



Pion-kaon femtoscopy in Au+Au collisions at STAR

Katarzyna Poniatowska*
For STAR Collaboration

*Warsaw University of Technology
Faculty of Physics



Outline

- HBT (Hanbury-Brown, Twiss) interferometry
- Space-time asymmetry
- Motivation
- Data selection
- Comparison of correlation functions for
 $\sqrt{s_{NN}} = (130, 39, 19.6, 7.7) \text{ GeV}$
- Comparison of asymmetry for
 $\sqrt{s_{NN}} = (130, 39, 19.6, 7.7) \text{ GeV}$
- Summary

HBT interferometry (Hanbury-Brown, Twiss)

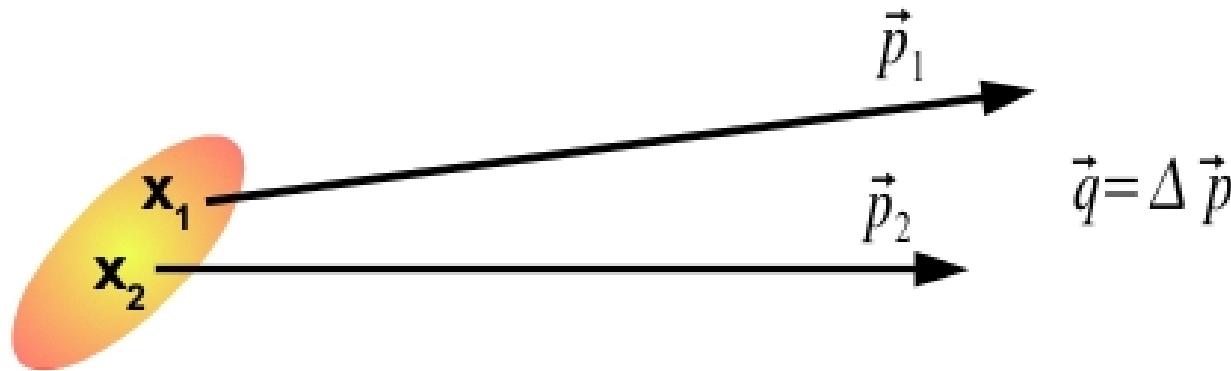
Interferometries:

1. An interference of two waves from different sources located in space. The detector records amplitude-phase diffraction signal.
2. **The interferometry of intensity occurs, when we detect two identical particles in different points in time and space (or with different energy and momentum)**

Method proposed in '50 by Hanbury-Brown and Twiss.

At first was used in astronomy to measure the angular size of stars (using emission of gamma quantum) => photons interferometry of intensity

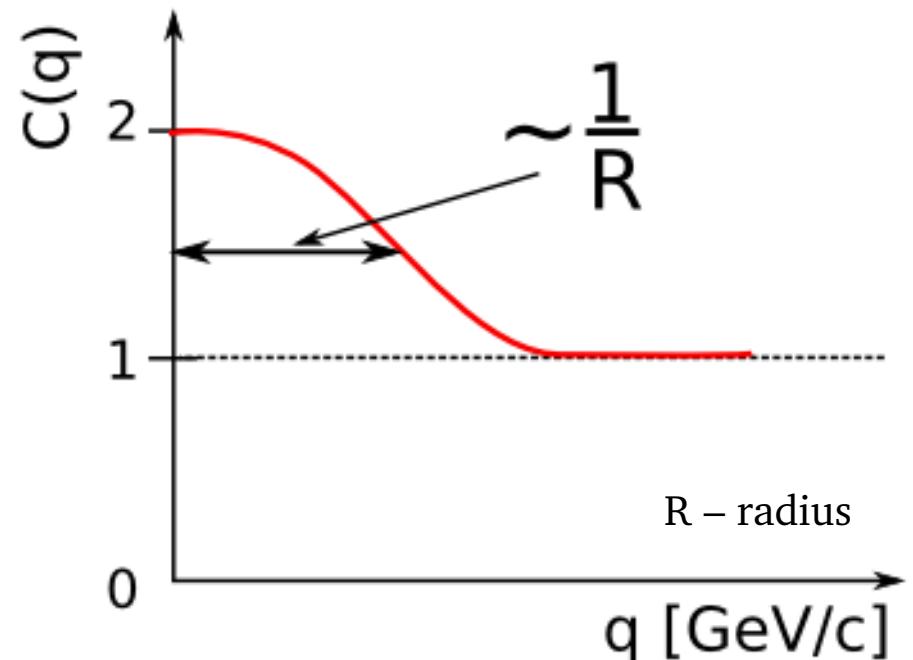
HBT



- Analyze many pairs of identical particles (p_1, x_1) and $(p_2, x_2) \rightarrow$ calculate q the difference momentum vector of those particles.
- Calculate the correlation function of pairs.

$$C(\mathbf{p}_1, \mathbf{p}_2) = \frac{P_2(\mathbf{p}_1, \mathbf{p}_2)}{P_1(\mathbf{p}_1)P_1(\mathbf{p}_2)}$$

- Calculate the source size.



$P_2(\mathbf{p}_1, \mathbf{p}_2)$ – the probability of observing two particles with momentum \mathbf{p}_1 , and \mathbf{p}_2 at the same time and the same place.
 $P_1(\mathbf{p}_1), P_1(\mathbf{p}_2)$ – the probability of observing two particles with momentum \mathbf{p}_1 , and \mathbf{p}_2 separately.

Space-time asymmetry

$\cos(\Psi) > 0$

Catching up

Long time of
effective
interaction.

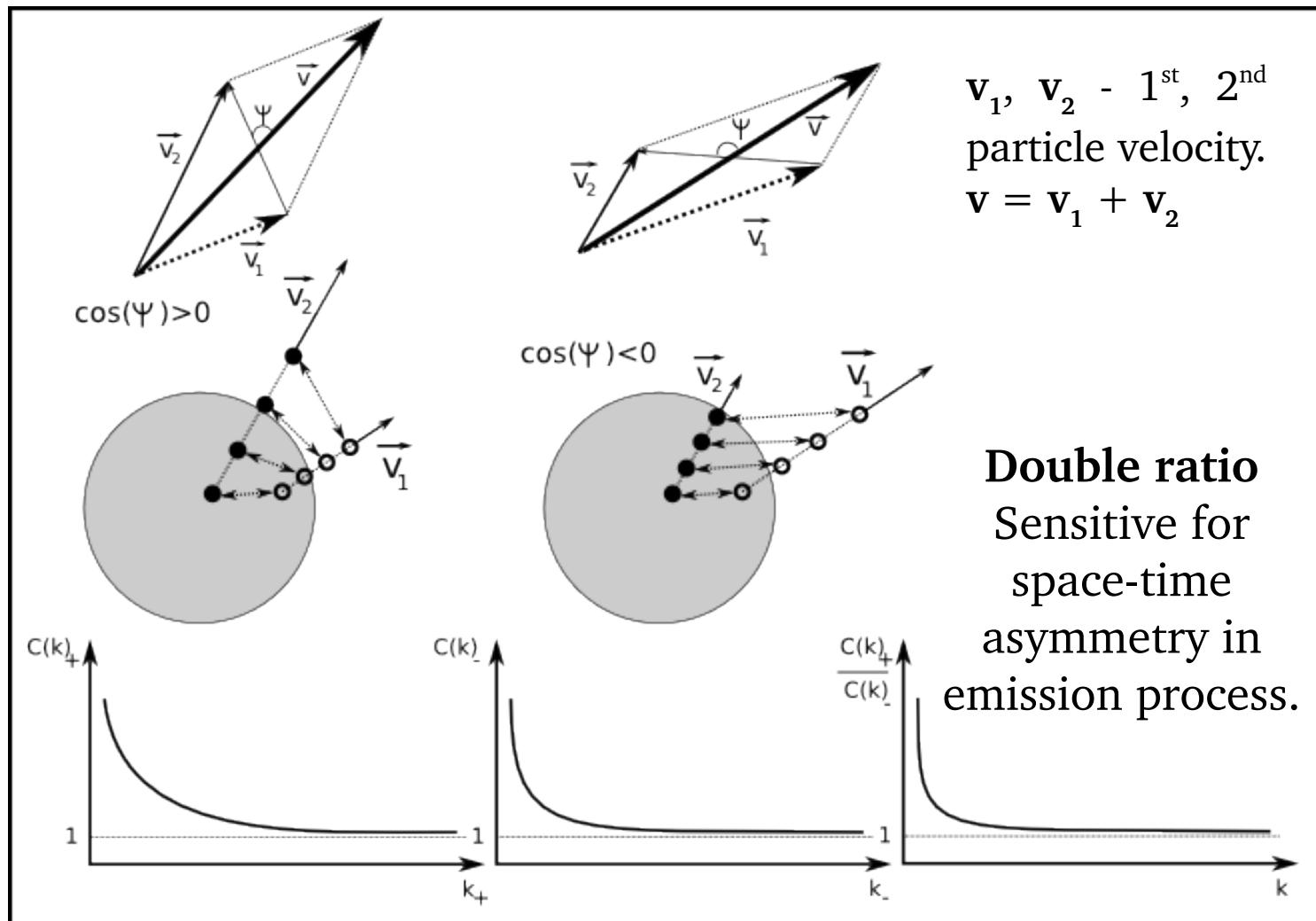
Strong
correlation.

