

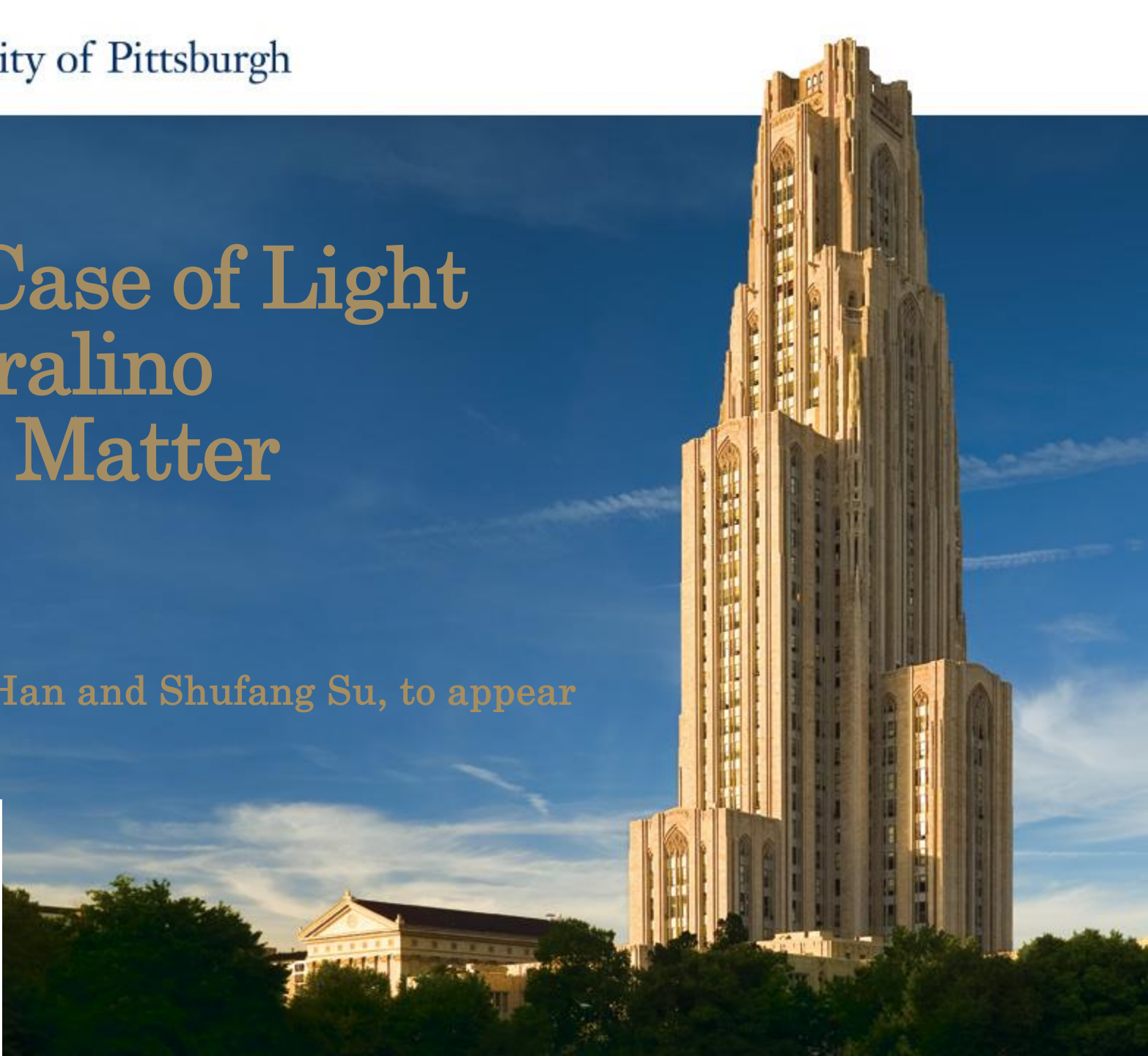


University of Pittsburgh

# The Case of Light Neutralino Dark Matter

Zhen Liu

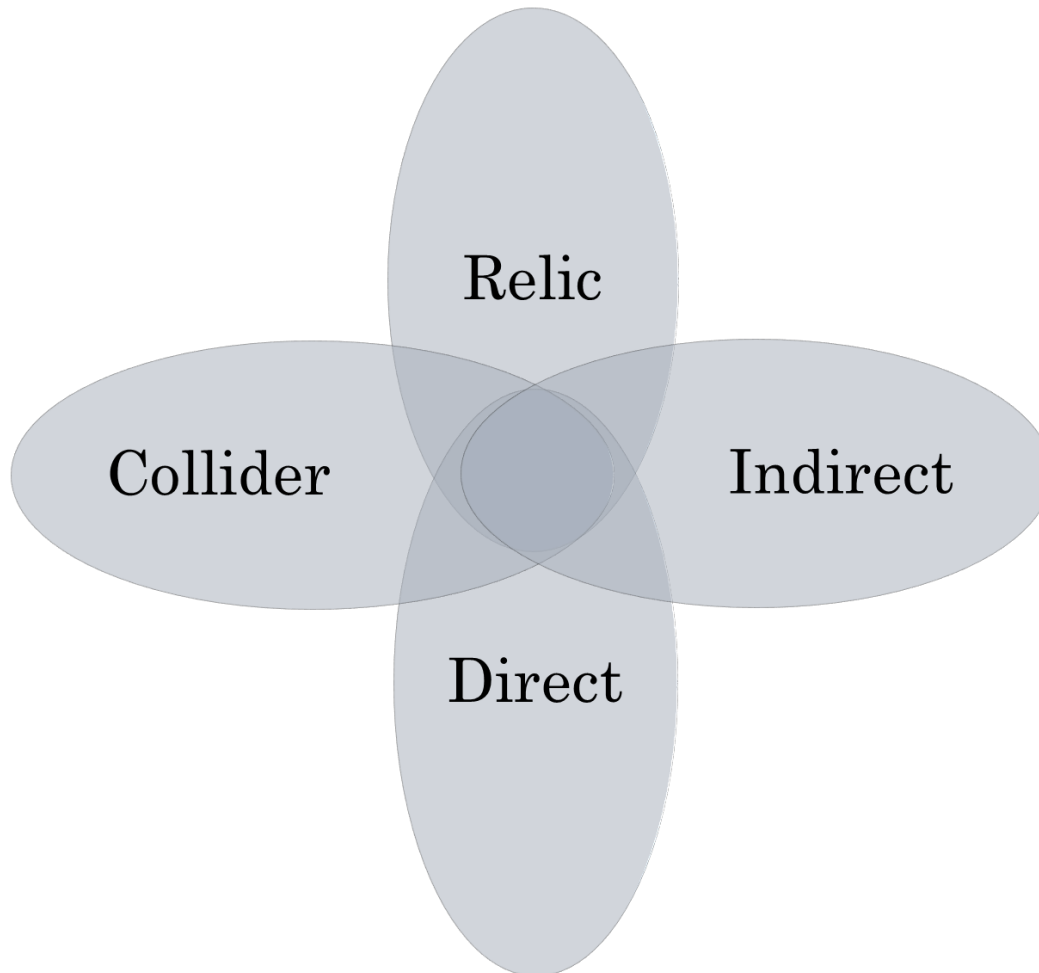
with Tao Han and Shufang Su, to appear



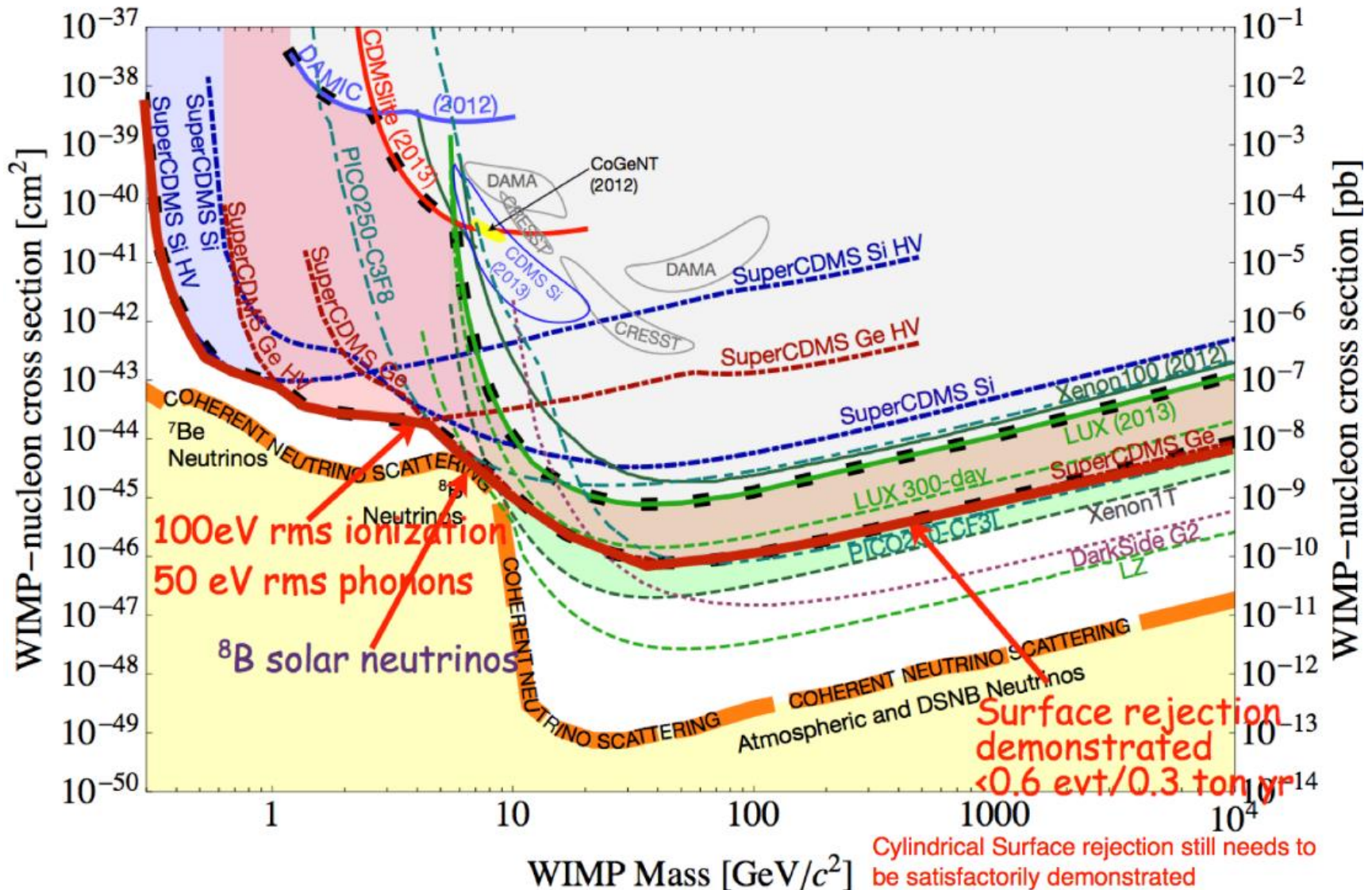
# Motivation

Particle DM is well-motivated!

