

# How to transfer experimental results to theorists?

Convener: Thomas Blake (Warwick U.)

Contributors: Konstantinos Petridis (Imperial College) and Danny van Dyk (Siegen U.)

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# Current Situation

## How is data used right now? - New Physics searches

- Altmannshofer, [Straub](#) [1308.1501] and within
  - ▶ Experimental errors Gaussian, measurements of same quantities by different experiments averaged (weighted average of symmetrised errors).
  - ▶ Form factor correlations included
- Beaujean, [Bobeth](#), [van Dyk](#) [1310.2478] and within
  - ▶ Experimental errors if symmetric treated as Gaussian, if  $> \text{few}\%$  asymmetry use LogGamma.
  - ▶ Correlation info for lattice FFs, but not for LCSR FFs nor LHCb data...
- Descotes, [Matias](#), Virto [1307.5683] and within
  - ▶ Experimental errors Gaussian.
  - ▶ For exclusive decays LHCb data only, no  $B_s$
  - ▶ Correlation info for data from “toys”
- Horgan, Liu, [Meinel](#), Wingate [1310.3887]
  - ▶ Experimental errors Gaussian, measurements of same quantities by different experiments averaged (weighted average of symmetrised errors).

# Current Situation

## How is data used right now? - Form factors

- Beaujean, Bobeth, van Dyk [1310.2478] and within
  - ▶ combination of  $B \rightarrow K^* \gamma$ ,  $B \rightarrow K^* \ell^+ \ell^-$  helpful to fix non-factorizable power corrections
  - ▶ constraints on FFs, power corrections
- Hambrock, Hiller, Schacht, Zwicky [1308.4379] and within
  - ▶ Fit FFs from large  $q^2$  data only
  - ▶ Experimental errors Gaussian
  - ▶ Only ratios of  $B \rightarrow K^*$  angular observables

# Binning of Angular Observables

- fine bins as used for  $B^+ \rightarrow K^+ \mu^+ \mu^-$  analysis appear OK

- basically  $1\text{GeV}^2$  steps, with slight adjustments
- $\phi$  cut out
- $J/\psi, \psi(2S)$  cut out
- some reservations about cutting out  $\phi$  (Sebastian)

Table 2: Differential branching fraction results for  $B^+ \rightarrow K^+ \mu^+ \mu^-$

$q^2$ range ( $\text{GeV}^2/c^4$ )	Differential branching fraction ( $\times 10^{-9}$ )		
	central value	stat error	syst error
$0.1 < q^2 < 0.98$	33.2	1.8	1.7
$1.1 < q^2 < 2.0$	23.3	1.5	1.2
$2.0 < q^2 < 3.0$	28.2	1.6	1.4
$3.0 < q^2 < 4.0$	25.4	1.5	1.3
$4.0 < q^2 < 5.0$	22.1	1.4	1.1
$5.0 < q^2 < 6.0$	23.1	1.4	1.2
$6.0 < q^2 < 7.0$	24.5	1.4	1.2
$7.0 < q^2 < 8.0$	23.1	1.4	1.2
$11.0 < q^2 < 11.8$	17.7	1.3	0.9
$11.8 < q^2 < 12.5$	19.3	1.2	1.0
$15.0 < q^2 < 16.0$	16.1	1.0	0.8
$16.0 < q^2 < 17.0$	16.4	1.0	0.8
$17.0 < q^2 < 18.0$	20.6	1.1	1.0
$18.0 < q^2 < 19.0$	13.7	1.0	0.7
$19.0 < q^2 < 20.0$	7.4	0.8	0.4
$20.0 < q^2 < 21.0$	5.9	0.7	0.3
$21.0 < q^2 < 22.0$	4.3	0.7	0.2
$1.1 < q^2 < 6.0$	24.2	0.7	1.2
$15.0 < q^2 < 22.0$	12.1	0.4	0.6

# Charmonium

- so far, veto windows  $J/\psi$  and  $\psi(2S)$
- for further studies, also give results *within* existing charmonium vetoes
  - ▶ angular observables  $J_n$  should be fine
  - ▶ use similar bin size as in rest of the phase space
  - ▶ experiment:  $J/\psi$  tail is problematic due to detector effects
  - ▶ experiment:  $\psi(2S)$  seems fine
- do not remove broad resonances, see previous session

# Correlation and Likelihood

- So far experimental results do not provide information on:
  - ▶ Correlations between observables and their uncertainties arising from experimental effects such as background or detector acceptance
  - ▶ Confidence level intervals beyond  $1\sigma$
- Particularly in light of recent results/deviations it is crucial to provide both
- How exactly? Case dependent?

