

Review of double beta decay experiments: Present status and near future

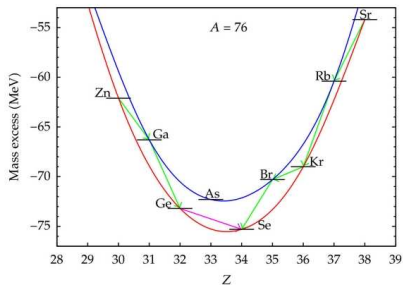
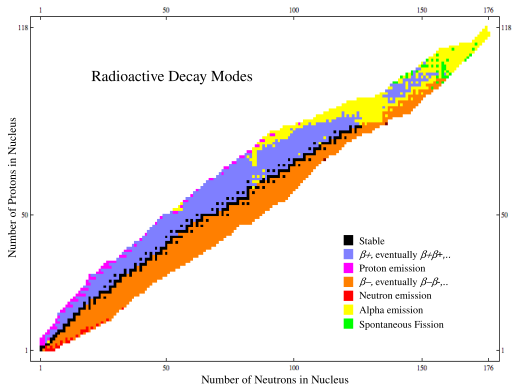
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13th International Workshop on Tau Lepton Physics
September 15-19, 2014, Aachen, Germany



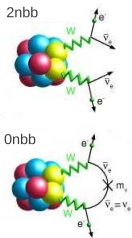
Searching for the double beta decay



Towards the valley of stability:

- **Second-order nuclear transition** that can occur between two odd-odd isobars:
 - single β decay **energetically forbidden**
 - Among competing channels ($\beta^-\beta^-$, $\beta^+\beta^+$, $EC\beta^+$, $ECEC$) $\beta^-\beta^-$ has highest rates
- **35 candidates** in Nature; examples:
 ^{76}Ge , ^{82}Se , ^{96}Zr , ^{100}Mo , ^{116}Cd , ^{130}Te , ^{136}Xe

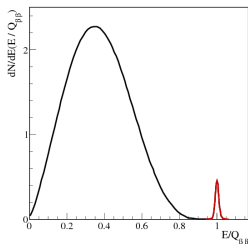
Double beta decay: $2\nu\beta\beta$ versus $0\nu\beta\beta$



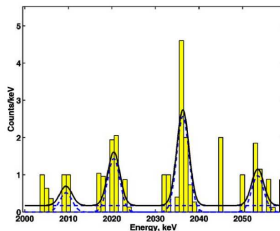
Role of neutrinos in $\beta\beta$ decay:

- 1 $2\nu\beta\beta$ decay
 - Allowed by Standard Model
 - Signature: β -like spectrum
 - Observed in 12 candidates: $O(T_{1/2})=10^{18}-10^{24}$ yr (Longest-lived decay processes observed in Nature)
- 2 $0\nu\beta\beta$ (or $0\nu\chi^0(\chi^0)\beta\beta$) decay
 - Not allowed by Standard Model
 - Signature: Full energy peak at $Q_{\beta\beta}$
 - One claim (in ^{76}Ge) by subgroup of Heidelberg-Moskow collaboration (HdM)

Expected experimental signature from $2\nu\beta\beta$ and $0\nu\beta\beta$ decays



Claim: observed signal in ^{76}Ge at $Q_{\beta\beta}=2039$ keV
 Klapdor-Kleingrothaus et al., Phys. Lett. B 586 (2004) 198



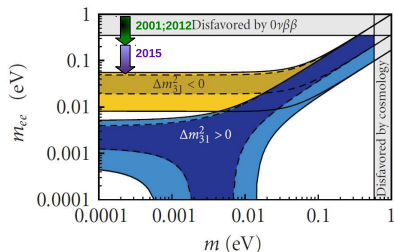
Motivations for $0\nu\beta\beta$ decay search

Observation of $0\nu\beta\beta$ decay helps answering 3 fundamental questions:

- 1 Is lepton number conservation **violated** ($\Delta L=2$)? Are neutrinos **their own anti-particles** ?
→ If yes, physics beyond the Standard Model of Elementary Particles with impact up to cosmology
- 2 What is the **absolute neutrino mass scale** ?
→ Neutrino-oscillations reveal only squared mass-differences of neutrino mass eigenstates
- 3 Is the **neutrino mass spectrum degenerate, normal or inverted** ?
→ Atm./solar/reactor neutrino oscillation experiments allow different scenarios (**Hierarchy problem**)
→ Measurement of 'effective' neutrino mass which is very rich of information:

$$\langle m_{ee} \rangle = \left| \sum |U_{ei}|^2 m^i e^{i\alpha_i} \right|$$

with U_{ei} neutrino mixing matrix, m^i neutrino mass eigenstate, α_i CP-violating Majorana phases



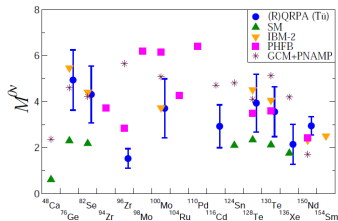
Observable: $0\nu\beta\beta$ decay rate at $Q_{\beta\beta}$, i.e. half-life $T_{1/2}$. If not observed, then quoting a lower limit of $T_{1/2}$ (90%C.L.).

