




Recent Developments in the Characterization of QGP



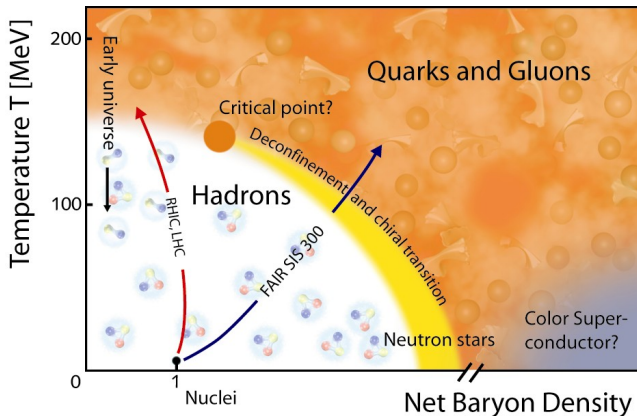
Sangyong Jeon

*Department of Physics
McGill University
Montréal, QC, CANADA*

- Charles Gale
- Sangyong Jeon
- *Björn Schenke*
(Formerly McGill, now BNL)
- *Clint Young*
(Formerly McGill, UMN now MSU)
- *Matt Luzum*
(Formerly McGill, now LBNL)
- *Gabriel Denicol*
- *Chun Shen* (Joining in October 2014)
- Sangwook Ryu
- Gojko Vujanovic
- Jean-Francois Paquet
- Michael Richard
- Igor Kozlov
- Khadija El Berhoumi
- Jean-Bernard Rose
- Chanwook Park

and many thanks to G.D. Moore, S. Caron-Huot, R. Venugopalan, P. Tribedy, U. Heinz, D. Srivastava, S. Bass, C. Nonaka, M. Mustafa, E. Frodermann, R. Fries, A. Majumder, L. Yaffe, P. Arnold, and others ...

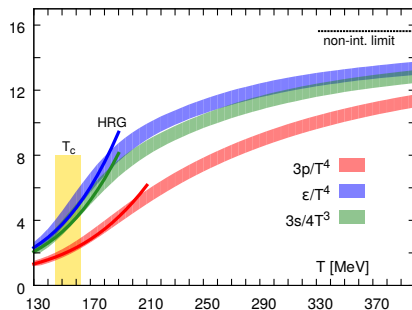
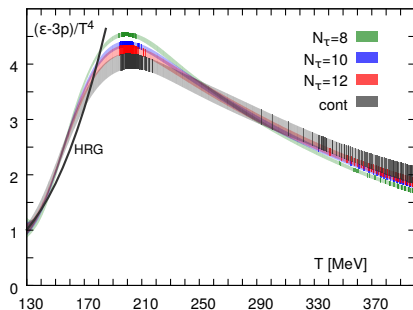
Many-body QCD



Picture credit: GSI (www.gsi.de)

QGP: WHAT do we know and **HOW** do we know about it?

LQCD evidence of QGP



HotQCD Collaboration arXiv:1407.6387

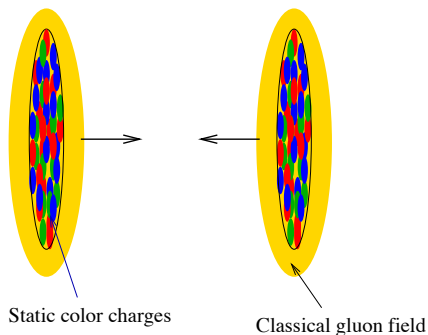
- QGP phase transition:
Rapid crossover with light u, d and heavier s
- $T_c \sim 155$ MeV

What is a QGP?

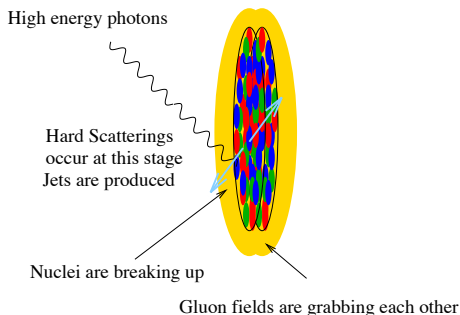
- Deconfined state of quarks and gluons
- $T \gtrsim T_c \sim \Lambda_{\text{QCD}} \sim 200 \text{ MeV}$
Or about 2 trillion kelvin. Sun's interior is only about 10 million kelvin.
- $\epsilon \gtrsim 3 \text{ GeV}/\text{fm}^3$
1 mm^3 of QGP contains about 5×10^{26} J. This can power a modern nation for about 10 million years.
- **Pressure** $\sim \epsilon/3$
This is about $10^{35} \text{ N}/\text{m}^2$.
Compare: $M_{\text{Earth}}g \approx 6 \times 10^{25} \text{ N}$.

Observable Consequences

- $T \sim \Lambda_{\text{QCD}} \implies$ Multiplicity, Spectra
- High density \implies Jet quenching
- High pressure \implies Hydrodynamic flow
 - Weakly coupled: $\eta/s \gg 1$
 - Strongly coupled: $\eta/s \ll 1 \approx$ Perfect (Ideal) fluid.
- Small $\langle \delta Q^2 \rangle / s \implies$ Tight unlike-sign correlation
- Critical point \implies Large momentum fluctuations
- Must survive during the evolution including re-equilibration in the hadronic phase



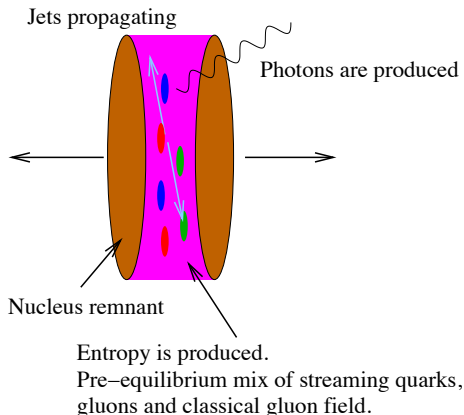
- Nuclear initial state: Static color charges (large x partons) + Classical gluon field (small x gluons) \implies Color Glass Condensate.



- On contact:

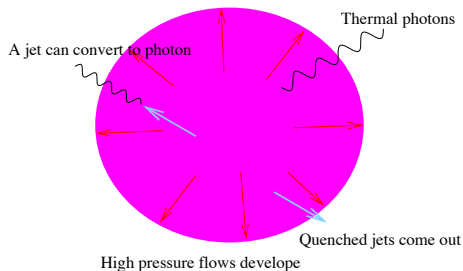
- Gluon fields collide and start to evolve towards QGP \implies Glasma.
- Jets, heavy quarks, prompt photons are produced by high p_T parton collisions

Creating QGP

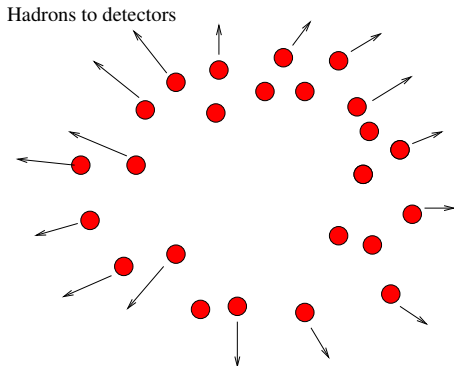


- Pre-Equilibrium

- Glasma evolving towards QGP
- Jets, heavy quarks, prompt photons propagating



- Hydrodynamic stage - QGP is formed
 - Approximate local equilibrium achieved \implies Viscous Hydrodynamics applies
 - Jets interact with QGP, quarkonia dissolves, thermal photons and dileptons produced



- After hadronization:

- System may be too dilute for hydrodynamics
- Chemical freeze-out first
- Kinetic freeze-out later

- Soft (Roughly in order of increasing complexity)
 - Charge fluctuations
 - Multiplicities
 - Hydrodynamic flows and viscosities
- Hard & EM probes
 - Jet Quenching
 - Electromagnetic probes

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 - *Jet Quenching* *
 - Electromagnetic probes

*: Cover if time permits

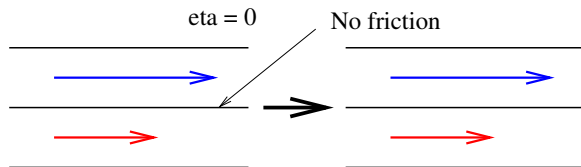
Hydrodynamic Flow and Viscosities

Shear Viscosity

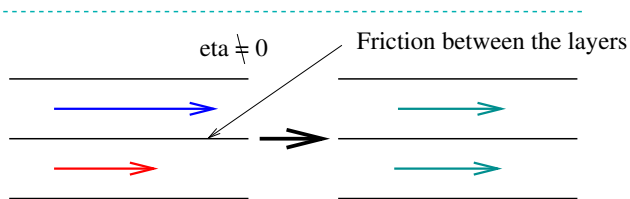
- QGP property to learn:
 - Specific shear viscosity: η/s

Shear viscosity in context

Differential hydrodynamic flow

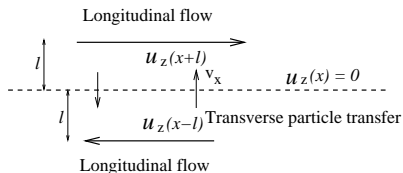


The relative velocity of the two layers does not change.



The velocities eventually become the same.

