

Space Charge Solver in pyORBIT compared to those in PATRIC and Madx

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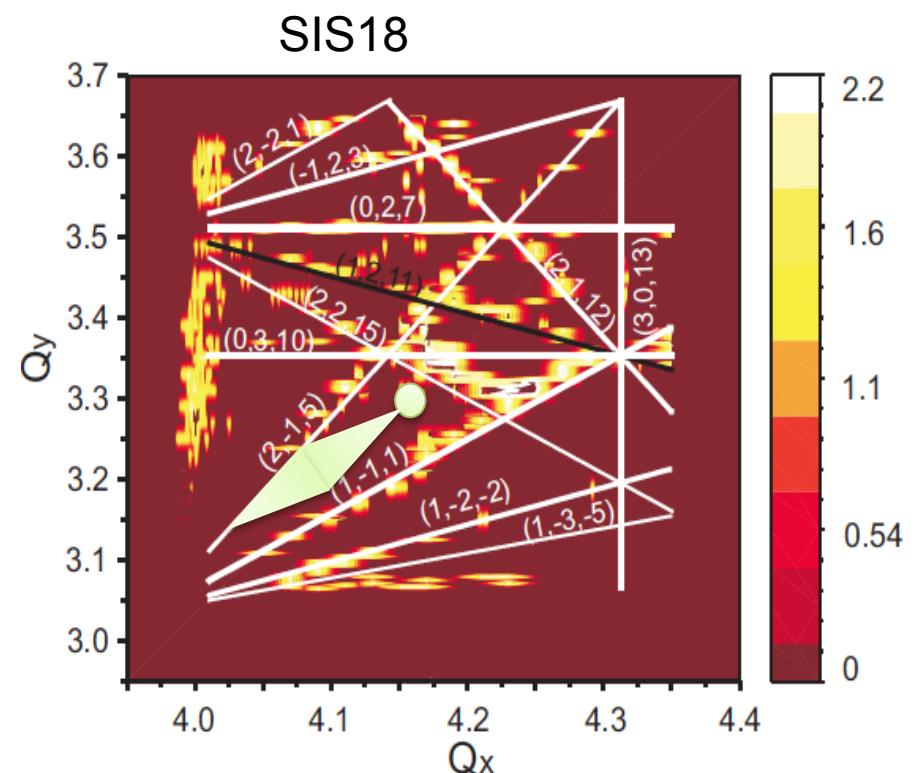
Outline

- Space charge (SC) effects in SIS18/100
- SC Solver
 - Frozen SC Solver (Madx)
 - PIC Solver
 - Longitudinal SC (pyorbit)
 - Transversal SC
 - PATRIC
 - pyORBIT
- SC matching
 - Transversal and longitudinal
- Tune footprint for Madx, pyORBIT and PATRIC
- Summary and Outlook

SIS18 Space charge effects

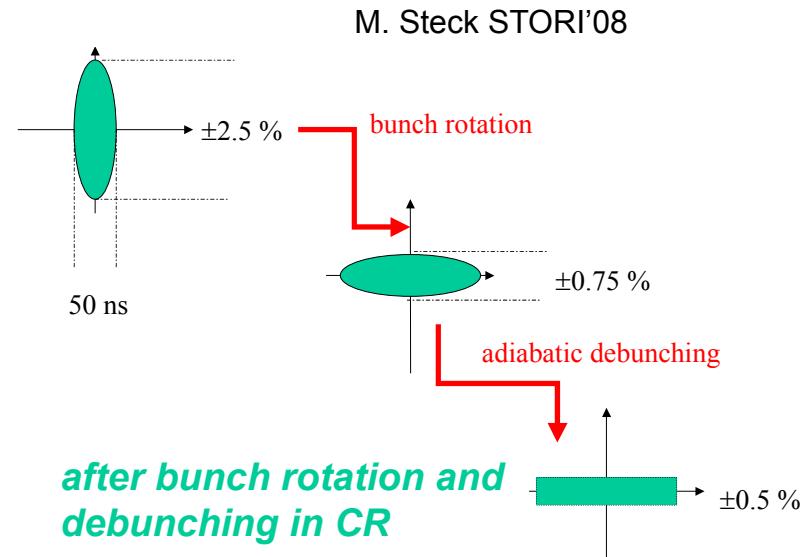
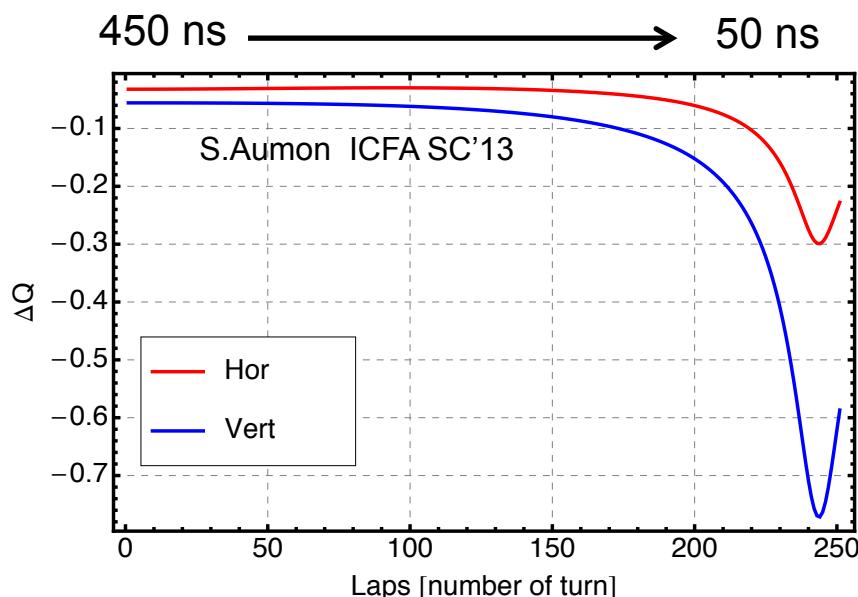
- Intensity goal: The space charge limit of $\Delta Q_y = 0.5$
- Injection energy: 11.4 MeV/u ($\beta_0 = 0.155$), Emittances: $\epsilon_{x,y} = 150 / 50 \text{ mm mrad}$
- Space charge induced voltage reduction by 40%
- Dual rf bucket with flattened bunch profile
- Codes demands:
 - Resolution of **synchrotron motion** (long bunches) -> 2.5 D SC solver
 - **Dual rf**
 - Matching routines for longitudinal and transversal space charge
- Studies:
 - Long term simulation after MTI
 - Comparison to frozen space charge simulation results

