

Exotics Searches at LHC

**Cigdem Issever
University of Oxford
ATLAS Experiment, CERN**

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Discussions with

- Henri Bachacou
- Bryan Lynn
- Christophe Grojean
- Glenn Starkman
- Steven Worm

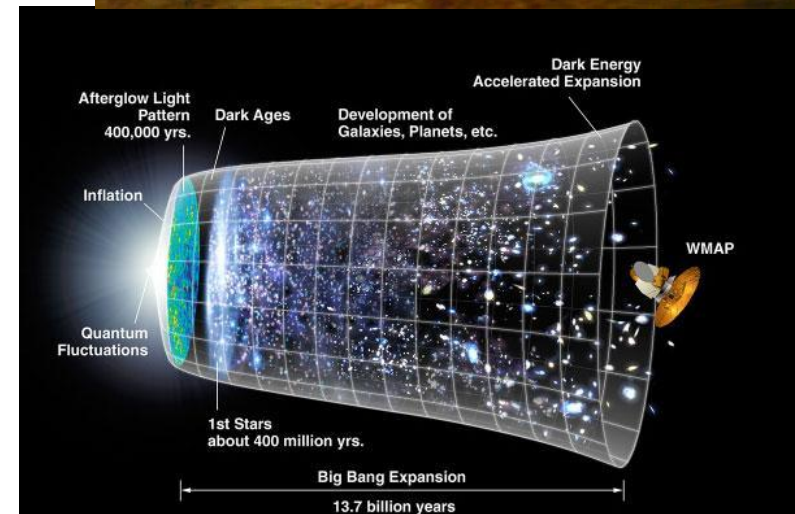


Outline....

- Why search for new physics?
- What are Exotics Searches?
- Examples of Searches

Why search for new physics?

- We are **reSEARCHers**
- We strive for **new understandings**
- Our goal: “create” **KNOWLEDGE**
We are “Wissenschaftler”



Inspiring

Humbling

FUN

and a LOT of work.....



Why look beyond *the Standard Model*?

■ Experimental Evidence

■ Non-baryonic dark matter (~27%)

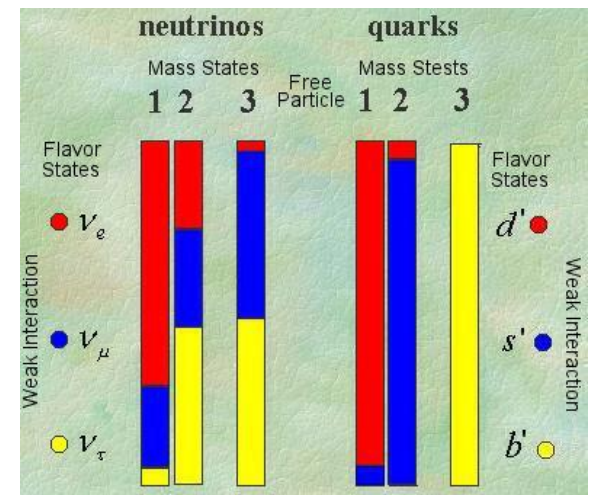
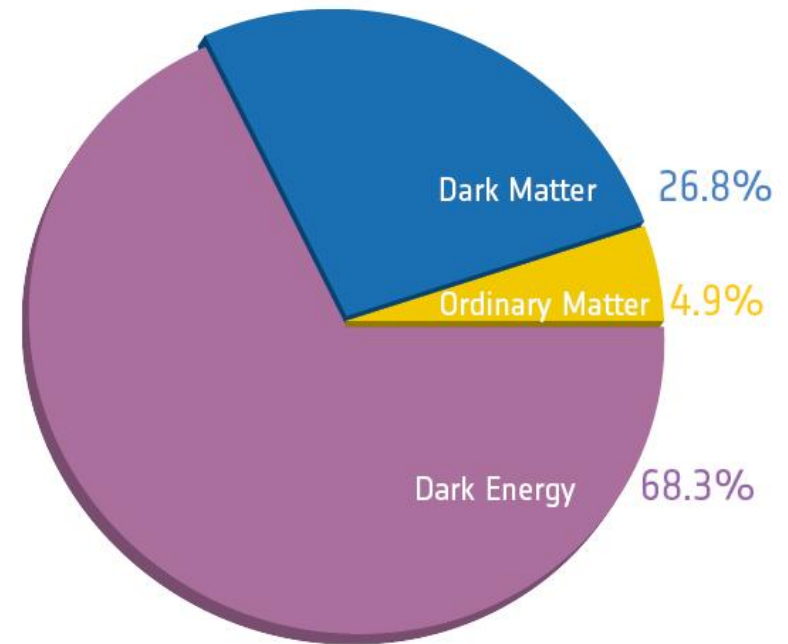
- Inferred from gravitational effects
- Rotational speed of galaxies
- Orbital velocities of galaxies in clusters
- [Cosmic Microwave Background](#)
-

■ Dark Energy (~68%)

- Accelerated Expansion of the Universe

■ Neutrinos have mass and mix

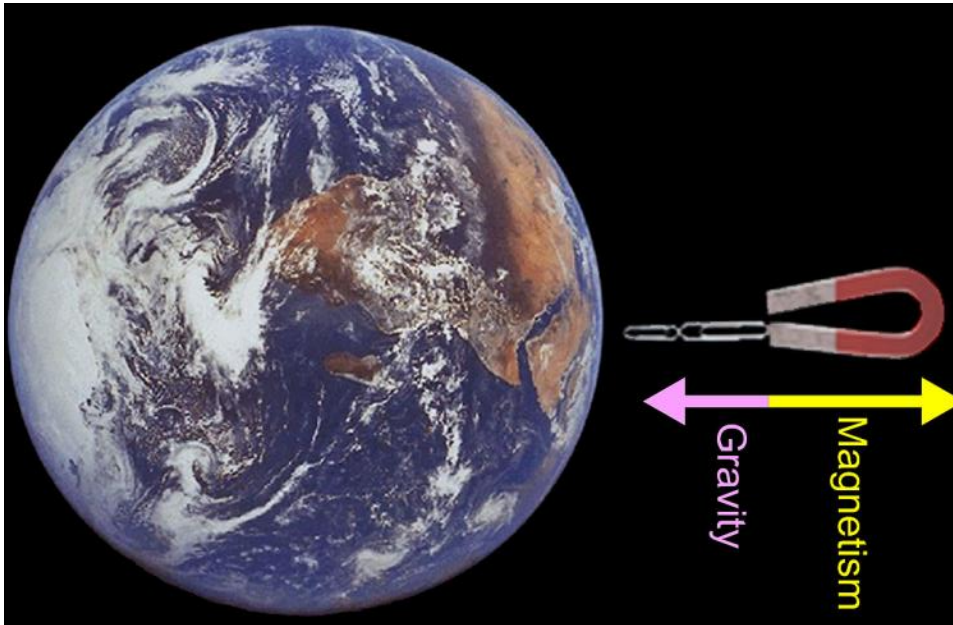
■ ...



Why look beyond the Standard Model?

■ Aesthetic/Theoretical Reasons

- Gravity is not included
- Family structure? Why 3?
- Hierarchy problem:
 - Why is gravity so weak?

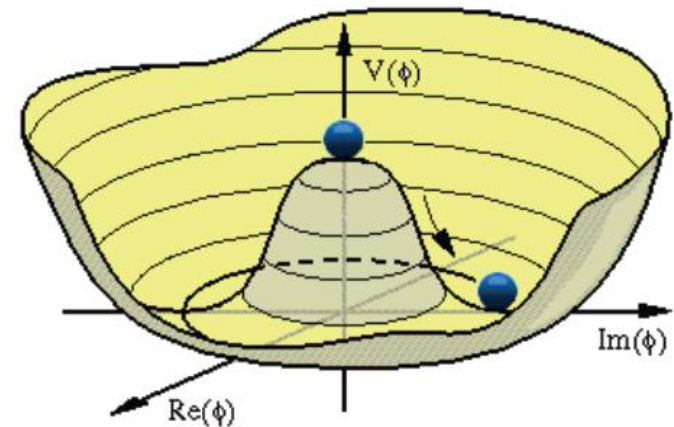


$$\begin{aligned}
 & -\frac{1}{2}\partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g_s^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e + \\
 & \frac{1}{2}ig_s^2(\bar{q}_i^\mu \gamma^\mu q_j^\mu)g_\mu^a + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu \bar{G}^a G^b g_\mu^c - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\
 & M^2 W_\mu^+ W_\mu^- - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2}M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2}\partial_\mu H \partial_\mu H - \\
 & \frac{1}{2}m_h^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \frac{1}{2c_w^2}M \phi^0 \phi^0 - \beta_h [\frac{2M^2}{g^2} + \\
 & \frac{2M}{g}H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-)] + \frac{2M^4}{g^4}\alpha_h - ig_{cw}[\partial_\nu Z_\mu^0(W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - Z_\nu^0(W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0(W_\nu^+ \partial_\nu W_\mu^- - \\
 & W_\nu^- \partial_\nu W_\mu^+) - ig_{sw}[\partial_\nu A_\mu(W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu(W_\mu^+ \partial_\nu W_\mu^- - \\
 & W_\mu^- \partial_\nu W_\mu^+) + A_\mu(W_\nu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\nu W_\mu^+)] - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \\
 & \frac{1}{2}g^2 W_\mu^+ W_\nu^- W_\mu^+ W_\nu^- + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + \\
 & g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w [A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - 2A_\mu Z_\mu^0 W_\nu^+ W_\nu^-] - g\alpha [H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-] - \\
 & \frac{1}{8}g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2] - \\
 & gMW_\mu^+ W_\mu^- H - \frac{1}{2}g\frac{M}{c_w^2}Z_\mu^0 Z_\mu^0 H - \frac{1}{2}ig[W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - \\
 & W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)] + \frac{1}{2}g[W_\mu^+ (H\partial_\mu \phi^- - \phi^- \partial_\mu H) - W_\mu^- (H\partial_\mu \phi^+ - \\
 & \phi^+ \partial_\mu H)] + \frac{1}{2}g\frac{1}{c_w}(Z_\mu^0 (H\partial_\mu \phi^0 - \phi^0 \partial_\mu H) - ig\frac{1}{c_w}MZ_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \\
 & ig_{sw}MA_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - ig\frac{1-2c_w^2}{2c_w}Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\
 & ig_{sw}A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4}g^2 W_\mu^+ W_\mu^- [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \\
 & \frac{1}{4}g^2 \frac{1}{c_w^2}Z_\mu^0 Z_\mu^0 [H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{s_w^2}{c_w^2}Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w^2}Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{2s_w}{c_w^2}(2c_w^2 - 1)Z_\mu^0 A_\mu \phi^+ \phi^- - \\
 & g^1 s_w^2 A_\mu A_\mu \phi^+ \phi^- - \bar{e}^\lambda (\gamma^\mu \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda \gamma^\mu \partial \nu^\lambda - \bar{u}_j^\lambda (\gamma^\mu \partial + m_u^\lambda) u_j^\lambda - \\
 & \bar{d}_j^\lambda (\gamma^\mu \partial + m_d^\lambda) d_j^\lambda + ig_{sw}A_\mu [-(\bar{e}^\lambda \gamma^\mu e^\lambda) + \frac{2}{3}(\bar{u}_j^\lambda \gamma^\mu u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma^\mu d_j^\lambda)] + \\
 & \frac{ig}{4c_w}Z_\mu^0 [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - \\
 & 1 - \gamma^5) u_j^\lambda) + (\bar{d}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 - \gamma^5) d_j^\lambda)] + \frac{ig}{2\sqrt{2}}W_\mu^+ [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) e^\lambda) + \\
 & (\bar{u}_j^\lambda \gamma^\mu (1 + \gamma^5) C_{\lambda\kappa} d_j^\kappa)] + \frac{ig}{2\sqrt{2}}W_\mu^- [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger \gamma^\mu (1 + \\
 & \gamma^5) u_j^\lambda)] + \frac{ig}{2\sqrt{2}}\frac{m_\lambda^\lambda}{M}[-\phi^+ (\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) e^\lambda) + \phi^- (\bar{e}^\lambda (1 + \gamma^5) \nu^\lambda)] - \\
 & \frac{g}{2}\frac{m_\lambda^\lambda}{M}[H(\bar{e}^\lambda e^\lambda) + i\phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda)] + \frac{ig}{2M\sqrt{2}}\phi^+ [-m_d^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + \\
 & m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \gamma^5) d_j^\kappa)] + \frac{ig}{2M\sqrt{2}}\phi^- [m_d^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 - \\
 & \gamma^5) u_j^\kappa)] - \frac{g}{2}\frac{m_\lambda^\lambda}{M}H(\bar{u}_j^\lambda u_j^\lambda) - \frac{g}{2}\frac{m_\lambda^\lambda}{M}H(\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2}\frac{m_\lambda^\lambda}{M}\phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \\
 & \frac{ig}{2}\frac{m_\lambda^\lambda}{M}\phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \bar{X}^+ (\partial^2 - M^2)X^+ + \bar{X}^- (\partial^2 - M^2)X^- + \bar{X}^0 (\partial^2 - \\
 & \frac{M^2}{c_w^2})X^0 + \bar{Y} \partial^2 Y + ig_{cw}W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \partial_\mu \bar{X}^+ X^0) + ig_{sw}W_\mu^+ (\partial_\mu \bar{Y} X^- - \\
 & \partial_\mu \bar{X}^+ Y) + ig_{cw}W_\mu^- (\partial_\mu \bar{X}^- X^0 - \partial_\mu \bar{X}^- X^+) + ig_{sw}W_\mu^- (\partial_\mu \bar{X}^- Y - \\
 & \partial_\mu \bar{Y} X^+) + ig_{cw}Z_\mu^0 (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) + ig_{sw}A_\mu (\partial_\mu \bar{X}^+ X^+ - \\
 & \partial_\mu \bar{X}^- X^-) - \frac{1}{2}gM[\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w^2}\bar{X}^0 X^0 H] + \\
 & \frac{1-2c_w^2}{2c_w}igM[\bar{X}^+ X^0 \phi^+ - \bar{X}^- X^0 \phi^-] + \frac{1}{2c_w}igM[\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \\
 & igMs_w[\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \frac{1}{2}igM[\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0]
 \end{aligned}$$

Another Reason:

A Higgs boson

- ONLY spin 0 elementary particle
- Couplings are NOT dictated by gauge symmetry
 - Hmm....
- Symmetry breaking
 - Underlying reason?
- Small mass possible if new physics
 - “Fine Tuning Problem”



Higgs is an EXOTIC particle.

Implications of the Higgs Discovery

- Last prediction from an experimentally well tested model.
 - No real guidance on the model market
 - New insights have to come from experiments
 - **Generic searches!**

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$$\mathcal{L}_H = (D_\mu H)^\dagger (D^\mu H) - \lambda (H^\dagger H)^2 + \lambda \boldsymbol{v}^2 H^\dagger H$$

Implications of the Higgs Discovery

- Last prediction from an experimentally well tested model.
 - No real guidance on the model market
 - New insights have to come from experiments
 - **Generic searches!**
 - If new boson is the SM Higgs
- Know now experimentally **scale of Standard Model.**

$$\mathcal{L}_H = (D_\mu H)^\dagger (D^\mu H) - \lambda (H^\dagger H)^2 + \lambda \boldsymbol{v}^2 H^\dagger H$$

- $\boldsymbol{v} = (\sqrt{2}G_F)^{-1/2} \sim 246 \text{ GeV} \sim 10^{-16} \text{ cm}$
- Search beyond this scale → **TeV and above!**

Many Examples for this in History

PERIODIC TABLE OF THE ELEMENTS

<http://www.periodni.com>

GROUP I A 2 II A 13 III A 14 IV A 15 VA 16 VIA 17 VIIA 18 VIIIA

PERIOD 1 2 3 4 5 6 7

RELATIVE ATOMIC MASS (1)

GROUP IUPAC

GROUP CAS

ATOMIC NUMBER

SYMBOL

ELEMENT NAME

1 1.0079 H HYDROGEN

2 4.0026 He HELIUM

3 6.941 Li LITHIUM

4 9.0122 Be BERYLLIUM

5 10.811 B BORON

6 12.011 C CARBON

7 14.007 N NITROGEN

8 15.999 O OXYGEN

9 18.998 F FLUORINE

10 20.180 Ne NEON

11 22.990 Na SODIUM

12 24.305 Mg MAGNESIUM

13 26.982 Al ALUMINUM

14 28.086 Si SILICON

15 30.974 P PHOSPHORUS

16 32.065 S SULPHUR

17 35.453 Cl CHLORINE

18 39.948 Ar ARGON

19 39.098 K POTASSIUM

20 40.078 Ca CALCIUM

21 44.956 Sc SCANDIUM

22 47.867 Ti TITANIUM

23 50.942 V VANADIUM

24 51.996 Cr CHROMIUM

25 54.938 Mn MANGANESE

26 55.845 Fe IRON

27 58.933 Co COBALT

28 58.693 Ni NICKEL

29 63.546 Cu COPPER

30 65.38 Zn ZINC

31 69.723 Ga GALLIUM

32 72.64 Ge GERMANIUM

33 74.922 As ARSENIC

34 78.96 Se SELENIUM

35 79.904 Br BROMINE

36 83.798 Kr KRYPTON

37 85.468 Rb RUBIDIUM

38 87.62 Sr STRONTIUM

39 88.906 Y YTTRIUM

40 91.224 Zr ZIRCONIUM

41 92.906 Nb NIOBIUM

42 95.94 Mo MOLYBDENUM

43 (98) Tc TECHNETIUM

44 101.07 Ru RUTHENIUM

45 102.91 Rh RHODIUM

46 106.42 Pd PALLADIUM

47 107.87 Ag SILVER

48 112.41 Cd CADMIUM

49 114.82 In INDIUM

50 118.71 Sn TIN

51 121.76 Sb ANTIMONY

52 127.60 Te TELLURIUM

53 126.90 I IODINE

54 131.29 Xe XENON

55 132.91 Cs CAESIUM

56 137.33 Ba BARIUM

57-71 La-Lu Lanthanide

72 178.49 Hf HAFNIUM

73 180.95 Ta TANTALUM

74 183.84 W TUNGSTEN

75 186.21 Re RHENIUM

76 190.23 Os OSMIUM

77 192.22 Ir IRIUM

78 195.08 Pt PLATINUM

79 196.97 Au GOLD

80 200.59 Hg MERCURY

81 204.38 Tl THALLIUM

82 207.2 Pb LEAD

83 208.98 Bi BISMUTH

84 (209) Po POLONIUM

85 (210) At ASTATINE

86 (222) Rn RADON

87 (223) Fr FRANCIUM

88 (226) Ra RADIUM

89-103 Ac-Lr Actinide

104 (261) Rf RUTHERFORDIUM

105 (268) Db DUBNIUM

106 (271) Sg SEABORGIUM

107 (272) Bh BOHRNIUM

108 (277) Hs HASSIUM

109 (276) Mt MEITNERIUM

110 (281) Ds DARMSTADTIUM

111 (289) Rg ROENTGIUM

112 (285) Cn COPERNICIUM

113 (...) Uut UNUNTRIUM

114 (287) Fl FLEROVIUM

115 (...) Uup UNUNPENTIUM

116 (291) Lv LIVERMORIUM

117 (...) Uus UNUNSEPTIUM

118 (...) Uuo UNUNOCTIUM

Legend:

- Metal
- Semimetal
- Nonmetal
- Alkali metal
- Alkaline earth metal
- Transition metals
- Lanthanide
- Actinide
- Chalcogens element
- Halogens element
- Noble gas

STANDARD STATE (25 °C, 101 kPa)

Ne - gas Fe - solid Hg - liquid - synthetic

Around **1900** reached atomic scale **$10^{-8} \text{ cm} \approx \hbar^2 / e^2 m_e$**



Quantum Mechanics



Quantum Electrodynamics

The Periodic Table of Particle Physics

The Standard Model and the Higgs boson

	Fermions			Bosons	
Quarks	u up	c charm	t top	γ photon	Force carriers
	d down	s strange	b bottom	Z Z boson	
Leptons	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	W W boson	
	e electron	μ muon	τ tau	g gluon	
				Higgs boson	

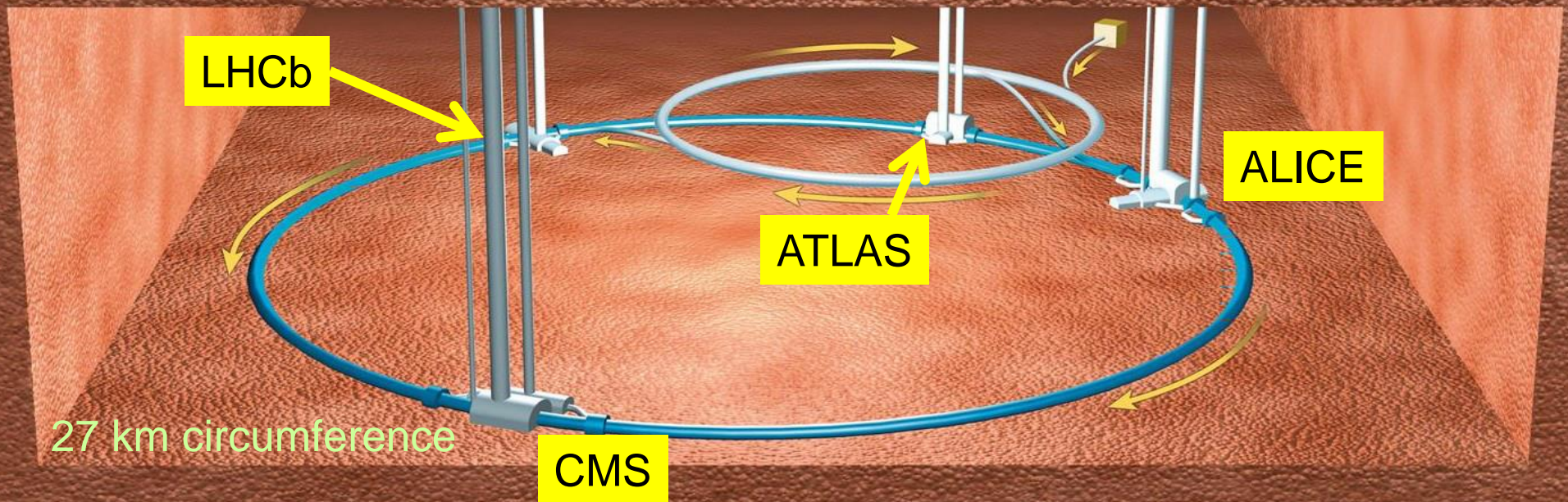
Source: AAAS

Today Very Special Time

LHC above energy scale of Standard Model:

$$>> \text{TeV}^{-1} \sim 10^{-17} \text{ cm}$$

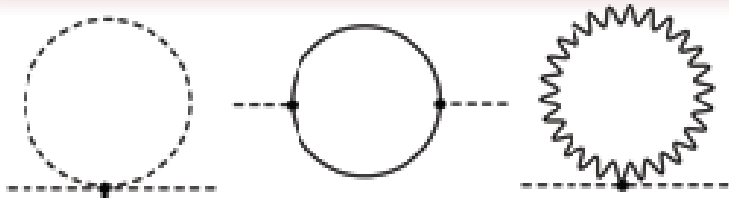
Probes New Physics



What else is there beside SUSY framework?

- SUSY is NOT a model
 - “Symmetry principle characterizing a BSM framework with an infinite number of models”....Lykken
- SUSY mass limits pushed to 1 TeV
 - SUSY becoming more “Exotic” the higher the mass limits get.
- SUSY is only one possible way.....
 - Many more ways to solve problems with Standard Model
 - What if nature has not chosen low scale SUSY?
 - Make sure to cover every feasible corner...

Fine Tuning Problem....



G. Servant $\Rightarrow \delta M_H^2 \propto \Lambda^2$

4 ways to solve it

- Supersymmetry

- Sparticles cancel particle contributions

- Extra Dimensions

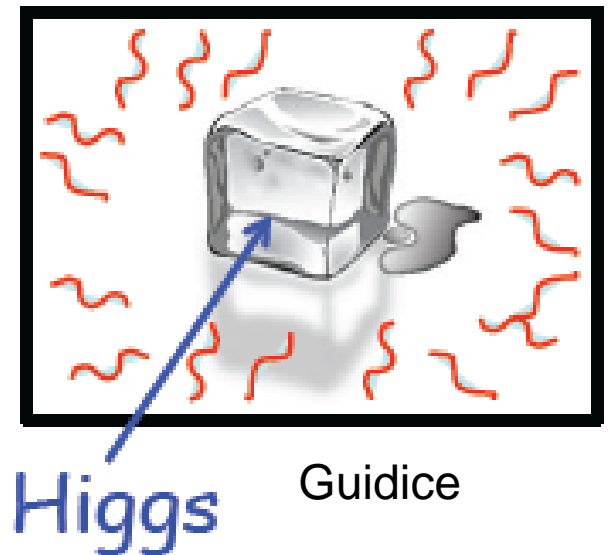
- Higgs is a vector in 5D

- Higgs is composite

- Strongly coupled new physics

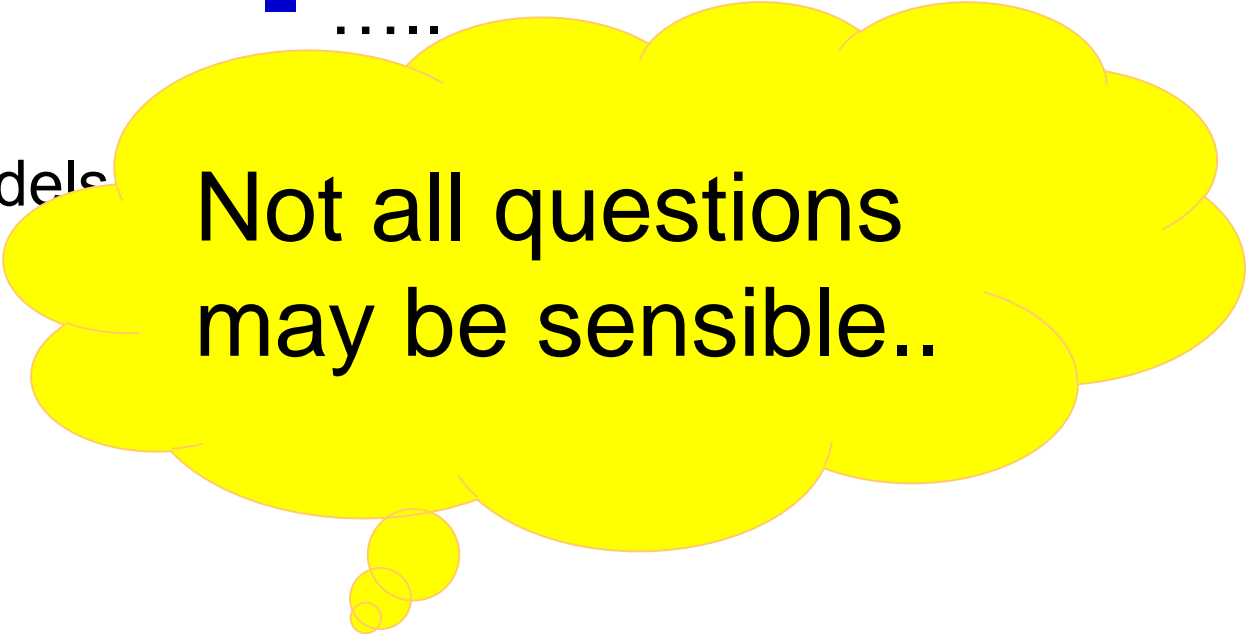
- There is no fine tuning problem in SM

- *Not everybody thinks SM has a fine tuning problem*



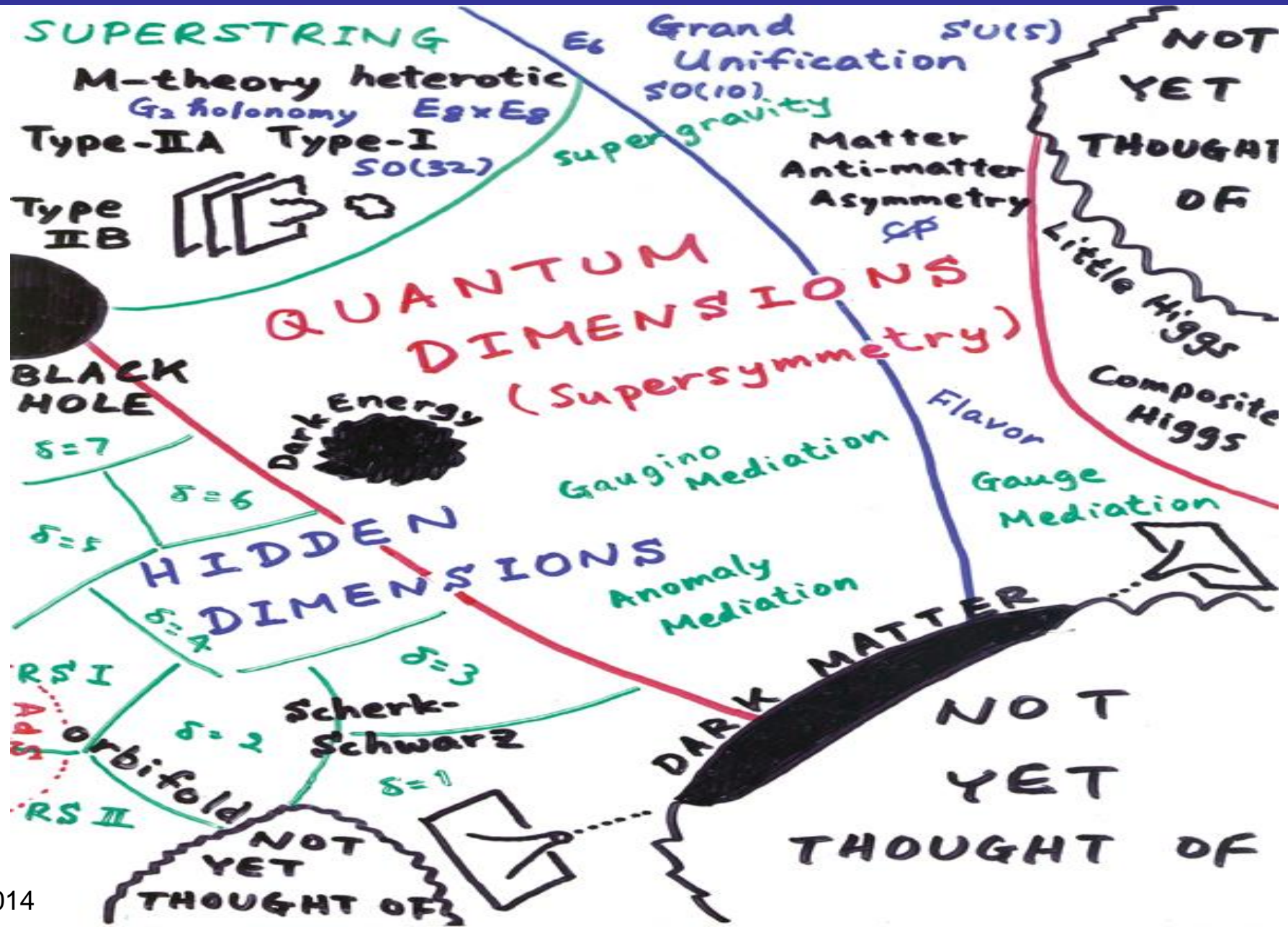
Models try to answer questions

- Hierarchy Problem
 - EWK force $\sim 10^{32}$ X Gravity?
 - Extra dimension models
- Fine Tuning Problem
 - SUSY
 - Composite Higgs
 - Extra dimension models
- What is Dark Matter?
 - SUSY
 - Extra dimensions....
- Family structure in SM?
- Running coupling constants?
 - GUT
- Have elementary particles a sub-structure?
-

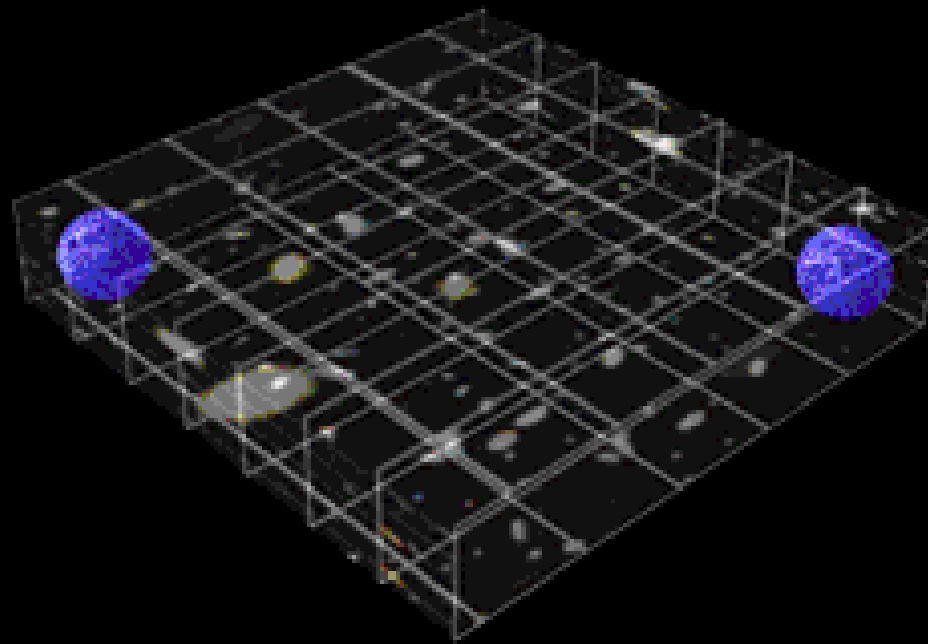


Not all questions
may be sensible..

Models



Extra Dimensions



Extra Dimensions are not a new idea!

- 1920's Kaluza&Klein unify electromagnetism with gravity
 - 1970 String Theory is born
 - QM of oscillating strings
 - Bosonic
 - 1971 SUSY enters the stage
 - 1974 Gravitons “pop out” of string theory
- 1984 Superstring Theory
 - 10, 11 or 26 dimensions needed
 - Compactified
 - SUSY needed for fermions
 - 1998 Large Extra Dim.
 - Nima Arkani-Hamed, Savas Dimopoulos, and Gia Dvali
 - Warped Extra Dim.(1999)

Geometrisation of Gravitational Field

$$G_{\mu\nu} + g_{\mu\nu}\Lambda = 8\pi G T_{\mu\nu}$$

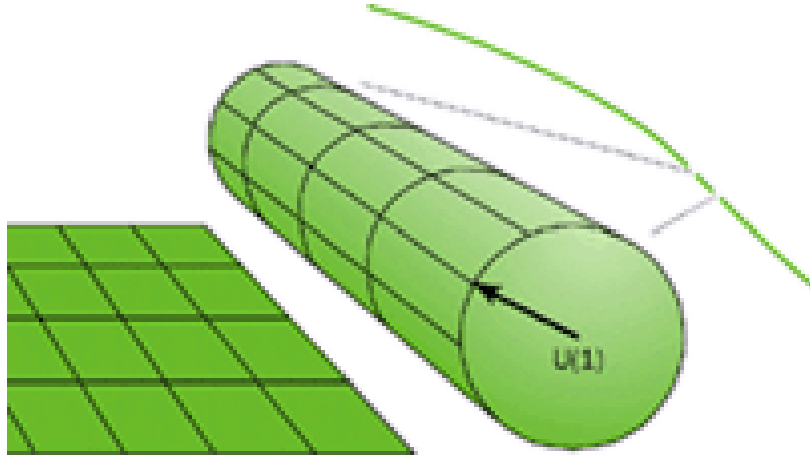
Geometry

Matter

Dynamically correlated

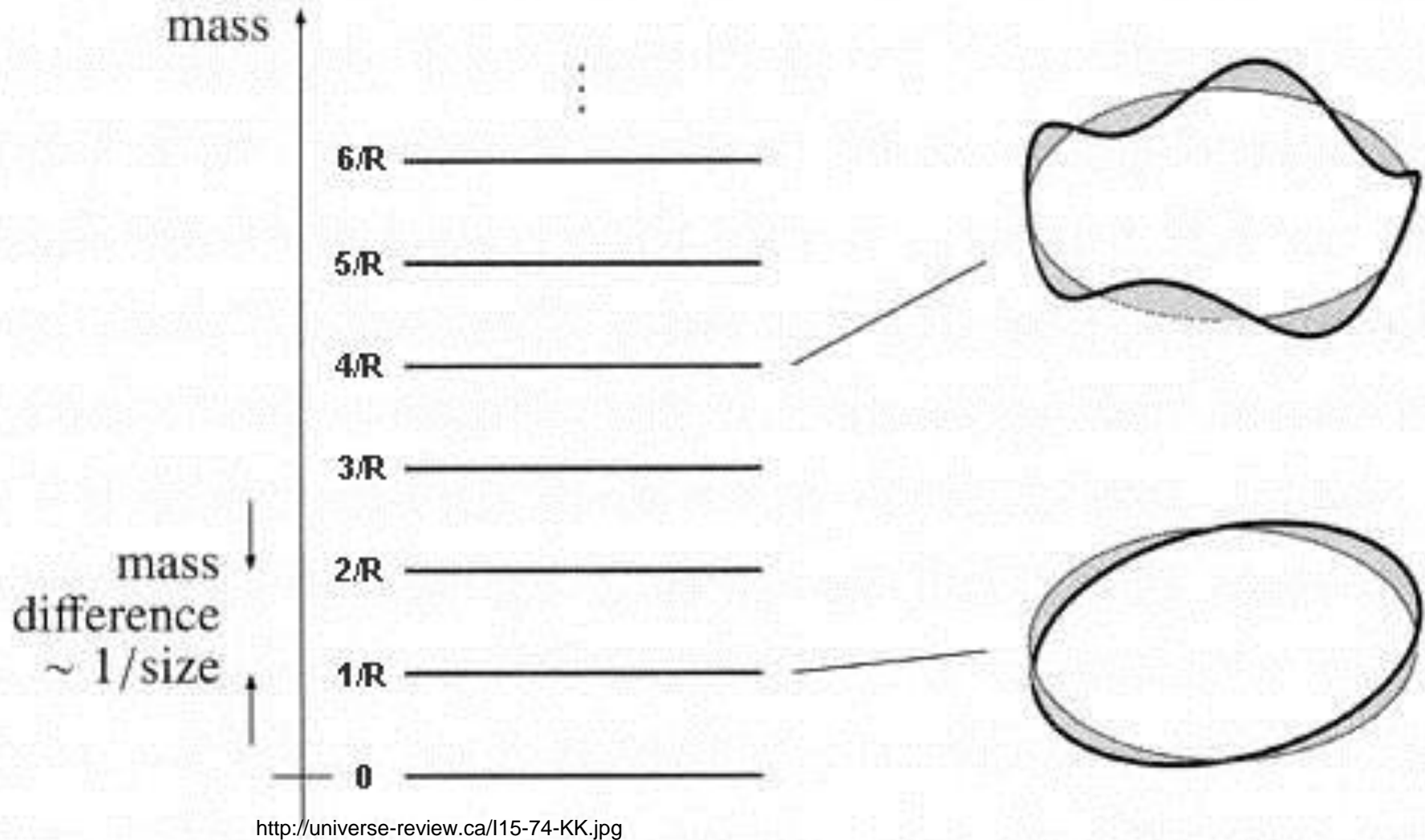
Kaluza-Klein Theory

- Geometrical unification of gravity and electromagnetism



- Formulate GR in 5D
 - 4D gravity + $U(1)$ gauge theory + scalar field (radion)
- Basis of string theory
- Problems:
 - Classic theory
 - Not chiral, fermions are vector like

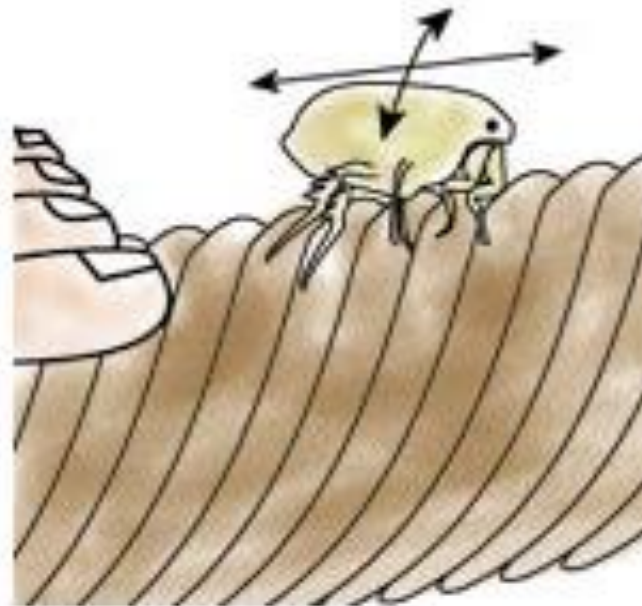
KK Particles



Where are they?

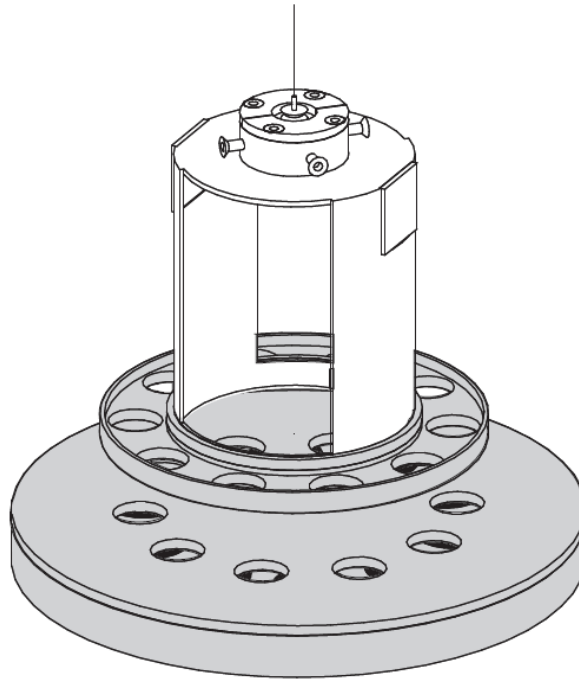
- ED may explain complexity of particle physics
- Where are they?

http://www.particleadventure.org/frameless/extra_dim.html



...but a flea can move
in two dimensions.

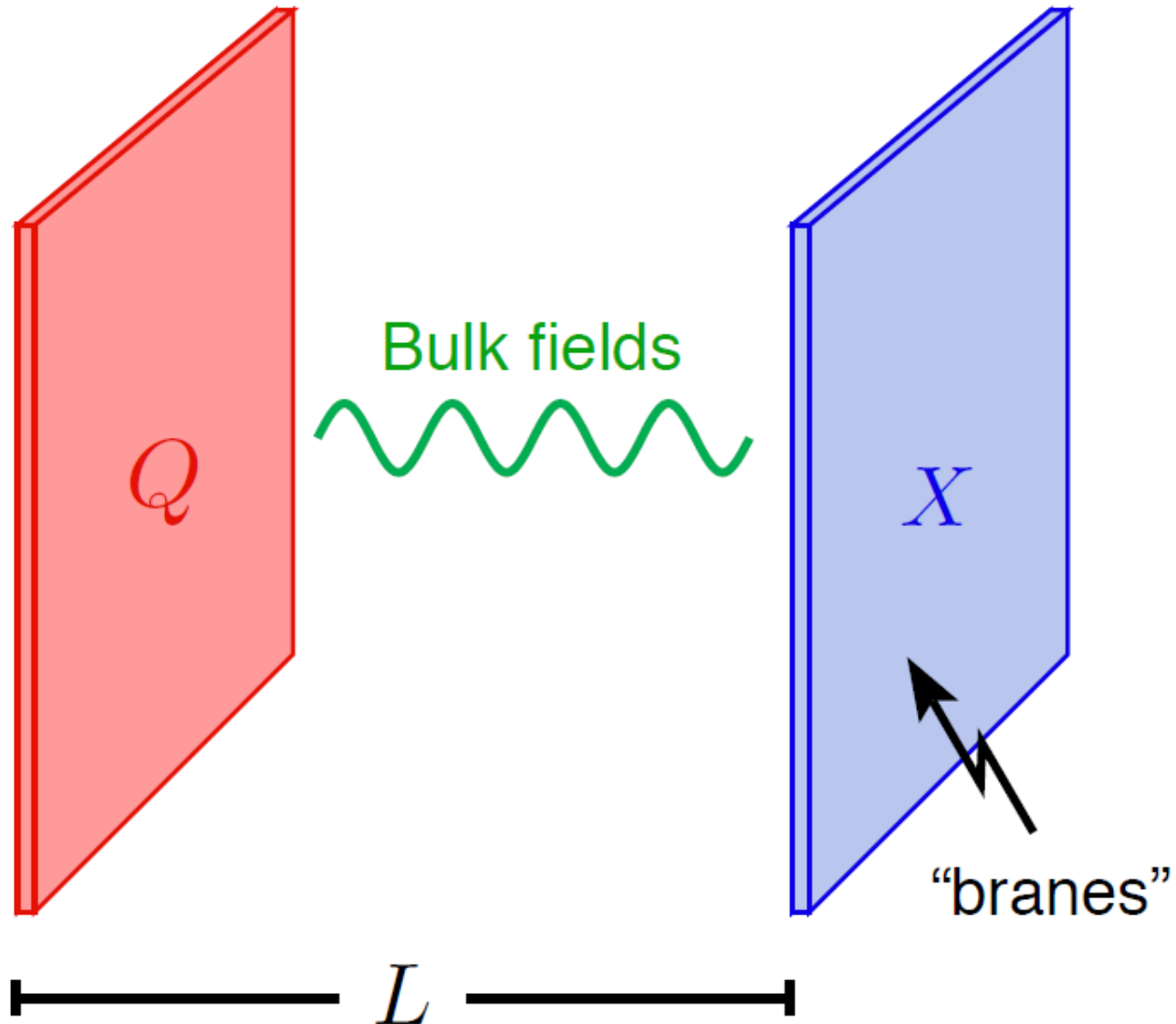
Table Top Experiments



$1/r^2$ -law valid for $R=44 \mu\text{m}$ ($M_D \sim 4 \text{ TeV}$) at 95%

E.G. Adelberger, Prog.Part.Nucl.Phys. 62 (2009) 102-134

Modern Extra Dimension Models



Gravity in Extra Dimension

At small distances gravity can be very strong,
up to 10^{38} times stronger:

$$\mathbf{F} \approx \frac{\mathbf{G}_D}{r^{n+2}}$$

$$\mathbf{M}_D^{n+2} = \frac{(2\pi)^n}{8\pi \mathbf{G}_D}$$

At large distances gravity seems weak

$$\mathbf{G}_D = \mathbf{G} L^n$$

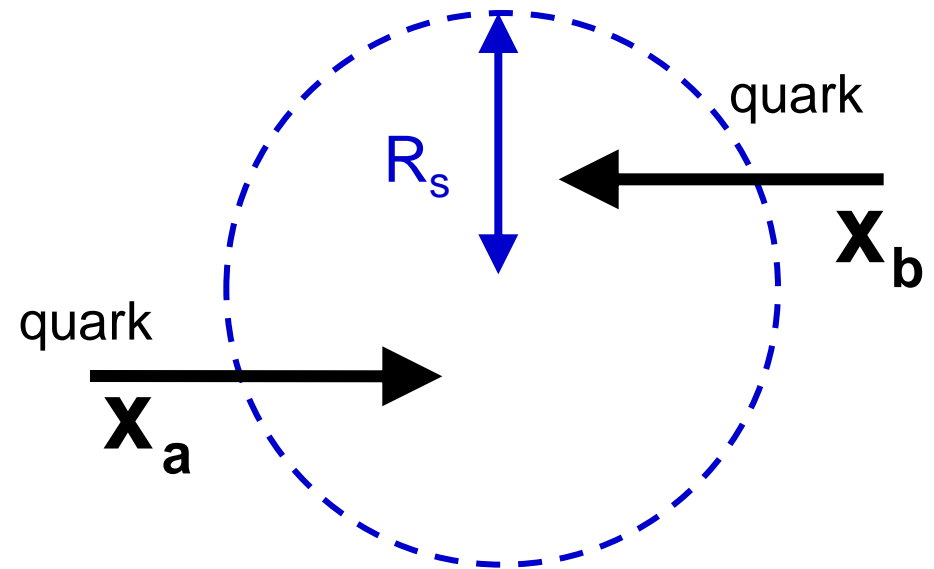
$$\mathbf{F} \approx \frac{\mathbf{G}_D}{L^n \cdot r^2} \approx \frac{\mathbf{G}}{r^2}$$

\mathbf{G} is “diluted” strength of gravity in our 3-dim. space.
 \mathbf{G}_D is the $(4+n)$ -dimensional Newton gravity constant.

Production of Black Holes at the LHC

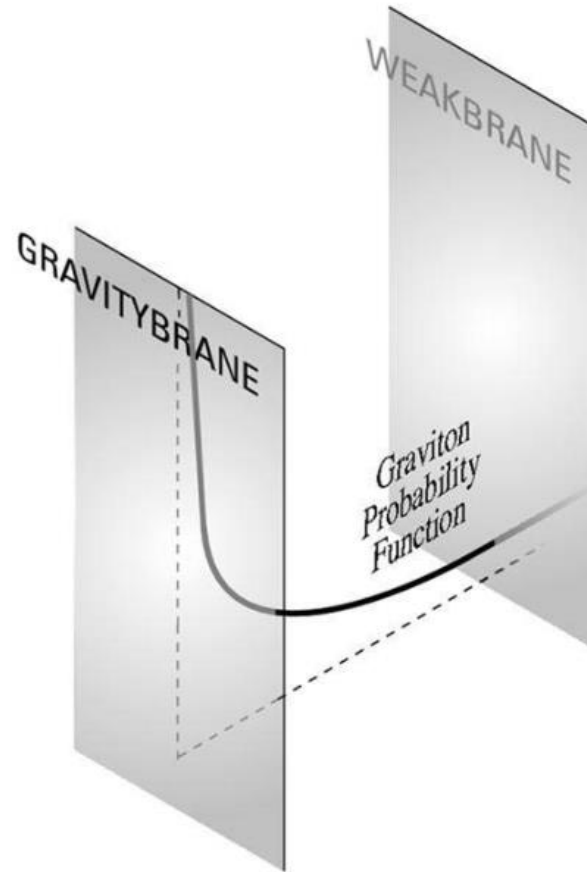
$$R_s = \frac{2 G^* L^n M}{c^2}$$

$$M = \sqrt{S X_a X_b} = \sqrt{\hat{S}}$$



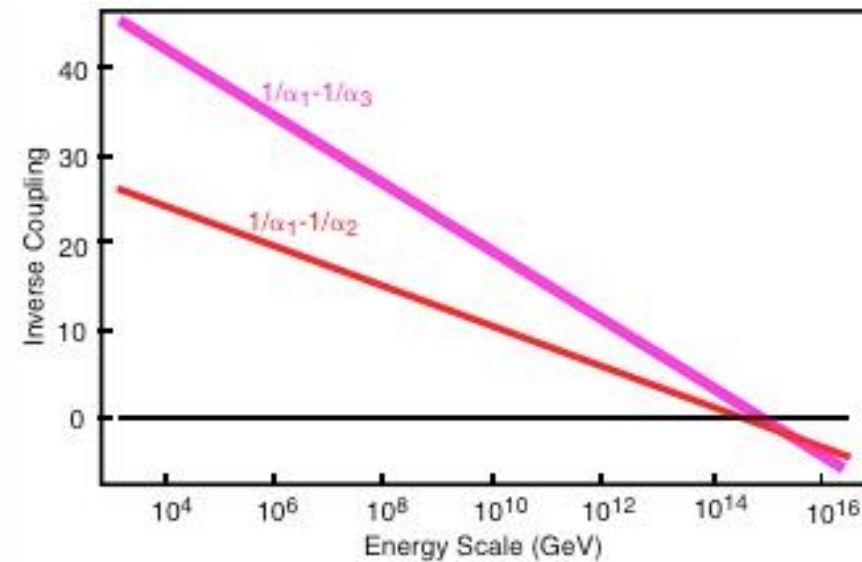
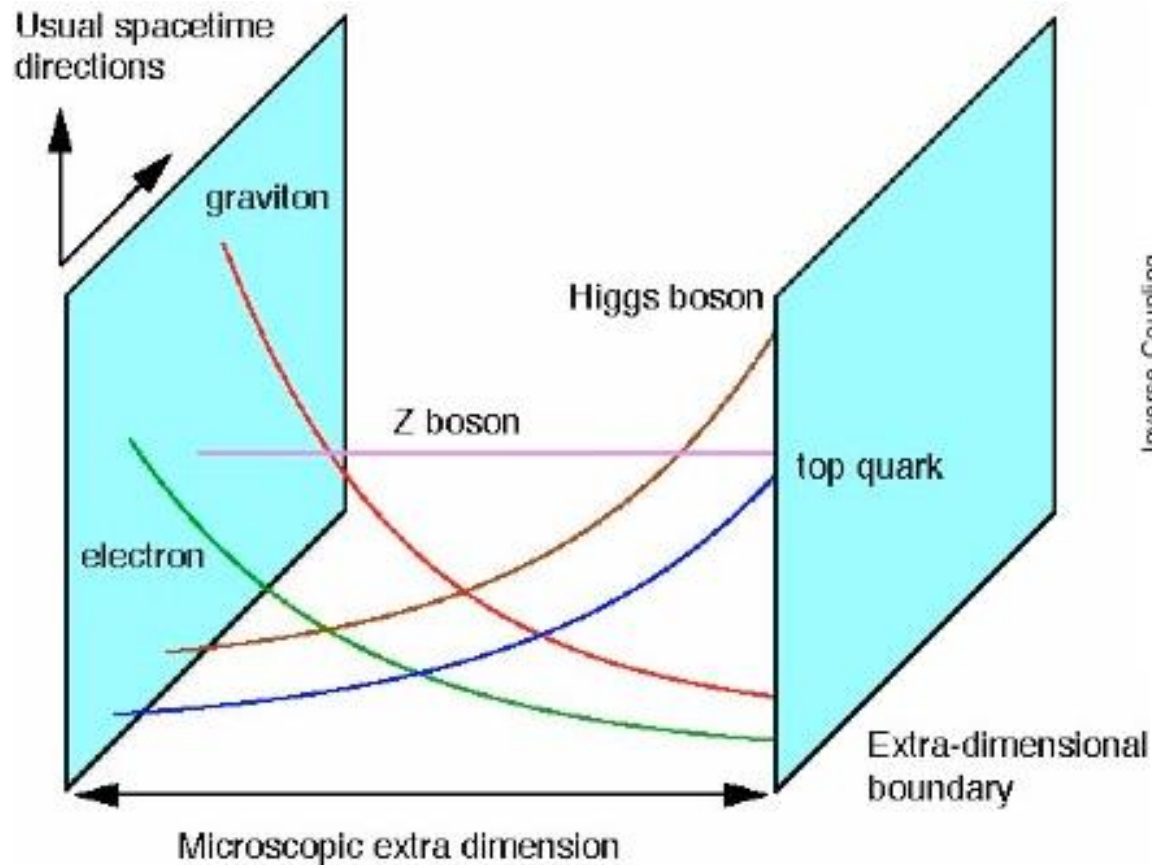
$$R_s^{2\text{quarks}} \leq 10^{-18} \text{ m}$$

Warped Extra Dimensions



$$ds^2 = \underline{a(y)^2} dx^\mu dx^\nu \eta_{\mu\nu} + dy^2, \quad \underline{a(y) = e^{-ky}}$$

Other Warped Extra Dimension Models



Supersymmetry

- Geometric interpretation using Superspace

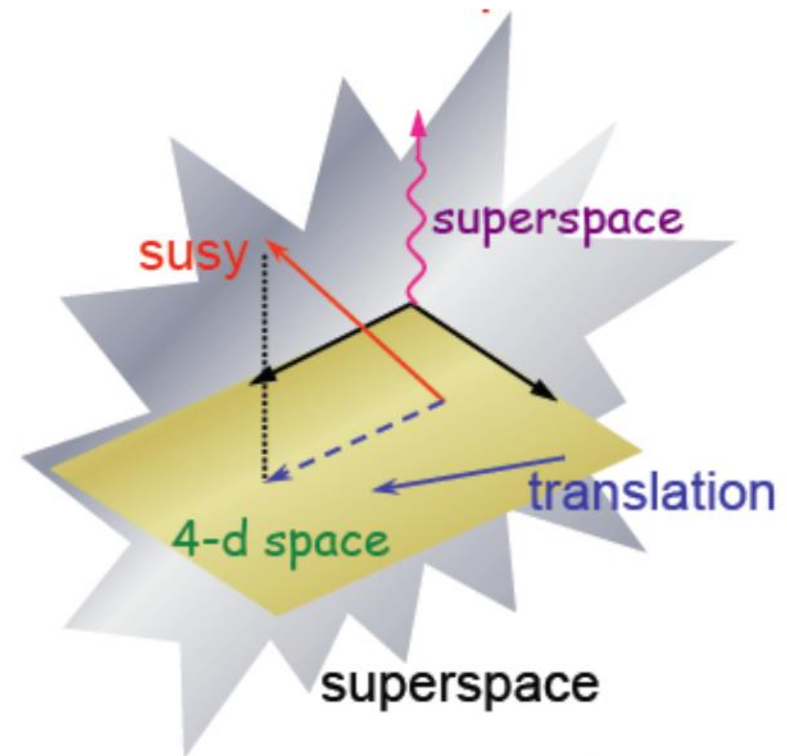
Guidice



4-d space-time

Poincaré

$$P = (t, x, y, z)$$

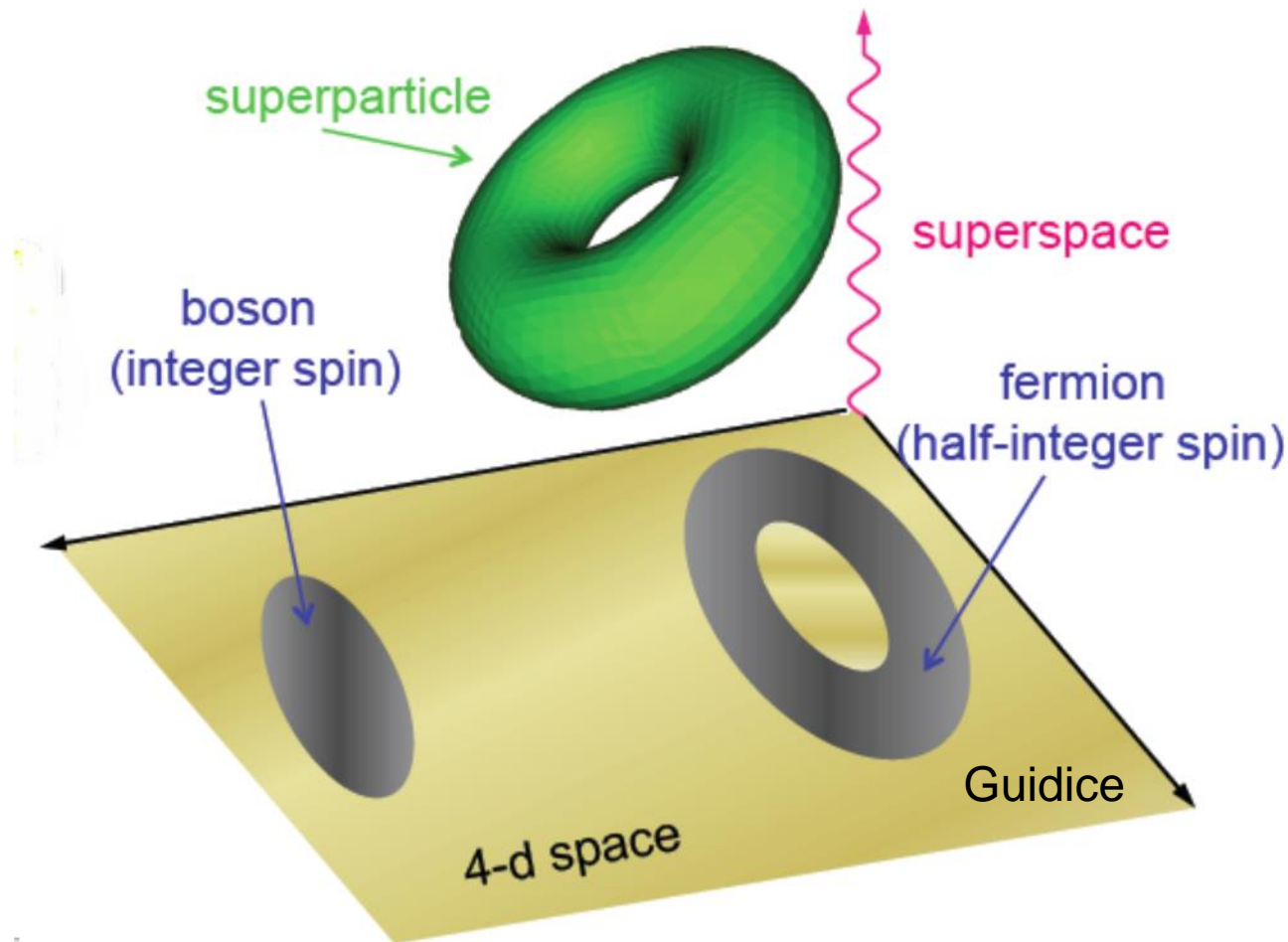


superspace

supersymmetry

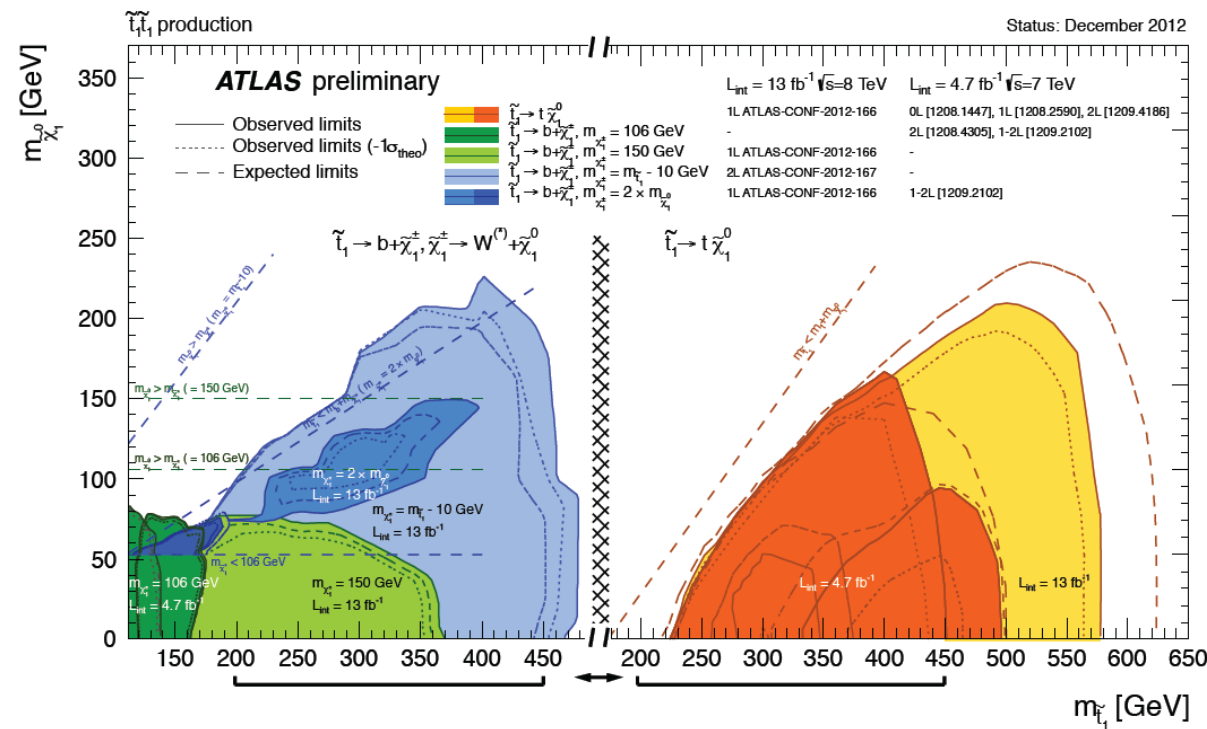
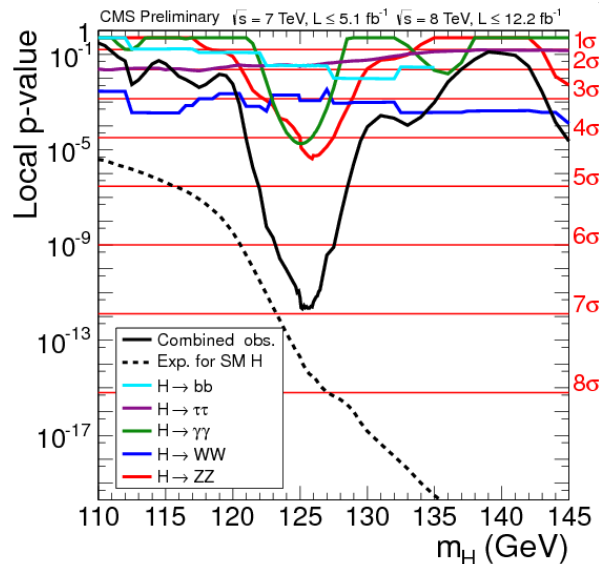
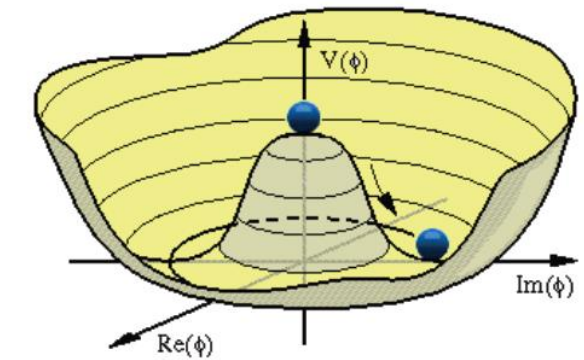
$$P = (t, x, y, z, \theta, \bar{\theta})$$

SUSY is Symmetry Group of Superspace

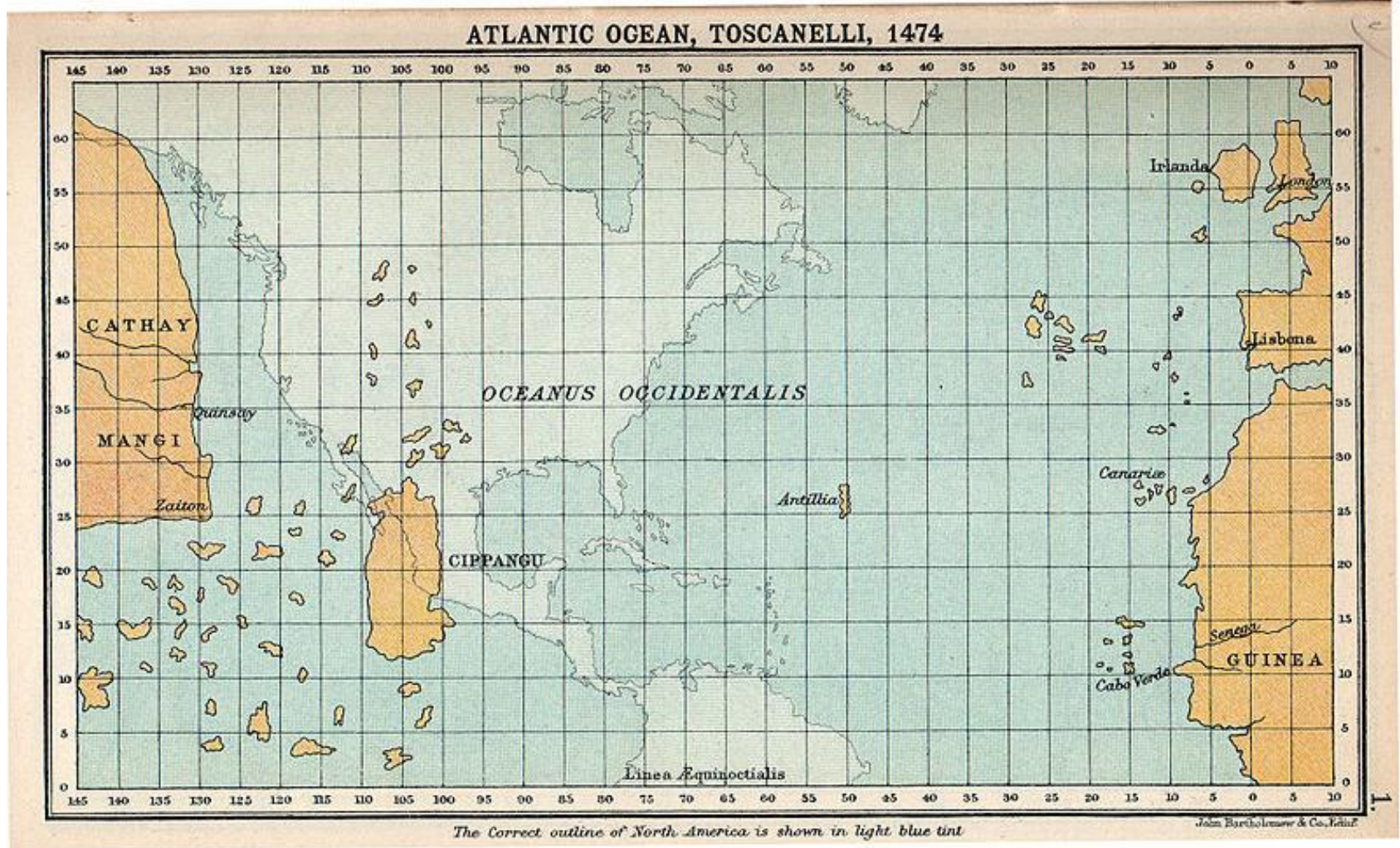


Role of Models in “most” Exotics Searches?

