Key concepts of particle physics

CERN Italian teachers’ programme,
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The questions addressed by Particle physics are the same that guided the development of Natural Philosophy in the course of History.

- How does the Universe work?
- Where does it come from?
- Where is it going?
- What are the ultimate components of matter?
- How do they “move”?
- What “moves them”?

The most ambitious among all sciences!
to indentify few fundamental principles, from which to derive the properties of all natural phenomena, both in the macrocosm (the sky, the Universe) and at the human scale.

What has changed in the course of history is the perception of the true complexity of things, the ability to carry out quantitative measurements, and the epistemological criteria establishing the completeness of a given explanation and understanding.

In common, the identification of two categories:

(a) The components of matter
(b) The forces that govern their behaviours
Components: air, water, fire, earth

Forces:
- air and fire pushed upwards
- earth and water pulled downwards

Judgement of correctness:
how come a tree falls in the water, but then gets pushed up and floats?

Reevaluation of the theory (Archimedes)
- all matter is pulled downwards, but with intensity proportional to its weight:
  A body immersed in water receives a push upwards equal to the weight of the displaced water

Air is lighter than the rock, therefore it floats on top of it. Warm air is lighter than cold air, and by it it’s pushed up.
Notice that there is no a-priori guarantee that Nature can be described by a limited number of principles, or that these apply everywhere and at all times.

For example Energy conservation had been put in doubt by the first quantitative studies of nuclear beta decays in the 1920-30’s.

The great success of modern physics lies in its incredibly accurate unified description of the full multitude of observed natural phenomena.
dark matter \( 23\% \)

non-luminous atoms (e.g. planets, dead stars, dust, etc), \( \sim 4\% \)

stars, neutrinos, photons \( \sim 0.5\% \)

dark energy \( 73\% \)
• The understanding of the **Big Bang** and of what preceded it requires the understanding of the behaviour of Nature in presence of gravitational fields of intensity similar to that of nuclear forces (Quantum Gravity).

• The sources of **Inflation** and of **Dark Matter** and **Dark Energy**, which, respectively, shaped and will determine the future of the large-scale structure of the Universe, have to be found within the spectrum of particles which will form our ultimate “*theory of everything*”. 
Level 0: what? how?

- Are there fundamental building blocks?
- If so, what are they?
- How do they interact?
- How do they determine the properties of the Universe?
ALL ORDINARY MATTER BELONGS TO THIS GROUP.

LEPTONS

- **electron**
  - Electric charge: -1.
  - Responsible for electricity and chemical reactions.

- **electron neutrino**
  - Electric charge: 0.
  - Rarely interacts with other matter.

QUARKS

- **up**
  - Electric charge: +2/3.
  - Protons have 2 up quarks.
  - Neutrons have 1 up quark.

- **down**
  - Electric charge: -1/3.
  - ... and one down quark.
  - ... and two down quarks.

- **muon**
  - A heavier relative of the electron.

- **muon neutrino**
  - Created with muons when some particles decay.

- **tau**
  - Heavier still.

- **tau neutrino**
  - Not-yet observed directly.

- **charm**
  - A heavier relative of the up.

- **strange**
  - A heavier relative of the down.

- **top**
  - Heavier still, recently observed.

- **bottom**
  - Heavier still.

ANTIMATTER

Each particle also has an antimatter counterpart ... sort of a mirror image.
Interactions (or “forces”)

• Responsible for:
  • Formation of **bound states** \((E<0)\):
    • Earth-Sun
    • Electron-Nucleus
  • **Scattering** \((E>0)\):
    • Motion of an electron in a metal
    • Propagation of light
    • Deflection of charged particles moving through an electromagnetic field
  • **Transmutations**:
    • Atomic transitions (emission of radiation as an electron changes orbit)
    • Decays \((n\rightarrow p \, e \, \text{neutrino}, \text{radioactivity})\)

*e.g.* protons in the LHC
The fundamental interactions:

**vector bosons, spin=$\hbar/2\pi$**

<table>
<thead>
<tr>
<th>FORCE</th>
<th>COUPLES TO:</th>
<th>FORCE CARRIER:</th>
</tr>
</thead>
<tbody>
<tr>
<td>Electromagnetism</td>
<td>electric charge</td>
<td>photon ($m=0$)</td>
</tr>
<tr>
<td>“weak” force</td>
<td>“weak” charge</td>
<td>$W^\pm$ ($m=80$), $Z^0$ ($m=91$)</td>
</tr>
<tr>
<td>“strong” force</td>
<td>“colour”</td>
<td>8 gluons ($m=0$)</td>
</tr>
</tbody>
</table>

**tensor boson, spin=$2\hbar/2\pi$**

<table>
<thead>
<tr>
<th>gravity</th>
<th>energy</th>
<th>graviton ($m=0$)</th>
</tr>
</thead>
</table>

**scalar boson, spin=0**

<table>
<thead>
<tr>
<th>mass</th>
<th><strong>Higgs ($m\sim125$)</strong></th>
</tr>
</thead>
</table>
Main conceptual results

- **Simplicity** (of the building blocks and their interactions): complexity emerges from the large variety of combinations of large aggregates of elementary objects (like the LEGO sets!)

- **Unity** (of the laws of interaction)

- **Unity** (of the elements):
  
  “a proton is a proton is a proton”

- **Uniqueness** (of the fundamental laws): independence from place, time and external conditions
Elementary particles are subject to the same fundamental principles that you teach in high school:

- “F=ma”
- causality (the cause precedes the effect)
- conservation of energy (E), momentum (p) and angular momentum (L) (invariance of physical laws under space and time translations)
- Einstein’s principle of special relativity
- quantum mechanics (wave-particle duality, uncertainty principle, energy quantization, etc...)
Principles and consequences of special relativity

- No signal/information can propagate faster than light
- The laws of physics are the same in each two reference frames in constant relative motion
- No preferred reference frame, no “center of the Universe”
- Light has the same speed in all frames
- Time is “relative”

\[ V \quad \frac{L'}{c} = T' \quad \frac{L}{c} > T' \]
Elementary particles have very tiny masses, and the forces present in the accelerators, as well as in the Universe, can easily accelerate them to speeds close to the speed of light. Relativistic effects are therefore essential, and the description of the behaviour of elementary particles should be consistent with the laws of special relativity.

In particular, any model of interactions should fulfill the principle that forces cannot be transmitted over distances instantaneously.
N.B.: in quantum mechanics waves and particles are different representations of the same object; therefore to the wave which transmits the signal of the interaction we should associate a particle.
Properties of the interactions

- **Locality** (interaction properties only depend on the properties of the participants at a space-time point)

- **Causality** (the effect follows the cause. It cannot manifest itself before the time it takes for light to cover the distance between cause and effect.)

- **Universality** (the interaction between two particles factorizes in terms the independent properties (e.g. charges) of the individual particles)
Simple ... but subtle!

What happens to energy conservation?!
Count fast!
Quantum mechanics

**Heisenberg uncertainty principle:**

An energy measurement performed within a short time $\Delta t$ can at best reach a precision $\Delta E \geq \frac{1}{\Delta t}$

Within this time lapse it’s impossible to determine whether energy is conserved or not, since we can’t measure it accurately enough. Therefore it’s possible to “cheat” nature, and allow the exchange of energy between the two particles.
Key concepts of particle physics

Lecture 2

CERN Italian teachers’ programme,
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Lepton Interactions ($l = e, \mu, \tau$)

