## Models with Small CC at Finite T

#### also Nearly Braneless Conformal Symmetry Breaking



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Work in progress with Don Bunk and Bithika Jain Pheno and exploration of: Bellazzini, Csáki, JH, Serra, Terning 1305.3919 2nd NPKI Workshop: Jeju Island



- Cosmological history modified in models where the CC is protected dynamically, i.e. by non-linearly realized conformal invariance
  - Phase transitions generally contribute dynamically to vacuum energy
     only end result is tiny
- How do such phase transitions proceed?
- Are results distinct from Creminelli, Nicolis, Rattazzi (hep-th/0107141)?
  - RS phase transition with Goldberger-Wise stablization (has hidden fine tuning)

# Dueling order parameters



## Scale Transformations

**Dilatations:** 

$$x \to x' = e^{-\alpha}x$$

Operators transform:  $\mathcal{O}(x) \to \mathcal{O}'(x) = e^{\alpha \Delta} \mathcal{O}(e^{\alpha} x)$ 



very small bird, normal leaf? kind of small bird, big leaf?

 $\Delta$  is non-perturbative quantum operator scaling dimension

Linearized transformation of action with sourced O:  $S \longrightarrow S + \sum_{i} \int d^4x \, \alpha g_i (\Delta_i - 4) \mathcal{O}_i(x)$ 

# Spontaneous breaking

CFT operator gets VEV:

$$\langle \mathcal{O}(x) \rangle = f^{\Delta}$$

Single corresponding goldstone boson: Low, Manohar '01  $\sigma(x) \to \sigma(e^{\alpha}x) + \alpha f$ 

Non-linear realization in effective theory:

$$f \to f \chi \equiv f e^{\sigma/f}$$

Restores symmetry to LEEFT

# The Dilaton Quartic

Most general terms invariant under dilatations:

$$\mathcal{L}_{\text{eff}} \approx -af^{4}\chi^{4} + \frac{f^{2}}{2}(\partial_{\mu}\chi)^{2} + \text{higher derivative terms}$$

$$\uparrow$$
**dilaton quartic**

**Obstruction to SBSI:** 

•  $a > 0 \rightarrow f = 0$  (no breaking)

Fubini '76

- $a < 0 \rightarrow f = \infty$  (runaway)
- $a = 0 \rightarrow f = anything (flat direction)$



 $f = \infty$ 

# Near-Marginal Deformation

$$\delta S = \int d^4 x \lambda(\mu) \mathcal{O}$$

Quartic has dependence on near marginal coupling:



slowly varying function of f

Deformation can stabilize f away from origin

$$V' = f^3 \left[ 4F(\lambda(f)) + \beta F'(\lambda(f)) \right] = 0$$

# The Dilaton Mass

Expanding the potential:

 $m_{dil}^{2} = f^{2}\beta \left[\beta F'' + 4F' + \beta' F'\right] \simeq 4f^{2}\beta F'(\lambda(f)) = -16f^{2}F(\lambda(f))$ 

small, so dilaton' is light, right?

**F** is the cosmological constant in f units:  $F_{NDA} \sim \frac{\Lambda^4}{16\pi^2 f^4} \sim 16\pi^2$ 

Need large  $\beta$  to find minimum  $V' = f^3 [4F(\lambda(f)) + \beta F'(\lambda(f))] = 0$ Theory not conformal at scale f - **no light dilaton**  $m_{dil}^2 \sim 256\pi^2 f^2 \sim \Lambda^2$  3 TeV <u>not suppressed</u> OR we can tune away the quartic to get a near flat-direction

# CPR idea

Contino, Pomarol, Rattazzi (Talk at Planck 2010)