We explore DM from **ALL DIRECTIONS** with unprecedented (and growing) precision!

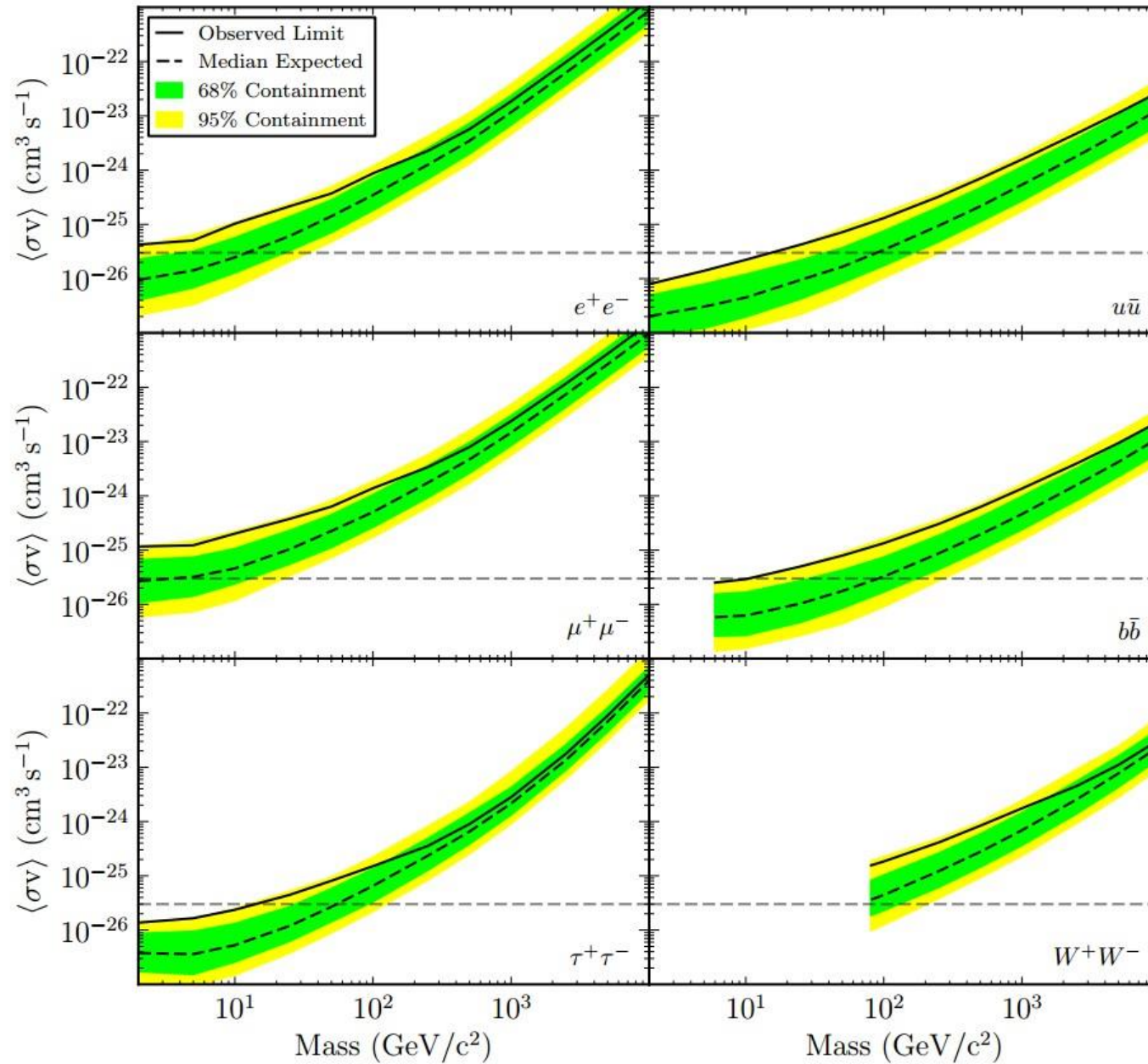


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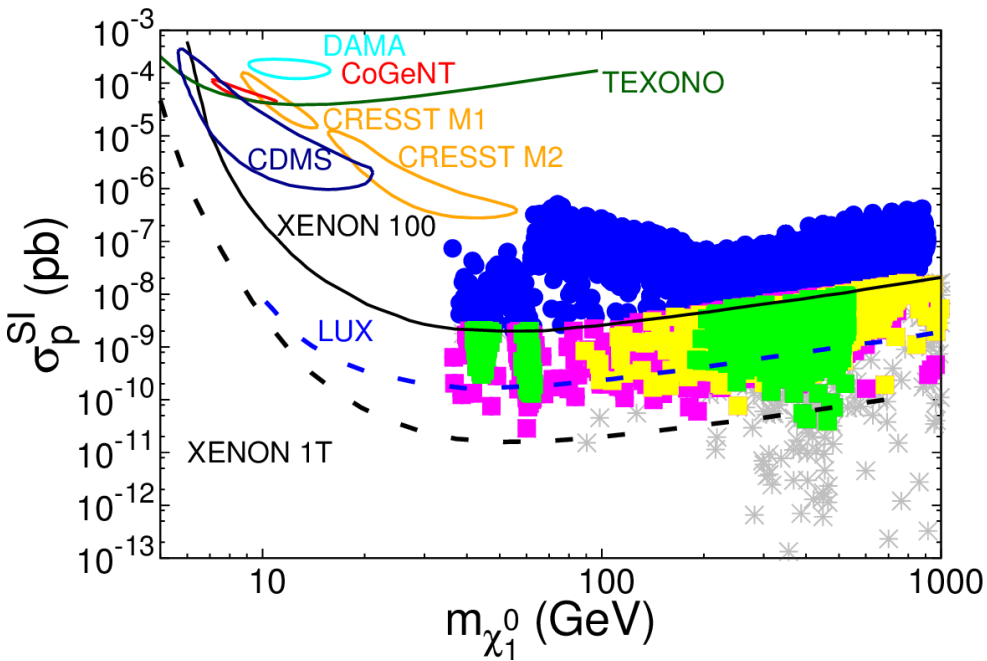
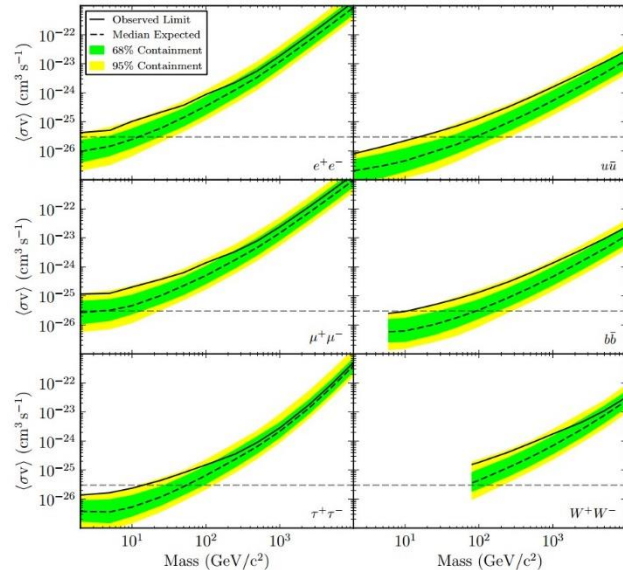
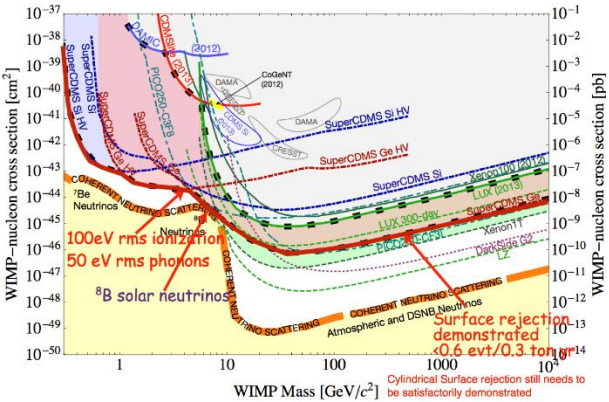