- $\nu l \gamma$: electric charge $-e$
- $\nu l Z$: weak charge $g_W$
- $\nu l W$: weak force
- $\nu l Z$: weak force
Quark Interactions

\[ u \xrightarrow{\text{Photon}} \alpha \frac{2}{3} e \]

\[ d \xrightarrow{\text{Photon}} \alpha \frac{-1}{3} e \]

\[ q' \xrightarrow{\text{W}} \propto g_W \]

\[ q \xrightarrow{\text{Z}} \]

\[ q \xrightarrow{\text{gluon}} \propto g_S = \text{strong coupling} \]
Example

Proton

Q = \frac{2}{3} e + \frac{2}{3} e - \frac{1}{3} e = e

Neutron

Q = \frac{2}{3} e - \frac{1}{3} e - \frac{1}{3} e = 0
Example, the pions

$\pi^+ = u \bar{d}$

$\pi^- = \bar{u} d$

$Q = \frac{2}{3} e + (\bar{d})(-\frac{1}{3}) e = e$

$Q = -\frac{2}{3} e + (\bar{u})(-\frac{1}{3}) e = -e$

where $\bar{q}$ is the antiquark of quark $q$

$\pi^0 = \bar{d} \bar{d} + u u$

$Q = 0$
Example, kaons

\[ K^+ = u \bar{s} \]
\[ K^- = \bar{u}s \]
\[ Q = \frac{2}{3}e + (-)(-\frac{1}{3})e = e \]
\[ Q = -2/3 e + (-1/3) e = -e \]

\[ \bar{K}^0 = d \bar{s} \]
\[ K^0 = \bar{d}s \]
\[ Q = -\frac{1}{3}e + (-)(-\frac{1}{3})e = 0 \]
\[ Q = (-)(-\frac{1}{3})e + (-\frac{1}{3}e) = 0 \]
Example: radioactivity

\[ N \rightarrow N + 1 \]

Diagram showing the process of radioactive decay, with particles such as protons (p), neutrons (n), and electrons (e⁻) interacting.
Example: kaon decays

\[ K^0 \rightarrow \pi^+ \, e \, \nu \]
Example: kaon decays

$K^0 \rightarrow \pi^+ \pi^-$
**Mass:**
- Composite particles -> dynamical origin, calculable: $M = E/c^2$, $E = T + U$
- Fundamental particles -> assigned parameter; origin ???
- Measurement:
  - in decays: $P = \sum p_i$, $M^2 = P^2$
  - in production: $M =$ minimum energy necessary for creation

**Charge:**
- Which type (electric, weak, strong)?
- Are there other charges?? What is the origin of charge??
- Measurement: interaction strength
  - lifetime of a particle before its decay
  - reaction probabilities (rate counting)

**Spin** (intrinsic angular momentum):
- Integer -> bosons, Semi-integer -> fermions
- Origin??
- Pauli principle (two identical fermions cannot occupy the same quantum state) at the origin of matter stability and diversity
- Measurement: angular distributions in scattering or decay processes
Examples of mass determination:

M = energy at production threshold

Production rate for $e^+e^- \rightarrow$ hadrons, as a function of the center of mass energy

The peaks represent the appearance of a new possible final state, made it possible by having enough CM energy to create it. It appears as a “resonance” in a “spectrum”.

\[
\begin{align*}
e^- & \quad \gamma^* & \quad \bar{q} \\
e^+ & \quad & \quad q \\
q & \quad & \quad e^- \\
\bar{q} & \quad \gamma^* & \quad e^+ \\
q & \quad & \quad \bar{q}
\end{align*}
\]
Decays and lifetimes

- If the couplings of a particle $A$ allow it to transform itself into a series of particles $B_1, \ldots, B_n$, and if $m_A > m_{B_1} + \ldots + m_{B_n}$, $A$ decays into $B_1 + \ldots + B_n$.

Only particles for which no decay channel is open can be stable. As of today, we only know of three such examples: the electron, the lightest neutrino and the proton (although there are theories in which the proton is predicted to decay with a lifetime of about $10^{34}$ years, as well as theories in which stable heavy particles explain the origin of dark matter).

