

# Using Neutrino Mass Sum Rules to *realistically* probe Flavour Models

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*Based on:*

JHEP **1312** (2013) 005, with Steve King & Alex Stuart

FLASY 2014, Brighton, 20-06-2014

# *Contents:*

1. Mass Sum Rules: Origin & Anatomy
2. Analysis & example cases
3. Global Results
4. Conclusions

# 1. Mass Sum Rules: Origin & Anatomy

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- **BASIC PROBLEM FOR FLAVOUR MODELS:**

if the mixing angles are not correctly predicted, we classify a model as “excluded”

→ thus: mixing angles are not well suited to distinguish models without precision

- an alternative observable would be the  $|m_{ee}|$  in neutrinoless double beta decay, which depends on the mass scale and on the phases

→ BUT: suffers from nuclear physics unknowns

☺ Fortunately, some models yield **SUM RULES!!!**

# 1. Mass Sum Rules: Origin & Anatomy

- initial observation: if  $M_\nu$  contains two decisive parameters (typically by two flavon couplings), then it yields a mass sum rule
- known example based on  $A_4$ :

$$M_R = \begin{pmatrix} 2\alpha_s + \alpha_0 & -\alpha_s & -\alpha_s \\ -\alpha_s & 2\alpha_s & -\alpha_s + \alpha_0 \\ -\alpha_s & -\alpha_s + \alpha_0 & 2\alpha_s \end{pmatrix} \Lambda \quad M_D = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} y\nu$$

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- (complex) masses & sum rule:

$$(\tilde{m}_1, \tilde{m}_2, \tilde{m}_3) = \left( \frac{1}{3\alpha_s + \alpha_0}, \frac{1}{\alpha_0}, \frac{1}{3\alpha_s - \alpha_0} \right) \frac{y^2 v^2}{\Lambda}$$



$$\frac{1}{\tilde{m}_1} - \frac{1}{\tilde{m}_3} = \frac{2}{\tilde{m}_2}$$

**MASS SUM RULE**

## 2. Analysis & example cases

### **DISCLAIMER:**

- since we wanted to perform a complete study, we had to revisit cases that had already been studied
- we had to partially update (to be frank: *correct*) these discussions
- nearly all the authors of these partially incorrect results are here...

**→ *Please don't throw any rotten eggs or tomatoes at me, we did it only in the best interest of science!!!***

## 2. Analysis & example cases

- **EARLIER WORKS BY:**

Altarelli, Barry, Bazzocchi, Chen, Dorame,  
Feruglio, Hagedorn, Hirsch, Meloni, Morisi,  
Peinado, Valle,...

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- AND, IN PARTICULAR, SINCE HE COMPLAINED THAT I WOULD NOT MENTION HIM IN MY TALK:

**WERNER RODEJOHANN**

→ *PIONEERING SYSTEMATIC PAPER ON MASS SUM RULES:*  
Barry & Rodejohann, Nucl. Phys. **B842**, 33 (2011)

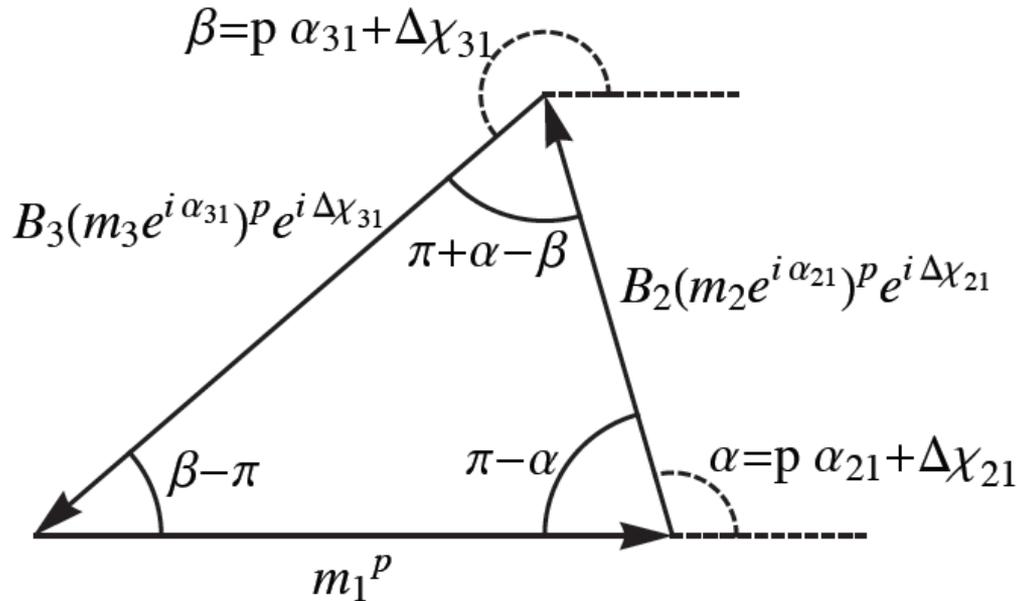
## 2. Analysis & example cases

- any mass sum rule can be written as follows:

$$m_1^p + B_2 (m_2 e^{i\alpha_{21}})^p e^{i\Delta\chi_{21}} + B_3 (m_3 e^{i\alpha_{31}})^p e^{i\Delta\chi_{31}} = 0.$$

