



JINR: State of Today and Perspectives

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(represented by A.Olshevskiy)

JINR, Dubna

Almost 50 years long way from the theoretical hypothesis on the nature and mechanism of generation of masses of the heavy vector fields of the Electroweak Interactions (W,Z bosons) to the direct experimental observation at the Large Hadron Collider in CERN marked in 2012 by the Nobel Prize



What is the next?

At the border of the XX-XXI centuries

Elementary Particle Physics, Nuclear Physics,
Astrophysics and Cosmology

are developing in a close interrelation as a
unique direction of the fundamental research on
the edge frontier of the modern nature science =

the Fundamental Physics

whose goal is search of the primary unique
universal laws of Nature which are governing
phenomena in both – the micro and macro
worlds and relate the deep structure of matter
with evolution of our Universe as a whole.

Main Puzzles of Fundamental Physics in the new Millennium

What gave first push to our Universe birth?

Physics of the Big Bang and evolution of the early Universe:

- *The Inflation hypothesis (exponentially fast expansion and generation of spectra of inhomogenities)
- *Evolution of the Hot Universe
 - *Nucleosynthesis
- *Acceleration of the Universe expanding

Very Broad Scientific International Cooperation in

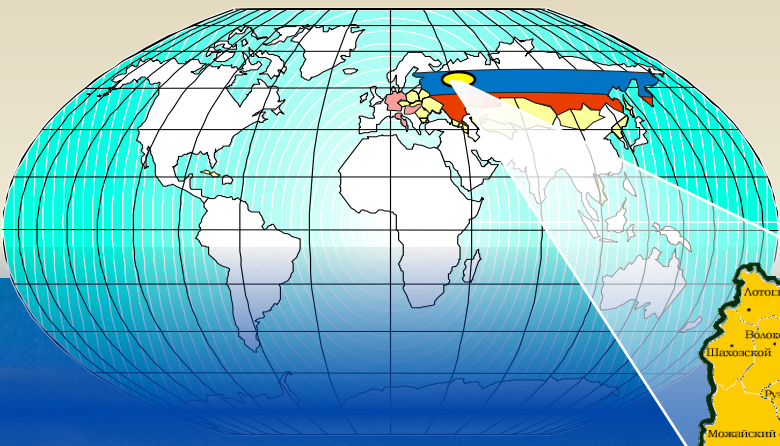
Energy & Intensity Frontiers

Europe, US & Latin America, Asia, South Africa

International labs: **CERN (est. 1954)** & **JINR (est. 1956)**

Russia

Where is JINR?



Moscow Region



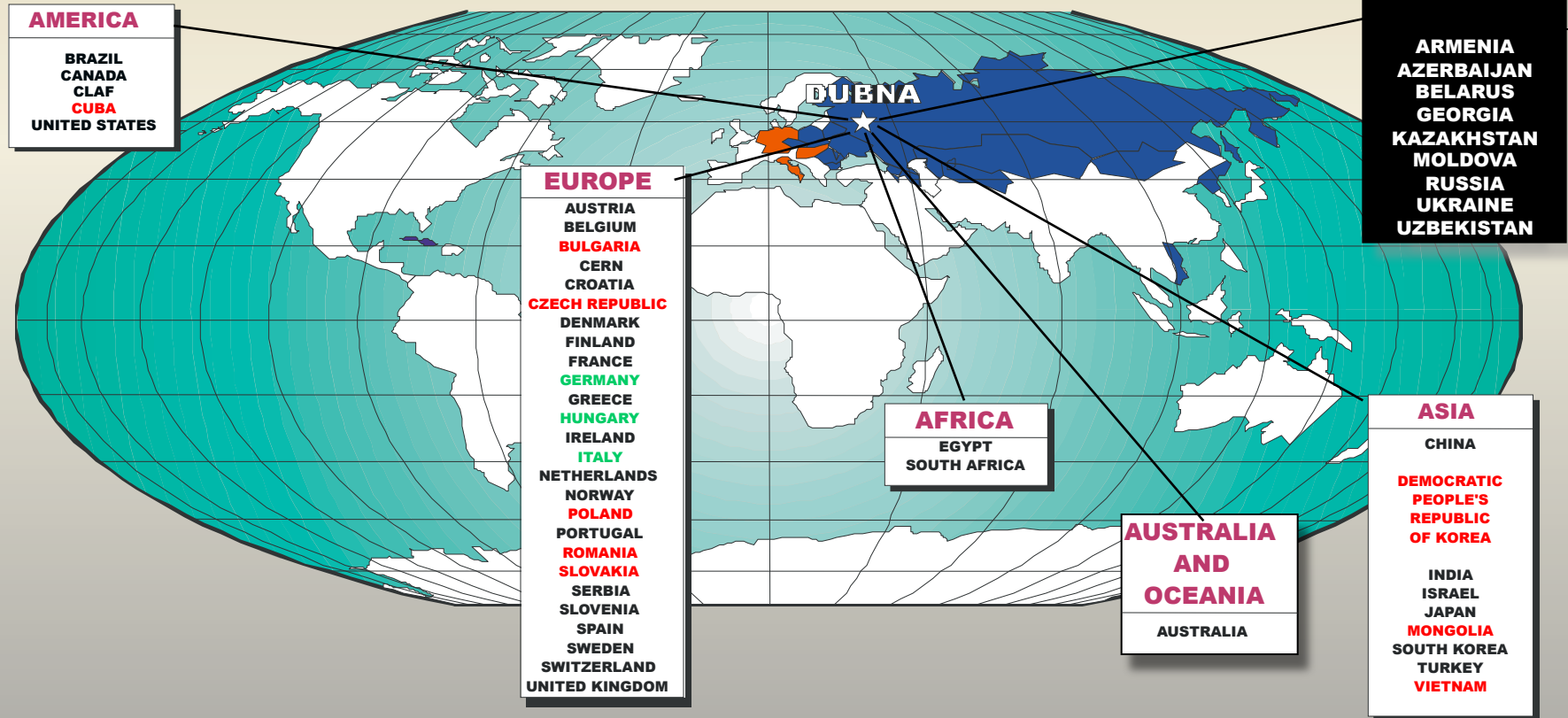
Dubna



JINR



JINR Cooperation



JINR's partners are about 700 institutions located in 60 countries, including about 300 institutions and universities from the JINR Member States

JINR's Basic Scientific Directions

- High Energy Physics & Relativistic Heavy Ion Physics
 - Nuclear Physics
 - Condensed Matter Physics
 - Radiation Biology & Astrobiology
-
- ***Main Supporting Activities***
 - Theory of Pphys., Nphys., CMphys.
 - Networking and computing
 - Physics instruments and methods
 - Training of young specialists



JINR Research Facilities

- Complex Nuclotron-NICA (Mega-Science Project):
High density & strong interacting matter search
- Pulsed reactor IBR-2: Condensed matter & Nuclear physics
-
- Accelerators Complex: New heavy elements synthesis
-
- Complex of computing & information technologies
- Accelerators complex: Applied and & medical sciences,
proton therapy

High Energy Physics @ JINR

From Synchrophasotron to Nuclotron to NICA

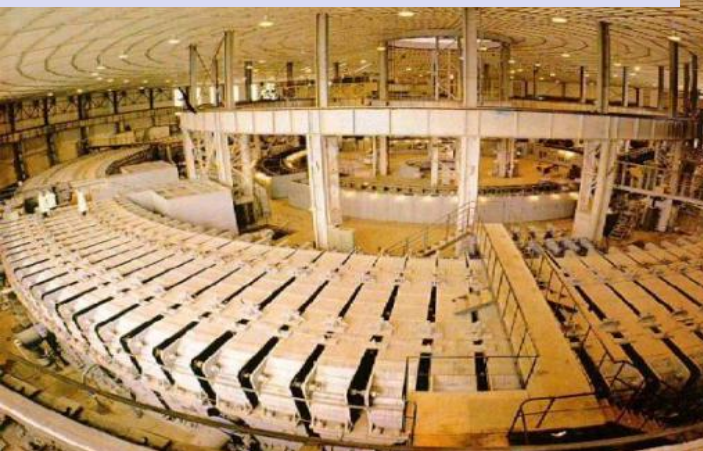
1957 - Synchrophasotron

10 GeV proton accelerator – world leader in energy.

Beginning of era of high-energy physics



V.Veksler – phase stability principle discovery



1993 –
Nuclotron

*First in the world
Superconducting
Synchrotron
of heavy ions*



A.Baldin – start of relativistic nuclear physics era



CERN-JINR School 2014

2019 – NICA

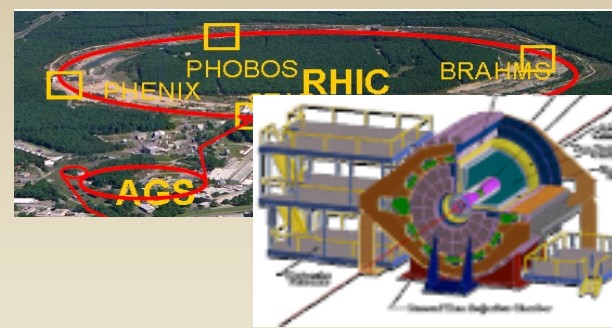
*Superconducting collider
of heavy ions*



Study of baryonic matter at extreme conditions

The JINR 7-year plan for 2010-2016 approved by CPP in 2009:

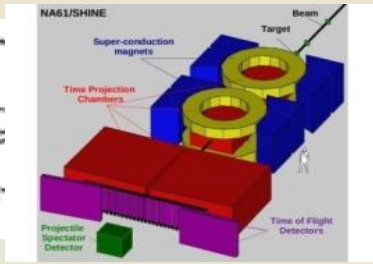
NICA – the JINR flagship project in HEP



2nd generation Heavy Ion experiments

STAR/PHENIX @ BNL/RHIC.

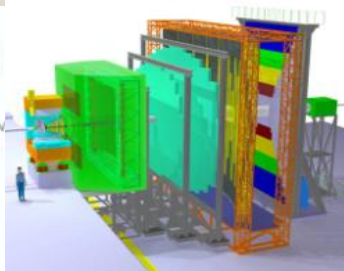
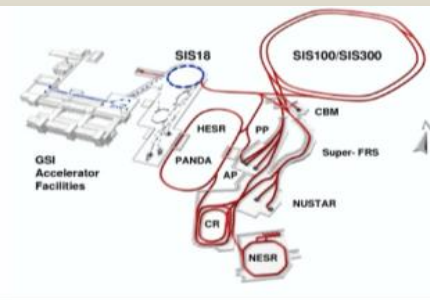
designed for high energy researches ($\sqrt{s_{NN}} > 20$ GeV),
low luminosity for LES program $L < 10^{26}$ cm⁻²s⁻¹ for Au⁷⁹⁺



NA61 @ CERN/SPS.

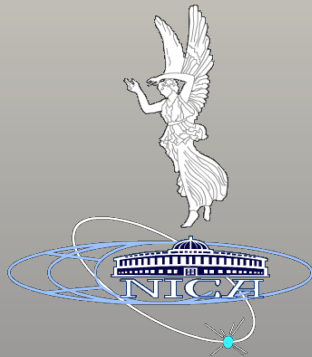
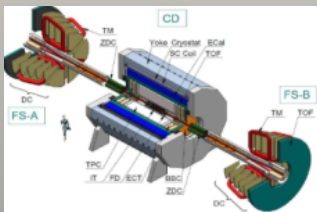
Fixed target, non-uniform acceptance, few energies
(10,20,30,40,80,160A GeV), poor nomenclature of
beam species

3rd generation Heavy Ion experiments



CBM @ FAIR/SIS-100/300

Fixed target, $E/A = 10-40$ GeV, high luminosity

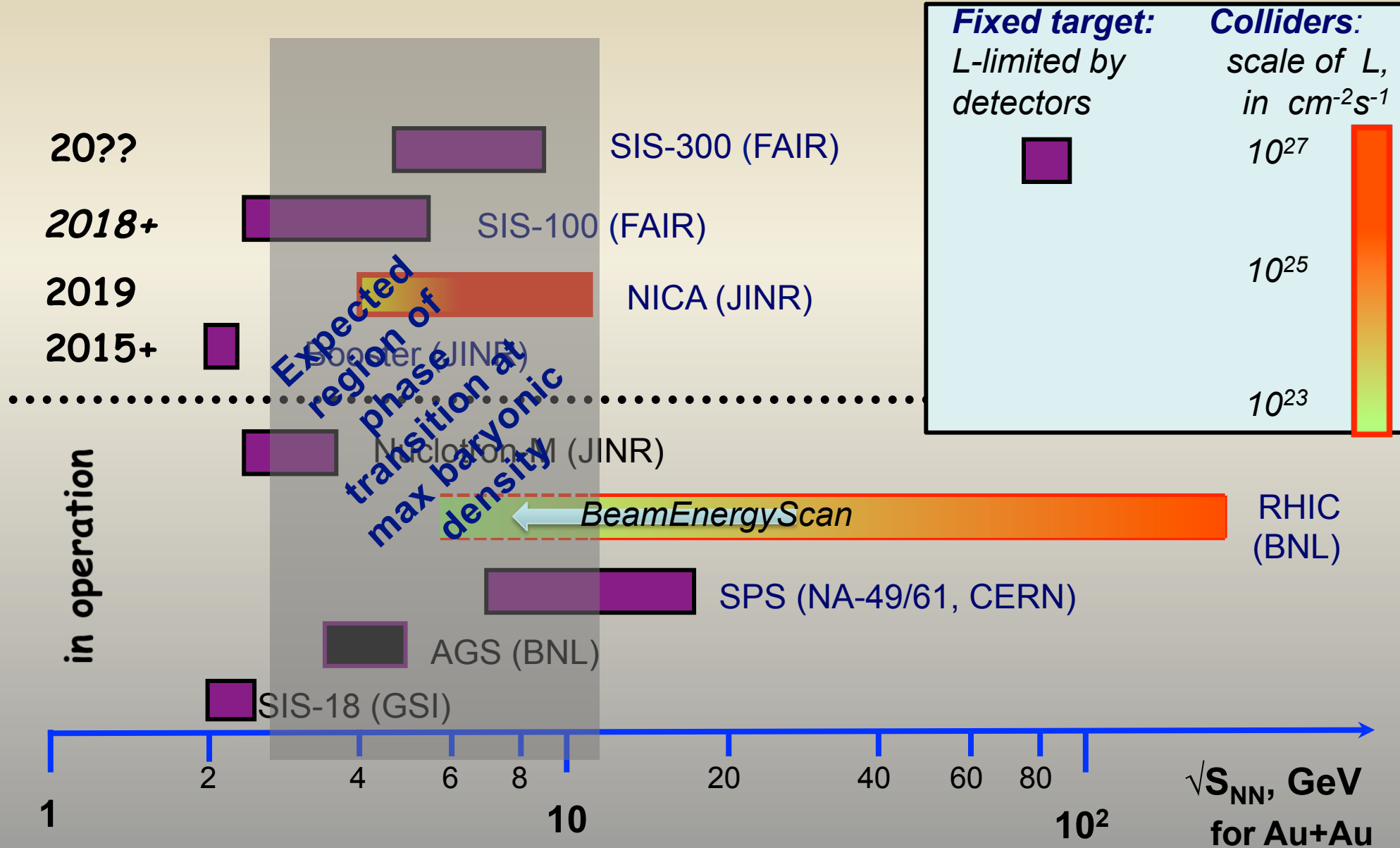


MPD & SPD @ JINR/NICA

Collider, small enough
 $\sqrt{s_{NN}} = 4-11$ GeV, a variety of colliding systems,
 $L \sim 10^{27}$ cm⁻²s⁻¹ for Au⁷⁹⁺