- Example:

```
\begin{align*}
\mu^+ & \rightarrow \nu_\mu \\
W^+ & \rightarrow e^+ + \nu_e \\
\end{align*}
```

- The stronger the couplings, and the larger the mass difference, the faster the decay:

$$N(t) = N(0) \, e^{-t/\tau}$$

where $\tau = \tau(M, g)$ is the **life time**
Example: counting the number of neutrinos

\[ \tau \alpha \frac{1}{\text{number of holes}} \sim \frac{1}{\text{number of decay channels}} \]

\[ \tau(Z) = \frac{1}{\Gamma(Z)} \alpha \frac{1}{\text{number of decay channels}} \]

\[ \Gamma(Z) = \sum_{q \ m_q < m_Z/2} \Gamma(Z \rightarrow qq) + \sum_{\ell \ m_\ell < m_Z/2} \Gamma(Z \rightarrow \ell^+\ell^-) + \sum_{\nu \ m_\nu < m_Z/2} \Gamma(Z \rightarrow \nu\nu) \]
The measurement of a width can tell us something about what is not directly seen! It’s like knowing that there is a leak in your tank if your car runs out of petrol while sitting in the parking lot!

More in general, the measurement of a width will give us the strength of the coupling of the decaying particle to the decay products. The width (lifetime) itself is therefore not an “intrinsic” property of a particle, but is a consequence of its mass and of its interactions with other particles.

\[
N_{\text{events}}(e^+e^- \rightarrow Z^0) \propto \left[ (S - M_Z^2)^2 + M_Z^2 \Gamma_Z^2 \right]^{-1}
\]

\[
\Gamma_Z \propto \sqrt{S} \text{= Energy}(e^+e^-)
\]

LEP \( e^+e^- \rightarrow Z^0 \) data, showing that the number of neutrino species \( N_\nu = 3 \)
Key concepts of particle physics

Lecture 3

CERN HST programme,
Sept 8-12 2014

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Level-1 questions: Why?

- Why 3 families of quarks and leptons?
- Why some particles have mass?
- Why $m(\text{neutrino}) \sim 10^{-7} \ m(e)$?
- Why is there a matter-antimatter asymmetry in the Universe?
- Why $F_{\text{gravity}} \sim 10^{-40} \ F_{\text{electric}}$?
- Are particles really pointlike? Strings?? Membranes?
- Why $D=3+1$?
- ..........
- Why something instead of nothing?
To firmly establish the “what”:
- discover the crucial missing element of the Standard Model, namely the **Higgs boson** => done!
- search for possible **new fundamental interactions**, too weak to have been observed so far
- search for possible **new generations** of quarks or leptons
- confirm/disprove the **elementary nature** of quarks/leptons
- discover direct evidence for the particle responsible for the **Dark Matter** in the Universe

To firmly establish the “how”: the observation of the Higgs boson, and the determination of its properties, will complete the dynamical picture of the Standard Model, confirming (hopefully!) our presumed understanding of “how” particles acquire a mass.

To seek new elements which can help us shed light on the most difficult question, namely **WHY?**
Inside the proton we can find, in addition to the component uud quarks, also gluons as well as quark-antiquark pairs.

If we probe the proton at energies high enough, we take a picture of the proton with a very sharp time resolution, and we can “detect” the presence of these additional components. In particular, the gluons and antiquarks present inside will participate in the reactions involving proton.

Notice that, if $\Delta t$ is small enough, even pairs of quark-antiquark belonging to the heavier generations (e.g. s-sbar, c-cbar) can appear!! The proton can contain quarks heavier than itself!!
Probing the proton structure

If the energy with which we probe the proton is small, the proton holds together, and it simply bounces off ....

From the detailed experimental study of this process, we learned that the proton behaves like an extended object, with a charge radius of $O(10^{-12} \text{ cm})$
Probing the proton structure

If the energy transferred by the probe is large enough, we can excite the proton, giving rise to a baryonic “resonance”

... which decays to a proton and a pion

The study of this process helped understanding the mature of the forces that hold together quarks
Quarks inside a proton: they can’t be separated nor extracted from it. If we try, the energy we need to inject in the system is transformed into a new quark-antiquark pair, which screens the individual quark.....
As the energy put into the system becomes larger and larger (w.r.t. the quark masses), it is possible to form multiple quark-antiquark pairs, and the proton breaks up into a multitude of particles.
Examples of reactions in proton collisions