→ complex equation that yields **TWO** conditions

→ geometrical interpretation as triangle:



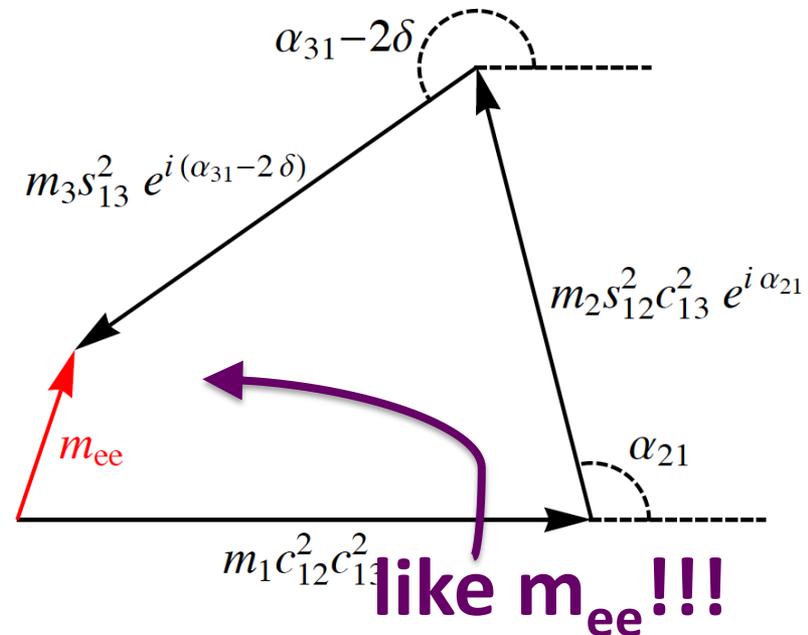
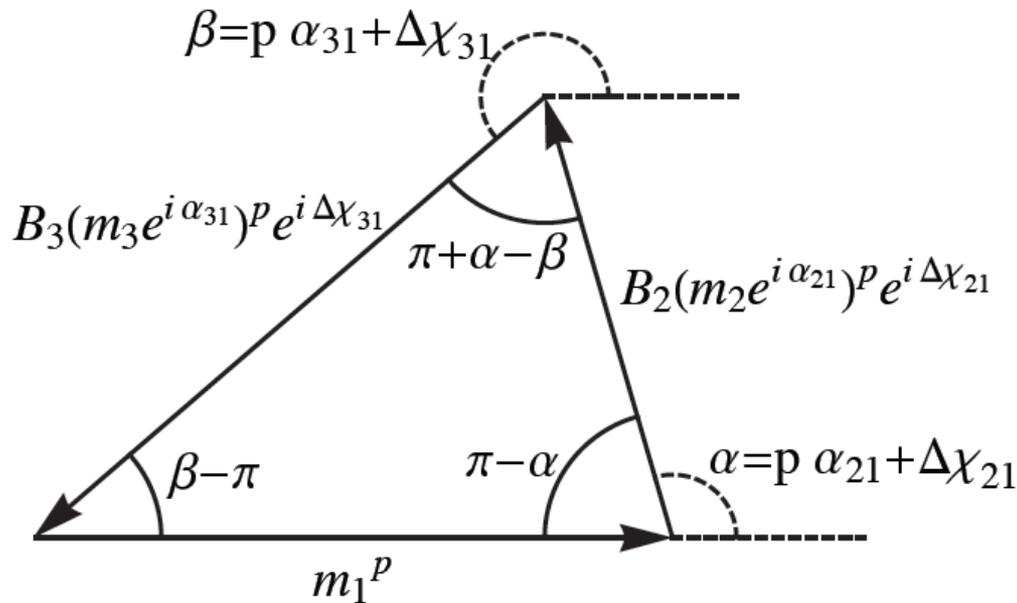
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## 2. Analysis & example cases

- sum rule based on  $S_4$  with inverse seesaw:

$$\boxed{\frac{1}{\sqrt{\tilde{m}_1}} + \frac{1}{\sqrt{\tilde{m}_2}} = \frac{2}{\sqrt{\tilde{m}_3}}}$$

Dorame *et al.*, Phys. Rev.  
**D86**, 056001 (2012)