$$\Delta Q_y^{sc} = -\frac{2NZ^2g_f}{\pi A \beta_0^2 \gamma_0^3 B_f (\varepsilon_y + \sqrt{\varepsilon_y \varepsilon_x})}$$



SIS100 Bunch compression

- Intense short ions beam required by experiments for **plasma physics** and **exotic elements productions**
-> 50 ns short U²⁸⁺ beam with 5×10^{11} particles per bunch at 1.5 GeV/u
- Space charge affects the optics
- Large transverse space charge tune shift



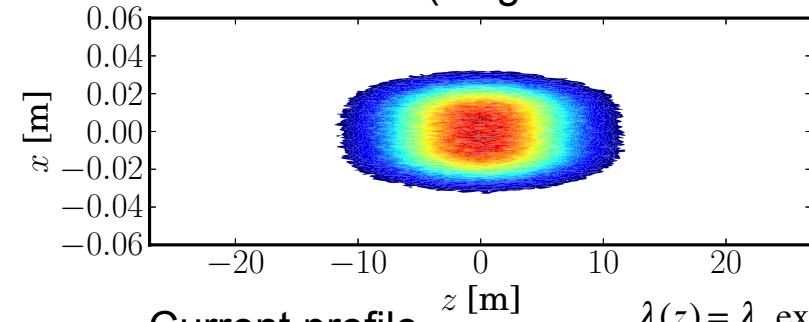
- Studies:
- Space charge effect on the transverse beam distribution (2D)
- Effects during bunch compression (magnets error, resonances etc..)

PATRIC: Space charge tune shift

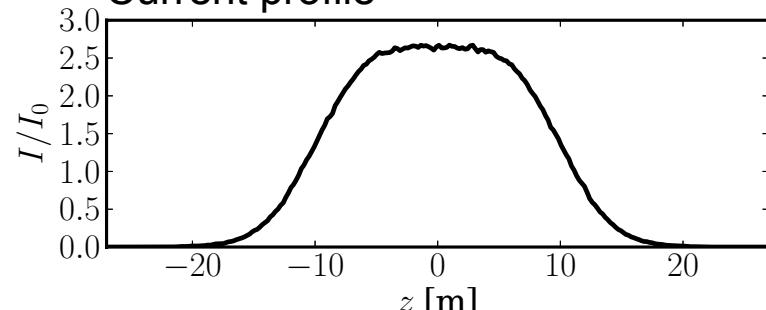


- ✓ Matched 3D bunch distribution (in a dual rf bucket)

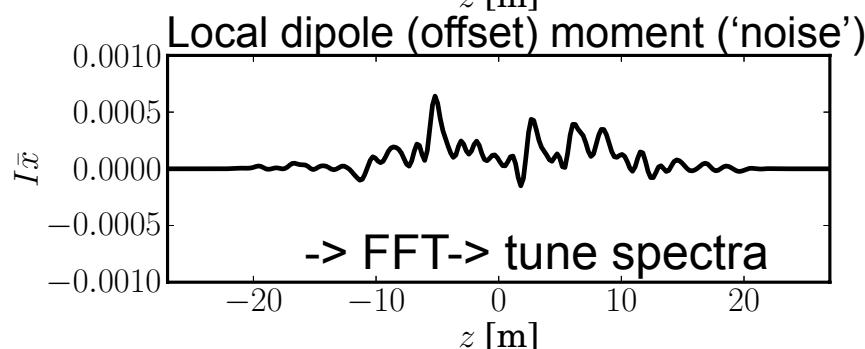
Bunch distribution (longitudinal-horizontal)



