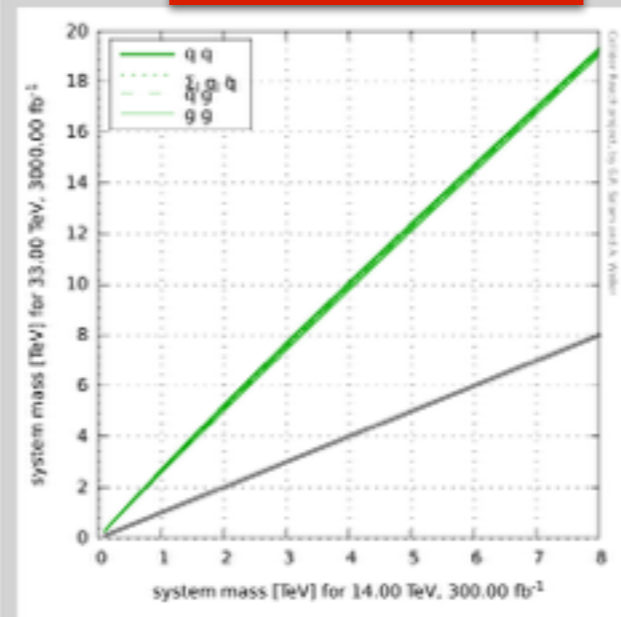
Collider 2: CoM energy TeV, integrated luminosity fb⁻¹

PDF:

linear plot

log-log plot

Plots

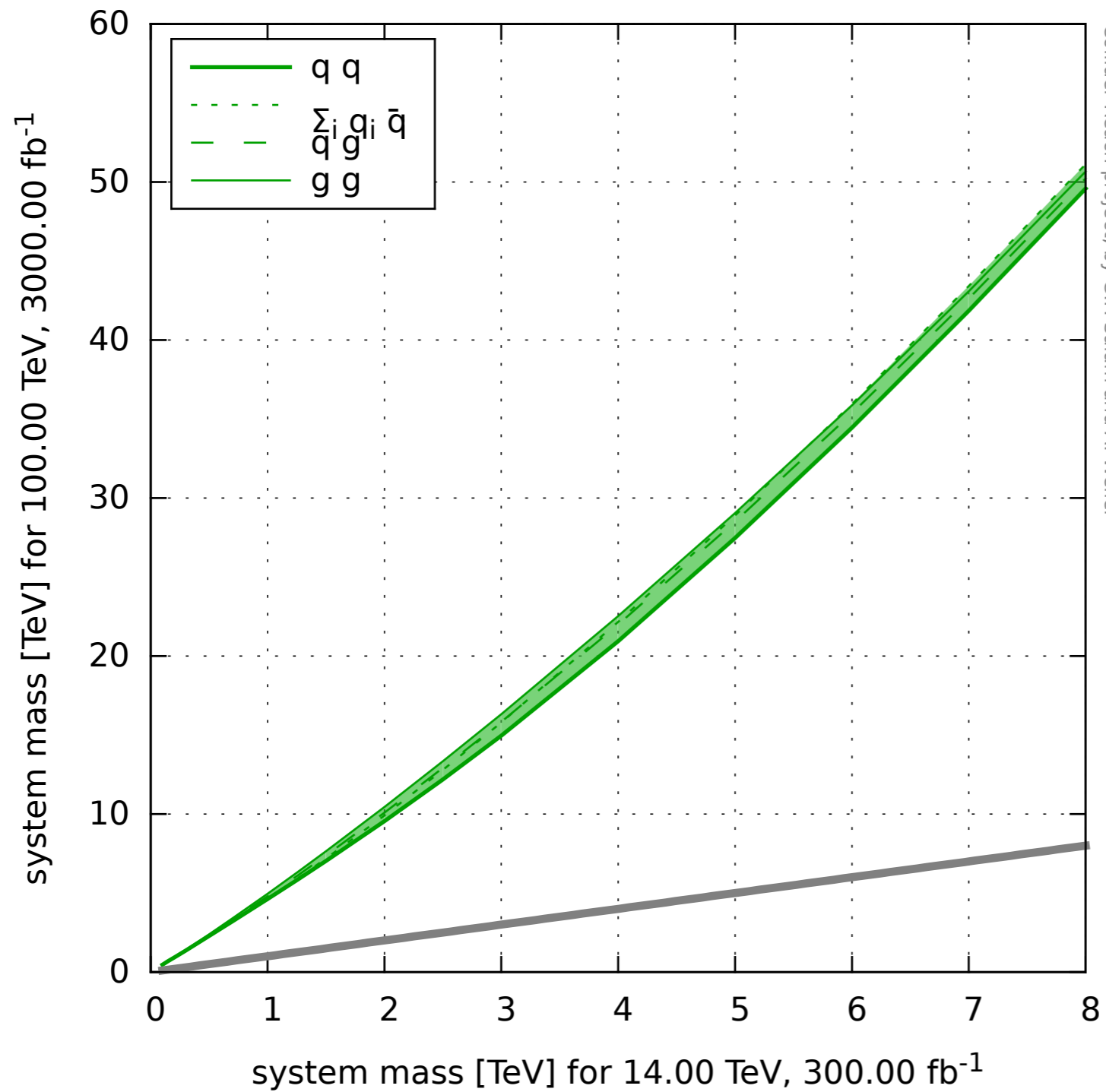


Download: [collider.pdf](#), [colliderloglog.pdf](#), plot generation [log file](#)

The PDF choice was CT10nlo.LHgrid

Original mass	gg	qg	allqq	qqbar
100.	283.	291.	298.	297.
125.	350.	359.	368.	367.
150.	416.	427.	438.	437.
200.	547.	562.	576.	575.
300.	806.	827.	848.	847.
500.	1317.	1350.	1386.	1382.
700.	1822.	1866.	1916.	1907.
1000.	2570.	2628.	2702.	2680.
1250.	3188.	3256.	3349.	3314.
1500.	3802.	3879.	3990.	3939.
2000.	5018.	5110.	5251.	5169.
2500.	6223.	6327.	6488.	6380.
3000.	7417.	7530.	7703.	7578.
4000.	9782.	9904.	10082.	9945.
5000.	12120.	12246.	12417.	12284.
6000.	14439.	14565.	14726.	14601.
7000.	16748.	16871.	17021.	16905.
8000.	19053.	19169.	19310.	19206.

14 TeV_{300 fb⁻¹} → 100 TeV_{3 ab⁻¹}



The PDF choice was CT10nlo.LHgrid

Original mass	gg	qg	allqq	qqbar
100.	469.	465.	462.	457.
125.	585.	579.	575.	568.
150.	702.	693.	687.	679.
200.	937.	923.	912.	902.
300.	1414.	1386.	1365.	1350.
500.	2394.	2332.	2279.	2261.
700.	3401.	3300.	3206.	3194.
1000.	4956.	4793.	4619.	4640.
1250.	6287.	6072.	5818.	5892.
1500.	7647.	7382.	7038.	7187.
2000.	10444.	10090.	9552.	9905.
2500.	13337.	12908.	12185.	12781.
3000.	16319.	15833.	14954.	15795.
4000.	22531.	21986.	20933.	22162.
5000.	29050.	28508.	27467.	28894.
6000.	35863.	35366.	34451.	35960.
7000.	43079.	42620.	41854.	43411.
8000.	50671.	50230.	49590.	51132.

