

# Heavy Quarkonium magnetic dipole transitions in pNRQCD

Mainly based on Phys. Rev. D87, 074024. A. Pineda and J. Segovia

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## Motivation

- ▶ 1st principle computation of heavy quarkonium properties from QCD. To give model independent predictions with model independent errors.
- ▶ Determination of Standard Model parameters:  $m_Q$ ,  $\alpha_s$ , ...

Potential Non-Relativistic QCD in the weak coupling regime is ideal for this.  
 $m \gg mv \gg mv^2$

$$\left( i\partial_0 - \frac{\mathbf{p}^2}{2m} - V_s^{(0)}(r) \right) \Phi(\mathbf{r}) = 0$$

+ corrections to the potential  
 + interaction with other low  
 energy degrees of freedom

$\left. \right\} \text{pNRQCD(Pineda, Soto)} \quad E \sim mv^2$

In the strict weak coupling regime the starting point is

$$V_s^{(0)} \simeq V^c \equiv -C_f \frac{\alpha_s(\mu)}{r}.$$

Define the leading order:  $\alpha_V \sim v$

$$V_s^{(0)} = -C_f \frac{\alpha_V(1/r)}{r} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} + a_2 \frac{\alpha_s^2(\mu)}{16\pi^2} + \dots \right)$$

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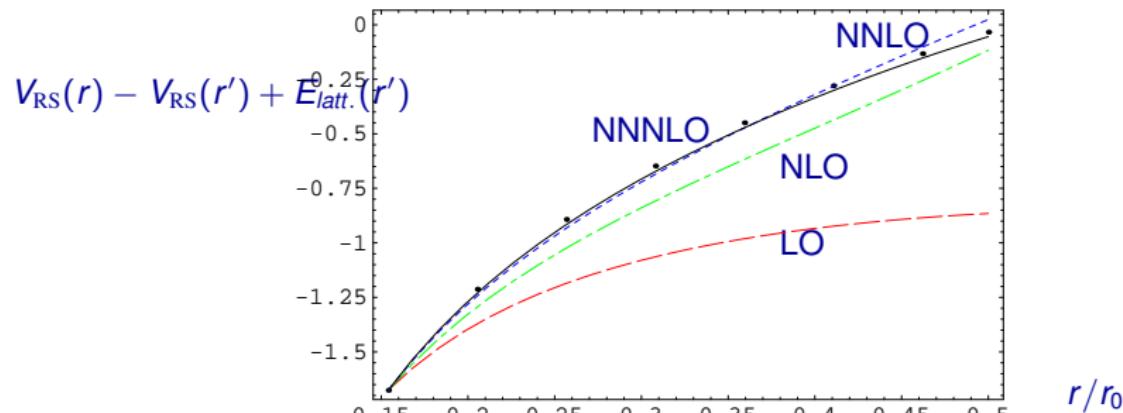
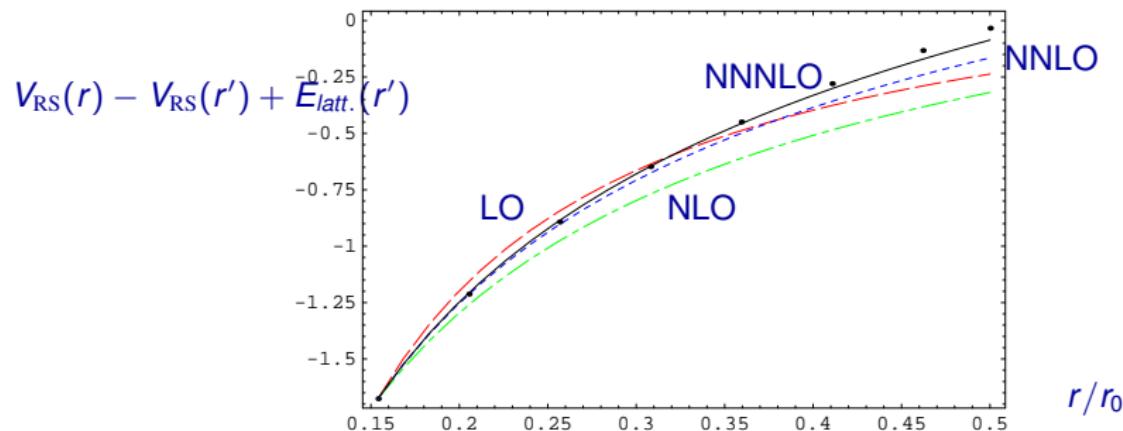
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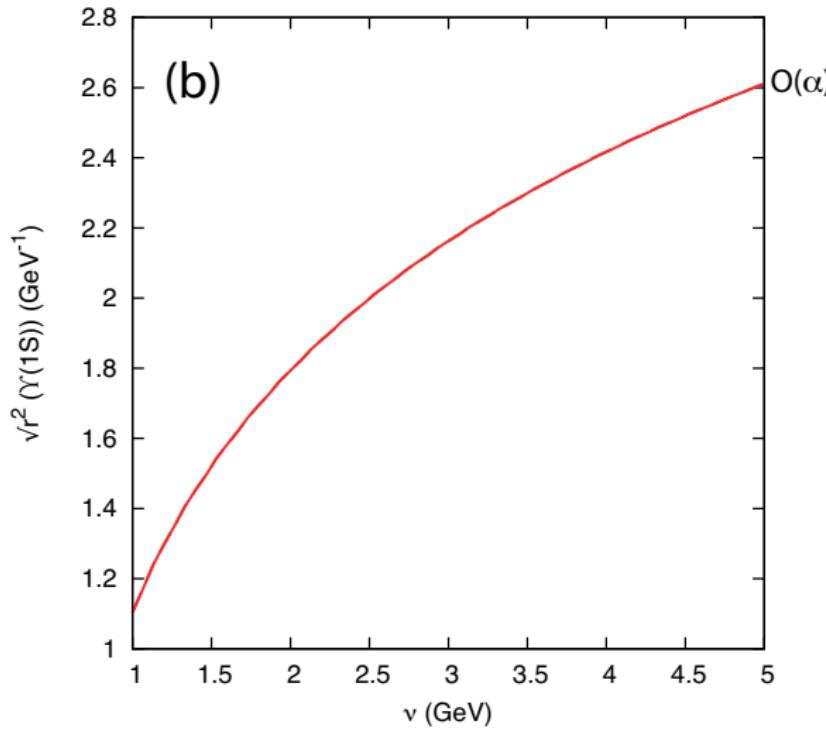
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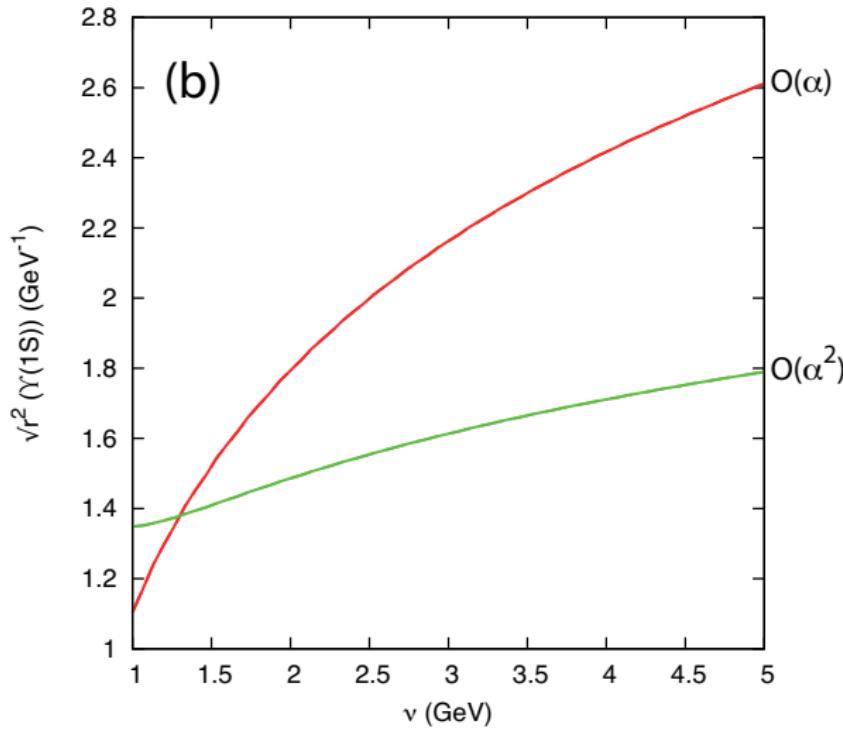