- **Best limit in the past** obtained by HdM (2001):
 $T_{1/2}^{0\nu} > 1.9 \times 10^{25}$ yr; $\langle m_{ee} \rangle \leq 0.35$ eV
- **KKDC claim (2004):**
 $T_{1/2}^{0\nu} = 1.17 \times 10^{25}$ yr; $\langle m_{ee} \rangle \sim (0.23-0.59)$ eV

Half-life correlation with effective Majorana neutrino mass

$$(T_{1/2})^{-1} = G^{0\nu} |M^{0\nu}|^2 \langle m_{ee} \rangle^2$$

with $G^{0\nu}$: phase space factor, $M^{0\nu}$: nuclear matrix element, $\langle m_{\beta\beta} \rangle = \left| \sum_j m_j U_{ej}^2 \right|$



$M^{0\nu}$ calculations:

- Improvements for NSM and QRPA:
 - Most QRPA discrepancies solved
 - Progress in understanding source of spread of NSM values
- New methods IBM, EDF, pHFB

$Q_{\beta\beta}$ values:

→ Penning-traps (e.g. ^{130}Te : 5% shift)

Cross sections for neutron reactions:

(e.g. $^{207}\text{Pb}(n, n'\gamma)$): DEP of 3062 keV $\simeq Q_{\beta\beta}$ of ^{76}Ge)

Theory's demand for a larger number of measurements with different isotopes

- Avoid (not well) known rare background events at $Q_{\beta\beta}$
- NME uncertainties $\leq 30\%$ for neutrino mass spectrum & CP violating phases
- Mechanisms: Light vs. heavy Majorana neutrino exchange, RHC,...

Determination of the half-life

$$T_{1/2} \propto \begin{cases} a \cdot \epsilon \cdot M \cdot T, & \text{background-free} \\ a \cdot \epsilon \cdot \sqrt{\frac{M \cdot T}{\Delta E \cdot B}}, & \text{if background is present} \end{cases}$$

with **a**: Abun./Enrich.; **M**: Mass; **ε**: act.volume; **ΔE**: e-res.; **T**: life-time; **B**: bkgd ([counts/kg/keV/yr] around region-of-interest)

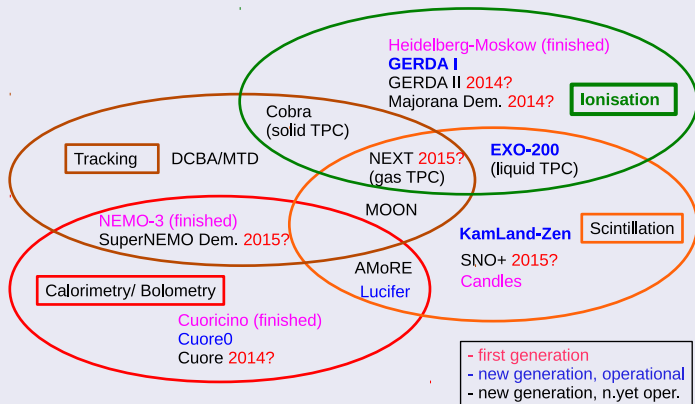
Isotope	$Q_{\beta\beta}$ [keV]	nat. abundance [%]	annual production [ton]	Experiment (oper./funded)	FWHM/E @ $Q_{\beta\beta}$ [%]	Mass [kg]
⁴⁸ Ca	4273.7	0.19	5×10^5	Candless		0.35
⁷⁶ Ge	2039.1	7.8	10	GERDA	0.1-0.2	15 → 35
				Majorana Dem.	0.1-0.2	30
⁸² Se	2995.5	9.2	250	SuperNEMO		7 → 100
				Lucifer		-
¹⁰⁰ Mo	3035.0	9.6	2.5×10^4	MOON		480
				AMoRe		100
¹¹⁶ Cd	2809.1	7.6	1.5×10^3	Cobra		64
¹³⁰ Te	2530.3	34.5	50	CUORE	0.2	10 → 200
¹³⁶ Xe	2457.8	8.9	1.5	EXO	4.0	175
				KamLand-Zen	9.8	330 → 1000
				NEXT		100
¹⁵⁰ Nd	3367.3	5.6	500	SNO+(orig.)		44