Current profile

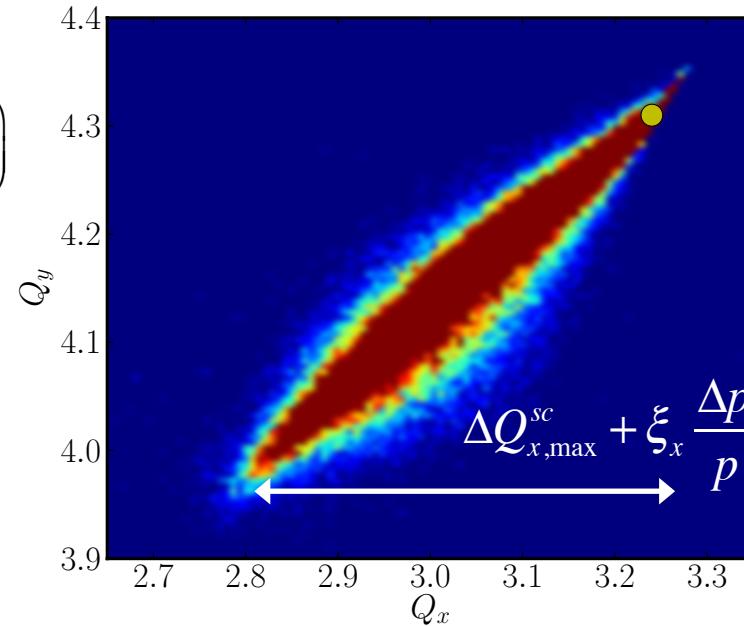


$$\lambda(z) = \lambda_0 \exp\left(-\frac{z^4}{2\sigma_l^4}\right)$$



-> FFT-> tune spectra

Tune footprint with space charge (SIS18)



O. Boine-Frankenheim, ICFA SC 2013

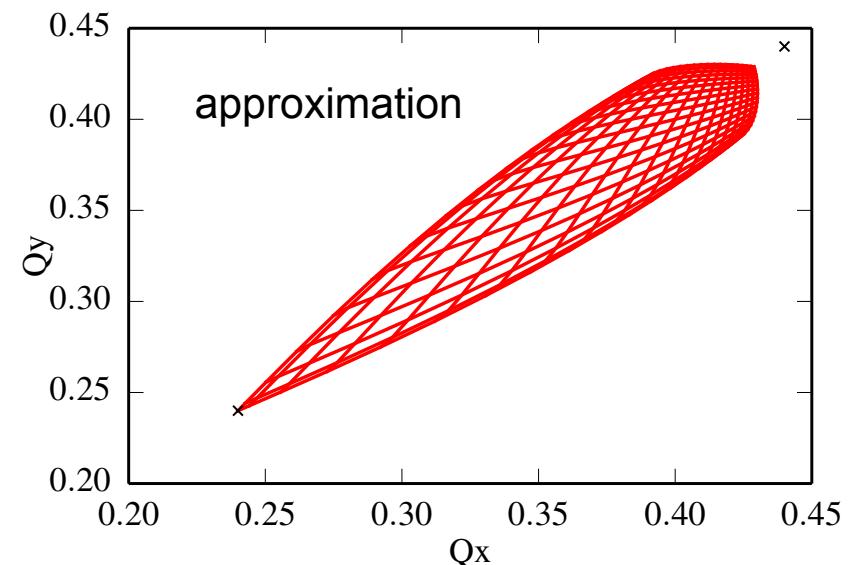
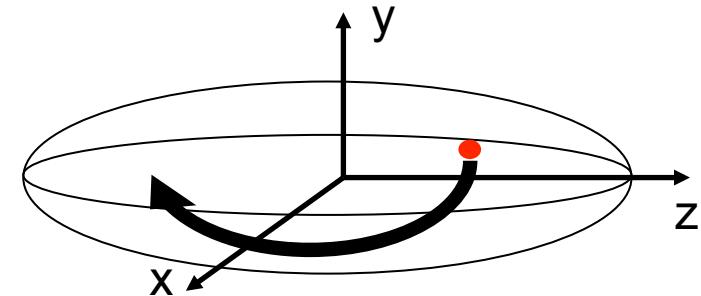


Frozen space charge

- The solution of Poisson equation for a 2D Gaussian can be calculated analytical

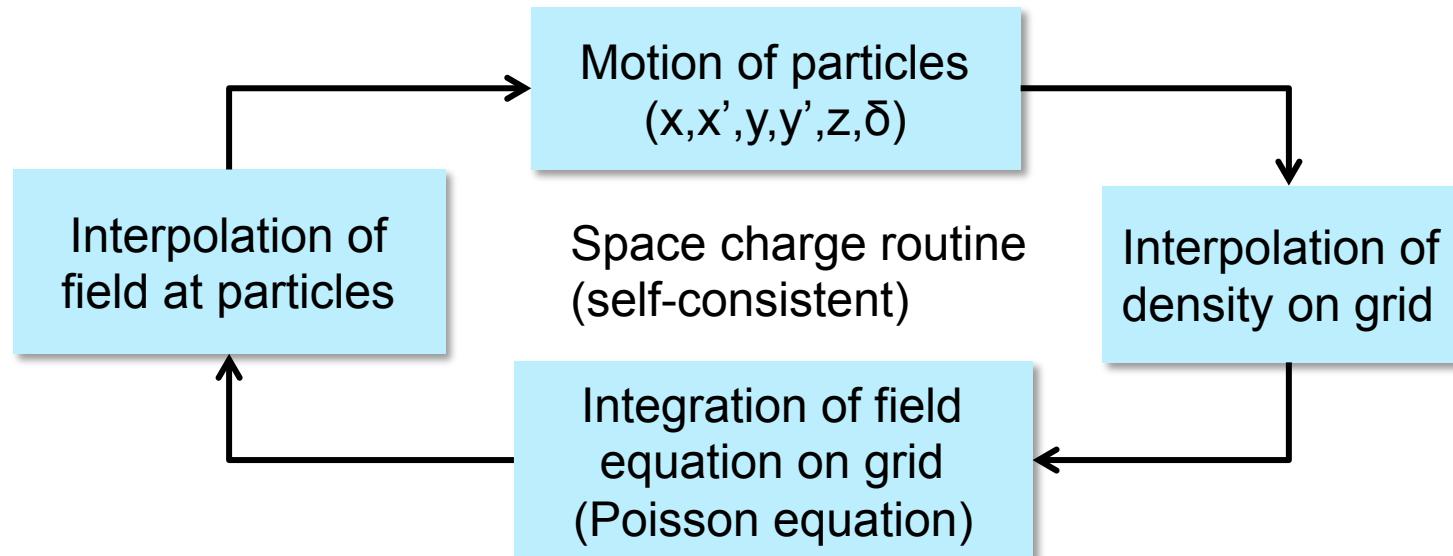
$$\nabla^2 \Phi = \frac{\rho}{\epsilon_0 \gamma^2} \quad \rho \sim e^{-\left(\frac{x^2}{2\sigma_x^2} + \frac{y^2}{2\sigma_y^2}\right)}$$