- in terms of our general parameters:

$$p = -1/2, \quad B_2 = 1, \quad B_3 = 2, \quad \Delta\chi_{21} = 0, \quad \text{and} \quad \Delta\chi_{31} = \pi$$

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- one can show that  $m_3 < m_1 < m_2$  implies:

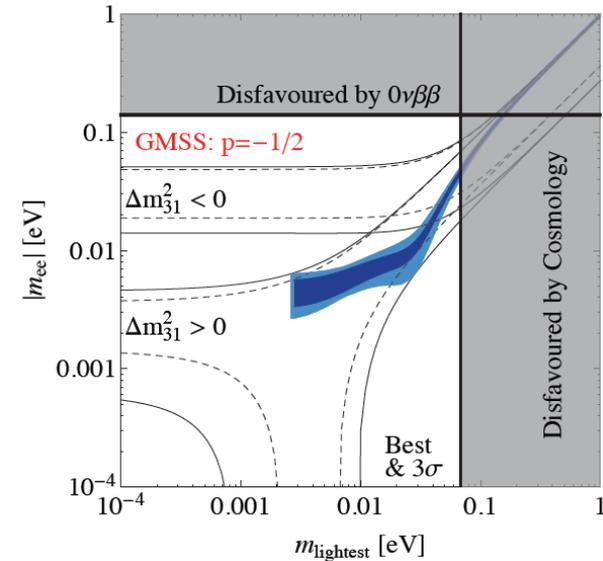
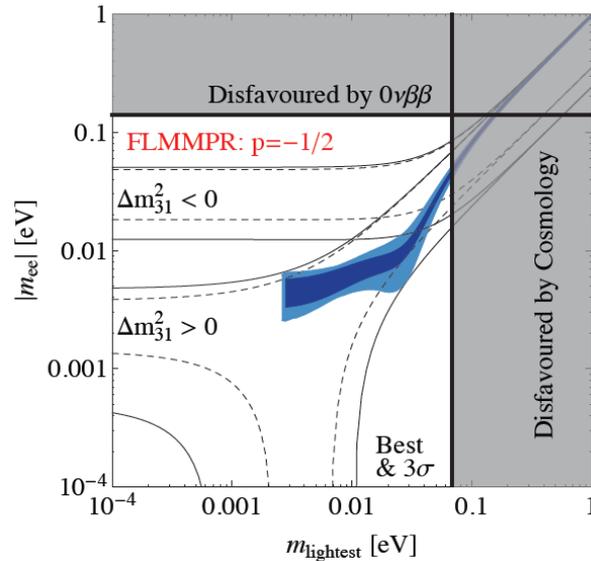
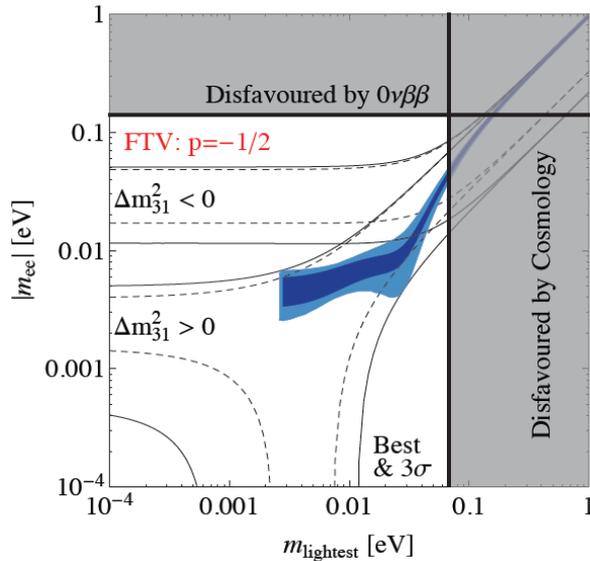
$$\underline{|\text{LHS}|} = \left| \frac{1}{\sqrt{\tilde{m}_1}} + \frac{1}{\sqrt{\tilde{m}_2}} \right| = \left| \frac{1}{\sqrt{m_1}} + \frac{e^{i(\phi_1 - \phi_2)/2}}{\sqrt{m_2}} \right| \leq \frac{1}{\sqrt{m_1}} + \frac{1}{\sqrt{m_2}}$$

$$\underline{\leq} \frac{1}{\sqrt{m_3}} + \frac{1}{\sqrt{m_3}} = \frac{2}{\sqrt{m_3}} = \underline{|\text{RHS}|} \quad \rightarrow \quad |\text{LHS}| < |\text{RHS}|$$