Quark-Quark scattering:
ATLAS EXPERIMENT

Run Number: 201006, Event Number: 55422459
Date: 2012-04-09 14:07:47 UTC

Leading jets 1.96 & 1.65 TeV. Invariant mass 3.81 TeV

Real-life example of jets produced at the LHC
Examples of reactions in proton collisions

Quark-Antiquark annihilation:

\[ u \; d\bar{u} \rightarrow W \]

In principle the “force carrier” of new interactions could be created in the same way, provided their mass is not too large.
Examples of reactions in proton collisions

gluon-gluon reactions:

\( gg \rightarrow \text{top antitop} \)
The Higgs boson
The depth of “Why?” questions is a measure of the maturity of the field. We can only approach “why” questions when we have a solid understanding of the “what”s and “how”s.

**Example: mass**

\[ m = E/c^2 \]  \( \Rightarrow \) for a composite system the mass is obtained by solving the dynamics of the bound state.

So \( m_p = 938 \text{ MeV} \) requires a “how” explanation, not a “why” one.

But what about elementary particles? Elementary \( \Rightarrow \) no internal dynamics.

Need to develop a new framework within which to understand the value of the electron mass.
The Higgs and particles’ masses

Light propagating in a medium is slowed down by its continuous interaction with the medium itself.

\[ \Rightarrow c_{\text{medium}} < c_{\text{vaccum}} \]

The time it takes to move across the medium is longer than if light were propagating in the vacuum,

Think of the Higgs field as being a continuum medium embedding the whole Universe. Particles interacting with it will undergo a similar “slow-down” phenomenon. Rather than “slowing down”, however, the interaction with the Higgs medium gives them “inertia” => mass.
The number “v” is a universal property of the Higgs field background. The quantity “λ” is characteristic of the particle moving in the Higgs field. Particles which have large λ will have large mass, with \( m \propto \lambda v \)

Now the question of “why does a given particle has mass \( m \)” is replaced by the question “why does a given particle couple with the Higgs field with strength \( \lambda \propto m / v \)”

However at least now we have a model to understand **how** particles acquire a mass.
Detecting the Higgs boson

Like any other medium, the Higgs continuum background can be perturbed. Similarly to what happens if we bang on a table, creating sound waves, if we “bang” on the Higgs background (something achieved by concentrating a lot of energy in a small volume) we can stimulate “Higgs waves”. These waves manifest themselves as particles*, the so-called Higgs bosons.

What is required is that the energy available be larger than the Higgs mass \(\Rightarrow\) LHC !!!!

* Even the sound waves in a solid are sometimes identified with “quasi-particles”, called “phonons”
Higgs: Four main production mechanisms

- **Gluon-gluon fusion**
  \[
  g + g \rightarrow t + H^0
  \]

- **Vector boson fusion**
  \[
  W^- + H^0
  \]

- **Radiation from vector bosons**
  \[
  W^\pm, Z + H^0
  \]

- **Radiation from top quarks**
  \[
  g + t \rightarrow H^0
  \]
  \[
  g + \bar{t} \rightarrow H^0
  \]
Higgs decays

\[ H \xrightarrow{f} \bar{f} \]

\[ H \xrightarrow{V(\ast)} \bar{f} \]

\[ f = \text{quarks or leptons} \]

\[ V = W \text{ or } Z \]

\[ H \xrightarrow{\text{top}} \gamma \gamma \]

\[ H \xrightarrow{W^+} W^{-} \gamma \gamma \]
Examples of reactions in proton collisions

Produzione di Higgs
Higgs candidate event, in ATLAS, with $H \rightarrow ZZ^{*} \rightarrow \mu^{+} \mu^{-} e^{+} e^{-}$
The things I didn’t talk about

- Supersymmetry
- Superstrings
- Dark energy
- Matter-antimatter asymmetry of the Universe
- Quark-gluon plasma