$\cos(\Psi) < 0$

Run away

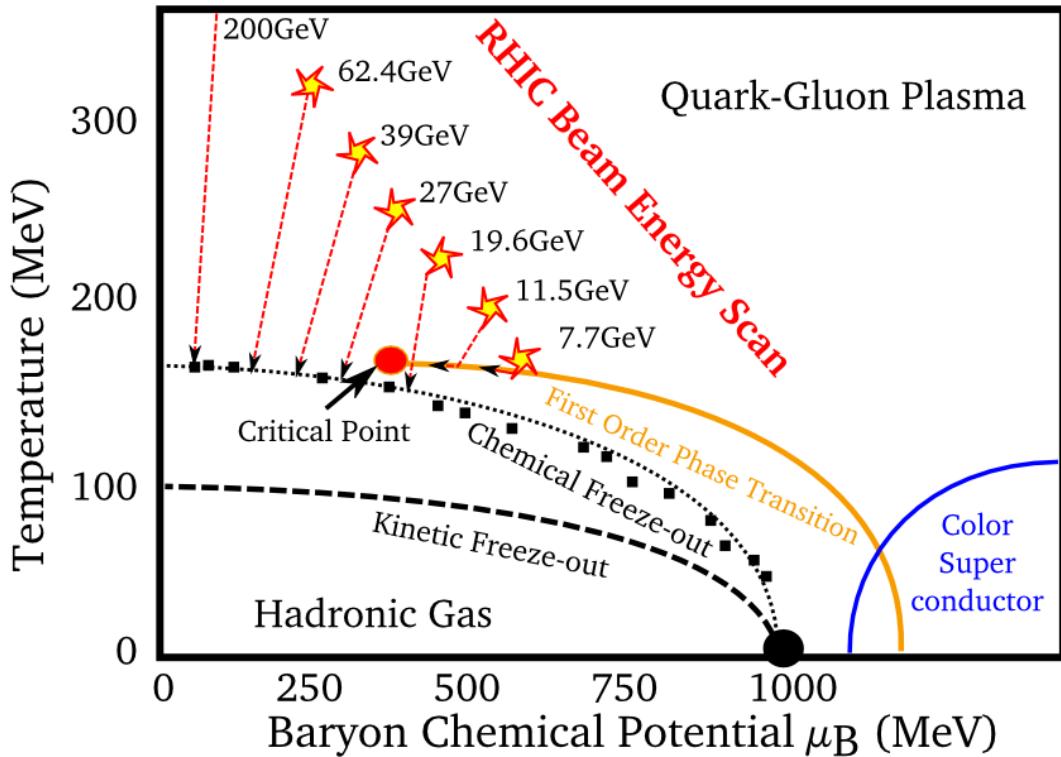
Short time of
effective
interaction.

Weak
correlation.



k^* - the magnitude of the three-momentum of the particles in the pair rest frame.

Motivation



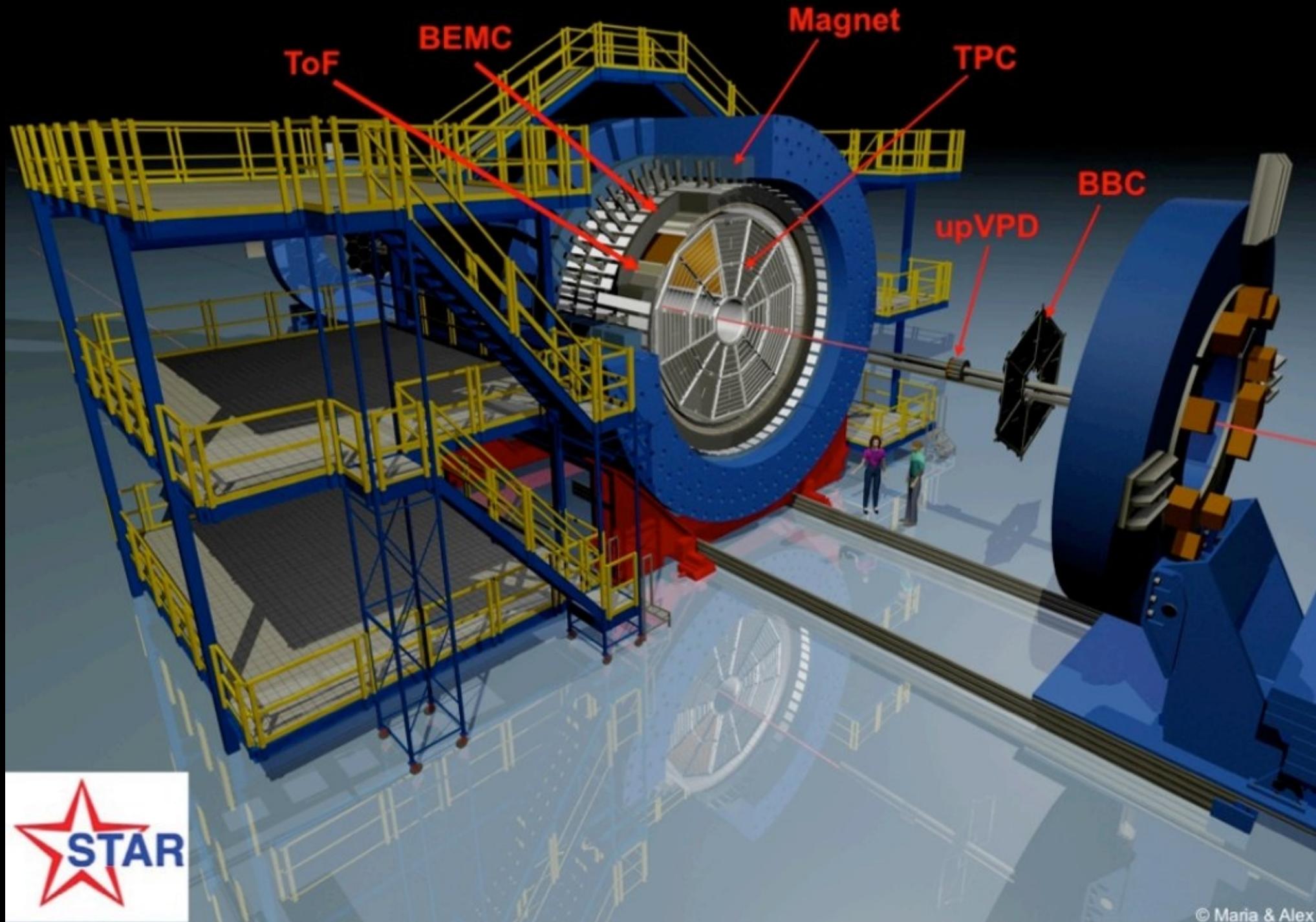
Analyze all Beam Energy Scan (BES) energies and find answers:

If or how pion-kaon source changes with energy?

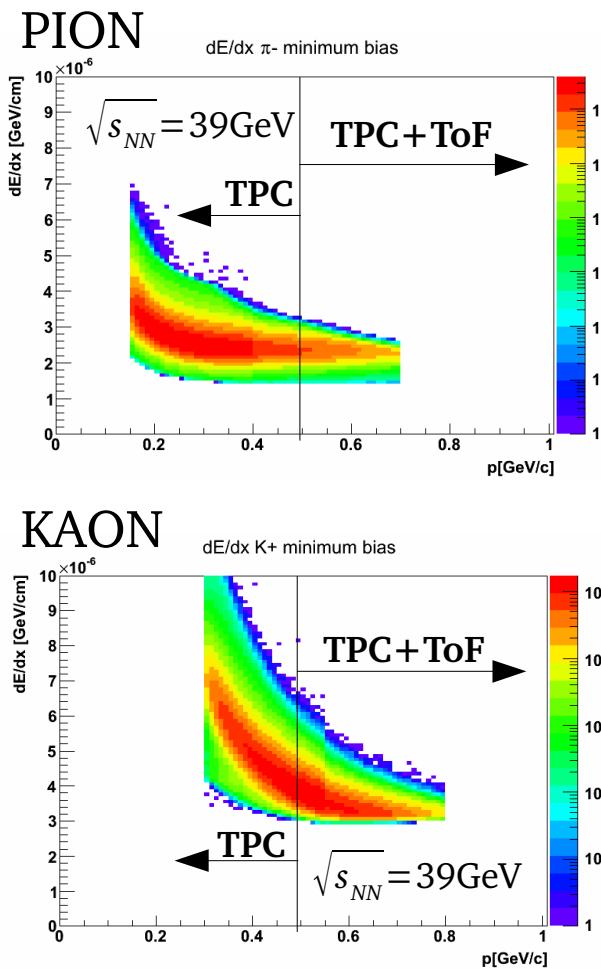
If or how pion-kaon asymmetry in emission process looks for all BES energies?

If or how the flow affects the pion-kaon system, consisting of particles of different masses?

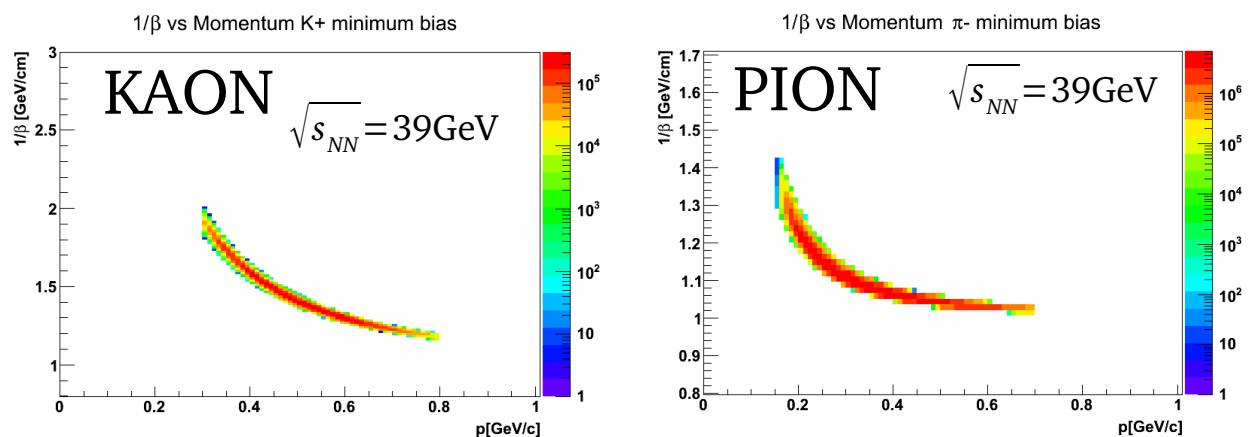
The Solenoidal Tracker At RHIC (STAR)