- No specific Model to guide us.
- No unified parameter phase space to map results

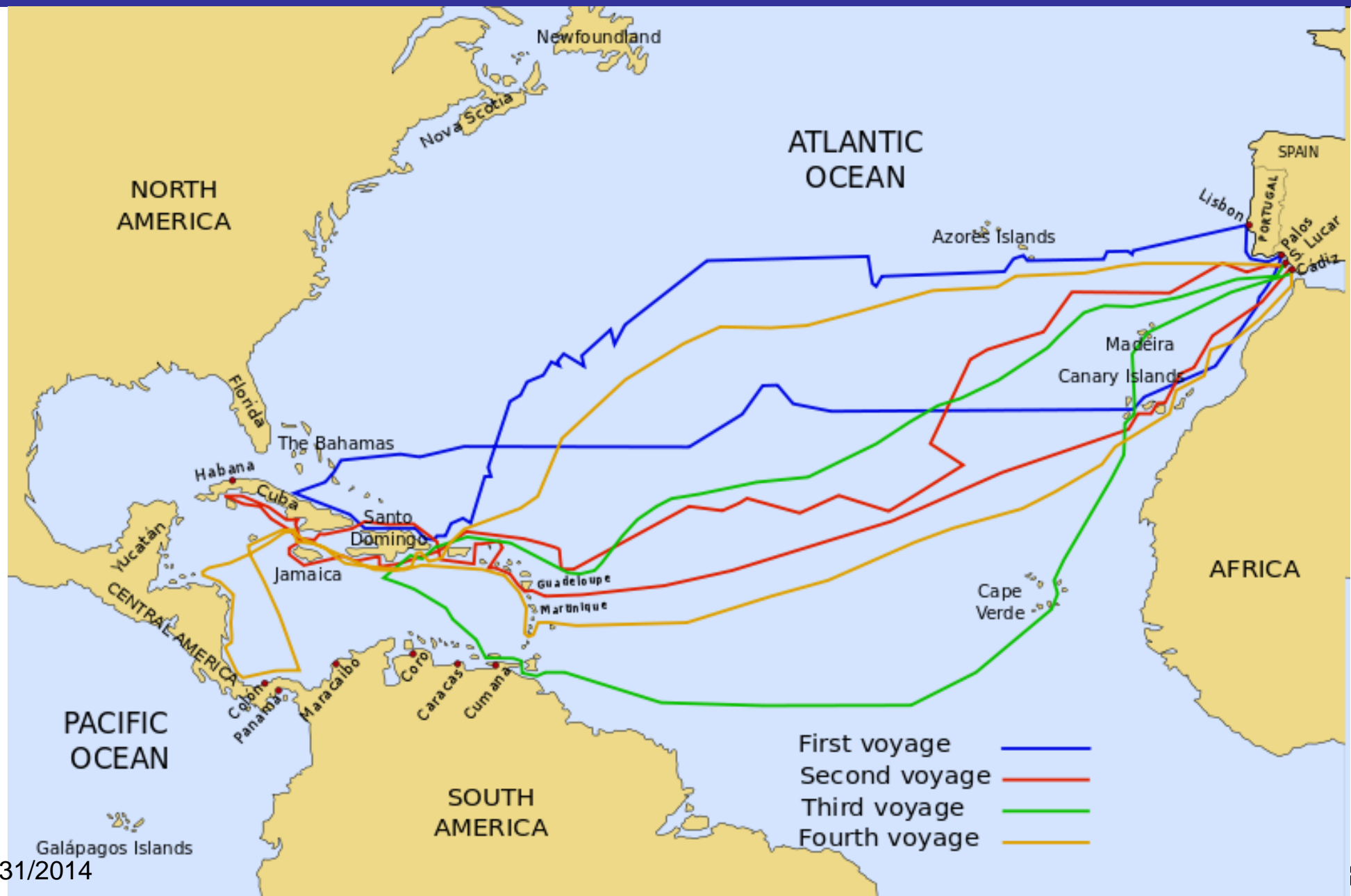


The Role of Models in “most” Exotics Searches



Toscanelli's model of the geography of the Atlantic Ocean, which directly influenced Columbus's plans

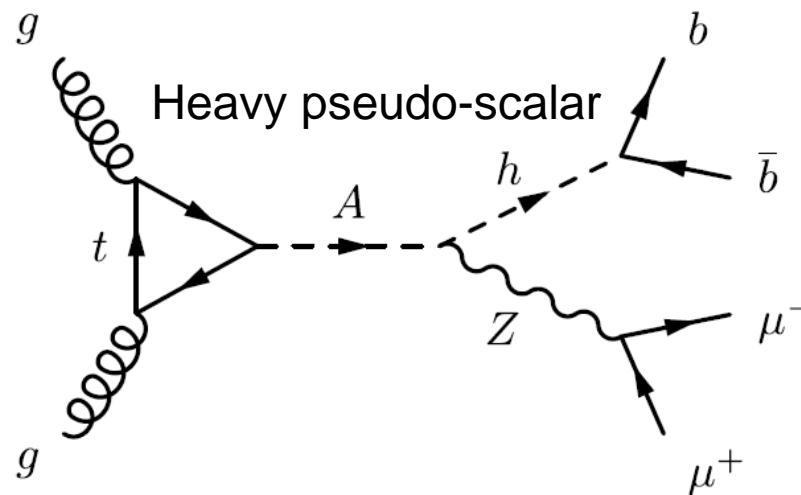
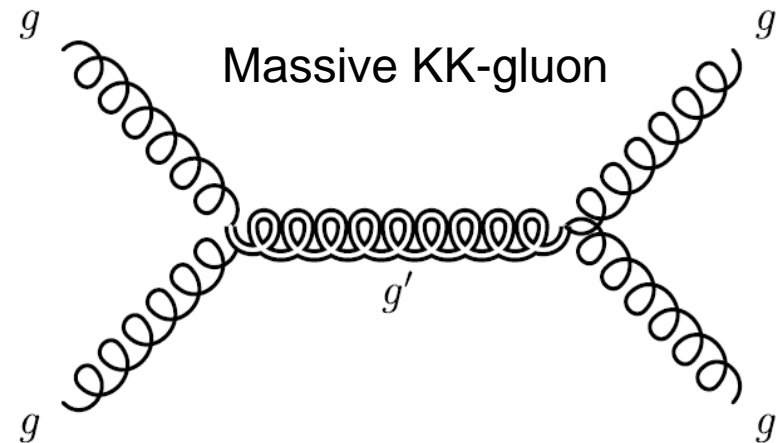
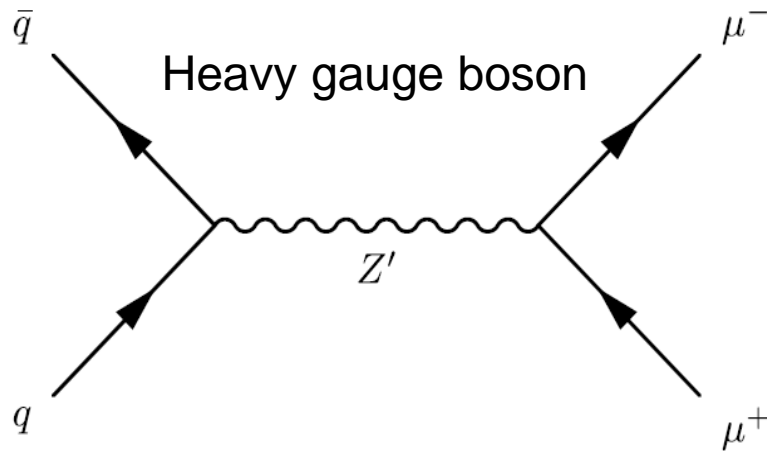
The Role of Models in “most” Exotics Searches



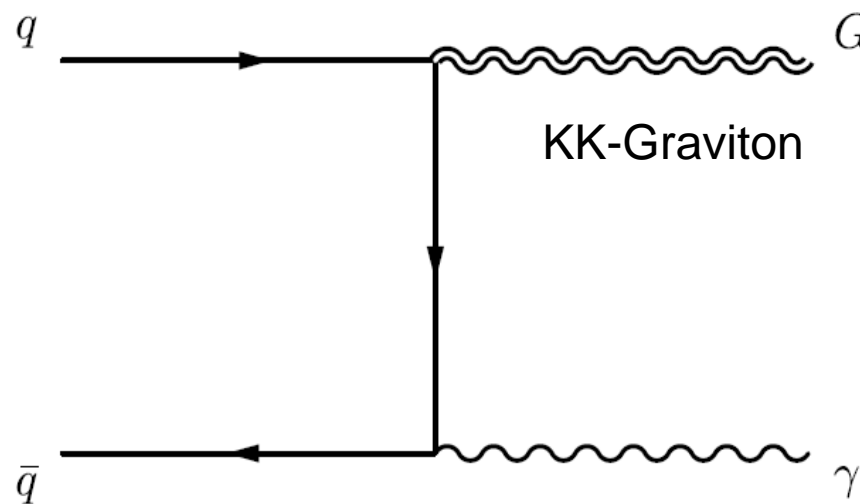
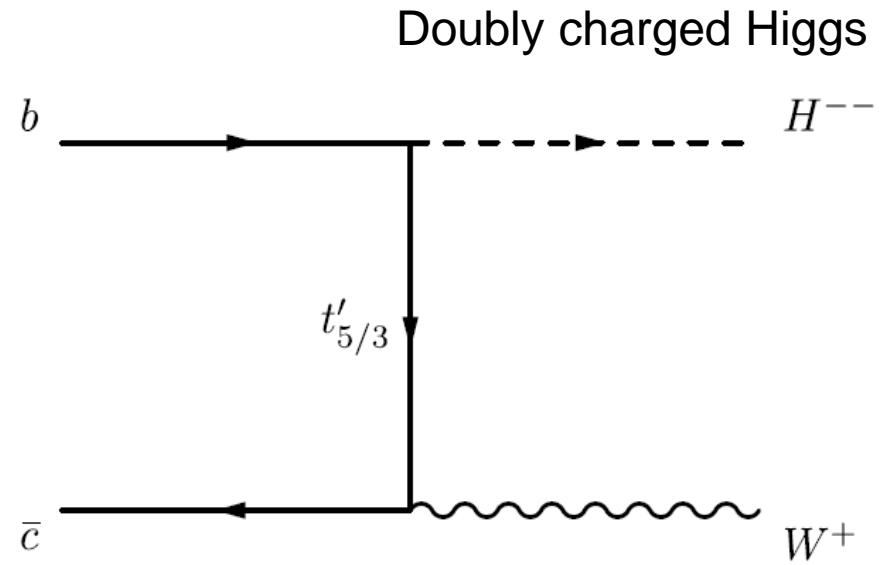
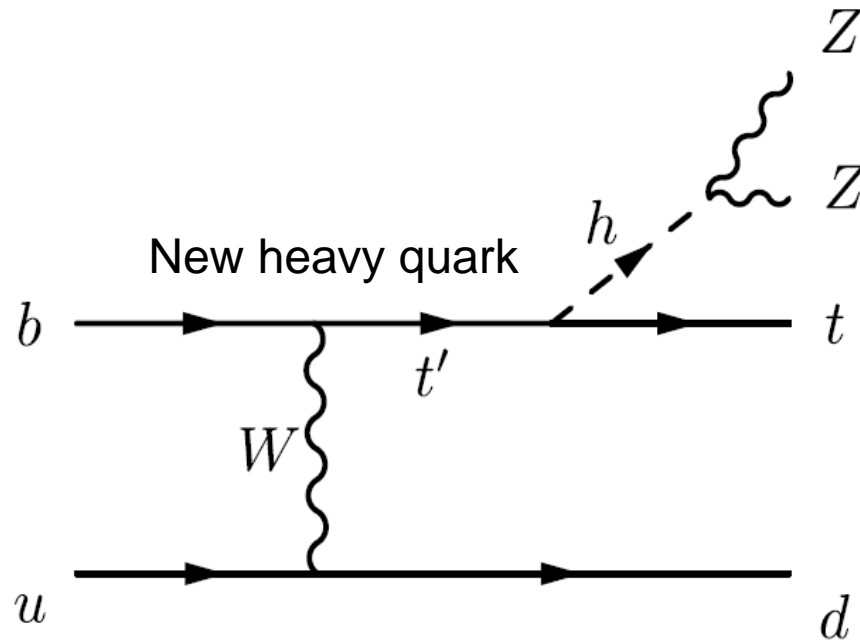
The Role of Models in “most” Exotics Searches

- Models used to quantify our reach.
 - How far did we get?
 - How do we compare to previous searches?
- We use so called Bench Mark Models
 - Used before by other experiments
- Simplified Models or generic resonances

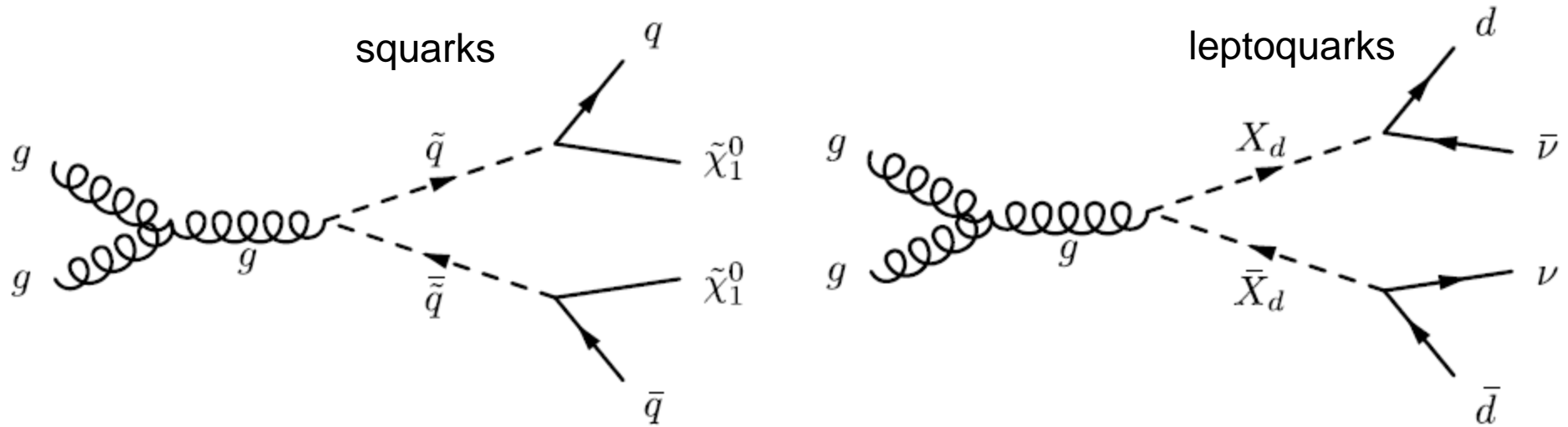
Exotics Search Signatures: s-channel Production



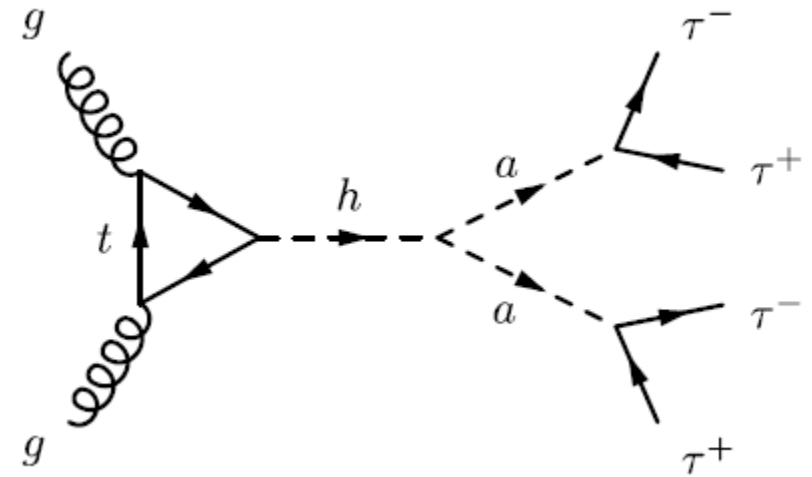
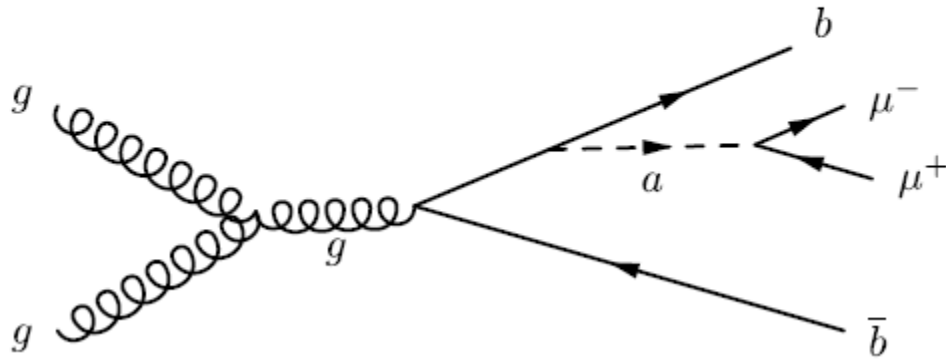
Exotics Search Signatures: Associate Production



Exotics Search Signatures: Pair Production



Exotics Search Signatures: BSMstrahlung



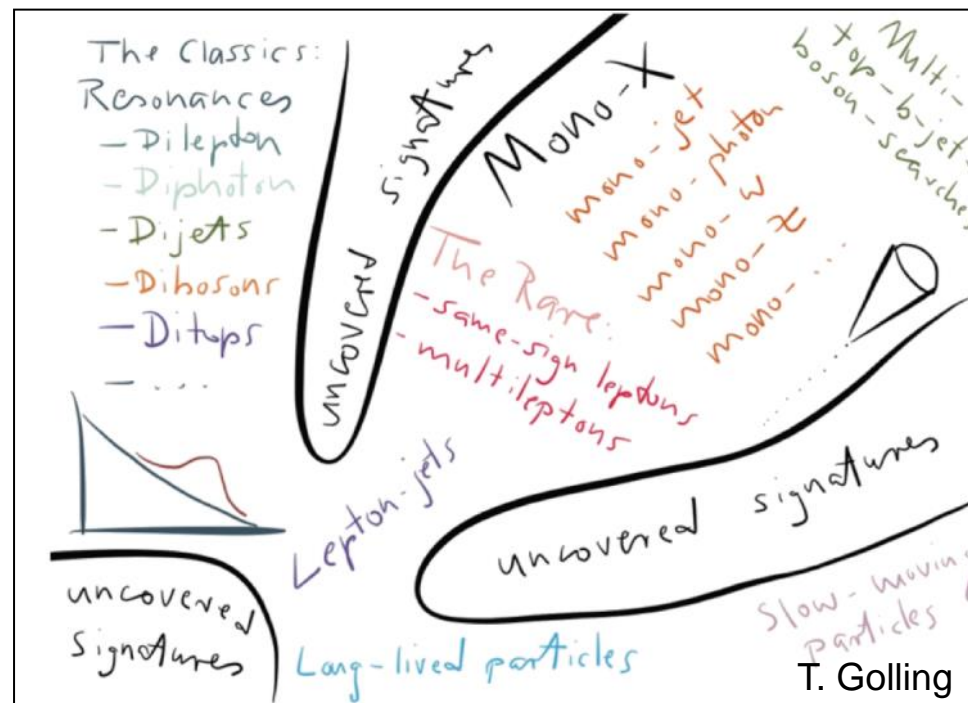
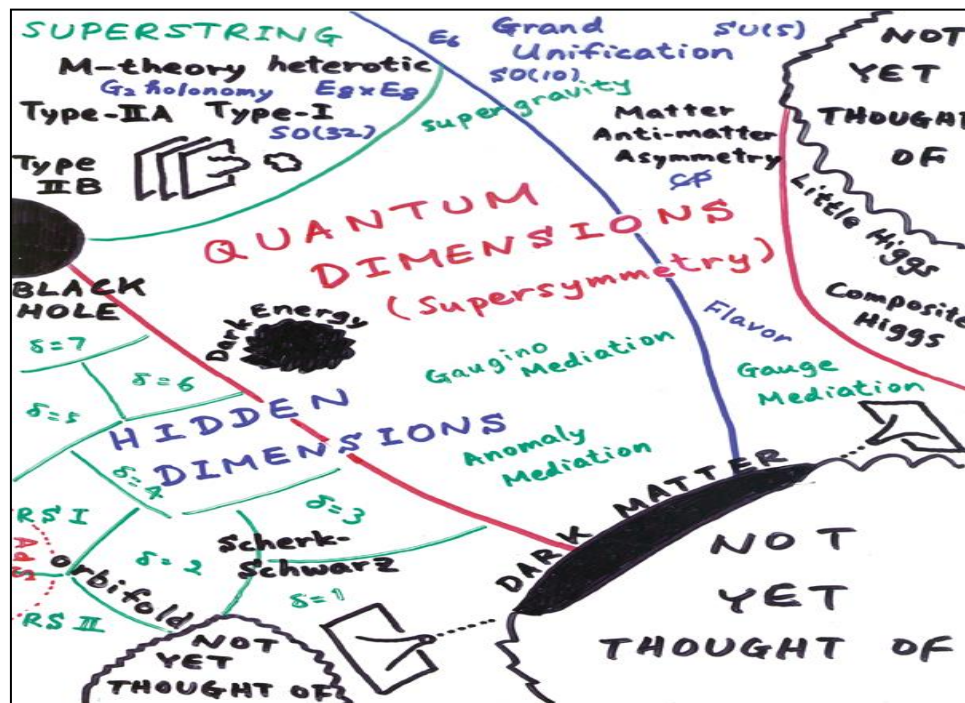
Pseudo-scalar

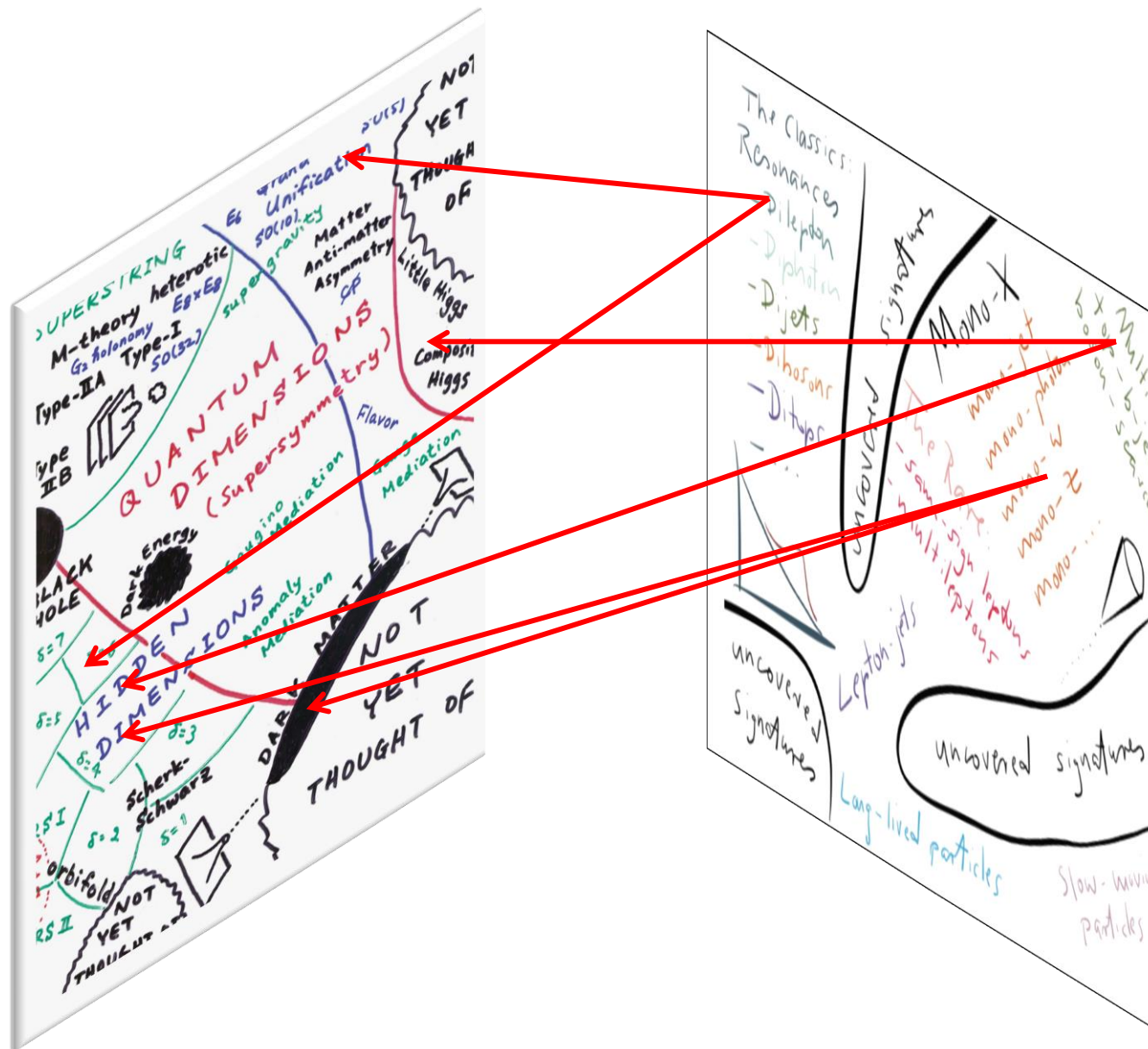
Signature Landscape

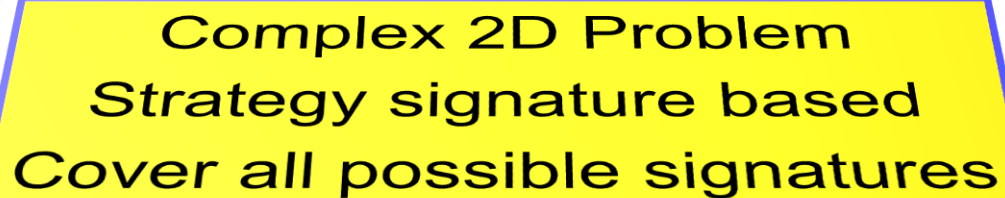
T. Golling



Models and Signatures

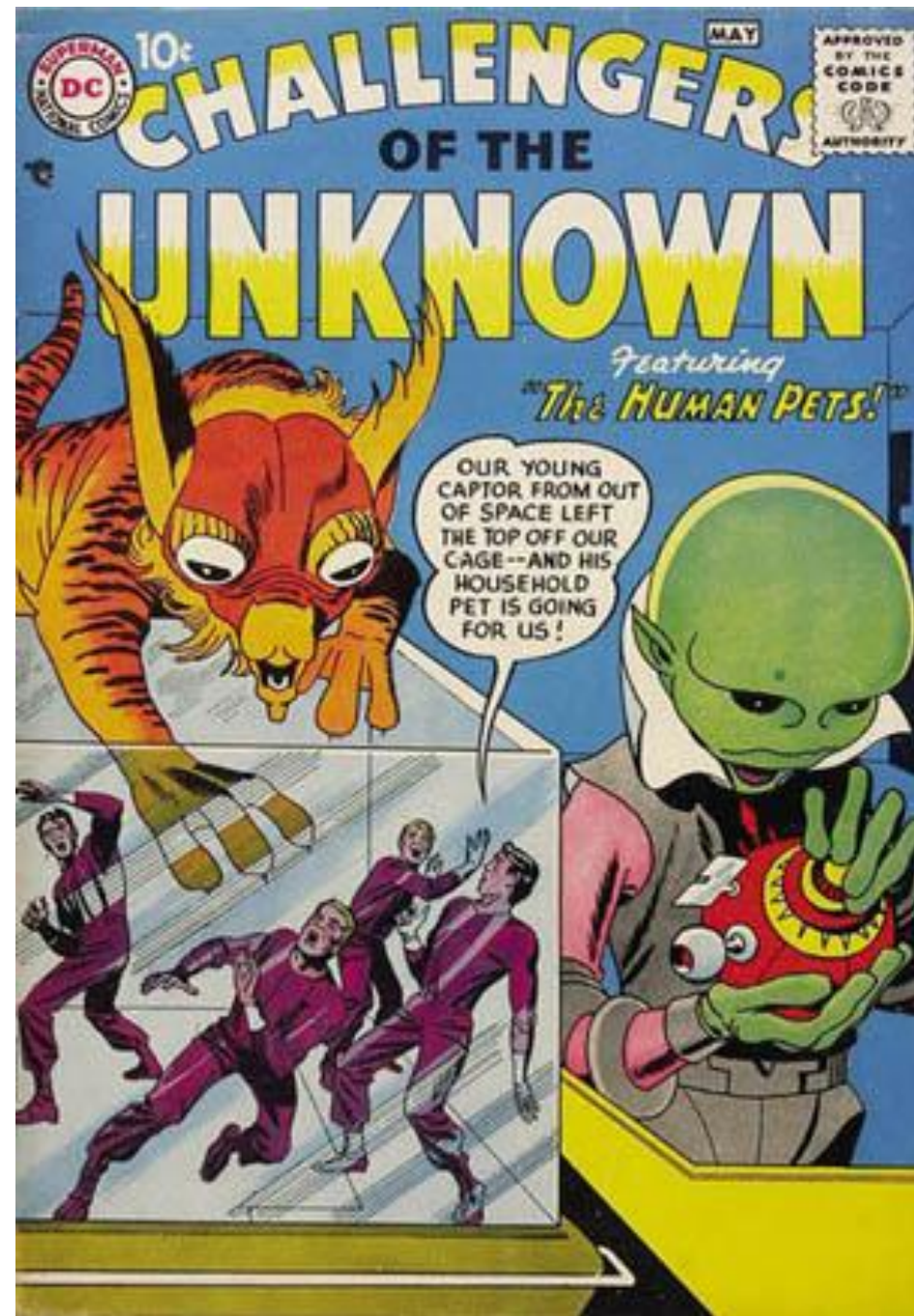






Exotics Searches

- Wide range of final states
- Wide range of models
- GENERIC
 - Look for resonances
 - Look for any disagreement from expectations
- Extremes
 - Experimentally
 - Theoretically



Basic Principles of Exotics Searches

- Identify your discriminant!
- Most important: Robust background estimation!
- Biases ?
 - 100% blind analysis → not appropriate at LHC
 - Control regions
- Trade-off between Signal and Background
 - Do NOT optimize towards a specific model
 - Selection cuts defined by triggers and background reduction.

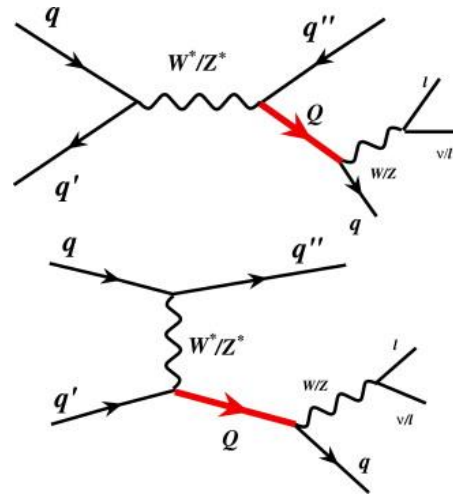
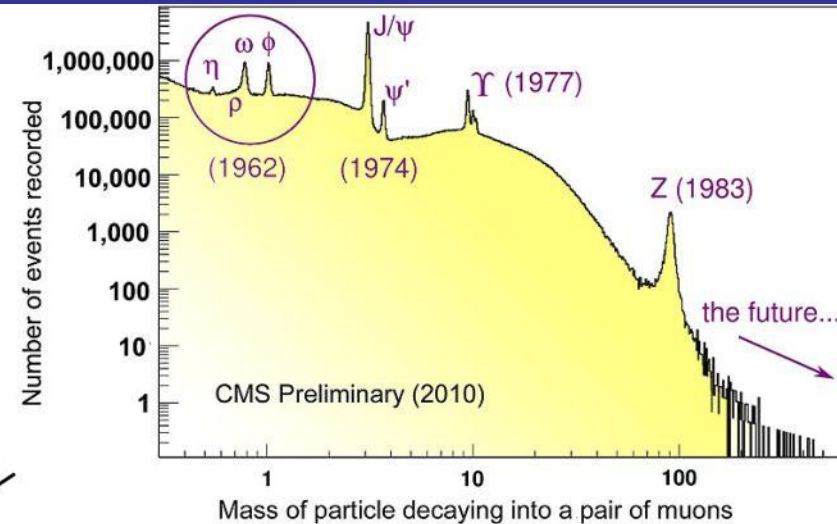
Basic Principles of a Search

- You have a background estimate...what now?
- Check if data agrees with this expectation.
- If it does not agree...
 - Is the significance increasing with more data?
 - Look at time dependences...
 - Cross checks....
 - Discovery if significance is greater than 5 sigma.
- If it does agree....
 - How far did we explore the new physics phase?
 - Use models to quantify the search reach.
 - Useable for others (publish acceptance and efficiencies)

Exotics Searches

■ Heavy resonances

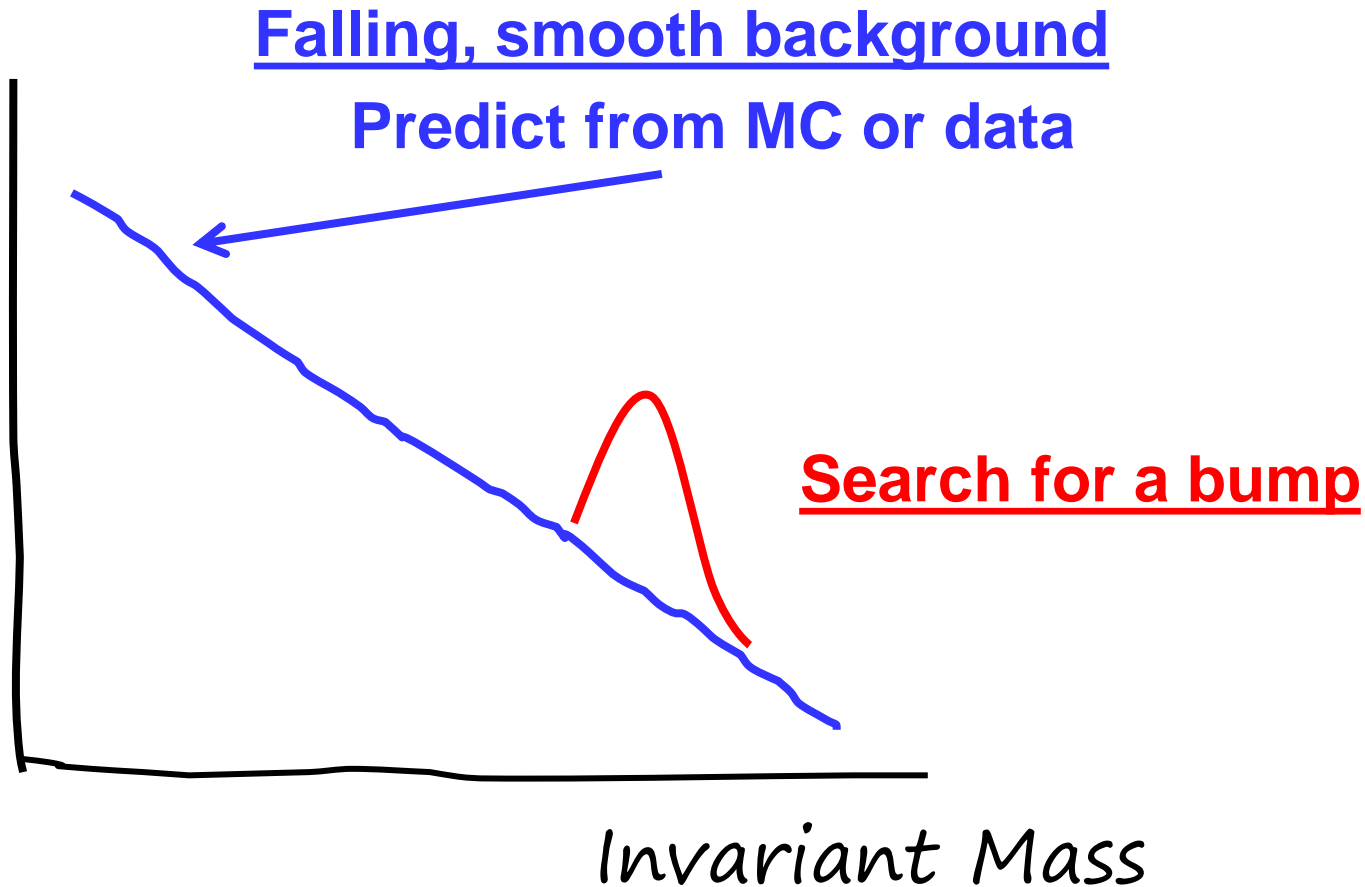
- Dileptons
- Dijets
- $T\bar{t}$
- HH



■ Vector-like quarks

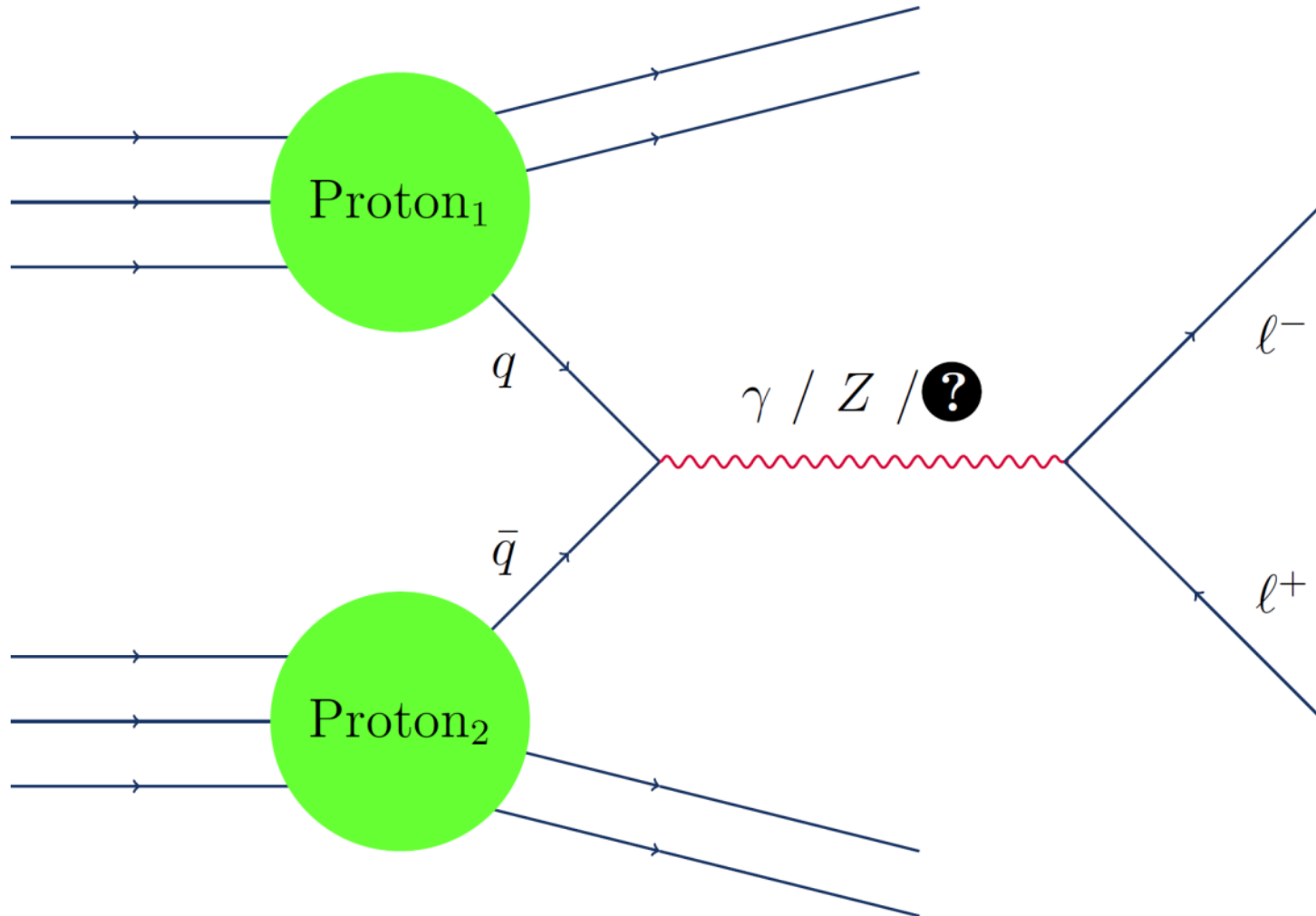
■ Dark matter and extra dimension

Resonance Searches



Dilepton Resonance Search

Noam Tal Hod
CERN-THESIS-2012-155

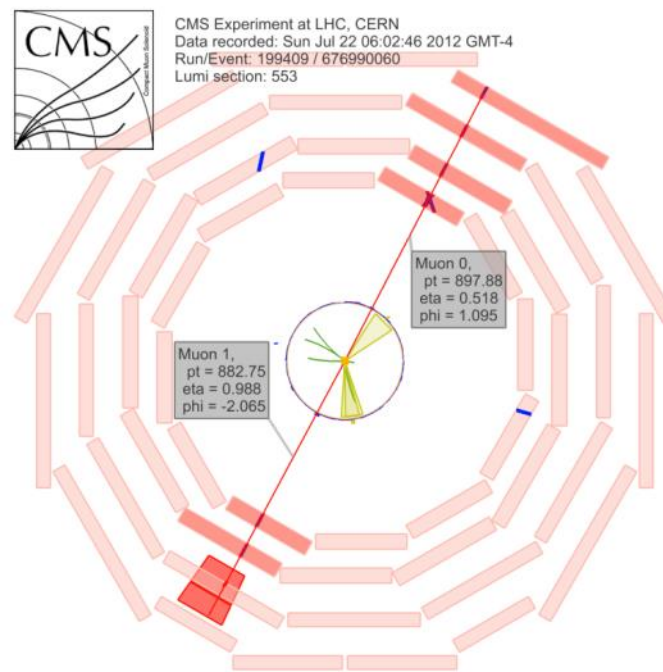


Dilepton Resonance Search

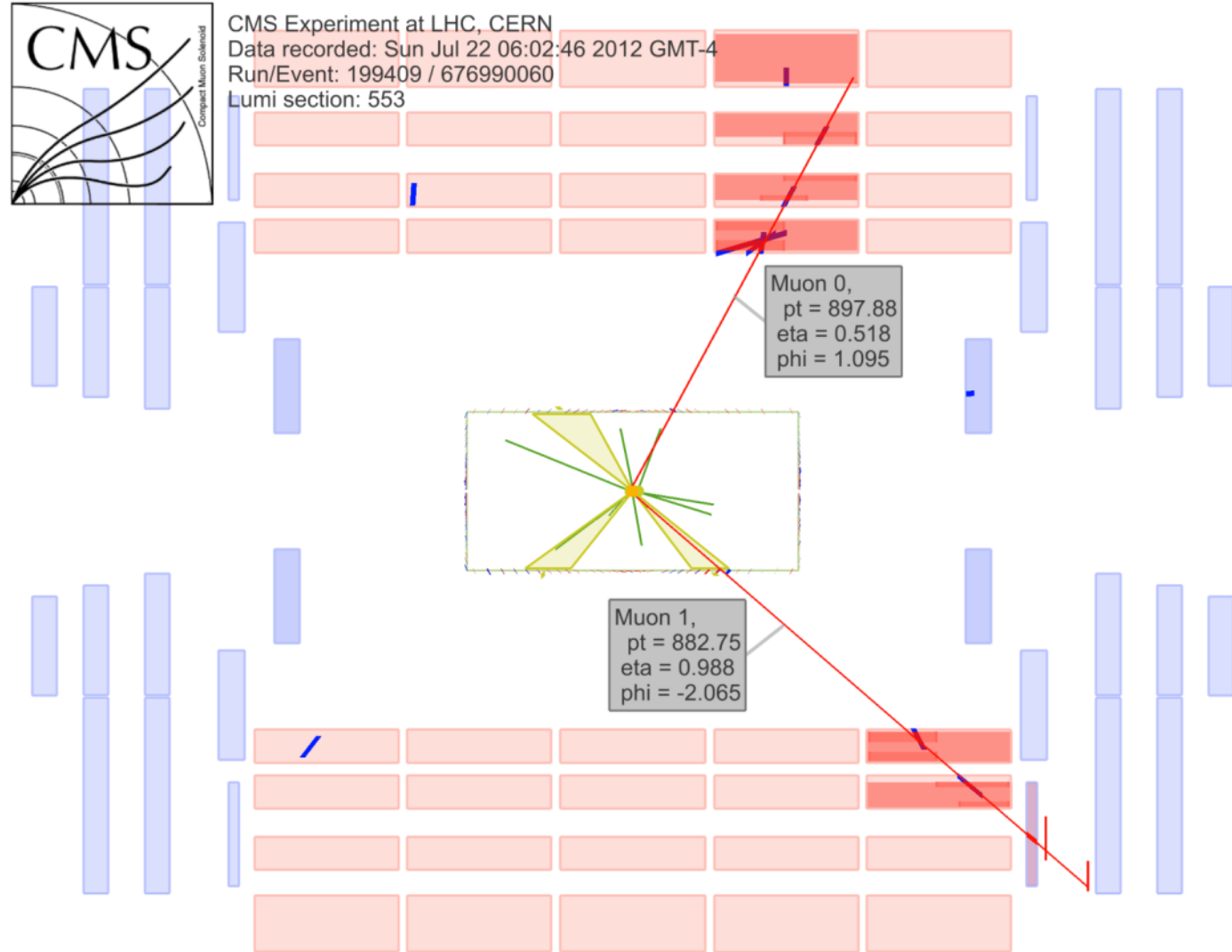
ATLAS-CONF-2013-017
PAS EXO-12-061

- Models:
 - Little Higgs \rightarrow heavy gauge boson(s) (Z'/W')
 - GUT-inspired theories \rightarrow heavy gauge boson(s) (Z'/W')
 - Strong and EWK force merged into one interaction
 - Described by higher symmetry group
 - Popular choices:
 - Left right symmetric models ($SO(10)$)
 - E_6 symmetry models
 - Sequential Standard Model (SSM)
 - Z' carbon copy of Z^0 just heavier
 - Z' decays into any SM lepton-antilepton pair
 - decay into gauge bosons is suppressed by hand
 - not gauge invariant, not very realistic but
 - reference model
 - Randall-Sundrum ED \rightarrow Kaluza-Klein graviton
 - Technicolor \rightarrow narrow techni-hadrons

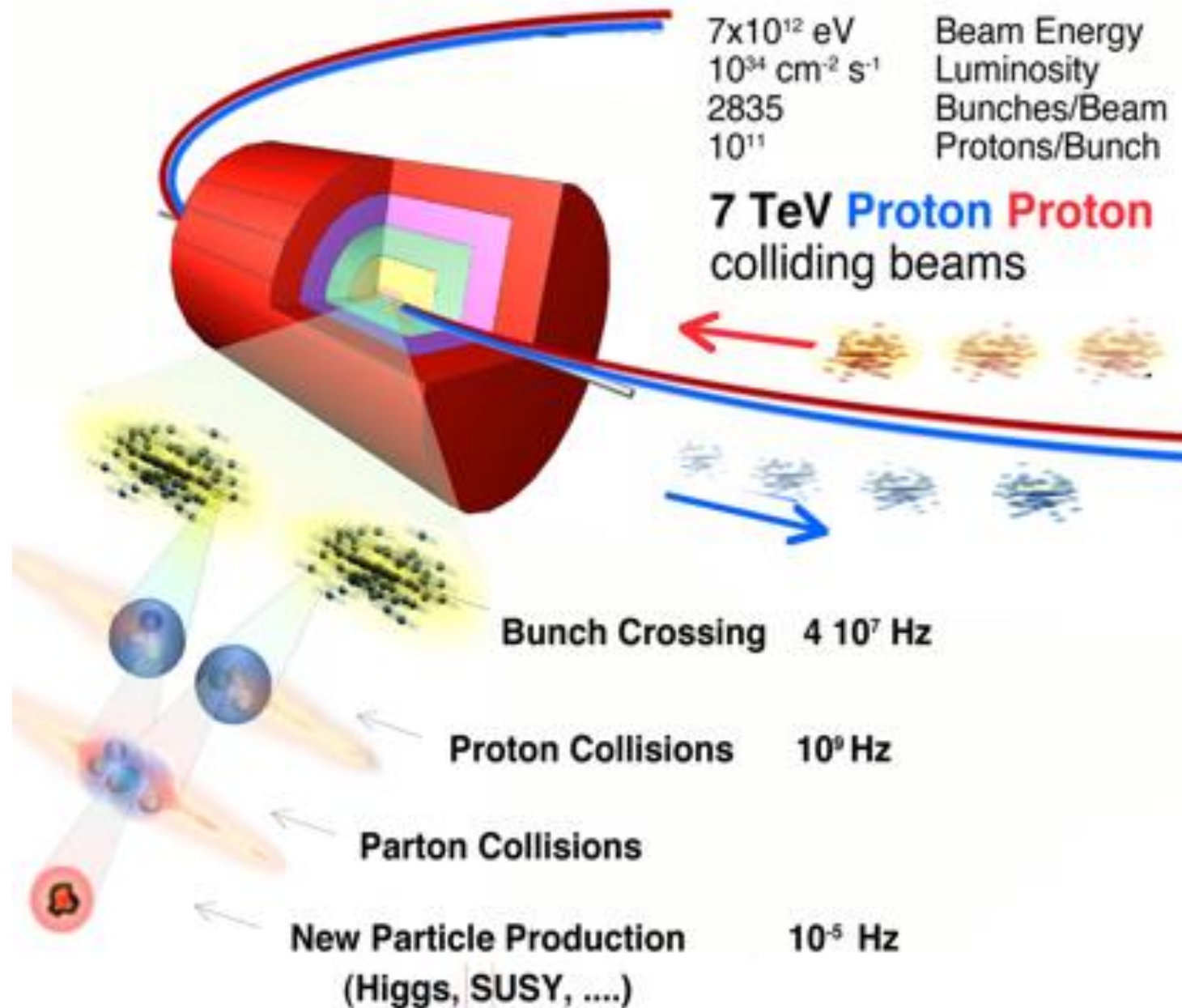
CMS Highest Dimuon Invariant Mass Event; 8 TeV



$m_{\text{inv}} = 1824 \text{ GeV}$



Proton-Proton Collisions



Luminosity

- Single most important quantity
 - Drives ability to observe new rare processes

$$L = \frac{f * n_{\text{bunch}} * N_p^2}{4\pi * \sigma_x * \sigma_y}$$

- revolving frequency $f = 11245.5/\text{s}$
- $n_{\text{bunch}} = 2808$
- $N_p = 1.15 \times 10^{11}$ Protons/Bunch
- Area of beams: $4\pi\sigma_x\sigma_y \sim 40 \mu\text{m}$

- Rate of physics processes per unit time $\sim L$

$$N_{\text{Obs}} = \int L dt * \epsilon * \sigma_{\text{process}}$$

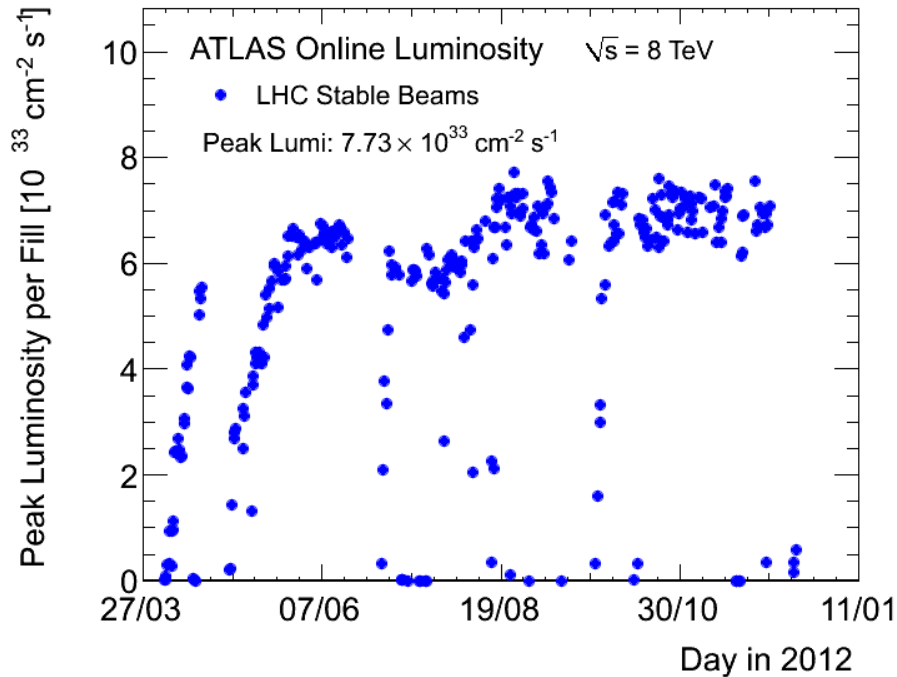
Efficiency; optimized by experimentalists

Cross section; given by nature; predicted by theory

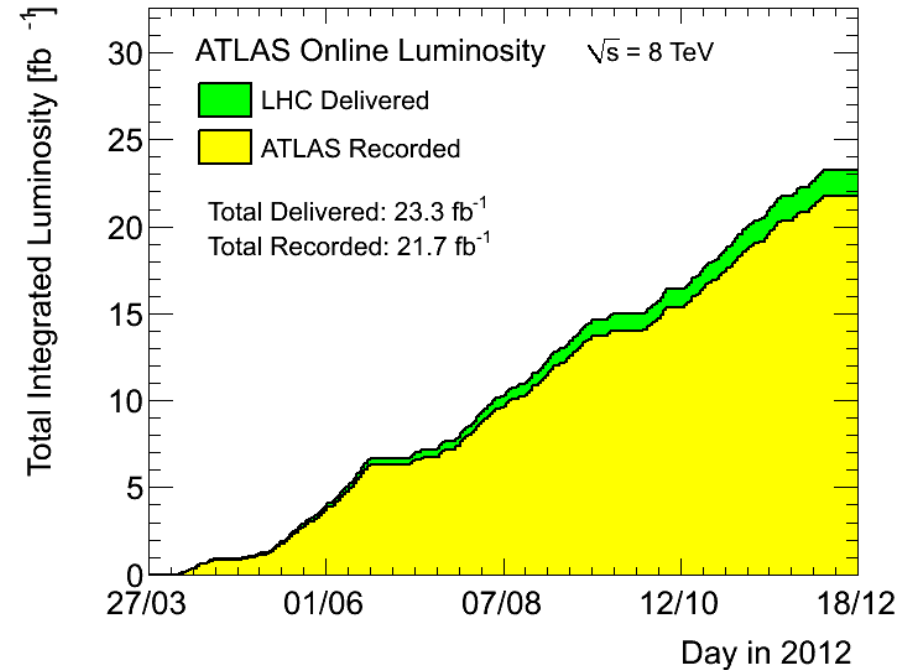
Maximize $N_{\text{obs}} \rightarrow \max \epsilon$ and L

Our data sample for 2012

Peak Luminosity in 2012



Integrated Luminosity in 2012



Delivered Integrated L: 23.3 fb^{-1}

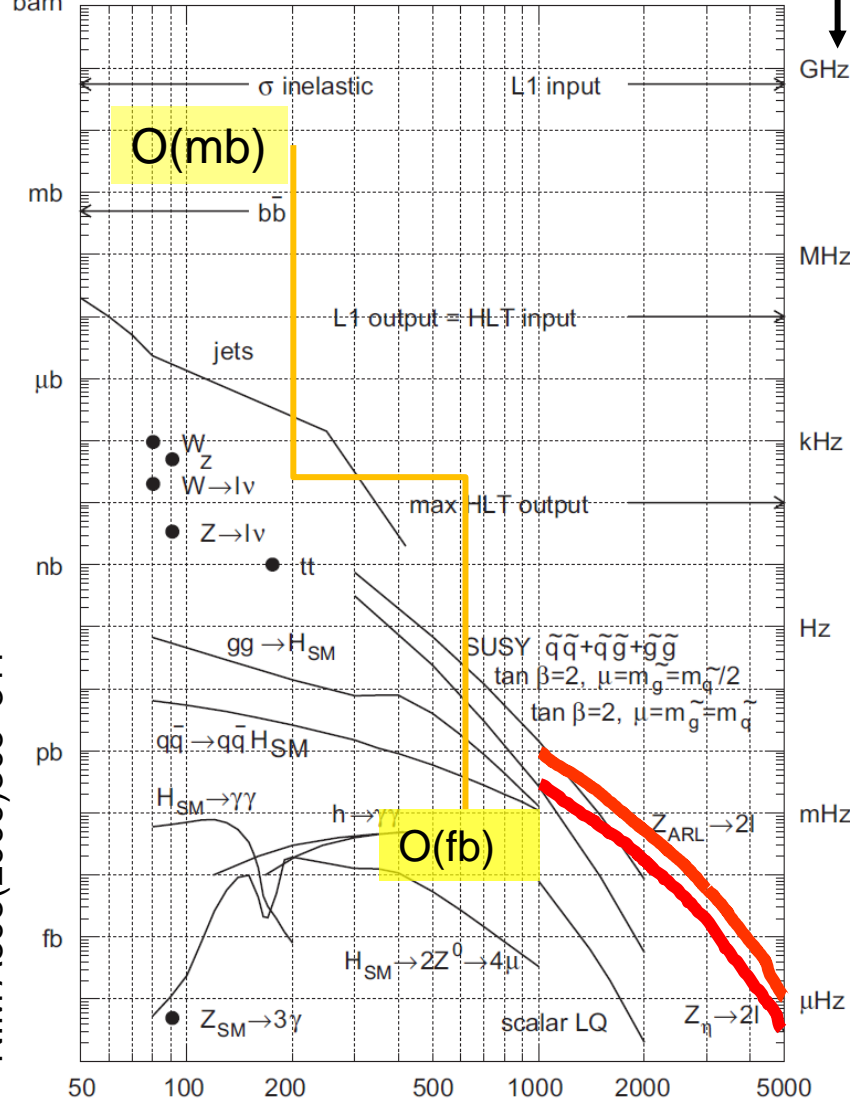
Recorded Integrated L: 21.7 fb^{-1}

$$1b = 10^{-24} \text{ cm}^2$$

$$1fb = 10^{-39} \text{ cm}^2$$

Rates of physics processes @ LHC

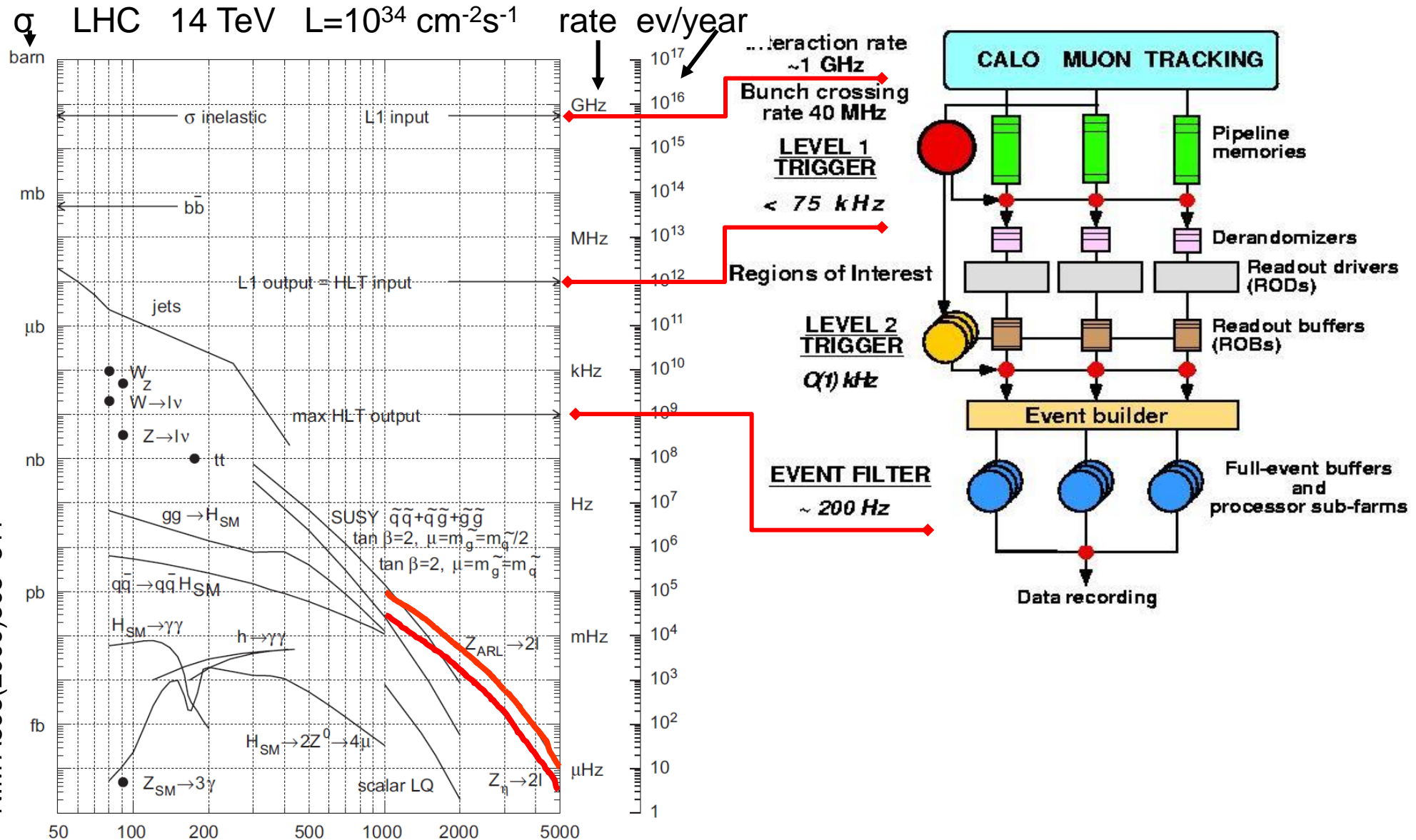
σ LHC 14 TeV $L=10^{34} \text{ cm}^{-2}\text{s}^{-1}$ rate ev/year



Interesting physics swamped by background

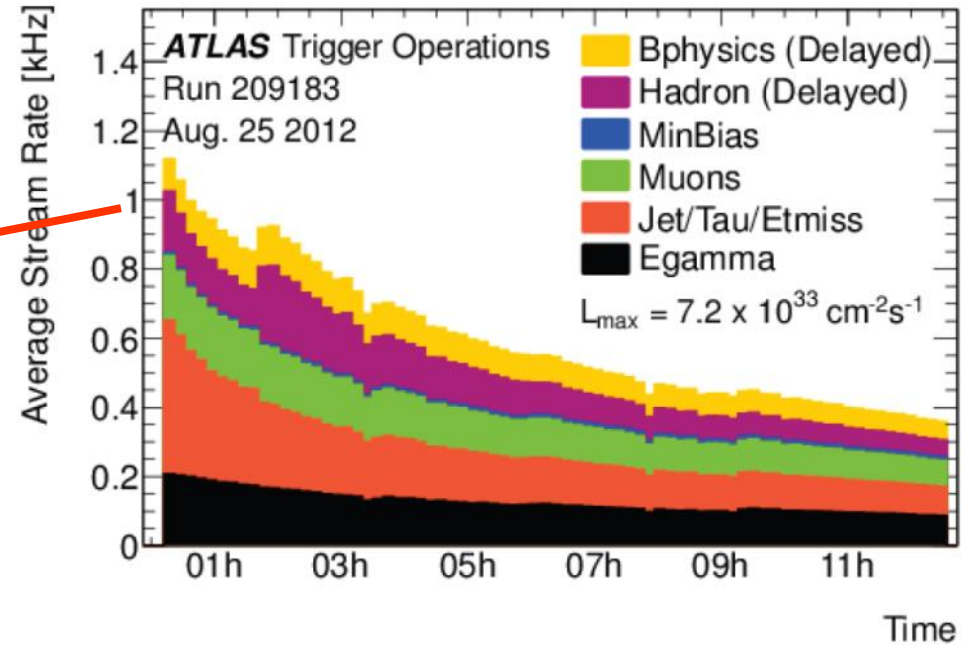
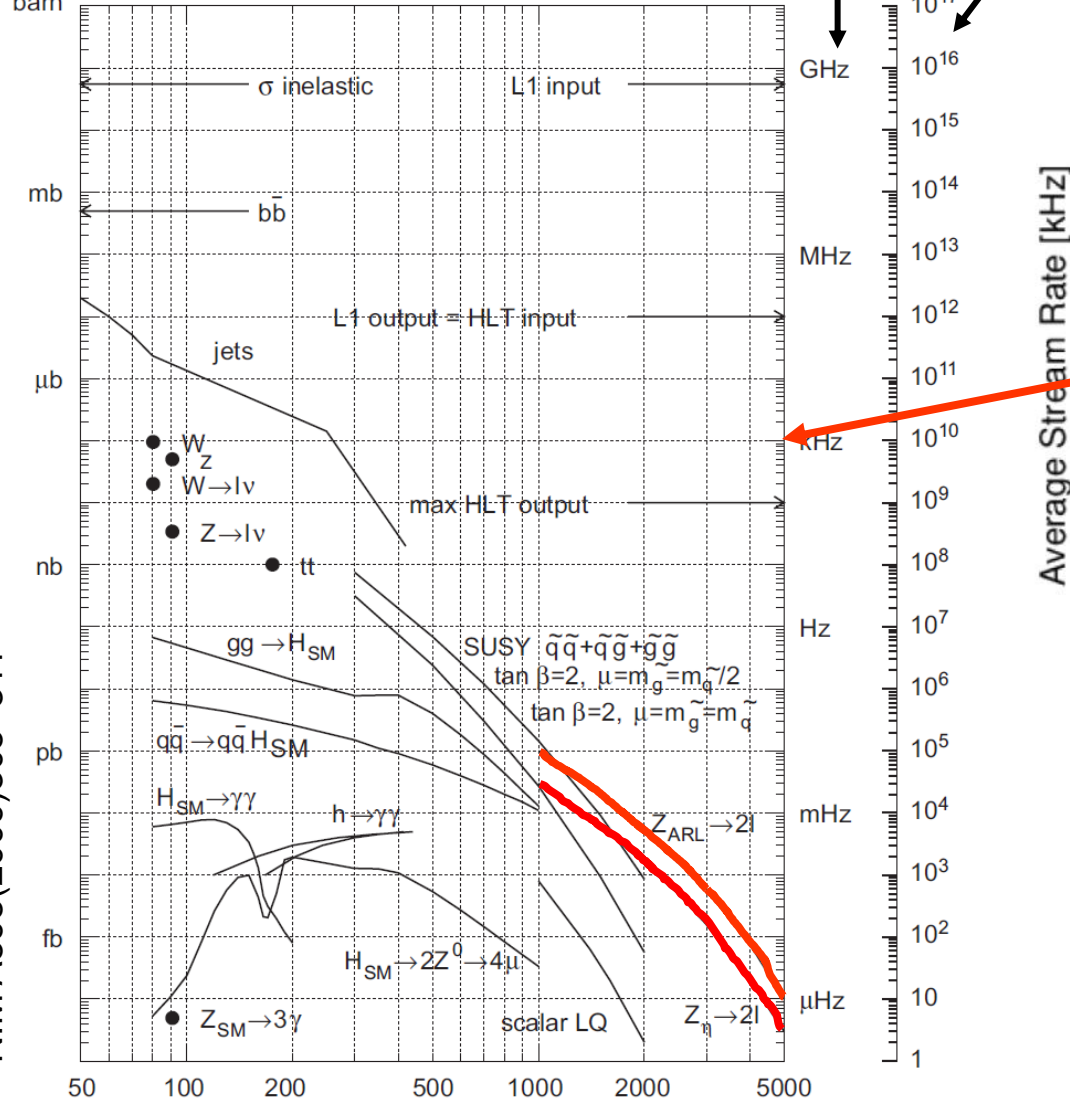
- Cross section for new physics:
 - $\sim 10^{12}$ times lower !!
- Need to filter \rightarrow TRIGGER SYSTEMS
- Carefully decide what to record
- You do not have another chance

Compare this to rates of physics processes



Compare this to rates of physics processes

σ LHC 14 TeV $L=10^{34} \text{ cm}^{-2}\text{s}^{-1}$ rate ev/year



NIM A598(2009)305-311

Dilepton Resonance Search: Trigger Strategy

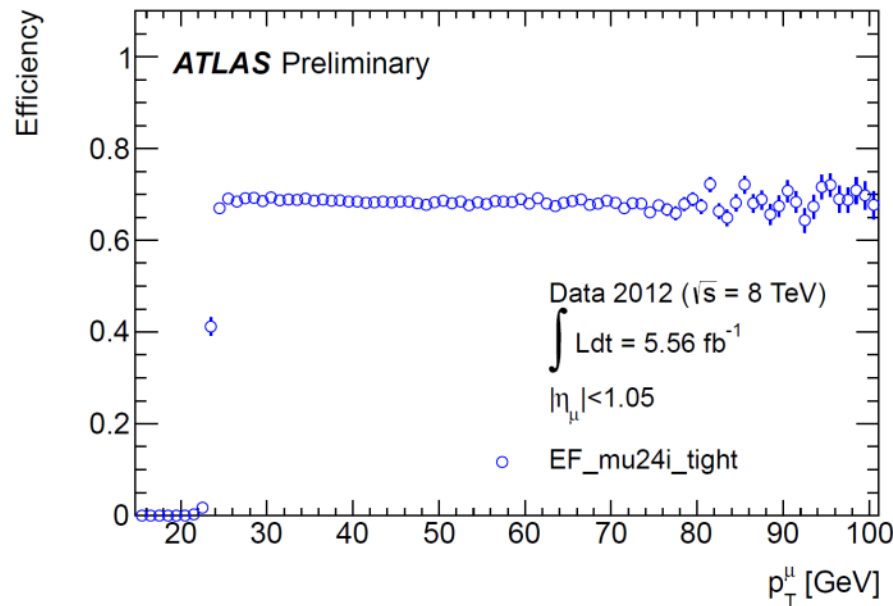
ATLAS

ee channel

- Diphoton trigger
- $E_T > 35$ GeV and $E_T > 25$ GeV

$\mu\mu$ channel

- Single muon triggers
- $E_T > 24$ GeV or $E_T > 36$ GeV



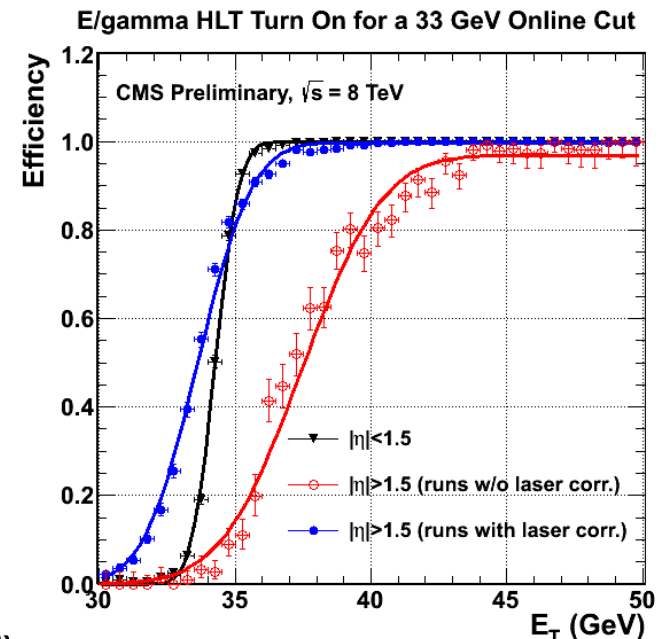
CMS

ee channel

- Dielectron trigger
- Both clusters w $E_T > 33$ GeV

$\mu\mu$ channel

- single muon trigger
- $E_T > 40$ GeV



CMS Di-Electron Event Zoomed into Inner Detector

PAS EXO-12-061

CMS Experiment at LHC, CERN
Data recorded: Tue Aug 21 07:30:43 2012 CEST
Run/Event: 201278 / 2107823234
Lumi section: 2053

Track $p_T > 3$ GeV

Electron 0,
 $p_T = 541.32$
 $\eta = -0.027$
 $\phi = 0.041$

Electron 1,
 $p_T = 587.69$
 $\eta = -1.941$
 $\phi = -3.094$

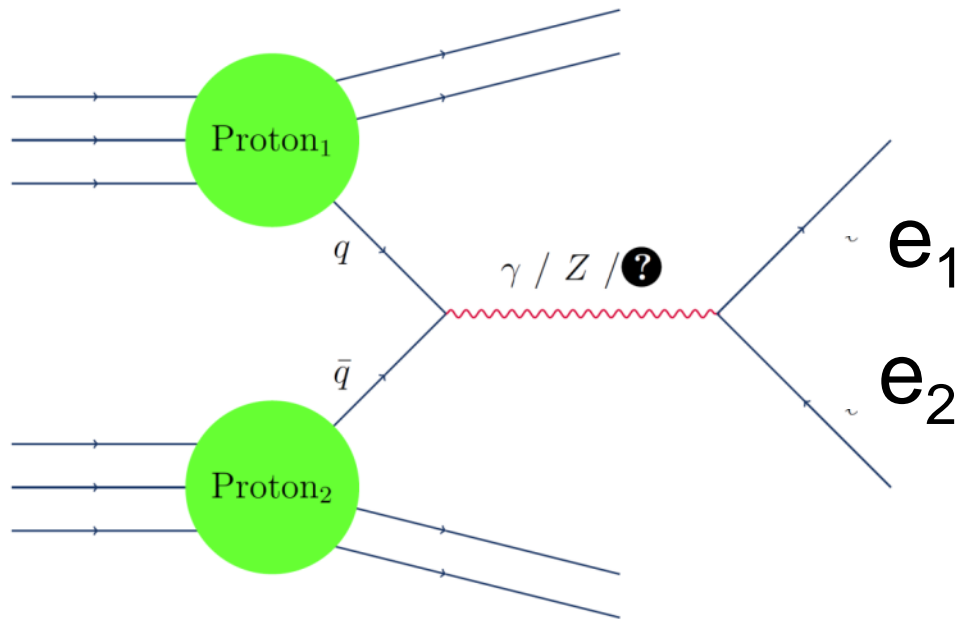
CMS barrel pixel detector

Multiple interaction vertices

CMS barrel silicon strip

Require ≥ 1 Vertex
ATLAS: + ≥ 2 tracks
CMS: + ≥ 4 tracks

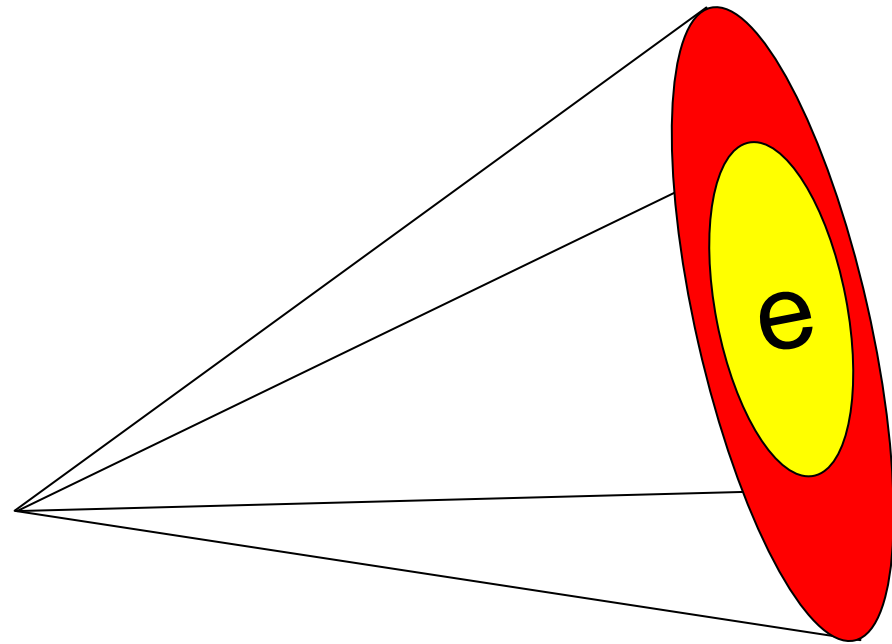
Selection for Di-Electron Channel



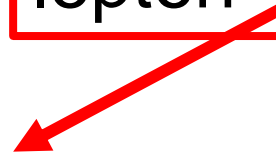
ATLAS	CMS
$E_T^1 > 40 \text{ GeV}$	$E_T^1 > 35 \text{ GeV}$
$E_T^2 > 30 \text{ GeV}$	$E_T^2 > 35 \text{ GeV}$

Problem: jets fake electrons
Use isolation to reduce fakes

Electron Isolation I_{conesize}



Energy/momentum around lepton



	ATLAS	CMS	
leading	$I_{\text{calo}}^{0.2} < 0.7\% \cdot E_T + 5 \text{ GeV}$	$I_{\text{tracker}}^{0.3} < 5 \text{ GeV}$	$I_{\text{Calo}}^{0.3} < 3\% \cdot E_T$
subleading	$I_{\text{calo}}^{0.2} < 2.2\% \cdot E_T + 6 \text{ GeV}$		

Acceptance x Efficiency after all Selections

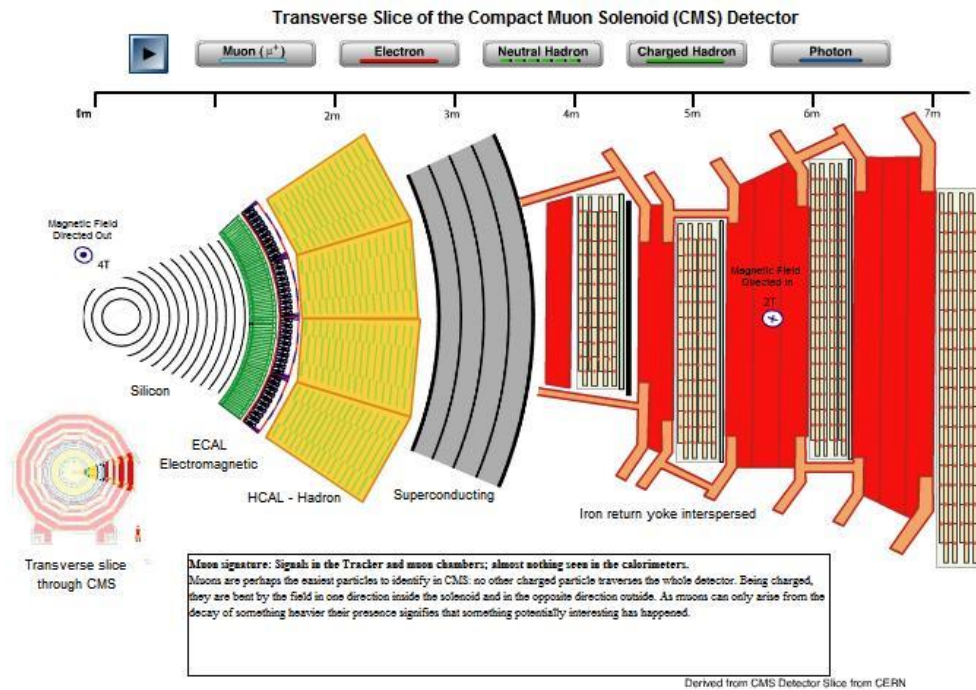
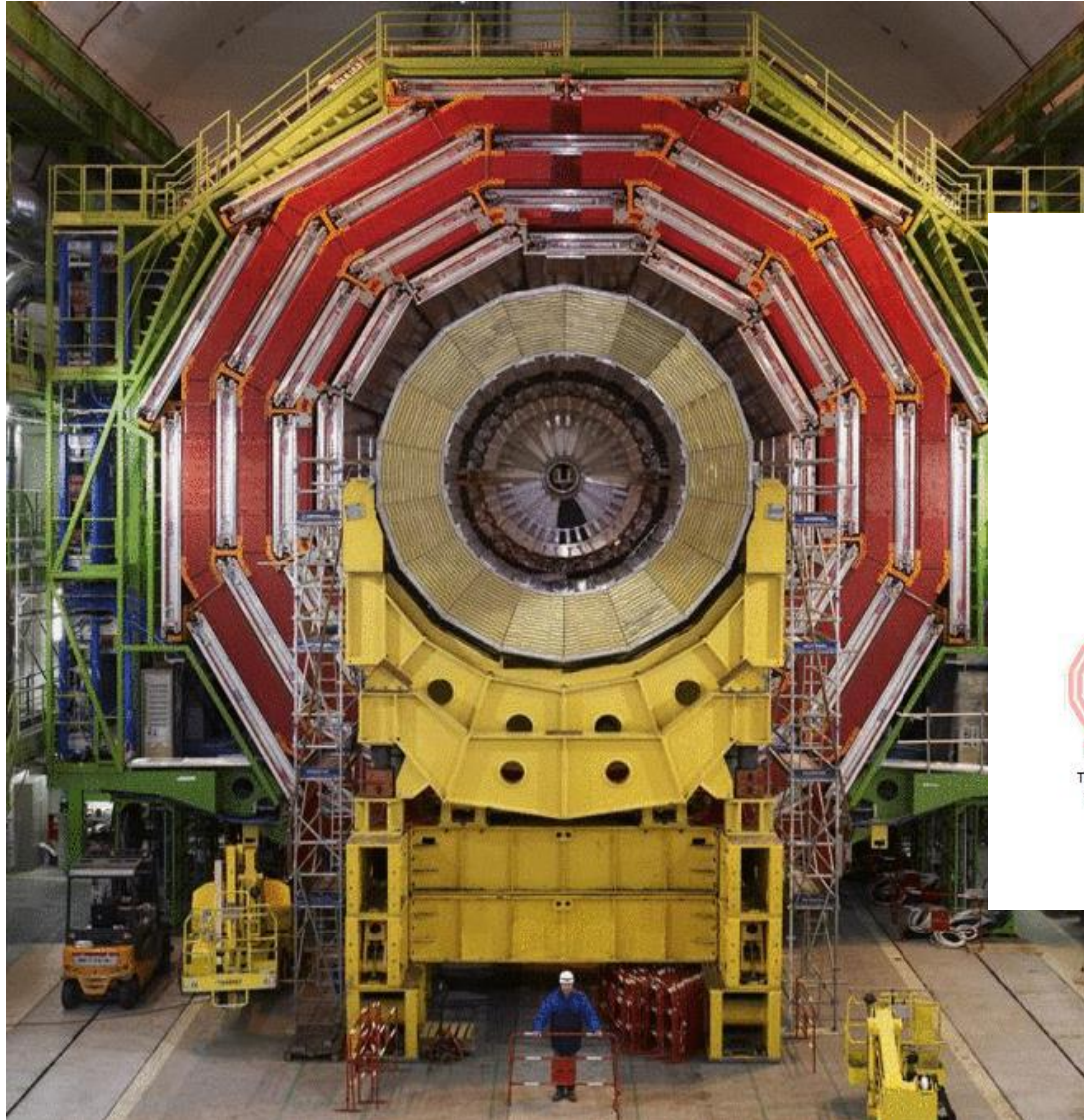
ATLAS