Kinetic Theory estimate (Why η/s is natural)



u_z : Flow velocity
 v_x : Average speed of microscopic particles

- Rough estimate (fluid rest frame, or $u_z(x) = 0$)
 - The momentum density: $T_{0z} = (\epsilon + \mathcal{P})u_0 u_z$ moves in the x direction with v_x

$$\begin{aligned} &\langle \epsilon + \mathcal{P} \rangle |v_x| u_0 (u_z(x - l_{\text{mfp}}) - u_z(x + l_{\text{mfp}})) \\ &\sim -\eta u_0 \partial_x u_z \end{aligned}$$

Here l_{mfp} : Mean free path

- Recall thermo. id.: $\langle \epsilon + \mathcal{P} \rangle = sT$

$$\eta \sim \langle \epsilon + \mathcal{P} \rangle l_{\text{mfp}} \langle |v_x| \rangle \sim sT l_{\text{mfp}} \langle |v_x| \rangle$$

Perturbative estimate

High Temperature limit: $\langle |v_x| \rangle = O(1)$

- $\eta/s \approx T l_{\text{mfp}} \approx \frac{T}{n\sigma} \sim \frac{1}{T^2\sigma}$ ($\sim \Delta p \Delta x \geq \hbar/2$)
- The only energy scale: T

$$\sigma \sim \frac{(\text{coupling constant})^\#}{T^2}$$

Hence

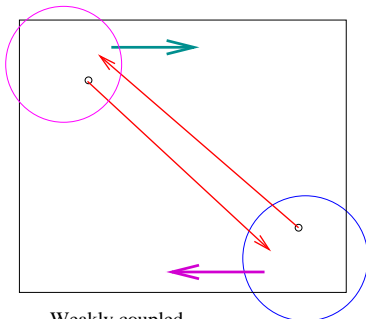
$$\frac{\eta}{s} \sim \frac{1}{(\text{coupling constant})^\#}$$

- In QCD, naively expect $\eta/s \sim \frac{1}{\alpha_s^2}$
- Coulomb enhancement (cut-off by m_D) leads to

$$\eta/s \sim \frac{1}{\alpha_s^2 \ln(1/\alpha_s)}$$

Interaction Strength and Viscosity

Weak coupling allows rapid momentum diffusion

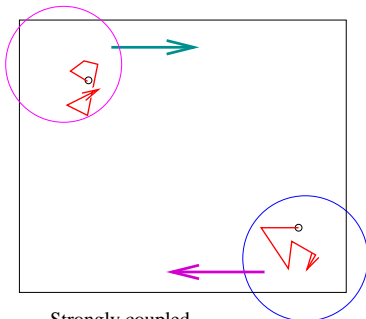


Weakly coupled
Long distance until next collision
Easy mixing

Large η/s : $u_\mu(x)$ changes due to pressure gradient and diffusion

Interaction Strength and Viscosity

Strong coupling does not allow momentum diffusion

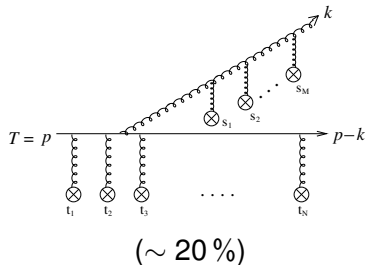
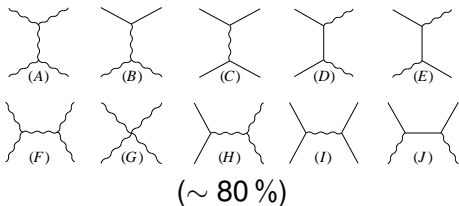


Strongly coupled
Very short distance until next collision
Mixing takes very long time

Small η/s : $u_\mu(x)$ changes due to pressure gradient only

Full leading order QCD calculation of η/γ

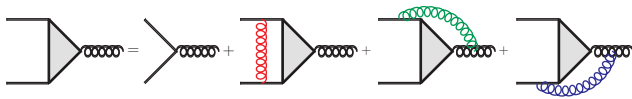
Relevant processes



Use kinetic theory (equiv. to Kubo when linearized)

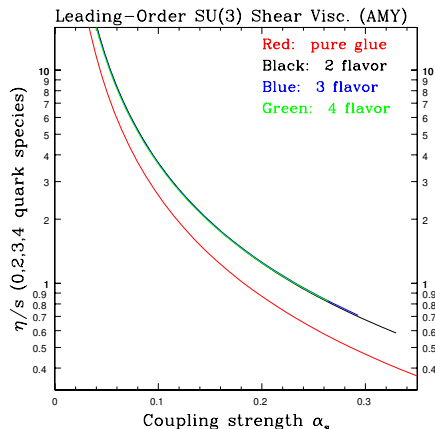
$$\frac{df}{dt} = \mathcal{C}_{2 \leftrightarrow 2} + \mathcal{C}_{1 \leftrightarrow 2}$$

Complication: $1 \leftrightarrow 2$ process needs resummation (LPM effect, AMY)
SD-Eq (Figure by G. Qin):



Full leading order QCD calculation of η/s

- Arnold-Moore-Yaffe (JHEP **0305**, 051 (2003)) [Updated figure from G.D. Moore]:



- Minimum $\eta/s \approx 0.6 \approx 7.5/4\pi$ for $\alpha_s \approx 0.3$
- Approximate formula

$$\eta/s \approx \frac{1}{15.4\alpha_s^2 \ln(0.46/\alpha_s)}$$

Shear viscosity in strongly coupled $\mathcal{N}=4$ SYM

Son, Starinets, Policastro, Kovtun, Buchel, Liu, ...

- Strong coupling limit, 4 ingredients
 - Kubo formula

$$\eta = \lim_{\omega \rightarrow 0} \frac{1}{2\omega} \int dt d^3x e^{i\omega t} \langle [T_{xy}(x), T_{xy}(0)] \rangle$$

- Gauge-Gravity duality

$$\sigma_{\text{abs}}(\omega) = \frac{8\pi G}{\omega} \int dt d^3x e^{i\omega t} \langle [T_{xy}(x), T_{xy}(0)] \rangle$$

- $\lim_{\omega \rightarrow 0} \sigma_{\text{abs}}(\omega) = A_{\text{blackhole}}$
- Entropy of the BH : $s = A_{\text{blackhole}}/4G$

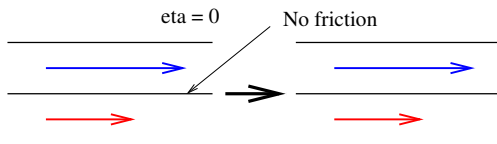
Therefore, (including the first order correction)

$$\frac{\eta}{s} = \frac{1}{4\pi} \left(1 + \frac{7.12}{(g^2 N_c)^{3/2}} \right)$$

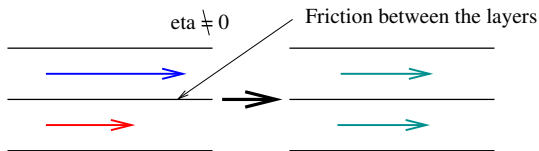
Correction is small if $g \gg 1$ (10% at $g = 2.4$).

To measure shear viscosity

Create this situation (Differential flow)



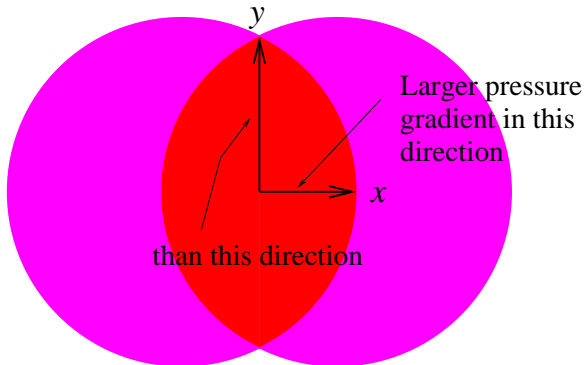
The relative velocity of the two layers does not change.



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Hydrodynamic response of the system to the initial spatial anisotropy

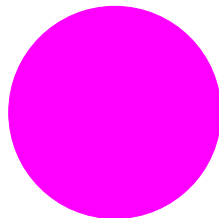
Beam direction: z



Momentum anisotropy develops

$$\epsilon = \frac{\langle y^2 - x^2 \rangle}{\langle x^2 + y^2 \rangle}$$

Spatial evolution

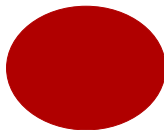
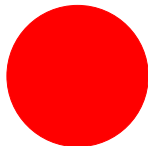


$\Delta P/\Delta x$ is greater than
 $\Delta P/\Delta y$

No more differential flow

Momentum space evolution

$$v_2 = \frac{\langle p_x^2 - p_y^2 \rangle}{\langle p_x^2 + p_y^2 \rangle}$$

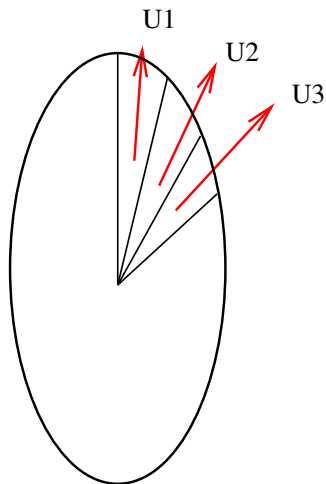


Initially spherical

Anisotropy is fixed

This happens within a few fm/c with $|v_{\text{fin.}}| \sim 0.6 c$.

Effect of viscosity



- $\eta = 0$ means $u_1 < u_2 < u_3$ is maintained for a long time
- $\eta \neq 0$ means that $u_1 \simeq u_2 \simeq u_3$ is achieved more quickly
- Shear viscosity smears out flow differences (it's a diffusion)
- Shear Viscosity **reduces** non-sphericity

- Information content of single particle spectra

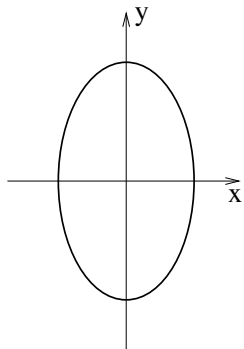
$$\frac{dN_i}{dy d^2p_T} = \frac{dN_i}{2\pi p_T dp_T dy} \left(1 + \sum_{n=1}^{\infty} 2v_{i,n}(p_T, y) \cos(n\phi) \right)$$

- “Flow”, $v_{i,n}(p_T)$, comes from the initial state energy density distribution via pressure

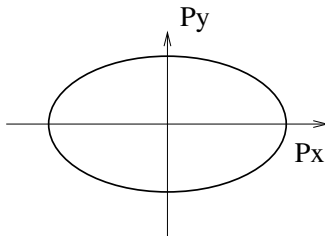
$$\varepsilon(\mathbf{x}_T, \eta) = \varepsilon(r_T, \eta) \left(1 + \sum_{n=1}^{\infty} 2\epsilon_n(r_T, \eta) \cos(n\phi) \right)$$