**$\rightarrow$  forbids inverted mass ordering**

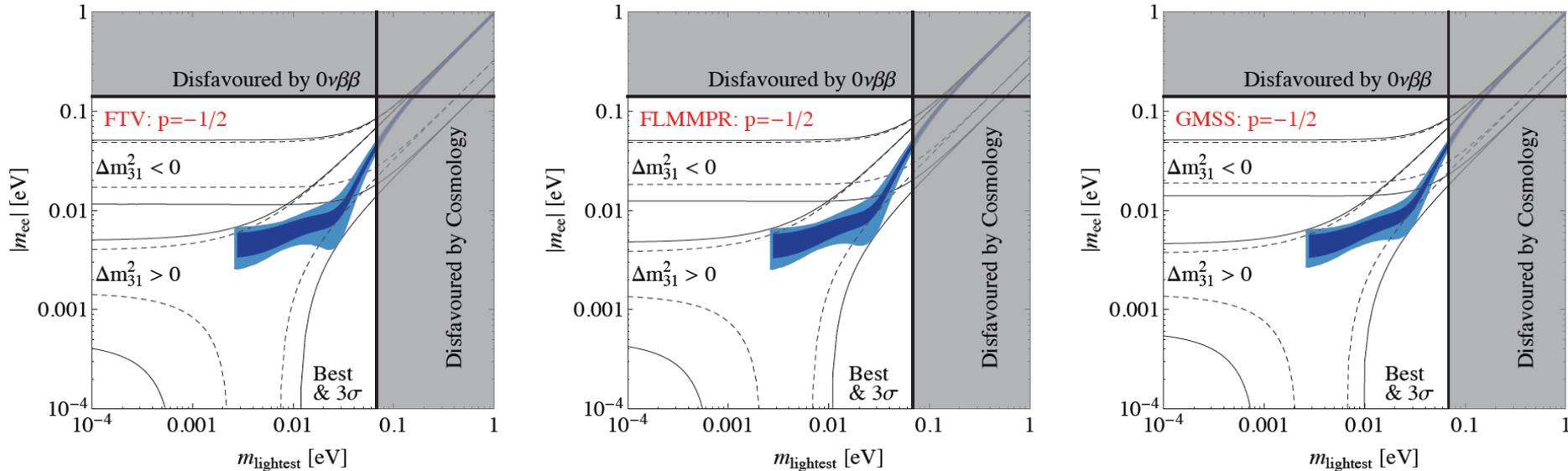
## 2. Analysis & example cases

- we did our numerics for all 3 global fits:



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➔ *plots look very similar for all three fits*

➔ *plots confirm our analytical results*

➔ *this is why we are so confident that we did things correctly*

# Advertisement: More sum rules

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$$\frac{1}{\tilde{m}_3} + \frac{2i(-1)^\eta}{\tilde{m}_2} = \frac{1}{\tilde{m}_1}$$

King, Luhn, Stuart: Nucl.  
Phys. **B867**, 203 (2013)

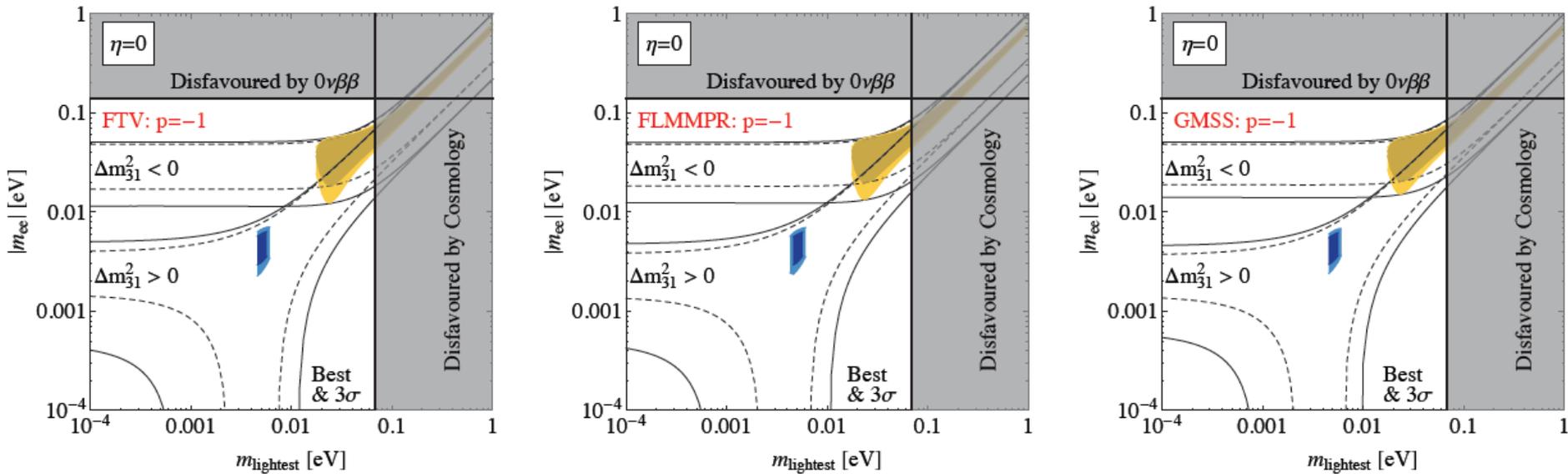
**model:  $\Delta(96)$  & seesaw type I**