CMS

$$A \times \epsilon(m = 2 \text{ TeV}) = \mathbf{73\%} \quad A \times \epsilon(m = 2.5 \text{ TeV}) = \mathbf{67\%}$$

Similar

Di-Muon Channel



Dilepton Resonance Search:: $\mu\mu$ selections

ATLAS

- Single muon triggers
- $p_T > 25 \text{ GeV}$
- $|\eta| < 2.4$
- Suppress cosmic rays
 - $|d_0| < 0.2 \text{ mm}$
 - $|z_0 - z(\text{vertex})| < 1 \text{ mm}$
- Suppress jets faking μ 's
 - $\sum p_T(\Delta R < 0.3) < 5\% \cdot p_T$
- Require opposite charge

CMS

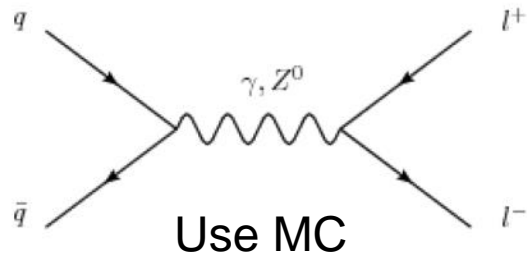
- Single muon trigger
- $p_T > 45 \text{ GeV}$
- $|\eta| < 2.4$
- Suppress cosmic rays
 - $|d_0| < 0.2 \text{ mm}$
 - $|z_0 - z(\text{vertex})| < 24 \text{ cm}$
- Suppress jets faking μ 's
 - $\sum p_T(\Delta R < 0.3) < 10\% \cdot p_T$
 - $|z_0 - z(\text{vertex})| < 0.2 \text{ mm}$
- Require opposite charge

Very different

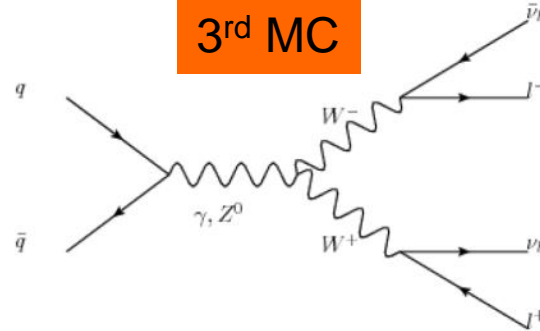
$$A \times \epsilon(m = 2 \text{ TeV}) = 46\% \quad A \times \epsilon(m = 2.5 \text{ TeV}) = 80\%$$

Dilepton Resonance Search: Backgrounds ee

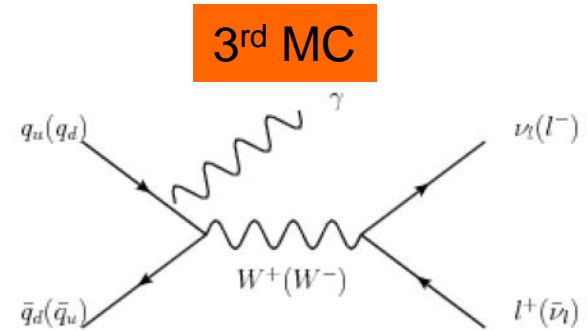
dominant & irreducible



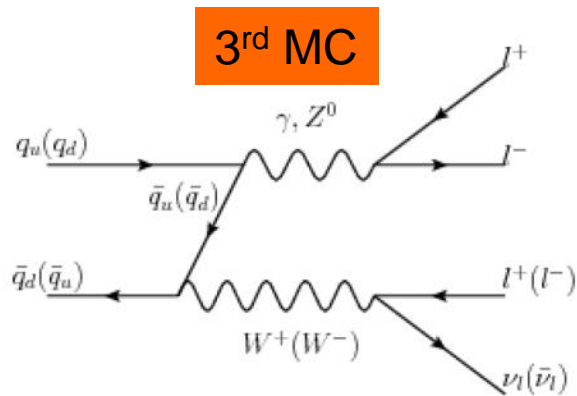
(a) Drell-Yan



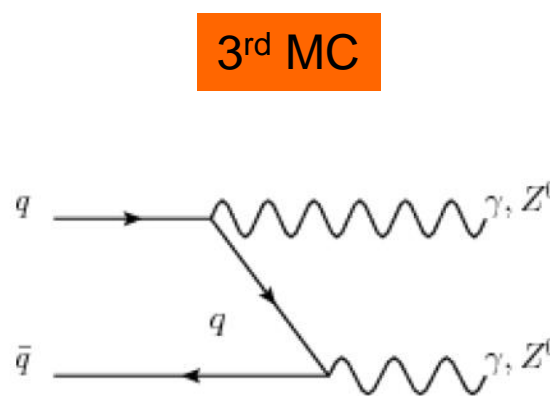
(b) WW



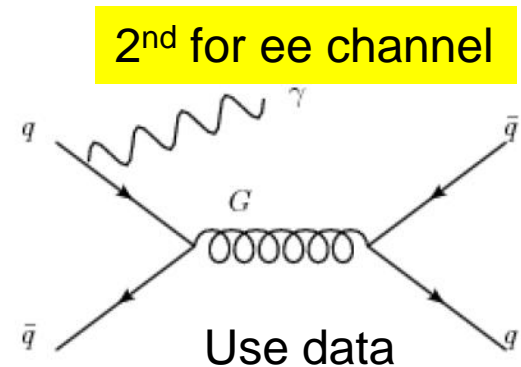
(c) W γ



(d) WZ, W γ



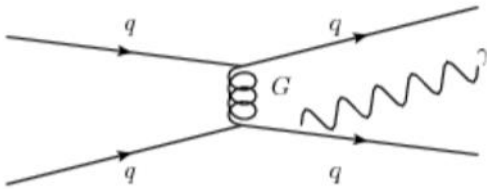
(e) ZZ, Z γ , $\gamma\gamma$



(f) Dijets (without the external photon line), γ +jets

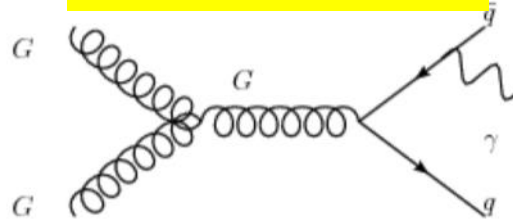
Dilepton Resonance Search: Backgrounds ee

2nd for ee channel data



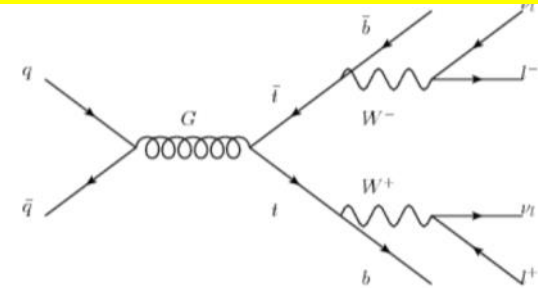
(g) Dijets (without the external photon line), γ +jets

2nd for ee channel data



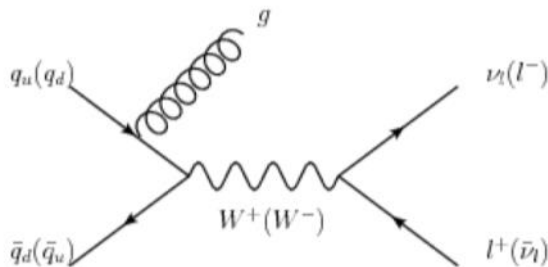
(h) Dijets (without the external photon line), γ +jets

2nd for ee channel semi-leptonic

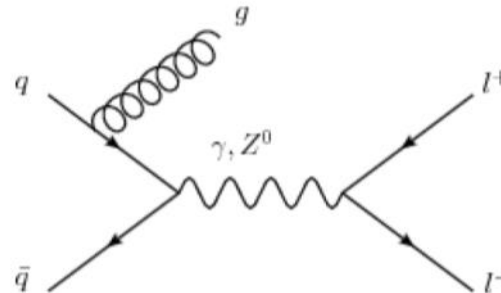


(i) $t\bar{t}$

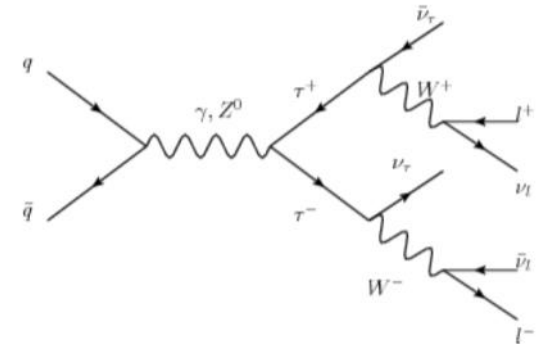
2nd for ee channel data



(j) W +jets



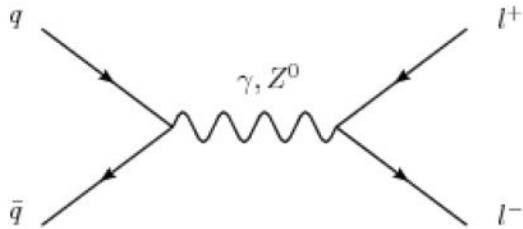
(k) Z +jets



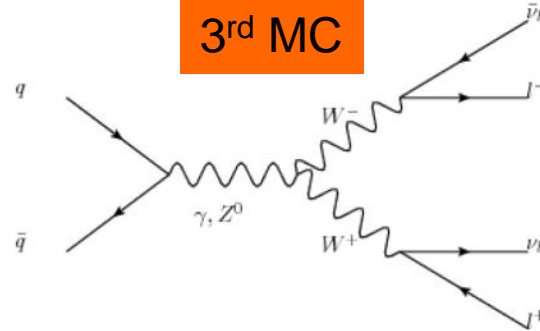
(l) DY to taus to leptons

Dilepton Resonance Search: Backgrounds $\mu\mu$

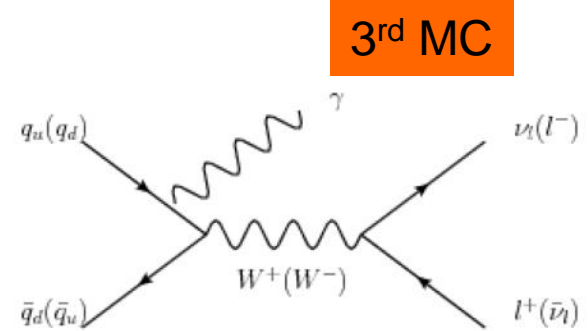
dominant & irreducible mc



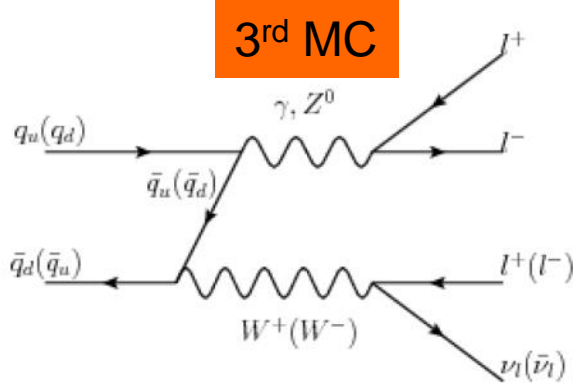
(a) Drell-Yan



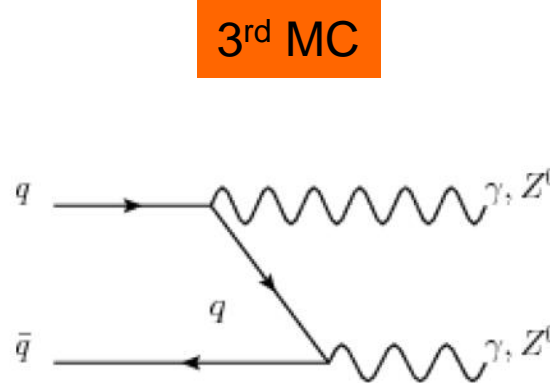
(b) WW



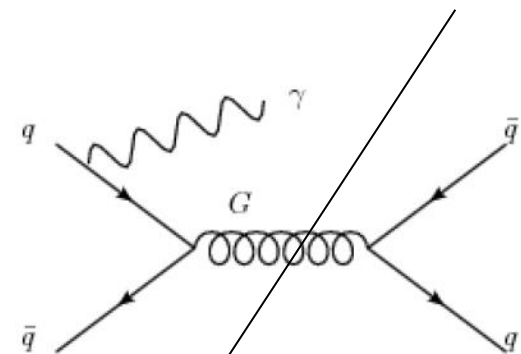
(c) W γ



(d) WZ, W γ

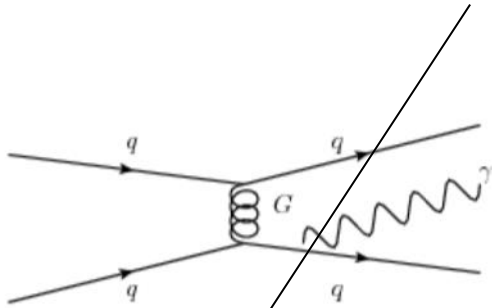


(e) ZZ, Z γ , $\gamma\gamma$

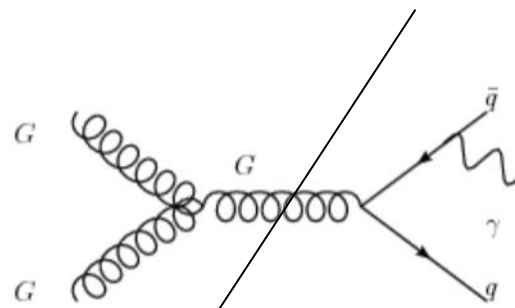


(f) Dijets (without the external photon line), γ +jets

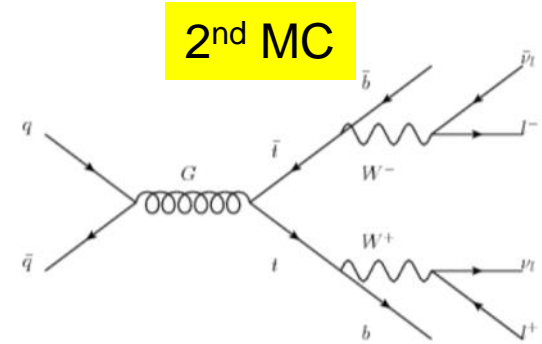
Dilepton Resonance Search: Backgrounds $\mu\mu$



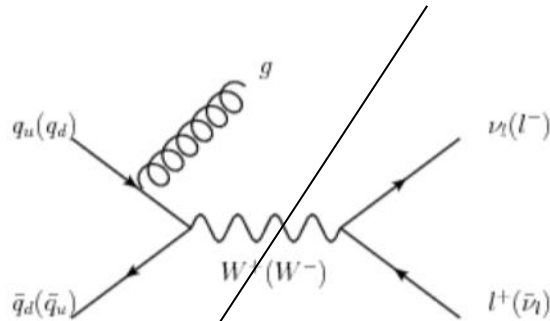
(g) Dijets (without the external photon line), γ +jets



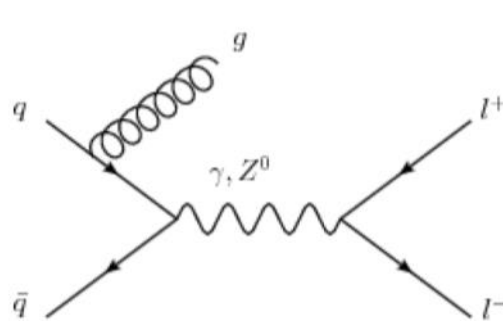
(h) Dijets (without the external photon line), γ +jets



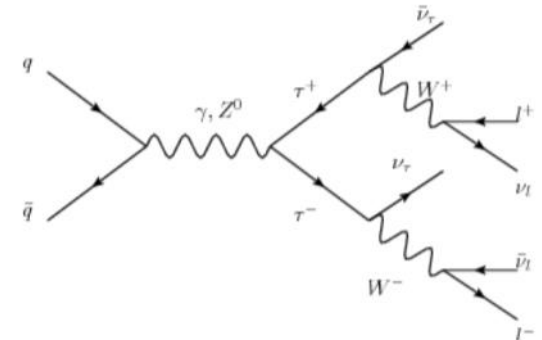
(i) $t\bar{t}$



(j) W +jets



(k) Z +jets



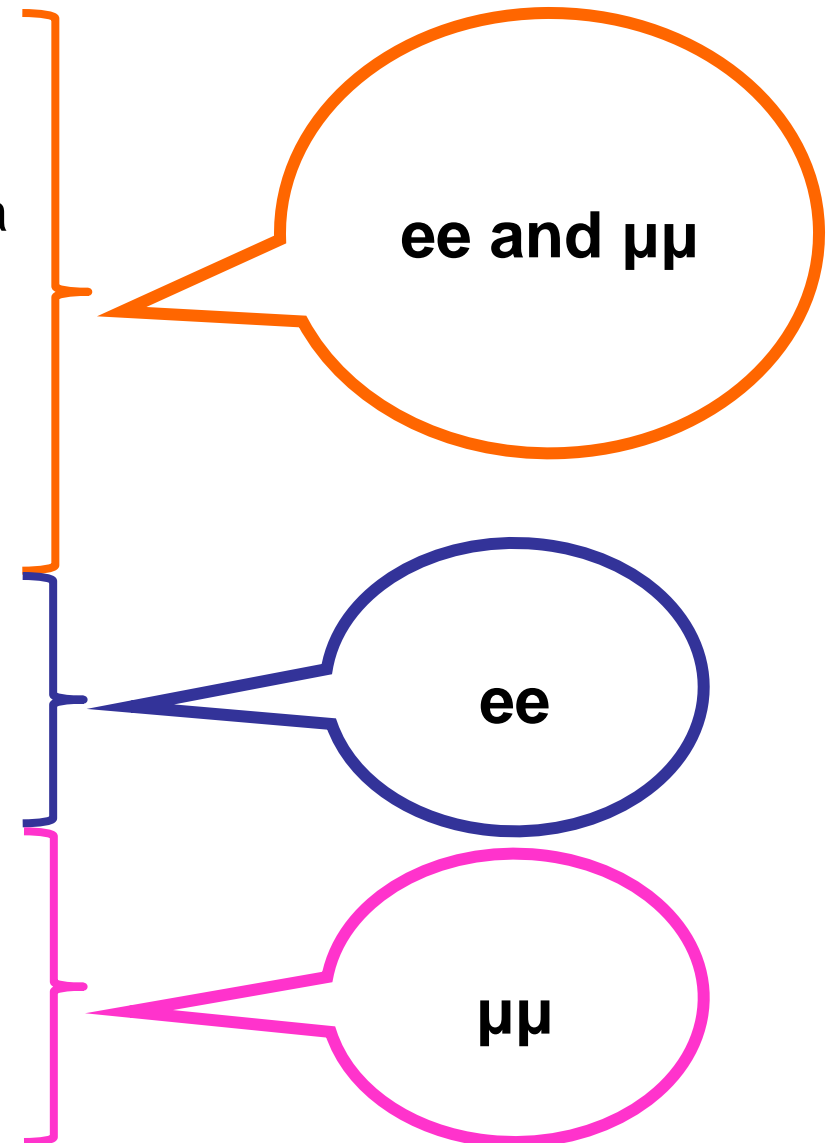
(l) DY to taus to leptons

2nd MC

Heavy Resonances Search: 8 TeV Dileptons

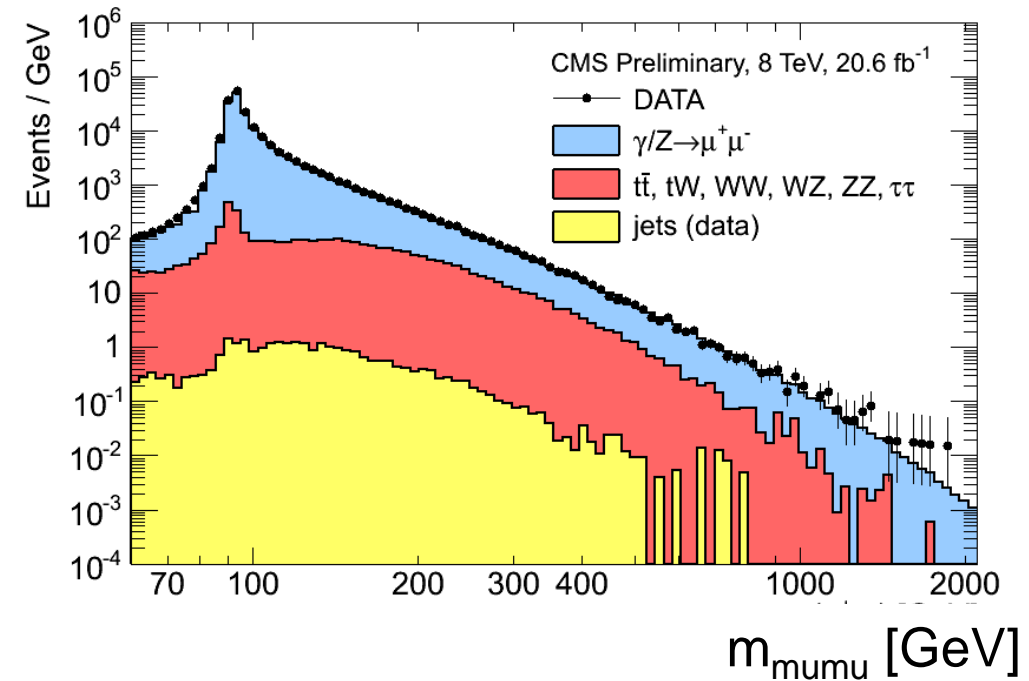
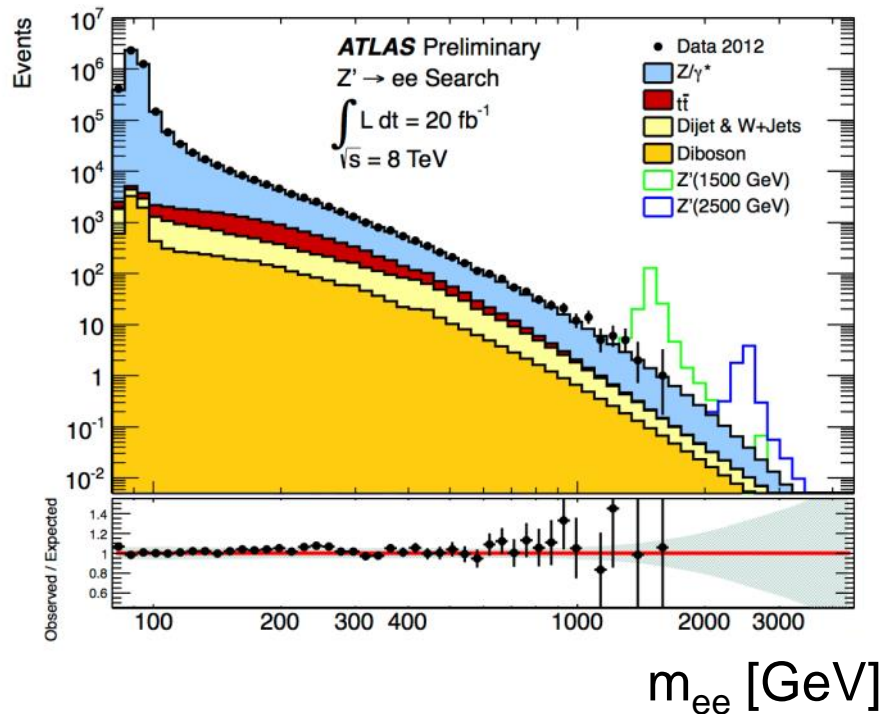
Backgrounds

- SM Drell-Yan: $\gamma^*/Z \rightarrow l^+l^-$
 - shape taken from Monte Carlo
 - normalisation taken from Z peak in data
- t-tbar:
 - where tt goes to e+e-, mu+mu-
 - est. from MC, cross-checked in data
 - also includes $Z \rightarrow \tau\tau$, WW, WZ
- Jet Background:
 - di-jet, W+jet events where the jets are misidentified as electrons/muons
- Cosmic Ray Background:
 - muons from cosmic rays
 - estimated <0.1 event after vertex and angular difference requirements



Dilepton Search: The Discriminant

ATLAS-CONF-2013-017
PAS EXO-12-061



Invariant mass reach of 1 - 2 TeV

Dilepton Resonance Search: Systematic Uncertainties

Source	Dielectrons		Dimuons	
	Signal	Background	Signal	Background
Normalization	5%	NA	5%	NA
PDF variation	NA	15%	NA	15%
PDF choice	NA	17%	NA	17%
Scale	NA	-	NA	-
α_s	NA	4%	NA	4%
Electroweak corrections	NA	3%	NA	3%
Photon-induced corrections	NA	4%	NA	4%
Efficiency	-	-	6%	6%
Resolution	-	-	-	3% (7%)
W + jet and multi-jet background	NA	9%	NA	-
Diboson and $t\bar{t}$ extrapolation	NA	5%	NA	4%
Total	5%	26%	8%	25% (26%)

Heavy Resonances Search: 8 TeV Dileptons

m_{ee} [GeV]	110 - 200	200 - 400	400 - 800	800 - 1200	1200 - 3000	3000 - 4500
Z/γ^*	119000 ± 8000	13700 ± 900	1290 ± 80	68 ± 4	9.8 ± 1.1	0.008 ± 0.005
$t\bar{t}$	7000 ± 800	2400 ± 400	160 ± 60	2.5 ± 0.6	0.11 ± 0.04	< 0.001
Diboson	1830 ± 210	660 ± 160	93 ± 33	4.8 ± 0.8	0.79 ± 0.26	0.005 ± 0.004
Dijet, W + jet	3900 ± 800	1260 ± 310	230 ± 110	8.6 ± 2.4	0.9 ± 0.6	0.004 ± 0.006
Total	131000 ± 8000	18000 ± 1100	1780 ± 150	84 ± 5	11.6 ± 1.3	0.017 ± 0.009
Data	133131	18570	1827	98	10	0

ATLAS-CONF-2013-017

$m_{\mu\mu}$ [GeV]	110 - 200	200 - 400	400 - 800	800 - 1200	1200 - 3000	3000 - 4500
Z/γ^*	111000 ± 8000	11000 ± 1000	1000 ± 100	49 ± 5	7.3 ± 1.3	0.033 ± 0.029
$t\bar{t}$	5900 ± 900	1900 ± 400	140 ± 60	2.7 ± 0.7	0.16 ± 0.08	< 0.001
Diboson	1520 ± 190	520 ± 140	62 ± 26	2.8 ± 1.0	0.38 ± 0.28	0.002 ± 0.003
Total	118000 ± 8000	13300 ± 1100	1160 ± 120	55 ± 5	7.8 ± 1.3	0.035 ± 0.029
Data	118701	13349	1109	48	8	0

What do you do now?

- Observed numbers consistent with background???
- Many ways to do it → Statistics Lectures/Tutorial
- One way e.g.:

- $P(n \geq n_{obs}) = 1 - f(n; s = 0; b) = 1 - \sum_{n=0}^{n_{obs}-1} \frac{b^n}{n!} e^{-b}$

- Probability, assuming $s = 0$, to observe as many events or more for a given expected background amount, b .

- For 800 – 1200 GeV bin in $\mu\mu$

- $b = 55, n_{obs}=48 \rightarrow P=84\%$

Heavy Resonances Search: 8 TeV Dileptons

ATLAS-CONF-2013-017

m_{ee} [GeV]	110 - 200	200 - 400	400 - 800	800 - 1200	1200 - 3000	3000 - 4500
Z/γ^*	119000 ± 8000	13700 ± 900	1290 ± 80	68 ± 4	9.8 ± 1.1	0.008 ± 0.005
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Total	131000 ± 8000	18000 ± 1100	1780 ± 150	84 ± 5	11.6 ± 1.3	0.017 ± 0.009
Data	133131	18570	1827	98	10	0

Analysis: $P(ee) = 18\%$

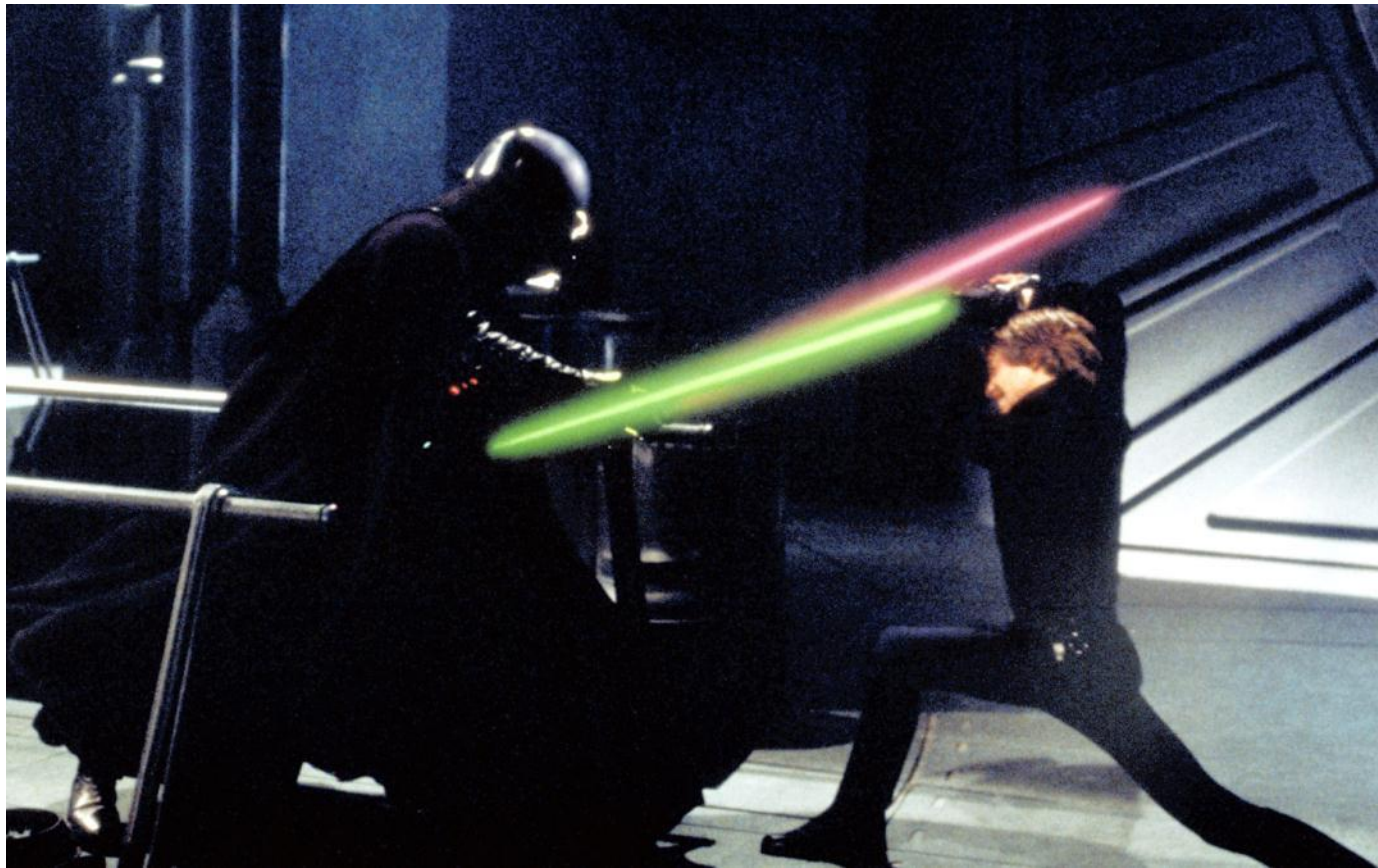
Analysis: $P(\mu\mu) = 98\%$

$m_{\mu\mu}$ [GeV]	110 - 200	200 - 400	400 - 800	800 - 1200	1200 - 3000	3000 - 4500
Z/γ^*	111000 ± 8000	11000 ± 1000	1000 ± 100	49 ± 5	7.3 ± 1.3	0.033 ± 0.029
$t\bar{t}$	5900 ± 900	1900 ± 400	140 ± 60	2.7 ± 0.7	0.16 ± 0.08	< 0.001
Diboson	1520 ± 190	520 ± 140	62 ± 26	2.8 ± 1.0	0.38 ± 0.28	0.002 ± 0.003
Total	118000 ± 8000	13300 ± 1100	1160 ± 120	55 ± 5	7.8 ± 1.3	0.035 ± 0.029
Data	118701	13349	1109	48	8	0

No deviation from expectation found.

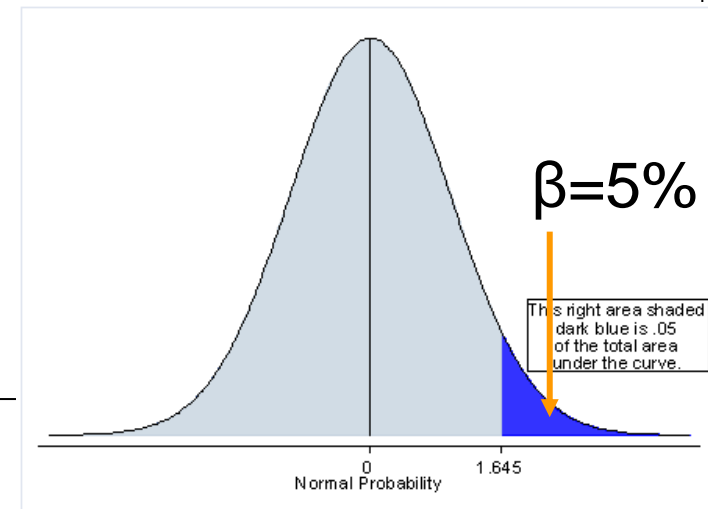
We did not find any deviation.....

- Quantify the sensitivity and reach of our analysis
- Again, many ways to do it....
 - “Religious” wars are being fought about this.....



Back of the envelope demonstration.....to get the idea

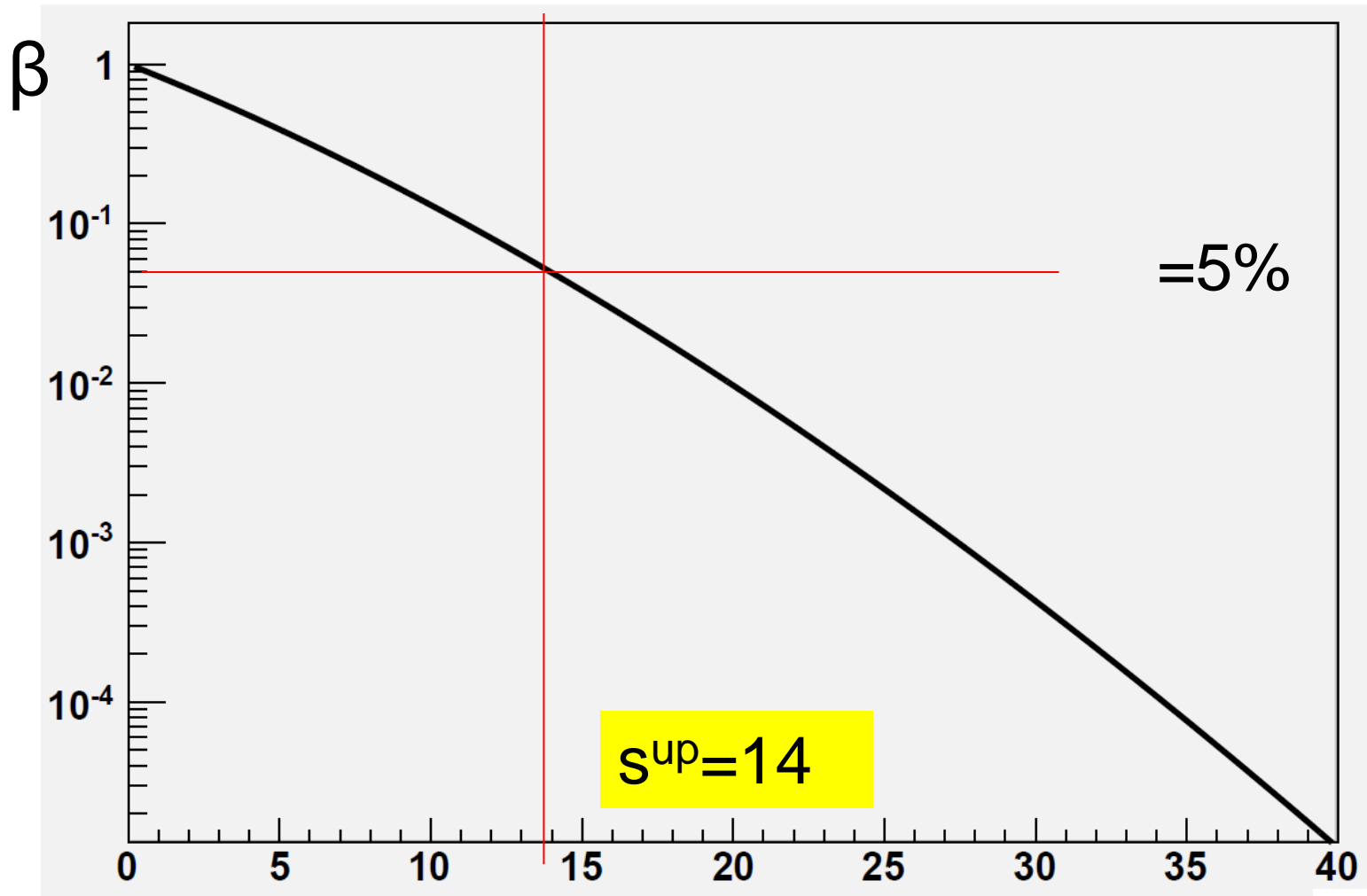
- $n_{\text{obs}} = s + b$
- We want an upper limit (bound on s) given we expect b background events and have observed n_{obs} events.
- Use Bayesian method with uniform prior density
- $\beta = e^{-s^{\text{up}}} \sum_{n=0}^{n_{\text{obs}}} (s^{\text{up}})^n / n!$ solve this numerical
- We ignore error on b
- We ignore systematic errors



■ $\beta = e^{-s^{up}} \sum_{n=0}^{n_{obs}} (s^{up})^n / n!$ solve this numerical

■ Back to our example

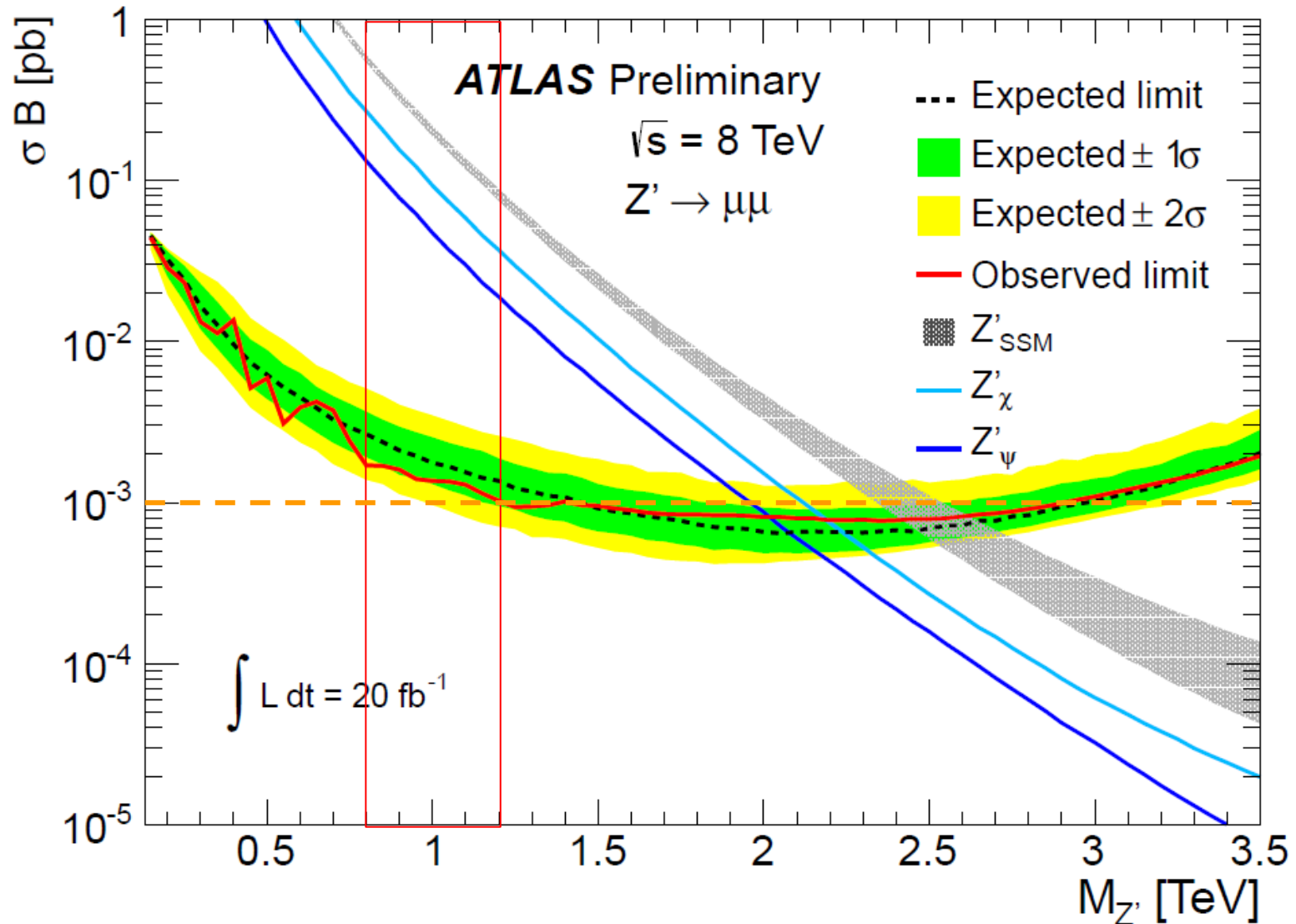
- $800 \text{ GeV} < m_{\mu\mu} < 1200 \text{ GeV}$
- We have observed $n_{obs} = 48$ events
- We expect $b=55$ background events
- Our Acceptance x Efficiency $\sim 50\%$
- We have analysed $L = 20 \text{ fb}^{-1}$ of data



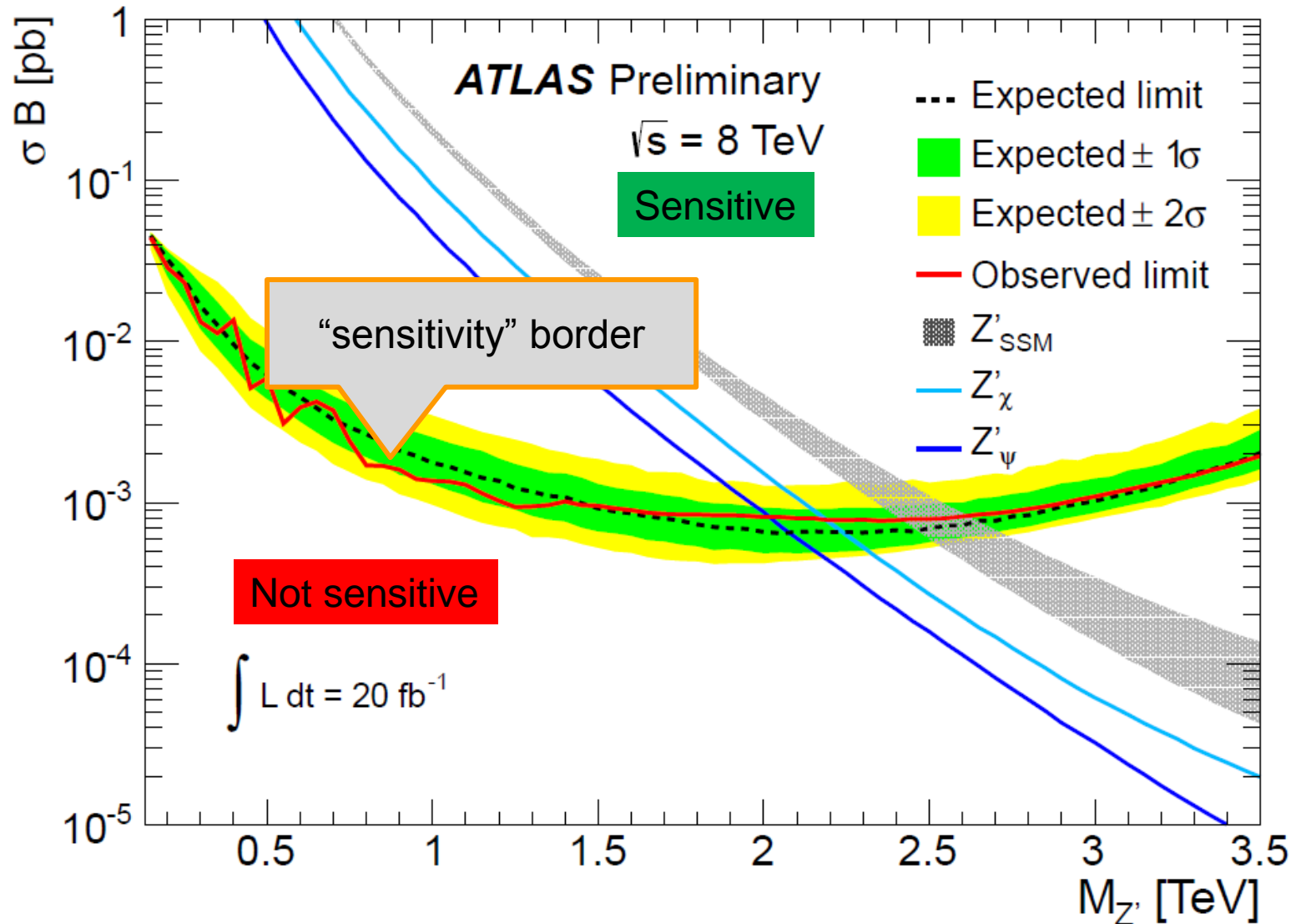
95% C.I. upper cross section limit
 $14/20\text{fb}^{-1} = 0.7\text{fb} \sim 1\text{fb} = 10^{-3} \text{ pb}$

S_{up}

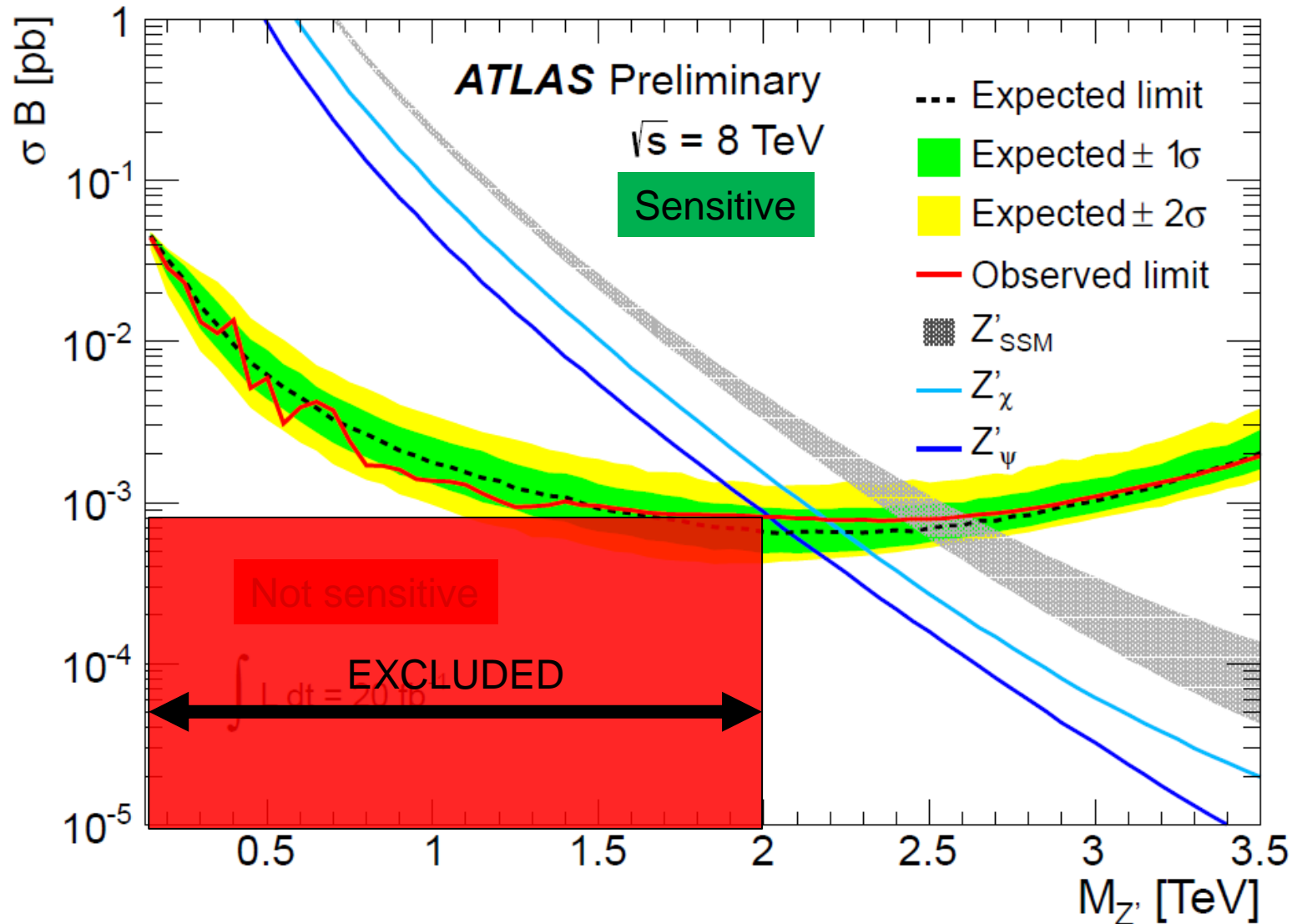
Let us compare with the published limit...



Let us compare with the published limit...

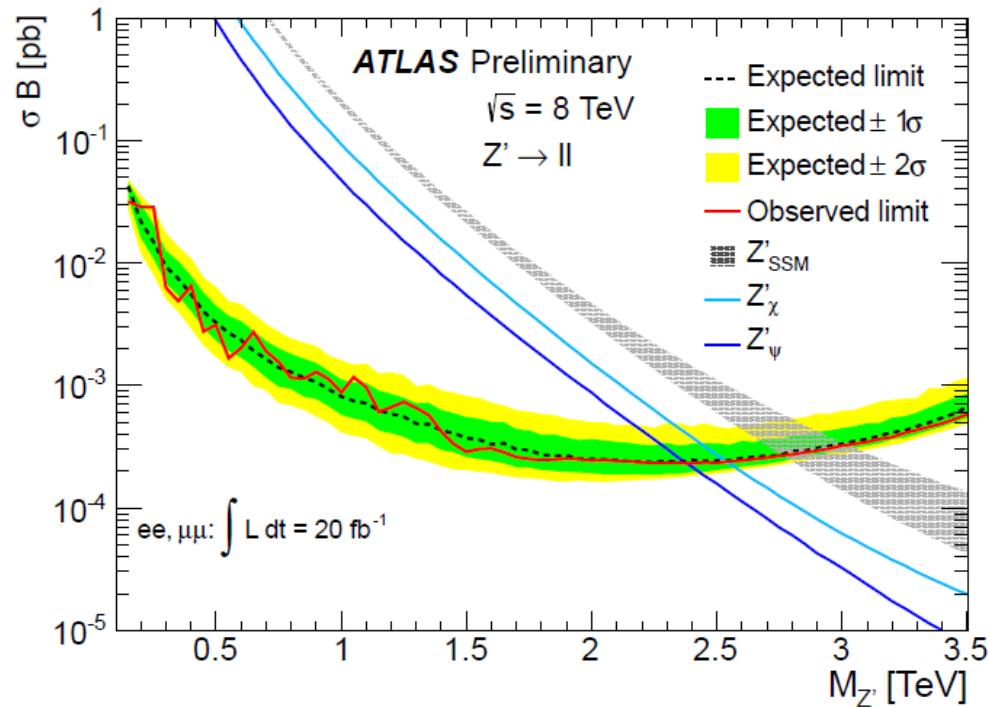


Let us compare with the published limit...



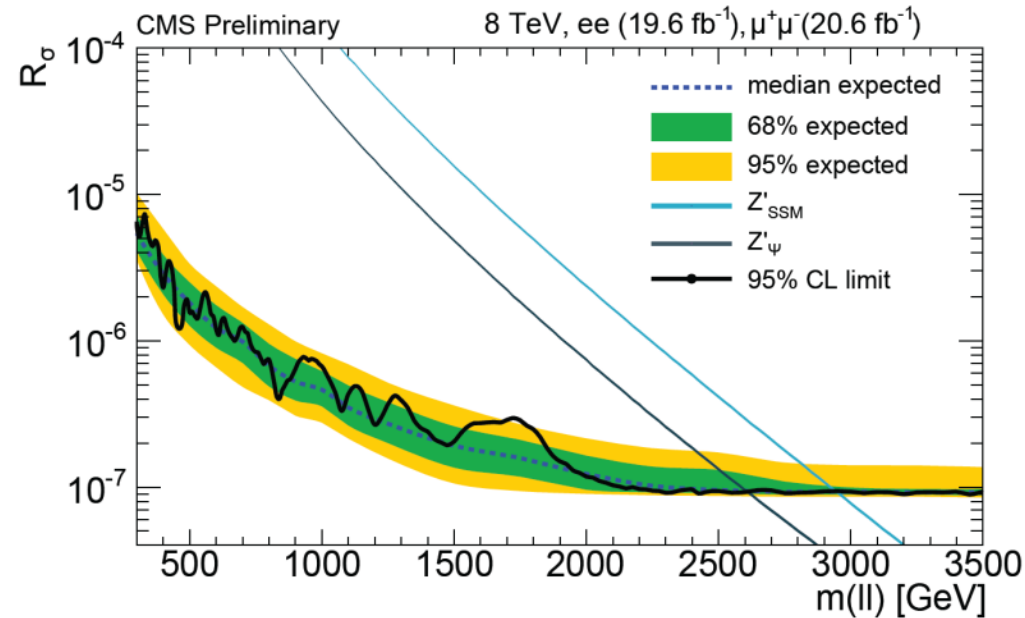
Limits for both channels combined

ATLAS



$Z'_{SSM} > 2.86 \text{ TeV @ 95\% C.L.}$

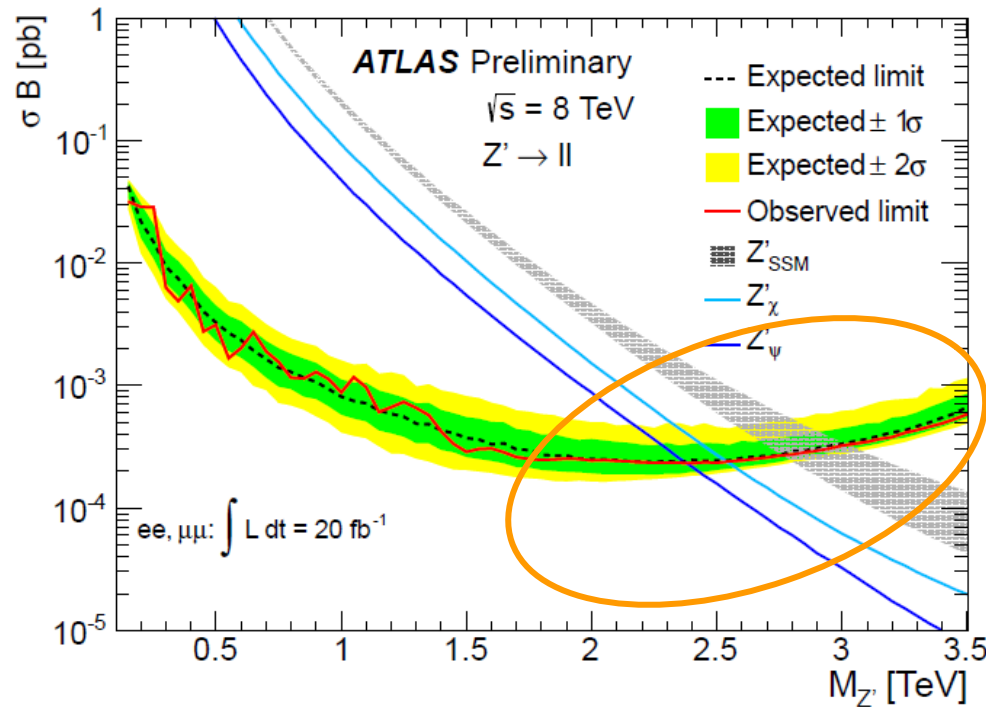
CMS



$Z'_{SSM} > 2.96 \text{ TeV @ 95\% C.L.}$

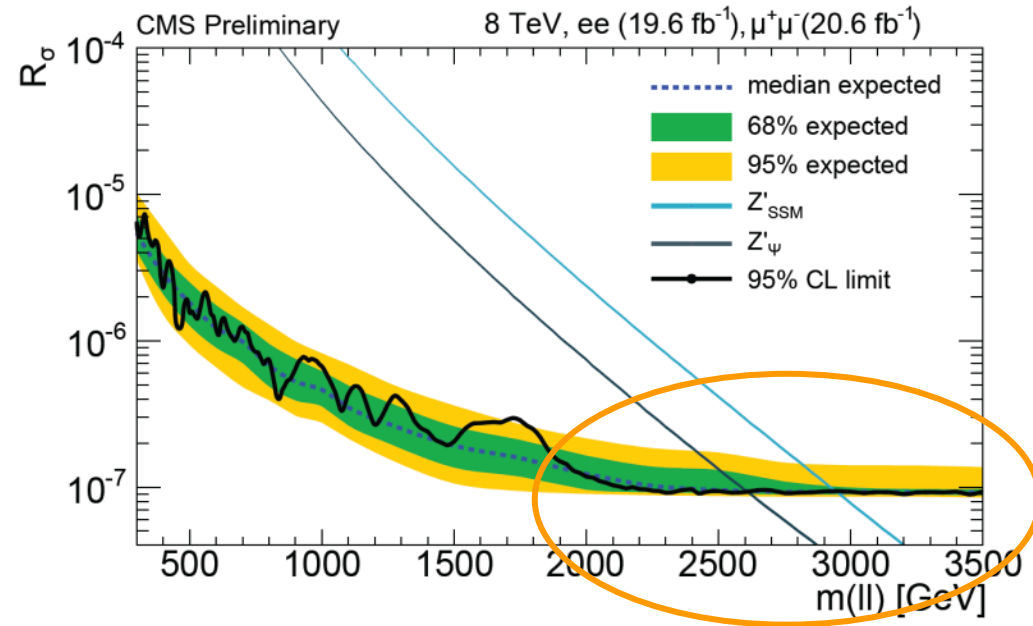
Let us discuss a bit the difference btw ATLAS/CMS

ATLAS



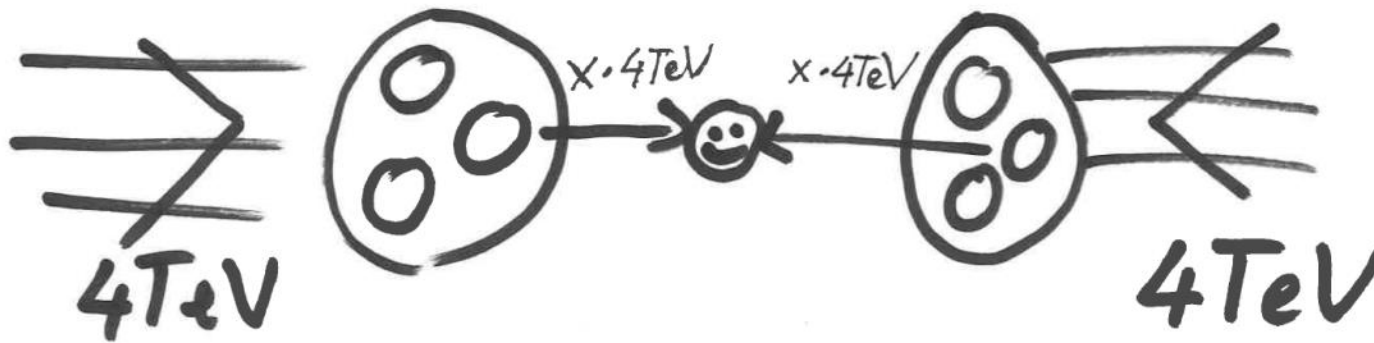
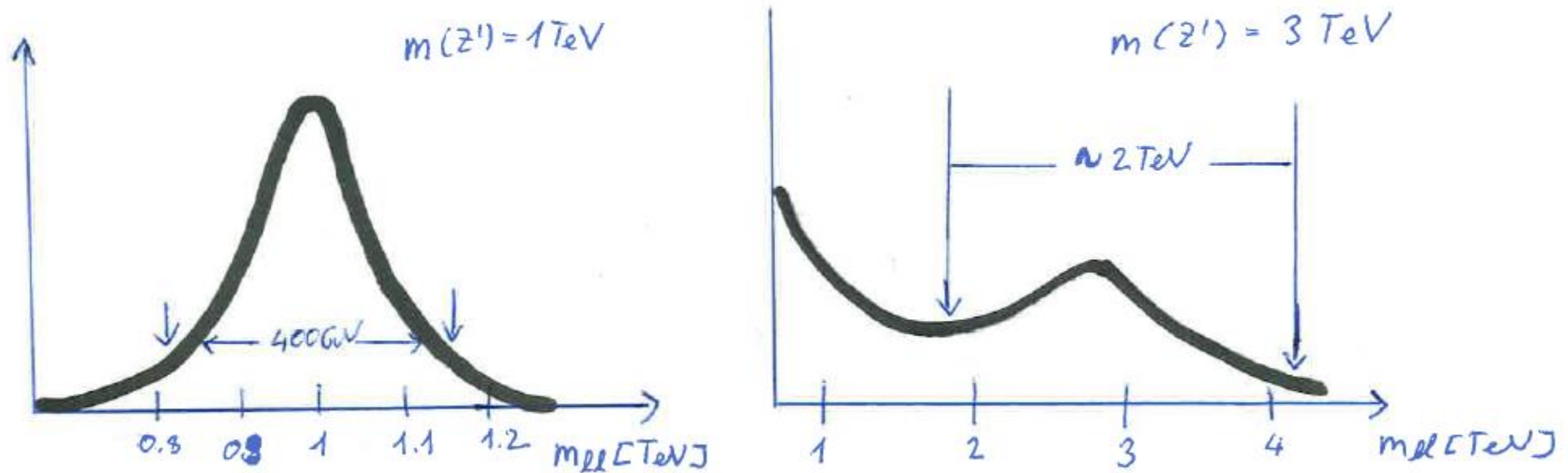
$Z'_{\text{SSM}} > 2.86 \text{ TeV} @ 95\% \text{ C.L.}$

CMS



$Z'_{\text{SSM}} > 2.96 \text{ TeV} @ 95\% \text{ C.L.}$

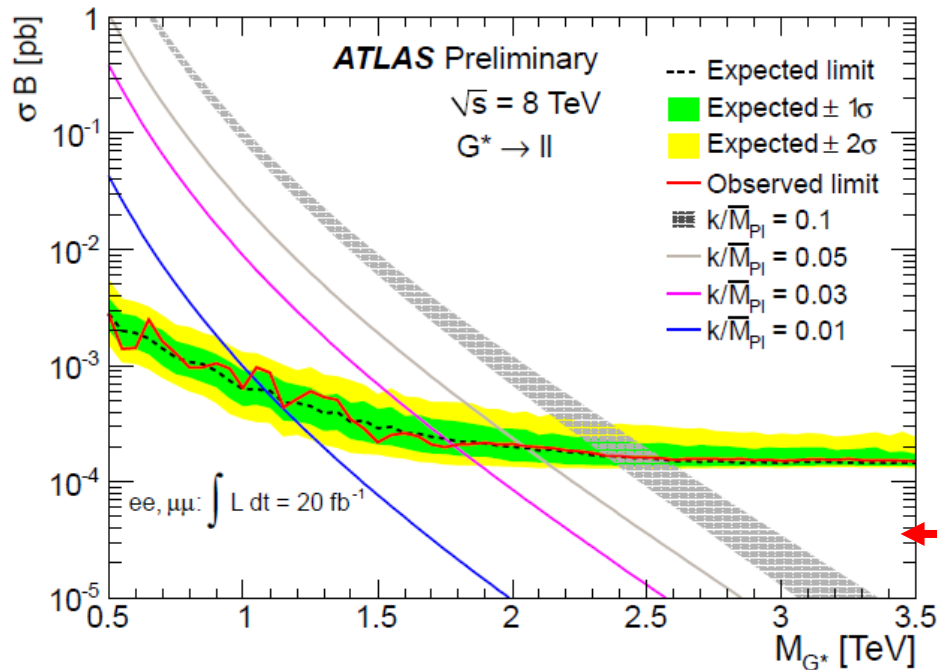
Signal Shapes and Parton Luminosities



ATLAS CMS Differences in the Limit Setting

ATLAS

- Uses signal templates for limits
- Loss of sensitivity at high masses
 - Parton luminosities
- Upper cross section limits model specific

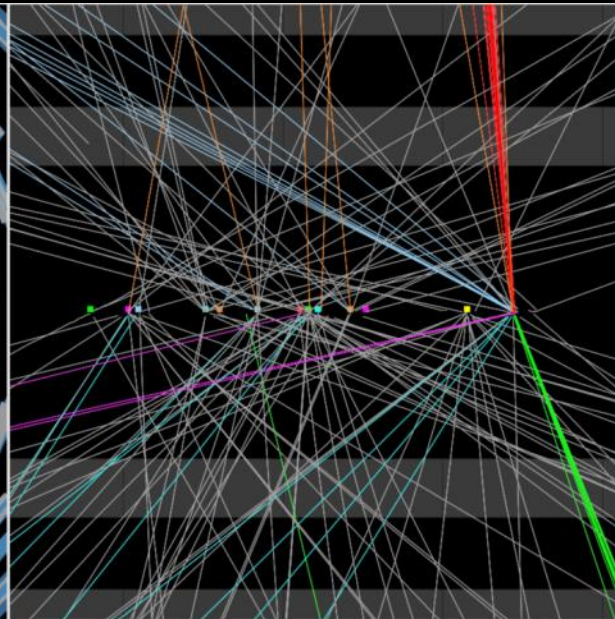
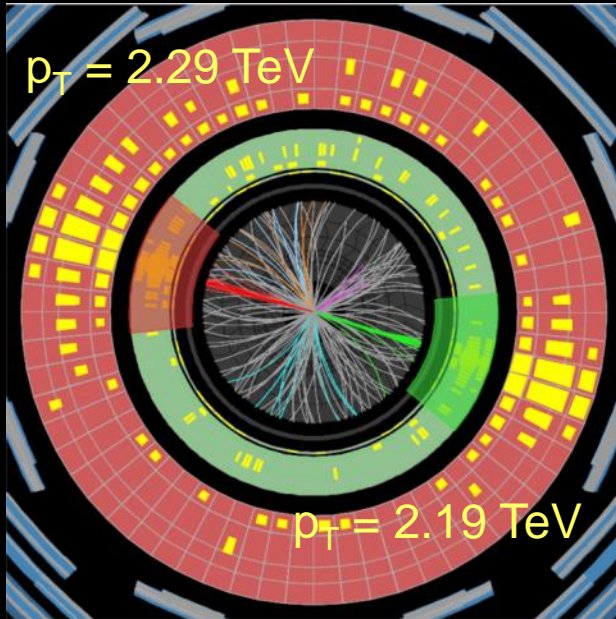


CMS

- Uses narrow resonance
 - For cross section upper limit
 - Cross section upper limits less model dependent
 - Give outside world description of what was done
- Take signal shapes within $\pm 40\%$ of the mass peak into account to compute theory curves
- Not sensitive to parton luminosities
- generic resonance search

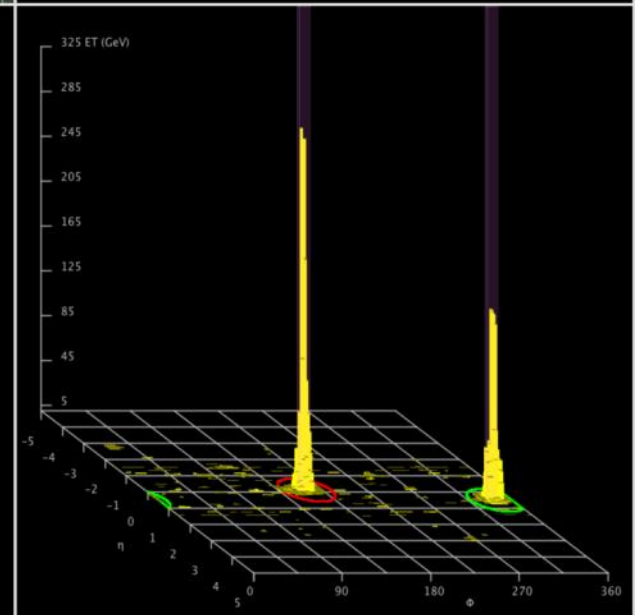
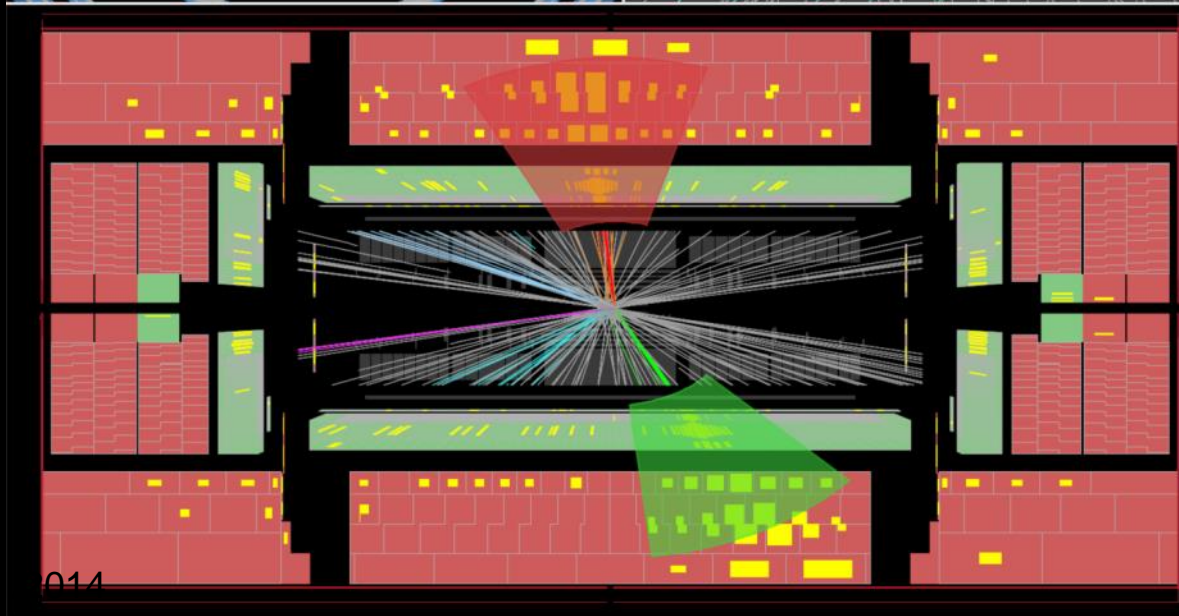
KK Graviton narrow resonance
Obs limit does not go up