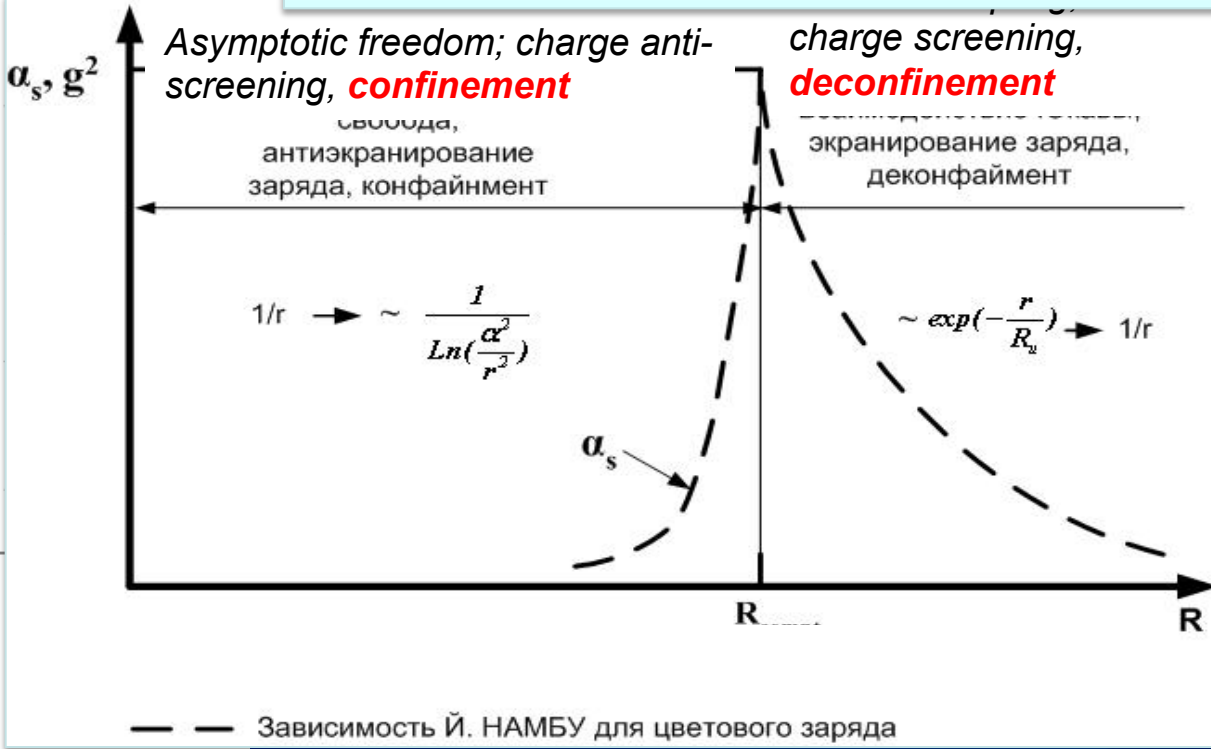
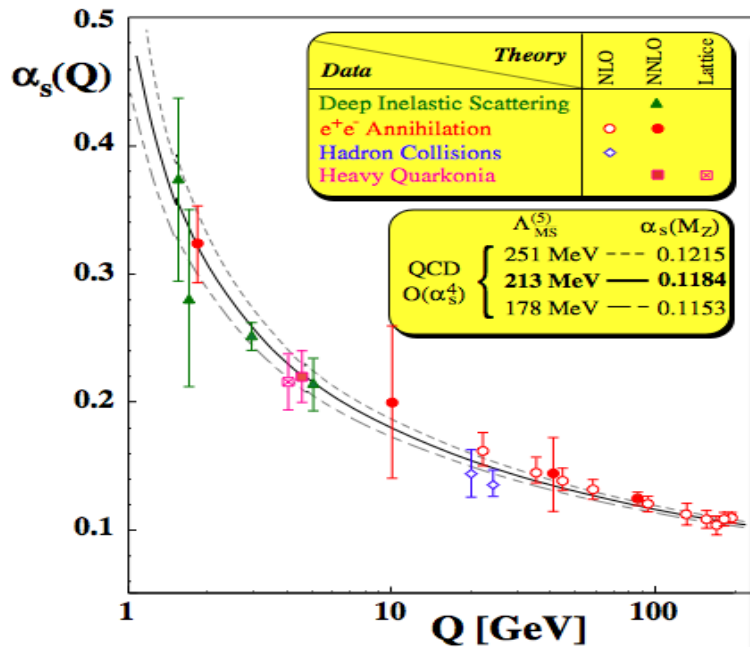
Common European
Research
infrastructure for
Heavy Ion High
Energy Physics: NICA
+ FAIR

Present and future HI machines



BASIS: Asymptotic freedom of quarks

N. Bogolyubov, D. Shirkov
 D.J. Gross, H.D. Politzer, F. Wilczek
 A. Tavkhelidze, B. Struminsky, V. Matveev

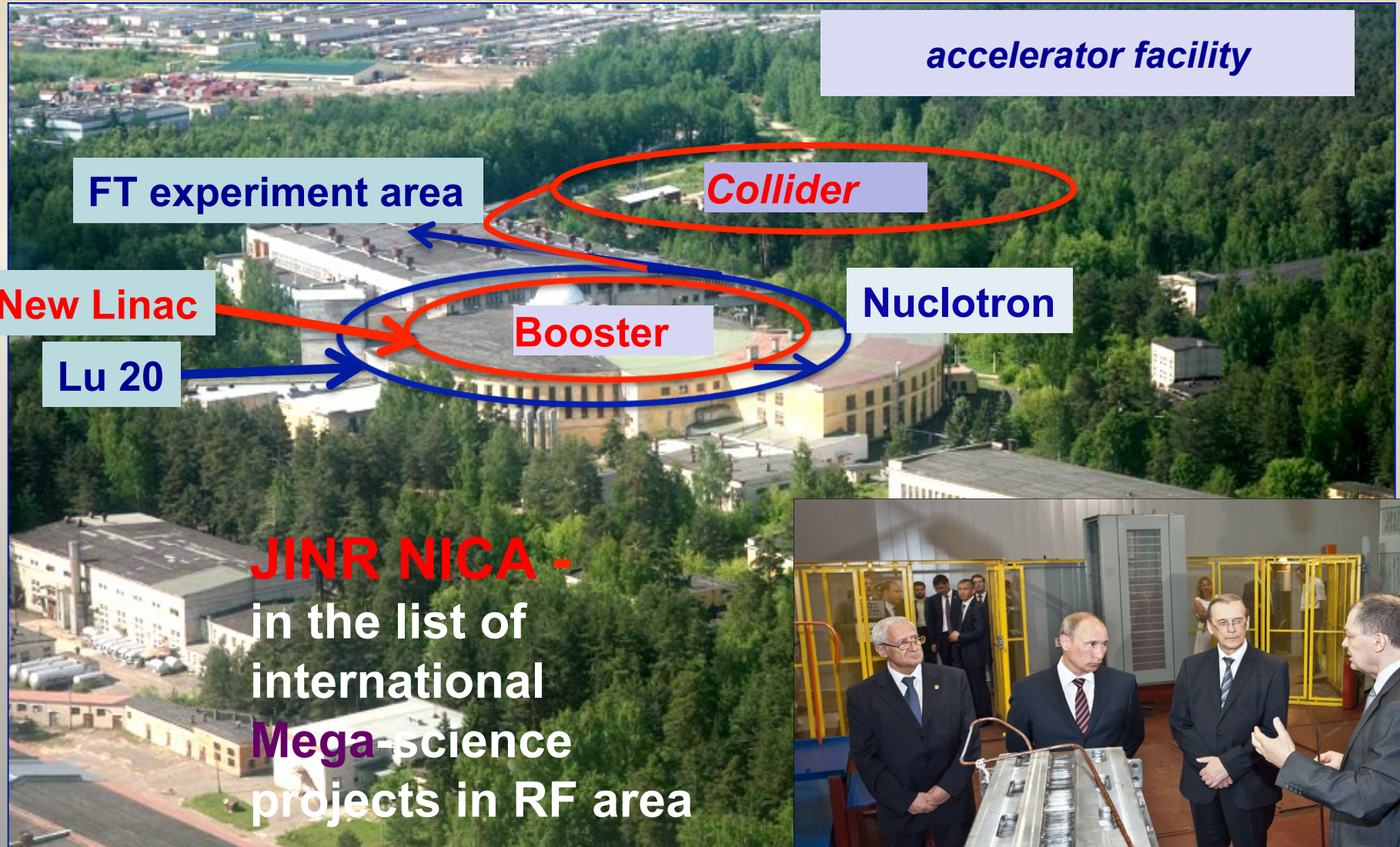


the access to “asymptotically free” regime in hard processes

and in **super-dense matter** (inter-particle distances $\sim 1/T$)

super-dense matter could be obtained in **heavy ion collisions**

Complex NICA @ JINR



Superconducting accelerator complex **NICA**

(**N**uclotron based **I**on **C**ollider **f**Acility)

KRION-6T+HILac (3MeV/u),
SPP and LU-20 (5MeV/u)

Fixed target
experiments area
(BM@N, b.205)

Booster (600
MeV/u)

Nuclotron
0,6-4,5 GeV/u

(SPD)

NICA Collider
(1-4,5 GeV/u, C=503 m)

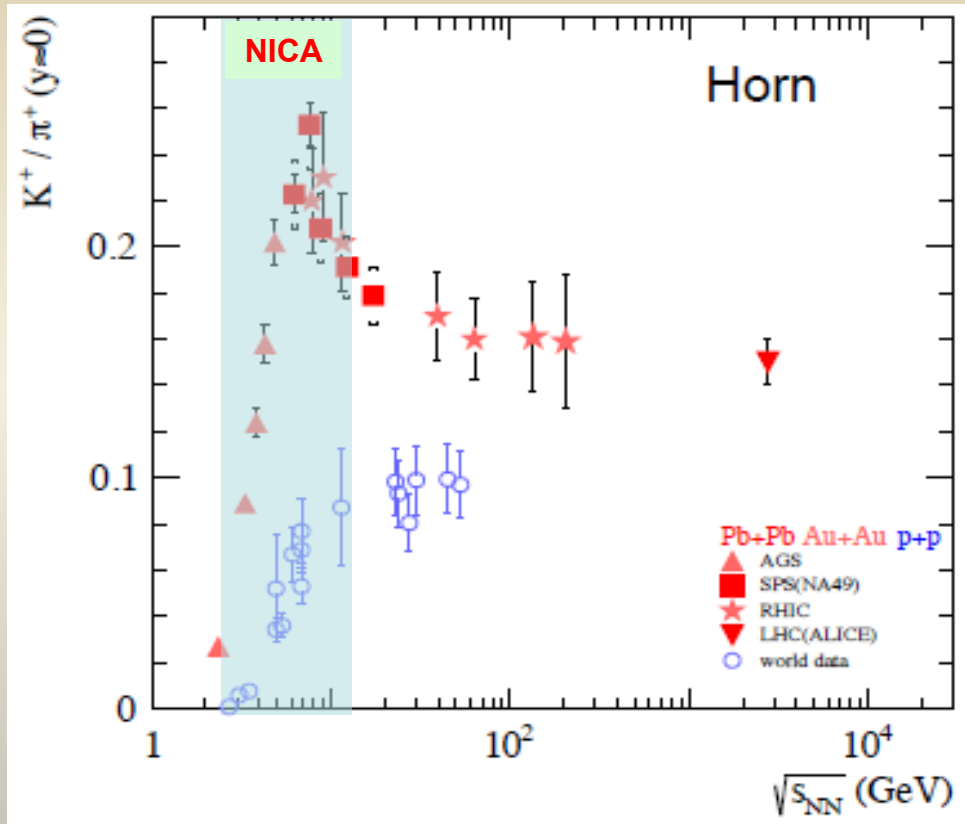
MultiPurpose
detector (MPD)

Booster synchrotron



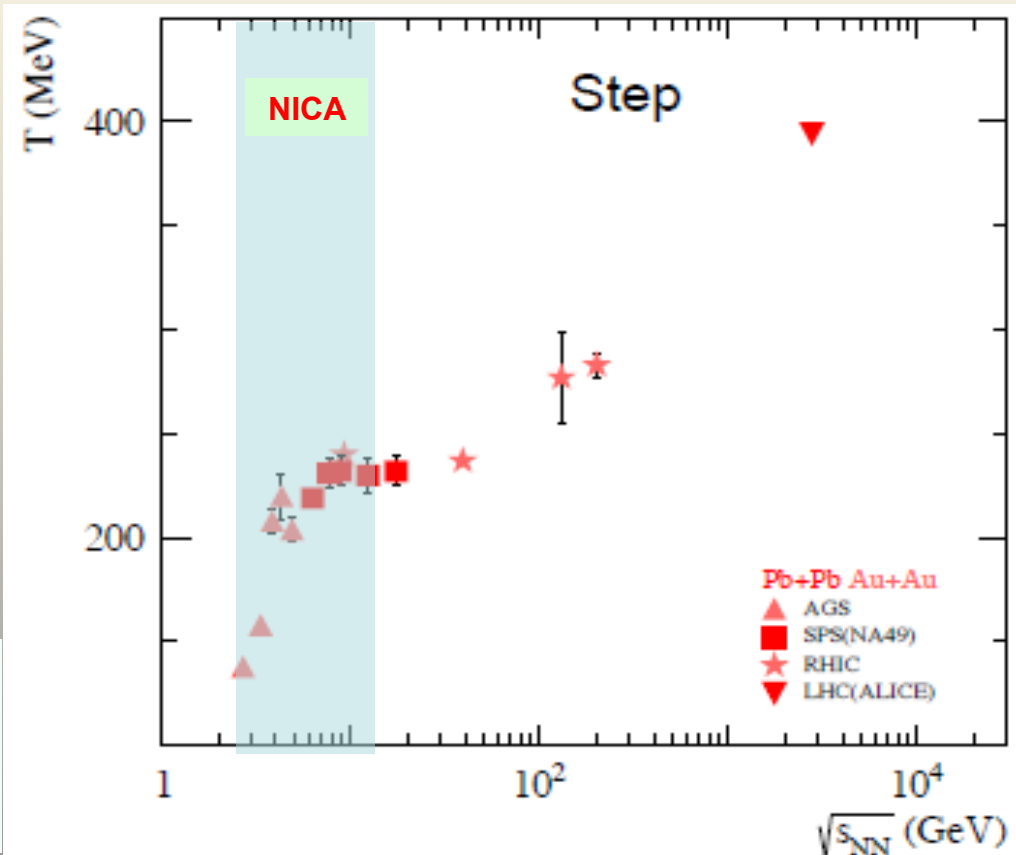
29.06.14

Towards the “deconfinement” & mixed phase



- Non-monotonic energy dependence of the K^+/π^+ ratio (“Horn”) – onset of deconfinement?