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$\tilde{m}_3^{-1} \pm 2i\tilde{m}_2^{-1} = \tilde{m}_1^{-1}$	NO, $3\sigma$	IO, $3\sigma$
$m_{\text{lightest}}$	$\sim [0.0045, 0.0061]$	$\gtrsim 0.018$
$ m_{ee} $ (FTV)	$[0.0023, 0.0071]$	$> 0.012$
$ m_{ee} $ (FLMMPR)	$[0.0024, 0.0070]$	$> 0.013$
$ m_{ee} $ (GMSS)	$[0.0024, 0.0068]$	$> 0.015$

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$$\tilde{m}_1^{-1} - 2\tilde{m}_2^{-1} - \tilde{m}_3^{-1} = 0$$

Altarelli, Barry, Chen, Feruglio,  
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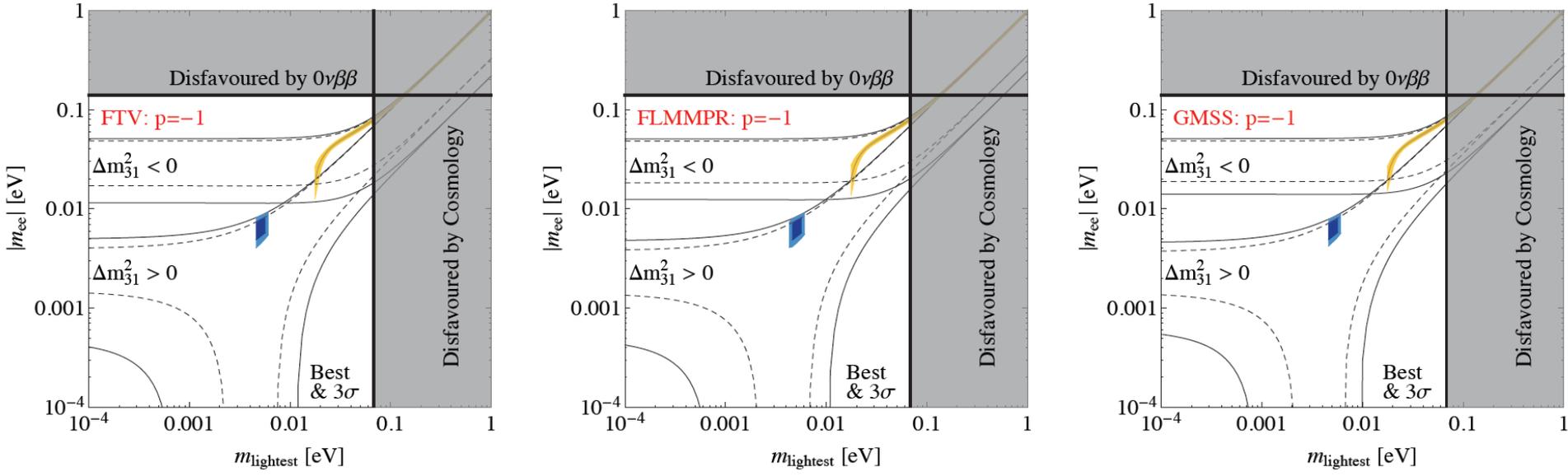
**models:  $A_4$  or  $T'$  & seesaw type I**

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$2\tilde{m}_2^{-1} + \tilde{m}_3^{-1} = \tilde{m}_1^{-1}$	NO, 3σ	IO, 3σ
$m_{\text{lightest}}$	$\sim [0.0045, 0.0061]$	$\gtrsim 0.018$
$ m_{ee} $ (FTV)	$[0.0040, 0.0092]$	$> 0.013$
$ m_{ee} $ (FLMMPR)	$[0.0041, 0.0091]$	$> 0.014$
$ m_{ee} $ (GMSS)	$[0.0042, 0.0089]$	$> 0.015$

# Advertisement: More sum rules

$$\tilde{m}_1 + \tilde{m}_2 = 2\tilde{m}_3$$

Lindner, Merle, Niro:  
JCAP **1101**, 034 (2011)