- Efficiency of $\epsilon_n \rightarrow v_n$ depends on η/s

- Elliptic Flow – $\cos(2\phi)$ component



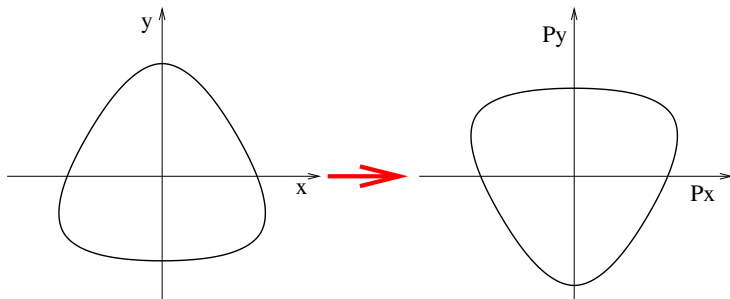
Spatial anisotropy



Pressure does
the conversion

Momentum anisotropy

- Triangular Flow – $\cos(3\phi)$ component



Spatial anisotropy

Pressure does
the conversion

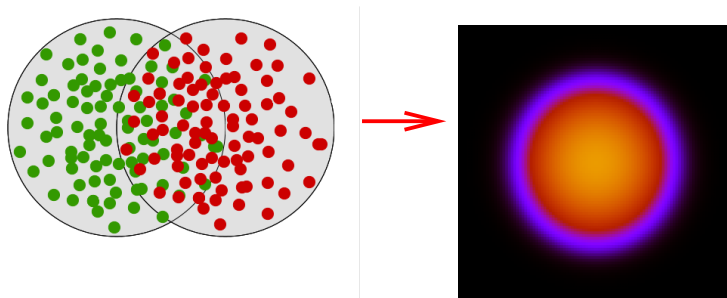
Momentum anisotropy

- In general: The higher n , the more sensitive to η/s

[Alver and Roland, Phys.Rev.C81:054905, 2010]

Physics from Hydrodynamics

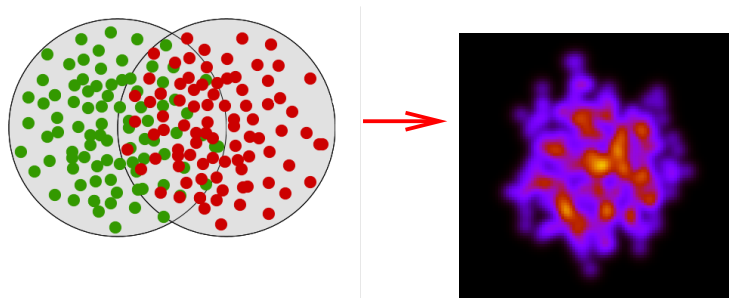
What determines the initial shape?



- Averaged smooth initial condition \implies Only v_{even} 's survive.

Physics from Hydrodynamics

What determines the initial shape?



- Fluctuating initial condition \implies All v_n are non-zero.
- Local fluctuations \implies Higher $v_n \implies$ More sensitive to dissipative smoothing



Middle

Prelude

MUSIC

MUSCl for Ion Collisions

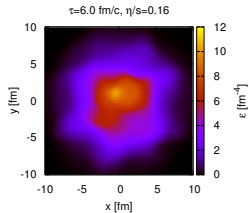
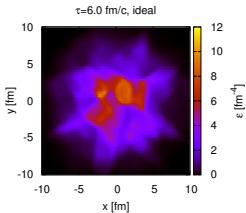
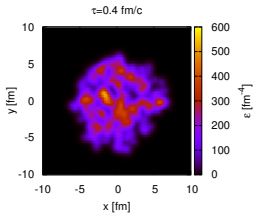
MUSCL: Monotone Upstream-centered Schemes for Conservation Laws

3+1D Event-by-Event Viscous Hydrodynamics

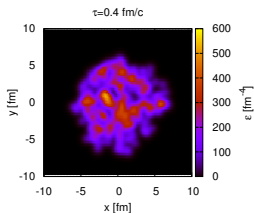
[Schenke, Jeon, Gale, Phys.Rev.Lett. 106 (2011) 042301, Phys.Rev. C82 (2010) 014903]

- 3+1D parallel implementation of [Kurganov-Tadmor Scheme](#) in (τ, η) with an additional baryon current (No need for a Riemann Solver. OK to take $\Delta t \rightarrow 0$.)
- Ideal and Viscous Hydro
- Event-by-Event fluctuating initial condition
- Sophisticated Freeze-out surface construction
- Full resonance decay (3+1D version of Kolb and Heinz)
- Many different equation of states including the newest from Huovinen and Petreczky
- New Development: UrQMD afterburner + Jet propagation (MARTINI)

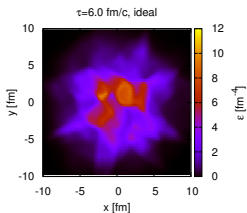
Ideal vs. Viscous



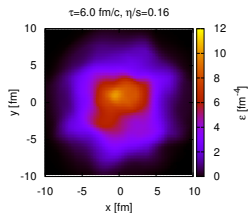
Effect of viscosity



Initial condition



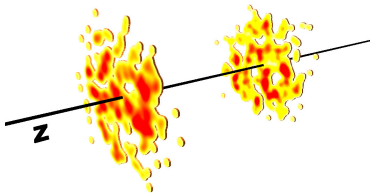
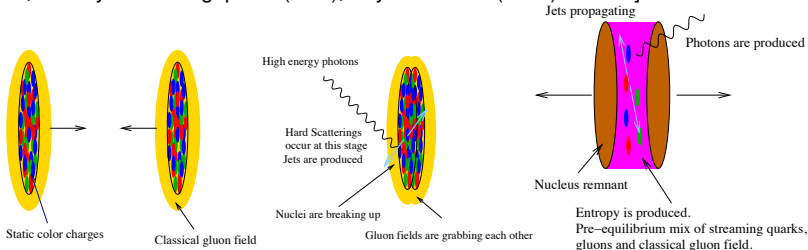
Ideal, $\tau = 6$ fm



Viscous, $\tau = 6$ fm

- Smoothing is evident
- Overall asymmetry is less affected
- The hot spots get much rounder
- v_3 is more affected than v_2

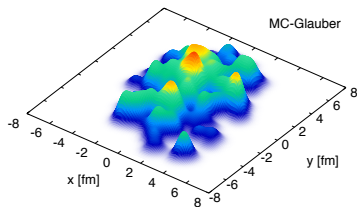
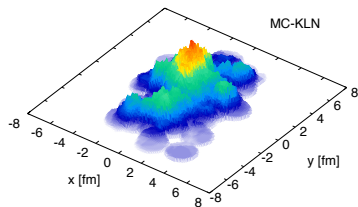
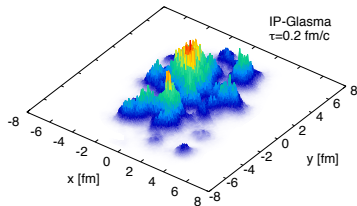
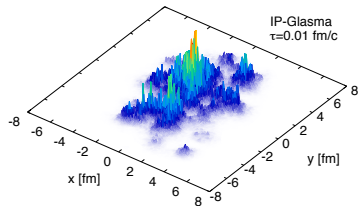
[Schenke, Tribedy and Venugopalan (STV), Phys.Rev. C86 (2012) 034908]



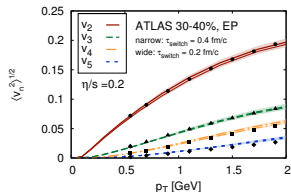
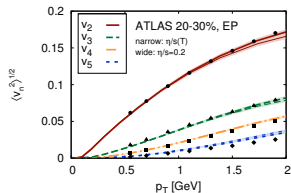
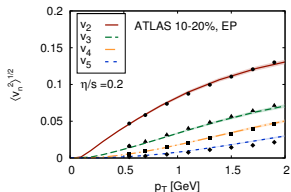
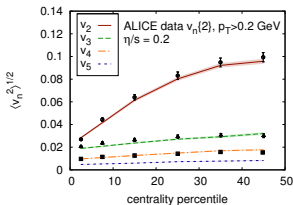
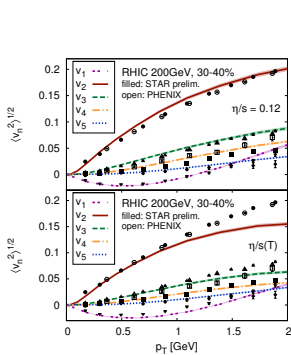
- Based on $[D_\mu, G^{\mu\nu}] = J^\nu$ with $J^\nu = \rho_A(x^-, \mathbf{x}_\perp)\delta^{+\nu} + \rho_B(x^+, \mathbf{x}_\perp)\delta^{-\nu}$
- $\rho_{A,B}(x^\pm, \mathbf{x}_\perp)$ distributions from IP-Sat
- Calculate the classical gluon field A^μ between the two receding disks (Glasma)

Initial Energy Density – IP-Glasma

Figures from [Schenke, Tribedy, Venugopalan, Phys.Rev.Lett.108.252301], [Gale, Jeon, Schenke, Int. J. of Mod. Phys. A, Vol. 28, 1340011 (2013)],



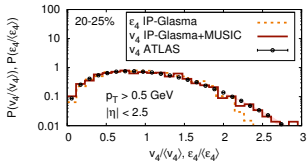
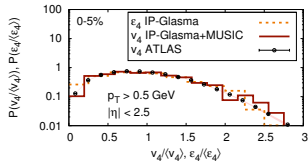
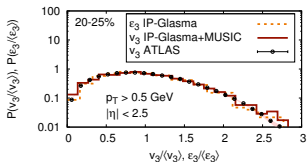
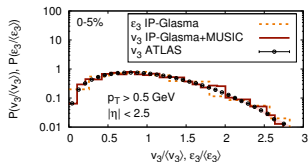
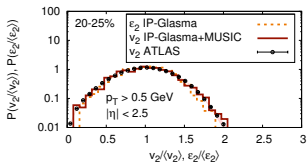
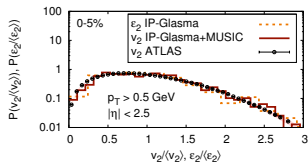
IP-Glasma RHIC and LHC:



Figures from [Gale, Jeon, Schenke, Tribedy and Venugopalan (GJSTV), Phys.Rev.Lett.110.012302]