Data selection

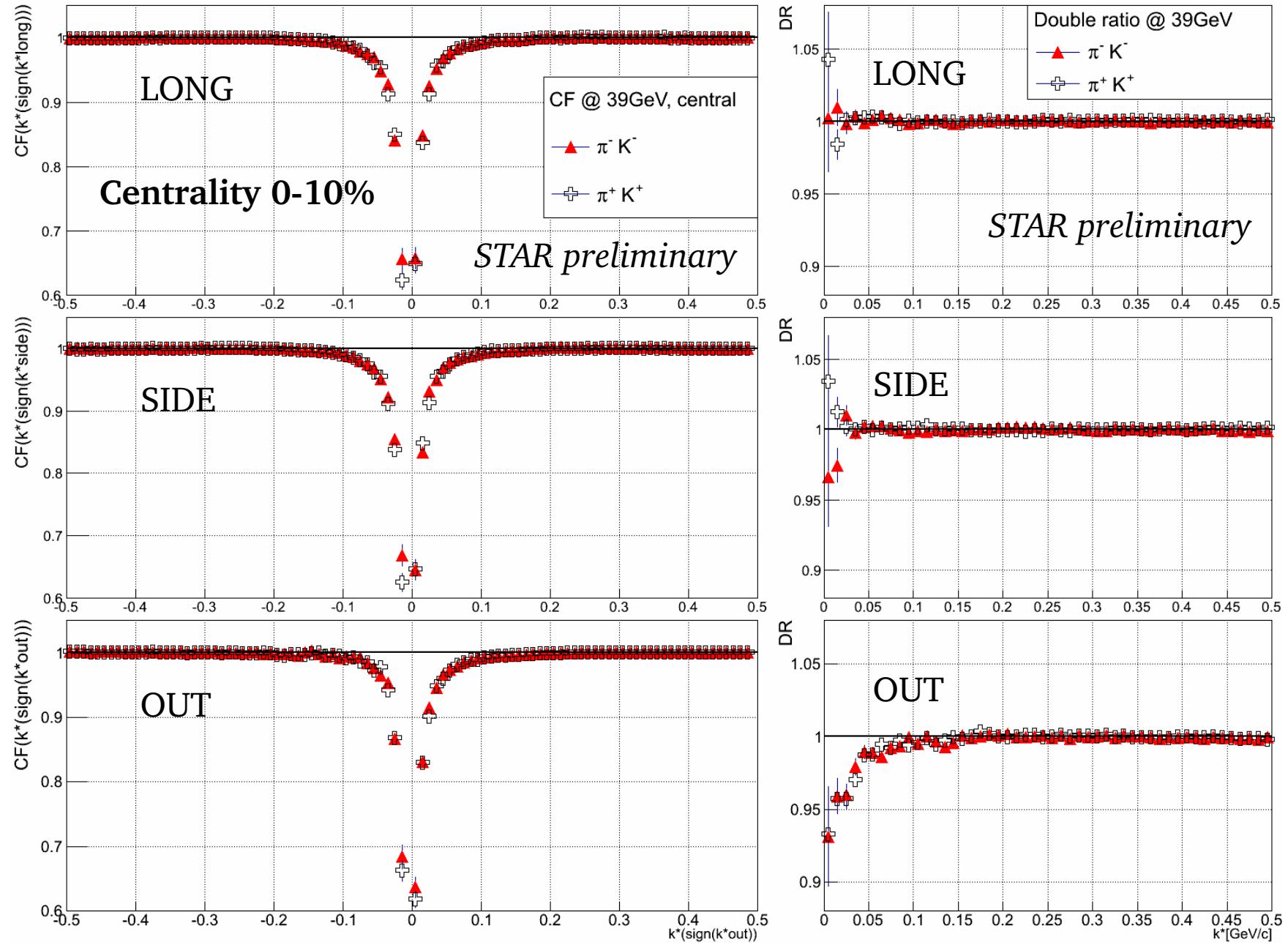


$\sqrt{s_{NN}}$	7.7 GeV	19.6 GeV	39 GeV	130 GeV
analyzed	4.1 mln events	17.5 mln events	91.2 mln events	0,32 mln events
π transverse momentum	[0.1, 1.2] GeV/c	[0.1, 1.2] GeV/c	[0.15, 0.7] GeV/c	[0.08, 0.25] GeV/c
K transverse momentum	[0.1, 1.2] GeV/c	[0.1, 1.2] GeV/c	[0.3, 0.8] GeV/c	[0.4, 0.7] GeV/c
N σ for π and K	< 3.0	< 3.0	< 3.0	< 3.0
πm^2	[0.0176, 0.022] GeV^2/c^4	[0.0176, 0.022] GeV^2/c^4	[0.017, 0.026] GeV^2/c^4	----
$K m^2$	[0.23, 0.26] GeV^2/c^4	[0.23, 0.26] GeV^2/c^4	[0.22, 0.27] GeV^2/c^4	----
$ \eta $	< 0.5	< 0.5	< 0.5	< 0.5
DCA	< 3cm	< 3cm	< 3cm	< 3cm

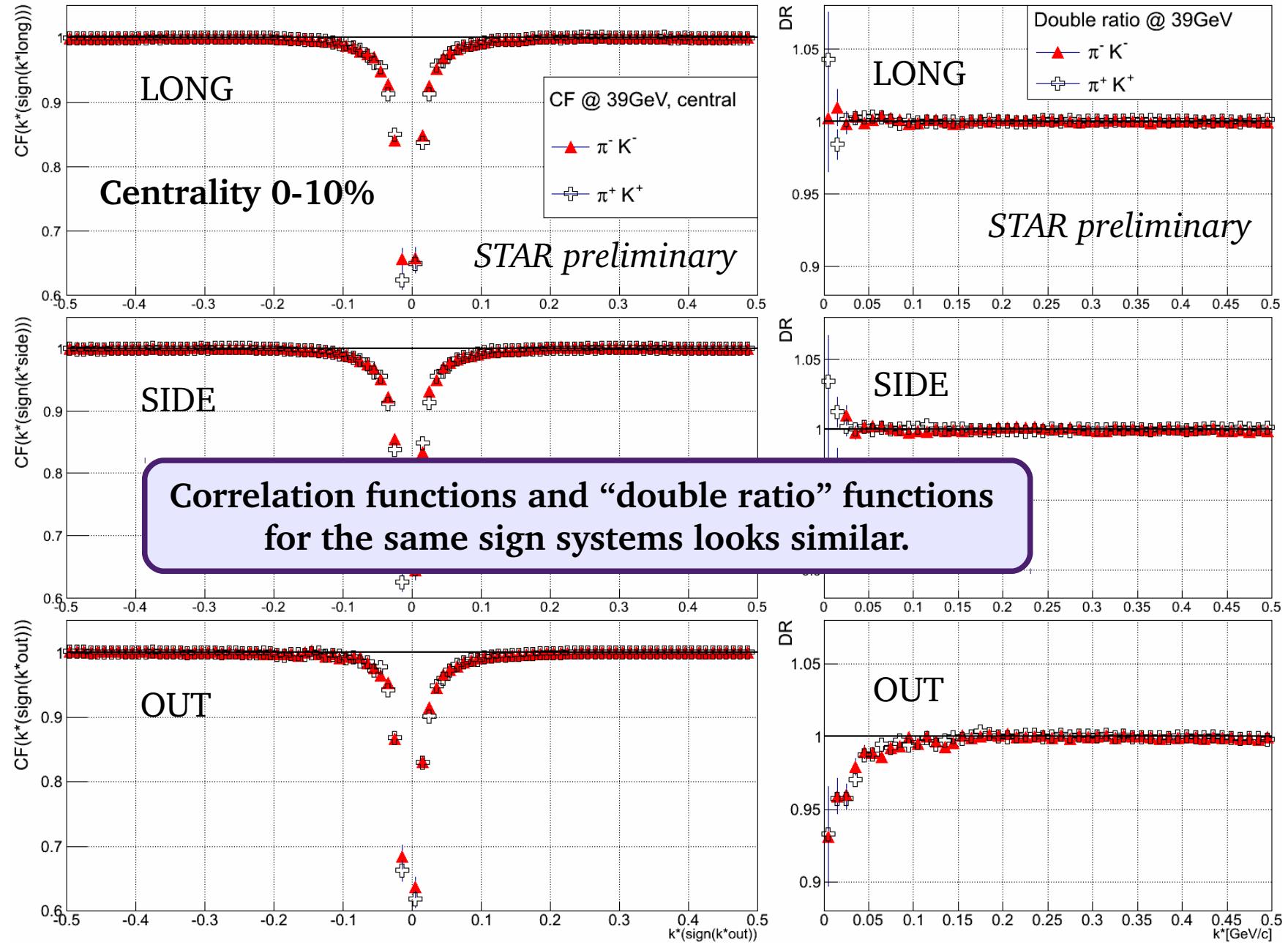


Centrality definition:
central: 0-10%
minimum bias: 0-80%

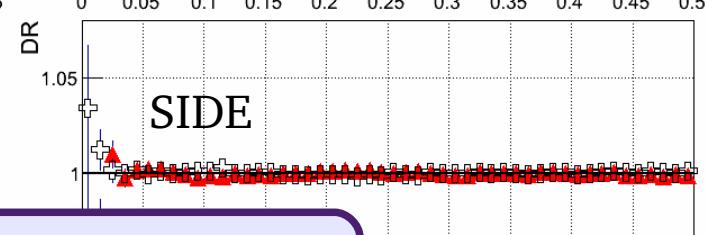
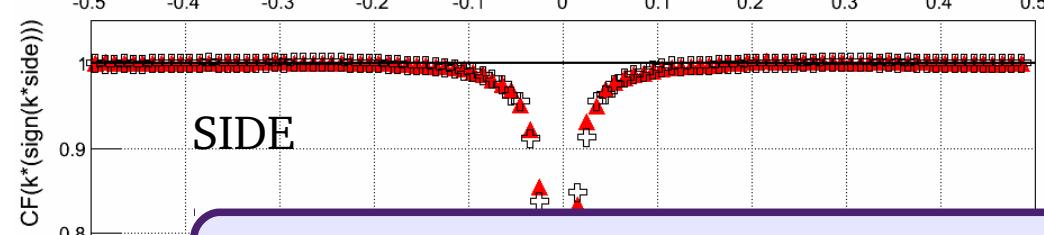
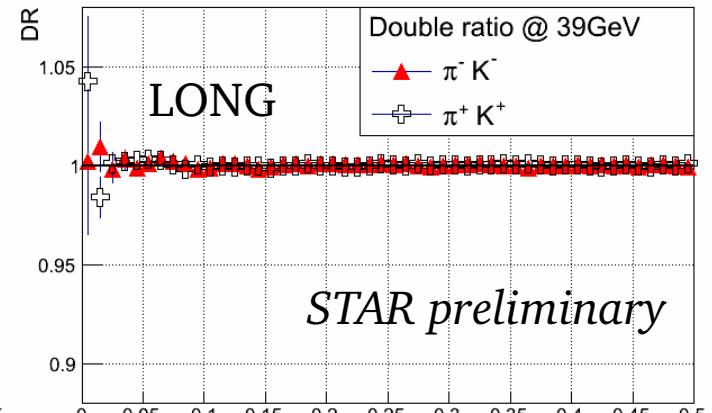
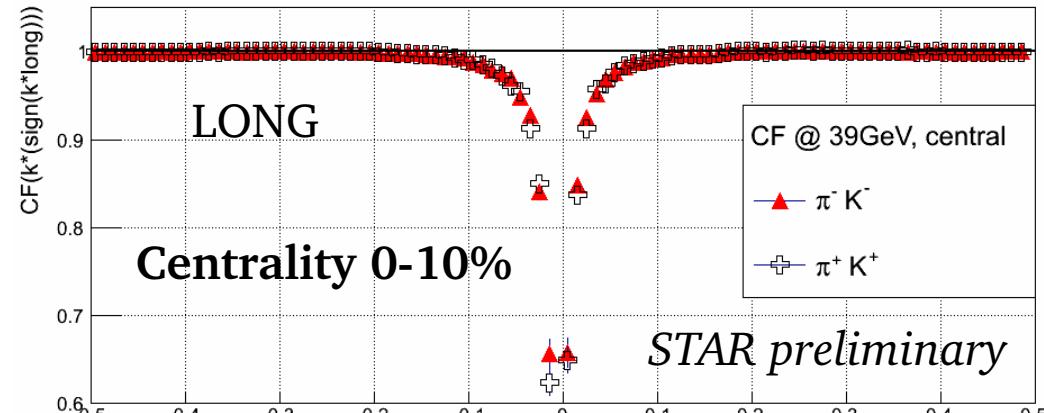
CF & DR for π^- K- and π^+ K+