Recent results and plans of CDMS, talk by Bernard Sadoulet

# Motivation



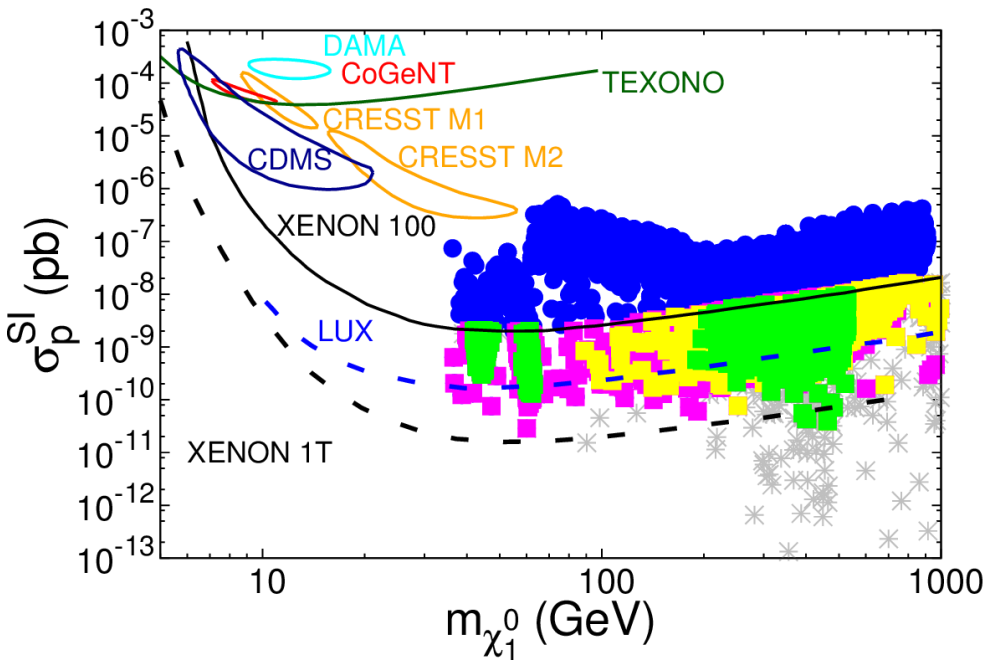
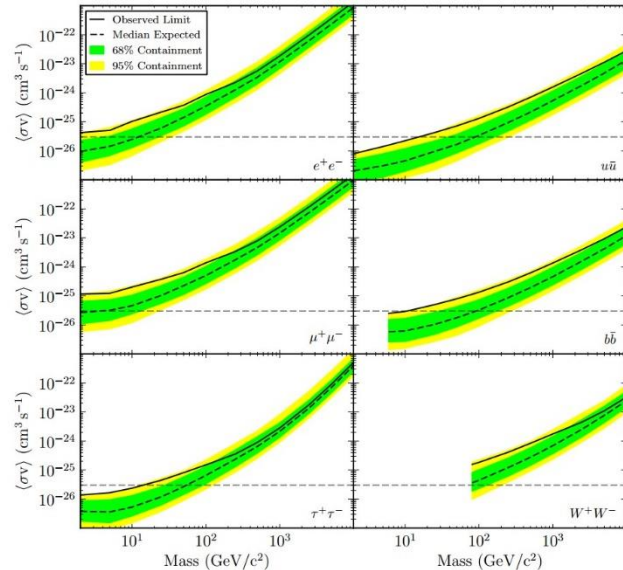
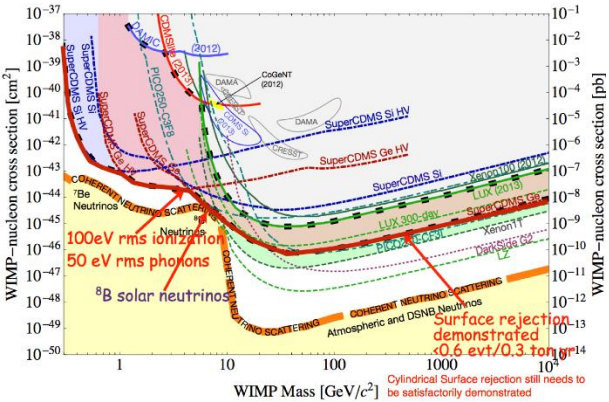
# Motivation



Collider Searches from Mono-jet, photon, Z, b... mono-everything!  
 Collider indirect probes with model assumed, Higgs precision/Higgcision.

T.Han, ZL, A. Natarajan 1303.3040

# Motivation



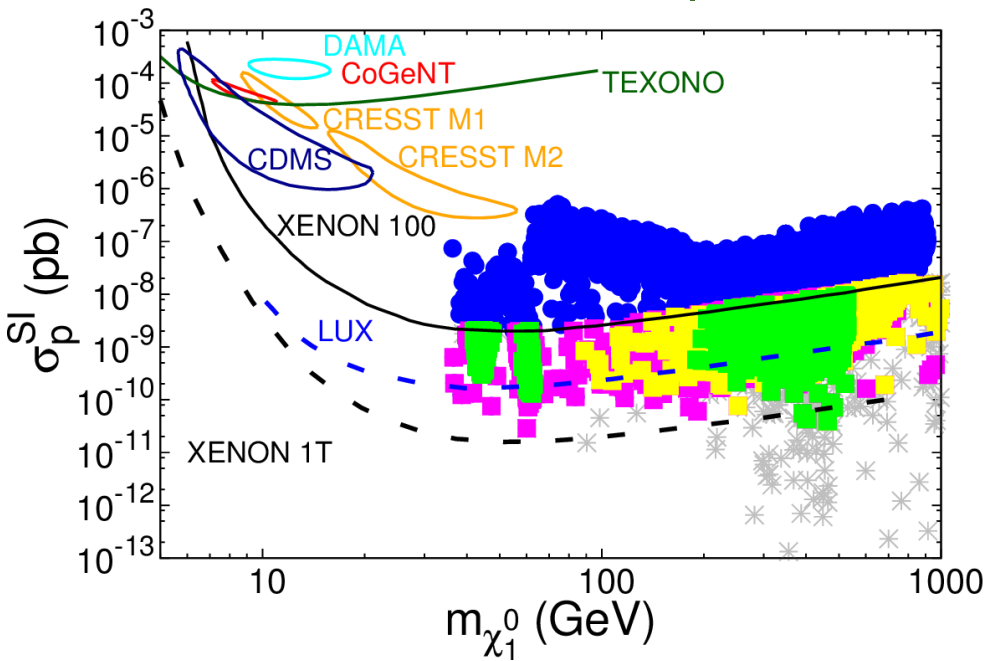
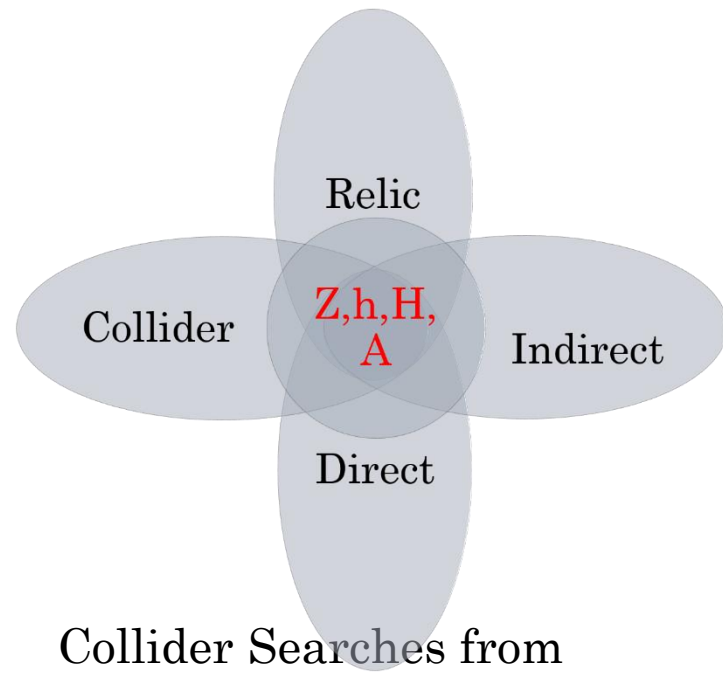
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“the most beloved” SUSY models.  
highly constrained  $\Leftrightarrow$  highly predictable.

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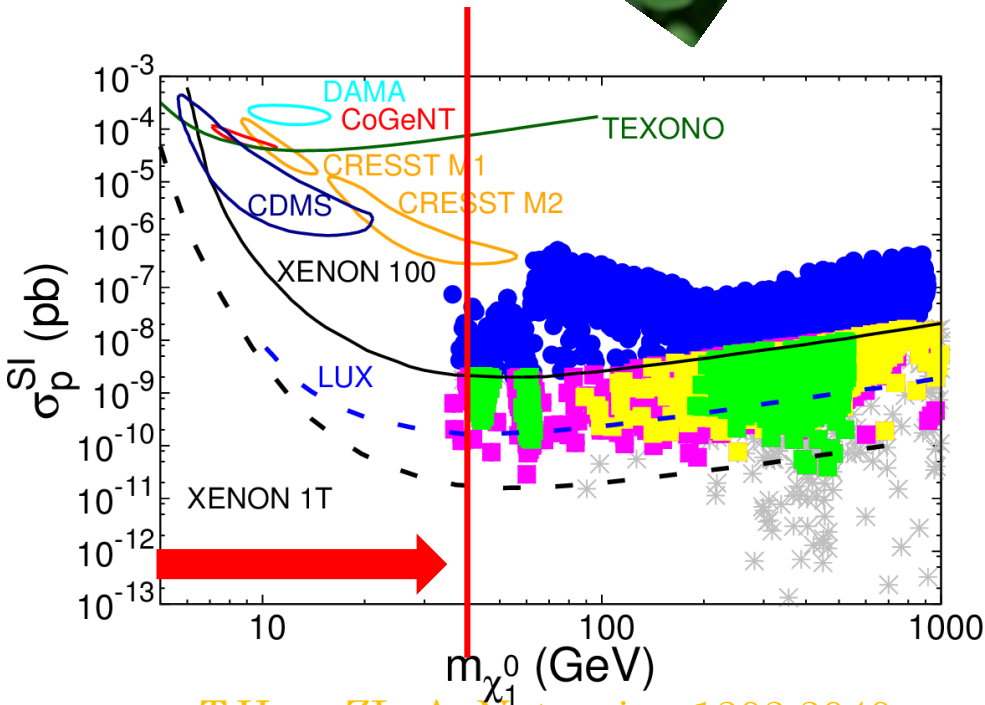