- Also the space charge tune spread as a function of the particle amplitude
- Tracking: The kick acting on the particle is computed from the electric field (analytical)
- Sometimes during tracking simulations the beam intensity and Gaussian beam size are adapted



See e.g. A. Burov, et. al., Transverse instabilities of coasting beams with space charge, Phys. Rev. ST-AB (2009)
M. Bassetti, et. al., Closed expression for the electrical field of a two-dimensional Gaussian charge, CERN-ISR-TH/80-06

PIC algorithms: Space charge routine



- The particle-to-particle and particle-to-environment interactions are calculated with PIC (Particle In Cell) algorithms
- Space charge forces or potential are obtained by solving the Poisson equation (2D: Solve the electric field equation)
- Solving the integral from on grid with FFT and boundary condition for arbitrary beam distribution

$$\nabla^2 \Phi = \frac{\rho}{\epsilon_0 \gamma^2}$$

$$\frac{\partial E_x}{\partial x} + \frac{\partial E_y}{\partial y} = \frac{\rho}{\epsilon_0}$$

$$\Phi(\vec{r}) = \int \frac{\rho(\vec{r} - \vec{r}')}{|\vec{r} - \vec{r}'|^2} d\vec{r}'$$

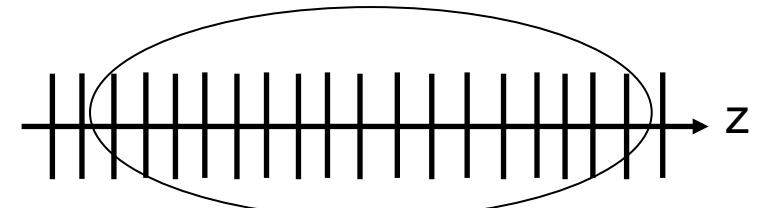
PyOrbit: Longitudinal impedance + space charge



- The longitudinal impedance is represented by its harmonic content in terms of the fundamental ring frequency

Particle are binned longitudinally (1D)

FFT of binned distribution
 $Z_n^{sc} + Z_n^{extern}$



Calculate and add kicks for each particles

```
void LSpaceChargeCalc::trackBunch(Bunch* bunch){  
    ...  
    for(int n = 1; n < nBins / 2; n++){  
        realPart = std::real(_zImped_n[n]);  
        imagPart = std::imag(_zImped_n[n]) + zSpaceCharge_n;  
        _z[n] = n * sqrt(realPart * realPart + imagPart * imagPart);  
        _chi[n] = atan2(imagPart, realPart);}  
    ...  
    for(int n = 1; n < nBins / 2; n++){  
        cosArg = n * (angle + OrbitConst::PI) + _fftphase[n] + _chi[n];  
        kick += 2 * _fftmagnitude[n] * _z[n] * cos(cosArg);}  
    return kick;  
}
```

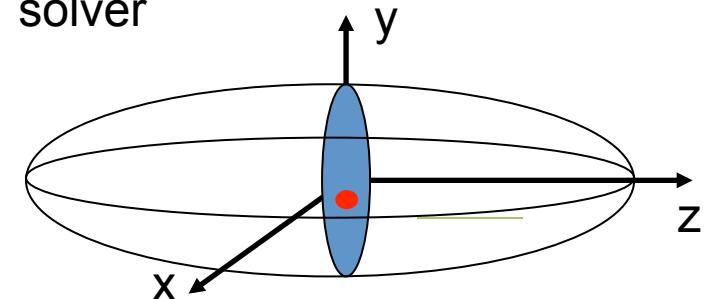
- Similar approach in PATRIC

Source: Talks by S. Cousineau and J.A. Holmes, ORNL

PyOrbit: Transverse Space Charge Solver



- Solving Poisson equation on 2D grid with fast FFT solver
 - Particles are binned in 2D rectangular grid (momentum-conserving interpolation scheme)
- Grid is adapted to beam size
- Kicks are weighted by local longitudinal density (non-uniform distribution)



```
void SpaceChargeForceCalc2p5D::trackBunch(Bunch* bunch, double length){  
    ...  
  
    forceSolver->findForce(rhoGrid, forceGridX, forceGridY);  
    forceGridX->interpolateBilinear(x,y,fx);  
    forceGridY->interpolateBilinear(x,y,fy);  
    ...  
  
    for(i=0, bunchSize,i++){  
        Lfactor = zGrid->getValue(z) * factor;  
        bunch->xp(i) += fx * Lfactor;  
        bunch->yp(i) += fy * Lfactor;  
    }  
}
```