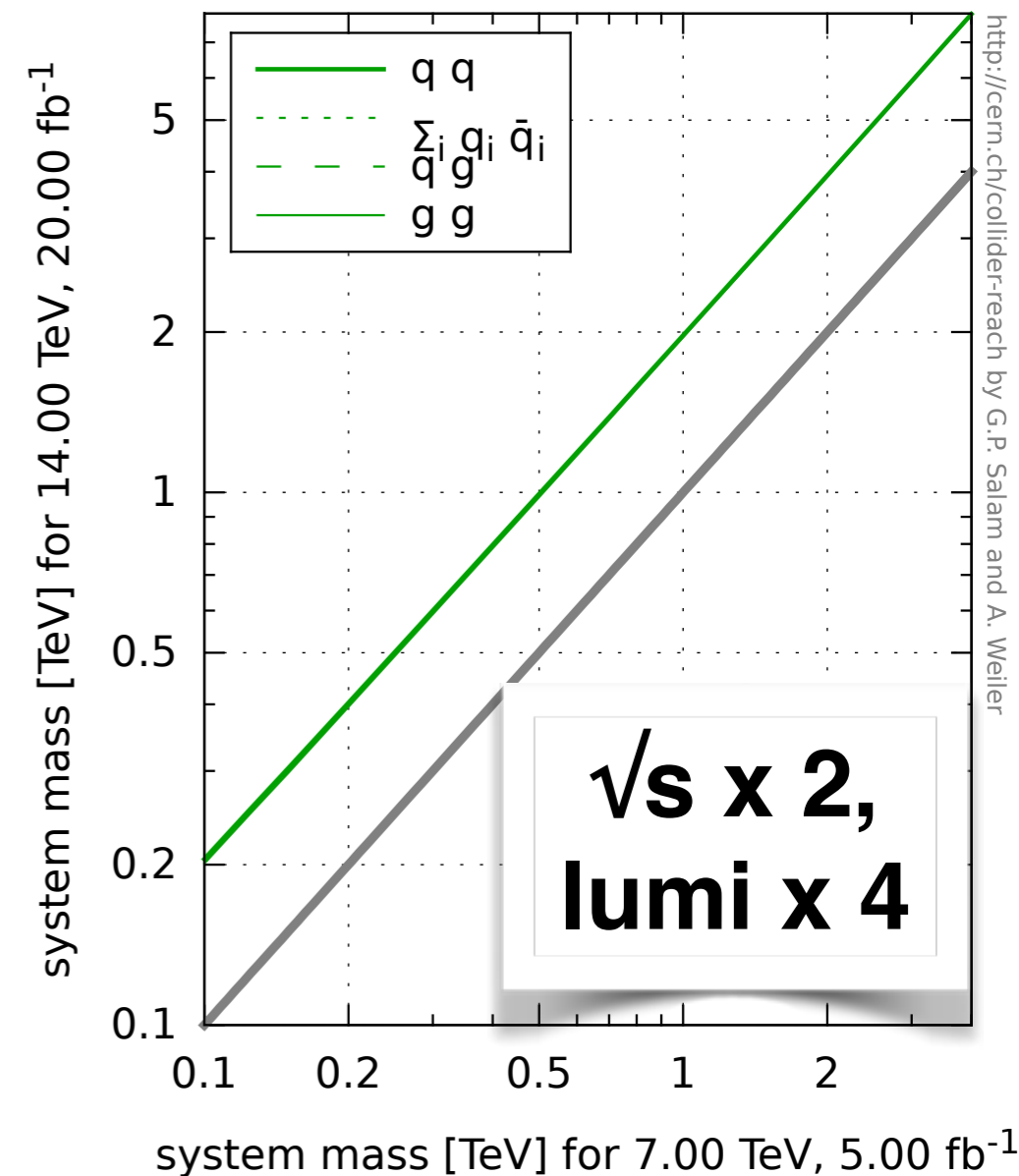
Rule of Thumb #1

Increase collider energy by **X**
& increase luminosity by a factor **X²**

→ **reach goes up by a factor X**

Because you keep same Bjorken-x & luminosity increase compensates for 1/mass² scaling of cross sections