E-by-E v_n distributions

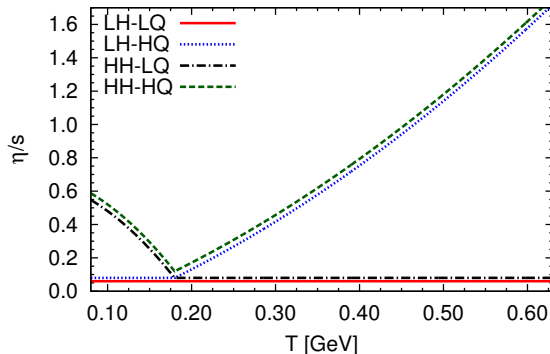
- E-by-E v_n is sensitive to E-by-E ϵ_n



Figures from [GJSTV, Nucl.Phys.A904-905 2013 (2013) 409c-412c]

Shear Viscosity

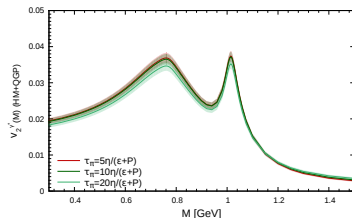
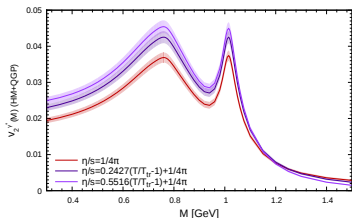
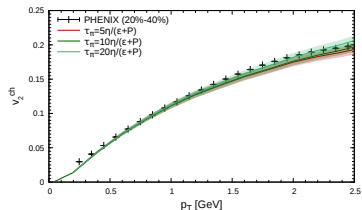
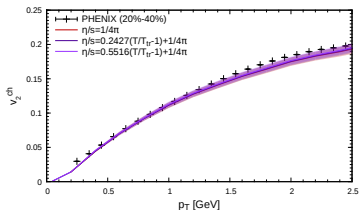
- RHIC: $\eta/s \approx 1/4\pi$
- LHC: $\eta/s \approx 2.5 \times 1/4\pi$
or
- Temperature dependent $(\eta/s)(T)$ (We used HH-HQ)



[Niemi, Denicol, Huovinen, Molnar, Rischke, Phys.Rev. C86 (2012) 014909]

Influence of $(\eta/s)(T)$ and τ_π

[Vujanovic, Paquet, Denicol, Luzum, Schenke, Jeon, Gale, arXiv:1408.1098, Proc. of QM14]



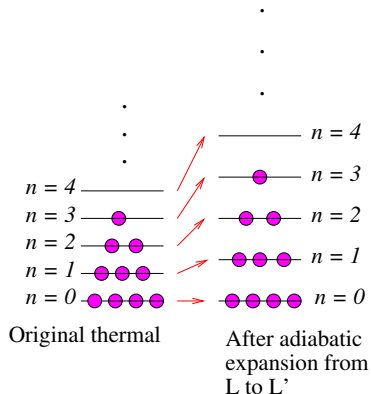
Hadronic and dilepton v_2 with differing $(\eta/s)(T)$

Hadronic and dilepton v_2 with differing τ_π

Bulk viscosity and ultra-central collisions

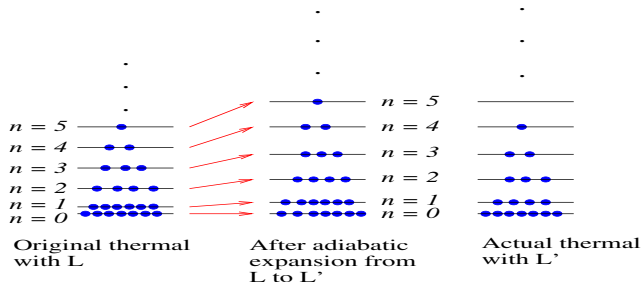
- QGP property to learn: ζ/s

Bulk viscosity in context



- Thermal occupation number in a box with the size L :
 $N_n(LT) = \exp(-2\pi n/LT)$ since $p_n = 2\pi n/L$.
- After an adiabatic expansion $L \rightarrow L'$, the original $\{N_n\}$ is still thermal with $T' = (L/L')T$
- Redistribution **not** necessary: $\zeta = 0$ for conformal theory

Bulk viscosity in context

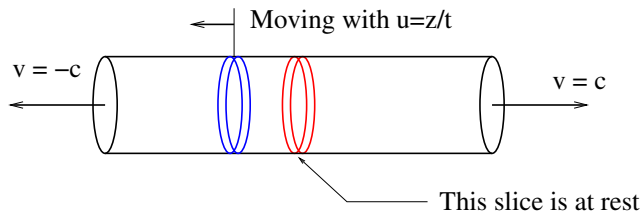


- Thermal occupation number in a box with the size L :

$$N_n(m/T, LT) = \exp\left(-\sqrt{m^2/T^2 + (2\pi n)^2/L^2 T^2}\right) \text{ since } \rho_n = 2\pi n/L.$$

- After an adiabatic expansion $L \rightarrow L'$, the original $\{N_n(m/T, LT)\}$ is **no longer** thermal: There is no choice of T' that will make the original $\{N_n\}$ look like another thermal distribution with L'
- Redistribution must occur: $\zeta \neq 0$ for non-conformal theory

Bulk viscosity in context



- Bjorken flow – Idealized, but valid near mid-rapidity.
- Hubble-like flow in the beam direction
- Expansion disrupts local equilibrium unless conformal
- Entropy must increase \implies Multiplicity must increase
- Longitudinal expansion slows down when $\zeta \neq 0$
- Pressure degrades

Bulk viscosity calculations

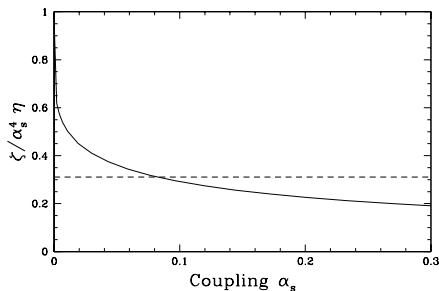
- Kubo formula

$$\zeta = \frac{1}{2} \lim_{\omega \rightarrow 0} \frac{1}{\omega} \int_{-\infty}^{\infty} dt e^{i\omega t} \int d^3x \langle [\hat{P}(x, t), \hat{P}(0, 0)] \rangle$$

Using energy conservation, can replace

$$\hat{P} \rightarrow \frac{1}{3} \hat{T}^{\mu}_{\mu} = \hat{P} - \frac{1}{3} \hat{\epsilon} \quad (T^{\mu}_{\mu} = 0 \text{ if conformal})$$

- Leading order pQCD [Arnold, Dugan, Moore, Phys.Rev.D74:085021,2006]



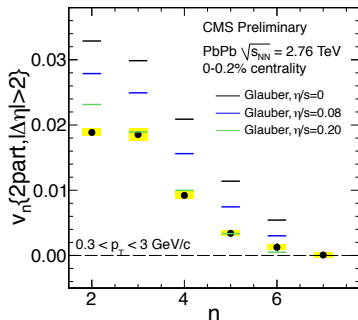
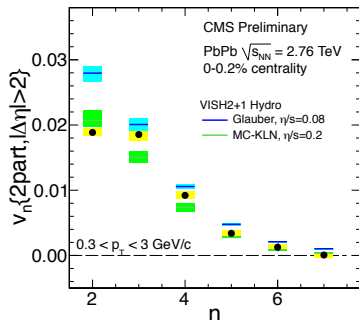
Dotted line: $\zeta = 15(1/3 - v_s^2)^2 \eta$
(Weinberg)

- Pure Glue: $\zeta \approx 50(1/3 - v_s^2)^2 \eta$
[Dusling, Schaefer, Phys.Rev. C85 (2012) 044909]

- Relaxation time approx:
 $\zeta \approx 70(1/3 - v_s^2)^2 \eta$
[Denicol, Jeon, Gale, arXiv:1403.0962]

- $v_s^2 = \frac{\partial P}{\partial \epsilon}$

MC-Glauber & MC-KLN



Figures from CMS-PAS HIN-12-011 and also [Harris, Kharzeev and Ullrich, 2012 CERN Courier 52 (no. 9) 17] and [Luzum and Ollitrault, Nuclear Physics A 904-905 (2013) pp. 377-380]

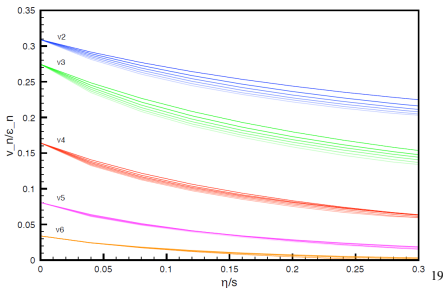
- Eccentricity $\epsilon_n \approx \epsilon_2$: Purely fluctuation driven
- The tension: Exp. $v_2 \approx v_3$ but Theo. $v_2 > v_3$
- Above calculations with $\zeta = 0$

Effect of bulk viscous pressure

$$\frac{\zeta}{s} = b \times \frac{\eta}{s} \left(\frac{1}{3} - c_s^2 \right)^2 \quad b=0, 15, 30, 45, 60, 75$$

MUSIC 2.0

0-1% - LHC

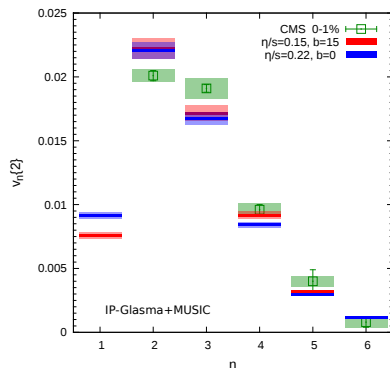
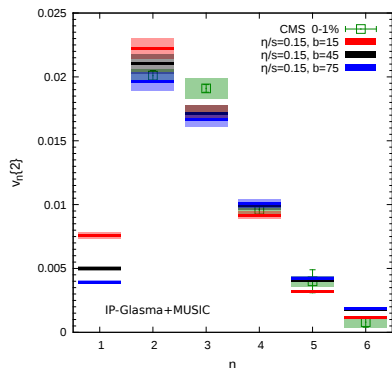


[Slide from G. Denicol's QM14 presentation]

- Higher harmonics \implies Less sensitive to ζ/s
- Ultra-central collision: $\epsilon_n \approx \epsilon_2$ – A good place for ζ/s to show up