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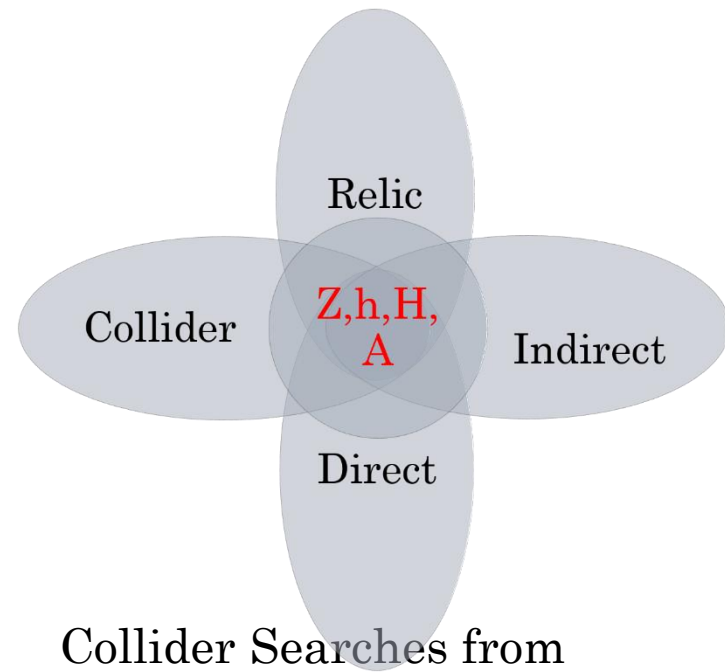
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T.Han, ZL, A. Natarajan 1303.3040



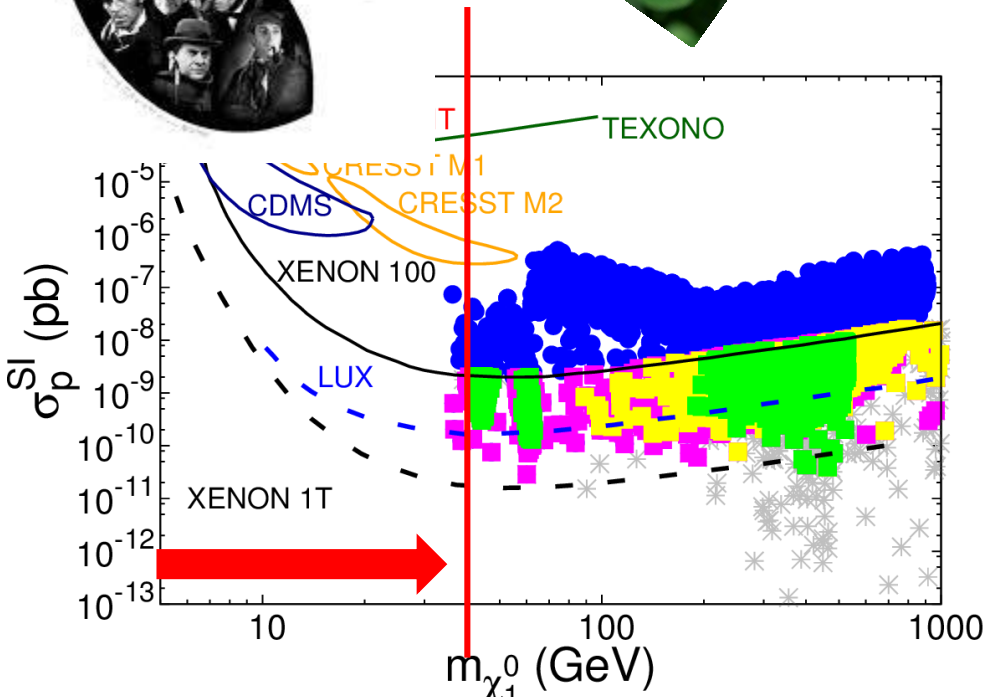
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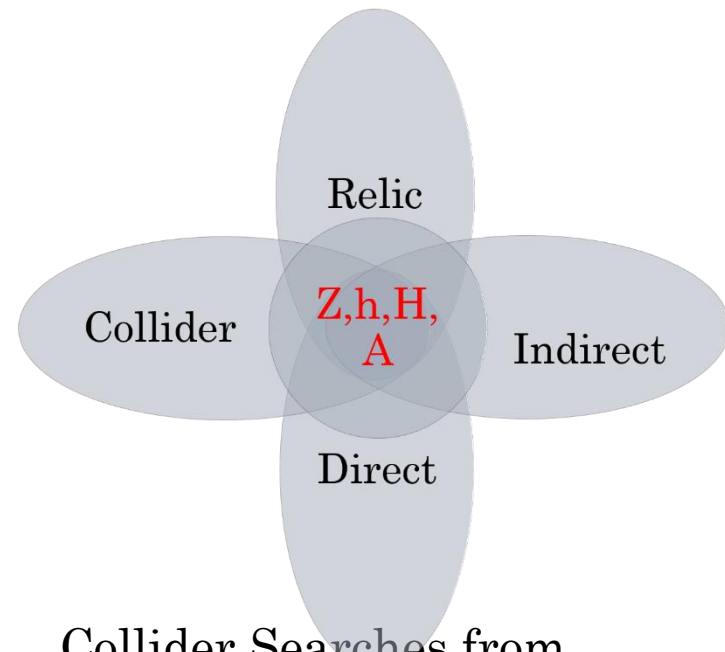
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# Motivation



T.Han, ZL, A. Natarajan 1303.3040



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# Neutralino LSP in the NMSSM

Neutralino LSP serves as good DM candidate in R-parity conserving SUSY models

A mixture of **Bino**, Wino, Higgsino and **Singlino**

$$\tilde{\chi}_1^0 = N_{11}\tilde{B} + N_{12}\tilde{w}^3 + N_{13}\tilde{h}_d + N_{14}\tilde{h}_u + N_{15}\tilde{S}$$
$$M_{\tilde{\chi}^0} = \begin{pmatrix} M_1 & 0 & -g_1 \frac{v_d}{\sqrt{2}} & g_1 \frac{v_u}{\sqrt{2}} & 0 \\ & M_2 & g_2 \frac{v_d}{\sqrt{2}} & -g_2 \frac{v_u}{\sqrt{2}} & 0 \\ & & 0 & -\mu & -\lambda v_u \\ & & & 0 & -\lambda v_d \\ & & & & 2\frac{\kappa}{\lambda}\mu \end{pmatrix}$$

# Higgs Sector in the NMSSM

## • Type II Two Higgs Doublet Model plus singlet $S$

$$W_{\text{NMSSM}} = Y_u \bar{u} H_u Q + Y_d \bar{d} H_d Q + Y_e \bar{e} H_d L + \lambda S H_u H_d + \frac{1}{3} \kappa S^3$$

$$V_{H,\text{Soft}} = m_{H_u}^2 H_u^\dagger H_u + m_{H_d}^2 H_d^\dagger H_d + M_S^2 |S|^2 + (\lambda A_\lambda (H_u^T \epsilon H_d) S + \frac{1}{3} \kappa A_\kappa S^3 + c.c.)$$

## • SSB

$$H_u = \begin{pmatrix} H_u^+ \\ H_u^0 \end{pmatrix} \rightarrow v_u / \sqrt{2} \quad H_d = \begin{pmatrix} H_d^0 \\ H_d^- \end{pmatrix} \rightarrow v_d / \sqrt{2} \quad S \rightarrow v_s / \sqrt{2} \quad (\mu = \lambda v_s / \sqrt{2})$$

$$v_u^2 + v_d^2 = v^2 = (246 \text{ GeV})^2$$

$$\tan \beta = v_u / v_d$$

after EWSB, 7 physical Higgses

CP-even Higgses:  $H_1, H_2, H_3$

CP-odd Higgses:  $A_1, A_2$

Charged Higgses:  $H^\pm$

For discussion about Low-mass Higgs states, see our work  
N.Christensen, T. Han, ZL and S. Su, 1303.2113

# Light Neutralino LSP

Light Wino/Higgsino unlikely due to

Underabundant/No light Chargino/Direct Detection  
Rate High

Bino and Singlino are **pretty inert**

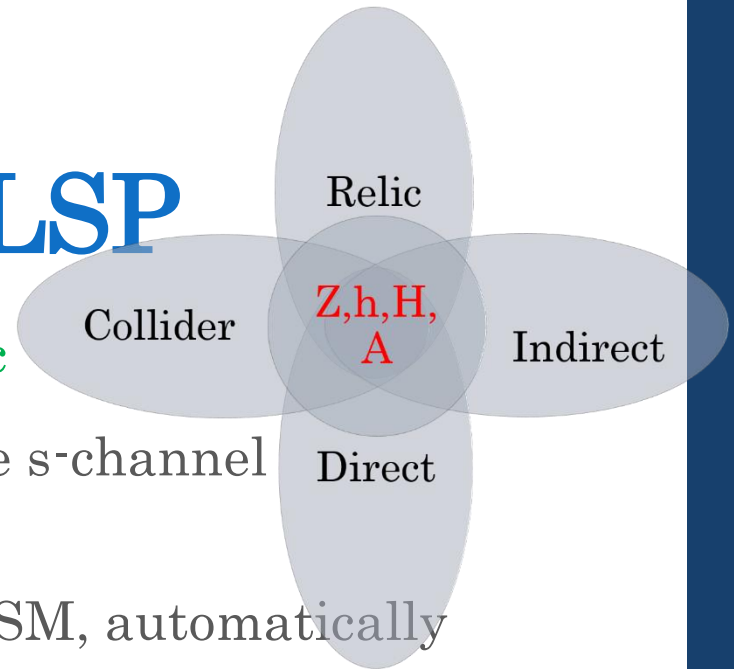
# Light Neutralino LSP

Ways out if assuming thermal relic

Famous “**Funnel**” regions to hit the s-channel resonance for the mediator.