# Bottomonium



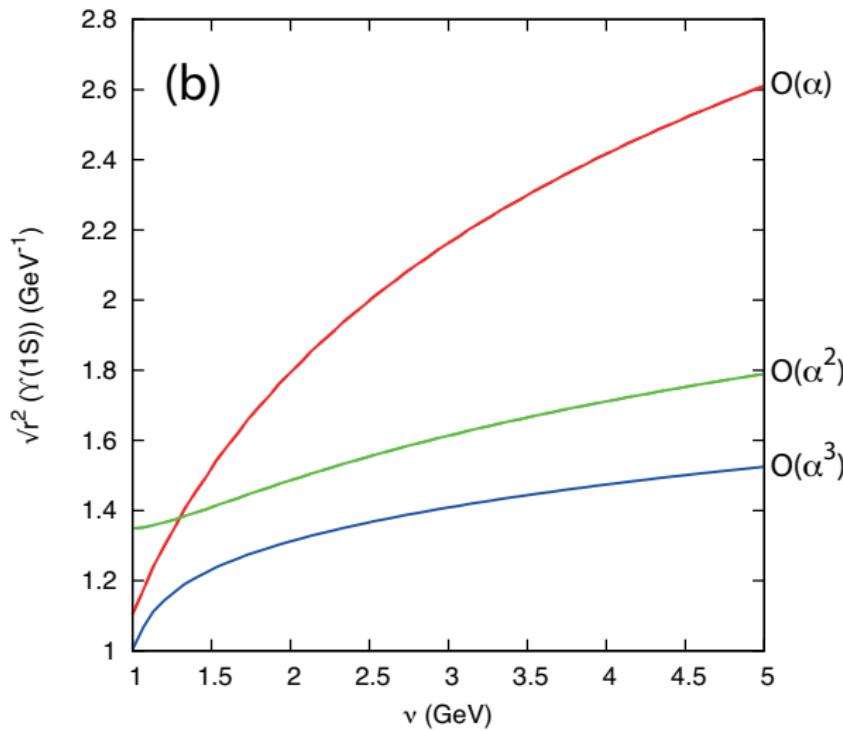
$$V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r}$$

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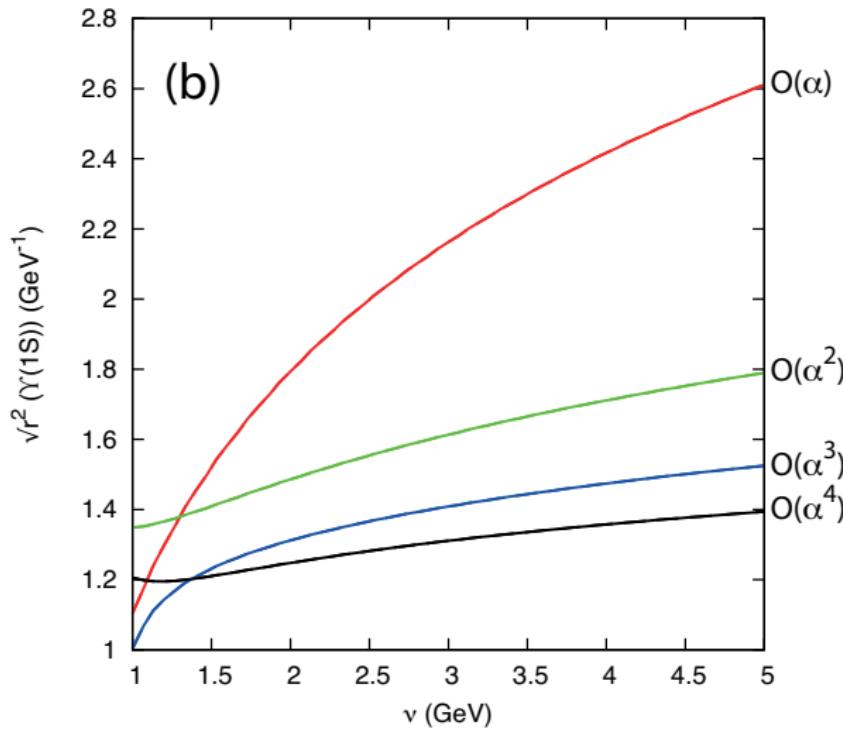
$$V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} \right)$$

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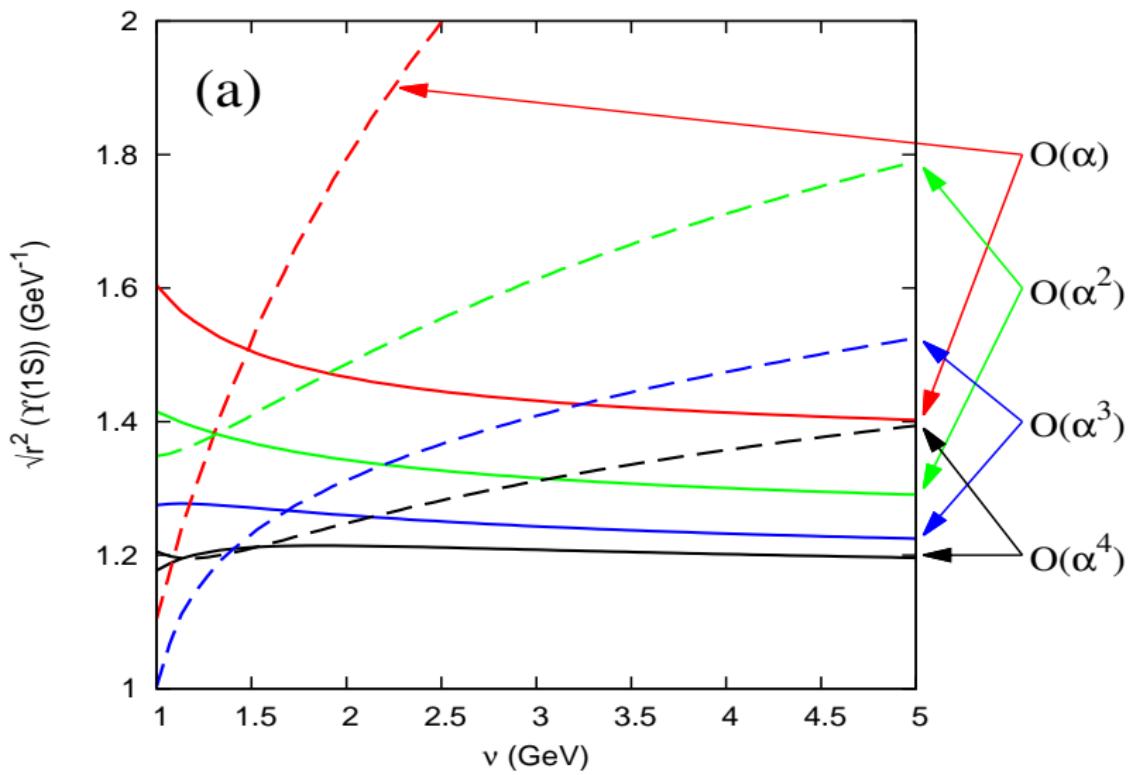


$$V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} + a_2 \frac{\alpha_s^2(\mu)}{(4\pi)^2} \right)$$