•  $F(\lambda)$  generically large, but if  $\lambda$  near marginal for range of  $\lambda$ , theory will scan over F with scale

$$\frac{d\lambda}{d\log\mu} = \beta(\mu) \equiv \epsilon \ll 1$$

 large F will not generate spontaneously broken scale invariance

- minimum when  $F \sim 0$
- dilaton mass proportional to E

### Various dynamical possibilities









### Naturally Light Dilaton, Small CC

5D scalar min.  
coupled to gravity: 
$$S = \int d^5 x \sqrt{g} \left[ \frac{1}{2} (\partial_M \phi)^2 - V(\phi) - \frac{1}{2\kappa^2} R \right]$$
  
+ brane potentials  
at yo and y<sub>1</sub>  
small  $\beta \Leftrightarrow$  slowly changing V( $\Phi$ ):  $V(\phi) = \Lambda_5 + \epsilon f(\phi)$   
Csáki, Bellazzini, JH, Serra, Terning 1305.3919  
Metric Ansatz:  
flat 4D slices  
 $ds^2 = e^{-2A(\tilde{y})} \eta_{\mu\nu} dx^{\mu} dx^{\nu} - d\tilde{y}^2$   
"True scale" coordinates:  $\mu = ke^{-y}$   
 $A(\tilde{y}) = y \quad G(y) = A'(\tilde{y}(y))^2$   
 $ds^2 = e^{-2y} \eta_{\mu\nu} dx^{\mu} dx^{\nu} - \frac{dy^2}{G(y)}$   
AdS/CFT: EOM capture  
running even when far  
from AdS  
Deviations from pure AdS encoded in G(y)

 $\mathbf{V}$ 

## Scalar-Einstein Equations

#### **Equations of motion:**

$$\begin{split} G &= \frac{\frac{-\kappa^2}{6}V(\phi)}{1 - \frac{\kappa^2}{12}\dot{\phi}^2} \\ \frac{\dot{G}}{G} &= \frac{2\kappa^2}{3}\dot{\phi}^2 \\ \ddot{\phi} &= \left(4 - \frac{1}{2}\frac{\dot{G}}{G}\right)\dot{\phi} + \frac{1}{G}\frac{\partial V}{\partial \phi} \end{split}$$

Can substitute G(y) in terms of scalar vev in last equation:

**Master Evolution Equation:** 

 $\ddot{\phi} = 4\left(\dot{\phi} - \frac{3}{2\kappa^2}\frac{\partial \log V(\phi)}{\partial \phi}\right)\left(1 - \frac{\kappa^2}{12}\dot{\phi}^2\right)$ 

$$\phi \sim \log \lambda$$

Captures running (and condensation) of sourced operators in ~CFT

## Dilaton Effective Potential

Bulk action total derivative (Bellazzini et. al. 1305.3919) integrates to pure boundary term, along with brane localized potentials and jump contributions

What is the behavior of the dilaton effective potential for various bulk scalar potentials? (various deformations of CFT)

Task = work out UV and IR asymptotics

How is spontaneously broken scale invariance manifested?

## Where's the IR brane?



## Constant V (exact CFT)



Asymptotics: 
$$\phi \approx \begin{cases} \phi_0 \pm \frac{1}{4} \sqrt{\frac{12}{\kappa^2}} \left(\frac{f}{\mu_0}\right)^4 + \mathcal{O}\left(\left(\frac{f}{\mu_0}\right)^{12}\right) & \mu_0 \gg f & \text{UV} \\ \phi_0 \pm \frac{1}{4} \sqrt{\frac{12}{\kappa^2}} \left[\log\left(2\left(\frac{f}{\mu_1}\right)^4\right)\right] & \mu_1 \ll f & \text{IR} \end{cases}$$
  
 $G = k^2 \begin{cases} 1 + \mathcal{O}\left(\left(\frac{f}{\mu_0}\right)^8\right) & \mu_0 \gg f & \text{UV} \\ \left(\frac{f}{\mu_1}\right)^4 + \mathcal{O}\left(\left(\frac{\mu_1}{f}\right)^4\right) & \mu_1 \ll f & \text{IR} & \text{A'} \neq \text{ constant!} \end{cases}$ 