Especially near the PQ-limit NMSSM, automatically light CP-odd singlet like Higgs. [Dark Light Higgs, P.Drapper, T.Liu, C.Wagner, L-T. Wang and H.Zhang](#)

Famous “**Co-annihilation**” regions to have sfermions in thermal equilibrium help LSP annihilate efficiently. [A.Arbey, M.Battaglia and F.Mahmoudi, 1205.2557](#)



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# Finding the Solutions

Using modified  
NMSSMTools4

Multiple way of  
scanning:

**General**

**Dedicated**

**Seeded**

	General	Sbottom	Stau	$H_1, A_1$ -funnel
$m_{A_{tree}}$	[0,3000]	—	—	—
$\tan \beta$	[1,55]	—	—	—
$\mu$	[100,500]	—	—	—
$ A_\kappa $	[0,1000]	—	—	—
$\lambda$	[0,1]	—	—	[0.01,0.6]
$\kappa$	[0,1]	either $\kappa \in [2, 30]\lambda/(2\mu)$ or $M_1 \in [2, 30]$ , or both		
$ M_1 $	[0,500]			
$M_{Q3}, M_{U3}$	[0,3000]	—	—	—
$ A_t $	[0,4000]	—	—	—
$M_{D3}$	[0,3000]	[0,80]	3000	
$ A_b $	[0,4000]	—	0	
$M_{L3}, M_{E3}$	[0,3000]	3000	[0,500]	3000
$ A_\tau $	[0,4000]	0	[0,2000]	0

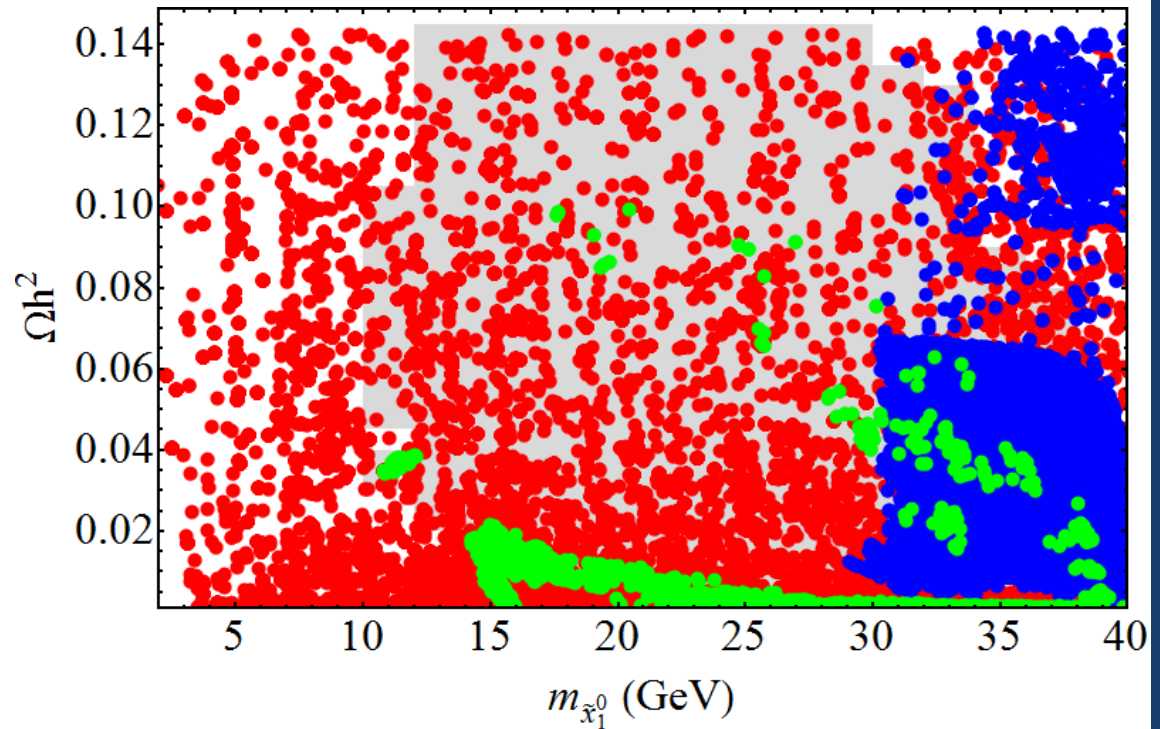
Log prior used  
for several  
dedicated scans

# Finding the Solutions

- Theoretical constraints such as Vacuum stability.
- Collider Higgs search limits from the LEP, the Tevatron and the LHC.
- LEP, Tevatron and LHC constrains on searches of supersymmetric particles, such as charginos, leptons and squarks;
- 2- $\sigma$  window of the SM-like Higgs boson mass: 122.7 – 128.7 GeV (including linearly added estimated theoretical uncertainties of  $\pm 2$  GeV).
- 2- $\sigma$  window of the SM-like Higgs bosons cross sections for  $\gamma\gamma$ ,  $ZZ$ ,  $W^+W^-$ ,  $\tau^+\tau^-$  and  $b\bar{b}$  different production modes.
- Z boson invisible width and hadronic width
- B-physics constrains, including  $b \rightarrow s\gamma$ ,  $B_s \rightarrow \mu^+\mu^-$ ,  $B \rightarrow \chi_s\mu^+\mu^-$  and  $B^+ \rightarrow \tau^+\nu_\tau$ , as well as  $\Delta m_s$ ,  $\Delta m_d$ ,  $m_{\eta_b(1S)}$  and  $\Upsilon(1S) \rightarrow a\gamma, h\gamma$ .