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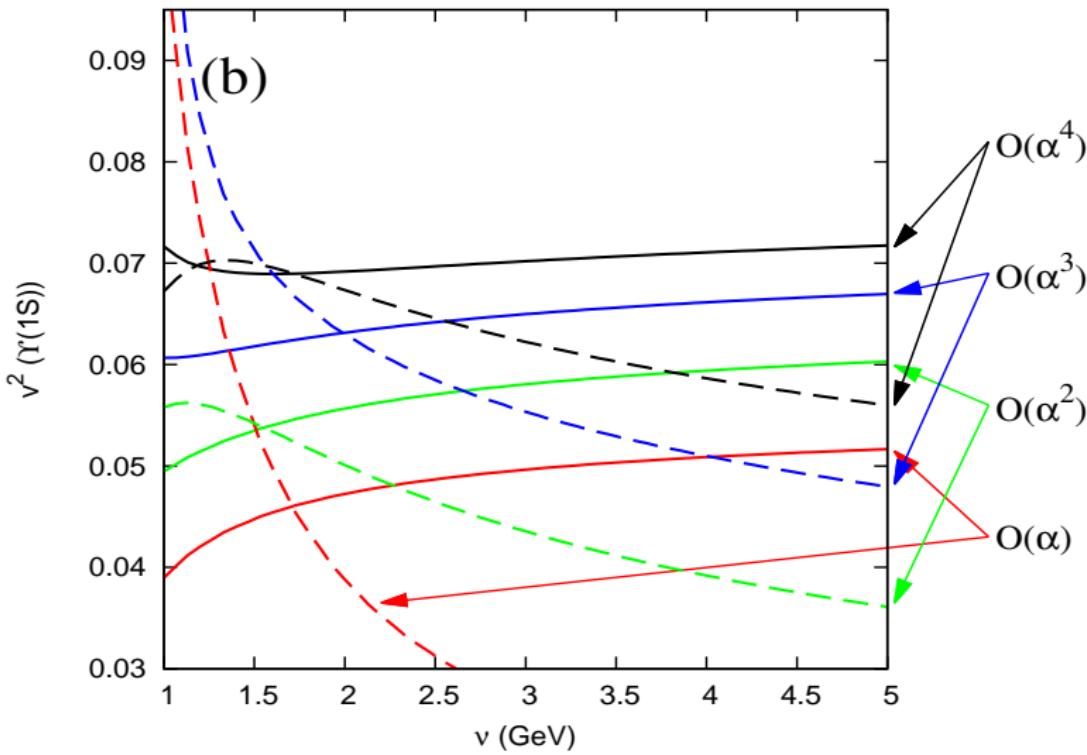


Figure:  $v_{10}^2 = \langle p^2 \rangle / m^2$  using the static potential  $V_{RS'}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$ . Dashed lines with  $\nu_r = \infty$ . Continuous lines with  $\nu_r = 0.7$  GeV. In both cases  $\nu_f = 0.7$  GeV.

# Charmonium

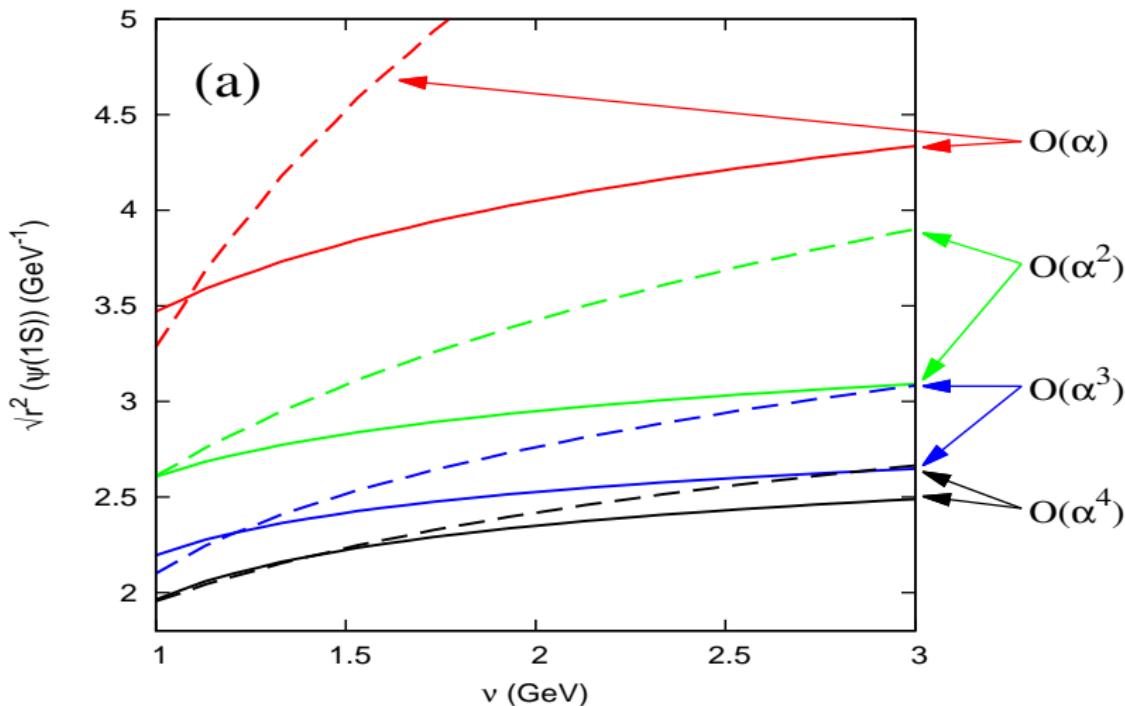


Figure:  $\sqrt{\langle r^2 \rangle_{10}}$  using the static potential  $V_{\text{RS'}}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$ . Dashed lines with  $\nu_r = \infty$ . Continuous lines with  $\nu_r = 0.7 \text{ GeV}$ . In both cases  $\nu_f = 0.7 \text{ GeV}$ .