Plateau in the apparent temperature of the kaon spectra (“Step”) – signal of the mixed phase?



NICA Mega-Science Project International Consortium



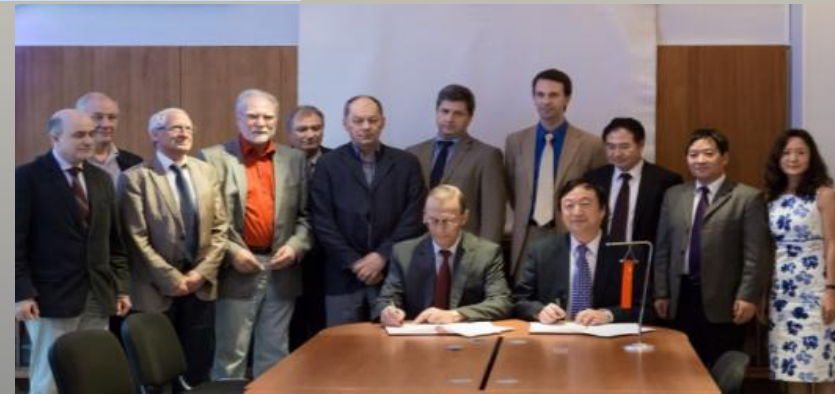
6 countries

- Protocol signed by:
- Belarus, *Bulgaria*, Germany, Kazakhstan
- Russia, Ukraine

Dubna, August 08, 2013.

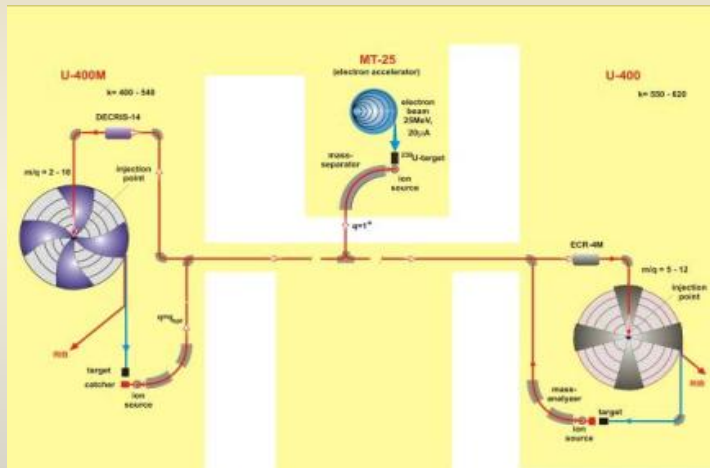


Scientific cooperation @ NICA projects

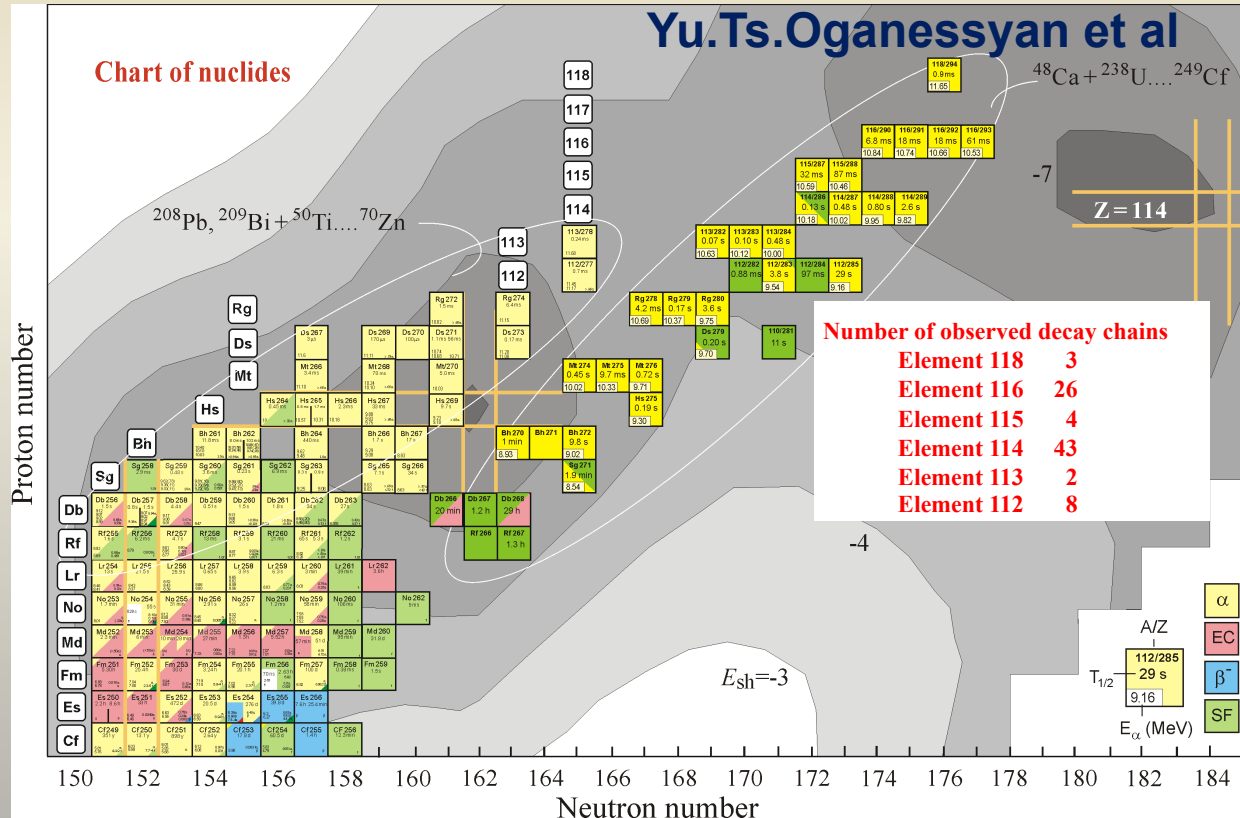


Super Heavy Elements (DRIBs-III project)

For the last decade JINR has become one of the leading scientific centres in the world in low energy heavy-ion physics.



**DRIBs (I,II,III) –
Dubna Radioactive
Ion Beams**

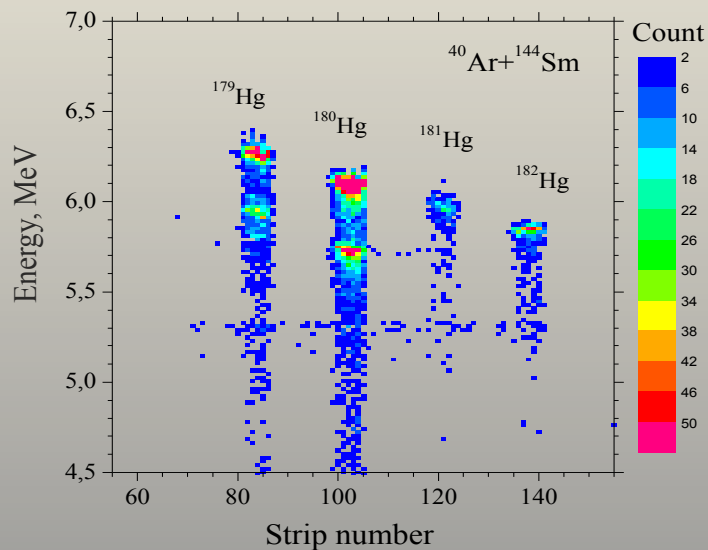


U400 and U400M isochronous cyclotrons are combined into the accelerator complex – project DRIBs which deals with the production of beams of exotic light neutron-deficient and neutron-rich nuclei in reactions with light ions.

Mass-Separator of Heavy Atoms - "MASHA"

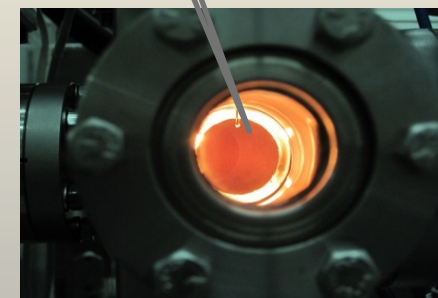
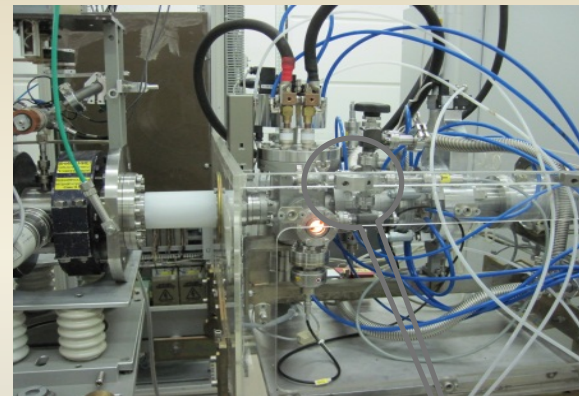


Mass-separator "MASHA"
at the beam of U-400M



Detection of mercury isotopes at
the focal plane

Hot catcher

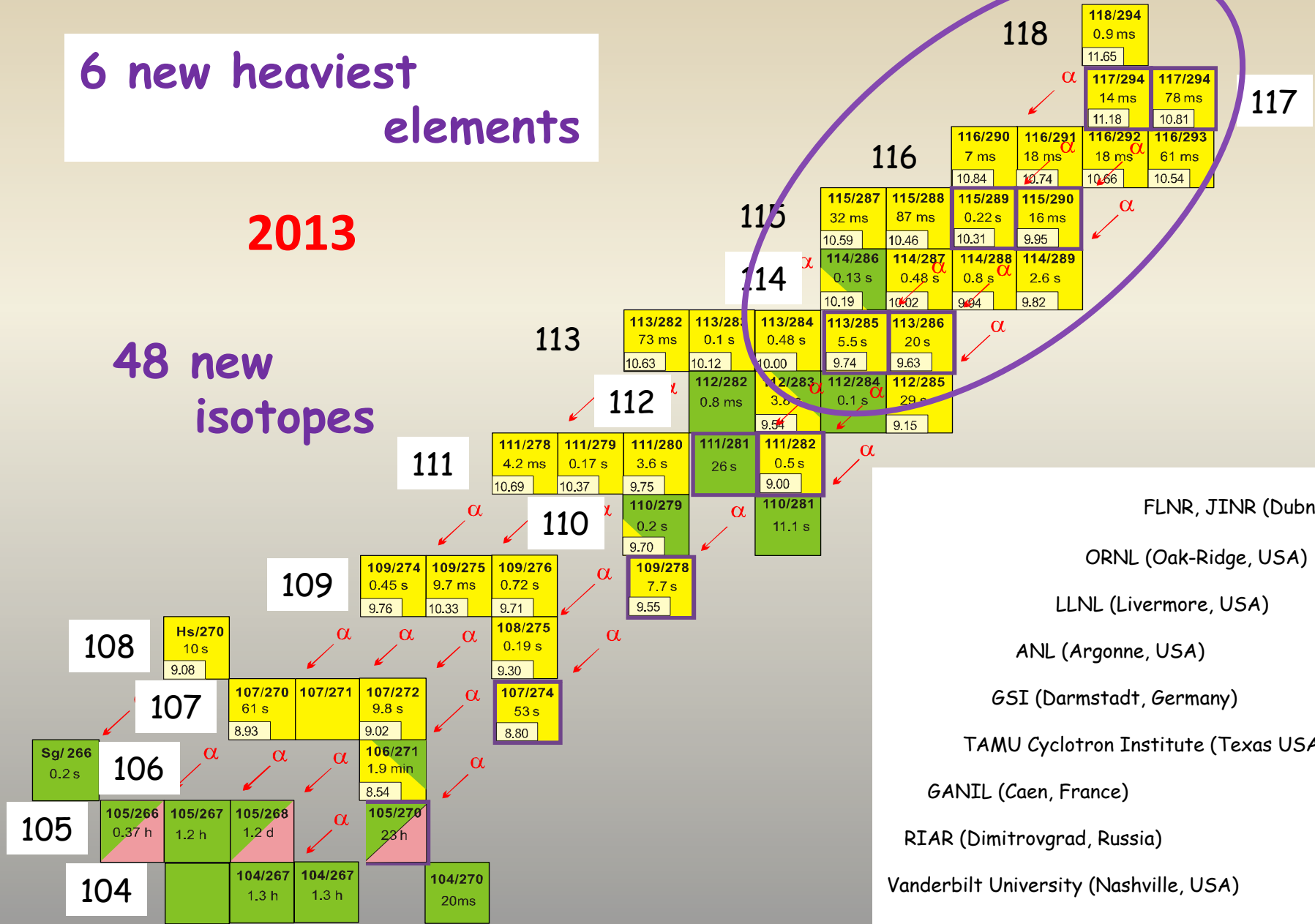


Mass measurement of ^{283}Cn
in reaction $^{48}\text{Ca} + ^{238}\text{U}$ started

6 new heaviest elements

2013

48 new isotopes

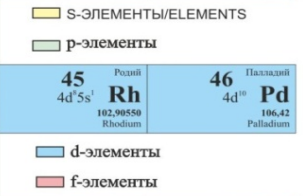
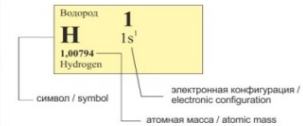


- FLNR, JINR (Dubna)
- ORNL (Oak-Ridge, USA)
- LLNL (Livermore, USA)
- ANL (Argonne, USA)
- GSI (Darmstadt, Germany)
- TAMU Cyclotron Institute (Texas USA)
- GANIL (Caen, France)
- RIAR (Dimitrovgrad, Russia)
- Vanderbilt University (Nashville, USA)