# Relic Abundance:

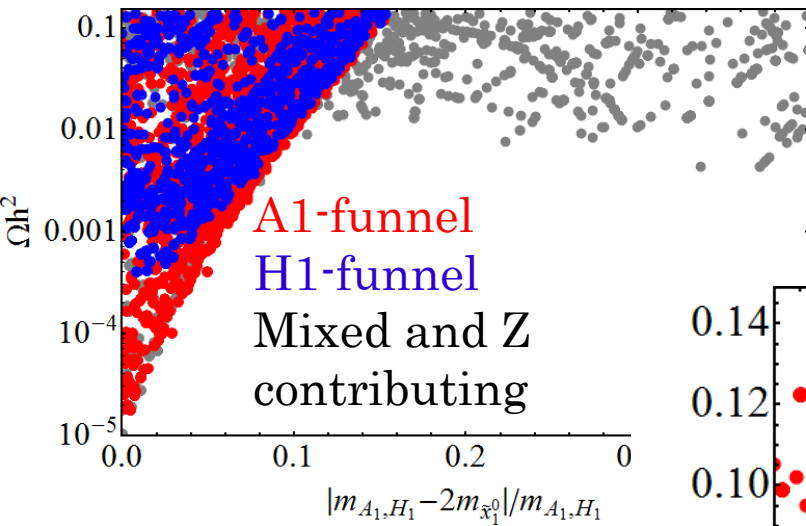


Red A1/H1-funnel

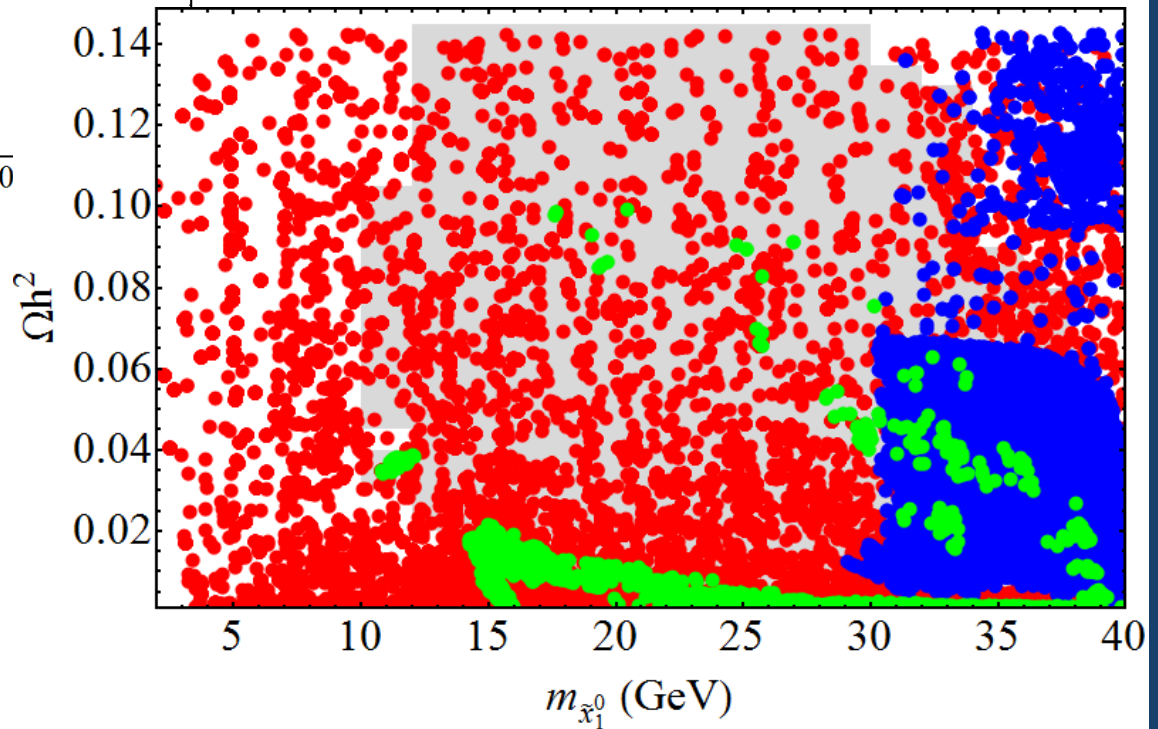
Blue Stau coannihilation

Green Sbottom coannihilation

Shaded Sbottom excluded by direct detection



# Relic Abundance:



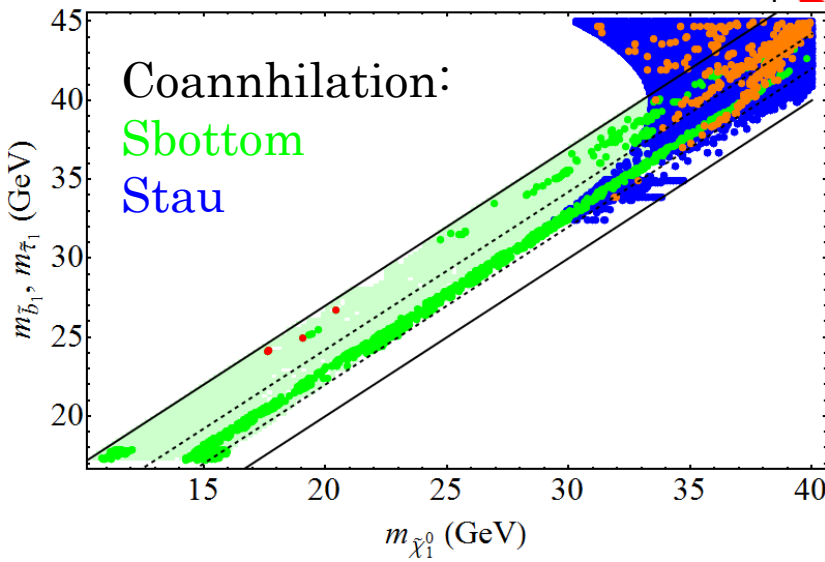
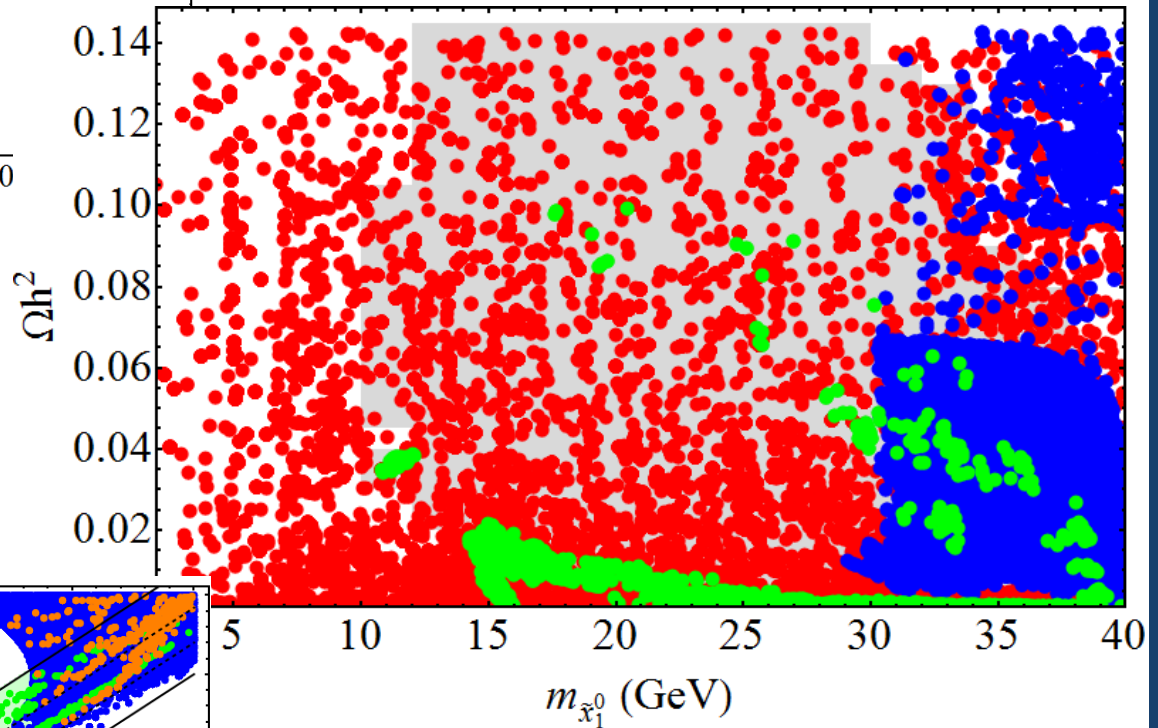
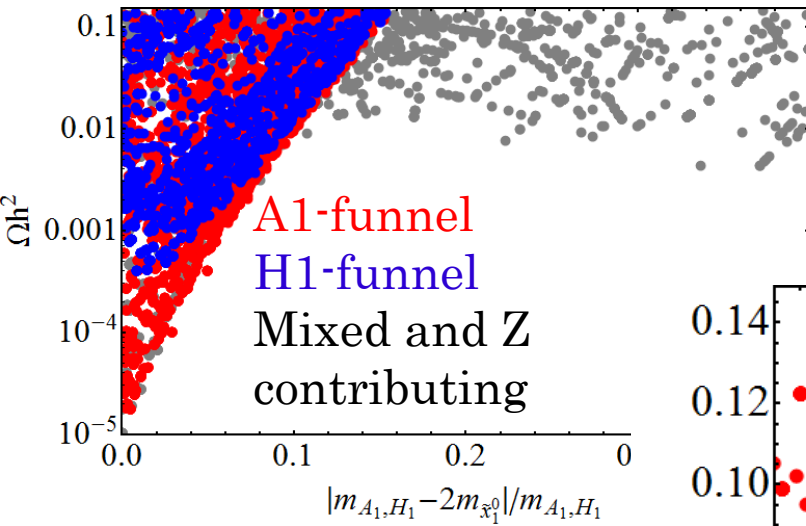
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# Relic Abundance:



LUX/SuperCDMS imposed:

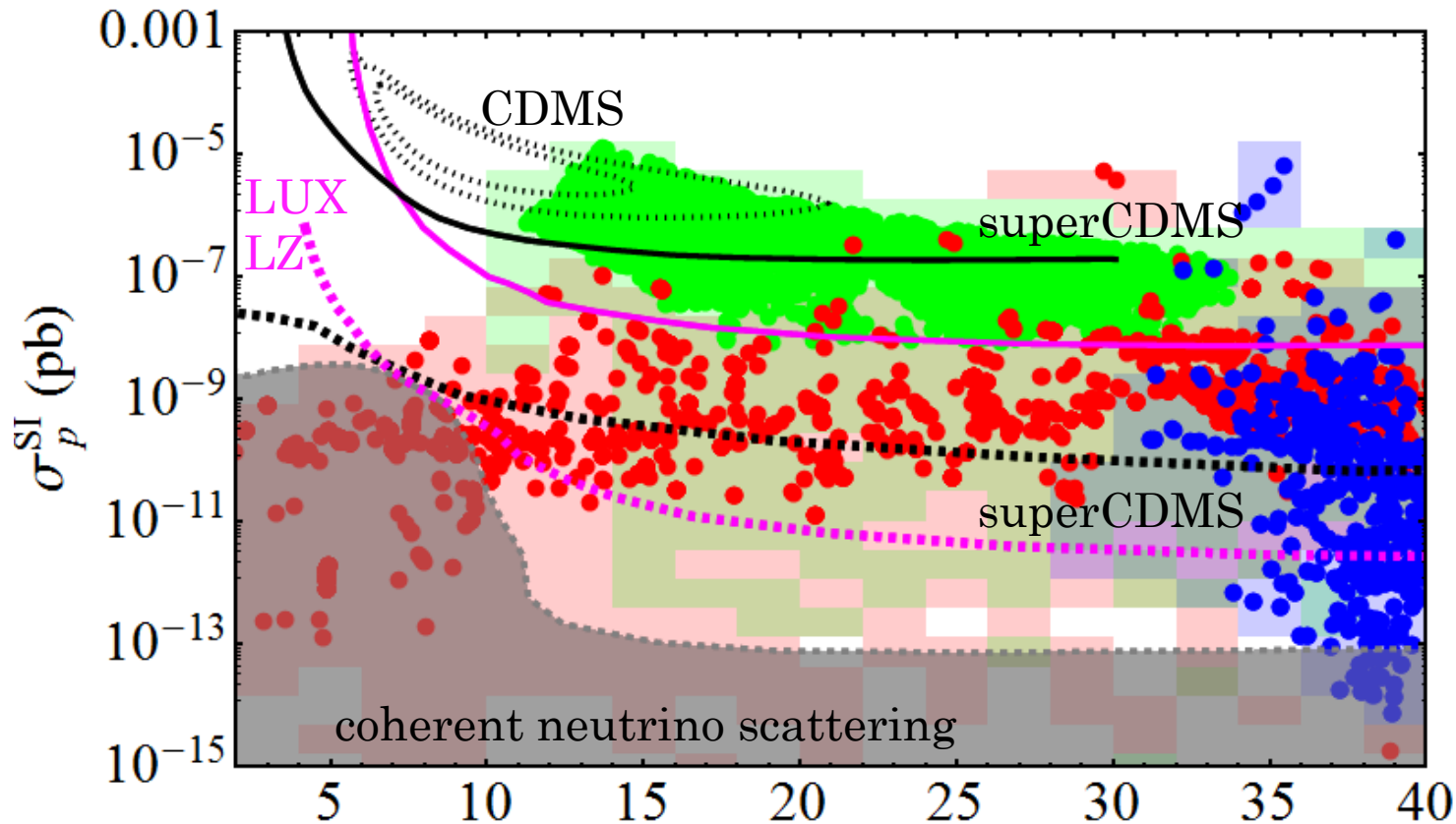
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# Direct Detection