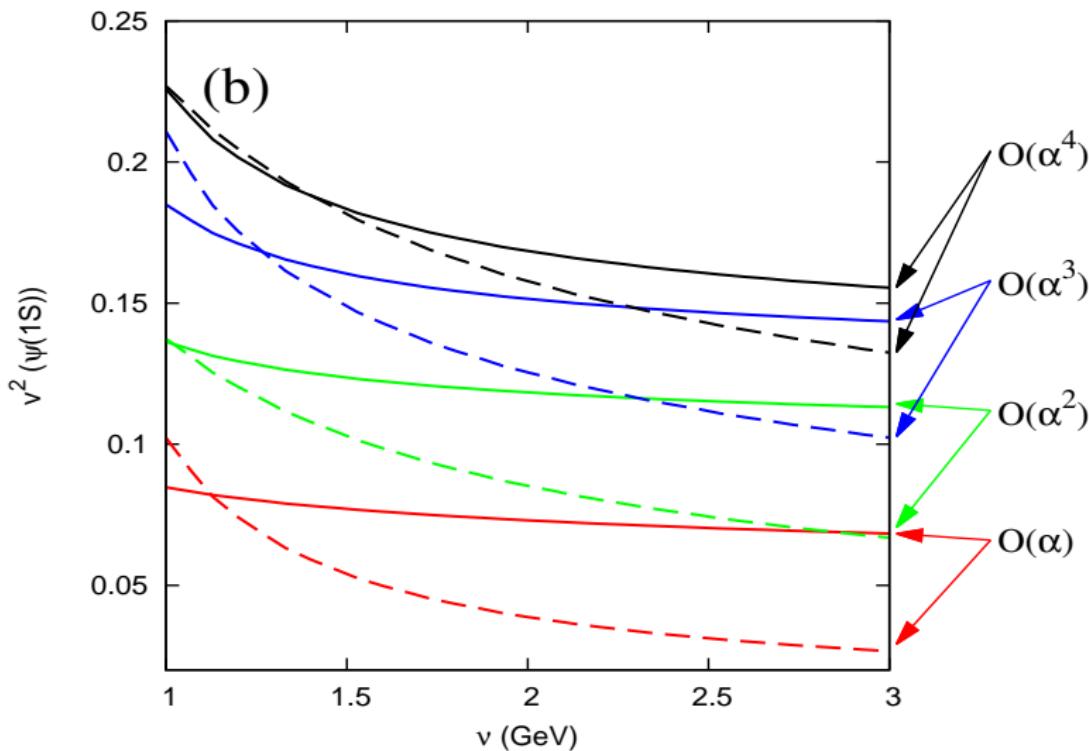


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	$bb(1S)$	$c\bar{c}(1S)$	$b\bar{b}(1P)$	$b\bar{b}(2S)$
$v$	0.26	0.43	0.25	0.24
$\sqrt{\langle r^2 \rangle} (\text{GeV}^{-1})$	1.2	2.2	2.1	2.9

Table: Estimates for  $v \equiv \sqrt{\langle p^2 \rangle / m^2}$  and  $\sqrt{\langle r^2 \rangle}$  for the heavy quarkonium states. For the  $b\bar{b}(2S)$  state the number we give for  $v$  is quite uncertain.

# Proposal: Reorganization of perturbation theory

Strict perturbation theory

$$H^{(0)} = \frac{\mathbf{p}^2}{m} + V^C \longrightarrow E_n^C, \quad \phi_n^C(\mathbf{r})$$

Coulomb potential

$$V_s^C = -C_f \frac{\alpha_s(\mu)}{r}$$

Relativistic +  $\delta V^{(0)}$  corrections:

$$\Delta H = \delta V^{(0)} + \frac{V^{(1)}}{m} + \frac{V^{(2)}}{m^2} + \dots$$

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Improved perturbation theory "acceleration of perturbation theory"

$$H^{(0)} = \frac{\mathbf{p}^2}{m} + V^{(0)} \longrightarrow E_n^{(0)}, \quad \phi_n^{(0)}(\mathbf{r})$$

Keep the static potential exactly

$$V_s^{(0)} = -C_f \frac{\alpha_s(1/r)}{r} \left( 1 + a_1 \frac{\alpha_s(1/r)}{4\pi} + a_2 \frac{\alpha_s^2(1/r)}{16\pi^2} + \dots \right)$$

Relativistic corrections:

$$\Delta H = \frac{V^{(1)}}{m} + \frac{V^{(2)}}{m^2} + \dots$$

Applied to inclusive electromagnetic decay ratios (Kiyo, Pineda, Signer).

Apply it to M1 radiative transitions.

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## Allowed transitions

Strict weak coupling: Brambilla, Jia, Vairo

$$\Gamma(n^3S_1 \rightarrow n^1S_0\gamma) = \frac{4}{3}\alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ 1 + 2\kappa - \frac{5}{3} \frac{\langle p^2 \rangle_n^C}{m^2} \right],$$

$$\Gamma(n^3P_J \rightarrow n^1P_1\gamma) = \frac{3\Gamma(n^1P_1 \rightarrow n^3P_J\gamma)}{2J+1} = \frac{4}{3}\alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ 1 + 2\kappa - d_J \frac{\langle p^2 \rangle_{n1}^C}{m^2} \right],$$

where  $d_0 = 1$ ,  $d_1 = 2$ ,  $d_2 = 8/5$ ,

$$k_\gamma = |\vec{k}| = \frac{M_H^2 - M_{H'}^2}{2M_H},$$

and the anomalous magnetic moment of the heavy quark reads

$$\kappa = \kappa^{(1)} \alpha_s(m)$$

Leading order= potential models

## Allowed transitions

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$$\kappa = \kappa^{(1)}\alpha_s(m) + \kappa^{(2)}\alpha_s^2(m)$$

We also implement the renormalon.

$$\Gamma_{\gamma(1S) \rightarrow \eta_b(1S)\gamma}$$

	LO	$\mathcal{O}(\alpha_s)$	$\mathcal{O}(\alpha_s^2)$	$\mathcal{O}(v^2)$	$\alpha_s \times \mathcal{O}(\alpha_s^2)$	$v \times \mathcal{O}(v^2)$
$\delta\Gamma$ (eV)	14.87	1.29	0.73	-1.71	0.15	-0.45

Table: *The leading and subleading contributions to  $\Gamma_{\gamma(1S) \rightarrow \eta_b(1S)\gamma}$ . The last two numbers are error estimates obtained by multiplying the subleading  $\mathcal{O}(\alpha_s^2)$  contribution by  $\alpha_s$  and the subleading  $\mathcal{O}(v^2)$  contribution by  $v$ .*

$$\Gamma_{\gamma(1S) \rightarrow \eta_b(1S)\gamma} = 15.18 \pm 0.45 (\mathcal{O}(v^3))^{-0.12}_{-0.05} (N_m)^{-0.04}_{+0.03} (\alpha_s)^{-0.20}_{+0.20} (m_{\overline{\text{MS}}}) \text{ eV ,}$$

which after combining the errors in quadrature reads

$$\Gamma_{\gamma(1S) \rightarrow \eta_b(1S)\gamma} = 15.18(51) \text{ eV .}$$

$$\Gamma_{J/\psi(1S) \rightarrow \eta_c(1S)\gamma}$$

	LO	$\mathcal{O}(\alpha_s)$	$\mathcal{O}(\alpha_s^2)$	$\mathcal{O}(v^2)$	$\alpha_s \times \mathcal{O}(\alpha_s^2)$	$v \times \mathcal{O}(v^2)$
$\delta\Gamma$ (keV)	2.34	0.33	0.16	-0.71	0.05	-0.30

Table: *The leading and subleading contributions to  $\Gamma_{J/\psi(1S) \rightarrow \eta_c(1S)\gamma}$ . The last two numbers are error estimates obtained by multiplying the subleading  $\mathcal{O}(\alpha_s^2)$  contribution by  $\alpha_s$  and the subleading  $\mathcal{O}(v^2)$  contribution by  $v$ .*

$$\Gamma_{J/\psi(1S) \rightarrow \eta_c(1S)\gamma} = 2.12 \pm 0.30 (\mathcal{O}(v^3))^{+0.21}_{-0.23} (N_m)^{-0.02}_{+0.02} (\alpha_s)^{-0.10}_{+0.11} (m_{\overline{\text{MS}}}) \text{ keV ,}$$