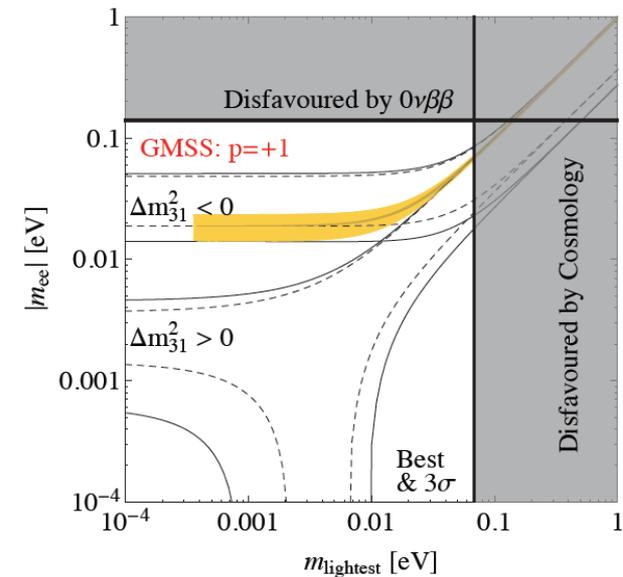
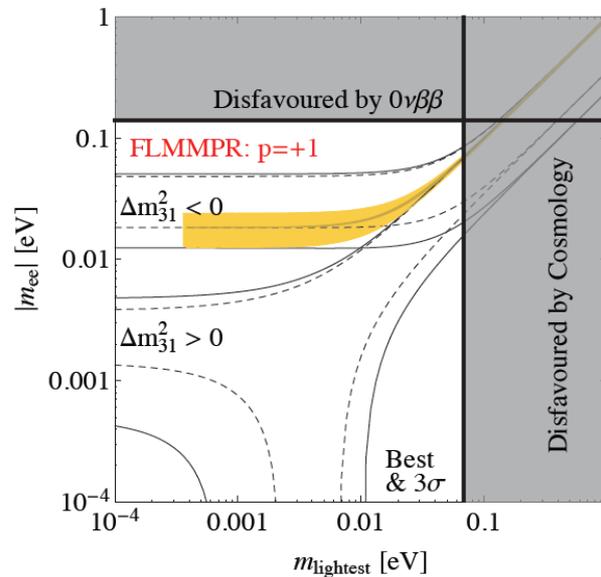
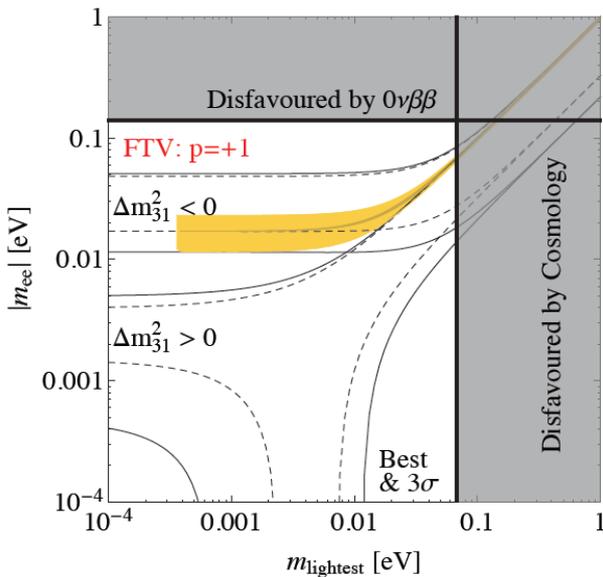
model:  $L_e$ - $L_\mu$ - $L_\tau$  & seesaw type II

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$\tilde{m}_1 + \tilde{m}_2 = 2\tilde{m}_3$	NO, $3\sigma$	IO, $3\sigma$
$m_{\text{lightest}}$	forbidden	$\gtrsim 0.00036$
$ m_{ee} $ (FTV)	forbidden	$> 0.012$
$ m_{ee} $ (FLMMPR)	forbidden	$> 0.013$
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# 3. Global Results