Tight/Loose relic Density required  $m_{\tilde{\chi}_1^0}$  (GeV)

For Solid/Shaded

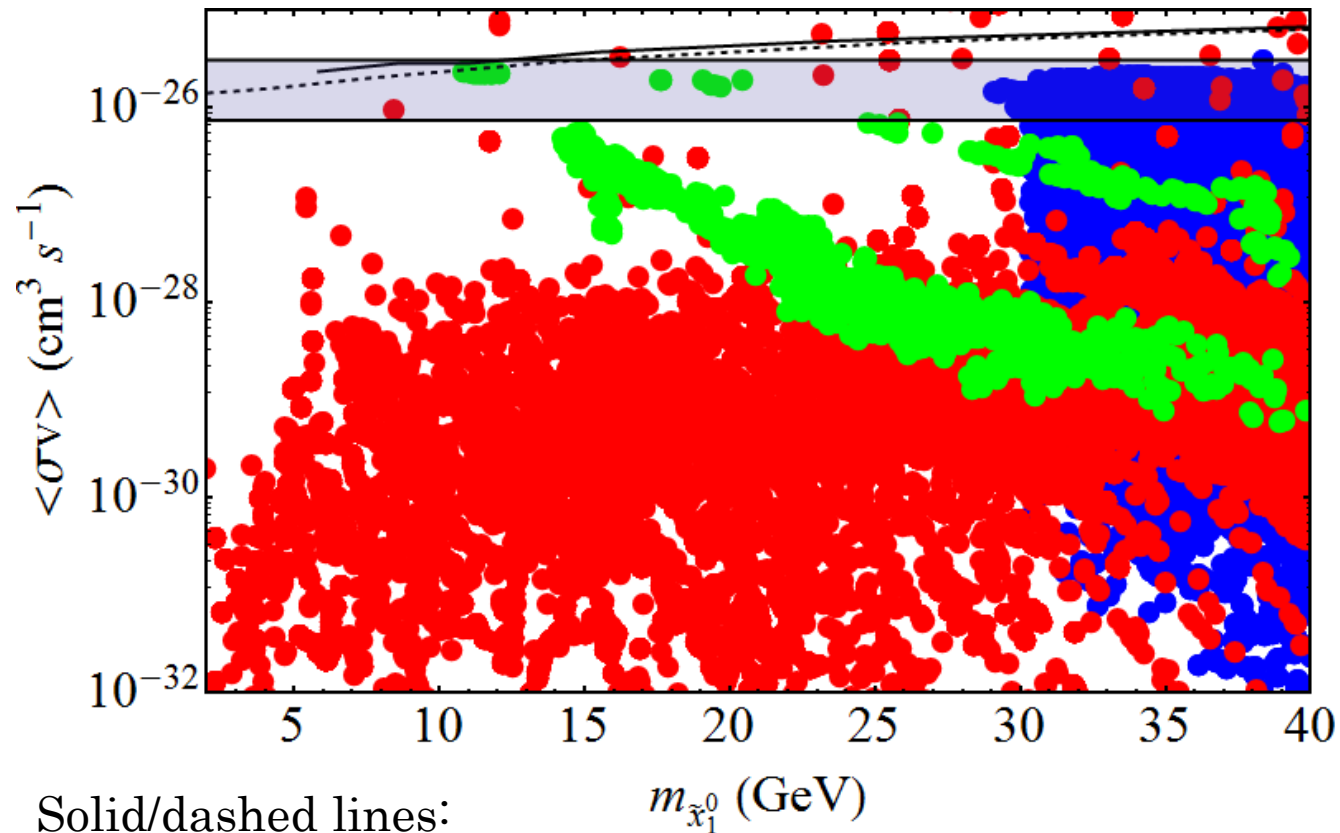
$$0.0947 (0.001) < \Omega_{\tilde{\chi}_1^0} h^2 < 0.142$$

Red A1/H1-funnel

Blue Stau coannihilation

Green Sbottom coannihilation

# Indirect Detection



Solid/dashed lines:

$m_{\tilde{\chi}_1^0}$  (GeV)

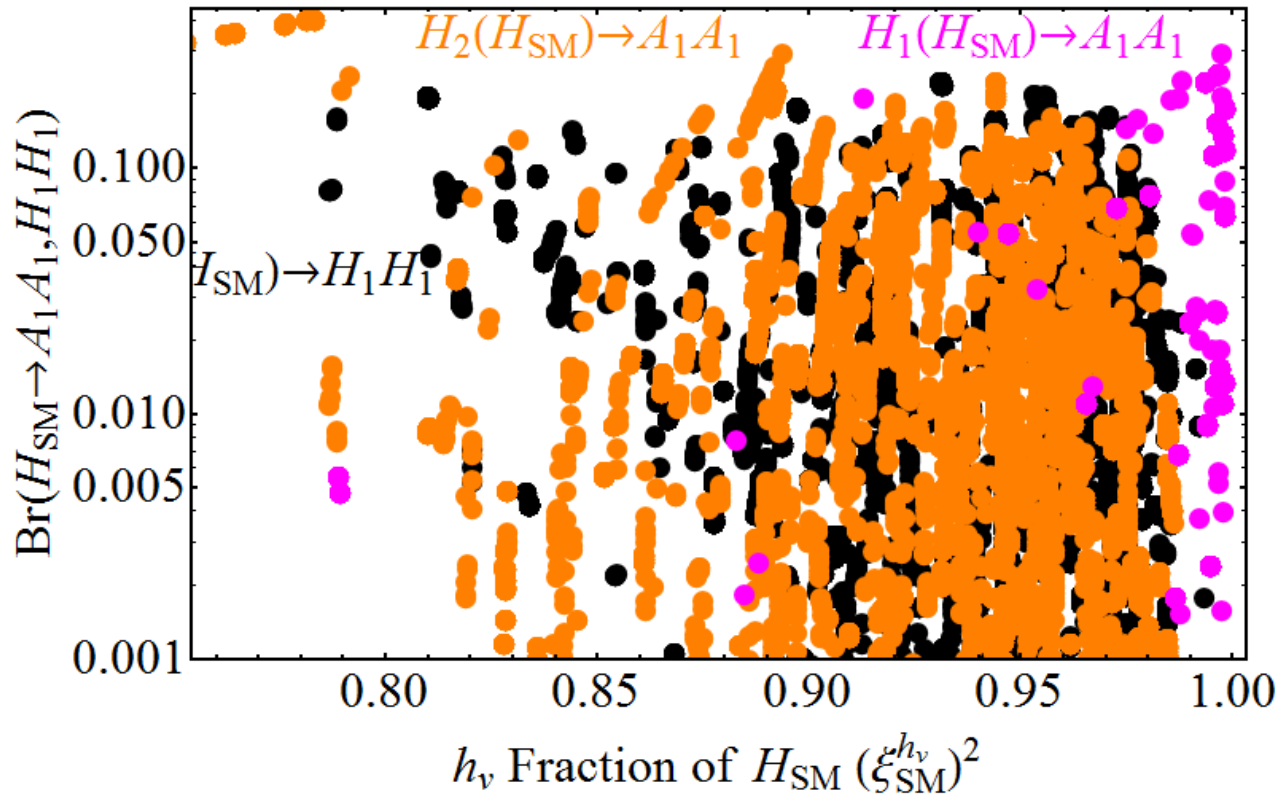
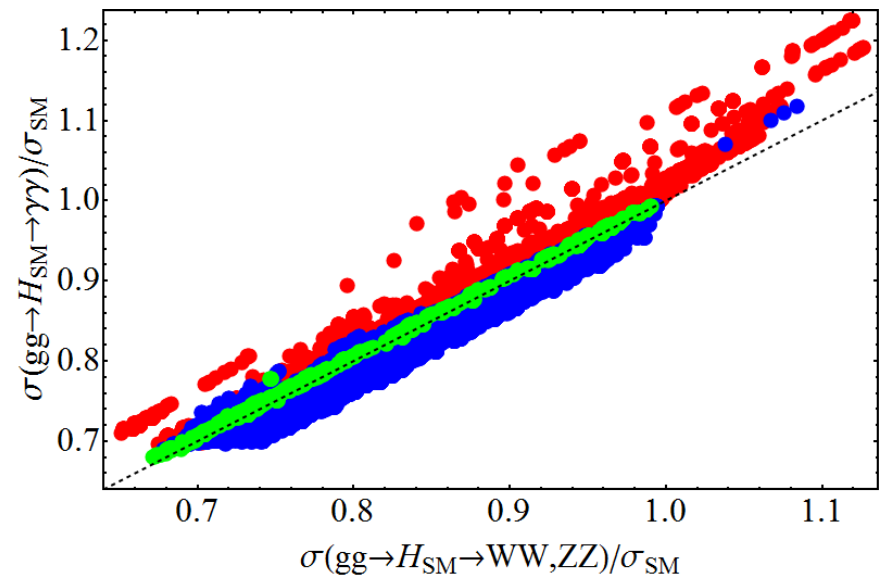
Combined Exclusion from Fermi-LAT from inner galaxy and dwarf galaxies assuming DM annihilate to bottom pair/ tau pair assuming NFW profile

Shaded:

Best fit for GeV Gamma Ray Excess for 35 GeV DM into bb.

# 126 GeV Higgs Boson!

Not much Enhanced Diphoton  
 Stau, sbottom loop does not  
 contribute much in our case as one  
 expected from limits of Br of Higgs.