PDF scaling variations are small effect



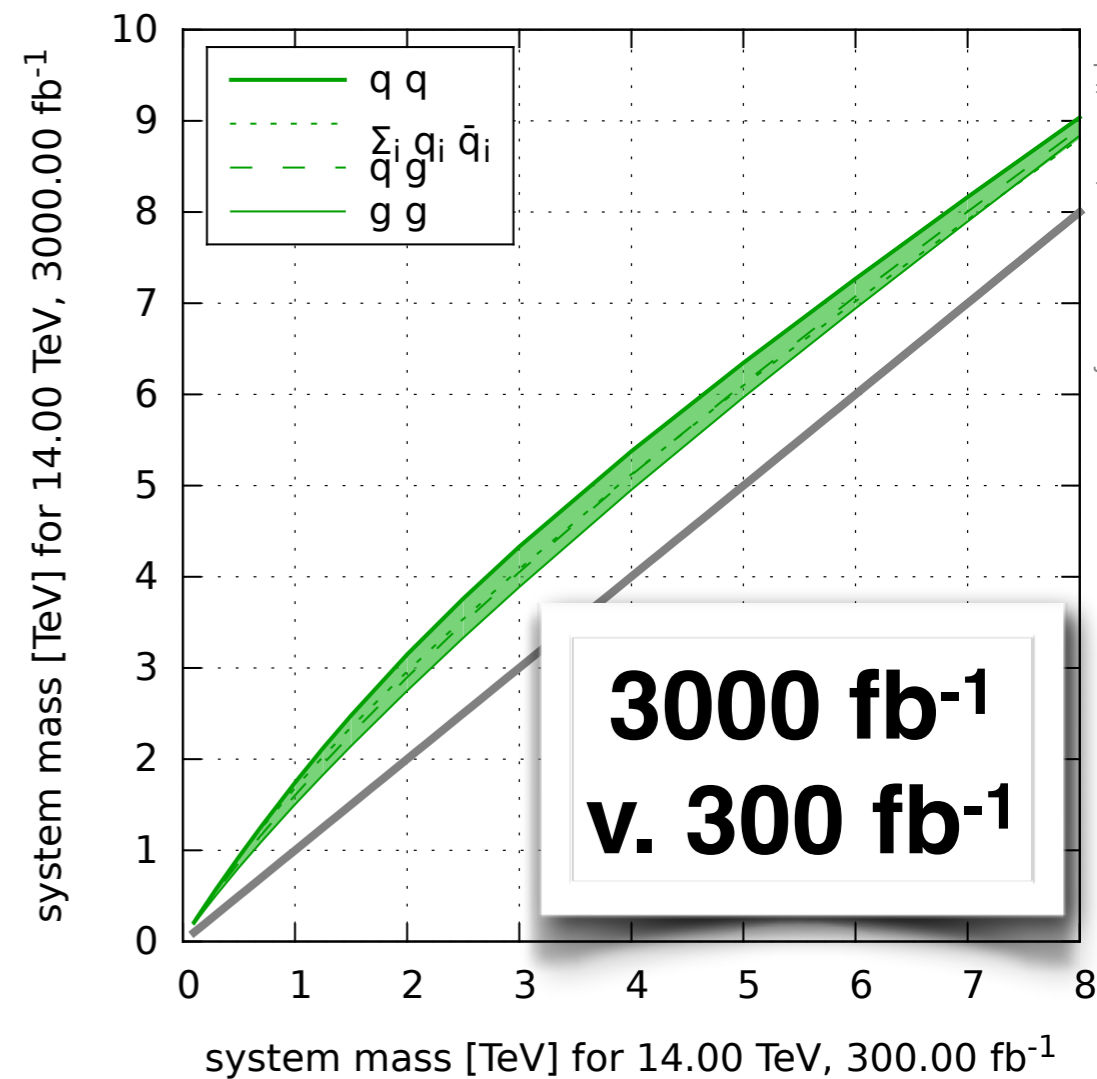
Rule of Thumb #2

(apparently not widely known previously)

Increase luminosity by factor 10
→ **reach increases by constant**
 $\Delta m \approx 0.07\sqrt{s}$

i.e. for $\sqrt{s}=14$ TeV, reach goes by up
1 TeV

No deep reason — a somewhat
random characteristic of large-x PDFs.
Only holds for $0.15 \lesssim M/\sqrt{s} \lesssim 0.6$



Consequence of rule #2

(may be a bit fragile & only for $S \approx B$)