Dijet Event Display with $m_{\text{inv}} = 4.69 \text{ TeV}$

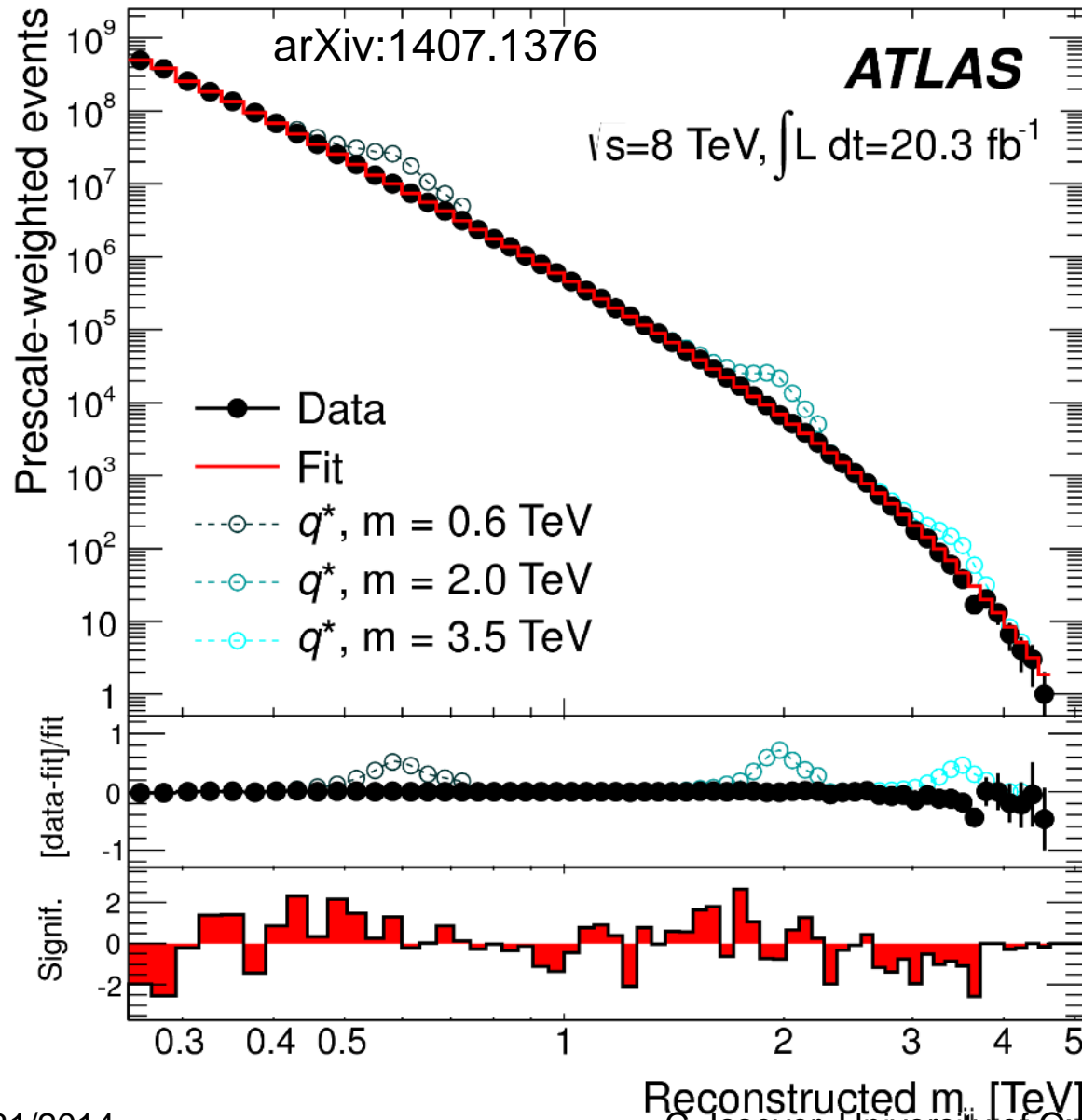


Run Number: 209580, Event Number: 179229707

Date: 2012-08-31 20:24:29 CEST

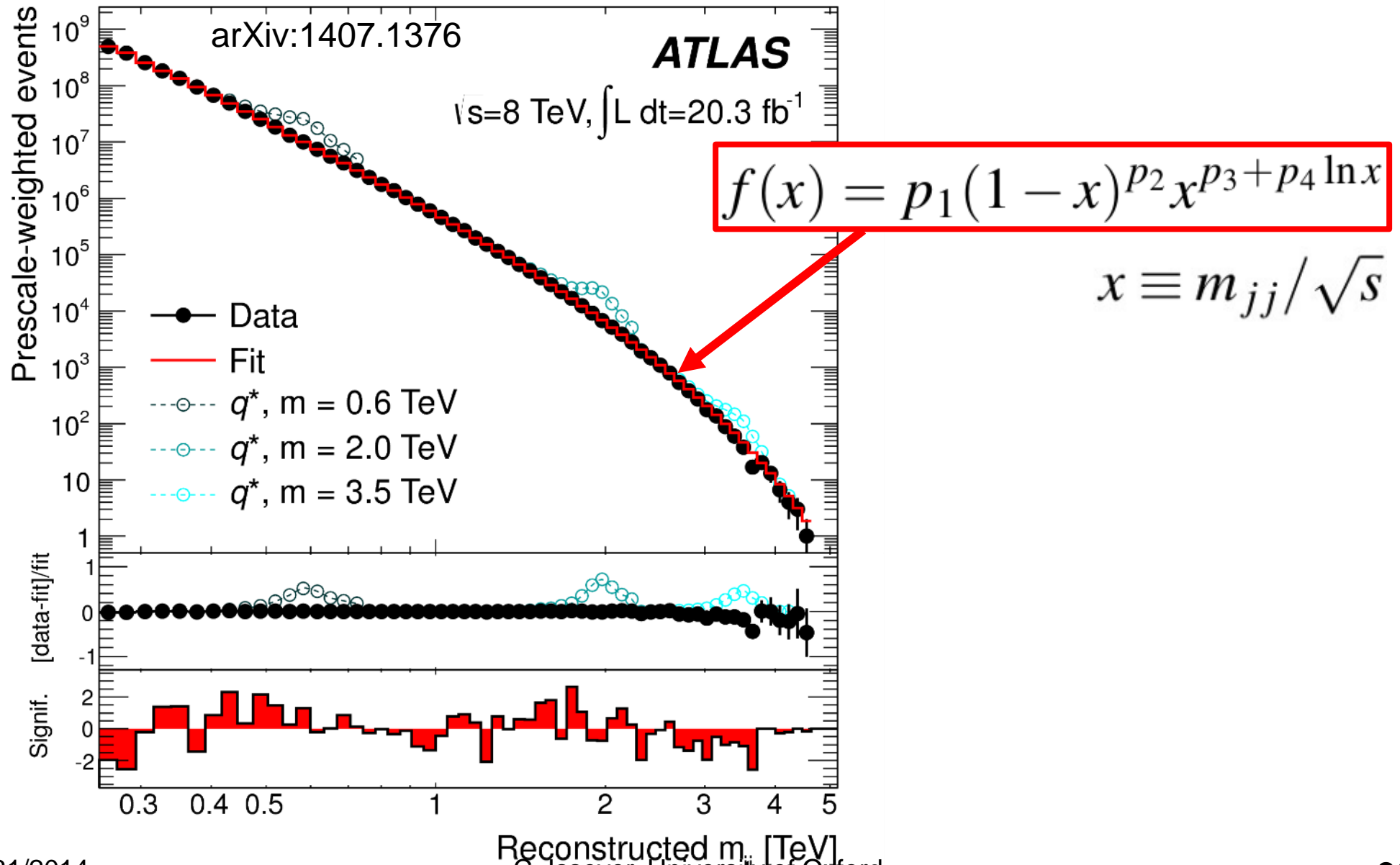


Heavy Resonance Search: 8 TeV Dijets

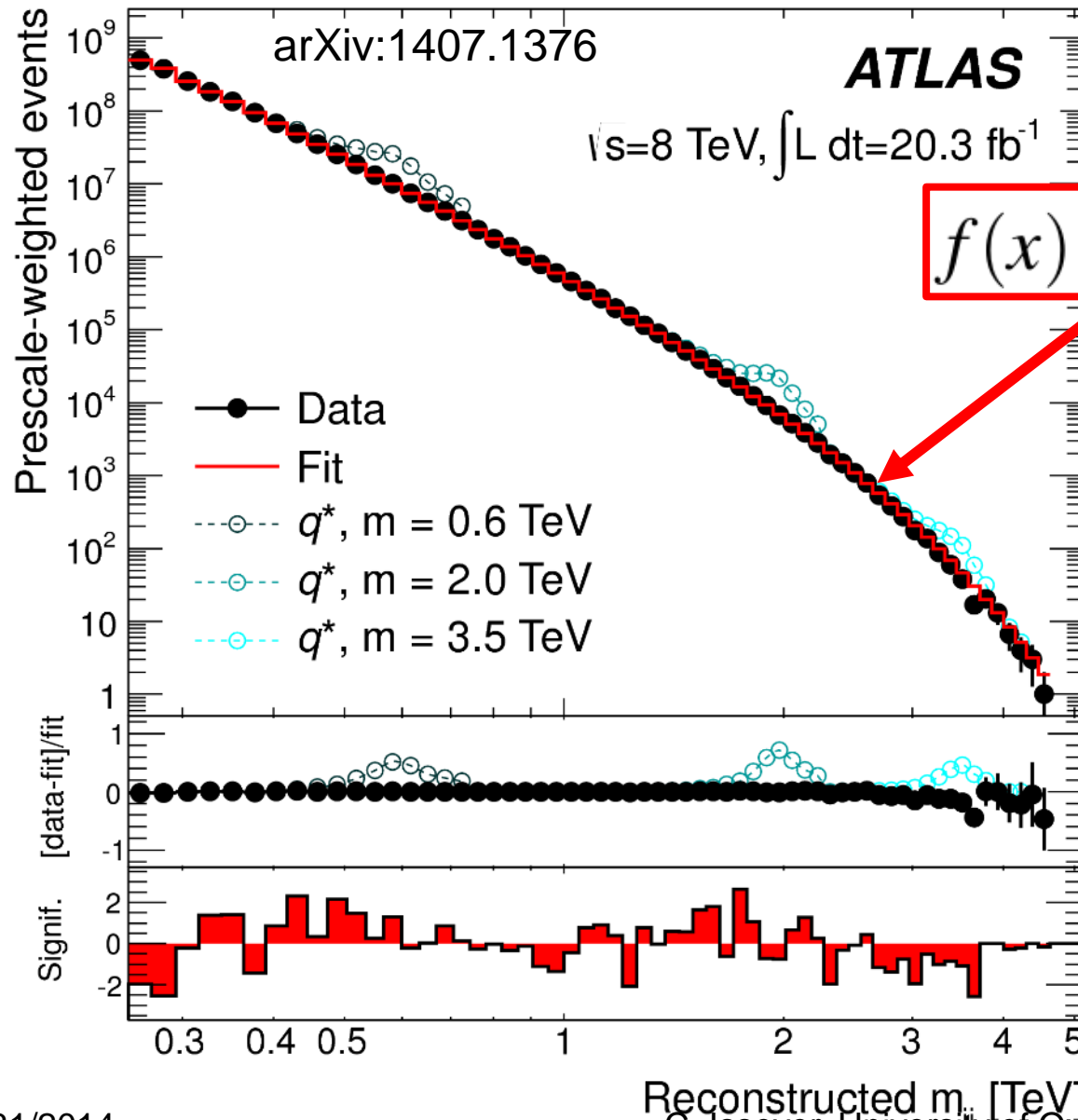


Probing
quark structure
~ 5 TeV

Heavy Resonance Search: 8 TeV Dijets



Heavy Resonance Search: 8 TeV Dijets



$$f(x) = p_1 (1 - x)^{p_2} x^{p_3 + p_4 \ln x}$$

$$x \equiv m_{jj} / \sqrt{s}$$

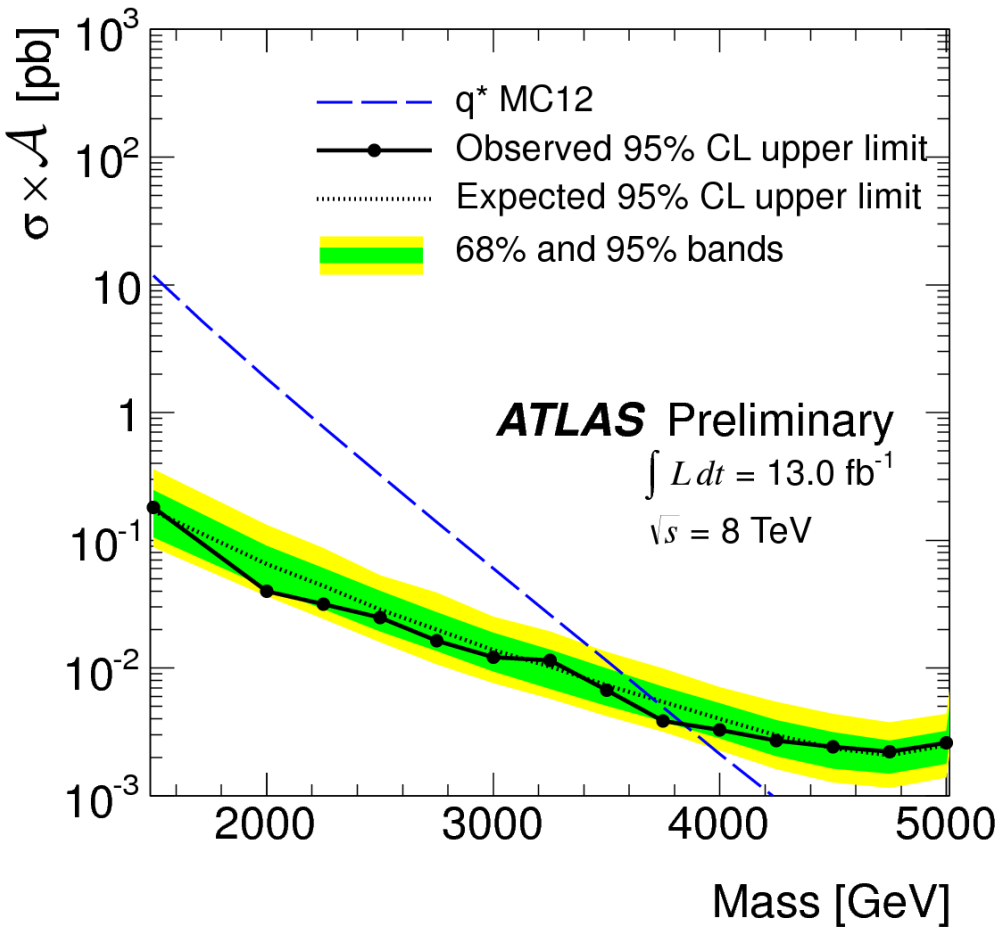
Global $\chi^2/\text{NDF}=79/56$

P-value = 0.27

Fit describes data

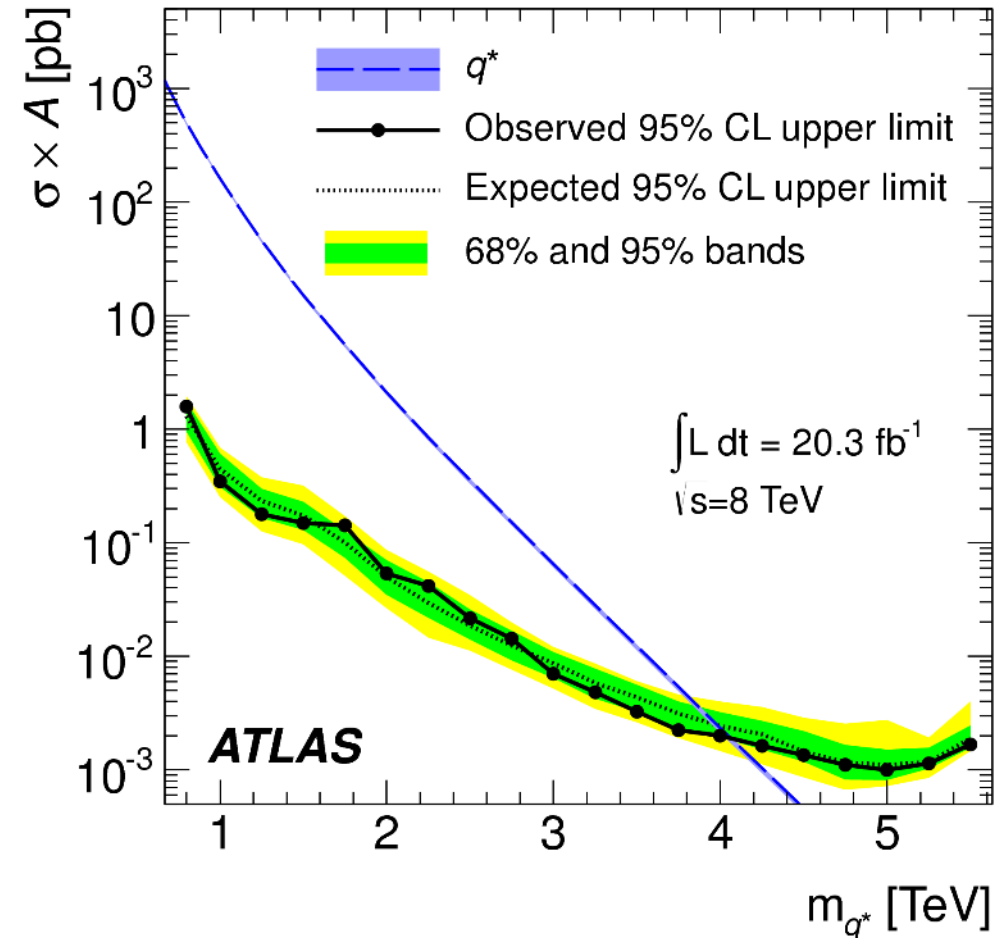
Heavy Resonance Search: 8 TeV Dijets

Previous Results: 13 fb⁻¹



$m > 3.84 \text{ TeV}$ at 95% CL

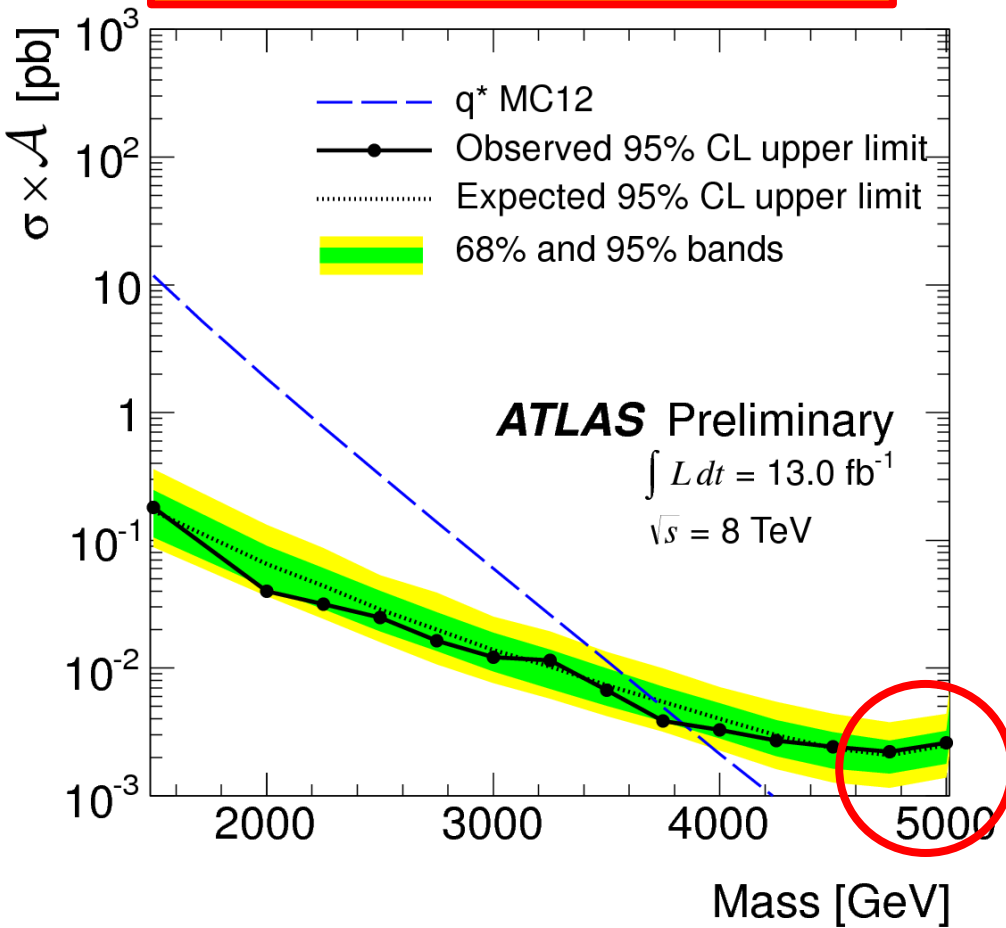
New Results: 20 fb⁻¹



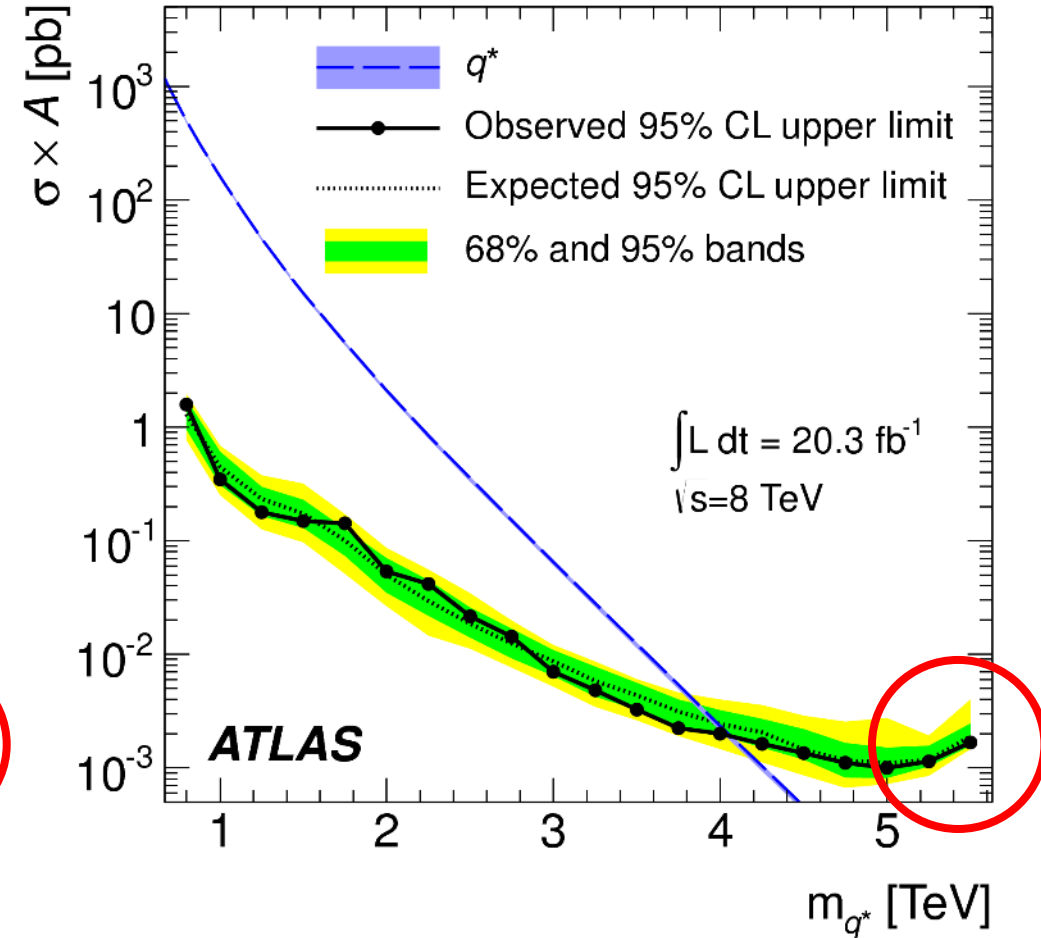
$m > 4.09 \text{ TeV}$ at 95% CL

Heavy Resonance Search: 8 TeV Dijets

Previous Results: 13 fb⁻¹

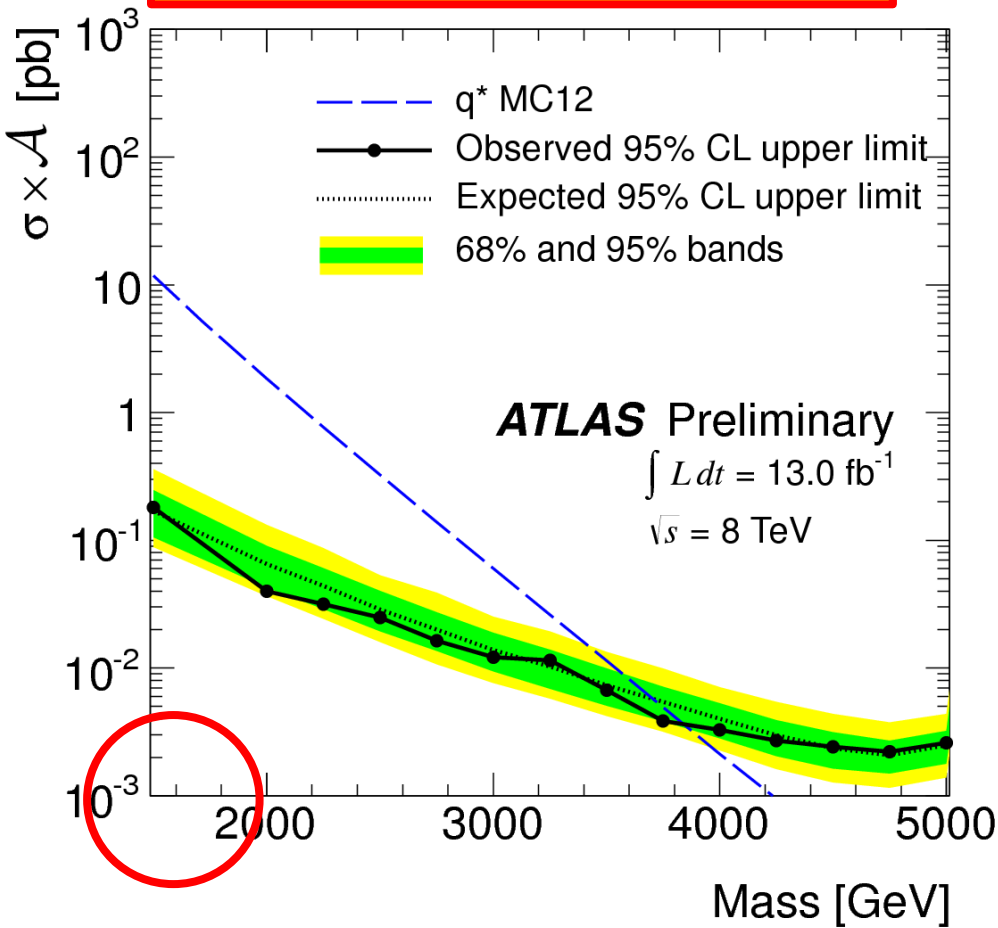


New Results: 20 fb⁻¹

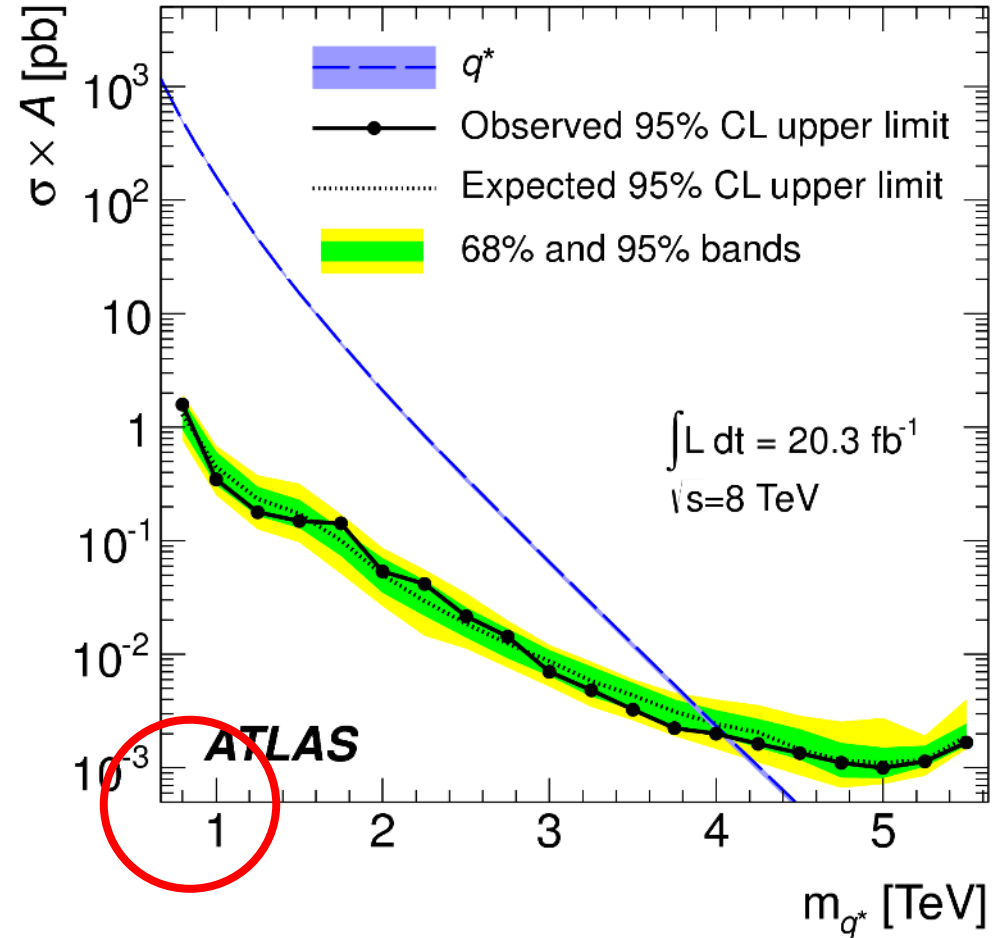


Heavy Resonance Search: 8 TeV Dijets

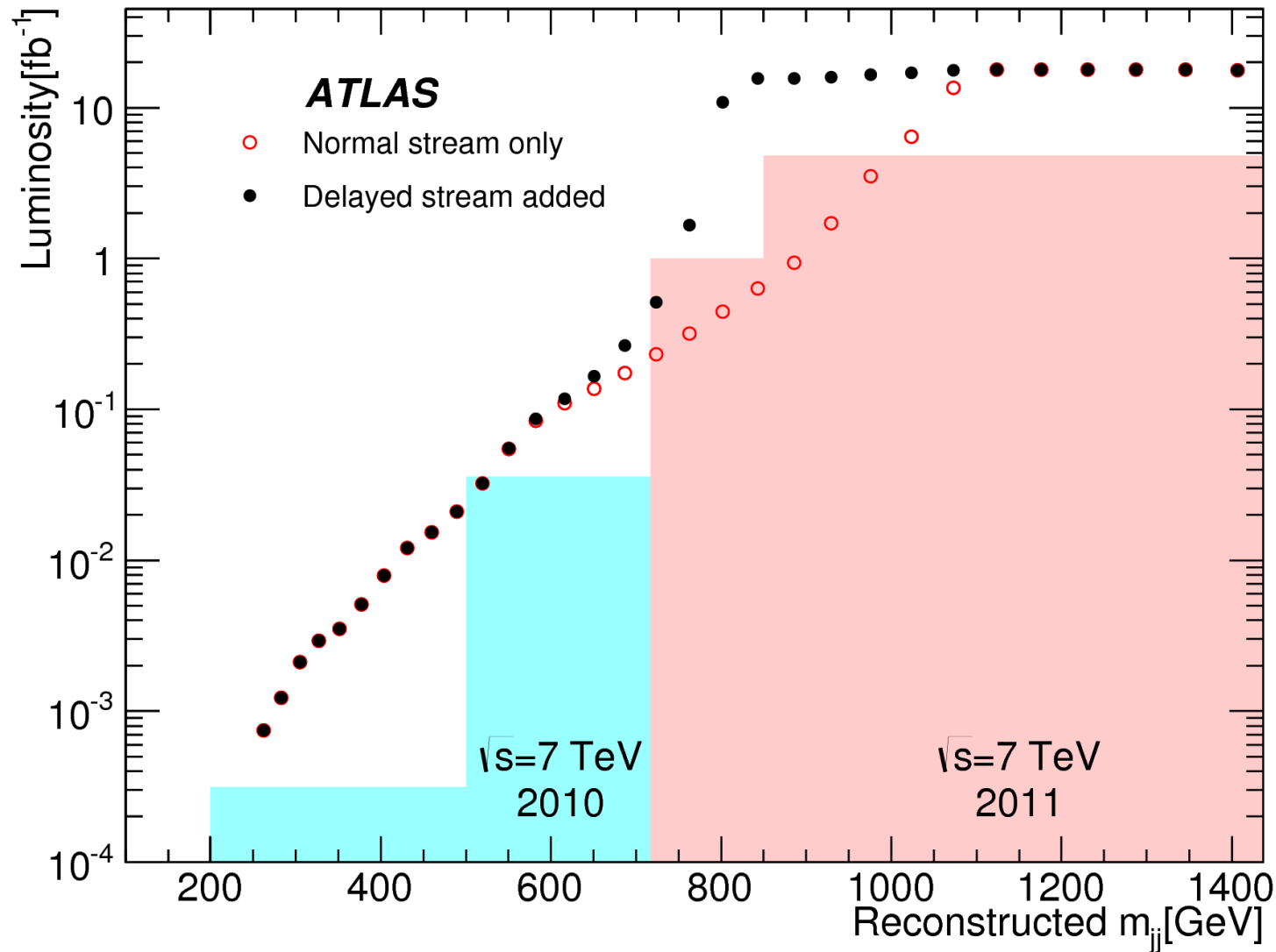
Previous Results: 13 fb⁻¹



New Results: 20 fb⁻¹

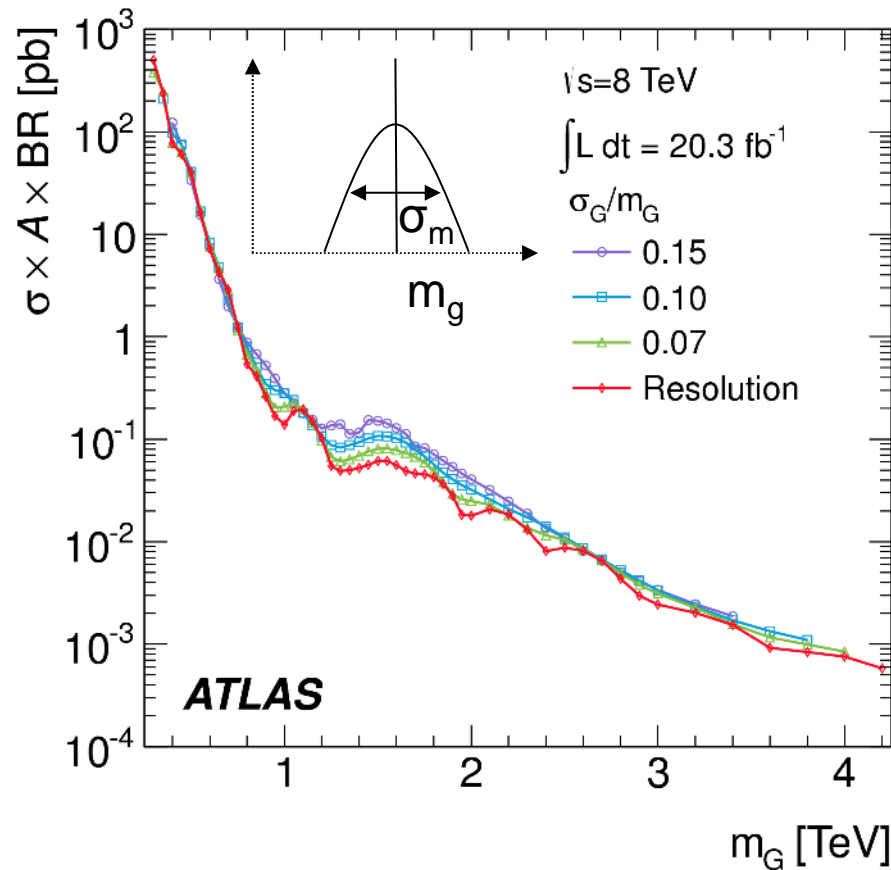


Extending Reach to low invariant masses

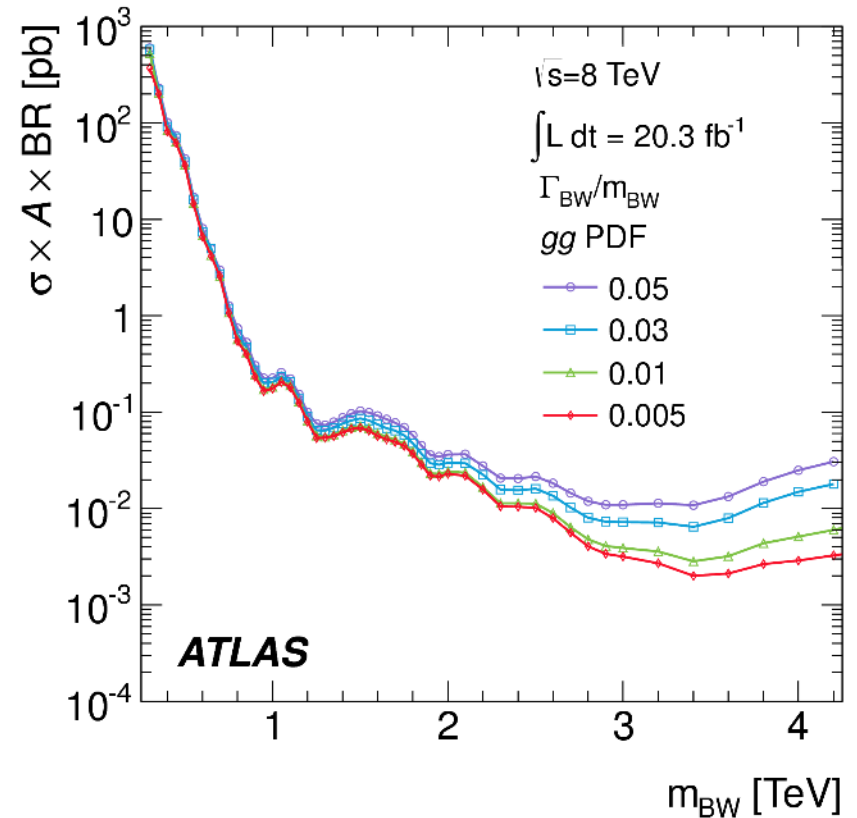


Heavy Resonance Search: 8 TeV Dijets

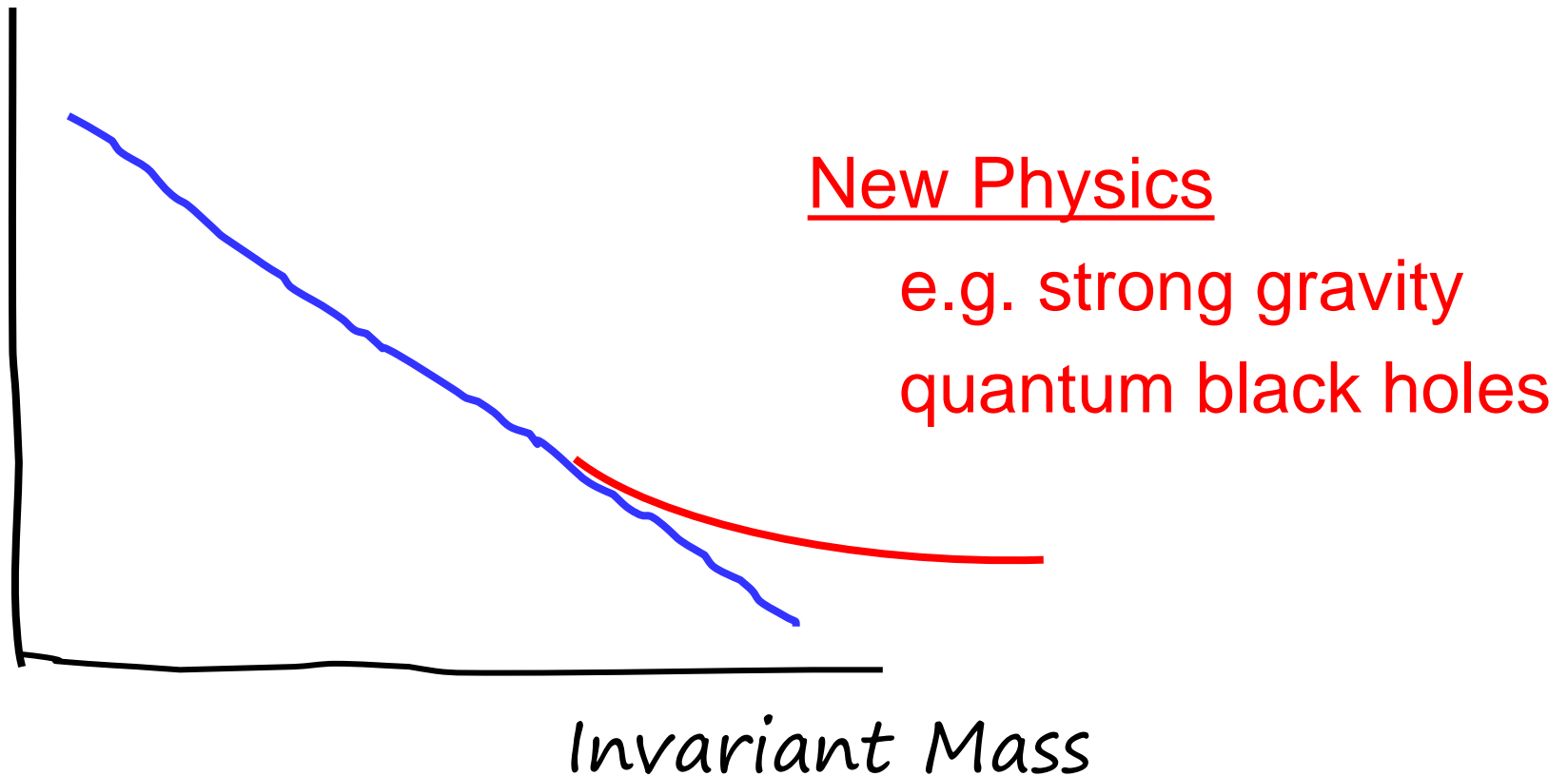
Gaussian resonance limits:
mean mass, m_G , and $3 \sigma_G$



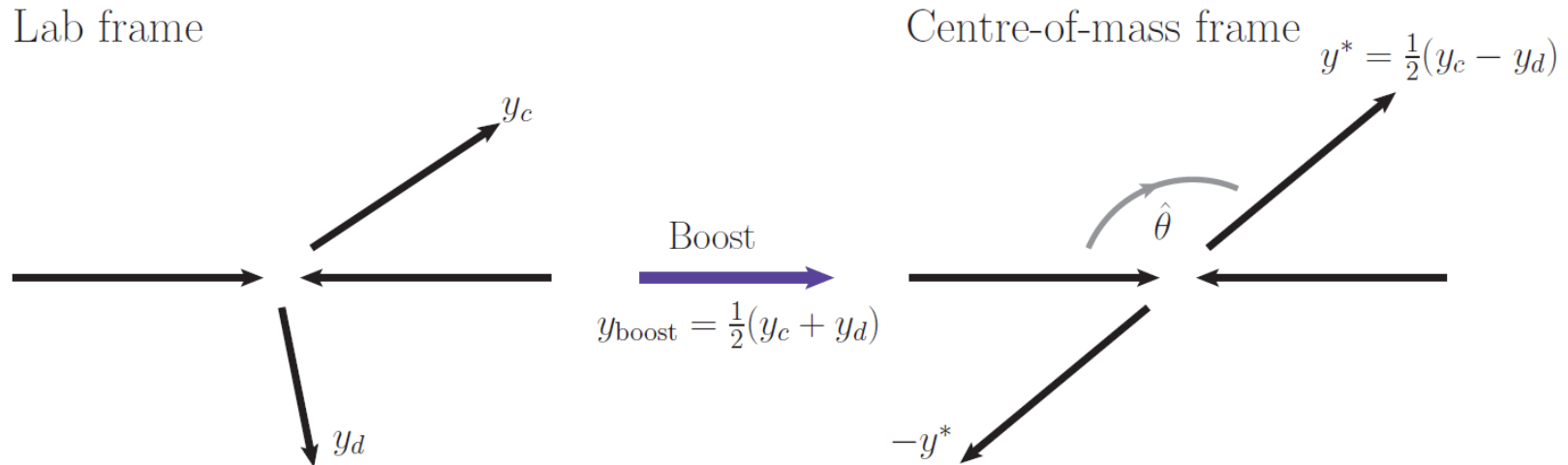
Breit Wigner x PDF



Search for Excess in Tail



Search for Threshold Effects: Dijet Angular



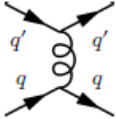
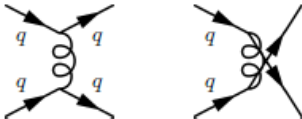

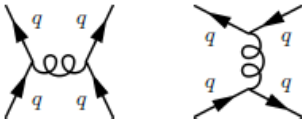
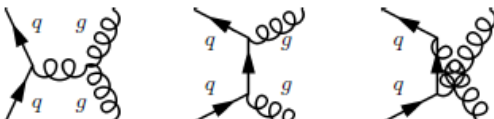
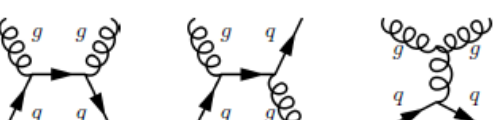
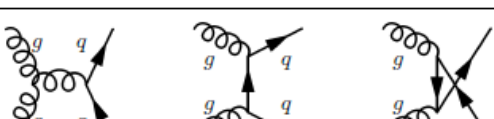
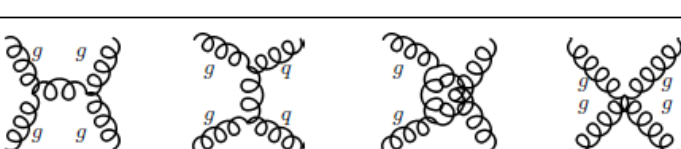
$$d\hat{\sigma}/d(\cos \hat{\theta}) \propto \sin^{-4}(\hat{\theta}/2) \quad \text{t-channel Spin-1 exchange}$$

$$\chi = \frac{1 + |\cos \hat{\theta}|}{1 - |\cos \hat{\theta}|} \sim \frac{1}{1 - |\cos \hat{\theta}|} \propto \frac{\hat{s}}{\hat{t}}$$

$$\frac{d\hat{\sigma}}{d\chi} \propto \frac{\alpha_s^2}{\hat{s}} \quad (\hat{s} \text{ fixed}) \quad \hat{s} = m_{jj}$$

Constant in χ for fixed m_{jj}

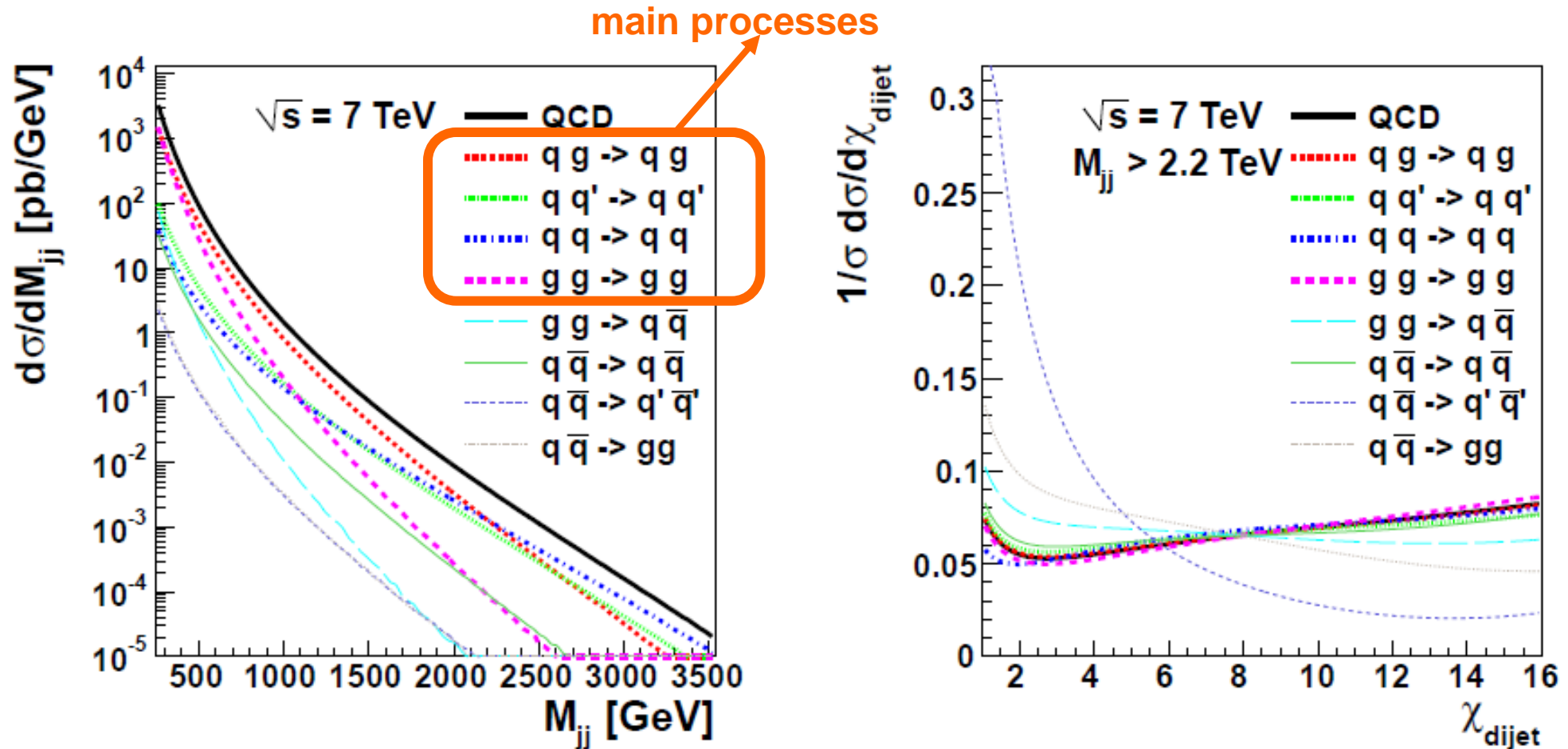
Search for Threshold Effects: Dijet Angular

$qq' \rightarrow qq' = q\bar{q}' \rightarrow q\bar{q}'$ $\frac{64}{9}\alpha_s^2 \left(\frac{s^2+u^2}{t^2} \right)$	
$qq \rightarrow qq$ $\frac{64}{9}\alpha_s^2 \left(\frac{s^2+u^2}{t^2} + \frac{s^2+t^2}{u^2} - \frac{2}{3} \frac{s}{ut} \right)$	
$q\bar{q} \rightarrow q'\bar{q}'$ $\frac{64}{9}\alpha_s^2 \left(\frac{t^2+u^2}{s^2} \right)$	
$q\bar{q} \rightarrow q\bar{q}$ $\frac{64}{9}\alpha_s^2 \left(\frac{s^2+u^2}{t^2} + \frac{t^2+u^2}{s^2} - \frac{2}{3} \frac{u^2}{st} \right)$	
$q\bar{q} \rightarrow gg$ $\frac{128}{3}\alpha_s^2 \left(\frac{4}{9} \frac{t^2+u^2}{t^2} - \frac{u^2+t^2}{s^2} \right)$	
$gg \rightarrow qq$ $16\alpha_s^2 \left(\frac{s^2+u^2}{t^2} - \frac{4}{9} \frac{s^2+u^2}{su} \right)$	
$gg \rightarrow q\bar{q}$ $\frac{8}{3}\alpha_s^2 \left(\frac{1}{3} \frac{t^2+u^2}{t^2} - \frac{3}{4} \frac{t^2+u^2}{s^2} \right)$	
$gg \rightarrow gg$ $72\alpha_s^2 \left(3 + \frac{t^2+u^2}{s^2} + \frac{s^2+u^2}{t^2} + \frac{s^2+t^2}{u^2} \right)$	

QCD is a bit more complicated.....

Andreas Dominik Hinzmann

Search for Threshold Effects: Dijet Angular

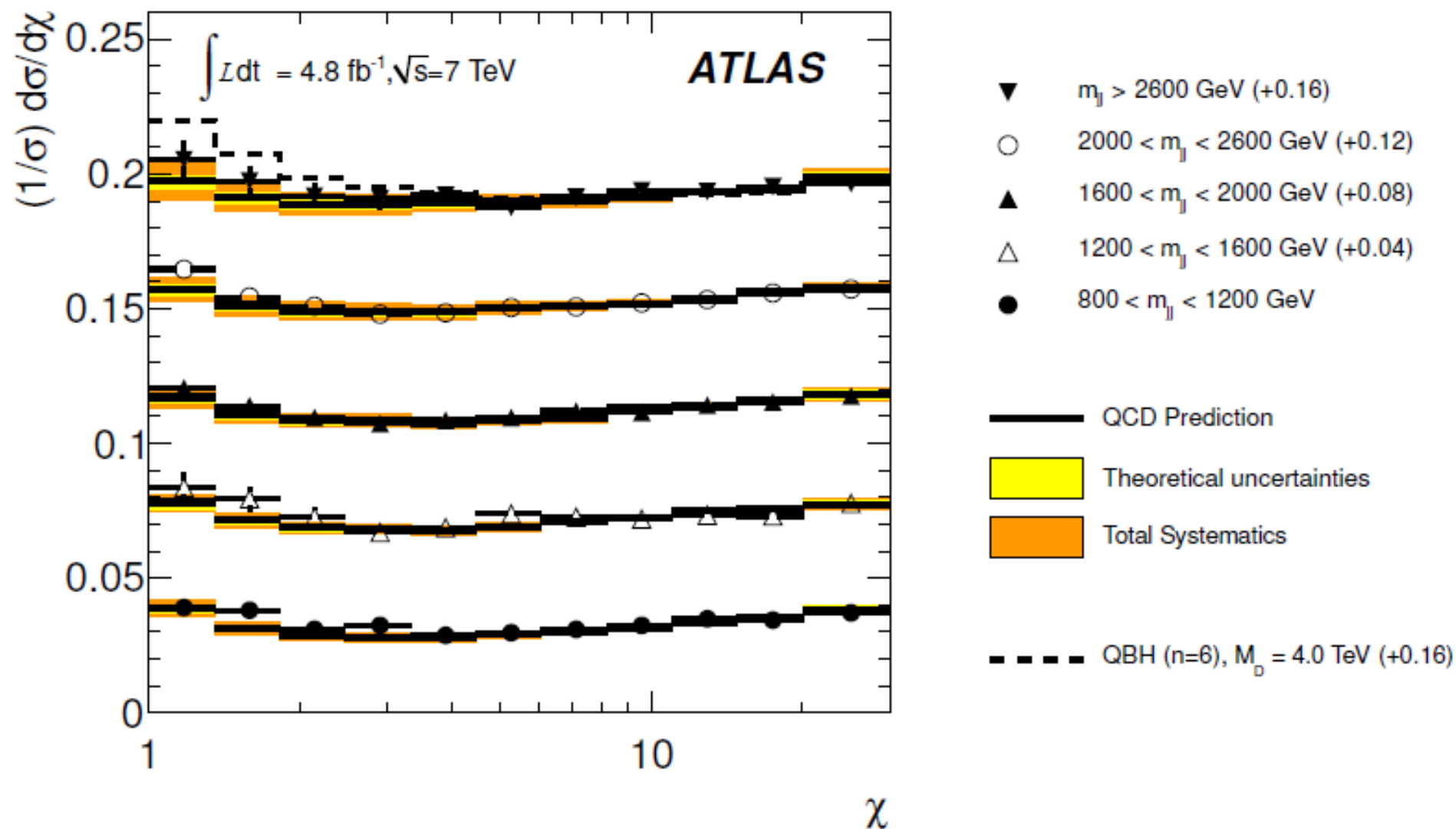


low M_{jj} $g g$ and $q g$ dominate
high M_{jj} $q q$ dominate

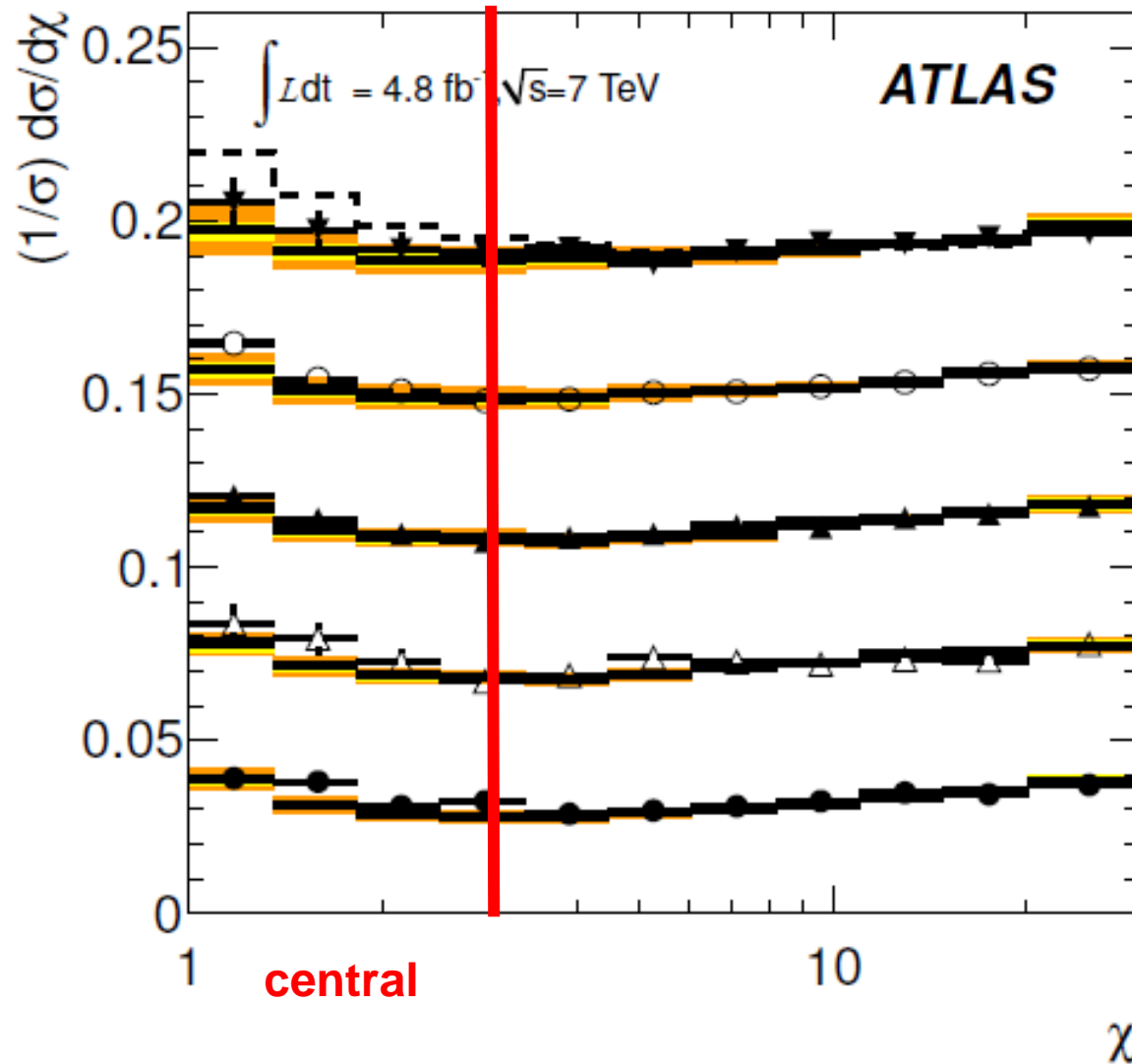
QCD \sim flat in χ

Search for Threshold Effects: Dijet Angular

[arXiv:1210.1718](https://arxiv.org/abs/1210.1718)



Finer binning in m_{jj} using $F_X(m_{jj})$

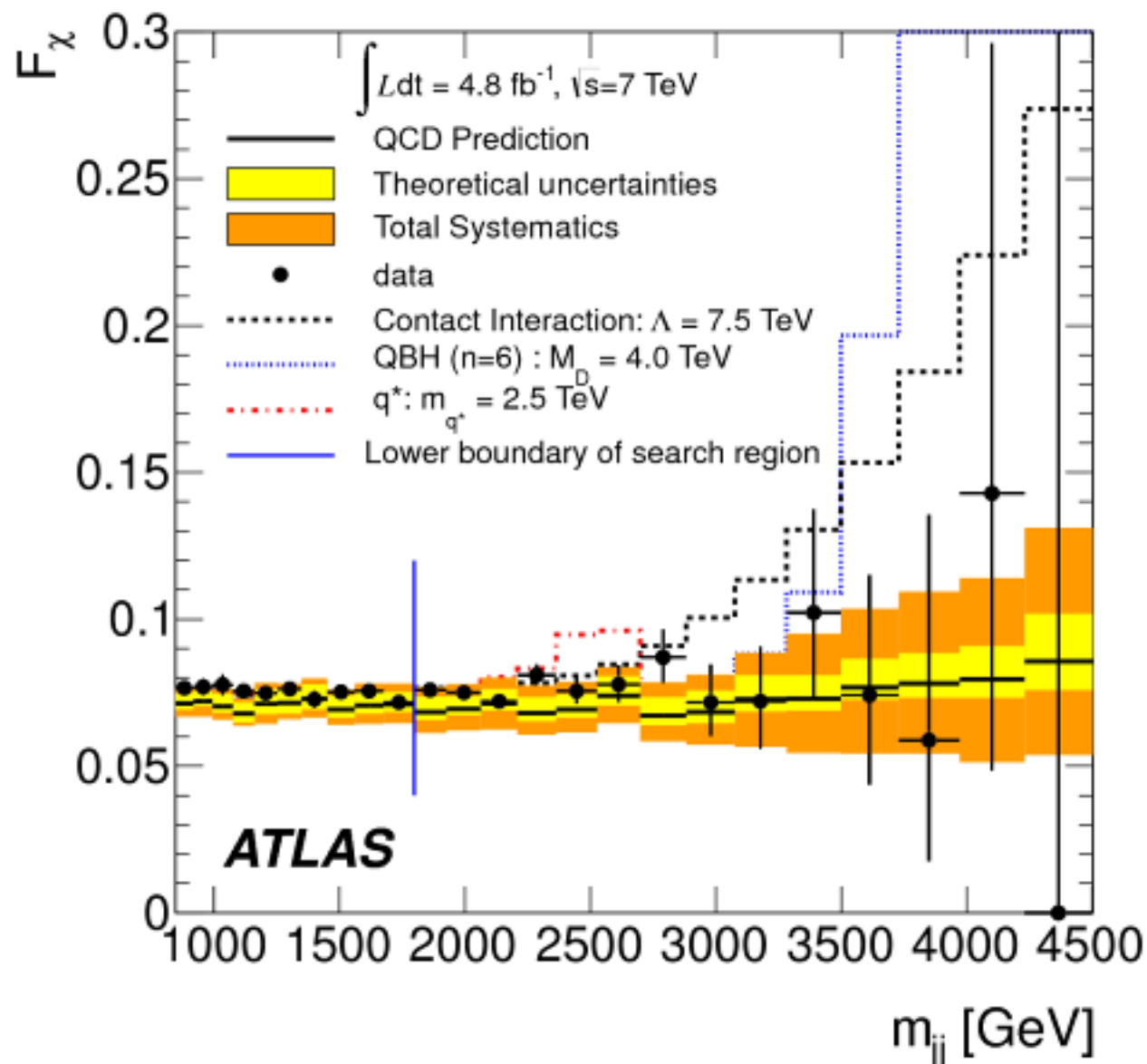


Finer binning in m_{jj} using $F_\chi(m_{jj})$

$$F_\chi(m_{jj}) \equiv \frac{dN_{\text{central}}/dm_{jj}}{dN_{\text{total}}/dm_{jj}}$$

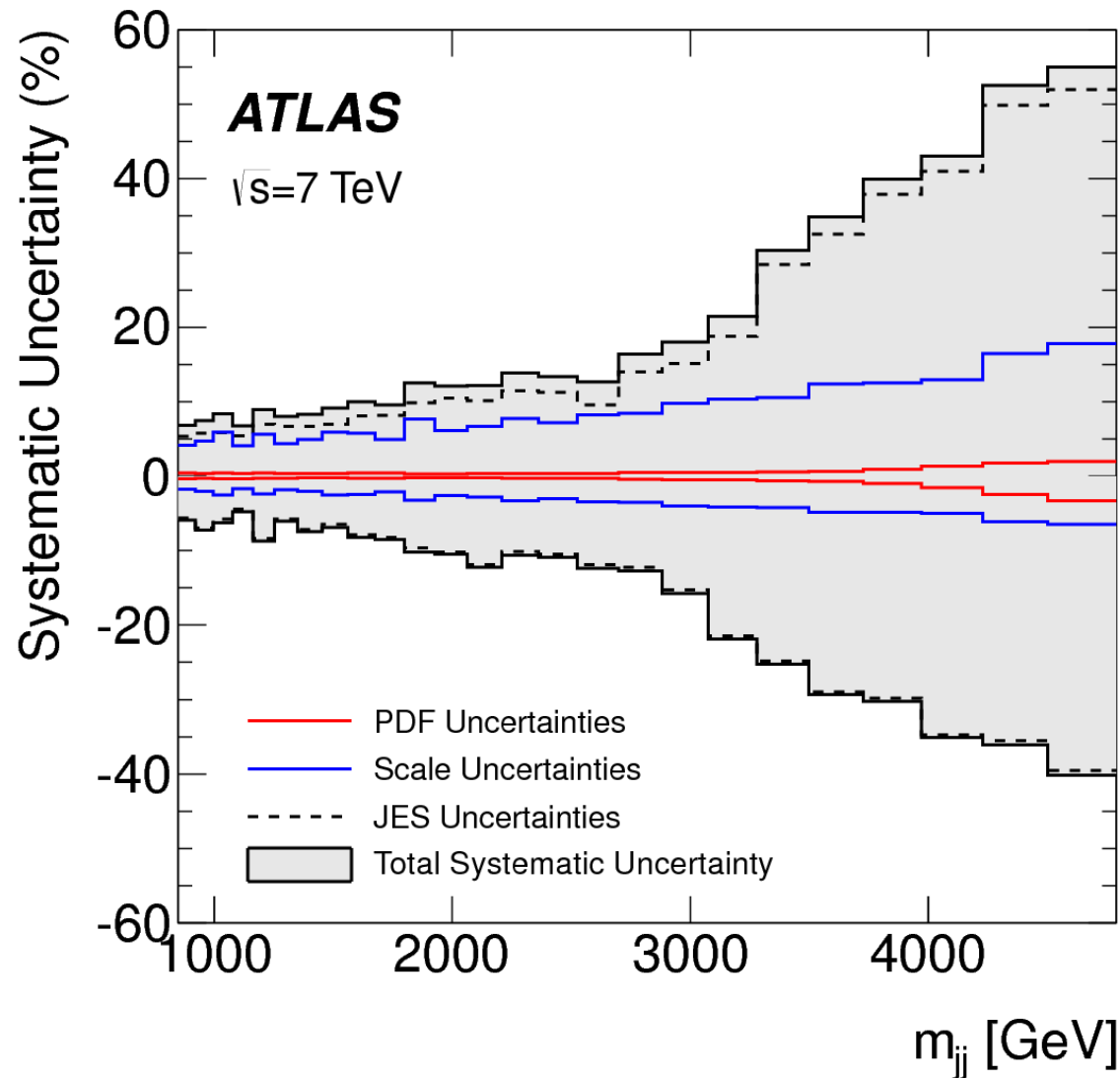
Search for Threshold Effects: Dijet Angular

[arXiv:1210.1710](https://arxiv.org/abs/1210.1710)



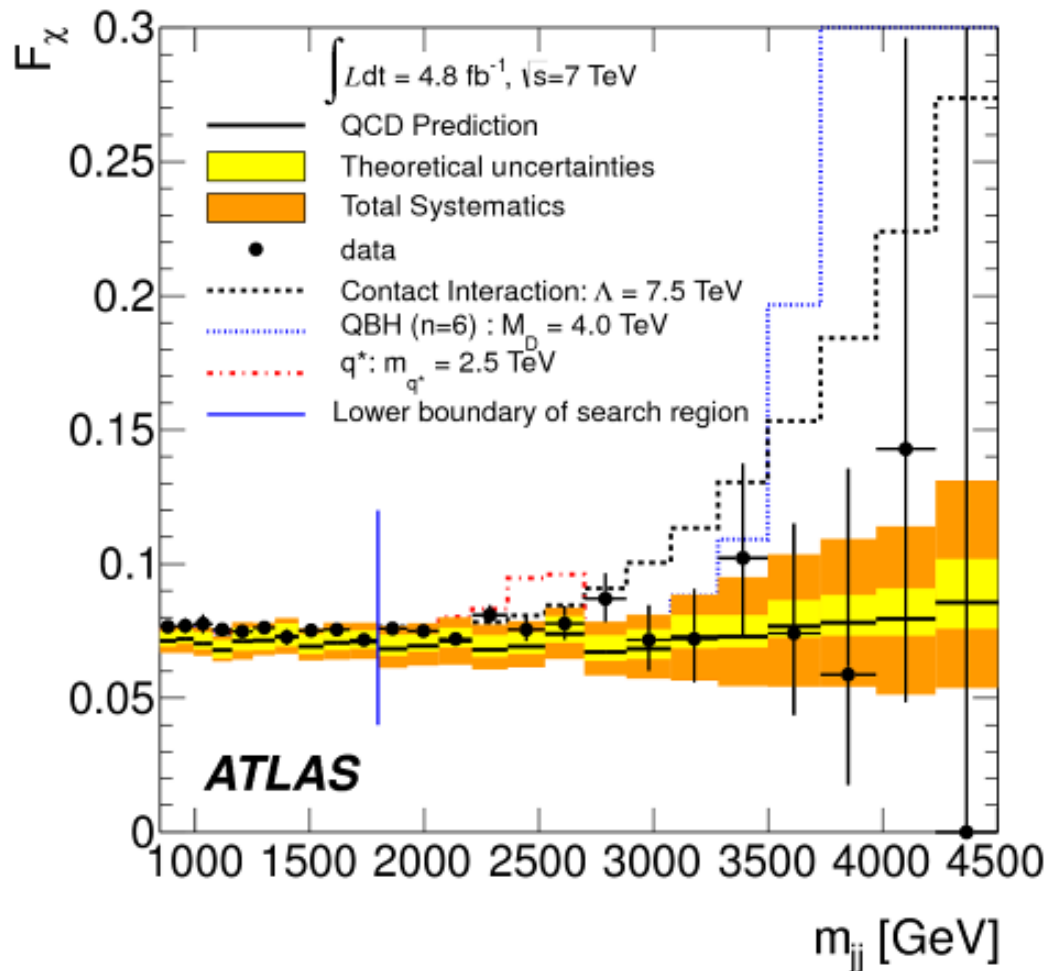
Search for Threshold Effects: Dijet Angular

[arXiv:1210.1718](https://arxiv.org/abs/1210.1718)



Search for Threshold Effects: Dijet Angular

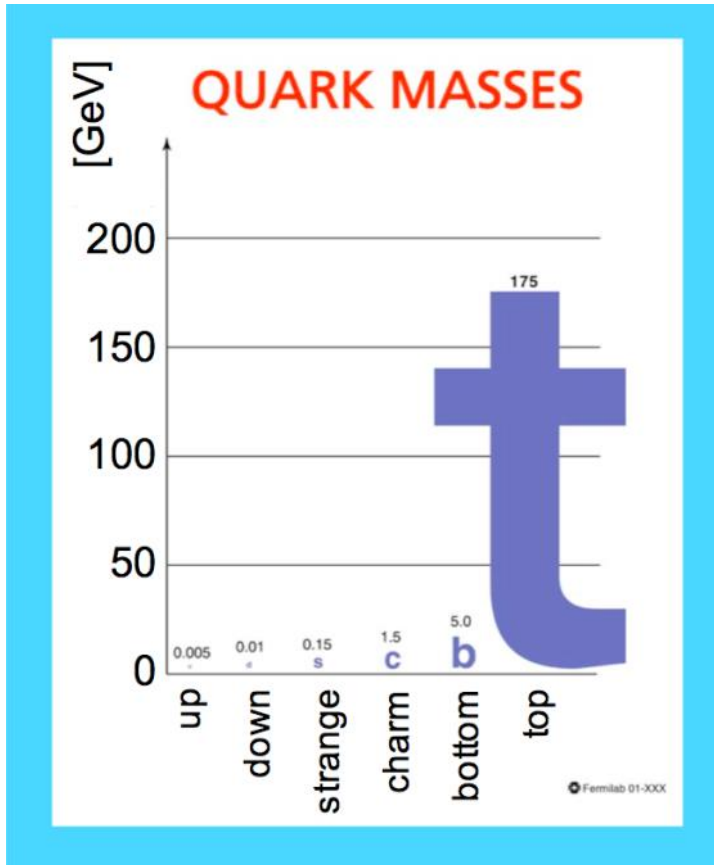
[arXiv:1210.1718](https://arxiv.org/abs/1210.1718)



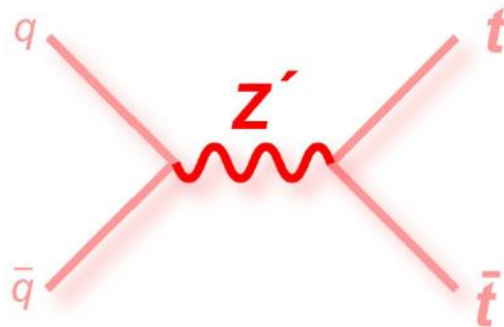
Models and Limits:

- Quark contact interaction (quark compositeness)
 - $\Lambda > 7.6$ TeV (7.7 TeV)
- Quantum Black holes
 - $M_D > 4.1$ TeV (4.2 TeV) $n=6$

New Physics Searches with high-pt top quarks

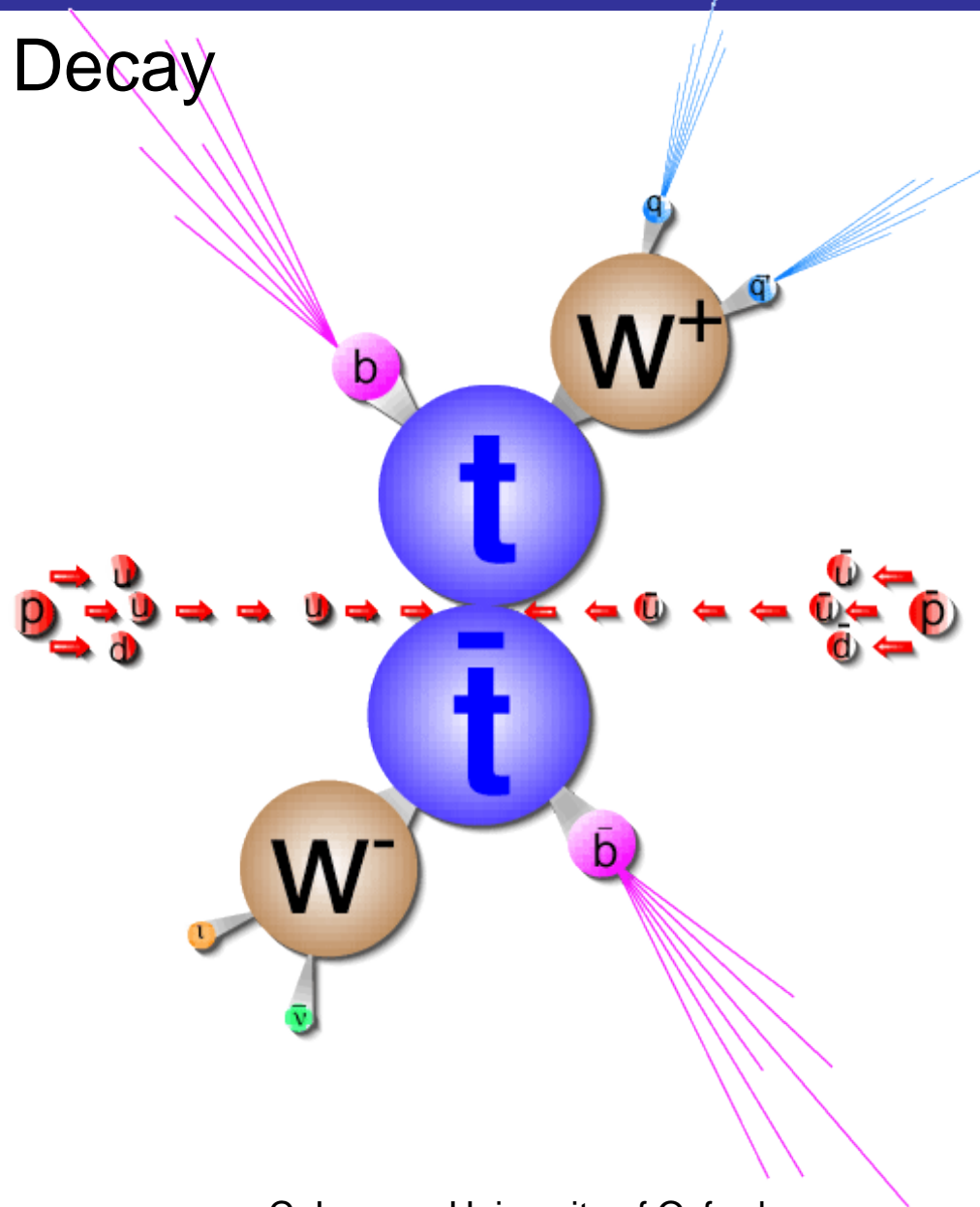


- Huge mass of top
 - Bizarre \rightarrow New Physics?
- Coupled to EWK symmetry breaking
- LHC is a top factory
- Heavy new particles
 - Couple strongly to top
 - Produce boosted tops