The effect of non-zero bulk viscosity

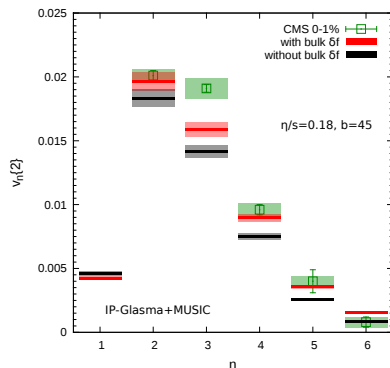
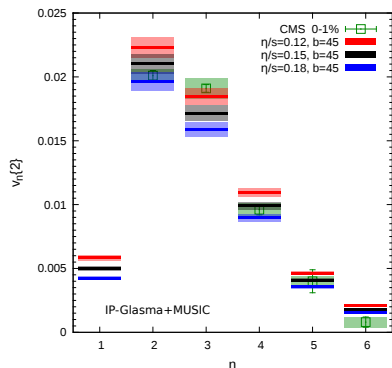
With $\zeta = b(1/3 - v_s^2)^2 \eta$



- Left: Effect of bulk viscosity with various values of b with η/s fixed
- Right: Trading bulk and shear

The effect of non-zero bulk viscosity

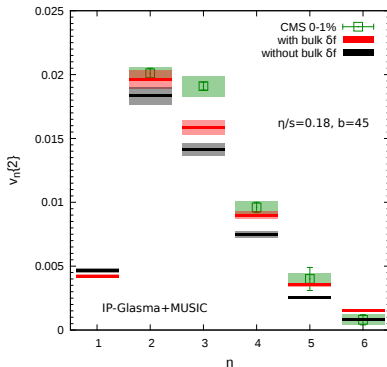
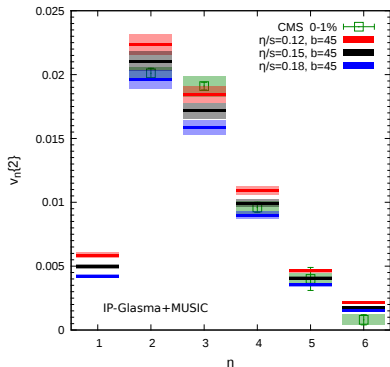
With $\zeta = b(1/3 - v_s^2)^2 \eta$



- Left: Effect of bulk viscosity with various values of η/s with b fixed
- Right: Effect of bulk corrections in δf .

The effect of non-zero bulk viscosity

With $\zeta = b(1/3 - v_s^2)^2 \eta$

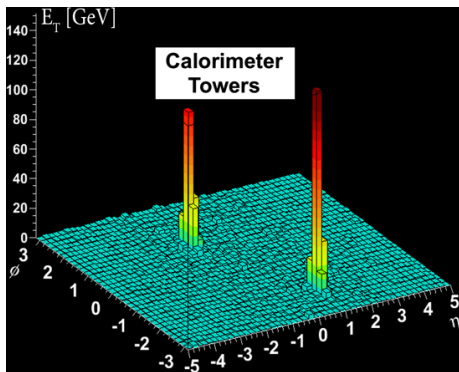


- Bulk viscosity improves the description
- Not perfect: Also need initial state nucleon-nucleon correlations
- Our best bet: $b \approx 45$

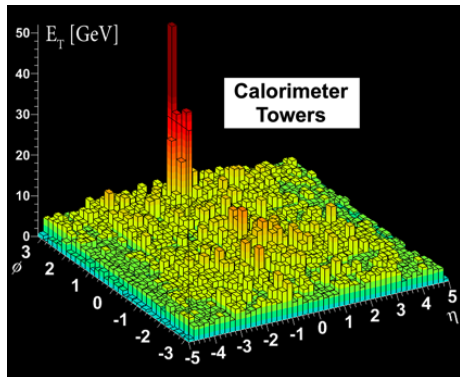
Jet Quenching Transport Coefficient \hat{q}

- QGP property to learn: Jet-medium interaction

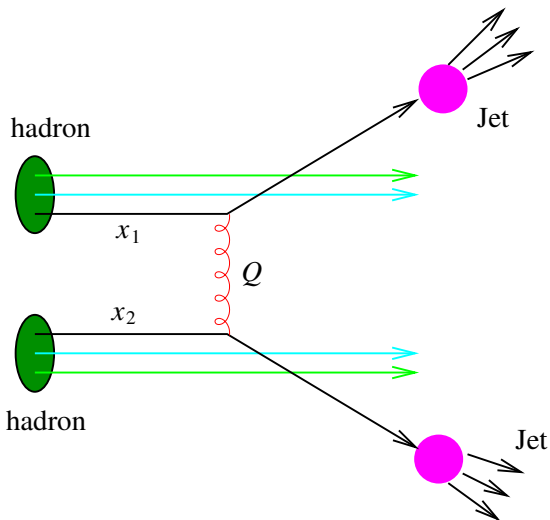
Jet Quenching



ATLAS: Two jets visible – Surface emission

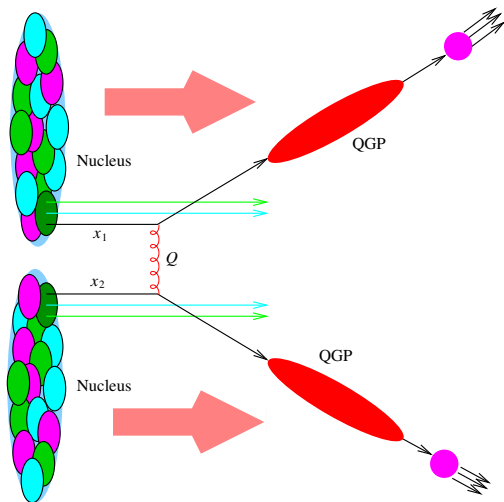


ATLAS: Mono jet. One jet escapes and its partner disappears in medium.



We understand this. More or less.

Schematic Understanding

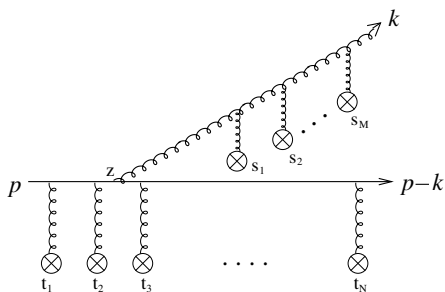


HIC Jet production scheme:

$$\begin{aligned} \frac{d\sigma_{AB}}{dt} &= \int_{\text{geometry}} \int_{abcd\mathbf{c}'} \\ &\times f_{a/A}(x_a, Q_f) f_{b/B}(x_b, Q_f) \\ &\times \frac{d\sigma_{ab \rightarrow cd}}{dt} \\ &\times \mathcal{P}(x_c \rightarrow x'_c | T, u^\mu) \\ &\times D(z'_c, Q) \end{aligned}$$

$\mathcal{P}(x_c \rightarrow x'_c | T, u^\mu)$: Medium modification of high energy parton property

Main radiative process in Energy Loss



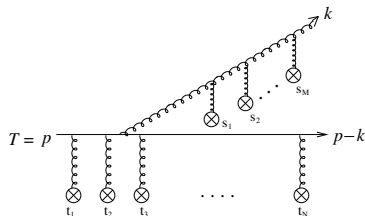
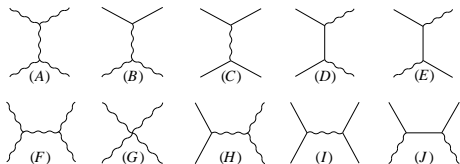
- Landau-Pomeranchuk-Migdal effect important - BDMPS-Z
- Relevant QGP property:

$$\hat{q} = \frac{\mu^2}{\lambda_{\text{mfp}}}$$

where μ is the average (spatial) momentum exchange and λ_{mfp} is the elastic collision mean free path.

Full leading order QCD calculation of energy loss rate

Relevant processes – The same as η/s calculation



Use kinetic theory to calculate the evolution of the jet

$$\frac{df}{dt} = \mathcal{C}_{2 \leftrightarrow 2} + \mathcal{C}_{1 \leftrightarrow 2}$$

Complication 1: $1 \leftrightarrow 2$ process needs resummation (LPM effect, AMY)

Complication 2: Needs to keep track of the daughters

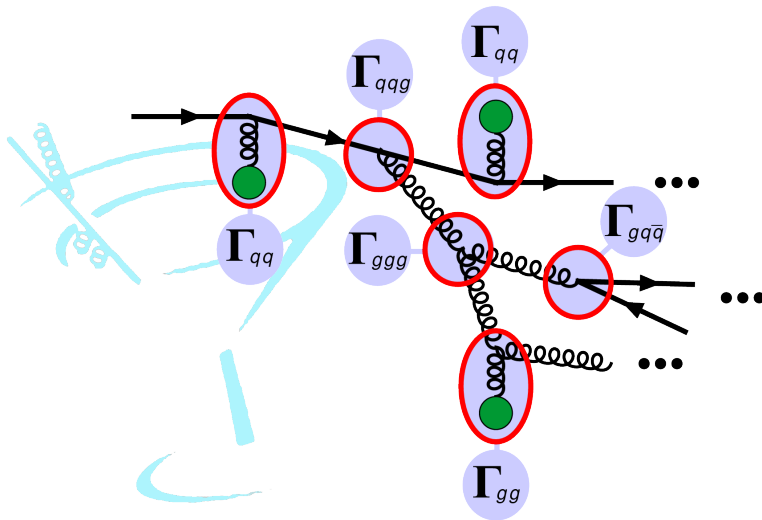
[c.f. E. Iancu's talk on Monday]

MARTINI



- Modular Algorithm for Relativistic Treatment of Heavy Ion Interactions
- Hybrid approach
 - Calculate Hydrodynamic evolution of the soft mode
 - Propagate jets in the evolving medium according to McGill-AMY

Parton propagation



An example path in MARTINI

Example – A MARTINI Run

MARTINI movie made by B. Schenke

Determining \hat{q} from Jet Quenching



- Miklos Gyulassy (Columbia University)
- Paul Romatschke (University of Colorado, Boulder)
- Steffen Bass, Berndt Mueller (Duke University)
- Michael Strickland (Kent State University)
- Xin-Nian Wang (co-spokesperson & project director) (LBNL)
- Ramona Vogt (LLNL)
- Ivan Vitev (LANL)
- Chales Gale, Sangyong Jeon (McGill University)
- Ulrich Heinz (co-spokesperson) (Ohio State University)
- Denes Molnar (Purdue University)
- Rainer Fries, Cheming Ko (Texas A & M University)
- Abhijit Majumder (Wayne State University)



