

Track-based alignment of the ATLAS Inner Detector: Run I performance & new extensions for the next physics data run

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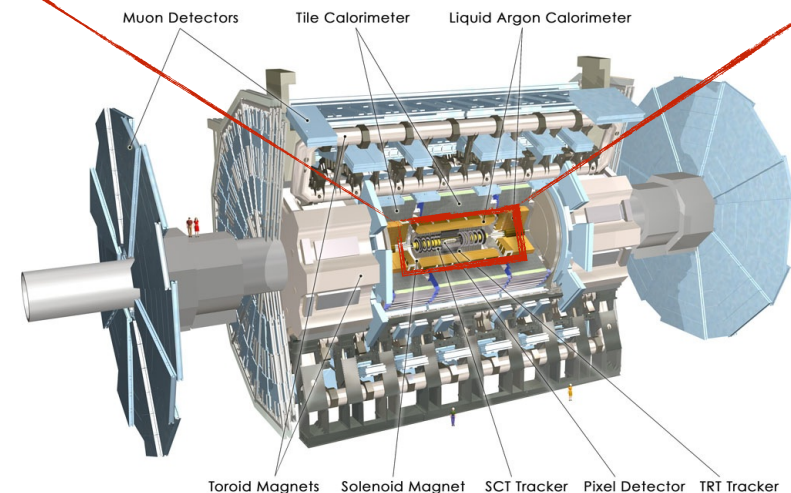
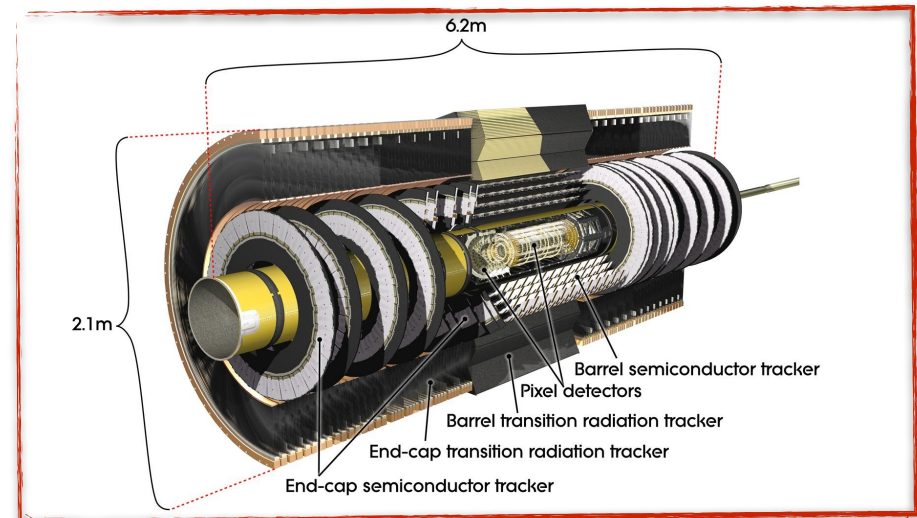
The ATLAS detector



Outline:

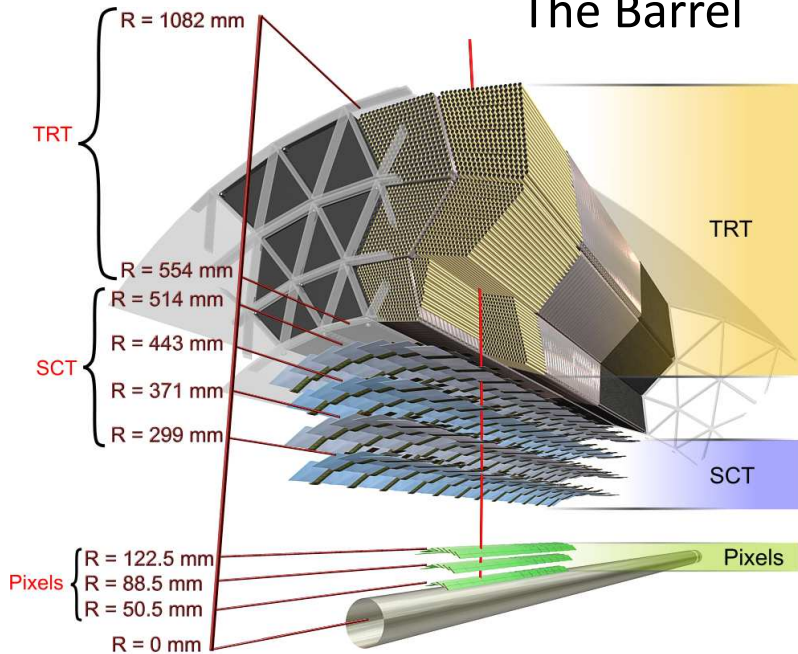
- Introduction
- Why is alignment a problem
- Methodology - track based alignment
- Run I results and performance
- Integration of IBL
- Looking ahead — Run II