# Correlation and Likelihood

## Take a typical tough case:

- Full angular fit of  $B \rightarrow K^*$  involves large number of parameters
  - ▶ 8 to 24 per  $B$  flavour and  $q^2$  region depending on parametrisation
- Cannot trivially sample the likelihood space
- Even if we could, likelihood parametrisation might not be ideal
  - ▶ e.g coefficients of amplitude ansatz
  - ▶ transforming likelihood to more user-friendly basis non-trivial
- Additionally fitting for  $J$ 's or amplitudes results in non-Gaussian likelihood with level of non-Gaussian behaviour depending on fitting strategy
  - ▶ Cannot blindly provide error matrix of fit either
  - ▶ Devise methods to quantify/correct non-Gaussian behaviour

# Correlation and Likelihood

## Easy and user friendly solution:

- Provide stripped down LHCb dataset (background subtracted?)
  - ▶ e.g ROOT n-tuple with angles,  $q^2$ ,  $B$  flavour, background fraction...
  - ▶ Provide continuous  $q^2$  data for large and low recoil region(?)
- Helper classes that:
  - ▶ Build likelihood based on pdf with  $J$ 's or amplitudes (or whatever else experimentalists use) with a full working example reproducing published result
  - ▶ Allows users to build their own likelihood with interfaces to EOS, SuperIso... (requires understanding of how data is used right now)
  - ▶ Provide tools that automatically add experimental nuisance parameters to a given likelihood



## Fitting the $B \rightarrow K^*$ Amplitudes - How?

- fit transversity amplitudes instead of angular observables at  $1\text{GeV}^2 \leq q^2 \leq 6\text{GeV}^2$
- parametrization:  $\lambda = \perp, \parallel, 0$  transversity states,  $\chi = L, R$  lepton chirality

$$A_{\lambda}^{\chi} = \frac{\alpha_{\lambda}^{\chi}}{q^2} + \beta_{\lambda}^{\chi} + \gamma_{\lambda}^{\chi} q^2$$

- amplitudes are complex  $\Rightarrow$  parameters  $\alpha, \beta, \gamma \in \mathbb{C}$
- 4 symmetry relations between amplitudes [Matias, Mescia, Ramon, Virto \[1202.4266\]](#)
- number of real-valued fit parameters  $N$

$$N = (3 \times 2 \times 2 - 4) \times 3 = 24$$

- only usable with full correlation information

## Fitting the $B \rightarrow K^*$ Amplitudes - Why?

- contains more information on  $q^2$  dependence than large bins
- other reasons?

# Fitting the $B \rightarrow K^*$ Amplitudes - Why Not?

- model bias, disregards  $A_S$ ,  $A_t$ , tensor amplitudes
  - ▶ not yet excluded (scalars: [Hurth, Mahmoudi \[1312.5267\]](#), tensors: [Bobeth, Hiller, van Dyk \[1212.2312\]](#))
  - ▶ 2014 LHCb measurement of  $B \rightarrow K \mu^+ \mu^-$  might exclude scalars and tensors
- transversity basis is only one basis of amplitudes
  - ▶ some groups prefer helicity basis: [Jäger, Camalich \[1212.2263\]](#)
- correlation information needed:  $24 \times 24$  *no S-wave contributions*
  - ▶ observables:  $18 \times 18$  per bin, with S wave
  - ▶ virtually no inter- $q^2$ -bin correlation
  - ▶ small bins provide also shape information

## Fitting the $B \rightarrow K^*$ Amplitudes - ToDo

- is parametrization sufficient? back of an envelope!

$$A(q^2) = N(q^2) \times \left( C_9 \pm C_{10} + \frac{\mathcal{T}(q^2)}{\xi(q^2)} \right) \xi(q^2)$$

- norm  $N$  (modulo prefactors)

$$N(q^2) \sim \frac{\sqrt{q^2 \lambda(M_B^2, M_K^2, q^2)}}{M_B^3} = N_0 \sqrt{q^2} + N_1 \sqrt{q^2}^3 + N_2 \sqrt{q^2}^5 + \dots$$

- form factor  $\xi$  (asymptotically)

$$\xi(q^2) = \frac{1}{q^2 - M_B^2} = \xi_0 + \xi_1 q^2 + \xi_2 q^4 + \dots$$

- correlator  $\mathcal{T}$  ( $C_7$  only)

$$\frac{\mathcal{T}(q^2)}{\xi(q^2)} = \frac{M_B^2}{q^2} C_7 + \dots$$

- so shouldn't amplitudes be parametrized as

$$A(q^2) \simeq \sqrt{q^2} \left( \frac{\alpha}{q^2} + \beta + \gamma q^2 \right) \quad ?$$