Extracting jet transport coefficient from jet quenching at RHIC and LHC,

JET Collaboration, Burke et al. Phys.Rev. C90 (2014) 014909

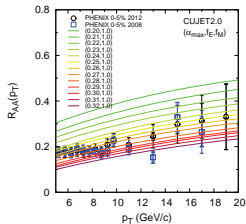
- Models
 - GLV-CUJET (Columbia)
 - HT-BW (Berkeley-Wuhan)
 - HT-M (Wayne State)
 - MARTINI (McGill)
 - McGill-AMY (McGill-Wuhan)

Determining \hat{q} from Jet Quenching

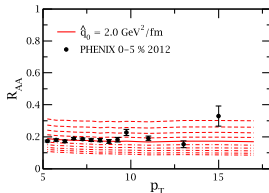
- All models based on many-body QCD
- Differences in dealing with the medium and multiple collisions
- All have adjustable parameters that can be related to \hat{q}

Determining \hat{q} from Jet Quenching

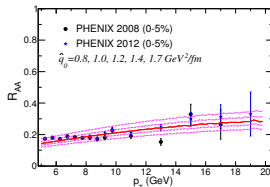
$$R_{AA} = \frac{dN_{AA}/dp_T dy}{N_{coll} dN_{pp}/dp_T dy} \text{ plots for RHIC}$$



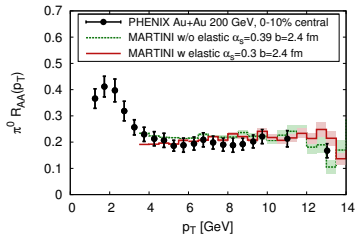
CUJET



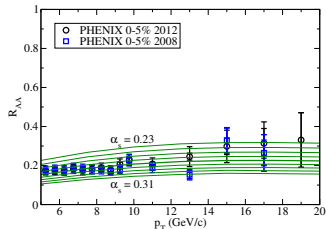
HT-M



HT-BW



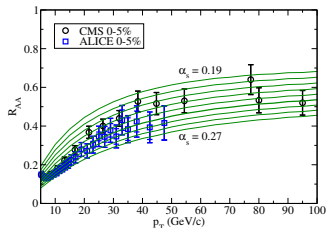
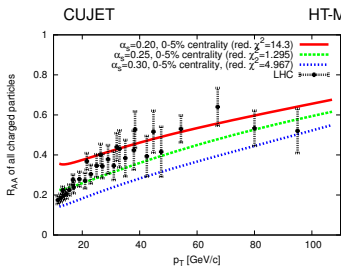
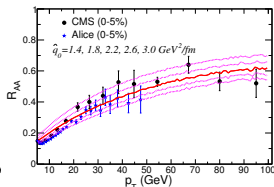
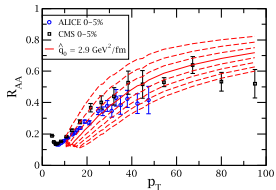
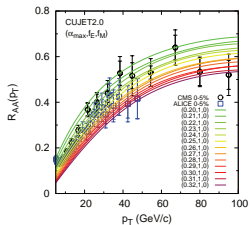
MARTINI



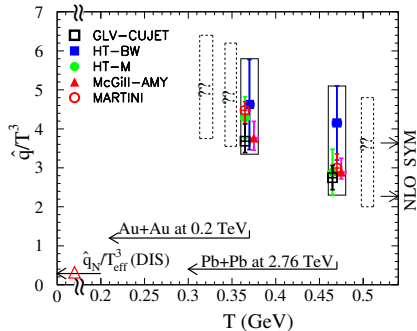
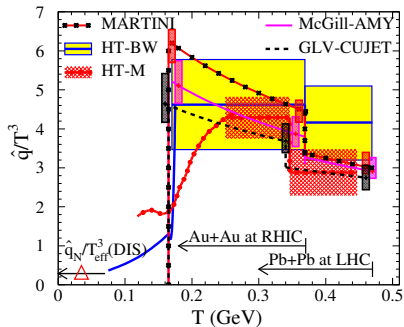
McGill-AMY

Determining \hat{q} from Jet Quenching

$$R_{AA} = \frac{dN_{AA}/dp_T dy}{N_{coll} dN_{pp}/dp_T dy} \text{ plots for the LHC}$$



Determining \hat{q} from Jet Quenching



- RHIC: $\hat{q}/T^3 \approx 3 - 6$
- LHC: $\hat{q}/T^3 \approx 2 - 5$

- High statistics data from RHIC and the LHC starting to permit precise evaluation of QGP properties such as $\eta/s, \zeta/s, \hat{q}$
- All parts of heavy ion collision calculations/simulations/event-generations (initial conditions, hydro evolution, hadronic afterburner, jet-medium interactions) are maturing
- Theory challenges:
 - Non-perturbative calculations of QGP properties with $\alpha_s \approx 0.3$ or $g \approx 2$
 - Thermalization mechanism ? [c.f. A. Kurkela's talk on Tuesday]
 - Hadronization mechanism ?

Hard Probes 2015

June 29 - July 03, 2015



McGill

Montréal

Backup Slides

Charge fluctuations

- QGP property to learn: Deconfined quarks and gluons?

$$\langle \delta Q^2 \rangle_{\text{th}} = q^2 \langle \delta N^2 \rangle_{\text{th}} \approx q^2 \langle N \rangle_{\text{th}}$$

- QGP: $q_u = 2/3$, $q_{d,s} = -1/3$, $q_g = 0$
- Hadrons: $q = 0, 1, 2, \dots$
- With $\langle N_{\text{parton}} \rangle \approx \langle N_{\text{hadron}} \rangle \sim (3/2) \langle N_{\text{ch}} \rangle$,

$$\frac{\langle \delta Q^2 \rangle_{\text{free QGP}}}{\langle N_{\text{ch}} \rangle} \sim 1/4$$

$$\frac{\langle \delta Q^2 \rangle_{\text{free } \pi}}{\langle N_{\text{ch}} \rangle} \sim 1$$

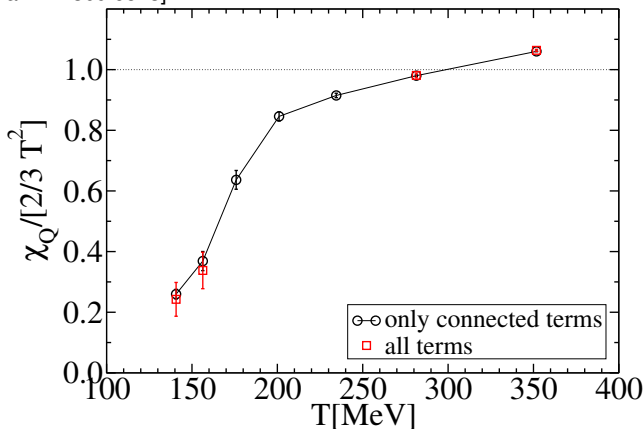
[Jeon & Koch, Phys.Rev.Lett. 85 (2000) 2076-2079,

Asakawa, Heinz & Muller, Phys.Rev.Lett. 85 (2000) 2072-2075

Similar conclusions from the balance function approach by S. Pratt]

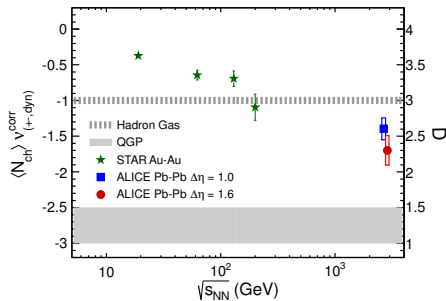
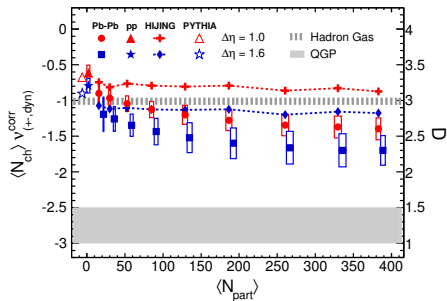
Charge susceptibility from LQCD

[Giudice et al. arXiv:1309.6523]



- $\langle \delta Q^2 \rangle = \chi_Q$
- In QGP phase, χ_Q is close to the free QGP limit

Observable – $D = 4 \langle \delta Q^2 \rangle / \langle N_{\text{ch}} \rangle$



[ALICE, Phys.Rev.Lett. 110 (2013) 15, 152301]

- Left: D as a function of centrality
- Right: D as a function of collision energy
- Conclusion: Smaller fluctuations seen but it seems to be contaminated by re-equilibration in the hadronic phase

Expected QGP Properties

Expected properties (3 light flavors)

- High number density

$$\begin{aligned}n &\approx (36 + 16) \int \frac{d^3 p}{(2\pi)^3} e^{-p/T} \approx 5 T^3 \\ &= 5 \left(\frac{T}{200 \text{ MeV}} \right)^3 \text{ fm}^{-3}\end{aligned}$$

- High energy density

$$\begin{aligned}\varepsilon &\approx (36 + 16) \int \frac{d^3 p}{(2\pi)^3} p e^{-p/T} \approx 15 T^4 \\ &= 3 \left(\frac{T}{200 \text{ MeV}} \right)^4 \text{ GeV/fm}^3\end{aligned}$$

- RHIC central collisions: $T \sim 390 \text{ MeV}$ or $\varepsilon \sim 40 - 45 \text{ GeV/fm}^3$
- LHC central collisions: $T \sim 470 \text{ MeV}$ or $\varepsilon \sim 90 - 95 \text{ GeV/fm}^3$

- High pressure (approximately conformal if coupling is weak)

$$P \approx \frac{1}{3} \epsilon$$

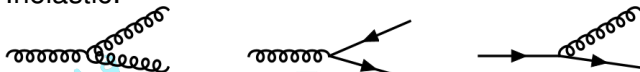
- Quarks carry fractional charges. About 30 % of entropy ($s = (\epsilon + P)/T$) is in the gluons. \implies Small charge/entropy ratio
- First order? Second order? Cross-over? – Critical point?