Motivations for a larger number of experiments and isotopes:

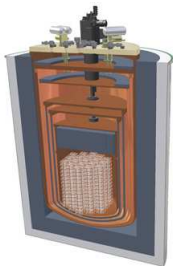
- **Reduction of background**: detector segmentation, particle tracking, pulse shape, $Q_{\beta\beta}$ above lines of ²²⁸Th
- **Advantages of single isotopes**: better ΔE, enrichment of isotope, scalability (detector size, world annual production, prize)
- **Avoid/reduce systematics**: independent techniques, measurements with $\leq 30\%$ precision

Isotopes and technologies for $0\nu\beta\beta$ decay search

- Selected best isotope candidates: 8 out of 35 (\leftarrow nat. abundance, $Q_{\beta\beta}$, $G^{0\nu} \propto (Z, Q_{\beta\beta}^5)$, chem. properties)
- Techniques and their advantages:



- Source=Detector:** large isotopic masses possible and scalable (10-100 kg now, 1 ton near future)
- Source \neq Detector:** Tracking particle momenta \rightarrow Topology of events and mechanisms
- Liquid scintillator detectors:** multi-functional, largest target masses, purification strategies
- Solid state detectors** achieve the best energy resolution: FWHM/E at $Q_{\beta\beta}=0.15\%$



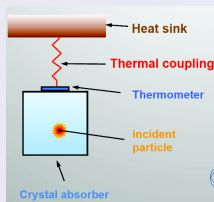
Location: LNGS, Assergi, Italy

Detectors, cryostat, and shield

- TeO_2 crystals cooled down to ~ 10 mK with He within a multi-layer copper cryostat
- Isotopic nat. abundance of ^{130}Te : 34.1% (no enrichment!)
- Energy resolution: 0.2% in ROI !
- Inner Roman lead layer and outer lead layer
- Ra barrier and neutron shield
- 1400 m overburden, corresponding 3500 m w.e.

Concept: **DBD Source = Absorber**

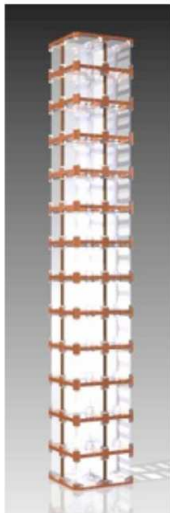
Bolometric technique



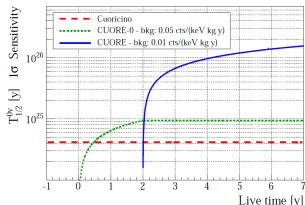
- TeO_2 absorbs energy deposition E by particle
- Energy deposition E registered by a thermistor (NTD Ge) as temperature increase:
Signal: $\Delta T = E/C$, C: capacity
Time constant = C/G ; G: thermal coupling
Need: \rightarrow low-heat C \rightarrow mK + diele.diamagn.mat.
- Very good energy resolution achievable:
 ~ 5 keV @ $Q_{\beta\beta}$ (2527 keV), corr. **$\text{FWHM}/E=0.2\%$**

Status and sensitivity of CUORE

Cuore-0 tower



Cryostat of Cuore-0



- **Cuore-0** (2012—2014):
 - 1 tower with 52 crystals, 39 kg $^{130}\text{TeO}_2$, 11 kg ^{130}Te
 - Data collection: 2011-2014, blinded
 - Achieved background: $\text{BI}=0.02$ cts/(keV·kg·yr)
 - Expected after 3 yr run:
 $\rightarrow T_{1/2}^{0\nu} > 5.9 \times 10^{24}$ yr, $\langle m_{\beta\beta} \rangle < 0.17\text{-}0.39$ eV
- **Cuore** (2015—2020):
 - **Detector mass:** 19 'Cuore-0' towers, 988 kg $^{130}\text{TeO}_2$, 206 kg ^{130}Te
 - **Status:** All towers assembled, radiopurity of all crystals measured
 - **Expected sensitivity:**
 - With $\text{BI}=0.01$ cts/(keV·kg·yr), expected after 5 yr run:
 $\rightarrow T_{1/2}^{0\nu} > 1.6 \times 10^{26}$ yr, $\langle m_{\beta\beta} \rangle < 0.04\text{-}0.10$ eV

Concept: DBD source = Detector



Location: LNGS, Assergi, Italy

Setup and background reduction:

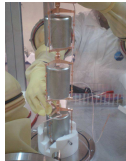
- **1.4 km overburden**, corresponding 3500 m w.e.
→ Reduction of cosmic-muon flux by six orders of magn. down to $\sim 1 \mu / (m^2 \cdot h)$ (PB)
- **Water tank and plastic scintillator**
 - R=5 m, h=9.0 m, 590 m³ ultra-pure water
 - water acts as neutron moderator/absorber (PB)
 - both components act as muon Cherenkov veto (AB)
- **Large volume cryostat:**
 - R=2 m, h=5.9 m, 64 m³ LAr
 - LAr acts as cooling medium for diodes
 - LAr attenuates external radiation (PB)
 - LAr scintillation light planned to be used an background rejection (AB)
- **GERmanium Detector Array:**
 - Operating bare diodes in LAr (-198°C) using low-mass, ultra radiopure Cu holders (PB)
 - Detectors enriched in ⁷⁶Ge
 - Coincidence modus and pulse shape tracing (AB)
 - **Phase I (2011-2013):** 1-string and 3-string arms with each 3 detectors
 - **Phase II (start: begin 2015):** One 7-string arm

PB = passive background rejection

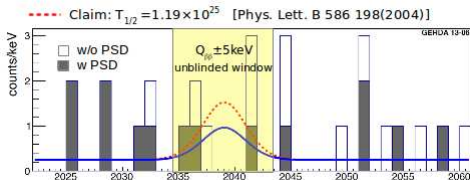
AB = active background rejection

Results from GERDA Phase I

String with
co-axial Ge diodes



String with
Broad Energy Ge diodes



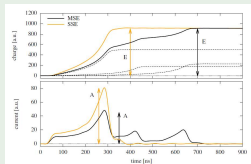
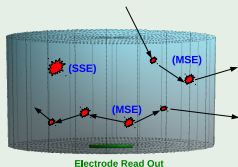
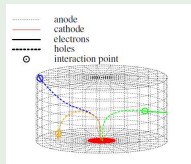
GERDA Phase I (2011-2013)

- **Physics goal fulfilled:** 21.6 kg·yr with $BI \approx 0.01$ cts/(kg·yr·keV) including pulse shape cuts (**unprecedented !**)
- **Detector performance:** excellent energy resolution at 0.2%, energy scale mostly very stable
- **Data analysis:** fully **blinded** (**unprecedented !**)
- **Physics results:**
 - $T_{1/2}^{0\nu} > 2.1 \times 10^{25}$ yr (90% C.L.)
 - combined Ge experiments: $T_{1/2}^{0\nu} > 3.0 \times 10^{25}$ yr (90% C.L.))
 - **HdM claim (2004) rejected at 99% level by GERDA alone**
 - Combined Ge experiments (HdM + IGEX + Gerda): $\langle m_{\beta\beta} \rangle < (0.2-0.4)$ eV

Upgrade to GERDA Phase II (planned start: begin 2015)

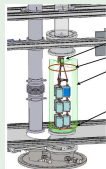
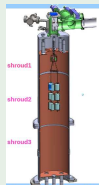
Assembly of new detectors:

- **Novel technology:** $30\text{ }^{76}\text{Ge}$ enriched BEGe-type detectors for $\sim 20\text{ kg}$
→ Improved sensitivity expected due to: increased target mass, better energy resolution, enhanced pulse shape background discrimination, improved read-out electronics

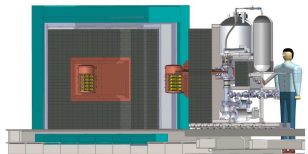


Exploitation of liquid argon scintillation light:

- **Background detection:** rejection of external background events via detection of **scintillation light** in liquid argon ($\lambda=128\text{ nm}$)
- **Hybrid solution for read-out:** Photomultiplier tubes + wavelength-shifter glass-fibres
- **Envisoned BI** (BEGe + light instr.): $\leq 0.001\text{ cts}/(\text{keV}\cdot\text{kg}\cdot\text{yr})$, after 3 yr of data collection:
→ $T_{1/2}^{0\nu} > 1.5\text{-}2 \times 10^{26}\text{ yr}$, $\langle m_{\beta\beta} \rangle < 0.08\text{-}0.12\text{ eV}$



Concept: **DBD source = Detector**



Location: Sanford UG lab, Lead, SD, USA (1.480 m overburden)

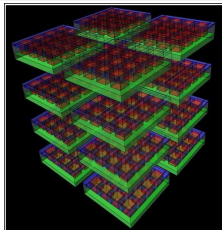
R&D, design, goals and status:

- **P-type point contact Ge detectors** (low noise, E-threshold at 0.5 keV → suitable also for dark matter search!)
- **40 kg** (out of them 30 kg enriched in ^{76}Ge) germanium diodes in two **large volume vacuum crystats** (**7 strings of 3-5 detectors each**) within ultra-low background shield made of ultra-radiopure copper and lead.
- Demonstrate sensitivity: $\leq 0.001 \text{ cts}/(\text{keV}\cdot\text{kg}\cdot\text{yr})$ (like GERDA Phase II)
- Installation/commissioning of 1.cryostat in **spring 2013**, of 2.cryostat in **fall 2014**; data collection start in **2015**

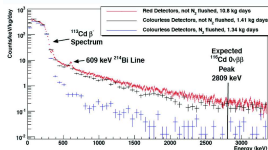
GERDA-Majorana cooperation:

- **Exchange** of knowledge and technologies, joint Geant4-based **simulation tool MaGe**
- **Intention for 100 kg-1 ton experiment**; goal sensitivity: $\langle m_{\beta\beta} \rangle$ **0.02-0.12 meV** (IH region)

Concept: DBD source = detector

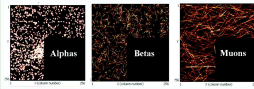


Location: LNGS, Assergi, Italy



Solid state TPC

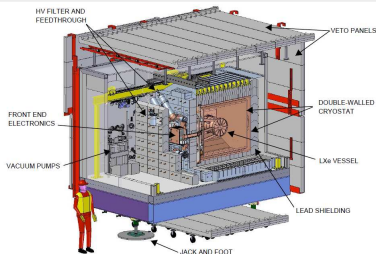
- Large Array of **CdZnTe** **semicond.** detectors with goal sensitivity: $T_{1/2} > 10^{26}$ yr ($\langle m_{\beta\beta} \rangle$ 0.05 eV)
- **Composition:** 5 $0\nu\beta\beta^- \beta^-$, + 4 EC/ β^+ , EC/EC candidates !
Best candidate: ^{116}Cd : $Q_{\beta\beta}=2809$ keV; \rightarrow enrichment
- **Two modular designs** (both operational at room temp.!):
 - **CoPlanar Grid Detectors (CPG):** little 'location' info (with PSA), $\Delta E < 2\%$, and simple read-out
 - **Pixelated Detectors (PD):** 3D 'location', particle ID if pixels small ($\sim 100 \mu\text{m}$), $\Delta E < 1\%$, but complex read-out



Status (2013/2014):

- Termination of coplaner grid with 64 $10 \times 10 \times 10$ cm³ detectors
- MC/tests for $20 \times 20 \times 15$ cm³ detectors, 10⁴ in final design
- MC/tests for shield options: multi-layer, liq. scint., water
- $2\nu\beta\beta$: 6 $T_{1/2}$ limits $> 10^{20}$ yr, 7 new upper limits
- $0\nu\beta\beta$: achieved background at ^{116}Cd ROI for CPG: 5 cts/(keV·kg·yr) (2012)

Concept: DBD source = Detector



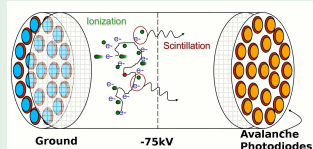
Detector design and background reduction

- LXe Vessel in ultra-radiopure copper cryostat filled with high-purity heat transfer fluid HFE7000
- Lead shield
- 4 plastic scintillators as active muon vetos
- 700 m overburden, corresponding 1600 m w.e.

Location: WIPP lab, Carlsbad, NM, USA

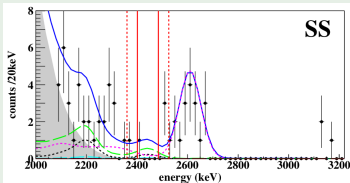
Detection principle

- Medium: 175 kg of LXe; ^{136}Xe enrichment: 80.6%
- Detection principle:



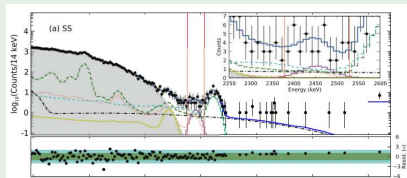
- Collection charge wires measure ionized electrons
- Large Area Avalanche Photodiodes (APDs) measure 178 nm scintillation light
- Gain information from:
 - Drifttime:
 - Position reconstruction res.: X,Y: 18 mm; Z: 6 mm
 - Distinguish single $\beta/\beta\beta$'s from multiple γ -ray clusters:
 - Ionisation vs. Scintillation:
 - Discrimination of α from $\beta/\beta\beta/\gamma$
 - Improve energy resolution down to 1.67%