Two new elements in the Table

Периодическая таблица элементов Д.И. Менделеева
D.I. Mendeleev's Periodic Table of Elements

период	группа	группы элементов																							
		a	I	б	a	II	б	a	III	б	a	IV	б	a	V	б	a	VI	б	a	VII	б	a	VIII	б
1	I	Водород H 1,00794 Hydrogen	1 1s ¹																					Гелий He 4,0026 Helium	2 1s ²
2	II	Литий Li 6,941 Lithium	3 2s ¹	Бериллий Be 9,012182 Beryllium	4 2s ²	Бор B 10,811 Boron	5 2p ¹	Углерод C 12,011 Carbon	6 2p ²	Азот N 14,00674 Nitrogen	7 2p ³	Кислород O 15,9994 Oxygen	8 2p ⁴	Фтор F 18,9984032 Fluorine	9 2p ⁵	Неон Ne 20,1797 Neon	10 2p ⁶								
3	III	Натрий Na 22,989768 Sodium	11 3s ¹	Магний Mg 24,3050 Magnesium	12 3s ²	Алюминий Al 26,981539 Aluminum	13 3p ¹	Кремний Si 28,0855 Silicon	14 3p ²	Фосфор P 30,973762 Phosphorus	15 3p ³	Сера S 32,066 Sulfur	16 3p ⁴	Хлор Cl 35,4527 Chlorine	17 3p ⁵	Аргон Ar 39,948 Argon	18 3p ⁶								
4	IV	Калий K 39,0983 Potassium	19 4s ¹	Кальций Ca 40,078 Calcium	20 4s ²	21 3d ¹ 4s ²	Скандий Sc 44,955910 Scandium	22 3d ² 4s ²	Титан Ti 47,88 Titanium	23 3d ³ 4s ²	Ванадий V 50,9415 Vanadium	24 3d ⁴ 4s ¹	Хром Cr 51,9961 Chromium	25 3d ⁵ 4s ²	Марганец Mn 54,93805 Manganese	26 3d ⁶ 4s ²	Железо Fe 55,847 Iron	27 3d ⁷ 4s ²	Кобальт Co 58,93320 Cobalt	28 3d ⁸ 4s ²	Никель Ni 58,6934 Nickel				
	V	29 3d ¹⁰ 4s ¹	Медь Cu 63,546 Copper	30 3d ¹⁰ 4s ²	Цинк Zn 65,39 Zinc	31 4p ¹	Галлий Ga 69,723 Gallium	32 4p ²	Германий Ge 72,61 Germanium	33 4p ³	Мышьяк As 74,92159 Arsenic	34 4p ⁴	Селен Se 78,96 Selenium	35 4p ⁵	Бром Br 79,904 Bromine	36 4p ⁶	Криптон Kr 83,80 Krypton								
5	VI	Рубидий Rb 85,4678 Rubidium	37 5s ¹	Стронций Sr 87,62 Strontium	38 5s ²	39 4d ¹ 5s ²	Иттрий Y 88,90585 Yttrium	40 4d ² 5s ²	Цирконий Zr 91,224 Zirconium	41 4d ⁵ 5s ¹	Нобий Nb 92,90638 Niobium	42 4d ⁵ 5s ¹	Молибден Mo 95,94 Molybdenum	43 4d ⁵ 5s ²	Технеций Tc [98] Technetium	44 4d ⁵ 5s ¹	Рутений Ru 101,07 Ruthenium	45 4d ⁵ 5s ¹	Рений Rh 102,90550 Rhodium	46 4d ¹⁰	Палладий Pd 106,42 Palladium				
	VII	47 4d ¹⁰ 5s ¹	Серебро Ag 107,8682 Silver	48 4d ¹⁰ 5s ²	Кадмий Cd 112,411 Cadmium	49 5p ¹	Индий In 114,818 Indium	50 5p ²	Олово Sn 118,710 Tin	51 5p ³	Сурьма Sb 121,757 Antimony	52 5p ⁴	Теллур Te 127,60 Tellurium	53 5p ⁵	Иод I 126,90447 Iodine	54 5p ⁶	Ксенон Xe 131,29 Xenon								
6	VIII	Цезий Cs 132,90543 Cesium	55 6s ¹	Барий Ba 137,327 Barium	56 6s ²	57 5d ¹ 6s ²	Лантан La 138,9055 Lanthanum	72 5d ⁶ 6s ²	Гафний Hf 178,49 Hafnium	73 5d ⁶ 6s ²	Тантал Ta 180,9479 Tantalum	74 5d ⁶ 6s ²	Вольфрам W 183,84 Tungsten	75 5d ⁶ 6s ²	Рений Re 186,207 Rhenium	76 5d ⁶ 6s ²	Осмий Os 190,23 Osmium	77 5d ⁶ 6s ²	Иридий Ir 192,22 Iridium	78 5d ⁶ 6s ¹	Платина Pt 195,08 Platinum				
	IX	79 5d ¹⁰ 6s ¹	Золото Au 196,96654 Gold	80 5d ¹⁰ 6s ²	Ртуть Hg 200,59 Mercury	81 6p ¹	Таллий Tl 204,3833 Thallium	82 6p ²	Свинец Pb 207,2 Lead	83 6p ³	Висмут Bi 208,98037 Bismuth	84 6p ⁴	Полоний Po [209] Polonium	85 6p ⁵	Астат At [210] Astatine	86 6p ⁶	Радон Rn [222] Radon								
7	X	Франций Fr [223] Francium	87 7s ¹	Радий Ra 226,025 Radium	88 7s ²	89 6d ⁷ 7s ²	Актиний Ac [227] Actinium	104 [261] Rutherfordium	Резерфордий Rf [261] Rutherfordium	105 [262] Dubnium	Дубний Db [262] Dubnium	106 [266] Seaborgium	Сиббериум Sg [266] Seaborgium	107 [267] Bohrium	Борий Bh [267] Bohrium	108 [269] Hassium	109 [285] Meitnerium	Мейтнерий Mt [285] Meitnerium	110 [289] Darmstadtium	Дармштадтий Ds [289] Darmstadtium					
	XI	111 [272] Roentgenium	Рентгений Rg [272] Roentgenium	112 [285] Copernicium	Коперниций Cn [285] Copernicium	113	Флеровий Fl [289] Flerovium	114	115	Ливерморий Lv [293] Livermorium	116	117	118												



Лантаноиды Lanthanides

Церий Ce 140,115 Cerium	Прасеодим Pr 140,90765 Praseodymium	Неодим Nd 144,24 Neodymium	Прометий Pm [145] Promethium	Самарий Sm 150,36 Samarium	Европий Eu 151,965 Europium	Гадолиний Gd 157,25 Gadolinium	Тербий Tb 158,92534 Terbium	Диспрозий Dy 162,50 Dysprosium	Гольмий Ho 164,93032 Holmium	Эрбий Er 167,26 Erbium	Тулий Tm 168,93421 Thulium	Иттербий Yb 173,04 Ytterbium	Лютеций Lu 174,967 Lutetium
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Актиноиды Actinides

Торий Th 232,0381 Thorium	Протактиний Pa 231,03688 Protactinium	Уран U 238,0289 Uranium	Нептуний Np 237,04817 Neptunium	Плутоний Pu 244,06422 Plutonium	Америций Am 243,06136 Americium	Кюрий Cm 247,07724 Curium	Берклий Bk 247,07125 Berkelium	Калифорний Cf 251,08320 Californium	Эйнштейний Es 252,08329 Einsteinium	Фермий Fm 257,10528 Fermium	Менделевий Md 288,106 Mendelevium	Нобелий No 289,10375 Nobelium	Лоуренсий Lr 262,10932 Lawrencium
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What is beyond 118 element?

Heaviest target: $^{249}\text{Cf} \rightarrow Z_{\text{max}} = 118$

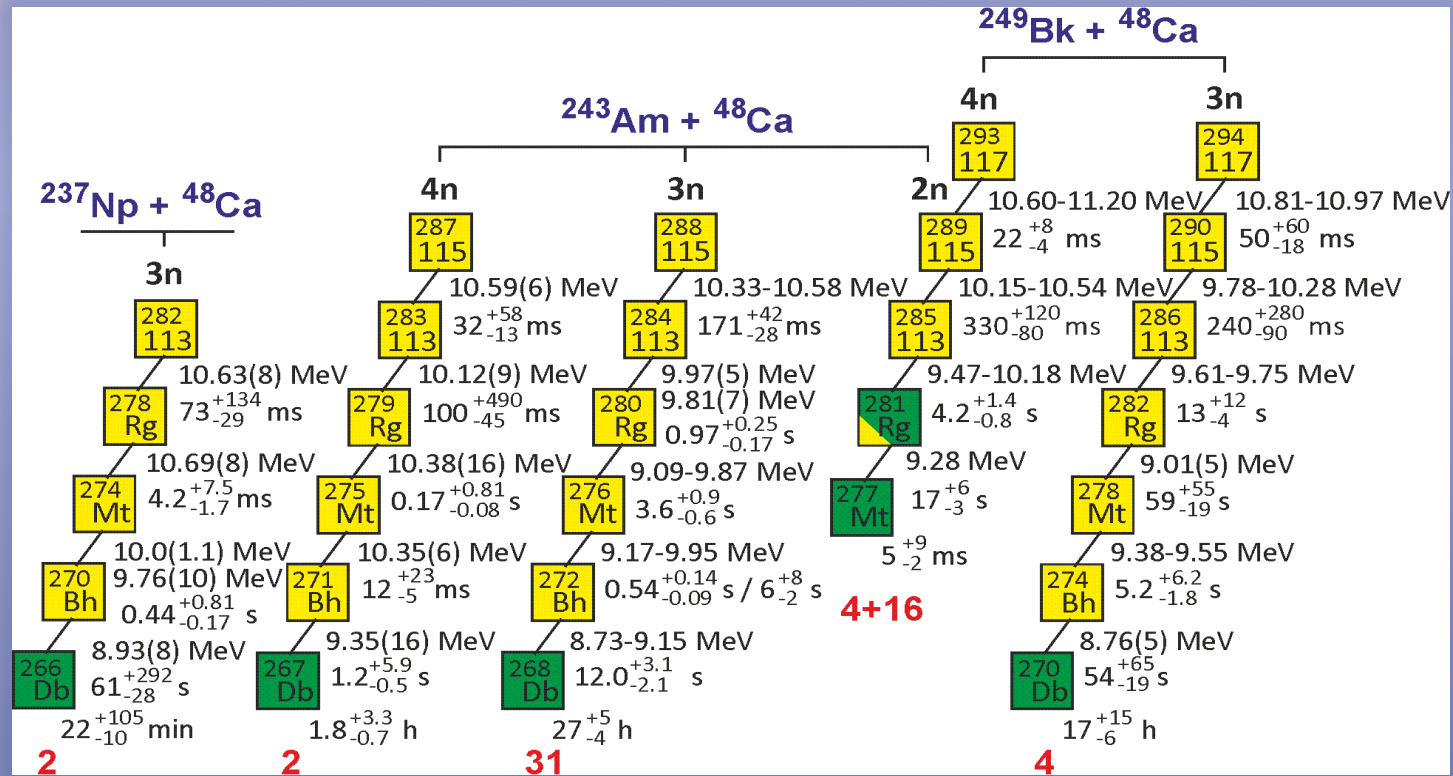


- Heavier projectiles (^{50}Ti , ^{54}Cr , ^{58}Fe , ^{64}Ni)
- Heavier targets (^{251}Cf , ^{254}Es -???)
- Symmetric reactions: $^{136}\text{Xe} + ^{136}\text{Xe}$, $^{136}\text{Xe} + ^{150}\text{Nd}$, $^{150}\text{Nd} + ^{150}\text{Nd}$;
- Nucleon transfer reactions ($^{136}\text{Xe} + ^{208}\text{Pb}$, $^{238}\text{U} + ^{248}\text{Cm}$).

Preparation to study $^{48}\text{Ca} + ^{251}\text{Cf}$ reaction
(Spring 2014)

ODD SUPERHEAVIES: 113, 115, 117

Application for recognition of these discoveries is sent to IUPAC in November 2013



Russian Information Agency has recently published the result of widely distributed questionnaire about most significant discoveries by Russian researchers over last two decades.

The list of those discoveries is opening with **SHE synthesis at JINR.**

Super-Heavy Elements Factory – «SHEF»