Inner detector: Pixel/SCT/TRT

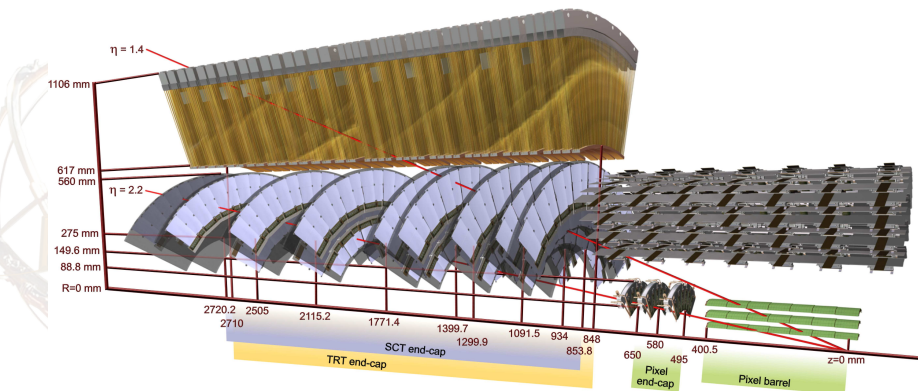


Introduction

The Barrel



The Endcaps



The Inner Detector:

Requirements:

- good p_T resolution ($<30\%$ at 500 GeV p_T)
- good impact parameter resolution \rightarrow b-tagging performance

Alignment challenges:

- huge number of d.o.f.
PIX: 1744 SCT: 4088 TRT: 350848
- Different scales of resolution
PIX: $O(10 \mu\text{m})$
SCT: $O(30 \mu\text{m})$
TRT: $O(140 \mu\text{m})$

Track based alignment algorithm

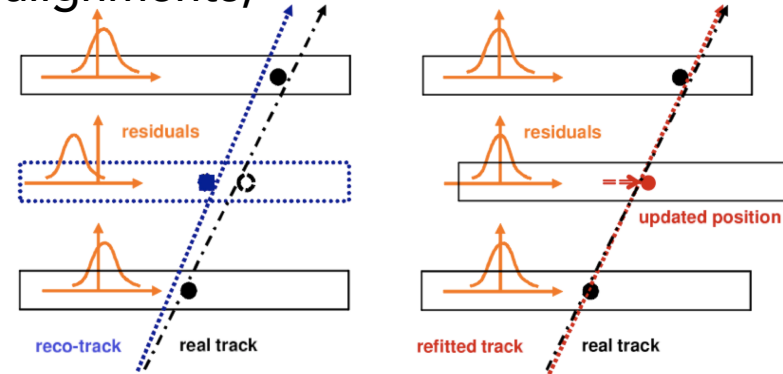


The χ^2 of the track fit is sensitive to detector misalignments, i.e. for a single hit:

$$\chi_i^2 = \left(\frac{|\mathbf{x}_i^{meas} - \mathbf{x}_i^{fit}|}{\sigma_i} \right)^2 = \left(\frac{r(\vec{\tau}, \mathbf{a})}{\sigma_i} \right)^2$$

The residual depends on:

- Track parameters τ
- Alignment parameters \mathbf{a} (3 translations + 3 rotations)



Collect large number of events (tracks) \rightarrow sum over all hits:

- The χ^2 is minimized when detector elements are at true position
- Approx. with Taylor expansion around $\mathbf{a}_0 \rightarrow$ correction $\Delta \mathbf{a}$

$$\frac{d\chi^2}{d\mathbf{a}} = 0$$

Global χ^2 : A simultaneous optimization (fit) of both track parameters and detector element positions is performed