1. publication (May 2012) 120 d with 98.5 kg of LXe:



- Observed background: 1(5) events within $1(2)\sigma$ around $0\nu\beta\beta$ ROI
 $\rightarrow \text{BI} = 0.0015 \text{ cts}/(\text{keV}\cdot\text{kg}\cdot\text{yr}) \rightarrow$ within specs
- Largest contaminant: ^{232}Th , ^{238}U , ^{137}Xe
- $T_{1/2}^{0\nu} > 1.6 \times 10^{25} \text{ yr}$,
 $\langle m_{\beta\beta} \rangle < 0.14\text{-}0.38 \text{ eV (90\% C.L.)}$

2. publication (February 2014) 478 d with 98.5 kg of LXe:

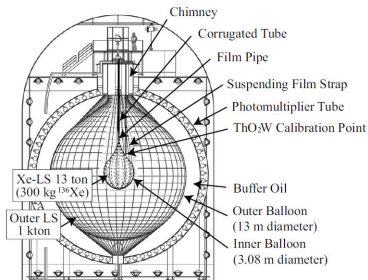


- Data set includes first 120 d
- Partial unblinding procedure applied
- Observed background: 31 ± 5 events within 2σ around $0\nu\beta\beta$
 $\rightarrow \text{BI} = 0.0017 \text{ cts}/(\text{keV}\cdot\text{kg}\cdot\text{yr}) \rightarrow$ within specs
- $T_{1/2}^{0\nu} > 1.1 \times 10^{25} \text{ yr}$,
 $\langle m_{\beta\beta} \rangle < 0.19\text{-}0.45 \text{ eV (90\% C.L.)}$

\rightarrow Background in 1. publication slightly **underfluctuating**,
 Reason for better sensitivity limit than in 2. publication, even though 4fold data set available.

The KamLAND-Zen experiment

KamLAND-Zen is 'embedded' within KamLAND
using **Xe-loaded LS**



Location: Kamioka, Japan

Overburden, 3.2 kt water tank, SSS

- 1000 m overburden, corresponding 2700 m.w.e.
- Water tank: neutron moderator and muon Cherenkov detector
- SSS: 1879 PMTs detecting scintillation light

1000 t KamLand LS (R=6.5 m)

- Dodecane 80%, PC 20%, PPO 1.36 g/l
- **Active shield:** ext. γ s, int. γ s from IB/Xe-loaded LS

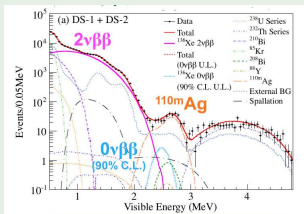
Inner balloon (R=1.54 m, 25 μ m thick)

- 13 t of Xe-loaded LS:
 - Decane 82.3%, PC 17.7%, PPO 2.7 g/l
 - Xe \sim 3 wt% (320 kg)
- ^{136}Xe enrichment: **90.9%**

Advantages of using a) Xe-loaded LS b) in KamLAND

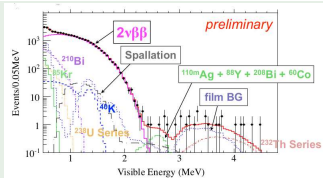
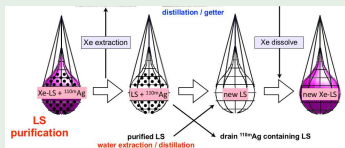
- Xe: soluble in LS, high isotopic enrichment, extraction and purification
- Use existing ultra-pure detector; low-energy anti-neutrino measurement continues (multipurpose)

Phase 1 (publ. Feb. 2013) Exposure: 86 kg·yr



- **Unexpected peak at 2.6 MeV** close to $Q_{\beta\beta}$ of ^{136}Xe ; → rate 'stable' or long-lived
→ search for exotic candidates; most probable: ^{110m}Ar (Fukushima fallout), ^{208}Bi , ^{88}Y , ^{60}Co
- $T_{1/2}^{0\nu} > 1.9 \times 10^{25}$ yr (90% C.L.) (slightly better than EXO-200)

LS purification & Phase 2 (publ. June 2014) 115 d, 348 kg of ^{136}Xe

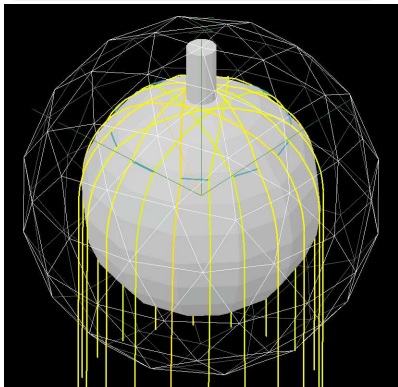


- $T_{1/2}^{0\nu} > 1.3 \times 10^{25}$ yr (90% C.L.) (preliminary)
- **Combined Phases 1+2:** $> 2.6 \times 10^{25}$ yr (90% C.L.)
 $\langle m_{\beta\beta} \rangle < 0.14\text{-}0.28$ eV (90% C.L.)