which, after combining the errors in quadrature, reads

$$\Gamma_{J/\psi(1S) \rightarrow \eta_c(1S)\gamma} = 2.12(40) \text{ keV .}$$

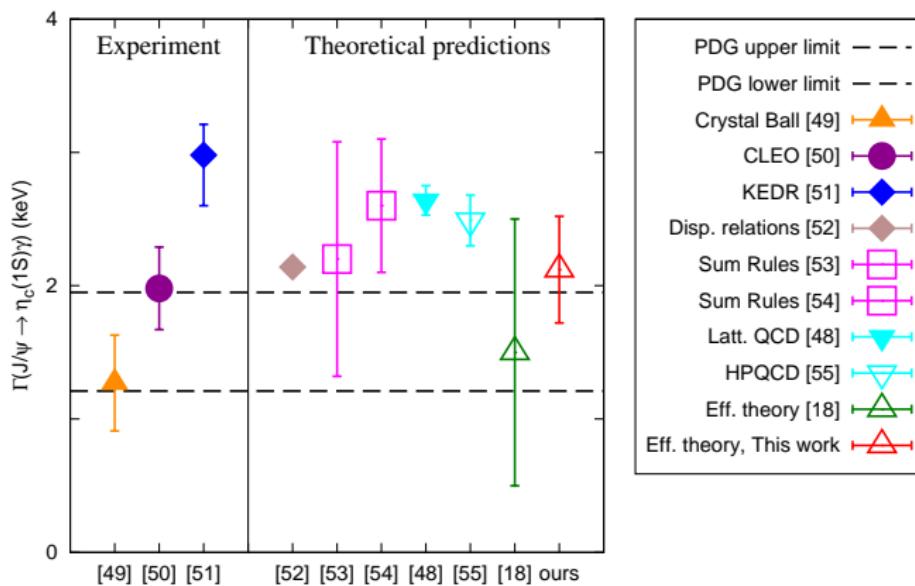


Figure: *Comparison of different theoretical and experimental predictions for  $\Gamma_{J/\psi \rightarrow \eta_c\gamma}$ .*

Hindered transitions. Strict weak coupling: Brambilla, Jia, Vairo

$$\Gamma(n^3S_1 \rightarrow n'{}^1S_0\gamma) \stackrel{n \neq n'}{=} \frac{4}{3} \alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ \frac{k_\gamma^2}{24} {}_{n'} \langle r^2 \rangle_n^C + \frac{5}{6} \frac{{}_{n'} \langle p^2 \rangle_n^C}{m^2} - \frac{2}{m^2} \frac{{}_{n'} \langle V_{S^2}(\vec{r}) \rangle_n}{E_n^C - E_{n'}^C} \right]^2,$$

$$\Gamma(n^1S_0 \rightarrow n'{}^3S_1\gamma) \stackrel{n \neq n'}{=} 4\alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ \frac{k_\gamma^2}{24} {}_{n'} \langle r^2 \rangle_n^C + \frac{5}{6} \frac{{}_{n'} \langle p^2 \rangle_n^C}{m^2} + \frac{2}{m^2} \frac{{}_{n'} \langle V_{S^2}(\vec{r}) \rangle_n^C}{E_n^C - E_{n'}^C} \right]^2,$$

$$V_{S^2}(\vec{r}) = \frac{4}{3} \pi C_f D_{S^2,s}^{(2)}(\nu) \delta^{(3)}(\vec{r})$$

$$D_{S^2,s}^{(2)}(\nu) = \alpha_s(\nu)$$

Hindered transitions. Improved weak coupling: Pineda, Segovia

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$$V_{S^2}(\vec{r}) = \frac{4}{3} \pi C_f D_{S^2,s}^{(2)}(\nu) \delta^{(3)}(\vec{r})$$

$$D_{S^2,s}^{(2)}(\nu) = \alpha_s(\nu) C_F^2(\nu) - \frac{3}{2\pi C_f} (d_{sv}(\nu) + C_f d_{vv}(\nu))$$

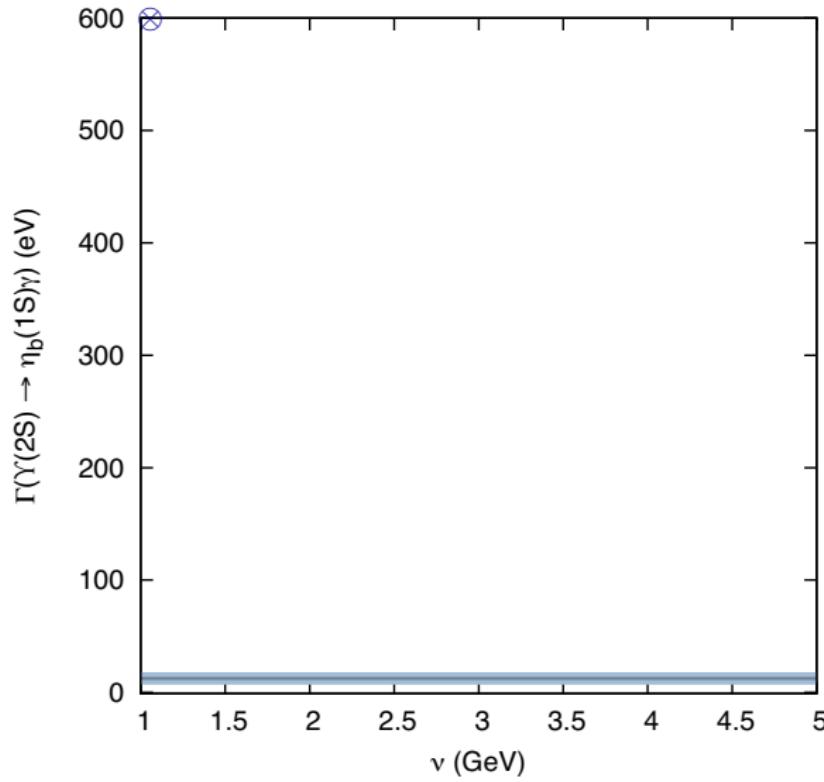
depends on the NRQCD Wilson coefficients. With LL accuracy they read

$$C_F(\nu) = z^{-C_A}, \quad d_{sv}(\nu) = d_{sv}(m),$$

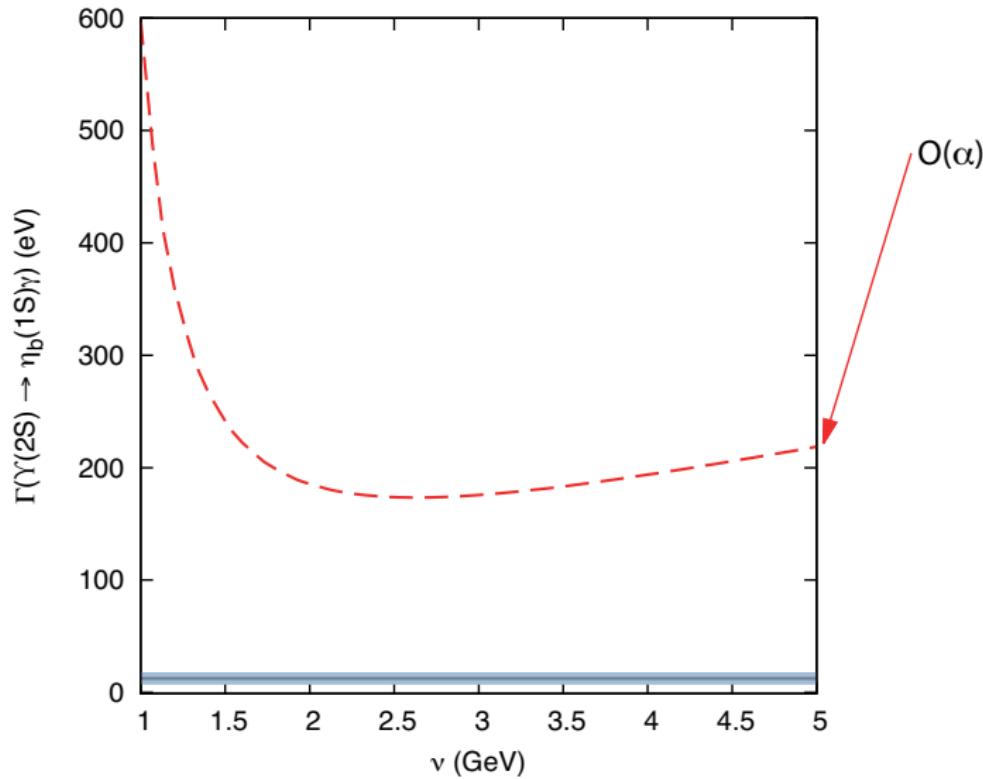
$$d_{vv}(\nu) = d_{vv}(m) + \frac{C_A}{\beta_0 - 2C_A} \pi \alpha_s(m) (z^{\beta_0 - 2C_A} - 1),$$