Top Quark Production and Decay

Semi-Leptonic Decay

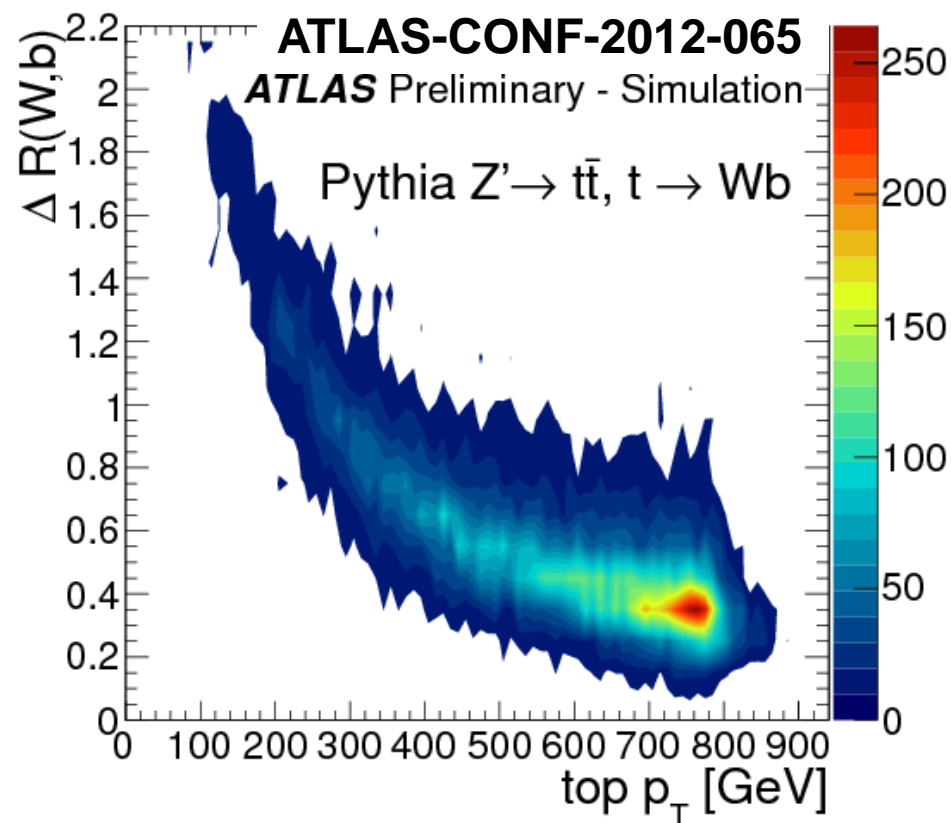


Boosted Regime

- Rule of thumb:

$$dR \sim \frac{2m}{p_T}$$

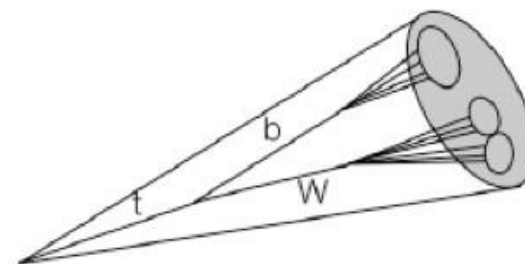
- top with $p_T > 350$ GeV
decay products within $R \sim 1$



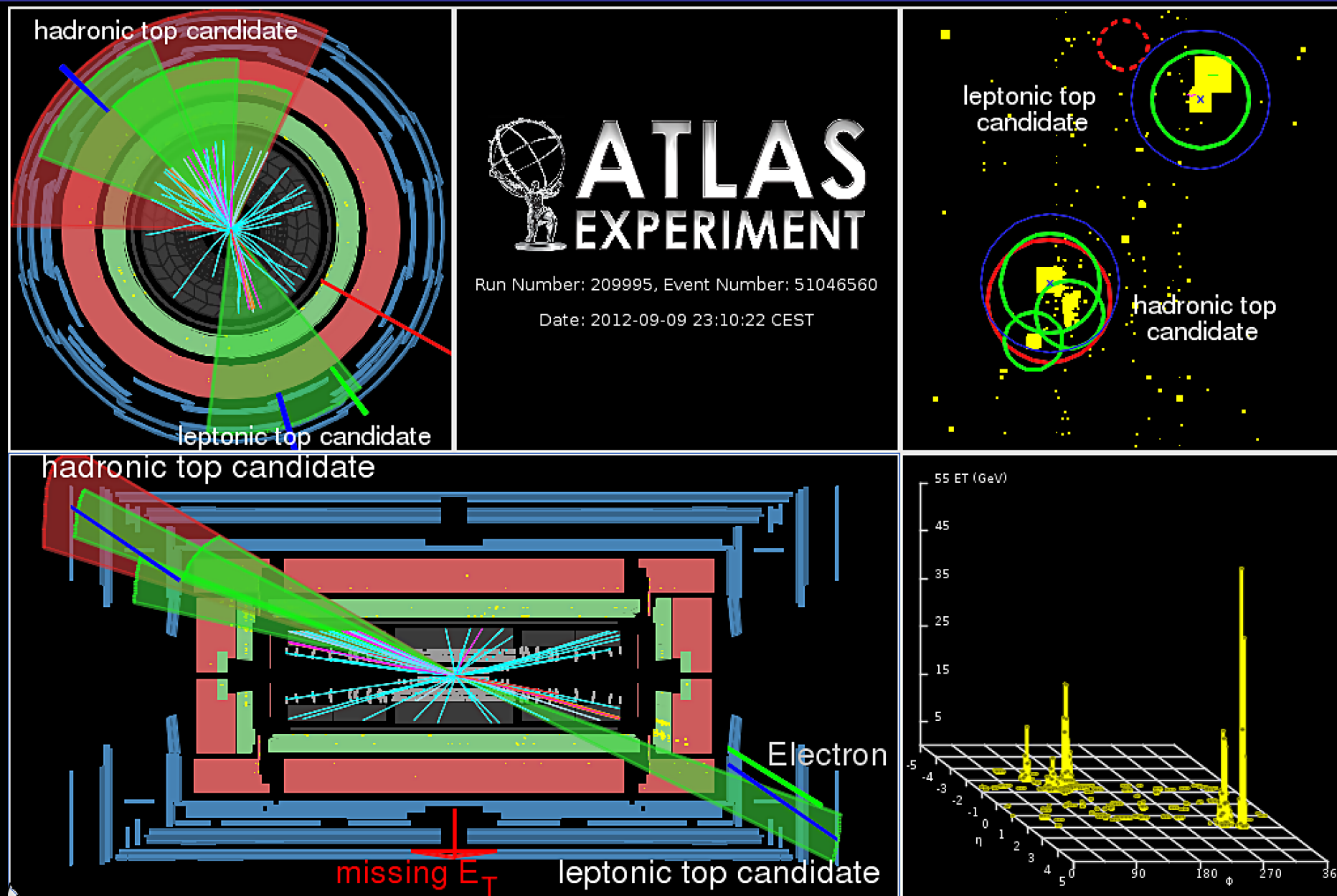
hadronic top
candidate

leptonic top
candidate

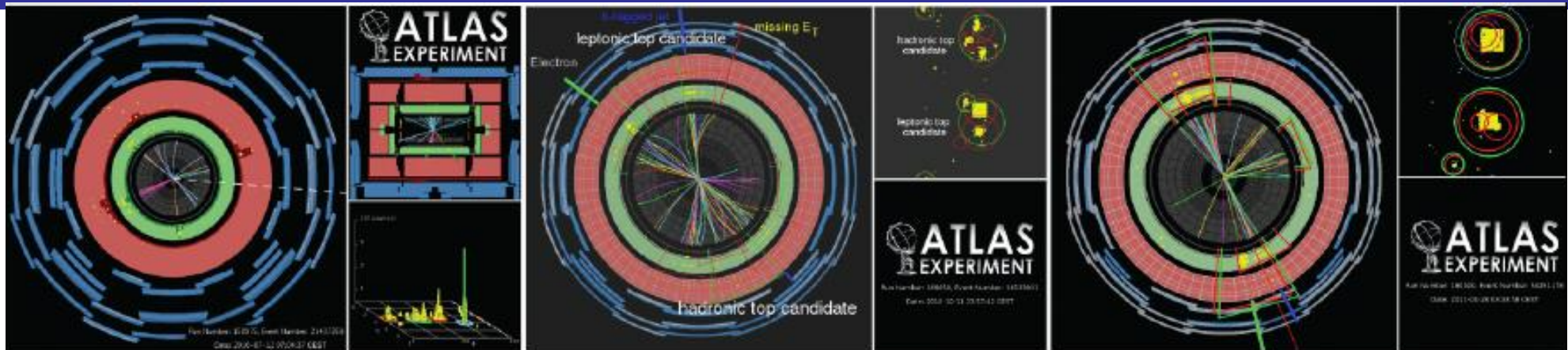
$R = 1$
 $m_j = 197$ GeV
 $E_T = 356$ GeV



Boosted Top Event Candidate with $m_{t\bar{t}} = 2.5$ TeV



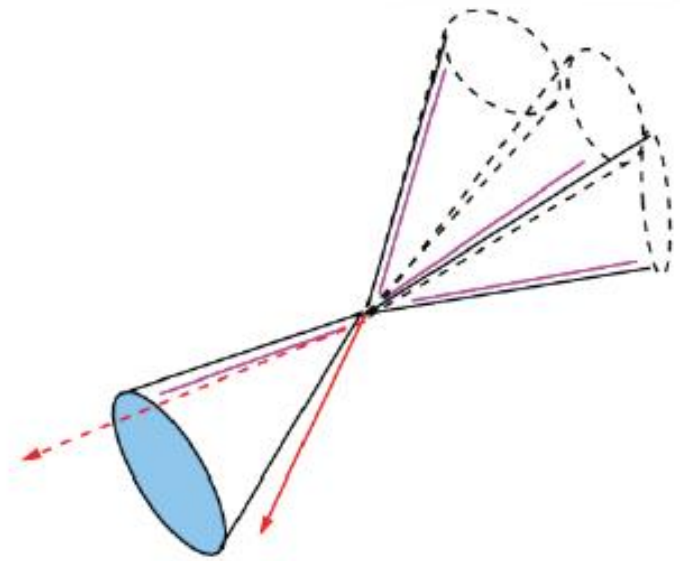
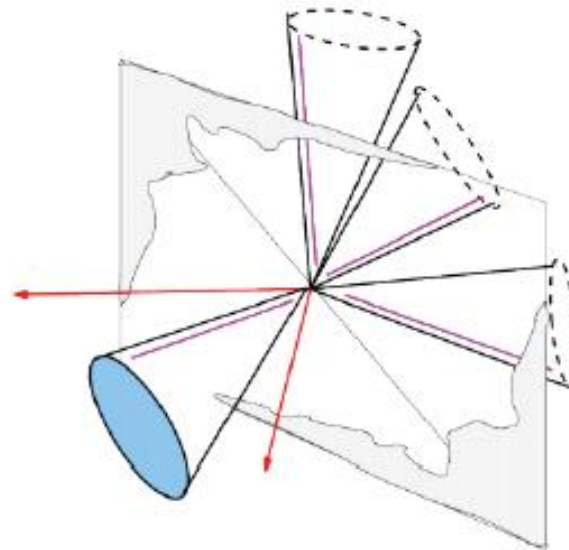
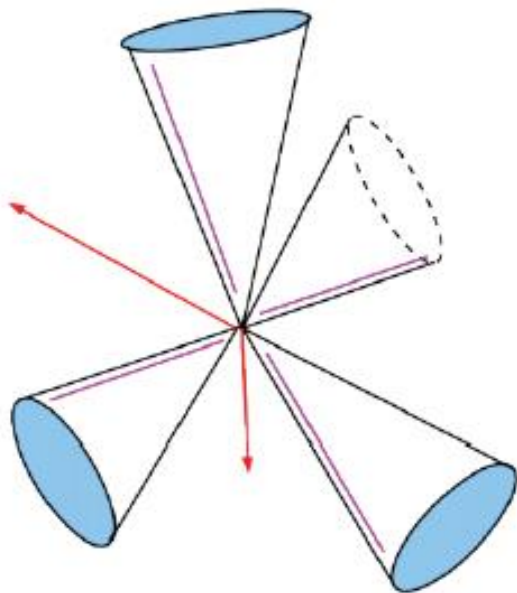
Top Reconstruction @ LHC: 3 Regimes



At rest: $M_{tt} < 500 \text{ GeV}$

Transition region:
 $500 \text{ GeV} < M_{tt} < 700 \text{ GeV}$

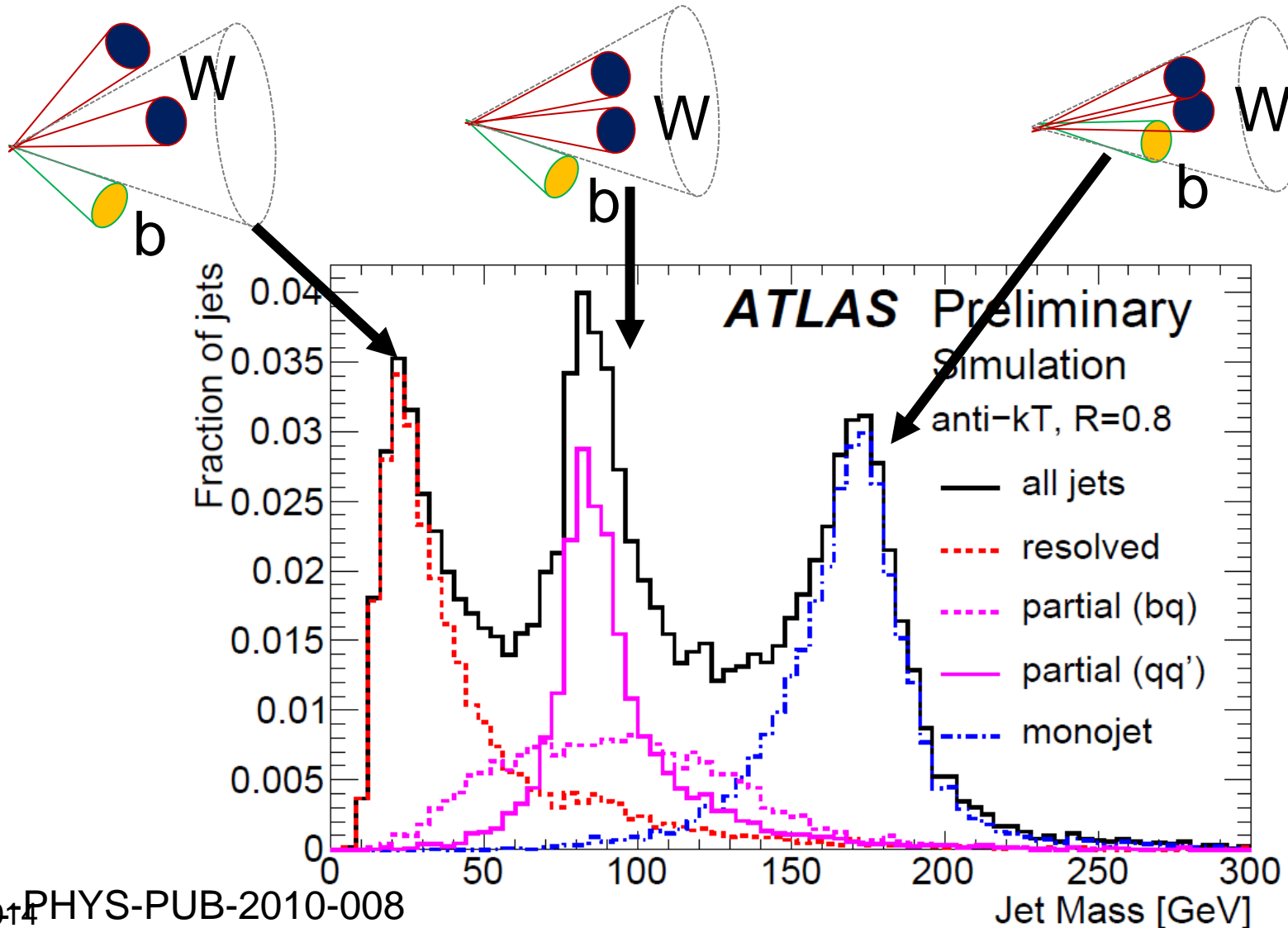
Mono-jet: $M_{tt} > 700 \text{ GeV}$



ATL-PHYS-PUB-2008-010

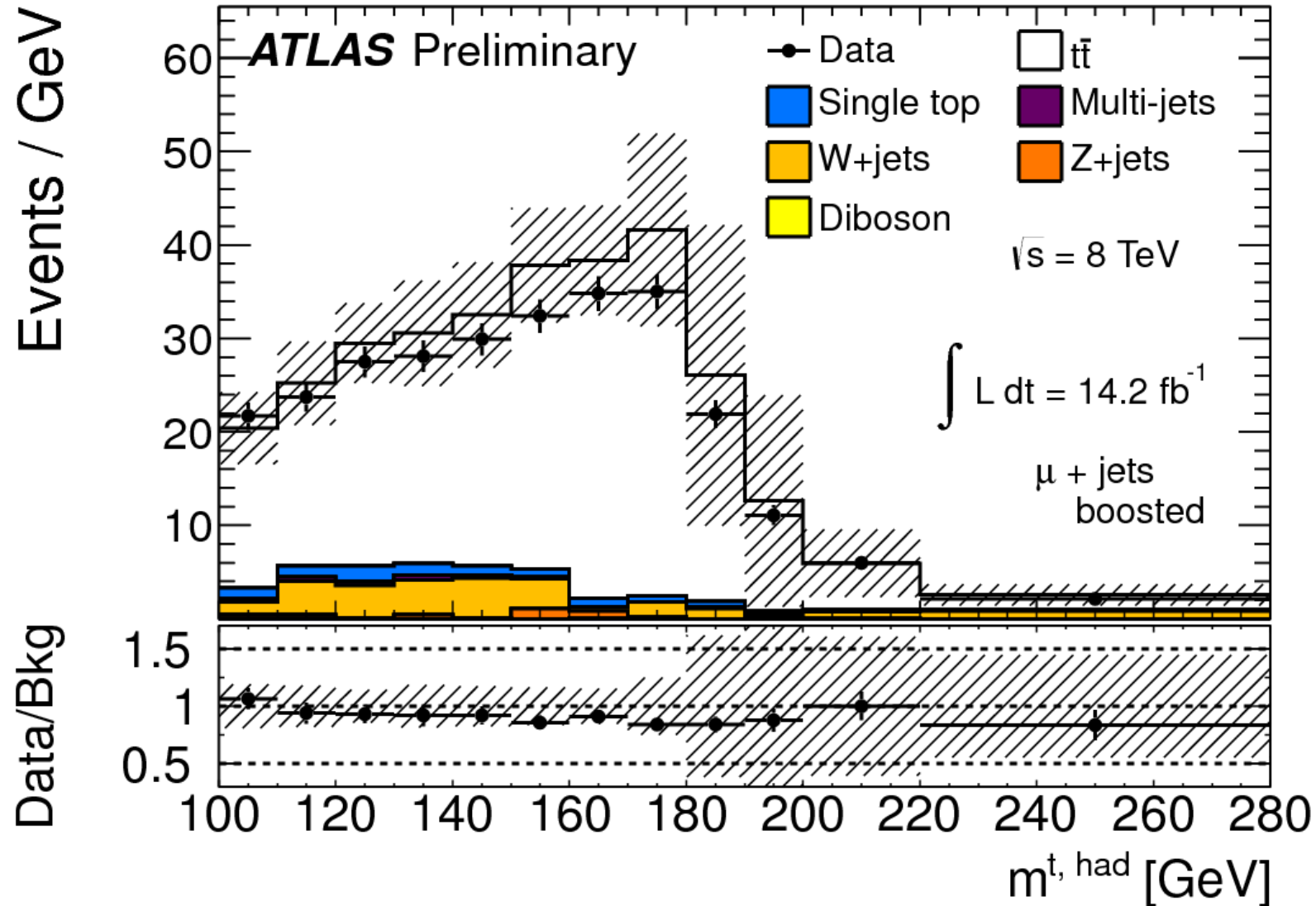
Jet Substructure: jet mass

- Use jet substructure to “tag” boosted tops



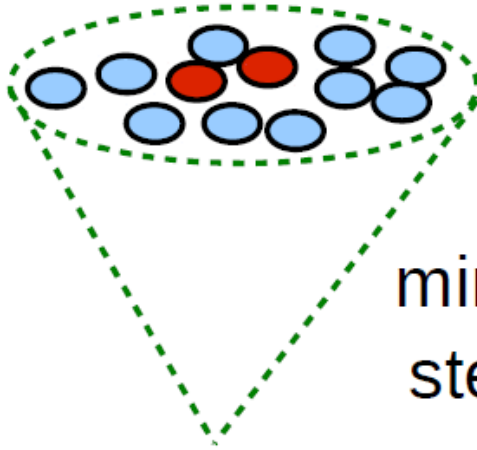
Jet Mass

ATLAS-CONF-2013-052



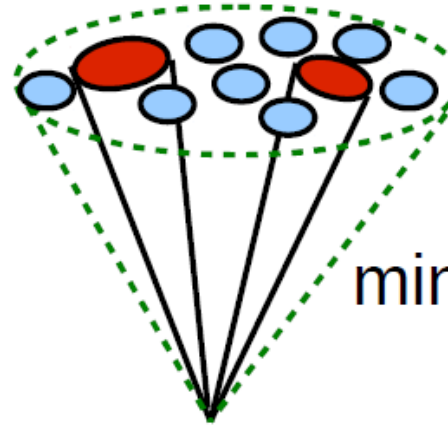
Jet Substructure: Splitting Scale

QCD



$\min p_T \times dR =$
steeply falling
spectrum

Boosted W boson

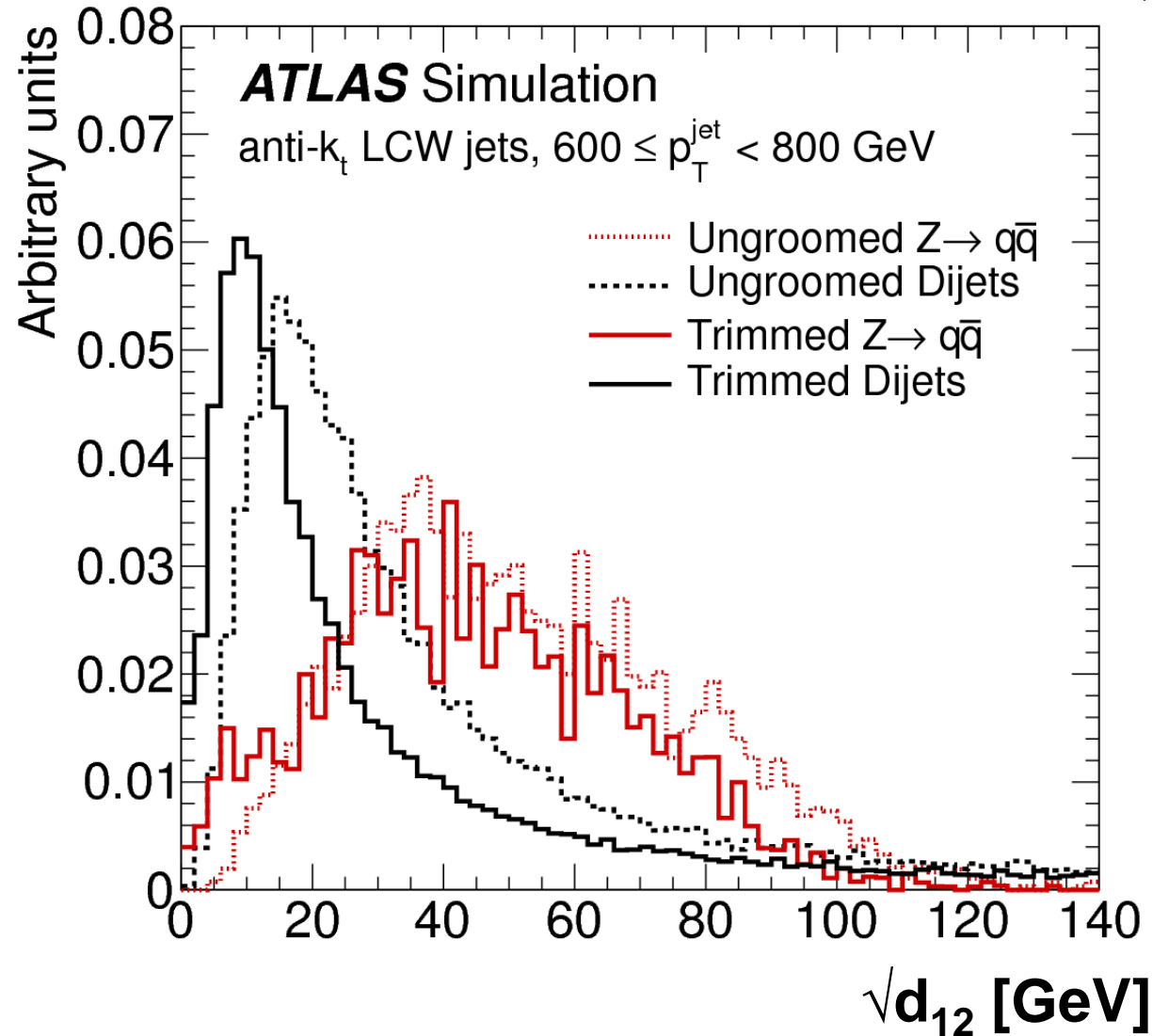


$\min p_T \times dR =$
 $M_{\text{jet}} / 2$

$$d_{ij} = \min(p_{Ti}^2, p_{Tj}^2) \times \Delta R_{ij}^2 / R^2$$

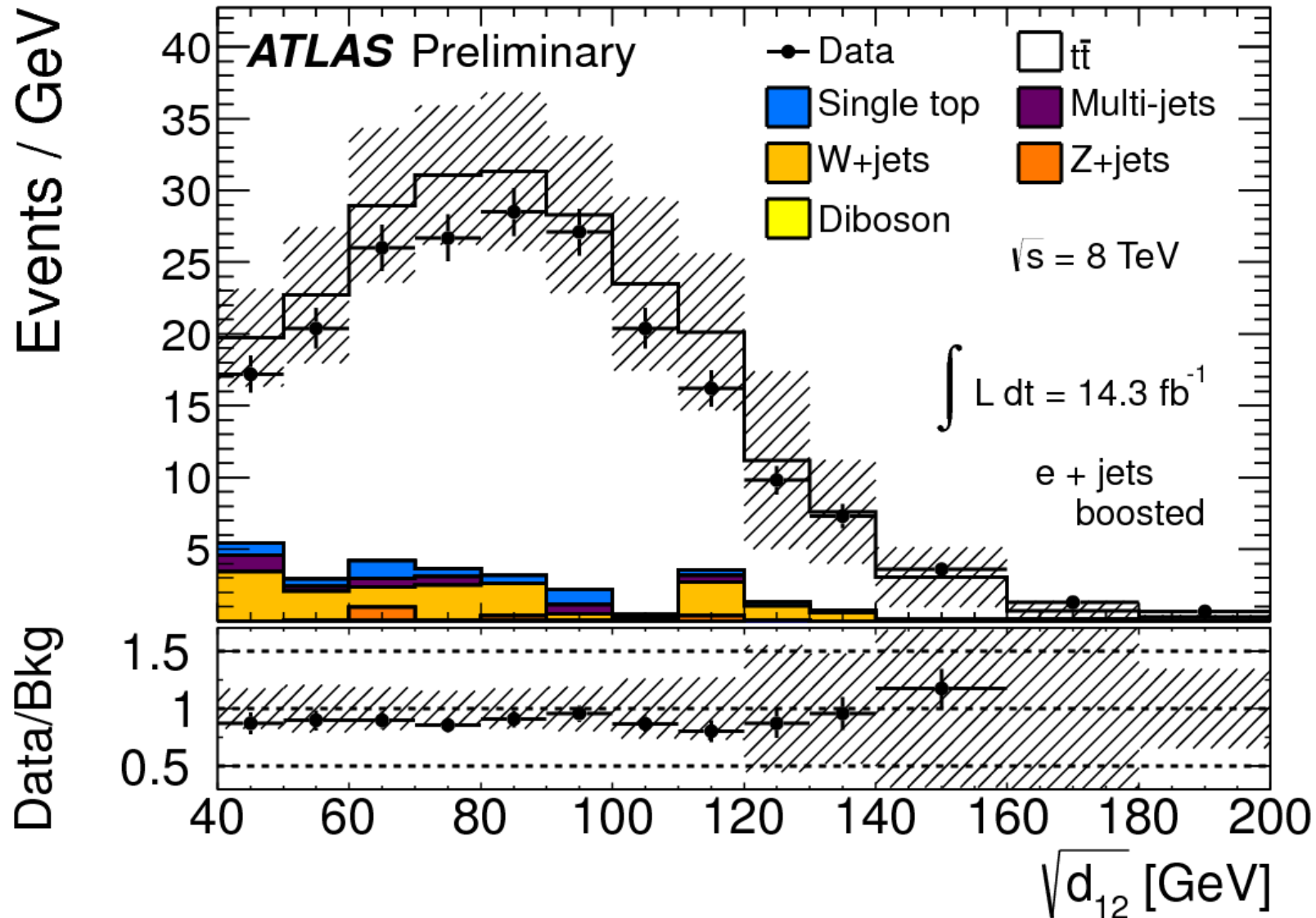
Jet Substructure: Splitting Scales

CERN-PH-EP-2013-069, arXiv:1306.4945



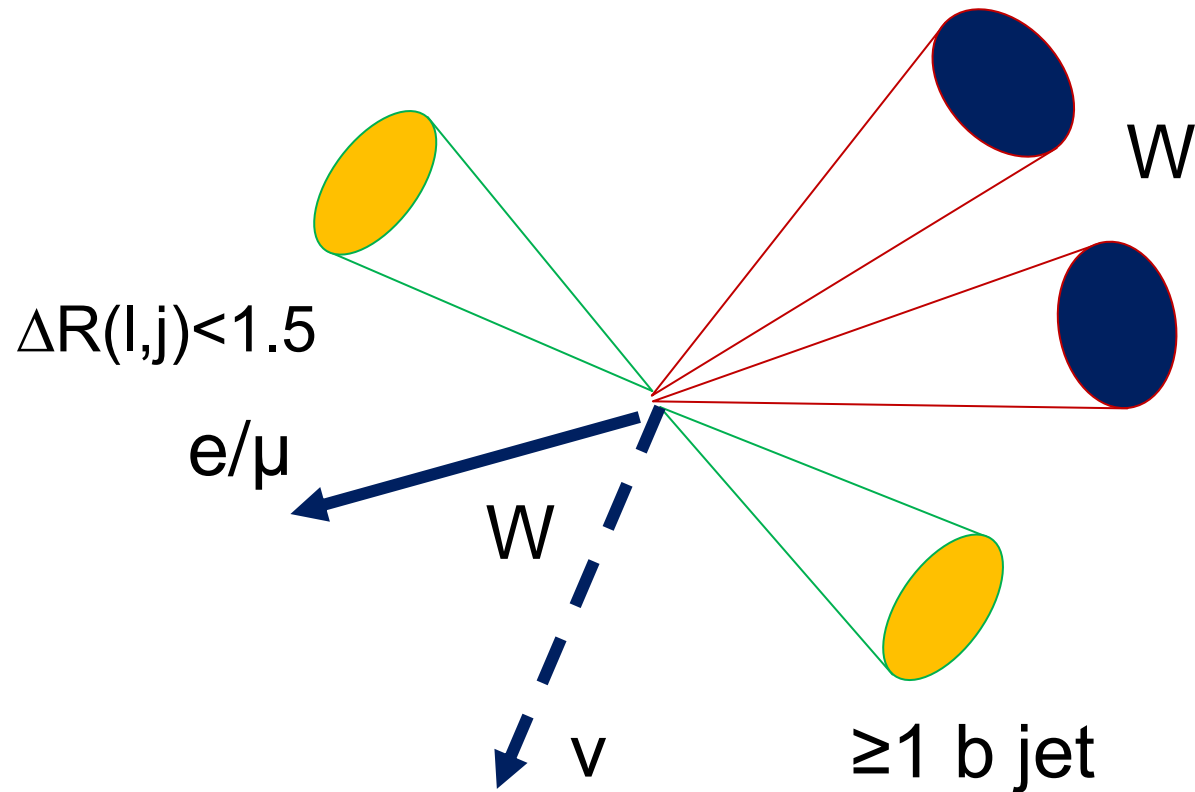
Jet Substructure: Splitting Scale

ATLAS-CONF-2013-052



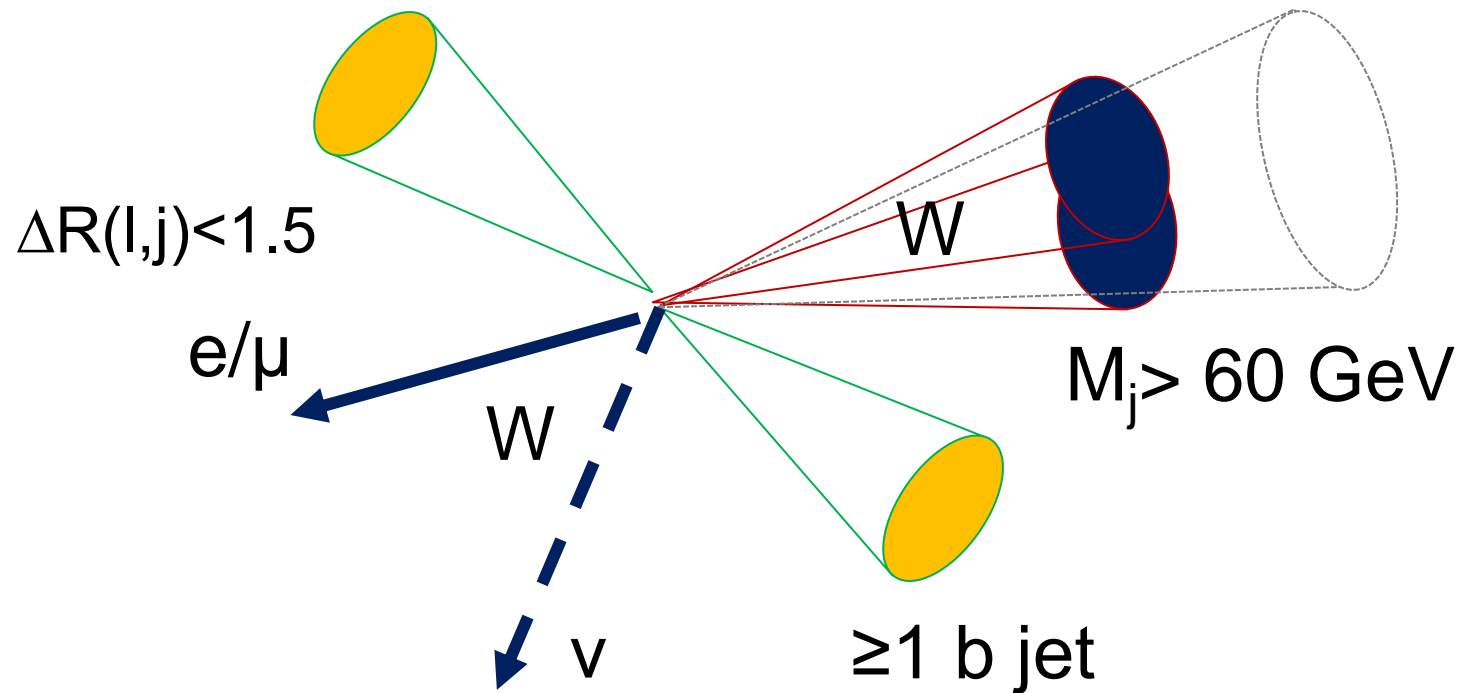
Resolved Selection

≥ 4 small jets, j , with $p_T > 25$ GeV, $|\eta| < 2.5$



Merged Selection

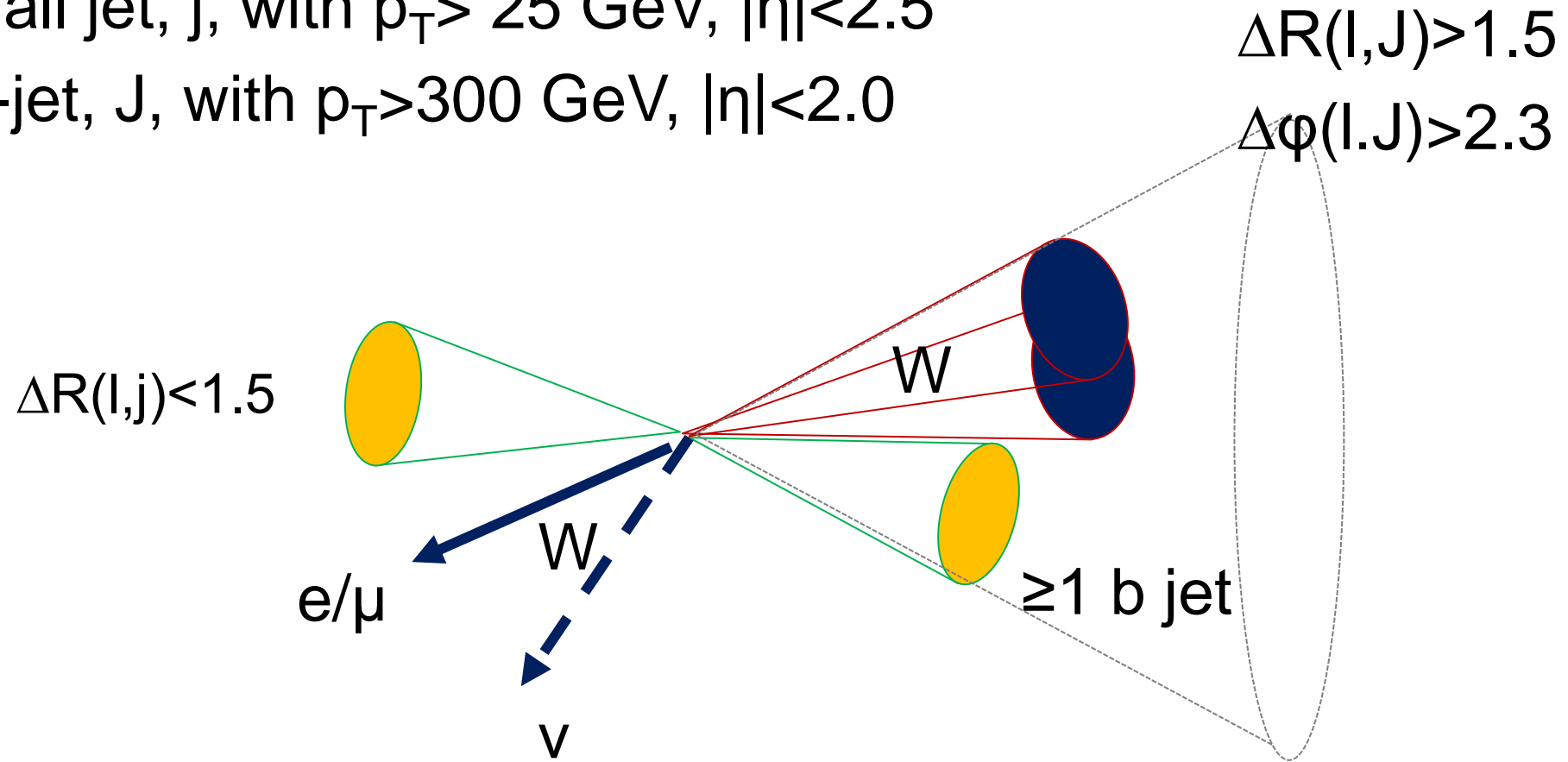
3 small jets, j , with $p_T > 25$ GeV, $|\eta| < 2.5$



Boosted Selection

≥ 1 small jet, j , with $p_T > 25$ GeV, $|\eta| < 2.5$

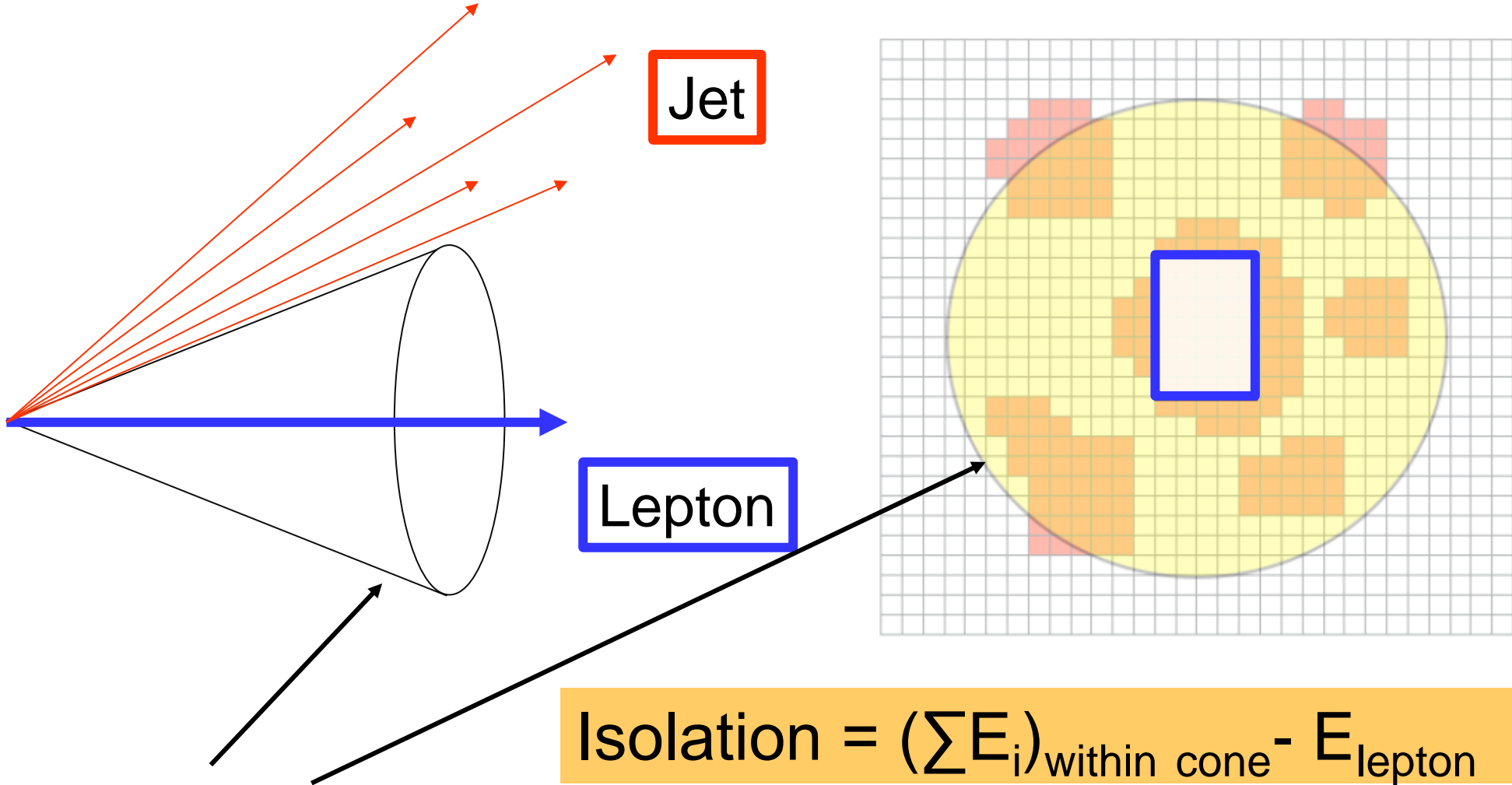
≥ 1 fat-jet, J , with $p_T > 300$ GeV, $|\eta| < 2.0$



$M_J > 100$ GeV

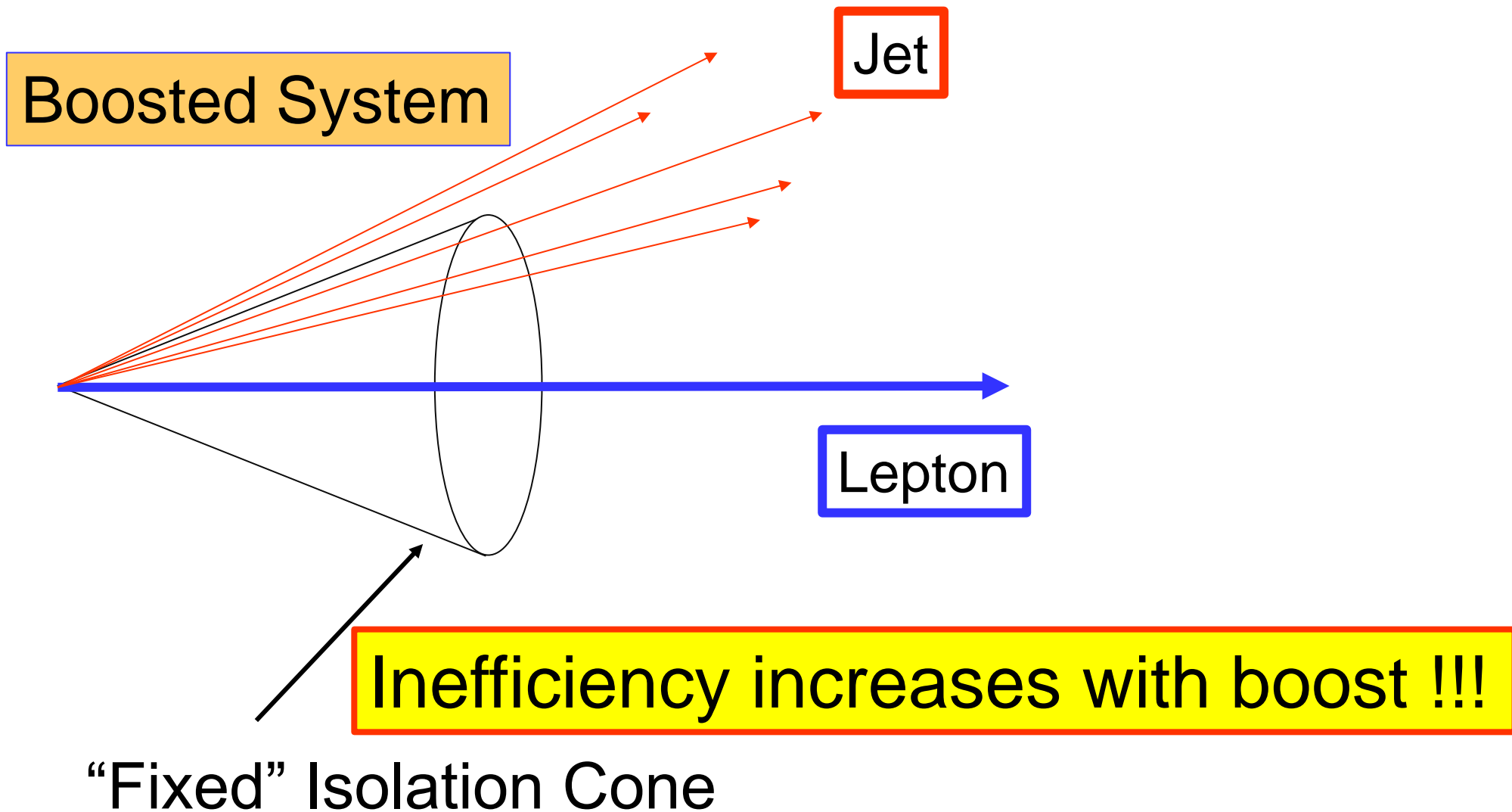
$\sqrt{d_{12}} > 40$ GeV

Fixed Cone Size Lepton Isolation



“Fixed” Isolation Cone

Fixed Cone Size Isolation



“Mini”-Isolation

$R = k/E_T$ or k/p_T
e.g. $k = 10 \text{ GeV}$

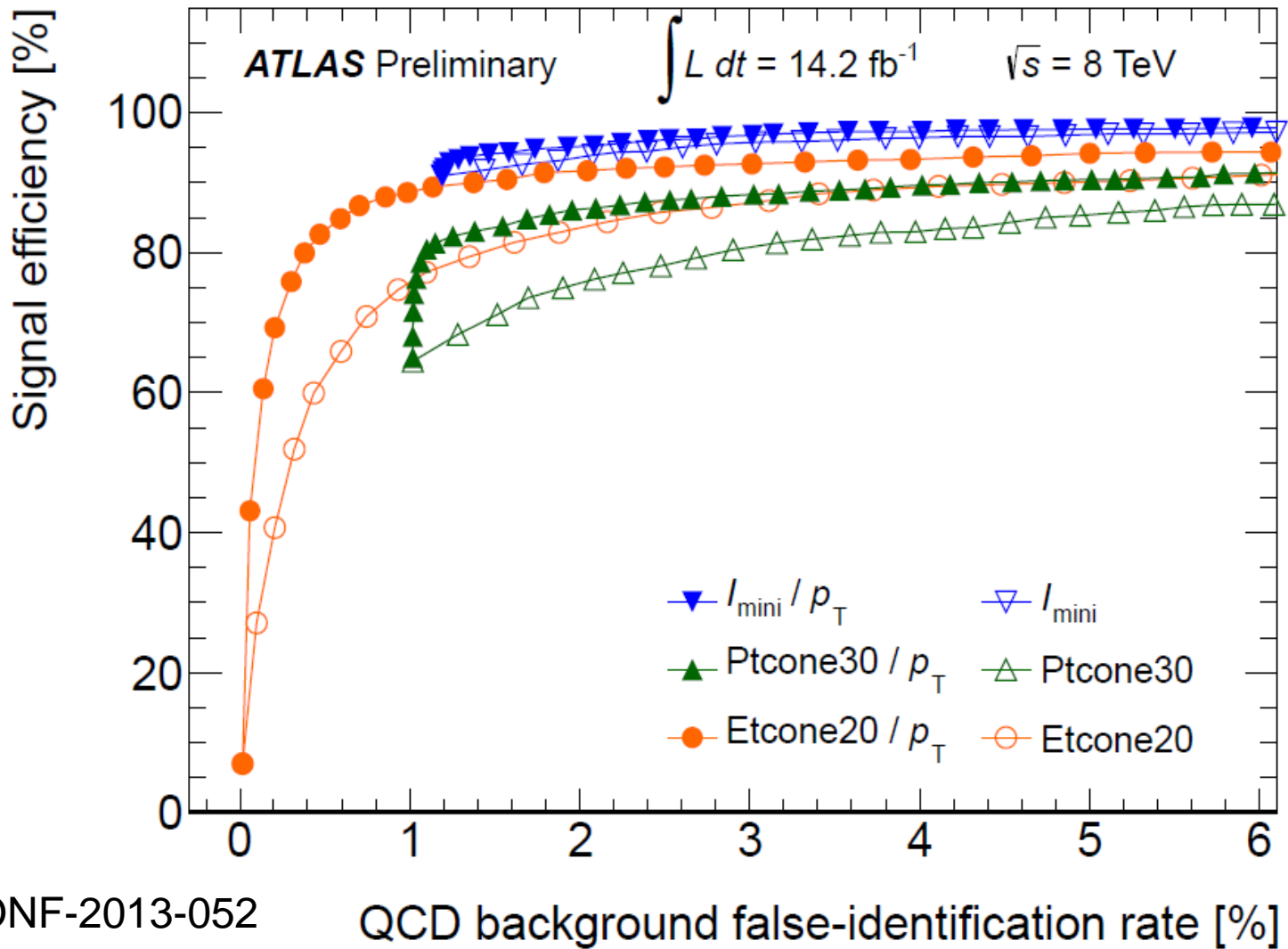
Jet

Lepton

$I_{\text{mini}} < 0.05 * E_T$ (for electrons)
 $I_{\text{mini}} < 0.05 * p_T$ (for muons)

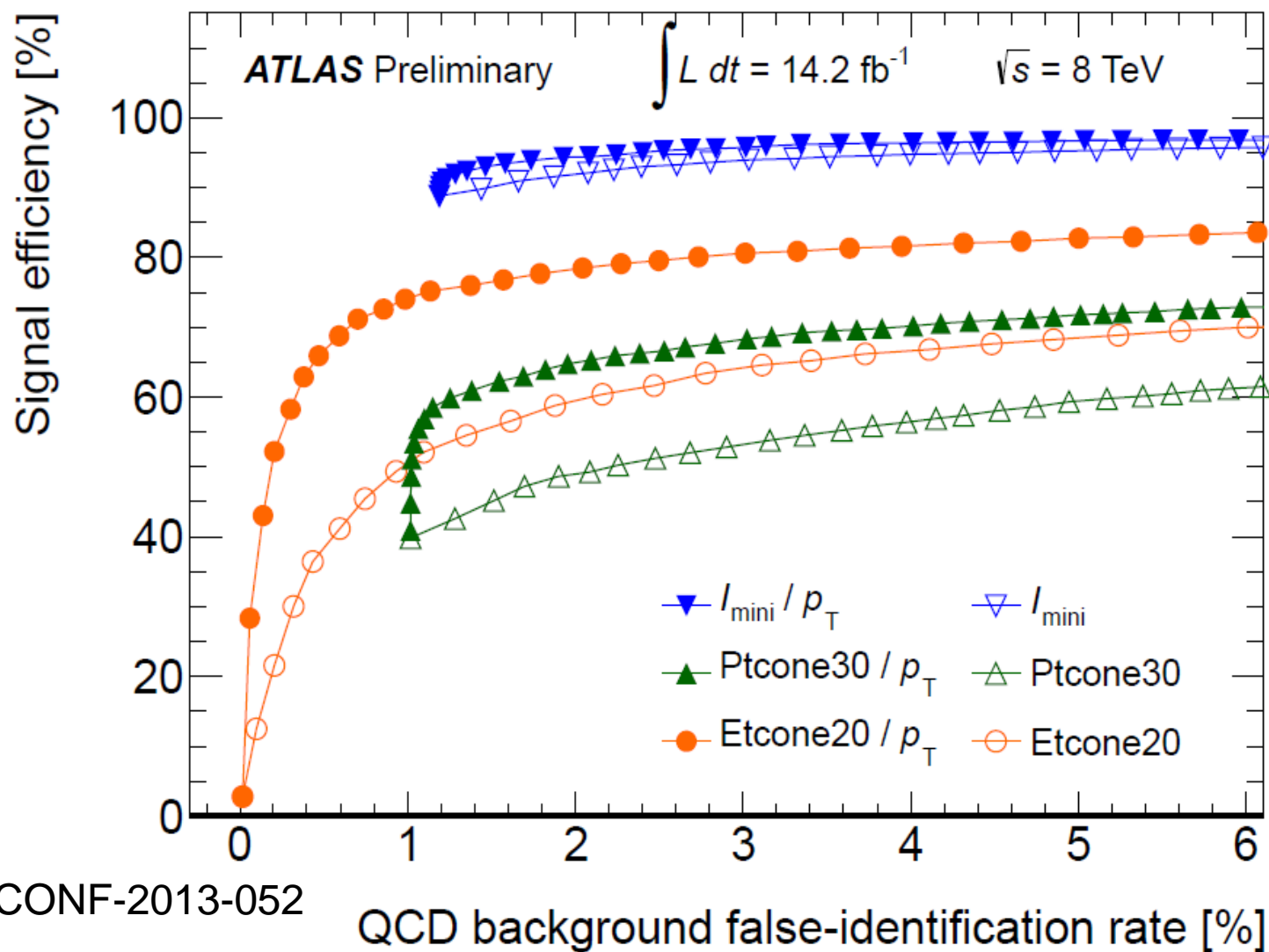
“Variable” Isolation Cone

Efficiency Comparisons



(b) 1.0 TeV Z'

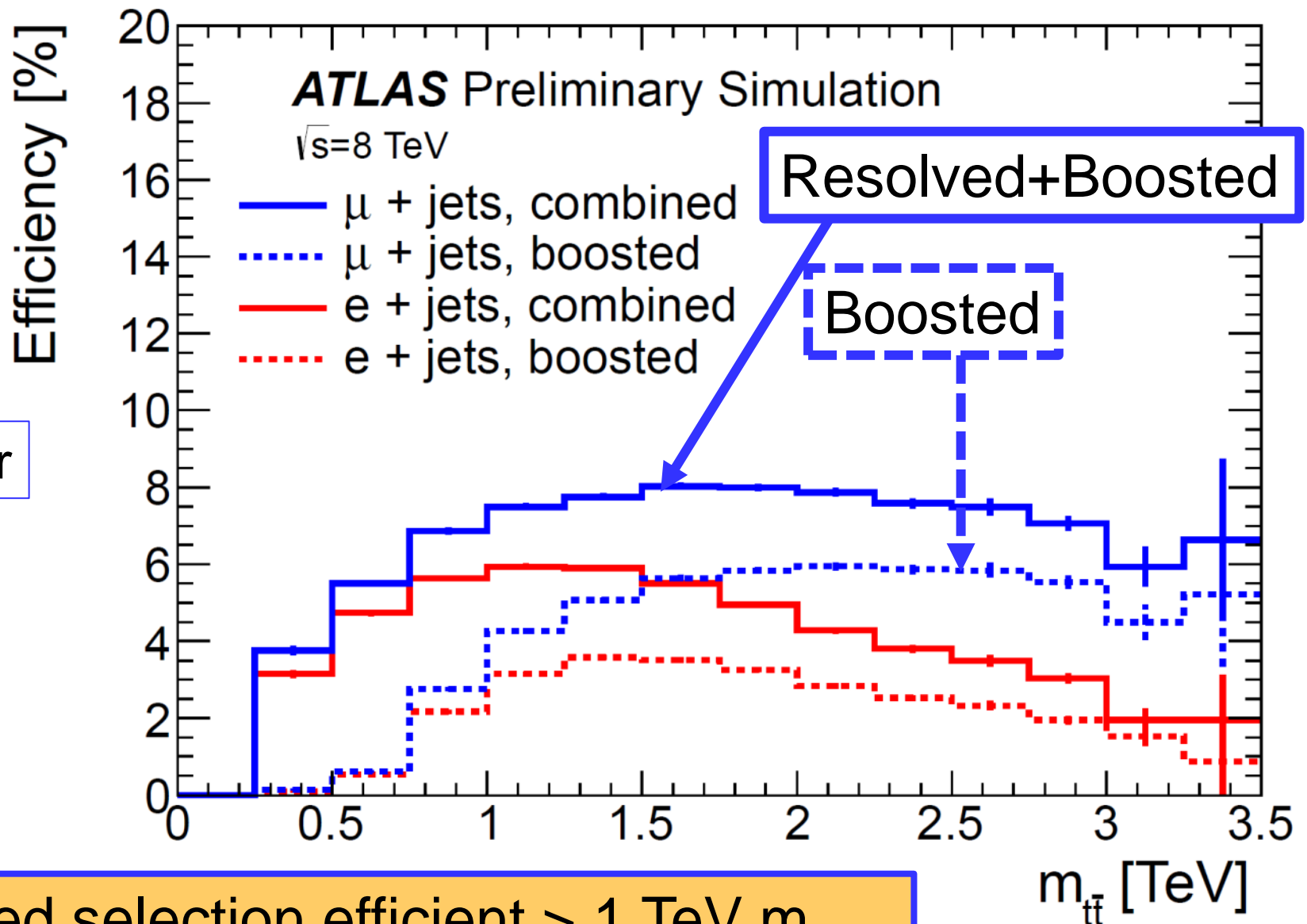
Efficiency Comparisons



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(d) 2.0 TeV Z'

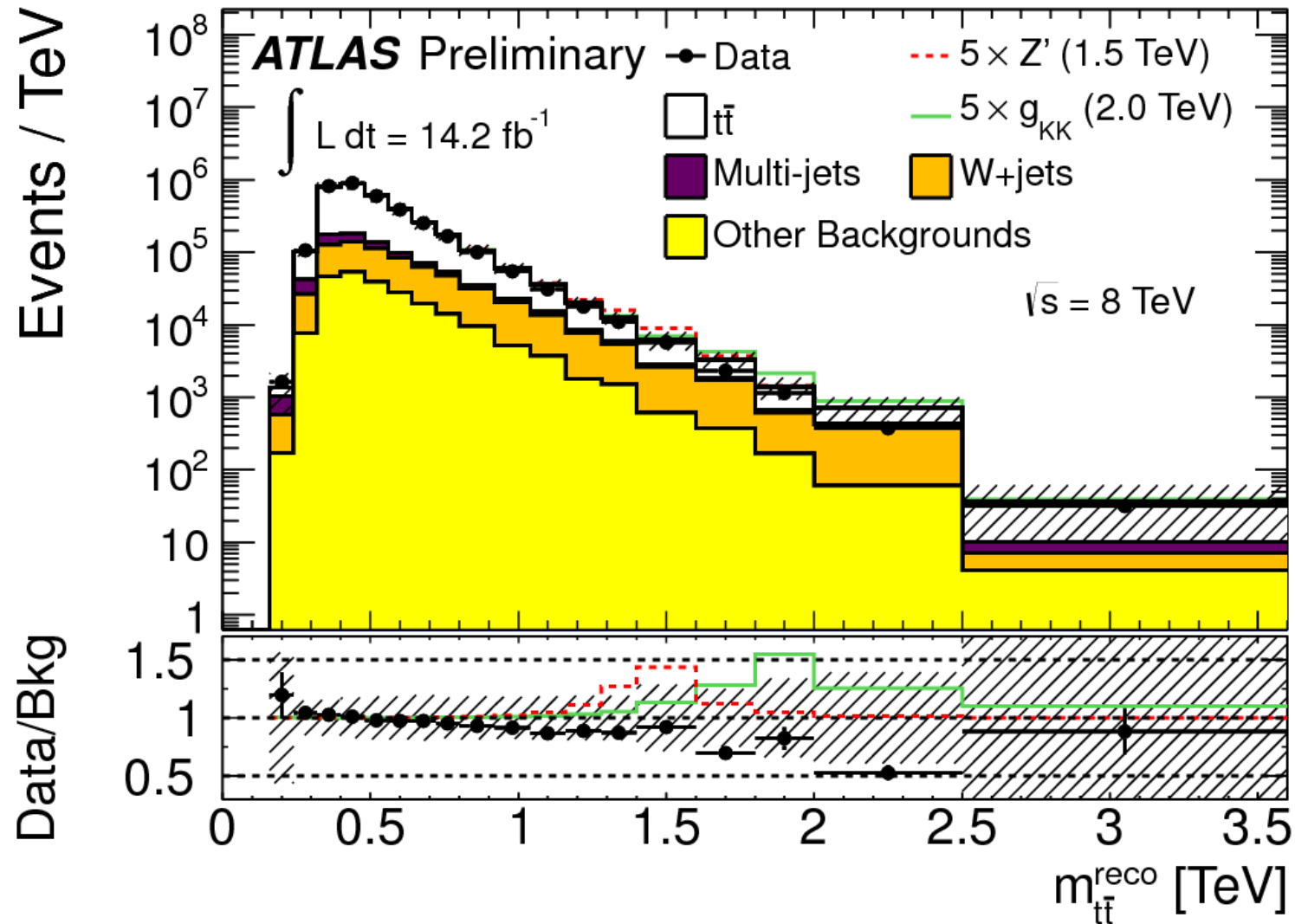
Geometrical Acceptance + Selection Efficiencies



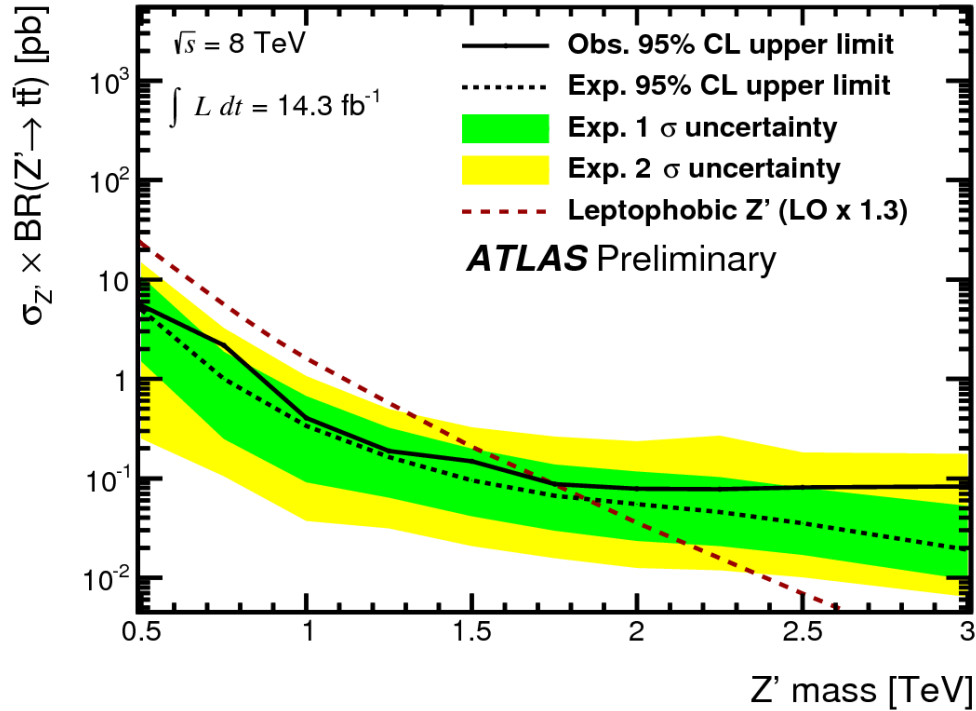
Boosted selection efficient > 1 TeV $m_{t\bar{t}}$

Discriminant distribution $m_{t\bar{t}}$

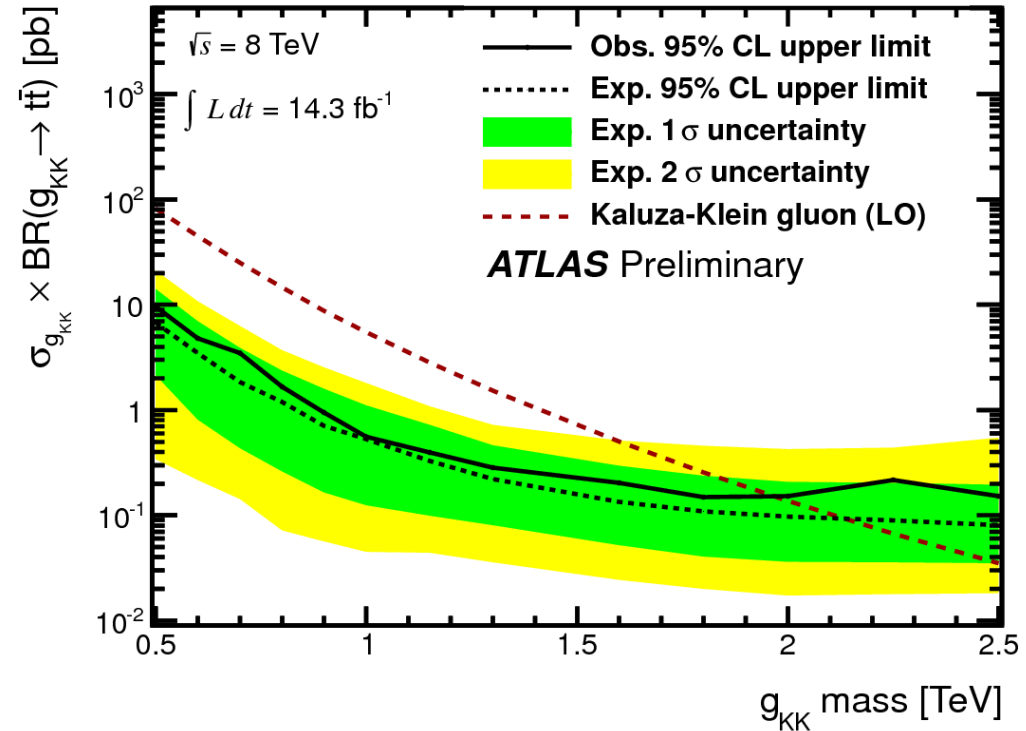
- $m_{t\bar{t}}$ resolved + boosted in e +jets and μ +jets



Heavy Resonances Search: $T\bar{t}$

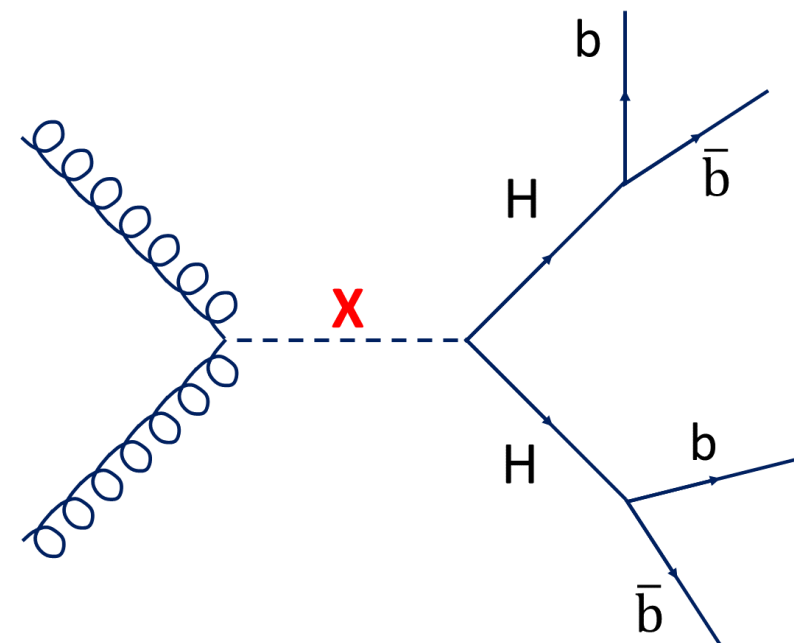
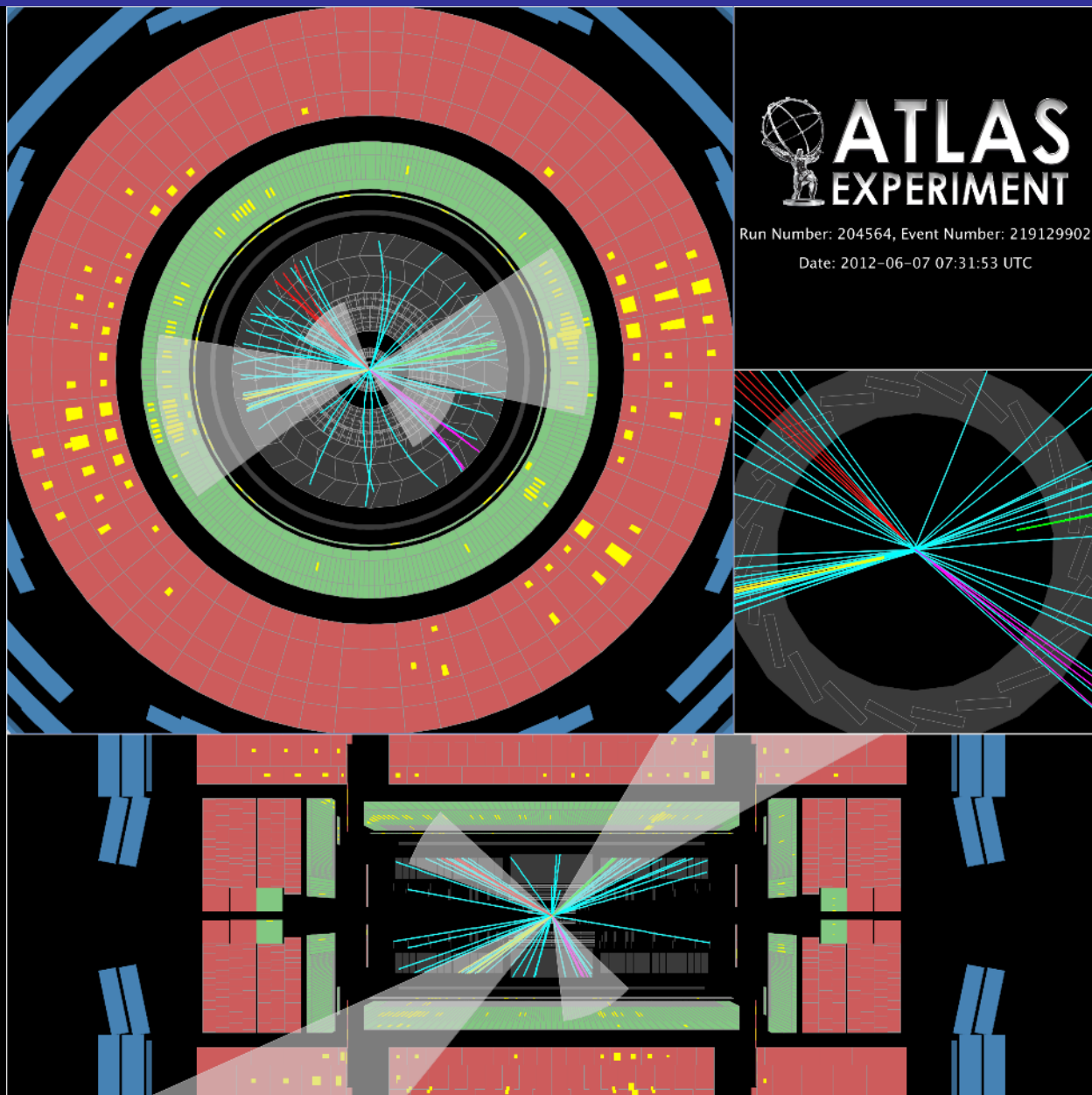