→ **Near future:** Up to 800 kg instead of 384 kg ^{136}Xe (first $\beta\beta$ 1-ton experiment ?)
→ **Long term:** KamLAND2-Zen (pressurized Xe, high energy resolution)

SNO+: another multipurpose experiment

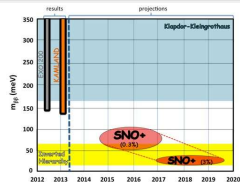
SNO+ 'recycles' upgraded SNO infrastructure
filled with **Te-loaded LS**



Location: Sudbury, Ontario, Canada

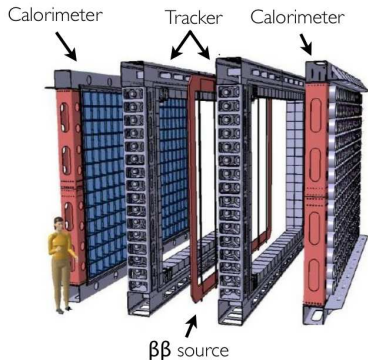
Detector setup

- 2000 m overburden, corresponding 6000 m w.e.
- Outer stainless steel frame (D=17.8 m), carrying 9500 PMTs and acrylic vessel AV (907 m³)
- Between PMTs and AV: 3 m thick ultrapure light water
- Upgrade of trigger/DAQ due to 45× more light
- Scintillator**: 780 ton of acrylic-compatible linear alkylbenzene (LAB), hold down with ropes (buoyancy) installed in January 2013. Water filling (before scin. filling) starting in 9 days !
- Phase I**: Original idea with ¹⁵⁰Nd rejected, March 2013 decision to change to ¹³⁰Te
Spring 2014: 1st batch of Te purchased, HV test for PMTs, laser scatt. syst. tested
 - ~0.3% of nat. Te (% ¹³⁰Te (800 kg)
 - $\langle m_{eff} \rangle = 0.05$ eV (after ~3 yr)
- Phase II**: ~3% of nat. Te → $\langle m_{eff} \rangle = 15$ meV (full IH)



SuperNEMO: a tracking calorimeter

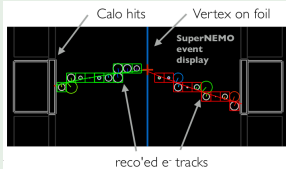
Concept: DBD Source \neq Absorber



Location: Modane, France (tbc)
1700 m overburden, corr. 4800 m w.e.

Detector description

- 3 main components:
 - Multiple $\beta\beta$ foil: ^{82}Se , ^{150}Nd , ^{48}Ca , ^{100}Mo ...
 - Multiwire chamber: track reco., vertex reco.
 - Calorimeters: energy measurement
- **Unique full reco. of $\beta\beta$ kinematics!**



Strategy

- NEMO-3:** 7 kg ^{100}Mo : $T_{1/2}^{0\nu} > 1.1 \times 10^{24}$ yr (90% C.L.)
- SuperNEMO Demonstrator** (planned start in 2015):
 - ^{82}Se main isotope; ener.res: 8% \rightarrow 4% at 3 MeV
 - Lower background, 10 \times higher exposure
- SuperNEMO:**
 - 20 Demonstrator modules using 100 kg ^{82}Se
 - Goal sensitivity: $T_{1/2}^{0\nu} > 10^{26}$ yr
 - $\langle m_{\beta\beta} \rangle < 0.04\text{-}0.1$ eV (90% C.L.)

Status in 2014:

- **Flagship of several new $0\nu\beta\beta$ decay experiments** under development, partly they have been completed. Several R&D projects for future upgrades or next generation experiments are in progress.
- In **2012/2013**, three of the new experiments (**EXO-200**, **KamLAND-Zen**, **GERDA**) reached again **sensitivities comparable/better** than the previous most sensitive experiment Heidelberg-Moskau (HdM).
- The **claim of observation** by a subgroup of the HdM collaboration in 2002 was **rejected** with high probability.
- A lot of **progresses in calculating** the $0\nu\beta\beta$ nuclear processes were achieved.