$$z = \left[ \frac{\alpha_s(\nu)}{\alpha_s(m)} \right]^{\frac{1}{\beta_0}} \simeq 1 - \frac{1}{2\pi} \alpha_s(\nu) \ln \left( \frac{\nu}{m} \right),$$

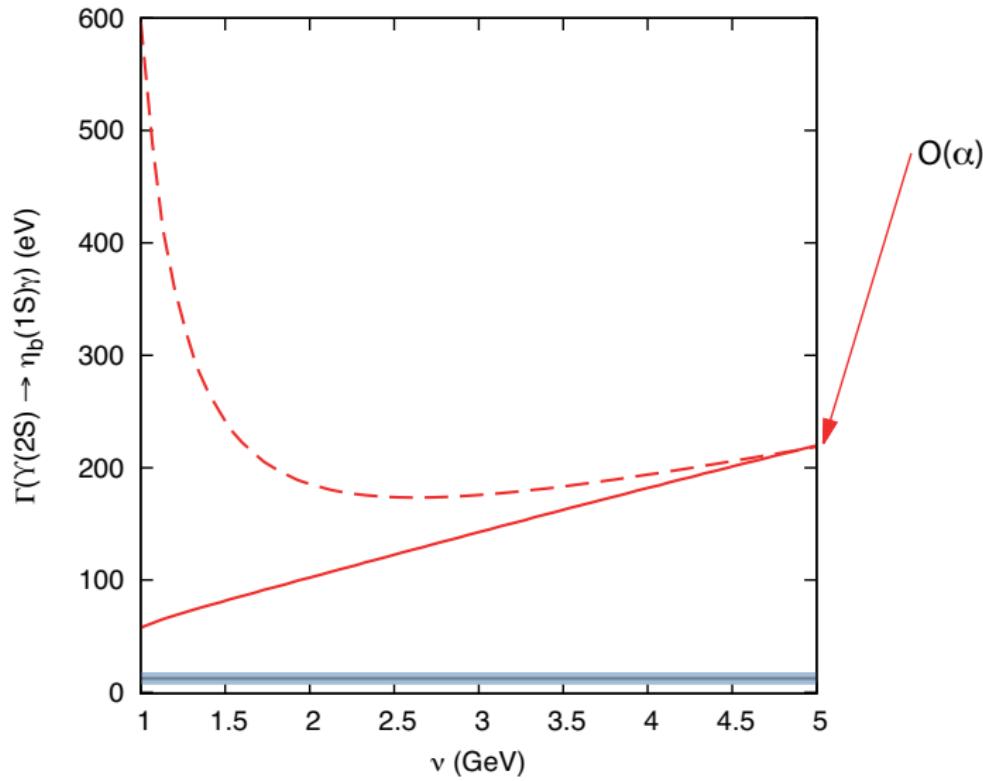
$$d_{sv}(m) = C_f \left( C_f - \frac{C_A}{2} \right) \pi \alpha_s(m), \quad d_{vv}(m) = - \left( C_f - \frac{C_A}{2} \right) \pi \alpha_s(m).$$



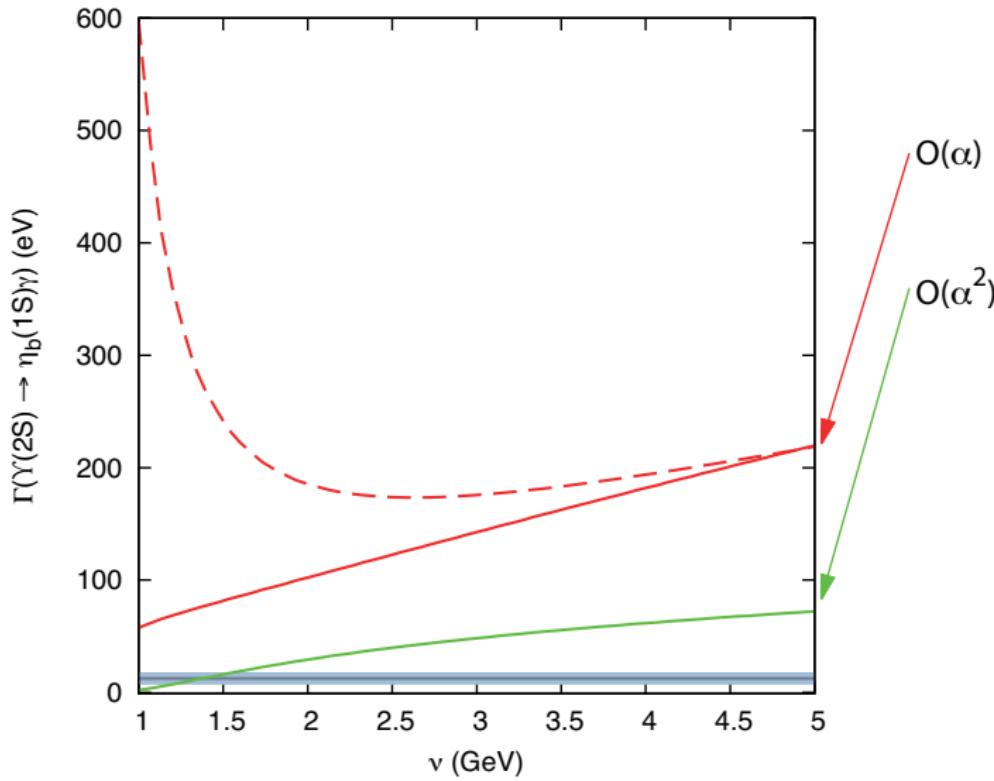
$$\text{no RG , } V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} , \mu = 1 \text{ GeV}$$