Production

Today: $\sim 5 \cdot 10^{19}$

With SHEF: $1.5 \cdot 10^{21}$

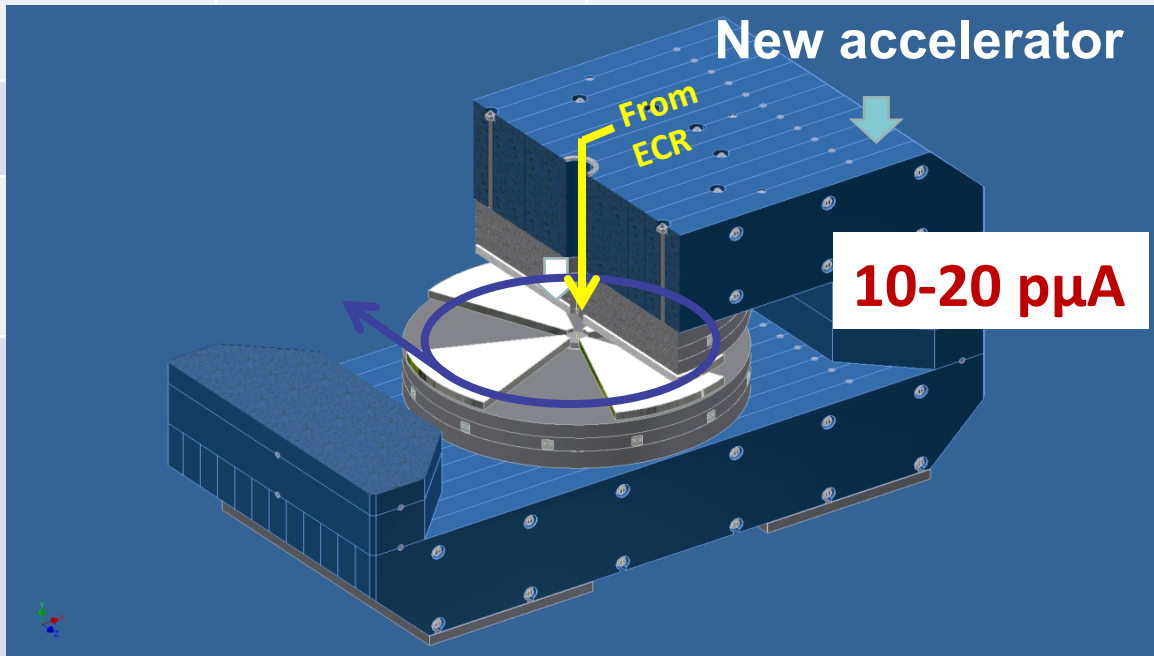
factor: 30

Increased

Intensity beam

&

Beam time

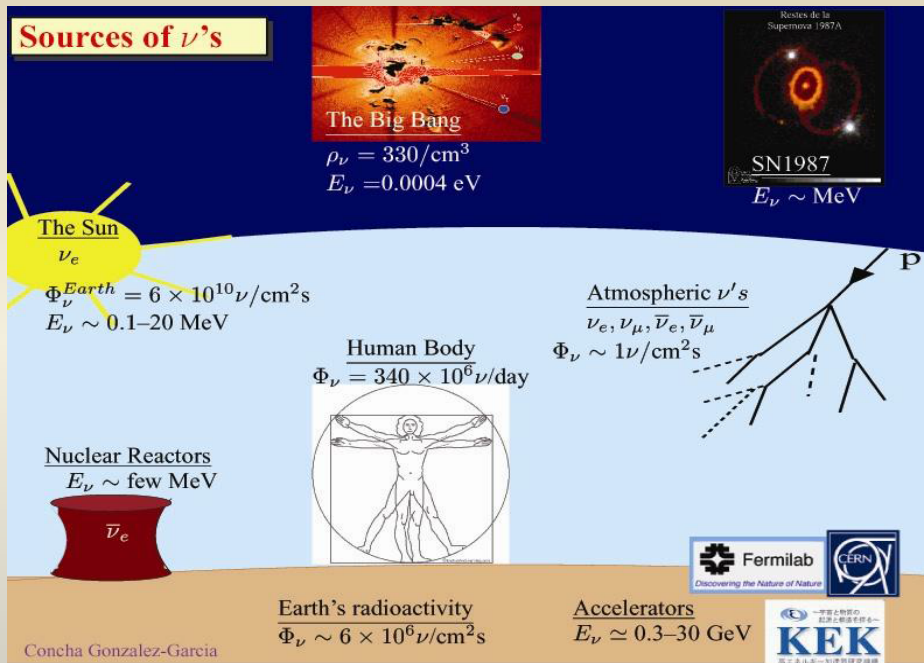


SHEF



~ 7000 h/year

Neutrino Physics



F. Reines, *Ann. Rev. Nucl. Sci.* **10** (1960) 1–26.

- ✓ fission **reactors** are the best source of low energy antineutrinos
- ✓ muon-neutrinos can be produced by allowing **pions to decay in flight**
- ✓ **the Sun** is a “most copious source of neutrinos”
- ✓ cosmic rays striking the Earth’s **atmosphere** should produce a significant flux of neutrinos from pion decay
- ✓ high-energy neutrinos produced by **astrophysical objects** would provide information not available from cosmic rays

B. Pontecorvo:

- ✓ proposed the first (radiochemical) **method of neutrino detection** and possibility of **reactor experiments**.
- ✓ was the first who came to an idea of **$\mu - e$ universality** of the weak interaction.
- ✓ proposed the experiment with accelerator neutrinos to prove that **V_μ and V_e are different particles**.
- ✓ was the first who came to idea of **neutrino oscillations** and proposed many experiments to check it.



New frontier

A lot more to learn

- Majorana or Dirac masses? $\implies 2\beta 0\nu$ -decays
- Absolute mass scale \implies Katrin + cosmology + $2\beta 0\nu$
- Direct or inverse hierarchy \implies Long baseline accelerator and reactor neutrino experiments
- CP-violation \implies Long baseline accelerator experiments
- Electromagnetic properties \implies Search for neutrino magnetic moment
- Sterile neutrinos?

JINR is an important player in nearly all of this

NEMO \implies SuperNEMO, GERDA \implies GERDA/Majorana
Daya Bay \implies JUNO, NO ν A
GEMMA
DANSS

60 years of reactor neutrino physics

2011/2012 -
The year of θ_{13}

Daya Bay
Double Chooz
Reno



2008 - Precision measurement of Δm_{12}^2 . Evidence for oscillation

2003 - First observation of reactor antineutrino disappearance

1995 - Nobel Prize to Fred Reines at UC Irvine

1980s & 1990s - Reactor neutrino flux measurements in U.S. and Europe

1956 - First observation of (anti)neutrinos



KamLAND

Past Reactor Experiments

Hanford
Savannah River
ILL, France
Bugey, France
Rovno, Russia
Goesgen, Switzerland
Krasnoyarsk, Russia
Palo Verde
Chooz, France



Chooz



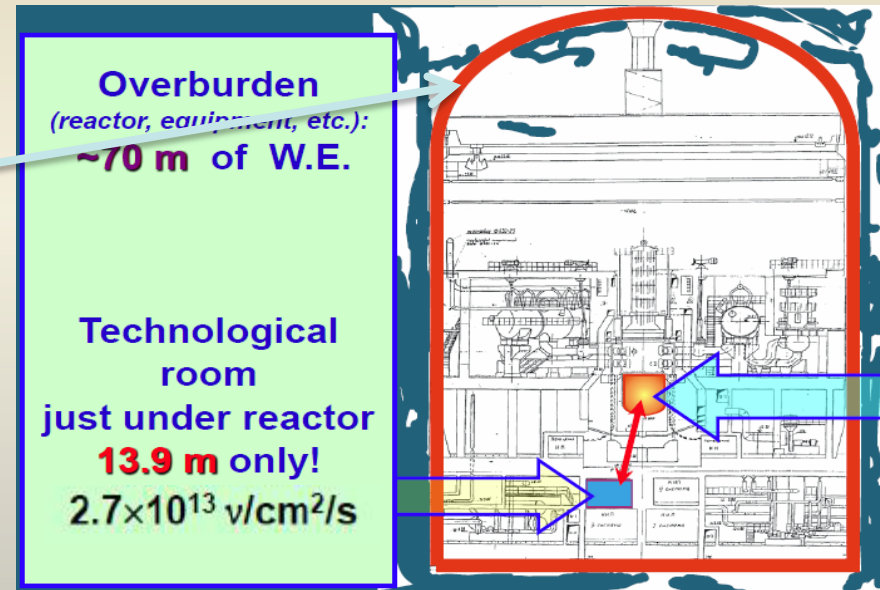
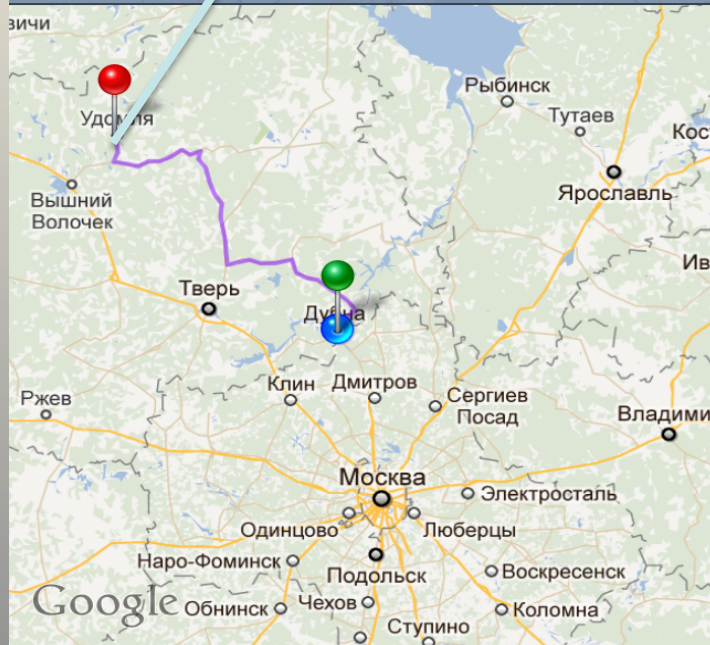
Chooz



Savannah River

1953 – first experiment at Hanford

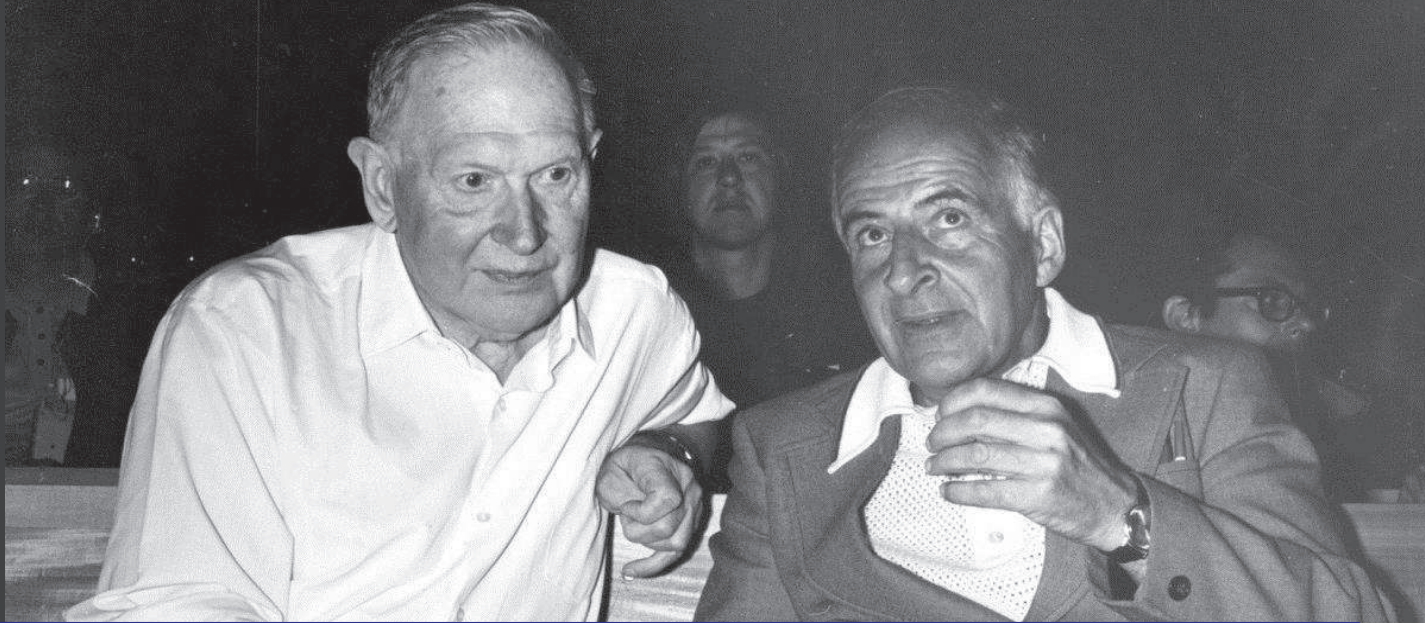
Experiments at Kalinin NPP



Fundamental and Applied Research:

- ✓ Search for Neutrino Magnetic Moment
- ✓ Measurement of Neutrino Fluxes and Spectra
- ✓ Search for Sterile Neutrino States

High energy neutrinos: Prolegomenon



M.Markov, **1960**:

„We propose to install detectors to determine the direction of charged particles deep in a lake or in the sea and with the help of Cherenkov radiation“ Proc. 1960 ICHEP, Rochester, p. 578.

Neutrino Astronomy

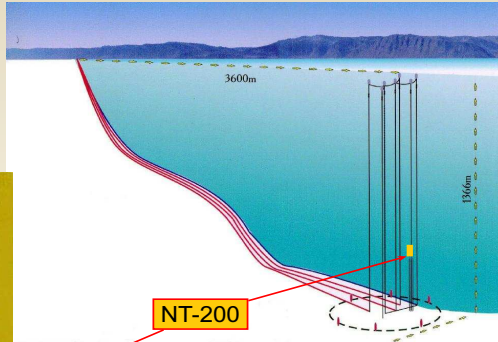
Among others, underwater/under-ice detectors are special

Main advantage: large scale.

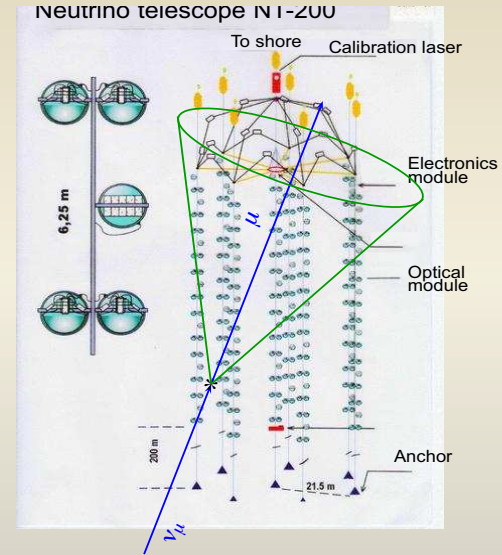
Multi-purpose

- Neutrino physics (in future) – mass hierarchy, CP-violation
 - With atmospheric neutrinos
 - Very long baseline experiments
- Search for dark matter: annihilations in the Earth and Sun
$$X + \bar{X} \rightarrow \pi^{\pm} + \dots \rightarrow \nu, \bar{\nu}$$
- Search for exotica (magnetic monopoles, quark nuggets, violation of Lorentz-invariance...)
- Cosmic ray physics
- Environmental studies
- Many more

Neutrino Astronomy



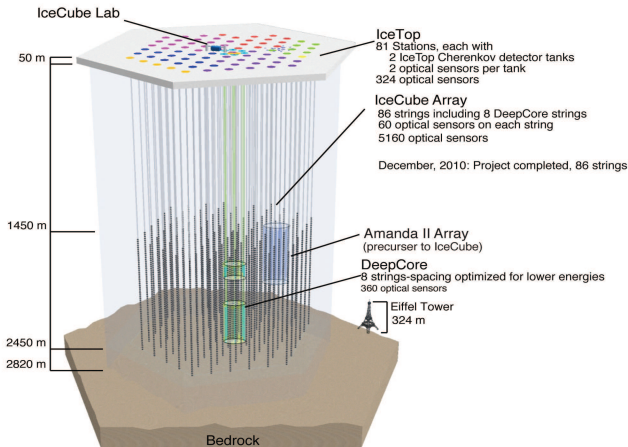
**PROOF OF PRINCIPLE:
BAIKAL**



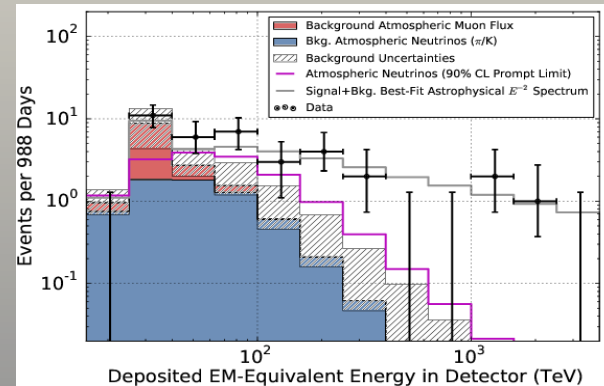
Newsmaker

The IceCube Neutrino Observatory

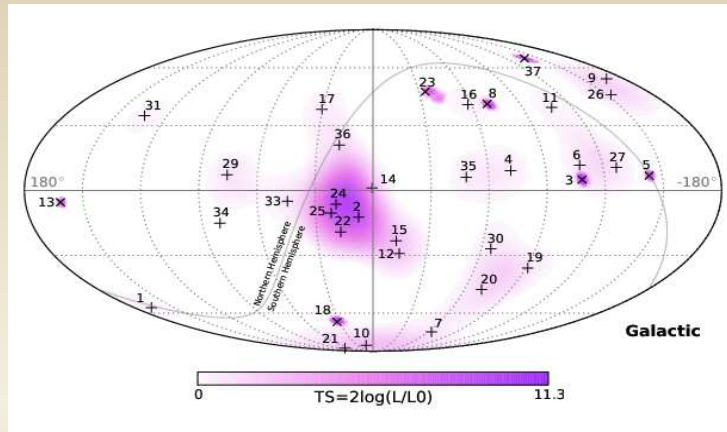
- Total of 86 strings and 162 IceTop tanks;
- Full operation with all strings since May 2011.