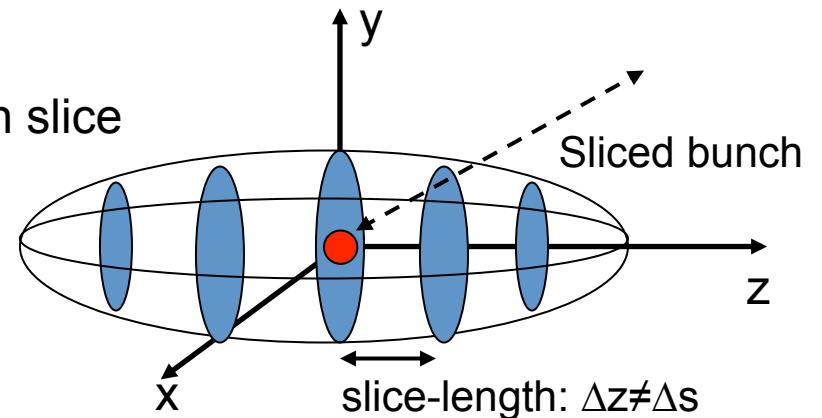
Source: Talks by S. Cousineau and J.A. Holmes, ORNL and pyORBIT source code

PATRIC: Transverse Space Charge Solver



- Bunch is sliced n times
- The 2D space charge field is computed for each slice
- Grid is static

$$\frac{\partial E_x}{\partial x} + \frac{\partial E_y}{\partial y} = \frac{\rho(x,y,\{z,s_m\})}{\epsilon_0} \quad (\text{fast 2.5D Poisson solver})$$



```

Pics.gatherXYZ(charge*qe/rho_xyz.get_dz(),rho_xyz); # 3D Grid

// send and receive density ghost grids to neighbor
MPI_Isend(rho_xyz.get_ghostl(),...)
MPI_Recv(rho_xy_tmp.get_grid(),...)
...

poisson_xyz(Ex3,Ey3,rho_xyz,gf1); # 2D Poisson solver for slice

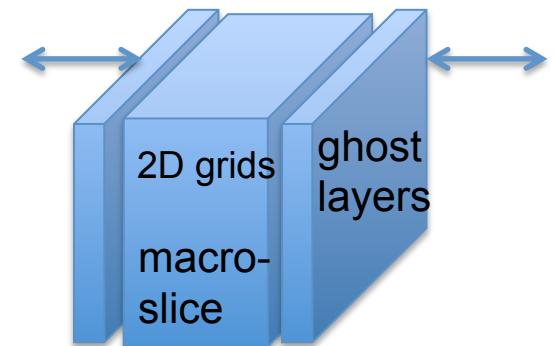
// send and receive efield ghost grids to neighbor slices
MPI_Isend(Ex3[0].get_grid(),...)
MPI_Recv(Ex3.get_ghostr(),...)
...

Pics.kick(Ex3,Ey3,ds)

```

- Corrected interpolation

MPI send/receive to neighbor slices



➤ Grid slices “feel” all other slices

Oliver's talk, ICAP 2012

Trans. SC matching in pyORBIT and PATRIC

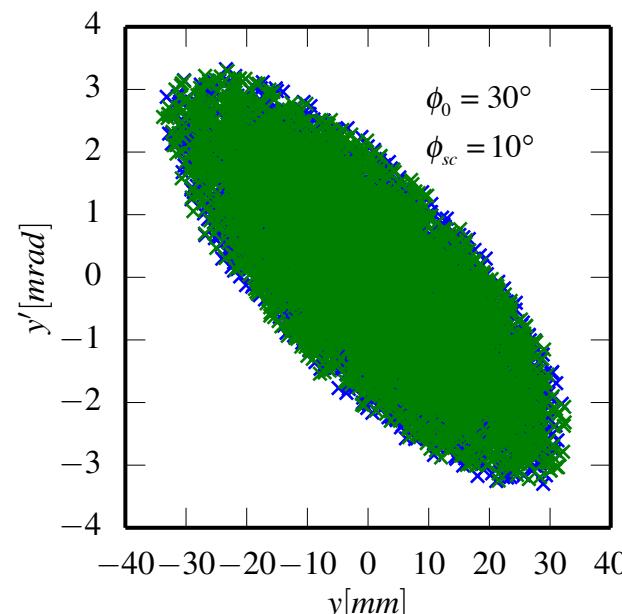
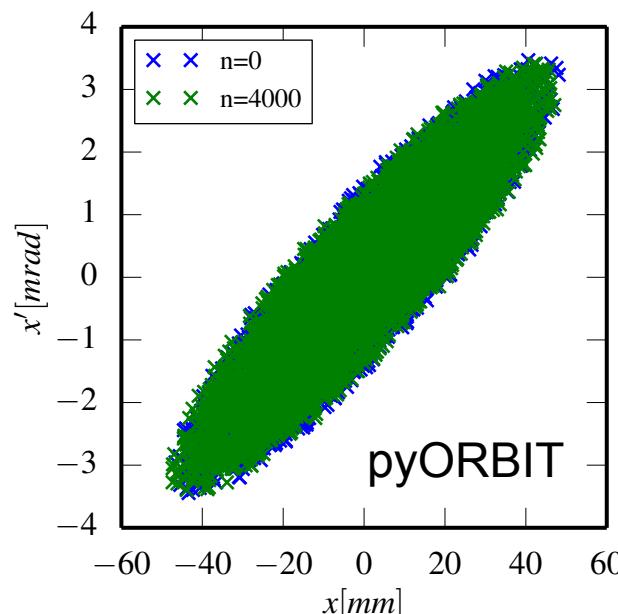
See e.g. Transverse matching with space charge (Venturini, Reiser, PRL 1998)