Exclusion is $2\text{-}\sigma$

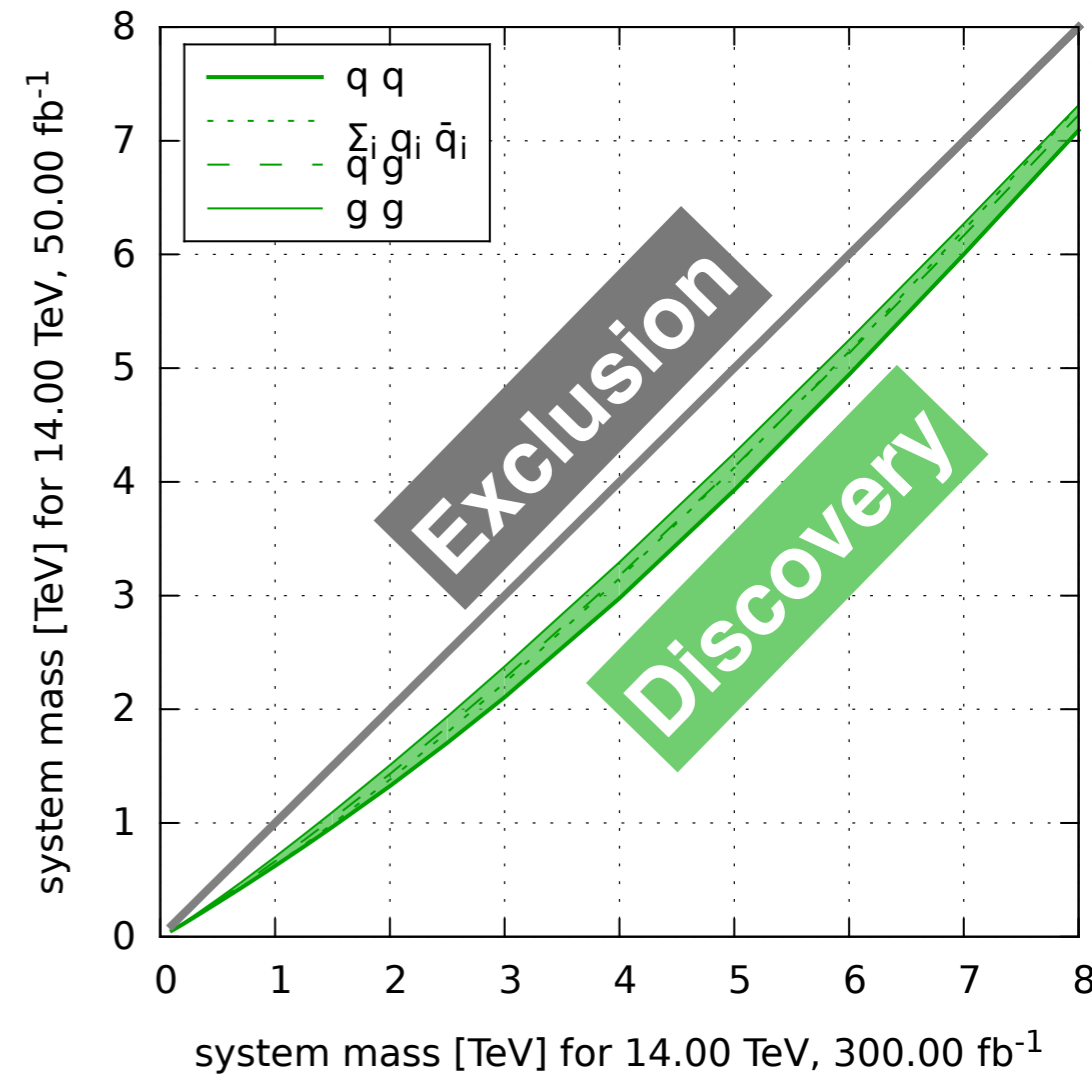
Discovery is $5\text{-}\sigma$

Need $(5/2)^2 = 6.25$ increase in lumi to go from one to the other.

Using rule #2:

discovery reach is about $0.05\sqrt{s}$
below exclusion reach

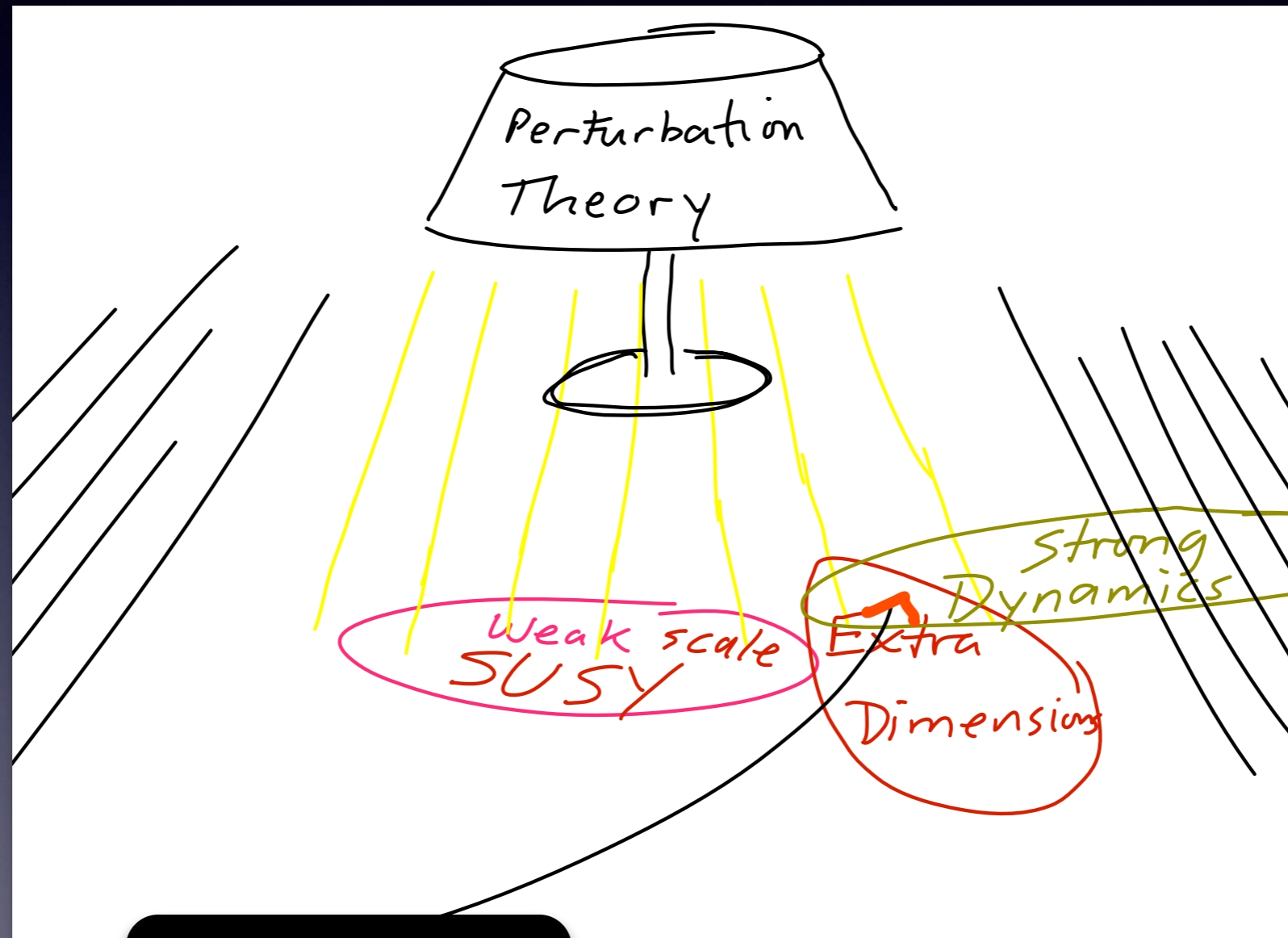
~ 0.8 TeV at 14 TeV



What else?