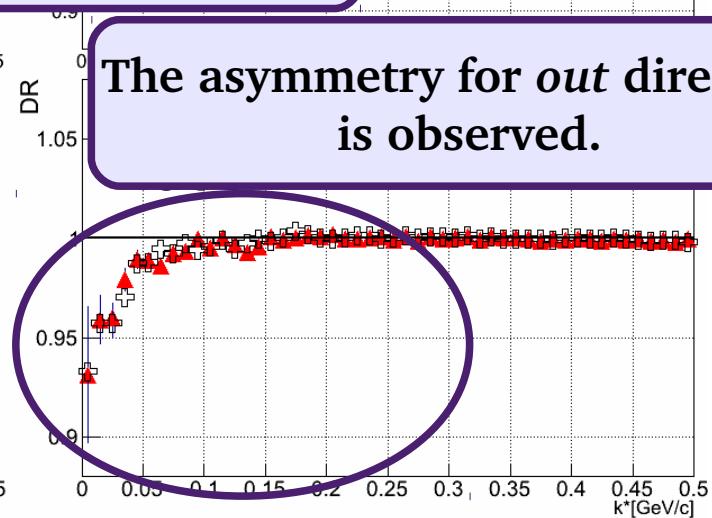
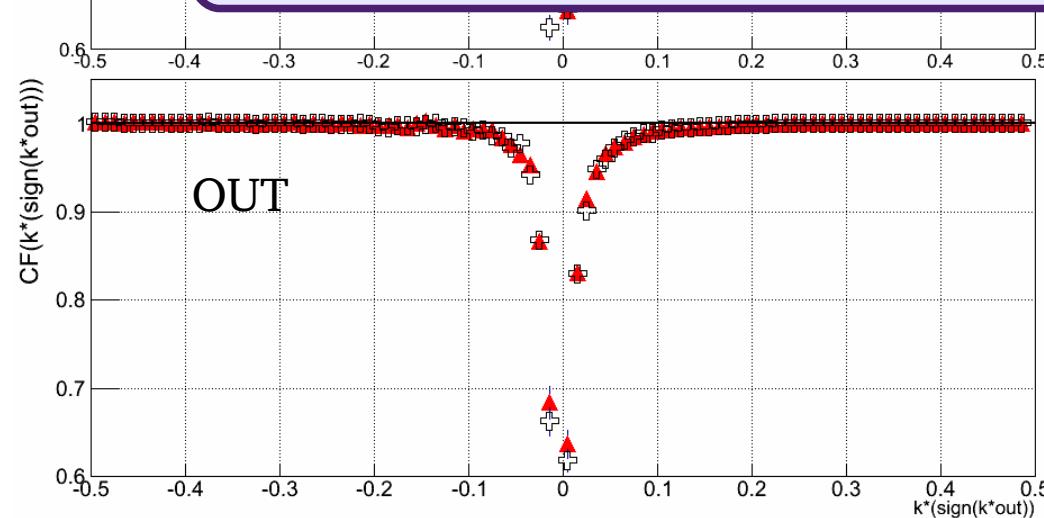
CF & DR for $\pi^- K^-$ and $\pi^+ K^+$



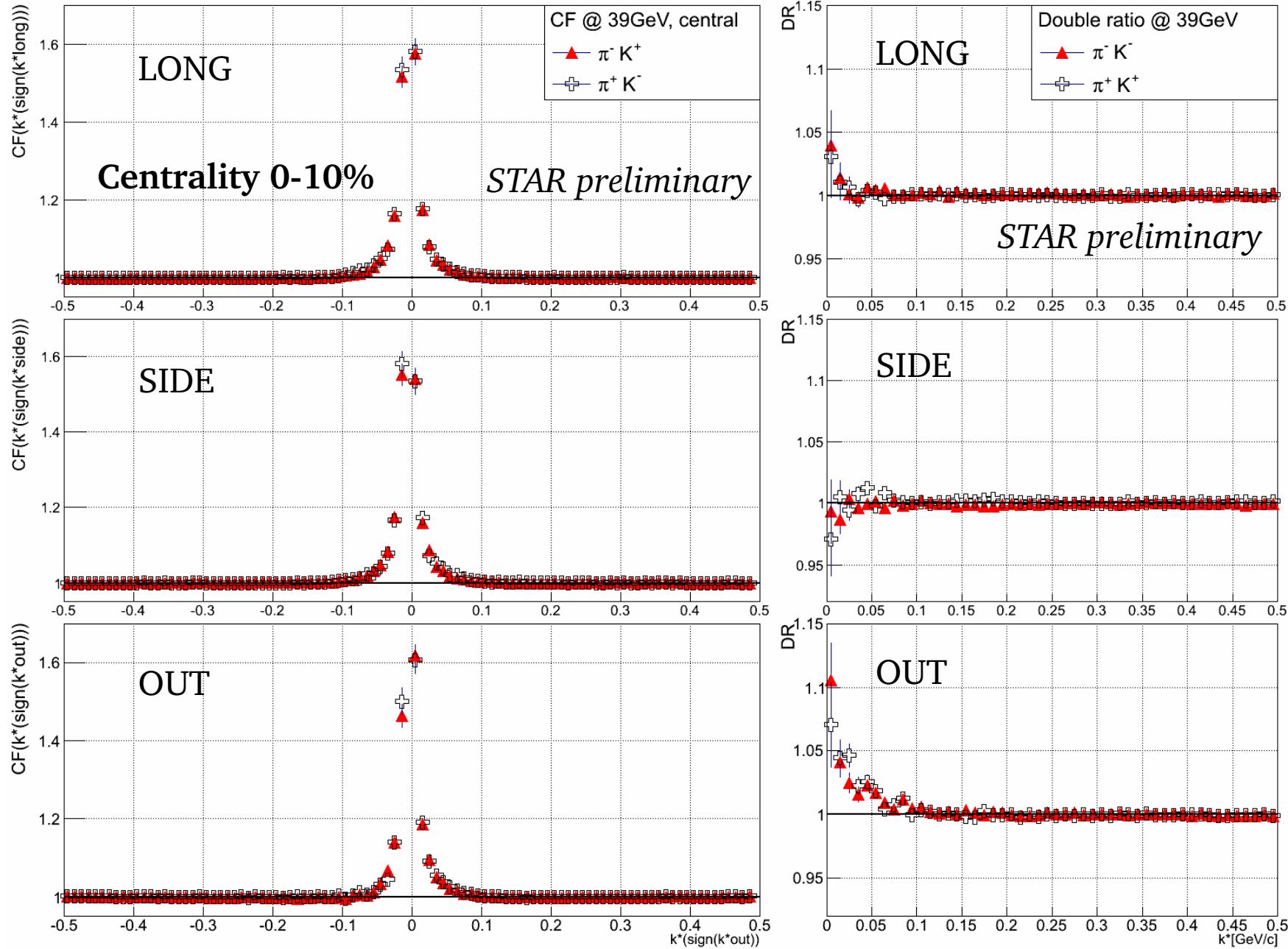
CF & DR for $\pi^- K^-$ and $\pi^+ K^+$



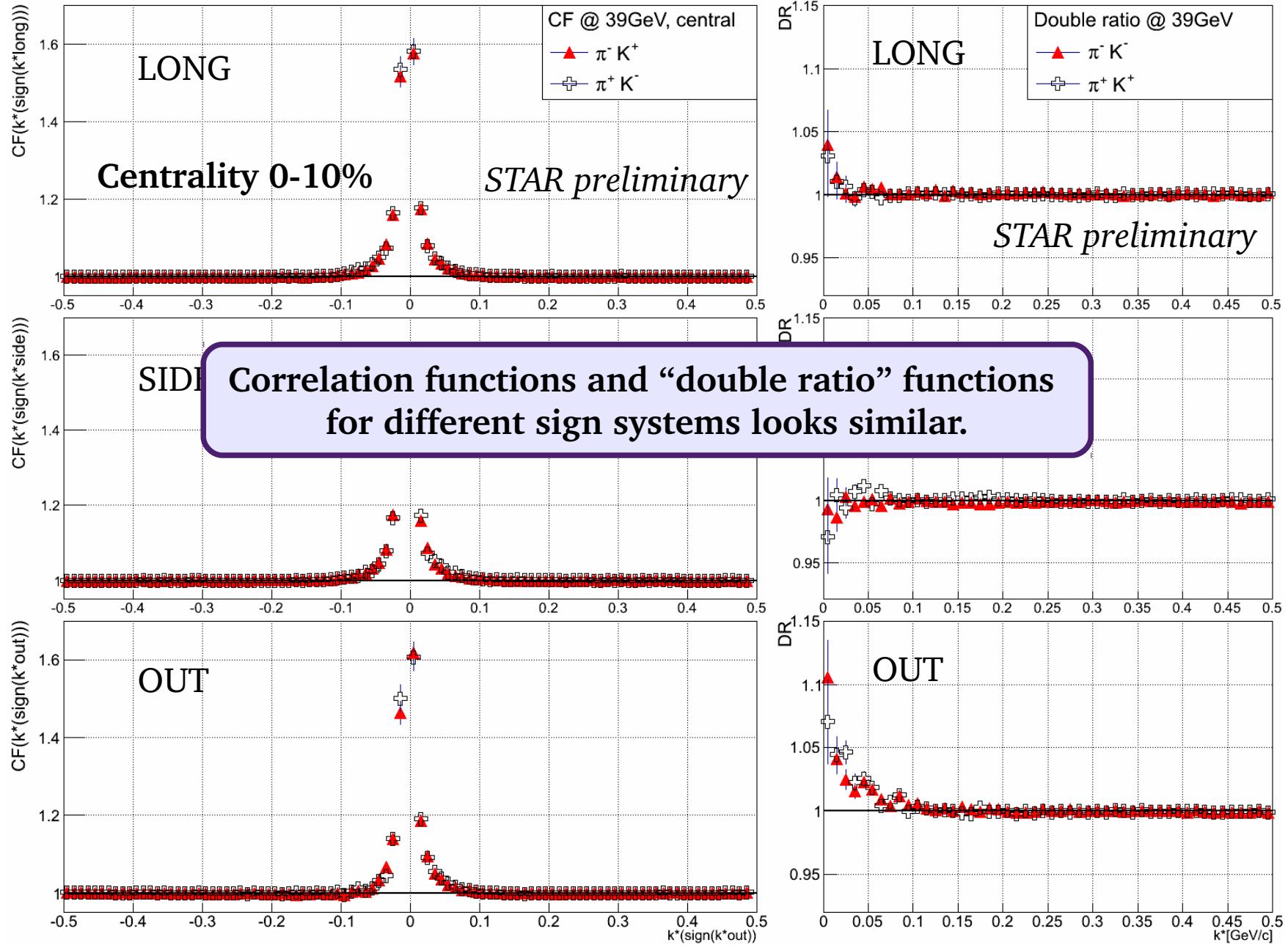
Correlation functions and “double ratio” functions
for the same sign systems looks similar.