- general findings:
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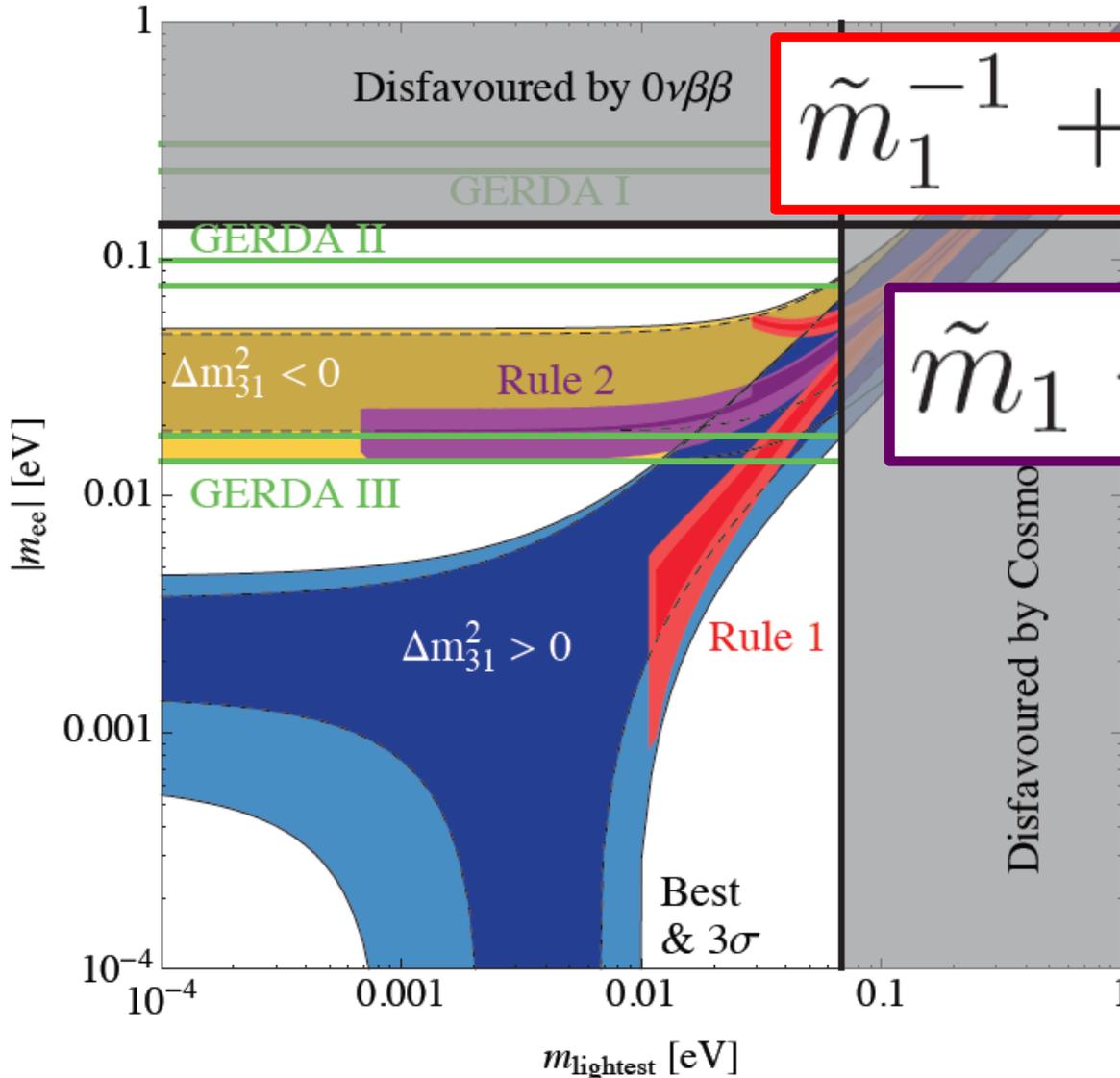
- general findings:
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  - phases **DO** play an important role:
    - $\delta$  *cannot* be neglected
    - phase re-definitions must be done consistently in both sum rule AND  $m_{ee}$

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- general findings:
  - neutrino mass sum rules probe ***whole classes of flavour models***
  - phases **DO** play an important role:
    - $\delta$  ***cannot*** be neglected
    - phase re-definitions must be done consistently in both sum rule AND  $m_{ee}$
  - sum rules generically yield **TWO** conditions
    - ➔ ordering and phase relations restricted
    - ➔ characteristic “signatures” in  $|m_{ee}|$
    - ➔ ***restrictions can be so strong that they even may overcome the nuclear physics uncertainties***

# 3. Global Results

- model distinction:



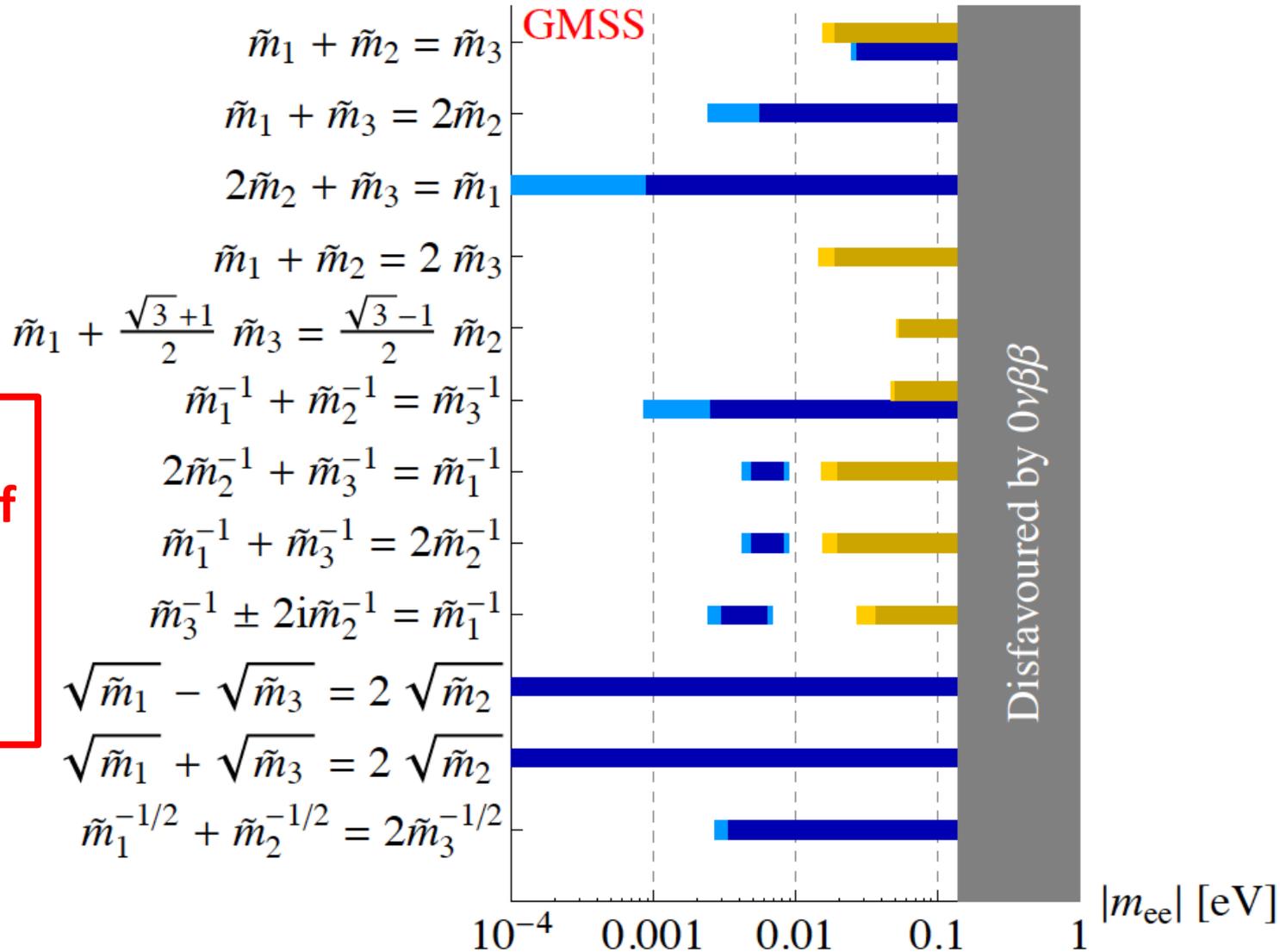
$$\tilde{m}_1^{-1} + \tilde{m}_2^{-1} = \tilde{m}_3^{-1}$$

$$\tilde{m}_1 + \tilde{m}_2 = \tilde{m}_3$$

Distinctive power  
in spite of nuclear  
physics  
uncertainties!!!

# 3. Global Results

- global analysis:



Complete classification of more than 50 known flavour models

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- we have done the most complete classification:
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  - *all three global fits*
  - *most recent experimental bounds & NME uncertainties*

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- if **your** model predicts a sum rule and **you** want it analysed (pick your fit, get a nice plot, etc.)...
  - ... **I am happy to serve as co-author!!! :-p**



**THANK  
YOU!!!**

# Effective theory of a doubly charged singlet scalar: complementarity of neutrino physics and the LHC

Stephen F. King\*, Alexander Merle† and Luca Panizzi‡

*Physics and Astronomy, University of Southampton  
Southampton, SO17 1BJ, United Kingdom*

June 18, 2014

## Abstract

We consider a rather minimal extension of the Standard Model involving just one extra particle, namely a single  $SU(2)_L$  singlet scalar  $S^{++}$  and its antiparticle  $S^{--}$ . We propose a model independent effective operator, which yields an effective coupling of  $S^{\pm\pm}$  to pairs of same sign weak gauge bosons,  $W^\pm W^\pm$ . We also allow tree-level couplings of  $S^{\pm\pm}$  to pairs of same sign right-handed charged leptons  $l_R^\pm l_R^\pm$  of the same or different flavour. We calculate explicitly the resulting two-loop diagrams in the effective theory responsible for neutrino mass and mixing. We propose sets of benchmark points for various  $S^{\pm\pm}$  masses and couplings which can yield successful neutrino masses and mixing, consistent with limits on charged lepton flavour violation and neutrinoless double beta decay. We discuss the prospects for  $S^{\pm\pm}$  production at the LHC at the benchmark points, including single and pair production of  $S^{\pm\pm}$  plus jets and missing energy. The model illustrates the complementarity between neutrino physics (including LFV) and the LHC, involving just one new particle, the  $S^{\pm\pm}$ .

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