First neutrinos of cosmic origin



Arrival directions



- Are there cosmic neutrinos indeed?
- Neutrinos from Galaxy or outside?
- Most likely from Galaxy, then

- Point sources?
- Galactic center (black hole)?
- Fermi bubbles?

Key: angular resolution \implies water, not ice!

There must be hard gammas as well: both

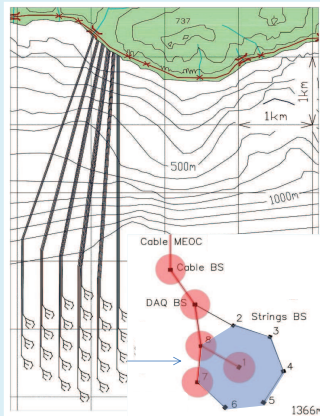
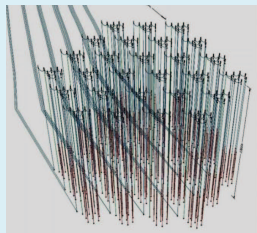
$$\pi^\mp \rightarrow \mu^\mp + \nu_\mu(\bar{\nu}_\mu), \text{ and } \pi^0 \rightarrow \gamma\gamma$$

HiScore, HAWK, CTA

All this is within reach of Baikal GVD

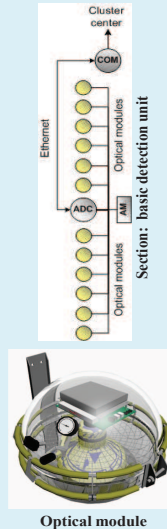
Gigaton Volume Detector (Lake Baikal)

10368 photo-sensors at 216 strings
 27 subarrays (clusters with 8 strings)
 String: 4 sections, 48 photo-sensors
 Active depths: 600 – 1300 m
 To Shore: 4 – 6 km
Instrumented water volume
 $V = 1.5 \text{ km}^3$ $S = 2 \text{ km}^2$
Angular resolution
 Muons: 0.25 degree
 Showers: 3.5-5.5 degree



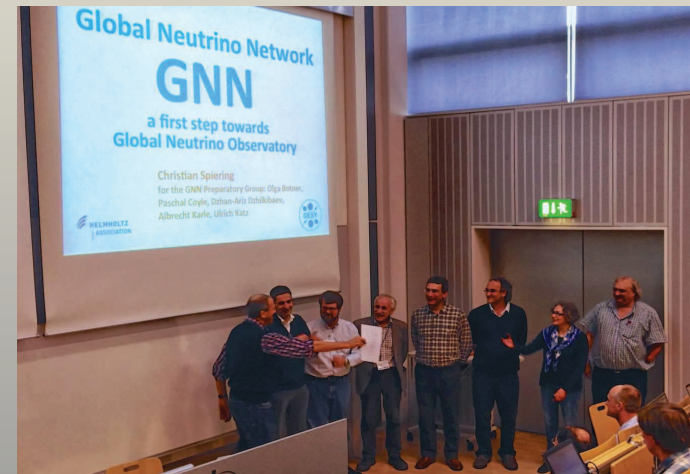
GVD array

1st GVD cluster: 8 strings
 ● - Installed strings and cable stations



Optical module

Good company: GVD, IceCube, KM3Net



October 15, 2013. München. MoU signing.

From left to right: Christian Spiering, Maarten de Jong, Tyce DeYoung, Zhan.-Arys Djilkibaev, Juan-Jose Hernandez-Rey, Paschal Coyle, Olga Botner, Uli Katz.



Future of Daya Bay is **JUNO**, aimed at neutrino mass hierarchy. Apart from the mass hierarchy, with JUNO it would be possible to look for SN-neutrinos, geo-neutrinos, sterile neutrinos and may be even CP-violation ...

White Book and D.Naumov's talk

Reach physics goals

- Supernova neutrinos (less than 20 events so far)
 - $\bar{\nu}_e + p \rightarrow n + e^+$, ~ 3000 correlated events
 - $\bar{\nu}_e + {}^{12}\text{C} \rightarrow {}^{12}\text{B}^* + e^+$, ~ 10 -100 correlated events
 - $\nu_e + {}^{12}\text{C} \rightarrow {}^{12}\text{N}^* + e^-$, ~ 10 -100 correlated events
 - $\nu_x + {}^{12}\text{C} \rightarrow {}^{12}\text{C}^* + \nu_x$, ~ 600 correlated events
 - $\nu_x + p \rightarrow \nu_x + p$, single events
 - $\nu_x + e^- \rightarrow \nu_x + e^-$, single events
- Geoneutrinos
 - 10 times more than recorded by BOREXINO and KamLAND
 - Difficult on systematics
 - Background to reactor antineutrinos

Tasks to be solved

- Large detector (20 kt of LS): design, mechanics, chemistry, stability
- Energy resolution $3\%/\sqrt{E}$ (1200 p.e./MeV)
 - Highly transparent LS
 - High light yield
 - High (80%) PMT coverage
 - High QE PMT (40-50%) \rightarrow a number of new problems to solve

JINR possible contribution:

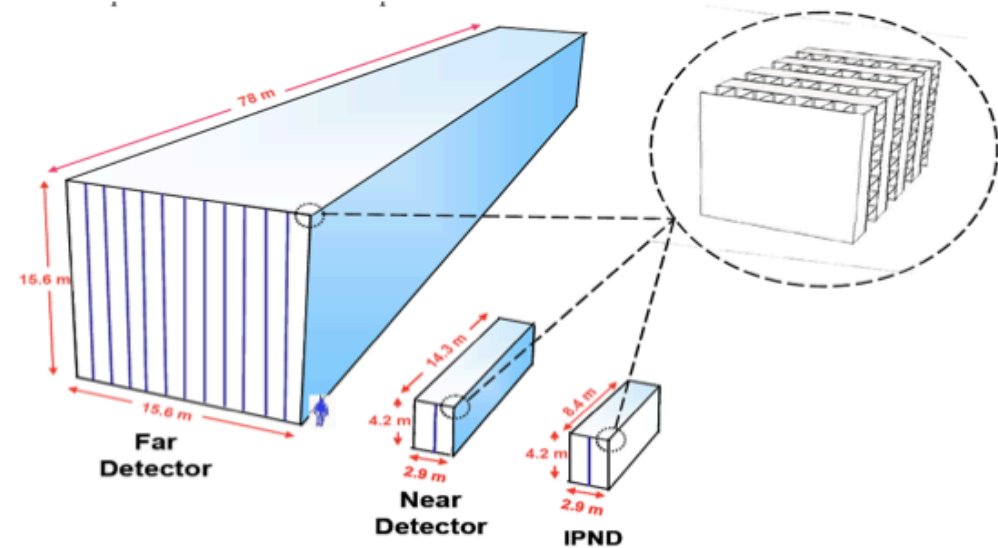
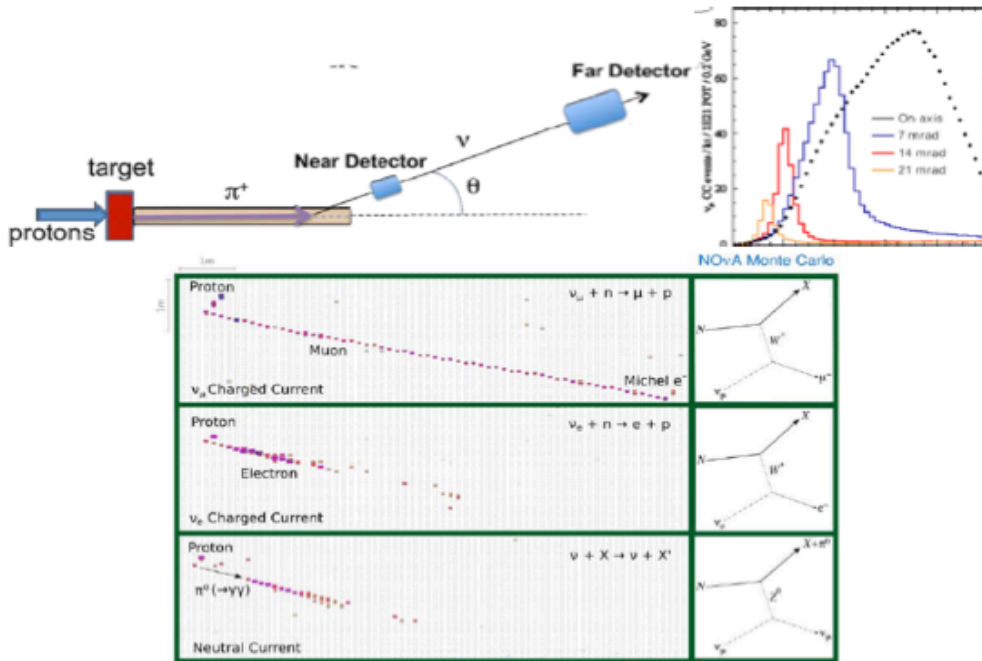
- Intelligent HV system
- PMT protection against Earth magnetic field
- Construction of a dedicated laboratory for large PMT tests and LS studies
- μ -veto based on OPERA plastic scintillator
- Detector design
- Simulation and reconstruction
- Data analysis



NOvA (NuMI Off-Axis ν_e Appearance) — a new generation accelerator long baseline experiment for study $\nu_\mu \rightarrow \nu_e$ oscillations.



The goal is to precisely measure the **parameters of the neutrino mixing matrix**, the **neutrino mass hierarchy** and **CP violation effects** in the lepton sector.



The NOvA apparatus consists of a **Near Detector** (220 ton) on the **Fermilab** site where the ν_μ -s are produced, and a **Far Detector** (14 kton) 810 km distant, both filled with liquid scintillator, have similar construction and situated 14 mrad **off-axes** to the ν beam. Detectors will be ready to reach full data-taking capability in 2014. The following **6 years** of data taking are optimized for running with ν and $\bar{\nu}$ beams.

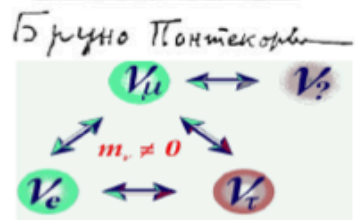


Neutrino history in Dubna — B. Pontecorvo and M. Markov!



About 55 years ago the idea of neutrino oscillation was born in Dubna!

Bruno Pontecorvo:
“The oscillations are very simple trick: 1–2–3 and ... all OK!”



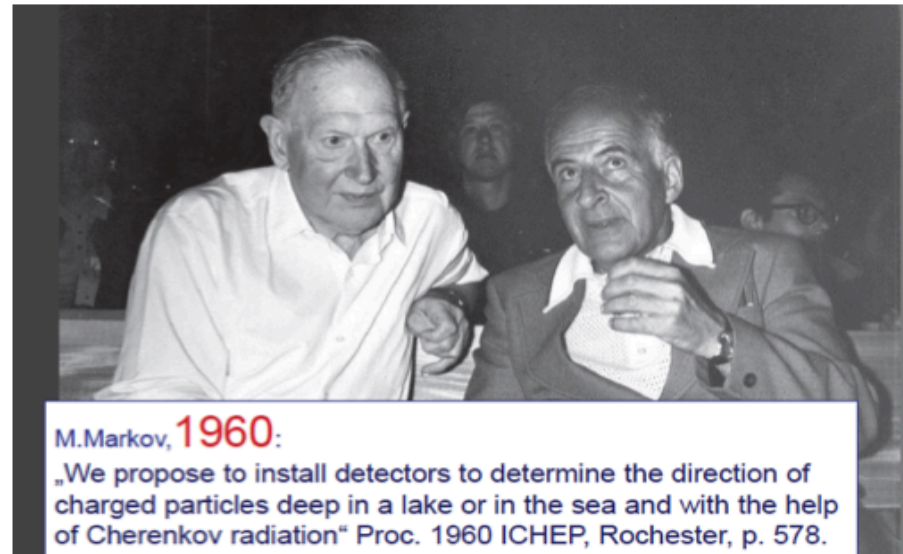
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Weak eigenstates Atmospheric CP phase Sub-dominant Solar Mass eigenstates
 $c_{ij} = \cos\theta_{ij}$, $s_{ij} = \sin\theta_{ij}$ θ_{13} oscillations

The beginning of neutrino astrophysics ...

M.A. Markov had proposed for the first time to define incoming directions of a cosmic ray charged particle in water by means of its Cherenkov radiation.

The first realization of this idea was the Baikal neutrino telescope NT-200. This is foundation of the IceCube success of today!



M. Markov, 1960:
„We propose to install detectors to determine the direction of charged particles deep in a lake or in the sea and with the help of Cherenkov radiation“ Proc. 1960 ICHEP, Rochester, p. 578.