$m(Z') > 1.8 \text{ TeV @95\% CI}$
 $\Gamma/m(Z') = 1.2\%$



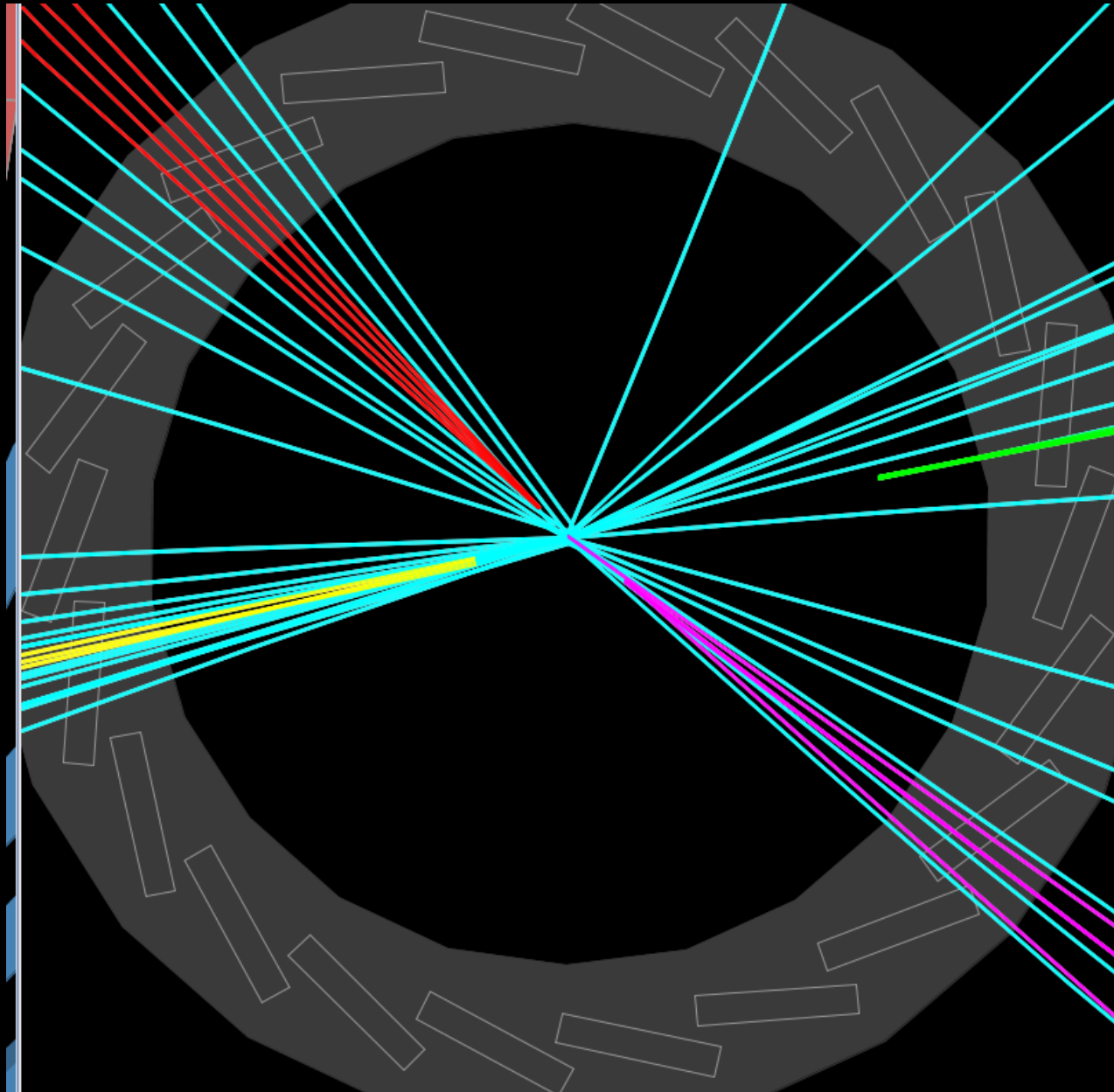
$m(g_{KK}) > 2.0 \text{ TeV @95\% CI}$
 $\Gamma/m(g_{KK}) = 15\%$

Resonance Searches with Higgs Pairs



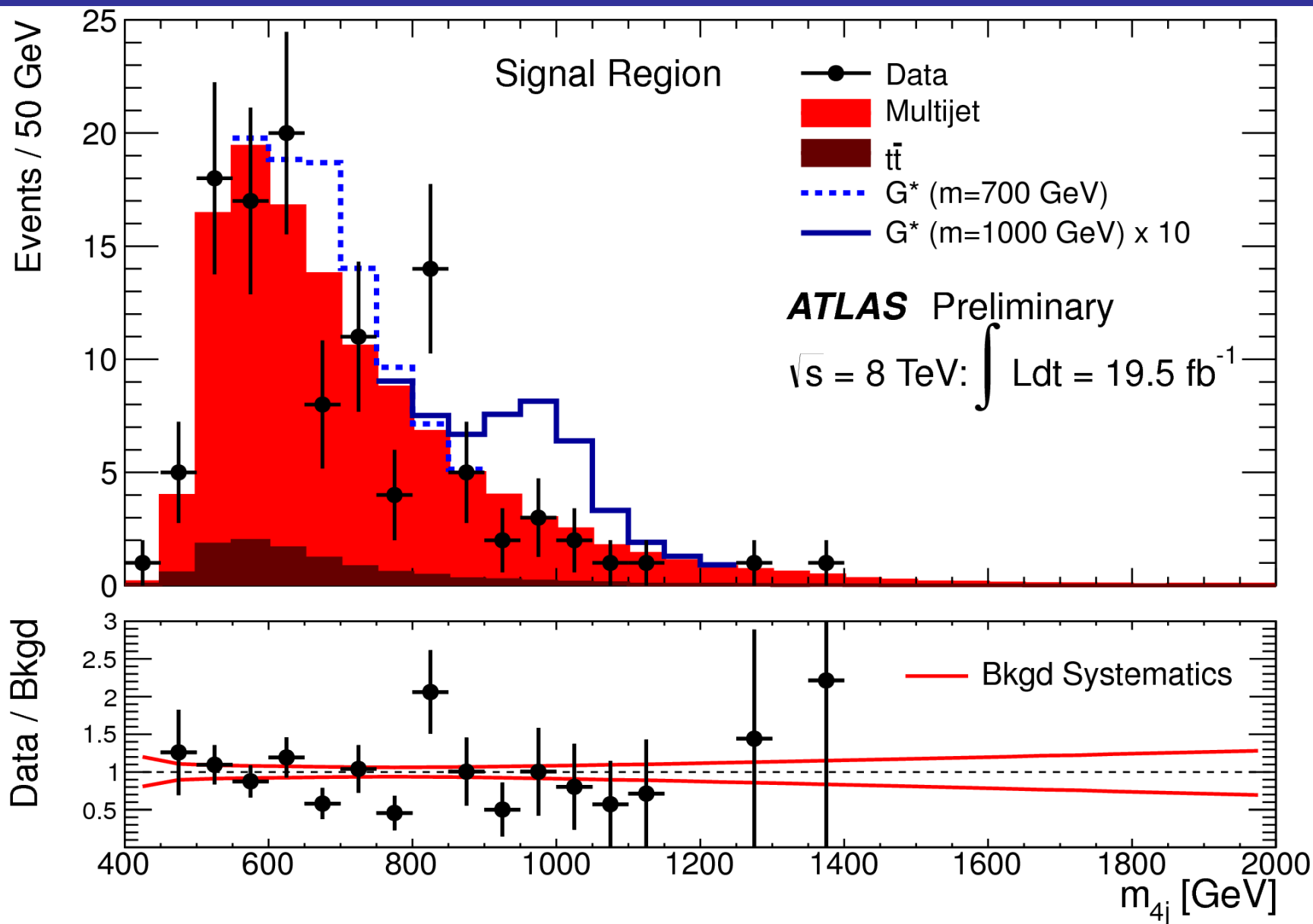
ATLAS-CONF-2014-00

Resonance Searches with Higgs Pairs

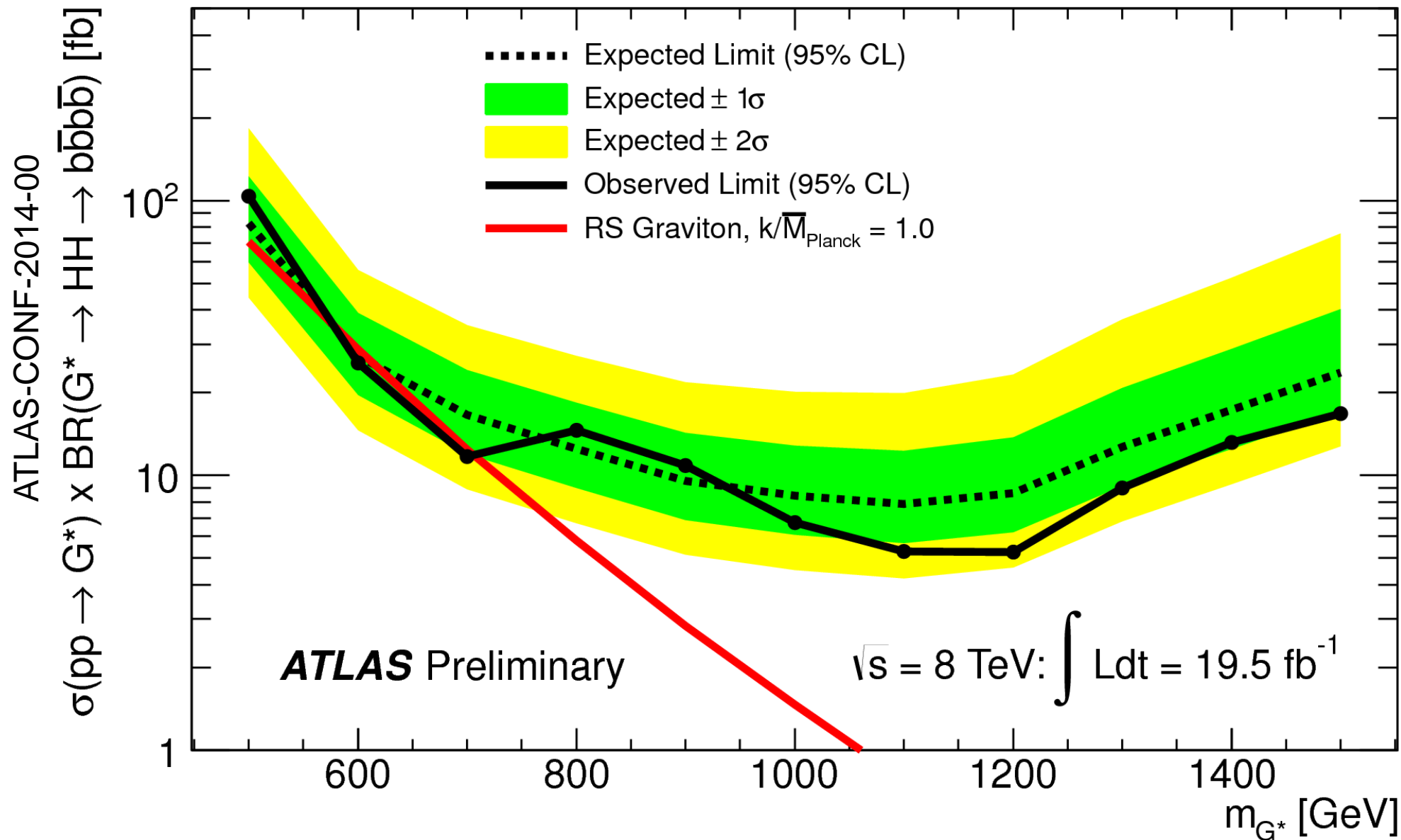


2H \rightarrow 4 b resonance search

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2H \rightarrow 4 b resonance search

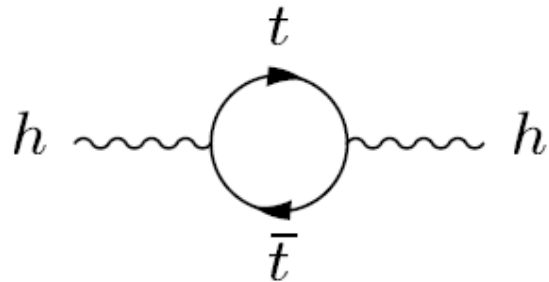


Heavy Quarks

Quarks	u	c	t	t'
	d	s	b	b'
Leptons	ν_e	ν_μ	ν_τ	ν'
	e	μ	τ	τ'
	I	II	III	IV

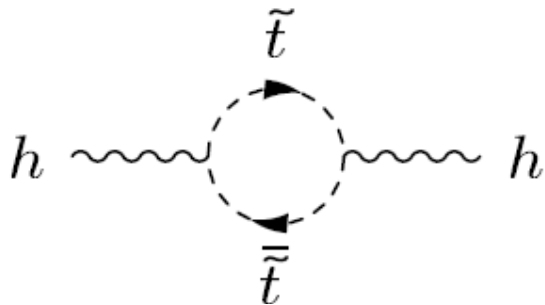
Fine Tuning Problem and SUSY

- Same problem with Higgs



$$\Delta\mu_{\text{top}}^2 = -6 \frac{h_t^2}{4\pi^2} \frac{1}{r_H^2} \sim (100 \text{ GeV})^2$$

125 GeV = (huge number)-(huge number) even more fine tuned!



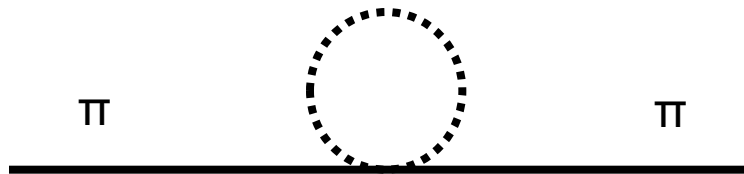
Add new particles (spin symmetry): SUSY

$$\Delta\mu_{\text{stop}}^2 + \Delta\mu_{\text{top}}^2 = -6 \frac{h_t^2}{4\pi^2} (m_{\tilde{t}}^2 - m_t^2) \log \frac{1}{r_H^2 m_{\tilde{t}}^2}$$

Composite Higgs

- But there is another way....look at QCD

Pion mass is not divergent.

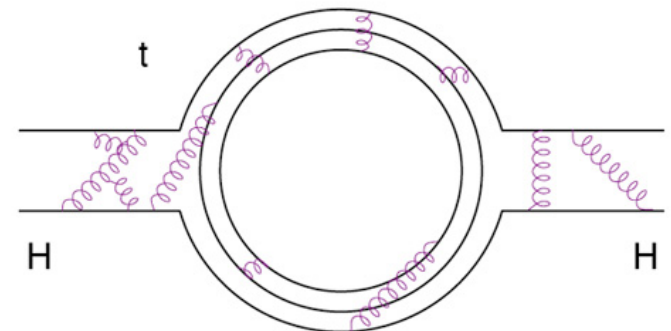


Why?

It is a composite particle!

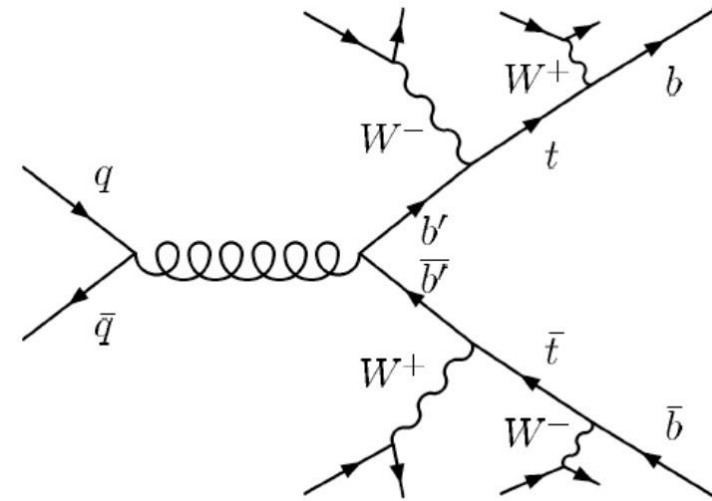
- **Assume Higgs is a composite particle**

- Changes couplings
- Introduces new partners to top quarks
- Vector-like quarks...
 - (both chiralities same under $SU(2) \times U(1)$)
- Solves fine-tuning problem....

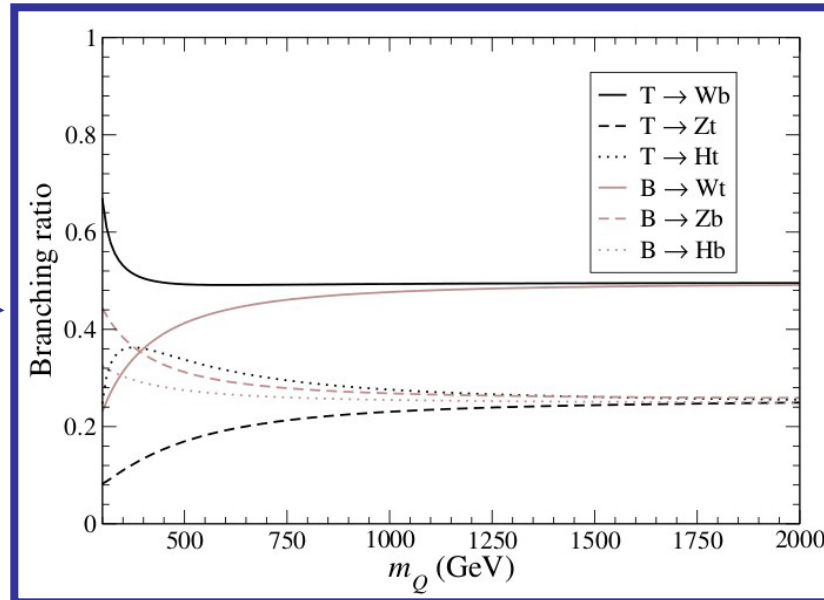


4th Generation and Heavy Quarks

- 4th generation would significantly enhance Higgs production cross section
 - (almost) excluded by observed Higgs cross-section
 - $t't' \rightarrow WbWb$ (100%): just like t - t bar but heavier
 - $b'b' \rightarrow WtWt$ (100%): just like t tbar but messier



- Beyond 4th generation: **Vector-Like Quarks** in Composite Higgs theories
 - More diverse phenomenology
 - T' : Decays to Wb , Zt , Ht
 - B' : Decays to Wt , Zb , Hb
- Loose constraints on CKM4 \rightarrow decays to light quarks possible!



4th Generation and Heavy Quarks

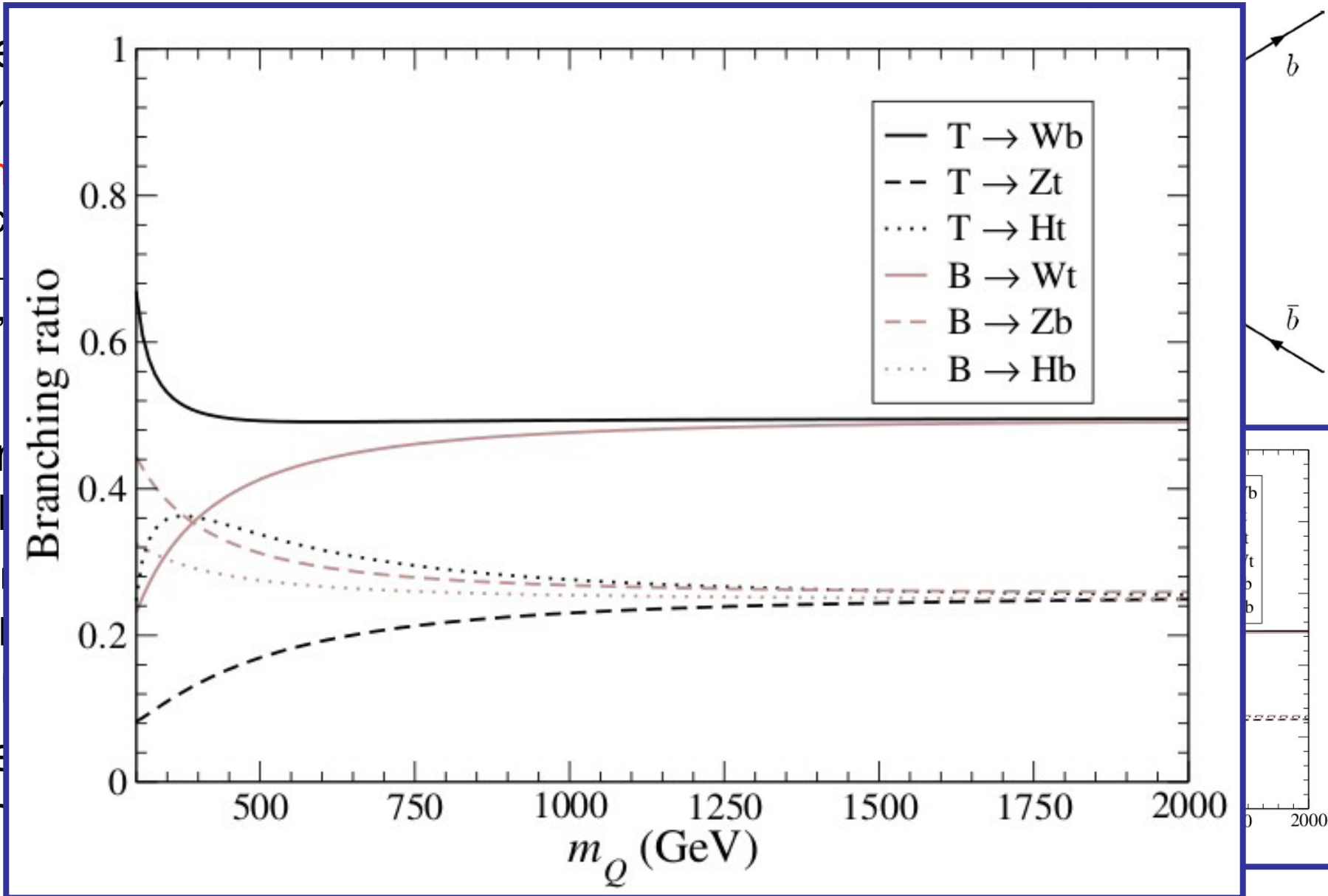
- 4th generation
enhanced

- (almost) equal
second generation
- $t't'$ -
 $b'b'$

- Beyond
Quark

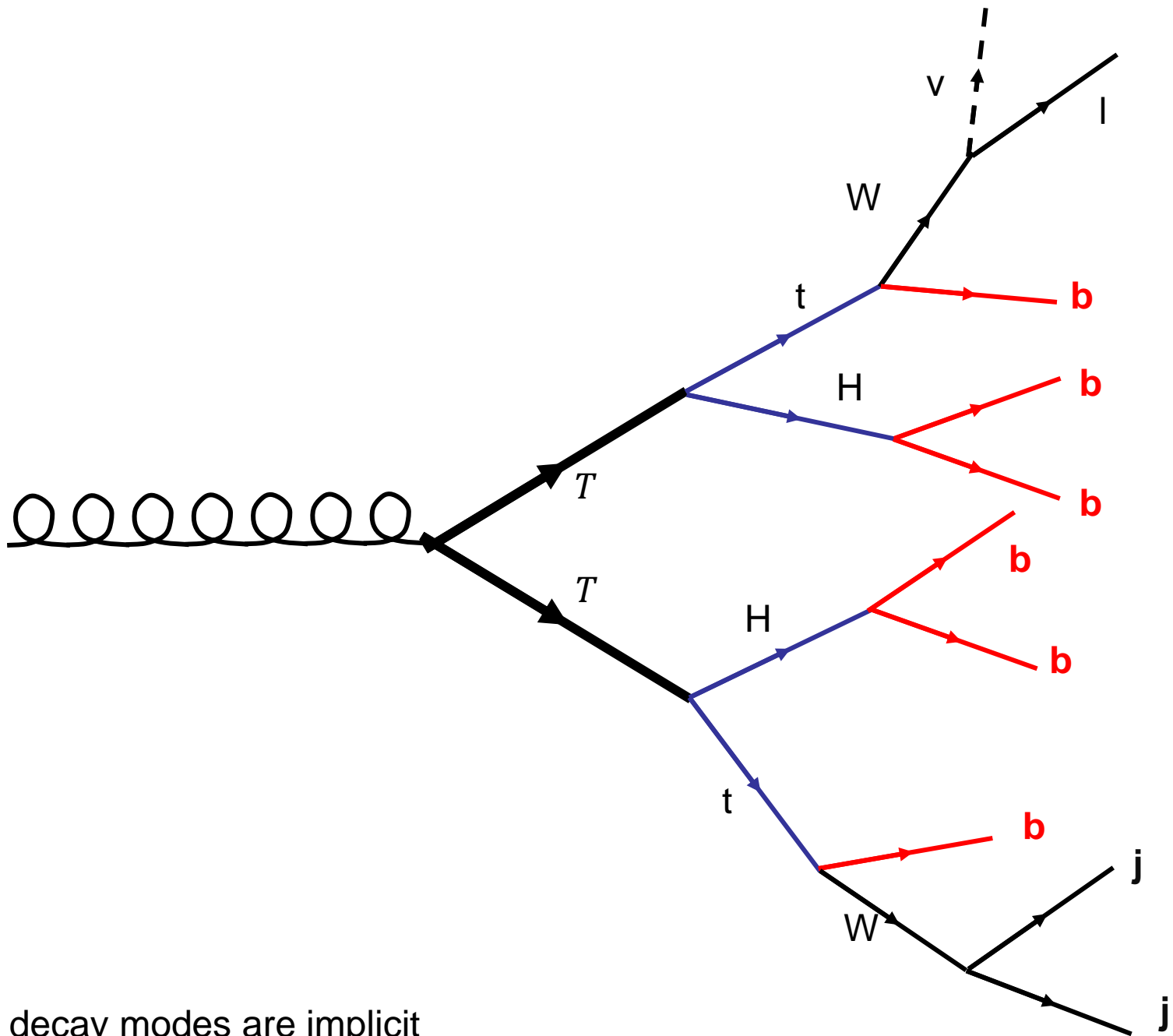
- More
 T' :
 B' :

- Loose
to light

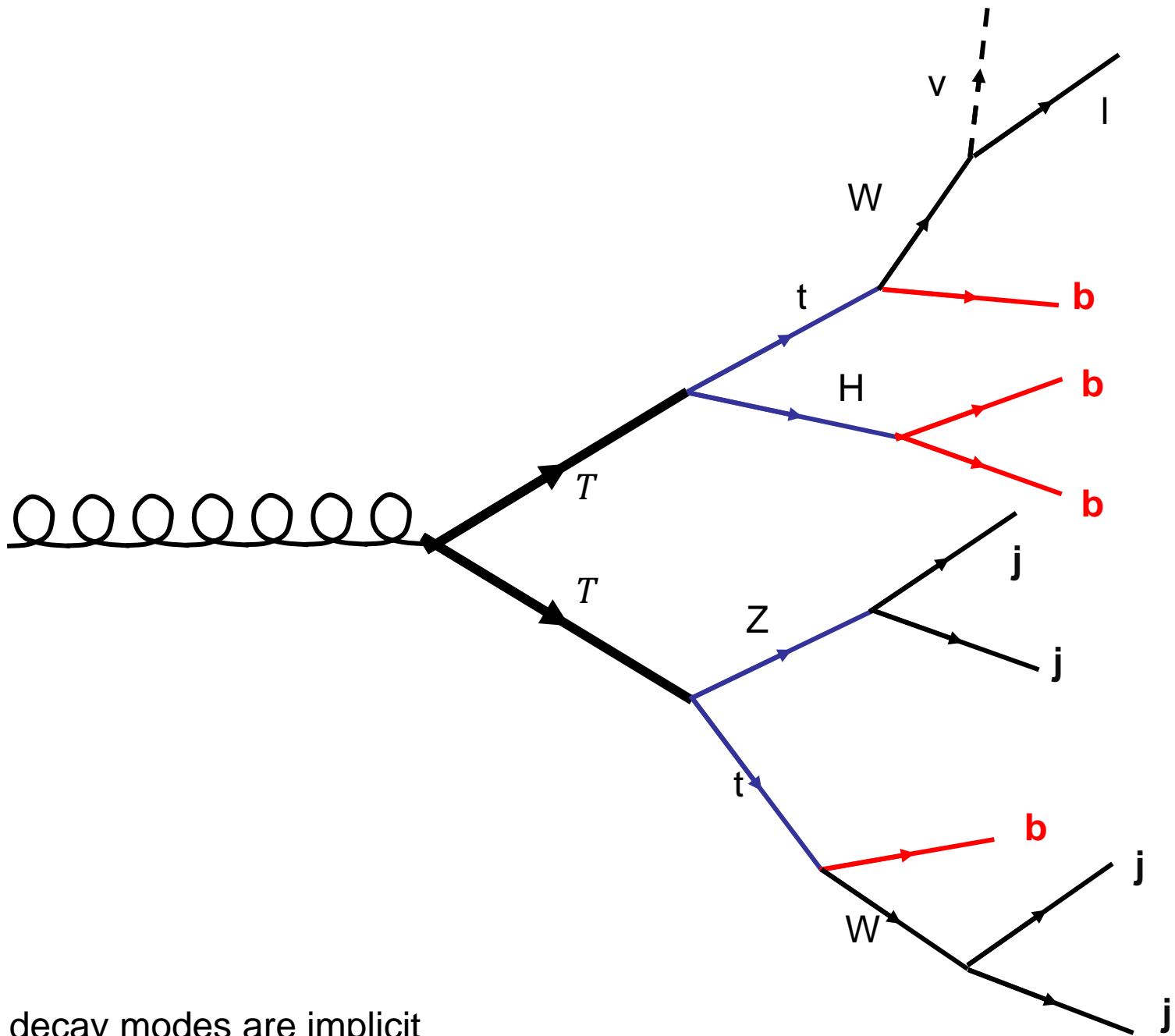


t' → H t

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Complex-conjugate decay modes are implicit



Complex-conjugate decay modes are implicit

Selections

$$E_T^{\text{miss}} > 20 \text{ GeV}$$

$$E_t^{\text{miss}} + m_T > 60 \text{ GeV}$$

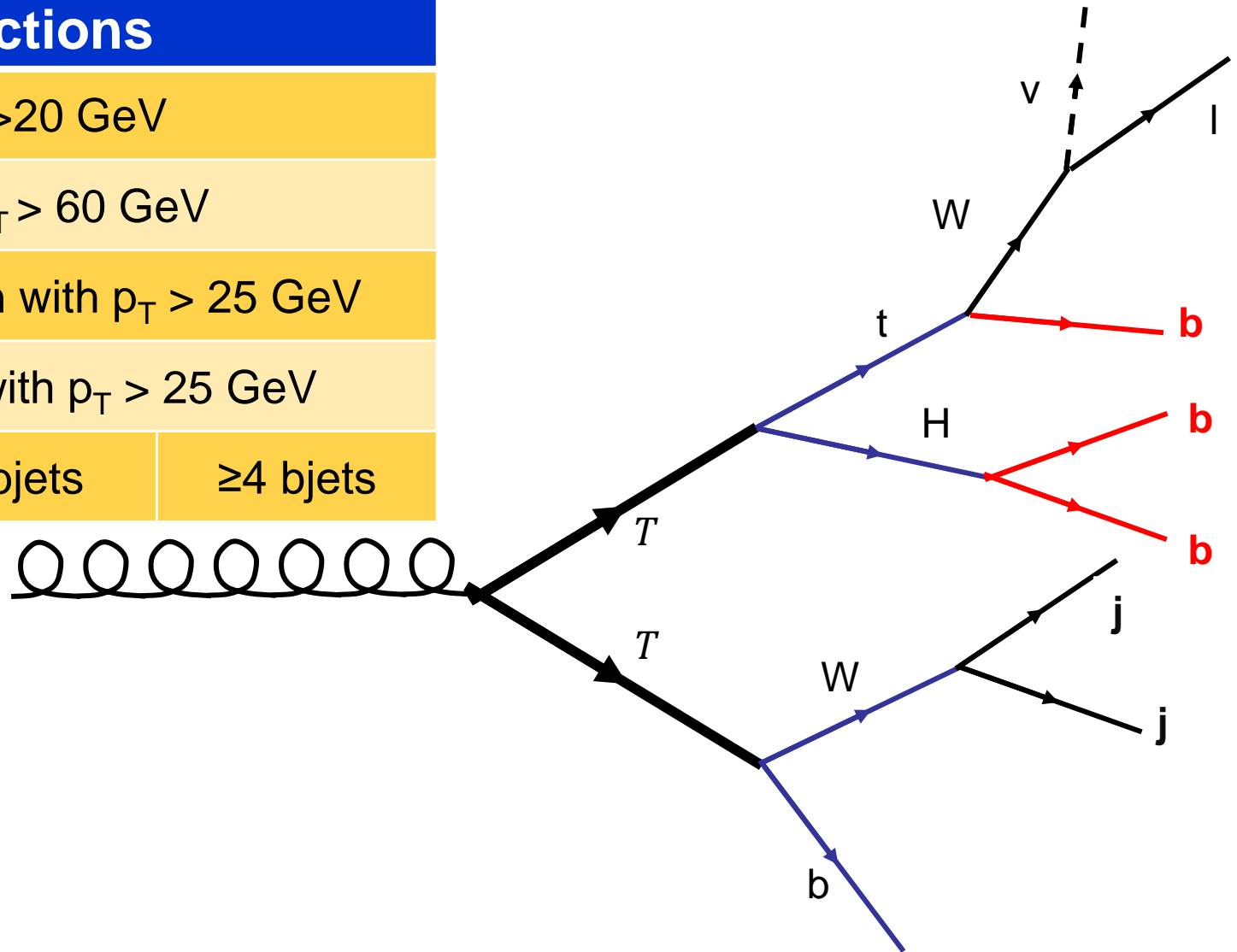
1 Isolated lepton with $p_T > 25 \text{ GeV}$

≥ 6 ak4 jets with $p_T > 25 \text{ GeV}$

2 bjets

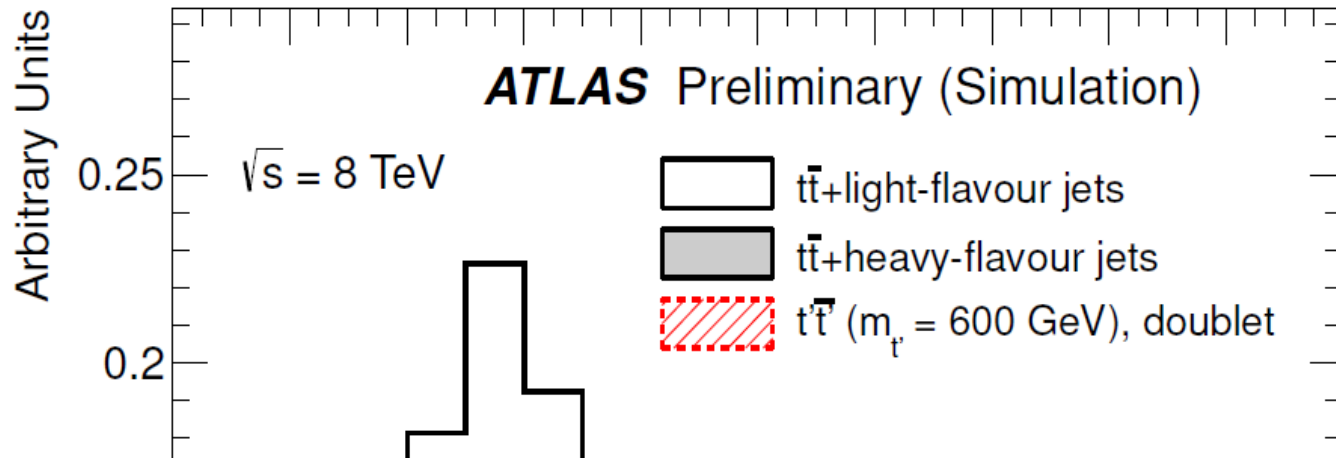
3 bjets

≥ 4 bjets

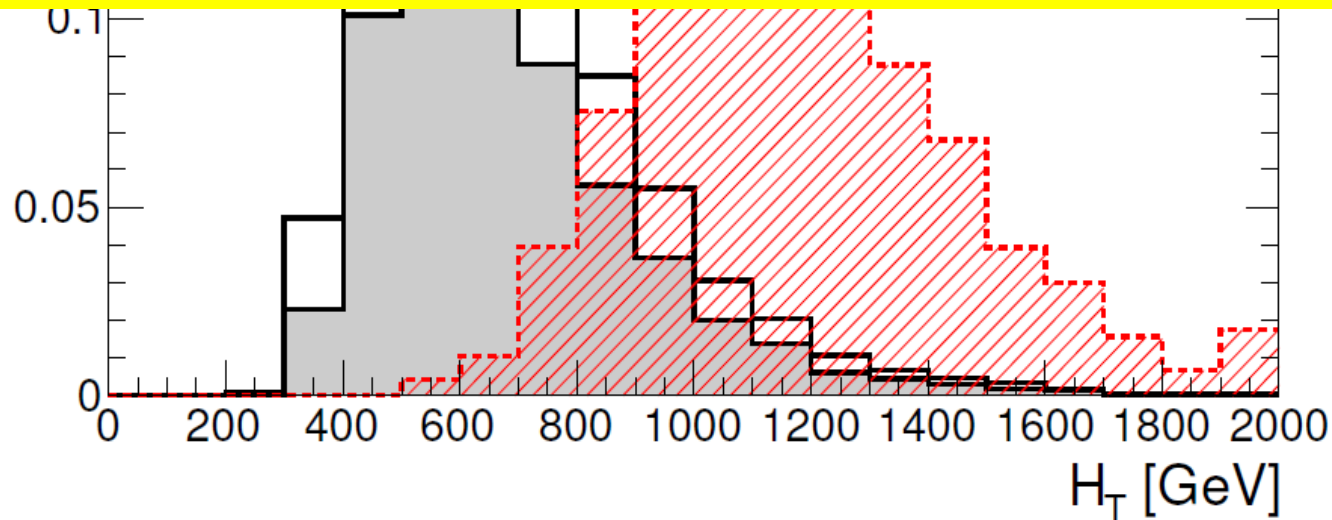


Complex-conjugate decay modes are implicit

Discriminant Variable H_T

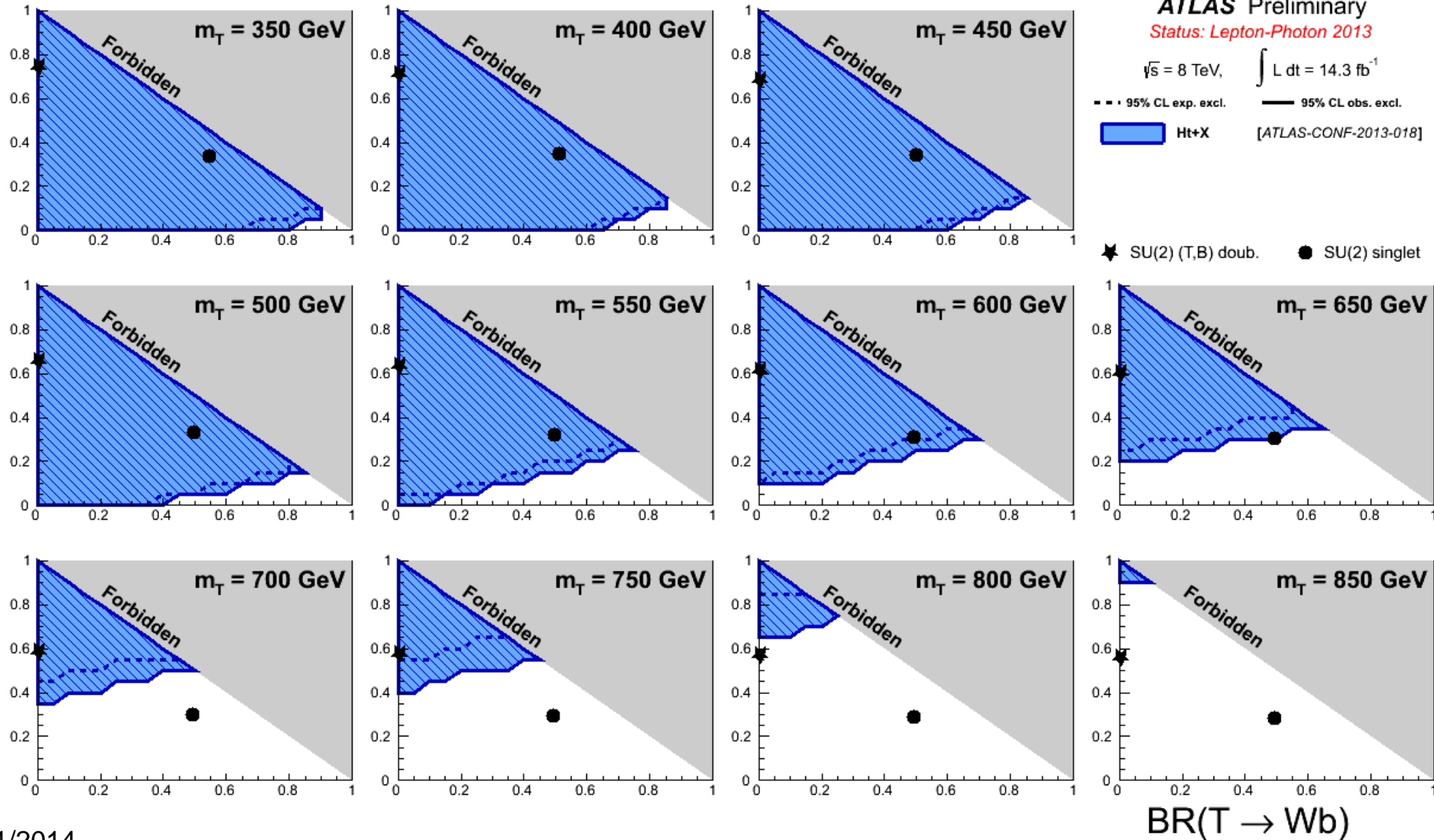


$$H_T = \sum_{\text{Scalar Sum}} P_{T,\text{lepton}} + E_{T,\text{miss}} + P_{T,\text{jets}}$$



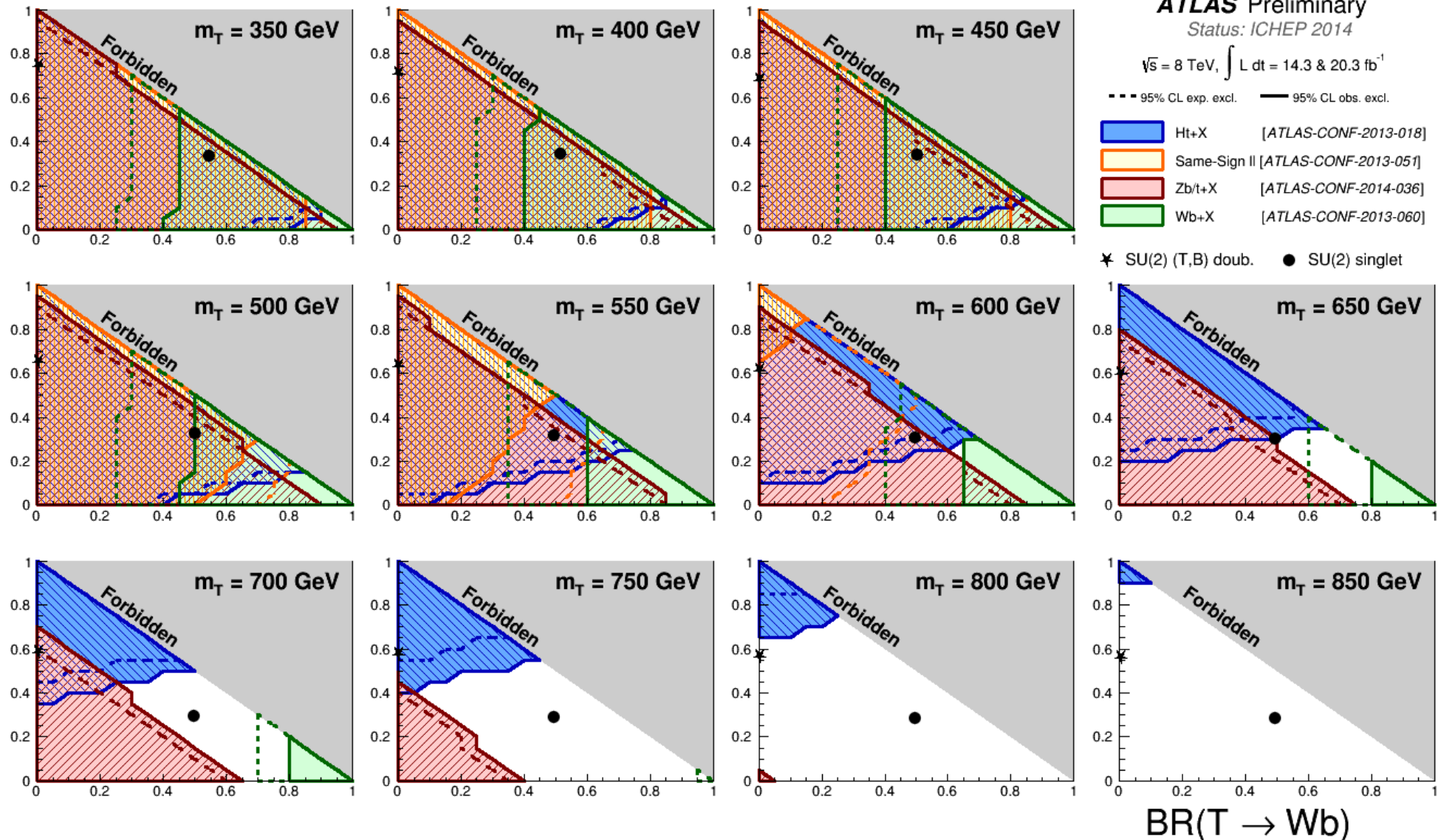
Exclusion Limits for Vector Like T Quark

BR($T \rightarrow Ht$)



Exclusion Limits for Vector Like T Quark

BR(T → Ht)



Inclusive Same-Sign Dilepton Search

[1210.4538](#)

- Model independent approach
 - Limit presented in terms of fiducial cross-section limit

$$\sigma_{95}^{\text{fid}} = \frac{N_{95}}{\epsilon_{\text{fid}} \times \int \mathcal{L} dt}$$

95% CL upper limit on yield
(given N_{obs} and N_{bkg})

Reconstruction and Selection efficiency
Within acceptance

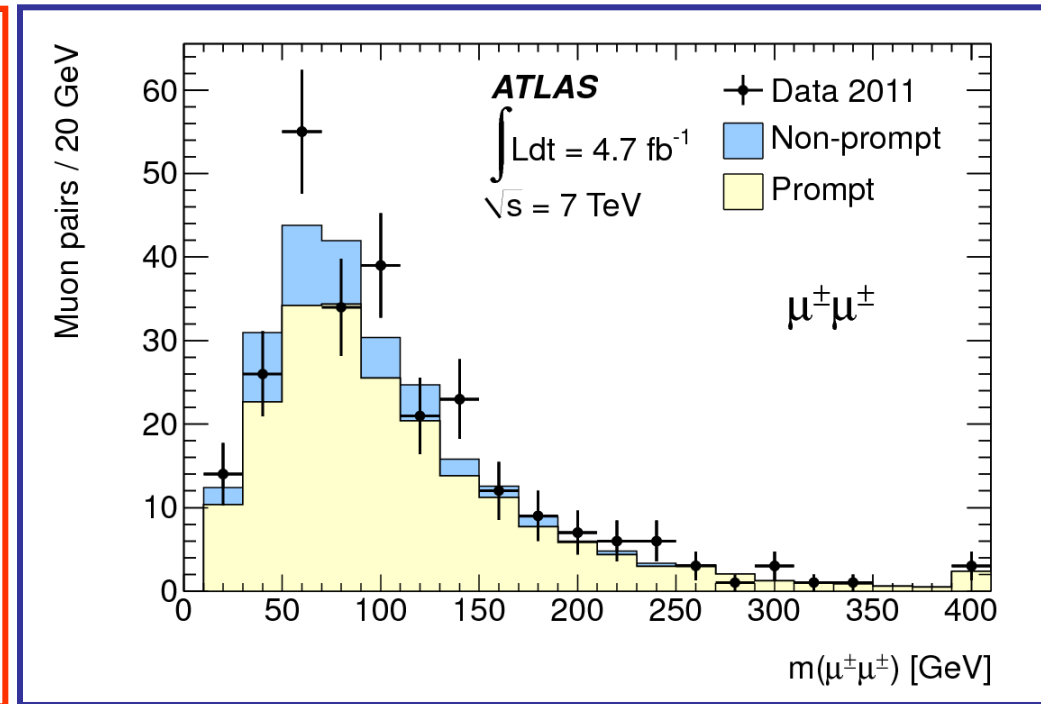
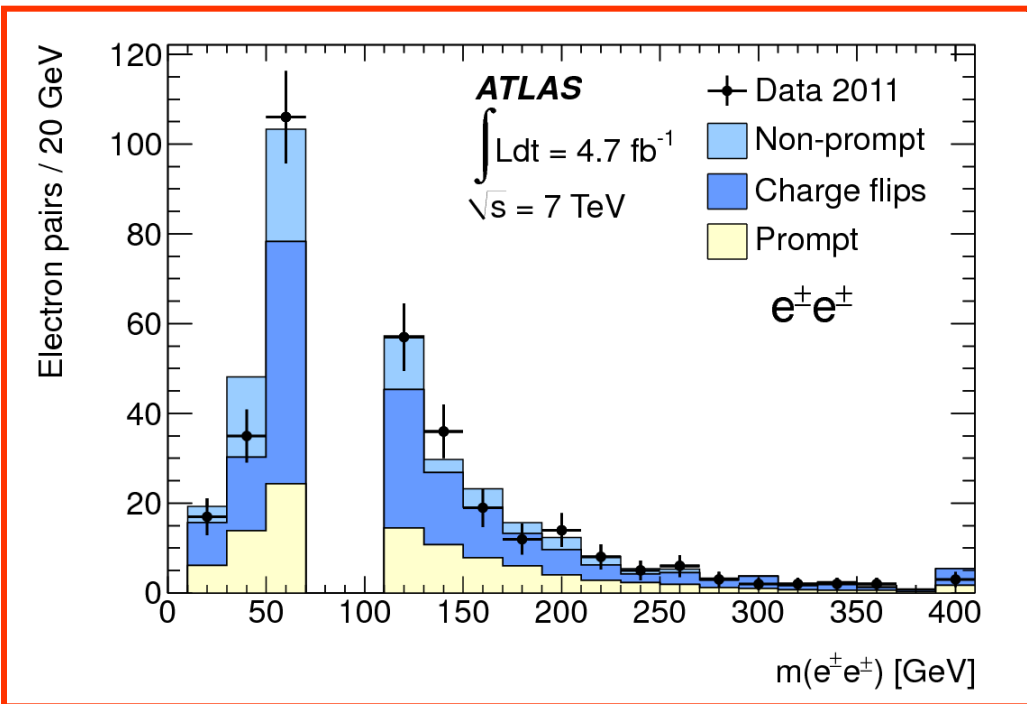
- σ^{fid} is (almost) model-independent
- Can turn σ^{fid} into σ^{total} with generator-level information only
- Caveat: not exactly model-independent → must be conservative

	Electron requirement	Muon requirement
Leading lepton p_T	$p_T > 25 \text{ GeV}$	$p_T > 20 \text{ GeV}$
Sub-leading lepton p_T	$p_T > 20 \text{ GeV}$	$p_T > 20 \text{ GeV}$
Lepton η	$ \eta < 1.37$ or $1.52 < \eta < 2.47$	$ \eta < 2.5$
Isolation	$p_T^{\text{cone}0.3}/p_T < 0.1$	$p_T^{\text{cone}0.4}/p_T < 0.06$ and $p_T^{\text{cone}0.4} < 4 \text{ GeV} + 0.02 \times p_T$

Particle-level definition
of acceptance

Inclusive Same-Sign Dilepton Search

[1210.4538](#)



Inclusive Same-Sign Dilepton Search

[1210.4538](#)

■ 95% upper limits

■ 1.7 fb and 64 fb

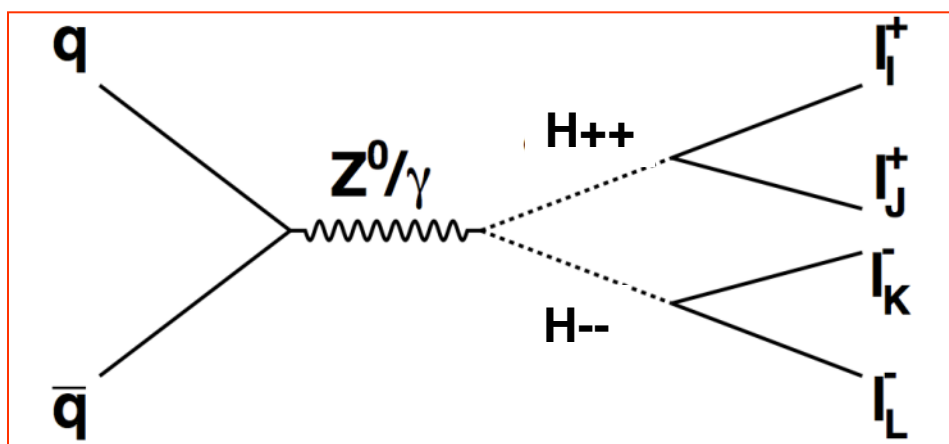
Fiducial cross section
upper limits

	e^-e^-	
$m > 15 \text{ GeV}$	$23.2^{+8.6}_{-5.8}$	25.7
$m > 100 \text{ GeV}$	$12.0^{+5.3}_{-2.8}$	18.7
$m > 200 \text{ GeV}$	$4.9^{+1.9}_{-1.2}$	4.0
$m > 300 \text{ GeV}$	$2.9^{+1.0}_{-0.6}$	2.7
$m > 400 \text{ GeV}$	$1.8^{+0.8}_{-0.4}$	2.3

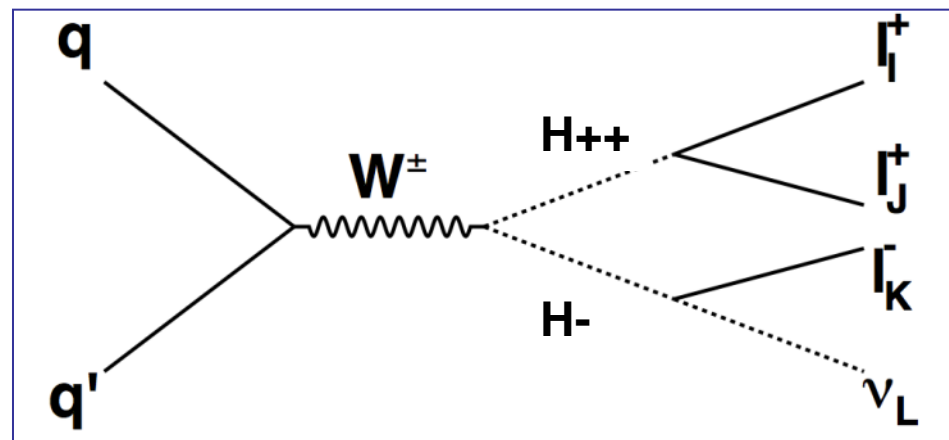
Mass	ee		eμ		μμ	
	exp	obs	exp	obs	exp	obs
	95% C.L. upper limit [fb]					
Mass range	expected observed $e^\pm e^\pm$		expected observed $e^\pm \mu^\pm$		expected observed $\mu^\pm \mu^\pm$	
$m > 15 \text{ GeV}$	46^{+15}_{-12}	42	56^{+23}_{-15}	64	$24.0^{+8.9}_{-6.0}$	29.8
$m > 100 \text{ GeV}$	$24.1^{+8.9}_{-6.2}$	23.4	$23.0^{+9.1}_{-6.7}$	31.2	$12.2^{+4.5}_{-3.0}$	15.0
$m > 200 \text{ GeV}$	$8.8^{+3.4}_{-2.1}$	7.5	$8.4^{+3.4}_{-1.7}$	9.8	$4.3^{+1.8}_{-1.1}$	6.7
$m > 300 \text{ GeV}$	$4.5^{+1.8}_{-1.3}$	3.9	$4.1^{+1.8}_{-0.9}$	4.6	$2.4^{+0.9}_{-0.7}$	2.6
$m > 400 \text{ GeV}$	$2.9^{+1.1}_{-0.8}$	2.4	$3.0^{+1.0}_{-0.8}$	3.1	$1.7^{+0.6}_{-0.5}$	1.7
	e^+e^+		$e^+\mu^+$		$\mu^+\mu^+$	
$m > 15 \text{ GeV}$	$29.1^{+10.2}_{-8.6}$	22.8	$34.9^{+12.2}_{-8.6}$	34.1	$15.0^{+6.1}_{-3.3}$	15.2
$m > 100 \text{ GeV}$	$16.1^{+5.9}_{-4.3}$	12.0	$15.4^{+5.9}_{-4.1}$	18.0	$8.4^{+3.2}_{-2.4}$	7.9
$m > 200 \text{ GeV}$	$7.0^{+2.9}_{-2.2}$	6.1	$6.6^{+3.5}_{-1.8}$	8.8	$3.5^{+1.6}_{-0.7}$	4.3
$m > 300 \text{ GeV}$	$3.7^{+1.4}_{-1.0}$	2.9	$3.2^{+1.2}_{-0.9}$	3.2	$2.0^{+0.8}_{-0.5}$	2.1
$m > 400 \text{ GeV}$	$2.3^{+1.1}_{-0.6}$	1.7	$2.4^{+0.9}_{-0.6}$	2.5	$1.5^{+0.6}_{-0.3}$	1.8
	e^-e^-		$e^-\mu^-$		$\mu^-\mu^-$	
$m > 15 \text{ GeV}$	$23.2^{+8.6}_{-5.8}$	25.7	$26.2^{+10.6}_{-7.6}$	34.4	$12.1^{+4.5}_{-3.5}$	18.5
$m > 100 \text{ GeV}$	$12.0^{+5.3}_{-2.8}$	18.7	$11.5^{+4.2}_{-3.5}$	16.9	$6.0^{+2.3}_{-1.9}$	10.1
$m > 200 \text{ GeV}$	$4.9^{+1.9}_{-1.2}$	4.0	$4.6^{+2.1}_{-1.2}$	4.5	$2.7^{+1.1}_{-0.7}$	4.4
$m > 300 \text{ GeV}$	$2.9^{+1.0}_{-0.6}$	2.7	$2.7^{+1.1}_{-0.6}$	3.5	$1.5^{+0.8}_{-0.3}$	1.7
$m > 400 \text{ GeV}$	$1.8^{+0.8}_{-0.4}$	2.3	$2.3^{+0.8}_{-0.5}$	2.5	$1.2^{+0.4}_{-0.0}$	1.2

Inclusive Same-Sign Dilepton Search: $H^{++/--}$ Limits

- Models explaining non-zero neutrino masses predict $H^{++/--}$
 - e.g. minimal type II seesaw model
 - additional scalar field
 - triplet (under $SU(2)_L$ with $Y=2$): $H^{++/--}, H^{+/-}, H^0$



pair production



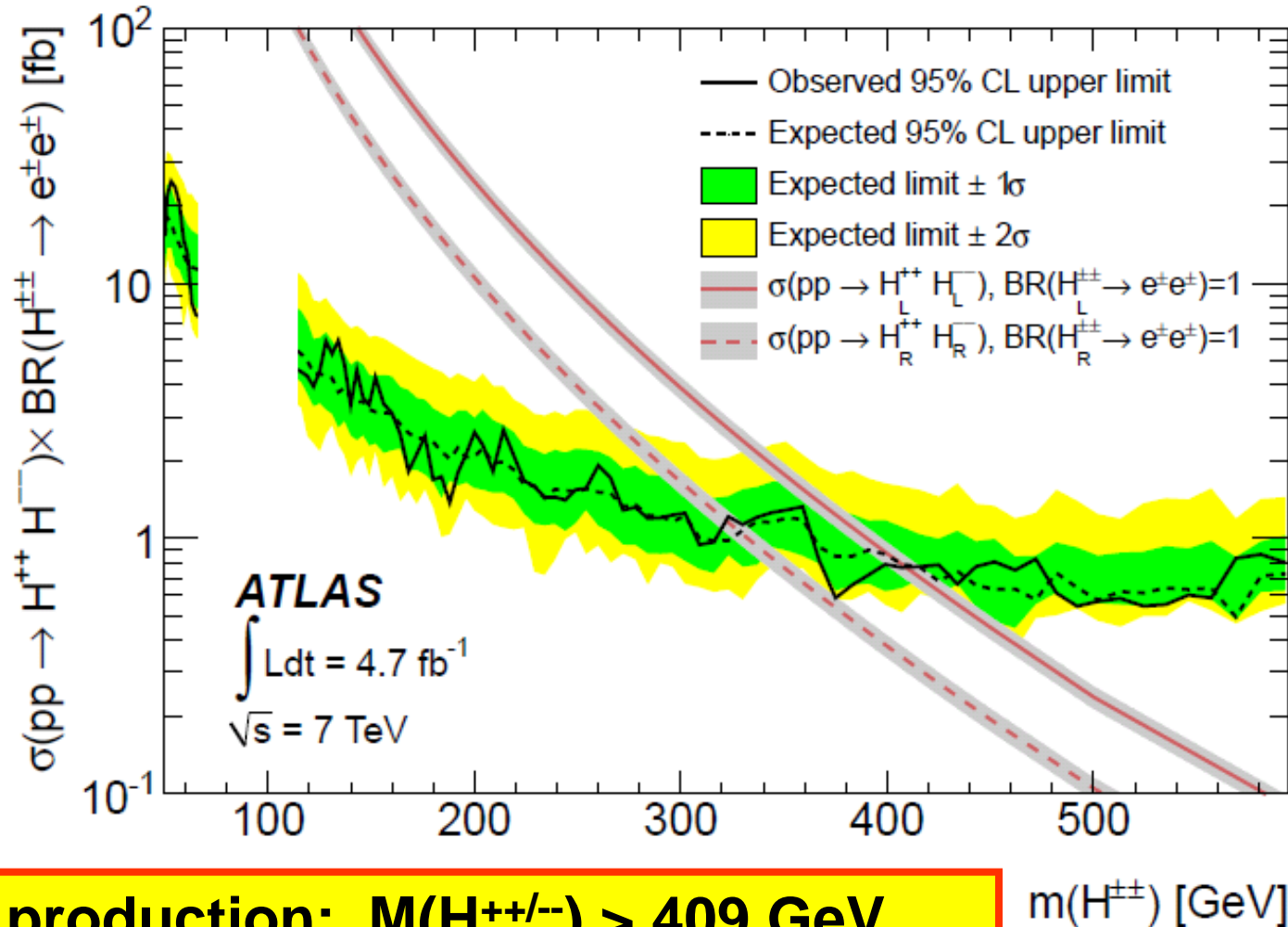
associate production

Signature: same-sign leptons

Doubly Charged Higgs Limits

[arXiv:1210.5070](https://arxiv.org/abs/1210.5070)

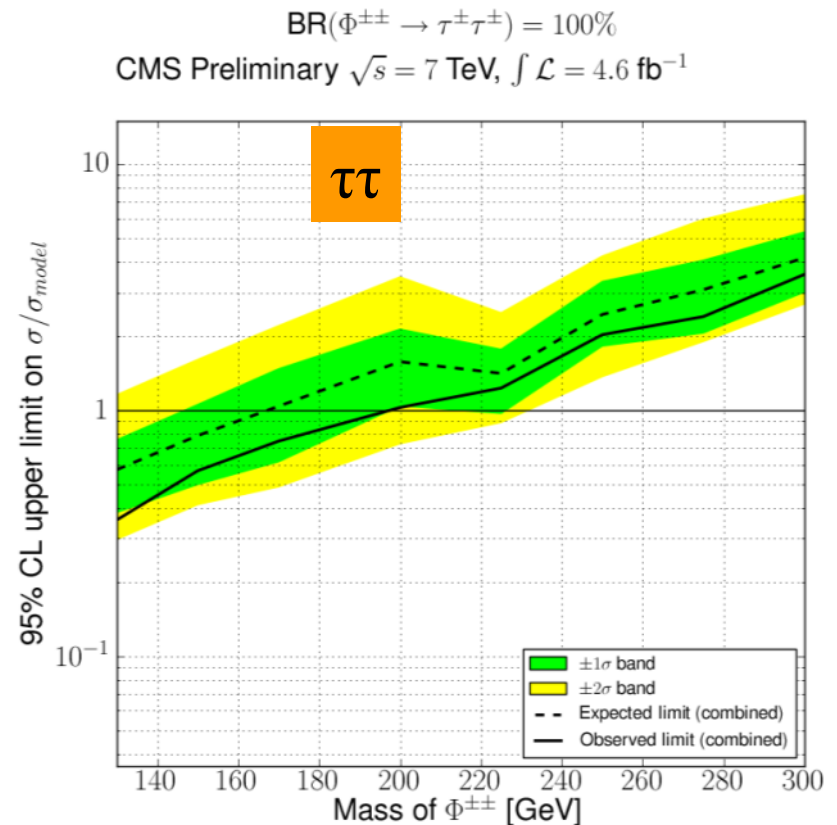
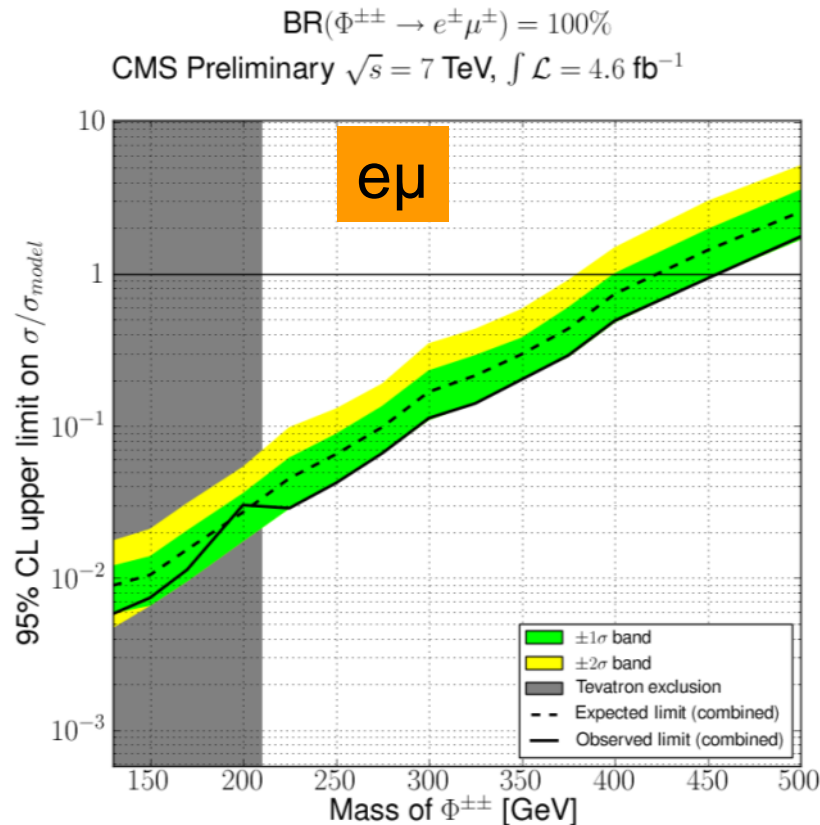
- Used e.g. limits on doubly charged Higgs



Doubly Charged Higgs Limits

- Example of more optimized search
- Includes also τ -channel and associate production.