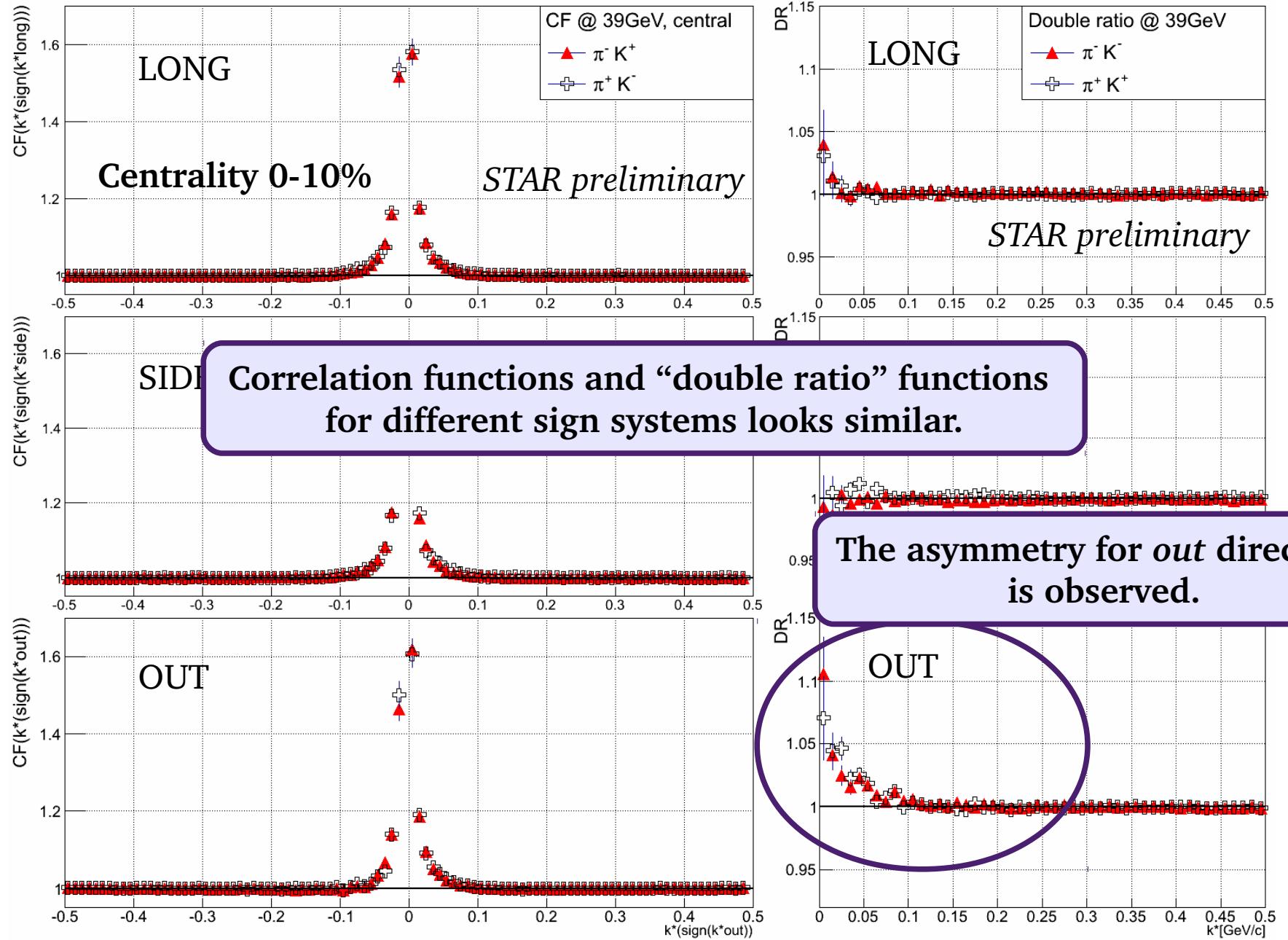
CF & DR for $\pi^- K^+$ and $\pi^+ K^-$



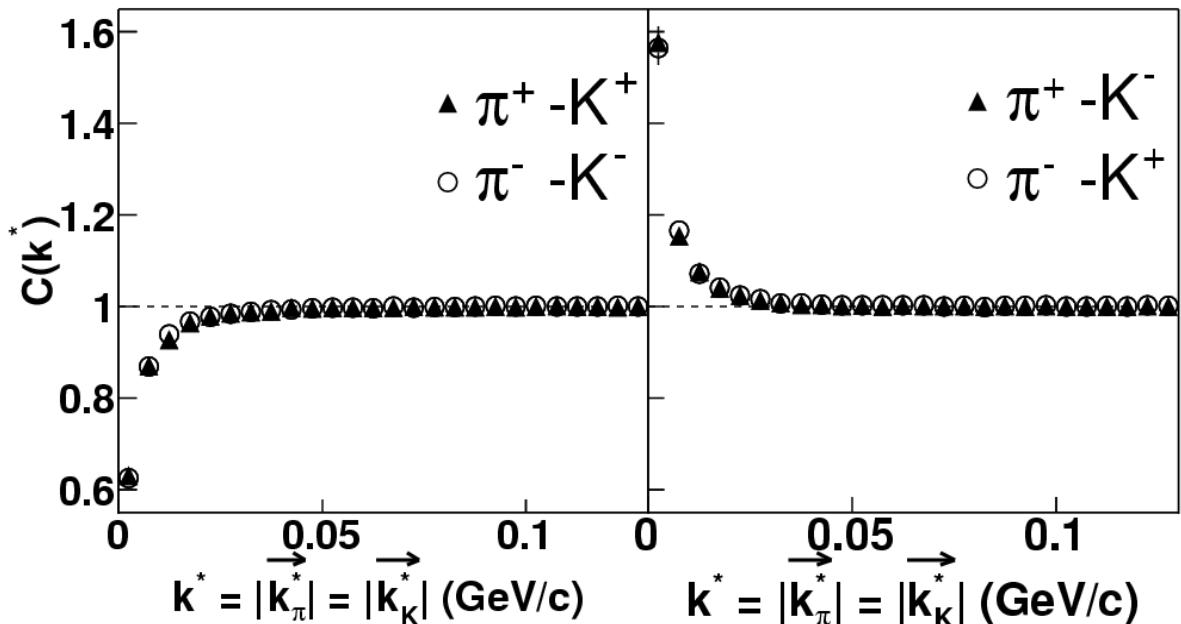
CF & DR for $\pi^- K^+$ and $\pi^+ K^-$



CF & DR for $\pi^- K^+$ and $\pi^+ K^-$

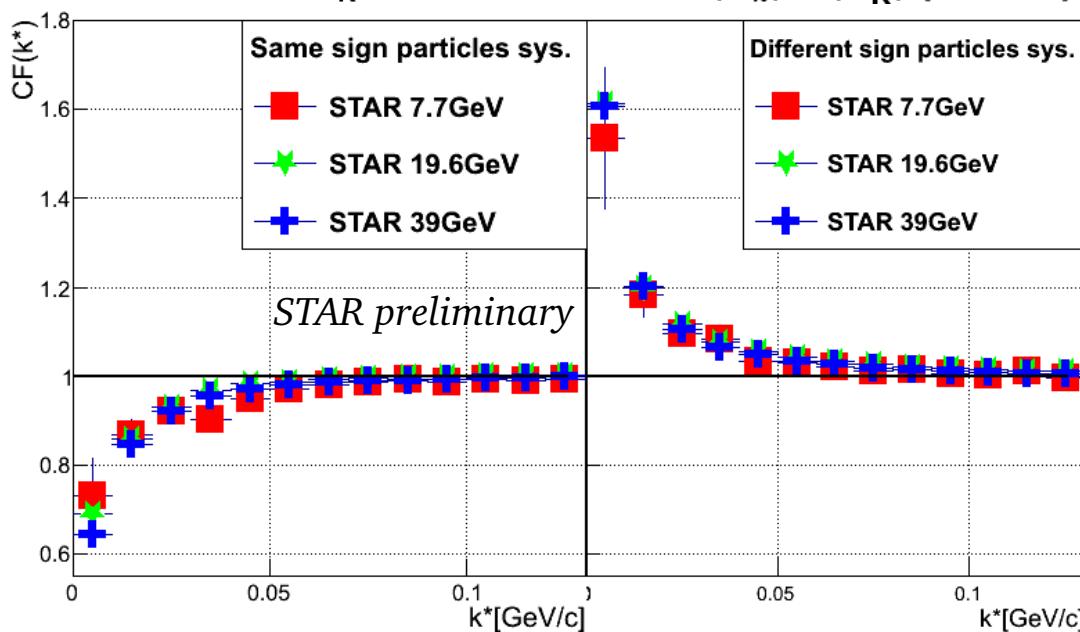


Comparison of Au+Au data for $\sqrt{s_{NN}}=(130, 39, 19.6, 7.7)\text{GeV}$



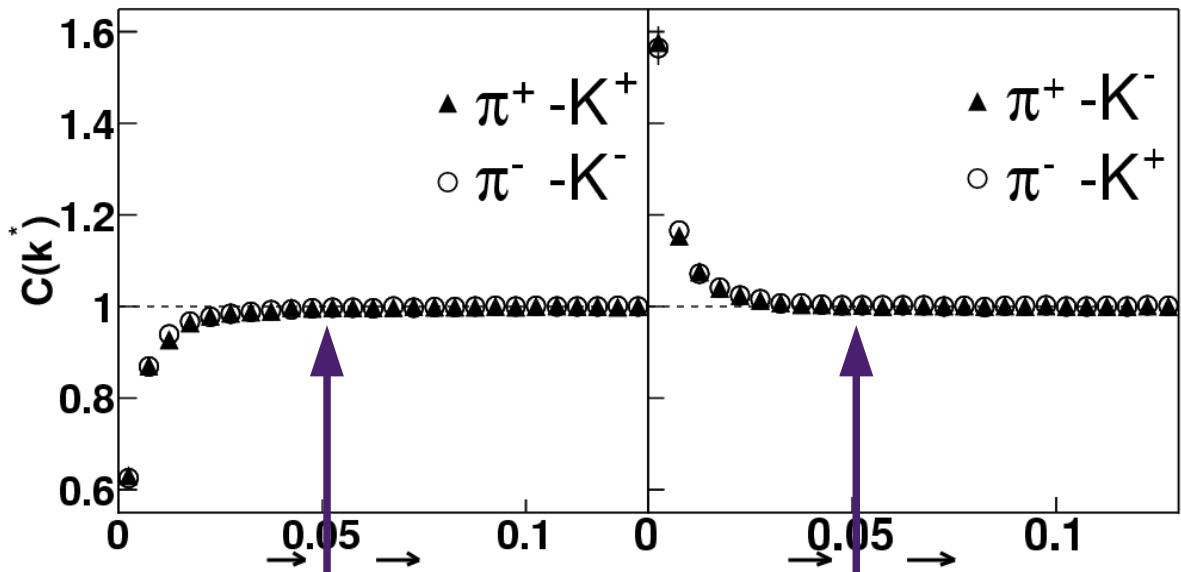
$\text{Au}+\text{Au}$ collision at $\sqrt{s_{NN}}=130\text{GeV}$ (central)

Phys. Rev. Lett. 91 (2003) 262302



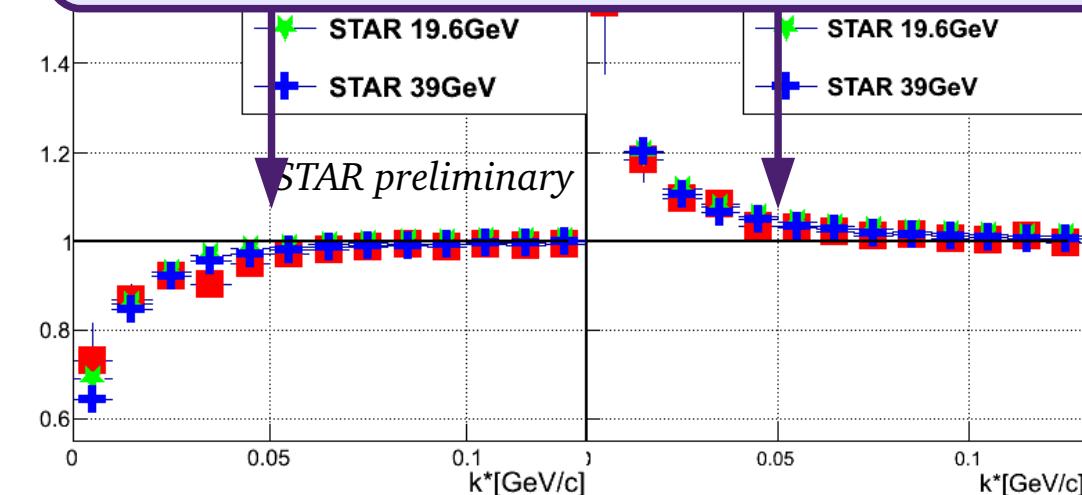
$\text{Au}+\text{Au}$ collision at
 $\sqrt{s_{NN}}=(39, 19.6, 7.7)\text{GeV}$
(minimum bias)

Comparison of Au+Au data for $\sqrt{s_{NN}}=(130, 39, 19.6, 7.7) GeV$