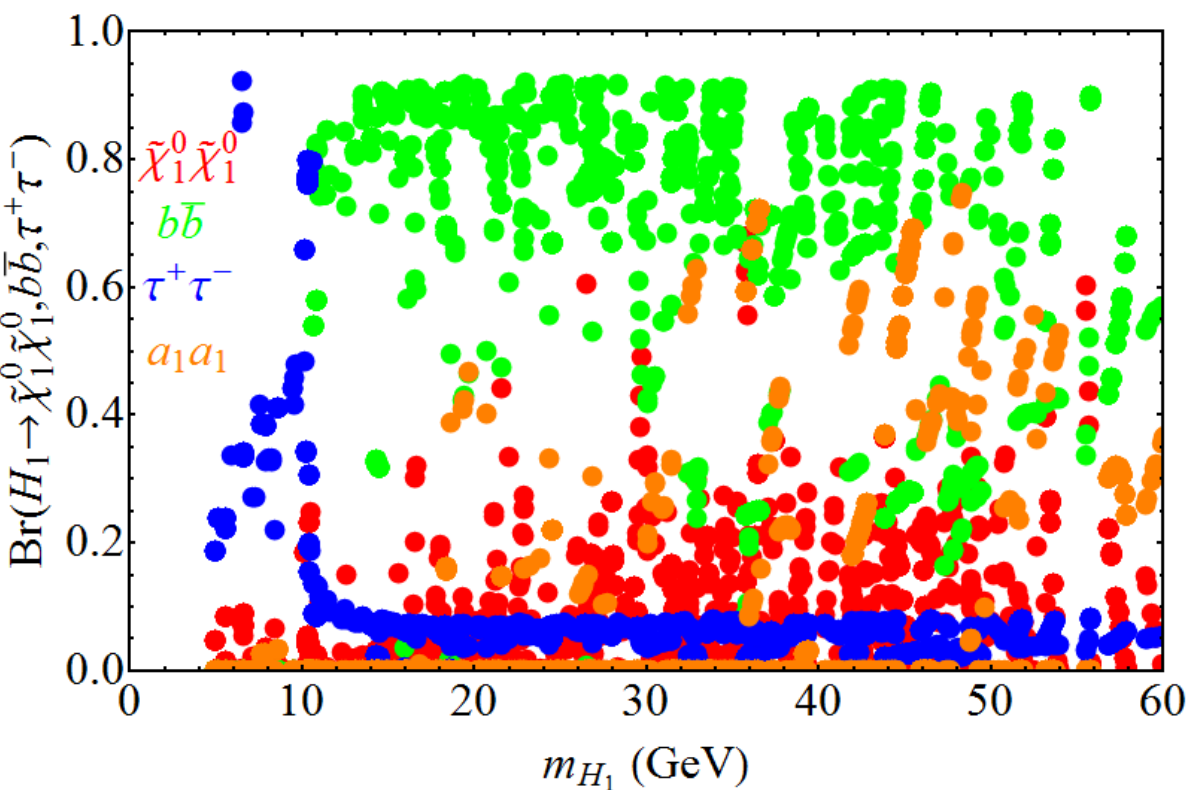
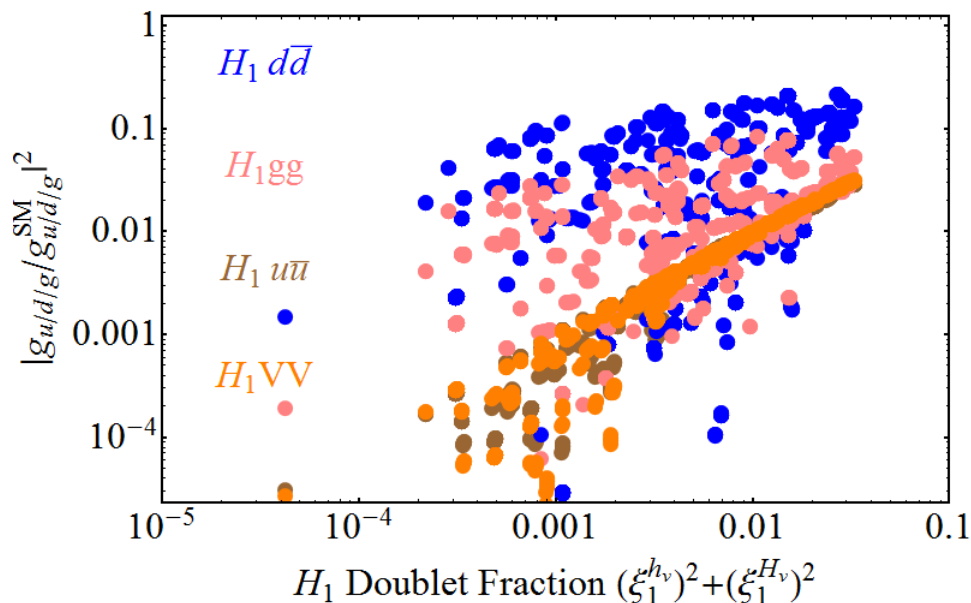
Exotic decays  
 Up to 30%

May have  
 triggering issue.  
 See discussions  
 by the Exotic  
 Higgs Decay  
 Working Group,  
 1312.4992

# Example: Light $h_1$ properties

Generally singlet-Higgs-like

<10% SM Higgs production rate for most modes for a given mass



Mainly bottom pairs  
tau pairs.

Decay to LSP pairs  
kinematically  
suppressed as well as  
required by efficient  
annihilation.

$a_1$  pairs possible

# Summary

We discuss

- $<40$  GeV neutralino DM in the NMSSM (sfermion coannihilation also viable in the MSSM)
- (Speculated) Viable Solutions found for
  - **a1,h1-funnel**
  - **Stau coannihilation**
  - **Sbottom coannihilation**
- Discuss the complementarity of different DM searches; (spin-independent, spin-dependent) direct searches, indirect searches (including Gamma-ray excess)
- Relevant collider signatures discussed/proposed; see more in the paper to come



**Thank You!**

Thank You!  
Hope you enjoy Pheno 14

**PHENO 2014**  
May 5-7 2014 University of Pittsburgh  
PITTSburgh Particle physics, Astrophysics & Cosmology Center  
<http://indico.cern.ch/e/pheno14> (PITT PACC)

*"Full Steam Ahead!"*

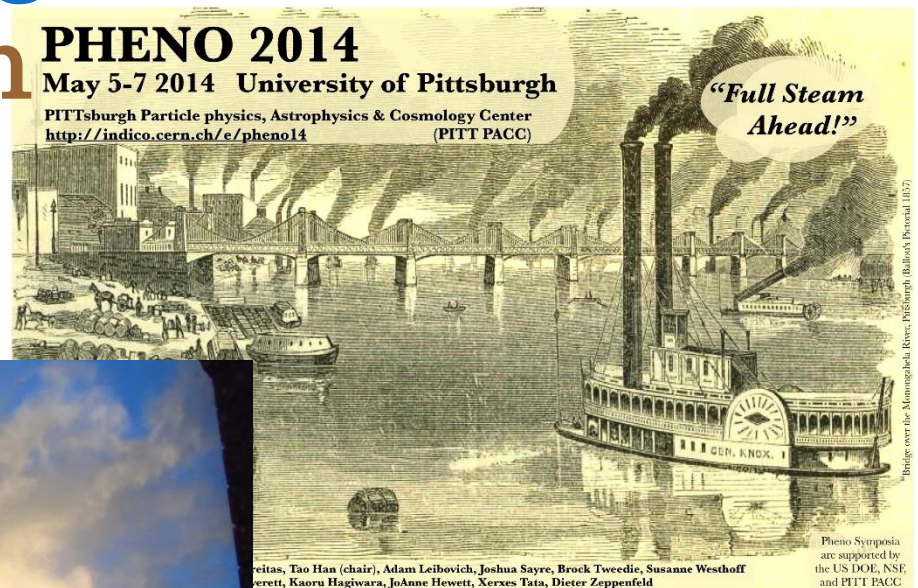
Bridge over the Monongahela River, Pittsburgh, Ballou's Pictorial 1837

**Organizers:**  
Cindy Cercone, Neil Christensen, Ayres Freitas, Tao Han (chair), Adam Leibovich, Joshua Sayre, Brock Tweedie, Susanne Westhoff  
**Program Advisors:** Vernon Barger, Lisa Everett, Kaoru Hagiwara, JoAnne Hewett, Xerxes Tata, Dieter Zeppenfeld

Pheno Symposia are supported by the US DOE, NSF, and PITT PACC

# Thank You!

# Hope you enjoy Pheno 14 and this charming City of Pittsburgh



# Tuning?

Showed up in many places

- 1) Resonance
- 2) Co-annihilation  $\Leftrightarrow$  small mass splitting
- 3) Z decoupling

# LEP Constraints from mono-photon

$>81.9$  95  
 none  $m_T - 26.3$  95  
<sup>3</sup> ABDALLAH 03M DLPH  $\Delta m > 15$  GeV, all  $\theta_T$   
<sup>3</sup> ABDALLAH 03M DLPH  $\Delta m > m_T$ , all  $\theta_T$

Cut		Selection		
		low $\Delta M$	very low $\Delta M$	ultra low $\Delta M$
hadronic calorimeter energy	<	12 GeV	10 GeV	10 GeV
$E_{BGO} - E_\gamma$	<	10 GeV	6 GeV	1 GeV
remaining calorimetric energy	<	12 GeV	8 GeV	6 GeV
muon momentum	<	8 GeV	No muon	No muon
$P_t$ track	<	10 GeV	4 GeV	none
transverse energy imbalance	>	0.1	0.2	0.3
longitudinal energy imbalance	<	0.85	none	none
number of tracks	<	10	7	5
number of BGO energy clusters	<	15	10	6
isolation angle of the photon	<	160°	none	none