- Solving of the two coupled beam envelopes equation with space charge

$$X'' + k_x X - \frac{2K}{X+Y} - \frac{\epsilon_x^2}{X^3} = 0$$

$$Y'' + k_y Y - \frac{2K}{X+Y} - \frac{\epsilon_x^2}{Y^3} = 0$$

```
envx,envxs,envy,envys,Dx,Dxs,s = match_root(lattice,emitx,emity,sigma_p,Ksc)
phasex,phasey = phase_advance(envx,envy,Dx,emitx,emity,dp,s) # check
```



See e.g.: Oliver's presentation tomorrow

Long. SC matching for arbitrary rf wave forms



See e.g. Hofmann-Pedersen, The bunch distribution can be matched to arbitrary rf bucket forms (1979)
O. Boine-F., *rf barrier compression with space charge*, Phys. Rev. ST-AB (2010)

'Hamiltonian':

$$H = \frac{v^2}{2} - \frac{qV_0}{m^* L} Y(z)$$

Potential function:

$$Y(z) = \int_0^z V(z) dz \quad V(z) = V_{rf}(z) + V_{sc}(z)$$

Single rf wave:

$$V_{rf}(z) = V_0 \sin(hz / R)$$

Dual rf wave:

$$V_{rf}(z) = V_0 \left(\sin(hz / R) + \frac{1}{2} \sin(2hz / R) \right)$$

Elliptic bunch distribution:

$$f(z, v) = f_0 \sqrt{1 - \frac{Y_{rf}}{Y_m} - \frac{v^2}{v_m^2}}$$

Potential at the bunch end: $Y_m = Y(z_m)$

Maximum velocity in the
bunch center:

$$\frac{v_m^2}{2} = \frac{qV_0}{m^* L} Y_m$$

Long. SC matching in pyORBIT and PATRIC

1. Set rf field and potential on 2D grid
($N_y = N_x = 2N_{\text{bin}}$)

```
machting.set_rf_dual(rf_field,Yrf_field,  
RFHNum,RatioVoltage)
```

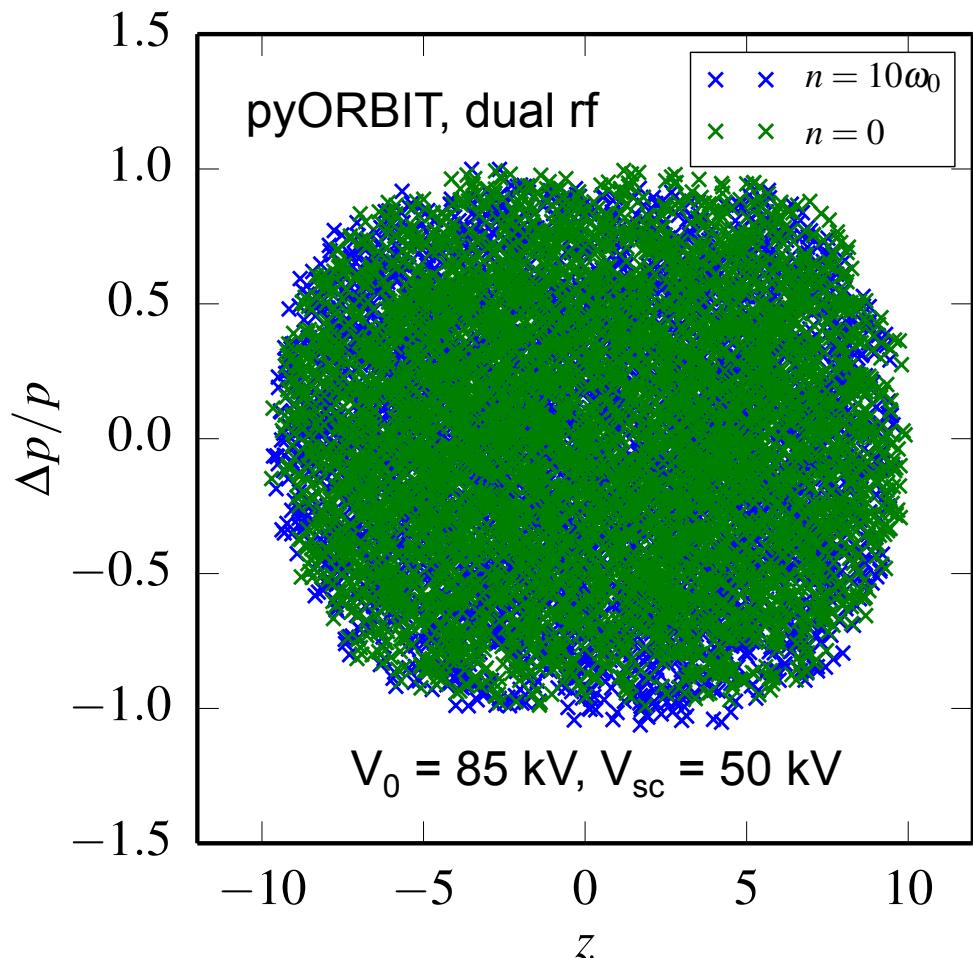
2. Determine matched voltage and beam distribution $f(z,v)$ on grid

```
V0, distf =  
machting.match_elliptic(Yrf_field, distf,  
zm, dpBunch, Zs, intensity)
```