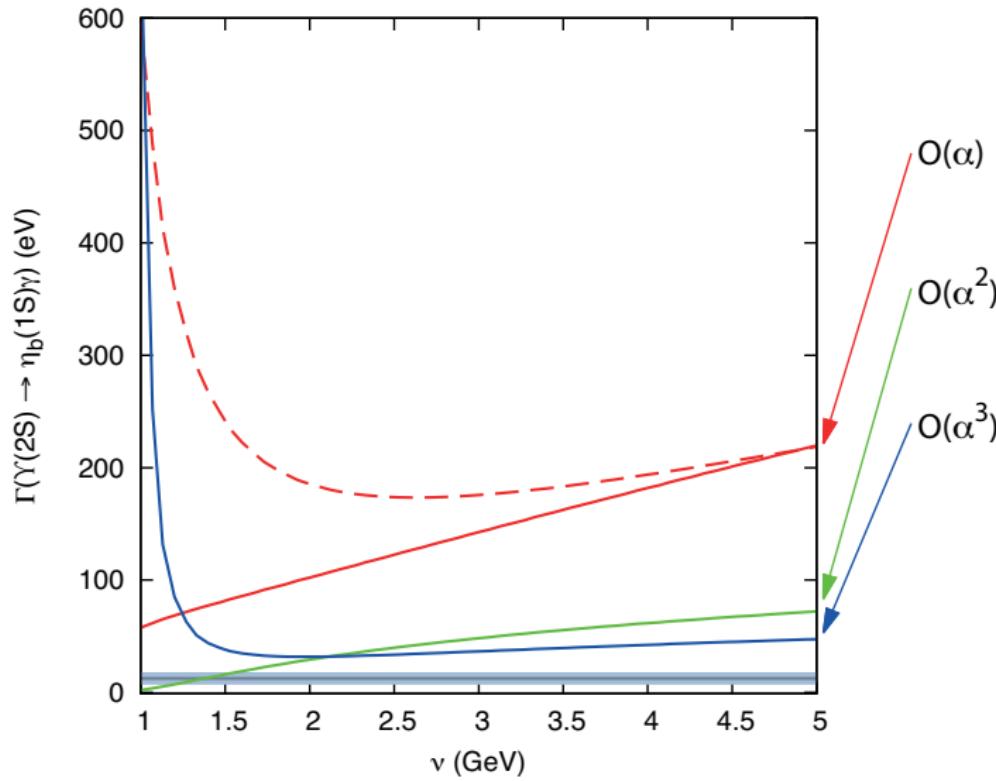
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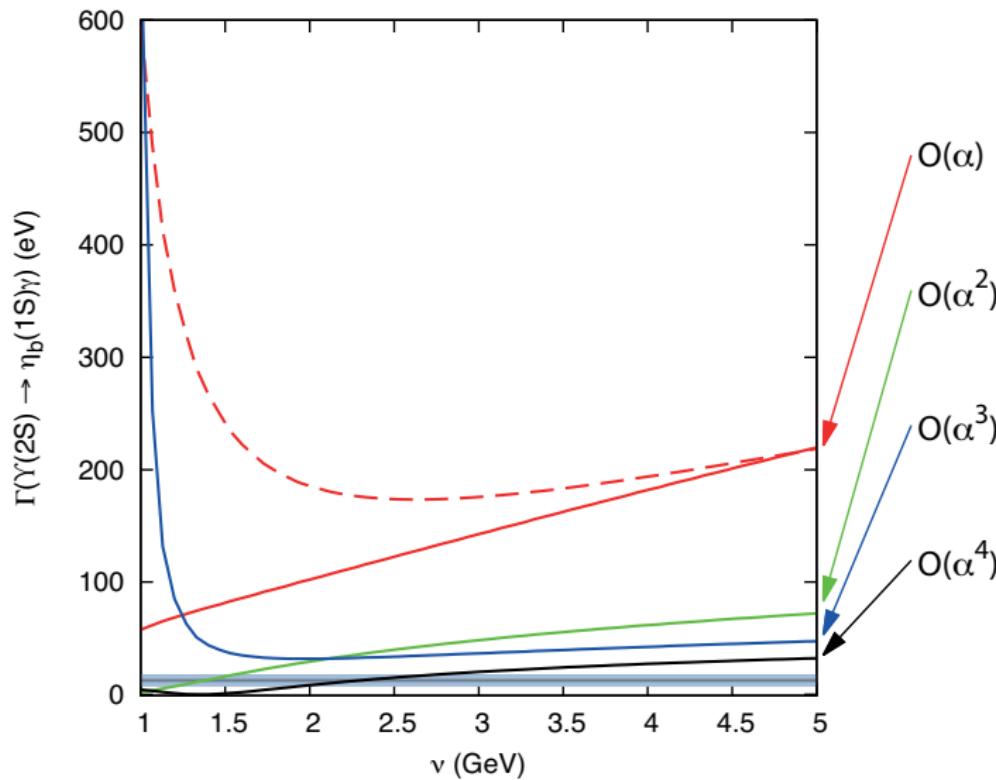
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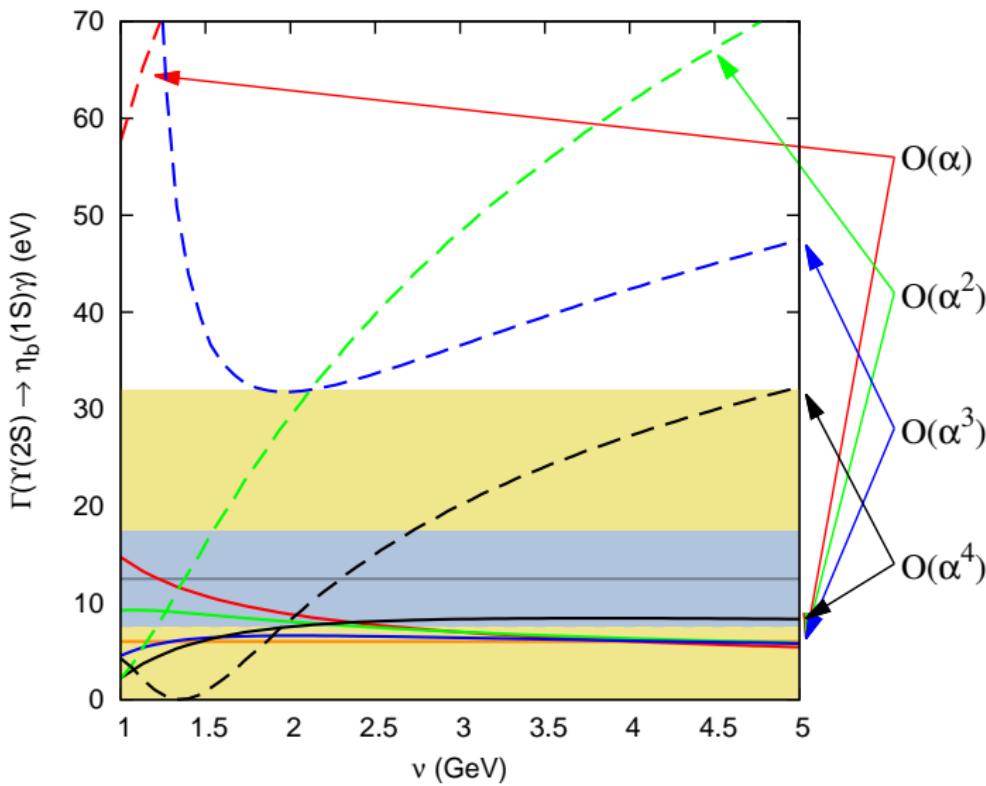
$$\text{RG , } V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} \right)$$



$$\text{RG , } V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} + a_2 \frac{\alpha_s^2(\mu)}{(4\pi)^2} \right)$$



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Prefactor (keV)	$\mathcal{A}(r^2)$	$\mathcal{A}(\vec{p}^2)$	$\mathcal{A}(V_{S^2})$	$\Gamma$ (eV)
10.3342	0.022	0.039	-0.042	6.3

Table: *The prefactor, the terms inside the brackets, and the total decay width  $\Gamma_{\Upsilon(2S) \rightarrow \eta_b(1S)\gamma}$ .*

$$\Gamma_{\Upsilon(2S) \rightarrow \eta_b(1S)\gamma}^{(\text{th})} = 0.006 \pm 0.006 (\mathcal{O}(v^5))^{+0.026}_{-0.006} (N_m)^{-0.001}_{+0.001} (\alpha_s)^{-0.000}_{+0.000} (m_{\overline{\text{MS}}}) \text{ keV}.$$

$$\Gamma_{\Upsilon(2S) \rightarrow \eta_b(1S)\gamma}^{(\text{th})} = 6^{+26}_{-06} \text{ eV}.$$

Matrix element.