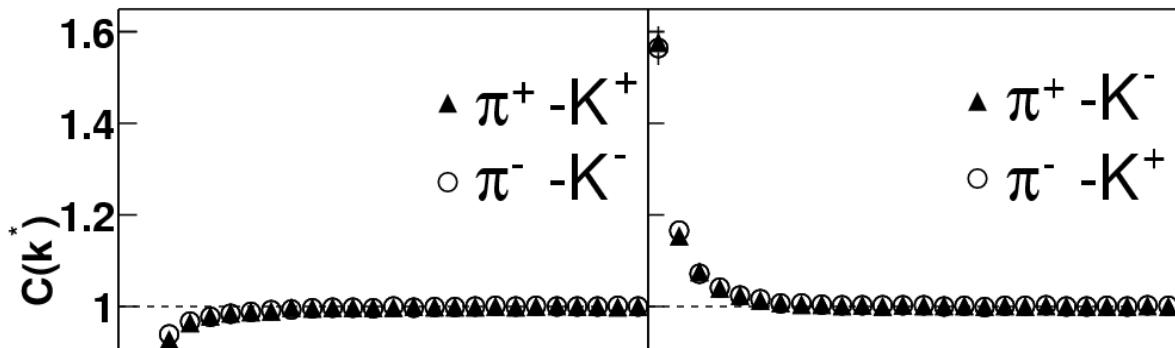
Au+Au collision at $\sqrt{s_{NN}}=130\text{GeV}$
(central)

Correlation function for $\sqrt{s_{NN}}=130\text{GeV}$ is narrower than CF for $\sqrt{s_{NN}}=(39, 19.6, 7.7)\text{GeV}$,
so the source size for $\sqrt{s_{NN}}=130\text{GeV}$ is bigger.



Au+Au collision at
 $\sqrt{s_{NN}}=(39, 19.6, 7.7) GeV$
(minimum bias)

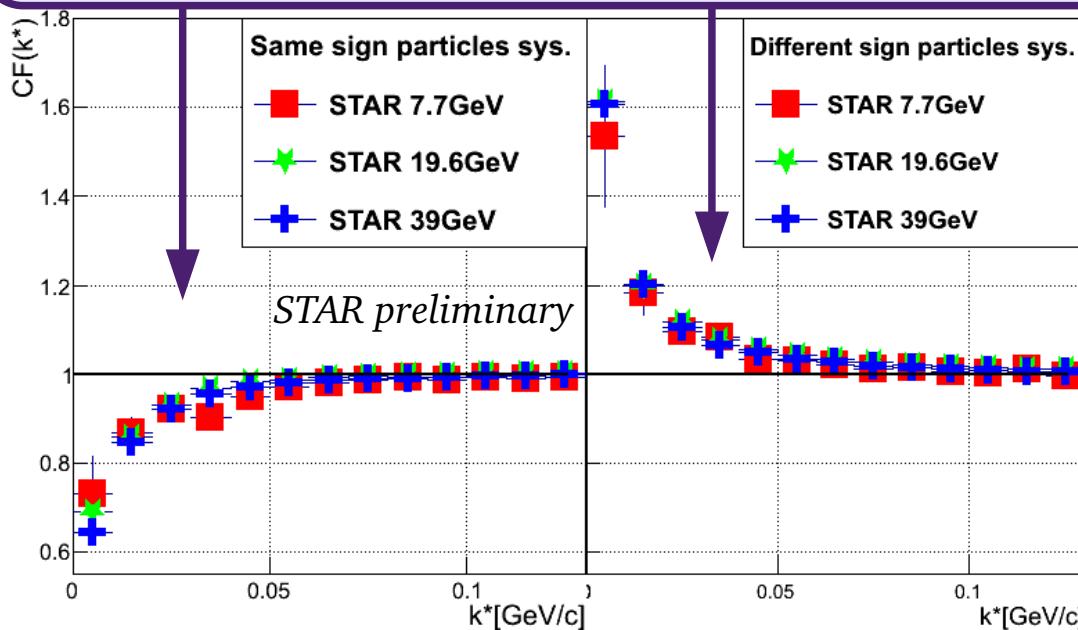
Comparison of Au+Au data for $\sqrt{s_{NN}}=(130, 39, 19.6, 7.7)\text{GeV}$



Au+Au collision at $\sqrt{s_{NN}}=130\text{GeV}$
(central)

Trends of correlation functions for all energies are the same.

The correlation functions for $\sqrt{s_{NN}}=(39, 19.6, 7.7)\text{GeV}$ looks the same within error bars.