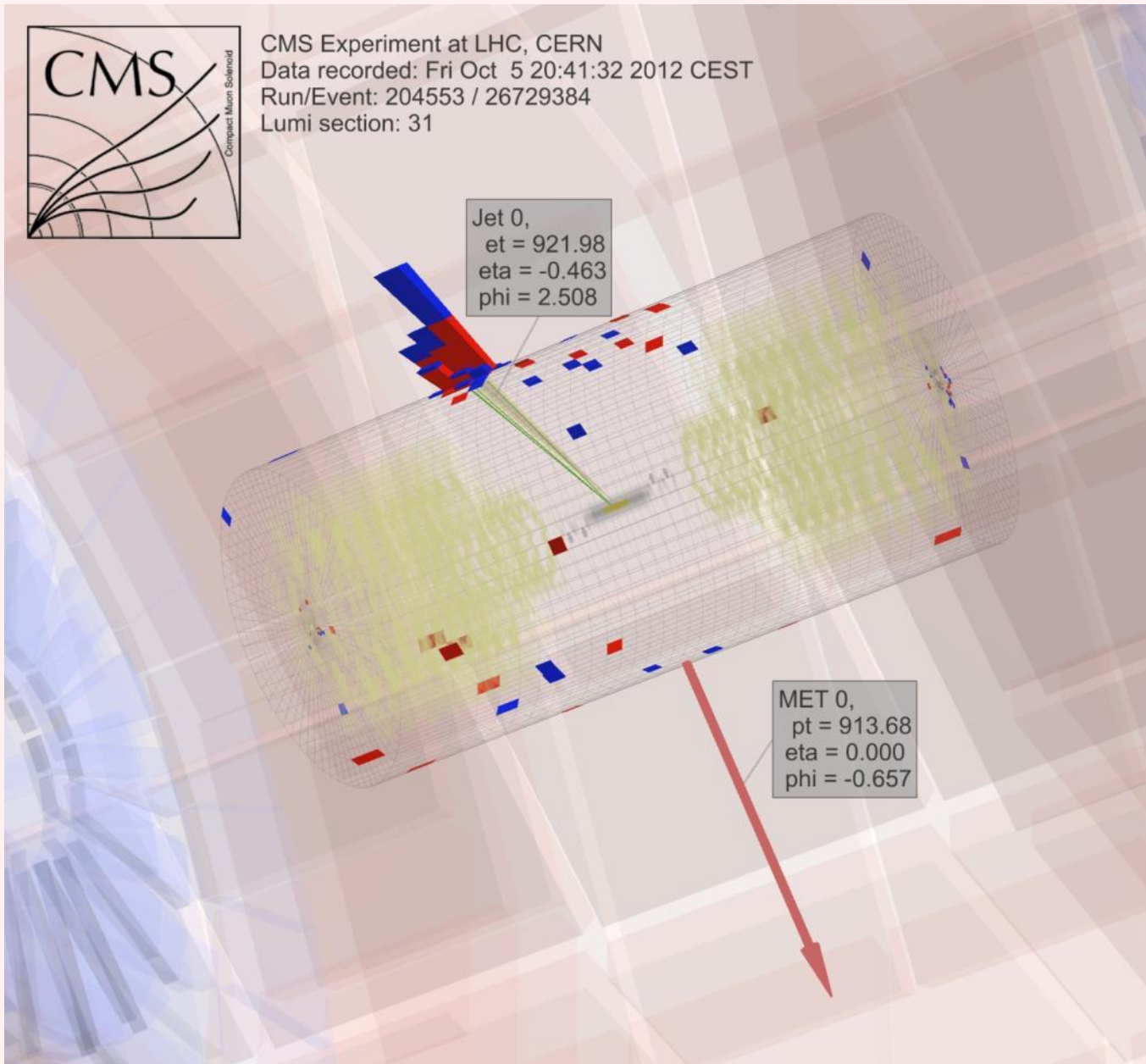
arXiv:1207.2666



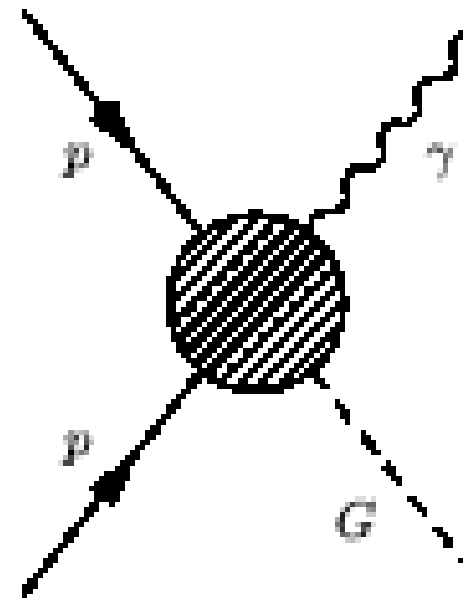
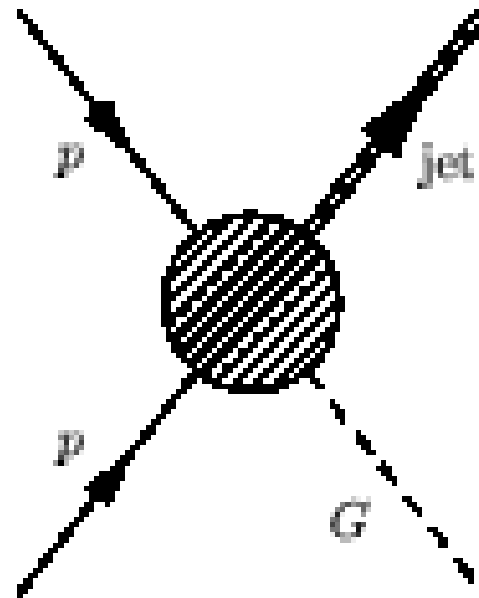
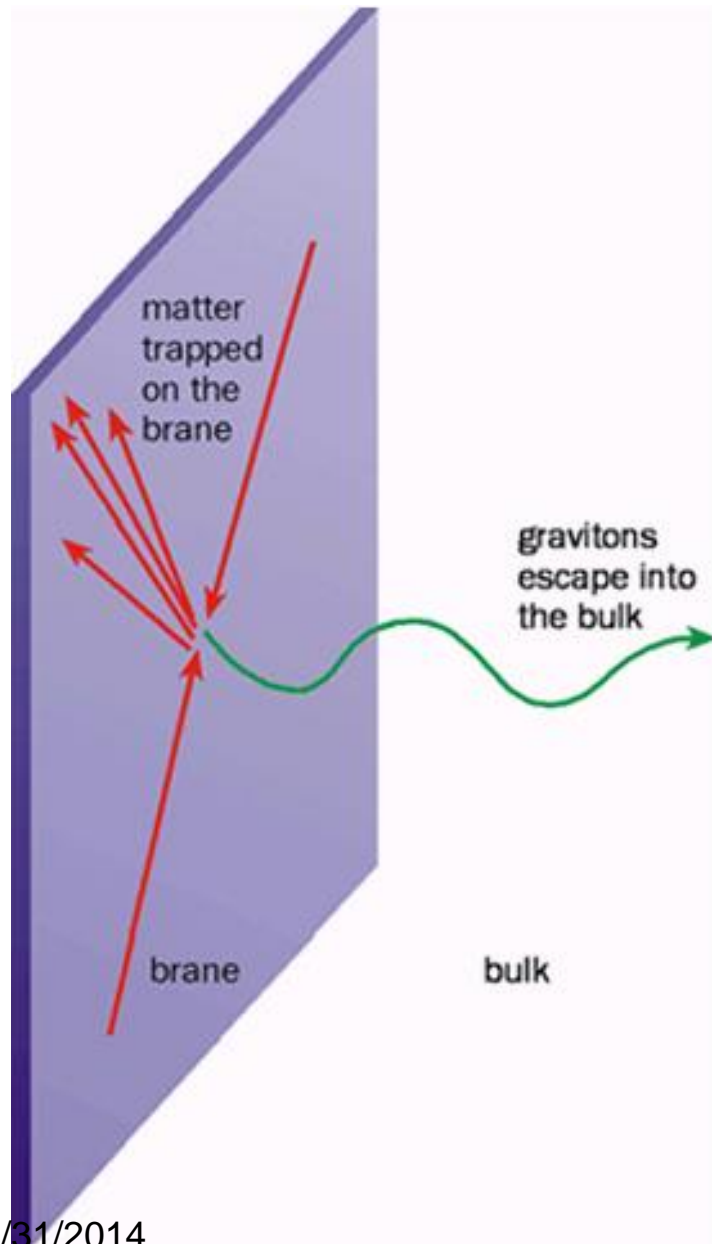
Combined $e\mu$: $M(H^{++/-}) > 455 \text{ GeV}$

Combined $\tau\tau$: $M(H^{++/-}) > 198 \text{ GeV}$

Mono Jet Event Display

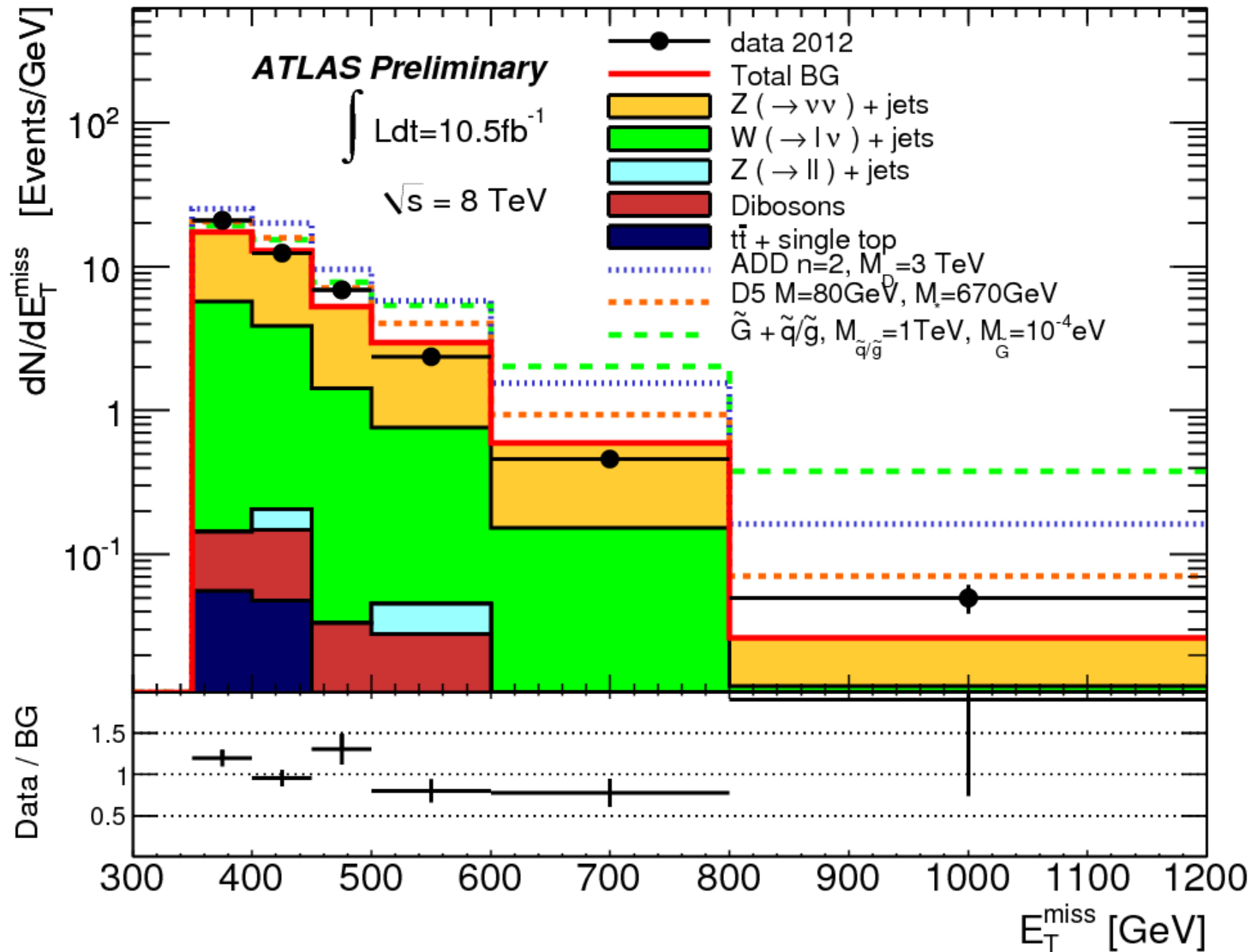


Graviton Production in Extra Dimensions

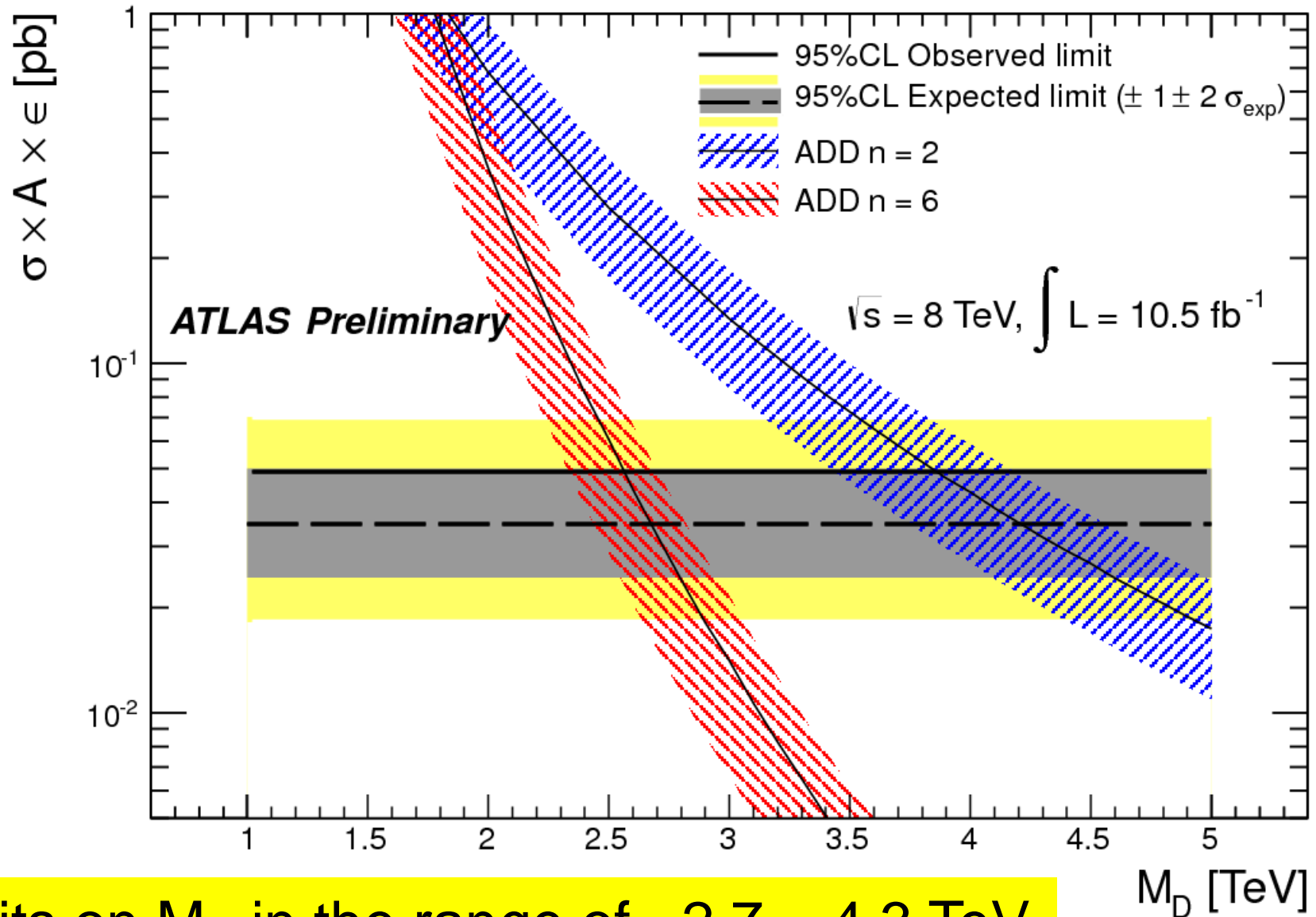


ME_T Distribution of Mono Jet Analysis

ATLAS-CONF-2012-147



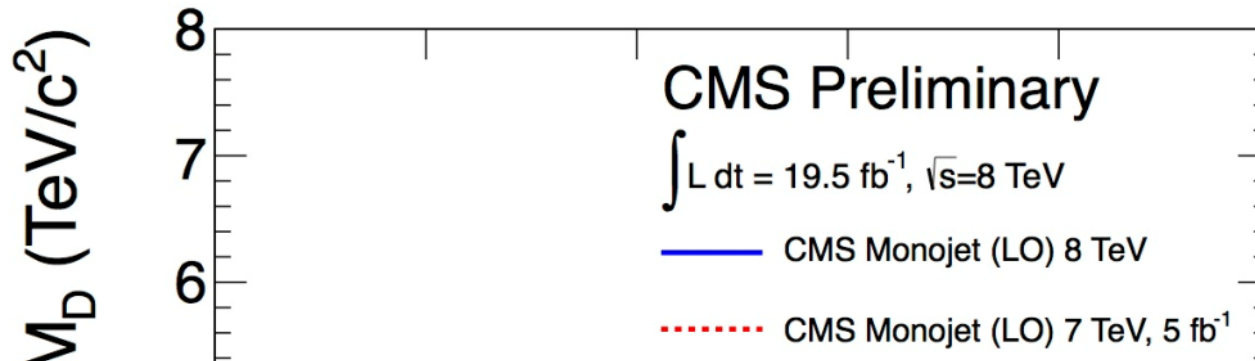
Exclusion Limits



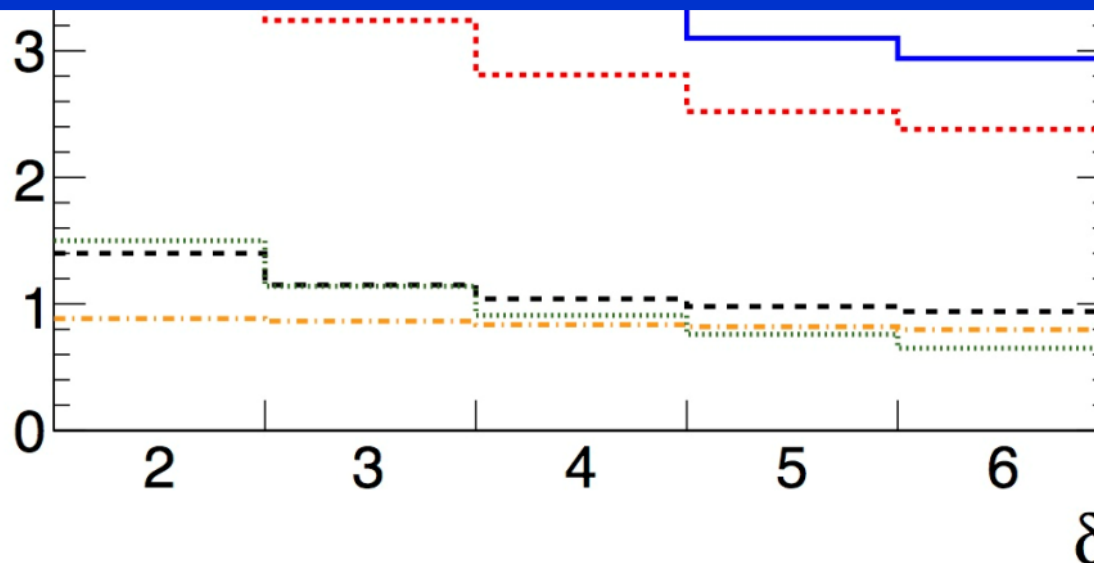
Limits on M_D in the range of $\sim 2.7 - 4.3 \text{ TeV}$

Exclusion Limits on M_D from CMS

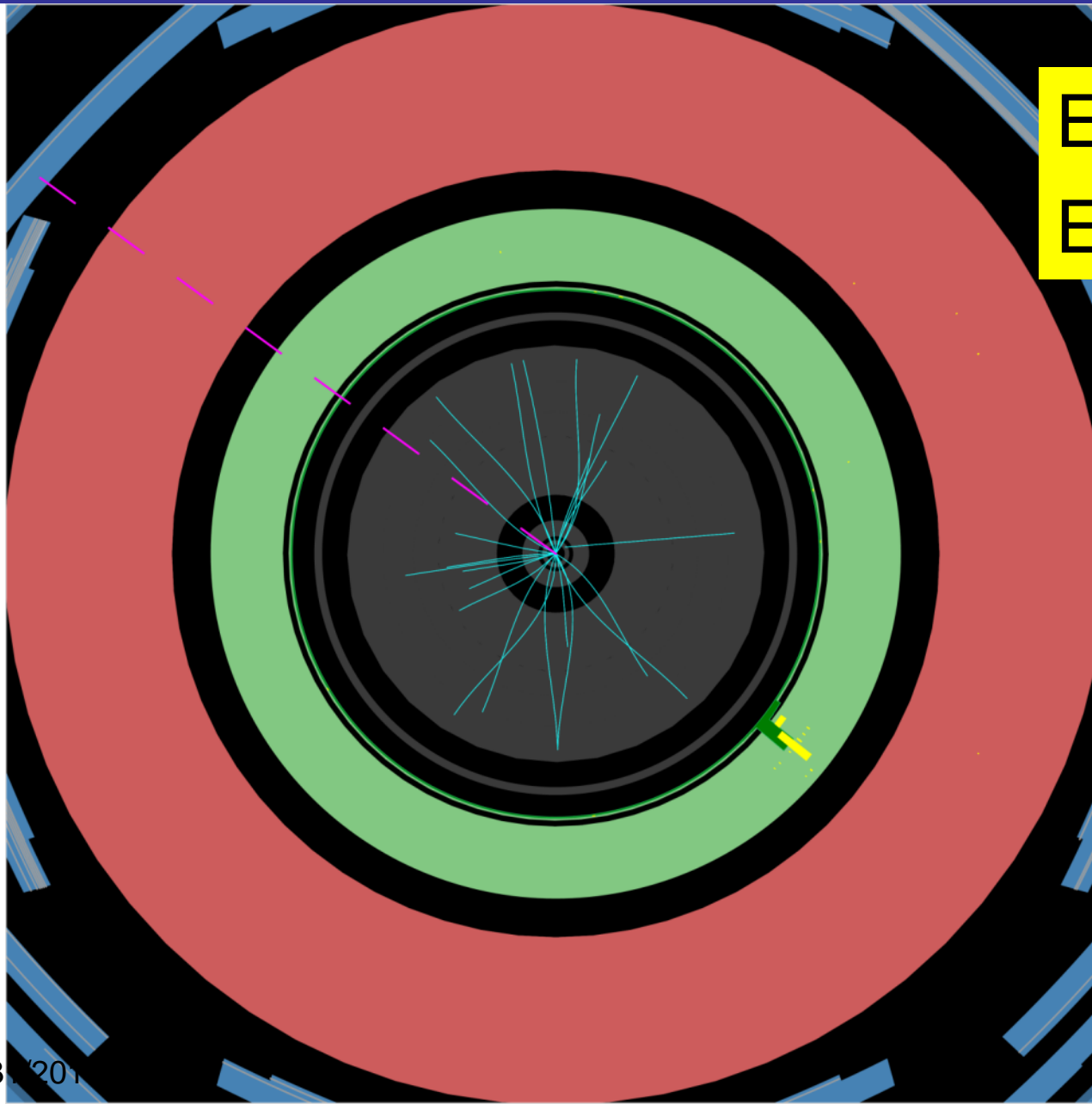
EXO-12-048 PAS



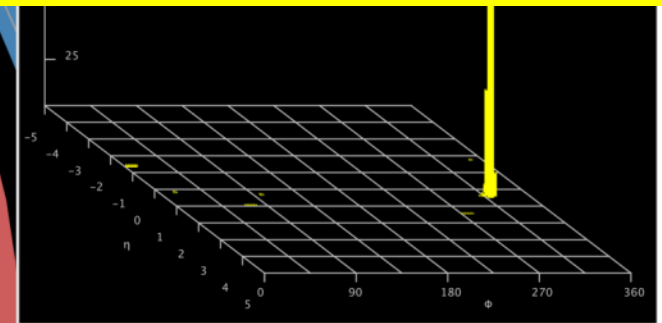
Semi-classical regime out of reach of the LHC
LHC operates in Quantum Gravitational regime



Mono Photon Searches for Extra Dimensions



$E_{\text{miss}} = 218.3 \text{ GeV}$
 $E_{\text{photon}} = 218 \text{ GeV}$



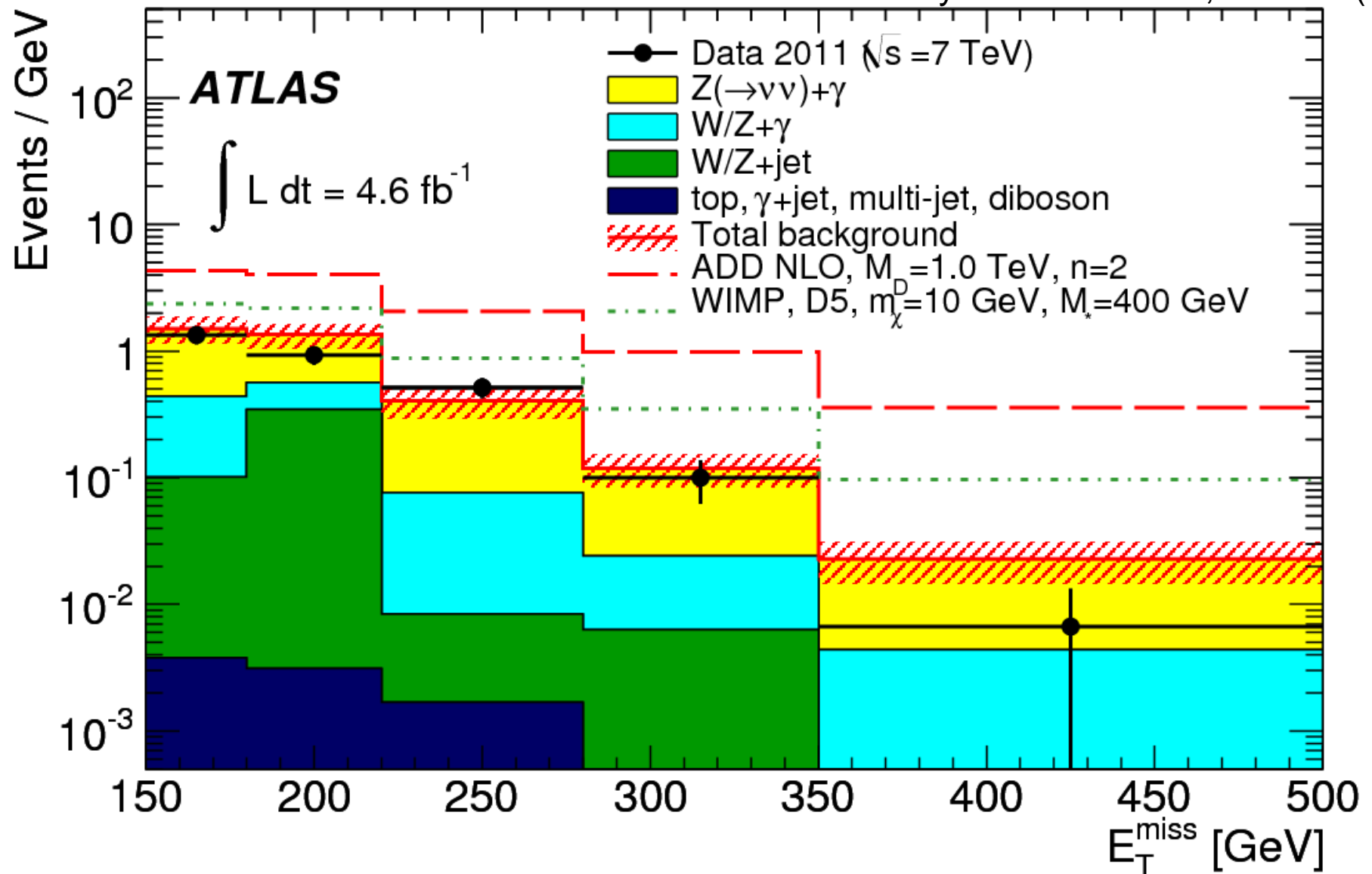
ATLAS
EXPERIMENT

Run Number: 179710, Event Number: 19174449

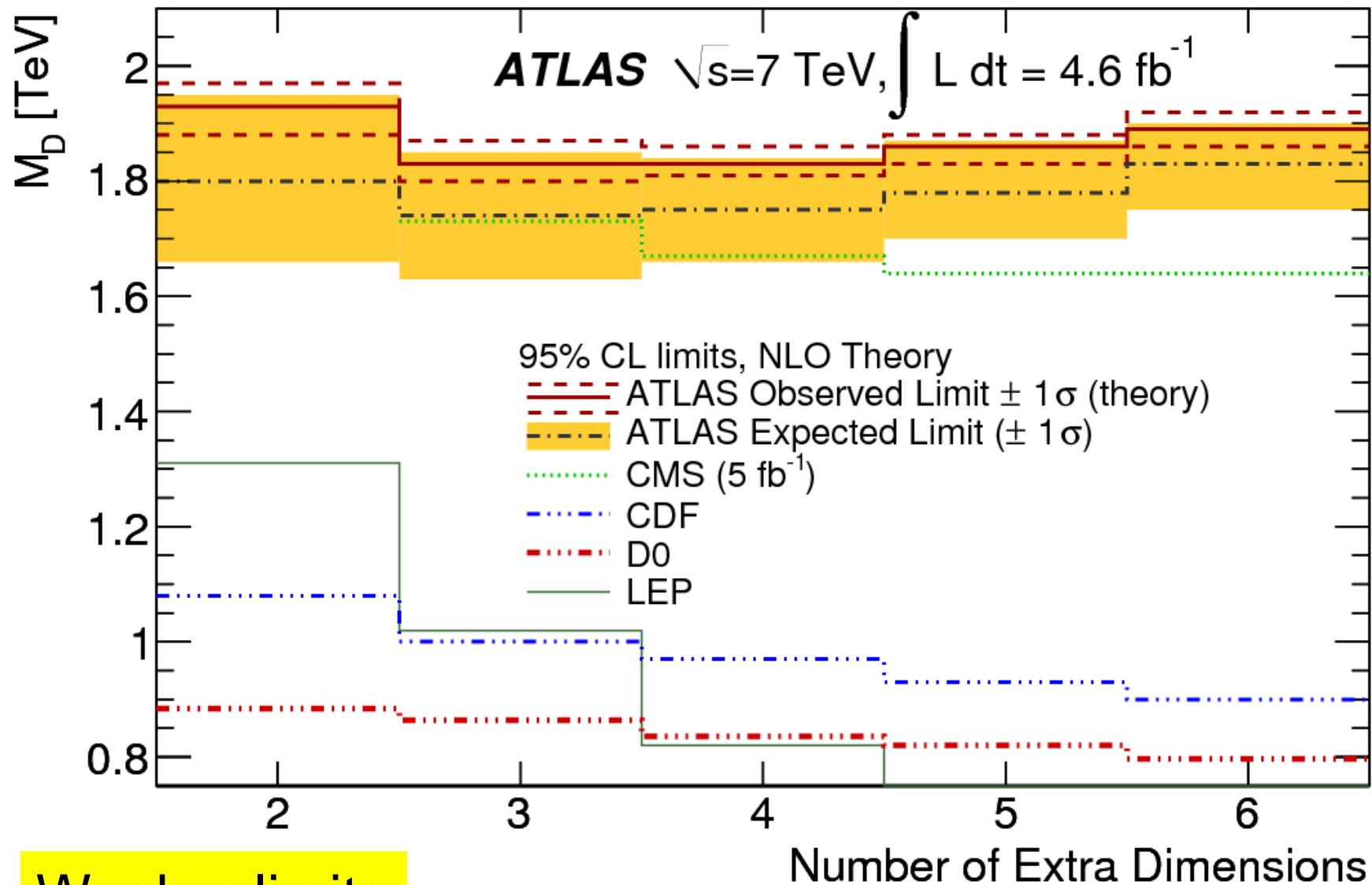
Date: 2011-04-15 03:48:32 CEST

The Discriminant

Phys. Rev. Lett 110, 011802 (2013)



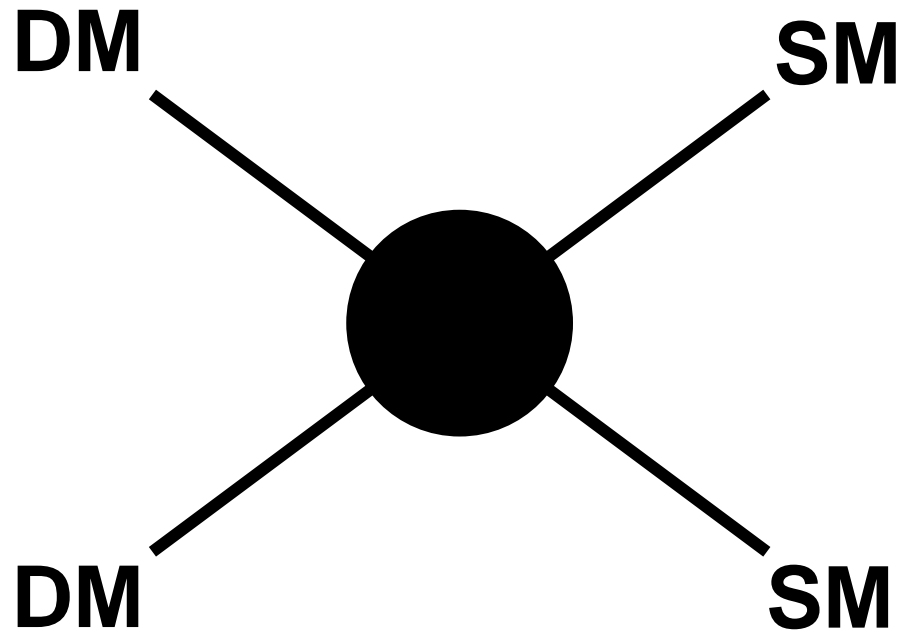
Limits on M_D in Mono Photon Search



Weaker limits

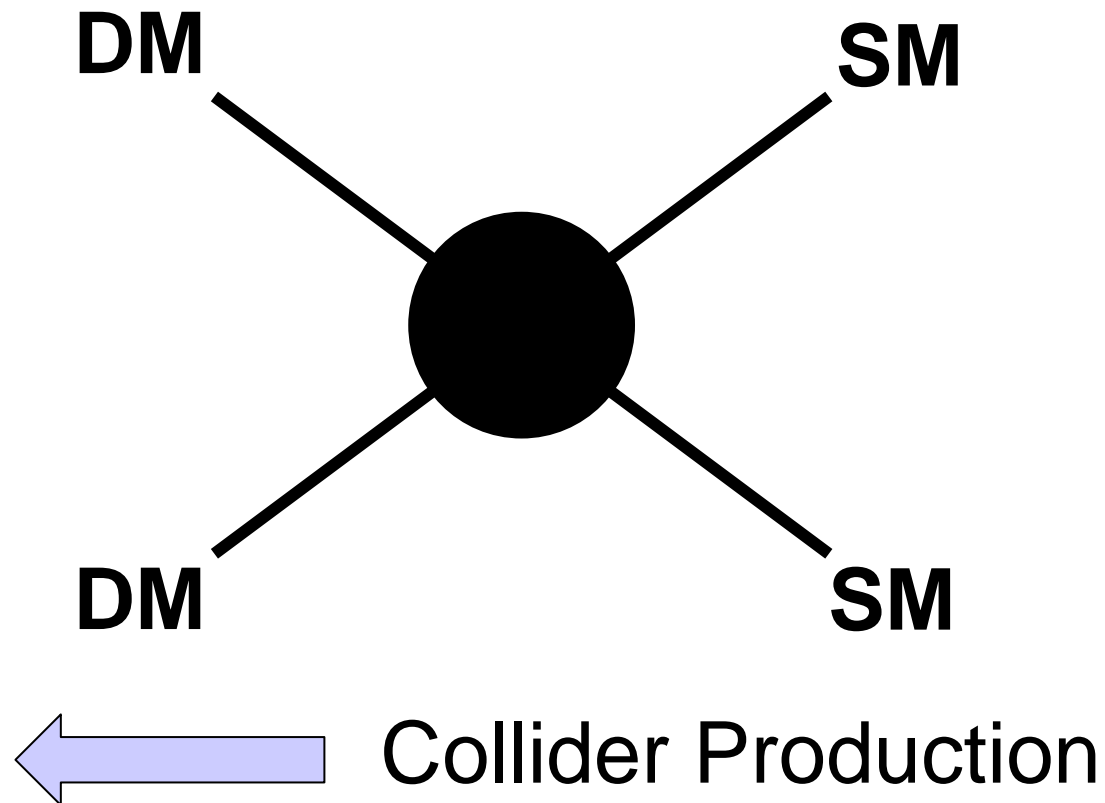
Dark Matter Detection

M. Fedderke



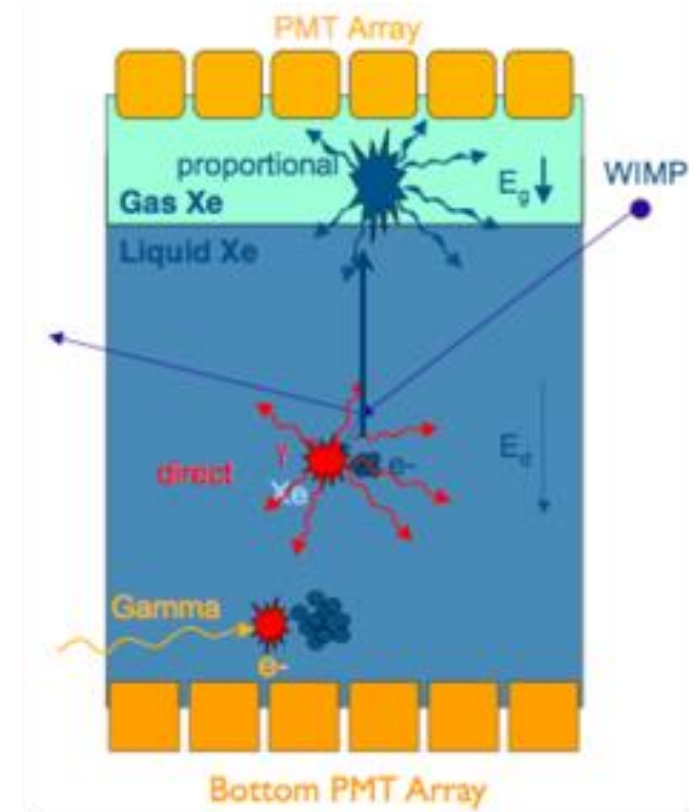
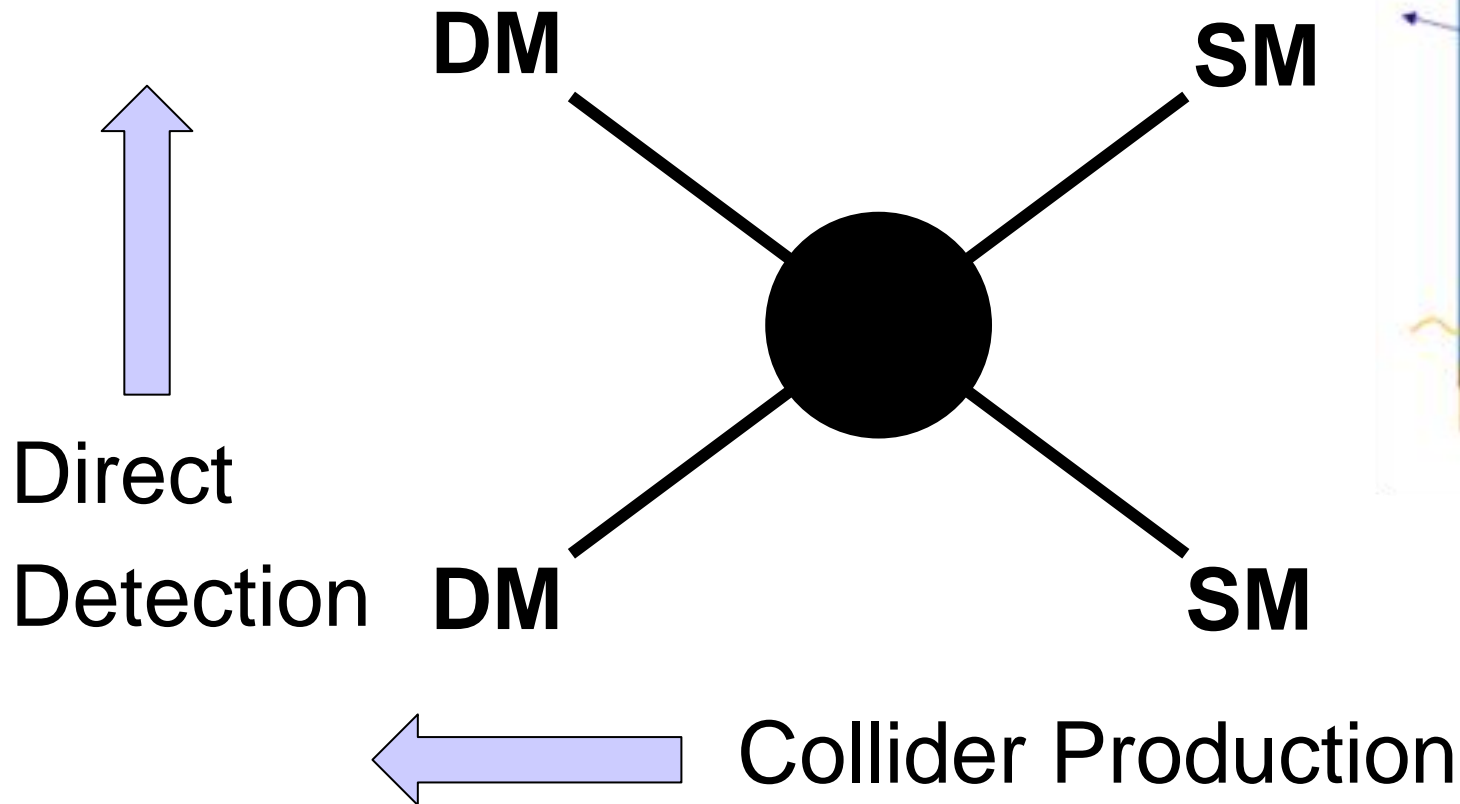
Dark Matter Detection

M. Fedderke



Dark Matter Detection

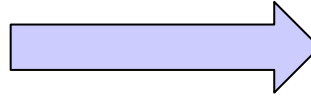
M. Fedderke



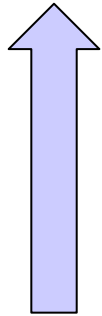
Dark Matter Detection

M. Fedderke

Indirect Detection

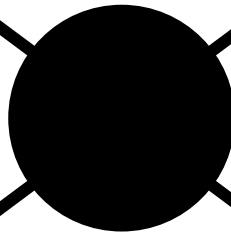


Direct
Detection



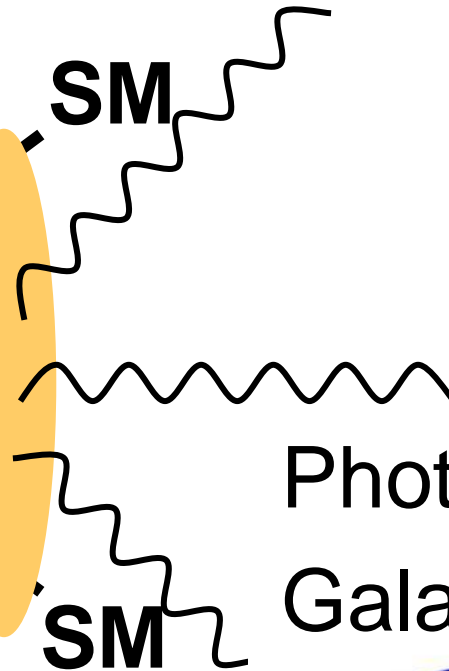
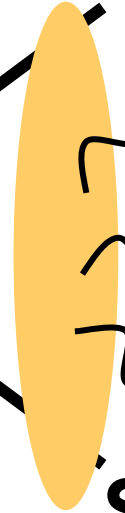
DM

DM

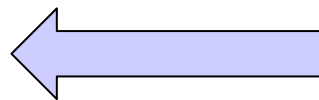


SM

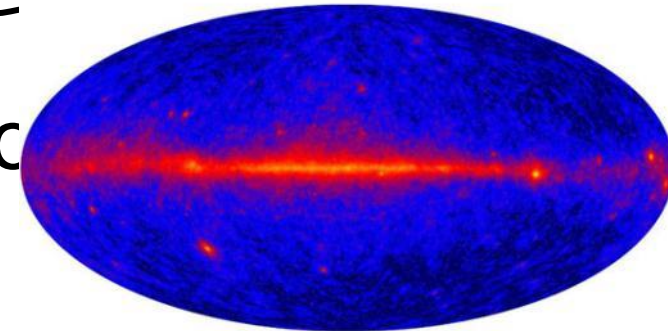
SM



Photons from
Galactic Centre



Collider Production

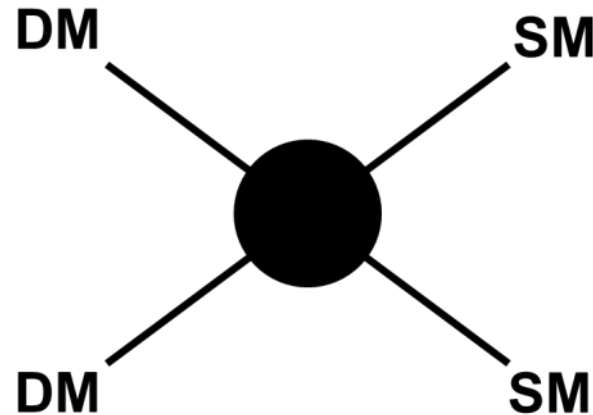


DM Interpretations of Mono-Object Analyses

Idea: Effective Theory

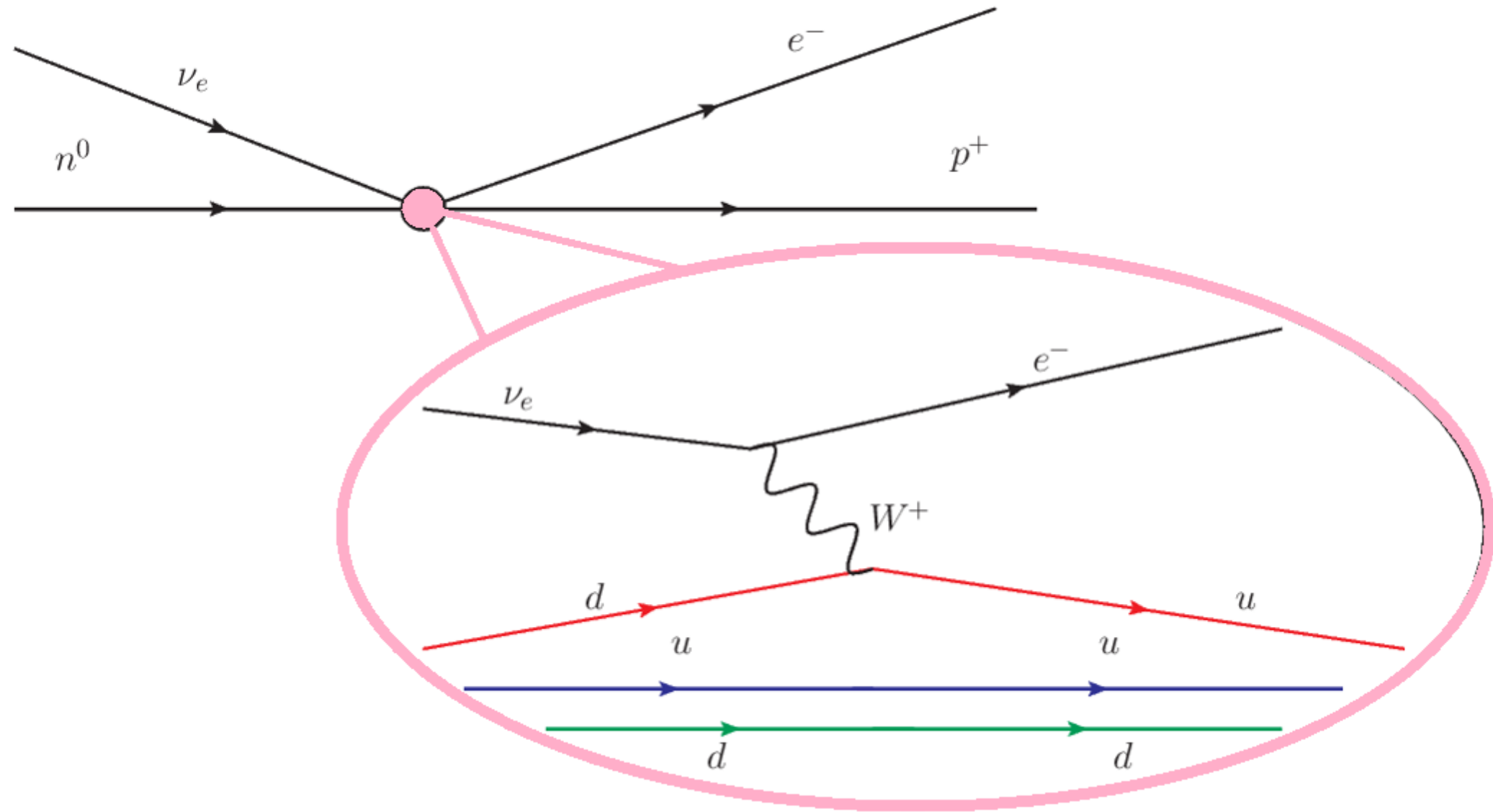
Johanna Gramling

- Heavy particle mediating interaction btw DM and SM



- too heavy to be on-shell \rightarrow can be integrated out
- interaction treated as contact interaction!

Like Fermi's Theory of Beta Decay



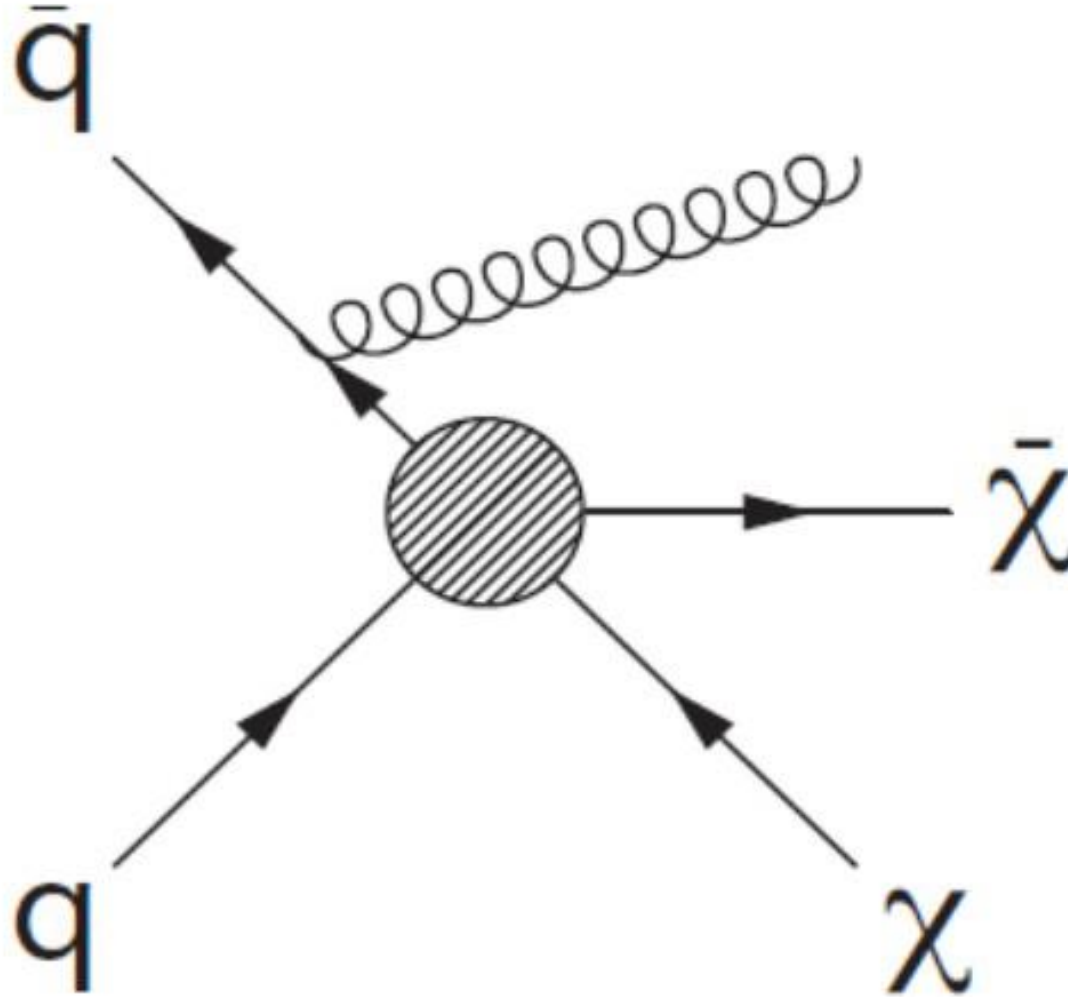
Advantage of Effective Theory

[arXiv:1008.1783](https://arxiv.org/abs/1008.1783)

- Model depends only on a few parameters
 - dark matter mass, m_χ
 - cut-off scale Λ or M_*
 - much easier than e.g. a full SUSY model
- Allows easy comparison to direct or indirect DM detection experiments
- DM
 - Fermion: Dirac or Majorana
 - Scalar: Complex or Real

$$\Lambda = \frac{m_M}{\sqrt{g_q g_\chi}}$$

Dark Matter Production at a Collider



Effective interactions coupling DM to SM quarks or gluons

1210.4491v2

Name	Initial state	Type	Operator
D1	qq	scalar	$\frac{m_q}{M_\star^3} \bar{\chi} \chi \bar{q} q$
D5	qq	vector	$\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \chi \bar{q} \gamma_\mu q$
D8	qq	axial-vector	$\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{q} \gamma_\mu \gamma^5 q$
D9	qq	tensor	$\frac{1}{M_\star^2} \bar{\chi} \sigma^{\mu\nu} \chi \bar{q} \sigma_{\mu\nu} q$
D11	gg	scalar	$\frac{1}{4M_\star^3} \bar{\chi} \chi \alpha_s (G_{\mu\nu}^a)^2$

characteristic set

Conditions of EFT

1. $g_q, g_\chi < 4\pi \rightarrow \frac{m_M}{4\pi} < \Lambda$ (to stay in perturbative regime)

2. $m_M > m_\chi$ (M can not be produced, but χ can)

$$\blacksquare \Lambda > \frac{m_M}{4\pi} > \frac{m_\chi}{4\pi}$$

Johanna Gramling

3. $m_M > Q_{TR}$

$$\blacksquare \Lambda > \frac{m_M}{4\pi} > \frac{Q_{TR}}{4\pi}$$

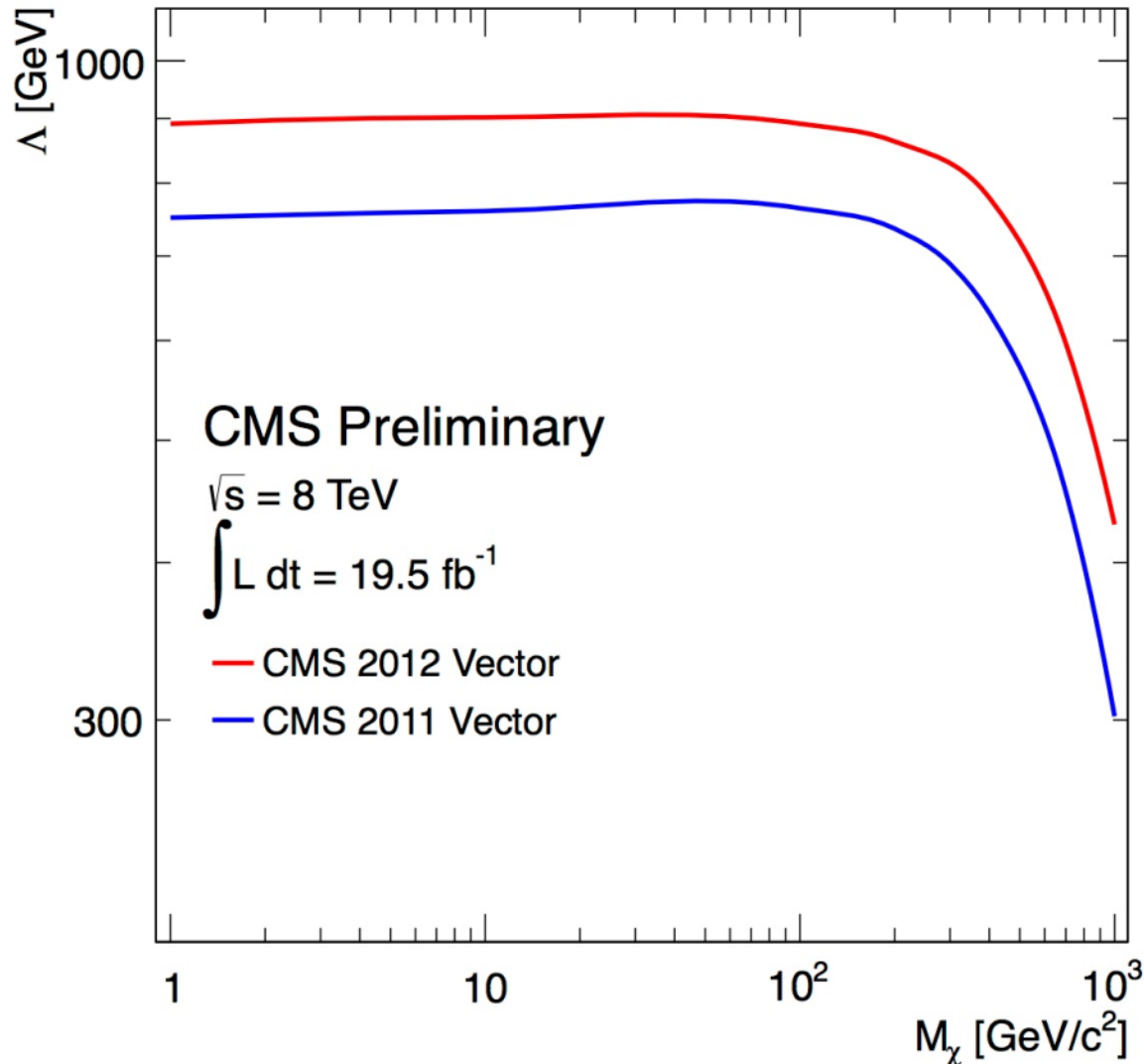
4. $Q_{TR} > 2m_\chi$ (DM pair-produced on-shell)

Combining 3 & 4 gives stronger constraint than 2!

$$\blacksquare \Lambda > \frac{Q_{TR}}{4\pi} > \frac{2m_\chi}{4\pi}$$

Spin Independent Limits on Λ

EXO-12-048 PAS



Let say $\sqrt{g_q g_\chi} = 1$

■ $\Lambda > Q_{TR} > 2m_\chi$

■ @LHC

■ $Q_{TR} \sim O(1 \text{ TeV})$

■ Limits on Λ

■ $< 1 \text{ TeV}$

■ Validity of EFT approach questionable

Johanna Gramling

Intensive Discussion about how to interpret Mono-X analyses

- G. Busonia, A. De Simone, E. Morgante, A. Riotto
 - “On the Validity of the Effective Field Theory for Dark Matter Searches at the LHC”, arXiv:1307.2253v1
 - Derive stronger bounds than currently used by LHC experiments

- New models:
 - A. DiFranzo, K. I. Nagao, A. Rajaraman, T.M.P. Tait,
 - “Simplified Models for Dark Matter Interacting with Quarks”, arXiv:1308.2679v1
 - S. Chang, R. Edezhath, J. Hutchinson, and M. Luty,
 - “Effective WIMPs”, arXiv:1307.8120v1
 - Yang Bai and Joshua Berger,
 - “Fermion Portal Dark Matter”, arXiv:1308.0612v2

Coming back to CMS Mono-Jet Search

EXO-12-048 PAS

Selections

≥ 1 good vertex

$> 20\%$ E_{jet} from charged hadrons

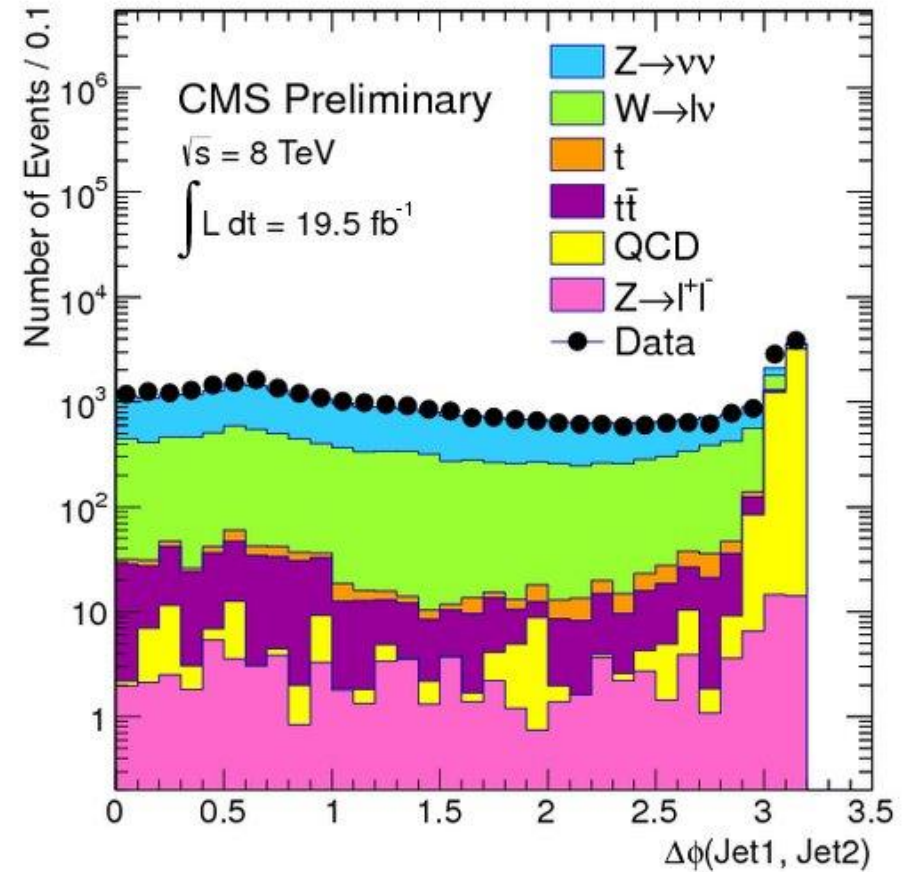
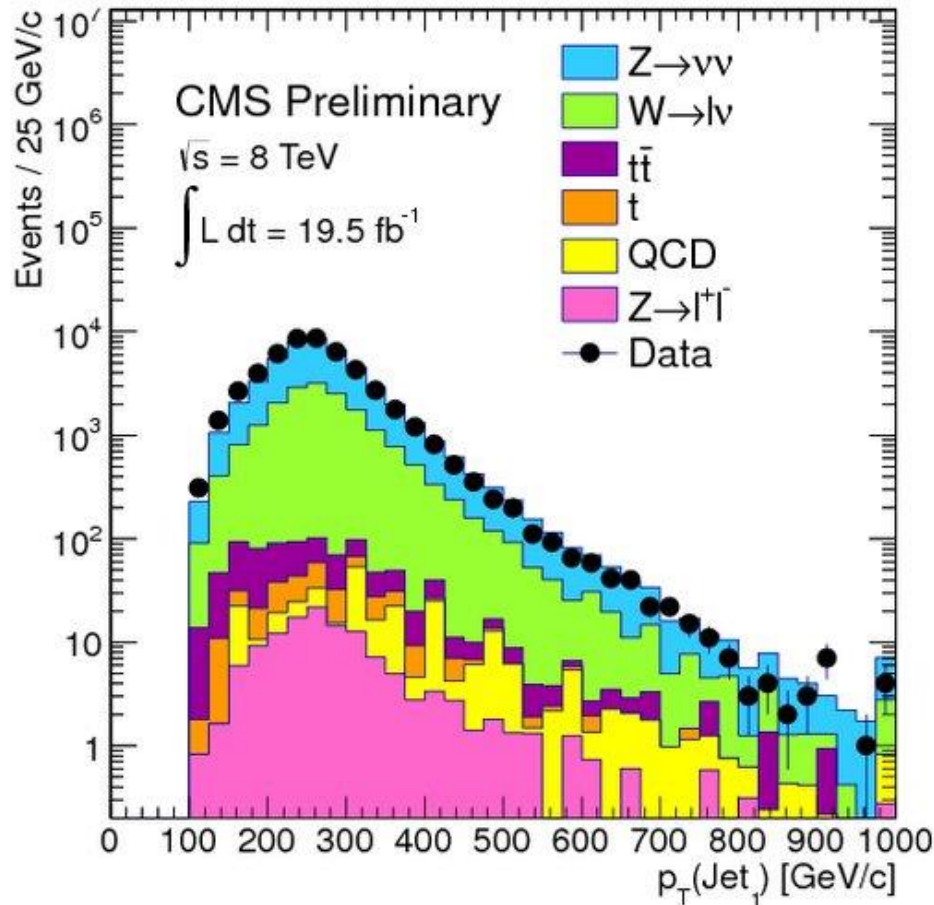
$< 70\%$ E_{jet} from neutral hadrons or photons

$p_{\text{T}}(\text{jet1}) > 110 \text{ GeV} \ \&\& \ |\eta_{\text{jet1}}| < 2.4$

no more than 2 jets with $p_{\text{T}} > 30 \text{ GeV}$ in $|\eta| < 4.5$
& except $\Delta\phi(j1, j2) < 2.5$

no isolated leptons

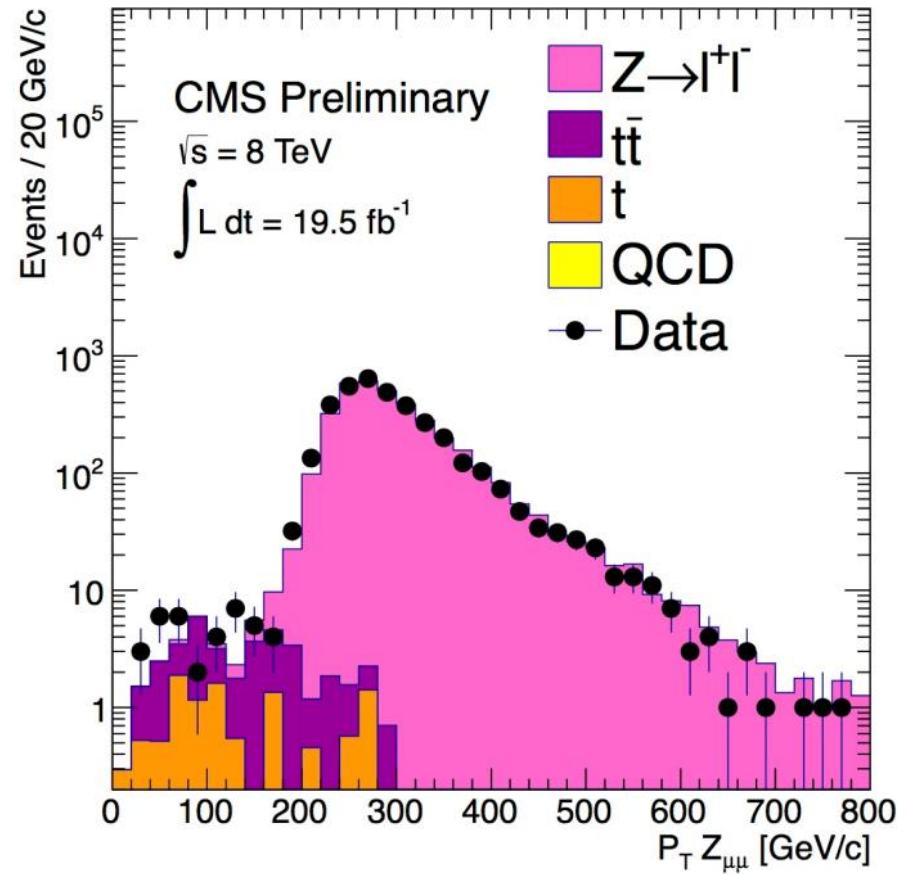
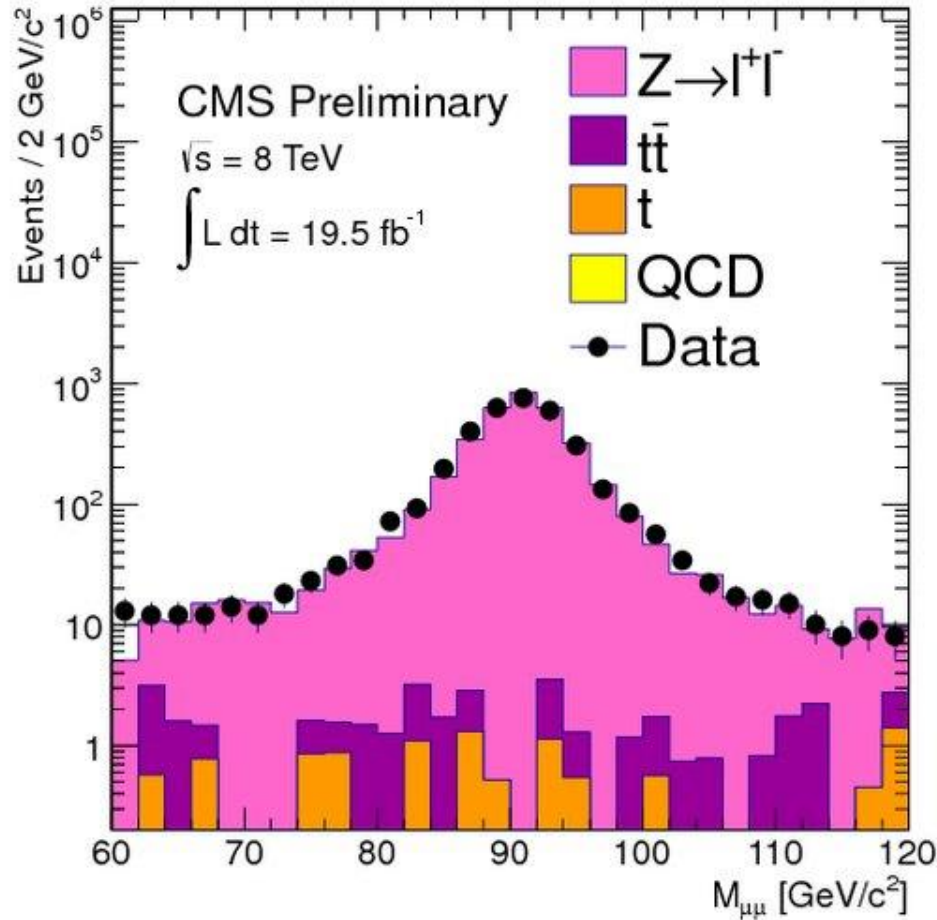
Selection Variable Distributions



Background: $Z(\nu\nu)+\text{jet}$

- Use data to estimate background
- Select $Z(\mu\mu)+\text{jet}$ applying all selections BUT lepton veto
- 2 μ with $p_T > 20 \text{ GeV}$ && $|\eta| < 2.1$
- ≥ 1 isolated μ
- $60 \text{ GeV} < m_{\mu\mu} < 120 \text{ GeV}$

Distribution of $Z(\mu\mu) + \text{jet}$ Sample

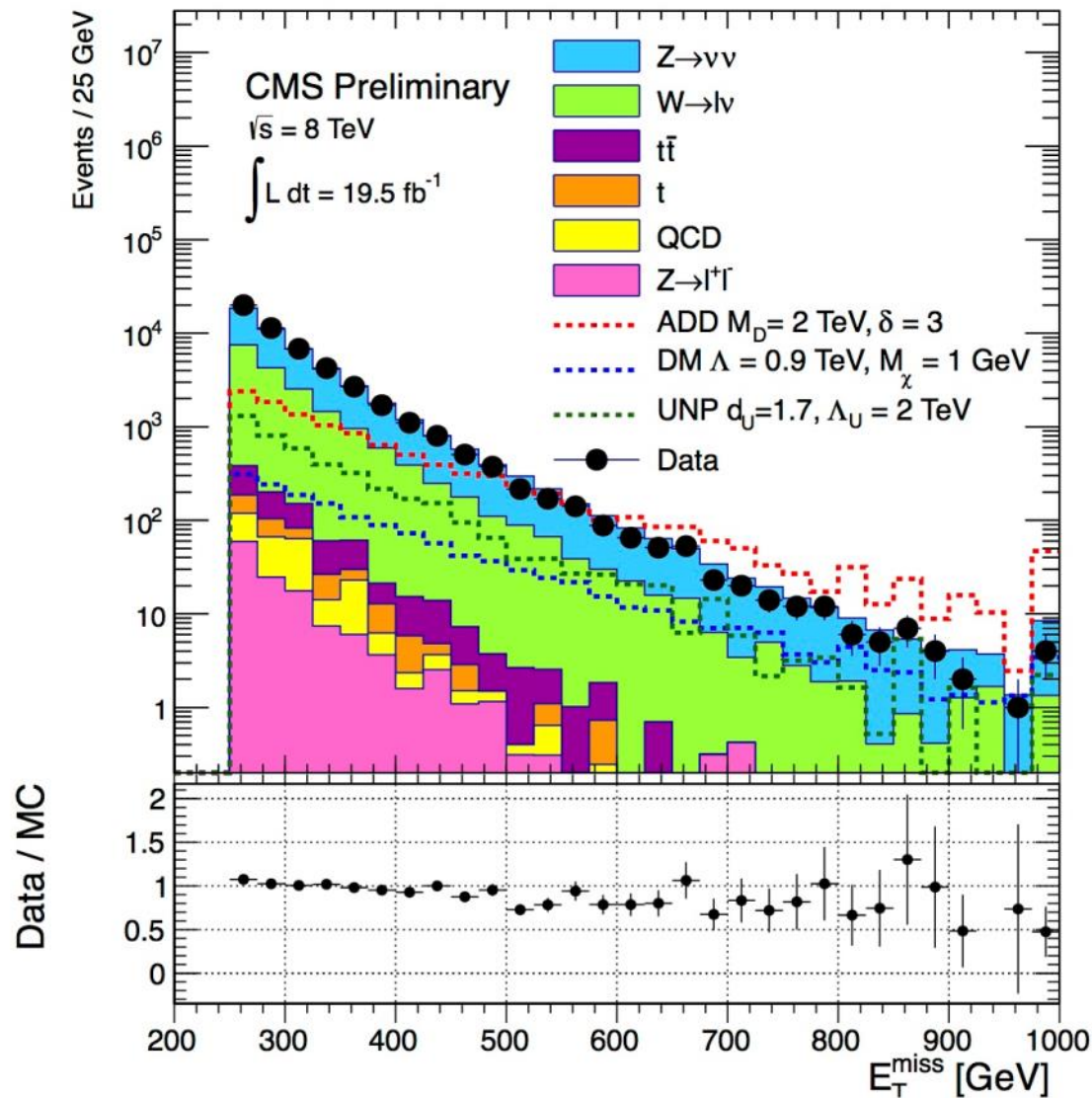


Background: $Z(\nu\nu)+\text{jet}$

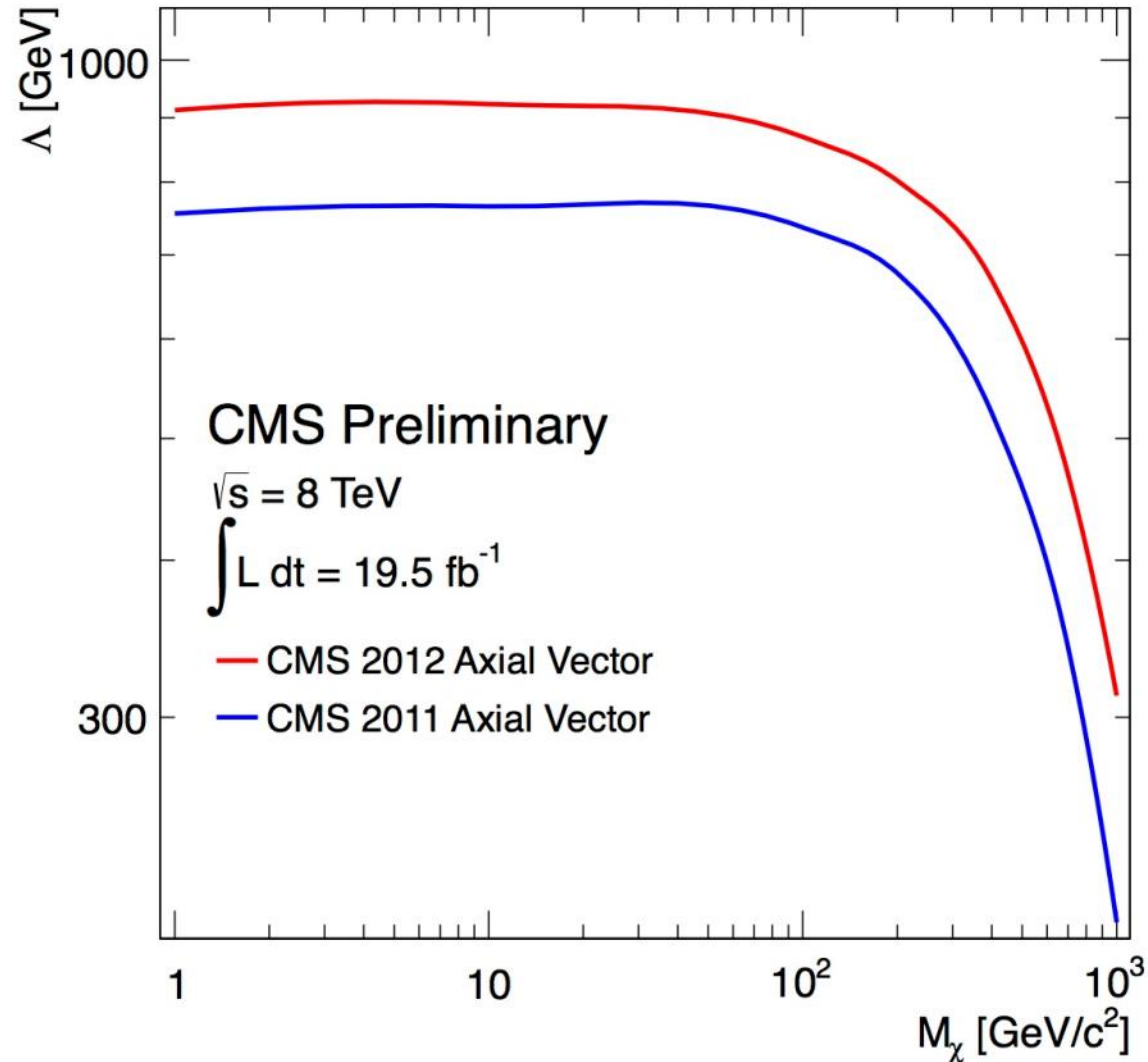
- Use data to estimate background
- Select $Z(\mu\mu)+\text{jet}$ applying all selections BUT lepton veto
- 2 μ with $p_T > 20 \text{ GeV}$ && $|\eta| < 2.1$
- ≥ 1 isolated μ
- $60 \text{ GeV} < m_{\mu\mu} < 120 \text{ GeV}$

$$N(Z(\nu\nu)) = \frac{N^{\text{obs}} - N^{\text{bgd}}}{A \times \epsilon} \cdot R \left(\frac{Z(\nu\nu)}{Z(\mu\mu)} \right)$$

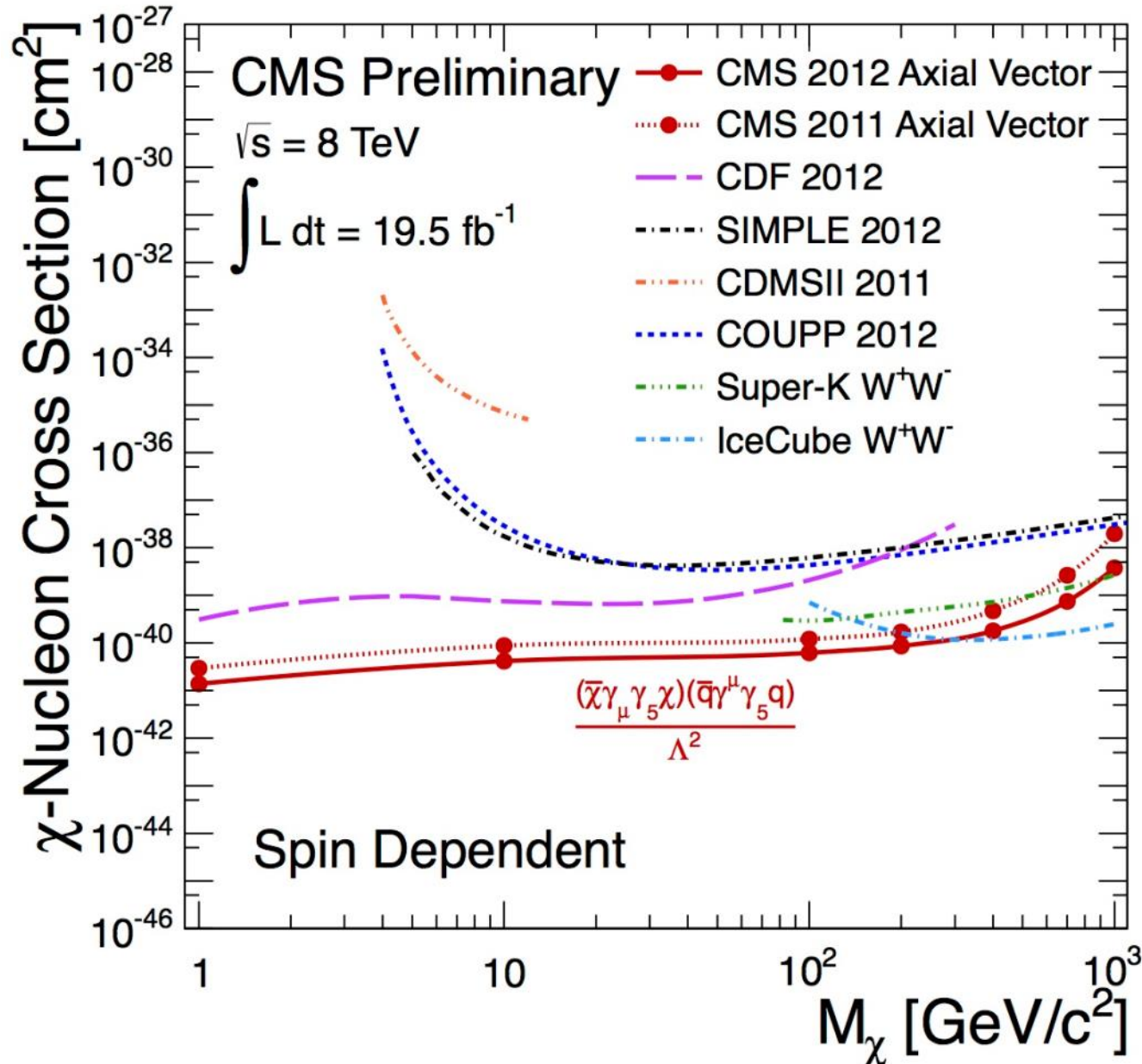
Missing E_T Distribution after all Selections