3. Generate random numbers sets x_i, v_i from $f(z_j, v_k)$

```
x,dp = machting.matchPIC(NPIC, distf)
```

Shown only 10% of computed macro-particles



-> also Gaussian distribution possible

Betatron tune calculation

- Average Phase advance (APA) method
 - Non instantaneous
 - Evaluation of phase difference for each particle after each turn (normalized coordinate system)
 - Lattice function must be known

- FFT method
 - Non instantaneous
 - FFT analysis of the average transvers oscillations
 - Resolution given by the number of average turns/cells

- One-Turn-Matrix (OTM) method
 - Quasi instantaneous
 - Reconstruction of OTM for each particle
 - Trace or eigenvalues of OTM gives fraction turn

$$\theta_{n,j} = \arctan \frac{x'_{n,j}}{x_{n,j}}$$

$$\tilde{Q}_j = \frac{\theta_{n-1} - \theta_n}{2\pi}$$

$$\Delta = \frac{1}{n}$$

$$FFT(x,y)$$

$$\Delta = \frac{1}{n}$$

$$x_j^{n+1} = M_j x_j^n$$

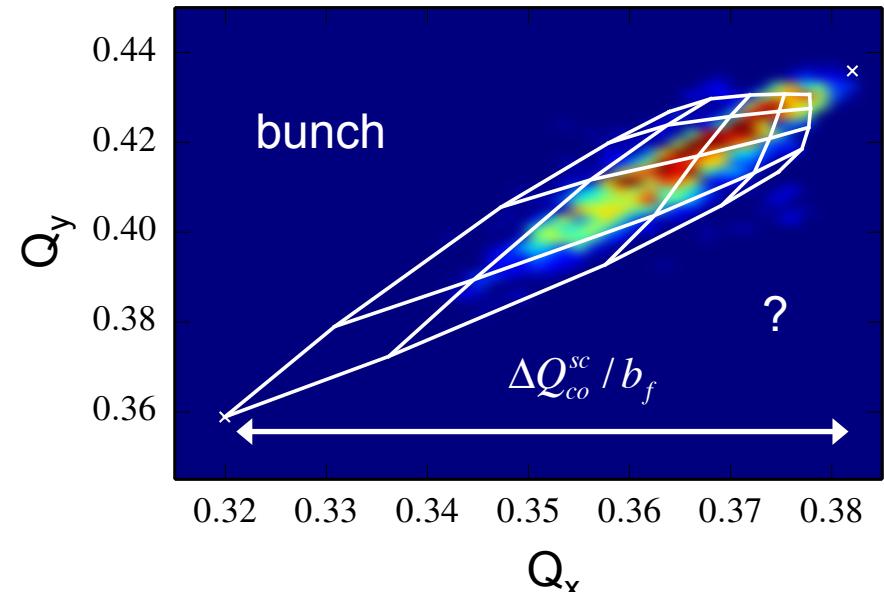
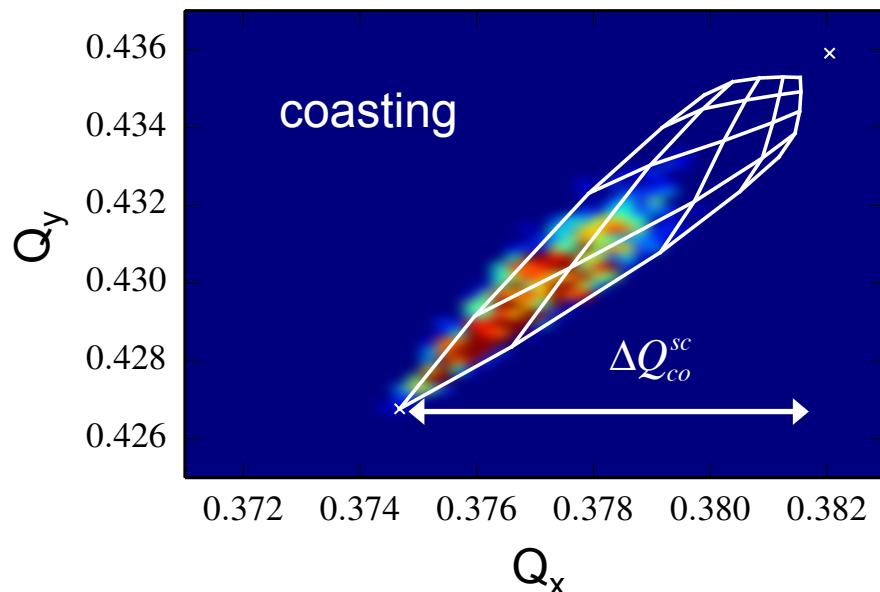
$$\tilde{Q}_j = \frac{1}{2\pi} \arccos \frac{\text{Tr}(M_j)}{2}$$