Designed and applied detector technologies

- **Former technologies** further **optimized**: e.g. Cuore, Gerda
- **Combine different technologies** for **active background suppression**;
technology transfer among low background rare-event physics projects, e.g. $0\nu\beta\beta \leftrightarrow$ dark matter)
- **Multifunctionality** of low background detectors: e.g. $KL \rightarrow KL\text{-Zen}$, $SNO \rightarrow SNO+$
- **Different technologies and isotopes** are **complementary**:
 - **Systematic uncertainties** (from experiment, background modelling and nuclear calculations) can be reduced/ruled out
 - **Discovery potential and sensitivity limits**:
 - Solid state detectors: low mass, high resolution \rightarrow signal discovery
 - Xe-experiments: high mass, low resolution \rightarrow push limits

Outlook:

- **Excluding $\langle m_{\beta\beta} \rangle = 0.1-0.5$ eV (degenerate hierarchy):**
 - This parameter space is **in the range of** EXO-200, KamLAND-Zen (^{136}Xe) and GERDA (^{76}Ge); within the next 5 years, other experiments (with partly other isotopes) will reach similar sensitivities.
 - So far, **no signal was found**. **The KKDC claim** of signal observation was **rejected with high probability**.
- **If range $\langle m_{\beta\beta} \rangle = 0.02-0.05$ eV holds (inverted hierarchy):**
 - **Necessity** for large scale enrichment, lower background reduction AND reliable identification of residual background
 - Possible experiments: nEXO, Gerda Phase III and Majorana, KamLand-Zen(2),...
 - Guesed start of commissioning of first experiments after 2020 (SNO+ with 3% Te)
 - **Discovery in 2-3 isotopes necessary** to confirm the observation
 - **An ultimate precision-experiments** (SuperNEMO- or NEXT-type) have to follow to improve NME and describe exchange mechanisms

2014+ will be exciting years for $\beta\beta$ decay experiments
searching for lepton number violation !

Example: GERDA's transition from Phase I to Phase II

$$T_{1/2} \propto f_{76} \cdot \epsilon \cdot \sqrt{\frac{M \cdot T}{\Delta E \cdot B}}, \quad (\text{if background is present})$$

- **M** (Mass): → new Ge detectors added, old 15 + new 20 kg
- **T** (DAQ livetime): → larger DAQ period, 20 to 100 kg·yr
- **f_{76}** (Enrichment): → ^{76}Ge enrichment from 8% to 90%
- **ϵ** (Detector efficiency): Precise characterisation of detector
- **ΔE** (Energy resolution): → novel detector technologies with better ΔE (BEGe vs. coax) from 0.2% to 0.15%
- **B** (Background suppression) (**most important!**):
 - **Goal**: BI from 0.01 to 0.001 cts/(kg·yr·keV)
 - Minimize exposure to cosmogenic radiation → less cosmogenic-induced prompt signals and delayed radioisotope decays
 - Enhanced pulse shape discrimination of BEGe detectors
 - Read-out scintillation light in LAr of bckg events

● Lucifer:

- **Isotopes:** 1. option ZnSe, enriched from 8.7 to 95% in ^{82}Se ;
2. option ZnMoO₄
- **Technique:** scintillation bolometry, → optimal for effective α -rejection
- **Goal of demonstrator (2010-2015):** prove feasibility of BI ~ 0.001 cts/kg/yr/keV
- **Crystal growth:** problematic, only two suppliers, mass yield transfer still too low

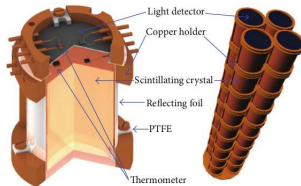
● Lumineu:

- **Isotope:** ^{100}Mo in ZnMoO₄
- **Technique:** scintillation bolometry; scintillation detected by small 170 μm thick Ge detectors
- **Status 2014:** prototype crystals with masses of 55-340 g produced and characterized

● AmoRE:

- **Begin of R&D:** 2009
- **Isotopes:** ^{100}Mo and ^{40}Ca in CaMoO₄
- **Status:** low-background, high transparent crystals with up to 0.6 kg produced
- **Planned sensitivity:** with 100 kg and 5 yr of data collection:
→ $T_{1/2}^{0\nu} > 3 \times 10^{26}$ yr, $\langle m_{\beta\beta} \rangle < 0.02\text{-}0.06$ eV

Design for Lucifer



Crystal for AmoRE R&D (2013)