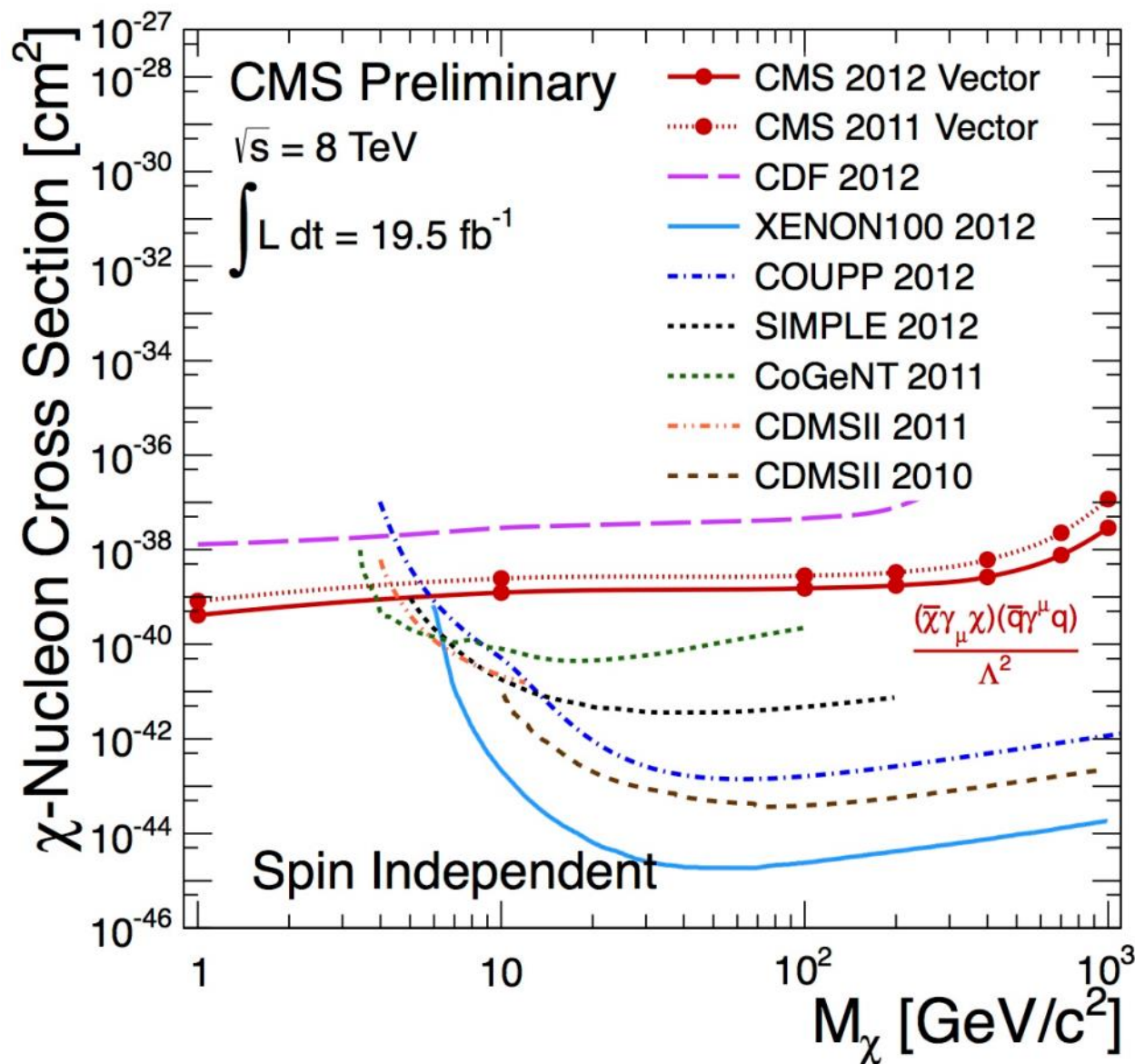
Spin Dependent Limits on Λ

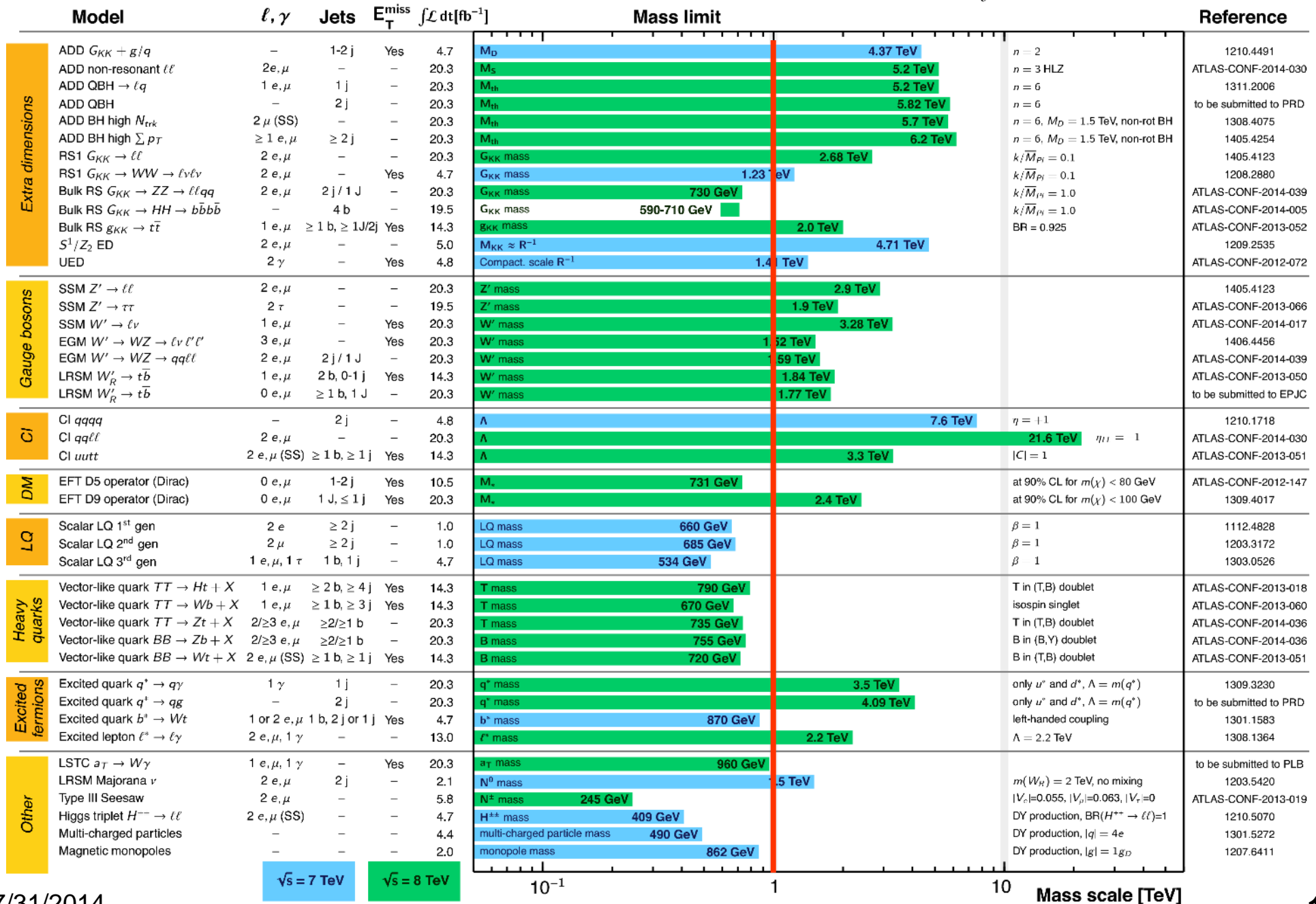


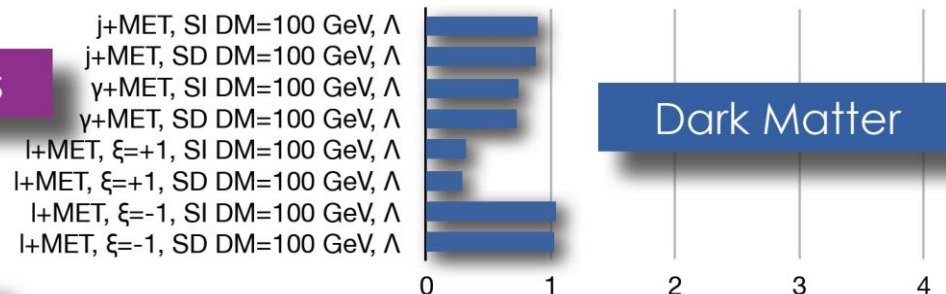
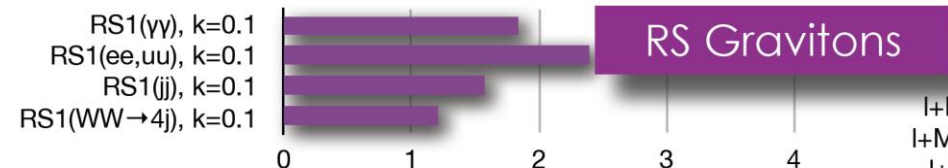
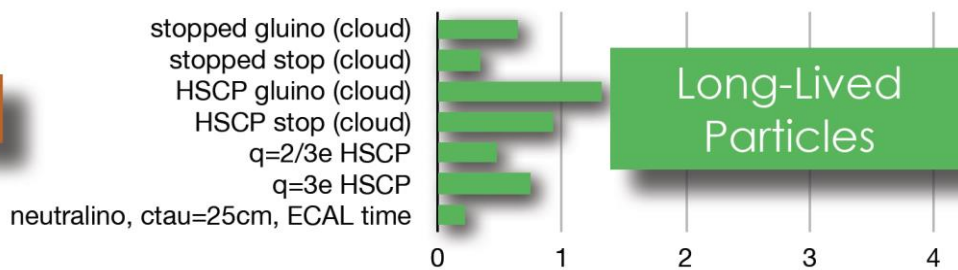
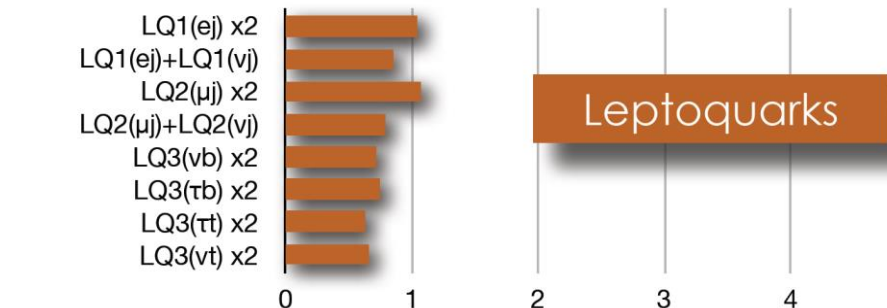
Darkmatter-Nucleon Cross Section Limit



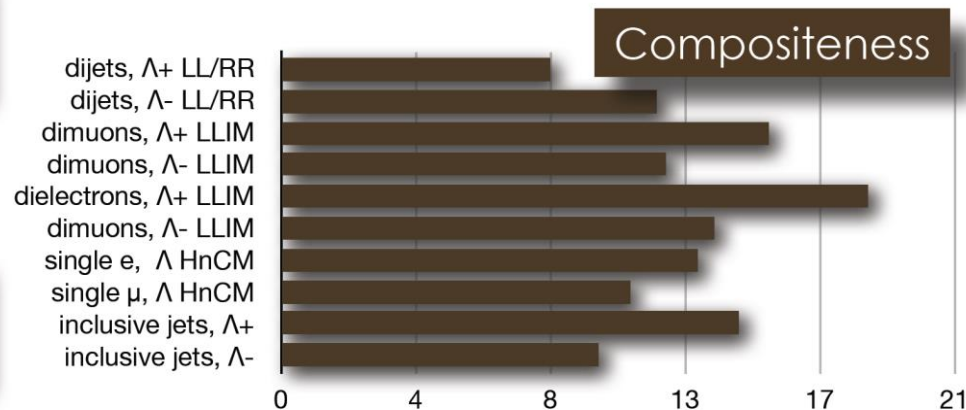
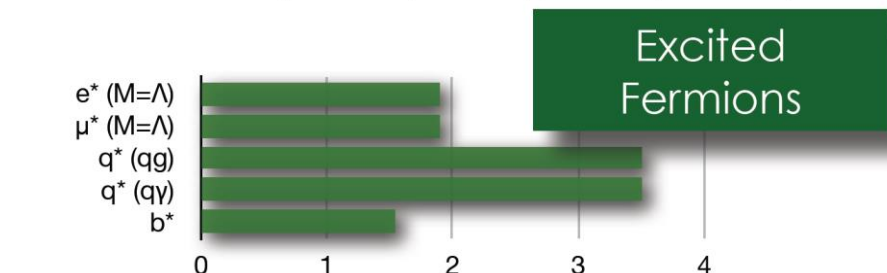
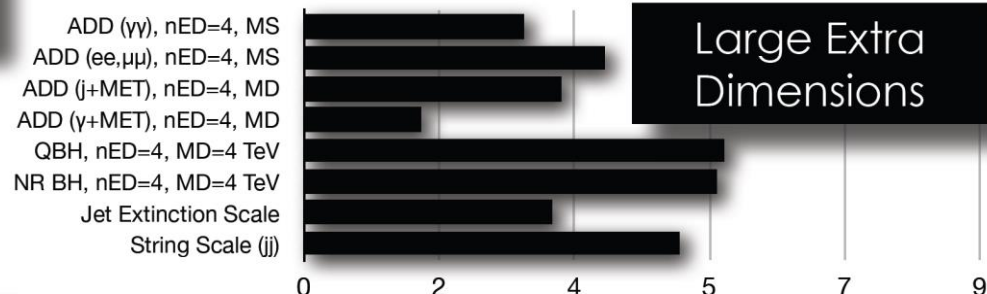
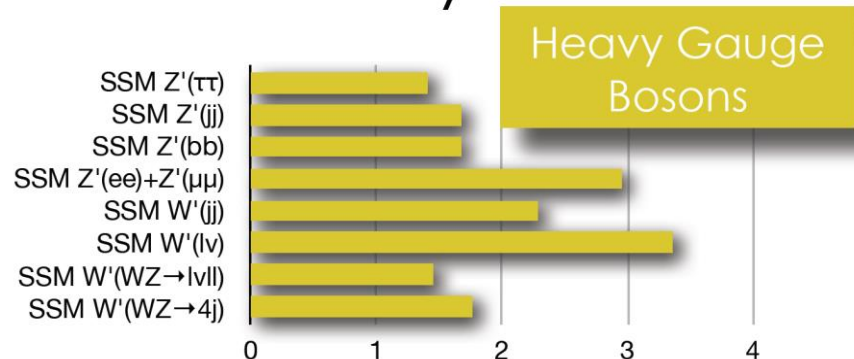
DM-Nucleon cross section upper limits







CMS Preliminary



Conclusions

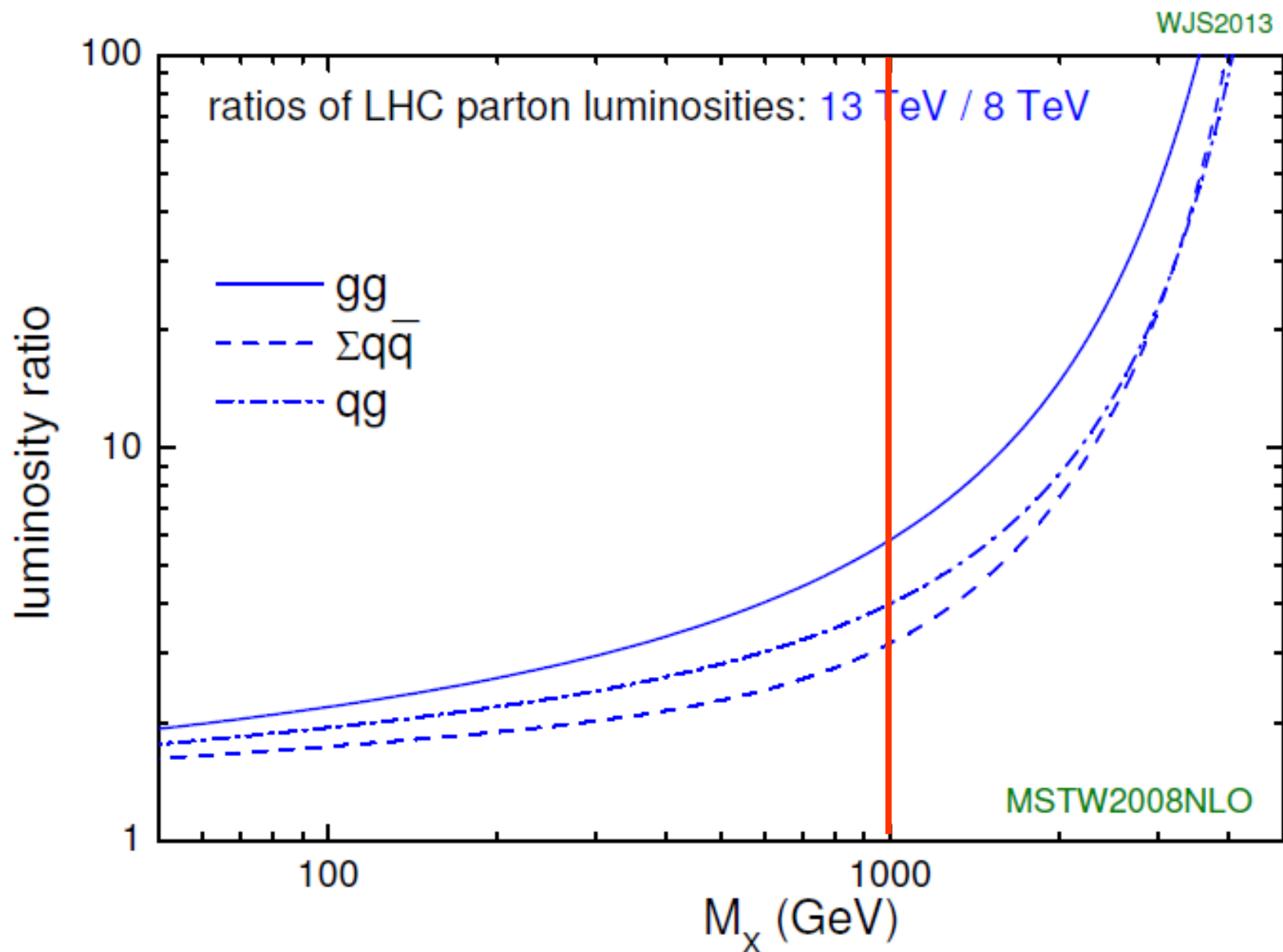
The Standard Model




C. Grojean, A. Weiler

There is new physics out there!

Expectation for Run2



Conclusions: Physics Priorities for HL-LHC

- If no new physics @ Run2
 - Precision measurements
 - Higgs
 - W,Z
 - Top
- 
- EWK Symmetry Breaking
- If new physics found @ Run2
 - Study its properties
 - Understand what we have found

Literature for Further Reading

■ Technicolor and related models

- [http://dx.doi.org/10.1016/0370-1573\(81\)90173-3](http://dx.doi.org/10.1016/0370-1573(81)90173-3)
- <http://dx.doi.org/10.1103/RevModPhys.55.449>
- <http://inspirehep.net/record/205523?ln=en>
- [http://dx.doi.org/10.1016/0146-6410\(83\)90005-4](http://dx.doi.org/10.1016/0146-6410(83)90005-4)

■ Extra Dimensions

- <http://arxiv.org/pdf/hep-ph/0302189.pdf>
- <http://arxiv.org/pdf/gr-qc/0312059.pdf>

■ Exotics new particles

- [http://dx.doi.org/10.1016/0370-1573\(89\)90071-9](http://dx.doi.org/10.1016/0370-1573(89)90071-9)
- <http://dx.doi.org/10.1142/S0217751X88000035>

■ GUT: [http://dx.doi.org/10.1016/0370-1573\(81\)90059-4](http://dx.doi.org/10.1016/0370-1573(81)90059-4)



Backup Slides

Standard Model Lagrangian

$$\mathcal{L}_{SM} = -\frac{1}{4g'^2} B_{\mu\nu} B^{\mu\nu} - \frac{1}{2g^2} \text{Tr}(W_{\mu\nu} W^{\mu\nu}) - \frac{1}{2g_s^2} \text{Tr}(G_{\mu\nu} G^{\mu\nu}) \\ + \bar{Q}_i i \not{D} Q_i + \bar{L}_i i \not{D} L_i + \bar{u}_i i \not{D} u_i + \bar{d}_i i \not{D} d_i + \bar{e}_i i \not{D} e_i$$

Above: Describes gauge fields and interactions

\not{D} determined by **gauge quantum numbers**

strange

Gravity is not included!!

	$SU(3)$	$SU(2)$	$U(1)$	chirality
Q	3	2	+1/6	left
U	3	1	+2/3	right
D	3	1	-1/3	right
L	1	2	-1/2	left
E	1	1	-1	right

Standard Model Lagrangian

$$\begin{aligned}\mathcal{L}_{SM} = & -\frac{1}{4g'^2}B_{\mu\nu}B^{\mu\nu} - \frac{1}{2g^2}\text{Tr}(W_{\mu\nu}W^{\mu\nu}) - \frac{1}{2g_s^2}\text{Tr}(G_{\mu\nu}G^{\mu\nu}) \\ & + \bar{Q}_i i \not{D} Q_i + \bar{L}_i i \not{D} L_i + \bar{u}_i i \not{D} u_i + \bar{d}_i i \not{D} d_i + \bar{e}_i i \not{D} e_i \\ & + (Y_u^{ij} \bar{Q}_i u_j \tilde{H} + Y_d^{ij} \bar{Q}_i d_j H + Y_l^{ij} \bar{L}_i e_j H + \text{h.c.})\end{aligned}$$

- Responsible for mass and mixing of quark masses
- Responsible for charged lepton masses
- Generation index: $i, j = 1, 2, 3$
- Why 3 families?
- No neutrino masses or mixing included

Standard Model Lagrangian

$$\begin{aligned}\mathcal{L}_{SM} = & -\frac{1}{4g'^2}B_{\mu\nu}B^{\mu\nu} - \frac{1}{2g^2}\text{Tr}(W_{\mu\nu}W^{\mu\nu}) - \frac{1}{2g_s^2}\text{Tr}(G_{\mu\nu}G^{\mu\nu}) \\ & + \bar{Q}_i i \not{D} Q_i + \bar{L}_i i \not{D} L_i + \bar{u}_i i \not{D} u_i + \bar{d}_i i \not{D} d_i + \bar{e}_i i \not{D} e_i \\ & + (Y_u^{ij} \bar{Q}_i u_j \tilde{H} + Y_d^{ij} \bar{Q}_i d_j H + Y_l^{ij} \bar{L}_i e_j H + \text{h.c.}) \\ & + (D_\mu H)^\dagger (D^\mu H) - \lambda (H^\dagger H)^2 - m^2 H^\dagger H + \underbrace{\frac{\theta}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} \text{Tr}(G_{\mu\nu} G_{\rho\sigma})}_{\theta \text{ term in QCD}}.\end{aligned}$$

Strong CP Problem in SM

- Why is $\theta < 1.2 \times 10^{-10}$???
- Natural value ~ 1

θ term in QCD

Periodic: $0 - 2\pi$

Violates T and CP

Standard Model Lagrangian

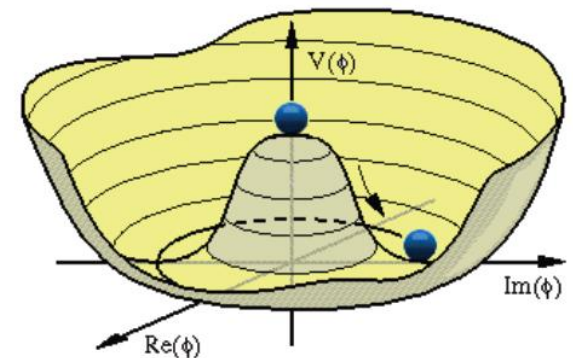
$$\mathcal{L}_{SM} = -\frac{1}{4g'^2} B_{\mu\nu} B^{\mu\nu} - \frac{1}{2g^2} \text{Tr}(W_{\mu\nu} W^{\mu\nu}) - \frac{1}{2g_s^2} \text{Tr}(G_{\mu\nu} G^{\mu\nu})$$

$$+ \bar{Q}_i i \not{D} Q_i + \bar{L}_i i \not{D} L_i + \bar{u}_i i \not{D} u_i + \bar{d}_i i \not{D} d_i + \bar{e}_i i \not{D} e_i$$

$$+ (Y_u^{ij} \bar{Q}_i u_j \tilde{H} + Y_d^{ij} \bar{Q}_i d_j H + Y_l^{ij} \bar{L}_i e_j H + \text{h.c.})$$

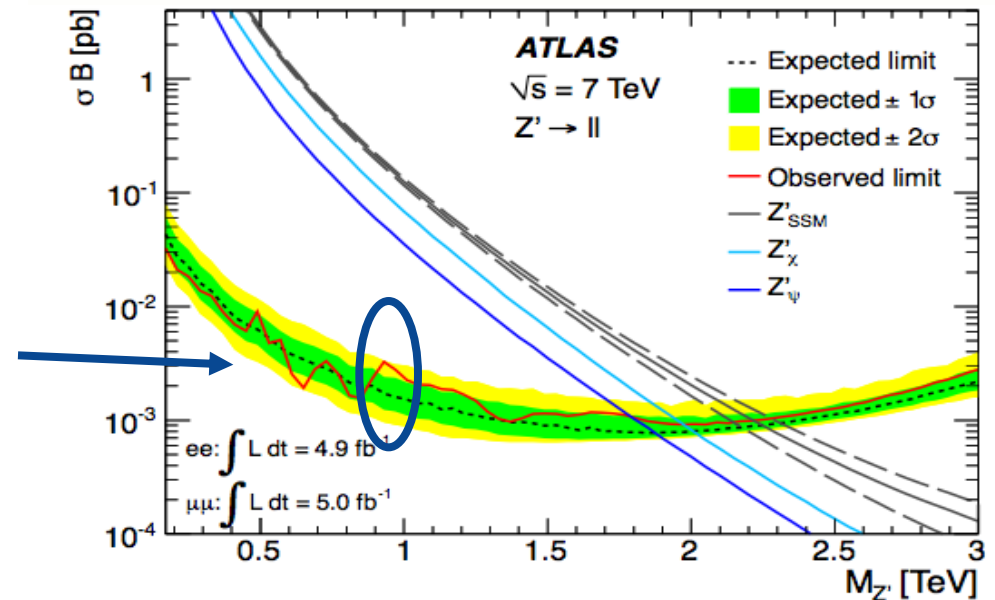
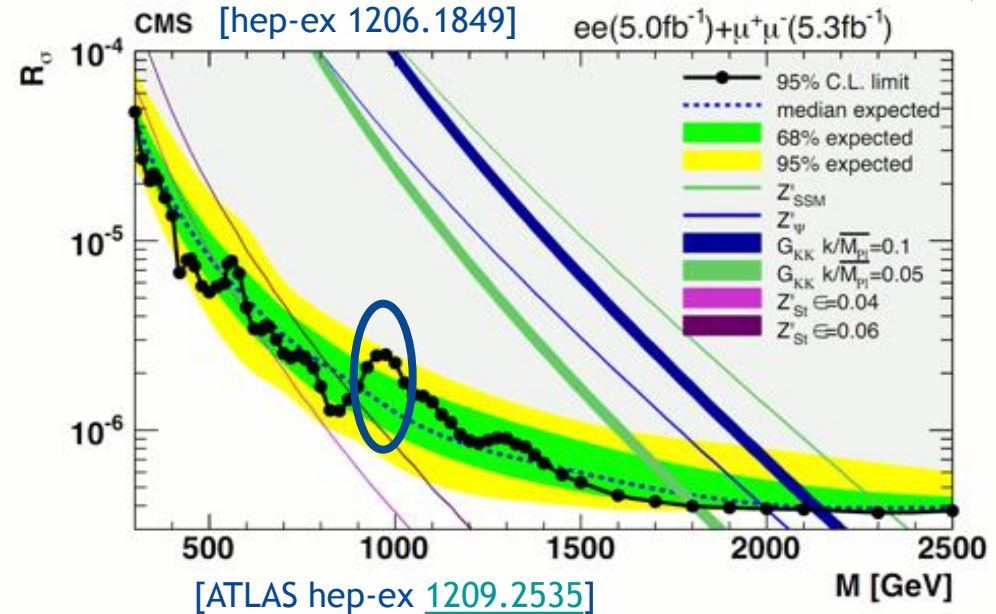
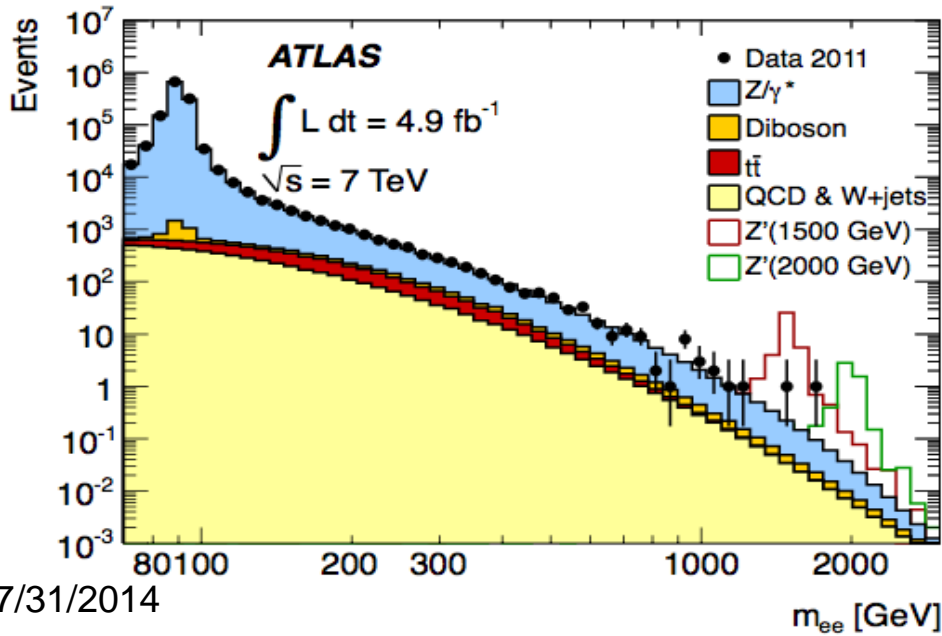
$$+ (D_\mu H)^\dagger (D^\mu H) - \lambda (H^\dagger H)^2 - m^2 H^\dagger H + \frac{\theta}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} \text{Tr}(G_{\mu\nu} G_{\rho\sigma}).$$

Higgs field



Z' in 2011 Data?

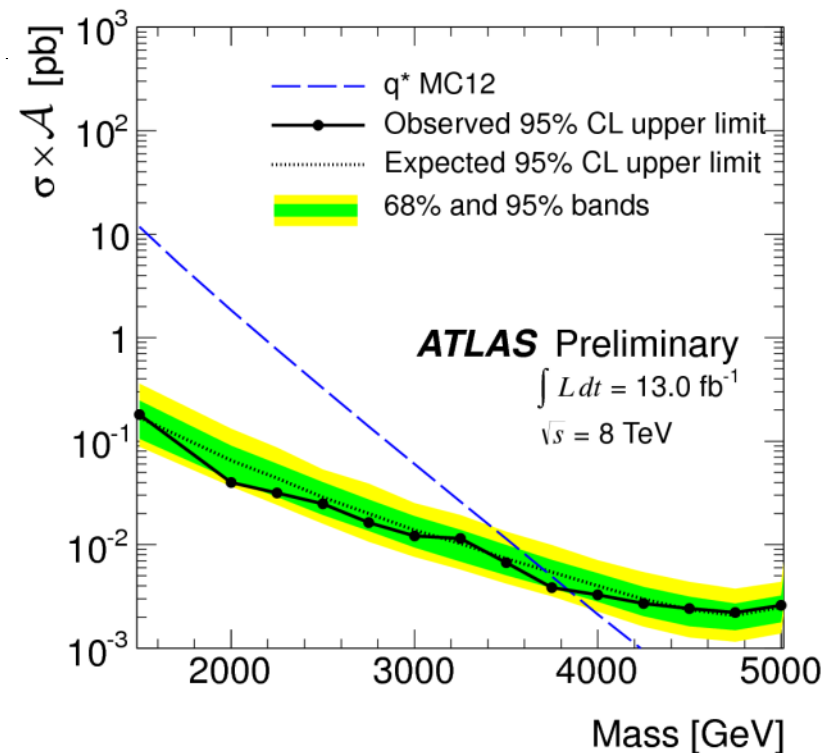
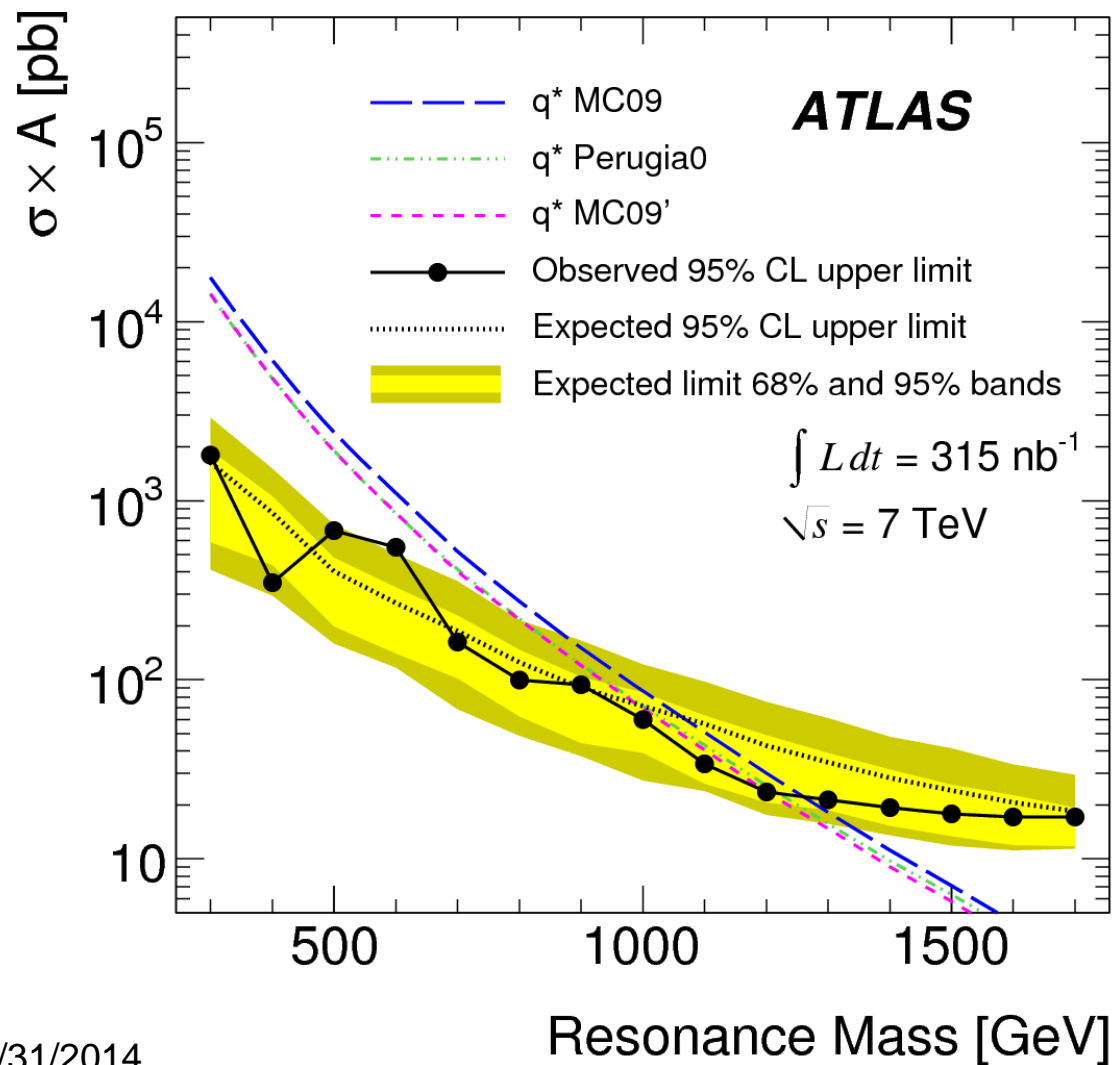
- Interesting features in dilepton spectra
 - around 2σ each for CMS & ATLAS in $e^+\mu^-$
 - similar in scale to 2011 Higgs excess



Mono Jet Signal Region Definitions

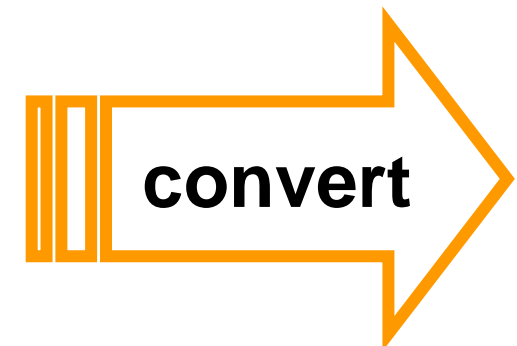
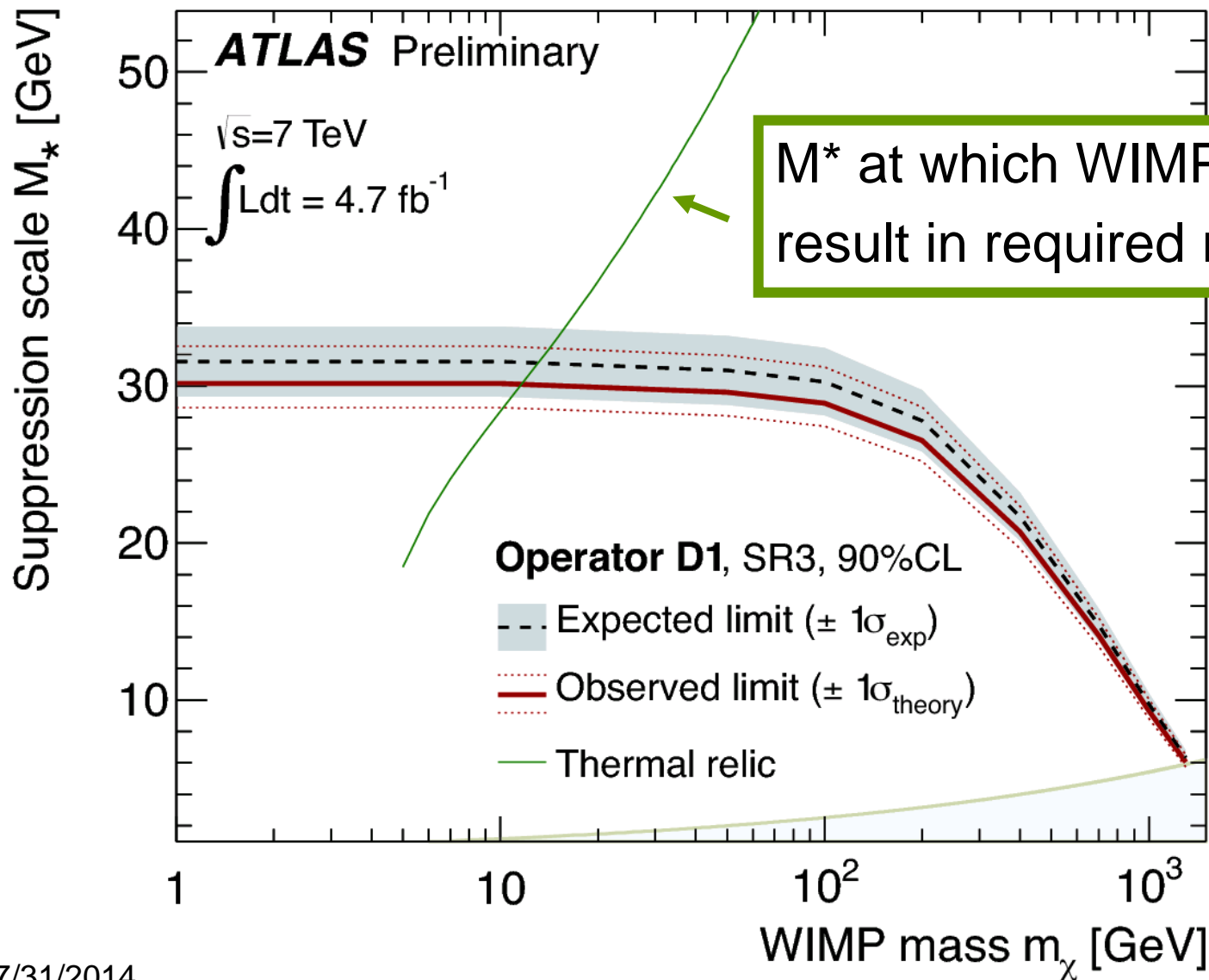
Signal regions	SR1	SR2	SR3	SR4
Common requirements	Data quality + trigger + vertex + jet quality + $ \eta^{\text{jet1}} < 2.0 + \Delta\phi(\mathbf{p}_T^{\text{miss}}, \mathbf{p}_T^{\text{jet2}}) > 0.5 + N_{\text{jets}} \leq 2 +$ lepton veto			
$E_T^{\text{miss}}, p_T^{\text{jet1}} >$	120 GeV	220 GeV	350 GeV	500 GeV

“Although the results of this analysis are interpreted in terms of the ADD model and WIMP pair production, the event selection criteria have not been tuned to maximize the sensitivity to any particular BSM scenario. To maintain sensitivity to a wide range of BSM models, four sets of overlapping kinematic selection criteria, designated as SR1 to SR4, are defined (table 2).”

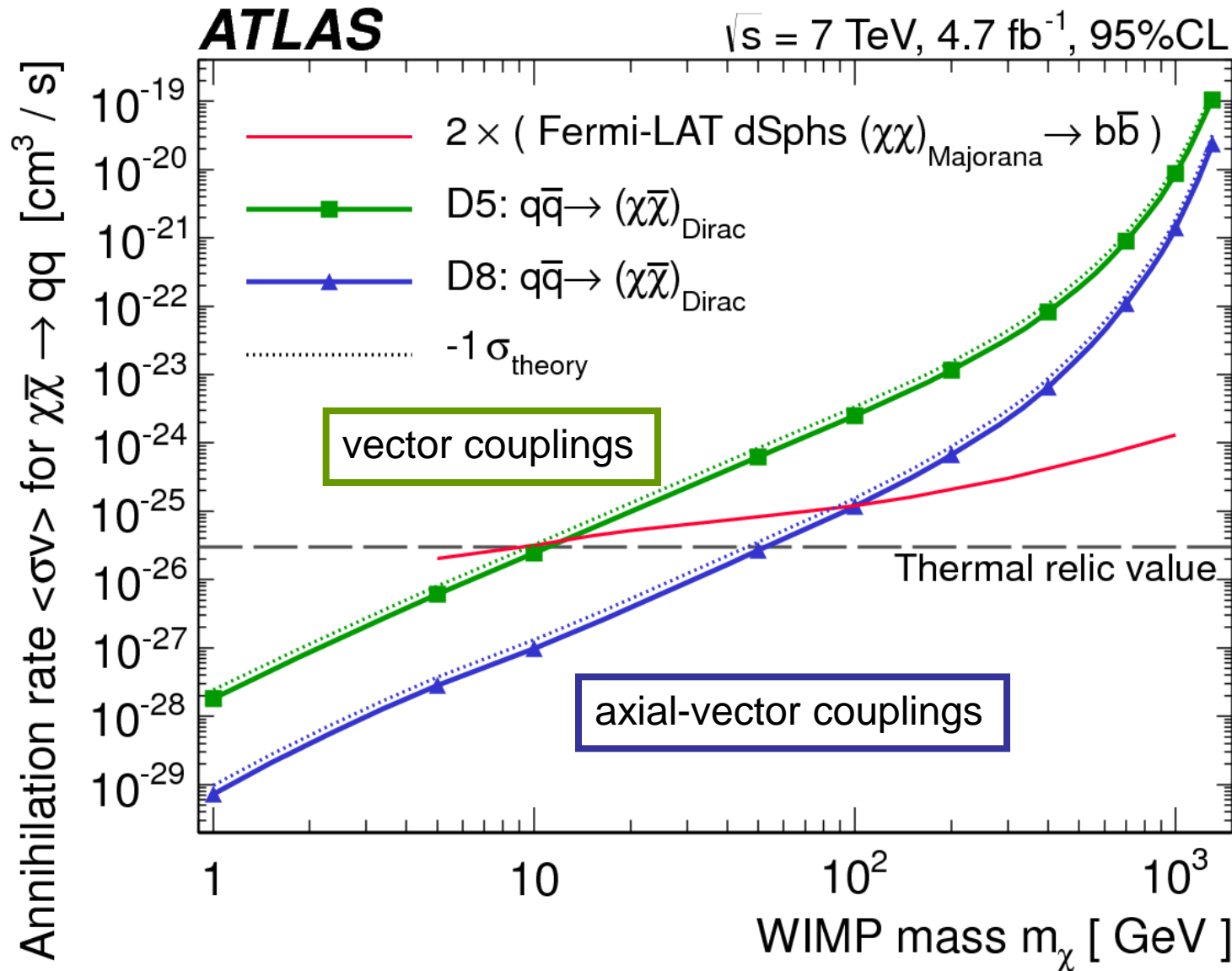


Limits on Dark Matter – Mono Jet

90% CL lower limits on M^*



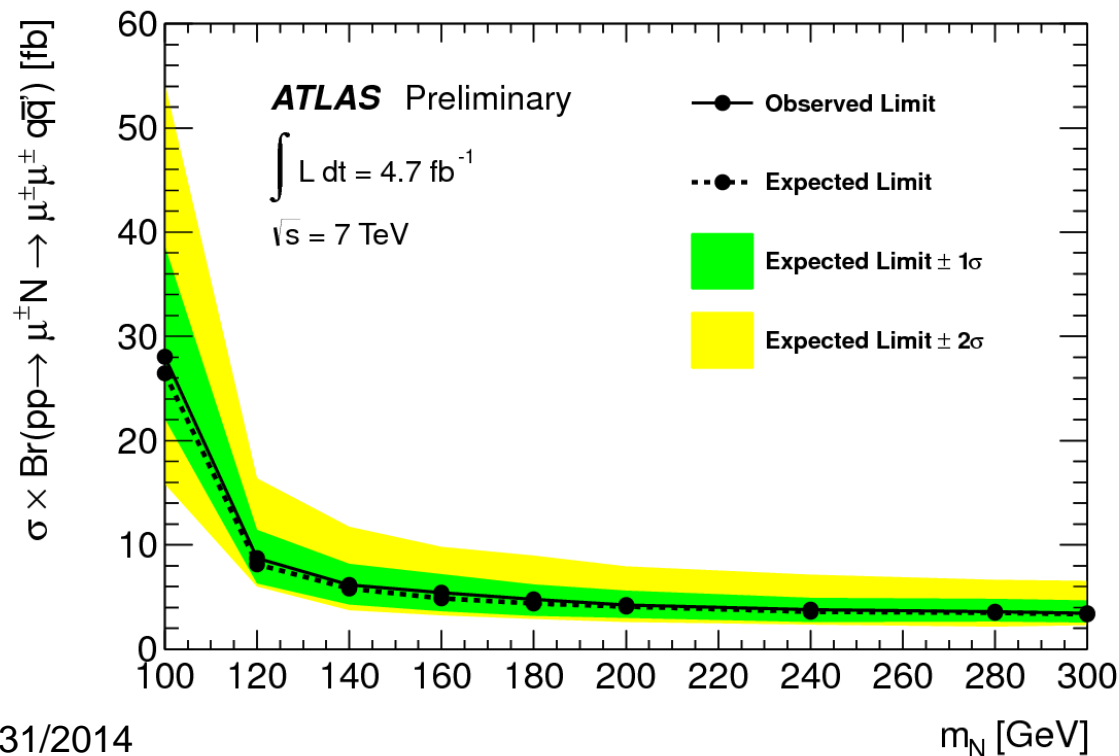
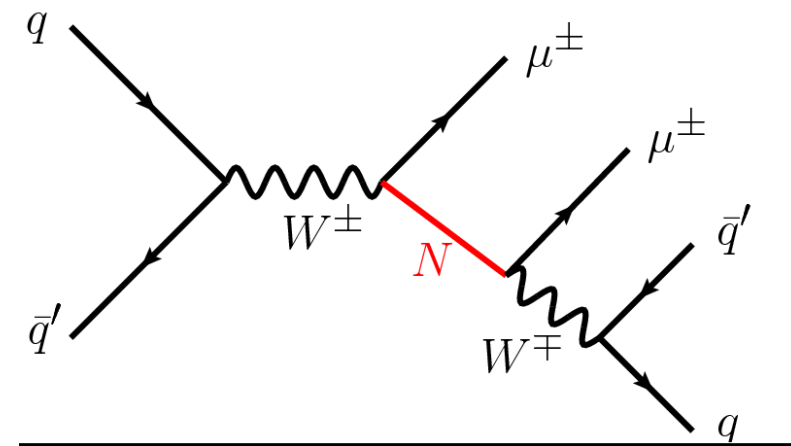
Limits on the annihilation rate of WIMPs



Majorana Neutrino Search in same-sign leptons

ATLAS-CONF-2012-139

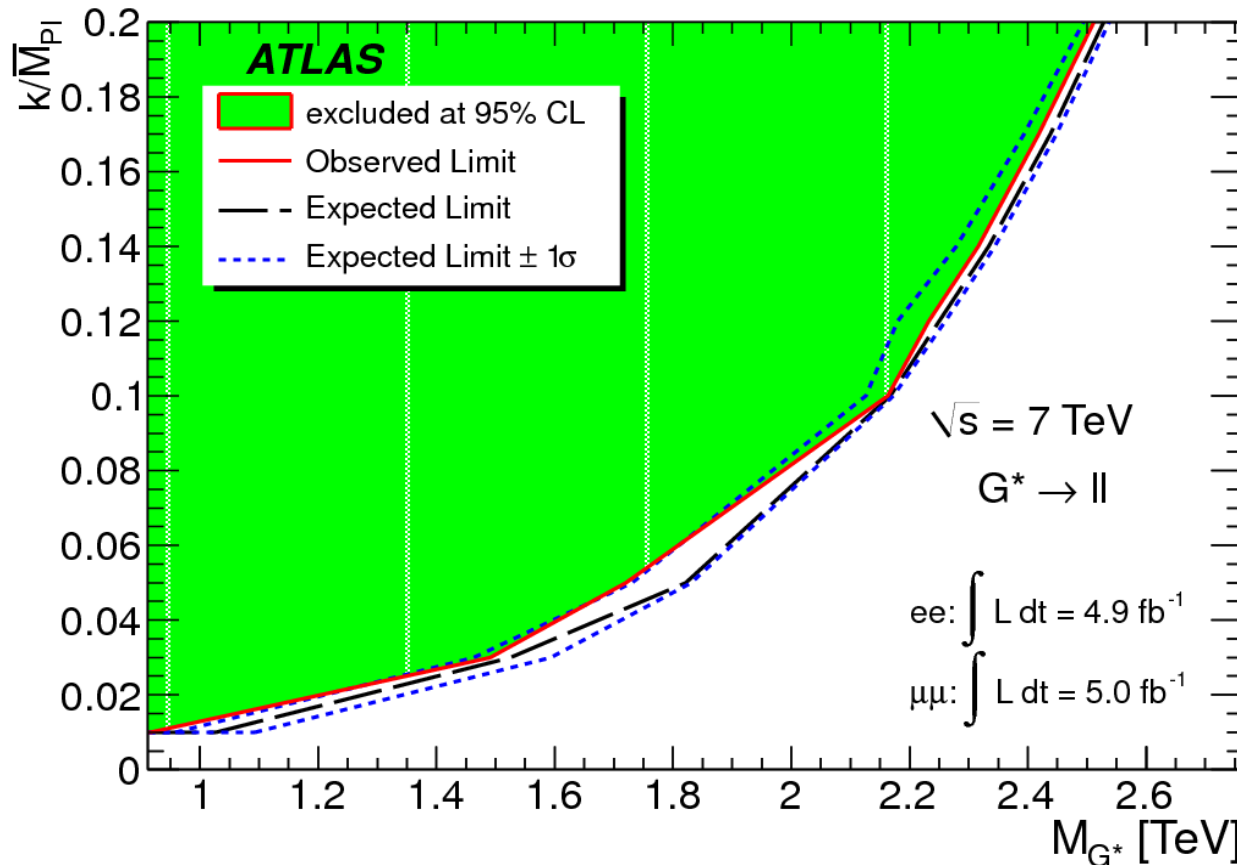
- Two same-sign muons
- ≥ 2 jets and low ME_T



observed limits range from 28 to 3.4 fb for heavy neutrino masses between 100 and 300 GeV

Search for Heavy Resonance: dilepton channel

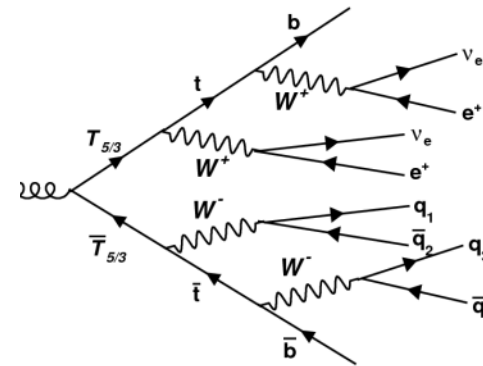
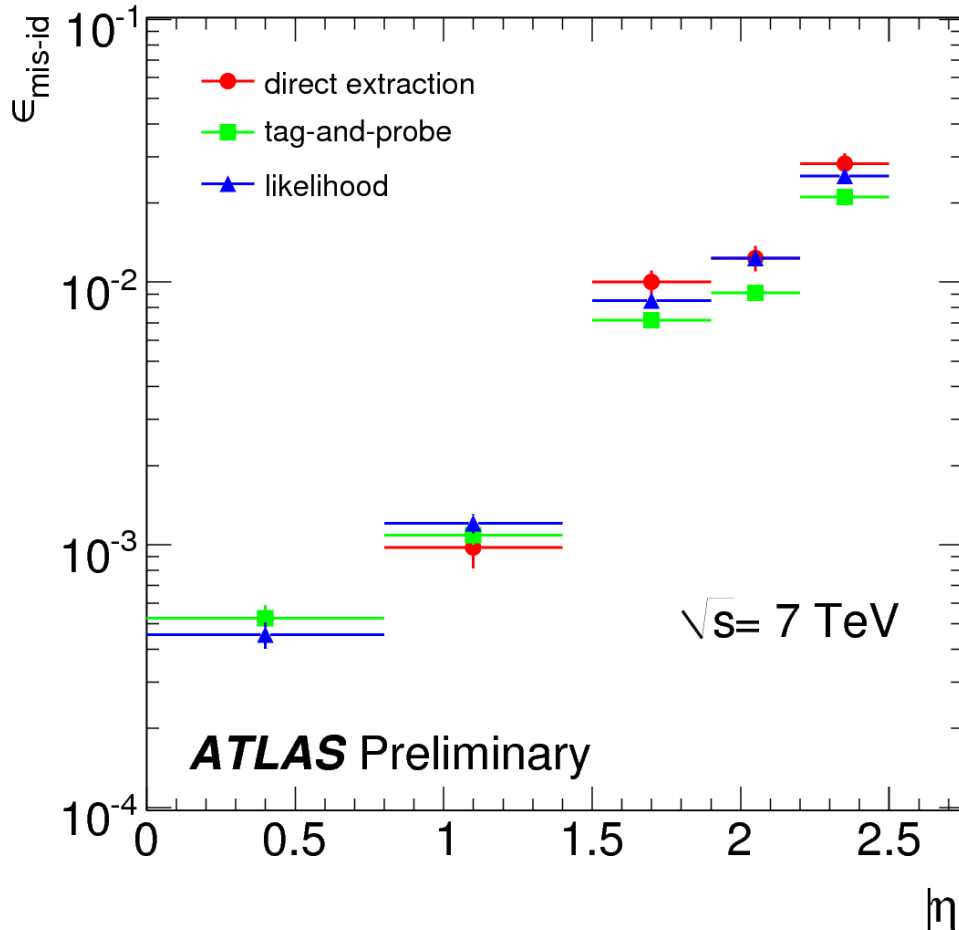
- Limits as a function of RS graviton mass and coupling
 $m(\text{RS graviton}, k/\text{MPI} = 0.1) > 2.16 \text{ TeV at } 95\% \text{ CL}$



Exotic Same-Sign Dilepton Signatures: b' , $T^{5/3}$

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Charge mis-id rate

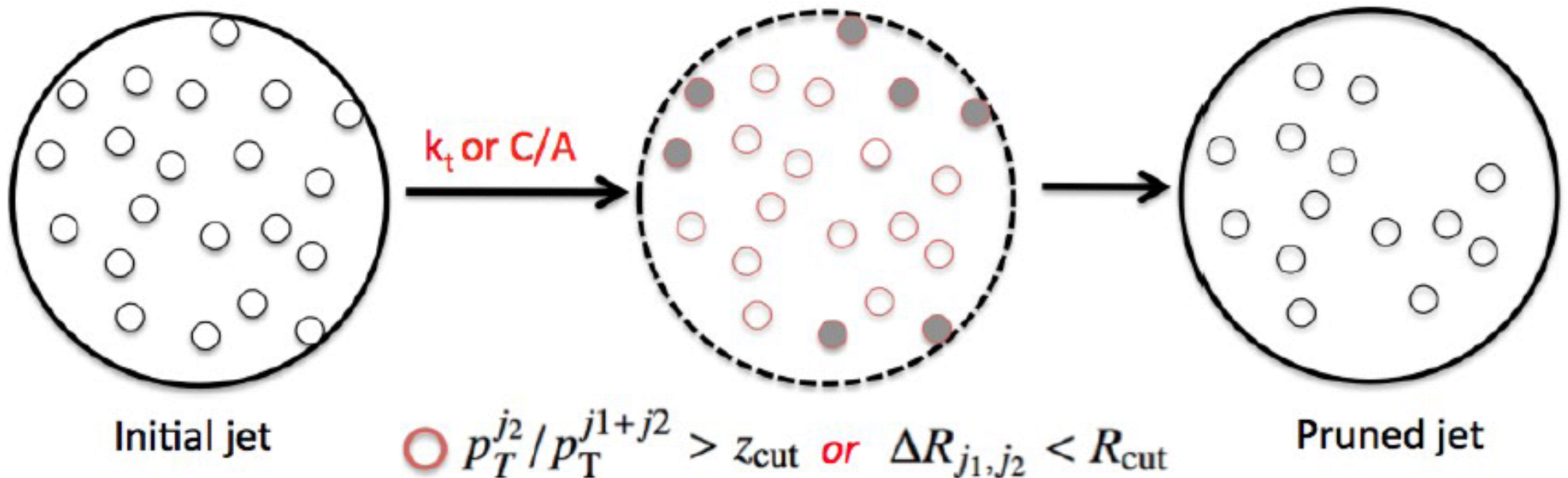


- 2 isolated same-sign leptons (e or μ)
- $ME_T > 40$ GeV
- ≥ 2 jets (≥ 1 b-tagged jet)
- large overall transverse momentum
- $H_T > 550$ GeV

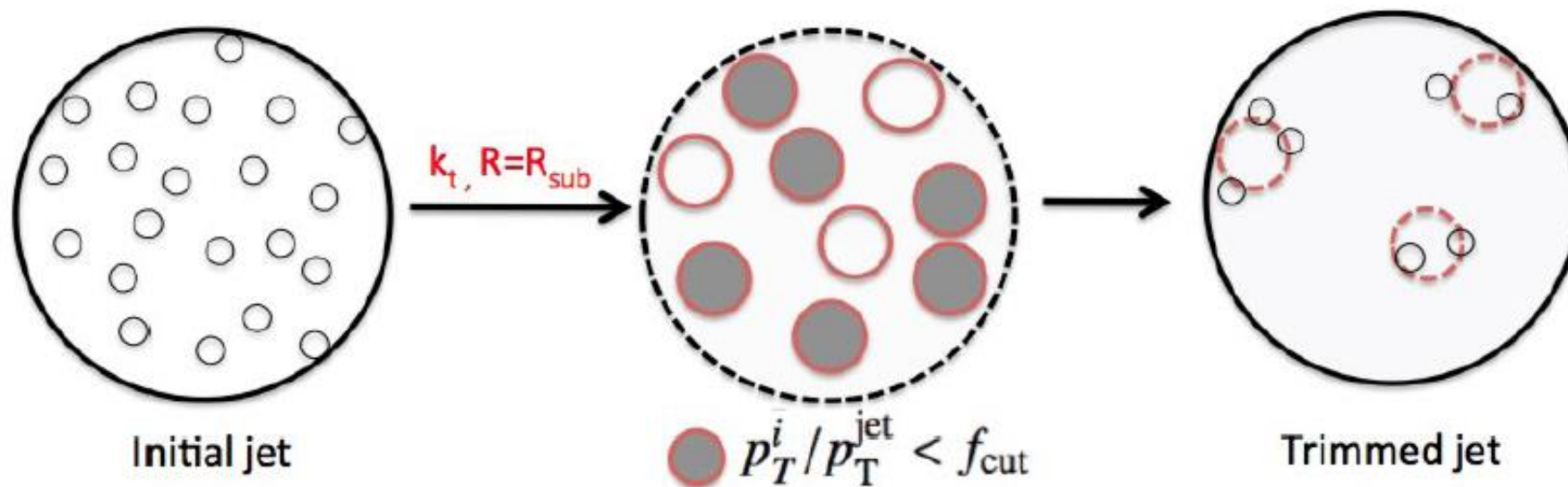
4 events observed
expected background of 5.6 ± 1.7

Jet Grooming

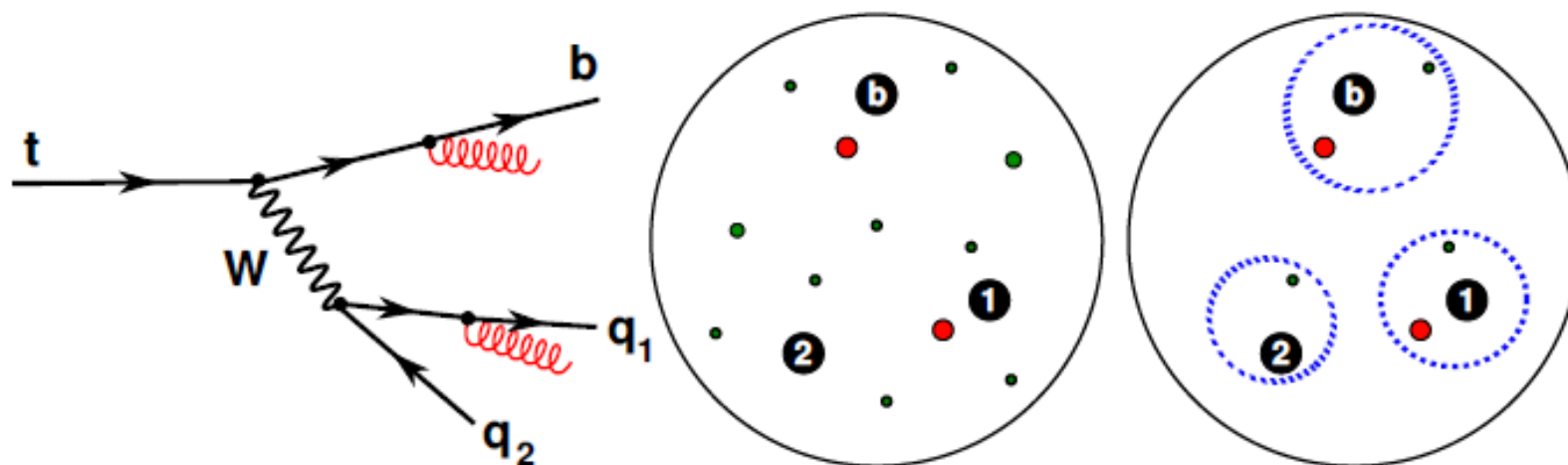
- “Pruning”:
- Start with a fat jet ($R \sim 1$ or more)
- Run k_t or C/A algorithm on clusters within the fat jet
- At each step, if merging of two clusters fails, remove cluster with smallest p_T



- “Trimming”:
- Start with a fat jet ($R \sim 1$ or more)
- Run k_t algorithm on clusters within the fat jet
- Keep only jets with $p_T > p_T(\text{fat jet}) \cdot f_{\text{cut}}$



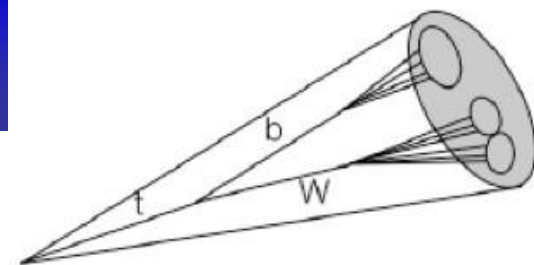
HEPTopTagger (Filtering)



- 1 Decompose until $m_{j_i} < 30 \text{ GeV}$ with mass drop requirement
 $m_{j_i} < \mu m_{\text{large jet}}$
- 2 Investigate 3 subjects and their constituents
- 3 Re-cluster using C/A with parameter
 $R = \min(0.3, \min_{ij} \Delta R(j_i, j_j)/2)$
- 4 Use only 5 hardest subjects of last step
- 5 Built exactly 3 subjects from the selected constituents

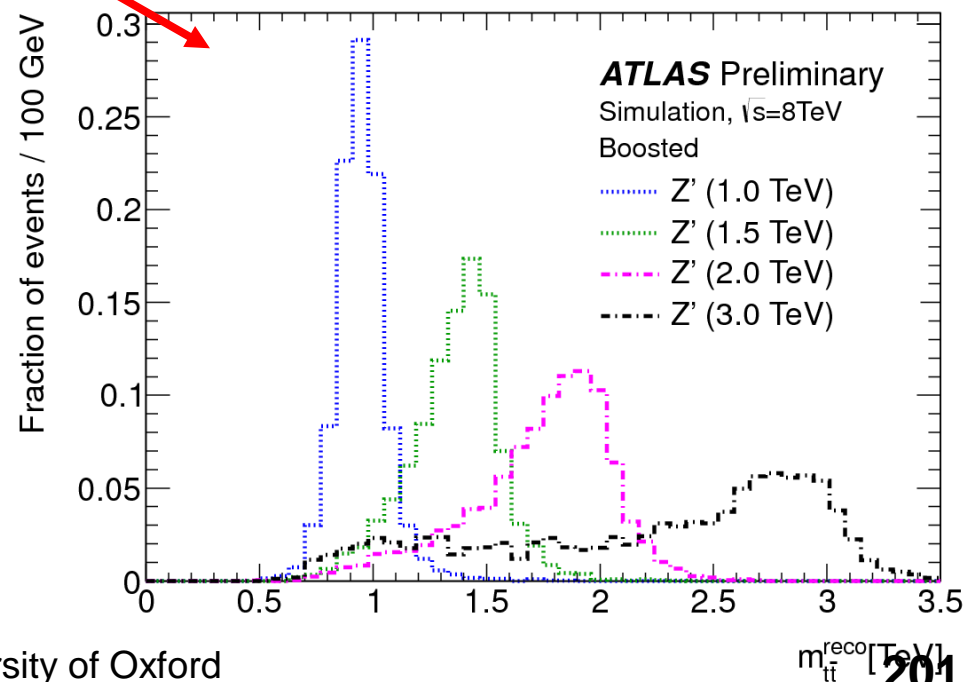
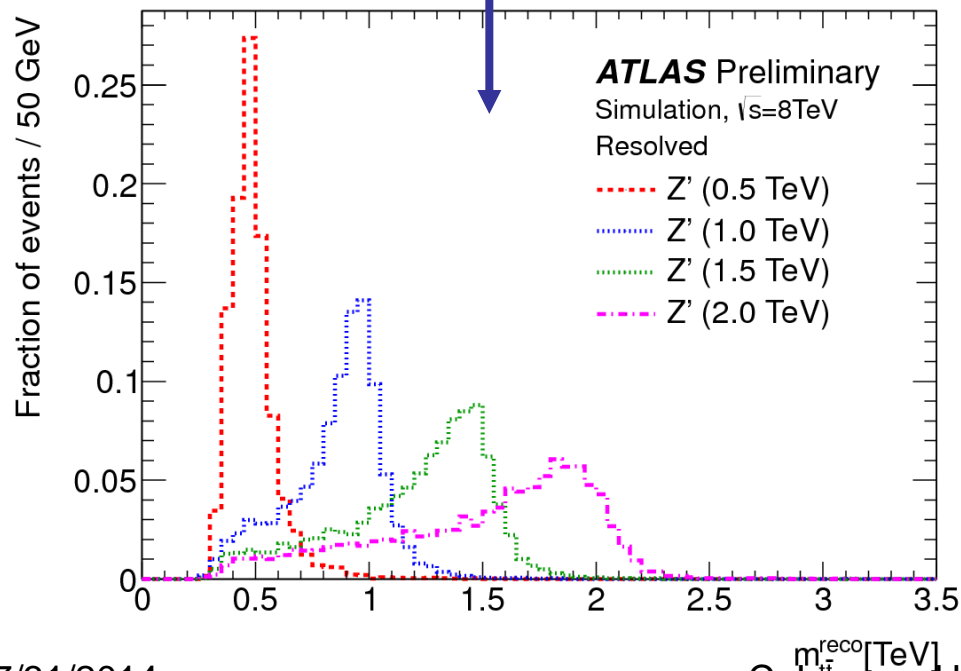
S. Fleischmann

Heavy Resonances Search: $Tt\bar{b}$



ATLAS-CONF-2013-052

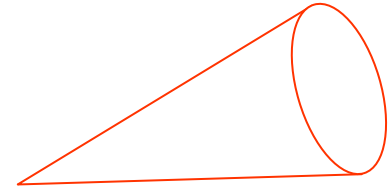
- Lepton+jets channel
- Models: e.g. bulk-RS (esp. KK gluons) and Leptophobic Z'
 - Large Branching Ratio to top-antitop
- Taking full advantage of boosted techniques
- Combining **resolved** and **boosted** reconstructions



Heavy Resonances Search: Object Selection

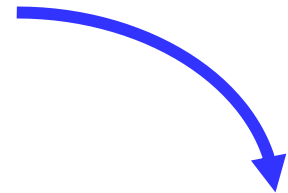
■ Jets

- Small jets: $p_T > 25 \text{ GeV} \ \&\& \ |\eta| < 2.5$
- Large jets: $p_T > 300 \text{ GeV} \ \&\& \ |\eta| < 2.0$
- Require that at least one of the small jets is b-tagged



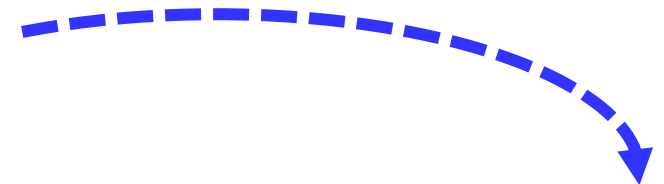
■ Electrons

- $p_T > 25 \text{ GeV} \ \&\& \ |\eta| < 1.37, 1.52 < |\eta| < 2.47$
- Mini Isolation: $I_{\text{mini}} < 0.05 E_T$
- z-impact parameter within 2mm of PV



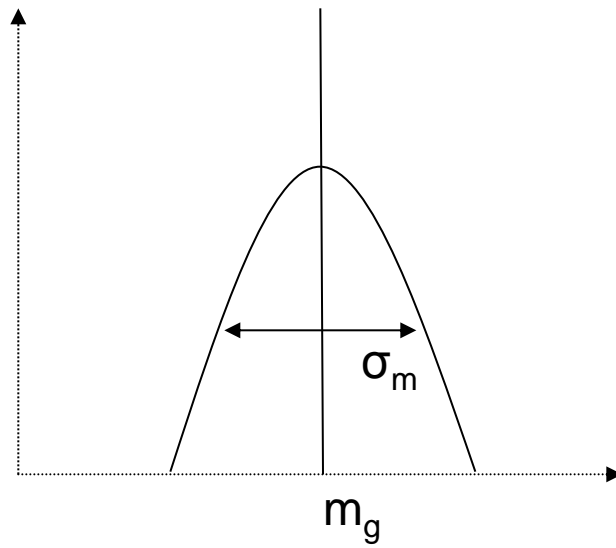
■ Muons

- $p_T > 25 \text{ GeV} \ \&\& \ |\eta| < 2.5$
- $I_{\text{mini}} < 0.05 p_T$
- z-impact parameter within 2mm of PV



Selections Continued

- Optimized for high-pt tops && reduce ttbar bkg
- High-pt single electron or muon trigger
- >1 primary vertex with ≥ 5 tracks of $p_T > 0.4$ GeV
- Electron channel
 - $ME_T > 30$ GeV && $m_T = \sqrt{2p_T ME_T (1 - \cos\Delta\phi)} > 30$ GeV
- Muon channel
 - $ME_T > 20$ GeV && $ME_T + m_T > 60$ GeV



Strong CP Problem of QCD

- QCD allows for CP violation
 - Has an effective strong CP violating term, Θ
 - $0 < \Theta < 2\pi$ possible ranges of values
 - CP violating interactions originating from QCD \rightarrow neutron electric dipole moment non zero
 - But neutron dipole moment measurements $\rightarrow \Theta \sim 0$
 - Not natural. Why?
- One solution: Peccei–Quinn mechanism
 - Introduce new symmetry
 - Θ becomes particle \rightarrow Axion
- Axions are predicted to change to and from photons in the presence of strong magnetic

Particle Accelerators

hep-ph/0201029, hep-ex/0605101, hep-ph/9909294, hep-ex/0710.3338,
hep-ex/0707.2524, Phys. Lett. B568 (2003) 35-47, ZEUS-prel-07-028

■ DESY:

- H1: $M_s^- > 0.78 \text{ TeV}$ and $M_s^+ > 0.82 \text{ TeV}$
- ZEUS: $M_s^- > 0.9 \text{ TeV}$ and $M_s^+ > 0.88 \text{ TeV}$

■ LEP:

- $M_D = 1.5 \text{ TeV}$ for $n = 2 \Leftrightarrow R = 0.2 \text{ }\mu\text{m}$
- $M_D = 0.75 \text{ TeV}$ for $n = 5 \Leftrightarrow R = 400 \text{ fm}$

■ CDF:

- $M_D = 1.33 \text{ TeV}$, $n = 2 \Leftrightarrow R = 0.27 \text{ }\mu\text{m}$
- $M_D = 0.88 \text{ TeV}$ for $n = 6 \Leftrightarrow R = 31 \text{ fm}$

■ D0 (ll, gg):

- $M_D = 1.23 \text{ TeV}$ lower limit

Astrophysical and Cosmological Constraints

hep-ph/0304029, hep-ph/0309173, hep-ph/0307228

- Places the most stringent lower limits on M_D in ADD
- Supernova cooling due to KK Graviton emission
 - SN 1987A did not emit more KK G than compatible with neutrino signal durations observed by Kamiokande and IMB places the limits: $M_D > 27 \text{ (2.4) TeV for } n = 2 \text{ (3)}.$
- Energetic Gamma Ray Experiment Telescope (EGRET)
 - Cosmic γ -ray-bkg:
 - $M_D > 70 \text{ (5) TeV for } n = 2 \text{ (3)}$
 - Neutron star halo of 100 MeV γ -rays:
 - $M_D > 97, 8, 1.5 \text{ TeV for } n = 2, 3, 4$
 - All neutron stars in the galactic bulge:
 - $M_D > 1130, 57, 7, 1.8 \text{ TeV for } n = 2, 3, 4, 5$
- Neutron star heating:
 - $M_D > 1760, 77, 9, 2 \text{ TeV for } n = 2, 3, 4, 5$
- Ultra high-energy cosmic-ray neutrinos:
 - lower bound $M_D = 1 \text{ to } 1.4 \text{ TeV, } n = 4 \text{ to } 7$