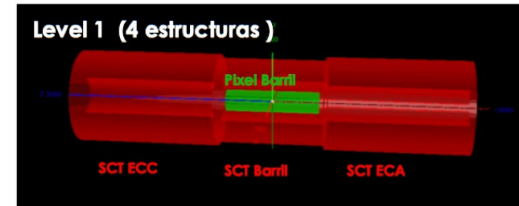
Local χ^2 : After fitting tracks, attempt is made to match detector positions accordingly (inherently iterative)

Alignment Levels

Multistage alignment following the assembly structure:

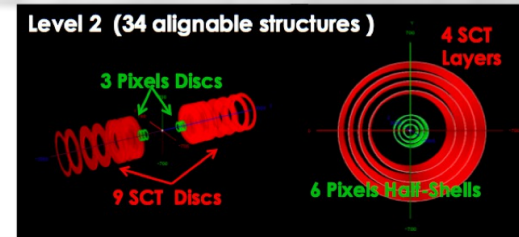
Level 1 — Rigid Body:

- run by run alignment to monitor environmental changes



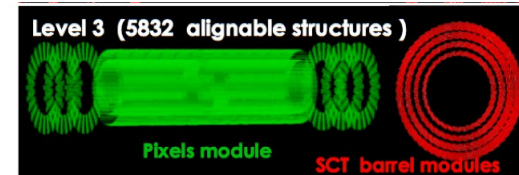
Level 2 — Disks/Staves:

- generally enough to achieve "good" precision -> residuals minimized



Level 3 — Si-Modules/TRT-straws:

- large statistics needed
—> used to eliminate "weak modes"



	structures			Corr. Size
	pixel	SCT	TRT	µm
Level 1	1	3	3	1000
Level 2	12	22	96	100
Level 3	1744	4088	350848	10

	strucures		DoFs
	pixel	SCT	#
Level 1	1	3	24
Level 2	12	22	204
Level 3	1744	4088	34992

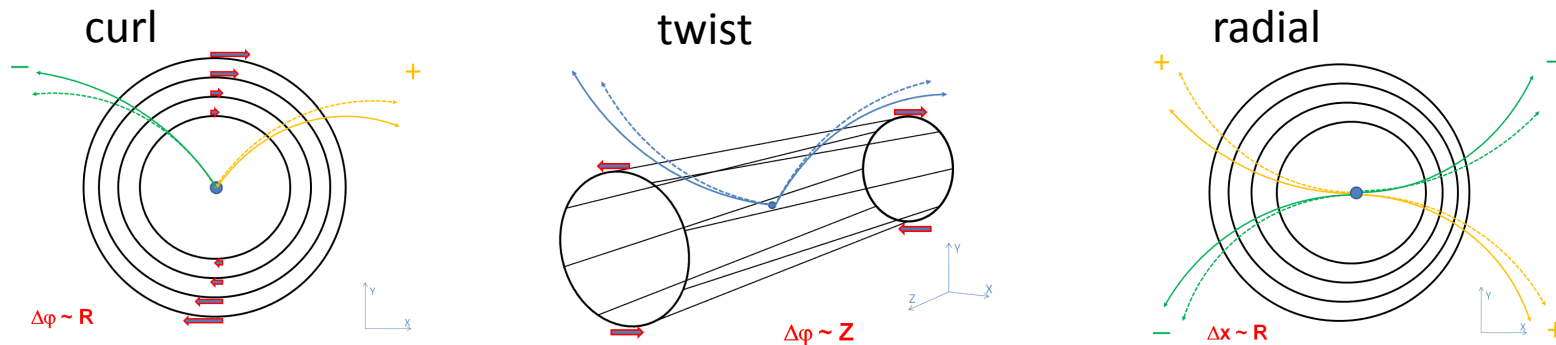
	TRT		DoFs
	Barrel	End cap	#
Level 1	1	2	18
Level 2	96	80	1056
straw	105088	245760	701696

Weak Modes & constraints



Global distortions which preserve the helical trajectory of tracks and leave the χ^2 unchanged are known as “weak modes”:

- The distortions are difficult to remove by the alignment algorithms



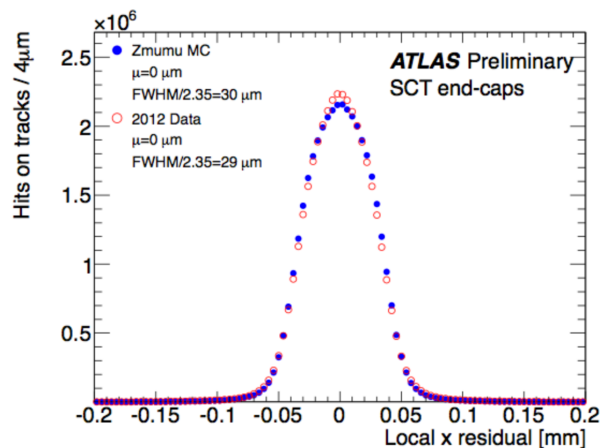
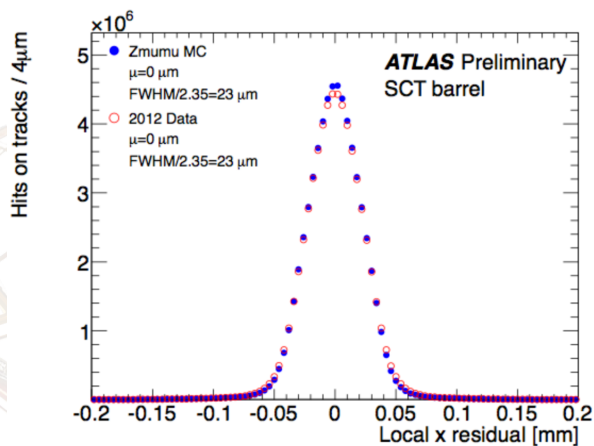
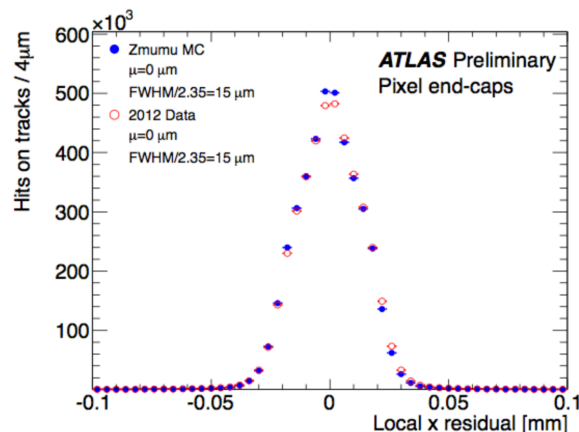
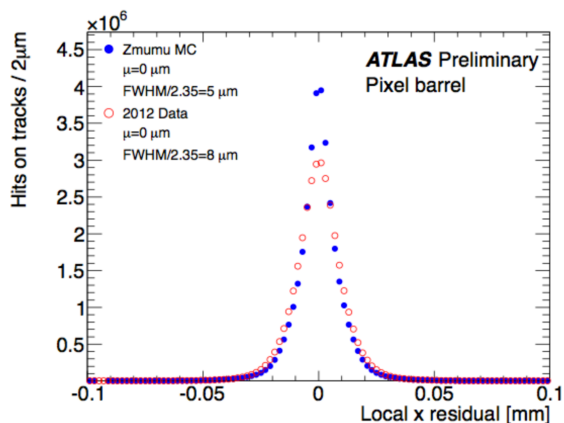
Extra data is necessary to constrain these modes:

- Beam spot constraint (used as additional track constraint)
- Mass of resonances (M_Z, K_S^0)
- Use of external detectors (E/P)
- Other event classes, i.e. Cosmic data (through-going tracks)

Alignment results: Residuals (Silicon)



- Residuals match well between data and MC
- Indicates the algorithms are working correctly
- Remaining difference most likely not due to misalignment alone (e.g. intrinsic detector resolution model)



Local x residual distributions:

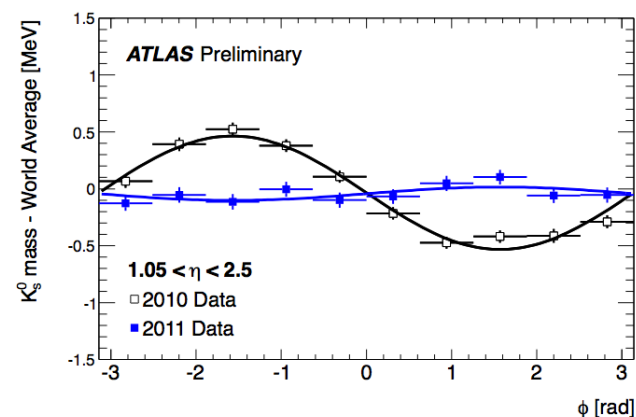
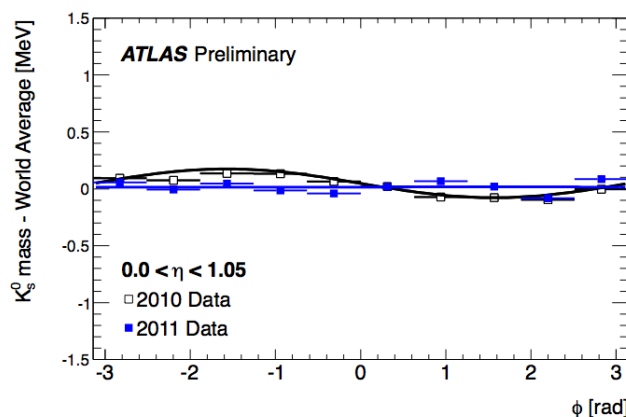
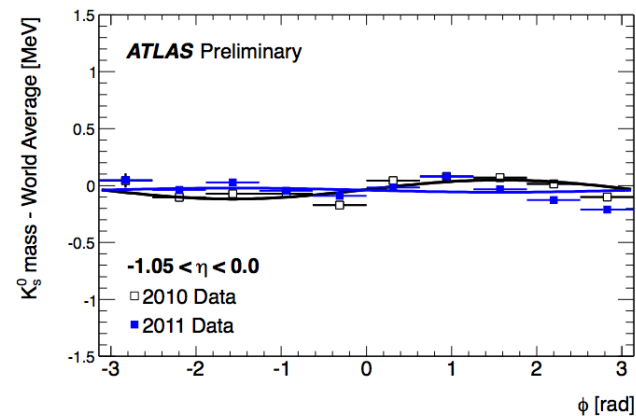
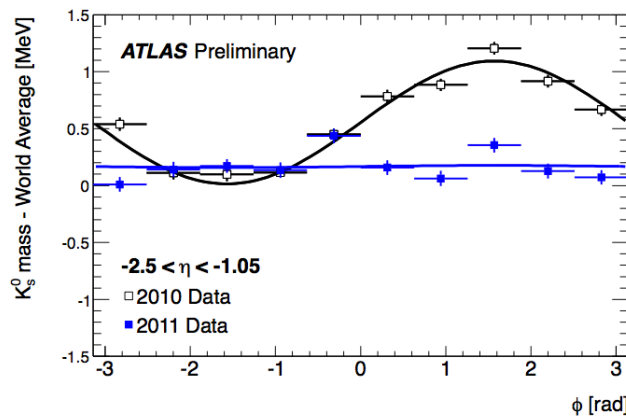
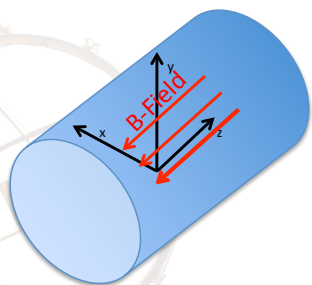
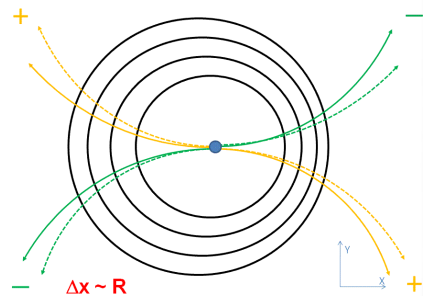
The Pixel & SCT local x residual distribution for the $Z \rightarrow \mu\mu$ data sample reconstructed with 2012 alignment constants, compared to MC simulations. Distributions are integrated over all hits-on-tracks

Alignment results: Radial distortions

Fit of the B-field direction from data:

- A tilt of the solenoid field was found as a bias in the K_s and J/ψ masses vs. ϕ
- Corrected by rotating the magnetic field in reconstruction by $+0.55$ mrad around the x-axis

lowest mode of radial distortion
 → charge symmetric

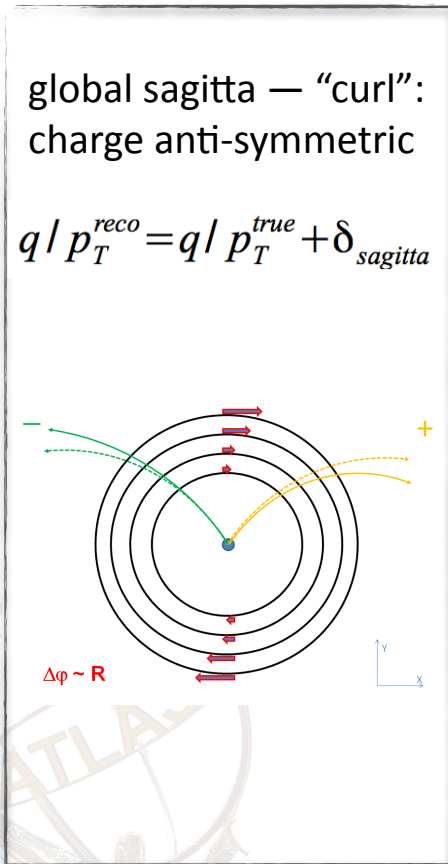


Alignment results: Sagitta bias



Momentum bias can be monitored:

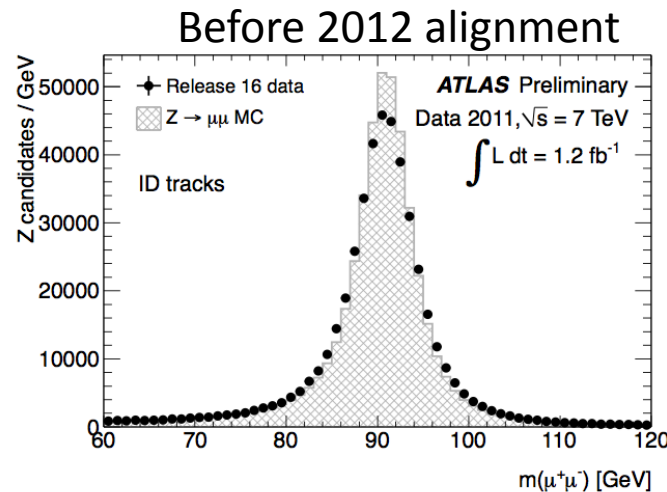
- Cosmic data
- Mass resonances ($Z \rightarrow \mu\mu$) & E/p for opposite charge tracks



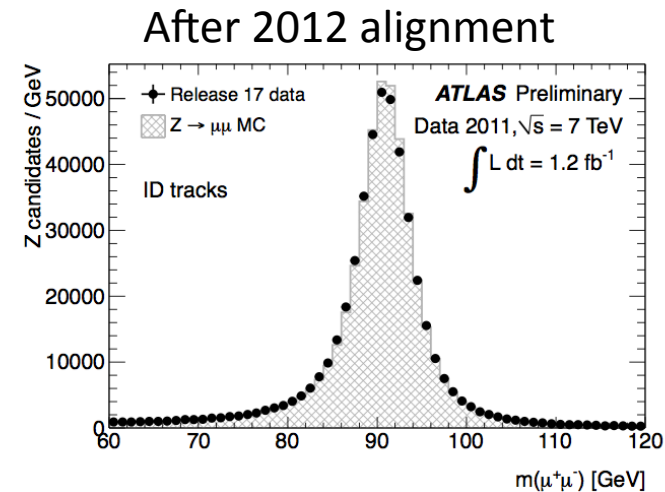
$$\langle E/p \rangle^\pm \rightarrow \langle E/p \rangle^\pm \pm \langle E_T \rangle \delta_{sagitta}$$

$$\langle E/p \rangle_{true}^+ = \langle E/p \rangle_{true}^- \longrightarrow$$

$$\delta_{sagitta} = \frac{\langle E/p \rangle_{rec}^+ - \langle E/p \rangle_{rec}^-}{2 \langle E_T \rangle}$$