MARTINI Backups

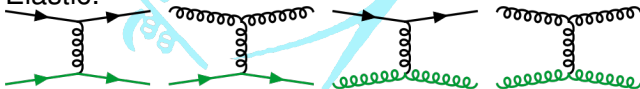
Parton propagation

Process include in MARTINI (all of them can be switched on & off):

- Inelastic:



- Elastic:



- Conversion:

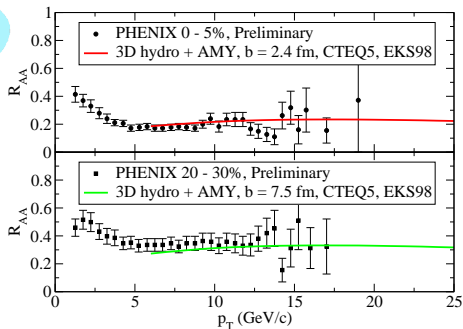
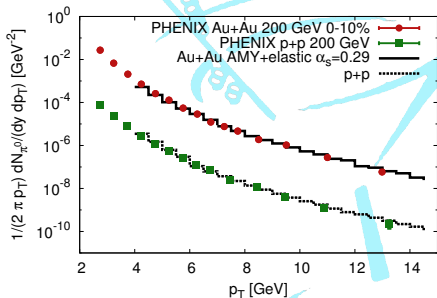


- Photon: emission & conversion

Pion production

[Schenke, Jeon and Gale, Phys. Rev. C 80, 054913 (2009)]

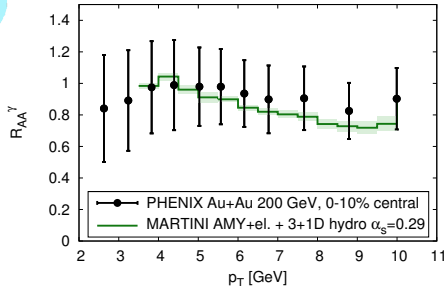
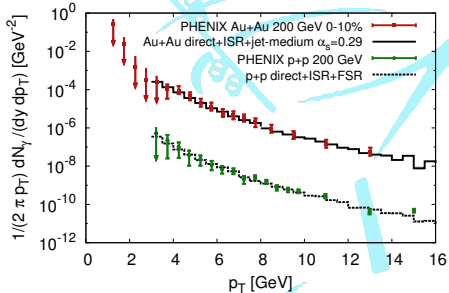
• π^0 spectra and R_{AA}



Photon production

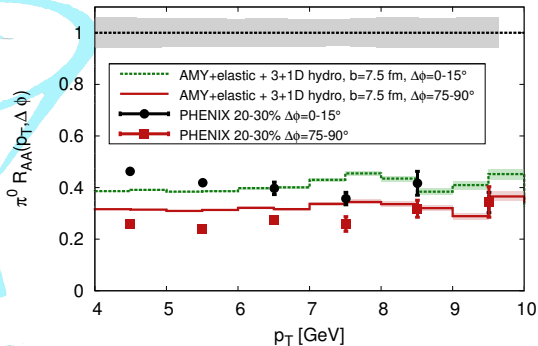
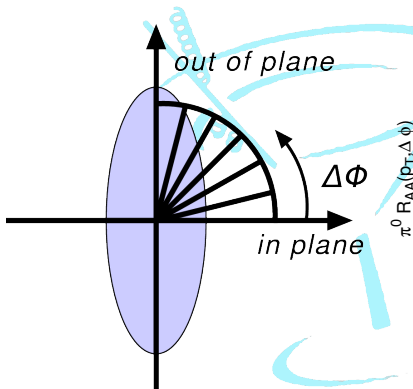
[Schenke, Jeon and Gale, Phys. Rev. C 80, 054913 (2009)]

● Spectra and R_{AA}^{γ}



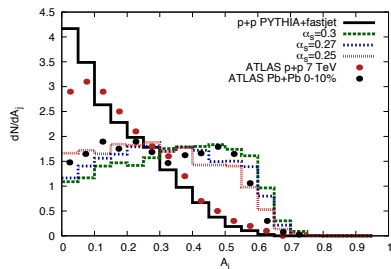
Azimuthal dependence of R_{AA}

- $R_{AA}(p_T, \Delta\phi)$

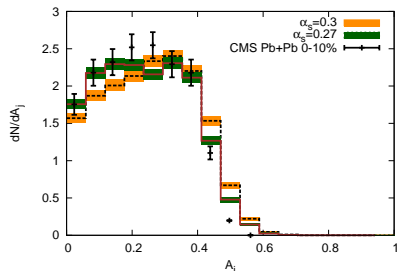


- $A = (E_t - E_a)/(E_t + E_a)$
- Right now ideal hydro with a smooth initial condition
- Full jet reconstruction with FASTJET

[Young, Schenke, Jeon, Gale, arXiv:1103.5769]



ATLAS, QM 2011



CMS, arXiv: 1102.1957 (2011)

Hydro Backups

Shear viscosity in context

- Energy-momentum conservation $\partial_\mu T^{\mu\nu} = 0$
- Local energy density and energy flow

$$T^{\mu\nu} u_\nu = -\varepsilon u^\mu$$

- Decomposition to non-dissipative and dissipative components

$$T^{\mu\nu} = \varepsilon u^\mu u^\nu + P(\varepsilon)(g^{\mu\nu} + u^\mu u^\nu) + \pi^{\mu\nu}$$

- Dissipative part ($\Delta^{\mu\nu} = g^{\mu\nu} + u^\mu u^\nu$, $\nabla^\mu = \Delta^{\mu\nu} \partial_\nu$)

$$\pi_{\text{NS}}^{\mu\nu} = -\eta \left(\nabla^\mu u^\nu + \nabla^\nu u^\mu - \frac{2}{3} \Delta^{\mu\nu} \nabla^\alpha u_\alpha \right) - \zeta \Delta^{\mu\nu} \nabla^\alpha u_\alpha$$

Minus signs dictated by increasing entropy

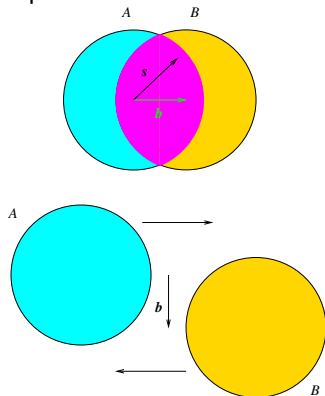
- Second order dissipative hydrodynamics

$$\Delta^{\mu\alpha} \Delta^{\nu\beta} u \cdot \partial \pi_{\alpha\beta} = -\frac{1}{\tau_\pi} \left(\pi^{\mu\nu} - 2\eta \nabla^{\langle\mu} u^{\mu\rangle} + \frac{4}{3} \tau_\pi \pi^{\mu\nu} (\partial_\alpha u^\alpha) \right)$$

[Baier, Romatschke, Son, Starinets, Stephanov JHEP 0804 (2008) 100]

Smooth (or Average) Initial Conditions

Optical Glauber

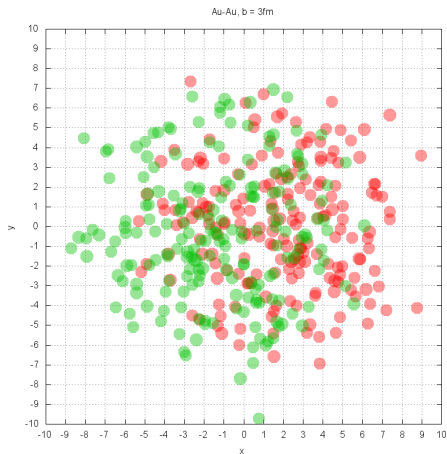


- Thickness functions: $T_{A,B}(\mathbf{s}) = \int_{-\infty}^{\infty} dz \rho_{A,B}(z, \mathbf{s})$
- Overlap function: $T_{AB}(\mathbf{b}) = \int d^2s T_A(\mathbf{s}) T_B(\mathbf{s} - \mathbf{b})$
- Number of collisions: $N_{\text{coll}}(\mathbf{b}) = AB T_{AB}(\mathbf{b}) \sigma_{\text{inel}}^{\text{NN}}$
- Number of participants:

$$N_{\text{part}}(\mathbf{b}) = A \int d^2s T_A(\mathbf{s}) \left(1 - [1 - T_B(\mathbf{s} - \mathbf{b}) \sigma_{\text{inel}}^{\text{NN}}]^B \right) + B \int d^2s T_B(\mathbf{s} - \mathbf{b}) \left(1 - [1 - T_A(\mathbf{s}) \sigma_{\text{inel}}^{\text{NN}}]^A \right)$$

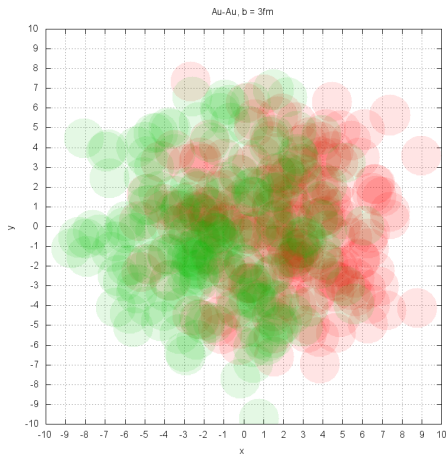
- Smooth distribution \implies No fluctuations.
No ϵ_{odd} nor V_{odd}

Fluctuating Initial Conditions



- 1 Sample nucleon positions

Fluctuating Initial Conditions



- 1 Sample nucleon positions
- 2 Decide what collides with what
- 3 Decide how much energy is deposited
- 4 Calculate the energy density in the hydro cells

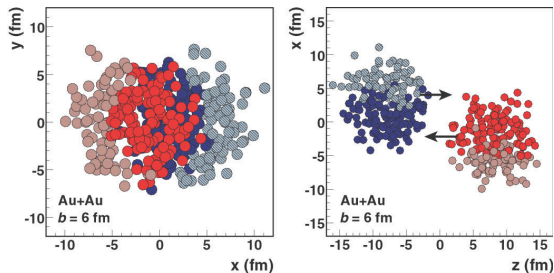
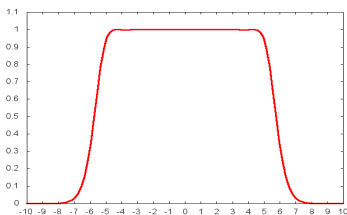
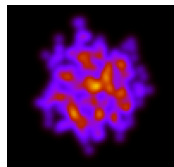