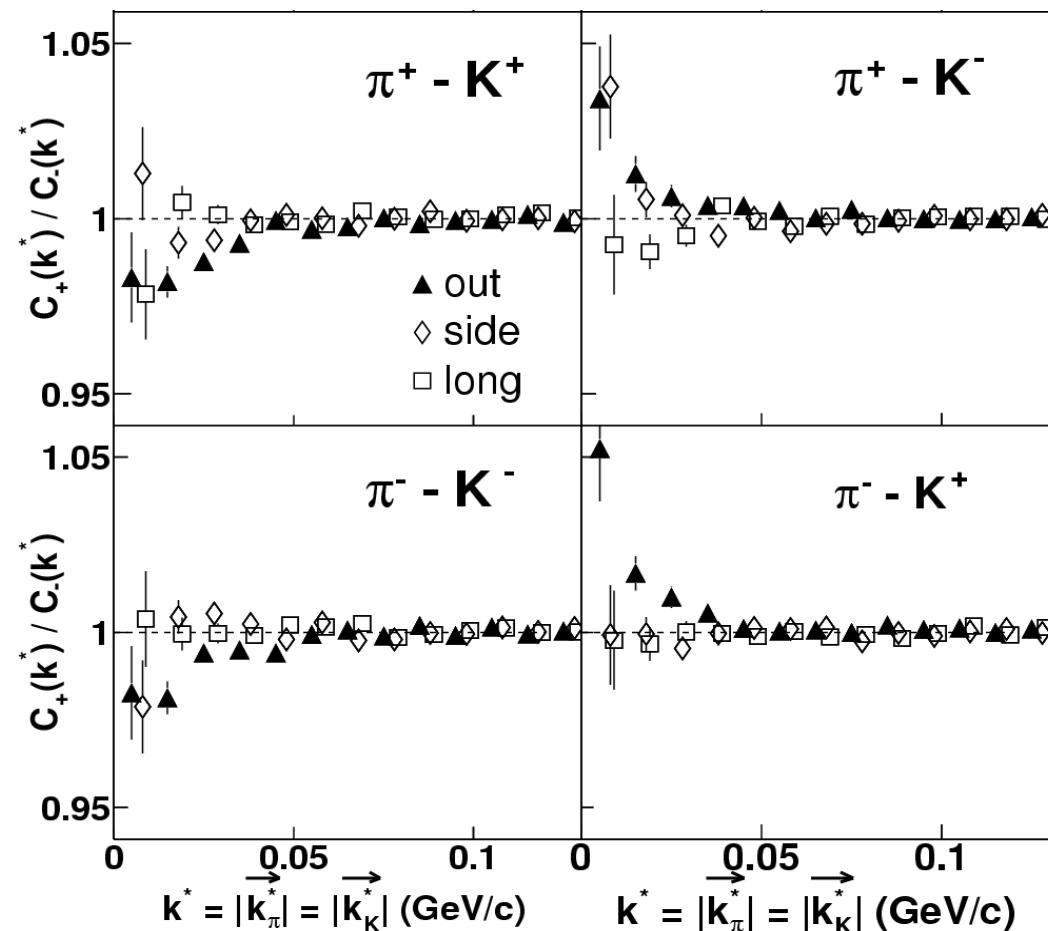


Au+Au collision at
 $\sqrt{s_{NN}}=(39, 19.6, 7.7)\text{GeV}$
(minimum bias)

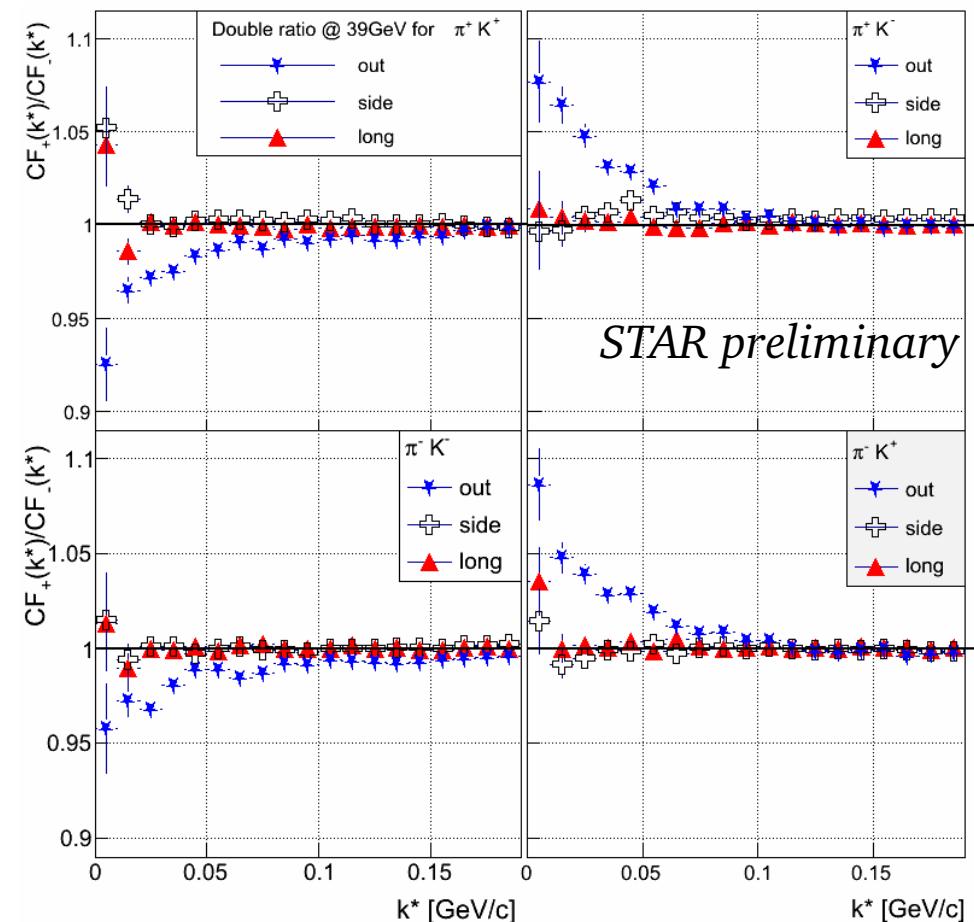
Comparison of Au+Au data for $\sqrt{s_{NN}}=130\text{GeV}$ and $\sqrt{s_{NN}}=39\text{GeV}$

Au+Au collision at $\sqrt{s_{NN}}=130\text{GeV}$
(central)

Phys. Rev. Lett. 91 (2003) 262302



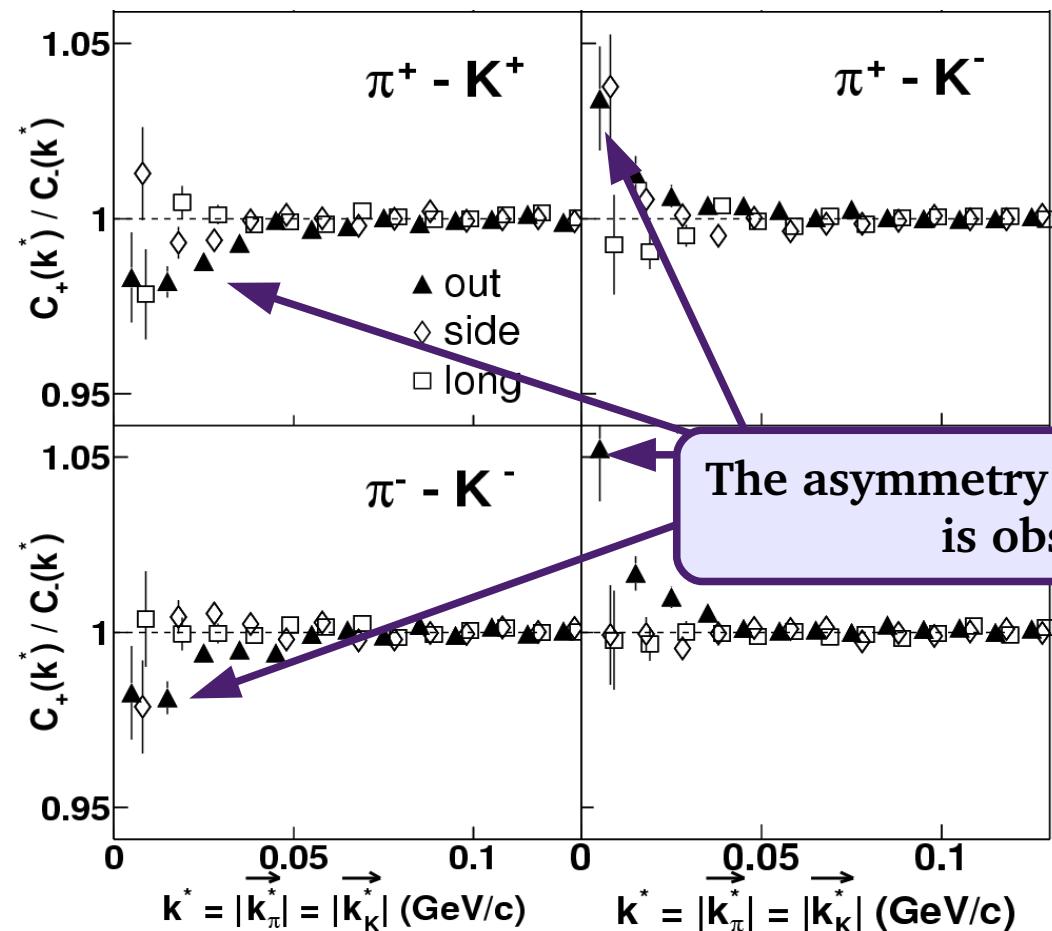
Au+Au collision at $\sqrt{s_{NN}}=39\text{GeV}$
(central)



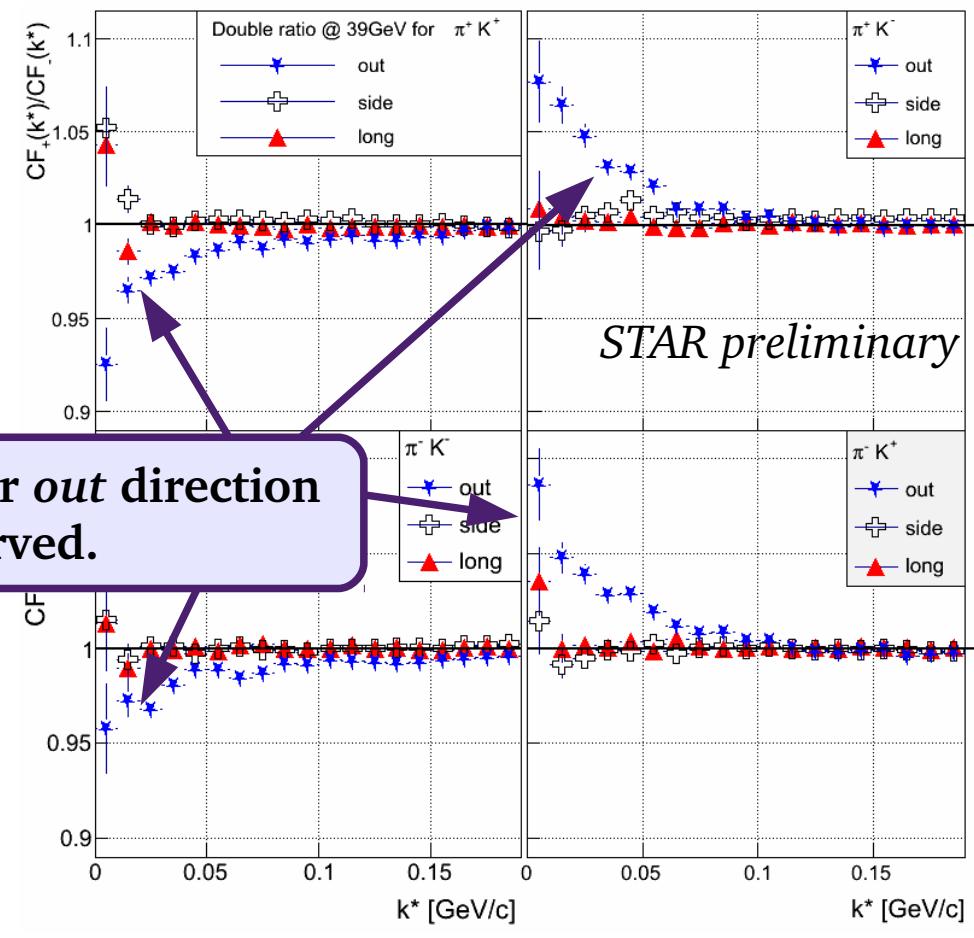
Comparison of Au+Au data for $\sqrt{s_{NN}}=130\text{GeV}$ and $\sqrt{s_{NN}}=39\text{GeV}$