- We are preparing for run II
- Maybe we won't see a natural resolution of the hierarchy problem
- Need to cover all bases: consider more exotic signatures, think hard about triggers

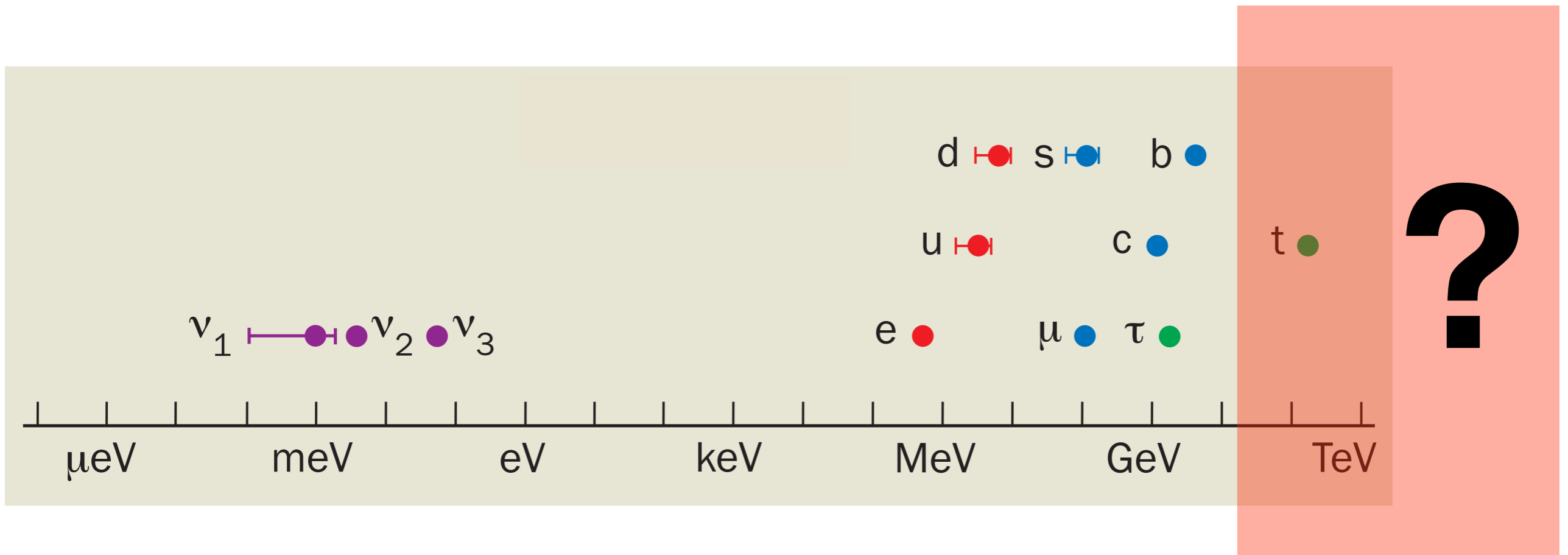
Looking under the lamp-post



AdS/CFT

Home-work

Find the TeV theory
beyond the SM



Conclusion

LHC₁₄ will be exciting (tuning $\propto E^2$). Let's be prepared and leave no stone unturned.



Implications of $m_H = 125 \text{ GeV}$

Potential is fully radiatively generated

Agashe et. al

$$V_{gauge}(h) = \frac{9}{2} \int \frac{d^4 p}{(2\pi)^4} \log \left(\Pi_0(p) + \frac{s_h^2}{4} \Pi_1(p) \right) \quad s_h \equiv \sin h/f$$

$$\Pi_0(p) = \frac{p^2}{g^2} + \Pi_a(p) \quad , \quad \Pi_1(p) = 2[\Pi_{\hat{a}}(p) - \Pi_a(p)]$$

$$\int d^4 p \Pi_1(p)/\Pi_0(p) < \infty$$

**Higgs dependent term
UV finite**

→ 'Weinberg sum rules'

$$\lim_{p^2 \rightarrow \infty} \Pi_1(p) = 0 \quad ,$$

$$\lim_{p^2 \rightarrow \infty} p^2 \Pi_1(p) = 0$$

UV finiteness requires at least two resonances

$$\Pi_1(p) = \frac{f^2 m_\rho^2 m_{a_1}^2}{(p^2 + m_\rho^2)(p^2 + m_{a_1}^2)} \quad \text{spin 1}$$