Figure from [Miller, Reygers, Sanders and Steinberg, Ann. Rev. Nucl. Part. Sci. 57 (2007) 205-243]

- Nucleon positions sampled from $T_{A,B}(\mathbf{s})$
- Collision if $d \leq \sqrt{\sigma_{\text{inel}}^{NN}/\pi}$



Profile in η



Transverse profile

- Energy density in a hydro cell positioned at \mathbf{x}_{\perp}^K

$$\epsilon_{\perp}(\mathbf{x}_{\perp}^K) = \sum_{J=1}^{N_w} \frac{\epsilon_0}{2\pi\sigma^2} e^{-(\mathbf{x}_{\perp}^K - \mathbf{x}_{\perp}^J)^2 / 2\sigma^2}$$

- The sum is over all wounded nucleons (usually truncated) with $\sigma \sim \tau_0$.
- η profile: Smooth. An Wood-Saxon or a stretch-Gaussian

QCD Estimates of η/s

- Danielewicz and Gyulassy [PRD **31**, 53 (1985)]:

- η/s bound from the kinetic theory:

$$\frac{\eta}{s} \gtrsim \frac{1}{12} \approx 0.08$$

- QCD estimate in the small α_S limit with $N_f = 2$ and $2 \rightarrow 2$ only (min. at $\alpha_S = 0.6$):

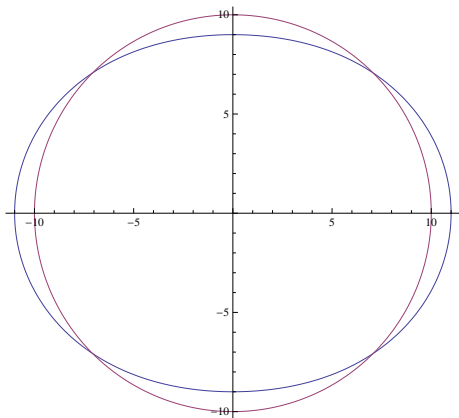
$$\eta \approx \frac{T}{\sigma_\eta} \approx \frac{0.57 T^3}{\alpha_S^2 \ln(1/\alpha_S)} \gtrsim 0.2s$$

- Baym, Monien, Pethick and Ravenhall [PRL **64**, 1867 (1990)]

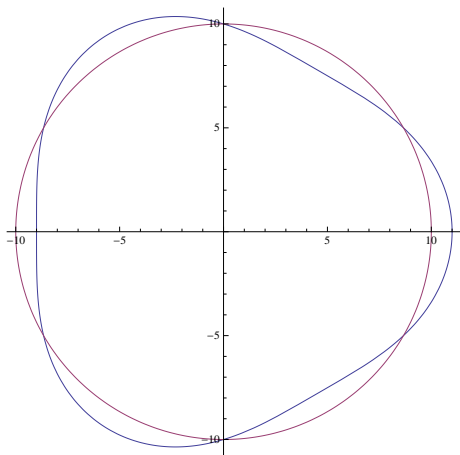
$$\eta \approx \frac{1.16 T^3}{\alpha_S^2 \ln(1/\alpha_S)} \gtrsim 0.4s \approx (5/4\pi)s$$

- M. Thoma [PLB **269**, 144 (1991)]

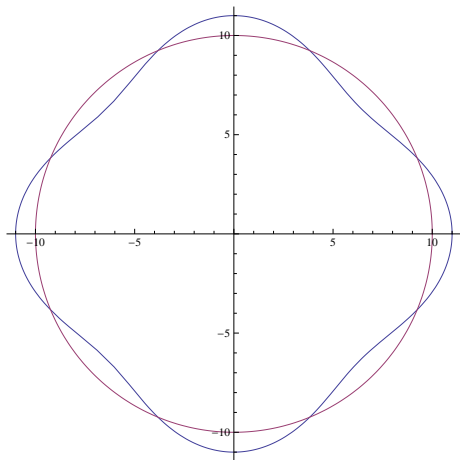
$$\eta \approx \frac{1.02 T^3}{\alpha_S^2 \ln(1/\alpha_S)} \gtrsim 0.4s \approx (5/4\pi)s$$



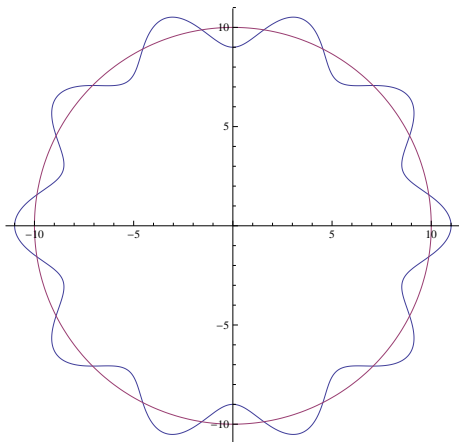
This causes elliptic flow. It is harder to destroy this than



this (v_3) ...



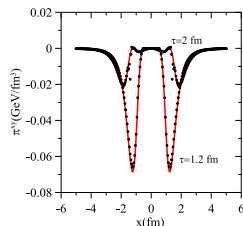
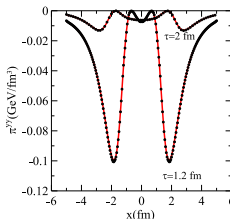
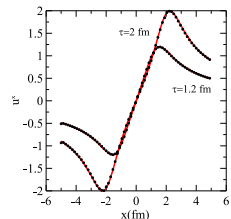
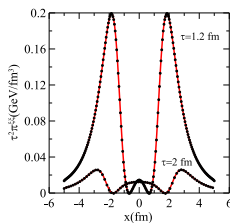
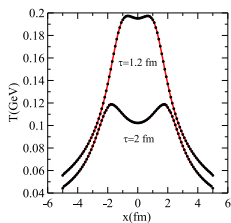
or this (v_4) ...



or this (v_{10}) ...

Confidence in viscous-hydro solutions

- Comparison with the viscous-Gubser flow and MUSIC
- Analytic ideal-hydro solution by [Gubser, Phys. Rev. D 82, 085027 (2010), Gubser and A. Yarom, Nucl. Phys. B 846, 469 (2011)]
- Semi-analytic Viscous-hydro solution by [Marrochio, Noronha, Denicol, Luzum, Jeon and Gale, arXiv:1307.6130].

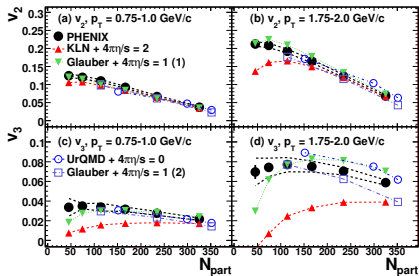
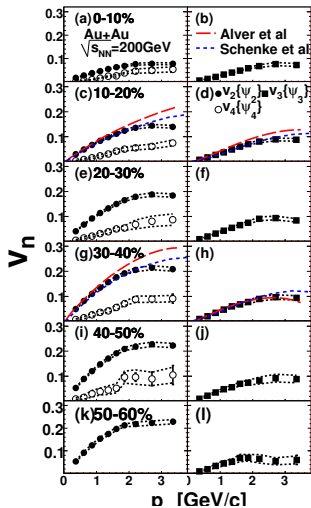


Figures from MNDLJG with

$$\eta/s = 0.2, \tau_R T = 5\eta/s.$$

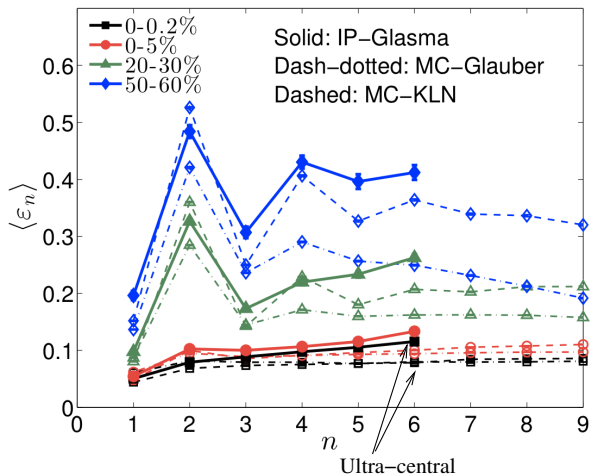
E-by-E MUSIC vs RHIC Data

Phenix (arXiv:1105.3928)



v_3 : Our predictions (Green Triangles)

Eccentricity distributions



In UC collisions: ϵ_n are almost purely fluctuation driven.

Figure from [Heinz and Snelling, Annu. Rev. Nucl. Part. Sci. 63 (2013) 123-151]

How Central is Ultra-Central?

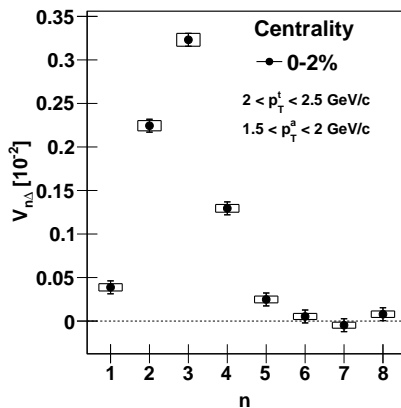
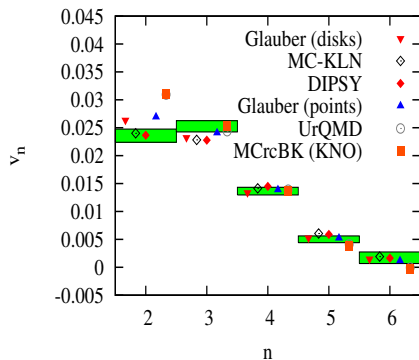


Figure from [ALICE, Phys.Lett. B708 (2012) 249-264]



[Green bars: 0 – 1 % ATLAS]. Figure from [Luzum and Ollitrault, Nuclear Physics A 904-905 (2013) pp. 377-380]

- As Centrality $\rightarrow 0$, $v_3 \rightarrow v_2$
- At 0 – 1 %, $v_3 \approx v_2$