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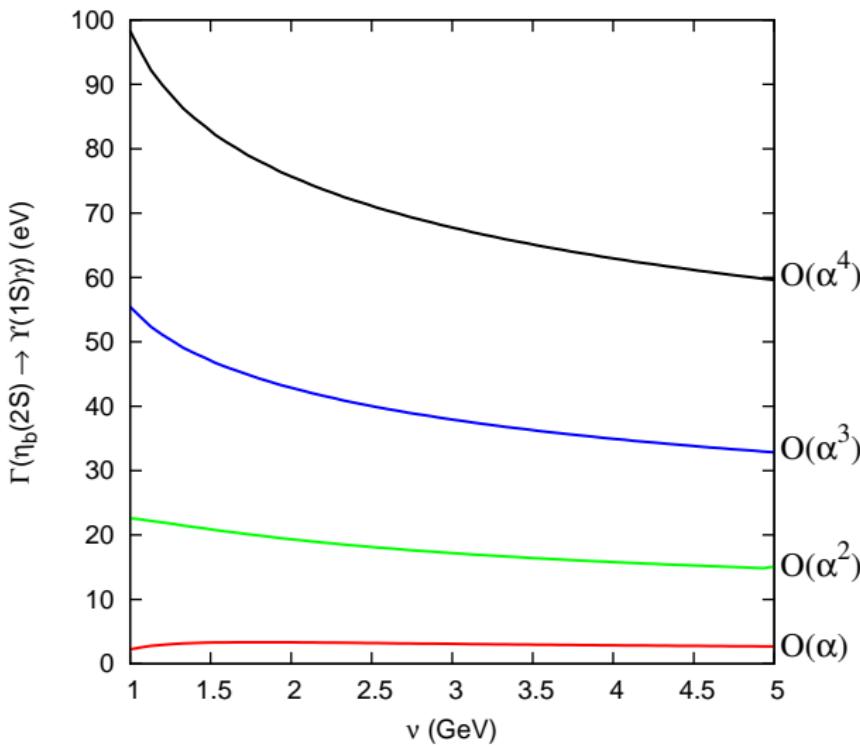


Figure: Plot of  $\Gamma_{\eta_b(2S) \rightarrow \gamma(1S)\gamma}$  using the static potential  $V_{RS'}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$  with  $\nu_r = \nu_f = 0.7$  GeV.

# CONCLUSIONS

Precision is  $k_\gamma^3/m^2 \times \mathcal{O}(\alpha_s^2, v^2)$  for the allowed transitions. The convergence for the  $b\bar{b}$  ground state was quite good, and also quite reasonable for the  $c\bar{c}$  ground state and the  $b\bar{b}$   $1P$  state.

$$\begin{aligned}\Gamma_{\Upsilon(1S) \rightarrow \eta_b(1S)\gamma} &= 15.18(51) \text{ eV}, \\ \Gamma_{J/\psi(1S) \rightarrow \eta_c(1S)\gamma} &= 2.12(40) \text{ keV}, \\ \Gamma_{h_b(1P) \rightarrow \chi_{b0}(1P)\gamma} &= 0.962(35) \text{ eV}, \\ \Gamma_{h_b(1P) \rightarrow \chi_{b1}(1P)\gamma} &= 8.99(55) \times 10^{-3} \text{ eV}, \\ \Gamma_{\chi_{b2}(1P) \rightarrow h_b(1P)\gamma} &= 0.118(6) \text{ eV}.\end{aligned}$$

Precision is  $k_\gamma^3/m^2 \times \mathcal{O}(v^4)$  for the forbidden transitions. Large logarithms associated with the heavy quark mass scale have also been resummed.

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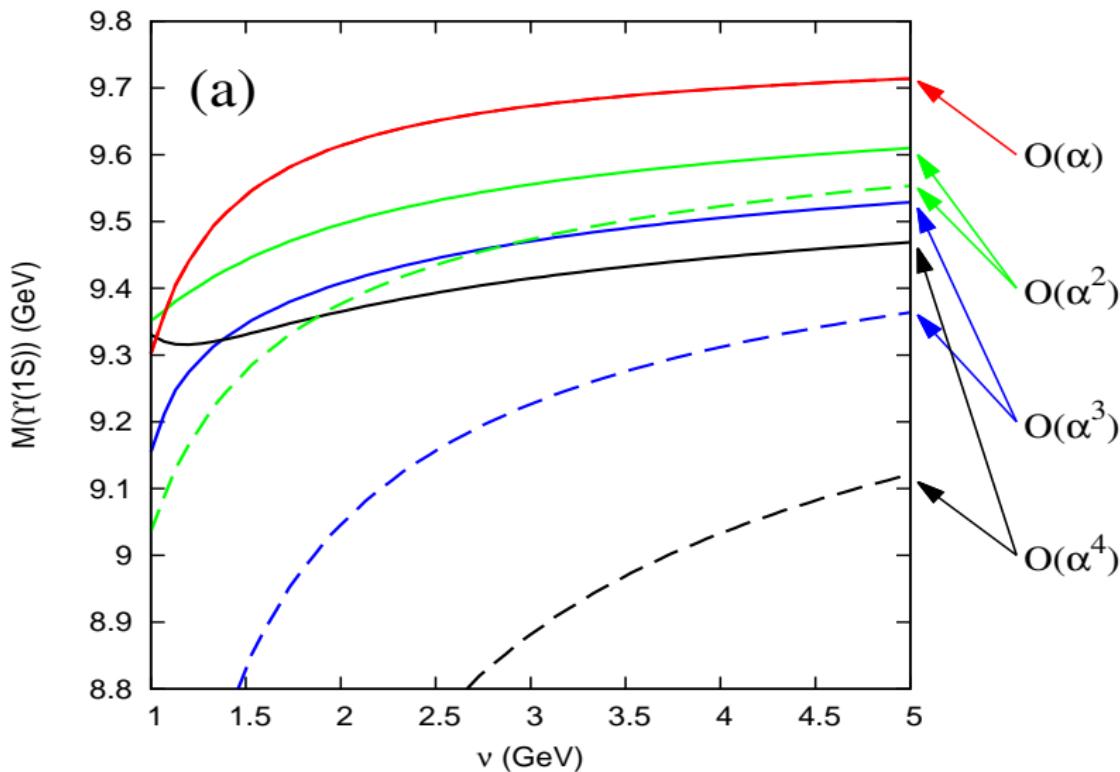


Figure:  $M_{10} = 2m_{b,RS'}(0.7\text{ GeV}) + E_{10}$  using the static potential  $V_{RS'}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$ . Dashed lines with  $\nu_f = 0$ . Continuous lines with  $\nu_f = 0.7$  GeV. In both cases  $\nu_r = \infty$  GeV.

$$V_{\text{RS}'}^{(N)}(r) = \begin{cases} (V^{(N)} + 2\delta m_{\text{RS}'}^{(N)})|_{\nu=\nu} \equiv \sum_{n=0}^N V_{\text{RS}',n} \alpha_s^{n+1}(\nu) & \text{if } r > \nu_r^{-1} \\ (V^{(N)} + 2\delta m_{\text{RS}'}^{(N)})|_{\nu=1/r} \equiv \sum_{n=0}^N V_{\text{RS}',n} \alpha_s^{n+1}(1/r) & \text{if } r < \nu_r^{-1}. \end{cases}$$

$$\delta m_X^{(N)}(\nu_f) = \nu_f \sum_{n=0}^N \delta m_X^{(n)}\left(\frac{\nu_f}{\nu}\right) \alpha_s^{n+1}(\nu)$$