Similarly for SO(5) fermionic contribution

Pomarol et al; Marzocca

$$m_h^2 \simeq \frac{N_c}{\pi^2} \left[\frac{m_t^2}{f^2} \frac{m_{Q_4}^2 m_{Q_1}^2}{m_{Q_1}^2 - m_{Q_4}^2} \log \left(\frac{m_{Q_1}^2}{m_{Q_4}^2} \right) \right]$$

similar result in deconstruction:
Matsedonskyi et al; Redi et al

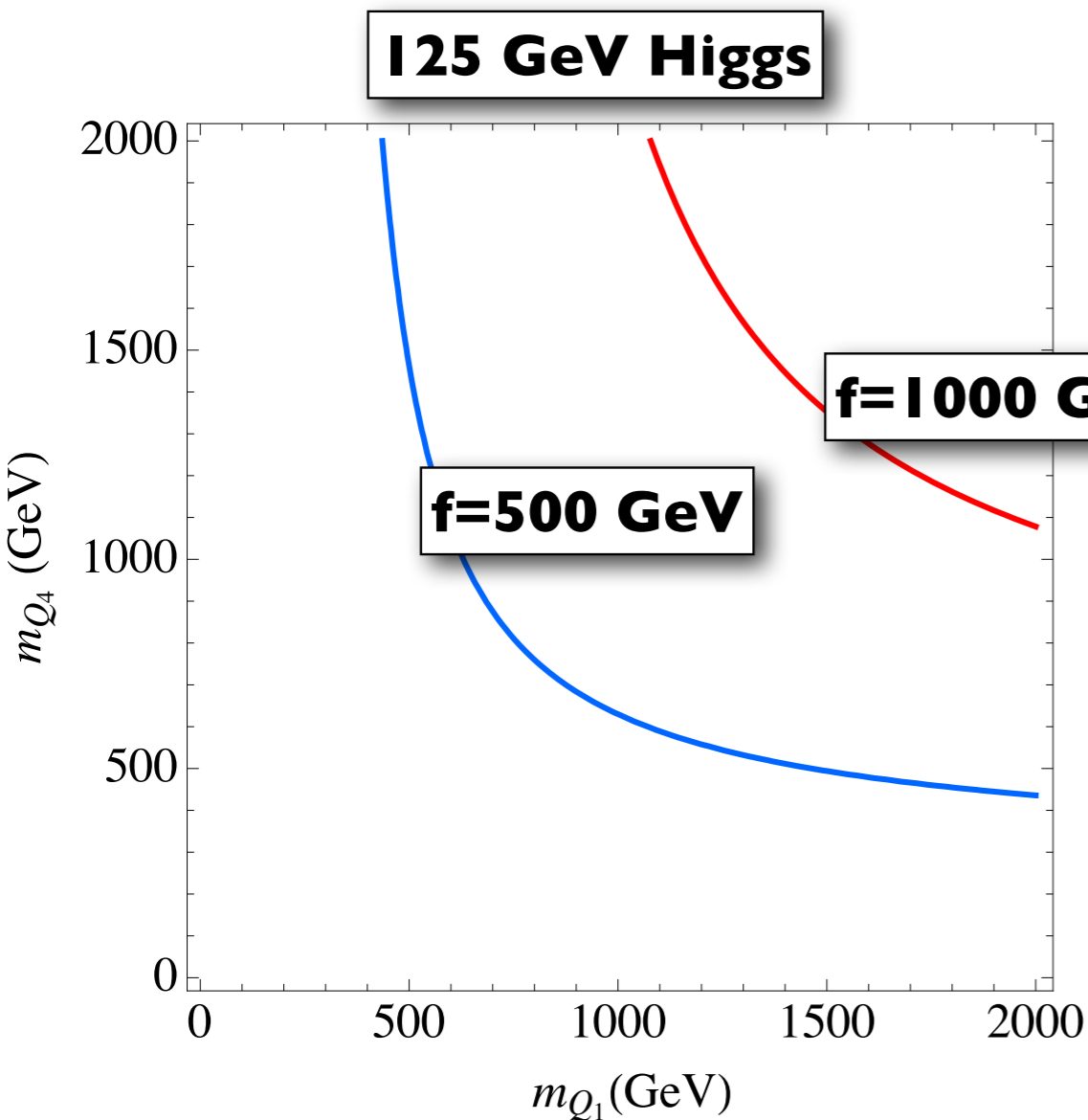
5 = 4 + 1 with EM charges 5/3, 2/3, -1/3
 Q_4 Q_1