### Plug into Holographic Dilaton Potential

$$V_{\text{cff}} = e^{-4y_0} \left[ V_0(\phi(y_0)) - \frac{6}{\kappa^2} \sqrt{G(y_0)} \right] + e^{-4y_1} \left[ V_1(\phi(y_1)) + \frac{6}{\kappa^2} \sqrt{G(y_1)} \right]$$

$$Substitute and re-arrange:$$

$$V_{\text{dilaton}} = \left( \frac{\mu_0}{k} \right)^4 \left[ V_0(\phi_0) - \frac{6k}{\kappa^2} \right] \pm \frac{1}{4} \sqrt{\frac{12}{\kappa^2}} \left( \frac{f}{k} \right)^4 \left[ \frac{\partial V_0}{\partial \phi}(\phi_0) + 2k \sqrt{\frac{12}{\kappa^2}} \right] + \mathcal{O} \left[ \left( \frac{f}{\mu_0} \right)^8 \right]$$

$$Bare C.C. \qquad \text{Dilaton quartic}$$

$$+ \text{ subleading terms of order } (\mu_1/f)^4$$

$$2 \text{ Needed Tunings:}$$

$$tune dilaton quartic to zero (force a flat direction)$$

$$fine-tune bare CC$$

$$Message:$$
soft wall geometric model of spontaneously broken scale invariance
$$IR \text{ brane plays subdominant role:}$$

$$iust cuts off growth of scalar field and curvature$$

### General bulk potentials: UV

#### Small Bulk Mass Term

$$V = -\frac{6k^2}{\kappa^2} \left(1 + \frac{\kappa^2}{3}\epsilon\phi^2\right)$$

Approximate (small backreaction) solution:

$$\phi = \phi_0 e^{\epsilon y} + \widetilde{\lambda} e^{(4-\epsilon)(y-y_c)}$$

$$G(y) = k \left( 1 + \mathcal{O}(e^{-8(y_c - y_0)}) \right)$$

#### UV contribution to dilaton effective potential:

$$V_{\rm UV} = \left(\frac{\mu_0}{k}\right)^4 \left[V_0(\phi_0 e^{\epsilon y_0}) - \frac{6k}{\kappa^2}\right] + \tilde{\lambda}_{\rm UV} \left(\frac{f}{k}\right)^{4-\epsilon} \left(\frac{\mu_0}{k}\right)^{-\epsilon} \frac{\partial V_0}{\partial \phi}(\phi_0 e^{\epsilon \phi})$$
  
Bare CC marginal "almost quartic"

### General bulk potentials: IR

$$\ddot{\phi} = 4 \left( \dot{\phi} - \frac{3}{2\kappa^2} \frac{\partial \log V(\phi)}{\partial \phi} \right) \left( 1 - \frac{\kappa^2}{12} \dot{\phi}^2 \right)$$
Parametrize:  

$$\dot{\phi} = \sqrt{\frac{12}{\kappa^2}} (1 - \delta v_{\rm IR})$$
IR Universality  
Recall:  $G(y) = \frac{\frac{-\kappa^2}{6} V(\phi)}{1 - \frac{\kappa^2}{12} \dot{\phi}^2} \approx -\frac{\kappa^2}{12} \frac{V(\phi)}{\delta v_{\rm IR}}$  In deep IR  
Expand scalar EOM  
in small  $\delta v$ :  
 $\delta v_{\rm IR} = -8 \left( 1 - \frac{1}{8} \frac{d}{dy} \log V(\phi) \right) \delta v_{\rm IR}$   
Integrable in asymptotic limit!  
 $\delta v_{\rm IR} = \tilde{\lambda} \frac{V(\phi(y))}{V(\phi(y_c))} e^{-8(y-y_c)}$   
 $G_{\rm IR}(y_1) = -\frac{\kappa^2}{12\tilde{\lambda}} V(\phi(y_c)) e^{8(y_1-y_c)}$   
 $G_{\rm IR} = e^{-4y_1} \left[ V_1(\phi(y_1)) + \frac{6}{\kappa^2} \sqrt{G(y_1)} \right] \rightarrow \sqrt{-\frac{\kappa^2}{12\tilde{\lambda}}} V(\phi_c) \left( \frac{f}{k} \right)^4$ 