See e.g: Bartolini, R, Bazzani, A, Tune evaluation in simulations and experiments, CERN SL/95-84 (AP)
 Luccio, A, D'Imperio N., Eigenvalues of the One-turn Matrix, C-A/AP'126

Madx: Tune footprint

- Frozen space charge
- Simple FODO lattice with SC
- ✓ Coasting beam
- Bunch beam
 - SC tune spread is a factor of two too small
 - Meaning of the options in the input files are unclear

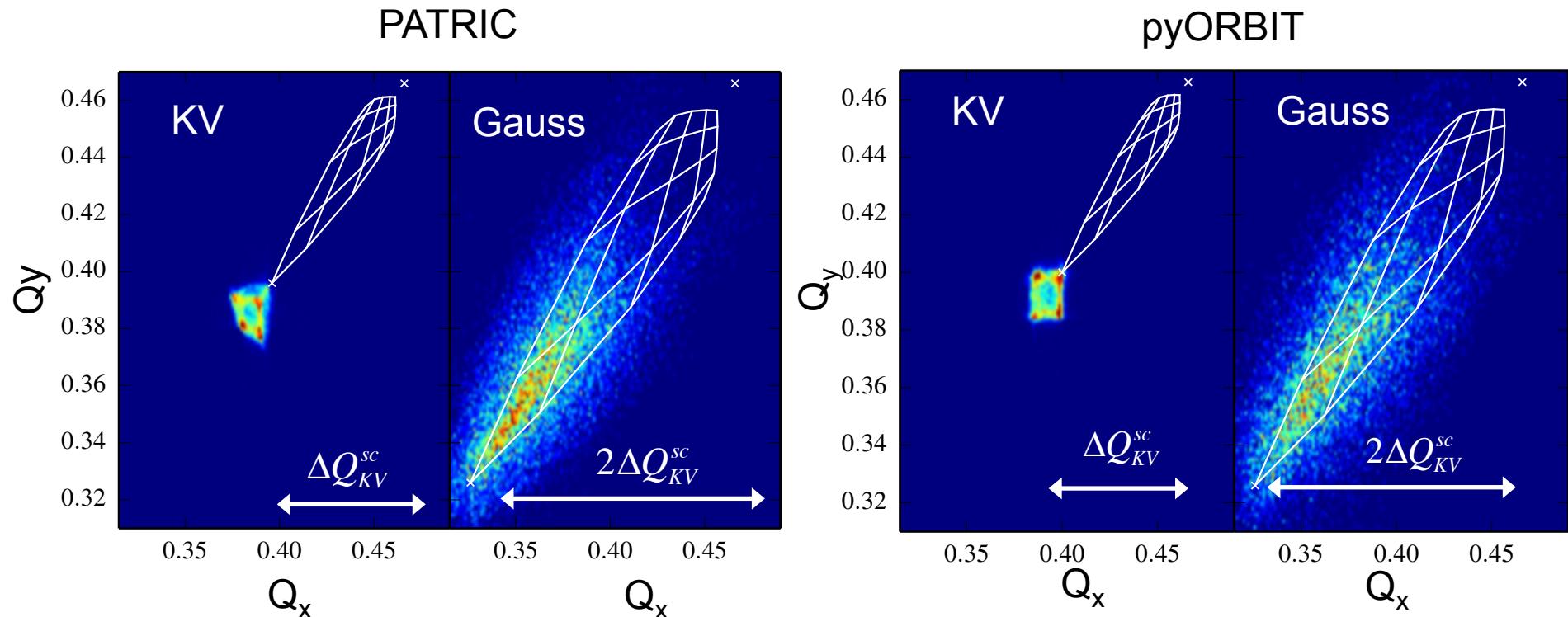
```
bb_i: beambeam, sigx:=sigxi, sigy:=sigyi, bbdirection=1;  
fodo0: line=(bb_1,dr,bb_2,qf,bb_3,dr,bb_4,dr,bb_5,qd,bb_6,dr);  
fodo_mult: line=(rfcav,12*fodo0);
```



Result from Stefan Sorge

PATRIC versus pyORBIT

- Tune footprint for 2D space charge solver
- In both codes constant focusing and coasting beam has been used
- Good agreement between the two codes



- In the early version of the presentation the tune footprint of pyORBIT simulation results with a Gauss distribution was distributed over the whole white diamond -> probable an effects of the lattice, constant focusing gives the expect tune footprint for all cases

- Space charge (SC) effects in SIS18/100
- SC Solver
 - Frozen SC Solver
 - Only Gaussian distribution
 - Solution of Poisson equation is solved analytical
 - PIC Solver
 - Arbitrary beam distribution and self-consistent
 - Poisson equation solved on Grid with FFT (different approaches)
- SC matching pyORBIT and PATRIC for trans. + long.
- Tune footprint for Madx, pyORBIT and PATRIC
 - More or less in good agreement
- SIS18/100: Bunch compression (“2D”)
- Long term “3D” simulation in SIS18 after MTI