→ solve for $m_h = 125$ GeV

Light Higgs implies light fermionic top partners

$$m_h^2 \simeq \frac{N_c}{\pi^2} \left[\frac{m_t^2}{f^2} \frac{m_{Q_4}^2 m_{Q_1}^2}{m_{Q_1}^2 - m_{Q_4}^2} \log \left(\frac{m_{Q_1}^2}{m_{Q_4}^2} \right) \right]$$

Pomarol et al; Marzocca



$$5 = 4 + 1$$

$Q_4 \quad Q_1$

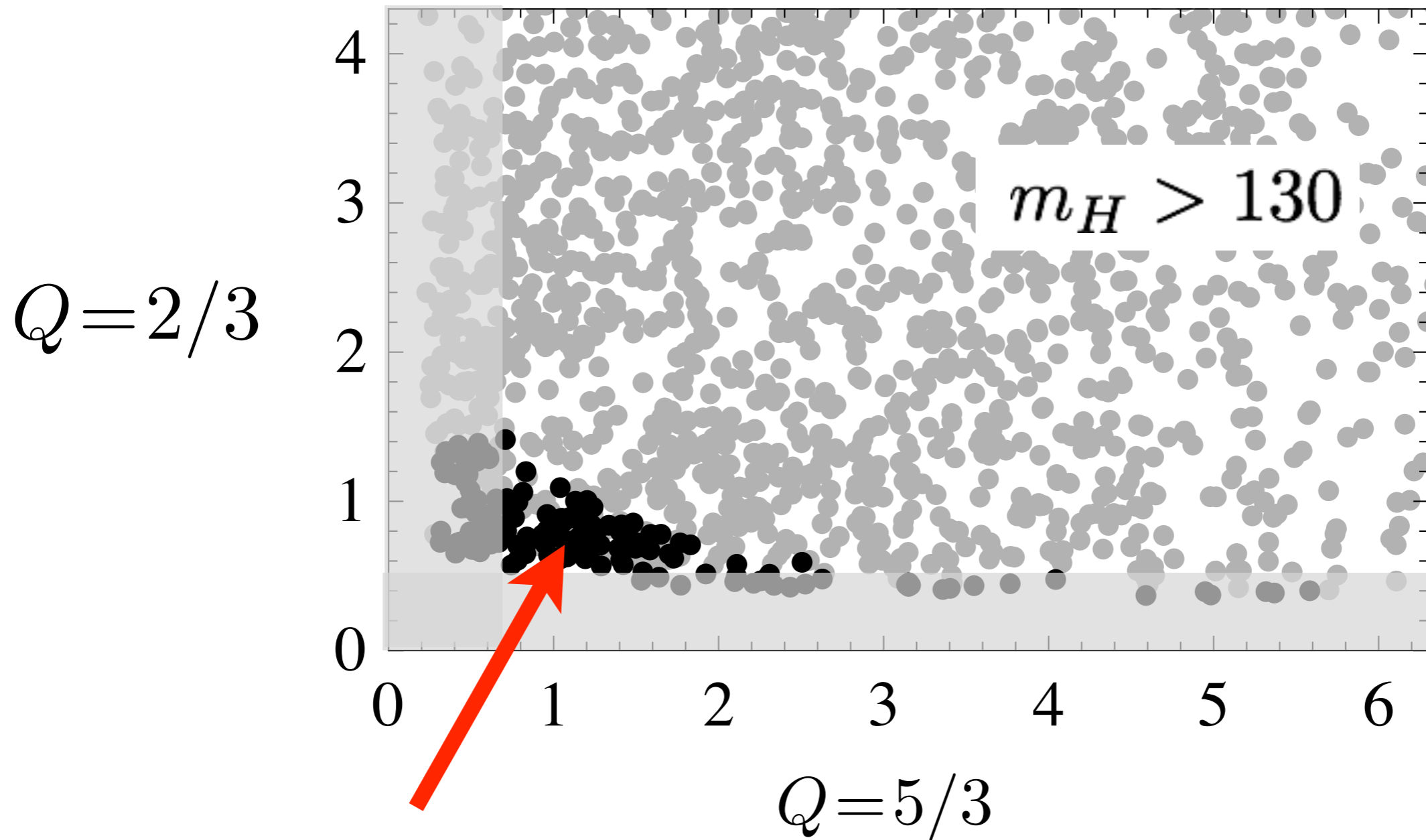
with EM charges $5/3, 2/3, -1/3$

Contino et al; Pomarol, Riva;
Matsedonskyi, Panico, Wulzer; Redi, Tesi;
Marzocca, Serone, Shu;

Scan over composite Higgs parameter space

$$\xi = 0.2$$

from 1204.6333



$$m_H = 115 \dots 130 \text{ GeV}$$