(a) All η regions



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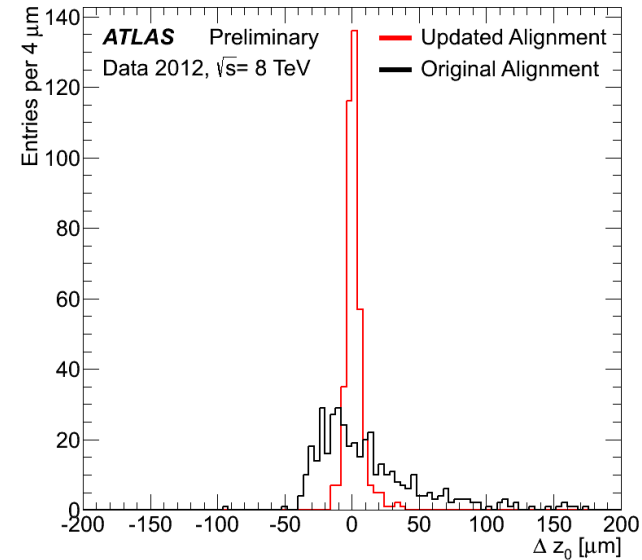
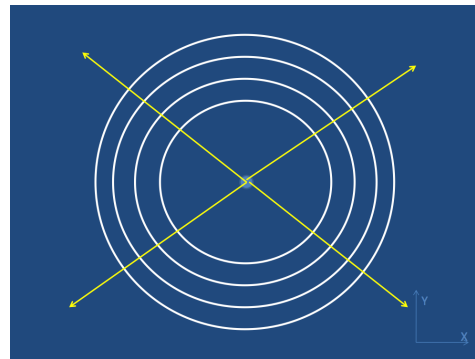
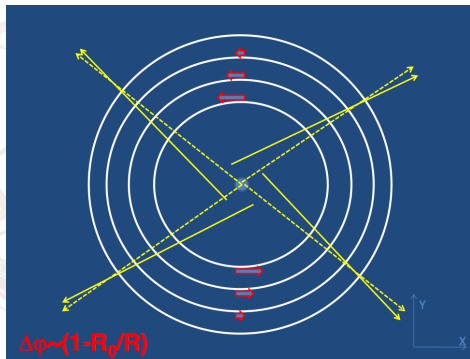
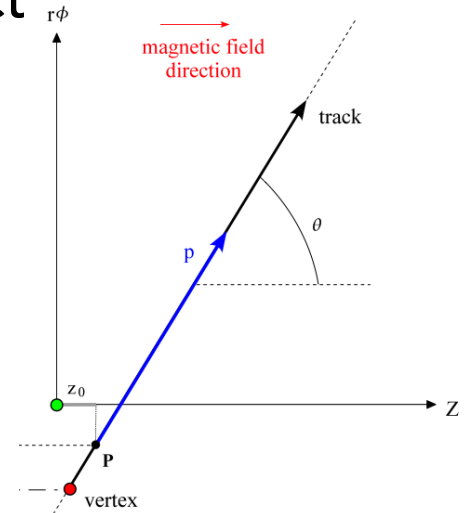
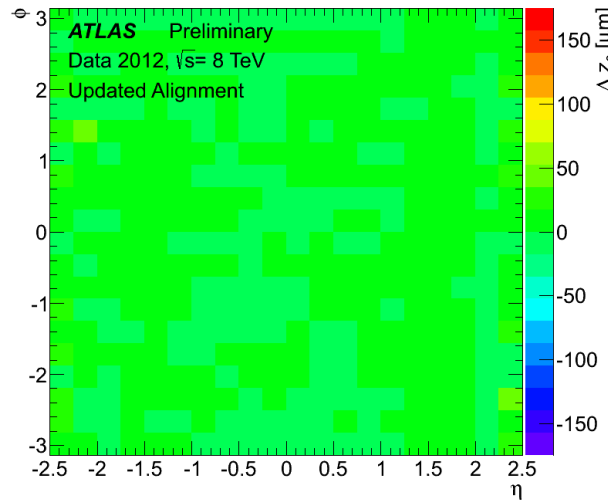
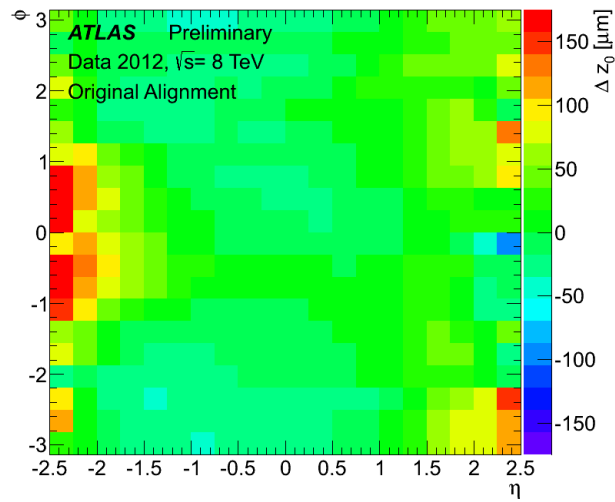
Alignment results: Impact parameter