Au+Au collision at $\sqrt{s_{NN}}=130\text{GeV}$
(central)

Phys. Rev. Lett. 91 (2003) 262302

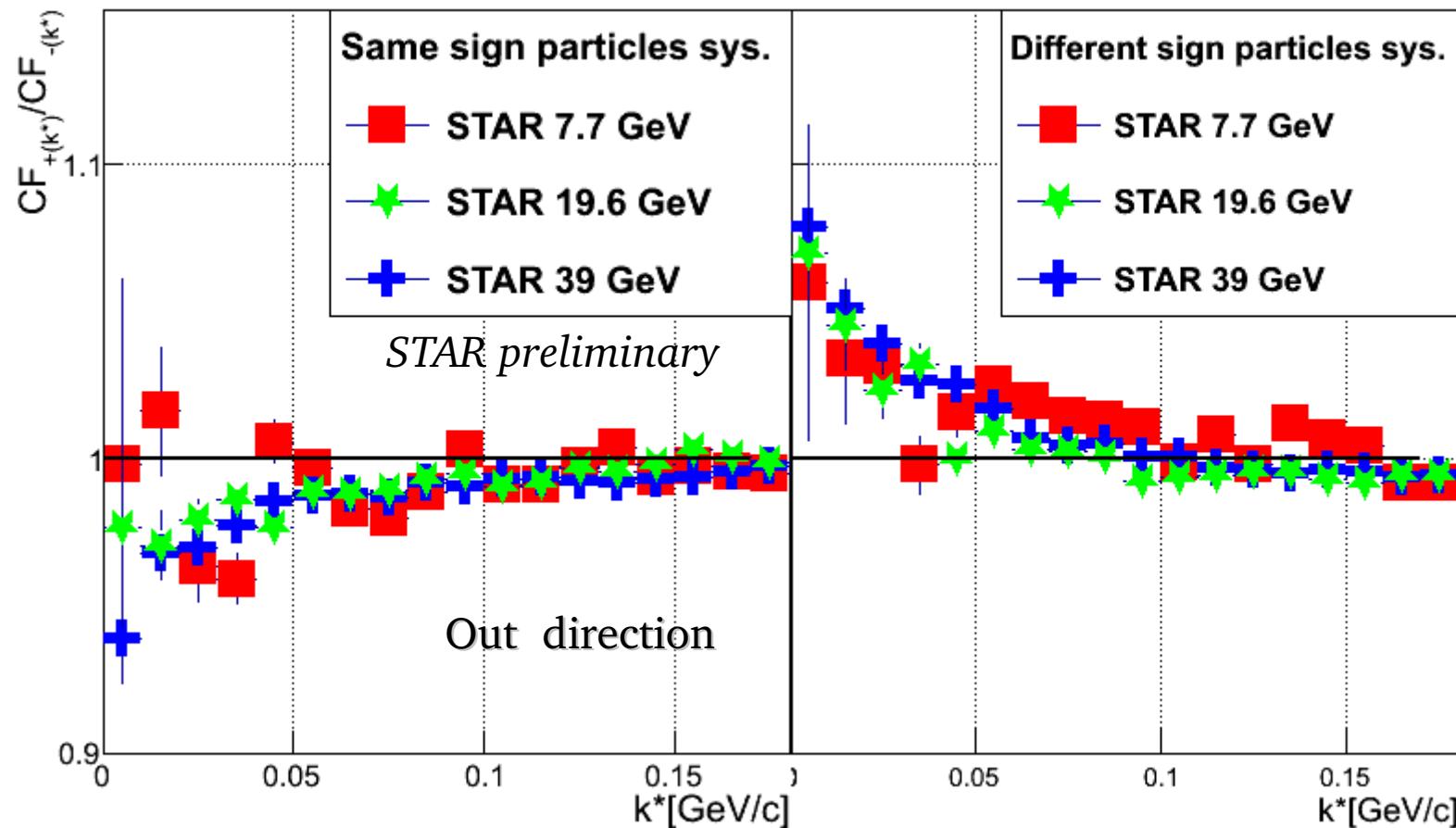


Au+Au collision at $\sqrt{s_{NN}}=39\text{GeV}$
(central)



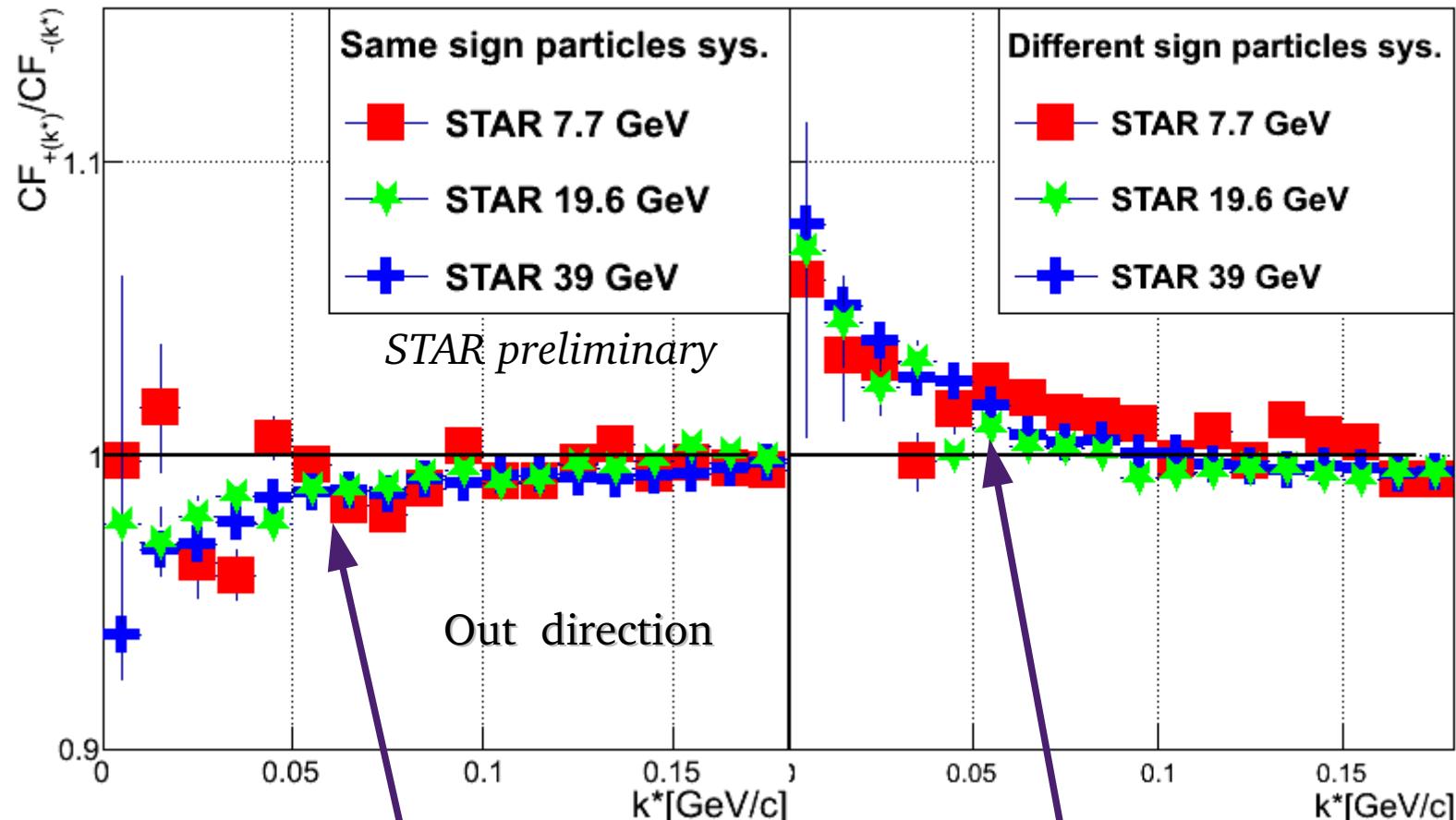
Comparison of Au+Au data for $\sqrt{s_{NN}} = (39, 19.6, 7.7) \text{ GeV}$

Minimum bias data.



Comparison of Au+Au data for $\sqrt{s_{NN}} = (39, 19.6, 7.7) \text{ GeV}$

Minimum bias data.



We observe the same tendency in asymmetry for all energies.

Summary

- Correlation functions and “double ratio” functions for minimum bias Au+Au collisions at $\sqrt{s_{NN}} = (39, 19.6, 7.7) \text{ GeV}$ are calculated.
- Trends of correlation functions and “double ratio” functions for Au+Au collisions at $\sqrt{s_{NN}} = (130, 39, 19.6, 7.7) \text{ GeV}$ are the same – results are consistent.
- Source size is bigger in Au+Au collisions at $\sqrt{s_{NN}} = 130 \text{ GeV}$ than at $\sqrt{s_{NN}} = (39, 19.6, 7.7) \text{ GeV}$.
- Due to limited statistics we can only say, that source size in Au+Au collision at $\sqrt{s_{NN}} = 7.7 \text{ GeV}$ and $\sqrt{s_{NN}} = 19.6 \text{ GeV}$ are close to source size in Au+Au collision at $\sqrt{s_{NN}} = 39 \text{ GeV}$.
- Pions are emitted closer to the system's center or/and later than kaons in systems created in collisions for $\sqrt{s_{NN}} = (130, 39, 19.6, 7.7) \text{ GeV}$.

Thank you for your attention!