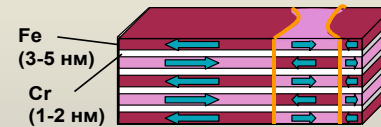


Opening of the monument to Bruno Pontecorvo and Venedict Dzhelepov at Dubna on 20 September 2013

Condensed Matter Physics

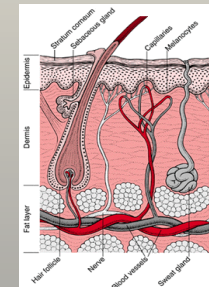
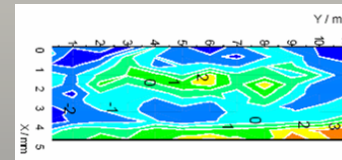
IBR-2 is included in the 20-year European strategic research program in the field of neutron scattering

Nanosystems and Nanotechnology



Biomedical research

New materials



IBR-2

Physics of high-temperature superconductivity

Geological texture research

Nanotechnology

Diagnostics. Earth science.

Upgraded IBR-2

Pulsed reactor with fast neutrons

mean power **2 MW**

pulse frequency **5 Hz**

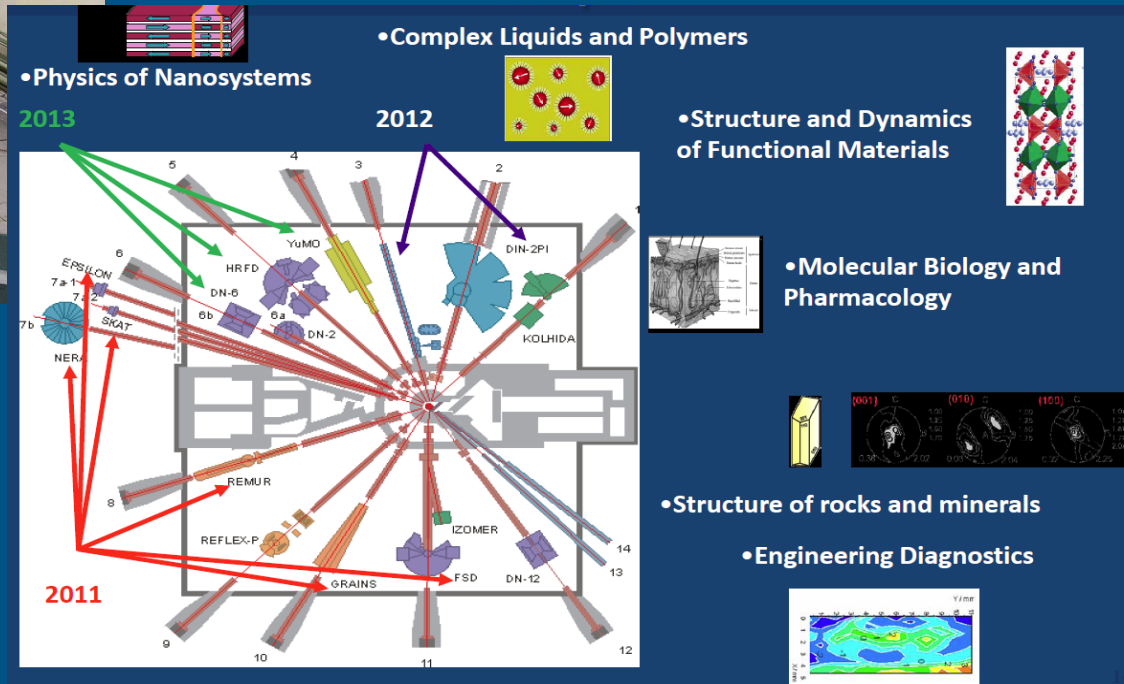
pulse width for fast neutrons **200 μ s**

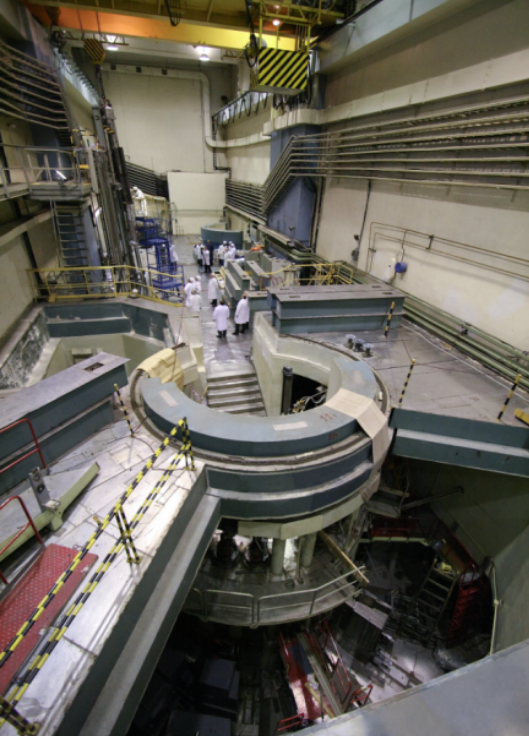
thermal neutrons flux density on the moderator surface: **10^{13} n/cm² /s**

maximum in pulse: **10^{16} n/cm² /s**



and spectrometer complex

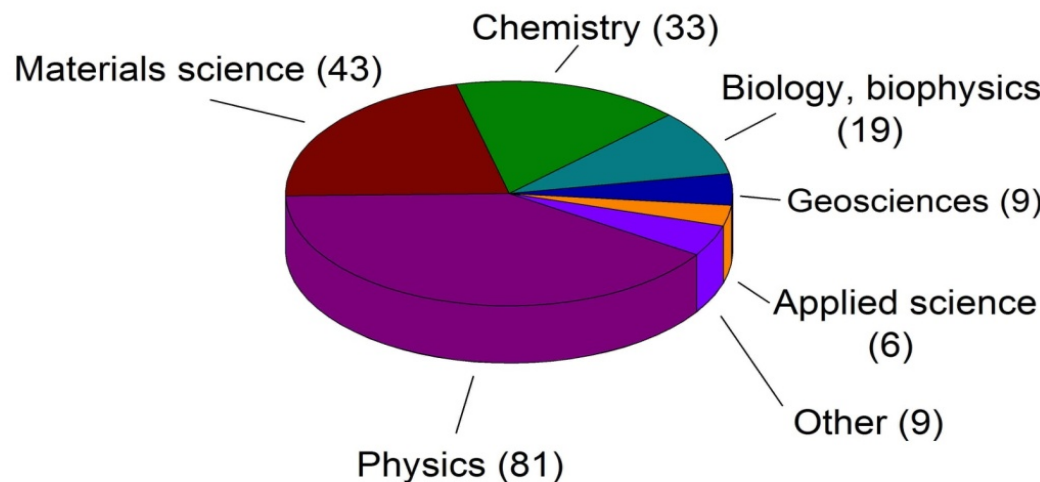
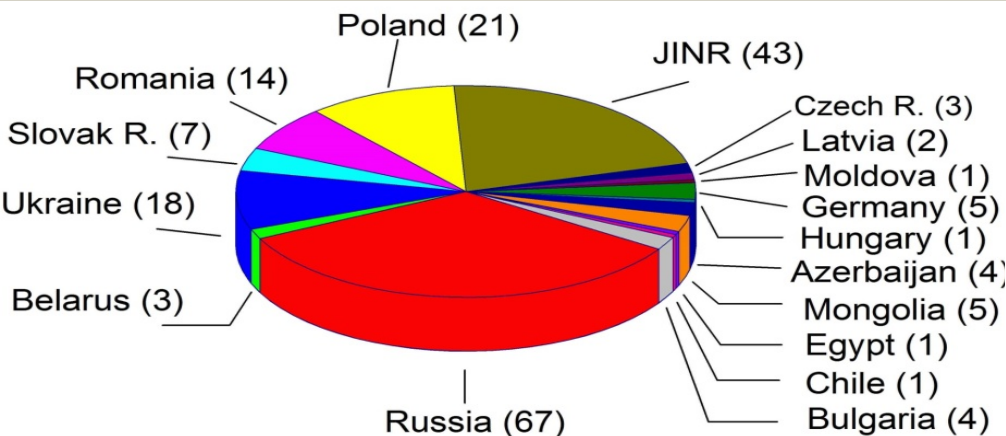




IBR-2

2578 hours for physical experiments, 12 cycles:
7 – water moderator, 6 – cryogenic moderator

- 195 proposals received for realization in 2013 during two calls
- 70% accepted for realization according to recommendations of Expert Committees
- 118 visits in the FLNP



THEORETICAL PHYSICS @ JINR

Main fields of research

- Theory of Elementary Particles and Fields
- Nuclear Theory, Nuclear Structure and Dynamics
- Theory of Condensed Matter and New Materials
- Modern Mathematical Physics
- Research and Education Project
“Dubna International School of
Theoretical Physics (DIAS-TH)”



Conferences and Schools

Every year (> 1000 participants)
DIAS-TH and Helmholtz Schools
(> 20 countries represented)

Educational Activity

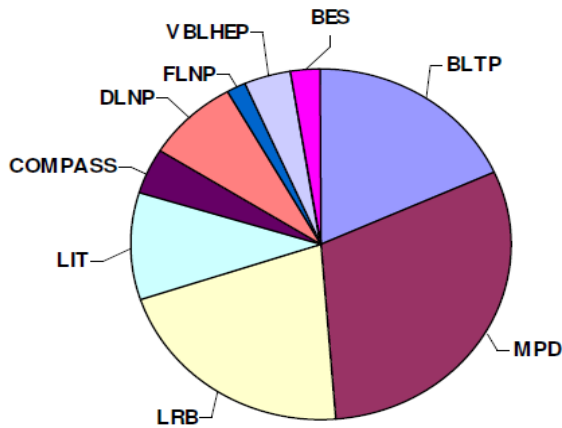
**Lectures courses at JINR UC,
DIAS-TH, Moscow U., Dubna U., MPTI, etc.**



JINR Central Information and Computing Complex

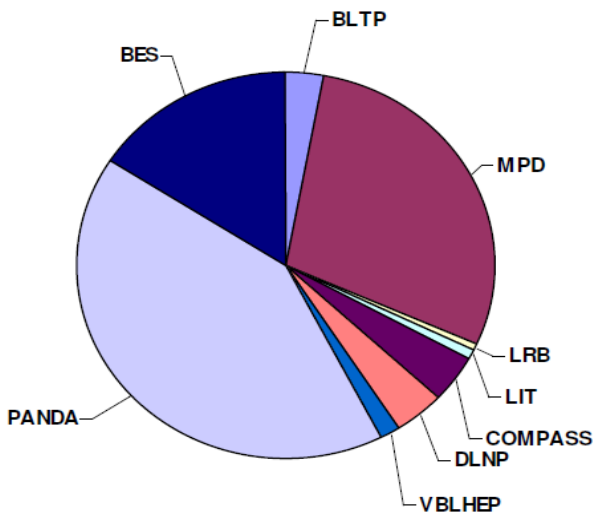
Grid users

Local users (no grid)



JINR CICC
Normalised CPU
time share: JINR
Laboratories and
experiments in
2013.

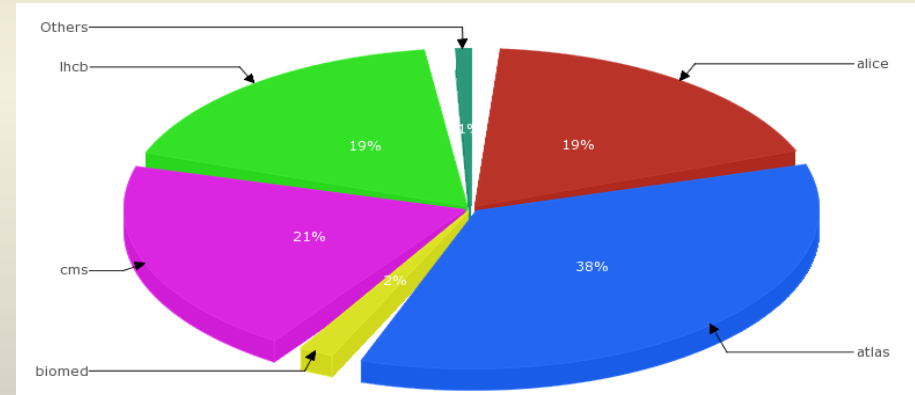
Total CPU time -
4 580 752



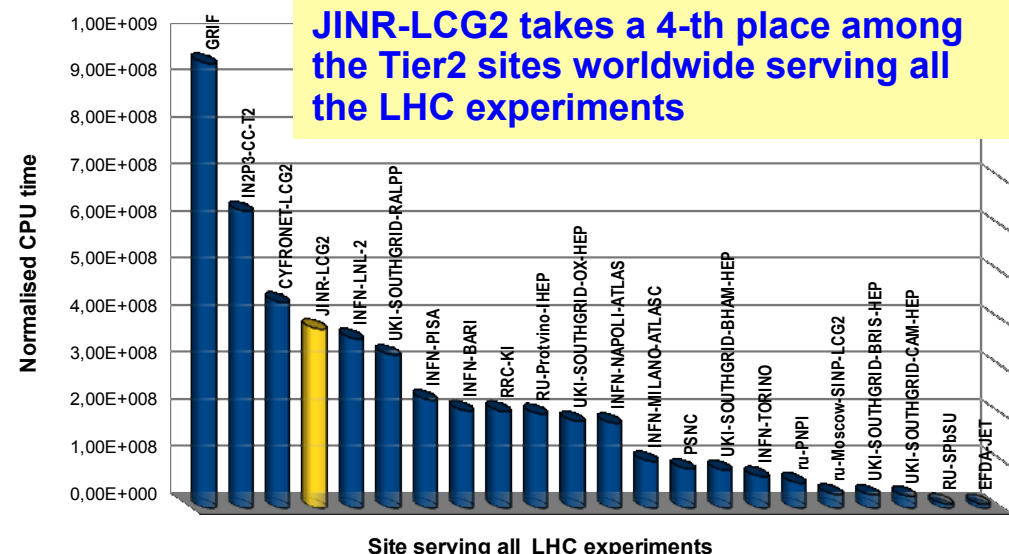
JINR CICC
Jobs executed
for JINR
Laboratories and
experiments in
2013:

226 832

JINR-LCG2 Normalised CPU time (2013)