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### **Resulting Dilaton Potential**

$$V_{\rm dilaton} = -\mu^{\epsilon} f^{4-\epsilon} + \lambda f^4$$

#### At min:

$$f \approx \mu \left(\frac{1}{\lambda}\right)^{1/\epsilon} \qquad V_{\text{dilaton}} = \frac{\epsilon}{4} \lambda f^4$$

#### CC is suppressed by $\epsilon!$

#### Up to kinetic term normalization of dilaton:

$$m_{
m dilaton}^2 \sim \epsilon f^2$$

Soft wall realization of small condensate contribution to CC in models with non-linearly realized SBSI

### Finite Temperature

Black hole in AdS

Same action - euclidean signature, compactified time  $t \in \{0, \beta = I/T\}$ 

$$ds^{2} = e^{-2y} \left[ h(y)dt^{2} + d\vec{x}^{2} \right] + \frac{1}{h(y)} \frac{dy^{2}}{G(y)}$$



### Black Brane



#### Near horizon geometry typically divergent (hairy) out-of-equilibrium

Equilibrium temp: adjust y<sub>1</sub> to remove conical singularity

$$T_{\rm eq} = \frac{e^{-y_1}}{2\pi} \sqrt{-\frac{\kappa^2}{6}} \dot{h}(y_1) V(\phi_h) \propto \frac{k}{2\pi} e^{-y_1}$$

# Constant V (exact CFT)

Evolution equations again integrable:

$$\frac{d}{dy}\log\dot{\phi} = \frac{d}{dy}\log\frac{\dot{h}}{h}$$

Solution:

$$\phi = \phi_0 + C_l \log h$$

Incompatible with horizon B.C.'s and finite  $\phi$  unless C<sub>I</sub>=0 Finite temperature ruins the flat direction! Constant  $\phi \rightarrow$  no condensate region - UNBROKEN CFT

#### AdS-Schwarzchild

$$h = 1 - e^{4(y - y_h)}$$
  $G = k^2$ 

Plugging into holo. potential (free energy):  $F_{\rm AdS-S} \propto T^4$ Usual result for unbroken CFT at finite T

### General Potentials

$$V = -\frac{6k^2}{\kappa^2} \left(1 + \frac{\kappa^2}{3}\epsilon\phi^2\right)$$

**Horizon B.C.'s:**  $\dot{\phi}\Big|_{y_h} = \frac{3}{2\kappa^2} \frac{\partial \log V}{\partial \phi}\Big|_{y_h} = \epsilon \phi\Big|_{y_h}$ 

In principle, there seems to be no great obstacle: horizon doesn't force trivial  $\phi$  = Constant

In practice, goal of numerical solutions with horizon at finite y<sub>h</sub> and condensate region has proven **extraordinarily stubborn condensate forces horizon to y=∞** 

Likely due to fragility of condensate under small perturbations (i.e. low temp) appearing as numerical stability issues

## Conclusions

- Models that make progress on the CC problem are few and far between
- We explore variants and pheno of holographic CPR:
  - Bellazzini, Csáki, JH, Serra, Terning 1305.3919
- Among this class of models are ones in which SBSI manifests as continuous geometries (soft wall SBSI) with IR brane playing lesser role as cutoff
- Should allow a continuous parametrizing of model at finite temperature
  - vital for understanding cosmological behavior of various models where CC problem has dynamical resolution
  - so far elusive, but ...

# Thank you!