Resolution of the transverse and longitudinal impact parameters (here: example of transverse):

- Sagitta bias corrected with Z \rightarrow $\mu\mu$ & E/p methods

Before 2012 alignment

After 2012 alignment

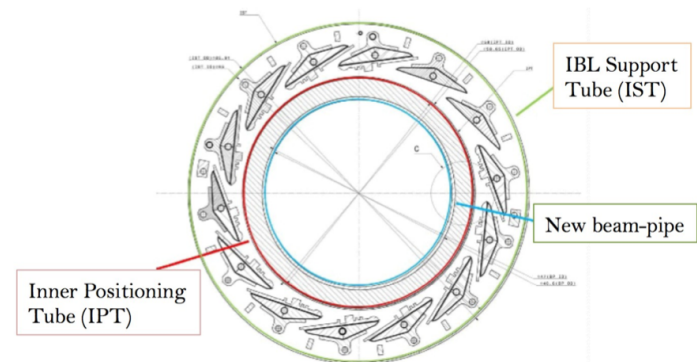
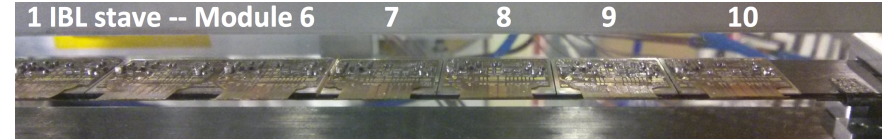


ID-upgrade for Run II — IBL



The new IBL poses a new challenge to the ID alignment

- New mechanically independent structure, integrated into Pixel detector
- 14 staves with average radius of 3.3 cm
- Pixel technology & novel 3D sensors
- Level 1 alignment has one additional independent structure
- Levels 2 & 3 more straight forward (additional layers & models)
- Software integration finished:
 - Monitoring
 - Data base structure

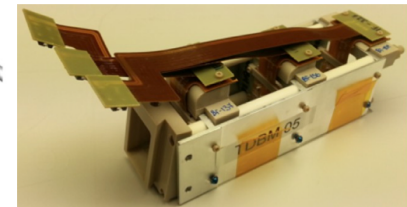
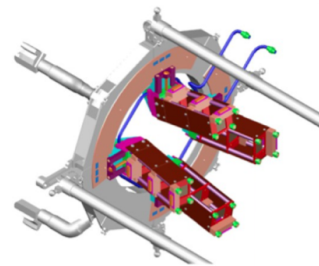


Run II alignment closure tests being performed:

- misalignment is introduced in MC
- Δa corrections recovered

Diamond Beam Monitor (DBM):

- Telescopes of diamond detectors in forward regions



—> see next talk for more details

Run I performance:

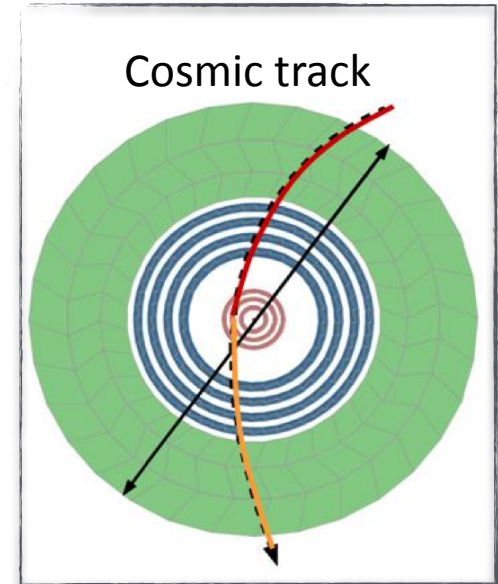
- Final alignment constants show very good agreement between data & MC in physics observables
- Physical detector movements tracked down
- Many weak modes have been eliminated

Run II preparations:

- Alignment framework prepared
- IBL fully integrated
- Investigate possibility to run level 2 alignment in Calibration loop
- More detailed modelling & integration of individual module distortions

Pre Run II alignment:

- First rough alignment with Cosmic data (expected 100 mHz event rate in IBL)
—> serves as input for first collision data alignment for Run II



Backup