About 5 million jobs were executed

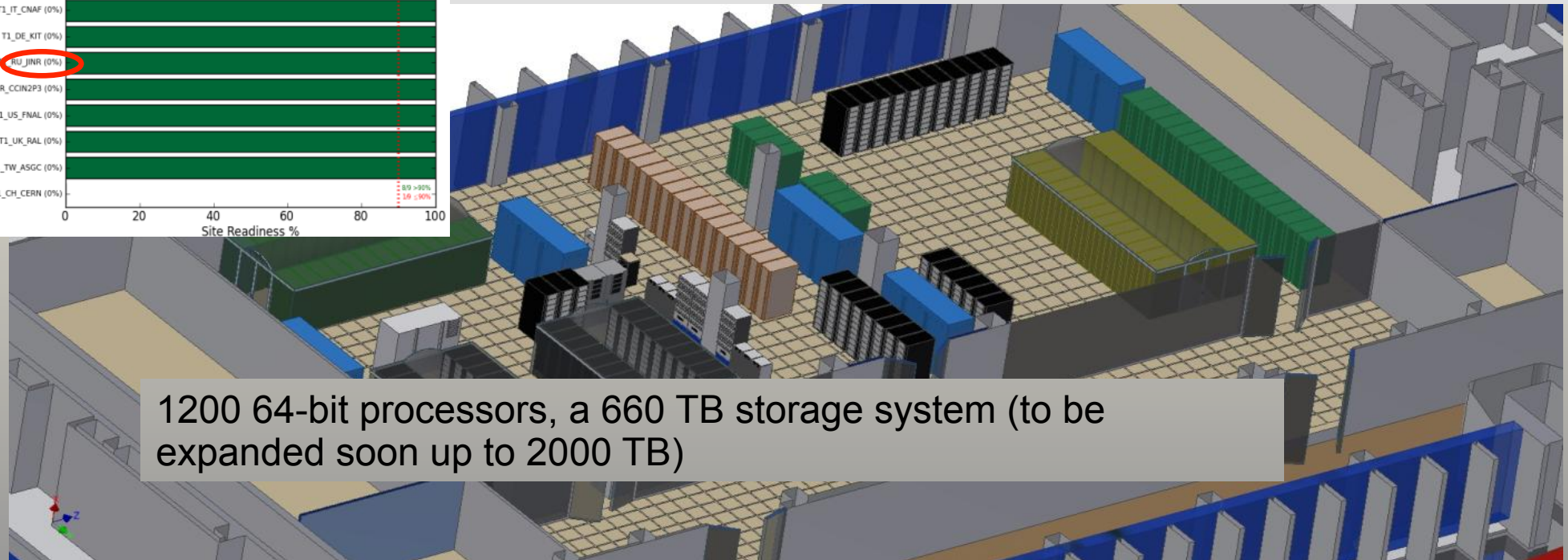
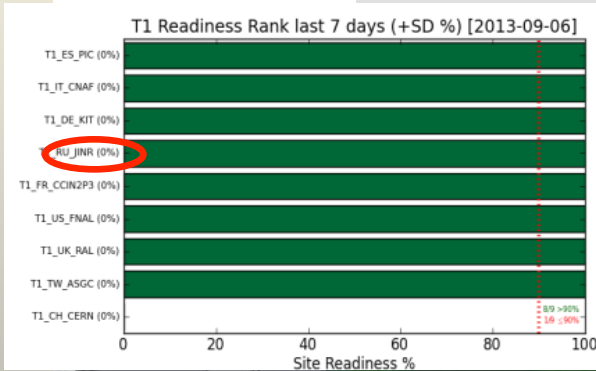


JINR-LCG2 takes a 4-th place among the Tier2 sites worldwide serving all the LHC experiments

Site serving all LHC experiments

CMS Tier-1 prototype

- Engineering infrastructure(uninterrupted power supply, climate-control);
- High-speed reliable network infrastructure with dedicated reserved data link to CERN (LHCOPN);
- Computing system and storage on the base of disk arrays and tape libraries of high capacity;
- Reliability and accessibility 100%





JINR Educational Program

At the moment 354 graduate and 37 postgraduate students take part in various JINR educational programs. The JINR University Center keeps updating the training programs according to the goals of the JINR seven-year plan.



International Student Practice (ISP)

In total 129 students from 9 JINR Member States were participated during three stages of ISP-2013: ARE, Belarus, Bulgaria, Czech Republic, Poland, Romania, Slovakia, South Africa, Ukraine



JINR Summer Students Program (SSP)

<http://students.jinr.ru/>

New!

In 2014 JINR UC launches the Summer Student Program. Main distinction SSP from ISP is a selection of participants on a competitive basis. A special web-site to coordinate this Program is already opened and we plan to start collecting applications for SSP-2014 by the mid of March.



JINR Outreach Activity



Programs at CERN and JINR started in November 2009 for the teachers from JINR Member States. Up to now 5 programs at CERN (193 participants) and 4 programs at JINR (176 participants)

New department “Development of the modern education programmes” was created at JINR University Center. One of the goals of this department is a creation of the educational programs to include current scientific data into the educational process, conduct virtual and online laboratory research based on information and communication technologies



UC continues organizing the video-conferences to promote achievements of modern science

CERN-JINR School 2014



Particle Physics (JINR outside)

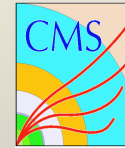
JINR @ CERN, BNL, Fermilab, GSI/FAIR, KEK, China

Bright ex.: **DISCOVERY OF HIGGS**–
THE STANDARD MODEL TRIUMPH –**HIGHEST**
ACHIEVEMENT OF MANKIND

JINR contribution: in **ATLAS**



& CMS



(experiments)

In theory:

- Transfer of spontaneous symmetry breaking phenomenon in condensed matter physics (N.N. Bogolyubov) to elementary particle physics.
- **New quantum number of quarks (color) introduction.**
- Renormalization principle in Quantum Field Theory.
- **Renormalization group and “Asymptotic freedom” phenomenon.**

Collaboration in Particle Physics

JINR @ CERN, BNL, Fermilab, GSI/FAIR, KEK, China

- I. **CERN (LHC):** LHC development – consolidation of SC magnets;
CMS, ALICE and ATLAS – data taking & analysis;
upgrade of all 3 detectors – moderate *additional resources*;
- II. **CERN (SPS):**
COMPASS – finished 1st phase. Detector modification to
measure GPD (DVCS) and polarized/unpolarized D-Y;
NA61 – neutrino and heavy-ion programs;
NA62 – measurement of extremely rare decays ($K^+ \rightarrow \pi^+ \nu \nu$) ;
DIRAC – lifetime measurement of $\pi\pi$ and πK atoms completed at PS;
collaboration formed to continue at SPS;
- III. **BNL (RHIC):**
STAR - energy scan HI program and physics with polarized beams
(important experience for future research at NICA)
- IV. **Fermilab:** **NovA** neutrino program – in progress, Mu2e ($\mu \rightarrow e$) – in discussion
- V. **GSI, FAIR (SIS-18/100/300):** HADES – on the beam
CBM, PANDA – in preparation
- VI. **J-PARC & KEK:** COMET ($\mu \rightarrow e$), in progress
- VII. **China:** Daya Bay Neutrino program new results in sin2Theta_13

CERN-JINR Partnership in Particle Physics



Done at Geneva on 28 January 2010, in two copies in the English language.

For the European Organization
for Nuclear Research (CERN)

Prof. Rolf Dieter Heuer

For the Joint Institute
for Nuclear Research (JINR)

Prof. Alexei N. Sissakian



28 January 2010, CERN Signing of the Agreement between CERN and JINR

3.2 Possible projects at the date of this Agreement include:

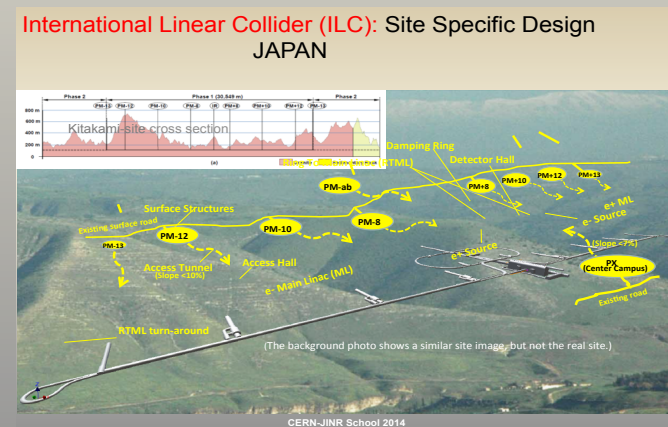
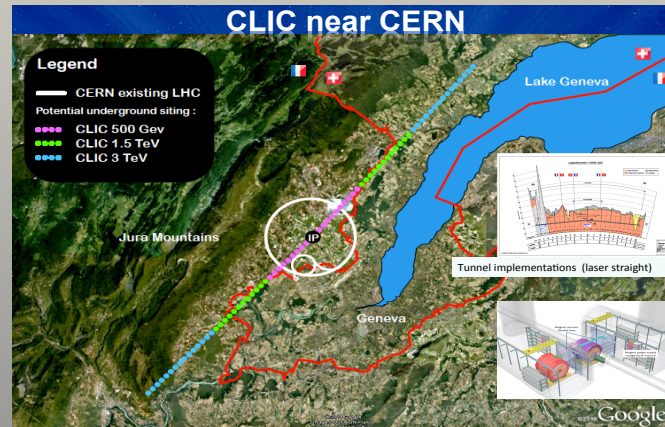
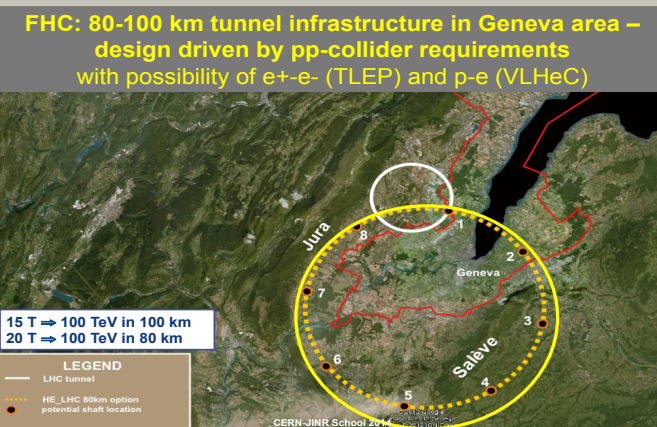
- the commissioning and operation of the Large Hadron Collider ("LHC") at CERN, including the ALICE, ATLAS and CMS experiments using the LHC;
- upgrades of the Nuclotron and the construction, commissioning and operation of the NICA collider project at JINR, including the MPD and SPD experiments using NICA;
- upgrades of the LHC injector chain, including the Linac4, SPL and PS2 projects;

Strategy for Particle Physics

Future major facilities in Europe, Russia and elsewhere require collaboration on a global scale.

CERN & JINR combine the framework within which to organize a global particle & heavy ion physics accelerator projects in Europe & Russia, and should also be the leading partners in global particle & heavy ion physics accelerator projects elsewhere. Possible additional contributions to such projects from CERN's & JINR's Member and Associate Member States should be coordinated with CERN & JINR.

Strategy comprises **high priority items** for large-scale scientific activities which could/should be organised as international projects



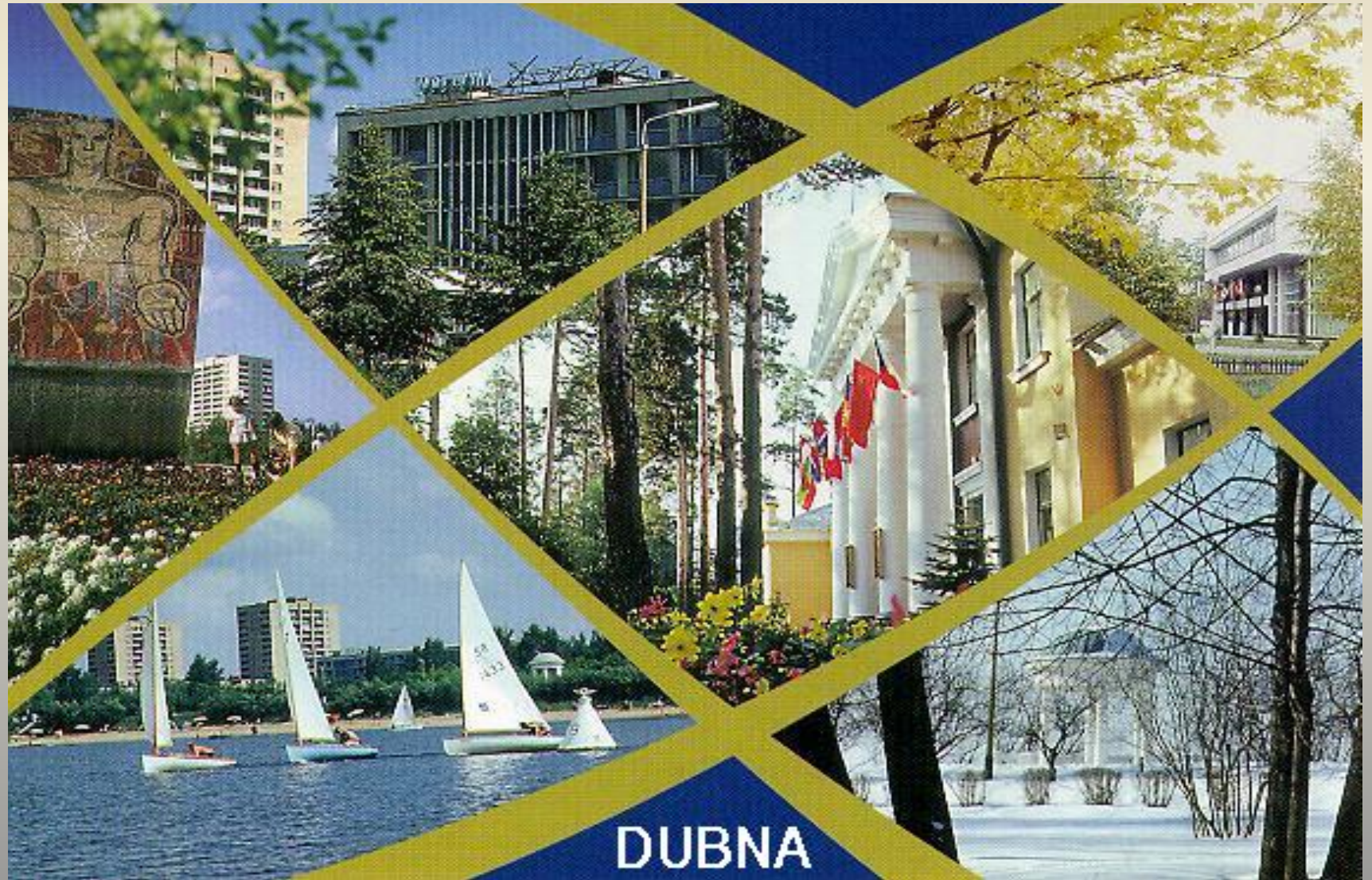
Summary: JINR World-Wide Collaboration

To realise our ambitious scientific plans based on the **home** and out of JINR large scale facilities and to fully exploit their excellent physics capabilities we suggest

- coherent and strong efforts of JINR member-states in all regions;
- long term stability and support in all regions;
- careful adjustment of schedules for these facilities

Global Partnership
with
long-term support

THANK YOU!



DUBNA

CERN-JINR School 2014